

Response and Uncertainty of the Parabolic Variance PVAR to Noninteger Exponents of Power Law

François Vernotte, Enrico Rubiola, Siyuan Chen

▶ To cite this version:

François Vernotte, Enrico Rubiola, Siyuan Chen. Response and Uncertainty of the Parabolic Variance PVAR to Noninteger Exponents of Power Law. IEEE Transactions on Instrumentation and Measurement, 2021, 70, pp.1-6. 10.1109/TIM.2021.3073721 . hal-02870806

HAL Id: hal-02870806

https://hal.science/hal-02870806

Submitted on 17 Jun 2020

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Response and Uncertainty of the Parabolic Variance PVAR to Non-Integer Exponents of the Power Law

François Vernotte, Siyuan Chen, Enrico Rubiola

Abstract-Oscillator fluctuations are described as the phase or frequency noise spectrum, or in terms of a wavelet variance as a function of the measurement time. The spectrum is generally approximated by the 'power law,' i.e., a Laurent polynomial with integer exponents of the frequency. This article extends the domain of application of PVAR, a wavelet variance which uses the linear regression on phase data to estimate the frequency, and called 'parabolic' because such regression is equivalent to a parabolic-shaped weight function applied to frequency fluctuations. In turn, PVAR is relevant in that it improves on the widely-used Modified Allan variance (MVAR) enabling the detection of the same noise processes at the same confidence level in a shorter measurement time. More specifically, we provide (i) the analytical expression of the response of the PVAR to the frequency-noise spectrum in the general case of noninteger exponents of the frequency, and (ii) a useful approximate expression of the statistical uncertainty.

Keywords—Allan variances; frequency stability; fractional noise; uncertainty assessment; degrees of freedom

I. Introduction

The fluctuations of an oscillator are generally described as the phase noise $\mathcal{L}(f)$, where f is the Fourier frequency, or as the two-sample variance $\sigma_{\nu}^2(\tau)$, where τ is the integration time. The latter takes different flavors, the most known of which are the Allan variance (AVAR) and the modified Allan variance (MVAR). The concepts related to $\mathcal{L}(f)$ were introduced in the 1960s to describe the fast fluctuations of oscillators for radars and frequency synthesis [1]. By contrast, $\sigma_{v}^{2}(\tau)$ was introduced to describe the fluctuations of Cs-beam clocks for timekeeping, with obvious focus on slow fluctuations [2], [3]. Traditionally, the boundary between these two choices was $\tau \approx 0.1...1$, or $f \approx 1 \dots 10$ Hz. The overlap was rather small, of the order of one decade. In fact, time counters could not be easily used at a sampling interval τ_0 smaller than ≈ 100 ms, limited by the slowness of the IEEE 488 BUS transferring ASCII data. By contrast, the measurement of $\mathcal{L}(f)$ at low Fourier frequencies was limited by the narrow dynamic range of the double balanced mixer used as the phase-to-voltage converter (no more than $\pm 20^{\circ}$), and of the analog to digital converters. The Fast Fourier Transform analyzers were so complex and expensive that they were avoided when possible. Interestingly, the two-sample variance is broadly equivalent to a one-octave filter centered at $f \approx 0.45/\tau$.

F. Vernotte and E. Rubiola are with (1) FEMTO-ST Institute, CNRS Lab no. 6174, Time and Frequency Department, (2) Observatory THETA, and (3) UBFC, all in Besançon, France. E. Rubiola is also with the Istituto Nazionale di Ricerca Metrologica INRiM, Divsion of Quantum Metrology and Nanotechnology, Torino, Italy. Siyuan Chen is with (1) Radioastronomy Station Nançay, PSL, Nançay, France, (2) LPC2E, Université d'Orléans, France, (3) FEMTO-ST Institute, Time and Frequency Department, (4) UBFC.

