

Experimental and Theoretical Study of the Chemical Network of the Hydrogenation of NO on Interstellar Dust Grains

Thanh Nguyen, Dahbia Talbi, Emanuele Congiu, Saoud Baouche, Amir Karton, Jean Christophe Loison, François Dulieu

▶ To cite this version:

Thanh Nguyen, Dahbia Talbi, Emanuele Congiu, Saoud Baouche, Amir Karton, et al.. Experimental and Theoretical Study of the Chemical Network of the Hydrogenation of NO on Interstellar Dust Grains. ACS Earth and Space Chemistry, 2019, 3 (7), pp.1196-1207. 10.1021/acsearthspacechem.9b00063. hal-02267729

HAL Id: hal-02267729

https://hal.science/hal-02267729

Submitted on 2 Dec 2019

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Experimental and theoretical study of the chemical network of the hydrogenation of NO on interstellar dust grains

Thanh Nguyen,*,† Dahbia Talbi,‡ Emanuele Congiu, \P Saoud Baouche, \P Amir Karton, \P Jean Christophe Loison, \P and Francois Dulieu*, \P

†Present address: Institute of Low Temperature Science, Hokkaido University, Sapporo, Hokkaido 060-0819, Japan

‡LUPM - Laboratoire Univers et Particules de Montpellier, 34095 Montpellier Cdex 05, France ¶Université de Cergy-Pontoise, Sorbonne Université, Observatoire de Paris, PSL Université, CNRS, LERMA, F-95000, Cergy-Pontoise, France

§School of Molecular Sciences, The University of Western Australia, Perth, Western Australia
6009, Australia

||ISM-CNRS, Universit de Bordeaux, 351 cours de la Libration, F-33405 Talence, France

E-mail: nguyenthanh@lowtem.hokudai.ac.jp; francois.dulieu@obspm.fr

Abstract

Nitrogen Monoxide (NO) is observed in the gas phase of molecular clouds. It may accrete on dust grains and there its hydrogenation should lead to hydroxylamine (NH₂OH), the same way that CO is transformed in methanol (CH₃OH) on the surface dust grains. NO hydrogenation has been said barrier-less, whereas CO hydrogenation proceed through quantum tunneling, and is thus slower. However, CH₃OH is widely observed and is considered as a proxy of complex organic molecules while hydroxylamine remains undetected. We aim at studying, analyzing, and understanding the

chemical network of NO hydrogenation on cold surfaces. Experiments are carried out using a new Ultra-Hight Vacuum (UHV) set-up named VENUS. NO molecules and H atoms are co-deposited on a golden mirror at different temperatures. Infrared spectroscopy as well as Temperature Programmed Desorption (TPD) are used to follow the NO reactivity, with both H and D, and in presence or absence of water substrate. Quantum calculations on water ice cluster models are computed separately.

During the hydrogenation of NO, 10 reactions proceed concurrently. They are identified and constrained by changing physical conditions in experiments or in calculations. Among them, we demonstrate that the HNO+H addition reaction has a barrier which is probably crossed via quantum tunneling at 10 K. Moreover, abstraction reactions are occurring although they are limited by H and O bonding with their environment. Chemical desorption should occur especially in absence of water which enhances the total production of hydroxylamine. The chemical network of the hydrogenation of NO has been re-investigated. Each of the 10 reactions are sorted by efficiency. We exclude the possibility of a chemical loop between NO and HNO, especially in presence of water. Therefore hydroxylamine remains the main product of the hydrogenation of NO on grains and the question of its non-detection in ices or in the gas phase, specifically in shocked regions where ice mantles should be sputtered, is still open.

Keywords: Astrochemistry, molecular clouds, NO hydrogenation, dust grains, energy barriers, quantum tunneling

1. INTRODUCTION

With the unprecedented capability of mapping at small scale and at high sensitivity astrophysical regions, the origin and the evolution of the molecular complexity is permanently questioned. As an example Ando et al. ¹ has shown that chemical composition of two molecular clouds, separated by only 10 pc in the nearby galaxy NGC253, exhibit a very different chemistry. In particular, in one clump both NO and CH₃OH are detected,

whereas another one presents a poor chemical composition. Progressively a long and subtle inventory of possible chemical links emerges from the observations. For example we pinpoint the possible correlation of HNCO and NH_2CHO^2 or the possible link of O bearing compounds³. Less representative but probably equally meaningful, the non detection of CH_2OH^4 or the detection of very low concentrations of the CH_3O molecule is an important but unsolved problem⁵. Because the chemical link in the gas phase with CH_3OH is firmly established⁶, how to explain their non detection in shocks where CH_3OH is very abundant?

