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## Non-collinear ordering of the orbital magnetic moments in magnetite

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The magnitude of the orbital magnetic moment [1-9] and its role as a trigger of the Verwey transition [10-17] in the prototypical Mott insulator, magnetite, remain contentious. Using 1s2presonant inelastic X-ray scattering magnetic linear angular distribution (RIXS-MLAD), we prove the existence of non-collinear orbital magnetic ordering and infer the presence of dynamical distortion creating a polaronic precursor for the metal to insulator transition. These conclusions are based on a subtle angular shift of the RIXS-MLAD spectral intensity as a function of the magnetic field orientation. Theoretical simulations show that these results are only consistent with non-collinear magnetic orbital ordering. To further support these claims we perform Fe K-edge X-ray magnetic circular dichroism (XMCD) in order to quantify the iron average orbital magnetic moment.

Magnetite ( $[Fe^{3+}]_A[Fe^{3+}, Fe^{2+}]_BO_4$ ) is the most abun- 47 insulators [23] ask for a new approach. 15 17 18 19 20 21 22 23 24 25 28 ferromagnetically coupled while the Fe ions in the same 62 be either antiparallel or non-collinear. 29 sublattice are ferromagnetically coupled (Fig. 1a). 30

31 32 33 be directly quantified by applying the sum rules on the 66 ordering that can tilt the orbital magnetic moment as  $_{24}$  L<sub>2,3</sub> XMCD signal [22]. In spite of the great success of  $_{67}$  large as  $82^{\circ}$  away from the spin magnetic moment. We 35 sum rules, the experimental and analysis procedures were 66 proposed a model for this non-collinear orbital magnetic 36 37 38  $_{41}$  coupling between the Fe A and B sublattices were also  $_{74}$  (111) Fe<sub>3</sub>O<sub>4</sub> single crystals at room temperature. The av-42 suggested [4] (for a unit formula of Fe<sub>3</sub>O<sub>4</sub>:  $Fe_B \mu_L = 75$  erage orbital magnetic moment projected along the mag-43  $1 \mu_{\rm B}$  and  $Fe_A \mu_L = -1 \mu_{\rm B}$ ). A summary of the orbital 76 netic field direction can be quantified by performing Fe K 44 and spin magnetic moments reported in literature using 77 pre-edge XMCD measurements [22]. Three main dichroic 45 various techniques is shown in Fig. S1. These large dis-78 features can be seen at  $E_I = 7112.7 \text{ eV}, E_{II} = 7114 \text{ eV}$ 46 crepancies regarding an essential quantity to many Mott  $_{70}$  and  $E_{III} = 7115.1 \,\mathrm{eV}$  with only significant angular de-

dant iron bearing mineral on Earth and it finds many 48 In this work we employed a combination of Fe K-edge applications in areas such as palaeomagnetism, medicine, 49 XMCD and 1s2p RIXS-MLAD measurements to investidata recording, and engineering [18]. Ever since Verwey's  $_{50}$  gate the orbital magnetic moment of Fe in Fe<sub>3</sub>O<sub>4</sub>. Experpioneering work [19], an immense amount of research has  $s_1$  iments at the Fe K-edge ( $1s \rightarrow 3d + 4p$  excitations) have been dedicated to  $Fe_3O_4$  in view of its importance as a  $s_2$  a probing depth of few  $\mu m$  and hence surface effects are reference for systems exhibiting the metal to insulator  $s_2$  negligible offering a valuable advantage over  $L_{2,3}$ -edge transition [11, 20, 21]. In Fe<sub>3</sub>O<sub>4</sub>, the Verwey transition <sup>54</sup> measurements. We quantified the average orbital magoccurs at  $T_V \sim 125$  K and results in a spontaneous change so netic moment by performing K-edge XMCD. The accuof both, the lattice symmetry and the electric conductiv-  $_{56}$  rately measured 1s2p RIXS-MLAD signal was used to ity. Above  $T_V$  Fe<sub>3</sub>O<sub>4</sub> has a cubic inverse spinel crystal 57 determine the average square orbital magnetic moment structure containing two different Fe sites. Fe<sup>3+</sup> ions re- 58 which complements the average quantity obtained from side in tetrahedral  $(T_d)$  interstices (the A sites) while 59 XMCD. A difference between the orbital magnetic moboth  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$  ions reside in octahedral  $(O_h)$  inter- **60** ment predicted by both experiments indicates compensastices (the B sites). The A and B sublattices are anti- 61 tion: the orbital magnetic moments of the Fe ions must

