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An explicit representation for the axisymmetric solutions of the free Maxwell equations

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Short title: *Explicit representation for axisymmetric free Maxwell fields*

Abstract

Garay-Avendaño & Zamboni-Rached (2014) defined two classes of axisymmetric solutions of the free Maxwell equations. We show that, by combining these two classes of solutions, one is able to describe in explicit form all time-harmonic axisymmetric free Maxwell fields, and hence, by summation on frequencies, all axisymmetric free Maxwell fields. This provides an explicit representation for these fields. It will be important, e.g., to have the interstellar radiation field in a disc galaxy modelled as an exact solution of the free Maxwell equations.

Keywords: Maxwell equations; axial symmetry; exact solutions; electromagnetic duality.

1 Introduction

Axially symmetric solutions of the Maxwell equations are quite important, at least as an often relevant approximation. For instance, axisymmetric magnetic fields occur naturally as produced by systems possessing an axis of revolution, such as disks or coils [1, 2], or astrophysical systems like accretion disks [3] or disk galaxies [4]. Axisymmetric solutions are also used to model

EM beams and their propagation (e.g. [5, 6]). In particular, non-diffracting beams are usually endowed with axial symmetry, see e.g. Refs. [7, 8, 9]. Naturally, one often considers time-harmonic solutions, since a general time dependence is got by summing such solutions. Two classes of time-harmonic axisymmetric solutions of the free Maxwell equations, mutually associated by EM duality, have been introduced recently [10]. The main aim was to “describe in exact and analytic form the propagation of nonparaxial scalar and electromagnetic beams.” However, as noted by the authors of Ref. [10], the analytical expression for a time-harmonic axisymmetric solution Ψ of the scalar wave equation, from which they start [Eq. (1) below], covers all such solutions [9] — thus not merely ones corresponding to nonparaxial scalar beams. The first class of EM fields defined in Ref. [10] is obtained by associating with any such scalar solution Ψ a vector potential \mathbf{A} by Eq. (10) below, and the second class is deduced from the first one by EM duality [10].

The aim of the present work is to show that, somewhat surprisingly, by combining these two classes one is able to describe *all* time-harmonic axisymmetric EM fields (and thus all axisymmetric EM fields). More precisely, we shall prove that any time-harmonic axisymmetric EM field is the sum of two EM fields, say $(\mathbf{E}_1, \mathbf{B}_1)$ and $(\mathbf{E}'_2, \mathbf{B}'_2)$, deduced from two time-harmonic axisymmetric solutions of the scalar wave equation, say Ψ_1 and Ψ_2 : the first EM field derives from the vector potential $\mathbf{A}_1 = \Psi_1 \mathbf{e}_z$, the second one is deduced by EM duality from the EM field that derives from the vector potential $\mathbf{A}_2 = \Psi_2 \mathbf{e}_z$. Because this method is based on determining two vector potentials from merely two scalar fields, this result was not necessarily expected. Section 2 presents and comments the results of Ref. [10]. (Equations (15)–(20) are new.) Section 3 gives the proof of the announced result, which is not immediate but uses standard mathematics with which one is familiar from classical field theory. Section 4 presents the Conclusion.

2 From scalar waves to Maxwell fields

2.1 Axially symmetric scalar waves

We adopt cylindrical coordinates ρ, ϕ, z about the symmetry axis, that is the z axis. Any *time-harmonic* axisymmetric solution of the *scalar* wave equation

(of d'Alembert) can be written [9, 10] as a sum of Bessel beams [7]:

$$\Psi_{\omega S}(t, \rho, z) = e^{-i\omega t} \int_{-K}^{+K} J_0\left(\rho\sqrt{K^2 - k^2}\right) e^{ikz} S(k) dk, \quad (1)$$

with ω the angular frequency, $K := \omega/c$,¹ and J_0 the first-kind Bessel function of order 0. (c is the velocity of light.) Thus the (axial) “wave vector spectrum” S is a (generally complex) function of the real variable $k = k_z$ ($-K \leq k \leq K$), that is the projection of the wave vector on the z axis. This function S determines the spatial dependency of the time-harmonic solution (1) in the two-dimensional space left by the axial symmetry, i.e. the half-plane ($\rho \geq 0, z \in]-\infty, +\infty[$).

Thus, any time-harmonic axisymmetric scalar wave Ψ can be put in the explicit form (1), in which no restriction has to be put on the “wave vector spectrum” S (except for a minimal regularity ensuring that the function Ψ is at least twice continuously differentiable: the integrability of S , $S \in L^1([-K, +K])$, would be enough for this). Of course, the *general* axisymmetric solution of the scalar wave equation can be got from (1) by an appropriate summation over a frequency spectrum: an integral (inverse Fourier transform) in the general case, or a discrete sum if a discrete frequency spectrum (ω_j) ($j = 1, \dots, N_\omega$) is considered for simplicity:

$$\Psi(t, \rho, z) = \sum_{j=1}^{N_\omega} \Psi_{\omega_j S_j}(t, \rho, z), \quad (2)$$

where, for $j = 1, \dots, N_\omega$, $\Psi_{\omega_j S_j}$ is the time-harmonic solution (1), corresponding with frequency ω_j and wave vector spectrum S_j . The different weights w_j which may be affected to the different frequencies can be included in the functions S_j , replacing S_j by $w_j S_j$.

