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### **Quantum Tunneling of Oxygen Atoms on Very Cold Surfaces**

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Any evolving system can change state via thermal mechanisms (hopping a barrier) or via quantum tunneling. Most of the time, efficient classical mechanisms dominate at high temperatures. This is why an increase of the temperature can initiate the chemistry. We present here an experimental investigation of O-atom diffusion and reactivity on water ice. We explore the 6–25 K temperature range at submonolayer surface coverages. We derive the diffusion temperature law and observe the transition from quantum to classical diffusion. Despite the high mass of O, quantum tunneling is efficient even at 6 K. As a consequence, the solid-state astrochemistry of cold regions should be reconsidered and should include the possibility of forming larger organic molecules than previously expected.

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Introduction.—Nuclear decay or chemical reactions may be described as the crossing through a potential barrier by quantum tunneling, or crossing over the same barrier by thermal hopping. Except for barrierless reactions, increasing the temperature initiates the chemistry. Actually, the quantum tunneling regime (like in nuclear decay) and the thermal activation (like in chemistry) are usually separated by orders of magnitude in temperature. Theoretically, the balance between classical thermal motion and quantum tunneling is a very active subject, especially because it impacts solid-state chemistry [1]. It is well established that a critical temperature exists below which tunneling is dominant [2,3]. Experimentally, such evidence is still missing in handbooks. Thanks to field ion microscopy, diffusion of single atoms on metals or on crystalline surfaces has been studied in detail for decades [4]. However, on amorphous surfaces, and especially on water ice substrates, the study of physisorbed light atoms presents enormous difficulties for any atomic microscopy technique that the field is still nearly unexplored. Yet, in the cold regions of the Universe, where temperatures are lower than 8 K [5], a rich chemistry is initiated on the surfaces of minuscule dust particles [6]. The species weakly bound to the surface are the pivotal media of this pristine chemistry, governed by the diffusion of reactive species [7]. So far, diffusion has only been partly explored experimentally for H atoms [8,9] and the role of amorphous structures in the diffusion properties is still an open question [10,11]. Nevertheless, the mobility of physisorbed species is key for the evolution

of the molecular complexity [12]. If species like O atoms freeze out on the surface of the grains, the chemistry is governed by H additions, leading to numerous saturated species (H<sub>2</sub>O, NH<sub>3</sub>, and CH<sub>4</sub>, which are chemical traps and end the chemical evolution). On the other hand, if other atoms (O, N, C,...) are mobile enough at low temperature, other additions may open up the field of the observed molecular complexity reached in the first stages of star formation, and that could lead to the formation of the building blocks of life (amino acids). Experimentally, very few studies have already involved physisorbed O atoms [13–15], whereas theoretically, calculations exist for ordered substrates such as graphite [16,17]. On cold surfaces, ozone reactivity has been the subject of experimental investigations for astrochemical [18,19] and atmospheric purposes [20]. The formation of  $O_2$  and  $O_3$  on amorphous silicates has been addressed recently [21] and will also be the subject of one of our future papers. We present here experimental evidence of the tunneling of physisorbed O atoms on different substrates (amorphous and crystalline water ice) via the study of ozone formation, and discuss the role of the morphology of the substrate in the quantum diffusion process.

Experiments and results.—Experiments have been performed using the FORMOLISM setup (described elsewhere [22,23]), an ultrahigh vacuum chamber coupled to a triply differentially pumped O beam aimed at the temperature controlled water ice samples. O atoms are obtained by dissociating  $O_2$  gas in a microwave discharge. The

dissociation fraction  $\mu$  is 71%. It corresponds to depositing 3 O<sub>2</sub> molecules and 14 O atoms. We have checked that atoms and molecules relax before adsorption by scanning the kinetic energy of the ionizing electrons of the mass spectrometer head intercepting the beam [24]. We detected no species with residual internal energy, i.e., having a ionization threshold below that of the ground state. No O<sub>3</sub> was present in the beam either. The beam flux was calibrated using temperature-programed desorption (TPD) by determining the O<sub>2</sub> exposure time required to saturate the  $O_2$  monolayer (ML) [25]. In this work, the exposures are expressed in terms of O<sub>2</sub> units, which means that 1 ML may also represent two layers of O atoms or 0.66 layers of pure ozone. The compact amorphous solid water (ASW) substrate was grown by vapor deposition on a 110 K substrate. We have also studied crystalline ice, made from an ASW substrate annealed up to the phase transition temperature around 140 K [25], and a "porous" ice template constituted of a 1.0 ML overlayer of porous water ice deposited at 10 K over an ASW substrate. This substrate has no pores but is topologically disordered and, particularly, presents already deep adsorption sites [26].