Guided by theoretical calculations, we show that the 63 In ferromagnets and ferrimagnets, the spin and or- of only possibility to explain both of our experimental rebital magnetic moments of the transition metal ions can 65 sults is the existence of a non-collinear magnetic orbital shown to be prone to huge uncertainties due to surface or ordering and inferred the presence of a dynamical distoreffects [1]. Orbital magnetic moments as small as  $0.01 \,\mu_{\rm B}$  70 tion related to the  $X_3$  phonon mode in the high tempera-[2] and as large as  $0.33 \,\mu_{\rm B}$  [3] were reported for Fe<sub>3</sub>O<sub>4</sub>. *n* ture phase. Our model imposes strong restrictions on the In addition, large orbital magnetic moment contributions 72 candidate mechanisms for the Verwey transition [10, 24]. that are of equal absolute values but with antiparallel 73 We investigated highly stoichiometric (001), (110) and so pendence at the first feature (Fig. 1b). The co-existence <sup>\$1</sup> of two Fe sites in Fe<sub>3</sub>O<sub>4</sub> complicates the direct analysis 82 of the XMCD signal and renders a full calculation important to assign spectral features to the specific Fe species. 83 Configuration interaction calculations taking into ac-84 count i)- intra-atomic Coulomb interaction, ii)- crystal 85 field, *iii*)- spin-orbit coupling, and *iv*)- exchange inter-86 action were performed using the quantum many-body 87 program Quanty [25–27]. Our theoretical simulations 88 show that the feature at  $E_I$  arises mainly from electric 89 quadrupole transitions (*i.e.*  $1s \rightarrow 3d$  excitations) at the 90 formal  $\mathrm{Fe}^{2+}$  ions while the features at  $E_{II}$  and  $E_{III}$  arise 91 mainly from electric dipole transitions (*i.e.*  $1s \rightarrow 3d + 4p$ excitations) at the  $\text{Fe}^{3+}$   $T_d$  ions due to onsite 3d - 4p93 orbital mixing as discussed by Westre et. al. [28]. The quadrupolar signals from the  $Fe^{3+}$  A and B sites nearly cancel out as shown in Fig. S8.

The experimental XMCD signal and its angular de-97 pendence can be best interpreted to arise from a par-98 tially quenched orbital magnetic moment at the formal  $\mathrm{Fe}^{2+}$  ions. An excellent agreement between the calculation and the experiment is observed (compare spectra <sup>102</sup> labelled Exp and Calc 1 in Fig. 1b). This partial quench-<sup>103</sup> ing is a result of the octahedral symmetry accompanied 104 by a small trigonal distortion  $(D_{\sigma} = 67 \pm 10 \,\mathrm{meV})$ . The Fe environment is not perfectly  $O_h$  since the point group 105 symmetry of the B site is rhombohedral  $(D_{3d} \equiv \bar{3}m)$ . 106 We found that the average orbital magnetic moment is 107  $0.26 \pm 0.03 \,\mu_{\rm B}$  per unit formula of Fe<sub>3</sub>O<sub>4</sub>. It is impor-108 tant to note that feature I is theoretically predicted to 136 the experimental dichroism signal (Fig. 2b). Theoretical 109 110 111 finite orbital magnetic moment in bulk  $Fe_3O_4$ . 112

113 114 115 117 118 119 120 and magnetic dichroism signals. We focus in this work on 148 to onsite 3d - 4p orbital mixing. 121 the effect of the magnetic dichroism on the linear angular 149 122 123 124 126 127 128 130 to analyze the RIXS spectra. 131