Nowadays these limitations are gone, and the overlap in the domain of application of $\mathscr{L}(f)$ and $\sigma_{\mathsf{y}}^2(\tau)$ is of 6-8 decades. Digital instruments can measure $\mathscr{L}(f)$ from 0.1-1 mHz [4], [5], [6], [7]. This is made possible by Software Defined Radio techniques (see [8], [9] for a general overview), which enable phase measurements not bounded to $\pm \pi$, and low sampling frequency by proper decimation of high-speed data. The CORDIC algorithm [10], [11] is the preferred choice to calculate $\varphi(t)$ from the digitized I/Q stream. Counters with picosecond resolution were available since the 1970s with the Nutt interpolator [12], but continuous time stamps at a sampling interval $\tau_0 \approx 100$ ns [13], [14] could be possible only thanks to Field Programmable Gate Arrays (FPGAs). The minimum τ is actually greater than τ_0 because trivial limitations intervene, but the practical limit is still of the order of several μ s. The conclusion is that assessing the *equivalence* between spectra and variances is more important than ever.

It is generally agreed that the phase noise of oscillators is well described by the 'power law' or 'polynomial law' model, which is the extension of the regular polynomial to the negative powers of the variable (Laurent polynomials). While the literature is shy about exceptions, we came across significant practical cases where the phase noise has a noninteger slope over a few decades. In other domains of physics, the term 'flicker noise' refers to a noise process whose spectrum is of the f^{β} type, where the exponent β is actually in [-1.2, -0.8] to [-1.5, -0.5] depending on the author [15], [16], [17]. Accordingly, we may find f^{β} phase noise in oscillators, and $f^{\beta-2}$ phase noise after the phase-to-frequency conversion known as the Leeson effect [18]. The fractionalorder frequency control, nowadays quite popular [19], [20], [21], is a good reason for non-integer slopes to be present in the spectrum of a locked oscillator or laser. Non-integer slopes also appear in other branches of frequency metrology. For example, theoretical predictions about millisecond pulsars suggest that the common FM noise could follow the $f^{-7/3}$ law [22], [23]. Finally, a continuous polynomial law is necessary in Bayesian statistical analysis, when we estimate the polynomial law from the measured spectrum [24]. Interestingly, the continuous law is needed as an intermediate step even when estimation targets integer exponents.

The response of $\sigma_y^2(\tau)$ to phase noise in the case of non-integer exponents of the power law was already solved for the Allan Variance (AVAR) and the Modified Allan Variance (MVAR) [25], while the Parabolic Variance (PVAR) was introduced later [26], [27]. In our opinion, MVAR is obsoleted by PVAR because PVAR is suitable to the same applications, and it enables the detection of the same noise phenomena,

in the same conditions at the same confidence level with a smaller data record [27], i.e., in a shorter measurement time.

This work stands on [25] and extends the results to PVAR providing conversion formulae, degrees of freedom and statistical uncertainty (Type A uncertainty, according to the definitions given by the International Vocabulary of Metrology 'VIM' [28]).

II. THE RESPONSE TO POLYNOMIAL SPECTRA

A. Basic Definitions and Tools

We consider a clock signal $V_0\cos[\omega_0t+\varphi(t)]$ of nominal frequency $\omega_0/2\pi$ and random phase $\varphi(t)$. It is understood that $\varphi(t)$ is not bound to $\pm\pi$, and that $|\dot{\varphi}(t)| \ll \omega_0$. The associated time fluctuation $\mathbf{x}(t) = \varphi(t)/\omega_0$ is usually referred to as *phase time*. The quantity $\mathbf{y}(t) = \dot{\mathbf{x}}(t)$ is the fractional frequency fluctuation.

According to the IEEE Standard 1139 [29], the phase noise is defined as $\mathcal{L}(f) = \frac{1}{2}S_{\varphi}(f)$, that is, half of the single-sided Power Spectral Density (PSD) of $\varphi(t)$. For our purposes, it is convenient to use the quantity

$$S_{\mathbf{y}}(f) = \frac{f^2}{(\omega_0/2\pi)^2} S_{\varphi}(f) , \qquad (1)$$

instead of $S_{\varphi}(f)$, which provides fully equivalent information. The associated polynomial law is usually written as

$$S_{\mathbf{y}}(f) = \sum_{\alpha = -2}^{2} \mathsf{h}_{\alpha} f^{\alpha} \;, \tag{2}$$

where the exponent α equals -2 for random walk FM noise, -1 for flicker FM noise, 0 for white FM noise, 1 for flicker PM noise, and 2 for white PM noise.

