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F. Kwabia-Tchana, Y. Attafi, L. Manceron, D. Doizi, J. Vander Auwera, et al.. Line intensities for the  $\nu 6$  and  $2\nu 3$  bands of methyl iodide (12CH3I). Journal of Quantitative Spectroscopy and Radiative Transfer, 2019, 222-223, pp.130-137. 10.1016/j.jqsrt.2018.10.001. hal-02340420

# HAL Id: hal-02340420 https://hal.science/hal-02340420

Submitted on 30 Oct 2019

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# Line intensities for the v<sub>6</sub> and 2v<sub>3</sub> bands of methyl iodide (<sup>12</sup>CH<sub>3</sub>I)

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Nb of figures: 6

Nb of Tables: 4

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#### Abstract

The goal of this study is to measure for the first time absolute line intensities for the  $v_6$  band of methyl iodide (CH<sub>3</sub>I) centered at 892.918 cm<sup>-1</sup>. High-resolution Fourier transform spectra were recorded at various pressures for the whole 500 – 1450 cm<sup>-1</sup> spectral range. Using these spectra, a large set of CH<sub>3</sub>I individual line intensities was measured for the  $v_6$  band. These experimental intensities were least squares fitted to derive the expansion of the  $v_6$  transition moment operator. The theoretical model used to describe the line positions and intensities accounts for the hyperfine structure in the 6<sup>1</sup> and ground states and for the vibration-rotation resonances that couple the 6<sup>1</sup> energy levels with those of the 3<sup>2</sup> and 2<sup>1</sup> vibrational states [A. Perrin *et al.*, J. Mol. Spectrosc. 324 (2016) 28 – 35]. As the 2 $v_3$  band is extremely weak, its associated transition moment operator was estimated from band strength available in the literature. A comprehensive list of line positions and intensities was generated for the  $v_6$  and  $2v_3$  bands of CH<sub>3</sub>I at 11 µm, which should be useful for the possible detection of this species by the future IASI-NG satellite instrument (Infrared Atmospheric Sounding Interferometer New Generation), now under preparation (https://iasi-ng.cnes.fr/en/IASI-NG/index.htm).

# *Keywords:* methyl iodide; CH<sub>3</sub>I; high-resolution Fourier transform spectroscopy; Coriolis resonances; line positions; line intensities; hyperfine structure

# A. Introduction

The goal of the present study is to generate a list of spectroscopic line parameters for methyl iodide (CH<sub>3</sub>I) in the 11  $\mu$ m region. This work focuses on line intensities for the v<sub>6</sub> band of <sup>12</sup>CH<sub>3</sub>I and follows our previous investigation that concerned line positions for the same band [1]. Although absent in the public access spectroscopic databases such as HITRAN (http://hitran.org/) [2] or GEISA [3], this band located at 892.916 cm<sup>-1</sup>, coincides with the 11  $\mu$ m transparency window of the atmosphere [4] and could therefore be a good candidate for the detection of atmospheric CH<sub>3</sub>I, for example, by the future IASI-NG satellite instrument

(Infrared Atmospheric Sounding Interferometer New Generation), now under preparation (https://iasi-ng.cnes.fr/en/IASI-NG/index.htm).

# Status of the v<sub>6</sub> line positions for <sup>12</sup>CH<sub>3</sub>I

As far as the line positions are concerned, methyl iodide was the subject of numerous studies. Only the literature relevant to the present investigation and the  $^{12}CH_3I$  isotopic species will be described here. A more complete review can be founded elsewhere [1, 5-13].