132 133 the horizontal ( $\phi = 0^{\circ}$ ) and vertical ( $\phi = 90^{\circ}$ ) config- 160 nantly from excitations to the  $t_{2q}$  orbitals, that are 90° 134 urations show a broad single pre-edge peak (Fig. S9a 161 angular shifted w.r.t. the second and third features cor- $_{135}$  and b), it is possible to identify three main features in  $_{162}$  responding to excitations dominantly into the  $e_q$  orbitals

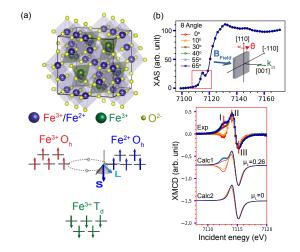


FIG. 1. (a) The unit cell of  $Fe_3O_4$  and the magnetic coupling between the Fe sites. Octahedral  $(O_h)$  Fe<sup>3+</sup> and Fe<sup>2+</sup> ions are antiferromagnetically coupled to the tetrahedral  $(T_d)$  Fe<sup>3+</sup> ions. (b) Fe K-edge measurements in  $Fe_3O_4$  single crystal. The top panel shows XAS results as a function of the sample azimuthal angle  $\theta$ . The bottom panel shows the corresponding XMCD experimental (dotted) and theoretical (solid) Fe preedge signal. Two model calculations are presented: i)- Calc 1 is the optimized calculation where a partially quenched orbital magnetic moment of  $0.26 \,\mu_{\rm B}$  per unit formula of Fe<sub>3</sub>O<sub>4</sub> was concluded, and *ii*)- Calc 2 is the theoretically expected XMCD signal for a fully quenched orbital magnetic moment scenario.

be completely suppressed in the case that  $\mu_L = 0 \mu_B$  (see 137 calculation of the RIXS dichroism signal shows that the Fig. 1b Calc 2) strongly supporting the presence of a  $_{138}$  quadrupolar contributions of the Fe<sup>3+</sup>  $T_d$  and  $O_h$  sites 139 nearly cancel out (see Fig. S10) and hence the three fea-To investigate possible non-collinearity of the orbital 140 tures labelled in Fig. 2b are dominantly attributed to the magnetic moment, we performed comprehensive 1s2p 141 formal Fe<sup>2+</sup> ions. This is consistent with the XMCD re-RIXS-MLAD measurements. This complements the av- 142 sults, where we found that the angular dependence is only erage projected result obtained from XMCD. The RIXS- 143 visible at the Fe<sup>2+</sup> ions. The calculated RIXS dichroic MLAD was measured by rotating the sample about the 144 plane of the formal  $Fe^{2+}$  ions reproduces the three main incident wave-vector direction  $(k_{in})$  aligned with the 145 spectral features. The calculation only misses a weak fea-[110] direction (refer to Fig. 2a). This implies that the 146 ture at incident energy  $\sim 7115.1 \,\mathrm{eV}$  which is associated RIXS-MLAD includes contributions from both structural 147 with the electric dipole transition at the Fe<sup>3+</sup> A site due

The full  $360^{\circ}$  experimental (theoretical) angular dedistribution of RIXS. We initially measured the RIXS- 150 pendence of the three main spectral features can be seen MLAD with the magnetic field *nearly* parallel to  $k_{in}$  as  $_{151}$  in Fig. 2c. The angular dependence is twofold and a a reference measurement. In this case the magnetic field 152 90° angular shift is observed between the first feature is oriented along a high symmetry crystallographic direc- 153 and both the second and third features. A first explation and the angle between the linear incident polariza- 154 nation of the general angular dependence can be protion  $(\epsilon_{in})$  and the magnetic field nearly does not change 155 vided by analysing the  $1s^13d^7$  intermediate states. Proas a function of the sample rotation. These choices sim- 156 jections of the intermediate states associated with nonplify the angular dependence and serve as a benchmark 157 zero transition matrix elements onto cubic crystal field 158 (*i.e.*  $O_h$ ) configurations were calculated using the pro-Although the experimental RIXS planes measured in 159 gram CTM4DOC [29]. The first feature arises domi-

[28]. We note that the angular dependence is anisotropic 218 An energy shift of  $\sim 0.2 \,\mathrm{eV}$  was found between the two 163 where the intensity of the third feature at  $\phi = 180^{\circ}$  is 219 subclasses.