2.2 Reminder on time-harmonic free Maxwell fields

In this subsection, we recall equations more briefly recalled in Ref. [10]. The electric and magnetic fields in SI units are given in terms of the scalar and

¹ Beware that instead $K := 2\omega/c$ in Ref. [10]. Our notation seems as natural and gives more condensed formulas.

(3-)vector potentials V and \mathbf{A} by

$$\mathbf{E} = -\nabla V - \frac{\partial \mathbf{A}}{\partial t}, \quad (3)$$

$$\mathbf{B} = \text{rot} \mathbf{A}. \quad (4)$$

These equations imply that \mathbf{E} and \mathbf{B} obey the first group of Maxwell equations. If one imposes the Lorenz gauge condition

$$\frac{1}{c^2} \frac{\partial V}{\partial t} + \text{div} \mathbf{A} = 0, \quad (5)$$

then the validity of the second group of the Maxwell equations in free space for \mathbf{E} and \mathbf{B} is equivalent to ask that V and \mathbf{A} verify d'Alembert's wave equation [11, 12]. Moreover, if one assumes a harmonic time-dependence for V and \mathbf{A} :

$$V(t, \mathbf{x}) = e^{-i\omega t} \hat{V}(\mathbf{x}), \quad \mathbf{A}(t, \mathbf{x}) = e^{-i\omega t} \hat{\mathbf{A}}(\mathbf{x}), \quad (6)$$

then the wave equation for \mathbf{A} becomes the Helmholtz equation: ²

$$\Delta \mathbf{A} + \frac{\omega^2}{c^2} \mathbf{A} = \mathbf{0}, \quad (7)$$

and the Lorenz gauge condition (5) rewrites as

$$V = -i \frac{c^2}{\omega} \text{div} \mathbf{A}. \quad (8)$$

If \mathbf{A} is time-harmonic [Eq. (6)₂] and obeys (7), then V given by (8) is time-harmonic and automatically satisfies the wave equation. The electric field (3) is then easily rewritten as

$$\mathbf{E} = i\omega \mathbf{A} + i \frac{c^2}{\omega} \nabla (\text{div} \mathbf{A}). \quad (9)$$

Thus, the data of a time-harmonic vector potential \mathbf{A} obeying the wave equation, or equivalently obeying Eq. (7), determines a unique solution of the free Maxwell equations, by Eqs. (4) and (9), and that solution is time-harmonic with the same frequency ω as for \mathbf{A} .

² Of course, just the same equation (7) applies to the ‘‘amplitude field’’ $\hat{\mathbf{A}}$.

2.3 Time-harmonic axisymmetric fields: from a scalar wave to a Maxwell field

To any axisymmetric time-harmonic solution $\Psi(t, \rho, z) = e^{-i\omega t} \hat{\Psi}(\rho, z)$ of the scalar wave equation, the authors of Ref. [10] associate a vector potential \mathbf{A} by

$$\mathbf{A} := \Psi \mathbf{e}_z, \quad \text{or} \quad A_z := \Psi, \quad A_\rho = A_\phi = 0. \quad (10)$$

(We shall denote by $(\mathbf{e}_\rho, \mathbf{e}_\phi, \mathbf{e}_z)$ the standard point-dependent orthonormal basis associated with the cylindrical coordinates.) Thus, in the way recalled in the foregoing subsection, they define a unique solution of the free Maxwell equations, which is time-harmonic. Their equations for the different components of this solution (\mathbf{E}, \mathbf{B}) are as follows:

$$B_\phi = -\frac{\partial A_z}{\partial \rho}, \quad E_\phi = 0, \quad (11)$$

$$E_\rho = i \frac{c^2}{\omega} \frac{\partial^2 A_z}{\partial \rho \partial z}, \quad B_\rho = 0, \quad (12)$$

$$E_z = i \frac{c^2}{\omega} \frac{\partial^2 A_z}{\partial z^2} + i\omega A_z, \quad B_z = 0. \quad (13)$$

These equations follow easily from Eqs. (4), (9) and (10), and from the axisymmetry of $A_z = \Psi(t, \rho, z)$, by using the standard formulas for the curl and divergence in cylindrical coordinates. Equations (11)–(13) provide an axisymmetric EM field whose electric field is radially polarized ($\mathbf{E} = E_\rho \mathbf{e}_\rho + E_z \mathbf{e}_z$), in short a “radially polarized” EM field.

An “azimuthally polarized” solution $(\mathbf{E}', \mathbf{B}')$ (in the sense that $\mathbf{E}' = E'_\phi \mathbf{e}_\phi$) of the free Maxwell equations can alternatively be deduced from the data Ψ , by transforming the solution (11)–(13) through the EM duality, that is:

$$\mathbf{E}' = c\mathbf{B}, \quad \mathbf{B}' = -\mathbf{E}/c. \quad (14)$$

The two axisymmetric time-harmonic solutions of the free Maxwell equations, thus associated with an axisymmetric time-harmonic solution Ψ of the scalar wave equation, will be called here the “GAZR1 solution” and the “GAZR2 solution”, respectively, because both were derived in Ref. [10].

