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► **To cite this version:**

Laura Gomez-Navarro, Emmanuel Cosme, Julien Le Sommer, Nicolas Papadakis, Ananda Pascual. Development of an Image De-Noising Method in Preparation for the Surface Water and Ocean Topography Satellite Mission. *Remote Sensing*, 2020, 12 (4), pp.734. 10.3390/rs12040734 . hal-02490206

**HAL Id: hal-02490206**

**<https://hal.science/hal-02490206v1>**

Submitted on 12 Nov 2020

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# Development of an image de-noising method in preparation for the Surface Water and Ocean Topography satellite mission

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## Abstract

In a near future, the Surface Water Ocean Topography (SWOT) mission will provide images of altimetric data at kilometeric resolution. This unprecedented 2-dimensional data structure will allow the estimation of geostrophy-related quantities that are essential for studying the ocean surface dynamics and for data assimilation uses. To estimate these quantities, i.e. compute spatial derivatives of the Sea Surface Height (SSH) measurements, the small-scale noise expected to affect the SWOT data must be smoothed out while minimizing the loss of relevant, physical SSH information. This paper introduces a new technique for de-noising the future SWOT SSH images. The de-noising model is formulated as a regularized least-square problem with a Tikhonov regularization based on the first, second, and third-order derivatives of SSH. The method is implemented and compared to other, convolution-based filtering methods with boxcar and Gaussian kernels. This is performed using a large set of pseudo-SWOT data generated in the Western Mediterranean Sea, from a  $1/60^\circ$  simulation and the SWOT simulator. Based on Root Mean Square Error and spectral diagnostics, our de-noising method shows a better performance than the convolution-based methods. We find the optimal parametrization to be when only the second-order SSH derivative is penalized. This de-noising reduces the spatial scale resolved by SWOT by a factor of 2, and at 10 km wavelengths the noise level is reduced by  $10^4$  and  $10^3$  for Summer and Winter respectively. This is encouraging for the processing of the future SWOT data.

*Keywords:* SWOT, De-noising, Variational regularization, western Mediterranean

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## 1. Introduction

The Surface Water Ocean Topography (SWOT) [1] mission will provide an unprecedented two-dimensional view of ocean surface topography at a pixel resolution of 2 km. The launch is scheduled for 2021. SWOT's wide-swath altimeter, based upon SAR interferometry technology, will measure Sea Surface Height (SSH) over a 120-km wide swath with a 20-km gap at the nadir. The satellite will also carry a conventional nadir altimeter. SWOT will evolve on two different orbits: the first 3 months of scientific data production will be dedicated to a fast-sampling phase, where the repeat cycle will be of 1 day. Then, the satellite will be moved to its nominal orbit with a 20.86-day repeat cycle. SWOT is a multi-disciplinary hydrology and oceanography mission, and here we focus on the latter.

The main oceanographic objective of SWOT is to observe the geostrophic fine-scale circulation at the global scale [2, 3]. The measurement system is designed to resolve ocean circulation patterns at scales down to 15 km, whereas the present-day constellation of conventional altimeters only resolves scales of 150-200 km and above [3]. In addition to potentially unexpected

16 discoveries, this order-of-magnitude gain in resolution will help quantifying several oceanic pro-  
17 cesses much more accurately than today. Among those processes are vertical motions, which  
18 are key to the vertical exchanges between the ocean surface and the atmosphere, and between  
19 the ocean surface and the deep ocean [4, 5, 6, 7, 8, 9]; and the dissipation of kinetic energy,  
20 which partly determines the climatic role of the global ocean [10, 11].

21  
22 The SWOT mission objectives will be reached if we can accurately estimate gridded maps of  
23 at least the first and second-order horizontal derivatives of SSH. Altimetry describes the upper  
24 ocean dynamics through geostrophy, which involves the horizontal SSH gradients. Geostrophy  
25 is a fairly good approximation of mesoscale dynamics, i.e. at scales larger than the first  
26 Rossby deformation radius (about 10-15 km in our region [12]), for which Rossby numbers are  
27 typically smaller than 1. Kinetic energy dissipation is driven by the horizontal strain rates of  
28 the ocean surface flow [e.g. 13]. Complete, gridded maps of SSH derivatives are required for  
29 climate studies and short-term operational applications. One way to make gridded maps from  
30 incomplete SSH observations (including SWOT, but not only) is to assimilate those data into  
31 dynamical models. The assimilation of SWOT is expected to be challenging because of the  
32 spatially correlated noise, and promising solutions to this rely upon the joint assimilation of  
33 SSH and its derivatives [14, 15]. All these considerations compel the scientific community to  
34 strive for getting accurate estimates of SSH derivatives.

35  
36 Unfortunately, SWOT data will very likely be contaminated by small-scale noise that pre-  
37 vents the direct computation of SSH derivatives. The noise expected to contaminate SWOT  
38 measurements gathers several components with different spatial coherences and different am-  
39 plitudes. Details are provided in the SWOT mission performance and error budget document  
40 [16]. To be prepared to exploit the future SWOT data, the SWOT simulator for ocean science  
41 has been developed to simulate realistic realizations of SWOT uncertainties [17]. Some are  
42 illustrated on Figure 2. Errors due to the satellite roll, the baseline dilation, and the path  
43 delay induced by atmospheric humidity, exhibit significant spatial correlations with different  
44 characteristic patterns. The system timing error presents errors invariant across-track, but with  
45 possible small-scale variations along-track. The KaRIn (Ka-band Radar Interferometer) noise  
46 is spatially uncorrelated, with higher amplitudes at nadir and near the edges of the swath. The  
47 path-delay component also exhibits small-scale variations due to sharp changes in air humidity.  
48 Efforts have already been undertaken to filter out SWOT small-scale noise by Gómez-Navarro  
49 et al. [18]. The authors show that the implementation of a diffusion-based filter allows to  
50 retrieve the dynamical spectral signature down to 40-60 km scales (20-30 km in terms of dy-  
51 namical pattern scales). However, the de-noising approach here is not specifically designed  
52 to retrieve SSH derivatives, and we believe there is room for improvement in the scales to be  
53 retrieved.

54  
55 This paper presents a method designed to remove the small-scale noise of the future SWOT  
56 data, which explicitly relies upon the regularity (bounded variations) of the first three orders  
57 of SSH derivatives. This de-noising method is rooted in image restoration techniques of the  
58 variational type [19, 20, 21, 22]. The range of image restoration techniques is extremely wide  
59 and diversified. Testing all existing methods is out of reach and irrelevant here. Our approach  
60 is then to acknowledge that our image is a smooth physical field with relatively smooth deriva-  
61 tives, and that the estimation of derivatives is an important issue. This consideration guides  
62 the design of the de-noising method presented in Section 2. The method involves a set of pa-  
63 rameters that must be adjusted. An essential task is to identify optimal sets of parameters.  
64 This study suggests a methodology to identify them. The experimental setup is described in  
65 Section 3. Sections 4 and 5 present the results and Section 6 summarizes the study, draws the

66 most relevant conclusions, discusses them, and suggests possible future research paths.

67

## 68 2. Variational de-noising of SWOT images with penalization of derivatives

### 69 2.1. Formulation of the image de-noising problem

70 The primary purpose of image de-noising here is to allow the computation of first and second-  
 71 order SSH spatial derivatives of SWOT data as accurately as possible. The two reasons, already  
 72 mentioned in the introduction, are: (i) these quantities represent geostrophic velocities and rel-  
 73 ative vorticity, respectively, whose estimation is central to the success of SWOT mission; and  
 74 (ii) these quantities can be needed to draw maximum benefits from the assimilation of SWOT  
 75 data into ocean circulation models [14, 15]. We therefore propose a method that explicitly  
 76 constrains these derivatives.

77

The proposed de-noising model is formulated as a regularized least-square problem with a Tikhonov regularization. The de-noised SWOT image  $h$  is searched for by minimizing the following cost function:

$$J(h) = \frac{1}{2} \|m \circ (h - h_{obs})\|^2 + \frac{\lambda_1}{2} \|\nabla h\|^2 + \frac{\lambda_2}{2} \|\Delta h\|^2 + \frac{\lambda_3}{2} \|\nabla \Delta h\|^2 \quad (1)$$

78 where  $\| \cdot \|$  represents the  $L_2$ -norm,  $h^{obs}$  is the original noisy image (i.e., our observation, the  
 79 pseudo-SWOT data),  $\nabla = (\partial/\partial x, \partial/\partial y)$  is the gradient operator, and  $\Delta = \partial^2/\partial x^2 + \partial^2/\partial y^2$   
 80 is the Laplacian operator. Letter  $m$  and sign  $\circ$  represent a mask and the entrywise matrix  
 81 product, respectively. They can be ignored for the present and the next sub-section: their role  
 82 is discussed in Section 2.3 below. The regularization terms impose regularity constraints on  
 83 geostrophic velocity, vorticity, and vorticity gradient, respectively. Parameters  $\lambda_1$ ,  $\lambda_2$  and  $\lambda_3$   
 84 must be prescribed. The search for their optimal values is reported in Section 3.3.

85

### 86 2.2. Resolution of the variational problem

The variational problem displayed in eq. 1 is solved using a gradient descent method [23]. The gradient of  $J$  is written:

$$\nabla J(h) = m \circ (h - h_{obs}) - \lambda_1 \Delta h + \lambda_2 \Delta \Delta h - \lambda_3 \Delta \Delta \Delta h \quad (2)$$

so that the solution can be reached after convergence of the following iterations:

$$h^{k+1} = h^k + \tau (m \circ (h_{obs} - h^k) + \lambda_1 \Delta h^k - \lambda_2 \Delta \Delta h^k + \lambda_3 \Delta \Delta \Delta h^k) \quad (3)$$

87 Stability of iterations is guaranteed if  $\tau \leq (1 + 8\lambda_1 + 64\lambda_2 + 512\lambda_3)^{-1}$ . In practice, it is  
 88 taken equal to this value. Two improvements on the method's implementation accelerate the  
 89 gradient descent: Firstly, iterations are started with a preconditioned image obtained by ap-  
 90 plying a Gaussian filter onto the original image, including inpainting as discussed in Sections  
 91 2.3 and 2.4 (note that  $h^{obs}$  remains the original, unfiltered image). Preconditioning consider-  
 92 ably speeds up the algorithm convergence, in particular for the inpainted regions. Secondly,  
 93 iterations are actually implemented with an acceleration of the scheme 3, based on the Fast  
 94 Iterative Shrinkage-Thresholding Algorithm (FISTA) [24], detailed in Appendix B. Iterations  
 95 are stopped when  $\|h^{k+1} - h^k\| < 10^{-9}$  or if  $k = 10^4$ . Those values have been fixed after a careful  
 96 search of a trade-off between accuracy and numerical efficiency.

97



98 The Laplacian operator is discretized with finite differences using the five-point stencils of  
99 the image pixels. As commonly done in image processing, the division by pixel size is ignored;  
100 this also reduces the probability of truncation errors due to operations with terms different by  
101 too many orders of magnitude. Pixels located at the boundaries, where the stencil is incomplete,  
102 must have a Laplacian value attributed yet; otherwise, the image would become smaller at each  
103 iteration of the gradient descent. The implementation of the Laplacian operator follows [22]  
104 and is detailed in Appendix A.

