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1 Fault plane modelling of the 2003 August 14 Lefkada Island (Greece)
2 earthquake based on the analysis of ENVISAT SAR interferograms

3
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24 **Abstract**
25

26 On 2003 August 14, a $M_w = 6.2$ earthquake occurred offshore the Lefkada Island in the
27 eastern Ionian Sea, one of the most seismically active areas in Europe. The earthquake caused
28 extended damages in the island, and a number of ground failures, especially along the north-
29 western coast. Seven ascending ENVISAT/ASAR images are used to process six co-seismic
30 interferograms. The ROI-PAC package is used for interferogram generation with the SRTM
31 DEM applied in a two-pass method. The formation of the co-seismic pairs is limited due to
32 the existence of one pre-seismic image only. Dense vegetation is covering the island, which is
33 an obstacle in getting good coherence, since C-band images are used. Nevertheless, ground
34 deformation, of more than 56 mm (two fringes) in the line of sight of the satellite, is detected
35 in all six co-seismic interferograms. By inversion of the data from the observed fringes, a best
36 fitting model of the activated fault is calculated assuming a dislocation in an elastic half
37 space. The inferred fault is a pure dextral strike-slip fault, dipping $59 \pm 5^\circ$ eastward, 16 ± 2
38 km long and 10 ± 2 km wide. It is located north of the fault of the $M_w = 6.5$ 2015 November
39 17 earthquake, and a 10-15 km gap remains between the two faults. The 2003 fault does not
40 reach the surface and its upper edge is at a depth of 3.5 ± 1 km. No evidence is found of slip
41 south of the Lefkada Island as suggested by some seismological studies.
42

43 **Keywords:** Earthquake dynamics, Seismicity and tectonics, Radar interferometry, Lefkada
44
45

46 **1. Introduction**
47

48 Greece is the country with the highest seismicity in Europe. The main crustal deformations
49 occur along the Hellenic Arc, as a result of the subduction of the African Nubian plate

50 beneath the Aegean plate (Le Pichon and Angelier, 1979; Ganas and Parsons, 2009a; Pérouse
51 et al., 2012). The Central Ionian region (Fig. 1) is one of the most active areas of shallow
52 seismicity in Greece (Benetatos et al., 2007) with an estimated full seismic coupling
53 according to Chousianitis et al. (2015). The Central Ionian Islands, Lefkada, Cephalonia,
54 Zakynthos, are located on the Aegean Plate between the Apulia – Eurasia continental collision
55 to the north and oceanic subduction to the south, with intensive surface deformations due to
56 ENE-WSW shortening (Ganas et al., 2013) with earthquake than can reach or exceed
57 magnitude 7 like in 1953 or 1983 (Papazachos and Papazachou, 1997). The plate boundary in
58 the Central Ionian Sea is defined by a major, right-lateral, strike-slip fault, the Cephalonia
59 Transform Fault (CTF) which is located at the western termination of the Hellenic subduction
60 zone, and is characterised by a bathymetric trough exceeding 3000 m depth. Strike-slip
61 faulting prevails in a large area to the east of the CTF reaching the western boundaries of the
62 Gulf of Corinth (Kiritzi et al., 2008; Ganas et al., 2009b; Serpetsidaki et al., 2014; Briole et
63 al., 2015; Chousianitis et al., 2015).

64
65 In terms of geology, the Ionian Islands are built with sedimentary rocks, mainly carbonates,
66 belonging to the so-called external Hellenides (Jacobshagen 1979; Underhill 1989). In
67 particular, the boundary between the two different geological zones, called Ionian and Paxos,
68 runs through this region and outcrops onshore Lefkada, forming the noteworthy Ionian Thrust
69 which is assumed to be the westernmost limit of External Hellenides chain (Lagios et al.,
70 2007). The main part of the Lefkada Island is made of a carbonate sequence of the Ionian
71 zone, while the SW part of the island consists of limestone of Paxos geological zone
72 (Bornovas 1964; Pavlides et al., 2004).

73
74 This tectonic regime predetermines the high seismic activity in the area of Lefkada Island.
75 The island was affected indeed by several significant events in the past (Table 1). The two
76 most recent strong earthquakes are the $M_w = 6.2$ 2003 August 14 earthquake analyzed in this
77 paper and the recent $M_w = 6.5$ 2015 November 17 earthquake (Ganas et al., 2016;
78 Chousianitis et al., 2016, Sokos et al., 2016). Both earthquakes occurred offshore the western
79 coast of Lefkada Island, causing severe damages around the whole island.

80
81

82 **2. Seismicity in the area and Lefkada fault location**

83
84 According to the historical documentation of the last centuries several strong earthquakes
85 occurred since the 17th century in the area of Lefkada Island. During the instrumental era five
86 events with magnitude higher than 6 were recorded: one in 1914, two in 1948 (Fig. 2, Table
87 1), the 2003 event studied here, and the recent earthquake of 2015.

88
89 Taking into account 23 microseismic events recorded with portable seismographs by Hatzfeld
90 et al. (1995) during a 7-week experiment in the summer of 1989, available macroseismic data,
91 as well as the bathymetry change to a NNE direction northern of Cephalonia Island, Louvari
92 et al. (1999) identify an active seismic segment of the CTF offshore west of the coast of
93 Lefkada. The authors suggest that this northern segment strikes in a NNE-SSW direction, dips
94 to ESE, has strike-slip motion combined with a small thrust component (Table 2), and has a
95 length of ~40 km starting from the northern part of the Cephalonia Island. They interpret the
96 thrust component of the $M = 5.8$ 1973 November 4 earthquake (Baker et al., 1997) as an
97 indication of the junction of the Albanian thrust belt with the Lefkada segment at the northern
98 end of the CTF. Pavlides et al. (2004) assumes a division of Lefkada segment into two smaller
99 segments separated by Sesoula Islet, west of Lefkada, striking at different directions, NNE-

100 SSW and NE-SW respectively for northern and southern sub-segment and behaving
101 independently.

102
103

104 **3. The 2003 August 14 earthquake**

105

106 On 2003 August 14 at 05:15 GMT a $M_w = 6.2$ earthquake occurred offshore the western coast
107 of Lefkada Island causing considerable damages (Fig. 3) in several settlements, all harbours
108 and the road network, as well as landslides, rock falls, soil liquefaction, subsidence, ground
109 cracks, especially along the north-western part of the island (Katopodi and Iosifidou, 2004;
110 Papathanassiou et al., 2005; Pavlides et al., 2004).

111

112 A maximum intensity of VIII in the town of Lefkada was observed, while elsewhere in the
113 island the intensity ranged from V to VII (Papadopoulos et al., 2003; Papathanassiou et al.,
114 2005). The character and location of the damages are similar to those related to the previous
115 strong earthquakes of similar magnitude in the area. Compared to the past events, the intensity
116 decreased from X-XI, possibly due to the strict construction rules established after the last
117 destructive earthquake of 1948. Nevertheless, ~120 houses were destroyed almost half of the
118 buildings of the island were damaged in different degree (Pavlides et al., 2004).

119

120 Up to 2003 August 31 more than 470 aftershocks ($M > 1.5$), concentrated in two clusters,
121 were recorded by the Patras Seismological Laboratory (Fig. 4), the strongest of which ($M_s =$
122 5.25) occurred eleven hours after the main shock. Papadopoulos et al. (2003), Karakostas et
123 al. (2004) and Papadimitriou et al. (2006) consider the southern cluster, located close to the
124 north-western coast of Cephalonia Island, rather representing a triggered seismicity on the
125 Cephalonia segment, due to stress increase from the main shock, than typical aftershock
126 activity, and it has to be noted that the same area was also activated after the 2014 Cephalonia
127 sequence. On the contrary, Benetatos et al. (2005; 2007), Zahradnik et al. (2005), Saltogianni
128 and Stiros (2015), Sokos et al. (2016) suggest a complicated fault geometry of Lefkada
129 segment, with two main sub-events and segmentation of the rupture in a double cluster.
130 According to this latter group of authors, the multiple source character of the main shock is
131 represented by two main sub-events along a $N12^\circ E$ to $N20^\circ E$ (depending on the authors) line
132 that ruptured Lefkada segment from north to south (Table 2, Fig. 4). The two sub-events are
133 separated in time and space at 14 s and ~40 km, respectively, have strike-slip focal
134 mechanisms and induced concentration of aftershocks in two clusters with mainly strike-slip
135 character along the NE-SW trending fault segment (Fig. 4). Thrust and reverse faulting
136 mechanisms also detected, mostly in the north and mainland parts of the island, indicate
137 possible activation of faults in adjacent areas, as the Ionian thrust zone.

138

139 The present investigation is concentrated on the deformations in the area of the main shock
140 (or first sub-event) and the following sequence of aftershocks on the Lefkada Island (Fig. 4).

141

142

143 **4. InSAR processing and analysis**

144

145 In the catalogue of the European Space Agency (ESA), for the period between January 2003
146 and January 2008, we found 13 ascending and 20 descending ENVISAT/ASAR (C-band)
147 passes from track 458 and track 322 respectively, over the area of interest (Fig. 5). In both
148 cases of satellite viewing, only one of these scenes is pre-seismic: 2003 March 21 for the

149 ascending and 2003 June 25 for descending (Table 5). This limits the generation of co-seismic
150 pairs with only one master image.

151
152 In spite of this limitation, the choice of the scenes used in our investigation is made in such a
153 way that the image pairs formed have perpendicular baselines less than 200 m with respect to
154 the pre-seismic image. This requirement is valid for six ascending images, which form 6
155 ascending co-seismic and 15 ascending post-seismic interferometric pairs (Fig. 6, Table 4 and
156 Table 5), and 8 descending scenes, forming 8 co-seismic and 23 post-seismic interferograms
157 (Table 4 and Table 5).

158
159 Standard processing flow of differential (2-pass) method was followed to generate the
160 interferograms, using the Jet Propulsion Laboratory/California Institute of Technology open
161 source package ROI_PAC 3.0 (Rosen et al., 2004). The raw images (level 0) were provided
162 by ESA as well as the Doris precise orbit files, used for removing the orbital effects. To
163 eliminate the topographic effect, a simulated interferogram is generated from the SRTM
164 digital elevation model (DEM) version 4 (Jarvis et al., 2008). This DEM has 90 m resolution
165 and a nominal vertical error of less than 16 m. Nevertheless, thorough studies conducted by
166 Mouratidis et al. (2010) indicate that the actual absolute and relative errors (standard
167 deviation) are well below 10 m. Multilooking (4 looks in range and 20 looks in azimuth) was
168 applied to minimize the loss of coherence and reduce the speckle noise in the produced
169 interferogram.

170
171 With this operation, the original pixel size of 7.8 m in range and 4.6 m in azimuth was
172 averaged to 31.2 m and 91.4 m respectively. The Goldstein-Werner filter (Goldstein and
173 Werner, 1998), optional in ROI_PAC package, was applied in the final stage of processing to
174 improve the quality of the interferogram. The results are geo-referenced to the coordinate
175 system of the used DEM (GCS/WGS84). Tacking into account the nominal accuracy of the
176 DEM ($h_{DEM} = 16$ m) and the altitude of ambiguity (h_a) for each interferogram, the maximum
177 expected error due to the topography is calculated (Table 4).

178
179 A main problem in our interferogram processing is the fact that the ENVISAT/ASAR images
180 are C-band, so the presence of vegetation has a strong influence on the coherence. The whole
181 island is indeed dense vegetated mainly with olive groves and pine forests, especially to the
182 east and south. This is probably the reason for coherence loss in all produced interferograms
183 (see the Appendixes). To illustrate better the regions where the received signal is more
184 reliable, a map of coherence is produced (Fig. 7) on a basis of all processed ascending
185 interferograms. The value of the coherence for every pixel from the map is calculated first as
186 a mean between the values of its neighbouring in the respective image and second the results
187 of the corresponding pixels from all interferogram are averaged. In this way the reoccurrence
188 in time and space of stable backscattered signal for each pixel is characterized. The areas with
189 better scattering capabilities on the produced map correspond to the not overgrown Stavrota
190 mountain (in the centre of the island), Kalavros to the north, big settlements like town of
191 Lefkada and Preveza, and plain areas around Voulkaria Lake in mainland north-eastern from
192 the island.

193
194 Examining the ascending data set, the interferogram which comprises the shortest time span
195 (174 days) and has only 35 m perpendicular baseline, namely 21.03.2003-12.09.2003
196 (interferogram No 1 – *ifg1* in Table 4), is expected to have the best quality. Lagios et al.
197 (2007) analyse exactly this interferometric pair and describe 1 fringe in the northernmost part
198 of the island and 2 fringes in the western part. However, in the current investigation a wider

199 set of data is studied and it is found that the interferogram *ifg5* (21.03.2003-5.11.2004), which
200 covers a longer period (592 days) and has a larger perpendicular baseline (88 m), shows much
201 better quality (Fig. 9 left) than the interferogram *ifg1* (APPENDIX A). The fringes of the
202 21.03.2003-5.11.2004 pair are more sharpened, while the interferogram pattern of pair *ifg1* is
203 strongly deformed. A possible explanation about this difference could be the lower percentage
204 of humidity and cloudiness during the acquisition of the slave image in the pair *ifg5* (Table 5),
205 than the same weather characteristics in the day of the slave image of pair *ifg1*. Although, the
206 local disturbances and the local temporal differences in humidity level have more important
207 role in the phase de-correlation, even the homogeneous change in the atmosphere (pressure,
208 humidity, and temperature) can be revealed by a contrasted relief, which modulates the
209 thickness of the troposphere that the signal must cross (Massonnet and Feigl, 1998) since the
210 topography induces tropospheric delays. In the case of Lefkada Island the elevation varies
211 from sea level to 1160 m in the central part of the island, which is roughly the place where the
212 strongest distortions occur. From the data of this study and other recent data over Lefkada
213 Island, we estimate that tropospheric disturbances could be of ~ 1 C-band fringe (28 mm of
214 delay) over the island.

215
216 No post-seismic deformation is observed in our interferograms (APPENDIX B) but we cannot
217 exclude the presence of a small amount of such deformation as it has been observed elsewhere
218 for strike-slip events of similar magnitude (e.g. Bürgmann et al., 1997).

219
220 The ascending co-seismic interferogram with the best quality (*ifg5*), covering the time period
221 between 2003 March 21 and 2004 November 5 (Fig. 9 left), shows at least 2 fringes in the
222 western part of the island. Since one fringe is equal to half the wavelength, which for
223 ENVISAT C-band is 28 mm, then these two fringes are equal to 56 mm of deformation of the
224 crust in the satellite line of sight (LoS). A verification in Table 4 shows that a very small part
225 of detected fringes could be attributed to errors due to the topography – 4 mm, which is only
226 1/7 of the fringe.