165 smaller than that at  $\phi = 0^{\circ}$ . This is related to the 220 166 167 168 169 170 171 172 173 be attributed to this fact. 174

We examined the coupling of the spin and orbital de-175 grees of freedom by displacing the magnetic field  $50^\circ$ 176 from the high symmetry [001] direction (refer to Fig. 3a). 177 The orientation of the magnetic field corresponds to 178 the  $\left[\frac{-\cos(40^\circ)}{\sqrt{2}}, \frac{\cos(40^\circ)}{\sqrt{2}}, \sin(40^\circ)\right]$  direction. Orienting the magnetic field in a low symmetry direction aligns the net 179 180 spin magnetic moment parallel to the field. If the orbital 181 magnetic moment is not fully quenched, it consequently 182 re-aligns towards the low symmetry direction. The final 183 orientation of the net magnetic moment depends on the 184 strength of the competing interactions such as magnetic 185 exchange, spin-orbit coupling and distortion. Hence, the 186 angular shift of the maximum intensity of the excitations 187 188 can be used to quantify magnetic-moment-induced dis-189 tortion of the electron cloud. Based on this concept, we <sup>190</sup> investigated the orbital magnetic moment of the formal 191  $\operatorname{Fe}^{2+}$  ions. A careful analysis of the full 360° angular <sup>192</sup> dependence exhibits a peculiar 10° angular shift of the maximum intensity between the second and third fea-193 tures in Fig. 3b. 194

195 presented in Fig. 3c. The model captures the essential 252 our observation. 196 aspects of the angular dependence and in particular the 253 197 198 199 200 201 202 203 204 205 206 respect to the local trigonal distortion varies between the 204 ray diffraction and optical reflectivity [41]. 208 four sites leading to anisotropic effects and generates four 265 209 210 211 212 213 <sup>214</sup> ing one subclass and sites 3 and 4 forming the other (see <sup>270</sup> trated in Fig. 1b and Fig. 4f), however the RIXS-MLAD <sup>215</sup> Fig. 4b). It is only when the dynamical distortion effect <sup>271</sup> measurement demonstrates that the average quantity is 216 is taken into consideration that the experimental RIXS- 272 not sufficient to describe the orbital magnetic moment <sup>217</sup> MLAD angular shift can be reproduces (see Fig. S13). <sup>273</sup> in Fe<sub>3</sub>O<sub>4</sub>. This is a result of the non-collinear orbital

