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### Precision efficiency calibration of a high-purity co-axial germanium detector at low energies

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#### Abstract

Following work done in the energy region above 100 keV, the high-precision calibration of a co-axial high-purity germanium detector has been continued in the energy region below 100 keV. Some of the previous measurements and Monte-Carlo simulations have been repeated with higher statistics and a new source measurement with <sup>169</sup>Yb has been added. For the energy range from 40 keV to 100 keV, an absolute precision for the detection efficiency of  $\pm 0.2\%$  has been reached, as previously obtained for energies above 100 keV. The low-energy behaviour of the germanium detector was further scrutinized by studying the germanium X-ray escape probability for the detection of low-energy photons. In addition, one experimental point, a  $\gamma$  ray at 2168 keV from the decay of <sup>38</sup>K, has been included for the total-to-peak ratios agreeing well with simulations. The same  $\gamma$  ray was also added for the single-and double-escape probabilities. Finally, the long term stability of the efficiency of the germanium detector was investigated by regularly measuring the full-energy peak efficiency with a precisely calibrated <sup>60</sup>Co source and found to be perfectly stable over a period of 10 years.

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Key words: gamma-ray spectroscopy, germanium detector, Monte-Carlo simulations PACS: 07.85.-m, 29.30.Kv, 23.20.-g

### 1. Introduction

Gamma-ray spectroscopy is a powerful tool for the study<sup>21</sup> of nuclear structure. In particular,  $\beta$ -decay studies are of-<sup>22</sup> 3 ten performed by measuring  $\gamma$ -ray intensities to deduce  $\beta$ - <sup>23</sup> 4 decay feeding probabilities. This applies in particular to <sup>24</sup> high-precision  $\beta$ -decay measurements as conducted for the <sup>25</sup> 6 purpose of weak-interaction studies. To determine the  $ft^{26}$ 7 value of super-allowed  $\beta$  decays of the  $0^+ \rightarrow 0^+$  type, high-<sup>27</sup> 8 precision measurements are required for the  $Q_{EC}$  value, the <sup>28</sup> 9 half-life, and the super-allowed branching ratio. The latter <sup>29</sup> 10 quantity is usually determined by measurements of the  $\gamma$ - <sup>30</sup> 11 12 decay probabilities of excited levels in the  $\beta$ -decay daugh-<sup>31</sup> ter nucleus, from which the  $\beta$ -decay feeding probabilities<sup>32</sup> 13 33 are determined. 14

Our group has a long standing experimental program <sup>34</sup> of measuring half-life values and branching ratios of these <sup>35</sup> super-allowed  $\beta$  decays [1–8] or of mirror decays [6,9,10], <sup>36</sup> where high-precision  $\gamma$ -ray spectroscopy is required. There- <sup>37</sup> fore, we have started in 2010 to precisely calibrate a highpurity (HP), n-type, co-axial germanium detector with a relative efficiency of 70%. The germanium crystal has an active volume of 278 cm<sup>3</sup> corresponding to about 1.48 kg of germanium. The most important part of this calibration work was published a few years ago [11]. In recent years, we continued this calibration work, in particular to improve the precision below 100 keV  $\gamma$ -ray energy, where we gave an absolute precision of only  $\pm 0.5\%$  in our previous work. This was mainly due to the fact that, at low energies, many background  $\gamma$  rays "pollute" the spectrum and we had not enough information to ascertain that we achieved the same precision as above 100 keV.

For a measurement recently performed with  $^{22}$ Mg, we need precise efficiencies in the energy range below 100 keV, which was the prime reason to extend the high-precision efficiency calibration of our detector down to 40 keV. In the present work, we describe measurements with a new source (<sup>169</sup>Yb). In addition, new measurements were performed with higher statistics as compared to our previous work with sources already used there (<sup>24</sup>Na, <sup>48</sup>Cr and

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<sup>40</sup> <sup>207</sup>Bi). Higher-statistics Monte-Carlo (MC) simulations
<sup>41</sup> have been performed for <sup>133</sup>Ba, <sup>152</sup>Eu and <sup>180</sup>Hf<sup>m</sup>. The
<sup>42</sup> new experimental data for <sup>24</sup>Na, <sup>48</sup>Cr, and <sup>169</sup>Yb were
<sup>43</sup> taken at ISOLDE, whereas the <sup>207</sup>Bi data come from a
<sup>44</sup> source measurement.