227
228 In the case of descending co-seismic set of data, the deformations after the occurrence of the
229 strong earthquake are detected with significantly lower quality, mainly because of the look
230 direction of the satellite towards the highest mountains, so that the deformation pattern
231 appeared distorted. The descending co-seismic interferogram with the best quality is the one
232 with the smallest temporal separation (*ifg22* – Fig. 8) and has a perpendicular baseline of 118
233 m (Table 4).

234 235 236 **5. Modelling**

237
238 The modelling and interpretation of the interferograms is made in three successive phases:
239 first we model the best ascending wrapped interferogram only, then we analyze the stacks of
240 unwrapped ascending and descending interferograms, and finally we perform a joint inversion
241 of ascending and descending data.

242
243 The fringes in the best ascending interferogram are sampled to 34 points which we found
244 sufficient to account for the level of quality of the data (using more sample points leads to the
245 same model). One value of slant range displacement is sampled every ~ 1 km (Fig. 9 left -
246 inset). The descending interferograms are not used in this first analysis, due to their lower
247 quality. The Harvard CMT solution is used as initial parameters in inversion procedure (Table
248 2) fixing the strike angle to 18° and dip angle to 59° . The applied inversion algorithm was

249 developed by Briole et al. (1986), it is based on the least-square minimization approach
250 proposed by Tarantola and Valette (1982). The Okada (1992) formulation for rectangular-
251 plane source in uniform elastic half-space is adopted for the dislocation source. The
252 parameters of the best fitting model (Fig. 9 right) are listed in Table 3. The epicentre is
253 located at the SW end of our modelled fault plane. If this is so the geodetic epicentre is offset
254 by about 6 km to the NE of the NOA epicentre.
255

256 The fringes obtained in the synthetic interferogram are in good agreement with the observed
257 ones, evident also by the residuals between the phase values of the fringe points from the
258 model and the same points in the real interferogram. The value of the mean is -0.9 mm and
259 the root-mean-square (r.m.s.) residual is 2.3 mm i.e. only 1/12 part of a fringe. The
260 concordance of the proposed modelled fringe points with the observed ones is manifested by
261 the distribution of the residual (Fig. 10 left), which is close to the normal one. Additionally,
262 since the number of the points is relatively small, a normal probability plot is presented (Fig.
263 10 right). This plot shows some outliers in both ends of the data set but still confirms the
264 normal distribution of the residuals.
265

266 In order to confirm the reliability of the proposed model of the fault, a series of inversions are
267 conducted with varying dip and length of the fault, taking into account the values proposed by
268 previous studies (Fig. 11). In both cases, while the different values for the dip and length are
269 tested, all the other parameters are left fixed. This investigation of the confidence of these
270 parameters shows that the best fitting model would be achieved at dip of 60° and length of the
271 faulting segment of 16 km. However, the r.m.s. difference for a model with dip of 60° and
272 59°, like the value proposed here, on the basis of Harvard CMT solution, is only 0.4 mm (1/70
273 of the fringe), which is not significant.
274

275 The sensitivity of the model to the dip alteration is expressed by “pressing” of the fringe’s
276 shape. In the case investigated here, the mean part of the outer fringe is not clearly outlined.
277 To increase the quality of the interferometric pattern and to improve the fringe shape, an
278 interferogram averaged from all the six co-seismic ascending interferograms is produced. As a
279 first step all of them are unwrapped using the SNAPHU software (Chen and Zebker, 2002)
280 (APPENDIX A). The stacking of all co-seismic ascending interferograms and the backward
281 wrapping (Fig. 12d) of the phases increased in some aspect the quality of the single wrapped
282 picture (Fig. 12a). The resulting field of deformation of the single interferogram (Fig. 12b)
283 shows maximum displacement toward the satellite in the range of 60 mm. The stacking of the
284 available interferograms (Fig. 12c) narrows the area of deformations and decreases its value
285 to ~50 mm.
286

287 When the re-wrapped interferogram is superimposed over the topography (Fig. 12e), high
288 dependency of the local relief is revealed. In spite of the more homogeneous field of
289 deformation which is represented by the unwrapped stacked interferogram compared to the
290 interferogram with the best quality, the averaged unwrapped pattern also combines all defects
291 of the interferograms used in the stacking. This is illustrated by cross-sections through the
292 unwrapped interferograms in Fig. 13. The A’-A” profile is traced parallel to the strike of the
293 fault and is presented in the upper part of the figure, while the B’-B” profile passes through
294 the island and represents the area with the most pronounced crust deformation (bottom of the
295 figure). The corresponding local relief is plotted to the bottom of the profiles.
296

297 It is seen how a processing error during the unwrapping procedure (APPENDIX A) caused
298 distortion of the field of deformation of *igf4* showed in the profile by dotted line (UWRI4 –

299 unwrapped real interferogram *ifg4*). If this interferogram is excluded of the averaging, the
300 middle part of the interferogram becomes more uniform (red line, UWSRI – unwrapped
301 stacked real interferogram), but the level of deformation in the outlying areas vary
302 significantly in the A'A" profile. The abrupt change of the deformation line appeared in a
303 place of a steep topographic depression at the boundary between the Ionian and Paxos
304 geological zones, which could influence strongly the unwrapping process. Nevertheless, the
305 averaged interferogram shows high level of concordance to the modelled one (MI in Fig. 13),
306 even better than the single interferogram used for the modelling (UWRI5). The received field
307 of crustal deformations, excluding the wrong *ifg4*, shows a maximum in the range of ~60 mm
308 towards the satellite (Fig. 12f). The shift between the real and the modelled data in the right
309 part of the A'A" profile is again a consequence of the unwrapping defect in the area of the
310 Ionian Thrust like in the case of *ifg4* but in the opposite direction. If the real data in this part
311 of the interferograms are shifted with the amount of exactly one fringe downward they will
312 correspond to the model. On the other hand, the left part of the B'B" profile shows a big
313 difference between the data and the model. Since the images used for producing these
314 interferograms are ascending, this area follows in the shadow of the high mountain in the
315 centre of the island looked from the satellite. Therefore, the signal in this area is in fact an
316 incoherent noise and the unwrapping procedure performs false information about the
317 deformation in this area. None conclusions can be made in this situation about the reliability
318 of the model in this area.

319
320 In order to find another confirmation about the agreement between the fields of deformation
321 revealed by the interferograms and predicted by the model, a comparison on a pixel level is
322 performed. The examination of the residuals between the real and the synthetic deformation
323 interferograms (Fig. 14) displays a high degree of correspondence between both patterns,
324 which is confirmed also by the histogram of the residuals (Fig. 15). It manifests very close to
325 the normal distribution, shifted by the mean value of 3.4 mm. The r.m.s. for the entire
326 examined pattern is 10.5 mm.

327
328 The stacking procedure is applied also to the descending set of data, so that the field of
329 deformation from the other point of satellite view could be gained. The stacked re-wrapped
330 interferogram shows one clear fringe (Fig. 16), but in the rest of the interferogram still high
331 level of noise presents (APPENDIX C).

332
333 Nevertheless, the averaged descending unwrapped interferogram will be used, together with
334 the ascending result, in a theoretical determination of the horizontal and vertical components
335 of the field of deformations. To define two of the three components of the LoS displacement,
336 an assumption for the third one should be adopted. Here, the northern component is equalised
337 for both satellite viewings, thus the displacement in up and east direction (Fig. 17) is extracted
338 using the parameters of the satellite geometry (Hanssen, 2001). The resulting pattern shows a
339 maximum vertical displacement of the Earth's crust in the range of ~50 mm in the western-
340 central part of the island and eastward displacements in the western coast in the maximum
341 range of ~60 mm.

342
343 Finally we make a model by jointly inverting ascending and descending data. The
344 combination of ascending and descending observations is potentially powerful to constrain
345 some of the fault parameters, in particular depth, location of the fault in the east-west axis,
346 and dip-slip component of motion. We use the two best interferograms (ascending *ifg5* and
347 descending *ifg22*, see Table 4), with the 34 picked values on two fringes visible on the
348 ascending interferogram (those already used before) and 20 picked values on the only clear

349 fringe of the descending one. The ratio 20/34 gives less weight to the descending data in the
350 inversion. This is consistent with the fact that, given its lower coherence, we are less
351 confident in this data than in the ascending one. Starting with the parameters of our preferred
352 model (IL16, Table 3) we first leave free all fault parameters in the inversion. Only two of the
353 fault parameters show significant changes: the depth of the upper edge of the fault that is
354 estimated deeper, at 4.5 km instead of 2.5 km, and the east coordinate of the fault that moves
355 westward in a correlated manner with the fault depth. The other parameters are not
356 significantly affected by the addition of the descending data. Another benefit of the joint
357 inversion of ascending and descending data is the fact that, because of the existence of two
358 opposite LoS, it is possible to seek for an optimum in the common mode offset to apply to the
359 picked value of the fringes. We tune the picked value so as to minimize the post-fit residuals
360 at a minimum value of 9 mm, and Table 7 and Table 8 contain the location and observed and
361 modelled value of our picked fringe points. In all inversions the component of dip-slip on the
362 fault plane remains close to zero, ranging from -18 mm (reverse) when the upper fault depth
363 was assigned to be 5 km to +21 mm (normal) when the upper fault depth was assigned to be 4
364 km. Table 9 contains the value of the two preferred models. In summary, the two main
365 constraints provided by the combination of the ascending and descending data are the
366 following: no significant component of dip-slip on the fault, an upper limit of the fault
367 possible deeper (4.5 km instead of 2.5 km with ascending only).

368
369

370 **6. Discussion and conclusions**

371

372 We examine fault size, geometry and its relation to earthquake magnitude and how this relates
373 to the historical seismicity record (Table 1). Our geodetic solution (Fig. 9) indicates that the
374 rupture length is ~ 16 km, a value close to that predicted by the empirical relationship of Wells
375 and Coppersmith (1994) (see Table 6), and much less than those predicted by the equations of
376 Papazachos and Papazachou (1997) and Konstantinou et al. (2005). The 2003 Lefkada event
377 occurred along one fault segment plane dipping to the east as suggested by the linear
378 inversion of the InSAR fringes. The earthquake was followed by numerous aftershocks on
379 neighbouring fault planes that are characterised as capable to produce moderate events
380 (Karakostas and Papadimitriou, 2010). The geodetically determined fault size is about half the
381 length of the Lefkada plate boundary segment defined by seismological data by Louvari et al.
382 (1999). In addition, Papadopoulos et al. (2003) have suggested that offshore earthquakes near
383 Lefkada follow the characteristic earthquake model, i.e. slip distribution and earthquake
384 magnitude are roughly constant. Our results indicate that, the 2003 earthquake ruptured only
385 the northern portion of the 50-km Lefkada segment of CTF, which is about 16 km long.
386 Assuming that the Lefkada segment of CTF comprises two seismic segments then, this
387 segmentation model predicts rupture lengths of the order of 15-20 km and maximum
388 earthquake magnitudes in the range 6.2-6.5. This suggestion has important consequences for
389 the seismic hazard of the area, as the earthquakes expected offshore Lefkada may reach $M_w =$
390 6.5 like the 2015 November 17 southern Lefkada earthquake (Ganas et al., 2016). For both
391 the 2003 earthquake (this study; Fig. 9 right) and the 2015 one (Ganas et al. 2016; Fig. 4),
392 InSAR constrains well the spatial extend of the rupture along the axis of the CTF. We infer
393 from the two studies that a 10-15 gap remains between the 2003 and the 2015 fault planes. A
394 gap of such length corresponds to a potential earthquake of $M_w \sim 5.7-6$. Some authors
395 (Saltogianni and Stiros, 2015; Sokos et al., 2016) have suggested that the 2003 event also
396 broke a fault segment south of the 2015 fault, in front of the northern Cephalonia, in a very
397 unusual manner. We found no evidence of such southern segment in our 2003 interferograms.
398 Moreover, this hypothetical jump of the 2003 rupture over the 2015 fault should also jump

399 over the 10-15 gap located between the 2003 and 2015 faults. This corresponds to a distance
400 of ~40 km between the second hypothetical sub-event and the main shock of 2003, itself well
401 modelled from the InSAR data and with a good agreement between geodetic and
402 seismological moments.

403
404 The fault model we propose has a geodetic moment $M_0 = 3.6 \cdot 10^{18}$ N m which corresponds to
405 $M_w = 6.3$, greater than the seismic moment reported by the Harvard CMT solution which is
406 $M_0 = 2.98 \cdot 10^{18}$ N m or $M_w = 6.2$. The difference might be due to the influence of the two
407 larger shallow aftershocks, that occurred the same day as the main shock, with $M_w = 5.1$ and
408 $M_w = 5.5$ respectively.

409
410 The interferometric pattern, produced by ENVISAR/ASAR images, defines field of
411 deformations caused by the 2003 Lefkada earthquake, in the range of 60-70 mm in the line of
412 sight of the satellite. An attempt to determine the horizontal and vertical component of
413 displacement is performed. An uplift of about 50 mm in the central area of deformation is
414 calculated and deformations with eastward direction in the range of 60 mm in the western
415 coast of the island are estimated.

416
417 The 2003 fault does not reach the surface. Its upper tip is located at a depth comprised
418 between 2.5 km (using ascending data only) and 4.5 km (using a combination of ascending
419 and descending). This is much deeper than the depth of the upper tip of the 2015 fault,
420 estimated to be ~1 km around the northern end of the fault according to Ganas et al. (2016).
421 The mode and timing of the deformation between that depth and the surface requires further
422 analysis. Is the uppermost part of the crust deforming continuously during the seismic cycle in
423 that segment or is there a possibility of very shallow and strong earthquakes? And why is the
424 2015 rupture shallower? In addition to SAR interferometry (which is insufficient for sampling
425 long time periods in areas of poor coherence), dense and permanent GPS will be important to
426 decipher the spatio-temporal behaviour of the uppermost crust along the Cephalonia
427 Transform Fault. The combination of ascending and descending InSAR observations indicates
428 that there is no significant component of dip-slip on the fault of this 2003 earthquake.