Relying on Fourier transform spectra, Alanko et al. [9], Paso and Alanko [10], and Haykal et al. [13] performed a very detailed analysis of the  $2v_3$  and  $v_6$  bands of  ${}^{12}CH_3I$  and identified a large set of transitions in both bands. However, assignments in the  $v_6$  band were complicated by the existence of hyperfine structures resulting from the large <sup>127</sup>I iodine nuclear quadrupole moment (nuclear spin I = 5/2). These hyperfine splittings are easily observable for  $v_6$  transitions involving J values not much larger than K (J  $\leq 2 \times K$ ). The energy level calculations performed by Paso and Alanko [10] accounted explicitly for the weak firstorder Coriolis resonance coupling the  $6^1$  and  $3^2$  energy levels, while the hyperfine structure was accounted for by a perturbative treatment. Later, Carocci et al. performed a new investigation of the rotational structure of the ground and 6<sup>1</sup> vibrational states of <sup>12</sup>CH<sub>3</sub>I using a Doppler-free double resonance technique [12]. The results of these rotational measurements were combined with a large set of literature data obtained by microwave techniques [5, 6, 8, 12] and from infrared investigations of the  $v_6$  band [10] to get complete sets of rotational and hyperfine parameters for both the ground and the 6<sup>1</sup> excited states of <sup>12</sup>CH<sub>3</sub>I [12]. However, the existence of a  $6^1$  and  $3^2$  Coriolis resonance was not mentioned in this latter study [12], and therefore not accounted for explicitly. Finally, our recent and extensive investigation [1] of the  $v_6$  and  $2v_3$  bands of  ${}^{12}CH_3I$  was performed using high resolution spectra recorded using Bruker IFS125HR Fourier transform spectrometers (FTS) available at the AILES beamline of the SOLEIL Synchrotron facility and at LISA in Créteil (France). The assignment procedure and linelist calculation performed in that work explicitly considered the hyperfine structure for the  $v_6$  and  $2v_3$  bands using literature values of the nuclear quadrupole and spin rotational constants in the ground,  $(6^1; \ell = 1)$  and  $(3^2; \ell = 0)$  states [8,12]. A large set of  $6^1$  and  $3^2$ experimental rotational energy levels was deduced from the observed rotational or hyperfine transitions, after removing from the latter the contribution of the individual line shift due to the hyperfine structure. These levels were satisfactorily reproduced using an effective Hamiltonian model involving the {6<sup>1</sup>, 3<sup>2</sup>, 2<sup>1</sup>} interacting states, thus accounting for the weak Coriolis type resonances coupling the 6<sup>1</sup> energy levels with those of the  $(3^2; \ell = 0)$  and  $(2^1; \ell = 0)$  states. Indeed, for series involving rather high rotational quantum numbers, additional Coriolis resonances coupling the  $(6^1; \ell = 1)$  and  $(2^1; \ell = 0)$  energy levels were observed for the first time [1]. Note that our recent study [1] investigated the strong v<sub>6</sub> band only, considering the v<sub>2</sub> band as a "perturber" of the v<sub>6</sub> band. An extended analysis using an effective model that involved the  $\{(2^1; \ell = 0), (3^16^1; \ell = 1), (5^1; \ell = 1)\}$  interacting states, thus including the v<sub>2</sub> band, had been performed in the frame of the very thorough investigation of the v<sub>2</sub>/v<sub>5</sub>/v<sub>3</sub>+v<sub>6</sub> band system of <sup>12</sup>CH<sub>3</sub>I and <sup>13</sup>CH<sub>3</sub>I [11].

# Status of the $v_6$ line intensities for ${}^{12}CH_3I$

To the best of our present knowledge, line intensity measurements were never carried out for the  $v_6$  or  $2v_3$  bands of  ${}^{12}CH_3I$ . Measurements of low-resolution band intensities have however been reported for the six fundamental bands of  ${}^{12}CH_3I$  [14,15], and absorption cross sections have been measured at Pacific Northwest National Laboratory (PNNL) [16]. *Ab initio* computed band intensities have been reported for  ${}^{12}CH_3I$  by Schneider and Thiel [17].

# B. Experimental procedure and wavenumbers calibration

# **B.1.** Experimental procedure

Seven absorption spectra of methyl iodide have been recorded in the range from 500 to  $1450 \text{ cm}^{-1}$  using the high-resolution Bruker IFS125HR FTS located at the LISA facility in Créteil (France). The instrument was equipped with a silicon carbide Globar source, a KBr/Ge beamsplitter and a liquid nitrogen cooled HgCdTe (MCT) detector. An optical filter with a bandpass of  $500 - 1450 \text{ cm}^{-1}$  was to improve the signal-to-noise ratio. The FTS was continuously evacuated below  $3 \times 10^{-4}$  hPa by a turbomolecular pump to minimize absorption by atmospheric gases. The diameter of the entrance aperture of the spectrometer was set to 1.5 mm to maximize the intensity of infrared radiation falling onto the MCT detector without saturation or loss of spectral resolution. Interferograms were recorded with a 40 kHz scanner frequency and a maximum optical path difference (MOPD) of 473.68 cm. According to the

Bruker definition (resolution = 0.9/MOPD), this corresponds to a resolution of  $0.0019 \text{ cm}^{-1}$ . The spectra were obtained by Fourier transformation of the interferograms using a Mertz phase correction with a 1 cm<sup>-1</sup> resolution, a zero-filling factor of 2 and no apodization (boxcar option).

The sample was purchased from Sigma Aldrich (FlukaChemie GmbH, 99% purity) and used without further purification. For all the measurements, the sample was contained in a Pyrex glass cell providing an absorption path length of  $12.4 \pm 0.1$  cm and equipped with wedged ZnSe anti-reflection coated windows. The sample pressure in the cell was measured using calibrated MKS Baratron capacitance manometers models 628D (13.332 hPa full scale) and 627D (133.32 hPa full scale) characterized by a stated reading accuracy of 0.12%. Considering the uncertainty arising from small variations of the pressure during the recording (~ 0.35%), we estimated the measurement uncertainty on the pressure to be equal to 0.5%. All the spectra were recorded at a stabilized room temperature of  $295 \pm 1$  K.

The following procedure was used to record the spectra. A background spectrum was first collected while the cell was being continuously evacuated. It was recorded at the same resolution as the sample spectra to ensure proper removal of the H<sub>2</sub>O or CO<sub>2</sub> absorption lines and of the weak channeling generated by the ZnSe cell windows. The infrared gas cell was then passivated several times with the CH<sub>3</sub>I sample. Finally, spectra were recorded for seven different sample pressures of methyl iodide. The seven pressures chosen and the number of interferograms recorded and averaged to yield the corresponding spectra are listed in Table 1. All the sample spectra were ratioed against the empty cell background spectrum. The root mean square (RMS) signal-to-noise ratio (S/N) in the ratioed spectra ranged between 100 and 300.