### 105 *2.3. Dealing with gaps in the image*

106 An inpainting method is implemented to deal with islands, continents, and the 20-km wide  
107 gap at the SWOT nadir, which all represent obstacles to the calculation of the second deriva-  
108 tives of images. Inpainting consists in filling the gaps consistently with the neighbouring water  
109 pixels. This is done by (i) extending images  $h^{obs}$  and  $h$  with pixels in the gaps, and (ii) filling  
110 mask  $m$  with ones in water pixels of the original image, and zeros in the gaps. Differential  
111 operators can then be applied to every image pixel, and the gradient descent iterations are  
112 carried out smoothly. Mask  $m$  is applied to the resulting image to obtain the final, filtered  
113 image with islands, continents, and the nadir gap.

114  
115 Inpainting should not be only considered as a trick to facilitate the gradient descent im-  
116 plementation, but also as an opportunity to fill the nadir gap for calibration, validation and  
117 reconstruction purposes. In the gaps, the image resulting from the iterations is determined only  
118 by the neighbouring water pixels and regularity constraints. The gap width (20 km) appears  
119 reasonably small in comparison with spatial scales of SSH variations in most parts of the mid-  
120 latitude, open ocean. The image values obtained at nadir may thus be comparable to those  
121 collected by the nadir instrument carried by SWOT, allowing calibration of the radar interfer-  
122 ometer, validation of data and reconstruction of SSH in gap-free images. Such opportunities  
123 will be explored in a future work.

### 124 *2.4. Comparison with convolution-based filters*

125 In Section 4, the image de-noising technique described above will be compared with standard-  
126 type filters, namely convolution-based filters. In our experiments we test the two commonly  
127 used boxcar and Gaussian convolution kernels, with a large range of parameters, and we shortly  
128 refer to the boxcar filter and the Gaussian filter. Their parameters are the box size (or foot-  
129 print) and the standard deviation for the Gaussian kernel (hereinafter referred to as  $\sigma$ ). Gaps  
130 in the SWOT swath (lands, islands, and nadir gap) are inpainted to facilitate filtering and to  
131 ensure the smoothness of SSH fields. Then, SSH values created in gaps are removed for the  
132 evaluation of the methods using the mask  $m$ . Inpainting is implemented as follows: (i) Image  
133 gaps are filled with zeros; (ii) both the filled image and the mask  $m$  are filtered with the same  
134 kernel; and (iv) the filtered filled image is divided entrywise by the filtered mask. Note that in  
135 an earlier study [18], a Laplacian diffusion filter was experimented. It is not reproduced here,  
136 since it is equivalent to the Gaussian filter implemented in this study.

## 137 **3. Experimental setup**

### 138 *3.1. Simulated SWOT dataset*

139 The input of our database is a 15-month North Atlantic simulation at a resolution of  $1/60^\circ$ .  
140 We use the NEMO3.6 ocean model coupled to LIM2 ice model, with atmospheric forcing from  
141 a global ocean reanalysis at  $1/4^\circ$  (GLORYS-v3) and ocean-atmosphere boundary conditions  
142 of Drakkar Forcing Set (DFS5.2), based on ERA-interim reanalysis. It has no high frequency  
143 forcing, thus does not include tides. The domain covers the North Atlantic from  $25^\circ\text{N}$  to  $66^\circ\text{N}$ .

144 The horizontal resolution is between 0.8 and 1.6 km (depending on latitude), and the grid has  
145 300 vertical levels. This NEMO model configuration is referred to as NATl60 and the source  
146 files and codes are available in [25]. The particular simulation used herein has been described  
147 in [26, 27, 28]. Lastly, the simulation time span is from mid-June 2012 to October 2013. [29]

148  
149 The SWOT simulator for Ocean Science (version 2.21) [17] is run to generate pseudo-SWOT  
150 scenes from the NATl60 simulation. The SWOT simulator first builds the SWOT observation  
151 grid, based on the provided satellite orbit. In this study, SWOT grid resolution is fixed at  
152 1 km. After this work started, the resolution of the basic SWOT level 2 SSH data products  
153 has been fixed to 2 km, but this small mismatch does not modify the general approach. After  
154 building the grid, the simulator reads SSH data from NATl60 and linearly interpolates them  
155 from model to SWOT grid (rendering the variable SSH\_model). In a last step, it computes  
156 random realizations of observation errors and adds them to the interpolated SSH data (ren-  
157 dering SSH\_obs). Observation errors considered at the moment are KaRIn errors, roll errors,  
158 phase errors, baseline dilation errors, timing errors, and errors due to signal alternation by  
159 atmospheric humidity. Among these errors, only the KaRIn noise is expected to be spatially  
160 uncorrelated. Technically, the SWOT simulator provides simulations of the noise-free SSH ob-  
161 served by SWOT, and of the noisy data that SWOT will actually yield (sum of the former  
162 and the noise:  $SSH\_obs = SSH\_model + noises/errors$ ). For the evaluation of image de-noising  
163 methods, it thus provides "true", noise-free images ( $h^{true}$ ) along with the realistic SWOT data  
164 ( $h^{obs}$ ) to process and compare with the truth.

165  
166 A set of 543,  $121 \times 200$  km<sup>2</sup> pseudo-SWOT scenes are generated in the western Mediter-  
167 ranean Sea, covering one Winter and two Summer seasons (choice limited by the model's time  
168 span). SWOT scenes are sampled from the fast-sampling phase satellite orbit, focusing on a  
169 cross-over region, i.e., where an ascending pass crosses a descending pass, therefore providing  
170 2 passes per day. The SWOT data simulation is carried out over three 3-month periods: July  
171 to September 2012 and 2013 (JAS12 and JAS13 hereafter), representing the Summer season,  
172 and February to April 2013 (FMA13) representing the Winter season. Summer periods provide  
173 92 (resp. 91) of ascending (resp. descending) passes; the Winter period provides 89 (resp.  
174 88) passes. The selected region belongs to the fast-sampling phase crossover in the western  
175 Mediterranean Sea. This is one of the regions selected for calibration/validation (Cal/Val) [30]  
176 in which *in situ* measurements have been made in the frame of SWOT [31]. To mitigate the  
177 computational complexity of the study and avoid the presence of continents and islands, limited  
178 subregions of the SWOT swaths are sampled. These subregions are 121 km-wide (the width  
179 of 2 SWOT swaths plus the gap) and 200 km-long. The region, the SWOT passes and the  
180 subregions are shown in Figure 1. It is worth noting that each scene is affected by a unique  
181 realization of the SWOT error.

182  
183 In this work, image de-noising techniques are first applied to the pseudo-SWOT scenes af-  
184 fected by the KaRIn noise only (SSH\_obs\_K), then to the scenes containing all errors (SSH\_obs).  
185 This approach allows to discriminate the effects and the performance of image de-noising in  
186 presence of the spatially correlated SWOT errors. A few realizations of the different compo-  
187 nents of the SWOT error are shown on Figure 2, where we can observe how most errors exhibit  
188 strong and long-range correlations, whilst the KaRIn error does not show any correlation at  
189 all.

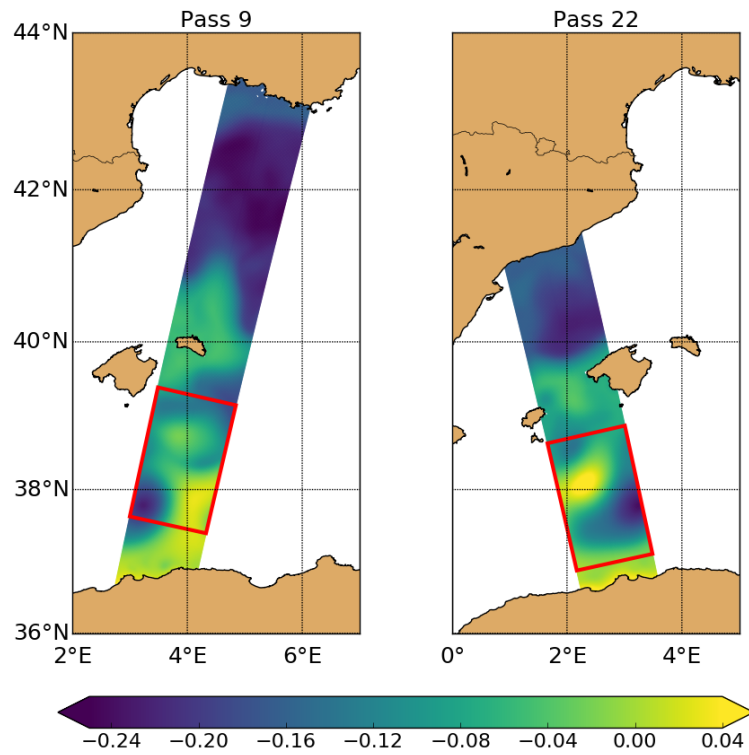


Figure 1: SSH\_model outputs [m] for cycle 1 of pass 9 (left) and 22 (right) of the JAS12 dataset. In red the subregions selected.