From a general perspective, the two-sample variance can be written as

$$\sigma_{\mathbf{y}}^{2}(\tau) = \frac{1}{2} \mathbb{E} \left\{ \left[\overline{\mathbf{y}}_{2} - \overline{\mathbf{y}}_{1} \right]^{2} \right\} , \tag{3}$$

where $\mathbb{E}\{\ \}$ is the mathematical expectation, and \bar{y}_1 and \bar{y}_2 are the two samples of y(t) averaged over contiguous time intervals of duration τ (hereafter the integration time). Our use of (3) differs from the general literature in that \bar{y}_2 and \bar{y}_1 are weighted averages. The uniform average gives AVAR, the triangular-weight average gives MVAR, and the parabolic-weight average gives PVAR. Other options are possible, for example the Hadamard and the Picimbono variances. Accordingly, (3) is rewritten as

$$\sigma_{\mathbf{y}}^{2}(\tau) = \mathbb{E}\left\{ \left[\int_{-\infty}^{\infty} \mathbf{y}(t) \ w(t) \ \mathrm{d}t \right]^{2} \right\} , \tag{4}$$

where w(t) is a wavelet-like function that describes $\overline{y}_2 - \overline{y}_1$, including the weight functions. The specific w(t), named $w_A(t)$ for AVAR, $w_M(t)$ for MVAR and $w_P(t)$ for PVAR are defined in [27, Fig. 3 and related text]. For example, the PVAR weight function is

$$w_P(t) = \frac{3\sqrt{2}\,t}{\tau^3}\,(|t| - \tau) \tag{5}$$

with $t \in [-\tau, \tau]$. This is a parabola, which we refer to as Ω , the most similar Greek letter. Since y(t) is the derivative of x(t), Eq. (4) can be rewritten as

$$\sigma_{\mathbf{y}}^{2}(\tau) = \mathbb{E}\left\{ \left[\int_{-\infty}^{\infty} \mathsf{x}(t) \, \dot{w}(t) \, \mathrm{d}t \right]^{2} \right\} , \tag{6}$$

where $\dot{w}(t)$ is the time derivative of w(t). Thus, it holds that

$$\dot{w}_P(t) = \frac{6\sqrt{2}}{\tau^3} \left(|t| - \frac{\tau}{2} \right) \tag{7}$$

for PVAR, with $t \in [-\tau, \tau]$.

In practice, the variance is calculated from a stream of N samples x_j regularly spaced by τ_0 , which gives the measurement time $\tau=m\tau_0$, integer m (hereafter, the normalized integration time). The expectation $\mathbb E$ is replaced with the average $\langle \ \rangle_M$ of M realizations of $\overline{\mathsf{y}}_2-\overline{\mathsf{y}}_1$, and $\sigma^2_{\mathsf{y}}(\tau)$ is replaced with $\mathrm{AVAR}(\tau)$ or $\mathrm{PVAR}(\tau)$

$$AVAR(\tau) = \frac{1}{2M} \sum_{i=0}^{M-1} \left[\overline{y}_{i+1} - \overline{y}_i \right]^2$$
 (8)

$$PVAR(\tau) = \frac{72}{Mm^{2}\tau^{2}} \sum_{i=0}^{M-1} \left[\sum_{k=0}^{m-1} \left(\frac{m-1}{2} - k \right) \left(x_{i+k} - x_{i+m+k} \right) \right]^{2}, \quad (9)$$

with M=N-2m, since w(t) spans over 2m samples. The main advantage of PVAR is that the weight applied to the x_j samples is equivalent to a linear regression, which features the least-squares fit of the slope. PVAR (τ) is therefore an estimator of the variance of the slope of the x_j samples over the duration τ . For a detailed description of PVAR and its properties see the original article [27]

B. Response of AVAR and PVAR to f^{α}

The response of a generic $\sigma_{v}^{2}(\tau)$ to $S_{v}(f)$ is

$$\sigma_{y}^{2}(\tau) = \int_{0}^{\infty} |H(f)|^{2} S_{y}(f) df,$$
 (10)

where H(f) is the transfer function, or

$$\sigma_{\mathbf{y}}^{2}(\tau,\alpha) = \int_{0}^{\infty} |H(f)|^{2} h_{\alpha} f^{\alpha} \, \mathrm{d}f \tag{11}$$

for the α -th term of the polynomial law (2). Using the subscript A for AVAR, $\left|H(f)\right|^2$ becomes

$$|H_A(f)|^2 = \frac{2\sin^2(2\pi f\tau)}{(\pi f\tau)^2}$$
, (12)

therefore

$$\text{AVAR}(\tau,\alpha) = \frac{\left(2^{-\alpha+1}-4\right)\Gamma(\alpha-1)\sin(\pi\alpha/2)}{(2\pi\tau)^{\alpha+1}}\,\mathsf{h}_\alpha\;. \eqno(13)$$

This is equivalent to [25, Eq. (14)] because we have not introduced the usual cutoff frequency f_H in (10).