#### **B.2.** Wavenumbers calibration

The spectra were calibrated by matching the measured positions of about 25 lines of residual  $CO_2$  observed therein to reference wavenumbers available in HITRAN [2] with a RMS deviation of 0.00013 cm<sup>-1</sup>. The absolute accuracy of the measured CH<sub>3</sub>I line positions was estimated as the square root of the sum of squares of the accuracy of the reference  $CO_2$  line positions in HITRAN (better than 0.0001 cm<sup>-1</sup>) plus twice the RMS deviation between the observed frequencies and values listed in HITRAN2016 for CO<sub>2</sub>, yielding in total 0.00028 cm<sup>-1</sup>.

# C. Line intensity measurements

### C.1. Retrieval of line intensities

The line intensities were measured using a multi-spectrum fitting program developed in Brussels [18, 19]. This program adjusts spectroscopic (e.g. line positions, intensities and widths) and spectrum-specific (e.g. baseline) parameters to best-fit synthetic spectra to the observed spectra using a Levenberg-Marquardt non-linear least squares fitting procedure. Each synthetic spectrum, interpolated 4 times with respect to the observed spectrum, is calculated as the convolution of the molecular transmission spectrum with an instrument line shape function, which includes the effects of the finite maximum optical path difference and of the finite source aperture diameter of the interferometer. In the present work, no distortion of the instrument line shape was observed: the nominal aperture diameter of 1.5 mm was therefore retained in the simulated spectra. The background in each spectrum was represented by a polynomial expansion up to the second order (a constant or an affine function was however found sufficient in most cases), and the profile of the lines was modeled using a Voigt function [20] with Gaussian width always held fixed to the value calculated for the Doppler broadening. The measurements were carried out on small spectral intervals, ranging from 0.11 to 0.7  $\text{cm}^{-1}$  and containing one to several lines, and involved the simultaneous fitting of all the spectra. For the line intensity measurements, we selected carefully the transitions for which the hyperfine structure is not observable: this means that the hyperfine broadening is weaker that the Doppler line width. The fitted line parameters were the positions, intensities and self-broadening coefficients (assumed to be proportional to the pressure of methyl iodide). Self-shift of the lines was ignored. The required initial values of the positions and intensities of lines belonging to the  $v_6$  band of methyl iodide considered in the present work were generated from our previous frequency analysis [1]. The selfbroadening coefficient of all the lines was initially set to 0.2 cm<sup>-1</sup> atm<sup>-1</sup>. Figure 1 presents an example of the results of a fit. It involved 14 lines observed in the spectral range 915.3 to 915.8 cm<sup>-1</sup>, pressures ranging from 3.406 to 28.70 hPa and a total of 44 fitted parameters. The upper and lower panels show the observed spectra and best-fit residuals, respectively. The absence of signatures out of the spectral noise in the residuals suggests that the Voigt profile is appropriate to fit the observed lines, whose S/N are about 100, as in the region presented in figure 1.

# C.2. Uncertainty analysis

Examination of the residuals of the fits shows that they are generally less than 1%. However, to estimate the accuracy of the measured intensities requires considering the uncertainties on the physical parameters, contributions from possible systematic errors [21] as well as the uncertainties derived from the fits, taken as the standard deviation. The various sources of error considered in the present work and their associated uncertainties expressed relative to the line intensities are given in Fig. 2 for 3 selected strong, medium and weak lines, representative of the 840 measured lines, i.e.  ${}^{R}R(26,0)$ ,  ${}^{P}P(30,4)$  and  ${}^{P}P(57,4)$ . Figure 2 shows that systematic errors and the standard deviations of the fits are the main sources of error. The dominant contributions to the systematic errors were found to arise from the location of the full-scale (100%) photometric level, electronic and detector nonlinearities. Baseline offset results in a corresponding intensity uncertainty. The effect of noise and small offset in these have been discussed in previous works [21]. The 0% level could be checked readily and the 100% level fitted on the spectra in regions without absorptions. An arbitrary, but conservative estimate of 2% of the line intensities has been retained here. For the 840 measured transitions, the average relative uncertainty on the line intensities is 1.5%. For each transition, we then calculated an upper limit for the overall uncertainty based on the maximum uncertainty of the individual experimental parameters, i.e.  $\varepsilon_{si}$  (sample purity),  $\varepsilon_t$  (temperature),  $\varepsilon_p$  (pressure),  $\varepsilon_{pl}$ (pathlength),  $\varepsilon_{\rm fit}$  (standard deviation from fit) and  $\varepsilon_{\rm sys}$  (systematic errors), assuming that these uncertainties are uncorrelated:

$$\varepsilon = \sqrt{\varepsilon_{si}^2 + \varepsilon_l^2 + \varepsilon_{pl}^2 + \varepsilon_{pl}^2 + \varepsilon_{fit}^2 + \varepsilon_{sys}^2} \qquad \text{Eq. (1)}$$

Using this equation and the experimental relative uncertainties involved, the estimated overall uncertainty (estimated accuracy) for each of the 840 line intensities measured was calculated. These uncertainties ("Unc") are provided in Table 2. On average, the estimated accuracy for line intensities is equal to 3%.

