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1 **Variations in eastern Mediterranean hydrology during the last climatic cycle as inferred from**  
2 **neodymium isotopes in foraminifera**

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15

16 **Abstract**

17 The Nd isotopic compositions ( $\epsilon\text{Nd}$ ) of mixed planktonic foraminifera have been analyzed in two  
18 sediment cores collected in the Nile deep-sea fan in order to reconstruct past  $\epsilon\text{Nd}$  of the Eastern  
19 Mediterranean Deep Water (EMDW) and to assess the relative contributions of Nile discharge and  
20 Modified Atlantic Water (MAW) inflow to the Eastern Mediterranean Sea hydrology, as well as  
21 their potential control on anoxic events over the last climatic cycle. The two foraminiferal  $\epsilon\text{Nd}$   
22 records are similar and display an increase in  $\epsilon\text{Nd}$  values during the African Humid Periods.  
23 Superimposed on this precession-forced variability (insolation received by the Earth at low  
24 latitudes), the record of variations in foraminiferal  $\epsilon\text{Nd}$  indicates a 2-unit decrease in  $\epsilon\text{Nd}$  during  
25 the interglacial Marine Isotope Stages (MIS) 5 and 1 compared to glacial MIS6, 4, 3 and 2. The  $\epsilon\text{Nd}$   
26 results suggest that the long-term glacial to interglacial changes in Nd isotopic composition of  
27 EMDW were not entirely induced by variations in Nile River discharge and Saharan dust inputs.  
28 Decreases in  $\epsilon\text{Nd}$  during MIS5 and MIS1 interglacials indicate an increase in the contribution of  
29 unradiogenic MAW to the eastern Mediterranean Sea related to high sea-level stands and greater  
30 seawater exchange between the North Atlantic and Mediterranean basins. In addition, radiogenic

31 seawater  $\epsilon\text{Nd}$  values observed during African Humid Periods (and sapropel events) are associated  
32 with an intensification of Nile discharge and an increase in residence time of deep-water masses in  
33 the eastern Mediterranean Sea, which induces an increase in the interaction between deep-water  
34 masses and radiogenic sediments along the margin of the eastern Mediterranean Sea. Results  
35 confirm that an intensification of the hydrological exchanges between the western and eastern  
36 Mediterranean basins during high sea-level stand and the subsequent higher proportion of  
37 Atlantic Water in the Levantine Basin may have preconditioned the eastern Mediterranean Sea to  
38 sapropel depositions during the last climatic cycle.

39 **Key words:** *Neodymium isotopes; foraminifera; paleo-hydrology; sapropels; eastern*  
40 *Mediterranean Sea; African monsoon.*

41

## 42 **1. Introduction**

43 The Mediterranean Sea is a semi-enclosed basin, where an excess of evaporation over  
44 precipitation and runoff induces a characteristic eastward increase in the salinity of the surface  
45 Atlantic water that enters through the Gibraltar Strait and flows to the eastern Mediterranean  
46 basin; this water effectively becomes denser and contributes to the formation of intermediate and  
47 deep waters. The resulting Mediterranean thermohaline circulation has been demonstrated to be  
48 highly sensitive to present-day and past climate changes in both high and low latitudes (Rossignol-  
49 Strick et al., 1982; Roether et al., 1996; Kallel et al., 1997; Pinardi and Masetti, 2000; Rohling et al.,  
50 2002; Emeis et al., 2003; Scrivner et al., 2004; Melki et al., 2009; Revel et al., 2010; Toucanne et al.,  
51 2015; Filippidi et al., 2016; Tesi et al., 2017). This variability is thought to be responsible for the  
52 deposition of organic-rich sapropel layers (characterized by  $>1\%$   $C_{\text{org}}$ ) in the Eastern  
53 Mediterranean Sea (EMS) (Murat and Got, 2000) and of Organic Rich Levels (ORL) in the Western  
54 Mediterranean Sea (WMS) (e.g. Rohling et al., 2015). However, paleo-hydrological processes and  
55 associated ventilation dynamics related to sapropel deposition are still debated because  
56 conventional water-mass proxies, such as  $\delta^{13}\text{C}$  analyzed on benthic foraminifera, cannot be used  
57 systematically due to the lack of benthic foraminifera in EMS sediments during time intervals  
58 when bottom waters were characterized by anoxic conditions.

59 Despite many unresolved problems, such as the processes responsible for the sluggish  
60 thermohaline circulation during deep-sea anoxic events, significant advances have been made

61 over recent decades in our understanding of sapropel formation since the pioneering work of  
62 Rossignol-Strick et al. (1982). There is now a common consensus that the formation of sapropels in  
63 the Mediterranean Sea is closely linked to: (1) reduced deep-water ventilation associated to  
64 freshwater input lowering surface water salinity, thus causing stratification of the water masses  
65 with limited or no oxygen renewal in deep water; or (2) nutrient-rich surface freshwater inputs,  
66 which significantly increase surface productivity and induce the mineralization of organic matter in  
67 deep water to a level that exceeds oxygen inputs through deep water mass renewal; or (3) a  
68 combination of both processes (Rohling, 1994; Cramp and O'Sullivan, 1999; Rohling et al., 2015).  
69 The leading role of stratification in most studied sapropels is revealed by the neodymium (Nd)  
70 isotopic composition of biogenic/autogenic fractions (e.g. Freydier et al., 2001; Cornuault et al.,  
71 2018; Wu et al., 2019), benthic foraminifera carbon isotopic records, faunal assemblages (e.g.  
72 Kuhnt et al., 2008; Schmiedl et al., 2010; Cornuault et al., 2016), redox-sensitive elemental  
73 compositions (e.g. Jilbert et al., 2010; Tachikawa et al., 2015; Tesi et al., 2017), and circulation  
74 models (e.g. Myers et al., 1998; Myers, 2002; Stratford et al., 2000; Bianchi et al., 2006; Grimm et  
75 al., 2015; Vadsaria et al., 2019).

76 Many studies have confirmed that periods of sapropel deposition are marked by high river runoff  
77 originating from the low-latitude monsoonal system, whereas time intervals between sapropel  
78 depositions are arid with reduced riverine runoff and increased wind-blown sediment supplies (e.g.  
79 Wehausen and Brumsack, 1999; Larrasoana et al., 2003; Zhao et al., 2012; Revel et al., 2014). A  
80 widely accepted explanation for the excess freshwater input during times of sapropel formation is  
81 related to heavy monsoonal precipitation in North Africa which was channeled by the Nile River  
82 and other North African paleo-rivers (e.g. Rossignol-Strick et al., 1982; Fontugne et al., 1994;  
83 Rohling et al., 2002; Emeis et al., 2003; Scrivner et al., 2004; Osborne et al., 2008, 2010; Revel et  
84 al., 2010; Wu et al., 2016, 2017). In North Africa, these humid periods have been attributed to the  
85 northward migration of the rain belt associated with the Inter-Tropical Convergence Zone (ITCZ)  
86 due to precession-driven insolation changes (Rossignol-Strick et al., 1982; DeMenocal et al., 2000;  
87 Gasse, 2000; Arbuszewski et al., 2013; Skonieczny et al., 2015, 2019). The last period of more  
88 intense rainfall compared to the present, the so-called African Humid Period (AHP: from ~14.8 and  
89 ~6 cal kyr BP; Shanahan et al., 2015; Bastian et al., 2017), is thought to be responsible for the  
90 formation of the organic-rich sapropel S1 in the EMS between 10.2 and 6.4 cal kyr BP (Mercone et  
91 al., 2000; De Lange et al., 2008). Thus, the strong correspondence between sapropel formation  
92 and periods of African monsoon intensification, when freshwater discharge by North African rivers

93 into the Mediterranean Sea increased, suggests that the reduction in sea surface salinity and  
94 deep-water convection could be the result of changes in the African monsoon system (Revel et al.,  
95 2015; Rohling et al., 2015). However, enhanced freshwater inputs from the northern margin of the  
96 Mediterranean Sea related to precipitations driven by westerly winds and the melting of glacial  
97 ice-sheets and linked to climate changes in the high latitudes of the Northern hemisphere (e.g.  
98 Kallel et al., 1997; Emeis et al., 2003; Melki et al., 2009; Toucanne et al., 2015; Filippidi et al., 2016;  
99 Tesi et al., 2017) would have also led to the reduction of deep-water formation in the EMS,  
100 limiting the oxygen supply to the deep water (Rohling, 1994). Furthermore, several studies have  
101 shown that the Mediterranean thermohaline circulation is sensitive to rapid climatic changes in  
102 the northern hemisphere (e.g. Heinrich and Dansgaard - Oeschger events) (Rohling et al., 1995;  
103 Kallel et al., 1997; Allen et al., 1999; Bartov et al., 2003; Martrat et al., 2004).

104 In addition, the narrow and shallow Siculo-Tunisian and Gibraltar Straits tend to limit exchanges  
105 between the eastern and western Mediterranean basins, and with the North Atlantic. It has been  
106 proposed that water exchanges through the Straits may have been significantly limited during the  
107 glacial low sea-level stands of the late Quaternary. A general circulation model has shown that  
108 Mediterranean water outflow was reduced by 50% during the Last Glacial Maximum (LGM)  
109 (Mikolajewicz, 2011), and several studies have pointed to a salinity increase in the eastern basin  
110 (Thunell and Williams, 1989; Myers et al., 1998a). In particular, Mikolajewicz (2011) has shown  
111 that the formation of deep waters in the eastern basin during the LGM was strengthened and that  
112 these waters were principally formed in the Aegean Sea and not in the Adriatic Sea as is the case  
113 today. It has also been proposed that rising global sea level during deglaciation led to greater  
114 exchanges between the basins and a drop in surface salinity, thus preconditioning the sapropel  
115 event by slowing down intermediate and deep convections (Grimm et al., 2015).

116 Studies of past changes in Mediterranean thermohaline circulation have mostly investigated the  
117 last glacial period, focusing on the deposition of sapropel S1 (10.2 – 6.4 cal kyr BP; Mercone et al.,  
118 2000) (e.g. Cacho et al., 2002; Sierro et al., 2005; Frigola et al., 2008; Schmiedl et al., 2010;  
119 Toucanne et al., 2012; Angue Minto'o et al., 2015; Jiménez-Espejo et al., 2015). Continuous  $\delta^{13}\text{C}$   
120 and  $\delta^{18}\text{O}$  records from epibenthic foraminifera of the last glacial – interglacial climatic cycle are  
121 rare and difficult to obtain for bottom suboxic or anoxic environments (e.g. sapropel events) in the  
122 EMS. This is due to the lack of a continuous population of epibenthic foraminifera species for  
123 which a calibrated vital effect for  $\delta^{13}\text{C}$  analyses is well established (Schmiedl et al., 2003, 2010).

124 Consequently, the hydrological pattern of the EMS is not well constrained for periods of sapropel  
125 deposition and for the time intervals that preconditioned such events.

126 In the present study, we have investigated the Nd isotopic composition ( $\epsilon\text{Nd}$ ) of the diagenetic Fe-  
127 Mn coatings precipitated on foraminifera shells. It has been demonstrated that  $\epsilon\text{Nd}$  measured on  
128 planktonic foraminifera shells represents mainly bottom seawater and/or pore water  $\epsilon\text{Nd}$   
129 (Tachikawa et al., 2013, 2014). Such  $\epsilon\text{Nd}$  obtained from foraminiferal shells have been previously  
130 used to provide information on water mass provenance and mixing in the ocean (Molina-Kescher  
131 et al., 2014; Wu et al., 2015b, 2019; Dubois-Dauphin et al., 2017; Cornuault et al., 2018). Modern  
132 Mediterranean seawater displays a wide range of  $\epsilon\text{Nd}$  values, from -11 to -5, with unradiogenic  
133 surface water entering the Mediterranean as Atlantic Water and more radiogenic intermediate  
134 and deep waters originating from the Levantine Basin (Henry et al., 1994; Tachikawa et al., 2004;  
135 Vance et al., 2004). With a residence time of 500-1000 years (Tachikawa et al., 2003; Siddall et al.,  
136 2008) and distinct local basin-scale sources,  $\epsilon\text{Nd}$  is assumed to behave quasi-conservatively, with  
137 great potential to fingerprint the provenance of Mediterranean water masses (Dubois-Dauphin et  
138 al., 2017; Cornuault et al., 2018; Wu et al., 2019).

139 The  $\epsilon\text{Nd}$  of mixed planktonic foraminifera from two sediment cores collected in the Levantine  
140 Basin were investigated in order to reconstruct past  $\epsilon\text{Nd}$  of the EMDW over the last 145 kyr.  
141 Combined with previous  $\epsilon\text{Nd}$  records, the new results allow us to assess the relative contributions  
142 of Nile discharge and inflow of Modified Atlantic Water (MAW) to the thermohaline circulation of  
143 the EMS over the last climatic cycle characterized by several anoxic events and the deposition of  
144 sapropels S1 to S5. This allows us to constrain the hydrology at the origin of these anoxic events  
145 under different environmental conditions (different sea levels, different freshwater supplies from  
146 the African rivers and different hydrological exchange conditions at the Gibraltar and the Siculo-  
147 Tunisian Straits).