Similarly, the transfer function associated to PVAR is

$$|H_P(f)|^2 = \frac{9\left[2\sin^2(\pi f \tau) - \pi \tau f \sin(2\pi f \tau)\right]}{2(\pi f \tau)^6} ,$$
 (14)

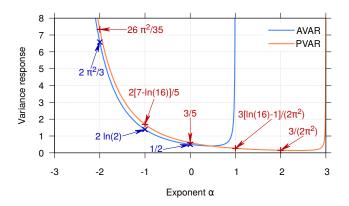


Figure 1. Continuous response of AVAR and PVAR compared to the known responses for $\alpha\in]-3,+3[.$ The responses of AVAR are not plotted for $\alpha\geq 1$ because this estimator diverges without the introduction of a high cut-off frequency.

which is Eq. (17) of [30]. Combining Eq. (11) and (14), we derive the response of PVAR

$$PVAR(\tau, \alpha) = 9 \times 2^{5-\alpha} \left[\alpha^2 - \alpha - 4 - 2^{\alpha} (\alpha - 3) \right] \times \frac{\Gamma(\alpha - 5) \sin(\pi \alpha/2)}{(2\pi\tau)^{\alpha+1}} \, h_{\alpha} \,. \tag{15}$$

Because PVAR converges for f^{α} from f^{-2} to f^{+2} FM noise, we can assume that (15) is valid for $\alpha \in]-3, +3[$.

Figure 1 shows AVAR and PVAR calculated from Eq. (13) and (15), as a function of α . For integer α , the results are the same as in [30, Table I].

III. DEGREES OF FREEDOM OF PVAR ESTIMATES

First, we have to find a simple expression for the number of degrees of freedom (dof) of PVAR estimates for integer power-law noises. Since an equation has been found for a white PM noise (see Eq. (24) in [30]), we assume that it can be generalized for other noise types to

$$\nu \approx \frac{35}{A(\alpha)m/M - B(\alpha)(m/M)^2}$$
 (16)

where $A(\alpha)$ and $B(\alpha)$ are coefficients that need to be determined. From [30, Eq. (24)], we already know that A(+2)=23 and B(+2)=-12. We determine the general $A(\alpha)$ and $B(\alpha)$ from massive Monte-Carlo simulations, and verify the results by comparing them to the dof computed for a continuous power-law.

A. Determination of the Coefficients from Monte-Carlo Simulations

The Monte-Carlo simulation was performed by computing 10 000 sequences of frequency deviations for each $\alpha \in \{-2,-1,0,+1,+2\}$ and for different data length $N \in \{128,2048,32768\}$, i.e. 150 000 simulated sequences in total. For a given α , N and τ , we derived the dof from the averages and the variances of the PVAR for the corresponding set of sequences by using the following well-known property of χ^2_{ν} distributions [30]:

$$\nu = 2 \frac{\mathbb{E}^2[\text{PVAR}(\tau)]}{\mathbb{V}[\text{PVAR}(\tau)]},\tag{17}$$

where $\mathbb{E}[\]$ and $\mathbb{V}[\]$ are the mathematical expectation and the variance of the argument. The least square fit results in

- $A(-2) \approx 34$, $A(-1) \approx 28$, $A(0) \approx 27$, $A(+1) \approx 27$, A(+2) = 23
- $B(\alpha) \approx 12$ for all α .

We have then modeled $A(\alpha)$ by the following 3^{rd} order polynomial and assumed that $B(\alpha) = B$ is constant:

$$A(\alpha) = 27 + \frac{1}{4}\alpha + \frac{5}{14}\alpha^2 - \frac{3}{4}\alpha^3$$

$$B = 12.$$
(18)

Using Eq. (16) and (18), we are now able to assess the dof of all PVAR estimates regardless of the normalized integration time m or the number of samples M.