As usual, it is implicit that, in Eqs. (11)–(13), B_ϕ , E_ρ and E_z are actually the *real parts* of the respective r.h.s. [as are also \mathbf{E} and \mathbf{B} in Eqs. (3), (4), and (9)]. Thus, with A_z given by (1) (as can always be the case in view of the general representativity of that equation), and using the fact that $dJ_0/dx = -J_1(x)$, we obtain:

$$B_{\phi\omega S} = \mathcal{R}e \left[e^{-i\omega t} \int_{-K}^{+K} \sqrt{K^2 - k^2} J_1 \left(\rho \sqrt{K^2 - k^2} \right) S(k) e^{ikz} dk \right], \quad (15)$$

$$E_{\rho\omega S} = \mathcal{R}e \left[-i \frac{c^2}{\omega} e^{-i\omega t} \int_{-K}^{+K} \sqrt{K^2 - k^2} J_1 \left(\rho \sqrt{K^2 - k^2} \right) ik S(k) e^{ikz} dk \right], \quad (16)$$

$$E_{z\omega S} = \mathcal{R}e \left[i e^{-i\omega t} \int_{-K}^{+K} J_0 \left(\rho \sqrt{K^2 - k^2} \right) \left(\omega - \frac{c^2}{\omega} k^2 \right) S(k) e^{ikz} dk \right], \quad (17)$$

where $K := \omega/c$. In the case with a (discrete) frequency spectrum, one just has to sum each component: (15), (16), or (17), over the different frequencies ω_j , with the corresponding values $K_j = \omega_j/c$ and spectra $S_j = S_j(k)$ ($-K_j \leq k \leq +K_j$) — as with a scalar wave (2):

$$B_\phi = \sum_{j=1}^{N_\omega} B_{\phi\omega_j} S_j, \quad (18)$$

$$E_\rho = \sum_{j=1}^{N_\omega} E_{\rho\omega_j} S_j, \quad (19)$$

$$E_z = \sum_{j=1}^{N_\omega} E_{z\omega_j} S_j. \quad (20)$$

3 From Maxwell fields to scalar waves

Now an important question arises: Do the GAZR solutions generate all axisymmetric time-harmonic solutions of the Maxwell equations (in which case, by summation on frequencies, they would generate all axisymmetric solutions

of the Maxwell equations)? That is: let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. Can one find a GAZR1 solution and a GAZR2 solution, whose sum give just that starting solution?

Note from Eqs. (11)–(13) and (14) that *the GAZR1 solution and the GAZR2 solution are complementary*: in cylindrical coordinates, the GAZR1 solution provides non-zero components B_ϕ, E_ρ, E_z , the other components E_ϕ, B_ρ, B_z being zero — and the exact opposite is true for the GAZR2 solution. In view of this complementarity, we can consider separately the two sets of components: B_ϕ, E_ρ, E_z on one side, and E_ϕ, B_ρ, B_z on the other side.

3.1 Sufficient conditions for the existence of the decomposition

For the “GAZR1” solution, which gives non-zero values to the first among the two sets of components just mentioned, we have the following result:

Proposition 1. *Let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. In order that a time-harmonic axisymmetric solution $(A_{1z}, B_{1\phi}, E_{1\rho}, E_{1z}; E_{1\phi} = B_{1\rho} = B_{1z} = 0)$, of the form (11)–(13), and having the same frequency ω as the starting solution $(\mathbf{A}, \mathbf{E}, \mathbf{B})$, be such that $B_{1\phi} = B_\phi, E_{1\rho} = E_\rho, E_{1z} = E_z$, it is sufficient that we have just*

$$B_{1\phi} = B_\phi. \quad (21)$$

Proof. Let $A_{1z}(t, \rho, z)$ be a time-harmonic axisymmetric solution of the wave equation, with frequency ω , and assume that $B_{1\phi}$ as defined by Eq. (11) [with A_{1z} in the place of A_z] is equal to B_ϕ , where \mathbf{B} is defined by Eq. (4). I.e., assume that

$$-\frac{\partial A_{1z}}{\partial \rho} = \frac{\partial A_\rho}{\partial z} - \frac{\partial A_z}{\partial \rho}. \quad (22)$$

Denoting by $\mathbf{A}_1 := A_{1z}\mathbf{e}_z$ the vector potential that provides the GAZR1 solution $(B_{1\phi}, E_{1\rho}, E_{1z}; E_{1\phi} = B_{1\rho} = B_{1z} = 0)$, let us compute $E_\rho - E_{1\rho}$ and $E_z - E_{1z}$. We have by Eq. (9):

$$\frac{\omega}{ic^2} (\mathbf{E} - \mathbf{E}_1) = \nabla (\text{div} \mathbf{A}') + \frac{\omega^2}{c^2} \mathbf{A}', \quad (23)$$

where

$$\mathbf{A}' := \mathbf{A} - \mathbf{A}_1 := \mathbf{A} - A_{1z}\mathbf{e}_z. \quad (24)$$