## 148 **2. Regional hydrological setting in terms of $\epsilon\text{Nd}$**

149 The Mediterranean Sea is an almost enclosed basin that exchanges surface and intermediate  
150 waters with the Atlantic Ocean through the Gibraltar Strait (sill depth  $\sim 300$  m) and surface water  
151 with the Black Sea through the Dardanelles Strait (sill depth  $\sim 100$  m). Driven by a negative water  
152 budget, an anti-estuarine circulation occurs in the Mediterranean Sea. Hence, the relatively fresh  
153 surface Atlantic Water (AW) (salinity  $\sim 36.5$ ), which has an unradiogenic  $\epsilon\text{Nd}$  signature of  $\sim -9.7$   
154 (Tachikawa et al., 2004), flows into the WMS and is then modified by mixing with the ambient

155 surface water to reach  $\sim -10.4$  in the Alboran Sea (Tachikawa et al., 2004, Spivack and Wasserburg,  
156 1988). During its eastward flow, AW mixes with the surrounding surface waters and underlying  
157 intermediate waters leading to the formation of the Modified Atlantic Water (MAW) that flows  
158 along the basin at 50-200 m water depth following a general cyclonic flow path with several eddies  
159 and meanders (Fig. 1). The  $\epsilon\text{Nd}$  values for MAW range from -10.8 to -9.0 in the western basin  
160 (Henry et al., 1994) and from -9.8 to -4.9 in the eastern basin (Tachikawa et al., 2004; Vance et al.,  
161 2004). Since evaporation exceeds precipitation and river runoff, the relatively fresh surface AW  
162 flowing into the WMS also becomes progressively saltier ( $\sim 38.5$ ) as it circulates eastward.

163 During winter time, intense cooling and strong wind-induced heat loss produce denser waters that  
164 sink via convection and form the intermediate waters in the Aegean Sea and the Levantine Basin  
165 and deep waters in the Gulf of Lions and the Adriatic Sea (Robinson et al., 2001; Schroeder et al.,  
166 2012). In particular, the Levantine Intermediate Water (LIW) is formed in the Cyprus-Rhodes area  
167 and it spreads westwards throughout the entire Mediterranean Basin at depths of between  $\sim 150$ -  
168 700 m (Lascaratos et al., 1993; Malanotte-Rizzoli et al., 1999). The LIW acquires its  $\epsilon\text{Nd}$  signature  
169 mainly from the partial dissolution of particles from the Nile River and the Egyptian/Libyan margin  
170 sediments of the EMS, which have  $\epsilon\text{Nd}$  values ranging from -7 to -2 (Weldeab et al., 2002a;  
171 Tachikawa et al., 2004; Ayache et al., 2016). This more radiogenic Nd signature is explained by  
172 sediment originating from the erosion of Ethiopian traps including Cenozoic basaltic silicate rocks  
173 ( $\epsilon\text{Nd} > 0$ ) (Garzanti et al., 2015). The LIW is thus characterized by a radiogenic  $\epsilon\text{Nd}$  of  $\sim -4.8$  in the  
174 eastern part of the EMS and by an east-west  $\epsilon\text{Nd}$  gradient ranging from -4.8 to -9.2, resulting from  
175 mixing with overlying and underlying water masses along its path (Henry et al., 1994; Tachikawa et  
176 al., 2004; Vance et al., 2004). In the Adriatic Sea, LIW is involved in the formation of Adriatic Deep  
177 Water (AdDW) that sinks into the deep EMS contributing, together with the Aegean Deep Water  
178 (AeDW), to the formation of the Eastern Mediterranean Deep Water (EMDW). The EMDW  $\epsilon\text{Nd}$   
179 values range from -7 to -6 with an average value of -6.5 (Tachikawa et al., 2004). The Western  
180 Mediterranean Deep Water (WMDW) is formed in the Gulf of Lions in winter, as a result of mixing  
181 between the relatively fresh surface water and the saline LIW; it then spreads into the Balearic  
182 Basin and Tyrrhenian Sea between  $\sim 2000$  and  $\sim 3000$  m (Millot, 1999; Schroeder et al., 2012) (Fig.  
183 1). The WMDW is characterized by an average  $\epsilon\text{Nd}$  value of  $-9.4 \pm 0.9$  (Henry et al., 1994;  
184 Tachikawa et al., 2004). Between the WMDW and the LIW (from  $\sim 700$  to  $\sim 2000$  m), the  
185 Tyrrhenian Deep Water (TDW) (Millot et al., 2006), which is produced by mixing between WMDW  
186 and EMDW, has an average  $\epsilon\text{Nd}$  value of  $-8.1 \pm 0.5$ .

187 The Blue Nile and Atbara rivers represent together 97% of the suspended sediment load and 68%  
188 of the freshwater supply of the total annual Nile discharge (Foucault and Stanley, 1989; Williams  
189 et al., 2000; Revel et al., 2015).  $\epsilon\text{Nd}$  signatures of sediments from the Nile Basin are characterized  
190 by contrasted signatures, ranging from radiogenic values ( $\epsilon\text{Nd} \approx 0$ ) for the Cenozoic Ethiopian  
191 traps to strongly unradiogenic values ( $\epsilon\text{Nd} \approx -30$ ) for the Precambrian Central Africa Craton  
192 (Garzanti et al., 2015). Sediments from the Bahr el Jebel (between Lake Albert and 10°N of  
193 latitude) are characterized by a  $\epsilon\text{Nd}$  value of -25, whereas Victoria-Albert Nile-derived fluvial  
194 muds by a range from -29 to -36 (Padoan et al., 2011). The Equatorial-White Nile sediment  
195 makes up approximately 3.5 % of the present-day total sediment discharge of the Nile and is fairly  
196 constant throughout the year (Blanchet et al., 2015; Garzanti et al., 2015), with a very  
197 unradiogenic signature (from -30 to -35). In contrast, the Blue Nile sediment, which is mainly  
198 transported during the humid phase, dominates the total sediment discharge of the Nile (72%)  
199 and is characterized by very radiogenic  $\epsilon\text{Nd}$  values (from -3 to 5; Padoan et al., 2011; Blanchet et  
200 al., 2013; Garzanti et al., 2015).

## 201 **2. Material and methods**

### 202 **2.1 Studied cores**

203 For this study, we have selected cores MD90-964 and MS27PT, which are bathed by the Eastern  
204 Mediterranean Deep Water (EMDW).

205 Core MD90-964 (33°02.75'N, 32°38.57'E; water depth 1375 m, length 32.12 m) was collected on  
206 the eastern part of the Nile deep-sea fan during the PROMETE III campaign on board R/V Marion  
207 Dufresne in September 1990 (Fig. 1). Core MD90-964 sediments consist of pale cream to yellowish  
208 brown foraminiferal and nannofossil marl ooze, interbedded with sapropelic layers that vary in  
209 thickness from 2 to 41 cm (Zhao et al., 2011). The  $\delta^{18}\text{O}$  of *Globigerinoides ruber*, clay mineralogy,  
210 total organic carbon (TOC) content and elemental intensities by X-ray fluorescence (XRF) have  
211 been studied previously by Zhao et al. (2011). The age model of core MD90-964 (Zhao et al., 2011)  
212 has been established by correlating the *G. ruber*  $\delta^{18}\text{O}$  record and the Mediterranean *G. ruber* stack  
213 compiled by Lourens (2004), using Gaussian filtering at both obliquity and precession, and cross-  
214 spectral analysis against a target curve that reflects characteristics of orbital parameters (ETP  
215 curve). According to the age model, the upper 7.30 m of core MD90-964 investigated in this study  
216 cover the last 145 kyr and present a linear mean sedimentation rate of 5 cm/kyr. Sapropels S1, S3,



217 S4 and S5 are identified by an increase of  $C_{org}$  content, which reaches up to 6 % during the  
218 deposition of sapropel S5 (Fig. 2; Zhao et al., 2011).

219 Core MS27PT (31°47.90'N, 29°27.70'E; water depth 1389 m, length 7.3 m) was retrieved on the  
220 western Nile delta, around 90 km from the mouth of the Rosetta Nile River, during the Mediflux  
221 MIMES cruise of the R/V Pelagia in 2004 (Fig. 1). The core site lies directly under the influence of  
222 the Nile freshwater discharge. For this study, we have investigated the upper 3.15 m of core  
223 MS27PT which consist of carbonate-rich facies with coarse quartz grains and clastic mud-rich  
224 facies where sapropel S1 has been identified (Revel et al., 2010) (Fig. 2). The age model of the  
225 studied interval of core MS27PT is based on 22 previously published AMS  $^{14}C$  dates (Revel et al.,  
226 2010, 2015; Bastian et al., 2017; Menot et al., 2020) (Fig. 2).

### 227 **3. Methods**

228 The samples investigated in this study consist of 15 to 30 mg of mono-specific planktonic  
229 foraminifera *G. ruber* and mixed planktonic foraminifera, hand-picked in the >150  $\mu m$  size fraction.  
230 Cleaning procedure and purification of Nd have been done in a class 100 clean laboratory using  
231 ultrapure reagents. All of the test samples were crushed between two glass slides to open the  
232 foraminiferal chambers. The calcite fragments were then ultrasonicated for 1 min before pipetting  
233 off the suspended particles with water to separate the waste. This step was repeated until the  
234 water became clear and free of clay. All samples were checked under a binocular microscope to  
235 ensure that all particles had been removed. Most of the samples were then analyzed after this  
236 preliminary physical cleaning step (hereafter referred as “uncleaned foraminifera”) (Wu et al.,  
237 2015b).

238 For this study, some physically cleaned samples were transferred to centrifuge tubes for an  
239 oxidative-reductive cleaning step (hereafter referred as “cleaned foraminifera”). The oxidative-  
240 reductive foraminiferal cleaning procedure followed that described by Vance and Burton (1999),  
241 using 10 ml reductive solution (1 M hydrous hydrazine, 16 M  $NH_4OH$ , 0.25 M citric acid in a ratio of  
242 1:6:3) and 5 ml oxidative solution (0.2 M  $NaOH$  and 30%  $H_2O_2$  in a 1:1 ratio) per sample to more  
243 efficiently remove authigenic Fe-Mn coatings and organic material. For the reductive step,  
244 samples were heated in a water bath at 80°C for 30 min, and were ultrasonicated every 3 min for  
245 10 s. After transferring the reductive cleaning solution to a centrifuge tube, the cleaned  
246 foraminifera were rinsed with Milli-Q water. The analytical procedure for the oxidative step was  
247 similar except that samples were ultrasonicated every 10 min for a period of 30 s.

248 All samples, including uncleaned foraminifera, underwent weak acid leaching for 5 min in 1 ml  
249 0.001 M HNO<sub>3</sub> with ultrasonication. After these cleaning steps, samples were transferred into a  
250 1.5 ml tube. 0.5 ml of Milli-Q water was first added to the tube, and then the foraminifera were  
251 dissolved using stepwise 100 µl 0.5 M HNO<sub>3</sub> until the dissolution reaction stopped. The dissolved  
252 samples were centrifuged, and the supernatant was immediately transferred to Teflon beakers to  
253 prevent leaching of any possible remaining phases. The dissolved foraminifera shell fractions and  
254 authigenic fractions contained in the reductive cleaning solution were dried using a hotplate for  
255 Nd extraction. Nd was purified using Eichrom TRU-Spec and Ln-Spec resins following the detailed  
256 analytical procedures described in Copard et al. (2010). The <sup>143</sup>Nd/<sup>144</sup>Nd ratios were measured  
257 using the ThermoScientific Neptune<sup>Plus</sup> Multi-Collector Inductively Coupled Plasma Mass  
258 Spectrometer (MC-ICP-MS), hosted at the Laboratoire des Sciences du Climat et de  
259 l'Environnement (LSCE) in Gif-sur-Yvette. For the Nd isotope analyses, sample and standard  
260 concentrations were matched at 10 ppb. Mass-dependent fractionation was corrected by  
261 normalizing <sup>146</sup>Nd/<sup>144</sup>Nd to 0.7219 and applying an exponential law. During the analytical sessions,  
262 every set of two samples was bracketed by analyses of the La Jolla Nd standard solution, which is  
263 characterized by certified values of 0.511858±0.000007 (Lugmair et al., 1983). The offset value  
264 between results and certified values of La Jolla standard was lower than 0.4 epsilon units (εNd) for  
265 all of the analyses presented in this study. The analytical errors reported herein correspond to the  
266 external 2 sigma standard deviation (based on repeated analyses of the La Jolla standard for the  
267 different analytical sessions) and range from 0.1 to 0.5 εNd (Table 1 and Table 2). The analytical  
268 blank values for Nd evaluated by using a quadrupole ICPMS were <4 pg, which represents less  
269 than 0.1% of the minimum Nd yield from foraminifera used in this study. As a result, no blank  
270 correction was applied. Results are expressed as εNd = [(<sup>143</sup>Nd/<sup>144</sup>Nd)<sub>sample</sub>/(<sup>143</sup>Nd/<sup>144</sup>Nd)<sub>CHUR</sub> - 1]  
271 \*10000, with the present-day (<sup>143</sup>Nd/<sup>144</sup>Nd)<sub>CHUR</sub> of 0.512638 (Jacobsen and Wasserburg, 1980).