Table 1

 $^{169}\mathrm{Yb}$  and  $^{180}\mathrm{Hf}^m$  source characteristics used to determine the detector efficiency. The characteristics of the other sources are given in [11].

Nuclide	$T_{1/2}$	$E_{\gamma}$ (keV)	$P_{\gamma}$	Reference
$^{169}\mathrm{Yb}$	32.016 d	50.4	1.464(21)	[12]
		58.0	0.3855(81)	
		63.1	0.4405(24)	
		93.6	0.02571(17)	
		109.8	0.1736(9)	
		118.2	0.01870(10)	
		130.5	0.1138(5)	
		177.2	0.2232(10)	
		198.0	0.3593(12)	
		307.7	0.10046(45)	
$^{180}\mathrm{Hf}^{m}$	5.53 h	57.5	0.4799(94)	[13]
		93.3	0.1751(14)	[14]
		215.3	0.8150(15)	[14]
		332.3	0.9443(05)	[14]
		443.1	0.8180(130)	[14]
		500.7	0.1421(28)	[13]

We took also data at ISOLDE for <sup>49</sup>Cr, which, according 45 to literature, should have had well determined branching 46 ratios for  $\gamma$  rays between 62 keV and 153 keV, ideal to 47 connect the region above 100 keV already well calibrated in  $^{70}$ 48 our previous work and the region below 100 keV. However,<sup>71</sup> 49 it turned out that the branching ratios were not at all in 50 agreement with the  $\gamma$ -ray rates we observed. Therefore, 72 51 instead of using <sup>49</sup>Cr to calibrate our detector, we used 73 52 this latter to propose a new measurement of the branching 74 53 ratios of this isotope [15]. 54 75

The general outcome of the present work is that the de-76 55 tector model developed in our previous work is also valid 77 56 with the same precision below 100 keV. Only for data points 78 57 below 40 keV we seem to start having systematic discrep-79 58 ancies between experimental data and simulations. It is not 80 59 clear whether this is a problem with our detector model<sup>\$1</sup> 60 or rather linked to electronics issues with the trigger effi-82 61 ciency. We refrain from speculating about the reason for 83 62 this discrepancy and assume that the efficiency of our de-84 63 tector is precise at a level of  $\pm 0.2\%$  down to a  $\gamma$ -ray energy 85 64 of 40 keV. 86 65

The detector model used for the simulations in the s7 present work is the same as in [11]. It is shown in fig- s8 ure 1. The main characteristics of the detector are given in s9 table 2. 90



Fig. 1. Detector model as used in CYLTRAN or GEANT4 simulations. The colour code is as follows: red: germanium; violet: germanium dead material; blue: aluminum; pink: brass. The upper green material (difficult to see, on the top of the germanium material) corresponds to screws from the detector holding structure, which, in CYLTRAN, had to be "distributed" over a hollow cylinder. The other colours correspond to isolating material and teflon. The source sits for all measurements at the origin of the coordinate system.

#### Table 2

Detector characteristics used in the modelling of the germanium detector.

	parameters
length of crystal	78.10 mm
radius of crystal	$34.84~\mathrm{mm}$
length of central hole	68.5  mm
radius of central hole	$6.1 \mathrm{mm}$
external dead zone	$8.5 \ \mu m$
internal dead zone	2.2 mm
back-side dead zone	0.8 mm
distance window - crystal	$5.7 \mathrm{~mm}$
entrance window thickness	$0.7 \mathrm{~mm}$