# D. Line intensity calculations

The integrated absorption cross section  $k_{\tilde{v}}^{N}$  [in cm<sup>-1</sup>/(molecule cm<sup>-2</sup>)] of a line of CH<sub>3</sub>I in natural abundance corresponding to a transition between a lower level *L* of energy *E*<sub>L</sub> and an upper level *U* of energy *E*<sub>U</sub> can be expressed as [2]:

$$k_{\tilde{\nu}}^{N} = I_{a} \frac{8\pi^{3}\tilde{\nu}}{4\pi\epsilon_{0} 3hc} \frac{1}{Z(T)} g_{nucl}^{I} g_{nucl}^{H} \left(1 - \exp\left(-\frac{hc\tilde{\nu}}{kT}\right)\right) \exp\left(-\frac{hcE_{L}}{kT}\right) R_{L}^{U} \qquad \text{Eq. (2)}$$

In Eq. (2), *T* is the temperature (in Kelvin),  $\tilde{v} = (E_U - E_L)/\text{hc}$  is the line position (in cm<sup>-1</sup>), I<sub>a</sub> = 0.98896 (8) is the isotopic abundance of <sup>12</sup>CH<sub>3</sub>I [22], and  $g_{\text{nucl}}^{\text{I}}$  and  $g_{\text{nucl}}^{\text{H}}$  are the statistical weights of the lower level respectively resulting from the iodine (I = 5/2) and hydrogen (I = 1/2) nuclear spins. For <sup>12</sup>CH<sub>3</sub>I,  $g_{\text{nucl}}^{\text{H}} = 4$ , irrespective of the A<sub>1</sub>:A<sub>2</sub>:E symmetries. As pointed out previously, the hyperfine structure due to the iodine quadrupole interaction may be observable for transitions involving low J values as compared to K, complicating measurements of individual line intensities. Therefore, the present least squares fit calculation only considered experimental intensities measured for lines free of any observable hyperfine structure. For such transitions,  $g_{\text{nucl}}^{\text{I}} = 6$  for all levels.

 $Z(T) = Z_{vib}(T) \times Z_{rot}(T)$  is the total internal partition function, written as the product of vibrational and rotational contributions. The vibrational part,  $Z_{vib}(T)$ , was computed in the Harmonic Oscillator Approximation (HOA) using the vibrational frequencies available in the literature [1, 7, 11, 23,24]. In this approximation, the  $Z_{vib}(T)$  and the rotational contribution,, $Z_{rot}(T)$  are assumed independent.  $Z_{rot}(T)$ , was calculated using the ground state rotational levels computed up to J=90 with the constants quoted in Ref. [12] and accounting for the nuclear spin statistical weights provided here above. The values of the rotational, vibrational and total partition functions thus calculated between 70 and 400 K are listed in Table A1 (available as supplementary material).

The present work only deals with cold bands originating from the CH<sub>3</sub>I nondegenerate ground state (0;  $\ell$  =0), identified as |0> in the following. In Eq. (2),  $R_L^U$  is the square of the matrix element of the transformed dipole moment operator  $\mu'_Z$ , which takes the following form:

$$\mathbf{R}_{\mathrm{L}}^{\mathrm{U}} = \left| \left\langle \left( \mathbf{v}'; \ell' \right), \mathbf{J}' \mathbf{K}' \middle| \boldsymbol{\mu}_{\mathrm{Z}}' \middle| 0, \mathbf{J}'' \mathbf{K}'' \right\rangle \right|^{2}$$
 Eq. (3)

 $(v'; \ell')$  are the vibrational quantum numbers in the upper state, which belongs to the  $\{(6^1; \ell = 1), (3^2; \ell = 0), (2^1; \ell = 0)\}$  set of interacting states. J''K'' and J'K' are the rotational

quantum numbers in the ground state and in the upper levels of the transition, respectively. The expansions of the  $|(v'; \ell'), J'K'\rangle$  and  $|0, J''K''\rangle$  upper and lower state rovibrational wavefunctions, which depend on the symmetry (A<sub>1</sub>, A<sub>2</sub> or E) of the considered energy level, are given in Eqs. (1) to (2) of Ref. [25]. Finally, the dipole moment operator  $\mu'_Z$  can be expanded as follows:

$$\mu'_{\rm Z} = \sum_{\rm v'\in B'} |0\rangle^{\rm v'} \mu_{\rm Z}^{\Delta\ell} \langle {\rm v'}; \ell'| \qquad \text{Eq. (4)}$$

