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A KINETIC APPROACH OF THE BI-TEMPERATURE EULER MODEL

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ABSTRACT. We are interested in the numerical approximation of the bi-temperature Euler equations, which is a non conservative hyperbolic system introduced in [3]. We consider a conservative underlying kinetic model, the Vlasov-BGK-Poisson system. We perform a scaling on this system in order to obtain its hydrodynamic limit. We present a deterministic numerical method to approximate this kinetic system. The method is shown to be Asymptotic-Preserving in the hydrodynamic limit, which means that any stability condition of the method is independent of any parameter ε , with $\varepsilon \rightarrow 0$. We prove that the method is, under appropriate choices, constant with the solution for bi-temperature Euler. Finally, our method is compared to methods for the fluid model (HLL, Suliciu).

1. INTRODUCTION

Plasma physics is renowned for being a multiscale problem ([11]). Phenomena range from plasma oscillations and charge separation at very small scales to fluid mechanics at macroscopic levels. Different physical models exist in order to describe a plasma according to the scale of interest. The kinetic model, constituted of the Vlasov-Boltzmann equations, coupled with Maxwell equations, are well suited for mesoscopic descriptions. In this model, each population of particles α (in the case of a plasma, these are electrons, ions and neutral atoms and molecules) is represented via the use of a distribution function f^α . This function, defined over the phase space, denotes the number of particles that, at a given time t at a given point x , possess a velocity v . Via an hydrodynamic limit, equations on the macroscopic conserved quantities (mass, momentum and energy) can be derived. These equations form the so-called bi-temperature Euler system (see [3] for the electrostatic case, and [7] for the transverse magnetic case). From a numerical point of view, these equations are less accurate than the underlying kinetic model, but they are cheaper to solve.

The bi-temperature Euler system is comprised of an equation over mass, an equation over momentum, and one equation over each species energy. These energy equations are non-conservative,

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accounting for the fact that, in a plasma, electrons and ions are exchanging energy through collisions and via electromagnetic interactions. Contrarily to the usual Euler system, for which the exact solution of any Riemann problem is known, these non-conservative terms prevent the knowledge of such solutions, especially when the solution contains a shock. This is a consequence that these non-conservative terms are written as products of distributions, which does not possess a clear definition. [15] presents a way of defining these products using a theory of paths. It is made clear that a certain amount of arbitrariness is needed in the choice of such paths. A major drawback of this property is that, when attempting numerical resolution of this system, every numerical scheme chooses its own path, and the Rankine-Hugoniot relation holding across shocks will be different for every scheme. This phenomenon is highlighted in [3]. More precisely, different schemes have different numerical viscosity, and the shock relations depend on such viscosities. Several solutions have been proposed in order to derive numerical methods that use the notion of paths to define jump relations ([23], [9]). Other approaches have been proposed for different kinds of non-conservative systems (see [10], [19], [22], [24], [1]). Finally, Bouchut's book [5] presents an overview of numerical methods for non-conservative Shallow-water equations.

In this paper, the aim is to propose a reference numerical method for such solutions using the underlying kinetic model. This kinetic model is conservative, and hence does not exhibit the drawback of the macroscopic model. Hence, the idea dwells in solving the Vlasov-BGK-Ampère system in the hydrodynamic limit, which will be presented in section 2.3. In order to be able to compare the results with the scheme applied to the bi-temperature Euler system, it is necessary to describe identical scales. By performing a scaling on the kinetic system, dimensionless parameters (scaled Debye length, scaled plasma frequency) are introduced in the system. Taking the hydrodynamic limit then amounts to taking the limit when these parameters tend to zero. In the general case, if a naive numerical approach is used, extremely restrictive stability conditions will appear, rendering the scheme unusable in a decent amount of computation time. Hence, an Asymptotic-Preserving (AP) scheme needs to be derived. Such a scheme possess stability conditions independent of these small parameters and is then able to compute solutions for both cases when parameters are of the order of one and when these parameters are arbitrarily small. The adjective Asymptotic Preserving has been introduced by Shi Jin in [21]. In the study of plasma simulation, numerous several Asymptotic Preserving methods have been proposed for various models, such as Euler-Maxwell, Euler-Poisson, Euler-Lorentz, Vlasov-Poisson ([18], [6], [17], [14], [13], [16], [4]).

The paper is organized as follows. Firstly, the macroscopic and kinetic models relative to our problem are introduced. A reduced 1D3V kinetic model is derived, and the relative properties of moment conservation and entropy dissipation are specified. Subsequently, a scaling of the kinetic model in order to exhibit the limiting behaviour of the system. Then, a Asymptotic-Preserving numerical method is presented. This method is able to compute solutions of the kinetic model in the asymptotic regime. Finally, test cases are presented in order to illustrate the derived method. In particular, the numerical results are compared with an HLL-inspired method and a Suliciu relaxation approach, both derived in [3].

2. DESCRIPTION OF THE MODELS

2.1. Bi-temperature Euler model. The one-dimensional bi-temperature Euler model describes the behaviour of a two species fluid (constituted of electrons and ions) on a macroscopic scale

(see [12]):

$$\begin{aligned}
(1a) \quad & \partial_t \rho + \partial_x(\rho u) = 0, \\
(1b) \quad & \partial_t(\rho u) + \partial_x(\rho u^2 + p^e + p^i) = 0, \\
(1c) \quad & \partial_t(\rho^i \epsilon^i + \frac{1}{2} \rho^i u^2) + \partial_x(u(\rho^i \epsilon^i + \frac{1}{2} \rho^i u^2 + p^i)) + u(c^i \partial_x p^e - c^e \partial_x p^i) \\
& \qquad \qquad \qquad = -\nu_{ei}(T^i - T^e), \\
(1d) \quad & \partial_t(\rho^e \epsilon^e + \frac{1}{2} \rho^e u^2) + \partial_x(u(\rho^e \epsilon^e + \frac{1}{2} \rho^e u^2 + p^e)) - u(c^i \partial_x p^e - c^e \partial_x p^i) \\
& \qquad \qquad \qquad = \nu_{ei}(T^i - T^e).
\end{aligned}$$

Superscripts e and i denote quantities related to electrons and ions, respectively. $\rho = \rho^e + \rho^i = m^e n^e + m^i n^i$ is the total density, m^α being the mass of a particle of species α , and n^α the corresponding concentration. u is the macroscopic velocity of the fluid. p^α, T^α and ϵ^α denote the partial pressures, temperatures and specific energies. $c^\alpha = \frac{m^\alpha}{m^\alpha + m^\beta}$ denotes the mass fractions. ν_{ei} is the temperature exchange rate between the two species, that can depend on space and time.

2.2. Macroscopic quantities. Define $f^\alpha(t, x, v) : \mathbb{R}_+ \times \mathbb{R}^3 \times \mathbb{R}^3 \mapsto \mathbb{R}_+$, the particle distribution function of species α , where t, x, v respectively denote the time, space and microscopic velocity variables. Integration with respect to $v \in \mathbb{R}^3$ will be denoted as follows, for any function g that depends on v such that $(1 + v^2)g \in L^1(\mathbb{R}^3)$:

$$\langle\langle g \rangle\rangle = \int_{\mathbb{R}^3} g dv.$$

Macroscopic quantities $n^\alpha(t, x), u^\alpha(t, x)$ and $T^\alpha(t, x)$ are defined as moments of f^α as follows:

$$\begin{aligned}
\langle\langle f^\alpha \rangle\rangle &= n^\alpha, \\
\langle\langle v f^\alpha \rangle\rangle &= n^\alpha u^\alpha, \\
\langle\langle m^\alpha \frac{v^2}{2} f^\alpha \rangle\rangle &= \frac{1}{2} m^\alpha n^\alpha (u^\alpha)^2 + \frac{3}{2} n^\alpha k_B T^\alpha.
\end{aligned}$$

Similarly, mixture macroscopic quantities are defined as:

$$\begin{aligned}
\langle\langle m^e f^e + m^i f^i \rangle\rangle &= \rho, \\
\langle\langle v(m^e f^e + m^i f^i) \rangle\rangle &= \rho u, \\
\langle\langle \frac{v^2}{2} (m^e f^e + m^i f^i) \rangle\rangle &= \frac{1}{2} \rho u^2 + \frac{3}{2} (n^e + n^i) k_B T.
\end{aligned}$$

Finally, we define electromagnetic quantities: $E \in \mathbb{R}$ is the electric field along x , q^α is the electric charge of species α . Note that $q^e = -Zq^i$, with Z the ionization rate. j and $\bar{\rho}$ are respectively the current and total electric charge, defined as:

$$(2) \quad \langle\langle q^e f^e + q^i f^i \rangle\rangle = \bar{\rho}, \quad \langle\langle v(q^e f^e + q^i f^i) \rangle\rangle = j.$$

2.3. Vlasov-BGK-Ampère system. Consider the Vlasov-BGK-Ampère model, for $t \in \mathbb{R}$, $x \in \mathbb{R}$, $v = (v_1, v_2, v_3) \in \mathbb{R}^3$ and $\alpha \in \{e, i\}$:

$$\begin{aligned}
(3a) \quad & \left\{ \begin{aligned} \partial_t f^\alpha + v_1 \partial_x f^\alpha + \frac{q^\alpha E}{m^\alpha} \partial_{v_1} f^\alpha &= \frac{1}{\tau^\alpha} (M(f^\alpha) - f^\alpha) + \frac{1}{\tau^{\alpha\beta}} (\bar{M}(f^\alpha) - f^\alpha), \end{aligned} \right. \\
(3b) \quad & \left\{ \begin{aligned} \partial_t E &= -\frac{j}{\varepsilon_0}, \end{aligned} \right. \\
(3c) \quad & \left\{ \begin{aligned} \partial_x E &= \frac{\bar{\rho}}{\varepsilon_0}, \end{aligned} \right.
\end{aligned}$$

where τ^α and $\tau^{\alpha\beta}$ are, respectively, the relaxation rate towards intra-species and inter-species equilibria ($\tau^{\alpha\beta} = \tau^{\beta\alpha}$). ε_0 is the dielectric permittivity of vacuum.