In all the experiments, the substrate is heated steadily (10 K/min) at the end of each deposition phase. Prior to each experiment, the sample is annealed to 100 K in order to stabilize the surface morphology before subsequent heating-cooling runs between 6.5 and 90 K.

Figure 1 shows the results of experiments performed by varying the  $O/O_2$  doses. Two desorption peaks are present:  $O_2$  desorption occurs between 35 and 50 K, and the ozone desorption is observed between 55 and 75 K (directly, or via the  $O_2^+$  fragments). We observe, at any coverage or temperature, ozone formation by depositing O and  $O_2$  mixtures on ASW at 10 K. O desorption is never observed. The shapes and positions of the  $O_3$  peaks are the same as those of  $O_3$  deposited from the gas phase and are only coverage dependent. We can thus exclude any second order

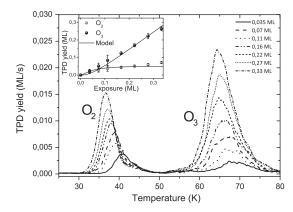


FIG. 1. TPD of  $O_2$  and  $O_3$  after deposition of various doses (0.04–0.4 ML) of an  $O/O_2$  mixture on ASW ice held at 10 K. Inset: Areas of the TPD peaks (in ML) vs the exposed  $O/O_2$  dose.

desorption effects, like it should be if  $O_3$  were formed on the onset of or during desorption. The circles in the inset of Fig. 1 represent the area under the TPD peaks as a function of deposited dose. The  $O_3/O_2$  ratio increases with coverage:  $O_2$  rapidly reaches an almost steady state while  $O_3$  raises quite linearly. The solid lines in the inset represent the model results (see below). These experiments suggest that the reactivity of pure oxygen species is limited to two reactions:

$$O + O \rightarrow O_2$$
 R1  $O + O_2 \rightarrow O_3$  R2.

The  $O + O_3 \rightarrow 2O_2$  reaction seems to be not competitive with the others, or it would not be possible to obtain an almost pure O<sub>3</sub> sample with increasing of the coverage. The two reactions R1 and R2 may arise from direct reactions between an impinging atom and an adsorbed species [Eley-Rideal mechanism (ER)] or may occur by diffusion of the species on the surface [Langmuir-Hinshelwood mechanism (LH)]. It is also possible that the hot atom mechanism (HAM) could have a role. Up to now, the HAM has been studied in conditions very different from ours: metallic surfaces, atoms with energy larger than 0.5 eV, or light atoms (H or D) (see Refs. [27,28] and references therein). The HAM is an initial diffusion of the impinging atoms due to their residual gas phase kinetic energy. We can consider this mechanism an extension of the ER mechanism as it occurs during the accommodation phase. In the case of O atoms, the kinetic energy of the beam is around 300 K and the sticking efficiency is more than 90%, which indicates a good transfer of energy to the water ice substrate. The binding energy is about 1000 K, and thermal accommodation occurs in a few site jumps, since the atoms that stick have lost about 300 K of the total energy upon their first impact. In what follows, we will consider the HAM mechanism included in the ER mechanism but with an enhanced cross section ( $\times$  3). The ER mechanism happens between a gas phase reactant and a surface reactant; by construction, it is not sensitive to the surface temperature and its efficiency depends on the coverage only. On the contrary, the LH mechanism (as well as diffusion) depends on the temperature of the surface. Therefore, during a TPD when the temperature increases, this mechanism could lead to the formation of other  $O_3/O_2$ molecules. We have attempted to check this possibility by following the evolution of the O<sub>3</sub> infrared absorption band intensity from 6.5 to 35 K. Because of a high detection limit (0.3 ML), this method could only be applied to the highest coverage experiment in Fig. 1, and even for this experiment the signal to noise ratio is too low to provide strong constraints (not shown). Within the experimental uncertainties, the O<sub>3</sub> infrared band does not vary during the TPD for the high coverage experiment, except at the temperature above which ozone begins to come off the surface ( $\sim 55$  K). This demonstrates that at least some O3 forms at the time of deposition, and the results are consistent with the theory