## 272 **4. Results**

### 273 **4.1 The εNd record of cores MS27PT and MD90-964**

274 The εNd values from core MS27PT, obtained on both monospecific samples of planktonic  
275 foraminifer *G. ruber* and mixed planktonic foraminifera, exhibit a wide range from -5.71±0.13 to -  
276 2.51±0.21 (Table 1 and Fig. 3). The εNd values of monospecific and mixed samples from the same  
277 sediment volume are in agreement within an uncertainty of 2 sigma (Table 1). It is worth noting  
278 that mean εNd values obtained from foraminifera cleaning solutions are similar or slightly more

279 radiogenic than cleaned foraminifera. This confirms that oxidative-reductive cleaning procedures  
280 applied to foraminiferal tests from our sampling sites are not effective to fully remove the Nd  
281 associated with the authigenic Fe-Mn coating (Wu et al., 2015b). Therefore,  $\epsilon\text{Nd}$  obtained in both  
282 reductively cleaned and non-reductively cleaned foraminifera are associated with bottom and/or  
283 pore-water  $\epsilon\text{Nd}$  values as it has been demonstrated elsewhere for a large compilation of  $\epsilon\text{Nd}$   
284 results obtained from foraminifera (Tachikawa et al., 2014).

285 For core MS27PT,  $\epsilon\text{Nd}$  displays lower values of  $-4.7\pm 0.2$  to  $-4.9\pm 0.1$  between 21 and 16.4 cal kyr BP  
286 (Fig. 3). It increases to more radiogenic values ( $-2.5\pm 0.2$  to  $-4.4\pm 0.3$ ) in the time interval between  
287 14.1 and 6.8 cal kyr BP, corresponding to the AHP which is associated with higher Nile flood  
288 discharges (DeMenocal et al., 2000; Revel et al., 2015; Castañeda et al., 2016). The highest value  
289 ( $-2.5\pm 0.2$ ) is recorded at 9.7 cal kyr BP. The  $\epsilon\text{Nd}$  values then steadily decrease from  $-2.5\pm 0.2$  to  
290  $-5.5\pm 0.3$  between 9.7 and 4.3 cal kyr BP. The time interval coeval with the period during which  
291 sapropel S1 was deposited (10.2-6.4 cal kyr BP) is associated with radiogenic  $\epsilon\text{Nd}$  values (from  
292  $-2.7\pm 0.2$  to  $-4.3\pm 0.2$ ). The time interval between 4.3 and 1.5 cal kyr BP presents unradiogenic  $\epsilon\text{Nd}$   
293 values (up to  $-5.9\pm 0.1$ ) which increase again thereafter to reach  $-4.0\pm 0.1$  at the core top.

294 For core MD90-964,  $\epsilon\text{Nd}$  values were obtained for samples of uncleaned mixed foraminifera from  
295 the Marine Isotope Stage (MIS) 6 to the Holocene (Fig. 4E). In general, glacial MIS6, MIS4, MIS3  
296 and MIS2 are characterized by high  $\epsilon\text{Nd}$  values, ranging from  $-3.8\pm 0.2$  to  $-2.8\pm 0.2$ . In contrast,  
297 interglacial MIS5 and the Late Holocene display lower  $\epsilon\text{Nd}$  values, from  $-4.5\pm 0.2$  to  $-2.8\pm 0.3$ .  
298 Superimposed on this long-term glacial-interglacial variation, the monsoon-precession induced  
299 signal is distinguished by more radiogenic values related to the African Humid Periods (from 123.5  
300 to 121.7 cal kyr BP, 102.3 to 99.3 cal kyr BP, 84.3 to 81.8 cal kyr BP; 55 cal kyr BP and 9.5 to 6.2 cal  
301 kyr BP) (e.g. Gasse, 2000; Zhao et al., 2012). These time intervals are also associated with an  
302 increase of  $C_{\text{org}}$  in core MD90-964 and are coeval with the deposition of sapropels S5 (128.0 to  
303 117.7 cal kyr BP), S4 (102.3 to 100.6 cal kyr BP), S3 (81.8 to 79.7 cal kyr BP) and S1 (10.1 to 5.7 cal  
304 kyr BP) (Fig. 4F) (Zhao et al., 2011). In particular,  $\epsilon\text{Nd}$  values systematically increase before the  
305 deposition of sapropels S4, S3 and S1.

306 For the last 20 cal kyr BP, the  $\epsilon\text{Nd}$  record obtained on core MD90-964 displays long-term  
307 variations that are consistent with those of core MS27PT. Interestingly, higher  $\epsilon\text{Nd}$  values from  
308  $\sim 14$  to  $\sim 6$  cal kyr BP are coeval with the timing of the AHP (Fig. 3). The time interval between 9 and  
309 6 cal kyr BP is marked by a steady decrease of  $\epsilon\text{Nd}$  during the Late Holocene (Fig. 3), whereas the

310 interval between 4 and about 1.5 cal kyr BP displays lower values before they increase again  
311 during the last 1.5 cal kyr BP.

#### 312 **4.2 Regional compilation of seawater $\epsilon$ Nd of the Levantine Basin**

313 Previous studies of the Nd isotopic compositions of authigenic oxy-hydroxides in the Levantine  
314 Basin have been conducted on leached bulk sediment (Freydier et al., 2001; Cornuault et al., 2018;  
315 Wu et al., 2019), fish debris/teeth (Wu et al., 2019) and on planktonic foraminifera (Scrivner et al.,  
316 2004; Vance et al., 2004; Osborne et al., 2010; Cornuault et al., 2018; Wu et al., 2019). Past  
317 seawater  $\epsilon$ Nd has been obtained from bulk sediment leachates using different analytical  
318 procedures such as bulk sediment leached with 1 M HCl (Freydier et al., 2001; Wu et al., 2019) and  
319 hydroxylamine hydrochloride of non-decarbonated samples (Tachikawa et al., 2004; Cornuault et  
320 al., 2018).  $\epsilon$ Nd analyses of foraminifera have previously been conducted on both uncleaned  
321 foraminifera (Cornuault et al., 2018) and samples treated with an oxidative-reductive leaching  
322 procedure (Scrivner et al., 2004; Vance et al., 2004; Osborne et al., 2010). All of these analytical  
323 procedures have been debated within the scientific community and have been deemed suitable  
324 for extracting the  $\epsilon$ Nd signature of the deep-water masses of the ocean.

325 Previous  $\epsilon$ Nd records based on cleaned foraminifera collected in the eastern Levantine Basin (ODP  
326 Site 967C - 34°04.270'N; 32°43.528'E; 2552.8 m; Scrivner et al., 2004), in the western Levantine  
327 Basin (OPD Site 971A - 33°42.818'N; 24°42.108'E; 2140.9 m; Osborne et al., 2008) and in the  
328 south-eastern Aegean Sea (core LC21 - 35°39.7'N; 26°35.0'E; 1520 m; Osborne et al., 2010) were  
329 used to reconstruct past  $\epsilon$ Nd of surface waters. However, it has been recently demonstrated that  
330  $\epsilon$ Nd of fossil planktonic foraminifera is not related to the ambient seawater at calcification depths,  
331 but instead reflects bottom and/or pore-water  $\epsilon$ Nd values due to the presence of residual  
332 authigenic Fe-Mn coatings precipitated onto the carbonate shells that cannot be removed by  
333 chemical cleaning procedures (Piotrowski et al., 2012; Roberts et al., 2012; Wu et al., 2015b; Xu et  
334 al., 2018).  $\epsilon$ Nd results obtained from cleaned foraminifera of core LC21 and ODP sites 967C and  
335 971A should be interpreted as bottom/pore water  $\epsilon$ Nd.

336 ODP Site 967C is also located within the EMDW (2551 m) and about 100 km away from core  
337 MD90-964 making it possible to directly compare their  $\epsilon$ Nd records for time intervals  
338 corresponding to sapropels S1 and S5 (Figs. 5 and 6).  $\epsilon$ Nd records for core MD90-964 and Site  
339 967C, display more radiogenic values during African Humid Periods than before and after these  
340 time intervals. For core MD90-964, the highest  $\epsilon$ Nd value is observed in the early African Humid

341 Period (at 13.2 cal kyr BP) before the deposition of sapropels S1. Except this highest  $\epsilon\text{Nd}$  value,  
342 both sapropels of core MD90-964 (S5 and S1) display a range from  $-3.5\pm 0.2$  to  $-4.8\pm 0.2$ ,  
343 characterized by slightly more radiogenic values compared to those obtained in the ODP Site 967C  
344 ( $-5.5\pm 0.5$  to  $-3.1\pm 0.2$ ; Scrivner et al., 2004).

345 For the last 20 cal kyr BP, core MS27PT is also characterized by similar long-term variations of  $\epsilon\text{Nd}$   
346 with an interval of more radiogenic values between 14 and 6 cal kyr BP followed by a progressive  
347 return to unradiogenic values from 6 to 4 cal kyr BP. For cores MS27PT and MD90-964, the last  
348 about 1.5 cal kyr BP show more radiogenic values. The good agreement observed in the long-term  
349 changes of the  $\epsilon\text{Nd}$  during the S5 (for core MD909-964 and ODP Site 967C) and the last 18 cal kyr  
350 BP (for core MS27PT, core MD909-964 and ODP Site 967) suggests that all these cores provide a  
351 regional Nd isotopic signature of the EMDW.

352 Core MD04-2722 located south of Cyprus at 1780 m water depth in the eastern Levantine Basin  
353 (Fig. 1) presents radiogenic LGM  $\epsilon\text{Nd}$  values (around -3), comparable to those observed for cores  
354 located along the Nile deep-sea fan.  $\epsilon\text{Nd}$  values decrease between 17.5 and 15 cal kyr BP followed  
355 by a time interval characterized by radiogenic isotopic composition between 14.5 and 5 cal kyr BP  
356 (Fig. 5). The time interval thereafter (between 5 and 3 cal kyr BP) is associated with a slight  
357 decrease in the  $\epsilon\text{Nd}$  values. Such long-term variations in  $\epsilon\text{Nd}$  are quite similar to those observed in  
358 cores MD90-964 and MSPT27 located upstream of the general counterclockwise deep-sea  
359 circulation of the eastern Mediterranean basin (Fig. 1). Superimposed on this long-term trend,  
360 core MD04-2722 is characterized by shifts of  $\epsilon\text{Nd}$  to unradiogenic values around 13.5 and 6.8 cal  
361 kyr BP that are not observed in other  $\epsilon\text{Nd}$  records of the Levantine Basin (MD90-964, MS27PT and  
362 ODP Site 967C), with the exception of one shift to a single unradiogenic value at 6.8 cal kyr BP  
363 which is also observed in the ODP Site 967C (Fig. 5). This may reflect different time resolutions or a  
364 potential local effect on Nd isotopic signature. This implies that long-term variations of  $\epsilon\text{Nd}$  below  
365 approximately 1300 m water depth reflect a regional  $\epsilon\text{Nd}$  signal of the Levantine Basin. We note  
366 a slight decrease of the  $\epsilon\text{Nd}$  range along the north-eastward circulation pattern of the deep-water  
367 masses that could be the result of a slight modification of the Nd isotopic signature of deep-water  
368 masses through lithogenic Nd input from the volcanic margin of the eastern Levantine Basin.

369 In contrast, the  $\epsilon\text{Nd}$  record of core BC07, located closer to core MD90-964 and at shallower depth  
370 (893 m), displays variations from  $-6.2\pm 0.3$  to  $-4.2\pm 0.3$  (Freydier et al., 2001) that differ somewhat  
371 from those of ODP Site 967C and core MD90-964 during the AHP (Fig. 5). While time intervals

372 before 15 cal kyr BP and after 6 cal kyr BP are characterized by similar  $\epsilon\text{Nd}$  variations as ODP Site  
373 967C and cores MD90-964 and MS27PT, the AHP is instead associated with more unradiogenic  
374  $\epsilon\text{Nd}$ , with values as low as  $-5.9\pm 0.3$  during the sapropel S1. This suggests that the  $\epsilon\text{Nd}$  record of  
375 water masses at  $\sim 900$  m differs from those obtained at greater water depths (below  $\sim 1300$  m) for  
376 a time interval between 14 and 6 cal kyr BP. The difference between  $\epsilon\text{Nd}$  records above and below  
377  $\sim 1300$  m reaches its maximum during the time interval of sapropel S1 deposition.

## 378 **5. Discussion**

379 The  $\epsilon\text{Nd}$  record obtained from uncleaned foraminifera of core MD90-964 allows us to establish,  
380 for the first time, the Nd isotopic signature of the EMDW of the Levantine Basin over the last  
381 climatic cycle and to extend our knowledge of the  $\epsilon\text{Nd}$  variability for sapropels S1 and S5 to those  
382 of S3 and S4. The  $\epsilon\text{Nd}$  record of core MD90-964 also displays glacial-interglacial variability with  
383 more radiogenic values during glacial MIS (Fig. 4).