The entropy-minimizing distribution function is the local Maxwellian distribution, denoted $M(f^\alpha)$, defined by:

$$M(f^\alpha)(t, x, v) = \frac{n^\alpha(t, x)}{(2\pi k_B \frac{T^\alpha(t, x)}{m^\alpha})^{\frac{3}{2}}} \exp\left(-\frac{|v - u^\alpha(t, x)|^2}{2k_B \frac{T^\alpha(t, x)}{m^\alpha}}\right).$$

$\overline{M}(f^\alpha)$, the exchange Maxwellian distribution, is defined as:

$$\overline{M}(f^\alpha)(t, x, v) = \frac{n^\alpha(t, x)}{(2\pi k_B \frac{T(t, x)}{m^\alpha})^{\frac{3}{2}}} \exp\left(-\frac{|v - u(t, x)|^2}{2k_B \frac{T(t, x)}{m^\alpha}}\right).$$

2.4. Reduced model.

2.4.1. *Definition of the reduced distribution functions.* The problem is assumed to be one dimensional in space and 3 dimensional in velocity (1D3V). Hence, the distribution functions f^α only depend on $x = x_1$ and $v = (v_1, v_2, v_3)$. This means that the function f^α is even in v_2 and v_3 , which can be expressed as:

$$(4a) \quad \left\{ \int_{\mathbb{R}^2} f^\alpha(t, x, v) v_2 dv_2 dv_3 = 0, \right.$$

$$(4b) \quad \left. \int_{\mathbb{R}^2} f^\alpha(t, x, v) v_3 dv_2 dv_3 = 0. \right.$$

Denote $u^\alpha = (u_1^\alpha, u_2^\alpha, u_3^\alpha)$, this assumption gives $u^\alpha = (u_1^\alpha, 0, 0)$. Hence, the notation $u^\alpha = u_1^\alpha$ is used. The number of variables is reduced, as in [2]. Define the marginals, for $\alpha = e, i$, by:

$$(5a) \quad \left\{ g_0^\alpha(t, x, v_1) = \int_{\mathbb{R}^2} f^\alpha(t, x, v) dv_2 dv_3, \right.$$

$$(5b) \quad \left. g_2^\alpha(t, x, v_1) = \int_{\mathbb{R}^2} f^\alpha(t, x, v) \frac{m^\alpha(v_2^2 + v_3^2)}{2} dv_2 dv_3, \right.$$

and denote integration with respect to v_1 , for any function g that depends on v_1 , as:

$$(6) \quad \langle g \rangle = \int_{\mathbb{R}} g dv_1.$$

2.4.2. *Reduced system of equations and its properties.* The idea is to describe the system of equations and the macroscopic quantities using these reduced functions. The system obtained after reduction reads as follows:

Proposition 1. *Under the assumptions (4) and (5), the system (3) can be reduced to the following one, with $\alpha = e, i$ and $p = 0, 2$:*

$$(7a) \quad \left\{ \partial_t g_p^\alpha + v_1 \partial_x g_p^\alpha + \frac{q^\alpha E}{m^\alpha} \partial_{v_1} g_p^\alpha = \frac{1}{\tau^\alpha} (M_p^\alpha - g_p^\alpha) + \frac{1}{\tau^{\alpha\beta}} (\overline{M}_p^\alpha - g_p^\alpha), \right.$$

$$(7b) \quad \left\{ \partial_x E|_{t=t_0} = \frac{1}{\varepsilon_0} \overline{\rho}|_{t=t_0}, \right.$$

$$(7c) \quad \left. \partial_t E = -\frac{1}{\varepsilon_0} j, \right.$$

where the different Maxwellian distributions are defined as:

$$\begin{aligned}
 (8a) \quad & M_0^\alpha(t, x, v_1) = \frac{n^\alpha(t, x)}{\sqrt{2\pi k_B \frac{T^\alpha(t, x)}{m^\alpha}}} \exp\left(-\frac{(v_1 - u^\alpha)^2}{2k_B \frac{T^\alpha(t, x)}{m^\alpha}}\right), \\
 (8b) \quad & M_2^\alpha(t, x, v_1) = k_B T^\alpha(t, x) M_0^\alpha(t, x, v_1), \\
 (8c) \quad & \overline{M}_0^\alpha(t, x, v_1) = \frac{n^\alpha(t, x)}{\sqrt{2\pi k_B \frac{T(t, x)}{m^\alpha}}} \exp\left(-\frac{(v_1 - u)^2}{2k_B \frac{T(t, x)}{m^\alpha}}\right), \\
 (8d) \quad & \overline{M}_2^\alpha(t, x, v_1) = k_B T(t, x) \overline{M}_0^\alpha(t, x, v_1).
 \end{aligned}$$

Moreover, g_0^α and g_2^α satisfy the following relations:

$$\begin{aligned}
 (9a) \quad & n^\alpha = \langle g_0^\alpha \rangle, \\
 (9b) \quad & n^\alpha u^\alpha = \langle g_0^\alpha v_1 \rangle, \\
 (9c) \quad & \frac{3}{2} n^\alpha k_B T^\alpha + \frac{1}{2} m^\alpha n^\alpha (u^\alpha)^2 = \left\langle g_0^\alpha \frac{m^\alpha v_1^2}{2} \right\rangle + \langle g_2^\alpha \rangle.
 \end{aligned}$$

Proof. Firstly, by integrating equation (3a) with respect to v_2 and v_3 , it comes:

$$\begin{aligned}
 (10) \quad & \partial_t \int_{\mathbb{R}^2} f^\alpha dv_2 dv_3 + v_1 \partial_x \int_{\mathbb{R}^2} f^\alpha dv_2 dv_3 = \frac{1}{\tau^\alpha} \left(\int_{\mathbb{R}^2} M(f^\alpha) dv_2 dv_3 - \int_{\mathbb{R}^2} f^\alpha dv_2 dv_3 \right) \\
 & + \frac{1}{\tau^{\alpha\beta}} \left(\int_{\mathbb{R}^2} \overline{M}(f^\alpha) dv_2 dv_3 - \int_{\mathbb{R}^2} f^\alpha dv_2 dv_3 \right).
 \end{aligned}$$

Define $M_0^\alpha, \overline{M}_0^\alpha : \mathbb{R}_t^+ \times \mathbb{R}_x \times \mathbb{R}_v \mapsto \mathbb{R}$ using (8).

The definition of g_0^α given in (5a) allows us to rewrite (10) as (7a) with $p = 0$. We have obtained an equation on the function g_0^α . Let us express the moments using g_0^α . The 0th-order moment can be expressed as:

$$\langle\langle f^\alpha \rangle\rangle = \langle g_0^\alpha \rangle = n^\alpha.$$

Similarly, the first order moment writes:

$$n^\alpha u^\alpha = \langle\langle f^\alpha v \rangle\rangle = \left\langle \int_{\mathbb{R}^2} f^\alpha \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix} dv_2 dv_3 \right\rangle.$$

According to (4a) and (4b), the second and third components of the first order moment are equal to zero. Then, we get, for the first component:

$$\left\langle v_1 \int_{\mathbb{R}^2} f^\alpha dv_2 dv_3 \right\rangle = \langle v_1 g_0^\alpha \rangle = n^\alpha u^\alpha.$$

The third moments reads as:

$$(11) \quad \langle\langle f^\alpha \frac{m^\alpha |v|^2}{2} \rangle\rangle = \left\langle g_0^\alpha \frac{m^\alpha v_1^2}{2} \right\rangle + \langle\langle f^\alpha \frac{m^\alpha (v_2^2 + v_3^2)}{2} \rangle\rangle.$$

Hence, another function g_2^α , defined by (5b), is introduced. Rewrite (11) as:

$$\frac{3}{2} n^\alpha k_B T^\alpha + \frac{1}{2} m^\alpha n^\alpha (u^\alpha)^2 = \left\langle g_0^\alpha \frac{m^\alpha v_1^2}{2} \right\rangle + \langle g_2^\alpha \rangle.$$

Finally, in order to close the system, an equation over g_2^α is needed. To obtain it, multiply (3a) by $\frac{m^\alpha(v_2^2+v_3^2)}{2}$, and integrate with respect to v_2 and v_3 :

$$\begin{aligned} & \int_{\mathbb{R}^2} \partial_t f^\alpha \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3 + \int_{\mathbb{R}^2} v_1 \partial_x f^\alpha \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3 \\ & + \frac{q^\alpha E}{m^\alpha} \int_{\mathbb{R}^2} \partial_{v_1} f^\alpha \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3 \\ & = \frac{1}{\tau^\alpha} \left(\int_{\mathbb{R}^2} M(f^\alpha) \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3 - \int_{\mathbb{R}^2} f \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3 \right). \end{aligned}$$

Then we introduce the function M_2^α of (8b) corresponding to:

$$M_2^\alpha = \int_{\mathbb{R}^2} M(f^\alpha) \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3.$$

We get equation (7a) with $p = 2$. \square

Remark 1. Equation (7b) means that, provided that Maxwell-Gauss equation is satisfied at $t = 0$ and if the Maxwell-Ampère equation is satisfied for all t , so will the Maxwell-Gauss equation.