## 2. Experimental results with $\gamma$ -ray sources and comparison with CYLTRAN simulations

In the present work, we have adopted the branching ratios for  ${}^{180}\text{Hf}^m$  proposed by Helmer *et al.* [14] and from ENSDF [13] in our analysis and in the decay scheme simulations, because we found them in much better agreement with our data than those used previously [11]. The branching ratios used in the present work for  $^{169}$ Yb and  $^{180}$ Hf<sup>m</sup> are given in Table 1. All other source characteristics are the same as presented in our previous paper. The commercial sources had activities on thin mylar foils such that the mylar itself had no influence on the  $\gamma$ -ray flux, even at very low energies. The sources produced on-line at ISOLDE were deposited with an energy of 30 keV at the surface of a small plastic cylinder (diameter 2 cm, thickness 5 mm). The high-energy  $\gamma$ -ray sources <sup>56</sup>Co and <sup>66</sup>Ga were produced by irradiating enriched <sup>56</sup>Fe (56 $\mu$ m thick) and <sup>*nat*</sup>Zn  $(5\mu m)$  targets with protons. All source measurements were performed under identical condition on a measurement setup with a source holder positioning the activities at exactly 15.000(5) cm from the entrance window of the germanium

91 detector.

As in our previous work [11], we obtained relative effi-146 92 ciencies for each  $\gamma$ -ray peak of each source by determining<sup>47</sup> 93 the number of counts in each photopeak and deviding this48 94 number by the branching ratio of the peak, the measure-149 95 ment time, and the dead time. The latter two can in prin-150 96 ciple be neglected, as these numbers are the same for alh<sub>51</sub> 97  $\gamma$ -ray peaks of a given source. The source activity is a free 98 parameter that allows to match the experimental data and 99 the MC simulations. Once the activity adjustment is done, 100 the difference of the individual experimental and MC effi-101 ciencies normalised by the experimental efficiencies is de-102 103 termined. This quantity is a measure for the agreement between experimental data and MC simulations and is plot-104 ted in figure 2 together with the absolute experimental ef-105 ficiencies as a function of the  $\gamma$ -ray energies. 106

The main errors come from the fit of the photopeaks and 107 the uncertainties with which the branching ratios of the 108 different sources are known. The peak integrals are deter-109 mined by a fit with a Gaussian plus a low-energy tail to take 110 incomplete charge collection into account and a background 111 function. This fit yields the number of counts in the peak 112 and its statistical uncertainty. To determine a systematic 113 uncertainty, we use three different background functions: i) 114 a second-order polynominal, ii) a linear step function (com-115 plement of an error function and a straight line), and (iii) 116 two independent linear functions left and right of the peak 117 "smoothly" connected under the peak. We chose for eachise 118 peak to be fit the two most appropriate background func-153 119 tions. The fit with the best  $\chi^2$  is the reference fit, whereas 54 120 we use the second best fit to determine a systematic un-155  $\,$ 121 certainty from the different integrals of the two fits. Morese 122 details can be found in our previous work [11]. 157 123

Part a) of figure 2 gives the absolute efficiency as as 124 function of energy established with 15 different radioactives9 125 sources, standard sources commercially available, but also 126 short-lived sources we produced at ISOLDE or IJCLab Or-161 127 say. In part b), we give the residuals between the experi-162 128 mental data and the MC simulations with the CYLTRANi63 129 code [16]. The vast majority of data lies within  $\pm 2\%$ . Theorem 130 residuals are consistent with no difference between the mea-165 131 surements and the simulations, which is expected as these 132 absolute normalisation is a free parameter for all sources67 133 except for the  $^{60}\mathrm{Co}$  source the activity of which has a pre-168 134 cision of 0.09%. To verify that the detection efficiency is 69 135 correctly described by our detector model, one has to in-170 136 spect the residuals for a given source and check that thereizi 137 is a zero difference between experimental and simulation<sub>172</sub> 138 data for low- and high-energy data points. 173 139

#### 140 2.1. Detection efficiency below 100 keV

141 A fine analysis of the residuals can give more insights<sub>78</sub> 142 about the precision of our detector model in different re-179 143 gions of energy. As a first step, we present in part c) and d)<sub>180</sub> 144 of figure 2 the residuals for  $\gamma$ -ray energies below and aboves1 500 keV, respectively. The data are the same as in part b) of the figure, in particular no change in normalisation was operated, whereas the fit is performed only on the respective parts of the data. As indicated in the figures, the residuals are in agreement with no difference between experiment and simulations. For the three residual plots, the  $\chi^2_{\nu}$  is close to unity.