We interpret the formation of these two subclasses as anisotropy in the detection in combination to a small 221 a result of a dynamical Jahn-Teller distortion at the Fe misalignment angle ( $\delta$ ) of the magnetic field relative to 222 B sites. The magnitude of the static trigonal distortion the rotation axis *i.e.* the detector position relative to  $_{223}$  lies close within the phonon energies of Fe<sub>3</sub>O<sub>4</sub> [30, 31], the sample as discussed in details in the Supplementary. 224 the magnetic exchange interaction and spin-orbit cou-In addition, our theoretical model assumes that the de- 225 pling, leading to a situation where electron-phonon intection system is a single point while in reality four Ge 226 teraction, dynamical Jahn-Teller and Kugel-Khomskii in-(440) analyzer crystals were used. The minor discrep- 227 teractions all play a role in determining the low energy ancy between the experiment and the calculations could 228 state. We treat this dynamical variation of the distor-229 tion in a first approximation as a small change in the 230 bond lengths over the four sites giving rise to a small en-231 ergy shift. This is a reasonable approximation because <sup>232</sup> the electronic structure adapts almost instantaneously to  $_{233}$  the crystallographic structure (*i.e.* the electronic motion <sup>234</sup> is much faster than the nuclear motion). In this case, the 235 effect of phonons could be simulated as a static distribu-<sup>236</sup> tion of bond lengths leading to a shift in energy between <sup>237</sup> the four sites. This is a common practice in XAS theory 238 as can be found in the paper by Nemeusat *et al.* [32] 239 where thermal fluctuations are simulated by a well cho-240 sen series of configurations. Although theoretical studies 241 that treat simultaneously the electronic and the lattice 242 degrees of freedom are required to comprehend the pre-243 cise effect of the dynamical distortion, we point out that 244 numerous theoretical works concluded the essential role <sup>245</sup> of the strong electron-phonon coupling in the presence of <sup>246</sup> strong electron correlations leading to dynamical Jahn-<sup>247</sup> Teller distortion and the creation of polarons [33-36]. In <sup>248</sup> particular, Piekarz et. al. [33, 34] identified the highly <sup>249</sup> dispersive  $X_3$  phonon mode as a primary order param-<sup>250</sup> eter of the Verwey transition which splits the four Fe B Theoretical calculation of the angular dependences are 251 sites into two subclasses. This agrees rather well with

We have undergone the task of simulating various 10° angular shift of the maximum intensity. The an- 254 X-ray spectroscopic measurements on the basis of our gular shift ( $\Omega$ ) quantified by fitting the angular depen- 255 model. In particular, we focused on comparing  $L_3$ dence to a  $\cos^2(\phi + \Omega)$  of the three features is reported 256 XMCD [2, 7] and  $L_3$  RIXS [7] measurements to our simuin Tab. S4. The anisotropy of the angular dependence is 257 lations. Our model can reproduce the experimental data not well reproduced, likely due to a small misalignment 255 and notably it captures the recently reported  $L_3$  RIXS of the magnetic field that has not been included in the 250 angular dependence well. The existence of this dynamcalculations (see Supplementary). It is now important to 200 ical distortion is furthermore supported by various exhighlight the key ingredients responsible for this angular 201 perimental work such as diffuse scattering experiments shift. The first factor is the static trigonal distortion. 262 using both neutrons [37] and X-rays [38], EXAFS [39], The relative orientation of the exchange interaction with 203 anomalous phonon broadening [40], and pump-probe X-

The presence of four non-equivalent Fe B sites in the non-equivalent Fe B sites. The theoretical RIXS-MLAD 266 high temperature phase has rather interesting implicafor the four sites are shown in Fig. S12. The second fac- 267 tions. Overall, we find that the average orbital magnetic tor is the effect of dynamical distortion that produces two 266 moment deduced by XMCD and RIXS-MLAD is the subclasses of the Fe B sites, namely, sites 1 and 2 form-  $_{269}$  same  $(0.26 \pm 0.03 \,\mu_{\rm B})$  per unit formula of Fe<sub>3</sub>O<sub>4</sub> as illus-

(a) (b) (C) I (arb. unit) [001]  $\Phi = 0^{\circ}$ 0.0 10 [-110] (90°) -Φ(0° Φ 712 [110] 180 710 -1-2-3225 708 270 Φ (deg) Norm, I (arb. unit) Φ= 90° [-110] 90 712 710 [1101 [001] 708 225 7112 7114 7116 270 Incident energy (eV) Φ (deg)

FIG. 2. Fe 1s2p RIXS-MLAD measurements. (a) A sketch of the scattering geometry employed. The magnetic field  $(B_{Field})$ is aligned nearly parallel to the incident wave-vector  $(k_{in})$  which corresponds to the [110] direction. (b) Experimental and theoretical dichroism RIXS planes computed as the difference between the RIXS plane at  $\phi = 90^{\circ}$  and at  $\phi = 0^{\circ}$ . The full experimental (dotted) and theoretical (solid) 360° RIXS-MLAD signals of the features labelled 1, 2 and 3 in the RIXS dichroic maps are shown in (c) respectively. The angular dependence signal is normalized as:  $RIXS - MLAD = \frac{RIXS(\phi) - Min[RIXS(\phi)]}{Max[RIXS(\phi) - Min[RIXS(\phi)]]}$ 