In order that the vector potential \mathbf{A} of the *a priori* given solution $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be axisymmetric, its components A_ρ , A_ϕ , A_z must depend only on t, ρ, z , i.e., be independent of ϕ . Therefore:

$$\operatorname{div}\mathbf{A} = \frac{1}{\rho} \frac{\partial(\rho A_\rho)}{\partial\rho} + \frac{\partial A_z}{\partial z}, \quad (25)$$

and, using this and (24):

$$\operatorname{div}\mathbf{A}' = \frac{1}{\rho} \frac{\partial(\rho A_\rho)}{\partial\rho} + \frac{\partial A'_z}{\partial z}. \quad (26)$$

Hence, in (23), we have

$$\begin{aligned} \nabla(\operatorname{div}\mathbf{A}') &= \nabla\left(\frac{\partial A_\rho}{\partial\rho} + \frac{A_\rho}{\rho} + \frac{\partial A'_z}{\partial z}\right) \\ &= \left(\frac{\partial^2 A_\rho}{\partial\rho^2} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial\rho} - \frac{1}{\rho^2} A_\rho + \frac{\partial^2 A'_z}{\partial\rho\partial z}\right)\mathbf{e}_\rho + \left(\frac{\partial^2 A_\rho}{\partial z\partial\rho} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial z} + \frac{\partial^2 A'_z}{\partial z^2}\right)\mathbf{e}_z. \end{aligned} \quad (27)$$

The radial component of the vector (23) is thus:

$$\frac{\omega}{ic^2}(\mathbf{E} - \mathbf{E}_1)_\rho = \frac{\partial^2 A_\rho}{\partial\rho^2} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial\rho} - \frac{A_\rho}{\rho^2} + \frac{\partial^2 A_z}{\partial\rho\partial z} - \frac{\partial^2 A_{1z}}{\partial\rho\partial z} + \frac{\omega^2}{c^2} A_\rho. \quad (28)$$

However, the vector potential \mathbf{A} obeys the Helmholtz equation (7), that is for the radial component (using the fact that $\frac{\partial A_\rho}{\partial\phi} = \frac{\partial A_\phi}{\partial\rho} \equiv 0$):

$$\begin{aligned} (\Delta\mathbf{A})_\rho + \frac{\omega^2}{c^2} A_\rho &\equiv \Delta A_\rho - \frac{A_\rho}{\rho^2} - \frac{2}{\rho^2} \frac{\partial A_\phi}{\partial\phi} + \frac{\omega^2}{c^2} A_\rho \\ &\equiv \frac{\partial^2 A_\rho}{\partial\rho^2} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial\rho} + \frac{\partial^2 A_\rho}{\partial z^2} - \frac{A_\rho}{\rho^2} + \frac{\omega^2}{c^2} A_\rho = 0. \end{aligned} \quad (29)$$

Inserting (29) into (28) gives us:

$$\frac{\omega}{ic^2}(\mathbf{E} - \mathbf{E}_1)_\rho = -\frac{\partial^2 A_\rho}{\partial z^2} + \frac{\partial^2 A_z}{\partial\rho\partial z} - \frac{\partial^2 A_{1z}}{\partial\rho\partial z} = -\frac{\partial}{\partial z} \left(\frac{\partial A_\rho}{\partial z} - \frac{\partial A_z}{\partial\rho} + \frac{\partial A_{1z}}{\partial\rho} \right). \quad (30)$$

Therefore, if Eq. (22) is satisfied, then we have $E_{1\rho} = E_\rho$.

Similarly, from (27), the axial component of the vector (23) is

$$\frac{\omega}{ic^2} (\mathbf{E} - \mathbf{E}_1)_z = \frac{\partial^2 A_\rho}{\partial z \partial \rho} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial z} + \frac{\partial^2 A_z}{\partial z^2} - \frac{\partial^2 A_{1z}}{\partial z^2} + \frac{\omega^2}{c^2} (A_z - A_{1z}). \quad (31)$$

On the other hand, the axial component of the Helmholtz equation (7) is

$$\begin{aligned} (\Delta \mathbf{A})_z + \frac{\omega^2}{c^2} A_z &\equiv \Delta A_z + \frac{\omega^2}{c^2} A_z \\ &\equiv \frac{\partial^2 A_z}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial A_z}{\partial \rho} + \frac{\partial^2 A_z}{\partial z^2} + \frac{\omega^2}{c^2} A_z = 0. \end{aligned} \quad (32)$$

If Eq. (22) is satisfied, we have

$$\frac{\partial^2 A_z}{\partial \rho^2} = \frac{\partial^2 A_\rho}{\partial \rho \partial z} + \frac{\partial^2 A_{1z}}{\partial \rho^2}. \quad (33)$$