Table 3

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Averages of residuals between the experimental efficiency and the simulated ones. For the present table we analysed different regions of energies and calculated weighted averages of the individual residuals.

Energy range	average residuals $(\%)$
$30$ - $500~{\rm keV}$	-0.031(099)
$30$ - $400~{\rm keV}$	-0.030(100)
$30$ - $300~{\rm keV}$	-0.188(125)
$30$ - $200~{\rm keV}$	-0.216(140)
$30$ - $100~{\rm keV}$	-0.401(288)
$30$ - $~60~{\rm keV}$	-0.593(462)
$40$ - $500~{\rm keV}$	0.003(100)
$40$ - $400~{\rm keV}$	-0.001(101)
$40$ - $300~{\rm keV}$	-0.148(127)
$40$ - $200~{\rm keV}$	-0.166(144)
$40$ - $100~{\rm keV}$	-0.197(316)
$40$ - $~60~{\rm keV}$	+0.032(614)

A more refined analysis is presented in table 3 for the low-energy part of our measurements. We selected different ranges of energies and determined whether or not the averages of the residuals are still in agreement with zero or not. Due to the energy selection, this analysis includes only part of the  $\gamma$  rays for the different sources. No change of the normalisation of source activities was made. If a deviation from zero is found, this means that for one or several sources less agreement is found e.g. for low energies.

This is the case for energies below 40 keV. Although the error bars increase significantly when using less and less data, there seems to be a general tendency of more and more negative average residuals for the low-energy part. This indicates that there is missing experimental decay strength, which could come from a problem with the trigger probability for these low  $\gamma$ -ray energies. Another possible explanation could be that the detector model we use is wrong with e.g. the detector entrance window or the detector dead layer being too thin. However, in order to match experiment and simulation, we would need to increase the window thickness by more than 50%, which is in contradiction with other measurements e.g. with an electron beam [11]. The dead layer would need to be increased by an even higher factor.

Therefore, we decided to limit our conclusions to energies larger than 40 keV. Above this energy we have very good agreement between model calculations and experiment. As we do not find any difference in agreement between the lowenergy part (40 - 100 keV) and the high-energy part treated in reference [11], we conclude that the absolute precision on

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Fig. 2. Full-energy efficiencies and residuals for the  $\gamma$ -ray detection efficiency for all multi- $\gamma$ -ray sources measured with the present detector. Compared to our previous publication [11], higher-statistics measurements have been performed for the <sup>24</sup>Na, <sup>48</sup>Cr and <sup>207</sup>Bi sources. As an additional source, <sup>169</sup>Yb has been added for its low-energy  $\gamma$  rays. Part a) shows the absolute efficiency of the germanium detector over the full range of 30 keV to 4 MeV. Part b) gives the residuals with respect to the MC calculations with the CYLTRAN code over the same energy range. In part c), we zoom on the energy range from 30 keV to 500 keV, whereas part d) shows the residuals in the region from 500 keV. The fit for this latter figure was made over the range 500 keV - 4000 keV. See text for details.

the efficiency of the germanium detector is  $\pm 0.2\%$  over the energy range of 40 keV to 4 MeV, with a relative efficiency precision of  $\pm 0.15\%$  over this energy range.

# 185 2.2. Total-to-peak ratio as well as single- and 186 double-escape ratios

As lead out in our previous work, the ideal source to determine total-to-peak ratios is an isotope with a single  $\gamma$ ray and no other decay radiation. <sup>38</sup>K comes close to that. It decays with a half-life of 7.65 min to basically a single level at 2167 keV (branching ratio of 99.8%) in <sup>38</sup>Ar, which subsequently emits a  $\gamma$  ray of this energy.

