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Ibn Ishāq and the Alfonsine Tables¹ (5 Jan)

Julio Samsó

Abstract

This paper seeks to show that astronomical tables (*zīj*es) belonging to the tradition of Ibn Ishāq al-Tunīsī (fl. 1193-1222) were available to the astronomers in the service of King Alfonso X (r. 1252-1284), who used them both in their astronomical works written in Castilian and for the production of the Parisian *Alfonsine Tables* which are, nowadays, extant in Latin.

Keywords: Ibn Ishāq, Alfonso X, *Alfonsine Tables*, astronomical tables, obliquity of the ecliptic, declination tables, mean solar, lunar and planetary motions.

No astronomical tables produced in the Maghrib seem to be extant² until the thirteenth and fourteenth centuries, during which an astronomical renaissance, strongly influenced by the Andalusian tradition, took place. It shared many of its characteristics and was its obvious continuation at a time in which Islamic astronomy in the Iberian Peninsula entered a period of decline. This can be seen through the appearance of an Andalusio-Maghribī group of *zīj*es which begins with the unfinished *zīj* of Abū l-‘Abbās ibn Ishāq al-Tamīmī al-Tūnīsī (fl. ca. 1193-1222). There is no biographical information about him, although we know that he lived in Tunis and Marrakech and that some of his tables were based on observations made in 619H/1222 CE.

The standard reference to Ibn Ishāq is a passage in Ibn Khaldūn’s (1332-1382) *Muqaddima* in which it is said that “the people of our times use the *zīj* attributed to Ibn Ishāq, the Tunisian astronomer of the beginning of the 7th /13th century”, based on his own observations and on the information transmitted by an unnamed Sicilian Jew.³ An analysis of the extant materials establishes that what he obtained from his Sicilian correspondent derived mainly from Andalusian astronomical literature.⁴ It seems clear that Ibn Ishāq compiled a set of tables for the computation of planetary longitudes, eclipses, equation of time, parallax and, probably, solar and lunar velocity. This is confirmed by the prologue to the *Minhāj* of Ibn al-Bannā’ (1256-1321), where the author states that his book is based on Ibn Ishāq’s *zīj*⁵ and that he discovered it in a collection of tables, noted down on cards, on which he based his [mean] motions of the planets and their equations.⁶ These tables were not accompanied by an elaborate collection of canons although they contained some kind of instructions for the use of a few tables. Ibn Ishāq’s original *zīj* seems to be lost, although quite a lot of information about it can be gathered from a set of five “recensions”

¹This research has been undertaken within ALFA, a European Research Council project (Consolidator grant 2016 agreement no. 723085) funded for 2017-2022.

²The well known astrologer Ibn Abī l-Rijāl (d. after 1038) apparently compiled a *zīj*, entitled *Hall al-‘aqad wa bayān al-raṣad* (“Untying of knots and explanation of observations”) which does not seem to be extant.

³Ibn Khaldūn, *Muqaddima*, ed. M. Quatremère, vol. III (Paris, 1858), pp. 107-8.

⁴Angel Mestres, “Maghribī Astronomy in the 13th Century: a Description of Manuscript Hyderabad Andhra Pradesh State Library 298”, in J. Casulleras & J. Samsó (eds), *From Baghdad to Barcelona. Studies in the Islamic Exact Sciences in Honour of Prof. Juan Vernet* (Barcelona, University of Barcelona, 1996), pp. 383-443.

⁵Ibn Khaldūn adds, in the aforementioned passage of the *Muqaddima*, that “Ibn al-Bannā’ summarized it in another [*zīj*] called «the method» (*al-Minhāj*)”.

⁶Juan Vernet, *Contribución al estudio de la labor astronómica de Ibn al-Bannā’* (Tetuán, 1952), pp. 13 (Arabic text) and 57 (Spanish translation).

of this work made in the Maghrib in the second half of the 13th and beginning of the 14th century. These recensions were prepared by three Maghribī astronomers:

- 1) An anonymous Tunisian astronomer who prepared, ca. 1266-1281, the recension extant in ms. Hyderabad Andra Pradesh State Library 298.⁷
- 2) Ibn al-Bannā' of Marrakech (1256-1321) in his extremely popular *Minhāj*.⁸
- 3-5) Recensions prepared by the Andalusī-Maghribī astronomer Ibn al-Raqqām (d. 1315): *al-Zīj al-Mustawfī* (Tunis, after 1280-81),⁹ *al-Zīj al-Shāmil* (Bougie, ca. 1290)¹⁰ and *al-Zīj al-Qawīm* (Tunis, after 1280-81, revised in Granada).¹¹ This scholar was an astronomer of Andalusian origin who worked most of his life in the Maghrib but established himself in Granada sometime after 1288-89. With the *Qawīm zīj*, which was used in that city, we see the return to al-Andalus of astronomical materials that had a clear Andalusian origin.