429
430

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432
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434 Agency in the framework of the Category-1 project ID 4484: "Investigation of the south-
435 eastern part of the Sofia complex graben (Bulgaria) by InSAR and GPS methods with
436 comparative analysis with the Lefkada Island (Greece)". The study was realised in the
437 framework of the bilateral project RILA (3/29, 2007-2008) between the Ministry of
438 Education, Youth and Science of Bulgaria and Ministry of National Education, France. The
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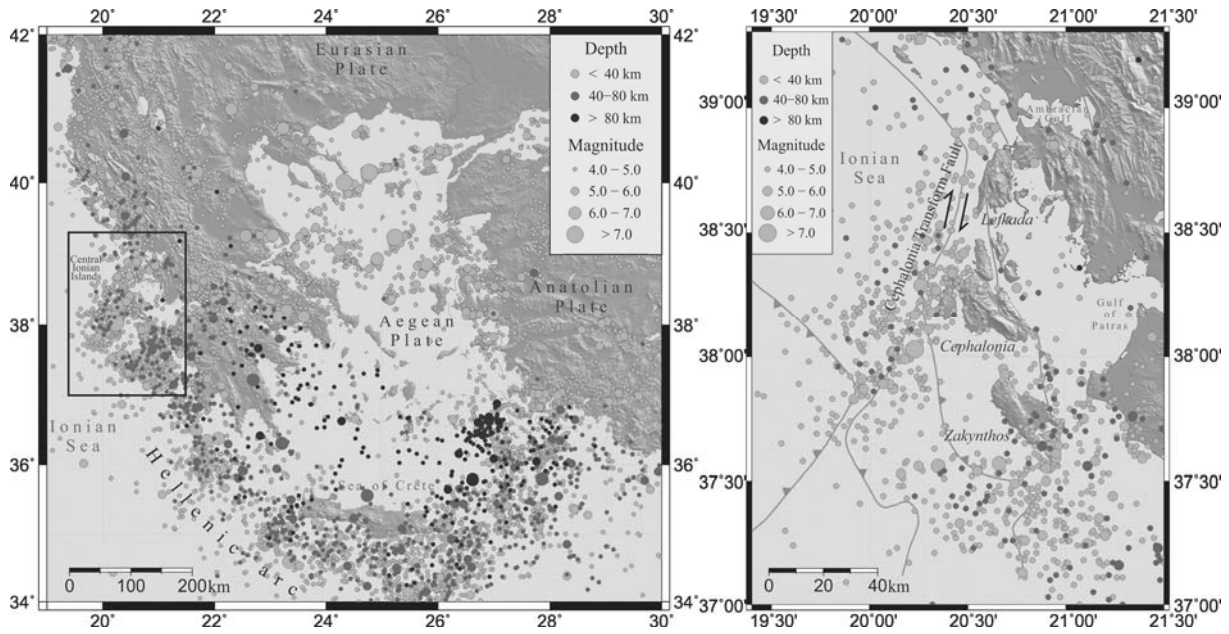
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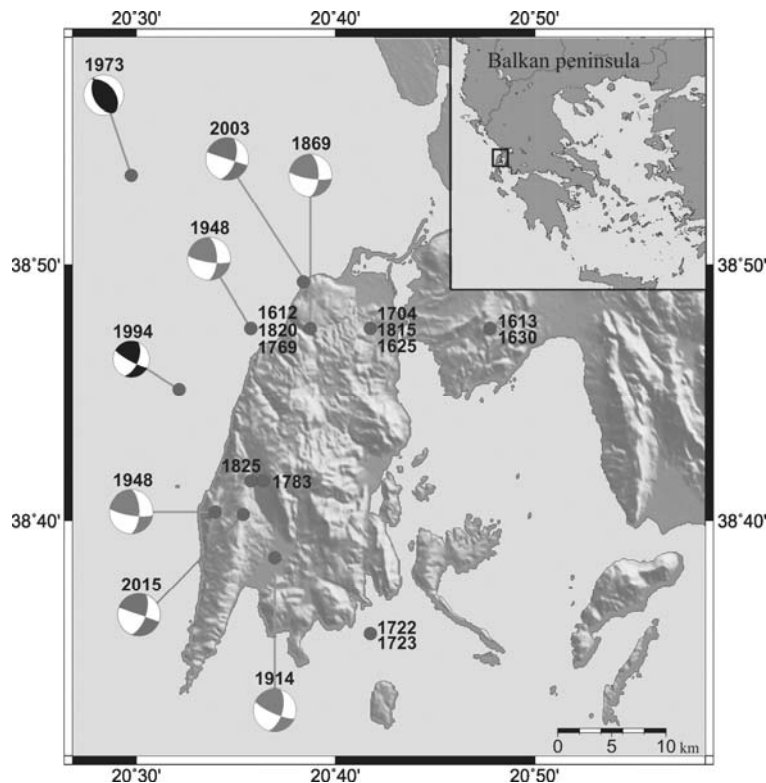
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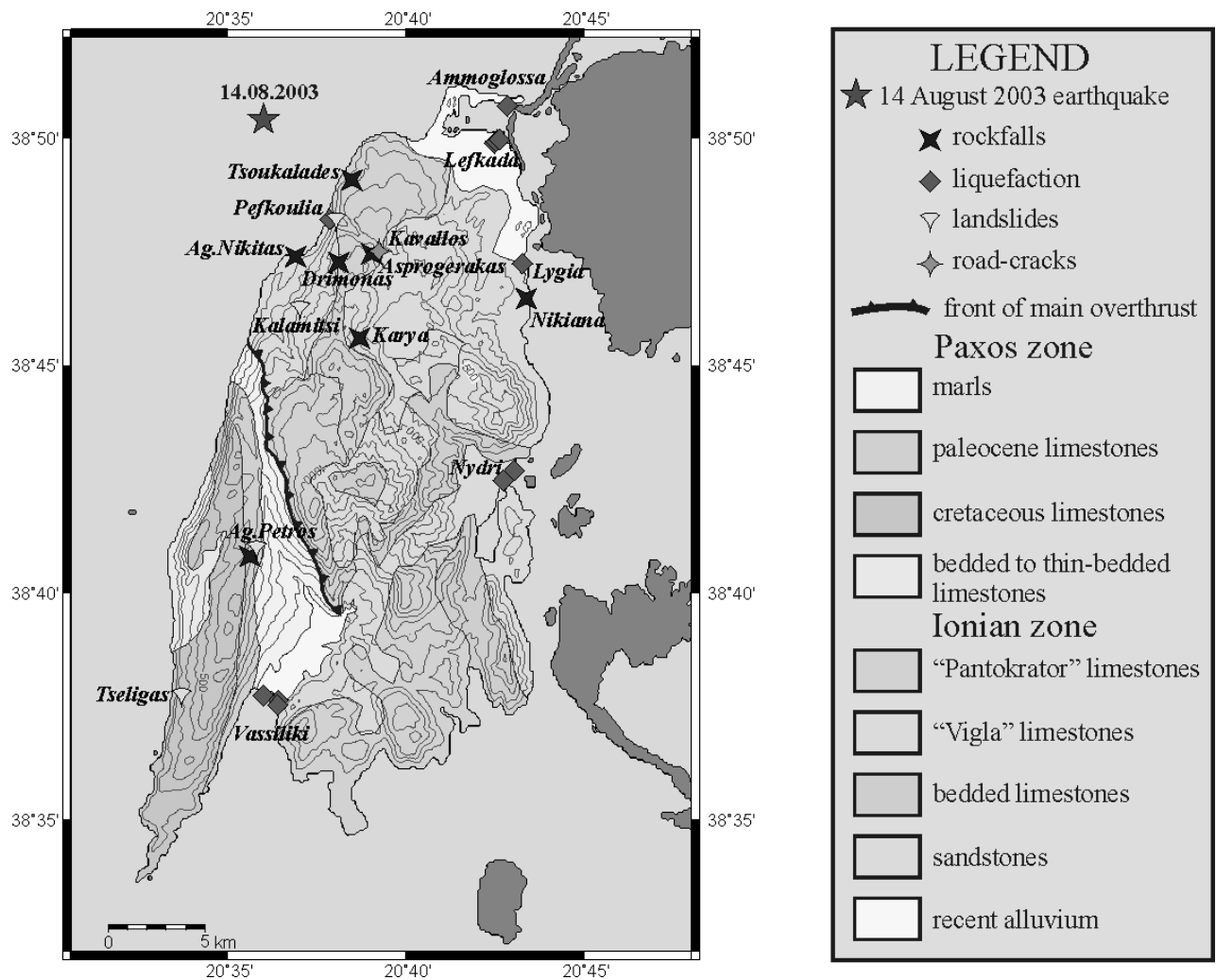
601 Fig. 1. Left: Seismicity ($M > 4$) in the Eastern Mediterranean since 1973 according to USGS.
 602 Rectangle – the region of Central Ionian Islands. Right: Seismicity in the area of the Central
 603 Ionian Islands. Fault geometry is adapted from Barrier et al. (2004).

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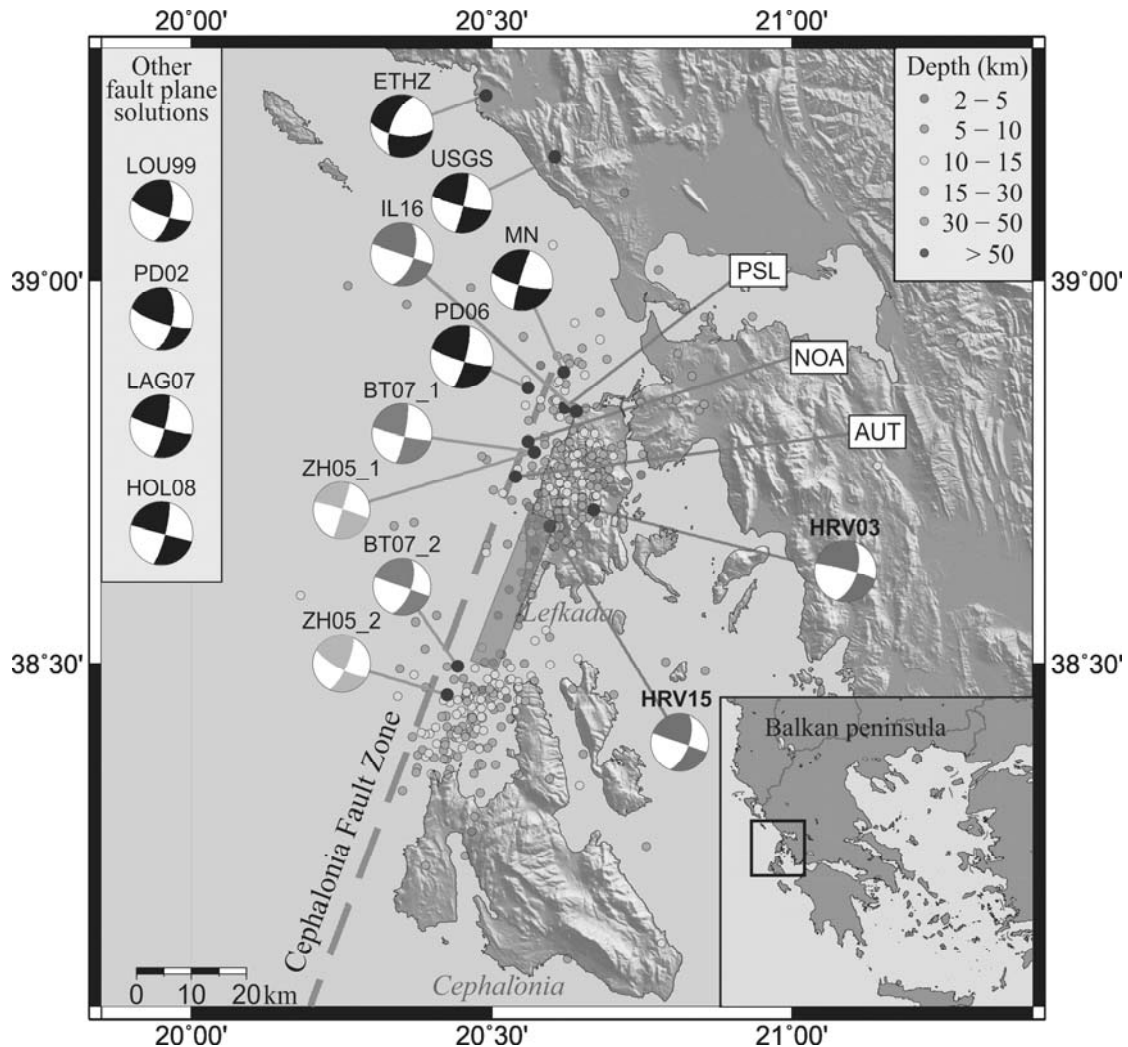
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606 Fig. 2. Historical seismicity in the area of Lefkada Island according to Table 1. In inset the
 607 location of the Lefkada Island in Eastern Mediterranean. In grey, the mechanisms of strong
 608 historical events are marked, while in black – the mechanisms of most recent earthquakes
 609 even though they were of smaller magnitudes. The mechanisms of the 2003 and 2015
 610 earthquakes are plotted in red.



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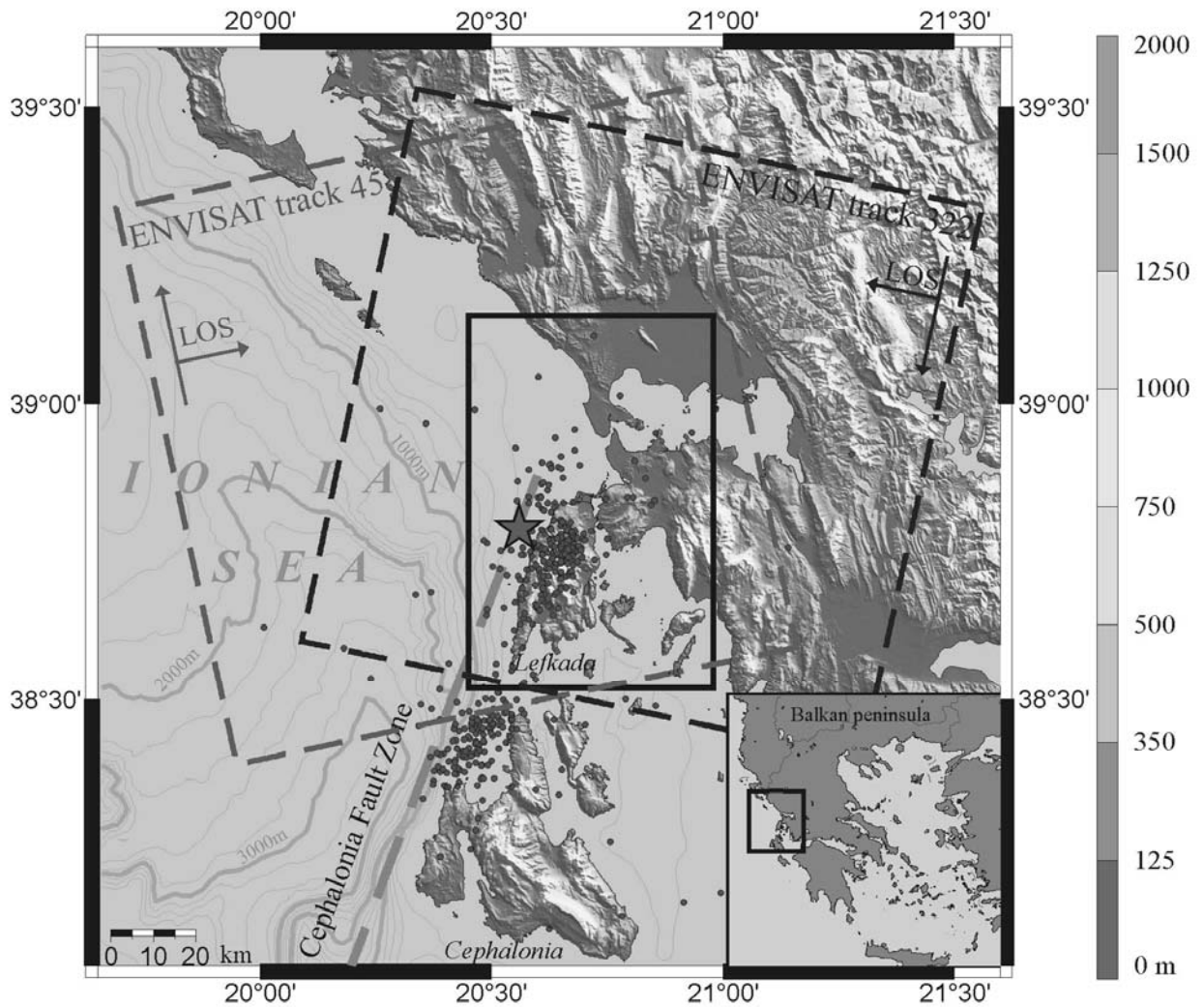
613 Fig. 3. Map of ground failures caused by the 14th August 2003 earthquake (after Pavlides et
614 al., 2004; Papathanassiou and Pavlides, 2007). On the background simplified geological map
615 of the Lefkada Island (Geological map of Greece 1963).