that thermally induced diffusion during the TPD is a secondary effect compared to diffusion and reactions at the deposition temperature.

To understand if diffusion effectively plays a role in the O<sub>3</sub>/O<sub>2</sub> formation, we have performed a second set of experiments in which we varied the deposition temperature of the substrate and the morphology of the water substrate itself, but fixing the initial  $O/O_2$  dose (coverage). As shown in Fig. 2, we observe that the O<sub>3</sub>/O<sub>2</sub> ratio increases with the temperature of the substrate. The evolution as a function of the exposure temperature indicates that the temperature of the surface during irradiation is a key parameter. In Fig. 2, we show the  $O_3$  and  $O_2$  yields for different temperatures and for three different ice substrates (ASW ice, crystalline ice, and ASW ice coated with 1 ML of porous water ice). There are differences between the three substrates, but they can be considered secondary if compared to the global trend. Even if the temperature dependence seems to be slow, at 25 K, the O<sub>3</sub> amount raises by 50% with respect to that at 6.5 K, while in the same temperature range the O<sub>2</sub> yield decreases by 75%. During the TPD, the adsorbates follow the same thermal history and should produce the same results. Therefore, the differences should originate at the time of the deposition phase. ER and HAM mechanisms are independent of the surface temperature; thus, the evolution of the  $O_3/O_2$  ratio is due to diffusion processes.

Basically, the  $O_3/O_2$  balance is due to the diffusion of O atoms. If the diffusion is extremely fast, each newly adsorbed atom scans the surface until it reacts with  $O_2$  to form  $O_3$ . If no  $O_2$  is present, O has to wait for another O atom to form  $O_2$ , which in turn will be transformed into  $O_3$  via another incoming and mobile atom. Therefore, almost all the O atoms and  $O_2$  molecules are transformed

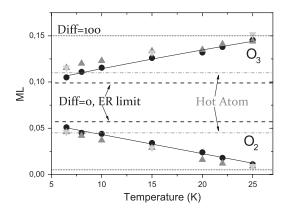


FIG. 2.  $O_2$  and  $O_3$  TPD peak areas after deposition of 0.21 ML of an  $O/O_2$  mixture ( $\mu = 71\%$ ) on three substrates: crystalline ice (circles), nonporous ASW ice (triangles pointing up), and porous water ice (triangles pointing down). Solid lines: model results. Dashed and dotted lines:  $O_2$  (lower part) or  $O_3$  (upper part) amount in the case of no diffusion (ER limit) and  $k = 100 \text{ s}^{-1}$ . Dash-dotted line: model results with HAM and no diffusion.

into  $O_3$  molecules. On the contrary, if the diffusion is slow, one O atom does not have enough time to scan the surface to meet an  $O_2$  molecule before another O comes. In this case, O atoms accumulate until the probability for an O atom to meet another O atom raises, and finally  $O_2$  formation is favored. In summary, Fig. 1 shows that an incoming O atom is more likely to find  $O_2$  molecules as the coverage increases. In Fig. 2, we show the increase of the  $O_3/O_2$  ratio with deposition temperature, which demonstrates the increase of diffusion with temperature.