In Eq. (32), we replace $\frac{\partial^2 A_z}{\partial \rho^2}$ by its value given on the r.h.s. above, and we replace $\frac{\partial A_z}{\partial \rho}$ by its value given by Eq. (22). This gives:

$$\frac{\partial^2 A_\rho}{\partial \rho \partial z} + \frac{\partial^2 A_{1z}}{\partial \rho^2} + \frac{1}{\rho} \frac{\partial A_\rho}{\partial z} + \frac{1}{\rho} \frac{\partial A_{1z}}{\partial \rho} + \frac{\partial^2 A_z}{\partial z^2} + \frac{\omega^2}{c^2} A_z = 0. \quad (34)$$

Using this equation in Eq. (31), we rewrite the latter as

$$\frac{\omega}{ic^2} (\mathbf{E} - \mathbf{E}_1)_z = -\frac{\partial^2 A_{1z}}{\partial \rho^2} - \frac{1}{\rho} \frac{\partial A_{1z}}{\partial \rho} - \frac{\partial^2 A_{1z}}{\partial z^2} - \frac{\omega^2}{c^2} A_{1z}. \quad (35)$$

We recognize the r.h.s as that of Eq. (32), though with the minus sign, and with A_{1z} in the place of A_z . I.e., Eq. (35) is just

$$\frac{\omega}{ic^2} (\mathbf{E} - \mathbf{E}_1)_z = -\Delta A_{1z} - \frac{\omega^2}{c^2} A_{1z}. \quad (36)$$

But this is zero, since A_{1z} is by assumption a time-harmonic solution of the wave equation, with frequency ω . Therefore, if Eq. (22) is satisfied, then we have $E_{1z} = E_z$, too. This completes the proof of Proposition 1. \square

It follows for the dual (“GAZR2”) solution:

Corollary 1. *Let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. In order that a time-harmonic solution $(A_{2z}, E'_{2\phi}, B'_{2\rho}, B'_{2z}; B'_{2\phi} = E'_{2\rho} = E'_{2z} = 0)$ with the same frequency, deduced from Eqs. (11)–(13) by the duality (14), be such that $E'_{2\phi} = E_\phi$, $B'_{2\rho} = B_\rho$, $B'_{2z} = B_z$, it is sufficient that we have just*

$$E'_{2\phi} = E_\phi. \quad (37)$$

Proof. The GAZR2 solution $(A_{2z}, E'_{2\phi}, B'_{2\rho}, B'_{2z}; B'_{2\phi} = E'_{2\rho} = E'_{2z} = 0)$ is deduced from the GAZR1 solution $(A_{2z}, B_{2\phi}, E_{2\rho}, E_{2z}; E_{2\phi} = B_{2\rho} = B_{2z} = 0)$, associated with the same potential A_{2z} , by the duality relation (14). Suppose that Eq. (37) is satisfied. With the starting solution $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ of the free Maxwell equations, we may associate another solution, by the inverse duality:

$$\tilde{\mathbf{B}} = \frac{1}{c}\mathbf{E}, \quad \tilde{\mathbf{E}} = -c\mathbf{B}. \quad (38)$$

The assumed relation (37) means that

$$B_{2\phi} = \tilde{B}_\phi. \quad (39)$$

Indeed, by applying successively (14)₁, (37), and (38)₁, we obtain:

$$B_{2\phi} = \frac{1}{c}E'_{2\phi} = \frac{1}{c}E_\phi = \tilde{B}_\phi. \quad (40)$$

In turn, the relation (39) means that we may apply Proposition 1 to the GAZR1 solution $(A_{2z}, B_{2\phi}, \dots)$ and the solution $(\tilde{\mathbf{A}}, \tilde{\mathbf{E}}, \tilde{\mathbf{B}})$.³ Thus, Proposition 1 tells us here that, since $B_{2\phi} = \tilde{B}_\phi$, we have also

$$E_{2\rho} = \tilde{E}_\rho \quad \text{and} \quad E_{2z} = \tilde{E}_z, \quad (41)$$

³ The explicit expression of the corresponding vector potential $\tilde{\mathbf{A}}$ as function of $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ is not needed: only the existence of an axisymmetric $\tilde{\mathbf{A}}$, such that $\tilde{\mathbf{B}} = \text{rot}\tilde{\mathbf{A}}$, is needed. Precisely, $\tilde{\mathbf{A}}$ is got as a solution of the PDE $\text{rot}\tilde{\mathbf{A}} = \tilde{\mathbf{B}}$. (Such a solution always exists in a topologically trivial domain, thus in particular if the domain is the whole space.) Hence $\tilde{\mathbf{A}}$ can indeed be chosen axisymmetric, i.e. with its components in cylindrical coordinates being independent of ϕ , because with this choice the independent variables of the PDE $\text{rot}\tilde{\mathbf{A}} = \tilde{\mathbf{B}}$ simply do not include ϕ — since $\tilde{\mathbf{B}} = \frac{1}{c}\mathbf{E}$ is axisymmetric, as is \mathbf{E} by assumption.