The presence of methanol in the gas phase and many other hydrogenated molecules is thought to be due to the sublimation of ice mantles. The identified species of the ice such as water, carbon dioxide, formaldehyde, and methanol are formed on the cold grains, by surface reactions combining H, O and CO.^{7–19} Under dense cloud conditions, just like CO molecules, NO molecules can also accrete and react with many species on the surface and with H in particular, leading to an increase in the chemical complexity of both O- and N-bearing species.

Even if the NO abundance is not as high as CO, it is however relatively abundant. In circumstellar oxygen-rich envelopes NO relative abundance is as high as 10^{-6} (Velilla Prieto et al. 20), whereas even in pre-stellar cores that undergo gas depletion it is still at the $1\text{-}3\times 10^{-8}$ level. 21 NO is an important gas phase molecule detected towards many dark and warm clouds with the abundance from 1×10^{-8} to 1×10^{-7} relative to H_2 abundance 22 , and is at the corner-stone of the N and O chemistry. Therefore the case of the non detection of hydroxylamine (NH₂OH) is an enigma, since its production by hydrogenation of NO on grains should occur without barrier 23 , and has other possible formation routes. 24 Taking into account the low reactivity of CO with H, the absence of NH₂OH in space and the omnipresence of CH₃OH is questioning our understanding of their synthesis, or their destruction pathways. 17,18,25

From experimental works 23,26,27 , a reactive network for the $\{NO+H\}$ reactive system

on cold surfaces has been proposed. Five reactions were able to satisfactorily explain the measurements:

$$H + H \xrightarrow{k1} H_2 \tag{1}$$

$$NO + H \xrightarrow{k2} HNO$$
 (2)

$$HNO + H \xrightarrow{k3} H_2NO$$
 (3)

$$H_2NO + H \xrightarrow{k4} NH_2OH$$
 (4)

$$HNO + NO \xrightarrow{k5} N_2O + OH$$
 (5)

followed by

$$OH + H \xrightarrow{k6} H_2O \tag{6}$$

Or alternatively to k5 and

$$HNO + HNO \xrightarrow{k5bis} N_2O + H_2O \tag{7}$$

However, in models is sometimes inserted the back reaction.

$$HNO + H \xrightarrow{k2B} NO + H_2$$
 (8)

Reaction (8) is an abstraction reaction, well known in methylated systems 14,19,28,29 , or even in sulfur bearing species. 30 The H + HNO reaction has been studied experimentally in the gas phase between 300 and 1170 K showing high rate constant (Glarborg et al. 31 and references therein). Theoretical calculations on the same reaction (Soto and Page 32 and Nguyen et al. 33) have shown either no barrier or a small one depending on the theoretical level. A low (or no) energy barrier for this reaction may be critical as it favors looping leading to H₂ formation. This kind of looping reactions k_2 and k_{2B} is a problem for astrochemical models including solid phase reactions (e.g. Vasyunin and Herbst, 34

Ruaud et al., 35 and Cuppen et al. 36), because their inclusion in the surface reaction network dissipates a large number of H atoms, deviating hydrogenation processes toward simple H_2 production. Therefore, the final degree of hydrogenation of the ice mantle is very sensitive to this balance of addition and abstraction mechanisms (private com. from Vasyunin et al. 37).

Thus it is important to put into the scope of our study the possibility of abstraction reactions. We add:

$$H_2NO + H \xrightarrow{k3B} HNO + H_2$$
 (9)

and

$$NH_2OH + H \xrightarrow{k4B} H_2NO + H_2 \tag{10}$$

There are 10 competing reactions on the surface, 8 actually belonging to the hydrogenation of NO. Reaction (6) is the completion of water, and reaction (1), the self-reaction of H, is the most important from the kinetic point of view. Because it's a fast reaction only limited by H diffusion, other reactions can't compete if they are too slow, and thus the system won't evolve chemically. This paper is intended to re-investigate the chemical network of the hydrogenation of NO by examining the whole question on two fronts, from both an experimental and a theoretical approach.

The article is organized as follows: we first present methods including the experimental setup, theoretical calculations. In the next section, we show and interpret experimental results, including the study of temperature dependence, the effect of different isotopes, the role of the water substrate and the theoretical results. In the following section, we conclude results in relation to the experiments and theory. In the discussion section, we compare approaches and discuss possible astrochemical implications before summarizing in the conclusion.

2. METHODS

2.1. Experimental setup

Experiments have been performed using the VENUS (VErs de NoUvelles Synthses) setup at the LERMA laboratory in the University of Cergy Pontoise.