In figure 3, we have updated the total-to-peak ratio plot 193 for our detector by including this data point for  $^{38}$ K taken 194 at ISOLDE. This additional data point is of interest, as it is 195 much higher in energy than the other data used up to now. 196 The new experimental data point is in good agreement with 197 our detector model and confirms that matter around the 198 detector has to be added to correctly describe the total-to-201 199 peak ratios in our MC simulations (see [11] for details). 202 200



Fig. 3. Total-to-Peak ratio for  $\gamma$ -ray energies ranging from 122 keV to 2170 keV. The good agreement between experiment and simulation already obtained up to 1300 keV in our previous work is extended to 2170 keV.

In a similar way, we also included the  $^{38}$ K data point in the plot of the single-escape and double-escape probabili-

ties for the 511 keV annihilation radiation (figure 4). Good35
agreement is achieved between experiment and MC simu-236
lations.

full-energy single double / full-Mar V full symbols: experiment 0 mbols: simulation <sup>24</sup>Na 0.1 <sup>38</sup>K ratio Cr ₿Ga 0.0 0.001 1000 1500 2000 2500 3000 3500 4000 4500 energy (keV)

Fig. 4. Ratios of the single-escape (red) and double-escape (blue) intensities over the full-energy intensities from source measurements with  $^{24}$ Na,  $^{38}$ K,  $^{56}$ Co,  $^{66}$ Ga, and  $^{88}$ Y. Full symbols represent experimental ratios, open symbols, slightly shifted to facilitate the reading of the figure, are from simulations with the CYLTRAN code. The additional point for  $^{38}$ K fits nicely in the systematics.

## 3. X-ray escape probability for low-energy photon 239 detection

An effect, which becomes important for photon energies<sup>41</sup> 208 below 100 keV, is the possibility of germanium X-ray es-242 209 cape. At low energies, the interaction of a photon takes<sup>43</sup> 210 place at the very entrance of the germanium crystal. When<sub>244</sub> 211 the photon interacts by photo-electric effect or by Comp-245 212 ton scattering, vacancies are created in the electronic shells<sup>246</sup> 213 of the germanium atoms, which are filled by electrons from 247 214 higher-lying shells thus creating X-rays. These X-rays areas 215 of low energy (of order 10 keV) and are usually absorbed<sub>49</sub> 216 by the germanium crystal. However, the closer the photon<sub>50</sub> 217 interaction occurs to the surface of the germanium crystab51 218 (i.e. the lower the photon energy), the higher is the prob-252 219 ability of the X-rays to escape from the detector thus cre-253 220 ating an event with less energy. Therefore, for low-energy<sub>254</sub> 221 X-ray or  $\gamma$ -ray peaks, a small satellite peak about 10 keV<sub>255</sub> 222 below the main peak can be observed, the intensity of which 256 223 increases with decreasing energy. 257 224

Helmer and co-workers [14] found that their MC simu-258 lations underestimate this effect by a factor of 1.16. After 559 scaling the escape probability by this factor, they found 50 satisfying agreement between simulations and experimen-261 tal measurements. The authors argued that the assumption 562 of a uniform detector surface as well as of complete charge 563 collection in the surface area might be questionable. 264

As we demonstrate in figure 5, we find agreement over265 the full energy range from 30 keV to 120 keV between the simulated escape probability and the experimentally deter-267



Fig. 5. X-ray escape probability for low-energy photons as measured with the germanium detector and simulated with the CYLTRAN code. The symbols give the experimental data points, whereas the full line stems from CYLTRAN simulations.

## 4. Long-term stability of the absolute efficiency with $^{60}\mathrm{Co}$ sources

Our germanium detector is kept cold all year long since more than 10 years (except e.g. for travel to experimental sites in Jyväskylä or at ISOLDE). The reason for that is to avoid that the lithium used for the p contacts diffuses in or out of the detector and thus changes the efficiency of the detector.

In order to verify the efficiency of our detector, we performed several measurements over the years with a  $^{60}$ Co source, which is calibrated in activity with a precision of better than 0.1%. Data were last taken with this source at the end of 2019. As  $^{60}$ Co has a half-life of 5.271 y, the activity has now decreased too much to still use this original source.