i.e., in view of (14)₂ and (38)₂:

$$-cB'_{2\rho} = -cB_\rho \quad \text{and} \quad -cB'_{2z} = -cB_z. \quad (42)$$

This proves Corollary 1. \square

3.2 Generality of the decomposition

Proposition 2. *Let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. There exists a time-harmonic axisymmetric solution A_{2z} of the wave equation, with the same frequency, such that the associated GAZR2 solution $(\mathbf{E}'_2, \mathbf{B}'_2)$, deduced from A_{2z} by Eqs. (11)–(13) followed by the duality transformation (14), satisfy*

$$E'_{2\phi} = E_\phi, \quad B'_{2\rho} = B_\rho, \quad B'_{2z} = B_z. \quad (43)$$

Proof. In view of Corollary 1, we merely have to prove that there exists a time-harmonic axisymmetric solution A_{2z} of the wave equation, such that Eq. (37) is satisfied. From Eqs. (9) and (25), we have

$$E_\phi = i\omega A_\phi. \quad (44)$$

Using this with Eqs. (11) and (14)₁, we may rewrite the sought-for relation (37) as:

$$-\frac{\partial A_{2z}}{\partial \rho} = \frac{i\omega}{c} A_\phi. \quad (45)$$

This equation can be solved by a quadrature:

$$A_{2z}(t, \rho, z) = h(t, z) + \int_{\rho_0}^{\rho} -\frac{i\omega}{c} A_\phi(t, \rho', z) d\rho'. \quad (46)$$

We thus have to find out if it is possible to determine the function h so that A_{2z} given by (46) obey the wave equation. Moreover, the unknown function A_{2z} must have a harmonic time dependence with frequency ω as has A_ϕ , i.e.,

$$A_{2z}(t, \rho, z) = \psi(\rho, z) e^{-i\omega t}, \quad A_\phi(t, \rho, z) = \hat{A}_\phi(\rho, z) e^{-i\omega t}, \quad (47)$$

so we must have $h(t, z) = e^{-i\omega t} g(z)$, too. Hence we may rewrite (46) as

$$\psi(\rho, z) = g(z) + \int_{\rho_0}^{\rho} -\frac{i\omega}{c} \hat{A}_\phi(\rho', z) d\rho', \quad (48)$$

and now the question is to know if g can be determined so that ψ obey the scalar Helmholtz equation, i.e. [cf. Eq. (32)]:

$$\mathcal{H}\psi := \Delta\psi + \frac{\omega^2}{c^2}\psi \equiv \frac{\partial^2\psi}{\partial\rho^2} + \frac{1}{\rho}\frac{\partial\psi}{\partial\rho} + \frac{\partial^2\psi}{\partial z^2} + \frac{\omega^2}{c^2}\psi = 0 \quad (49)$$

— knowing that A_ϕ or \hat{A}_ϕ does obey the ϕ component of the vector Helmholtz equation (7), i.e.:

$$\begin{aligned} \left(\Delta\hat{\mathbf{A}}\right)_\phi + \frac{\omega^2}{c^2}\hat{A}_\phi &\equiv \Delta\hat{A}_\phi - \frac{\hat{A}_\phi}{\rho^2} + \frac{2}{\rho^2}\frac{\partial\hat{A}_\rho}{\partial\phi} + \frac{\omega^2}{c^2}\hat{A}_\phi \\ &\equiv \frac{\partial^2\hat{A}_\phi}{\partial\rho^2} + \frac{1}{\rho}\frac{\partial\hat{A}_\phi}{\partial\rho} + \frac{\partial^2\hat{A}_\phi}{\partial z^2} - \frac{\hat{A}_\phi}{\rho^2} + \frac{\omega^2}{c^2}\hat{A}_\phi = 0. \end{aligned} \quad (50)$$

We have from Eqs. (45) and (47):

$$\frac{\partial\psi}{\partial\rho} = -\frac{i\omega}{c}\hat{A}_\phi, \quad (51)$$

hence

$$\frac{\partial^2\psi}{\partial\rho^2} = -\frac{i\omega}{c}\frac{\partial\hat{A}_\phi}{\partial\rho}. \quad (52)$$

And we get from (48):

$$\frac{\partial\psi}{\partial z} = \frac{dg}{dz} - \int_{\rho_0}^{\rho} \frac{i\omega}{c} \frac{\partial\hat{A}_\phi}{\partial z}(\rho', z) d\rho', \quad (53)$$

whence

$$\frac{\partial^2\psi}{\partial z^2} = \frac{d^2g}{dz^2} - \int_{\rho_0}^{\rho} \frac{i\omega}{c} \frac{\partial^2\hat{A}_\phi}{\partial z^2}(\rho', z) d\rho'. \quad (54)$$

Entering Eqs. (51), (52) and (54) into (49)₁, we obtain:

$$\begin{aligned} \mathcal{H}\psi &= -\frac{i\omega}{c}\frac{\partial\hat{A}_\phi}{\partial\rho} - \frac{i\omega}{c}\frac{\hat{A}_\phi}{\rho} + \frac{d^2g}{dz^2} - \int_{\rho_0}^{\rho} \frac{i\omega}{c} \frac{\partial^2\hat{A}_\phi}{\partial z^2}(\rho', z) d\rho' \\ &\quad + \frac{\omega^2}{c^2} \left(g - \frac{i\omega}{c} \int_{\rho_0}^{\rho} \hat{A}_\phi(\rho', z) d\rho' \right). \end{aligned} \quad (55)$$