Regarding the conservation properties of the BGK operators in equations (7a), the following proposition holds:

Proposition 2. The intra-species BGK operators defined by the Maxwellians (8a) and (8b) conserve the partial masses, momenta and energies according to the following identities, for $\alpha = e, i$:

$$(12) \quad \left\langle (M_0^\alpha - g_0^\alpha) \begin{pmatrix} 1 \\ v_1 \end{pmatrix} \right\rangle = 0, \quad \left\langle (M_0^\alpha - g_0^\alpha) \frac{m^\alpha(v_1)^2}{2} \right\rangle + \langle (M_2^\alpha - g_2^\alpha) \rangle = 0.$$

The exchange Maxwellians (8c) and (8d) define inter-species BGK operators that enjoy the following conservation properties:

$$\begin{aligned} \langle \overline{M}_0^\alpha - g_0^\alpha \rangle &= 0, \quad \sum_\alpha \langle (\overline{M}_0^\alpha - g_0^\alpha) m^\alpha v_1 \rangle = 0, \\ \sum_\alpha \left\langle (\overline{M}_0^\alpha - g_0^\alpha) \frac{m^\alpha(v_1)^2}{2} \right\rangle &+ \sum_\alpha \langle \overline{M}_2^\alpha - g_2^\alpha \rangle = 0. \end{aligned}$$

In order to obtain the entropy of the reduced system, consider the following definition of the Maxwellian distribution of the non-reduced system:

$$H^\alpha(M(f^\alpha)) = \min_{f^\alpha \geq 0} (H^\alpha(f^\alpha) \text{ s.t. } g_0^\alpha(v_1) = \int_{\mathbb{R}^2} f^\alpha(v) dv_2 dv_3, g_2^\alpha(v_1) = \int_{\mathbb{R}^2} f^\alpha(v) \frac{m^\alpha(v_2^2+v_3^2)}{2} dv_2 dv_3),$$

with

$$(13) \quad H^\alpha(f^\alpha) = \langle f^\alpha \log f^\alpha \rangle.$$

The resolution of this constrained minimization problem leads to the following expression of $M(f^\alpha)$:

$$(14) \quad M(f^\alpha)(g_0^\alpha, g_2^\alpha) = \frac{(g_0^\alpha)^2 m^\alpha}{2\pi g_2^\alpha} \exp\left(-\frac{g_0^\alpha m^\alpha (v_2^2+v_3^2)}{2g_2^\alpha}\right).$$

Hence, by injecting (14) in (13), it comes

$$H^\alpha(g_0^\alpha, g_2^\alpha) = \left\langle g_0^\alpha \log \frac{(g_0^\alpha)^2}{g_2^\alpha} - 2g_0^\alpha \right\rangle.$$

Hence, define the following notations for the partial entropies of the reduced model:

$$(15) \quad H^\alpha(g_0^\alpha, g_2^\alpha) = \langle \eta^\alpha(g_0^\alpha, g_2^\alpha) \rangle,$$

with

$$\eta^\alpha(g_0^\alpha, g_2^\alpha) = g_0^\alpha \log \frac{(g_0^\alpha)^2}{g_2^\alpha} - 2g_0^\alpha,$$

for $\alpha = e, i$, and the total entropy

$$(16) \quad H(g_0^e, g_2^e, g_0^i, g_2^i) = \sum_{\alpha} H^\alpha(g_0^\alpha, g_2^\alpha).$$

Firstly, this entropy is shown to be convex:

Proposition 3. *For $\alpha = e, i$, H^α is a strictly convex function. Consequently, H is strictly convex as well.*

Proof. Compute the gradient of η^α :

$$(17) \quad \nabla \eta^\alpha = \begin{pmatrix} \partial_{g_0^\alpha} \eta^\alpha \\ \partial_{g_2^\alpha} \eta^\alpha \end{pmatrix} = \begin{pmatrix} \log \frac{(g_0^\alpha)^2}{g_2^\alpha} \\ -\frac{g_0^\alpha}{g_2^\alpha} \end{pmatrix}.$$

Compute the Hessian:

$$(18) \quad \nabla^2 \eta^\alpha = \begin{pmatrix} \frac{2}{g_0^\alpha} & -\frac{1}{g_2^\alpha} \\ -\frac{1}{g_0^\alpha} & \frac{g_0^\alpha}{(g_2^\alpha)^2} \end{pmatrix}.$$

By positivity of the distribution functions, $\text{Tr}(\nabla^2 \eta^\alpha) > 0$ and $\det(\nabla^2 \eta^\alpha) = \frac{1}{(g_2^\alpha)^2} > 0$. Hence, the Hessian is positive definite and η^α is strictly convex. Consequently, H^α is strictly convex. \square

The following property on entropy dissipation holds:

Proposition 4. *The reduced system (7) satisfies the H-theorem for the entropy defined by (16):*

$$(19) \quad \frac{\partial H}{\partial t} + \frac{\partial G}{\partial x} \leq 0,$$

with G the entropy flux associated to H :

$$G(g_0^e, g_2^e, g_0^i, g_2^i) = \sum_{\alpha} G^\alpha(g_0^\alpha, g_2^\alpha) = \sum_{\alpha} \langle v_1 \eta^\alpha(g_0^\alpha, g_2^\alpha) \rangle.$$

Proof. Multiply equation (7a) by $\partial_{g_p^\alpha} \eta^\alpha(g_0^\alpha, g_2^\alpha)$, sum over p and integrate over v_1 :

$$\begin{aligned} \frac{\partial H^\alpha}{\partial t} + \frac{\partial G^\alpha}{\partial x} &= \frac{1}{\tau^\alpha} \langle (M_0^\alpha - g_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (M_2^\alpha - g_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \rangle \\ &+ \frac{1}{\tau_{\alpha\beta}} \langle (\bar{M}_0^\alpha - g_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (\bar{M}_2^\alpha - g_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \rangle. \end{aligned}$$

Then, summing over α gives:

$$(20) \quad \begin{aligned} \frac{\partial H}{\partial t} + \frac{\partial G}{\partial x} &= \sum_{\alpha} \frac{1}{\tau^{\alpha}} \langle (M_0^{\alpha} - g_0^{\alpha}) \partial_{g_0^{\alpha}} \eta^{\alpha} + (M_2^{\alpha} - g_2^{\alpha}) \partial_{g_2^{\alpha}} \eta^{\alpha} \rangle \\ &+ \sum_{\alpha} \frac{1}{\tau^{\alpha\beta}} \langle (\overline{M}_0^{\alpha} - g_0^{\alpha}) \partial_{g_0^{\alpha}} \eta^{\alpha} + (\overline{M}_2^{\alpha} - g_2^{\alpha}) \partial_{g_2^{\alpha}} \eta^{\alpha} \rangle. \end{aligned}$$

Consider every term of the right-hand side consecutively.

$$(21) \quad \begin{aligned} \langle (M_0^{\alpha} - g_0^{\alpha}) \partial_{g_0^{\alpha}} \eta^{\alpha} + (M_2^{\alpha} - g_2^{\alpha}) \partial_{g_2^{\alpha}} \eta^{\alpha} \rangle &= \langle (M_0^{\alpha} - g_0^{\alpha}) (\partial_{g_0^{\alpha}} \eta^{\alpha}(g_0^{\alpha}, g_2^{\alpha}) - \partial_{g_0^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha})) \rangle \\ &+ \langle (M_2^{\alpha} - g_2^{\alpha}) (\partial_{g_2^{\alpha}} \eta^{\alpha}(g_0^{\alpha}, g_2^{\alpha}) - \partial_{g_2^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha})) \rangle \\ &+ \langle (M_0^{\alpha} - g_0^{\alpha}) \partial_{g_0^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha}) \rangle \\ &+ \langle (M_2^{\alpha} - g_2^{\alpha}) \partial_{g_2^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha}) \rangle. \end{aligned}$$

By convexity of η^{α} , the sum of the first two terms of the right-hand side of (21) is negative. The remaining terms are

$$\begin{aligned} &\langle (M_0^{\alpha} - g_0^{\alpha}) \partial_{g_0^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha}) \rangle + \langle (M_2^{\alpha} - g_2^{\alpha}) \partial_{g_2^{\alpha}} \eta^{\alpha}(M_0^{\alpha}, M_2^{\alpha}) \rangle \\ &= \left\langle (M_0^{\alpha} - g_0^{\alpha}) \log \left(\frac{(M_0^{\alpha})^2}{M_2^{\alpha}} \right) \right\rangle - \left\langle (M_2^{\alpha} - g_2^{\alpha}) \frac{M_0^{\alpha}}{M_2^{\alpha}} \right\rangle \\ &= \frac{1}{k_B T^{\alpha}} \langle (M_0^{\alpha} - g_0^{\alpha}) \log(M_0^{\alpha}) \rangle - \frac{1}{k_B T^{\alpha}} \langle (M_2^{\alpha} - g_2^{\alpha}) \rangle \\ &= \frac{1}{k_B T^{\alpha}} \log \left(\frac{n^{\alpha}}{(2\pi k_B \frac{T^{\alpha}}{m^{\alpha}})^{\frac{1}{2}}} \right) \langle (M_0^{\alpha} - g_0^{\alpha}) \rangle + \left\langle (M_0^{\alpha} - g_0^{\alpha}) \left(\frac{v u^{\alpha}}{k_B \frac{T^{\alpha}}{m^{\alpha}}} - \frac{(u^{\alpha})^2}{2k_B \frac{T^{\alpha}}{m^{\alpha}}} \right) \right\rangle \\ &- \frac{1}{k_B T^{\alpha}} \left\langle (M_0^{\alpha} - g_0^{\alpha}) \frac{m^{\alpha} (v^{\alpha})^2}{2} + (M_2^{\alpha} - g_2^{\alpha}) \right\rangle \\ &= 0, \end{aligned}$$

according to (12). By an analogous computation on the remaining last term in (20), (19) is obtained. \square

2.5. Scaling of the equations. The regime of interest is the hydrodynamic limit of this model, which corresponds to the scale of the phenomena described by the bi-temperature Euler equations (see [3]). To do so, it is necessary to perform a scaling on this model, in order to exhibit the behaviour of the equations when the macroscopic characteristic scale is much larger than the ones involved by the kinetic equations (collisional scale, Debye scale). Denote $t_0, L, v_{\text{th}}, q, m$ respectively the reference time, length, velocity, electric charge and particle mass. Denoting with a tilde the dimensionless quantities, it comes:

$$\tilde{t} = \frac{t}{t_0}, \quad \tilde{x} = \frac{x}{L}, \quad \tilde{v}_1 = \frac{v_1}{v_{\text{th}}}, \quad \tilde{q}^{\alpha} = \frac{q^{\alpha}}{q}, \quad \tilde{m}^{\alpha} = \frac{m^{\alpha}}{m}.$$

From this scaling, define also the following dimensionless quantities:

$$\tilde{g}_p^{\alpha} = \frac{v_{\text{th}}}{n_0} g_p^{\alpha}, \quad \tilde{E} = \frac{q t_0}{v_{\text{th}} m} E, \quad \tilde{j} = \frac{j}{q v_{\text{th}} n_0}, \quad \tilde{\rho} = \frac{\bar{\rho}}{q n_0}, \quad \tilde{T} = \frac{k_B}{m v_{\text{th}}^2} T.$$

where n_0 is the reference density.