Model and discussion.—As discussed in the previous section, only two reactions (R1 and R2) occur on the surface via ER (HAM) or LH mechanisms. We have modeled our experiments by adapting a classical set of differential equations [29] (see the Supplemental Material for details [30]). We assume a diffusion-limited reactivity (no reaction barrier), including both ER and LH mechanisms (same efficiency), or HAM, which is estimated by enhancing the ER mechanism. Anyway, O atoms are of high enough mass to be able to transfer via each single collision a relevant fraction of their kinetic energy, and so in a few jumps O atoms are thermalized. Possible adjustments due to thermally induced diffusion during TPDs are also considered in order not to exclude the possibility of incomplete reactions during the exposure phase, even if we already noticed that it leads to minor effects. There is only one physical free parameter to adjust, k, which represents the effective surface diffusion. Therefore, the  $O_3/O_2$  ratio at 25 K is 45 times bigger than that at 6.5 K. Using only one adjustable parameter k, we reproduce perfectly all our data sets (solid lines in Figs. 1 and 2). Other alternative scenarios (barrier to reaction, low diffusivity, pure LH or ER mechanisms, and HAM) have been tested without the same success. In Fig. 2, the boundaries of the ER mechanism and HAM are represented by two constant lines. These limits are not sensitive to the surface temperature. Actually, it is possible to fit the coverage dependency with several different hypotheses (or parameters), but it is not possible to have both temperature and coverage dependencies satisfied at once. The plain circles of Fig. 3 show the diffusion law for O as a function of the temperature, obtained from the change in the balance of O2 and O3 production, assuming a diffusion dominated process. The diffusion coefficient increases by a factor of 50, but the logarithm scaling of the figure tends to flatten this aspect of the experimental results. The trend is somehow surprising because the measured diffusion barrier does not follow an Arrhenius law. The empty circles in Fig. 3 represent a typical Arrhenius law with a diffusion barrier of 450 K. Therefore, a pure thermal diffusion does not represent well our data. The diffusion is better simulated using the quantum tunneling of a square barrier, as described in Messiah's book [3]. We use two physical parameters: the width a and the height of the barrier  $E_a$ . As already described [31], diffusion includes two components, quantum tunneling that

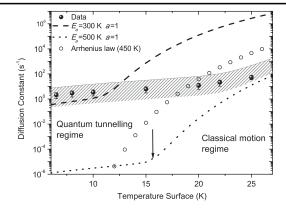


FIG. 3. Diffusion coefficient of O atoms on amorphous ice as a function of the surface temperature. Plain circles: experimental values. Open circles: Arrhenius law ( $E_a=450~\rm K$ ). Dashed and dotted lines: best fits of the diffusion using both quantum and thermal diffusion ( $a=1~\rm \AA,~E_a=500\text{--}300~\rm K$ ). Pinstripe zone:  $a=0.7\pm0.05~\rm \AA,~E_a=520~\rm K\pm60~\rm K$ . The arrow indicates the point on the dotted curve where the quantum to classical transition occurs.

dominates at low temperatures and thermal diffusion predominant at higher temperatures. The arrow in Fig. 3 represents the critical temperature where the transition occurs. In our experimental data, we observe that the quantum to classical regime occurs at around 20 K. We find the best results for  $a=0.70\pm0.05$  Å and  $E_a=520\pm60$  K. These two parameters have different effects on the diffusion curves, as shown in Fig. 3 (dashed and dotted lines). The pinstriped region in Fig. 3 represents the validity zone of the solution we found, since it is possible to partly compensate for the variation of the height of the barrier by changing its width. Adopting a square barrier is an extreme assumption, and the apparent low value of the height we found may be due to this unrealistic shape.

The main results of our study are the following: first, O-atom diffusion is governed by quantum tunneling up to 20 K. Our experiments show nicely the transition from the quantum world to the classic world. The thickness of the barrier (0.7 Å) may also be due to the amorphous nature of the ice. Indeed, one can consider that the diffusion on an amorphous surface is dominated by the fastest jumps between adsorption sites and that the diffusion barriers limiting the diffusion are the weakest ones [32]. Therefore, the apparent low value of the width is probably an estimation of the lower limit of the distribution of barrier widths of this disordered substrate. We now want to discuss the signification of a width barrier of 0.7 Å. On water ice, the mean distance of two adsorption sites is 3 Å, while the de Broglie wavelength  $(\lambda_h)$  associated with O atoms varies from 2.8 Å at 6.5 K to 1.5 Å at 20 K. It corresponds to the typical size where quantum effects are important. We can compare the 3 Å mean distance with the double (two-atom diffusion reaction) of  $\lambda_b$ . For temperatures lower than 22 K, quantum effects should thus be important. This is actually what we observe. The slightly lower value (0.7 instead of 1 Å guessed in the literature) could also be due to the disordered nature of the surface that distributes distances between adjacent adsorption sites and so reduces some of them. Therefore, cold O atoms would easily overlap two neighboring sites.