We have therefore calibrated a new  $^{60}$ Co source with a precision of 0.14%, which will be used in the future. This source has the same mounting as the old source. The calibration was performed with respect to the old precision source in several runs during 2018. A first verification of the detector efficiency was performed with the new source beginning of 2020 yielding agreement with the initial efficiency (see figure 6).

All efficiency values determined over the last 10 years are in agreement with each other and with the CYLTRAN simulations used as a reference. The numerical values of the efficiencies determined for the two  $\gamma$  rays of <sup>60</sup>Co are given in table 4. We can thus conclude that, when keeping the detector cold, a germanium detector can have an efficiency stable over periods of 10 years and more.

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mined one. No scaling factor is needed. We conclude therefore that, for our detector, CYLTRAN treats correctly this X-ray escape probability.



Fig. 6. Long-term stability of the absolute detection efficiency of the germanium detector. The efficiency is stable over a period of  $^{94}$ 10 years. The last data point has been measured with a new  $^{60}\mathrm{Co}$ source calibrated with respect to our old source in August-October 2018 by means of the germanium detector presently described. This source, with an activity of 19.948(28) kBq on August 2, 2018, has<sup>296</sup> a precision on its activity ( $\Delta A/A = 0.14\%$ ) slightly worse than our previous source ( $\Delta A/A = 0.09\%$ ). The precision of the last data point in the plot is limited by the detection statistics. The full lines  $^{297}$ represent the calulated efficiencies from the CYLTRAN code with the dashed lines being the error bars. 298

#### Table 4

Efficiency values as determined in different measurements over the  $h_{eff}^{300}$ Ellicitie y values as determined in dimensional  $^{301}$  last 10 years for the two  $\gamma$  rays from  $^{60}$ Co. The data taking in  $^{302}_{302}$ 2014 corresponds to a measurement where the method of coincident summing with a strong  $^{60}$ Co source of unknown activity was used  $^{303}_{304}$ (see [11]). The 2020 measurement was performed with a new source (see [11]). The 2020 measurement was performed  $^{305}$  calibrated with respect to our old high-precision  $^{60}$ Co source (see  $^{305}_{306}$ text) 307

year	$\epsilon(1173~{\rm keV})~(\%)$	$\epsilon(1332~{\rm keV})~(\%)$
2011	0.21753(18)	0.19966(13)
2012	0.21751(19)	0.19949(16)
2013	0.21763(42)	0.19976(42)
2014	0.21860(70)	0.19960(70)
2018	0.21699(48)	0.19950(44)
2018	0.21764(31)	0.19998(29)
2019	0.21757(34)	0.19901(35)
2020	0.21750(70)	0.19930(64)

#### 5. Conclusions 268

We have continued the efficiency calibration of a germa-269 nium detector for high-precision  $\gamma$ -ray spectroscopy. This 270 detector is used primarily for branching-ratio measure-271 ments of super-allowed  $0^+ \rightarrow 0^+$  decays where precisions 272 of the branching ratios of the order of 0.2% or better are 273 needed. With the present study, we achieved a precision 274 of the absolute detection efficiency for energies down to 275 40 keV of 0.2% and of the relative detection efficiency of 276 0.15%. 277

During recent years, we had access to short-lived sources produced at ISOLDE, which we used to add in particular measurements below 100 keV. We thus added  $^{169}$ Yb as a new source, which has well-known  $\gamma$  rays between 50 keV and 300 keV. In addition, we improved previous measurements by increasing the detection statistics or by increasing the statistics of our MC simulations.

Measurements and simulations of the X-ray escape probability as a function of photon energy allowed us to futher confirm the good understanding of the low-energy response of our detector.

Finally, we added a high-energy source, <sup>38</sup>K. This isotope has a single high-energy  $\gamma$  ray which allowed us to verify the total-to-peak ratio above 2 MeV. This source also yielded an additional data point for the single- and double-escape probability for annihilation radiation.

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