The derivation of these sources from Ibn Ishāq's *zīj* can be established due to the fact that the five works share the same sidereal mean motion and equation tables for computation of planetary longitudes which, according to Ibn al-Bannā', derived from Ibn Ishāq's original *zīj*. They all follow the ideas of Ibn al-Zarqālluh/Azarquiel (d. 1100) on trepidation, cyclical variation of the obliquity of the ecliptic, motion of the solar apogee, and correction of the Ptolemaic lunar model.

It seems clear that the Alfonsine astronomers had access to some version of Ibn Ishāq's *zīj*.¹² This can easily be proved from evidence furnished by two small treatises contained in the *Libro del saber de astrología*: the *Cuadrante para rectificar* and the *Libro del reloj de la piedra de la sombra*. The author of the *Cuadrante*, Ishāq b. Sīd, mentions a value of the obliquity of the ecliptic (ϵ) of 23;32,29° "for our time" as well as declination values for 30° and 60°.¹³

30°	11;31,11°
60°	20;14,13° (rec. value 20;14,14°)

⁷ Mestres, "Maghribī Astronomy" and his unpublished PhD dissertation entitled *Materials andalusins en el zīj de Ibn Ishāq al-Tūnisī*, presented at the University of Barcelona in 1999. It contains an edition of all the canons, as well as of quite a number of numerical tables.

⁸ Vernet, *Contribución*; J. Samsó & E. Millás, "Ibn al-Bannā', Ibn Ishāq and Ibn al-Zarqālluh's Solar Theory" in Samsó, *Islamic Astronomy and Medieval Spain* (Aldershot, Variorum, 1994), no. X; J. Samsó & E. Millás, "The Computation of Planetary Longitudes in the *zīj* of Ibn al-Bannā'", *Arabic Sciences and Philosophy* 8 (1998), 259-86 (repr. in Samsó, *Astronomy and Astrology in al-Andalus and the Maghrib* (Aldershot, Ashgate Variorum, 2007), no. VIII).

⁹ J. Samsó, "Ibn al-Raqqām's *al-Zīj al-Mustawfī* in MS Rabat National Library 2461", in Nathan Sidoli and Glen Van Brummelen (eds), *From Alexandria through Baghdad. Surveys and Studies in the Ancient Greek and Medieval Islamic Mathematical Sciences in Honor of J.L. Berggren* (Berlin-Heidelberg, Springer, 2014), pp. 297-328.

¹⁰ E.S. Kennedy, "The Astronomical Tables of Ibn al-Raqqām, A Scientist of Granada". *Zeitschrift für Geschichte der arabisch-islamischen Wissenschaften* 11 (1997), 35-72; M. 'Abd al-Rahmān, *Hisāb atwāl al-kawākib fī l-Zīj al-Shāmil fī tahdhīb al-Kāmil li-Ibn al-Raqqām*, unpublished Ph.D. thesis presented at the University of Barcelona, 1996.

¹¹ Kennedy, "Astronomical Tables".

¹² Some source belonging to this tradition was available in Toledo towards the middle of the 14th c.: see M. Castells, "Una tabla de posiciones medias planetarias en el *zīj* de Ibn Waqār (Toledo, ca. 1357)", in Casulleras & Samsó, *From Baghdad to Barcelona*, pp. 445-452.

¹³ Ms. Villa Amil 156 fol. 171 v (Biblioteca de la Universidad Complutense, Madrid); M. Rico y Sinobas, *Libro del Saber de Astronomía del Rey D. Alfonso X de Castilla, copilados, anotados y comentados por...* 5 vols. (Madrid, 1863-67), 3: 296. 23;33° is the value of the obliquity of the ecliptic mentioned in the Castilian canons of the Alfonsine tables: see J. Chabàs & B.R. Goldstein, *The Alfonsine Tables of Toledo* (Dordrecht, Kluwer, 2003), p. 60.

This value of ϵ reappears, rounded into $23;32,30^\circ$, in the *Libro del relogio*, where we find a complete declination table¹⁴ with this maximum value, which seems to be related to Ibn al-Zarqālluh’s model for the computation of ϵ , although with Ibn Ishāq’s parameters. In it, the pole of the ecliptic rotates around a polar epicycle with radius $0;10^\circ$. The centre of this epicycle is kept, in a fixed position, on a parallel of declination, whose distance from the pole of the equator is $23;42,30^\circ$. As a result, the value of ϵ will reach a maximum of $23;42,30^\circ + 0;10^\circ = 23;52,30^\circ$, and a minimum of $23;42,30^\circ - 0;10^\circ = 23;32,30^\circ$.¹⁵