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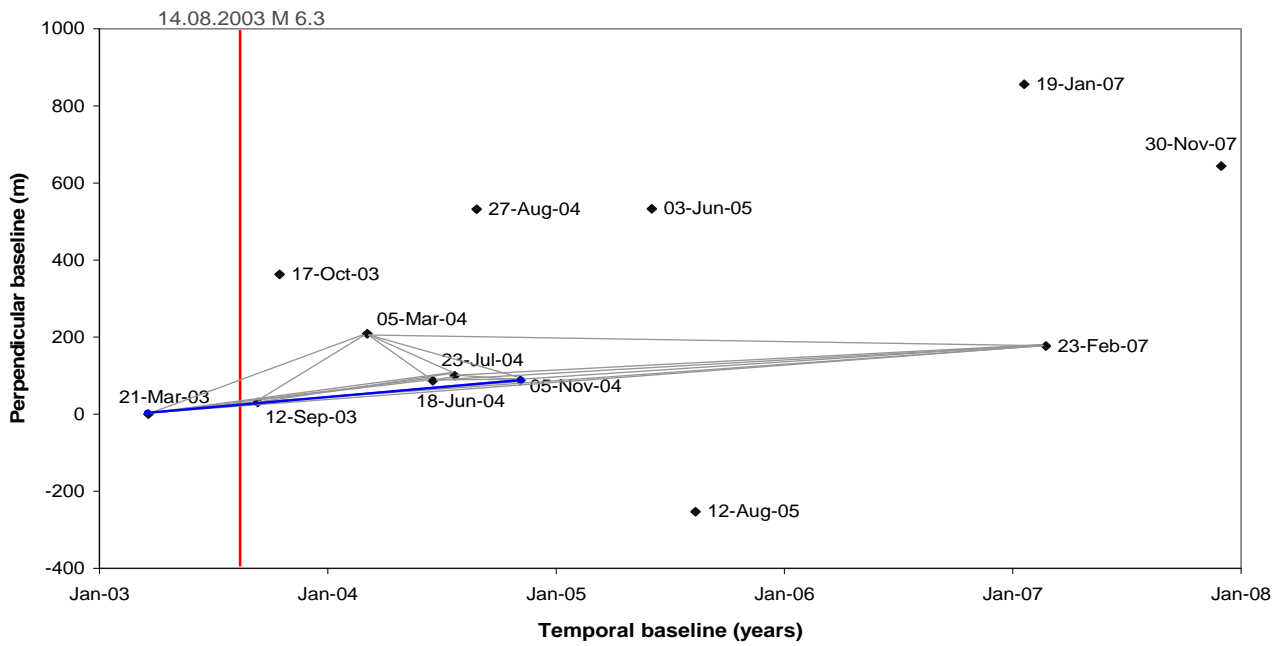
619 Fig. 4. The 14th August 2003 Lefkada earthquake from different sources (Table 2) – the
 620 notation is the same as the references notations in the table. Other fault plane solutions (Table
 621 3) are also shown (the notation is the same as the references notations in the table): LOU99 –
 622 typical focal mechanism obtained by waveform modelling of micro-earthquakes; PD02 –
 623 stress evolutionary model; LAG07 – InSAR forward modelling; HOL08 – Okada model
 624 based on co-seismic GPS measurements. Aftershock sequence from PSL is introduced with
 625 coloured circles according to the depth. In the present investigation the Harvard solution for
 626 the 2003 event (in red – HRV03) is taken as a basic in inversion process. The final solution of
 627 our study is shown in brown (IL16) referred to the centre of the fault predicted by the fault
 628 model. The Harvard solution for the 2015 event is shown (HRV15) with relocated epicentre
 629 and modelled fault plane (red rectangle) from Ganas et al. (2016).



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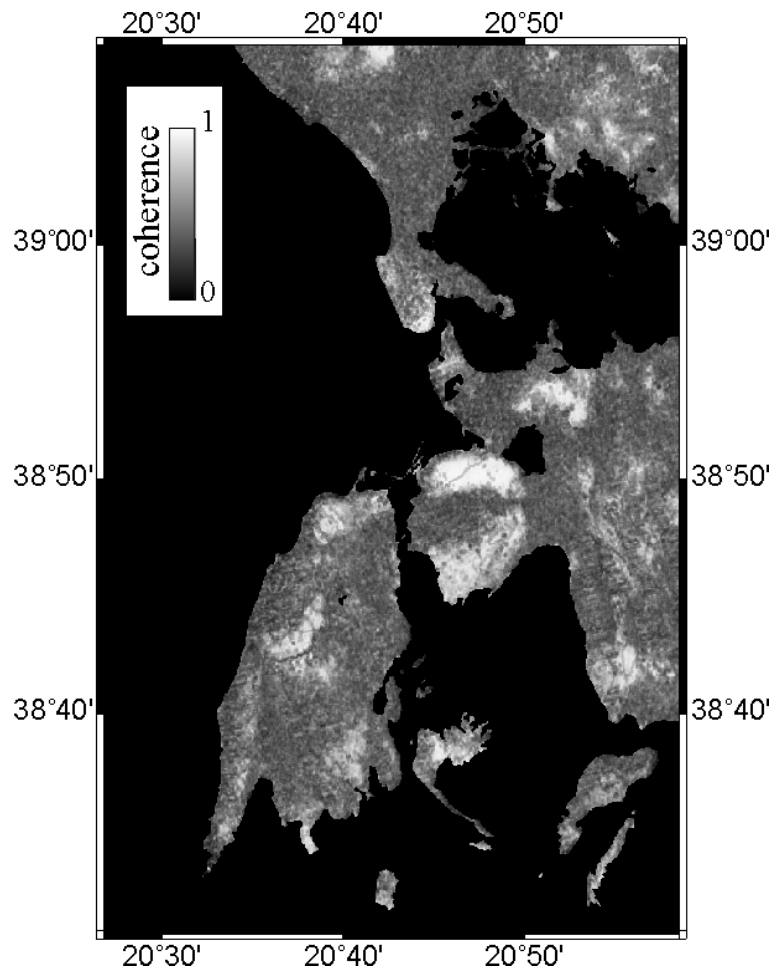
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632 Fig. 5. ENVISAT/ASAR ascending footprint, track 458, frame 774, represented by red
 633 dashed rectangle, descending footprint, track 322, frame 2821 – by blue dashed rectangle.
 634 Location of 14th August 2003 earthquake from NOA is shown with red star and aftershock
 635 sequence by PSL with red dots. The long red arrow represents the satellite flight direction and
 636 short arrow is the satellite looking direction (LOS); black rectangle – the area of investigation
 637 presented in the next figures; orange dashed line – generalized view of Cephalonia Fault
 638 Zone; isobaths – generated from ETOPO1 (Amante and Eakins, 2009); topography – SRTM
 639 v4.



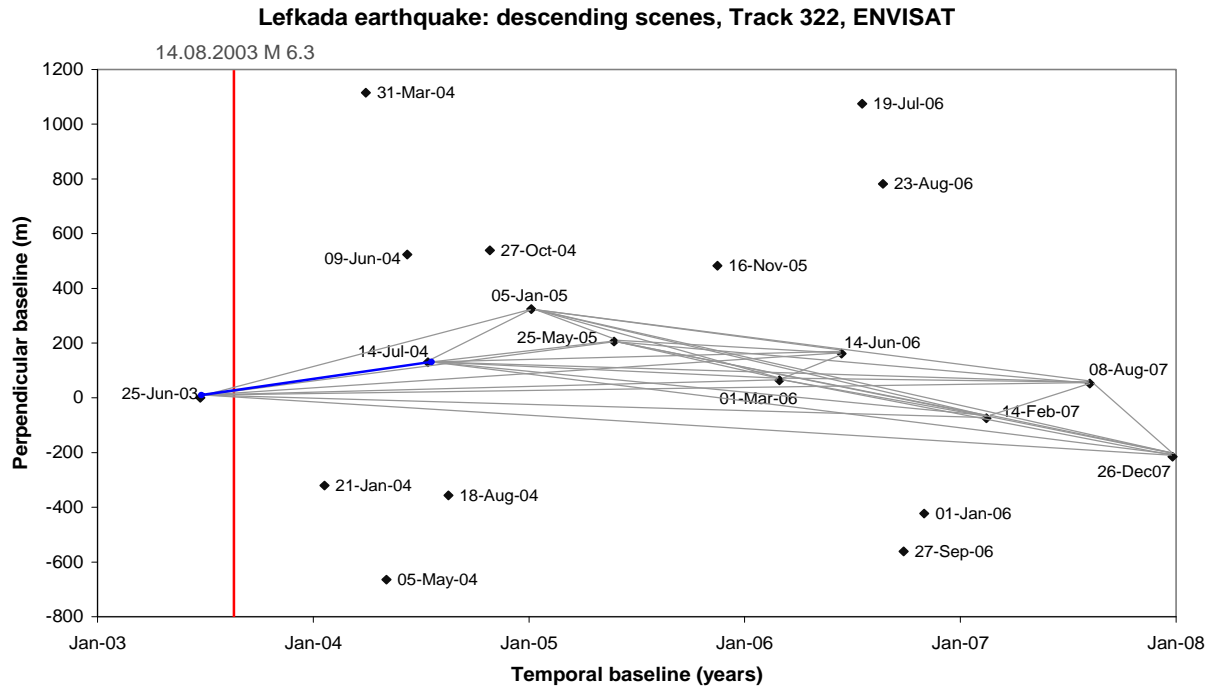
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Fig. 6. Available ascending scenes covering the Lefkada Island: with the red line – time of the 14th August 2003 earthquake, with grey lines – the generated interferograms, with blue line – couple with the best quality (*ifg5*: 21.03.2003 – 05.11.2004).



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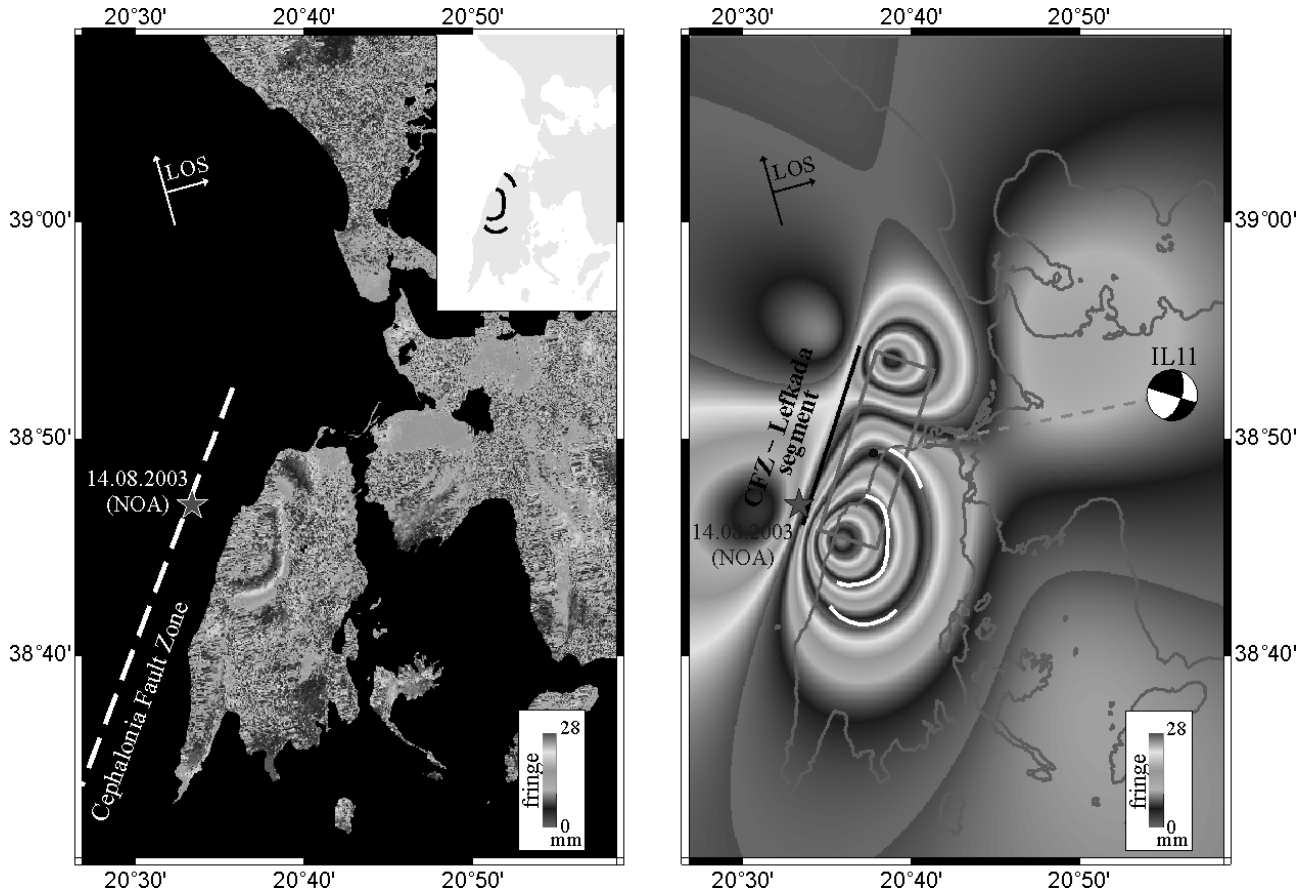
Fig. 7. Coherence map generated from 21 ascending interferograms, showing the mean level of coherence on the territory of the island. The parts which return better signals are in the central and northern part of the area as well as in the closest mainland.