Therefore, the scalar Helmholtz equation (49)₂ rewrites as

$$\frac{d^2g}{dz^2} + \frac{\omega^2}{c^2}g = \frac{i\omega}{c} \left[\frac{\partial\hat{A}_\phi}{\partial\rho} + \frac{\hat{A}_\phi}{\rho} + \int_{\rho_0}^{\rho} \left(\frac{\partial^2\hat{A}_\phi}{\partial z^2} + \frac{\omega^2}{c^2}\hat{A}_\phi \right) d\rho' \right], \quad (56)$$

or, using (50):

$$\frac{d^2g}{dz^2} + \frac{\omega^2}{c^2}g = \frac{i\omega}{c} \left[\frac{\partial \hat{A}_\phi}{\partial \rho} + \frac{\hat{A}_\phi}{\rho} + \int_{\rho_0}^{\rho} \left(-\frac{\partial^2 \hat{A}_\phi}{\partial \rho'^2} - \frac{1}{\rho'} \frac{\partial \hat{A}_\phi}{\partial \rho'} + \frac{\hat{A}_\phi}{\rho'^2} \right) d\rho' \right]. \quad (57)$$

An integration by parts gives us:

$$\int_{\rho_0}^{\rho} \left(-\frac{\partial^2 \hat{A}_\phi}{\partial \rho'^2} + \frac{\hat{A}_\phi}{\rho'^2} \right) d\rho' = - \left[\frac{\partial \hat{A}_\phi}{\partial \rho'} + \frac{\hat{A}_\phi}{\rho'} \right]_{\rho_0}^{\rho} + \int_{\rho_0}^{\rho} \frac{1}{\rho'} \frac{\partial \hat{A}_\phi}{\partial \rho'} d\rho', \quad (58)$$

so Eq. (57) rewrites as

$$\frac{d^2g}{dz^2} + \frac{\omega^2}{c^2}g = \frac{i\omega}{c} \left[\frac{\partial \hat{A}_\phi}{\partial \rho}(\rho_0, z) + \frac{\hat{A}_\phi(\rho_0, z)}{\rho_0} \right]. \quad (59)$$

As is well known and easy to check, this very ordinary differential equation can be solved explicitly by the method of variation of constants. (The general solution g of (59) depends linearly on two arbitrary constants.) And by construction, any among the solutions g of (59) is such that, with that g , the function ψ in Eq. (48) obeys the scalar Helmholtz equation (49). This proves Proposition 2. \square

Corollary 2. *Let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. There exists a time-harmonic axisymmetric solution A_{1z} of the wave equation, with the same frequency, such that the associated GAZR1 solution $(\mathbf{E}_1, \mathbf{B}_1)$, deduced from A_{1z} by Eqs. (11)–(13), satisfy*

$$B_{1\phi} = B_\phi, \quad E_{1\rho} = E_\rho, \quad E_{1z} = E_z. \quad (60)$$

Proof. Let A_{1z} be a time-harmonic axisymmetric solution of the wave equation, and consider:

(i) the GAZR1 solution defined from A_{1z} by Eqs. (11)–(13) (thus with A_{1z} , $B_{1\phi}$, $E_{1\rho}$, E_{1z} , ... instead of A_z , B_ϕ , E_ρ , E_z , ... respectively).

(ii) the GAZR2 solution defined from the same A_{1z} by applying the duality (14) to the said GAZR1 solution:

$$E'_{1\phi} := cB_{1\phi}, \quad B'_{1\rho} := -\frac{1}{c}E_{1\rho}, \quad B'_{1z} := -\frac{1}{c}E_{1z}, \quad (61)$$

$$E'_{1\rho} := cB_{1\rho} = 0, \quad E'_{1z} := cB_{1z} = 0, \quad B'_{1\phi} := -\frac{1}{c}E_{1\phi} = 0. \quad (62)$$

On the other hand, consider the free Maxwell field $(\mathbf{E}', \mathbf{B}')$ deduced from the given time-harmonic axisymmetric solution (\mathbf{E}, \mathbf{B}) of the free Maxwell equations by the same duality relation:

$$\mathbf{E}' := c\mathbf{B}, \quad \mathbf{B}' := -\mathbf{E}/c. \quad (63)$$

Just in the same way as it was shown in Note 3, we know that a vector potential \mathbf{A}' such that $\mathbf{B}' = \text{rot}\mathbf{A}'$ does exist and can be chosen to be axisymmetric (and is indeed chosen so) — as are \mathbf{E} and \mathbf{B} , and hence \mathbf{E}' and \mathbf{B}' . Clearly, the sought-for relation (60) is equivalent to

$$E'_{1\phi} = E'_{\phi}, \quad B'_{1\rho} = B'_{\rho}, \quad B'_{1z} = B'_{z}. \quad (64)$$

Therefore, the existence of A_{1z} as in the statement of Corollary 2 is ensured by Proposition 2. \square