The upper plot of Fig. 2 compares the dof obtained by the Monte-Carlo simulations and by Eq. (16) and (18) for all integer types of noise. The agreement is confirmed by the lower plot which shows that the discrepancies are within $\pm 10\%$ except for the very first values of m (m = 1, 2).

The model provided by (16) and (18) can be applied to the classical power law, with integer α . Next, we check its validity as an extension for real $\alpha \in]-3,3[$ by computing the dof of PVAR.

B. Verification for Continuous Polynomial-Law Noise

The dof can be computed from Eq. (17). The mathematical expectation of the response of PVAR is given by (15), and the

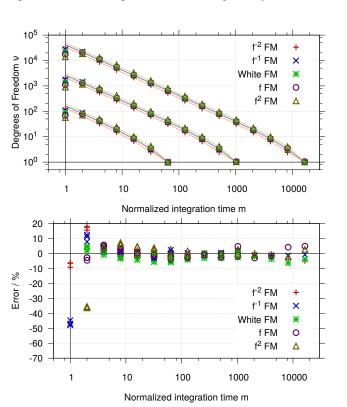


Figure 2. Above: comparison of the empirical dof (crosses) and the approximations (lines) given by Eq. (16) and (18) for all types of noise. The lines are drawn from point to point without interpolation and the last point is set to 1. Below: relative difference (in %) between the empirical dof and the approximations.

variance can be computed from (21) and (22) of [30]

$$\mathbb{V}\left[\text{PVAR}(\tau)\right] = \frac{2}{M^2} \sum_{i=0}^{M-1} \sum_{j=0}^{M-1} \left[\frac{72}{m^4 \tau^2} \right]$$

$$\sum_{k=0}^{m-1} \sum_{l=0}^{m-1} \left(\frac{m-1}{2} - k \right) \left(\frac{m-1}{2} - l \right)$$

$$\left\{ 2R_{\mathsf{x}} [(i+k-j-l)\tau_0] - R_{\mathsf{x}} [(i+k-j-m-l)\tau_0] \right\}$$

$$-R_{\mathsf{x}} [(i+m+k-j-l)\tau_0] \right\}^{-2}$$
(19)

where $R_{\rm x}(\tau)$ is the autocorrelation function of the phase-time ${\rm x}(t)$, i.e., $R_{\rm x}(\tau)=\mathbb{E}\left\{{\rm x}(t){\rm x}(t+\tau)\right\}$. We use the following continuous expression of $R_{\rm x}(\tau)$ versus the power-law exponent α (see [31], [25]):

$$R_{\mathsf{x}}(m\tau_0) = \frac{\mathsf{h}_{\alpha}}{2(2\pi)^{\alpha}\tau_0^{\alpha-1}} \frac{\Gamma(m-\alpha/2+1)\Gamma(\alpha-1)}{\Gamma(m+\alpha/2)\Gamma(\alpha/2)\Gamma(1-\alpha/2)}.$$
(20)

Because this expression involves the Γ function with argument of the order of m, the computations is practically limited to N=128 samples (notice that $\Gamma(128)\simeq 3\times 10^{213}$). This difficulty is avoided using the property that $\Gamma(z)=(z-1)\Gamma(z-1)$ for z>1, and the recursive formula

$$\frac{\Gamma(m-\alpha/2+1)}{\Gamma(m+\alpha/2)} = \frac{\Gamma(3-\alpha/2)}{\Gamma(2+\alpha/2)} \prod_{j=0}^{m-3} \frac{m-j-\alpha/2}{m-1-j+\alpha/2}$$

where the arguments of Γ are greater or equal to 1 for $\alpha \in [-2,+2]$. Therefore, the autocorrelation function can be computed for large N as

$$R_{\mathsf{x}}(m\tau_0) = \frac{\mathsf{h}_{\alpha}}{2(2\pi)^{\alpha}\tau_0^{\alpha-1}} \frac{\Gamma(3-\alpha/2+1)\Gamma(\alpha-1)}{\Gamma(2+\alpha/2)\Gamma(\alpha/2)\Gamma(1-\alpha/2)} \times \prod_{j=0}^{m-3} \frac{m-j-\alpha/2}{m-1-j+\alpha/2}. \tag{21}$$

We used this equation to compute the theoretical variance of $PVAR(\tau)$ versus the continuous variable α , and we deduced the dof from (17).