where  ${}^{v'}\mu_Z^{\Delta\ell}$  is the transformed dipole moment operator corresponding to the transition  $|0\rangle \rightarrow |v';\ell'\rangle$ . As  $v_6$  transitions access the  $6^1$  upper vibrational state involved in the  $\{(6^1;\ell=1),(3^2;\ell=0),(2^1;\ell=0)\}$  triad of interacting states, the transition moment operators of the parallel  $2v_3$  and  $v_2$  bands also contribute to the  $v_6$  line intensities, together with their  $v_6$  counterpart. For the  $v_6 \Delta \ell = \pm 1$  perpendicular band, the expansion of the transformed transition dipole moment operator can be written as:

$${}^{6}\mu_{\rm Z}^{\Delta\ell=1} = {}^{6}\mu_{0}^{\Delta\ell=1}\varphi_{\rm x} + {}^{6}\mu_{4}^{\Delta\ell=1}\left\{i\varphi_{\rm y},J_{\rm z}\right\} + {}^{6}\mu_{5}^{\Delta\ell=1}\left\{\varphi_{\rm z},iJ_{\rm y}\right\} \qquad \text{Eq. (5)}$$

This expression is an extension of the asymmetric-top molecules model [26]. Written in the phase convention of Refs. [27-29], it does not differ formally from the approach used in Refs. [28,29]. In Eq. (5),  $\varphi_x$ ,  $\varphi_y$ , and  $\varphi_z$  are the  $Z_x$ ,  $Z_y$ , and  $Z_z$  components of the direction cosines between the Z laboratory fixed axis and the *x*, *y* and *z* molecular axes, and  ${}^6\mu_q^{\Delta\ell=1}$  is the q<sup>th</sup> parameter in the expansion of  ${}^6\mu_Z^{\Delta\ell=1}$ . Because of the absence of information on a possible centrifugal distortion dependence of the line intensities in the  $2v_3$  and  $v_2$  bands, the expansion of the transition moment operators of the parallel  $2v_3$  and  $v_2$  bands,  ${}^{33}\mu_Z^{\Delta\ell=0}$  and  ${}^2\mu_Z^{\Delta\ell=0}$  respectively, was restricted to its 0-th order term:

$$^{33}\mu_Z^{\Delta\ell=0} = {}^{33}\mu_0^{\Delta\ell=0} \phi_z$$
 Eq. (6a)

In Eq. (5), operators with the  $(\Delta \ell = \pm 1; \Delta K = \pm 1)$  selection rules are only considered for the  $v_6$  perpendicular band while the zero-order term in Eq. (6) corresponds to the  $(\Delta \ell = 0; \Delta K = 0)$  rule for the  $2v_3$  or  $v_2$  parallel bands. Indeed, it proved useless to account in the expansion of the  $v_6$ ,  $2v_3$  or  $v_2$  transition moment operators for higher order terms. Some of these symmetry-allowed [28, 29] terms involve different  $(\Delta \ell; \Delta K)$  selection rules than those considered in Eqs. (5) and (6).

As already stated, individual line intensities are not available for the  $2v_3$  or  $v_2$  parallel bands. In spite of this, the relative strengths of the  $v_6/2v_3/v_2$  bands can be estimated to equal  $1/0.0126/2.352 \ (\pm 10\%)$  using the existing experimental band intensities [16]. The zero-order terms in the expansion of the  $v_6/2v_3/v_2$  transition moment operators are therefore in the following ratio (in absolute values):

$$\left|{}^{6}\mu_{0}^{\Delta\ell=\pm1}\right| / \left|{}^{33}\mu_{0}^{\Delta\ell=0}\right| / \left|{}^{2}\mu_{0}^{\Delta\ell=0}\right| \approx 1/0.122/1.53 \ (\pm10\%)$$
 Eq. (7)

The set of 840 line intensities measured for the  $v_6$  band was introduced in least squares fit calculations to determine the parameters involved in the expansion of  ${}^6\mu_Z^{\Delta\ell=1}$  [Eq. (5)]. As the Coriolis resonances between the  $6^1$ ,  $3^2$  and  $2^1$  vibrational levels only result in local perturbations involving a strong band ( $v_6$ ) and weak bands ( $2v_3$  or  $v_2$ ), the question is whether or not the  $2v_3$  or  $v_2$  transition moment operators contribute significantly to the intensities of the  $v_6$  band. Three sets of computations were therefore performed. The first two involved setting the  $v_2$  and  $2v_3$  bands transition moment zero order parameters (i.e.  ${}^2\mu_0^{\Delta\ell=0}$  and  $^{33}\mu_0^{\Delta\ell=0}$  , respectively) in turn to the value predicted from  $\,^6\mu_0^{\Delta\ell=\pm 1}$  as discussed here above, while the last computation was done setting  ${}^2\mu_0^{\Delta\ell=0}$  and  ${}^{33}\mu_0^{\Delta\ell=0}$  to zero. It appeared that these three computations do not differ significantly from each other, either in terms of the quality of the fit or in the values of the derived  ${}^6\mu_q^{\Delta\ell=1}$  parameters. This could be expected because the set of 840 experimental  $v_6$  line intensities used for the present calculation does not include lines affected by the Coriolis resonances involving  $v_2$  and  $2v_3$ . The latter lines were indeed too weak to be observed. It was therefore decided to perform the final computation of the line intensities setting the  $v_2$  transition moment operator to zero and the  ${}^{33}\mu_0^{\Delta\ell=0}$  parameter to the fixed value  ${}^{33}\mu_0^{\Delta\ell=0} \approx 0.122 \times {}^6\mu_0^{\Delta\ell=\pm 1}$  deduced as described here above. The resulting values of the  ${}^6\mu_q^{\Delta\ell=\pm 1}$  and  ${}^{33}\mu_0^{\Delta\ell=0}$  parameters are quoted in Table 2, together with the statistical analysis of the line intensity computation. Figure 3 gives an overview of the results of the calculation, while the complete list of measured and calculated line intensities is provided in Table A2 (available as supplementary material).