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651 Fig. 8. Available descending scenes covering the Lefkada Island: with the red line – time of
652 the 14th August 2003 earthquake, with grey lines – the generated interferograms, with blue
653 line – couple with the best quality (*ifg22*: 25.06.2003 – 14.07.2004).

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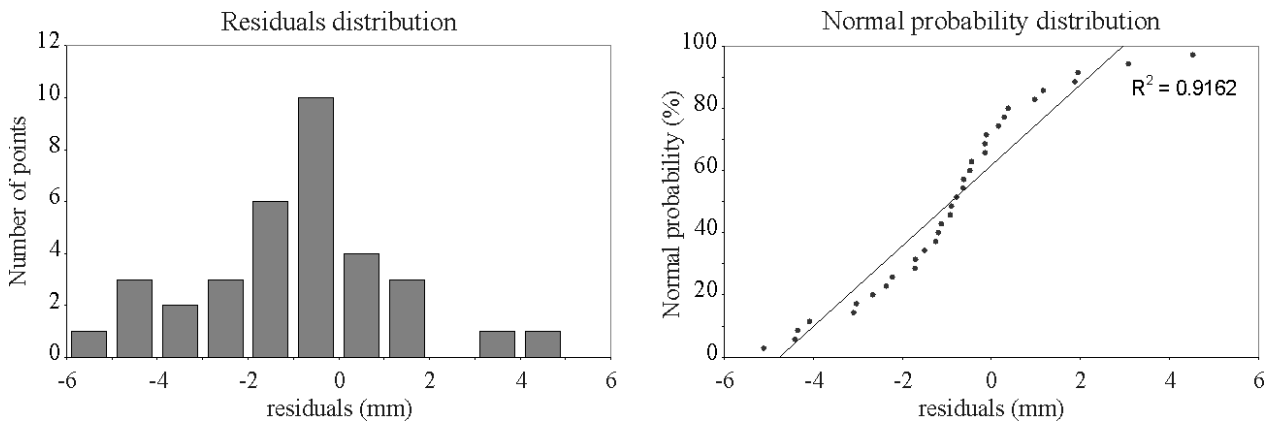
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657 Fig. 9. SAR interferometry results for the Lefkada earthquake. Left: Differential ascending
 658 interferogram – *ifg5* (BL 88m). The star shows the main shock according to NOA; the arrows
 659 – direction of ascending flight and side of looking of the satellite. An approximation of the
 660 Cephalonia Fault Zone is presented by white dashed line. Inset: The observed fringes. Right:
 661 Synthetic ascending interferogram obtained from inversion of the received observed fringes
 662 (left). The rectangle represents the approximate model of the fault segment.

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666 Fig. 10. Left: Residual distribution for the 34 points chosen from the observed fringes. Right:
 667 Normal probability plot for the same residual values.

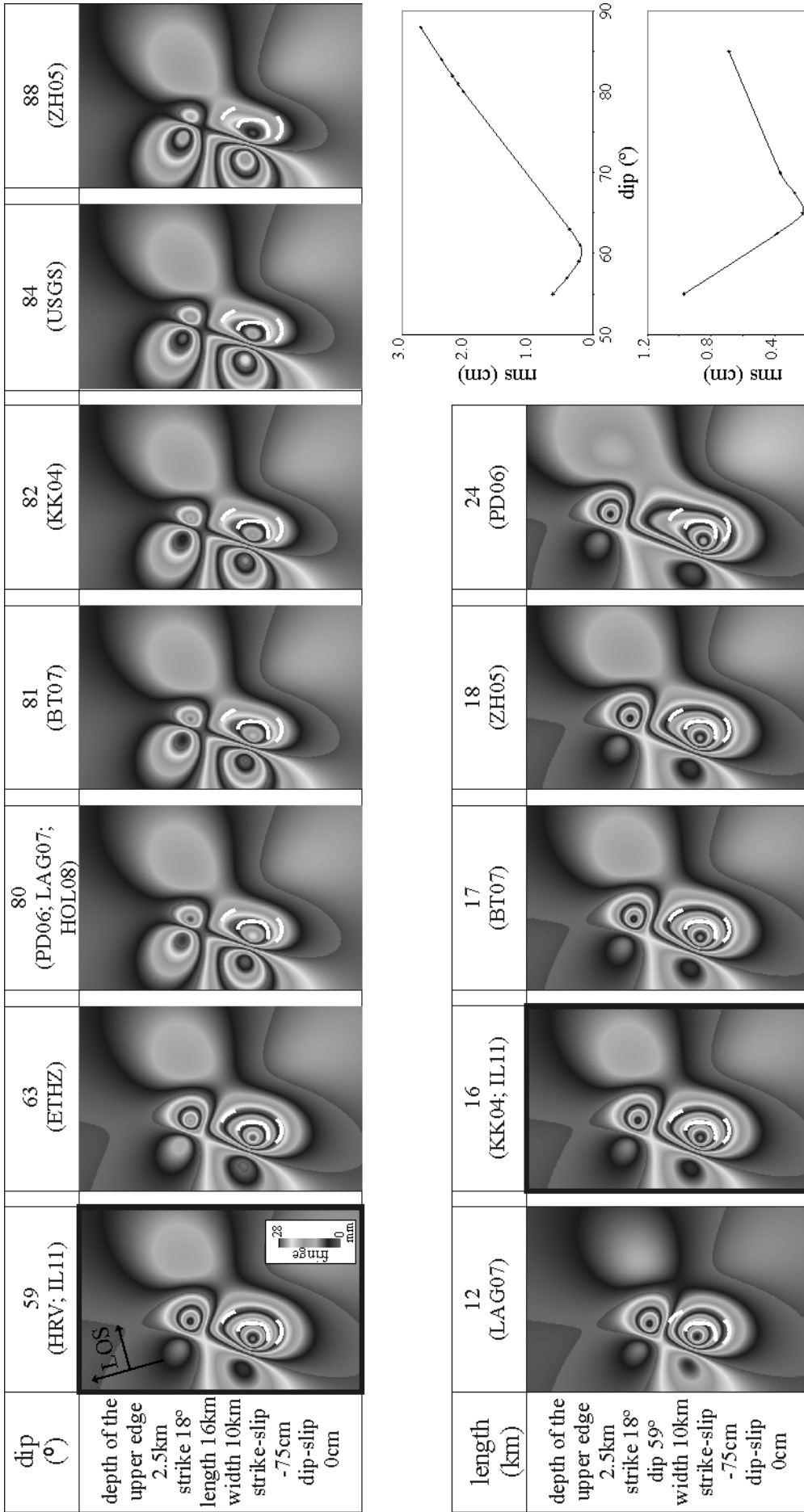
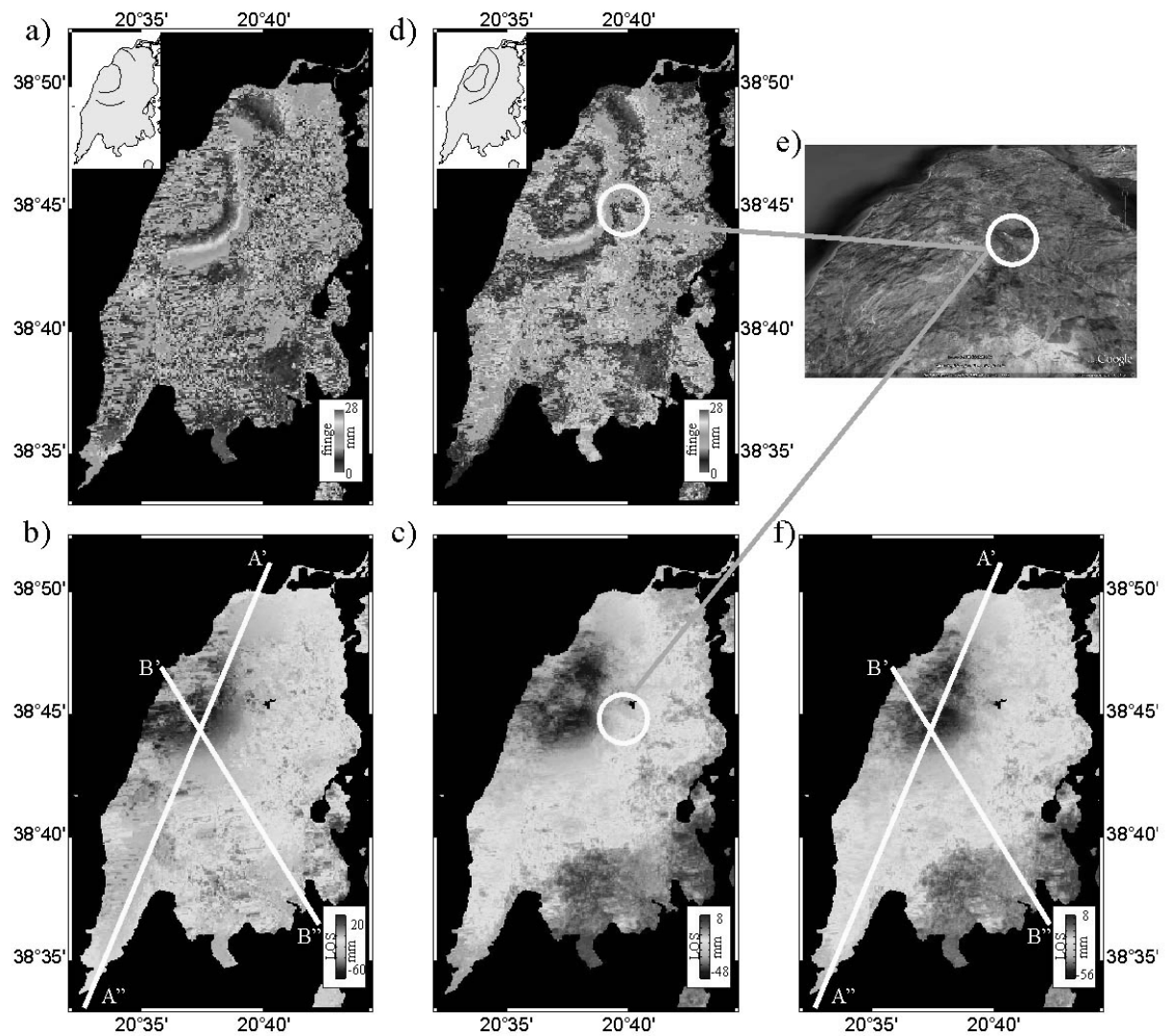


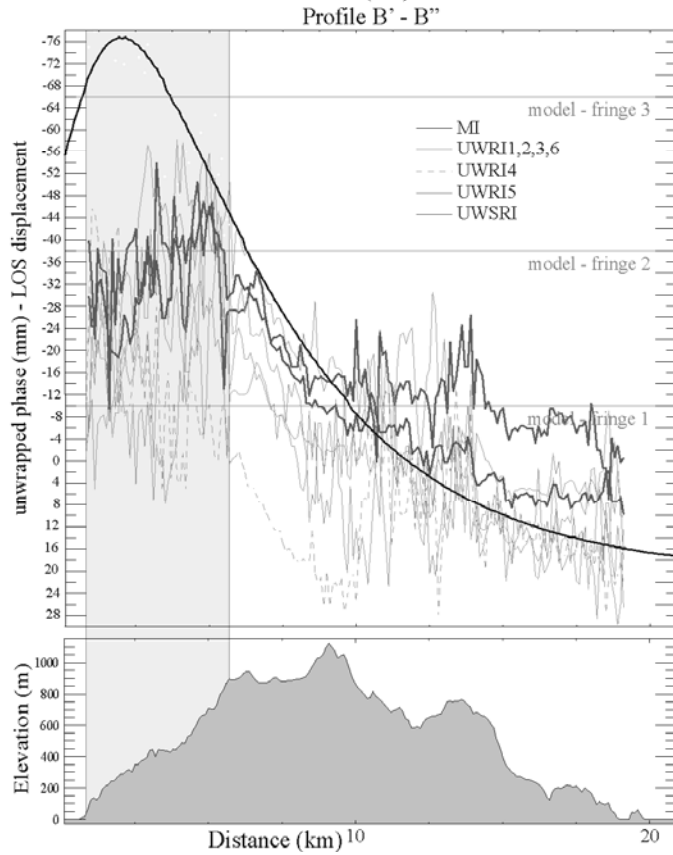
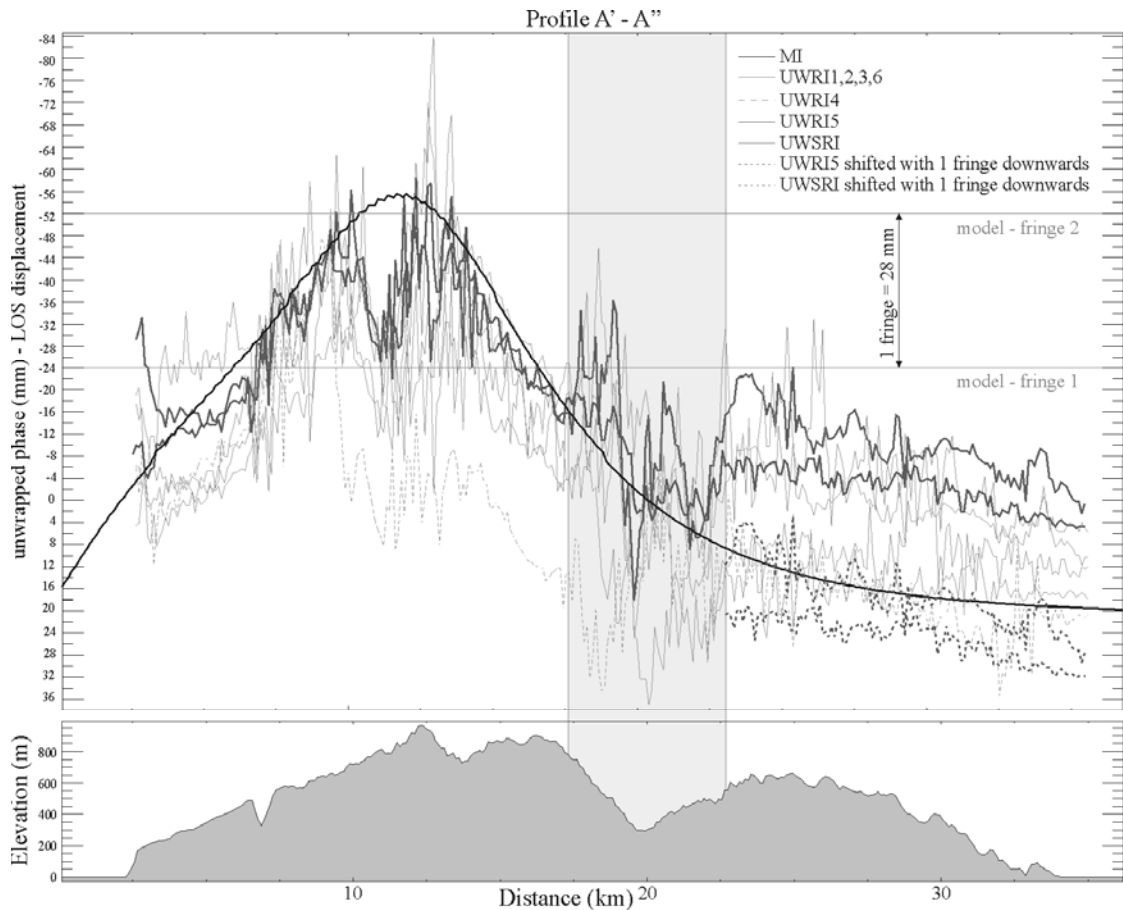
Fig. 11. Dip and length confidence study by comparison with the values proposed by previous studies. The notations are the same as for Table 2 and Table 3. To the right: rms misfits plot for the dip and the length respectively, inverted by the InSAR data. All other parameters are fixed (see the left of the figure)