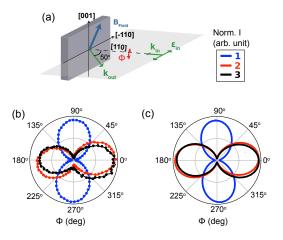


FIG. 3. Fe 1s2p RIXS-MLAD measurements. (a) A sketch of the scattering geometry employed. The magnetic field  $(B_{Field})$  is aligned to the  $\left[\frac{-\cos(40^\circ)}{\sqrt{2}}, \frac{\cos(40^\circ)}{\sqrt{2}}, \sin(40^\circ)\right]$  direction tion. The angular dependence is measured by rotating the lated (solid) angular dependence of the three main features (labeled 1, 2 and 3) are shown in panels (b) and (c).

274 ordering arising from the interplay between trigonal dis- 300 ments sensitive to the non-averaged quantity. This or-

275 tortion effects (static and dynamical), spin-orbit coupling  $_{276}$  and exchange interaction at the formal  $\mathrm{Fe}^{2+}$  ions. The  $_{277}$  orbital magnetic moment per  $Fe^{2+}$  ion is predicted to have a strong dependence on the magnetic field in con-278 trary to the spin magnetic moment which is collinear to 279 the magnetic field. Fig. 4 illustrates the dependence of 280 the orbital magnetic moments on the orientation of the 281 magnetic field when we rotate it about the [110] direc-282 tion for the four sites independently. Large non-collinear 283 orbital contributions that tilt as much as  $82^{\circ}$  away from 284 the magnetic spin moment orientation are present. Fur-285 286 thermore, the collinear contribution per site ranges from 0 to 150% of the average quantity as a function of the 287 orientation of the field. Remarkably, the average orbital 288 magnetic moment for the four sites remains nearly con-289 stant (Fig. 4f). 290

The large discrepancies regarding the orbital magnetic 291 <sup>292</sup> moment of Fe in  $Fe_3O_4$  can now be understood in light of sample about the [110] direction ( $\phi$  rotation).  $\phi = 0^{\circ}$  is de- 293 the large non-collinear contribution, the site dependency fined when the incident polarization vector  $(\epsilon_{in})$  is aligned to 294 and the magnetic field angular dependence. Experiments the [-110] direction. The experimental (dotted) and calcu- 295 sensitive to the effective orbital magnetic moment yield 296 different results to those sensitive to the projected av-297 erage quantity, or the average of the squared projected 298 quantity. Moreover, variations as a function of the ori-299 entation of the magnetic field are expected for experi-

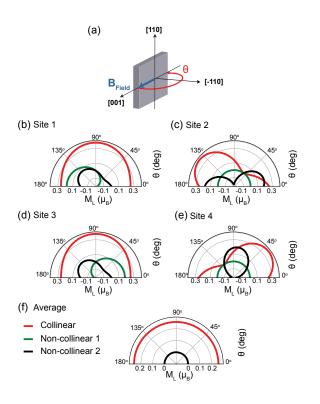


FIG. 4. The angular dependence of the orbital momentum  $(M_L)$  of the four  $\mathrm{Fe}^{2+}$  ions as a function of the rotation of 348 the magnetic field  $(B_{Field})$  about the [110] orientation. (a) 349 Sketch of the rotation geometry. The angular dependence of the orbital momentum projected along the direction of  $B_{Field}$ 351 and two perpendicular non-collinear contributions are shown 352 in panel (b), (c), (d) and (e). The average orbital magnetic 353 moments of the four  $Fe^{2+}$  ions are shown in (f).