384 A prerequisite for interpreting such seawater  $\epsilon\text{Nd}$  variations through time is the characterization  
385 of present-day Nd isotopic composition of the main water masses circulating in the eastern  
386 Mediterranean basin. Thus, we can evaluate potential temporal changes in the  $\epsilon\text{Nd}$  of the end-  
387 members during the last climatic cycle, and assess the potential influences of lithogenic Nd input  
388 and regional “boundary exchange” on  $\epsilon\text{Nd}$  of deep-water masses.

### 389 **5.1 Possible impact of changes in lithogenic Nd input on the $\epsilon\text{Nd}$ of the Eastern** 390 **Mediterranean Deep Water in the Levantine Basin**

391 At the present time, the EMS is characterized by contrasting seawater  $\epsilon\text{Nd}$  values ( $\sim -10$  to  $-5$ ) with  
392 strong zonal and vertical gradients which are attributed to a mixing between the inflowing MAW  
393 ( $\sim -10$ ) and the more radiogenic underlying LIW ( $\sim -5$ ) and EMDW ( $\sim -6.5$ ) (Tachikawa et al., 2004)  
394 (Fig 1b). The latter acquires its  $\epsilon\text{Nd}$  signature from boundary exchange with radiogenic basaltic  
395 material originating from the Ethiopian traps, which is transported to the EMS by the Nile river  
396 ( $\epsilon\text{Nd} = -3$  to  $+3$ , Tachikawa et al., 2004), and from the volcanic arc of the Aegean Sea ( $\epsilon\text{Nd} \sim -2.5$ ),  
397 with the most radiogenic material being located at the eastern border of the Levantine Basin ( $\epsilon\text{Nd}$   
398  $= +3.5$  to  $+6$ ) (Ayache et al., 2016).

399 In this regard, previous studies have explored past variations of seawater  $\epsilon\text{Nd}$  in the Levantine  
400 Basin as a balance between unradiogenic aeolian Saharan dust and radiogenic Nile river discharge  
401 (Scrivner et al., 2004; Revel et al., 2015; Cornuault et al., 2018; Wu et al., 2019), including the

402 possibility of a significant contribution from paleo-rivers along the African margin (Osborne et al.,  
403 2010), or a predominant contribution through the mechanism of boundary exchange (Ayache et  
404 al., 2016; Vadsaria et al., 2019; Wu et al., 2016). Tachikawa et al. (2004) have suggested that the  
405 Nd flux from Nile River freshwater has a minor influence on the present-day Mediterranean  
406 seawater  $\epsilon\text{Nd}$ , with its Nd concentration being low and the water input 10 times smaller than the  
407 unradiogenic AW flowing from the Western Basin to the shallow waters of the Levantine Basin.  
408 Several previous studies have demonstrated that detrital material in the Nile deep-sea fan results  
409 from the mixing of sediments derived from Saharan/Libyan dusts and Nile sediments (Weldeab et  
410 al., 2002a; Scheuven et al., 2013; Garzanti et al., 2015; Revel et al., 2015). Cores MD90-964 and  
411 MS27PT, located on the Nile deep-sea fan, are then suitable for assessing lithogenic Nd input to  
412 the EMDW.

#### 413 **5.1.1 Contribution of the Nile river**

414 Past humid periods, corresponding to the sapropels, were systematically accompanied by higher  
415 deposition of iron/smectite-rich sediments in the Nile deep-sea fan, reflecting enhanced physical  
416 erosion and transport of sediments from the Ethiopian Highlands (Krom et al., 1999, 2002;  
417 Weldeab et al., 2002b; Revel et al., 2010, 2014, 2015; Langgut et al., 2011; Zhao et al., 2012). In  
418 contrast, past arid periods are associated with a lower contribution from the Blue Nile and a  
419 higher relative proportion of sediments derived from the White Nile, along with a peak in the  
420 aeolian dust component (Revel et al., 2015; Zhao et al., 2012). Large smectite contents recorded in  
421 cores MD90-964 (Zhao et al., 2012) (Fig. 4D) and MS27PT, combined with Sr and Nd isotopic  
422 compositions (Revel et al., 2014) indicate the dominant contribution of the Blue Nile (smectite) to  
423 the Nile discharge, at least during the last climatic cycles. The low proportion of sediment derived  
424 from the White Nile is mainly due to the Sudd marshes in Southern Sudan which efficiently trap  
425 sediments from tributaries of the White Nile (Padoan et al., 2011; Garzanti et al., 2015). In  
426 addition, during the African Humid Periods, it has been suggested that large lakes formed in the  
427 region extending from Sudd to the confluence at Khartoum (Barrows et al., 2014). These lakes  
428 likely further limited the northward transport of unradiogenic Nd sediments to the Nile deep-sea  
429 fan. Thus, there is no argument to support the hypothesis that changes in the relative contribution  
430 of the Equatorial-White Nile and Blue Nile tributaries triggered the seawater  $\epsilon\text{Nd}$  variations  
431 observed in cores MD90-964 and MSPT27.

#### 432 **5.1.2 Contribution of African dust**

433 The Saharan dust transported to the EMS is characterized by an unradiogenic  $\epsilon\text{Nd}$  values (-15 to -  
434 11; Padoan et al., 2011; Scheuvens et al., 2013), potentially contributing to the unradiogenic Nd  
435 isotopic signature of surface and intermediate water masses (e.g. Tachikawa et al 2004; Ayache et  
436 al., 2016). However, the present-day distribution of  $\epsilon\text{Nd}$  in the surface and intermediate water  
437 masses of the Mediterranean Sea is mainly attributed to the mixing of the AW and the LIW and  
438 suggests a negligible contribution of Nd lithogenic input from aeolian dusts (Wu et al., 2019).

439 In addition, core MD90-964, located at a distal position on the deep-sea fan, is characterized by  
440 lower mean sedimentation rates (5 cm/kyr) compared to core MS27PT (12 cm/kyr) and may be  
441 associated with a higher proportion of Saharan dust components. Variations in the mineralogical  
442 ratio of smectite/illite obtained on the clay fraction of core MD90-964 have been used to track the  
443 relative proportions of Saharan dust (characterized by high illite contents) and Nile contributions  
444 (dominated by smectite, 90%, derived from the Blue Nile and Atbara River draining the Ethiopian  
445 Plateau flood basalts) (Zhao et al., 2011) (Fig. 4D). The transport of Saharan dust to the Levantine  
446 Basin, reconstructed from the smectite/illite ratio, displays slight glacial-interglacial variations with  
447 elevated dust inputs during glacial periods. Such results are in agreement with previous studies  
448 reporting an increase in dust input to the Mediterranean Sea during this time span induced by the  
449 southward migration of the ITCZ (Weldeab et al., 2002b; Revel et al., 2010; Ehrmann et al., 2016,  
450 2017; Lamb et al., 2018) (Fig. 4D). Such glacial increases of unradiogenic  $\epsilon\text{Nd}$  dust inputs to the  
451 eastern Mediterranean Basin do not appear to have significantly modified the seawater  $\epsilon\text{Nd}$   
452 record obtained in core MD90-964, since glacial periods (MIS6 and MIS2-4) are systematically  
453 associated with more radiogenic  $\epsilon\text{Nd}$  values than those of interglacial periods (MIS5 and MIS1)  
454 (Figs. 4D and 4E).

455 In addition, the seawater  $\epsilon\text{Nd}$  record obtained in core BC07, which was collected at a shallower  
456 depth (893 m), could have been slightly more affected by the dissolution of Saharan dust relative  
457 to the other deeper cores. Such higher contribution of dust dissolution in the upper water column  
458 has been demonstrated in previous studies (e.g. Ayache et al., 2016). Yet, this core does not  
459 exhibit variations in seawater  $\epsilon\text{Nd}$  that might have been induced by a glacial increase in dust fluxes.  
460 The time interval corresponding to the AHP (from  $\sim 14.8$  and  $\sim 6$  cal kyr BP), when a decrease of  
461 aeolian dusts has been demonstrated (DeMenocal et al., 2000; Ehrmann et al., 2013), is in fact  
462 associated with more unradiogenic  $\epsilon\text{Nd}$  (Fig. 5). A negligible contribution of the dissolution of  
463 African dust on past seawater  $\epsilon\text{Nd}$  since the last glacial period has been also proposed by



464 Cornuault et al. (2018) and Wu et al. (2019) for several seawater  $\epsilon\text{Nd}$  records obtained from cores  
465 located between  $\sim 800$  and  $\sim 3400$  m water depth in the western part of the EMS.

### 466 **5.1.3 Contribution of African paleo-rivers**

467 A vast fossil river channel network has been identified in the Libyan Sahara, which was active  
468 during African Humid Periods (Rohling et al., 2002). These paleo-rivers carried alteration products  
469 originating from basalts of the Tibesti mountains, ( $\epsilon\text{Nd} = 5$  to  $7$ ; Allegre et al., 1981). Some lake  
470 mollusks analyzed in Wadi Behar Belema and Wadi Quoquin (two of these paleo-rivers) have  
471 shown that the freshwater was significantly radiogenic ( $\epsilon\text{Nd} = -2.2$  to  $-1.8$ ) compared to the  
472 Saharan dust (Osborne et al., 2008). Therefore, it has been suggested that these Libyan paleo-  
473 rivers contributed an additional source of radiogenic Nd to the eastern Mediterranean Basin  
474 during African Humid Periods (Scrivner et al., 2004). Seawater  $\epsilon\text{Nd}$  records for core ODP site 971A  
475 (proximal to the mouths of these paleo-rivers compared to more distal ODP Site 967C or core  
476 LC21), were used to support this hypothesis during sapropel S5 (Osborne et al., 2008, 2010).  
477 However, for different reasons, variations in the foraminiferal  $\epsilon\text{Nd}$  obtained from cores located in  
478 the eastern EMS cannot be entirely explained by changes in the Nd isotopic composition of the  
479 water masses induced by lithogenic input from the African paleo-rivers. First, Sr and Nd isotopic  
480 signatures of the detrital fraction of sediments close to the mouth of these Libyan paleo-rivers  
481 (core CP10BC), have shown maximum sediment discharges during the sapropel S1b time interval  
482 (Wu et al., 2016) whereas  $\epsilon\text{Nd}$  records display more radiogenic values at the beginning of the AHP  
483 ( $\sim 14.8$  cal kyr BP), reaching a maximum at  $\sim 13.5$  cal kyr BP for core MD90-964 and at  $\sim 9.7$  cal kyr  
484 BP for core MS27PT (Fig. 5). Second, the sapropel time interval in cores MD90-964 and MS27PT, as  
485 well as in several cores previously investigated throughout the EMS (Wu et al., 2019), is associated  
486 with a continuous decrease in  $\epsilon\text{Nd}$  values during the deposition of sapropel S1 indicating a more  
487 radiogenic signature for the S1a than for the S1b time interval (Fig. 5). Finally, core CP10, which  
488 should be susceptible to a significant influence from the discharge of Libyan paleo-river sediments  
489 due to its proximal location does not in fact exhibit any influence on its seawater  $\epsilon\text{Nd}$  record (Wu  
490 et al., 2019). Therefore, while we cannot fully rule out a local influence of detrital discharge from  
491 the Libyan paleo-rivers on seawater  $\epsilon\text{Nd}$ , its contribution to Levantine deep-water masses can be  
492 considered negligible.

493 In conclusion, the process of “boundary exchange” between water masses and volcanic sediment  
494 along the margin of the eastern and northern Levantine Basin is one of the major sources of

495 lithogenic Nd to the Mediterranean Sea (Ayache et al., 2016), but neither changes in the inputs of  
496 lithogenic Nd from the Nile River nor from the paleo-rivers of the North African margin can explain  
497 the glacial-interglacial variability observed in the seawater  $\epsilon\text{Nd}$  records of core MD90-964.

## 498 **5.2 Paleo-hydrological implications of the $\epsilon\text{Nd}$ records during glacial-interglacial** 499 **transitions and Terminations I and II**

500 The  $\epsilon\text{Nd}$  records for the EMS display significant glacial and interglacial variations, ranging from  
501  $-2.0 \pm 0.2$  to  $-4.3 \pm 0.2$ , with more radiogenic values during glacial MIS6, 4, 3 and 2 (Fig. 4) that may  
502 have been induced by the result of an increase in the residence time of deep-water masses in the  
503 EMS and/or a decrease in the proportion of unradiogenic MAW flowing into the EMS.