The experiments were carried out under ultra-high vacuum (UHV) conditions with pressure from 1×10^{-10} to 5×10^{-10} mbar. Two different beams are used to deposit the reactants to the sample surface, which is constituted of a copper mirror coated with gold. The sample temperature can be controlled from 7 K to 350 K by using a regulated resistive heater. Absorption spectra of adsorbate (reactants and products) are recorded through Fourier Transform-Reflection Adsorption Infrared Spectroscopy (FT-RAIRS). Quadrupole mass spectrometer (QMS) is used 1) for probing the reactants beams before they accrete to the surface; 2) for measuring the products during the thermally programmed desorption (TPD).

VENUS has 4 different beam lines, but only two are used in this study. The top beam was used to inject the NO molecules onto the surface. The source pressure is 2.05×10^{-4} mbar, which corresponds to a flux at the surface level of 2×10^{12} molecules/cm²/s. The atomic hydrogen was sent to the surface through the right beam. Its flux is around 8×10^{12} atoms/cm²/s. H/D atoms were generated by dissociating of H_2/D_2 molecules within a microwave discharge of 75W/40W. The dissociation efficiency of H_2/D_2 was 75 \pm 5%. We use FT-RAIRS to probe the new features during the co-deposition duration. After the deposition, the products are detected with the TPD technique, which consists in measuring with the QMS the different products desorbing during the linear increase of the surface temperature (0.2 K/s).

In order to calibrate the mono-layer of NO (ML, 1ML = 10^{15} molecules/cm²), we performed series of NO TPD (like in Noble et al. 11) with different doses deposited at 10 K, then heated to 80 K. Thus, we determined that 1 ML of NO is obtained after 8 to 10 minutes

of deposition. All the experiments presented here consist in co-deposition experiments, which means that NO and H(D) were sent simultaneously on the substrate (varied in composition, gold or water) at different temperatures, kept constant during the deposition. At the end of the co-deposition, the hydrogenation is continued for 10 minutes.

We note $\{NO+H\}$, the reactive system, to make the difference with the NO+H (\longrightarrow HNO) reaction. Actually as mentioned earlier the $\{NO+H\}$ system include at least ten reactions, and at higher temperatures, some desorption pathways can also be activated.

To understand the possible role of the substrate, we have performed similar codeposition experiments, but on a water substrate. The substrate was previously grown at a surface temperature of 10 K by background deposition of 5 ML of H₂O. So the substrate is porous and exhibits a larger surface area due to its 3D structure, but remains almost completely accessible to deposit.³⁸

2.2. Theoretical approach

To guide the interpretation of our experimental results, we carried out Density Functional Theory (DFT) calculations to locate the transition states on the H + HNO, H + H₂NO and H + H₂NOH reactions paths in the gas phase. The hybrid-meta GGA M06-2X³⁹ functional was used for that purpose in conjunction with the aug-cc-pVTZ atomic basis set. While the M06-2X highly nonlocal functional is well suited to describe noncovalent bonds (transition states, loose complexes) it may slightly underestimate barriers. ^{39,40} Therefore for the low energy transition states involved in the H + HNO reactions (barriers below 1000 K) we recalculated these barriers using the CCSD(T)⁴¹ method on geometries optimized at the above M06-2X level, and the explicitly correlated (F12b) version of the coupled-cluster method (CCSD(T)-F12b)⁴² combined to the cc-pVQZ basis set with F12 optimized (cc-pVQZ-F12)⁴³ and QCISD/6-311++G(d,p) geometries. Finally, because of the constant decrease of the energy barriers for the H + HNO hydrogen abstraction reaction we used the Weizmann-4 thermochemical protocol^{44,45} which calculates the CCSDTQ5/CBS energy

using geometries and zero point energies optimized at the CCSD(T)/cc-pVQZ level.

To model the ice surface of the experiment we used a cluster of water molecules. To define the size of this cluster we investigated HNO-XH₂O complexes with X varying from 4 to 6. Our full geometry optimization of these complexes demonstrated that the orientation of the HNO molecule with respect to the clusters was not affected by their size. Indeed 4 water molecules were sufficient to account for all the hydrogen bonds that bind HNO to the water clusters and therefore to account for the surfaces interactions between HNO and the water clusters. This was verified too for H₂NOH-XH₂O complexes with X=4 and 5. Limiting to 4 water molecules the size of our water cluster allowed us to study the hydrogenation of NO catalyzed by the ice at a theoretical level equivalent to the one used for the hydrogenation of HNO in the gaze phase, including reaction path investigations, frequency analyses and zero point energy calculations. Combining a 6-311++G(d,p) basis set to the hybrid-meta GGA M06-2X functional set we used an ultrafine integration grid for the optimization of our NO-4H₂O, HNO-4H₂O, H₂NO-4H₂O and H₂NOH-4H₂O minima and transition state. This was required regarding the loose structure of our complexes and some of their very low frequency modes. To locate the transition state for the addition of H on the HNO-4H₂O complex, an extensive investigation of the H-HNO-4H₂O hyper-surface was required. This was achieved by optimizing the H-HNO-4H₂O complex for each fixed distance of approach H...HNO-4H₂O. The transition state was determined by locating the maximum on the calculated optimized potential energy curve.