#### E. Synthetic spectrum

A comprehensive list of line positions and intensities (at 296 K) was generated for the  $v_6$  and  $2v_3$  bands of  ${}^{12}$ CH<sub>3</sub>I using the vibrational energies, the rotational and coupling constants given in Table 7 of Ref. [1] for the upper interacting states and the rotational constants provided in Ref. [12] for the ground state, together with the values of the  ${}^{6}\mu_{q}^{\Delta\ell=\pm 1}$  and  ${}^{33}\mu_{0}^{\Delta\ell=0}$ 

transition moment constants given in Table 2. The calculations were performed for a "pure" isotopic sample of <sup>12</sup>CH<sub>3</sub>I using an intensity cut off of  $2x10^{-26}$ cm<sup>-1</sup>/(molecule cm<sup>-2</sup>). This list is provided as supplementary material of this article. Table 3 provides some details on this linelist.

As already mentioned, the large values of the <sup>127</sup>I iodine nuclear quadrupole moment leads to easily observable hyperfine structures for v<sub>6</sub> transitions involving  $J \le 2 \times K$  (such a splitting is much smaller for parallel bands such as  $2v_3$ ). A given [J', K'] - [J'', K''] vibrationrotation transition of the v<sub>6</sub> or  $2v_3$  band is split into several hyperfine components with F'' =  $J'' \pm 1/2$ ,  $J'' \pm 3/2$ ,  $J'' \pm 5/2$  and  $\Delta F = \pm 1$ . The theoretical methods used to account for these effects are well described in the literature [30,31]. In the present work, the shift <sup>Hyp</sup> $\delta = ([J', K', F'] -$ [J'', K'', F'']) - ([J', K'] - [J'', K'']) of the position of a given [J', K', F'] - [J'', K'', F'']hyperfine component relative to the position of its [J', K'] - [J'', K''] counterpart was computed using the semi-perturbative treatment described in Ref. [1]. The line intensity of each hyperfine component was deduced from its corresponding [J', K'] - [J'', K''] hyperfinefree vibration-rotation transition by replacing in Eq. (1)  $g_{nucl}^{I} = 6$  by  $J''F''g_{nucl}^{I}$  with [30]:

$$\int_{J'F'}^{J'F'} g_{nucl}^{I} = (2F'+1)(2F''+1) \begin{cases} J' & F' & 5/2 \\ F'' & J'' & 1 \end{cases}^{2}$$
Eq. (8)

Obviously, summation over F' and F" leads to  $\sum_{F'F'} \int_{J'F'}^{J'F'} g_{nucl}^{I} = 6$ .

However, accounting for the hyperfine structure increases dramatically the size of the linelist for the  $v_6$  and  $2v_3$  bands. To limit the size of the linelist, the detailed hyperfine structure was generated only when it proved necessary for the correct modeling of the CH<sub>3</sub>I spectrum at 11  $\mu$ m.

For the present linelist, we decided to account for the hyperfine structure provided it leads to a  $^{Hyp}\delta$  hyperfine shift larger (in absolute value) than a pre-defined  $^{Hyper}$ Limit threshold limit of 0.002 cm<sup>-1</sup>. This lower limit more or less corresponds to the observed width of the lines in the experimental spectra studied in this work, involving the combined contributions from the Doppler width [ $\gamma_D$  (half width at half maximum) ~ 0.0005 cm<sup>-1</sup> at T = 296 K in the studied spectral region], from pressure broadening [32] and from the experimental line shape of the FTS.

# F. Validation of the line intensity calculation

Detailed comparisons of observed and calculated spectra were already presented on Figs. 1 to 6 of Ref. [1]. Another detailed comparison is shown in Fig. 4. It is clear that for low J values, accounting for the hyperfine structure leads to a better agreement between the observed and calculated spectra.