504 Using a regional ocean-atmosphere coupled climate model, Vadsaria et al. (2019) have shown that  
505 a reduction in the formation of deep water in the Eastern basin is associated with an increase in  
506 deep-water  $\epsilon\text{Nd}$  due to a longer interaction between the water masses and the radiogenic  
507 sediments along the eastern margins of the EMS (Ayache et al., 2016). However, since the  
508 ventilation of the EMDW was enhanced during glacial MIS2, 4 and 6, due to saltier and denser  
509 surface water (Thunell and Williams, 1989; Béthoux et al., 1990; Myers et al., 1998a), the more  
510 radiogenic  $\epsilon\text{Nd}$  values for the EMDW during glacial periods cannot be attributed to a modification  
511 of the residence time of deep-water masses. Moreover, several models have shown that exchange  
512 between the Atlantic Ocean and the Mediterranean Sea may have been reduced by as much as  
513 half during the glacial low sea-level stands, resulting in a higher longitudinal salinity gradient in the  
514 Mediterranean Sea (Bryden et al., 1994; Myers et al., 1998b; Mikolajewicz, 2011; Grimm et al.,  
515 2015). Periods of glacial low sea level may have caused a 50% reduction in water exchange  
516 between the eastern and the western Mediterranean basins through the Sicilo-Tunisian Strait  
517 (Mikolajewicz, 2011; Grimm et al., 2015). Consequently, taking into account that glacial periods  
518 were associated with enhanced convection of deep water in the EMS, the more radiogenic  $\epsilon\text{Nd}$   
519 observed for the glacial EMDW can be attributed to a decrease in the contribution of the Modified  
520 Atlantic Water (MAW) to the eastern Mediterranean Basin. In this case, therefore, there is an  
521 entrainment of radiogenic  $\epsilon\text{Nd}$  surface water to the deep basin at the center of production of  
522 deep-water masses. However, we cannot exclude that the decrease in zonal water exchange  
523 between the WMS and the EMS during the glacial period favored also a longer contact time  
524 between the EMDW and the radiogenic sediment draping the eastern Mediterranean continental

525 margin during this time. This hypothesis would need to be tested further using a regional  
526 modeling approach.

527 However, we have tried to estimate the reduction of MAW to the EMS necessary to explain an  
528 increase of 1.5  $\epsilon$ Nd unit observed in core MD90-964 during glacial time relative to the low values  
529 of the MIS1 and 5 (out of sapropel and African Humid Periods) using a simple box-model for Nd in  
530 the EMS established by Wu et al. (2019). We have used the present-day input of Nile river and  
531 dust to the EMS (Nile $\times$ 1, Dust $\times$ 1; Fig.5 of Wu et al., 2019). Results suggest that the observed glacial  
532 increase of  $\epsilon$ Nd could be attributed to a decrease of MAW by 55 % relative to today. Such  
533 variations are likely overestimated as the radiogenic Nd contribution from Nile could have also  
534 been reduced during glacial period. Nevertheless, the calculated reduction of MAW inflow into the  
535 EMS during glacial low sea-level stands agrees with the 50 % reduction in water exchange  
536 between the eastern and the western Mediterranean basins through the Siculo-Tunisian Strait  
537 quantified by Mikolajewicz, (2011) for the LGM.

538 The  $\epsilon$ Nd record of core MD90-964 covers the last Terminations I and II. The significant decrease in  
539 the  $\epsilon$ Nd value (from  $-3.2\pm 0.2$  to  $-4.4\pm 0.2$ ) during Termination I begins at around 18 cal kyr BP and  
540 is coeval with a similar decrease in  $\delta^{13}\text{C}$  obtained from benthic foraminifera (Incarbona et al.,  
541 2011; Sprovieri et al., 2012; Toucanne et al., 2012; Cornuault et al., 2016) (Fig. 6). The decrease of  
542 the  $\epsilon$ Nd during Termination II observed in core MD90-964 is also associated with a decrease of  
543 benthic  $\delta^{13}\text{C}$ , which is particularly well marked in the Ionian Basin, close to the Siculo-Tunisian  
544 Strait (Incarbona et al., 2011). The decrease in  $\delta^{13}\text{C}$  obtained on benthic foraminifera has been  
545 linked to a reduction in deep-water convection in the Mediterranean Sea, but the influence of an  
546 increase in surface productivity cannot be excluded (Incarbona et al., 2011; Cornuault et al., 2016).  
547 During both terminations, the decrease in ventilation suggested by the benthic foraminifera  $\delta^{13}\text{C}$   
548 record is not associated with more radiogenic values, as expected if the seawater residence time  
549 in the EMS had increased. On the contrary, the  $\epsilon$ Nd record shows a decrease suggesting a higher  
550 contribution of unradiogenic AW to the Eastern Basin linked to the rise of the relative sea level  
551 and enhanced water mass exchange between the eastern and western Mediterranean basins. The  
552 timing of reduced ventilation and increase in Atlantic Water contribution to the eastern  
553 Mediterranean basin is consistent with the ocean-biogeochemical model results obtained by  
554 Grimm et al. (2015), which indicate that deep-water anoxia requires a long prelude of deep-water  
555 stagnation of about 6 kyr, with no particularly strong eutrophication. The timing and duration of

556 the stagnation prelude agree with the mechanisms for ORL1 (organic-rich layer) deposition in the  
557 western Mediterranean Sea proposed by Sierro et al. (2005).

558 In particular, the decreases in both benthic  $\delta^{13}\text{C}$  and  $\epsilon\text{Nd}$  before the sapropels S1 and S5 are  
559 coeval with the end of Heinrich Stadial 1 and 11 (HS1 and HS11), respectively. It has been  
560 postulated that an enhanced inflow of less salty seawater in the Mediterranean Sea during the  
561 HS1 time interval favored a reduction in deep-water convection in the western basin and caused  
562 the ORL1 deposition between 14.5 and 8.2 cal kyr BP (Cacho et al., 2002; Martínez-Ruiz et al.,  
563 2003). However, this time interval is still associated with deep-water convection allowing the  
564 transfer of unradiogenic  $\epsilon\text{Nd}$  from MAW to the EMDW at the center of deep-water production in  
565 the EMS. Our new  $\epsilon\text{Nd}$  record from core MD90-964 suggests similar mechanisms during  
566 Termination II, with a potential reduction in the salinity of the Atlantic Water flowing into the  
567 Mediterranean Sea during the HS11 (Kandiano et al., 2014; Jiménez-Espejo et al., 2015; Grant et  
568 al., 2016) and an increase in the proportion of MAW in the deep-water masses of the EMS, 5 to 6  
569 kyr prior to the deposition of sapropel S5.

### 570 **5.3 Variability of the EMS during the African Humid Periods**

571 The  $\epsilon\text{Nd}$  record of core MD90-964 reveals that African Humid Periods (maximum of summer  
572 insolation) of the last climatic cycle are generally associated with more radiogenic  $\epsilon\text{Nd}$  values of  
573 the EMDW (Fig. 4). The sample from core MS27PT dated at  $\sim 9.7$  cal kyr BP displays the most  
574 radiogenic  $\epsilon\text{Nd}$  value ( $-2.5 \pm 0.2$ ), which corresponds to the maximum of summer insolation  
575 received by the Earth at  $20^\circ\text{N}$  (Figs. 4 and 6). The planktonic foraminifera  $\delta^{18}\text{O}$  record from core  
576 9509, located under the River Nile plume in the southeastern Levantine Basin (Fig. 1) (Almogi-  
577 Labin et al., 2009), shows the most negative values ( $-1\text{‰}$ ) at  $\sim 9.3$  cal kyr BP, suggesting a large  
578 input of freshwater from the Nile during this time. Such variations may therefore involve (i) a  
579 more efficient exchange of Nd from seawater and river sediment plumes associated with an  
580 intensification of the Nile River sediment discharge during the African Humid Periods (Revel et al.,  
581 2010; Cornuault et al., 2016) and/or (ii) a longer residence time of water masses in the EMS during  
582 time intervals of stratification of the water column that could have led to increased boundary-  
583 exchange processes at the continental margin (Lacan and Jeandel, 2005).

584 Recent studies have shown that river sediment discharges can modify the Nd isotopic composition  
585 of surrounding water masses (Singh et al., 2012; Chen et al., 2013; Goswami et al., 2014; Osborne  
586 et al., 2014; Rousseau et al., 2015; Wu et al., 2015a), as recently observed in the Bay of Bengal for

587 a water depth greater than 2000 m (Yu et al., 2018). However, the  $\epsilon\text{Nd}$  values of core BC07,  
588 collected at a shallower water-depth (893 m) in the lower portion of the LIW, are less radiogenic  
589 by 2  $\epsilon\text{Nd}$  units during the sapropel S1 time interval relative to the other  $\epsilon\text{Nd}$  records available  
590 nearby (core MD90-964). This suggests that lithogenic Nd from Nile river sediment plumes did not  
591 play a major role in  $\epsilon\text{Nd}$  variations of intermediate and deep-water masses in the eastern  
592 Mediterranean Basin during the African Humid Period. However, Nile river sediment plumes could  
593 partially explain the  $\epsilon\text{Nd}$  variations during the deposition of sapropel S1 given that core MS27PT,  
594 located closer to the Nile river mouth, is characterized by higher  $\epsilon\text{Nd}$  amplitude compared to what  
595 observed in core MD90-964.

596 Results from the recent regional circulation model developed by Vadsaria et al. (2019), using a  
597 five-fold increase in Nile runoff during the deposition of sapropel S1, show more radiogenic  $\epsilon\text{Nd}$   
598 values (by 2  $\epsilon\text{Nd}$  units) for the deep-water masses below 1200 m, associated with a sluggish  
599 circulation, stagnation of deep water and consequent longer interaction with the surrounding  
600 radiogenic margins. This result agrees well with  $\epsilon\text{Nd}$  records obtained for sapropels S1 to S5 (cores  
601 MS27PT and MD90-964) that display an increase of up to 2.5  $\epsilon\text{Nd}$  units (Figs. 5 and 6). In addition,  
602 the less radiogenic values observed in core BC07 during sapropel S1 are also consistent with model  
603 results for intermediate waters (near the LIW) (Vadsaria et al., 2019). Modeling outputs have  
604 shown a modification of the sea surface circulation pattern favorable to the propagation of  
605 unradiogenic western Mediterranean Sea waters into the intermediate waters of the Levantine  
606 Basin. The resulting  $\epsilon\text{Nd}$  decrease observed at intermediate water depth is related to less vigorous  
607 circulation that also induces a reduction of the exchange with high-radiogenic material from the  
608 east of the EMS. Consequently, distinct variations of  $\epsilon\text{Nd}$  records for intermediate- and deep-  
609 waters observed in cores BC07 and MD90-964 during the interval of sapropel S1 deposition are  
610 supported by models and suggest deep-water stagnation and less vigorous circulation of the LIW.

611 The  $\epsilon\text{Nd}$  record of core MS27PT shows an increase around 4 kyr before the deposition of sapropel  
612 S1 (from 10.5 to 6.5 cal kyr BP, Bar-Matthews et al., 2000). This is in agreement with the recent  
613 Nile sediment discharge records that indicate an increasing discharge at ~15 cal kyr BP that could  
614 have induced a freshening of the Mediterranean Sea and a slow-down of the circulation (Revel et  
615 al., 2010). Our new results allow us to investigate the potential for a time lag between seawater  
616  $\epsilon\text{Nd}$  rise and sapropel onset for the previous sapropels S3, S4 and S5. The smectite/illite  
617 mineralogical ratio of core MD90-964 (Fig. 4) enables estimating the Nile discharge over the last

618 climatic cycle (Zhao et al., 2011). This mineralogical ratio increases at about 15 cal kyr BP, in  
619 agreement with an intensification of Nile river discharge prior to the deposition of sapropel S1  
620 (Revel et al., 2010). Taking into consideration the relatively low time resolution of the record,  
621 similar time lags can be observed between the seawater  $\epsilon\text{Nd}$  record and the deposition of  
622 sapropels S3 and S4 (Fig. 4). The seawater  $\epsilon\text{Nd}$  values systematically increase before the intervals  
623 of sapropels deposition but these shifts to more radiogenic values are not always associated with  
624 variations in the smectite/illite ratio. The lower time resolution of the  $\epsilon\text{Nd}$  record of core MD90-  
625 964 around sapropel S5 does prevent us from clearly identifying a similar time lag.

626 The time interval coeval with sapropel S5 is characterized by lower  $\epsilon\text{Nd}$  values relative to those of  
627 sapropels S1 and S4, and to a lesser extent to those of sapropel S3. ODP Site 967C, which is at a  
628 more distal position relative to the Nile River mouth, also displays comparable  $\epsilon\text{Nd}$  values during  
629 the S5 mid-point, with slightly less radiogenic values (of about 0.5  $\epsilon\text{Nd}$  units) at the limits of S5 (Fig.  
630 6). African monsoon rainfall has been more intense during the time interval of the S5 deposition  
631 than during sapropels S1 or S4 (Bar-Matthews et al., 2000), likely leading to more intense Nile  
632 sediment and freshwater discharges. The lower  $\delta^{18}\text{O}$  *G. ruber* values obtained from core MD90-  
633 964 during S5 compared to S1 points to a more significant freshwater discharge from the Nile to  
634 the studied site (Fig. 2C). This suggests that the higher contribution of unradiogenic MAW in this  
635 period of particularly high relative sea level (Fig. 6A), may have overtaken the influence of an  
636 increase in radiogenic Nd from the Nile input and/or of a lesser vigorous deep-water circulation in  
637 the EMS during the deposition of sapropel S5. This would explain why the development of  
638 sapropel S5 is not related to higher  $\epsilon\text{Nd}$  values compared to other sapropels, such as S1 and S4.  
639 For sapropel S4, the  $\epsilon\text{Nd}$  record may indicate enhanced Nile river discharge, or more likely, a  
640 reduction in EMS-WMS exchange induced by lower sea level during MIS5c compared to MIS5e (Fig.  
641 6A). Further modeling studies are needed to estimate the relative contribution of the MAW to the  
642 eastern Mediterranean Basin under the environmental conditions prevailing during the different  
643 sapropel depositions and particularly during the high sea-level of S5.