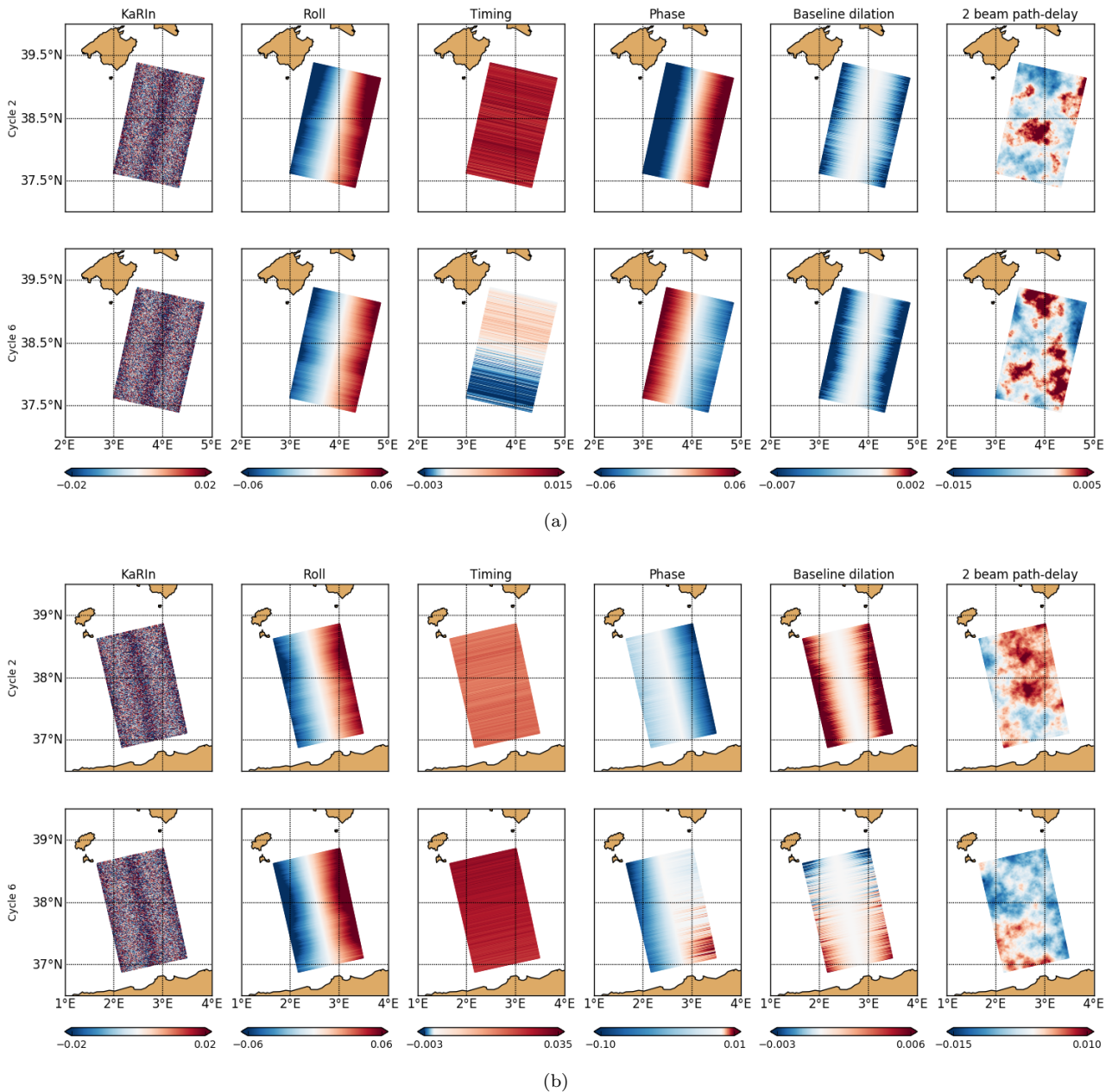


Figure 2: Examples of noises and errors [m] added by this SWOT simulator version 2.21 to our study region fast-sampling phase for JAS12 pass 9 (a) and 22 (b). Note that these simulations are performed without the 20-km gap at nadir.

### 3.2. Diagnostics for evaluation

The quantitative evaluation of de-noising methods is carried out computing Root Mean Square Errors (RMSE) and Mean Spectral Ratios (MSR). RMSE for a single de-noised SWOT field  $h$  is computed as the Euclidean distance to the corresponding original, noise-free field  $h^{\text{true}}$ :

$$RMSE(h) = \sqrt{\frac{1}{N} \sum_{i=1}^N (h_i - h_i^{\text{true}})^2} \quad (4)$$

where  $N$  is the number of pixels and  $i$  a pixel index. Single image RMSEs are then averaged out by season for the 3 seasons considered, and are computed for SSH,  $|\nabla\text{SSH}|$  and  $\Delta\text{SSH}$ . Thus, the test of a de-noising method with a specific set of parameters results in 9 RMSE values. To evaluate the improvement after the application of the different de-noising techniques and

198 parameters, we also calculate the percentage of the initial RMSE left. We calculate this RMSE  
 199 residual ( $RMSE_r$ ) as:

$$RMSE_r(h) = \frac{RMSE(h)}{RMSE(h^{obs})} \times 100, \quad (5)$$

200 where  $h$  is the de-noised field and  $h^{obs}$  the original noisy field (SSH\_obs\_K or SSH\_obs).  
 201

202 The spatial spectra of the de-noised SWOT SSH are compared with the spectra of the noise-  
 203 free and the noisy SWOT SSH. For each pass, we calculate the cross-track averaged, along-track  
 204 power spectrum. The spectra are then averaged out over each season, leading to one spectrum  
 205 per season. Information on the wavenumber spectrum calculations is given in Appendix C.  
 206 Again, to evaluate the improvement after the application of the different de-noising techniques  
 207 and parameters, we compare the noise-free and de-noised fields. To do so, the Mean Spectral  
 208 Ratio (MSR) is computed from the power spectral densities (PSD) of SSH. For each season,  
 209 MSR is computed as:

$$MSR = \sqrt{\frac{1}{\sum_{j=1}^{N_k} \delta k_j} \sum_{j=1}^{N_k} \left( \left( \log_{10} \left( \frac{PSD_j(h^{true})}{PSD_j(h)} \right) \right)^2 \times \delta k_j \right)}, \quad (6)$$

210 where  $N_k$  is the number of wavelengths considered;  $PSD_j(h^{true})$  and  $PSD_j(h)$  are the power  
 211 spectral density values at wavelength  $j$  for the original, noise-free SWOT field and the de-  
 212 noised SWOT field, respectively. The considered wavelengths span the interval from 9 km, the  
 213 approximate effective resolution of NATl60, to 200 km, the size of images along-track. MSR is  
 214 defined above so that the best score is 0.  
 215

### 217 3.3. Exploring parameters of the de-noising methods

218 For all de-noising methods, a wide range of parameters are tested to identify optimal pa-  
 219 rameters according to the diagnostics presented in Section 3.2. The convolution-based methods  
 220 use a single parameter that can easily be compared with the image dimensions in pixels. For  
 221 the boxcar kernel, the tested parameter values go from 3 to 200 km, and correspond to the  
 222 size of the box in pixels (1 km in our case). For the Gaussian kernel, the tested parameters go  
 223 from 0.25 to 300 and correspond to the standard deviation, in pixels (we test up to a big sigma  
 224 to have a highly oversmoothed image to reach the limit of the method). On the contrary, the  
 225 geometric interpretation of the parameters of the variational method is not straightforward,  
 226 and a wide exploration of the parameter space must be undertaken. However, due to computa-  
 227 tion time limitations, this cannot be performed in a strictly systematic manner. The adopted  
 228 procedure is detailed below.

#### 229 3.3.1. Orders of magnitude of the cost function terms

230 The orders of magnitude of the terms  $\|\nabla h\|^2$ ,  $\|\Delta h\|^2$  and  $\|\nabla \Delta h\|^2$  composing the cost func-  
 231 tion (eq. 1) are estimated to coarsely scale the parameters  $\lambda_1$ ,  $\lambda_2$  and  $\lambda_3$ . The rationale is, for  
 232 one of these terms (with its weight) to have some impact on the solution, it must be of an order  
 233 of magnitude not too different from the background term  $\|m \circ (h - h^{obs})\|^2$ . Figure 3 shows  
 234 the seasonal evolution of the derivative terms, computed from the model in a  $2^\circ \times 2^\circ$  region  
 235 containing the SWOT passes used in this study. The relative ratios between  $\|\nabla h\|^2$ ,  $\|\Delta h\|^2$   
 236 and  $\|\nabla \Delta h\|^2$  are approximately 1000:10:1. Therefore, if we want to include all three terms in  
 237 the cost function, the ratios between  $\lambda_1$ ,  $\lambda_2$  and  $\lambda_3$  should coarsely be 1:100:1000. Those ratios  
 238 must be only considered as a guideline to start the investigation, not a strict rule. Note that the  
 239 order of magnitude of the background term after minimization of the cost function is thought

240 to be in the range 1 to 100 in the same region. This has been estimated using the noise-free field.  
 241

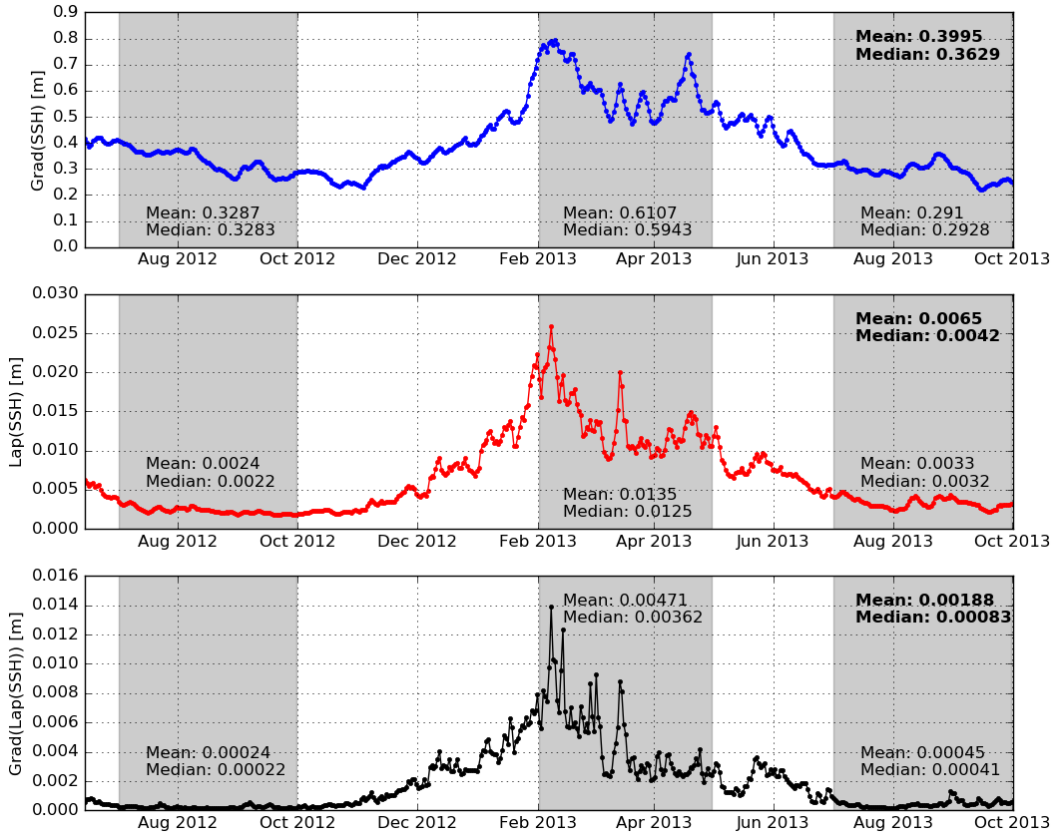


Figure 3: Seasonal variations of the cost function terms  $\|\nabla h\|^2$ ,  $\|\Delta h\|^2$  and  $\|\nabla \Delta h\|^2$ , from top to bottom. Shaded areas indicate the JAS12, FMA13 and JAS13 periods from left to right, respectively. The mean and median values are printed for each period and for the whole year (upper right corners, in bold).

### 242 3.3.2. Finding optimal sets of parameters

243 First, we created an exponential series of values to be tested for the three lambdas, consistently with the previously estimated relative ratios. For  $\lambda_1$ , the series is chosen as  $\{4^n, n = 0, \dots, 7\}$ . For  $\lambda_2$  and  $\lambda_3$ , the series are  $\{10 \times 4^n, n = 0, \dots, 7\}$  and  $\{100 \times 4^n, n = 0, \dots, 7\}$ , respectively. With these, six scenarios of DP de-noising are investigated, including one, two, or three penalization terms in the cost function 1. Three scenarios out of the six considered include a single penalization term (mono-parametric) of order 1, 2 and 3, successively. The other scenarios are made of terms of orders 1 and 2, 2 and 3 and the last one includes the three orders. For conciseness, particularly in the next section, we refer to the variational method with the first order term only as the  $\lambda_1$ -method. We similarly refer to the  $\lambda_2$ -method and to the  $(\lambda_1 + \lambda_2)$ -method when the first two penalization terms are considered, and so on.