Accounting for Proposition 2 and for Corollary 2, and remembering the “complementarity” of the GAZR1 and GAZR2 solutions, we thus can answer positively to the question asked at the beginning of this section:

Theorem. *Let $(\mathbf{A}, \mathbf{E}, \mathbf{B})$ be any time-harmonic axisymmetric solution of the free Maxwell equations. There exist a unique GAZR1 solution $(\mathbf{E}_1, \mathbf{B}_1)$ and a unique GAZR2 solution $(\mathbf{E}'_2, \mathbf{B}'_2)$, both with the same frequency as has $(\mathbf{A}, \mathbf{E}, \mathbf{B})$, and whose sum gives just that solution:*

$$\mathbf{E} = \mathbf{E}_1 + \mathbf{E}'_2, \quad \mathbf{B} = \mathbf{B}_1 + \mathbf{B}'_2. \quad (65)$$

Remark. Thus, the uniqueness of the representation concerns the electric and magnetic fields. It of course does not concern the potentials $\mathbf{A}_1 = A_{1z}\mathbf{e}_z$ and $\mathbf{A}_2 = A_{2z}\mathbf{e}_z$ that generate respectively $(\mathbf{E}_1, \mathbf{B}_1)$ and $(\mathbf{E}'_2, \mathbf{B}'_2)$.

4 Conclusion

In Ref. [10], two classes of axisymmetric solutions of the free Maxwell equations have been introduced, and it has been shown that they allow one to

obtain in explicit form nonparaxial EM beams. It has been proved here that, by combining these two classes one can define a method that allows one to get *all* time-harmonic axisymmetric free Maxwell fields — and thus, by the appropriate summation on frequencies, all axisymmetric free Maxwell fields. This method results immediately from the Theorem just above, which theorem is not an obviously expected result, and whose proof is not immediate. We thus have now a constructive method to obtain all axisymmetric free Maxwell fields. Namely, considering a discrete frequency spectrum $(\omega_j)_{j=1,\dots,N_\omega}$ for simplicity: there are $2N_\omega$ functions, $k \mapsto S_j(k)$, and $k \mapsto S'_j(k)$ ($j = 1, \dots, N_\omega$), such that the components B_ϕ, E_ρ, E_z of the field are given by Eqs. (18), (19), (20) respectively — while the components E_ϕ, B_ρ, B_z are given by these same equations applied with the primed spectra S'_j , followed by the duality transformation (14).

In a forthcoming work, we apply this to model the interstellar radiation field in a disc galaxy as an (axisymmetric) exact solution of the free Maxwell equations. In this application, it is very important that, due to the present work, one knows that any axisymmetric free Maxwell field can be got in this way.

References

- [1] Garrett M.W., Axially symmetric systems for generating and measuring magnetic fields. Part I, J. Appl. Phys., 1951, 22, 1091–1107.
- [2] Boridy E., Magnetic fields generated by axially symmetric systems, J. Appl. Phys., 1989, 66, 5691–5700.
- [3] Wang J.C.L., Sulkanen M.E., Lovelace R.V.E., Self-collimated electromagnetic jets from magnetized accretion disks: the even-symmetry case, Astrophys. J., 1990, 355, 38–43.
- [4] Beck R., Wielebinski, R., Magnetic fields in the Milky Way and in galaxies, In: Oswalt T.D., Gilmore G. (Eds.), Planets, Stars and Stellar Systems Vol. 5, Springer, Dordrecht, 2013, 641–723
- [5] Nesterov A.V., NizieV V.G., Propagation features of beams with axially symmetric polarization, J. Opt. B: Quantum and Semiclassical Optics, 2001, 31, 215–219.

- [6] Borghi R., Ciattoni A., Santarsiero M., Exact axial electromagnetic field for vectorial Gaussian and flattened Gaussian boundary distributions, *J. Opt. Soc. Am. A*, 2002, 19, 1207–1211.
- [7] Durnin J., Exact solutions for nondiffracting beams. I. The scalar theory, *J. Opt. Soc. Am. A*, 1987, 4, 651–654.
- [8] Durnin J., Miceli J.J., Jr, Eberly J.H., Diffraction-free beams, *Phys. Rev. Lett.*, 1987, 58, 1499–1501.
- [9] Zamboni-Rached M., Recami E., Hernández-Figueroa H.E., Structure of non-diffracting waves and some interesting applications, In: Hernández-Figueroa H.E., Zamboni-Rached M., Recami E. (Eds.), *Localized Waves*, John Wiley & Sons, Hoboken, 2008, 43–77
- [10] Garay-Avedaño R.L., Zamboni-Rached M., Exact analytic solutions of Maxwell’s equations describing propagating nonparaxial electromagnetic beams, *Appl. Opt.*, 2014, 53, 4524–4531.
- [11] Landau L.D., Lifshitz E.M., *The classical theory of fields*, 3rd English edition, Pergamon, Oxford, 1971, 108–109
- [12] Jackson J.D., *Classical electrodynamics*, 3rd edition, John Wiley & Sons, Hoboken, 1998, 239–240