## 644 **6. Conclusions**

645 Based on the foraminiferal  $\epsilon\text{Nd}$  record of cores MD90-964 and MS27PT, located in the eastern  
646 Levantine Basin, we provide evidence of significant glacial-interglacial variations in the inflow of  
647 Atlantic Water to the EMS during the last climate cycle (last 145 kyr). We have demonstrated for  
648 the first time that  $\epsilon\text{Nd}$  values for the EMDW are systematically associated with more radiogenic

649 values during glacial Marine Isotope Stages. Such long-term glacial to interglacial variations in  $\epsilon\text{Nd}$   
650 values cannot be solely the result of changes in Nile river discharge and Saharan dust inputs.  
651 Decreases in  $\epsilon\text{Nd}$  values during MIS5 and MIS1 interglacials have been attributed to an increase in  
652 the contribution of unradiogenic MAW to the EMS related to high sea-level stands and enhanced  
653 seawater exchange between the North Atlantic and the Mediterranean basins. Termination I and II  
654 are associated with a decrease in seawater  $\epsilon\text{Nd}$  in line with a decrease in the  $\delta^{13}\text{C}$  obtained from  
655 benthic foraminifera, suggesting a sluggish deep-sea ventilation in the EMS related to a higher  
656 contribution of Atlantic Water in the EMS through the Siculo-Tunisian Strait during sea-level rise.

657 Superimposed on this long-term glacial-interglacial variation, a monsoon–precession induced  
658 signal is distinguished in  $\epsilon\text{Nd}$  records by more radiogenic values related to African Humid Periods  
659 (and sapropel events). These periods of radiogenic  $\epsilon\text{Nd}$  of the EMDW have been associated with  
660 an intensification of Nile discharge and an increase in the residence time of deep-water masses in  
661 the EMS, leading to an increase in the contact time between deep-water masses and radiogenic  
662 sediments along the continental margin of the EMS.

663 Overall, our  $\epsilon\text{Nd}$  records combined with previous  $\epsilon\text{Nd}$  values obtained in the EMS reinforce the  
664 hypothesis that a drastic reduction in the hydrological exchanges between the western and  
665 eastern Mediterranean basins, and a subsequent higher proportion of Atlantic Water during sea  
666 level rise may have preconditioned sapropel deposition in the EMS during the last climatic cycle,  
667 as proposed by Grimm et al. (2015).

668

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675 significantly helped to improve this manuscript.

676

677

678 **Figure captions:**

679 Figure 1: (A) Bathymetric map of the Mediterranean Sea showing locations of core MD90-964  
680 (33°02.75'N; 32°38.57'E; water depth 1375 m, length 32.12 m), core MS27PT (31°47.90'N,  
681 29°27.70'E; water depth 1389 m, length 7.3 m) (blue dots) and all the other sites discussed in this  
682 study (black dots) (Freydier et al., 2001; Scrivner et al., 2004; Osborne 2008, 2010; Cornuault et al.,  
683 2018). The simplified modern Mediterranean water circulation is broadly indicated. Signatures of  
684 Nd isotopic composition are displayed for the Nile River and aeolian dust (Grousset et Biscaye,  
685 2005; Padoan et al., 2011; Scheuvens et al., 2013). The potential pathways of North-African paleo-  
686 rivers are also depicted in light gray (Rohling et al., 2002; Wu et al., 2017). (B) Longitudinal  
687 transect of annual mean salinity of the Mediterranean Sea (i.e. the red shading shown in (A); data  
688 from WOA13).  $\epsilon$ Nd values are represented by colored dots (Henry et al., 1994; Tachikawa et al.,  
689 2004; Vance et al., 2004). The map and section were generated using Ocean Data View software  
690 (Schlitzer, 2015).

691 Figure 2: Age models of cores MD90-964 and MS27PT: (A) age-depth curve for core MS27PT  
692 (continuous line) interpolated from AMS  $^{14}\text{C}$  dates (Revel et al., 2015; Bastian et al. 2017) and (B)  
693 age-depth curve for core MD90-964 (dashed line) (Zhao et al., 2011). (C)  $\delta^{18}\text{O}$  analyzed on  
694 planktonic foraminifera *G. ruber* of core MD90-964 (continuous line) and MedStack composite  
695 reference (dashed line; Lourens, 2004). (D) Organic carbon concentration of core MD90-964  
696 sediment (Zhao et al., 2011). (E)  $\delta^{18}\text{O}$  analyzed on *G. ruber* of core MS27PT. (F) Organic carbon  
697 concentration of core MS27PT sediment. Black triangles correspond to the radiocarbon dates  
698 obtained for core MS27PT. The Marine Isotope Stages (MIS) and time intervals of sapropel layers  
699 are also reported.

700 Figure 3: (A) Nd isotopic composition ( $\epsilon$ Nd) obtained from foraminifera of core MS27PT:  $\epsilon$ Nd  
701 values obtained from cleaning solution of *G. ruber* (solid green triangle), reductively cleaned *G.*  
702 *ruber* (solid blue triangle), non-reductively cleaned *G. ruber* (solid red triangle), reductively  
703 cleaned mixed planktonic foraminifera (solid blue rectangle), cleaning solution of mixed planktonic  
704 foraminifera (solid green rectangle), non-reductively cleaned mixed planktonic foraminifera (solid  
705 red rectangle); (B)  $\epsilon$ Nd obtained on non-reductively cleaned mixed planktonic foraminifera (solid  
706 black square) of core MD90-964 for the last 23 cal kyr BP. The Marine Isotope Stages (MIS) and  
707 time intervals of sapropel S1 and AHP are also reported.



708 Figure 4: Variations in (A) the Relative Sea Level (RSL) (Grant et al., 2014) and (B) the summer  
709 insolation (June and July) received by the Earth at 20°N, calculated using Analyseries software  
710 (Paillard et al., 1996) for the last 160 kyr. (C) Variations of the  $\delta^{18}\text{O}$  obtained from *G. ruber* of core  
711 MD90-964 (Zhao et al., 2011); (D) Variations of the smectite/illite ratio of core MD90-964 (Zhao et  
712 al. 2011); (E) Variations of the Nd isotopic composition ( $\epsilon\text{Nd}$ ) obtained from uncleaned mixed  
713 planktonic foraminifera of core MD90-964 (this study); (F) Variations of the Total Organic Carbon  
714 (wt%) of sediments from core MD90-964 (Zhao et al., 2011). The Marine Isotope Stages (MIS) and  
715 time intervals of Sapropel S1 to S5 deposition are also reported.

716 Figure 5: Comparison between  $\epsilon\text{Nd}$  records obtained in this study (cores MD90-964 and MS27PT)  
717 and published records for the last 25 kyr. Core MD04-2722: Cornuault et al. (2018); cores BC07:  
718 Freydier et al (2001); SL29: Wu et al. (2019); ODP Site 967C: Scrivner et al. (2004).

719 Figure 6: (A) Variations in Relative Sea Level (RSL) (Grant et al., 2014) and summer insolation (June  
720 and July) received by the Earth at 20°N, calculated using Analyseries software (Paillard et al., 1996)  
721 for last 160 kyr. (B)  $\epsilon\text{Nd}$  record obtained from ODP Site 967C (Osborne et al., 2010); (C)  $\epsilon\text{Nd}$  record  
722 obtained from non-reductively cleaned mixed planktonic foraminifera of core MD90-964 (this  
723 study); (D)  $\epsilon\text{Nd}$  record obtained from non-reductively cleaned mixed planktonic foraminifera of  
724 core MS27PT (this study); (E) compilation of  $\delta^{13}\text{C}$  obtained from benthic foraminifera from  
725 Mediterranean cores. Records of  $\delta^{13}\text{C}$  obtained from benthic foraminifera are published as  
726 followed: MD04-2722 by Cornuault et al. (2016); MD01-2472 by Toucanne et al. (2012); ODP Site  
727 963 by Incarbona et al. (2011); MD95-2043 by Cacho et al. (2007).

728

729 **Table captions:**

730 Table 1: Results of Nd isotopic composition obtained from planktonic foraminifera samples from  
731 core MS27PT. Nd isotopic composition were obtained from reductively cleaned foraminifera,  
732 cleaning solution, or uncleaned foraminifera. Foraminifera are monospecific (*G. ruber*) or  
733 multispecific (mixed planktonic foraminifera). The age of sediments have been determined by  
734 linear interpolation between 22 AMS  $^{14}\text{C}$  dates of foraminifera (Revel et al., 2010, 2015; Bastian et  
735 al., 2017; Menot et al., 2020).

736 Table 2: Results of Nd isotopic composition obtained from planktonic foraminifera samples from  
737 core MD90-964. The age of sediments have been obtained by the correlation between the *G.*

738 *ruber*  $\delta^{18}\text{O}$  record and the Mediterranean *G. ruber* stack compiled by Lourens (2004) (Zhao et al.,  
739 2011).

740

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Depth (cm)	Age (cal kyr BP)	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$	$\epsilon\text{Nd}$	$\pm 2\sigma$	Samples
0.5	0.7	0.512432	$\pm 0.000008$	-4.03	$\pm 0.17$	uncleaned mixed planktonic foraminifera
2.5	1.3	0.512399	$\pm 0.000008$	-4.67	$\pm 0.16$	uncleaned mixed planktonic foraminifera
6	2.4	0.512345	$\pm 0.000007$	-5.71	$\pm 0.13$	cleaned mixed planktonic foraminifera
6	2.4	0.512351	$\pm 0.000007$	-5.61	$\pm 0.13$	cleaned <i>G. ruber</i>
6	2.4	0.512332	$\pm 0.000018$	-5.97	$\pm 0.35$	cleaning solution, mixed planktonic foraminifera
6	2.4	0.512335	$\pm 0.000008$	-5.91	$\pm 0.15$	uncleaned mixed planktonic foraminifera
9.5	3.0	0.512354	$\pm 0.000010$	-5.53	$\pm 0.20$	uncleaned mixed planktonic foraminifera
10	3.1	0.512357	$\pm 0.000008$	-5.49	$\pm 0.16$	cleaned mixed planktonic foraminifera
10	3.1	0.512360	$\pm 0.000012$	-5.42	$\pm 0.23$	uncleaned mixed planktonic foraminifera
10	3.1	0.512359	$\pm 0.000007$	-5.44	$\pm 0.13$	uncleaned mixed planktonic foraminifera
14.5	4.3	0.512358	$\pm 0.000009$	-5.46	$\pm 0.18$	uncleaned mixed planktonic foraminifera
18	5.1	0.512415	$\pm 0.000014$	-4.36	$\pm 0.27$	uncleaned mixed planktonic foraminifera
18	5.1	0.512430	$\pm 0.000013$	-4.06	$\pm 0.25$	uncleaned <i>G. ruber</i>
22	5.9	0.512417	$\pm 0.000009$	-4.31	$\pm 0.18$	uncleaned mixed planktonic foraminifera
29	6.9	0.512455	$\pm 0.000011$	-3.57	$\pm 0.22$	uncleaned <i>G. ruber</i>
95	8.7	0.512450	$\pm 0.000012$	-3.67	$\pm 0.23$	uncleaned mixed planktonic foraminifera
102	8.9	0.512471	$\pm 0.000017$	-3.26	$\pm 0.33$	uncleaned mixed planktonic foraminifera
102	8.9	0.512490	$\pm 0.000015$	-2.89	$\pm 0.29$	uncleaned <i>G. ruber</i>
120	9.1	0.512468	$\pm 0.000009$	-3.31	$\pm 0.17$	uncleaned mixed planktonic foraminifera
200	9.7	0.512498	$\pm 0.000009$	-2.72	$\pm 0.18$	uncleaned mixed planktonic foraminifera
205	9.8	0.512509	$\pm 0.000011$	-2.51	$\pm 0.21$	cleaned mixed planktonic foraminifera
205	9.8	0.512529	$\pm 0.000008$	-2.12	$\pm 0.16$	cleaning solution, mixed planktonic foraminifera
273	12.3	0.512452	$\pm 0.000012$	-3.63	$\pm 0.23$	uncleaned mixed planktonic foraminifera
280	12.7	0.512482	$\pm 0.000013$	-3.05	$\pm 0.25$	uncleaned mixed planktonic foraminifera
293	14.1	0.512450	$\pm 0.000009$	-3.66	$\pm 0.17$	cleaned mixed planktonic foraminifera
293	14.1	0.512474	$\pm 0.000007$	-3.20	$\pm 0.14$	cleaned <i>G. ruber</i>
293	14.1	0.512459	$\pm 0.000023$	-3.49	$\pm 0.44$	cleaning solution, mixed planktonic foraminifera
293	14.1	0.512483	$\pm 0.000017$	-3.02	$\pm 0.34$	cleaning solution, <i>G. ruber</i>
299	16.4	0.512398	$\pm 0.000010$	-4.68	$\pm 0.19$	uncleaned mixed planktonic foraminifera
303	20.6	0.512390	$\pm 0.000007$	-4.83	$\pm 0.14$	cleaned mixed planktonic foraminifera
303	20.6	0.512386	$\pm 0.000007$	-4.91	$\pm 0.14$	cleaned mixed planktonic foraminifera
315	25.8	0.512411	$\pm 0.000007$	-4.44	$\pm 0.13$	cleaned <i>G. ruber</i>