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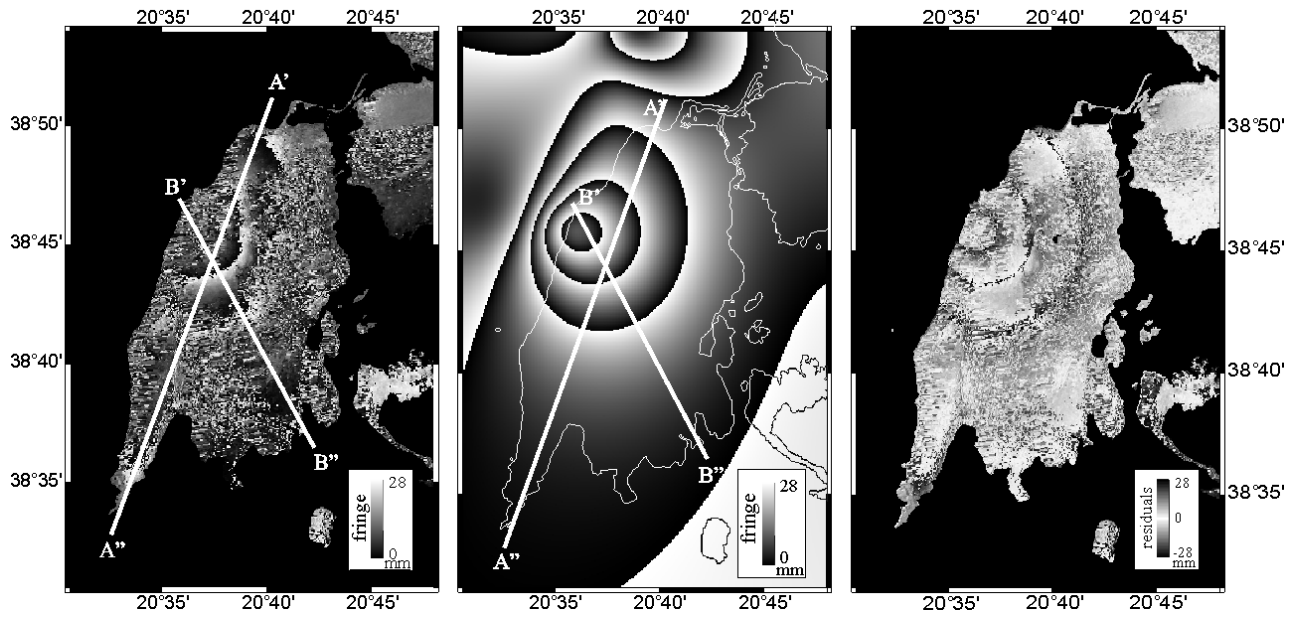
672 Fig. 12. Stacking of ascending interferograms. (a) The ascending interferogram with the best
673 quality (*ifg5*). (b) The same but unwrapped. (c) Average of the six ascending co-seismic
674 unwrapped interferograms. (d) Re-wrapped stacked interferogram and draped (e) over the
675 *Google Earth* topography of the island. (f) Stacking of five unwrapped interferograms
676 (without *ifg4*).

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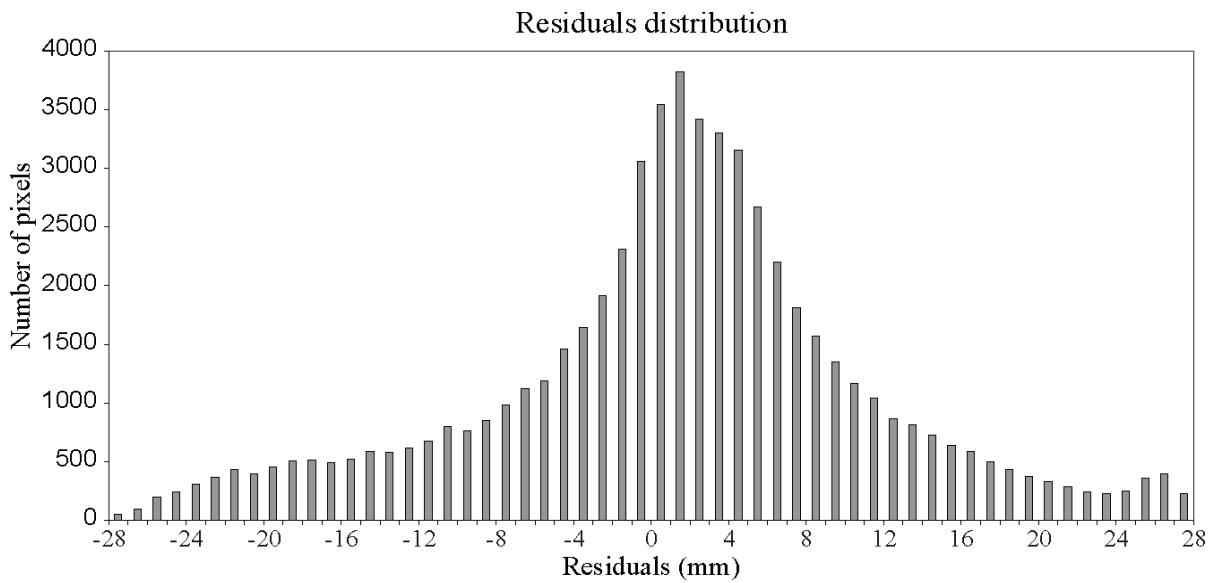
Fig. 13. Unwrapped phases over the profiles A' - A'' (up) and B' - B'' (down) (see Fig. 12).



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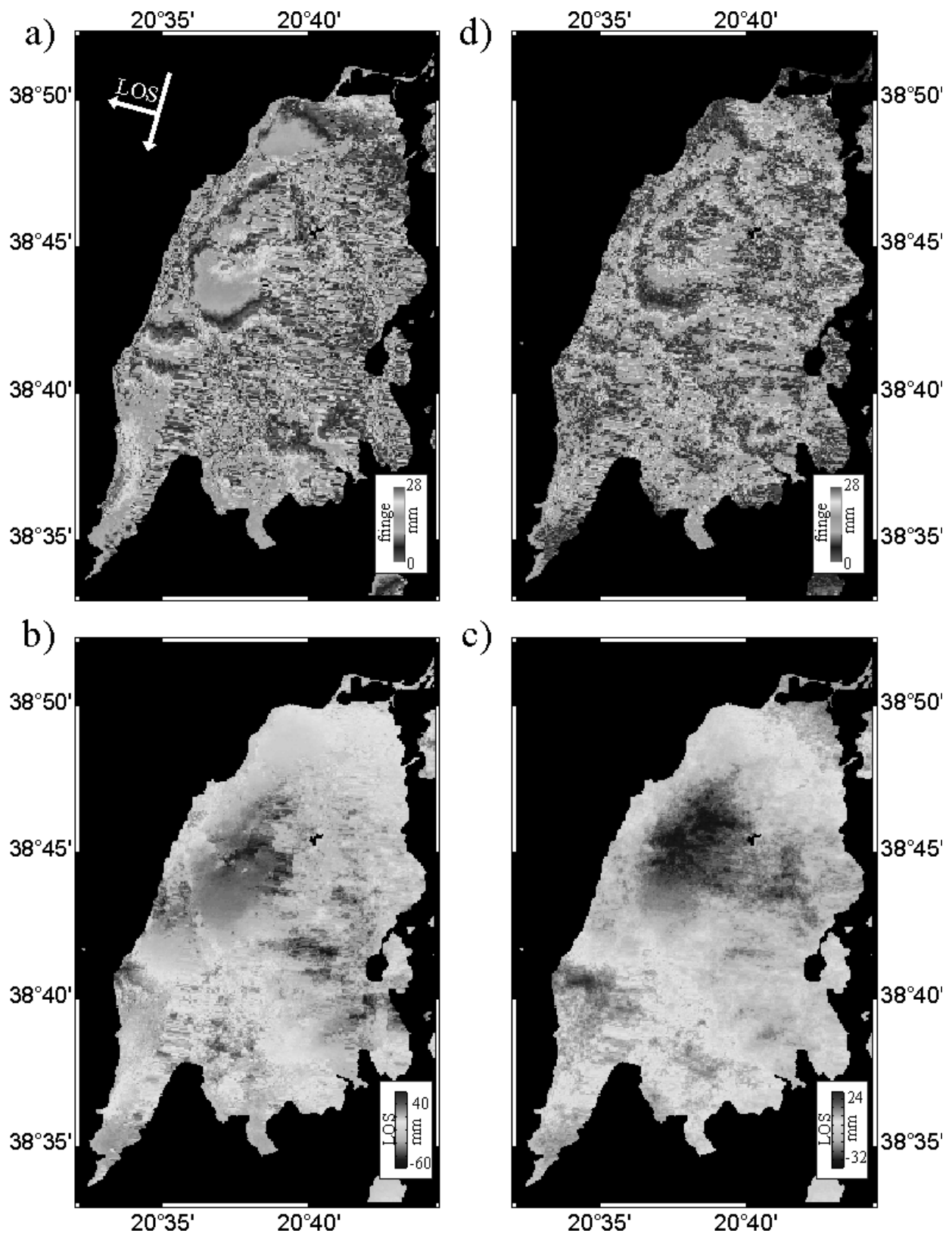
683 Fig. 14. Residual interferometric pattern (right) compared to the real (left) and synthetic
684 (centre) ascending interferograms

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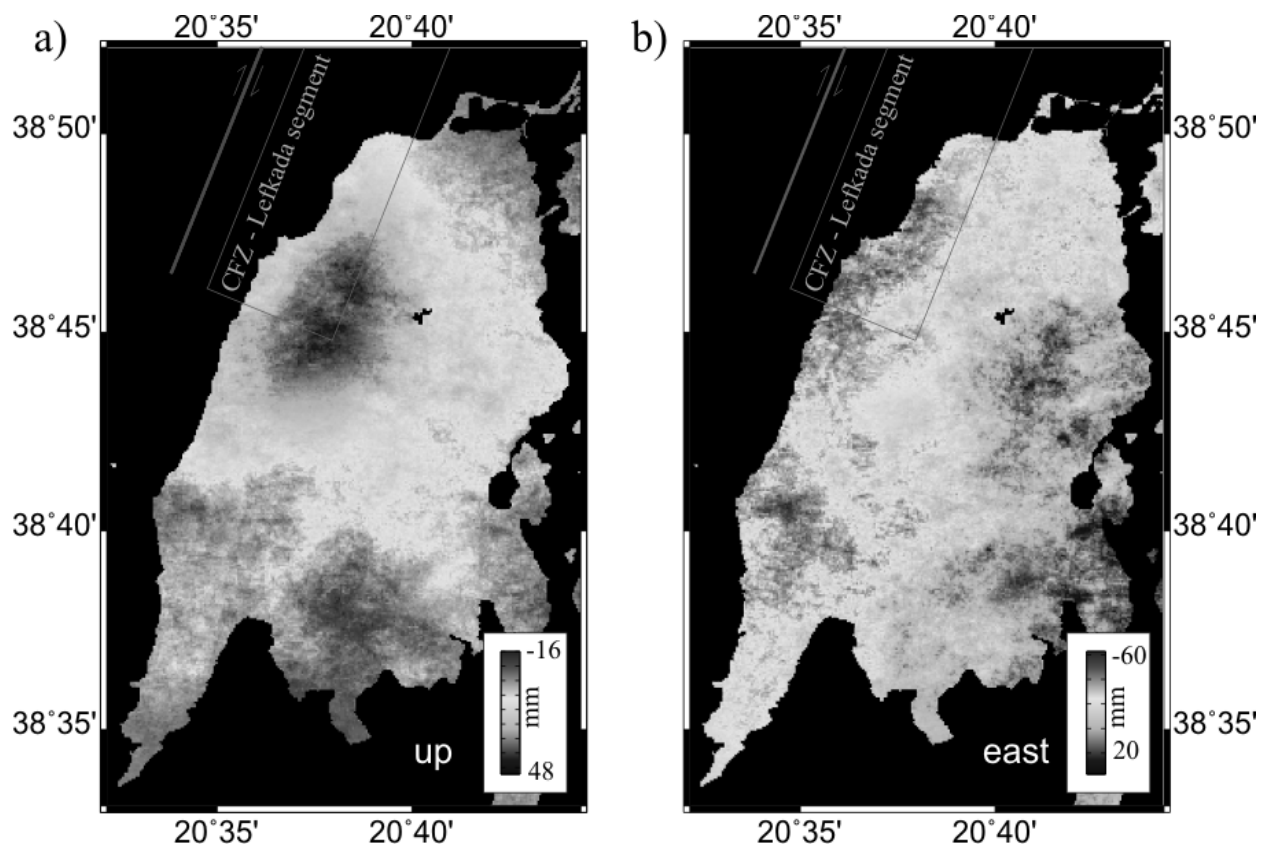
688
689 Fig. 15. Pixel residuals distribution

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693 Fig. 16. Stacking of the descending interferograms. (a) The descending interferogram with the
694 best quality (*ifg22*). (b) The same but unwrapped. (c) Average of the eight descending co-
695 seismic unwrapped interferograms. (d) Re-wrapped stacked descending interferogram.