Using the mean motion tables that regulate the position of the pole of the ecliptic in the polar epicycle, one can easily obtain $23;32,30^\circ$ for the end of 1193, the year in which the beginning of Ibn Ishāq’s astronomical activity has been established. In spite of this, it seems that this parameter might have an observational origin. According to Abū ‘Abd Allāh Muḥammad al-Ḥabbāk (d. after 1562), Ibn Ishāq himself stated that $23;32,30^\circ$ had been obtained by an anonymous astronomer from Miknāsa (Meknès, Morocco) through an observation made in 602/1205-6.¹⁶ This obviously means that Ibn Ishāq believed that the obliquity of the ecliptic had reached its minimum value in his own time and it implies that the declination table of the Alfonsine treatise on the sundial (*Libro del relogio*) is due to Ibn Ishāq himself and that it belongs to his primitive *zīj*. In 1992 Mercè Comes remarked that the solar declination table (no. 54 of the Hyderabad ms.) of the anonymous Tunisian recension of Ibn Ishāq’s *zīj* is the same as the one appearing in the Alfonsine sundial treatise, as both share the same small computational errors.¹⁷ Finally, it is interesting to note that the same parameter appears in al-Sanjufīnī’s *zīj*, compiled in Tibet in 1366.¹⁸

The influence of Ibn Ishāq’s *zīj* is not limited to the Castilian Alfonsine texts but can also be found in the Parisian *Alfonsine Tables* (*PAT*), especially in the mean motion parameters explicitly given in all the editions. It is interesting to compare these parameters to those underlying the sidereal tables of the tradition of Ibn Ishāq (Tunisian recension, Ibn al-Bannā’, Ibn al-Raqqām).¹⁹ Those corresponding to the mean motions in anomaly of Venus and Mercury (not affected by precession) are practically identical, although the same thing cannot be said about the mean lunar motion in anomaly (see Table 1). The mean motions in longitude of the Sun, Moon and planets show similar differences which vary between $0;0,0,7,31^\circ$ and $0;0,0,8,34^\circ$. Previous scholarship has not noticed this relation between the mean motions in Ibn Ishāq and the *PAT*.

Planet	Mean motion (PAT)	Mean motion (Ish)	PAT – Ish
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¹⁴ Ms. Villa Amil 156 fol. 178r; Rico, *Libros*, 4: 6.

¹⁵B.R. Goldstein, “On the Theory of Trepidation According to Thābit b. Qurra and al-Zarqālluh and its Implications for Homocentric Planetary Theory”, *Centaurus* 10 (1964), 232-247.

¹⁶Rachid Saidi, *Natā’iy al-afkār fī šarḥ Rawḍat al-Azhār de Abū ‘Abd Allāh Muḥammad al-Ḥabbāk. Comentario de la ur̄yūza Rawḍat al-azhār fī ‘ilm waqt al-layl wa l-nahār de Abū ‘Abd Allāh al-Īdirī. Edición y estudio*, unpublished Ph.D. thesis presented at the University of Barcelona, 2013, available at <http://hdl.handle.net/2445/62864>. See p. 75.

¹⁷Mercè Comes, “À propos de l’influence d’al-Zarqālluh en Afrique du Nord : l’apogée solaire et l’obliquité de l’écliptique dans le *zīj* d’Ibn Ishāq”, *Actas del II Coloquio Hispano-Marroquí de ciencias históricas*, Madrid, Instituto de Cooperación con el Mundo Árabe, 1992, pp. 147-159, see pp. 152-3. Reprinted in Mercè Comes, *Coordenadas del cielo y de la tierra* (Barcelona, Universidad de Barcelona, 2013), pp. 139-152, see pp. 144-5.

¹⁸E.S. Kennedy, “Eclipse Predictions in Arabic Astronomical Tables Prepared for the Mongol Viceroy of Tibet”, *Zeitschrift für Geschichte der arabisch-islamischen Wissenschaften* 4, 60-88. Reprinted in Kennedy, *Astronomy and Astrology in the Medieval Islamic World* (Aldershot, Variorum, 1998); E.S. Kennedy & J.P. Hogendijk, “Two Tables from an Arabic Astronomical Handbook for the Mongol Viceroy of Tibet” in Eric Leichty (ed.), *A Scientific Humanist. Studies in Memory of Abraham Sachs* (Philadelphia: University Museum, 1988), pp. 233-242. Reprinted in Kennedy, *Astronomy and Astrology*, no. XIV.

¹⁹ These mean motion parameters have been calculated independently by ‘Abd al-Rahmān (Ibn al-Raqqām) and Mestres (Tunisian recension), in their PhD theses, as well as by Samsó & Millás (Ibn al-Bannā’). See Samsó & Millás, “Ibn al-Zarqālluh’s Solar Theory” pp. 3-6; Samsó & Millás, “Planetary Longitudes” p. 262.