253  
 254 For each scenario, a two-step procedure is implemented to identify an optimal set of parameters. In a first step, de-noising of the full set of images is performed with all possible combinations of parameters permitted by the scenario and the parameter series defined previously. RMSEs and MSR are computed for all the combinations. In a second step, refined series of parameters are created in the neighborhood of the combination of parameters that yields the minimum RMSE and MSR scores. Image de-noising is then carried out again with all possible combinations of these series.

## 4. Optimal de-noising method

In this Section, the optimal de-noising method is searched for based on the RMSE and MSR scores described in Section 3. We investigate the KaRIn-noise-only scenario, then the all noises scenario, and finally have a closer look at the method identified as optimal. As it becomes clear in what follows, the notion of optimality does not only refer to qualitative measures. The design of a single index summarizing the performance of the method for the different RMSEs is indeed subjective. Moreover, we take into account the ease of implementation and parameterization as a criteria in the final decision. Minimum values of RSME and MSR for each season, method, and variable are reported in tables 1 and 2 for the KaRIn-only and all noises scenarios, respectively. RMSE scores are actually expressed as the percentage of the original RMSEs, i.e. those of the original, noisy data. For SSH,  $|\nabla\text{SSH}|$  and  $\Delta\text{SSH}$  RMSE and MSR of each de-noising configuration and parameterization, the scores do not necessarily correspond to the same optimal parameter (box size,  $\sigma$  or  $\lambda$ ).

### 4.1. RMSE and MSR scores with KaRIn noise only

For all variables ( $h$ ,  $\nabla h$ , and  $\Delta h$ ), all seasons and all de-noising methods, minimum RMSEs are smaller in Summer than in Winter (table 1). This is expected because the oceanic surface features in Winter are smaller than in Summer [32], so their observation is more affected by the KaRIn noise. Also, smaller structures are more affected by the smoothing due to the de-noising.

For all three seasons and all three variables, RMSEs and MSRs from the convolution-based methods and from the  $\lambda_1$ -method are larger than RMSEs and MSRs from all other variational methods. Also, the  $\lambda_3$ -method provides MSRs significantly higher than the other variational methods. None of these methods is the optimal de-noising one in this KaRIn-only noise configuration, and are not further discussed in the following.

In terms of both RMSEs and MSRs, and among the methods still on course, no method outperforms the others systematically and distinctly. For all three variables, RMSEs are close to each other, with differences less than a very few percents. MSRs are a bit more scattered, but without any clear predominance of a specific method in all seasons. However, the  $\lambda_2$ -method exhibits the lowest MSR values in Summer, and the second lowest value in Winter, close to the  $\lambda_2 + \lambda_3$ -method.

Finally, this analysis persuades us to further examine the  $\lambda_2$ -method for the KaRIn-only scenario (see Section 4.3). This choice is supported by the RMSE and MSR analysis above, which shows that other methods do not beat it clearly, but also by the fact that it is much easier to parametrize a single-parameter method rather than a two or three-parameter method.

### 4.2. RMSE and MSR scores with all noises

Normalized minimum RMSEs for  $h$  and  $|\nabla h|$  are higher than in the KaRIn-only scenario, by factors of 6 – 12 for  $h$ , and 1.5 – 4 for  $|\nabla h|$  (table 2). This is obviously due to the spatially correlated component of the noise (see Figure 2), which is not filtered out by any of the methods used here. Other approaches must be used to remove the correlated noise in order to obtain more accurate estimates.

Contrary to  $h$  and  $\nabla h$ , RMSEs for  $\Delta h$  are comparable with those obtained in the KaRIn-only case. They are 5% higher only. This slight increase in RMSE is the signature of the nonlinear (quadratic, more precisely) component of the correlated error, due to the baseline dilation [17, 16]. The other components are constant, linear or piecewise linear, thus are removed

308 by the second-order derivatives.

309

310 Considering only RMSEs on  $\Delta h$ , except for the boxcar and the  $\lambda_1$ , no method performs  
311 significantly better than the others, and RMSEs are higher in Winter than in Summer. This  
312 is similar to the KaRIn-only scenario. The Gaussian filter performs comparatively better than  
313 in the KaRIn-only scenario.

314

315 In terms of MSRs, the methods involving  $\lambda_2$  perform significantly better than the others,  
316 including the  $\lambda_3$ - and the Gaussian methods. These last two exhibit MSR larger than the others  
317 by factors of 1.5 to 4. In Winter, the  $\lambda_2$ -method is a little less effective than the multi-parameter  
318 methods, with a MSR twice as large.

319

320 The de-noising experiments with all noises, like those with the KaRIn noise only, lead us to  
321 favor the  $\lambda_2$ -method. The reasons are similar: based on RMSEs and MSRs, the method com-  
322 pares favorably with the others, and a single-parameter method is much easier to parametrize.  
323 The only result speaking against this choice is the MSR in Winter. Considering the score value  
324 though, and after the examination of the wavenumber spectra (see Figure 7), this point hardly  
325 justifies the disqualification of the  $\lambda_2$ -method.

#### 326 *4.3. A focus on the second-order variational method*

327 This Section investigates the sensitivity of the  $\lambda_2$ -de-noising to the parameter value. Figure  
328 4 shows the RMSEs for  $h$ ,  $|\nabla h|$ , and  $\Delta h$ , and the MSR for  $h$  as functions of  $\lambda_2$ . On each  
329 graph, the three seasons are shown for both KaRIn-only (solid lines) and all noises (dashed  
330 lines) scenarios, making a total of 6 curves.

331

332 Except for  $h$  and  $|\nabla h|$  RMSEs in the all noises scenario, all RMSE and MSR curves exhibit a  
333 clear minimum point, which indicates the existence of an optimal, or a range of close-to-optimal  
334  $\lambda_2$  values for the de-noising. Optimal values are larger in Summer than in Winter. This is very  
335 likely because small-scale dynamics are amplified in Winter [10]. Large  $\lambda_2$  values tend to over-  
336 smooth the SSH field in Winter, leading to higher residual errors. The seasonal difference in  
337 optimal  $\lambda_2$  values is particularly evident with MSR, with 100 in Winter, and 350 in Summer.  
338 RMSEs for  $h$  and  $\nabla h$  in the all noises scenario are dominated by the correlated SWOT errors,  
339 which remain present after de-noising. Consistently with the analysis of the previous section,  
340 those RMSEs are much higher in the all noises than in the KaRIn-only scenario.

341

342 In Summer, it is possible to identify a range of  $\lambda_2$  values that are close-to-optimal for the  
343 four scores concomitantly. The same holds in Winter for the two scores not dominated by the  
344 correlated errors (RMSEs on  $\Delta h$  and MSR). On Figure 4, horizontal error bars indicate the  
345 range of  $\lambda_2$  values that provide scores higher than the minimum by less than 5%. Subjectively  
346 based on this information, we propose  $\lambda_2$  intervals of [300 – 400] in Summer and [100 – 120]  
347 in Winter. MSR results for the two Summer seasons indicate slightly different optimal values,  
348 suggesting that the choice of a  $\lambda_2$  is inevitably subject to a part of subjectivity if no additional  
349 information on the ocean surface dynamics is available. Not detailed here, the results from the  
350 other (single or multiple-parameter) configurations of variational de-noising also exhibit such  
351 overlaps of intervals, except for  $\lambda_1$ .



Table 1: Scores summary of the different de-noising methods for the just KaRIn dataset (h=SSH).

Season	De-noising		RMSE <sub>r</sub>		Minimum MSR	
	Method	SSH	$ \nabla\text{SSH} $	$\Delta\text{SSH}$		
JAS12	Boxcar	12.43	0.094	0.300	0.2010	
	Gaussian	12.88	0.084	0.251	0.1111	
	DP	1	12.55	0.084	0.279	0.2028
		2	08.71	0.050	0.247	0.0143
		3	09.06	0.051	0.247	0.1021
		1 + 2	08.72	0.050	0.247	0.0192
		2 + 3	08.68	0.049	0.247	0.0205
		1 + 2 + 3	08.66	0.049	0.246	0.0259
FMA13	Boxcar	15.04	0.177	0.511	0.1066	
	Gaussian	13.60	0.153	0.437	0.0746	
	DP	1	15.41	0.173	0.483	0.1498
		2	10.92	0.115	0.420	0.0178
		3	10.86	0.113	0.416	0.0682
		1 + 2	10.92	0.115	0.420	0.0208
		2 + 3	10.79	0.113	0.416	0.0168
		1 + 2 + 3	10.82	0.113	0.416	0.0255
JAS13	Boxcar	11.98	0.086	0.326	0.1796	
	Gaussian	12.81	0.076	0.277	0.0911	
	DP	1	12.78	0.083	0.309	0.2031
		2	08.96	0.053	0.274	0.0216
		3	09.11	0.053	0.273	0.1010
		1 + 2	08.97	0.053	0.274	0.0394
		2 + 3	08.84	0.052	0.272	0.0243
		1 + 2 + 3	08.84	0.052	0.272	0.0269

Table 2: Scores summary of the different de-noising methods for the all noises dataset (h=SSH).