Applying this scaling to our model, and removing the tildas for the sake of readability, it comes, for $p = 0, 2$ and $\alpha = e, i$:

$$(22a) \quad \begin{cases} \partial_t g_p^\alpha + v_1 \partial_x g_p^\alpha + \frac{q^\alpha}{m^\alpha} E \partial_{v_1} g_p^\alpha = \frac{1}{\varepsilon^\alpha} (M_p^e - g_p^e) + \frac{1}{\varepsilon^{ei}} (\overline{M}_p^e - g_p^e), \\ \partial_t E = -\frac{j}{\beta^2}, \\ \partial_x E = \frac{\overline{\rho}}{\beta^2}, \end{cases}$$

where $\varepsilon^\alpha = \frac{\tau^\alpha}{t_0}$, $\varepsilon^{\alpha\beta} = \frac{\tau^{\alpha\beta}}{t_0}$ and $\beta = \frac{\lambda_{De}}{L}$, $\lambda_{De} = (\frac{\varepsilon_0 k_B T_0}{q^2 n_0})^{\frac{1}{2}}$, with $T_0 = \frac{m v_{th}^2}{k_B}$ the reference temperature. λ_{De} is the so-called Debye length, which is the characteristic length of charge separation in a plasma. In the regime of interest, this quantity is extremely small compared to the reference length of the domain L . Similarly, the same-species mean free path related quantities ε^α are small compared to the reference length L . Hence, reaching the hydrodynamic limit consists in taking the limit of the model when ε^α and β tend to 0. Notice that, however, $\varepsilon^{\alpha\beta}$ is finite. In the limit, the Maxwell equations (22b-22c) become $j = 0$ and $\rho = 0$, which can be rewritten as:

$$(23) \quad u^e = u^i, \quad n^e = Zn^i.$$

The equations (23) are called the quasi-neutrality constraints. This form of the system of equations is the singular quasi-neutral limit of our initial model, where two evolution equations degenerate into algebraic relations. The purpose of this work is to derive an Asymptotic-preserving (AP) numerical scheme, that is to say with a stability condition that is independant of every non-dimensional small parameters ε^α , $\varepsilon^{\alpha\beta}$ and β .

2.6. Entropy inequality for hydrodynamic limits. In [3], the existence of an entropy-flux pair is shown for the bi-temperature Euler system (1). In this section, an entropy dissipation inequality in the limit is derived for the reduced model presented in Proposition 1.

Firstly, denote $\mathcal{U} = (\rho, \rho u, \varepsilon^e, \varepsilon^i)$ the unknowns of the bi-temperature model and $(H^{\text{macro}}(\mathcal{U}), G^{\text{macro}}(\mathcal{U}))$ the corresponding macroscopic entropy-flux pair. Note the following properties regarding the link between kinetic and macroscopic entropies:

$$(24) \quad H(M_0^e, M_2^e, M_0^i, M_2^i) = H^{\text{macro}}(\mathcal{U}), \quad G(M_0^e, M_2^e, M_0^i, M_2^i) = G^{\text{macro}}(\mathcal{U}).$$

See [3] for further details on these properties.

Proposition 5. *Let \mathcal{U} be a limit solution of the kinetic model from Proposition 1, i.e moments of solution distribution functions (g_0^α, g_2^α) in the hydrodynamic limit. Then, the following inequality on entropy holds:*

$$(25) \quad \frac{\partial H^{\text{macro}}}{\partial t} + \frac{\partial G^{\text{macro}}}{\partial x} \leq -\frac{\nu^{ei}}{k_B T^i T^e} (T^i - T^e)^2,$$

with

$$\nu^{\alpha\beta} = \frac{3k_B n^\alpha n^\beta}{2\tau^{\alpha\beta}(n^\alpha + n^\beta)}.$$

Proof. Multiply equation (7a) by $\partial_{g_p^\alpha} \eta^\alpha(g_0^\alpha, g_2^\alpha)$, sum over p and integrate over v_1 . It comes, for $\alpha = e, i$:

$$\begin{aligned} \frac{\partial H^\alpha}{\partial t} + \frac{\partial G^\alpha}{\partial x} &= \frac{1}{\varepsilon^\alpha} \langle (M_0^\alpha - g_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (M_2^\alpha - g_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \rangle \\ &\quad + \frac{1}{\varepsilon^{\alpha\beta}} \langle (\overline{M}_0^\alpha - g_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (\overline{M}_2^\alpha - g_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \rangle. \end{aligned}$$

Using the entropy dissipation properties of the BGK operator shown in Section 2.4.2:

$$\frac{\partial H^\alpha}{\partial t} + \frac{\partial G^\alpha}{\partial x} \leq \frac{1}{\varepsilon^{\alpha\beta}} \langle (\overline{M}_0^\alpha - g_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (\overline{M}_2^\alpha - g_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \rangle.$$

By formally passing to limit $\epsilon^\alpha \rightarrow 0$, it comes $g_p^\alpha = Mg_p^\alpha$, $n^e = n^i$ and $u = u^e = u^i$. Hence, summing over α and using (24), the inequality becomes

$$\frac{\partial H^{\text{macro}}(\mathcal{U})}{\partial t} + \frac{\partial G^{\text{macro}}(\mathcal{U})}{\partial x} \leq \frac{1}{\epsilon^{\alpha\beta}} \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha + (\bar{M}_2^\alpha - M_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha \right\rangle.$$

The right hand side reads

$$\begin{aligned} & \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha(M_0^\alpha, M_2^\alpha) + (\bar{M}_2^\alpha - M_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha(M_0^\alpha, M_2^\alpha) \right\rangle \\ &= \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \log\left(\frac{(M_0^\alpha)^2}{M_2^\alpha}\right) - (\bar{M}_2^\alpha - M_2^\alpha) \frac{M_0^\alpha}{M_2^\alpha} \right\rangle \\ &= \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \log\left(\frac{M_0^\alpha}{k_B T^\alpha}\right) - \frac{1}{k_B T^\alpha} (\bar{M}_2^\alpha - M_2^\alpha) \right\rangle \\ &= \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \log\left(\frac{n^\alpha}{(2\pi k_B \frac{T^\alpha}{m^\alpha})^{\frac{1}{2}}}\right) - (\bar{M}_0^\alpha - M_0^\alpha) \frac{(v-u)^2}{2k_B \frac{T^\alpha}{m^\alpha}} - (\bar{M}_0^\alpha - M_0^\alpha) \log(k_B T^\alpha) \right\rangle \\ &\quad - \sum_{\alpha} \left\langle \frac{1}{k_B T^\alpha} (\bar{M}_2^\alpha - M_2^\alpha) \right\rangle. \end{aligned}$$

The first and third term are equal to zero according to the conservation properties. The second term is equal, after computation, to :

$$- \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \frac{(v-u)^2}{2k_B \frac{T^\alpha}{m^\alpha}} \right\rangle = - \sum_{\alpha} \frac{1}{k_B T^\alpha} \left\langle m^\alpha (\bar{M}_0^\alpha - M_0^\alpha) \frac{v^2}{2} \right\rangle.$$

The total right hand side is then equal to

$$\begin{aligned} & \sum_{\alpha} \left\langle (\bar{M}_0^\alpha - M_0^\alpha) \partial_{g_0^\alpha} \eta^\alpha(M_0^\alpha, M_2^\alpha) + (\bar{M}_2^\alpha - M_2^\alpha) \partial_{g_2^\alpha} \eta^\alpha(M_0^\alpha, M_2^\alpha) \right\rangle \\ &= - \sum_{\alpha} \frac{1}{k_B T^\alpha} \left(\left\langle m^\alpha (\bar{M}_0^\alpha - M_0^\alpha) \frac{v^2}{2} \right\rangle + \left\langle \bar{M}_2^\alpha - M_2^\alpha \right\rangle \right) \\ &= - \frac{3}{2} \sum_{\alpha} n^\alpha \frac{\bar{T} - T^\alpha}{T^\alpha}. \end{aligned}$$

According to the definition of T , it comes:

$$\bar{T} - T^\alpha = \frac{n^\alpha}{n^\alpha + n^\beta} (T^\beta - T^\alpha).$$

Hence,

$$- \frac{3}{2\tau^{\alpha\beta}} \sum_{\alpha} n^\alpha \frac{\bar{T} - T^\alpha}{T^\alpha} = - \frac{3}{2\tau^{\alpha\beta}} \frac{n^\alpha n^\beta}{n^\alpha + n^\beta} \frac{(T^\beta - T^\alpha)^2}{T^\alpha T^\beta},$$

and we get (25). □

3. NUMERICAL TREATMENT

In this section, numerical treatment of system (22) is addressed. The method to compute the evolution of the particle distribution is presented. For the sake of simplicity, in the Ampère equation and BGK operator, every non-dimensional quantity ϵ^e, ϵ^i and β are chosen equal to τ . Reaching the hydrodynamic limit is then equivalent to the limit $\tau \rightarrow 0$.

3.1. Vlasov-BGK equations: general method. A time-splitting method is used on the Vlasov-BGK equations, in order to separate the Vlasov equations from the BGK operators. More precisely, the following equations are going to be consecutively solved:

$$(26) \quad \partial_t g_p^\alpha + v_1 \partial_x g_p^\alpha + \frac{q^\alpha}{m^\alpha} E \partial_{v_1} g_p^\alpha = 0,$$

the Vlasov-Maxwell equations for $\alpha = e, i$, coupled with the Ampère equation:

$$(27) \quad \partial_t E = -\frac{j}{\tau^2}.$$

Then, the effects of the inter-species collisions are computed:

$$(28) \quad \partial_t g_p^\alpha = \frac{1}{\varepsilon_{\alpha\beta}} (\overline{M}_p^\alpha - g_p^\alpha),$$

and finally, intra-species collisions:

$$(29) \quad \partial_t g_p^\alpha = \frac{1}{\tau} (M_p^\alpha - g_p^\alpha),$$

where $\tau \rightarrow 0$.

A uniform discretization of the phase space is considered. Concerning the velocity discretization, in the present case, only static grids are considered, contrarily to [8]. A suitable subset of \mathbb{R} that contains the most part of the distribution function has to be chosen in order to perform correct numerical evaluations of the moments.