301 dering of the orbital magnetic moment is predicted to 357 be short-ranged due to the dynamical distortions at the 358 302  $_{303}$  high temperature phase of  $Fe_3O_4$ . The combination of  $^{359}$ <sup>304</sup> 1s2p RIXS-MLAD and XMCD provides a powerful tool <sup>360</sup> 305 to quantify site-selectively non-collinear magnetic ordering with bulk sensitivity. Finally, we show that the or-306 363 bital degree of freedom is an important precursor for the 364 307 Verwey transition in  $Fe_3O_4$  given the fact that it is cou-308 pled to a primary order parameter. 309

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- [1] E. Goering, M. Lafkioti, S. Gold, and G. Schütz, J. Magn. 331 Magn. Mater. **310**, e249 (2007). 332
- E. Goering, S. Gold, M. Lafkioti, and G. Schütz, Euro-[2]333 phys. Lett. 73, 97 (2006). 334
- [3] D. J. Huang, C. F. Chang, H.-T. Jeng, G. Y. Guo, H.-J. 335 Lin, W. B. Wu, H. C. Ku, A. Fujimori, Y. Takahashi, 336 and C. T. Chen, Phys. Rev. Lett. 93, 077204 (2004).
  - E. Goering, Phys. Status Solidi B 248, 2345 (2011). [4]
  - E. Arenholz, G. van der Laan, R. V. Chopdekar, and [5]Y. Suzuki, Phys. Rev. B 74, 094407 (2006).
  - [6] H. Y. Huang, Z. Y. Chen, R. P. Wang, F. M. De Groot, W. B. Wu, J. Okamoto, A. Chainani, J. S. Zhou, H. T. Jeng, G. Y. Guo, et al., Nat. Commun. 8, 15929 (2017).
  - [7] H. Elnaggar, R. P. Wang, S. Lafuerza, E. Paris, A. C. Komerak, H. Guo, Y. Tseng, D. Mcnelly, F. Frati, M. W. Haverkort, et al., arXiv:1811.04836 [cond-mat.strel] (2018).
  - [8] Y. Li, P. A. Montano, B. Barbiellini, P. E. Mijnarends, S. Kaprzyk, and A. Bansil, J. Phys. Chem. Solids 68, 1556 (2007).
  - [9] J. A. Duffy, J. W. Taylor, S. B. Dugdale, C. Shenton-Taylor, M. W. Butchers, S. R. Giblin, M. J. Cooper, Y. Sakurai, and M. Itou, Phys. Rev. B 81, 134424 (2010).
- [10]I. Leonov, A. N. Yaresko, V. N. Antonov, M. A. Korotin, 354 and V. I. Anisimov, Phys. Rev. Lett. 93, 146404 (2004).
- M. Coey, Nature 430, 155EP (2004). 356 111
  - [12]Y. Tokura and N. Nagaosa, Science 288, 462 (2000).
  - [13]P. G. Radaelli, New J. Phys. 7, 53 (2005).
  - [14]J. Schlappa, C. Schüßler-Langeheine, C. F. Chang, H. Ott, A. Tanaka, Z. Hu, M. W. Haverkort, E. Schierle, E. Weschke, G. Kaindl, et al., Phys. Rev. Lett. 100, 026406 (2008).
  - [15]S. B. Wilkins, S. Di Matteo, T. A. W. Beale, Y. Joly, C. Mazzoli, P. D. Hatton, P. Bencok, F. Yakhou, and V. A. M. Brabers, Phys. Rev. B 79, 201102 (2009).
- A. Tanaka, C. F. Chang, M. Buchholz, C. Trabant, 366 [16]E. Schierle, J. Schlappa, D. Schmitz, H. Ott, P. Metcalf, L. H. Tjeng, et al., Phys. Rev. Lett. 108, 227203 (2012).
  - A. Tanaka, C. F. Chang, M. Buchholz, C. Trabant, [17]E. Schierle, J. Schlappa, D. Schmitz, H. Ott, P. Metcalf, L. H. Tjeng, et al., Phys. Rev. B 88, 195110 (2013).
  - [18] D. Dunlop and O. Özdemir, Rock magnetism (Cambridge Univ. Press., 1997)
  - E. J. W. Verwey, Nature 144, 327 (1939).
  - [20]N. F. Mott, Rev. Mod. Phys. 40, 677 (1968).
- 377 [21] M. Imada, A. Fujimori, and Y. Tokura, Rev. Mod. Phys.