Let us define $P_{\nu}(\alpha,m,M)=\frac{35M}{m\nu}$. From Eq. (16), we see that $P_{\nu}(\alpha,m,M)\approx A(\alpha)-Bm/M$. The top plot of Fig. 3 shows $P_{\nu}(\alpha,m,M)$ computed from (19) (crosses) and approximated from (18) (solid lines) versus the noise power-law α for $m\in\{4,11,32\}$ (we prefer to plot $P_{\nu}(\alpha,m,M)$ instead of ν for a better visualisation). The agreement is quite good for m=11 and m=16, but there is a notable difference for m=4 and $\alpha<-1$. The lower plot of Fig. 3 shows that this discrepancy is of $\approx 20\%$ maximum, but it remains within $\pm 5\%$ in most cases (all α for m>8, and all m for $\alpha>-1$). This agreement is satisfactory to get an acceptable assessment of the PVAR uncertainties since the relative uncertainties are proportional to $1/\sqrt{\nu}$: they are therefore always below 10% and mostly within $\pm 2.5\%$.

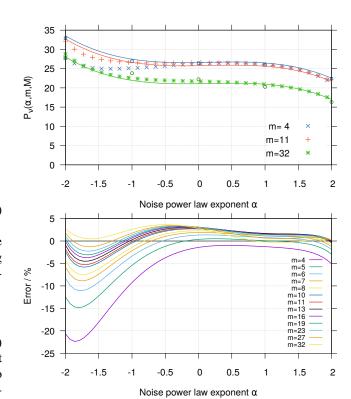


Figure 3. Above: comparison of $P_{\nu}(\alpha,m,M)$ computed from (19) $(\times,+,*)$ and approximated by (16) (solid lines) for N=128 data. The blue squares and the green circles are respectively the values obtained for m=4 and m=32 from the Monte-Carlo simulations. Below: Error (in %) between the approximated values of $P_{\nu}(\alpha,m,M)$ and the computed values.

C. The Case of the Largest Integration Time

The approximation given by Eq. (16) and (18) is close enough to the empirical dof ν for $m \leq N/4$. Moreover, we know that $\nu=1$ for m=N/2. This is enough to draw Fig. 2 since no interpolation is performed between the last 2 points, i.e. m=N/4 and m=N/2. On the other hand, we note that the approximation diverges beyond N/4 (dashed lines in the upper plot of Fig. 4) , if intermediate values of m are computed. However, it is important to assess the uncertaintiy within this interval, particularly if N is not a power of 2.

We fill this gap by interpolating the dof within round $(2^{3/20}N/4) \leq m \leq \text{round}(2^{-3/20}N/2)$ (rounding is necessary to ensure that m is an integer), i.e. between $m_1 \approx \text{round}(1.11N/4)$ and $m_2 \approx \text{round}(0.901N/2)$, with the following semi-logarithmic fit

$$\nu(m) = a\ln(m) + b \tag{22}$$

with

$$a = \frac{\nu(m_1) - 1}{\ln(m_1) - \ln(m_2)} \tag{23}$$

$$b = \frac{\ln(m_1) - \nu(m_1) \ln(m_2)}{\ln(m_1) - \ln(m_2)}.$$
 (24)

For $m \ge m_2$, the dof are set to 1.

To focus on the result of the semi-logarithmic fit an enlargement of the highest 2 decades of m, i.e. $m \in [4096, 8192]$ for N=32768 data is shwon in the top of Fig. 4. The bottom

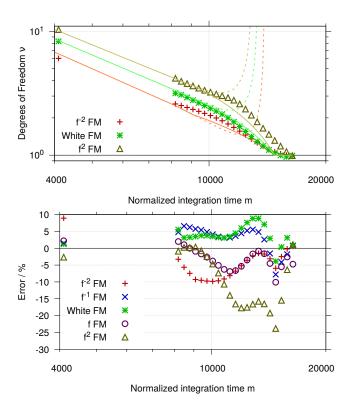


Figure 4. Above: comparison of the empirical dof (crosses) and the semi-logarithmic fits (solid lines) for $N/4 \leq m \leq N/2$ and for random walk FM, white FM and white PM. The dashed lines represent the approximations given by Eq. (16) and (18). In this example N=32768 samples and the logarithmic increment of the m-values is $2^{1/20}$ within [N/4,N/2]. Below: Error (in %) between the empirical dof and the semi-logarithmic fits for all types of noise.