However, in this model, two particle distributions are considered, in order to follow the dynamics of both electrons and ions via a kinetic description. The major difference between these two species is the value of the mass of a particle that can have a ratio of the order of one thousand. Consequently, the distribution functions have very different supports. Thus, if only one velocity grid is employed, then this grid has to be large enough to encompass the largest distribution and has to be fine enough to describe the smallest one, which is numerically very costly. Hence, two velocity grids, one for each species, are considered. The boundaries of the domains are chosen as follows: for a given initial condition, let us denote the initial moments by n_0^α , u_0^α and T_0^α . For $\alpha = e, i$, we have:

$$(30) \quad V^\alpha = \left[\min_x (u_0^\alpha(x)) - l^\alpha \max_x \sqrt{k_B \frac{T_0^\alpha(x)}{m^\alpha}}, \max_x (u_0^\alpha(x)) + l^\alpha \max_x \sqrt{k_B \frac{T_0^\alpha(x)}{m^\alpha}} \right],$$

where l^α is chosen to fit the initial maxwellian distribution. Note that this grid is the same for all points in space and is constant in time.

Notations for the discretization are chosen as follows: Δx denotes the discretization step of the domain $[L_{\min}, L_{\max}] \subset \mathbb{R}$ in N_x cells. For $\alpha = e, i$, the velocity domains $[V_{\min}^\alpha, V_{\max}^\alpha] \subset \mathbb{R}$ are both discretized using N_v cells and the discretization steps are denoted Δv^e and Δv^i . Then, we use indices k and j to denote quantities computed at the k -th cell in space and j -th cell in velocity. Moreover, we denote Δt the time step and we use a superscript n to denote the quantity computed at time $t^n = t^0 + n\Delta t$.

Step 1: Computation of g and E

The following discretization is performed, for $\alpha = e, i$:

$$(31) \quad \tilde{g}_{0,k,j}^{\alpha,n+1} = g_{0,k,j}^{\alpha,n} - \frac{\Delta t}{\Delta x} (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n}) - \frac{\Delta t}{2\Delta v^\alpha} \frac{q^\alpha E_k^{n+1}}{m^\alpha} (g_{0,k,j+1}^{\alpha,n} - g_{0,k,j-1}^{\alpha,n}),$$

coupled with

$$(32) \quad E_k^{n+1} = E_k^n - \frac{\Delta t}{\tau} j_k^{n+1}.$$

Note that the term $\partial_{v_1} g_0$ has been discretized using a centred scheme. For the space discretization, ϕ_0 represented the flux function of the spatial discretization which will be specified in the next section.

Then, the following procedure is applied to obtain $\tilde{g}_2^{\alpha, n+1}$:

$$(33) \quad \tilde{g}_{2,k,j}^{\alpha, n+1} = g_{2,k,j}^{\alpha, n} - \frac{\Delta t}{\Delta x} (\phi_{2,k+\frac{1}{2},j}^{\alpha, n} - \phi_{2,k-\frac{1}{2},j}^{\alpha, n}) - \frac{\Delta t}{\Delta v^\alpha} (\psi_{2,k,j+\frac{1}{2}}^{\alpha, n} - \psi_{2,k,j-\frac{1}{2}}^{\alpha, n}),$$

where

$$\phi_{2,k+\frac{1}{2},j}^{\alpha, n} = \frac{v_j^\alpha}{2} (g_{2,k+1,j}^{\alpha, n} + g_{2,k,j}^{\alpha, n}) - \frac{|v_j^\alpha|}{2} (g_{2,k+1,j}^{\alpha, n} - g_{2,k,j}^{\alpha, n}),$$

and

$$\psi_{2,k,j+\frac{1}{2}}^{\alpha, n} = \frac{q^\alpha E_k^{n+1}}{2m^\alpha} (g_{2,k,j+1}^{\alpha, n} + g_{2,k,j}^{\alpha, n}) - \frac{|q^\alpha E_k^{n+1}|}{2m^\alpha} (g_{2,k,j+1}^{\alpha, n} - g_{2,k,j}^{\alpha, n}).$$

Equations (31) are stable under the CFL condition :

$$(34) \quad \Delta t \leq \min \left(\frac{1}{\frac{\|V^\alpha\|_\infty}{\Delta x} + \frac{\|q^\alpha E^{n+1}\|_\infty}{m^\alpha \Delta v^\alpha}} \right).$$

Moments of $\tilde{g}_p^{e, n+1}$ and $\tilde{g}_p^{i, n+1}$ are then computed using the midpoint quadrature formula. $\forall k \in \{1, \dots, N_x\}$:

$$\begin{aligned} \tilde{n}_k^{\alpha, n+1} &= \sum_{j=1}^{N_v} \tilde{g}_{0,k,j}^{\alpha, n+1} \Delta v^\alpha, \\ \tilde{u}_k^{\alpha, n+1} &= \frac{1}{\tilde{n}_k^{\alpha, n+1}} \sum_{j=1}^{N_v} \tilde{g}_{0,k,j}^{\alpha, n+1} v_j \Delta v^\alpha, \\ \tilde{T}_k^{\alpha, n+1} &= \frac{1}{3\tilde{n}_k^{\alpha, n+1} k_B} \left[\sum_{j=1}^{N_v} m^\alpha \tilde{g}_{0,k,j}^{\alpha, n+1} v_j^2 \Delta v^\alpha + 2 \sum_{j=1}^{N_v} \tilde{g}_{2,k,j}^{\alpha, n+1} \Delta v^\alpha - m^\alpha \tilde{n}_k^{\alpha, n+1} \tilde{u}_k^{\alpha, n+1} \right]. \end{aligned}$$

In order to compute the mixture Maxwellian distribution in the next step, mixture quantities must also be computed:

$$\begin{aligned} \tilde{n}_k^{n+1} &= \tilde{n}_k^{e, n+1} + \tilde{n}_k^{i, n+1}, \\ \tilde{u}_k^{n+1} &= \frac{\sum_\alpha \tilde{\rho}_k^{\alpha, n+1} \tilde{u}_k^{\alpha, n+1}}{\sum_\alpha \tilde{\rho}_k^{\alpha, n+1}}, \\ \tilde{T}_k^{n+1} &= \frac{1}{\tilde{n}_k^{n+1} k_B} \left[\sum_\alpha \frac{1}{3} \tilde{\rho}_k^{\alpha, n+1} ((\tilde{u}_k^{\alpha, n+1})^2 - (\tilde{u}_k^{n+1})^2) + \sum_\alpha \tilde{n}_k^{\alpha, n+1} k_B \tilde{T}_k^{\alpha, n+1} \right]. \end{aligned}$$

Step 2: inter-species relaxation

We apply a backward Euler method to the equations (28):

$$\hat{g}_{p,k,j}^{\alpha, n+1} = \frac{1}{\Delta t + \tau^{ei}} \left[\Delta t \widetilde{M}_{p,k,j}^{\alpha, n+1} + \tau^{ei} \tilde{g}_{p,k,j}^{\alpha, n+1} \right]$$

which is unconditionally stable. $\widetilde{M}_{p,k,j}^{\alpha, n+1}$ is the mixture Maxwellian defined by the moments $\tilde{n}_k^{\alpha, n+1}$, \tilde{u}_k^{n+1} and \tilde{T}_k^{n+1} , which are conserved through this step. Hence, the mixture Maxwellian is known beforehand, which allows this implicitation. Partial velocities and temperatures (denoted in the next section with a hat) are then recomputed using the previous formulae.

Step 3: intra-species relaxation

The next step consists in the relaxation towards Maxwellian equilibrium. Since the BGK operator preserves moments, we have the following properties: $\forall k \in \{1, \dots, N_x\}$,

$$(35) \quad \hat{n}_k^{\alpha, n+1} = n_k^{\alpha, n+1}, \quad \hat{u}_k^{\alpha, n+1} = u_k^{\alpha, n+1}, \quad \hat{T}_k^{\alpha, n+1} = T_k^{\alpha, n+1}.$$

Similarly to the previous section, this property allows us to use an implicit scheme, the backward Euler method: $\forall k \in \{1, \dots, N_x\}$, $\forall j \in \{1, \dots, N_v\}$ and $p = 0, 2$,

$$(36) \quad g_{p,k,j}^{\alpha, n+1} = \hat{g}_{p,k,j}^{\alpha, n+1} + \frac{\Delta t}{\tau} (M_{p,k,j}^{\alpha, n+1} - g_{p,k,j}^{\alpha, n+1}),$$

where

$$M_{0,k,j}^{\alpha, n+1} = \frac{n_k^{\alpha, n+1}}{\sqrt{2\pi k_B \frac{T_k^{\alpha, n+1}}{m^\alpha}}} \exp\left(-\frac{(v_j - u_k^{\alpha, n+1})^2}{2k_B \frac{T_k^{\alpha, n+1}}{m^\alpha}}\right),$$

$$M_{2,k,j}^{\alpha, n+1} = k_B T_k^{\alpha, n+1} M_{0,k,j}^{\alpha, n+1},$$

which are expressed using quantities known at the end of step 2, according to the properties (35). Taking the limit $\tau \rightarrow 0$ in (36) leads to:

$$g_{p,k,j}^{\alpha, n+1} = M_{p,k,j}^{\alpha, n+1}.$$

In the next section, enforcement of quasineutrality constraints (23) will be addressed. More precisely, specific spatial fluxes will be given and a reformulation of the Maxwell-Ampère equation will be performed in order to derive an Asymptotic-Preserving method, valid in the limit $\tau \rightarrow 0$. Note that if these constraints are enforced through the hyperbolic step of the scheme (step 1), then they will be trivially preserved by both step 2 and 3. Hence, quasineutrality only has to be enforced through step 1.

3.2. Modified-viscosity upwind scheme. The space divergence is discretized explicitly, via the following modified upwind scheme:

$$(37) \quad \phi_{0,k+\frac{1}{2},j}^{\alpha, n} = \frac{v_j^\alpha}{2} (g_{0,k+1,j}^{\alpha, n} + g_{0,k,j}^{\alpha, n}) - \frac{|V_{\max}|}{2} (g_{0,k+1,j}^{\alpha, n} - g_{0,k,j}^{\alpha, n}),$$

where $V_{\max} = \max(|V_{\min}^e|, |V_{\min}^i|, |V_{\max}^e|, |V_{\min}^e|)$.