Season	De-noising		RMSE <sub>r</sub>		Minimum MSR	
	Method	SSH	$ \nabla\text{SSH} $	$\Delta\text{SSH}$		
JAS12	Boxcar	90.31	0.171	0.303	0.2024	
	Gaussian	90.10	0.156	0.264	0.1181	
	DP	1	87.60	0.159	0.281	0.1922
		2	90.57	0.174	0.261	0.0307
		3	90.22	0.156	0.265	0.1359
		1 + 2	87.61	0.158	0.262	0.0328
		2 + 3	90.22	0.156	0.261	0.0391
		1 + 2 + 3	87.40	0.156	0.261	0.0395
FMA13	Boxcar	90.88	0.250	0.511	0.1274	
	Gaussian	90.76	0.221	0.435	0.0515	
	DP	1	89.89	0.237	0.484	0.1415
		2	91.11	0.226	0.432	0.0314
		3	90.96	0.226	0.432	0.0868
		1 + 2	89.90	0.223	0.435	0.0160
		2 + 3	91.00	0.226	0.430	0.0177
		1 + 2 + 3	89.82	0.220	0.430	0.0203
JAS13	Boxcar	89.73	0.137	0.328	0.1792	
	Gaussian	89.18	0.126	0.289	0.1152	
	DP	1	84.30	0.131	0.310	0.1895
		2	90.36	0.142	0.287	0.0254
		3	89.66	0.127	0.290	0.1251
		1 + 2	84.19	0.131	0.287	0.0237
		2 + 3	89.66	0.127	0.286	0.0285
		1 + 2 + 3	83.75	0.127	0.286	0.0267

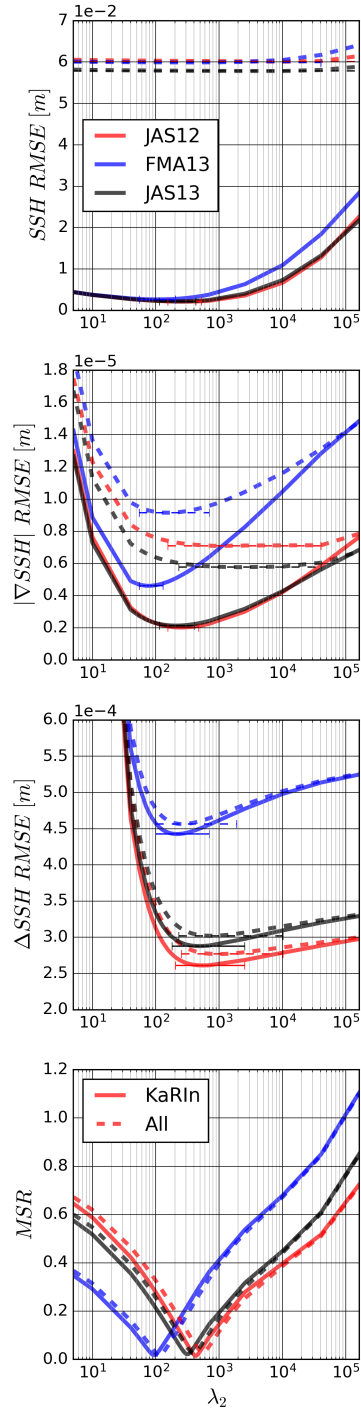


Figure 4: Scores of RMSE and MSR of  $\lambda_2$ -method from just KaRIn (solid line) and all noises (dashed line) for all 3 seasons. Horizontal error bars in the RMSE plots show the the range of  $\lambda_2$  values that provide scores higher than the minimum RMSE by less than 5%.

## 5. Retrieved SWOT fields and spatial spectra

Figures 5 and 6 illustrate the de-noising with the  $\lambda_2$ -method on SWOT passes, and are presented in the same format:  $h$  on the top panel,  $|\nabla h|$  on the central panel, and  $\Delta h$  on the bottom panel. The first panel shows, from left to right: the original, noise-free  $h$  field,  $h$  with KaRIn noise only,  $h$  with all sources of noise, the de-noised KaRIn-only  $h$ , and the de-noised all noises  $h$ . The second and third panels show the corresponding  $|\nabla h|$  and  $\Delta h$ , respectively. Figures 5a and 5b exhibit Summer time scenes with high and low correlated SWOT noise, respectively; Figures 6a and 6b are similar for Winter.

In all cases, de-noising leads to correct orders of magnitude for all fields, and particularly for  $|\nabla h|$  and  $\Delta h$ . As expected and already shown by [18, 33], the original SWOT data affected by small-scale noise does not provide any useful information about SSH derivatives. The de-noising method corrects this efficiently, and makes it possible to identify the main structural characteristics of the fields.

A strong spatially-correlated noise shows strong signatures on  $h$ , moderate signatures on  $|\nabla h|$ , and low signatures on  $\Delta h$ , except at the outer boundaries of the swath. The low signature on  $\Delta h$  was already observed in the RMSEs, and is due to the specific spatial structure of this noise. Most components are linear in the across-track direction. In the along-track direction, error correlations are high (Figure 2). Therefore, the correlated noise has a low effect on the second-order derivatives. The remaining noise at the outer boundaries is due to the finite difference method used to compute the derivatives described in Appendix A.

Although the resulting fields of  $\Delta h$  fall in correct orders of magnitude and capture the structure of the true fields at the scale of the swath, the kilometeric-scale fronts and filaments are smoothed out by the de-noising. Solving this issue would require the development of more sophisticated de-noising techniques, or a post-processing of the present result including, for example, some ocean dynamics through data assimilation techniques. This will be a natural step forward, since the first motivation for developing a de-noising technique constraining  $\Delta h$  is precisely the combined assimilation of  $h$  and its first two derivatives, as stated in the introduction.

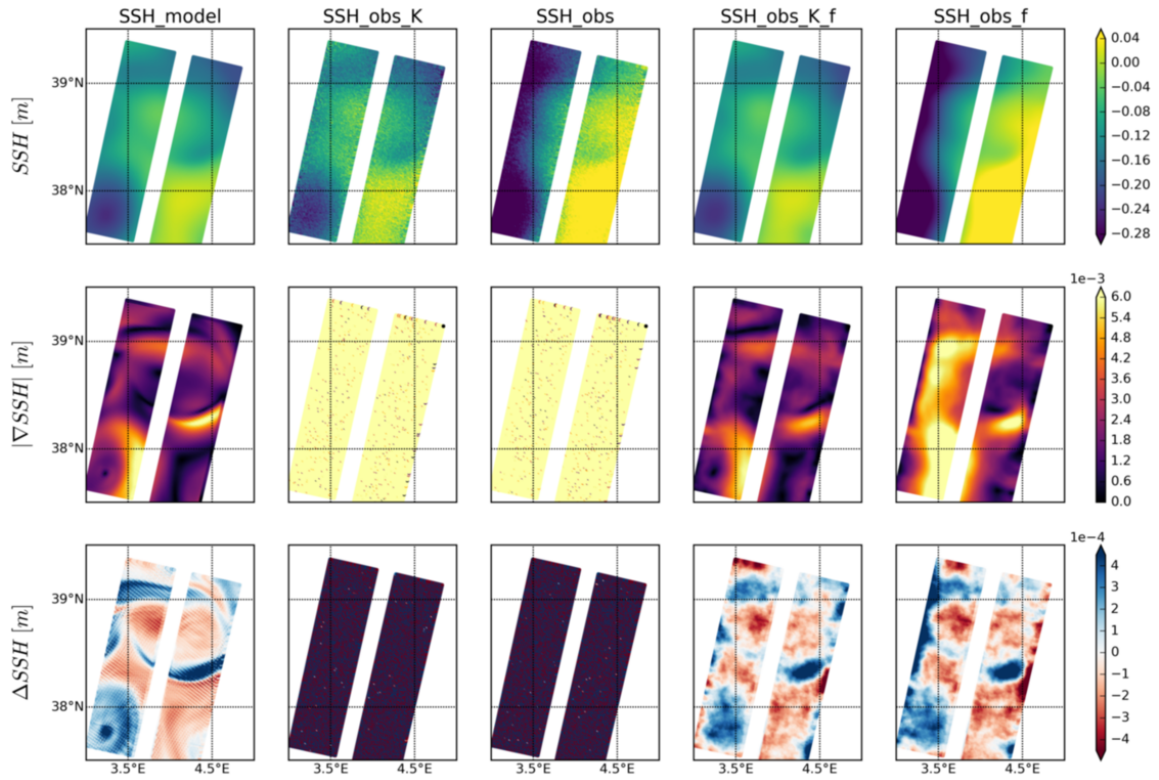
Figure 7 shows power spectral densities (PSD) of  $h$ . The rows distinguish the just KaRIn noise added and the all noises cases. The columns are for Summer 2012, Winter 2013, and Summer 2013. On each graph, the spectra are shown for the noise-free data (SSH\_model), the noisy data (SSH\_obs), the de-noised data (SSH\_obs.f), the pre-de-noising noise (noise) and post-de-noising noise (noise.f). The de-noised data have been obtained with the  $\lambda_2$ -method with parameter values chosen in the intervals identified in Section 4, and indicated on each graph.

From this spectral viewpoint, the de-noised data matches the noise-free data well at all scales down to  $\sim 15$  km. In the noisy data, the noise amplitude approaches the signal amplitude at wavelengths of 50 km in Summer and 40 km in Winter, and dominates the signal at shorter wavelengths. This is efficiently corrected by the de-noising. The process seems more efficient in Winter than in Summer, probably because of higher PSDs in Winter related to more intense ocean surface processes.

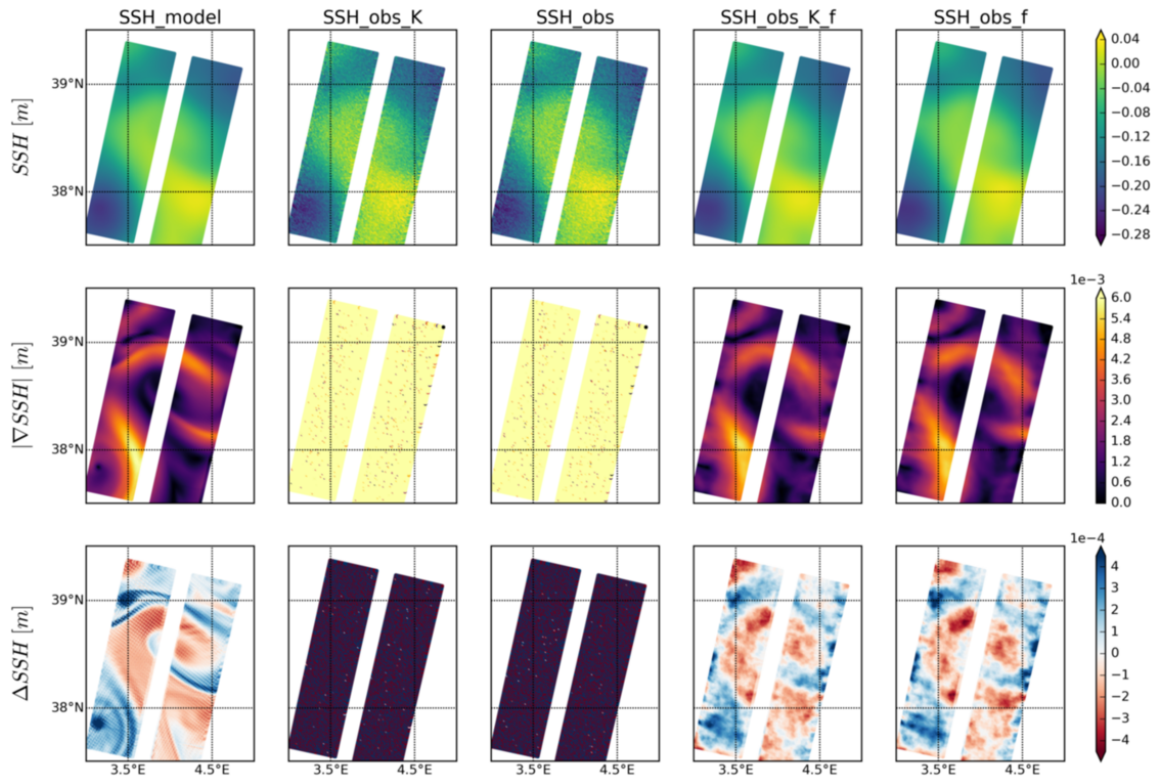
Following the definition proposed by Wang et al. [34] for the spatial scale resolved by SWOT, the de-noising reduces this scale by a factor of 2, leading to resolved scales of between approximately 20 and 30 km. Wang et al. [34] define the spatial scale resolved by SWOT by

404 the wavelength at which the SWOT noise spectrum intersects the spectrum of the true signal  
405 (SSH\_model here). Figure 7 indicates resolved scales of 50, 40, and 50 km in the JAS12, FMA13  
406 and JAS13 scenarios respectively, in both just KaRIn and all noises cases. After de-noising,  
407 the resolved scales are reduced to 25, 20, 20 km in the KaRIn-only case and 30, 20, 30 in the  
408 all noises case. Even below these scales, the noise left is very low, and within the variability of  
409 SSH\_model (red envelope in Figure 7). At wavelengths near 10 km, the noise is reduced by  $10^4$   
410 in the JAS scenarios and  $10^3$  in the FMA scenario.