plot shows the error between the fit and the dof computed from the Monte-Carlo simulations. Most of these errors are within $\pm 10\%$, except for white FM. In this case of white FM, the error is between +5% and -20%, and up to -24% for $m=14\,766$. However, this fit is sufficient to ensure an estimation of the PVAR uncertainty for the highest τ within $\sim 10\%$ at worst.

IV. CONCLUSION

We have determined the response of PVAR for continuous power-law noise spectra from a theoretical calculation. Using Monte-Carlo simulations, we have obtained a simplified expression providing the dof of the PVAR estimates within 10 %. We have proven that this expression remains valid for non-integer power-law noise types. Finally, we have shown that a simple interpolation is efficient to fit the dof for the highest octave of integration times. These results generalize the use of the PVAR to process signals with a non-integer powers in the polynomial-law spectrum. This can be used to analyze the timing of milliseconds pulsars and to estimate the non-integer exponent of a red noise, if it is detected.

ACKNOWLEDGEMENTS

This work is funded by the ANR Programme d'Investissement d'Avenir (PIA) under the FIRST-TF network (ANR-10-

LABX-48-01), the Oscillator IMP project (ANR-11-EQPX-0033-OSC-IMP) and the EUR EIPHI Graduate School (ANR-17-EURE-00002), and by grants from the Région Bourgogne Franche Comté intended to support the PIA.