This method is the classical upwind scheme, where the numerical viscosity $|v_j|$, that depends directly of the microscopic velocity, has been replaced by the smallest velocity that ensures stability for all equations. The discrete system writes as follows:

$$(38) \quad \tilde{g}_{0,k,j}^{\alpha, n+1} = g_{0,k,j}^{\alpha, n} - \frac{\Delta t}{\Delta x} (\phi_{0,k+\frac{1}{2},j}^{\alpha, n} - \phi_{0,k-\frac{1}{2},j}^{\alpha, n}) - \frac{\Delta t}{2\Delta v^\alpha} \frac{q^\alpha E_k^{n+1}}{m^\alpha} (g_{0,k,j+1}^{\alpha, n} - g_{0,k,j-1}^{\alpha, n}),$$

and

$$(39) \quad E_k^{n+1} = E_k^n - \frac{\Delta t}{\tau} j_k^{n+1}.$$

This level of implicitness allows for a reformulation of the equation on the electric field. More precisely, the equation used to compute the electric field in (39) is not valid in the limit $\tau \rightarrow 0$. To solve this problem, we follow the idea developed in [20]. By multiplying equation (38) by

$q^\alpha v_j^\alpha \Delta v^\alpha$ and summing over j , it comes:

$$(40) \quad \begin{aligned} q^\alpha \sum_j v_j^\alpha \widetilde{g}_{0,k,j}^{\alpha,n+1} \Delta v^\alpha &= q^\alpha \sum_j v_j^\alpha g_{0,k,j}^{\alpha,n} \Delta v^\alpha - q^\alpha \frac{\Delta t \Delta v^\alpha}{\Delta x} \sum_j v_j^\alpha (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n}) \\ &\quad - \frac{\Delta t}{2} \sum_j v_j^\alpha \left(\frac{(q^\alpha)^2 E_k^{n+1}}{m^\alpha} (g_{0,k,j+1}^{\alpha,n} - g_{0,k,j-1}^{\alpha,n}) \right). \end{aligned}$$

According to the definition of the current $j_k^n = \sum_\alpha q^\alpha \sum_j v_j^\alpha g_{0,k,j}^{\alpha,n} \Delta v^\alpha$, taking (40) into account with $\alpha = e, i$, it comes

$$\begin{aligned} \widetilde{j}_k^{n+1} &= j_k^n - \frac{\Delta t}{\Delta x} \sum_\alpha q^\alpha \sum_j \Delta v^\alpha v_j^\alpha (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n}) \\ &\quad - \frac{\Delta t E_k^{n+1}}{2} \sum_\alpha \frac{(q^\alpha)^2}{m^\alpha} \sum_j v_j^\alpha (g_{0,k,j+1}^{\alpha,n} - g_{0,k,j-1}^{\alpha,n}). \end{aligned}$$

By injecting this equation in (39), we get the following expression for the electric field:

$$E_k^{n+1} = \frac{\tau E_k^n - \Delta t j_k^n + \frac{\Delta t^2}{\Delta x} \sum_\alpha q^\alpha \sum_j \Delta v^\alpha v_j^\alpha (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n})}{\tau - \frac{\Delta t^2}{2} \sum_\alpha \frac{(q^\alpha)^2}{m^\alpha} \sum_j v_j^\alpha (g_{0,k,j+1}^{\alpha,n} - g_{0,k,j-1}^{\alpha,n})},$$

which is valid for all values of τ . In particular, we are interested in the limit $\tau \rightarrow 0$, which is:

$$E_k^{n+1} = \frac{\frac{1}{\Delta x} \sum_\alpha q^\alpha \sum_j \Delta v^\alpha v_j^\alpha (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n})}{\sum_\alpha \frac{(q^\alpha)^2}{m^\alpha} n_k^{\alpha,n}},$$

where it has been assumed that $j_k^n = 0$. Such a choice for E_k^{n+1} ensures that $\widetilde{j}_k^{n+1} = 0$, which is one of the two quasi-neutrality constraints (23). The choice of ϕ_0 (37) that has been made ensures that the other constraint $\widetilde{\rho}_k^{n+1} = 0$ is also verified. More precisely, compute the discrete equation on $\widetilde{\rho}_k^{n+1}$. $\forall k \in \{1, \dots, N_x\}$, provided that

$$\sum_j (\psi_{0,k,j+\frac{1}{2}}^{\alpha,n} - \psi_{0,k,j-\frac{1}{2}}^{\alpha,n}) = 0,$$

the electric charge is

$$\sum_j q^\alpha g_{0,k,j}^{\alpha,n+1} \Delta v^\alpha = \sum_j q^\alpha g_{0,k,j}^{\alpha,n} \Delta v^\alpha - \frac{\Delta t}{\Delta x} \sum_j (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n}) \Delta v^\alpha.$$

So, the following equation on the total electric charge is obtained:

$$\bar{\rho}_k^{n+1} = \bar{\rho}_k^n - \frac{\Delta t}{\Delta x} \sum_\alpha q^\alpha \sum_j (\phi_{0,k+\frac{1}{2},j}^{\alpha,n} - \phi_{0,k-\frac{1}{2},j}^{\alpha,n}) \Delta v^\alpha.$$

According to (37), we get:

$$(41) \quad \bar{\rho}_k^{n+1} = \bar{\rho}_k^n - \frac{\Delta t}{2\Delta x} \sum_\alpha q^\alpha \sum_j (v_j^\alpha (g_{0,k+1,j}^{\alpha,n} - g_{0,k-1,j}^{\alpha,n}) - |V_{\max}| (g_{0,k+1,j}^{\alpha,n} - 2g_{0,k,j}^{\alpha,n} + g_{0,k-1,j}^{\alpha,n})) \Delta v^\alpha.$$

By rearranging the terms, we obtain:

$$(42) \quad \bar{\rho}_k^{n+1} = \bar{\rho}_k^n - \frac{\Delta t}{2\Delta x} (j_{k+1}^n - j_{k-1}^n) + \frac{\Delta t}{2\Delta x} \sum_\alpha q^\alpha \sum_j |V_{\max}| (g_{0,k+1,j}^{\alpha,n} - 2g_{0,k,j}^{\alpha,n} + g_{0,k-1,j}^{\alpha,n}) \Delta v^\alpha$$

By assuming that $\bar{\rho}_k^n = j_{k+1}^n = j_{k-1}^n = 0$, it comes:

$$(43) \quad \bar{\rho}_k^{n+1} = \frac{\Delta t}{2\Delta x} V_{\max} (\bar{\rho}_{k+1}^n - 2\bar{\rho}_k^n + \bar{\rho}_{k-1}^n) = 0.$$

Hence, quasineutrality is conserved at all time. However, applying this modification to our scheme amounts to overevaluating the numerical viscosity by its highest value. This will create an important amount of dissipation in the numerical solutions obtained with this scheme. Hence, the discretization of the space domain needs to be very precise to compensate this viscosity, which will result in a longer computation time.

4. NUMERICAL RESULTS

In this section, numerical resolution of Riemann problems with our method is presented. Such problems are characterized by initial conditions constituted of two constant states separated by a discontinuity. In the study of such problems, an exact solution can be provided for the conserved quantities (density, momentum and total energy). However, it is not true for non-conserved quantities, and no reference results exist.

Comparative results obtained by an HLL-type scheme and a Suliciu relaxation method on the bi-temperature Euler system are displayed along the solutions from the kinetic scheme that has been derived in this paper (see [3] for more details on these methods).

The spatial domain is chosen as the interval $[0,1]$, supplemented with homogeneous Neumann boundary conditions. Physical constants (k_B, Z, q^i) are chosen equal to one. All test cases are done in the hydrodynamic limit $\tau \rightarrow 0$. Concerning the choice of l^α in (30), a convergence test has been performed on the initial conditions of the test cases in order to obtain a appropriate value that would optimize computational cost. Hence, in all test cases, $l^\alpha = 8$, for $\alpha \in \{e, i\}$.

A mass ratio of 10 ($m_e = 0.1$ and $m_i = 1$) is considered. The greater the mass ratio, the greater the value of V_{\max} in (37). Hence, a high mass ratio implies an important amount of numerical viscosity. Then, we are restricting ourselves to a mass ratio of 10, which is already demanding in terms of number of discretization points. However, in a proper physical context, mass ratios of the order of 1000 should be considered.

Shock tube with identical initial temperatures. The Riemann problem of the shock tube with identical ionic and electronic temperature is given by the following initial condition:

$$\begin{cases} n^\alpha(t_0, x) = 1, & u^\alpha(t_0, x) = 0, & T^\alpha(t_0, x) = 1 & \text{if } x \in [0, 0.5], \\ n^\alpha(t_0, x) = 0.125, & u^\alpha(t_0, x) = 0, & T^\alpha(t_0, x) = 1 & \text{if } x \in [0.5, 1]. \end{cases}$$

Parameters are chosen as $N_x = 120000$, $N_v = 40$ and $l = 8$. The inter-species collision relaxation time is $\tau^{ei} = 0.1$. Results are computed at time $t = 0.1$. For fluid methods, 4000 discretization points are used.

The results are provided in figure 1 and 2.

The quantities described by conservative equations (densities and velocities) are consistent with the fluid methods HLL and Suliciu. Moreover, quasi-neutrality is shown to be achieved, since electronic and ionic densities are superposed, as well as electronic and ionic velocities.

For the partial temperatures presented in figure 2, the intermediate state reached between the rarefaction wave and the contact discontinuity is identical for all schemes, which is expected since this part of the solution is smooth. However, the states obtained between the contact discontinuity and the shock are different for every scheme. Figures 3 display a zoom on this part of the solution. The results can be seen to be different for all three methods. The behaviour of the different schemes should be investigated further.