411

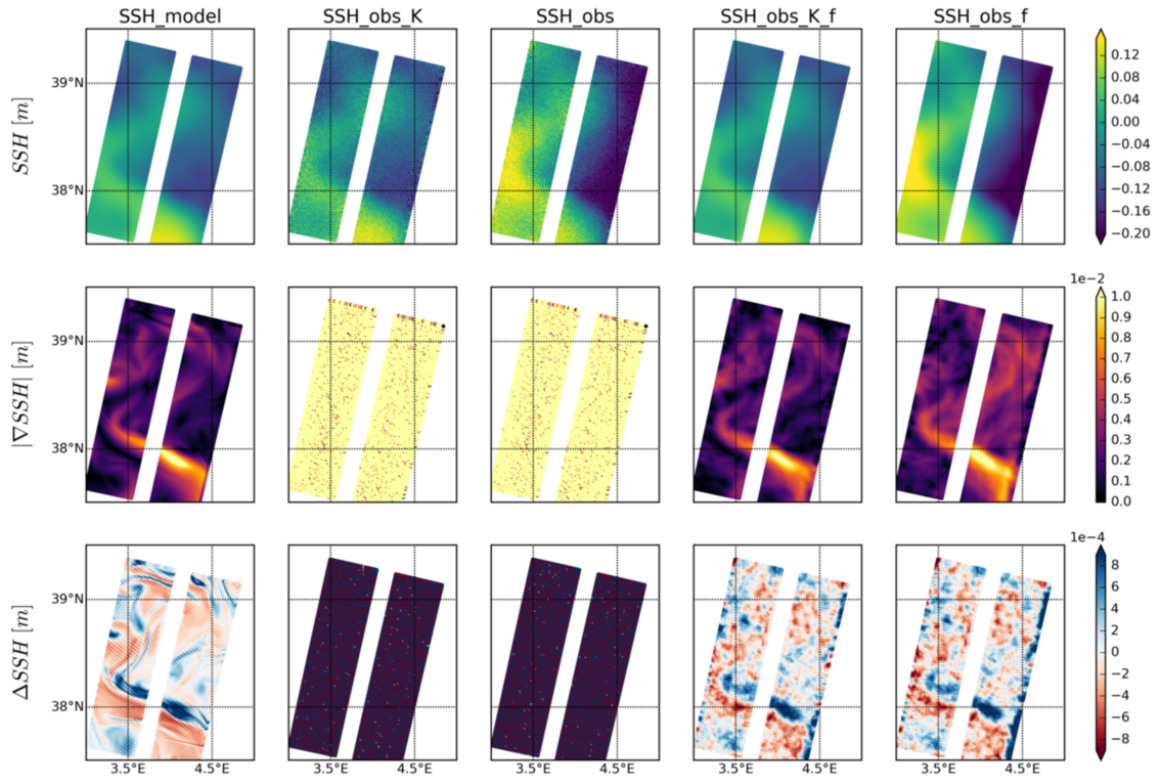


(a)

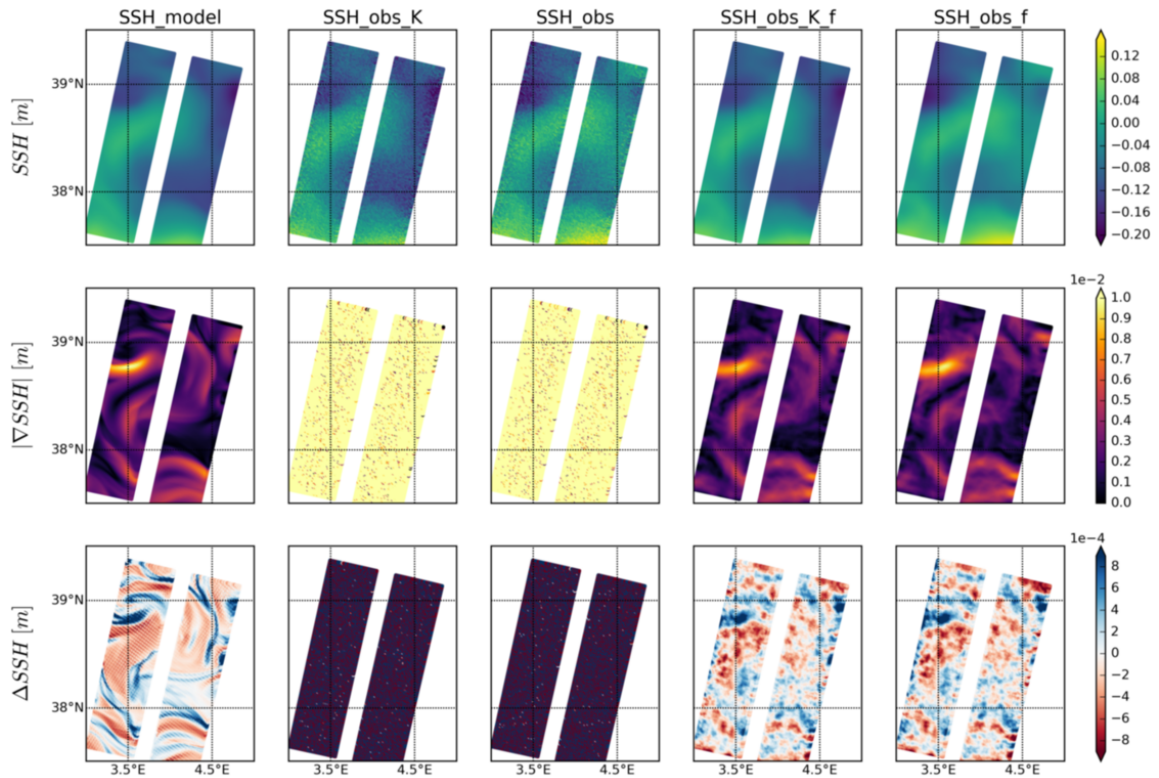


(b)

Figure 5: Fields of pass 09, cycle 2 (a) and 6 (b) of JAS12 dataset compared to the fields filtered with  $\lambda_2 = 430$ . From top to bottom: SSH, gradient of SSH and laplacian of SSH. From left to right: model interpolated to SWOT grid (SSH\_model), SSH\_model + KaRIn noise (SSH\_obs\_K), SSH\_model + all noises (SSH\_obs), filtered SSH\_obs\_K (SSH\_obs\_K\_f) and filtered SSH\_obs (SSH\_obs\_f).



(a)



(b)

Figure 6: Fields of pass 09, cycle 2 (a) and 6 (b) of FMA13 dataset compared to the fields filtered with  $\lambda_2 = 95$ . From top to bottom: SSH, gradient of SSH and laplacian of SSH. From left to right: model interpolated to SWOT grid (SSH\_model), SSH\_model + KaRIn noise (SSH\_obs\_K), SSH\_model + all noises (SSH\_obs), filtered SSH\_obs\_K (SSH\_obs\_K\_f) and filtered SSH\_obs (SSH\_obs\_f).



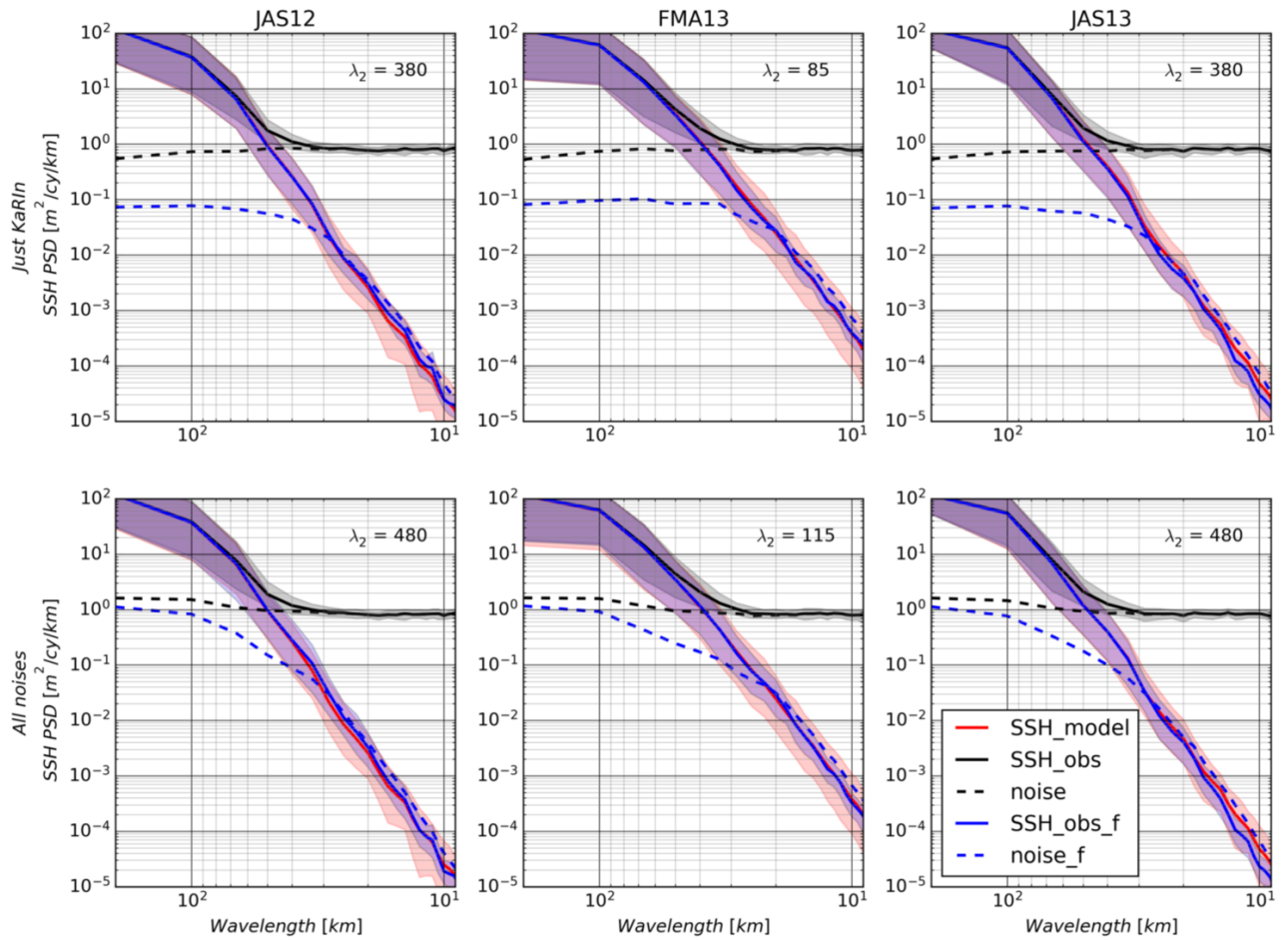


Figure 7: Spatial spectra of the model interpolated data (SSH\_model) are shown in red and of the pseudo-SWOT data (SSH\_obs) in black. Blue lines indicate the filtered pseudo-SWOT spectra (SSH\_obs\_f) obtained with the optimal  $\lambda_2$  found with the MSR score. The dashed lines are the noise spectra of SSH\_obs (noise) and SSH\_obs\_f (noise\_f). Shaded areas show values between the 5th and 95th percentiles, showing the PSD variability. Top row shows pseudo-SWOT data with just KaRIn noise added and bottom row with all noises. Columns represent the different seasonal datasets from left to right: Summer 2012 (JAS12), Winter 2013 (FMA13) and Summer 2013 (JAS13).