70, 1039 (1998). 378

- P. Carra, B. T. Thole, M. Altarelli, and X. Wang, Phys. 404 [22]379 Rev. Lett. 70, 694 (1993). 380
- D. I. Khomskii and M. V. Mostovoy, J. Phys. A: Math. 406 [23]381 Gen. 36, 9197 (2003). 407 382
- [24]H. Uzu and A. Tanaka, J. Phys. Soc. Jpn. 77, 074711 408 383 (2008).409 384
- M. W. Haverkort, M. Zwierzycki, and O. K. Andersen, 410 [25]385 Phys. Rev. B 85, 165113 (2012). 386
- [26] Y. Lu, M. Höppner, O. Gunnarsson, and M. W. 412 387 Haverkort, Phys. Rev. B 90, 085102 (2014). 413 388
- [27] M. W. Haverkort, G. Sangiovanni, P. Hansmann, 414 389 A. Toschi, Y. Lu, and S. Macke, EPL 108, 57004 (2014). 415 390
- [28] T. E. Westre, P. Kennepohl, J. G. DeWitt, B. Hedman, 416 391
- K. O. Hodgson, and E. I. Solomon, J. Am. Chem. Soc. 417 392 119, 6297 (1997). 393 418
- [29] M. U. Delgado-Jaime, K. Zhang, J. Vura-Weis, and 419 394 F. M. F. De Groot, J. Synchrotron Radiat. 23, 1264 420 395 (2016).396 421
- [30] L. V. Gasparov, D. B. Tanner, D. B. Romero, H. Berger, 422 397
- 398 G. Margaritondo, and L. Forró, Phys. Rev. B 62, 7939 423 (2000).399 424
- [31] B. Handke, A. Kozlowski, K. Parlinnski, J. Przewoznik, 425 400
- T. Slezak, A. I. Chumakov, L. Niesen, Z. Kakol, and 426 401
- J. Korecki, Phys. Rev. B 71, 144301 (2005). 402

[32] R. Nemausat, D. Cabaret, C. Gervais, C. Brouder, 403 N. Trcera, A. Bordage, I. Errea, and F. Mauri, Phys. Rev. B 92, 144310 (2015).

405

411

427

- [33]P. Piekarz, K. Parlinski, and A. M. Oles, Phys. Rev. Lett. 97, 156402 (2006).
- [34] P. Piekarz, K. Parlinski, and A. M. Oles, Phys. Rev. B 76, 165124 (2007).
- S. Borroni, G. S. Tucker, F. Pennacchio, J. Rajeswari, [35]U. Stuhr, A. Pisoni, J. Lorenzana, H. M. RØnnow, and F. Carbone, New J. Phys. **19**, 103013 (2017).
- [36]J. Cumby and J. P. Attfield, Nat. Commun. 8, 14235 EP (2017).
- [37]Y. Yamada, N. Wakabayashi, and R. M. Nicklow, Phys. Rev. B 21, 4642 (1980).
- A. Bosak, D. Chernyshov, M. Hoesch, P. Piekarz, [38]M. Le Tacon, M. Krisch, A. Kozłowski, A. M. Oleś, and K. Parlinski, Phys. Rev. X 4, 011040 (2014).
- [39]G. Subías, J. García, and J. Blasco, Phys. Rev. B 71, 155103 (2005).
- [40]M. Hoesch, P. Piekarz, A. Bosak, M. Le Tacon, M. Krisch, A. Kozłowski, A. M. Oleś, and K. Parlinski, Phys. Rev. Lett. 110, 207204 (2013).
- [41] S. de Jong, R. Kukreja, C. Trabant, N. Pontius, C. F. Chang, T. Kachel, M. Beye, F. Sorgenfrei, C. H. Back, B. Bräuer, et al., Nat. Mater. 12, 882EP (2013).