The results of the present line intensity study can also be compared with the existing low-resolution band intensity measurements. Table 4 lists the integrated band intensities  $S_{band}$  measured [15], determined from measured absorption cross sections [16] or calculated *ab initio* [17]. Integrated band intensities include contributions from isotopic species other than<sup>12</sup>CH<sub>3</sub>I and from various hot bands. Comparison of these integrated band intensities with the sum of the individual calculated line intensities reported in the present work therefore require application of the following empirical expression [33]

$$S_{\text{band}} \approx Z_{\text{Vib}}(T) \sum_{k} S_{k}$$
 Eq. (9)

where  $S_k$  is the intensity of line *k* in the linelist established for a pure isotopic sample of <sup>12</sup>CH<sub>3</sub>I which is described in section E. Eq. (9) assumes that the intensities of the v<sub>6</sub> band of all isotopologues of methyl iodine are the same, which seems to be reasonable as the v<sub>6</sub> band is only weakly perturbed. The integrated intensities of the v<sub>6</sub> and 2v<sub>3</sub> bands thus obtained for the present work, i.e.  $S(v_6) \approx 1.38 \times 10^{-18} \text{ cm}^{-1}$  (molecule cm<sup>-2</sup>) (±15%) and  $S(2v_3) \approx 2.04 \times 10^{-20} \text{ cm}^{-1}$  (molecule cm<sup>-2</sup>) (±20%), are also given in Table 4. It can be noted that the present line intensity measurements lead to an integrated band intensity in the v<sub>6</sub> region (750 – 1025 cm<sup>-1</sup>) that is about 8% smaller than the values reported by Dickson *et al.* [15] and Sharpe *et al.* [16]. This result will have to be checked by future line intensity studies. The *ab initio* prediction for the v<sub>6</sub> band intensity [17] is about twice as large as the experimental values [14, 15]. However, recent computations performed for CH<sub>3</sub>Cl [34, 35] clearly indicate that the quality of *ab initio* predictions was significantly improved since 1987.

Figures 5 and 6 compare the PNNL (Pacific Northwest National Laboratory) absorption cross sections with the results of the present calculations for the  $v_6$  and  $2v_3$  bands, respectively. As detailed in Ref. [16], PNNL absorption cross sections are for a sample (CH<sub>3</sub>I in the present case) diluted in 1 atm of nitrogen at 296 K and were derived from Fourier transform spectra recorded at a resolution of 0.112 cm<sup>-1</sup> (0.9/MOPD). As methyl iodide exhibits significant self-broadening, some of the sharper lines are actually broader than instrumental limits. To account for the hot bands contribution, our calculated line intensities were multiplied by  $Z_{Vib}(296K) = 1.1158$ . The comparison of Fig. 5 clearly shows that, on average, our computed intensities (cold  $v_6$  and hot bands) are weaker that the PNNL intensities by a factor of about 1.087. A closer look at the residuals (not shown) indicate that the signatures observed result from contributions from hot bands and possible line mixing effects.

# G. Conclusion

In this work, individual line intensities were measured for the first time for the  $v_6$  band of  ${}^{12}$ CH<sub>3</sub>I. Using a theoretical model that accounts for the resonances coupling the 6<sup>1</sup> energy levels with those of the 3<sup>2</sup> and 2<sup>1</sup> vibrational states, it was possible to determine the transition moment operator of the  $v_6$  band. The transition moment of the very weak 2v<sub>3</sub> band was determined using the integrated intensity of the band, estimated from the absorption cross sections measured at Pacific Northwest National Laboratory [16]. A linelist was generated for the  $v_6$  and  $2v_3$  bands, accounting for the hyperfine structure when it proved necessary. Comparisons with integrated band intensities and absorption cross sections available in the literature show that the present line intensities are about 8% lower.

### ACKNOWLEDGEMENTS

The DECA-PF project is sponsored by the French government `Investments for the future' program through the grant ANR-11-RSNR-0003 supervised by the French National Research Agency (ANR) under the `Research in Nuclear Safety and Radioprotection' (RSNR) research initiative.

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#### Table 1

Pressure of CH<sub>3</sub>I (in hPa) and number of interferograms averaged to yield the corresponding spectrum (# scans). All the spectra were recorded with an absorption path length of 12.4  $\pm$  0.1 cm, at a stabilized room temperature of 295  $\pm$  1 K, a resolution (equal to 0.9 divided by the maximum optical path difference) of 0.0019 cm<sup>-1</sup> and an entrance aperture diameter of the interferometer equal to 1.5 mm. The absolute uncertainty on the pressure is equal to 0.5% of the value given.

#	P(CH₃I)	# scans
S1	3.406 (17)	840
S2	6.433 (32)	828
S3	8.611 (43)	900
S4	11.418 (57)	940
S5	14.28 (7)	880
S6	21.28 (11)	900
S7	28.70 (15)	1040

# Table 2:

(A) Transition moment constants for the  $2v_3$  and  $v_6$  bands.

Band	Constants		Value in Debye
2v <sub>3</sub>	$^{33}\mu_0^{\Delta\ell=0}$	$arphi_{ m z}$	$\pm 6.84 \times 10^{-3}$ \$
$v_6$	$^6\mu_0^{\Delta\ell=\pm 1}$	$\varphi_x$	$6.14936(700)  imes 10^{-2}$
	$^6\mu_4^{\Delta\ell=\pm 1}$	$\left\{ i \varphi_{y}, J_{z} \right\}$	$-1.4331(990) \times 10^{-4}$
	$^6\mu_5^{\Delta\ell=\pm 1}$	$\left\{ \varphi_{\mathrm{z}},\mathrm{iJ}_{\mathrm{y}}\right\}$	$-7.96(150) \times 10^{-5}$

(B) Statistical analysis for the line intensity calculation:

Number of lines	840		
Percentage of lines			
0% ≤δ< 8%	65.6 %		
8% ≤ð< 16%	23.0%		
16% ≤δ< 30%	11.4%		
$\delta = \left  I_{Obs} - I_{Calc} \right  / I_{Obs} \text{ in \%}$			

 $\varphi_x$ ,  $\varphi_y$  and  $\varphi_z$  stands for the direction cosines  $\Phi_{Zx}$ ,  $\Phi_{Zy}$  and  $\Phi_{Zz}$ , respectively.