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Fig. 17. Theoretical displacements in Up and East directions.

Date (Y.M.D)	lat (°)	long (°)	M	depth (km)	Io	strike (°)	dip (°)	rake (°)	Mo (* 10 ¹⁸ Nm)	length (km)	slip (cm)	Ref.*
1612.05.26	38.8	20.6	6.6		VIII							2 in 1
1613.10.12	38.8	20.8	6.4		VIII							2 in 1
1625.06.28	38.8	20.7	6.6		IX							2 in 1
1630.07.02	38.8	20.8	6.7		IX							4 in 3
1704.11.22	38.8	20.7	6.3		IX							2 in 1
1722.06.05	38.6	20.7	6.4		VIII							2 in 1
1723.02.22	38.6	20.7	6.7		IX							2 in 1
1769.10.12	38.8	20.6	6.7		IX							2 in 1
1783.03.23	38.7	20.61	6.7		X							2 in 1
1815.03.17	38.8	20.7	6.3		VIII							2 in 1
1820.02.21	38.8	20.6	6.4		IX							2 in 1
1825.01.19	38.7	20.6	6.5		X							2 in 1
1869.12.28	38.80	20.65	6.4	3-15	X	1	60	165		26	35	4 in 5
1914.11.27	38.65	20.62	6.3	3-15	IX	16	60	165		23	29	6 in 5
1948.04.22	38.68	20.57	6.5	3-15	IX	1	60	165		30	42	4 in 5
1948.06.30	38.80	20.60	6.4	3-15	IX	1	60	165		26	35	4 in 5
1973.11.04	38.89	20.44	5.8	23	VII+	324	50	81	0.61			8 in 7
1994.02.25	38.76	20.54	5.5	9		22	58	168	0.215			7

702
703 **Table 1.** Historical and instrumental seismicity review. References: 1 – Papathanassiou et al.
704 (2005); 2 – Papazachos et al. (2000); 3 – Fokaefs and Papadopoulos (2004); 4 – Papazachos
705 and Papazachou (1997); 5 – Papadimitriou (2002); 6 – Papazachos et al. (1998); 7 – Louvari
706 et al. (1999); 8 – Baker et al. (1997). The intensity values are according to Papathanassiou et
707 al. (2005) except the event of 1630, whose value is taken from Fokaefs and Papadopoulos
708 (2004).

Ref.*	lat (°)	long (°)	Mw	dept h (km)	strike (°)	dip (°)	rake (°)	lengt h (km)	width (km)	Mo (* 10 ¹⁸ Nm)	Slip (cm)	
HRV	38.70	20.67	6.2	15	18	59	-174	-	-	2.98	-	
USGS	39.16	20.60	6.1	21	13	84	172	-	-	2.0	-	
MN	38.88	20.62	6.2	24	196	85	-166	-	-	2.6	-	
ETHZ	39.24	20.49	6.32	31	97	63	-30	-	-	3.72	-	
NOA	38.79	20.56	6.4	12	-	-	-	-	-	-	-	
PSL	38.83	20.62	ML 5.85	15.5	-	-	-	-	-	-	-	
AUT	38.74	20.54	6.3	11	-	-	-	-	-	-	-	
ZH05:	ms	38.84	20.60									
ZH05_1	s1	38.78	20.57	5.7	13	17	88	-177	18	9	0.47	-
ZH05_2	s2	38.46	20.43	5.8	13	24	74	164	-	-	0.49	-
BT07:	av			6.2	13	16	72	178	23	-	2.23	42
BT07_1	s1	38.78	20.57	6.0	11	12	81	174	17	10	1.3	34
BT07_2	s2	38.50	20.44	5.8	15	20	63	-179	13	10	0.65	23
PD06		38.86	20.56	6.3	9	15	80	170	24	2-12	2.9	40

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711 **Table 2.** Published source parameters for the 2003 August 14 Lefkada earthquake (compared
712 also in Fig. 4). ms – main shock, s1 – sub-event 1, s2 – sub-event 2, av – averaged.

713 * HRV – CMT Harvard ; USGS – U.S. Geological Survey; MN – MedNet (Mediterranean
714 Very Broadband Seismographic Network – INGV) in Zahradník et al. (2005); ETHZ – Swiss
715 Federal Institute of Technology Zurich; NOA – National Observatory of Athens, Institute of
716 Geodynamics; PSL – Patras Seismological Laboratory; AUT – Aristotle University of
717 Thessaloniki; ZH05_1 and ZH05_2 – 1 and 2 sub-events in Zahradník et al. (2005); BT07_1
718 and BT07_2 – 1 and 2 sub-events in Benetatos et al. (2007); PD06 - Papadimitriou et al.
719 (2006).

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Ref.	lat (°)	long (°)	Mw	dept h (km)	strik e (°)	dip (°)	rake (°)	L (km)	W (k m)	Mo (* 10 ¹⁸ Nm)	slip (cm)
LOU99	-	-	-	-	14	65	167	40	15	-	-
PD02 □	38.71	20.58	-	15	11	60	+165	-	-	-	-
KK04	-	-	-	3-12	-	-	-	16	-	-	~60
LAG07	-	-	-	13	15	80	174*	12	11	-	-45
HOL08	38.60	20.50	6.3*	7	15	80	180*	60	10	2.97	16.5
model in this study (IL16)	38.83	20.64#	6.3	6.8 #	18	59	180	16	10	3.60	75
Uncertainties of our model	0.01	0.01		1.	3.	5.	8.	2.	2.		15

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723 **Table 3.** Other fault plane solutions for Lefkada Segment of Cephalonia Fault Zone
724 (compared also in Fig. 4). The coordinates are for the centre of the fault. * estimated here; □ -
725 after Papazachos et al. (1998); # - centre of the fault; L – length; W –width; LOU99 – Louvari
726 et al. (1999); PD02 – Papadimitriou (2002); KK04 – Karakostas et al. (2004); LAG07 –
727 Lagios et al. (2007); HOL08 – Hollenstein et al. (2008).

Ni	Pass	Date 1	Date 2	Days	B_{\perp} (mm)	h_a (m)	e_{\max}^h (mm)
<i>ifg</i> 1	A	21.03.2003	12.09.2003	174	-35	268	2
<i>ifg</i> 2	A	21.03.2003	05.03.2004	349	-209	45	10
<i>ifg</i> 3	A	21.03.2003	18.06.2004	453	-87	108	4
<i>ifg</i> 4	A	21.03.2003	23.07.2004	489	-104	90	5
<i>ifg</i> 5	A	21.03.2003	05.11.2004	592	-88	107	4
<i>ifg</i> 6	A	21.03.2003	23.02.2007	1432	-177	53	8
<i>ifg</i> 7	A	12.09.2003	05.03.2004	175	-179	52	9
<i>ifg</i> 8	A	12.09.2003	18.06.2004	280	-56	167	3
<i>ifg</i> 9	A	12.09.2003	23.07.2004	315	-70	134	3
<i>ifg</i> 10	A	12.09.2003	05.11.2004	419	-57	165	3
<i>ifg</i> 11	A	12.09.2003	23.02.2007	1258	-146	64	7
<i>ifg</i> 12	A	05.03.2004	18.06.2004	104	122	77	6
<i>ifg</i> 13	A	05.03.2004	23.07.2004	140	109	86	5
<i>ifg</i> 14	A	05.03.2004	05.11.2004	243	121	77	6
<i>ifg</i> 15	A	05.03.2004	23.02.2007	1082	32	293	2
<i>ifg</i> 16	A	18.06.2004	23.07.2004	35	-18	521	1
<i>ifg</i> 17	A	18.06.2004	05.11.2004	139	-6	1563	0
<i>ifg</i> 18	A	18.06.2004	23.02.2007	978	-92	102	4
<i>ifg</i> 19	A	23.07.2004	05.11.2004	104	18	521	1
<i>ifg</i> 20	A	23.07.2004	23.02.2007	943	-77	122	4
<i>ifg</i> 21	A	05.11.2004	23.02.2007	839	-92	102	4
<i>ifg</i> 22	D	25.06.2003	14.07.2004	385	-118	79	6
<i>ifg</i> 23	D	25.06.2003	05.01.2005	560	-300	31	14
<i>ifg</i> 24	D	25.06.2003	25.05.2005	700	-201	47	9
<i>ifg</i> 25	D	25.06.2003	01.03.2006	980	-48	195	2
<i>ifg</i> 26	D	25.06.2003	14.06.2006	1085	-154	61	7
<i>ifg</i> 27	D	25.06.2003	14.02.2007	1330	79	119	4
<i>ifg</i> 28	D	25.06.2003	08.08.2007	1505	-45	208	2
<i>ifg</i> 29	D	25.06.2003	26.12.2007	1645	212	44	10

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730 **Table 4.** Details for the ascending (A) and descending (D) interferograms processed for this
731 study, where Ni – number of interferogram, B – perpendicular baseline (m), Days – offset
732 between acquisition of the two images; h_a – altitude of ambiguity; e_{\max}^h – maximum error due
733 to the topography. The bold lines show the best ascending (A) and descending (D)
734 interferograms.

Scene date	Pass	Mean temperature (°C)	Mean pressure at sea level (hPa)	Mean humidity (%)	Wind (m/s)	Cloudiness (%)	Cloud base altitude (m)	Precipitation
21.03.2003	A	+10.6	1016	74	4	20	1750-3000	None
12.09.2003	A	+22.6	1013	75	5	40	800	None
05.03.2004	A	+ 9.4	1014	54	6	80	800-2250	None
18.06.2004	A	+22.8	1004	72	3	50	800	Rain or Drizzle
23.07.2004	A	+27.3	1014	71	3	10	3000	None
05.11.2004	A	+18.1	1018	54	8	10	800-2250	None
23.02.2007	A	+11.8	1011	90	3	30	450-2250	None
25.06.2003	D	+26.6	1014	78	3	20	1750-3000	None
14.07.2004	D	+23.3	1011	57	6	20	800	None
05.01.2005	D	+7.4	1028	75	1	30	3000	None
25.05.2005	D	+20.4	1014	85	1	30	800	None
01.03.2006	D	+11.6	1013	75	0	70	800	Mist
14.06.2006	D	+19.3	1019	80	5	20	800	None
14.02.2007	D	+12.1	1009	72	6	40	800	27 mm
08.08.2007	D	+24.3	1013	82	3	20	450-3000	None
26.12.2007	D	+10.0	1025	94	3	50	250-800	15 mm, Fog, Rain and Drizzle

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Table 5. ENVISAT images used in the research and the weather conditions during their acquisitions according to the weather archives for Aktion Airport (38°37'N 20°46'E) (<http://meteo.infospace.ru>; <http://www.tutiempo.net>). The bold lines correspond to the dates of acquisition of the images used for the best ascending and descending interferograms.

Source of empirical relationship	L (Surface rupture length) - km	W (Down-dip rupture width) - km	U (Average displacement) - cm
Konstantinou <i>et al.</i> (2005)	$\log L = -1.49 + 0.47 M_w$ L = 29	$\log W = -1.07 + 0.34 M_w$ W = 11	
Papazachos & Papazachou (1997)	$\log L = -1.85 + 0.51 M_w$ L = 23	$\log W = -0.13 + 0.19 M_w$ W = 12	$\log u = -3.71 + 0.82 M_w$ u = 24
Well & Coppersmith (1994)	$\log L = -3.55 + 0.74 M_w$ L = 13	$\log W = -0.76 + 0.27 M_w$ W = 9	
this study	L = 16	W = 10	u = 75

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Table 6. Comparison between Length-Magnitude empirical relations and calculated fault dimensions

Long (°)	Lat (°)	UTM_E (km)	UTM_N (km)	LoS observed (mm)	LoS modelled (mm)
20.5765	38.6244	463.16	4275.20	24.4	7.2
20.5793	38.6332	463.41	4276.17	24.4	8.4
20.5853	38.6428	463.93	4277.24	24.4	10.2
20.5874	38.6558	464.11	4278.67	24.4	12.8
20.5884	38.6646	464.20	4279.64	24.4	14.9
20.5891	38.6709	464.26	4280.34	24.4	16.7
20.5909	38.6794	464.42	4281.29	24.4	19.6
20.5916	38.6844	464.48	4281.83	24.4	21.4
20.5926	38.6929	464.57	4282.78	24.4	24.9
20.5975	38.6984	465.00	4283.39	24.4	27.8
20.6084	38.7001	465.96	4283.57	24.4	28.9
20.6236	38.7042	467.28	4284.03	24.4	29.9
20.6335	38.7097	468.15	4284.64	24.4	30.8
20.6415	38.7166	468.85	4285.40	24.4	31.8
20.6493	38.7246	469.53	4286.28	24.4	32.5
20.6524	38.7345	469.81	4287.38	24.4	34.5
20.6588	38.7433	470.36	4288.35	24.4	33.8
20.6602	38.7524	470.49	4289.35	24.4	34.5
20.6630	38.7601	470.73	4290.21	24.4	33.6
20.6637	38.7692	470.79	4291.21	24.4	32.9
20.6608	38.7758	470.54	4291.94	24.4	33.5
20.6584	38.7857	470.33	4293.04	24.4	32.4
20.6516	38.7934	469.74	4293.89	24.4	32.9
20.5889	38.7416	464.25	4288.16	52.4	53.1
20.5971	38.7367	464.96	4287.61	52.4	50.9
20.6084	38.7347	465.95	4287.40	52.4	49.3
20.6193	38.7356	466.91	4287.50	52.4	47.6
20.6316	38.7370	467.99	4287.65	52.4	44.1
20.6351	38.7419	468.29	4288.20	52.4	44.7
20.6394	38.7469	468.66	4288.74	52.4	44.2
20.6397	38.7521	468.69	4289.32	52.4	45.3
20.6383	38.7593	468.57	4290.11	52.4	47.2
20.6354	38.7686	468.32	4291.15	52.4	48.8
20.6294	38.7799	467.79	4292.39	52.4	49.7

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Table 7. The 34 picked values on the two visible ascending fringes in the best ascending interferogram. The origin for the LoS was estimated during the joint inversion process. The coordinates are also given in UTM 34.