Shock tube with different initial temperatures. This test case is similar to the previous one, the only difference being the initial left ionic and electronic temperatures. The following initial

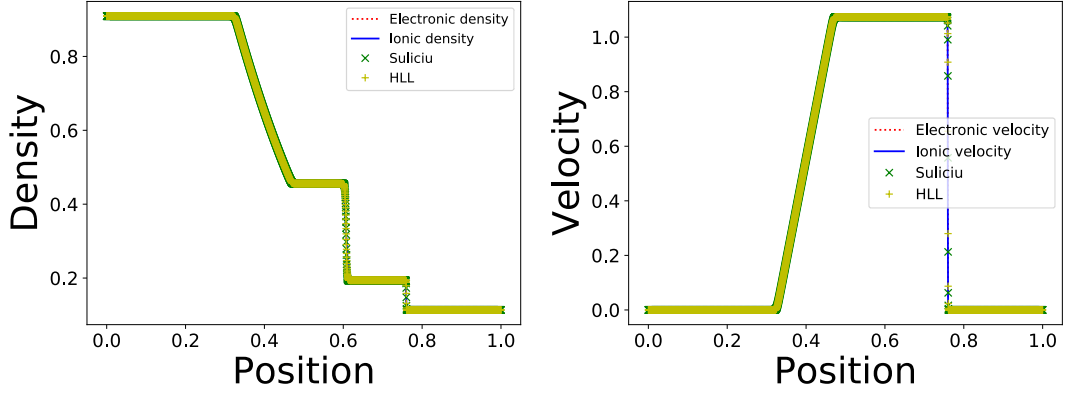


FIGURE 1. Density and velocity solutions of shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

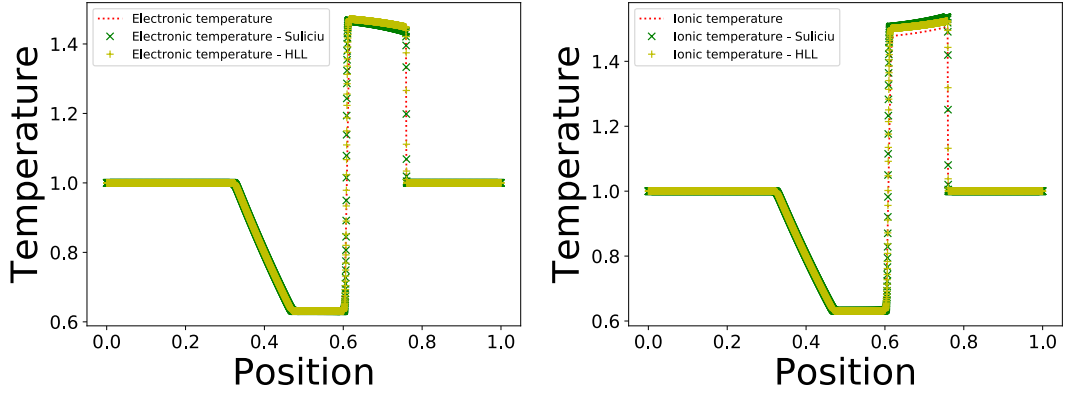


FIGURE 2. Electronic and ionic temperatures of a shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

conditions are:

$$\begin{cases} n^\alpha(t_0, x) = 1, & u^\alpha(t_0, x) = 0, & T^\alpha(t_0, x) = 1 & \text{if } x \in [0, 0.5], \\ n^\alpha(t_0, x) = 0.125, & u^\alpha(t_0, x) = 0, & T^e(t_0, x) = 2 & T^i(t_0, x) = 3 & \text{if } x \in [0.5, 1]. \end{cases}$$

Parameters are chosen as $N_x = 120000$, $N_v = 40$ and $l = 8$. The inter-species collision relaxation time is $\tau^{ei} = 0.1$. Results are computed at time $t = 0.1$. The result are provided in figure 4 and 5. The behaviour of the result if very similar to the previous one. Quasi-neutrality is achieved and temperatures exhibit different jump relations across discontinuities for each method. Rarefaction wave. This test case is constituted of two rarefaction waves going in opposite directions. It is given by the following data:

$$\begin{cases} n^\alpha(t_0, x) = 1, u^\alpha(t_0, x) = -1, T^\alpha(t_0, x) = 1 & \text{if } x \in [0, 0.5], \\ n^\alpha(t_0, x) = 1, u^\alpha(t_0, x) = 1, T^\alpha(t_0, x) = 1 & \text{if } x \in [0.5, 1]. \end{cases}$$

We have $N_x = 120000$, $N_v = 40$ and $l = 8$. The inter-species collision relaxation time is $\tau^{ei} = 0.1$. Results are computed at time $t = 0.1$.

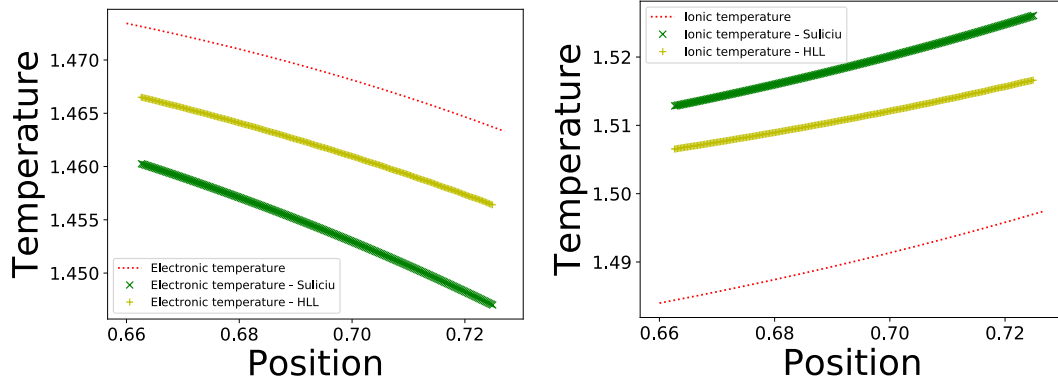


FIGURE 3. Different jump relations through shocks of electronic and ionic temperatures for a shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

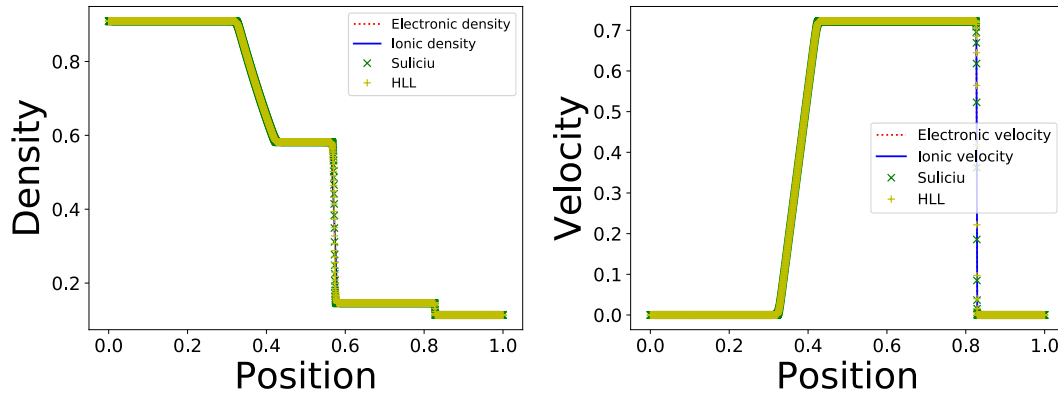


FIGURE 4. Density and velocity solutions of shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

The result are provided in figure 7 and 8. In this test case, the solution is smooth. There are no shock waves, and then, all methods are expected to give the same results. Figure 9 shows that this is what happens.

Shock wave. This test case is constituted of two shock waves. It is given by the following data:

$$\begin{cases} n^\alpha(t_0, x) = 1, u^\alpha(t_0, x) = 1, T^\alpha(t_0, x) = 1 & \text{if } x \in [0, 0.5], \\ n^\alpha(t_0, x) = 1, u^\alpha(t_0, x) = -1, T^\alpha(t_0, x) = 1 & \text{if } x \in [0.5, 1]. \end{cases}$$

We have $N_x = 120000$, $N_v = 40$ and $l = 8$. The inter-species collision relaxation time is $\tau^{ei} = 0.1$. Results are computed at time $t = 0.1$.

The results are provided in Figure 10 and 11. Quasi-neutrality is achieved on the conserved quantities. The solution contains two shock waves and different jump relations can be observed for each numerical method. In Figure 12, a zoom is performed on the constant values reached by each method between the two shocks. This shows the different behaviour between the three methods.

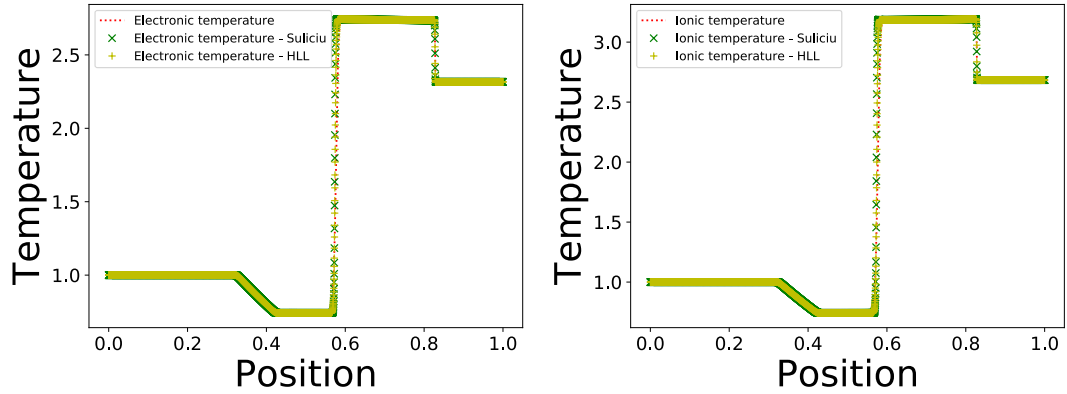


FIGURE 5. Electronic and ionic temperatures of a shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

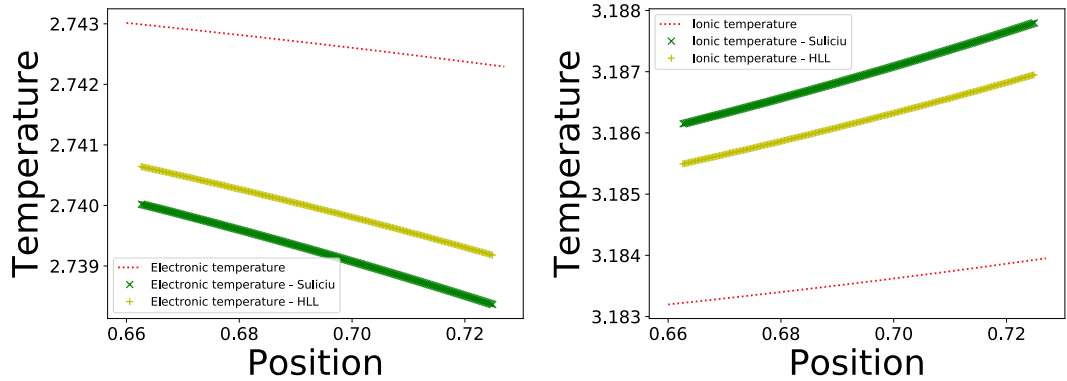


FIGURE 6. Different jump relations of electronic and ionic temperatures for a shock tube test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

5. CONCLUSION

In this article, a kinetic numerical method able to provide reference results for a non-conservative hyperbolic system, the bi-temperature Euler system, is proposed. This method is able to enforce all the desired properties that are relevant for comparisons with methods applied directly to the hyperbolic system. Such methods lack an unambiguous definition of the non-conservative products and exhibit different Rankine-Hugoniot relation when the solution contains shock waves. However, the proposed method introduces a important amount of numerical viscosity, which renders the method costly. A second-order method able to circumvent such a drawback is a current on-going work, which should be the subject of a subsequent article.