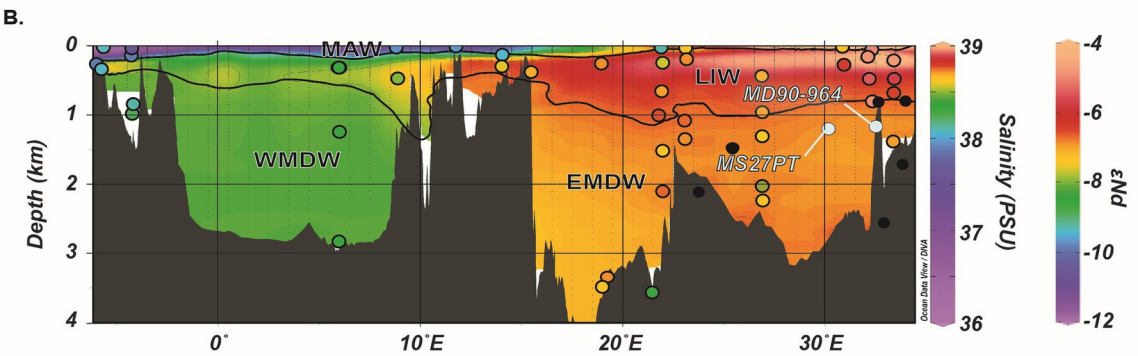
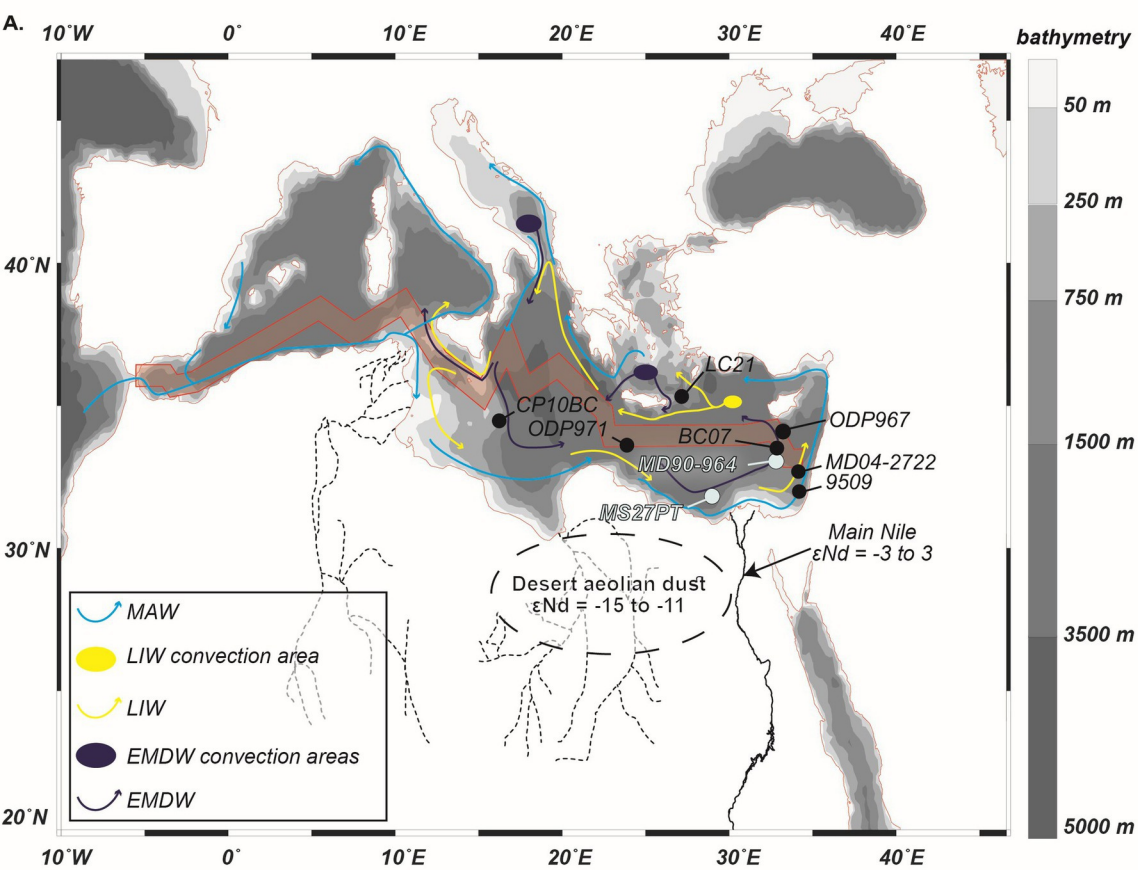
Table 1

Depth (cm)	Age (kyr)	$^{143}\text{Nd}/^{144}\text{Nd}$	$\pm 2\sigma$	$\epsilon\text{Nd}$	$\pm 2\sigma$
4	0.5	0.512441	$\pm 0.000011$	-3.84	$\pm 0.21$
8	0.9	0.512413	$\pm 0.000009$	-4.39	$\pm 0.18$
16	1.9	0.512410	$\pm 0.000008$	-4.45	$\pm 0.17$
28	3.3	0.512410	$\pm 0.000011$	-4.45	$\pm 0.21$
40	4.7	0.512438	$\pm 0.000010$	-3.89	$\pm 0.19$
52	6.2	0.512454	$\pm 0.000010$	-3.58	$\pm 0.19$
72	8.5	0.512450	$\pm 0.000011$	-3.66	$\pm 0.22$
80	9.5	0.512457	$\pm 0.000009$	-3.53	$\pm 0.18$
108	13.2	0.512476	$\pm 0.000009$	-3.17	$\pm 0.18$
128	15.8	0.512430	$\pm 0.000010$	-4.07	$\pm 0.19$
152	19.0	0.512464	$\pm 0.000009$	-3.40	$\pm 0.18$
160	20.0	0.512462	$\pm 0.000010$	-3.43	$\pm 0.19$
172	22.8	0.512537	$\pm 0.000009$	-1.96	$\pm 0.17$
200	29.5	0.512470	$\pm 0.000008$	-3.27	$\pm 0.15$
224	35.2	0.512483	$\pm 0.000008$	-3.03	$\pm 0.16$
236	38.0	0.512462	$\pm 0.000010$	-3.43	$\pm 0.19$
280	45.0	0.512478	$\pm 0.000010$	-3.12	$\pm 0.19$
304	48.8	0.512474	$\pm 0.000010$	-3.20	$\pm 0.20$
324	51.9	0.512473	$\pm 0.000009$	-3.22	$\pm 0.18$
380	60.8	0.512497	$\pm 0.000011$	-2.75	$\pm 0.21$
384	61.4	0.512483	$\pm 0.000009$	-3.02	$\pm 0.18$
420	67.1	0.512484	$\pm 0.000010$	-3.00	$\pm 0.19$
452	72.1	0.512442	$\pm 0.000010$	-3.82	$\pm 0.20$
484	77.2	0.512430	$\pm 0.000010$	-4.06	$\pm 0.20$
504	80.4	0.512440	$\pm 0.000008$	-3.85	$\pm 0.16$
512	81.8	0.512456	$\pm 0.000007$	-3.55	$\pm 0.15$
524	84.3	0.512459	$\pm 0.000008$	-3.50	$\pm 0.16$
556	91.0	0.512412	$\pm 0.000008$	-4.42	$\pm 0.15$
580	96.0	0.512453	$\pm 0.000010$	-3.60	$\pm 0.20$
592	98.5	0.512455	$\pm 0.000009$	-3.58	$\pm 0.18$
596	99.3	0.512484	$\pm 0.000010$	-3.00	$\pm 0.19$
600	100.2	0.512462	$\pm 0.000011$	-3.42	$\pm 0.22$
604	101.0	0.512469	$\pm 0.000009$	-3.30	$\pm 0.18$
608	102.3	0.512495	$\pm 0.000015$	-2.79	$\pm 0.29$
628	108.7	0.512439	$\pm 0.000009$	-3.88	$\pm 0.17$
632	110.0	0.512331	$\pm 0.000008$	-5.99	$\pm 0.39$
636	110.9	0.512454	$\pm 0.000009$	-3.59	$\pm 0.16$
652	114.5	0.512431	$\pm 0.000016$	-4.03	$\pm 0.18$
656	115.4	0.512432	$\pm 0.000010$	-4.01	$\pm 0.32$
660	116.3	0.512421	$\pm 0.000009$	-4.23	$\pm 0.19$
664	117.2	0.512427	$\pm 0.000008$	-4.12	$\pm 0.17$
668	118.1	0.512415	$\pm 0.000009$	-4.35	$\pm 0.16$

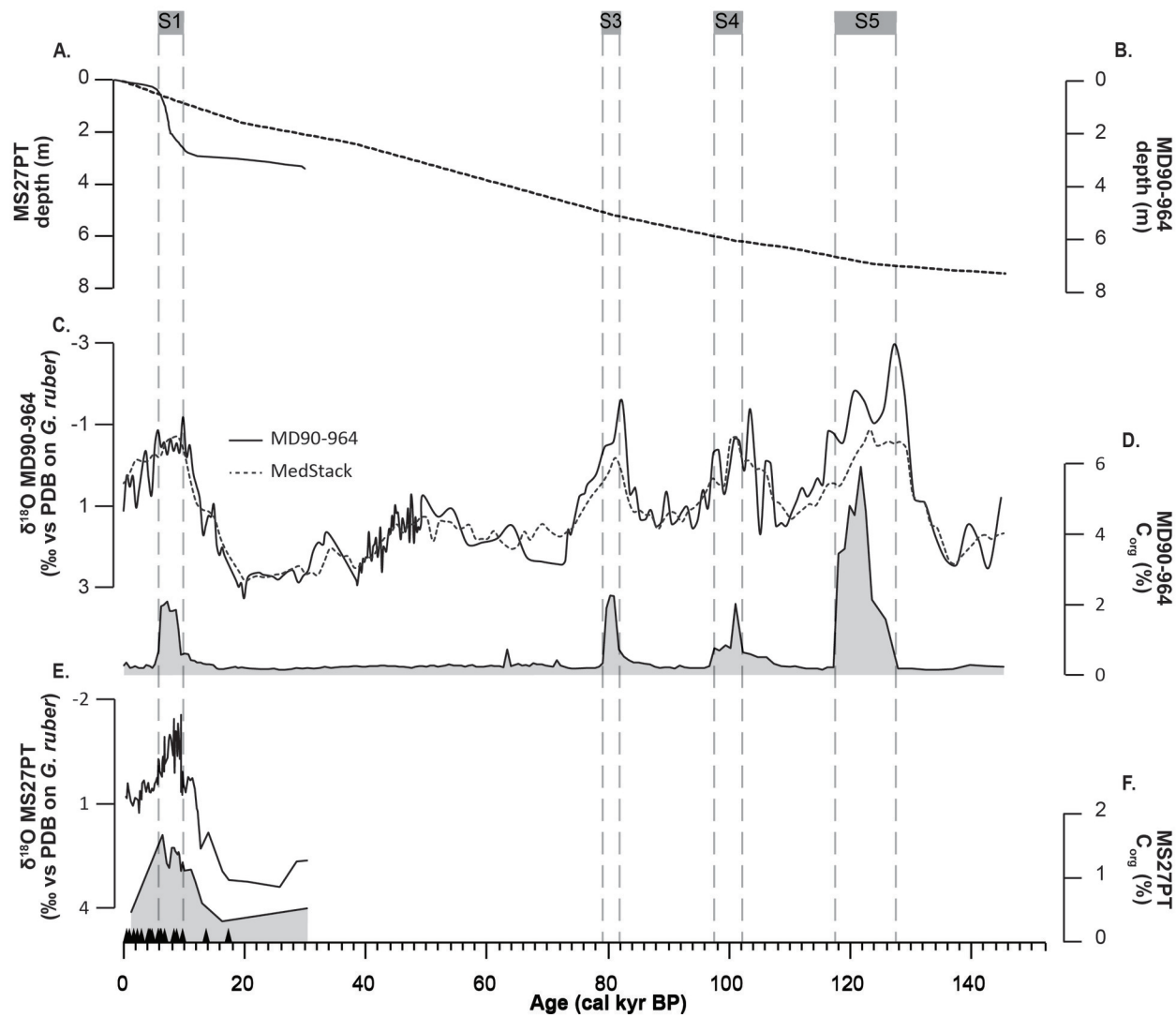
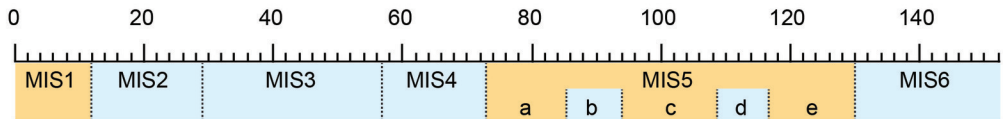
672	119.0	0.512397	$\pm 0.000010$	-4.70	$\pm 0.17$
676	119.9	0.512420	$\pm 0.000010$	-4.25	$\pm 0.20$
680	120.8	0.512405	$\pm 0.000011$	-4.54	$\pm 0.20$
684	121.7	0.512429	$\pm 0.000009$	-4.07	$\pm 0.22$
688	122.6	0.512436	$\pm 0.000009$	-3.94	$\pm 0.17$
692	123.5	0.512429	$\pm 0.000013$	-4.07	$\pm 0.18$
700	128.0	0.512399	$\pm 0.000011$	-4.66	$\pm 0.24$
704	130.3	0.512421	$\pm 0.000009$	-4.23	$\pm 0.21$
708	132.5	0.512418	$\pm 0.000012$	-4.29	$\pm 0.18$
712	134.8	0.512446	$\pm 0.000013$	-3.75	$\pm 0.23$
716	137.0	0.512483	$\pm 0.000008$	-3.02	$\pm 0.26$
720	139.8	0.512448	$\pm 0.000009$	-3.71	$\pm 0.16$
728	145.4	0.512454	$\pm 0.000011$	-3.59	$\pm 0.17$