1 Debye =  $3.22564 \times 10^{-30}$  Coulomb m. The quoted errors are one standard deviation.

<sup>\$</sup> Maintained at this value during the calculation.

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# Table 3

Results of the intensity computation at 296 K (for a pure isotopic sample of  $^{12}CH_3I$ )

Band	Number of lines	Number of lines	Band intensity	Int <sub>Max</sub>	Sigma Min	Sigma Max	J <sub>Max</sub>	K <sub>Max</sub>
	Нур	No Нур	× 10 <sup>-19</sup>	× 10 <sup>-23</sup>				
v <sub>6</sub>	16573	12394	12.4	250	705.15	1102.43	88	24
2v <sub>3</sub>	6331	5227	0.183	3.8	986.53	1076.68	88	21
Total	23157	17890	1.26					

"Hyp" and "No Hyp" identify the number of lines generated when accounting and not accounting for the hyperfine structure, respectively. "Band intensity" and " $Int_{Max}$ " (maximum value of the computed line intensities) are in cm<sup>-1</sup>/(molecule cm<sup>-2</sup>). "Sigma Min" and "Sigma Max" are the minimum and maximum values of the computed line positions (in cm<sup>-1</sup>), respectively.

Table 4

Status of the previous band intensity measurements or calculations.

Results of the present line intensity computation at 296 K for the quoted cold bands ( $v_6$  and  $2v_3$  respectively) together with their associated hot bands (see Eq. 9) and for a natural sample of CH<sub>3</sub>I.

	Integrated band intensity	Method	Ref.
Band	in cm <sup>-</sup> ¹/(molecule×cm <sup>-2</sup> )		
	0.150(2)×10 <sup>-17</sup>	Low resolution	[15]
V <sub>6</sub>		integrated intensities	
	0.150(2)×10 <sup>-17</sup>	Low resolution	[16]
$v_6$		integrated intensities	
$\nu_6$	0.322×10 <sup>-17</sup>	Ab initio calculation	[17]
	0.138×10 <sup>-17</sup>	Computation of the $v_6$ line intensities for the whole	This
$\nu_6$		band <sup>s</sup> and estimation of the hot band contribution	work
		using Eq. (9)	
211	$0.19 \times 10^{-19}$	Low resolution	[16]
2ν3		integrated intensities	
	$\sim 0.20 \times 10^{-19}$	Computation of the $2v_3$ line intensities for the whole	This
2v <sub>3</sub>		band and estimation of the hot band contribution	work
		using Eq. (9)	

\$ For the  $v_6$  band, the line intensity parameters were derived through a least square fit performed on the set of individual line by line intensity measurement (see Table 2).

\* For the 2v<sub>3</sub> band, the line intensity parameters were estimated from the PNNL band intensity data.

# List of figures

# Figure 1

Results of the multi-spectrum analysis applied to 7 spectra (S1 to S7) of a small part of the  $v_6$  band of CH<sub>3</sub>I. The top panel shows the observed spectra and the lower panel presents the best-fit residuals corresponding to spectra S1 to S7. For the spectral range in Figure 1, the transitions considered for the line intensity computation are listed in Annex 2. Please note that the 915.538 cm<sup>-1</sup> line, which corresponds to a blended line cluster was not taken into account for the line intensity computation.



Comparison between various sources of relative uncertainty on the measured intensities of 3 selected lines of high, medium and low intensity:  ${}^{R}R(26,0)$ ,  ${}^{P}P(30,4)$  and  ${}^{P}P(57,4)$ .



Residuals (Obs–Calc)/Obs (in %) of the line intensity calculation for  $v_6$  line intensities plotted as a function of the experimental line intensities.



Measured line intensity (in cm<sup>-1</sup>/(molecule.cm<sup>-1</sup>))

Portion of the CH<sub>3</sub>I spectrum near the origin of the <sup>R</sup>Q<sub>K=3</sub> lines structure of the v<sub>6</sub> band in the 909.1 cm<sup>-1</sup> region. The observed spectrum is spectrum S3 in Table 1. The calculated spectrum was generated considering the self-broadening coefficients of the lines measured in this work. For the low J lines, it is obvious that accounting for the hyperfine structure ["Calc. (with HPFS)"] leads to a better agreement with observation than when this effect is not considered during the calculation ["Calc. (no HPFS)"].



Comparison with the Pacific Northwest National Laboratory (PNNL) [16] absorption cross sections in the  $v_6$  band region (see text for details).



Same as Fig. 5, applied to the  $2v_3$  band region.