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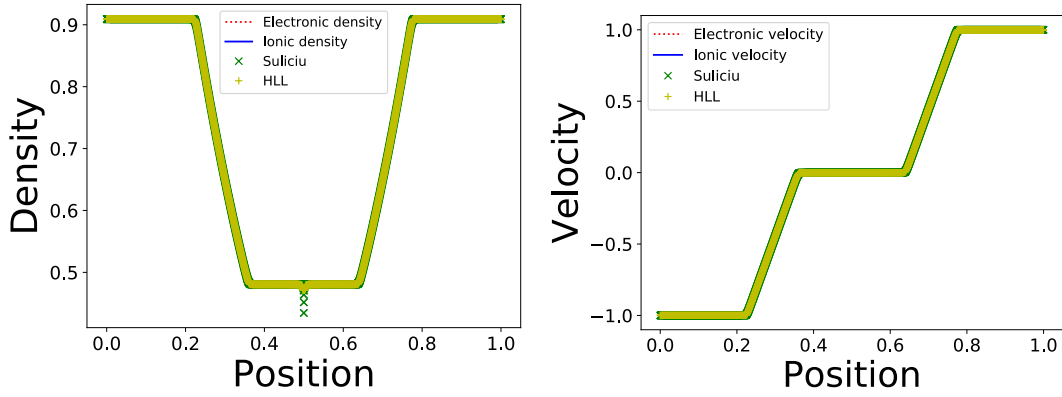


FIGURE 7. Density and velocity solution of a rarefaction wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

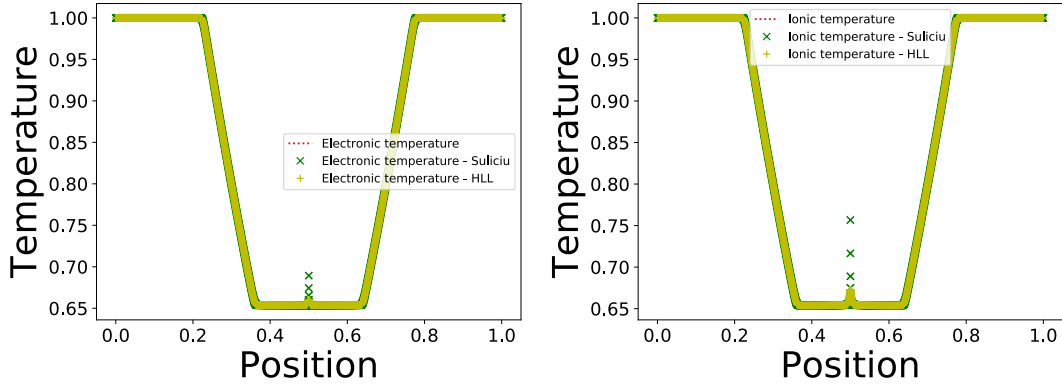


FIGURE 8. Temperature solutions of a rarefaction wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

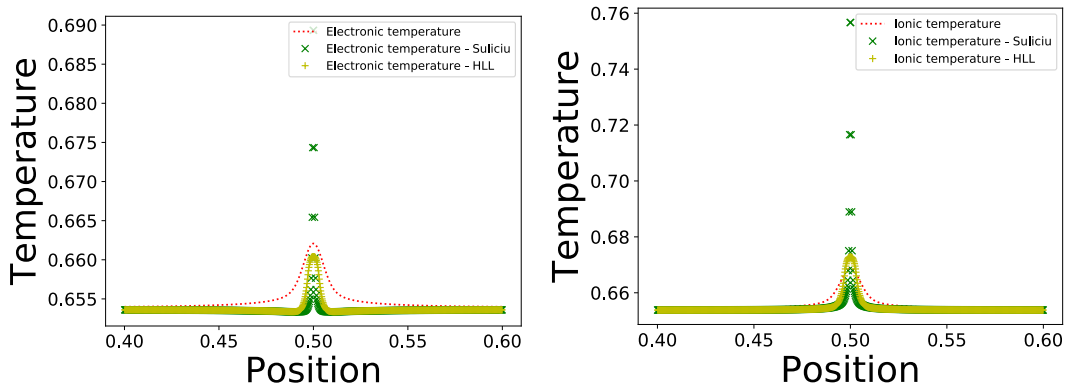


FIGURE 9. Zoom on temperature solutions of a rarefaction wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

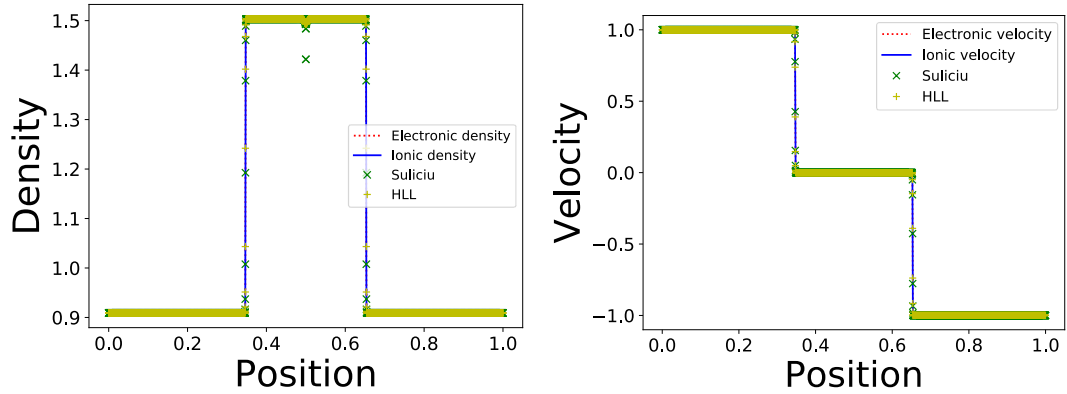


FIGURE 10. Density and velocity solutions of a shock wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

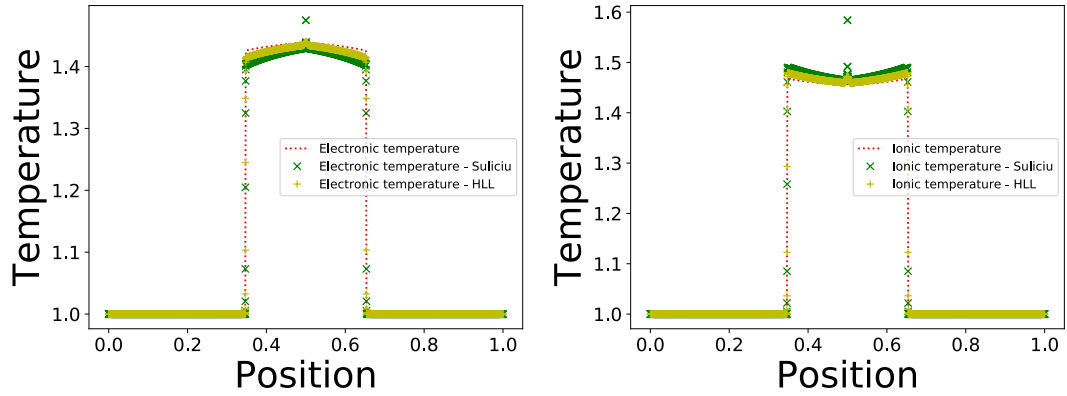


FIGURE 11. Electronic and ionic temperature of a shock wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

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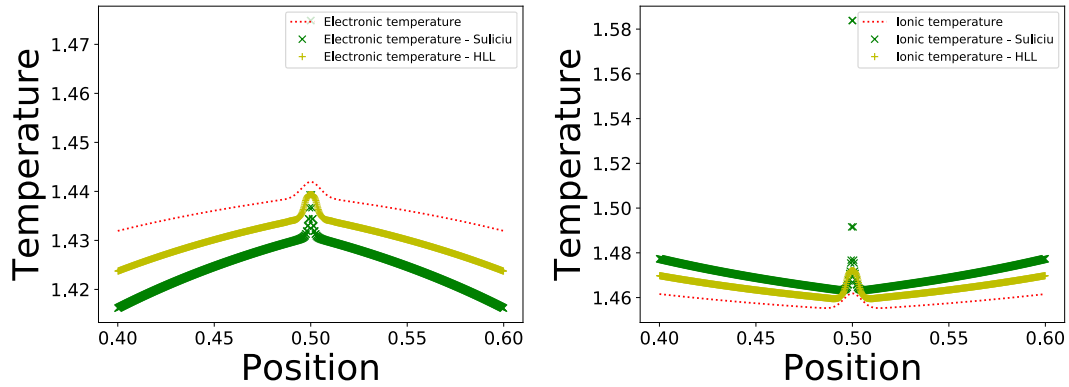


FIGURE 12. Electronic and ionic temperature of a shock wave test case with a mass ratio of 10 with 120000 space points, 40 velocity points and a domain length of 8

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