## 6. Discussion and Conclusions

Several objectives of the SWOT mission will be met only if the small-scale noise affecting the data can be efficiently removed. Small-scale noise, in particular the spatially uncorrelated KaRIn instrument noise, prevents the computation of horizontal SSH derivatives. This limits both the direct estimation of relevant oceanic variables on the SWOT swath, and the use of SWOT data to build gridded products of altimetry.

To remove the small-scale SWOT noise, we propose a de-noising method that performs better than conventional convolution-based methods both in terms of RMSE (physical space diagnostic) and spectra. The method, which originates from image processing applications, is based on the regularization of the SWOT SSH data by the penalization of its derivatives of orders 1 to 3 in a variational, optimization framework. This approach is chosen because it is in close connection with the oceanic variables of interest, namely geostrophic velocity and vorticity. After a thorough evaluation based on a large number of simulated SWOT scenes, the variational de-noising method exhibits better performance than standard, boxcar and Gaussian filters. We find the method performs best when only the second-order derivative ( $\lambda_2$ ) is considered in the cost function. Only one parameter needs to be set, which makes the parameterization of the method as simple as a convolution-based method. We find that this parameter can be set smaller or larger in function of the characteristics of our field: the higher the intensity of the signal, the lower the penalization and thus the value of the parameter (as we find in the FMA13  $\lambda_2$  values in contrast to JAS12). Also, if the noise level in our fields is higher (all noises scenario), the more we need to penalize and the larger the parameter value. In other words, the higher the signal to noise ratio (SNR) the less we need to penalize our field, and so the smaller  $\lambda_2$ .

The method will require further investigations before operational applications, since we have focused our attention to one particular region (the Western Mediterranean Sea), with an ocean circulation free of tidal forcing, and a prescribed Significant Wave Height (SWH) of 2 m. The present study shows that in one single region, the range of optimal parameters changes with the season, due to seasonal changes in the ocean surface dynamics. Similar conclusions are certainly expected with respect to regional and dynamical regimes. The NATL60 simulation used here does not include tidal forcing. The behavior and efficiency of the de-noising method may be questioned in presence of tidal motions and particularly tide-generated internal waves. Finally, the SWH prescribed in the SWOT simulator to compute the KaRIn error amplitude is prescribed to 2 m. As the SWH varies geographically and according to the atmospheric regime, KaRIn errors smaller or larger than those computed for the present study with the SWOT simulator can be expected [34]. The first two aspects (geographic variations of ocean dynamics and internal tides) are presently under study using data from several high-resolution simulations that include tidal forcing: the HYbrid Coordinate Ocean Model (HYCOM) [35], the Massachusetts Institute of Technology general circulation model (MITgcm) [36], and the recent extended NATL60 –eNATL60– simulation, not yet published).

The method should also benefit from additional developments to reconstruct more realistic fields of relative vorticity on the SWOT swath, and could ultimately lead to the estimation of vertical velocities. The de-noising process inevitably smoothes out the very fine-scale, elongated structures usually visible in surface relative vorticity fields [10, e.g.]. Restoring these structures should be investigated, perhaps using appropriate image processing techniques [37, 38], or methods already developed in the oceanographic community such as Lagrangian advection [39, 40]. Dynamical models could also be used in a data assimilation framework.

To conclude, the de-noising method opens the way to several relevant applications using the

462 SWOT data, possibly including SWOT data validation, assimilation, and SSH mapping. We  
463 mention SWOT data validation due to the in-painting capability of the variational de-noising  
464 method, i.e. the fact that the process naturally fills the 20-km gap of the SWOT swath (the  
465 gap is in-painted, and emptied again after de-noising to restore SWOT data in the original  
466 shape). In other words, the SWOT KaRIn data are interpolated on the track of the SWOT  
467 nadir altimeter. This is obviously relevant for data comparison and validation. De-noising is  
468 also interesting to pre-process the SWOT data before their assimilation in ocean circulation  
469 models. This actually was a primary motivation for the method development. Computing  
470 spatial derivatives of the SWOT data allows the implementation of data assimilation methods  
471 that account for SWOT error correlations [14, 15]. Alternatively, the relative vorticity derived  
472 from the de-noising can be directly assimilated. This option has not been explored yet to our  
473 knowledge. Finally, the de-noising can also be combined with other techniques to improve the  
474 assimilation. We particularly think about the technique recently developed by Metref et al.  
475 [41] to significantly reduce the impact of the geometrically structured, highly correlated SWOT  
476 errors (roll, phase, timing, and baseline errors).

477

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## 590 Appendix A. Calculation of Laplacian

Laplacian are computed using finite differences, following the method proposed by [22]. We note  $h$  the image of size  $N_x \times N_y$ . In a first step, the two components of the gradient are computed as ( $i = 1, \dots, N_x; j = 1, \dots, N_y$ ):

$$\begin{aligned} (\nabla h)_{i,j}^x &= h_{i+1,j} - h_{i,j} & \text{if } i < N_x \\ &= 0 & \text{if } i = N_x \\ (\nabla h)_{i,j}^y &= h_{i,j+1} - h_{i,j} & \text{if } j < N_y \\ &= 0 & \text{if } j = N_y \end{aligned}$$

In a second step, Laplacian is computed as the divergence of the gradient. Divergence of vector  $\mathbf{a} = (a^x, a^y)$  is computed as:

$$\text{div}(\mathbf{a}) = b_{i,j}^x + b_{i,j}^y$$

where:

$$b_{i,j}^x = \begin{cases} a_{i,j}^x - a_{i-1,j}^x & \text{if } 1 < i < N_x \\ a_{i,j}^x & \text{if } i = 1 \\ -a_{i-1,j}^x & \text{if } i = N_x \end{cases}$$

and

$$b_{i,j}^y = \begin{cases} a_{i,j}^y - a_{i,j-1}^y & \text{if } 1 < j < N_y \\ a_{i,j}^y & \text{if } j = 1 \\ -a_{i,j-1}^y & \text{if } j = N_y \end{cases}$$

591 The scheme implemented at the boundaries preserves the image size, contrary to what a stan-  
592 dard five-point stencil Laplacian operator would do. Preservation of image size is essential in  
593 the gradient descent iterations to end up with a final image of size similar to the initial image.

## 594 Appendix B. FISTA

To speed up the gradient descent iterations, the Fast Iterative Shrinkage-Thresholding Algorithm (FISTA) algorithm [24] is implemented. Setting  $t_0 = 1$  and introducing an auxiliary variable  $y$  initialized as  $y^0 = h^0$ , the iterative algorithm of Eq. 3 becomes:

$$\begin{aligned} h^{k+1} &= h^k + \tau (m \circ (h_{obs} - y^k) + \lambda_1 \Delta y^k - \lambda_2 \Delta \Delta y^k + \lambda_3 \Delta \Delta \Delta y^k) \\ t_{k+1} &= (1 + \sqrt{1 + 4t_k^2})/2 \\ y^{k+1} &= h^{k+1} + \frac{t_k - 1}{t_{k+1}} (h^{k+1} - h^k) \end{aligned} \tag{B.1}$$

## 595 Appendix C. Calculation of spatial spectra

596 The spatial spectra used as one of the scores for the de-noising parameterizations are cal-  
597 culated as follows:

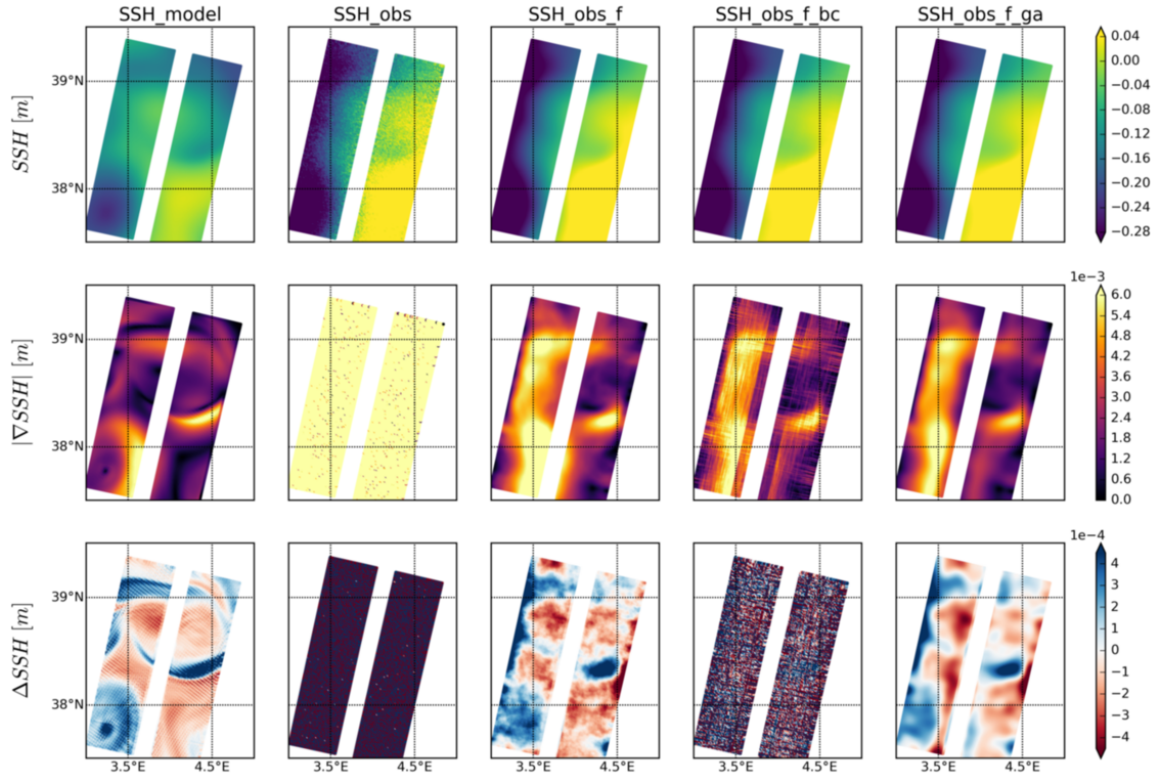
- 598 1. Apply a linear detrending;
- 599 2. Remove the spatial mean;
- 600 3. Apply a Tukey window with a 0.5 fraction of the window inside the cosine tapered region;
- 601 4. Compute the 1D spatial Fourier power spectra along-track for each SSH swath across-track
- 602 dimension.