Second, comparing the different morphologies of the ice allows us to conclude that the topological disorder of the substrate does not deeply affect the diffusion regime. This probably means that the lower limits of the distributions of barrier widths are not too different from each other for the substrates considered here. However, we find that the diffusion on the crystalline surface is faster than that on amorphous ice. This could be explained as follows: (i) the wave packet describing the adsorbed atom diffuses more quickly in the periodic potential of a crystalline surface than in the amorphous one [32]. (ii) Another study [11] showed that the diffusion and reactive properties are greatly changed by the occurrence of deep and shallow sites. In our case, we find that the presence of deep sites (a porous substrate) does not significantly reduce the effective diffusion. The trapping sites may force one O atom to stay in one adsorption site, but they cannot prevent another atom from reaching it, especially if tunneling dominates.

The third conclusion is that we find that the diffusion of oxygen at 10 K is  $k = 5 \text{ s}^{-1}$ , which corresponds in classical Arrhenius formalism to the thermal diffusion barrier of 300 K. The diffusion is faster than previously guessed by astronomers (diffusion barrier of 400 K [6] or even 900 K [33]). However, it is not very far from that of O diffusion in matrices  $240 \pm 80 \text{ K}$  [34], which could be considered an upper value for surface diffusion. Based upon arguments of polarizability scaling [12] and the mean value of Ref. [8], we would expect a value of  $275 \pm 30$  K, which is consistent with what is derived here. From our model, we can indirectly conclude that the maximal barrier height for the O<sub>2</sub> + O reaction is 190 K. Actually, if the barriers were higher, there would be noticeable discrepancies in the  $O_2/O_3$  ratio, which we do not observe. It is probably even smaller, as estimated in a previous study [34]. The implications for solid-state astrochemistry are of major importance. It was usually thought that the chemistry was mostly driven by H diffusion, and therefore the final products were mostly hydrogen saturated species such as H<sub>2</sub>O, NH<sub>3</sub>, CH<sub>4</sub>, and CH<sub>3</sub>OH. We can affirm now that O addition chemistry is competitive with H additions because of the comparable budget of O atoms and H atoms in dense and UV protected interstellar environments [35]. Furthermore, if we scale the diffusion for other atoms such as C and N, on the basis of their polarizability alone—the main parameter for physisorption—there is a reasonable range of temperatures ( $\leq 15$  K) where the mobility of O, C and N is activated. For this reason, the production of O, C, and N bearing molecules can grow, avoiding saturated chemical traps. This could also be one

of the source of complex nonvolatile organic compounds observed in meteorites, such as amino acids.

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- [1] T. P. M. Goumans and S. Andersson, Mon. Not. R. Astron. Soc. **406**, 2213 (2010).
- [2] M. J. Gillan, J. Phys. C 20, 3621 (1987).
- [3] A. Messiah, *Quantum Mechanics* (Wiley, Amsterdam, 1973).
- [4] G. Ehrlich, Surf. Sci. **299–300**, 628 (1994).
- [5] L. Pagani, A. Bacmann, S. Cabrit, and C. Vastel, Astron. Astrophys. 467, 179 (2007).
- [6] A. G. G. M. Tielens and W. Hagen, Astron. Astrophys. 114, 245 (1982).
- [7] H. M. Cuppen and E. Herbst, Astrophys. J. 668, 294 (2007).
- [8] E. Matar, E. Congiu, F. Dulieu, A. Momeni, and J. L. Lemaire, Astron. Astrophys. 492, L17 (2008).
- [9] N. Watanabe, Y. Kimura, A. Kouchi, T. Chigai, T. Hama, and V. Pirronello, Astrophys. J. Lett. 714, L233 (2010).
- [10] R. Smoluchowski, J. Phys. Chem. 87, 4229 (1983).
- [11] A. Wolff, I. Lohmar, J. Krug, Y. Frank, and O. Biham, Phys. Rev. E 81, 061109 (2010).
- [12] A.G.G.M. Tielens and L.J. Allamandola, in *Cool Interstellar Physics and Chemistry* (Pan Stanford, Singapore, 2011).
- [13] F. Dulieu, L. Amiaud, E. Congiu, J.-H. Fillion, E. Matar, A. Momeni, V. Pirronello, and J. L. Lemaire, Astron. Astrophys. 512, A30 (2010).
- [14] D. Jing, J. He, J. Brucato, A. De Sio, L. Tozzetti, and G. Vidali, Astrophys. J. 741, L9 (2011).
- [15] M. D. Ward and S. D. Price, Astrophys. J. 741, 121 (2011).