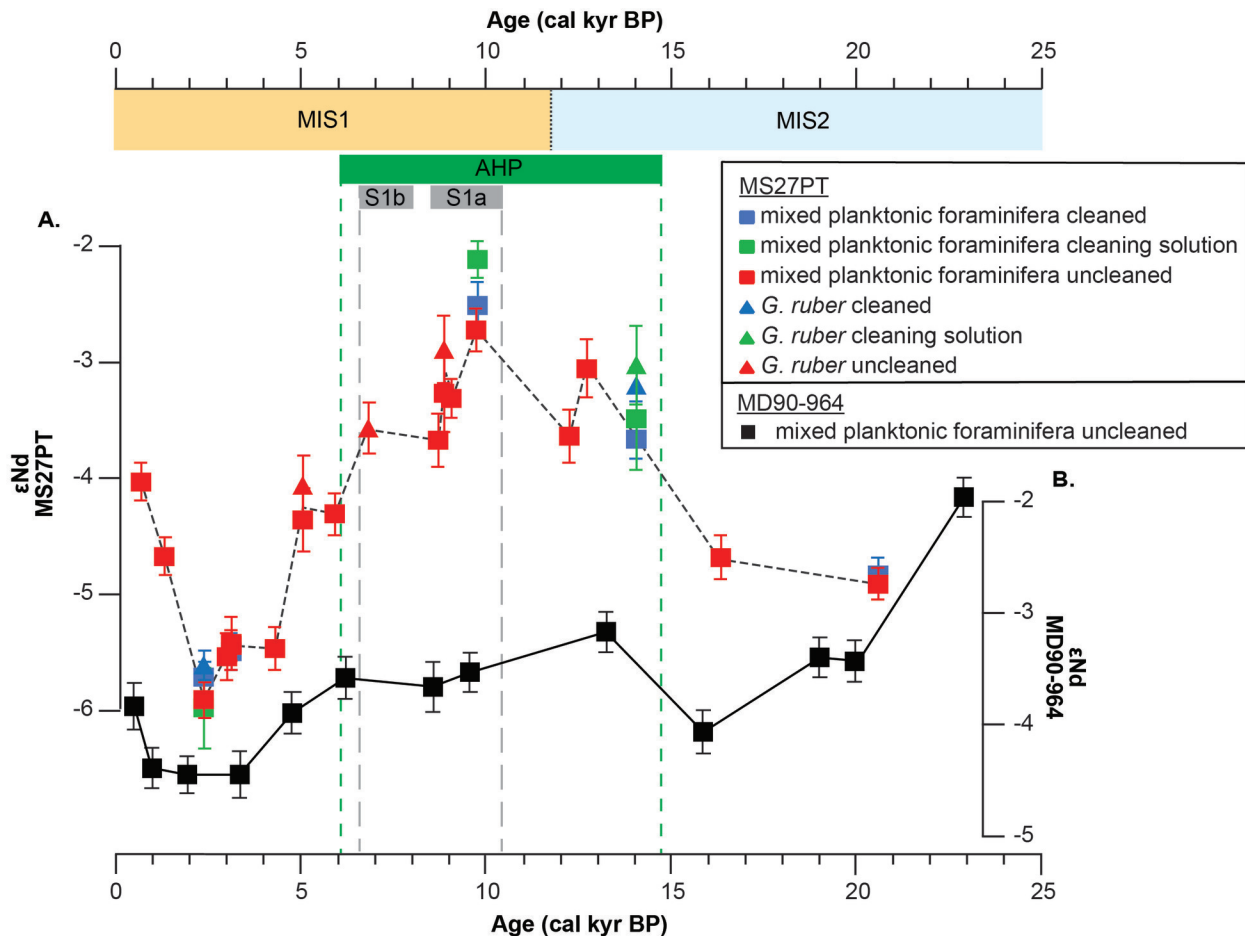
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*Table 2*

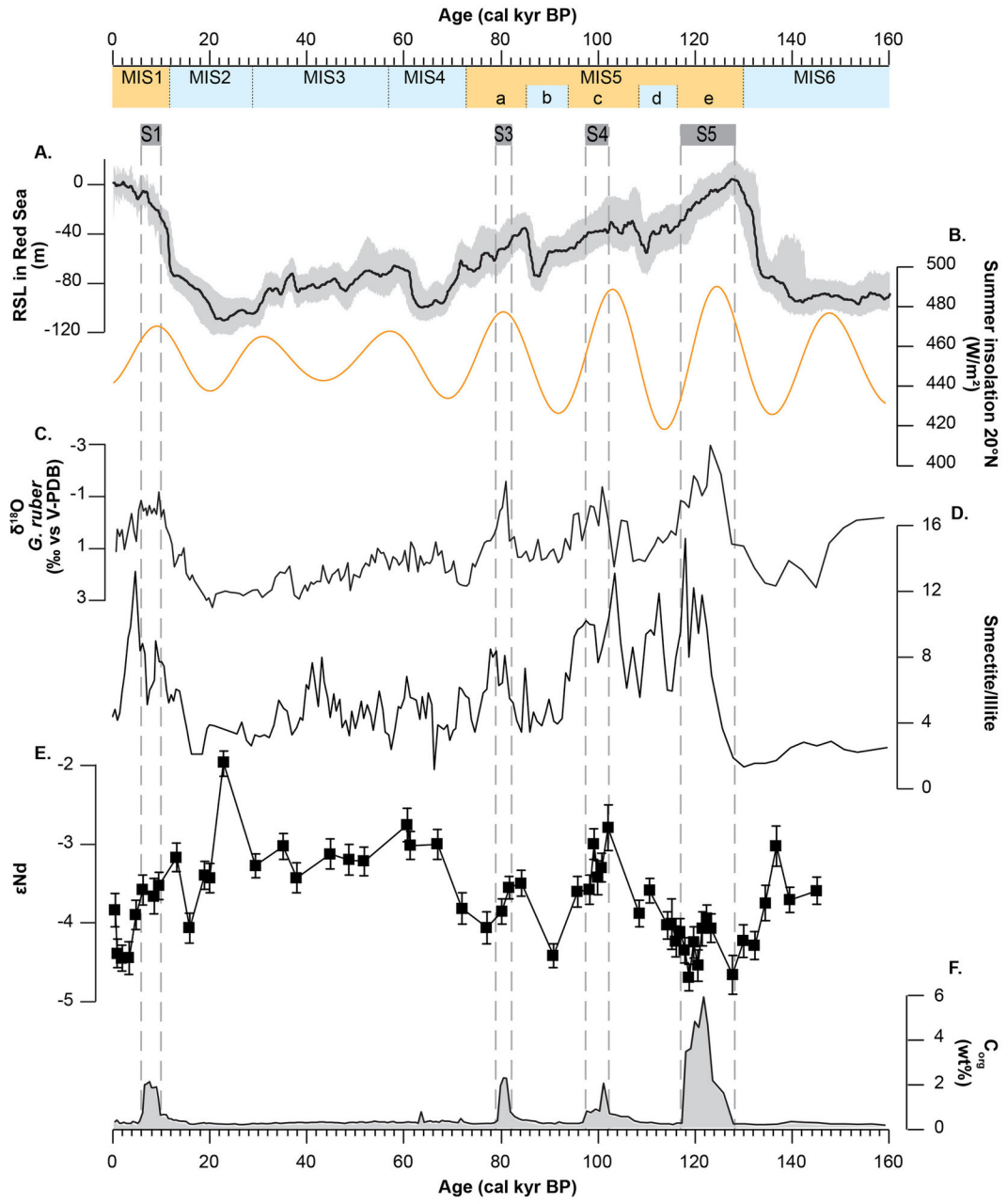


Age (cal kyr BP)



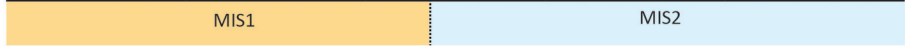




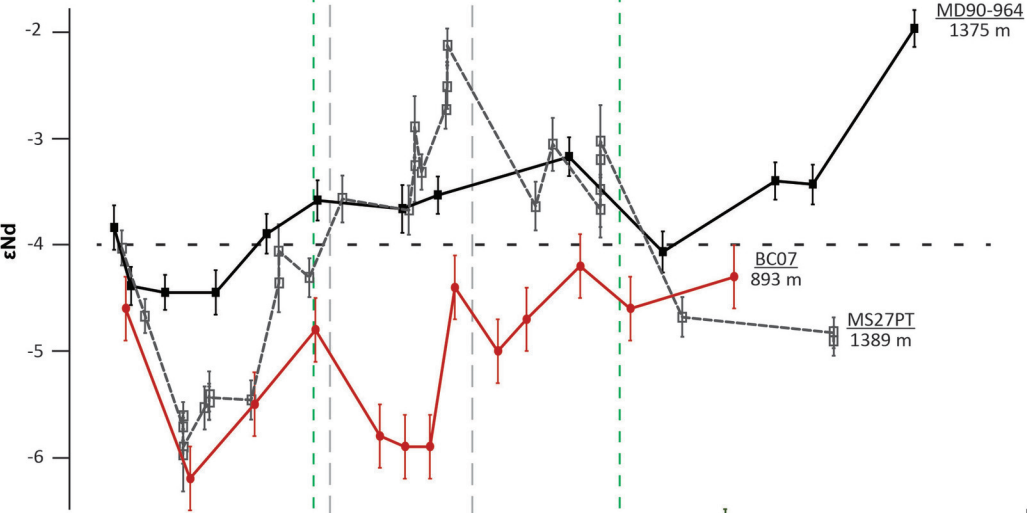


Age (cal kyr BP)

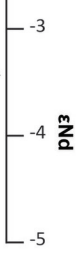
0 5 10 15 20 25



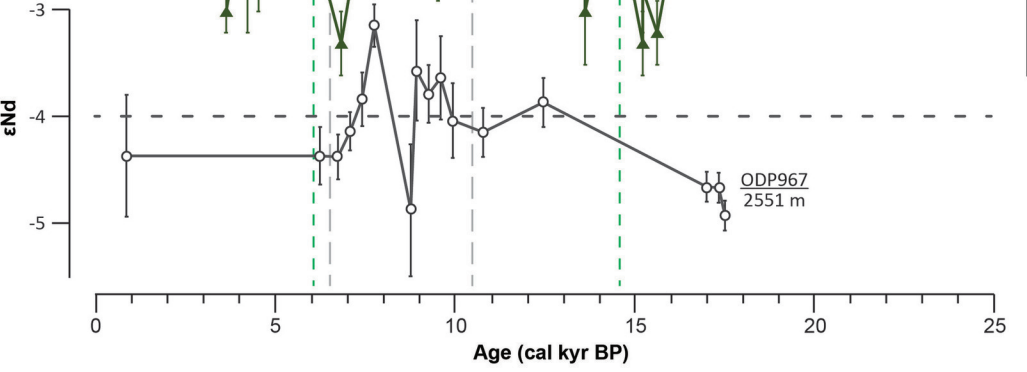
A.



B.



C.



Age (cal kyr BP)

0 5 10 15 20 25