## 603 Appendix D. Qualitative figures of different methods

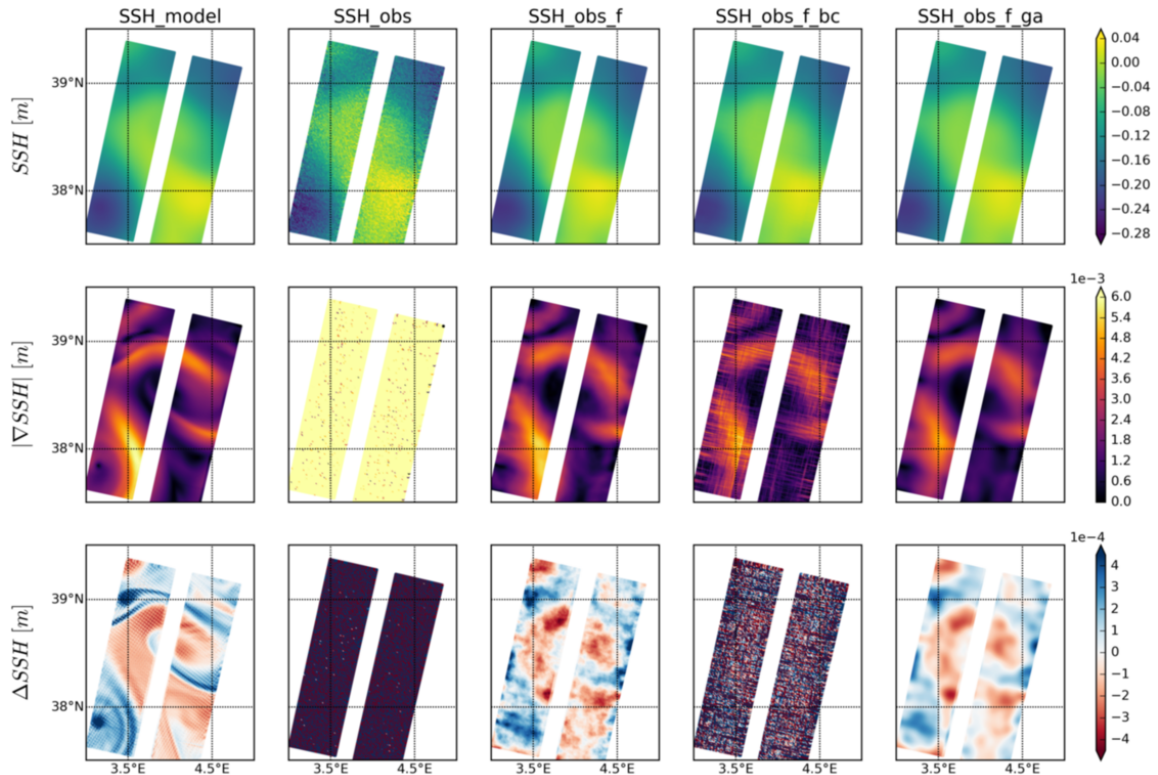
604 To better illustrate the advantage of our de-noising approach, we show in Figures D.8  
605 and D.9 the fields provided by the boxcar and Gaussian methods, corresponding to the  $\lambda_2$   
606 experiments presented in Figures 5 and 6. We only show the all noises scenario. Boxcar  
607 derivatives fields are very noisy, as it is specially visible for the laplacian fields. With the  
608 Gaussian method, the laplacian is less noisy than with our method, but the gradient is over-  
609 smoothed.

## 610 Appendix E. Softwares

- 611 • Standard image techniques: For both boxcar and Gaussian kernel python's `scipy.ndimage`  
612 module was used with the following specific functions:
  - 613 – Boxcar filter: `scipy.ndimage.generic_filter()`
  - 614 – Gaussian: `scipy.ndimage.gaussian()`
- 615 • Variational regularization method: <https://github.com/LauraGomezNavarro/SWOTmodule>



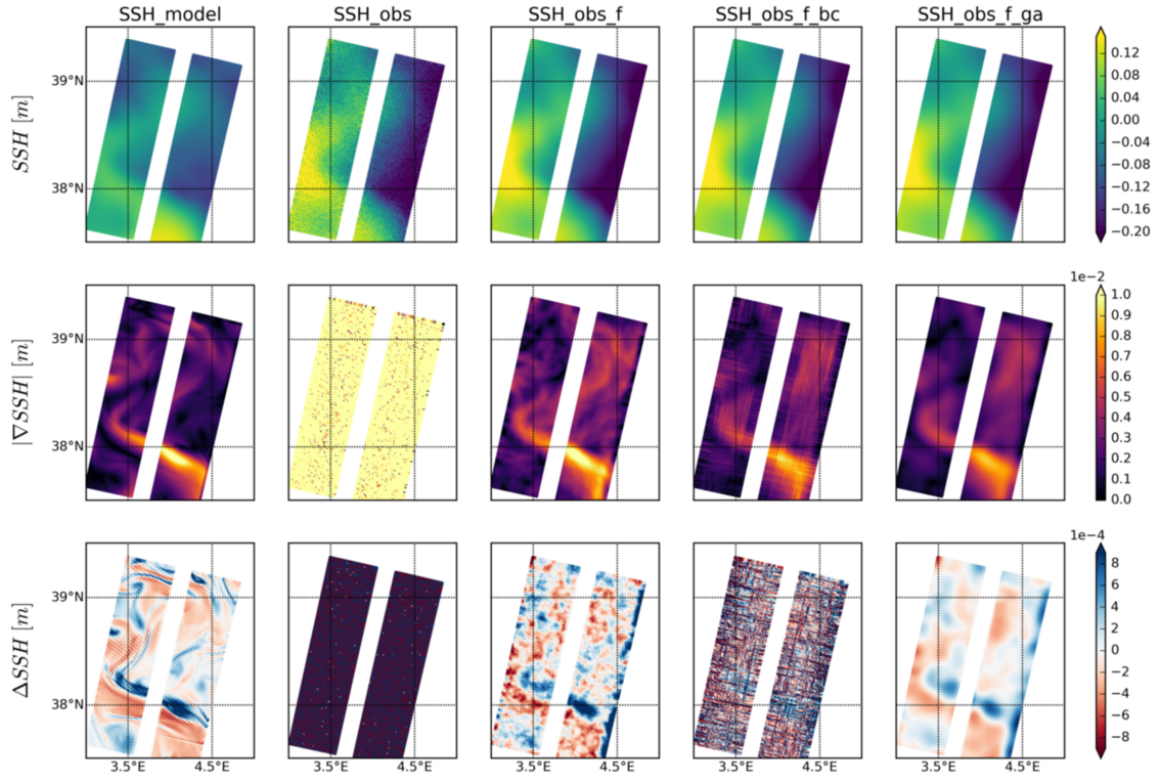
(a)



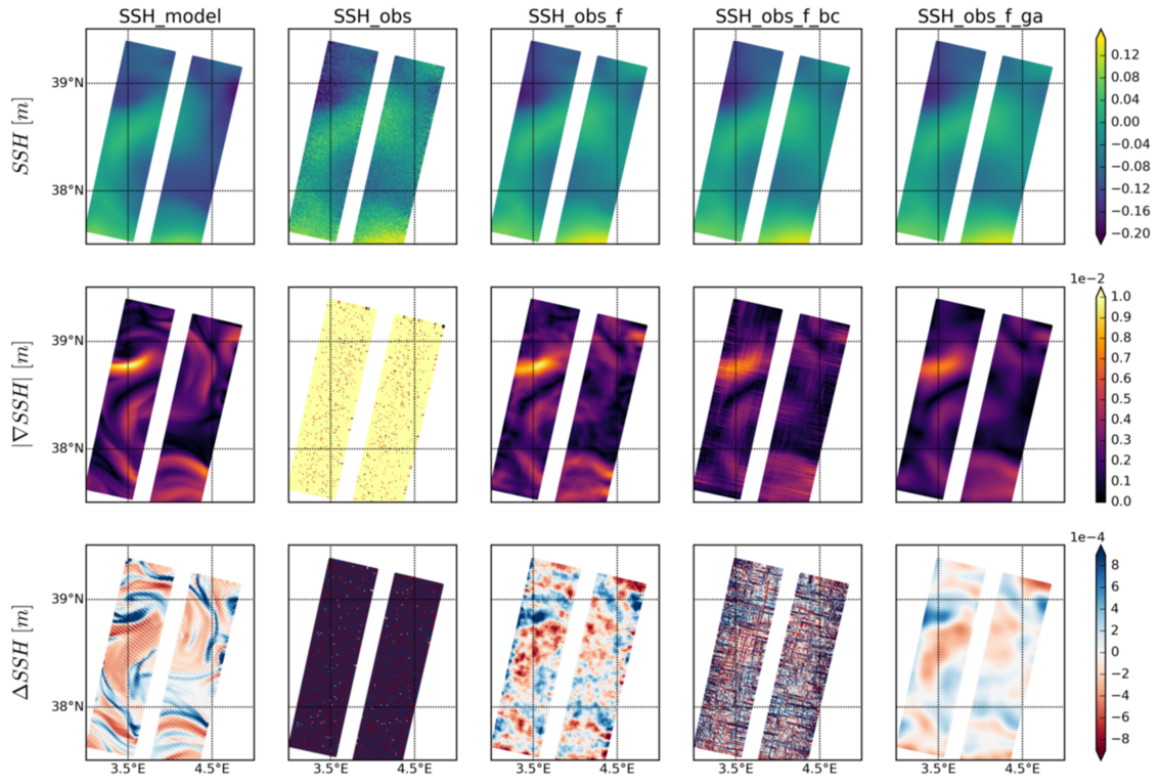
(b)

Figure D.8: Fields of pass 09, cycle 2 (a) and 6 (b) of JAS12, all noises dataset. From left to right: comparison between the SSH\_model, SSH\_obs, SSH\_obs filtered with our approach and  $\lambda_2 = 430$  (SSH\_obs\_f), with the optimal boxcar (SSH\_obs\_f\_bc) and with the optimal Gaussian (SSH\_obs\_f\_ga) methods. From top to bottom: SSH, gradient of SSH and laplacian of SSH





(a)



(b)

Figure D.9: Fields of pass 09, cycle 2 (a) and 6 (b) of FMA13, all noises dataset. From left to right: comparison between the SSH\_model, SSH\_obs, SSH\_obs filtered with our approach and  $\lambda_2 = 95$  (SSH\_obs.f), with the optimal boxcar (SSH\_obs.f.bc) and with the optimal Gaussian (SSH\_obs.f.ga) methods. From top to bottom: SSH, gradient of SSH and laplacian of SSH