- [16] H. Bergeron, N. Rougeau, V. Sidis, M. Sizun, D. Teillet-Billy, and F. Aguillon, J. Chem. Phys. 112, 11921 (2008).
- [17] R. Larciprete, P. Lacovig, S. Gardonio, A. Baraldi, and S. Lizzit, J. Phys. Chem. C 116, 9900 (2012).
- [18] H. Mokrane, H. Chaabouni, E. Accolla, M. Congiu, F. Dulieu, M. Chehrouri, and J. L. Lemaire, Astrophys. J. Lett. 705, L195 (2009).
- [19] C. Romanzin, S. Ioppolo, H. M. Cuppen, E. F. van Dishoeck, and H. Linnartz, J. Chem. Phys. 134, 084504 (2011).
- [20] C. Janssen and B. Tuzson, J. Phys. Chem. A 114, 9709 (2010).
- [21] D. Jing, J. He, J. R. Brucato, G. Vidali, L. Tozzetti, and A. De Sio, Astrophys. J. 756, 98 (2012).
- [22] L. Amiaud, F. Dulieu, J.-H. Fillion, A. Momeni, and J. L. J. L. Lemaire, J. Chem. Phys. 127, 094702 (2007).
- [23] E. Congiu, H. Chaabouni, C. Laffon, P. Parent, S. Baouche, and F. Dulieu, J. Chem. Phys. 137, 054713 (2012).
- [24] E. Congiu, E. Matar, L.E. Kristensen, F. Dulieu, and J. L. Lemaire, Mon. Not. R. Astron. Soc. 397, L96 (2009).
- [25] J. A. Noble, E. Congiu, F. Dulieu, and H. J. Fraser, Mon. Not. R. Astron. Soc. 421, 768 (2012).
- [26] J.-H. Fillion, L. Amiaud, E. Congiu, F. Dulieu, A. Momeni, and J.-L. Lemaire, Phys. Chem. Chem. Phys. 11, 4396 (2009).
- [27] L. C. Dinger and J. Kuppers, J. Chem. Phys. 114, 5338 (2001).
- [28] M. Tomellini, Physica (Amsterdam) **392A**, 875 (2013).
- [29] N. Katz, I. Furman, O. Biham, V. Pirronello, and G. Vidali, Astrophys. J. 522, 305 (1999).
- [30] See Supplemental Material at <a href="http://link.aps.org/supplemental/10.1103/PhysRevLett.111.053201">http://link.aps.org/supplemental/10.1103/PhysRevLett.111.053201</a> for the model used to analyze the experimental data.
- [31] S. Cazaux and A. Tielens, Astrophys. J. **604**, 222 (2004).
- [32] L. Karssemeijer, A. Pedersen, H. Jonsson, and H. Cuppen, Phys. Chem. Chem. Phys. **134**, 084504 (2012).
- [33] S. Cazaux, V. Cobut, M. Marseille, M. Spaans, and P. Caselli, J. Astrophys. Astron. **74**, 522 (2010).
- [34] A. V. Benderskii and C. A. Wight, J. Chem. Phys. **104**, 85 (1996).
- [35] P. Caselli, O. Stantcheva, T. Shalabiea, V. Shematovich, and E. Herbst, Planet. Space Sci. 50, 1257 (2002).