Sun	0;59,8,19,37,19,13,56°	0;59,8,11,28,26,22,5°	0;0,0,8,8,52,51,51°
Moon long.	13;10,35,1,15,11,4,35	13;10,34,52,46,51,38	0;0,0,8,28,19,26,35
Moon anom.	13;3,53,57,30,21,4,13	13;3,53,56,17,52,4	0;0,0,1,12,29,0,13
Node	0;3,10,38,7,14,49,10	0;3,10,46,41,0,4	-0;0,0,8,33,45,14,50
Saturn	0;2,0,35,17,40,21	0;2,0,27,46,42,42	0;0,0,7,30,57,39
Jupiter	0;4,59,15,27,7,24	0;4,59,7,36,25,11	0;0,0,7,50,42,13
Mars	0;31,26,38,40,5	0;31,26,31,9,5	0;0,0,7,31
Venus anom.	0;36,59,27,23,59,31	0;36,59,27,23,59,25	0;0,0,0,0,0,6
Mercury anom.	3;6,24,7,42,40,52	3;6,24,7,42,40,49	0;0,0,0,0,0,3

Table 1: Mean motions of the Parisian Alfonsine Tables and Ibn Ishāq's *zīj*.

The correction of, approximately, 0;0,0,8° for the mean motions in longitude corresponds to the value of precession which allows the conversion of mean sidereal motions into mean tropical motions. We can easily compute this value using the *PAT* which combine constant precession and trepidation:

According to the *PAT*, the argument of trepidation, and hence trepidation itself, attained 0° on 16 May 16 CE (JDN 1,727,038).²⁰ As the Alfonsine epoch used in the *PAT* is the 31 May 1252 (JDN 2,178,502):

$$2,178,502 - 1,727,038 = 451,464 \text{ days}$$

- 1) *PAT* constant precession: 0;0,0,4,20,41,17,12°/ day
 $0;0,0,4,20,41,17,12^\circ \cdot 451,464 = 9;4,52^\circ$
- 2) *PAT* periodic term: the position of the Head of Aries on 31 May 1252 is:
 $0;0,0,30,24,49,0^\circ \cdot 451,464 = 63;34,4^\circ$
 $\Delta\lambda = \arcsin(\sin 63;34,4^\circ \cdot \sin 9^\circ) = 8;3,9^\circ$
 ($\Delta\lambda$ being the increase of longitude due to the trepidation component)
- 3) Addition of the two terms:
 $\Delta\lambda' = 9;4,52^\circ + 8;3,9^\circ = 17;8,1^\circ$,
 ($\Delta\lambda'$ being the full increase in precession)
- 4) Daily increase of $\Delta\lambda'$:
 $17;8,1^\circ : 451,464 = 0;0,0,8,12^\circ/\text{day}$

It seems, therefore, that the Alfonsine tropical mean motions derive from the tradition of Ibn Ishāq which, as we have seen, is also the source of the declination table.²¹ This is, in my opinion, extremely important, as it provides further evidence to dismiss Poulle's hypothesis,²² according to which the Latin Alfonsine Tables were totally independent of the work undertaken by the astronomers of King Alfonso X and were actually produced by astronomers (John of Lignères, John of Murs and John of Saxony) in Paris between 1320 and 1330. It seems difficult to conceive that Ibn Ishāq's *szīj* could have been available to the Parisian astronomers. The fact that most Alfonsine mean motions derive clearly from Ibn Ishāq's parameters shows the Toledan origin of the Parisian Alfonsine mean motion tables.

²⁰J. Samsó & F. Castelló, "An Hypothesis on the Epoch of Ptolemy's Star Catalogue According to the Authors of the Alfonsine Tables", *Journal for the History of Astronomy* 19 (1988), pp. 115-120. Reprinted in Samsó, *Islamic Astronomy and Medieval Spain*, no. XX.

²¹The same dependance cannot be established by a comparison of the mean positions and longitudes of the solar and planetary apogees at epoch.

²²Poulle's hypothesis was contested by Chabás & Goldstein in their book on *The Alfonsine Tables of Toledo*. See Emmanuel Poulle, "Les Tables Alphonsines sont-elles d'Alphonse X?", in Mercè Comes, Roser Puig & Julio Samsó (eds), *De astronomia Alphonsi Regis* (Barcelona, University of Barcelona, 1987), pp. 51-69; Poulle, "The Alfonsine Tables and Alfonso X of Castille" *Journal for the History of Astronomy* 19 (1988), pp. 97-113. Repr. in E. Poulle, *Astronomie planétaire au Moyen Âge latin* (Aldershot, Variorum, 1996), no. V.