REFERENCES

- A. R. Chi, Ed., Short Term Frequency Stability. NASA SP-80, 1965, Proc. IEEE-NASA Symp. on the Definition and Measurement of Short-Term Frequency Stability, Goddard Space Flight Center, Greenbelt, MD, Nov. 23–24, 1964.
- [2] L. S. Cutler and C. L. Searle, "Some aspects of the theory and measurement of frequency fluctuations in frequency standards," *Proc. IEEE*, vol. 54, no. 2, pp. 136–154, Feb. 1966.
- [3] D. W. Allan, "Statistics of atomic frequency standards," Proc. IEEE, vol. 54, no. 2, pp. 221–230, Feb. 1966.
- [4] J. Miles, TimePod 5330A Programmable Cross Spectrum Analyzer, Operation and Service, Miles Design LLC, 2017. [Online]. Available: http://www.miles.io/TimePod_5330A_user_manual.pdf
- [5] Jackson Labs Technologies, Inc., "PhaseStation," http://www.jackson-labs.com/, accessed Nov. 30, 2020.
- [6] Rohde & Schwarz GmbH, "FSWP phase noise analyzer and VCO tester," https://www.rohde-schwarz.com, accessed Nov. 30, 2020.
- [7] G. Feldhaus and A. Roth, "A 1 MHz to 50 GHz direct down-conversion phase noise analyzer with cross-correlation," in *Proc. Europ. Freq. Time Forum*, York, UK, 2016.
- [8] J. Mitola, P. Marshall, K.-C. Chen, M. Mueck, and Z. G. Zvonar, "Sofrware defined radio — 20 years later: Part 1," *IEEE Comm Mag*, vol. 53, no. 9, Sep. 2015.
- [9] —, "Sofrware defined radio 20 years later: Part 2," *IEEE Comm Mag*, vol. 54, no. 1, Jan. 2016.
- [10] J. E. Volder, "The CORDIC trigonometric computing technique," IRE Transact. Electron. Comput., vol. 8, no. 3, pp. 330–334, Sep. 1959.
- [11] P. K. Meher, J. Valls, J. Tso-Bing, K. Sridharan, and K. Maharatna, "50 years of CORDIC: Algorithms, architectures, and applications," *IEEE Transact. Circuits Sys. I*, vol. 56, no. 9, pp. 1893–1907, Sep. 2009.
- [12] R. Nutt, "Digital time intervalometer," *Rev. Sci. Instrum.*, vol. 39, no. 9, pp. 1342–1345, Sep. 1968.
- [13] Carmel Instruments LLC, "Time interval analyzers and frequency counters," https://www.carmelinst.com, accessed Nov. 30, 2020.
- [14] Guidetech Inc., "Continuous time interval analyzers," http://www.guidetech.com, accessed Nov. 30, 2020.
- [15] P. Dutta and P. M. Horn, "Low-frequency fluctuations in solids: 1/f noise," Rev. Mod. Phys., vol. 53, no. 3, pp. 497–516, Jul. 1981.
- [16] M. B. Weissman, "1/f noise and other slow, nonexponential kinetics in condensed matter," Rev. Mod. Phys., vol. 60, no. 2, pp. 537–571, Apr. 1988.
- [17] E. Milotti, "Linear processes that produce 1/f or flicker noise," Phys. Rev. E, vol. 51, no. 4, pp. 3087–3103, Apr. 1995.
- [18] E. Rubiola, Phase Noise and Frequency Stability in Oscillators. Cambridge, UK: Cambridge University Press, Nov. 2008.
- [19] Y. Chen, I. Petráš, and X. Dingyü, "Fractional order control, a tutorial," in *Proc. American Control Conf.*, St. Louis, MO, USA, Jun. 2009, pp. 1397–1411.
- [20] I. Petráš, Fractional-Order Nonlinear Systems. Springer, 2011.
- [21] B. S. Sheard, M. B. Gray, D. E. McClelland, and D. A. Shaddock, "Laser frequency stabilization by locking to a LISA arm," *Phys. Lett. A*, vol. 320, no. 1, pp. 9–21, Jan. 2003.
- [22] E. S. Phinney, "A practical theorem on gravitational wave backgrounds," arXiv:astro-ph/0108028, Aug. 2001. [Online]. Available: https://arxiv. org/abs/astro-ph/0108028
- [23] S. Chen, A. Sesana, and W. Del Pozzo, "Efficient computation of the gravitational wave spectrum emitted by eccentric massive black hole binaries in stellar environments," *Monthly Notices Royal Astronomical* Soc., vol. 470, no. 2, pp. 1738–1749, May 2017.
- [24] S. Chen, F. Vernotte, and E. Rubiola, "Applying clock comparison methods to pulsar timing observations," *Monthly Notices Royal Astronomical Soc.*, submitted (see arXiv:2011.01912).
- [25] T. Walter, "Characterizing frequency stability: A continuous power-law model with discrete sampling," *IEEE Trans. Instrum. Meas.*, vol. 43, no. 1, pp. 69–79, Feb. 1994.
- [26] E. Benkler, C. Lisdat, and U. Sterr, "On the relation between uncertainties of weighted frequency averages and the various types of Allan deviations," *Metrologia*, vol. 55, no. 4, pp. 565–574, Aug. 2015.

- [27] F. Vernotte, M. Lenczner, P.-Y. Bourgeois, and E. Rubiola, "The parabolic variance (PVAR), a wavelet variance based on the least-square fit," *IEEE Trans. Ultras. Ferroelec. Freq. Contr.*, vol. 63, no. 4, pp. 611–623, Apr. 2016, special Issue of the 50th anniversary of the Allan variance.
- [28] Joint Committee for Guides in Metrology (JCGM), *International Vocabulary of Metrology Basic and General Concepts and Associated Terms (VIM), document JCGM 200:2012*, 3rd ed., 2012. [Online]. Available: https://www.bipm.org/en/publications/guides/
- [29] E. S. Ferre-Pikal, IEEE Standard Definitions of Physical Quantities for Fundamental Frequency and Time Metrology–Random Instabilities (IEEE Standard 1139-2008), IEEE, New York, Feb. 2009.
- [30] F. Vernotte, M. Lenczner, P.-Y. Bourgeois, and E. Rubiola, "The parabolic variance (PVAR), a wavelet variance based on least-square fit," *IEEE Transactions on Ultrasonics, Ferroelectrics and Frequency Control*, vol. 63, no. 4, pp. 611–623, 2016.
 [31] N. J. Kasdin, "Discrete simulation of colored noise and stochastic
- [31] N. J. Kasdin, "Discrete simulation of colored noise and stochastic processes and $1/f^{\alpha}$ power law noise generation," *Proc. IEEE*, vol. 83, no. 5, pp. 802–827, May 1995.