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Long (°)	Lat (°)	UTM_E (km)	UTM_N (km)	LoS observed (mm)	LoS modelled (mm)
20.6485	38.7945	469.46	4294.01	54.4	41.8
20.6358	38.7901	468.35	4293.52	54.4	51.4
20.6227	38.7804	467.21	4292.45	54.4	65.3
20.6136	38.7738	466.40	4291.72	54.4	71.7
20.6072	38.7625	465.85	4290.47	54.4	74.9
20.5998	38.7523	465.20	4289.35	54.4	70.0
20.5978	38.7375	465.02	4287.70	54.4	59.6
20.6003	38.7262	465.24	4286.46	54.4	52.4
20.6042	38.7144	465.58	4285.15	54.4	45.4
20.6151	38.7078	466.54	4284.42	54.4	43.9
20.6268	38.7084	467.56	4284.49	54.4	46.2
20.6366	38.7119	468.42	4284.88	54.4	49.1
20.6458	38.7205	469.22	4285.83	54.4	54.1
20.6574	38.7331	470.24	4287.23	54.4	58.2
20.6630	38.7453	470.73	4288.56	54.4	59.6
20.6665	38.7587	471.04	4290.05	54.4	57.8
20.6686	38.7689	471.22	4291.18	54.4	53.5
20.6651	38.7813	470.91	4292.55	54.4	47.5
20.6615	38.7907	470.60	4293.58	54.4	40.7
20.6545	38.7945	469.98	4294.01	54.4	39.5

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Table 8. The 20 picked values on the visible fringe in the best descending interferogram. The origin for the LoS was estimated during the joint inversion process. The coordinates are also given in UTM 34.

Ref.	lat (°)	long (°)	depth (km)	strike (°)	dip (°)	rake (°)	L (km)	W (km)	Mo (* 10 ¹⁸ Nm)	slip (cm)
model in this study (IL16)	38.83	20.64	6.8	18	59	180	16	10	3.60	75
alternative model using the combination of ascending and descending interferograms	38.83	20.63	8.8	18	59	180	16	10	3.50	73

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Table 9. Comparison of the fault models from ascending interferograms only (IL16, see Table 3) and from a combination ascending of descending interferograms. The coordinates are for the centre of the fault.

785 **APPENDIX A.** ENVISAT ascending co-seismic interferograms

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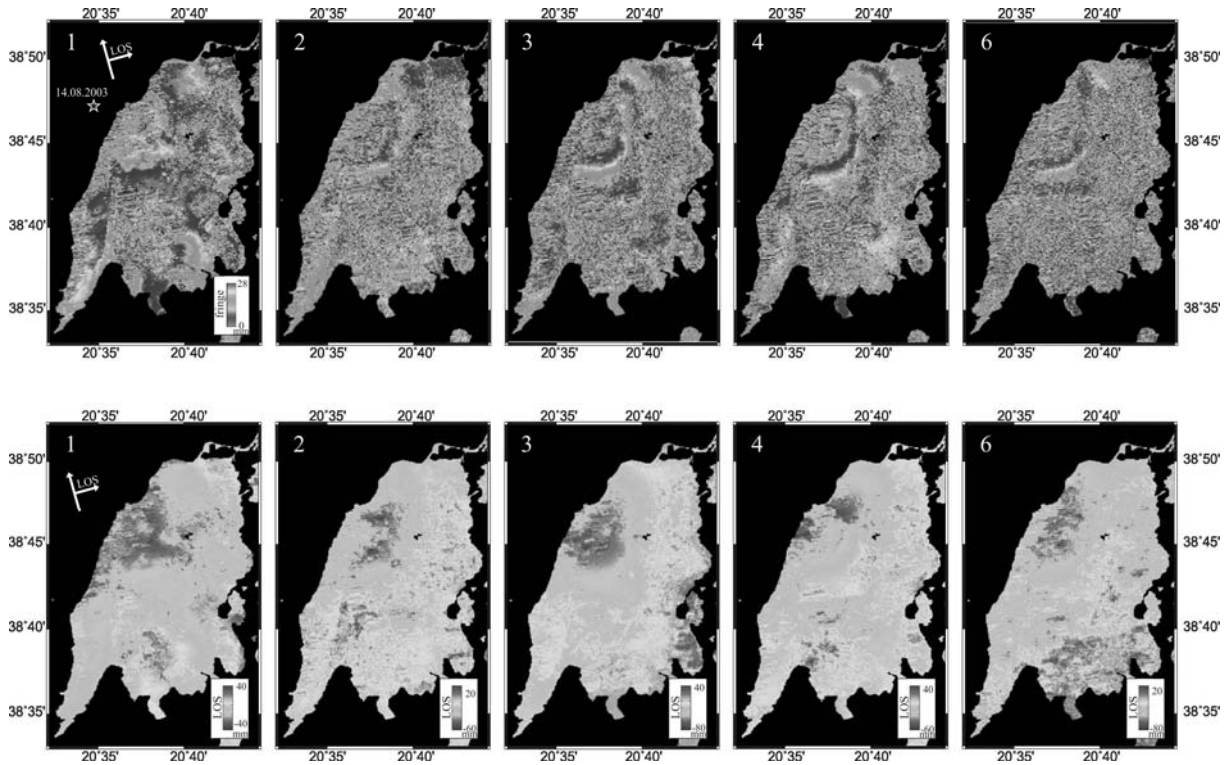
787 Wrapped (top) and unwrapped (bottom) ascending co-seismic interferograms (numbers
788 correspond to the interferogram number in Table 4)

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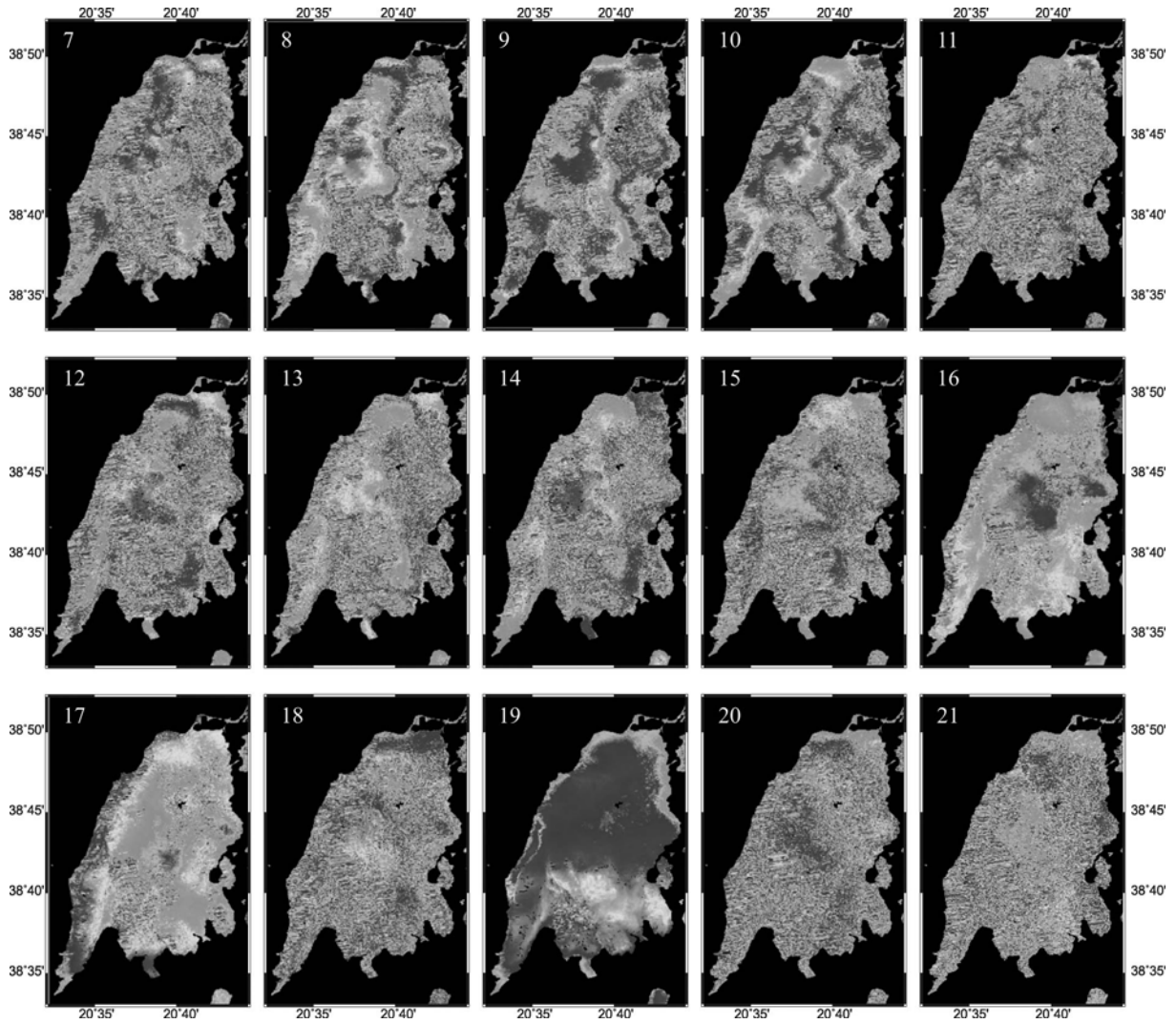


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795 **APPENDIX B.** ENVISAT ascending post-seismic interferograms

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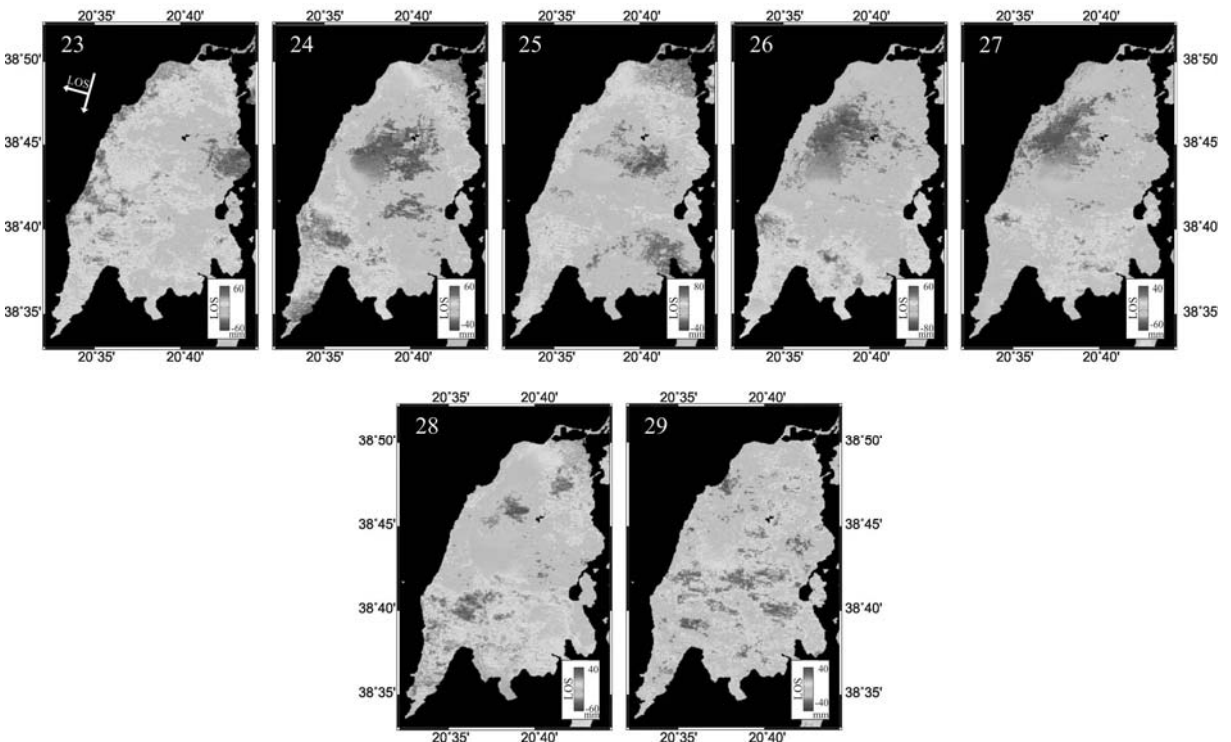
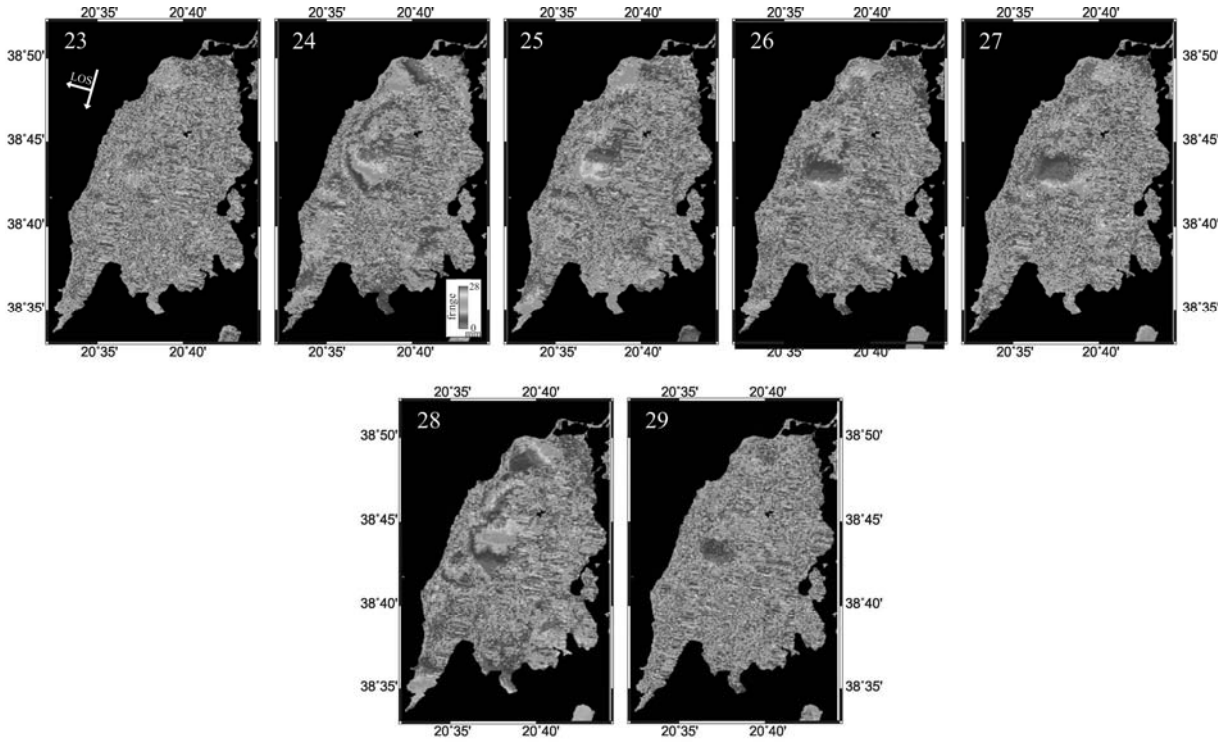
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APPENDIX C. ENVISAT descending co-seismic interferograms

Wrapped (top) and unwrapped (bottom) descending co-seismic interferograms (numbers correspond to the interferogram number in Table 4)



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