Finally, we determined the binding energies to the 4 water cluster, of the NO, HNO, H_2NO and NH_2OH molecules as the difference between the energy of the molecules bond to the water cluster and the sum of the energies of the water cluster and of the molecules. We also calculate the binding energy using our model developed in Wakelam et al.. ⁴⁶ All calculations have been done using either Gaussian 09^{47} or $MOLPRO^{48}$ quantum chemistry codes.

3. RESULTS AND DISCUSSION

3.1. Theoretical results

3.1.1. The H + HNO abstraction reaction

Experimental results on the H + HNO reaction in the gas phase have pointed out for a high rate constant with a low barrier in the entrance valley. 31,49 Previous MP2 and QCISD theoretical calculations on the abstraction pathway have led to a small barrier of 2 kJ/mol,³³ shown to vanish at the MRCI level.³² Our present theoretical investigations using the Weizmann-4 thermochemical protocol see Table 2 in the Supporting information confirm that there is no barrier for this abstraction reaction. Since this approach has been found to give thermochemical properties (such as atomization energies) with sub-kJ/mol 95% confidence intervals from highly accurate experimental determinations, 50 we can confidently confirm that this reaction is barrier-less in the gas phase a statement that cannot be generalized to the same reaction on the ice. Indeed our theoretical investigations show that HNO is bound to the 4 water cluster through two hydrogen bonds (Table 4 in the supporting information) involving respectively the oxygen and hydrogen atoms of HNO. Such a position with respect to the cluster prevents the hydrogen abstraction reaction H + HNO-4H₂O to occur. Indeed, in the gas phase (Table 2 in the Supporting information) the abstraction pathway requires a quasi-linear H...HNO approach (with an HHN angle close to 180°). Such an approach, as confirmed by our attempt to constrain the abstraction reaction on HNO adsorbed on the water cluster is very repulsive, forcing the H atom to follow the valley corresponding the addition reaction path.

3.1.2. The H + HNO addition reaction

For the addition of H on the nitrogen atom our results from Table 2 in the Supporting information show that this reaction presents a barrier of 6.6 kJ/mol at the DFT level (12.8 kJ/mol at the CCSDTQ5/CBS level). At the same DFT level, this barrier is almost

unchanged when HNO is on the 4 water cluster i.e. (6.1 kJ/mol-Table 5 in the Supporting information); the difference is less than 10%. The 4 water cluster has almost no catalytical effect on the reactivity of H with HNO bound to the water cluster. The addition of H on the oxygen atom of HNO presents a very high energy barrier in the gas phase (Table 2 in the Supporting information). We did not investigate this reaction with HNO on the 4 water cluster because of this high barrier, which we do not expect to be reduced to the level that this reaction could become efficient on the ice.

3.1.3. The $H + H_2NO$ reaction

Even though we expected the addition of H on H_2NO to form H_2NOH to be barrier-less in the gas phase we checked this point by investigation this reaction path at the both M06-2X/aug-cc-pVTZ and QCISD/6-311++G(d,p) levels . We also investigated the addition of H on H_2NO bound to the 4 water cluster and similarly we saw no evidence for the existence of a transition state for the formation of the $H_2NOH-4H_2O$ complex. It is interesting to notice here that the OH bond of H_2NOH is involved in a hydrogen bond with one of the oxygen of the water cluster (Table 4 in the Supporting information) with a consequence on its reactivity that will be discussed latter.

3.1.4. The $H + H_2NOH$ reaction

For this reaction, only the gas-phase mechanism have been considered. Our DFT calculations (Table 3 in the Supporting information) show that the two possible abstraction reactions for the formation of H_2NO and NHOH can compete, having similar energy barriers. On the ice one would expect the formation of NHOH to be favored because of the orientation of H_2NOH with respect to the surface of the ice.

3.2. Experimental results

3.2.1. Completeness of the reactions of the $\{NO + H\}$ system before the TPD

Because the TPD diagnostic requires the heating of the sample, it is important to know if all the initial reactants are consumed during the co-deposition phase. The use of the infrared spectroscopy gives precious indications because it can be performed during the reacting phase so before any thermal transformation of the sample. It is sometimes called in situ measurements. Figure 1 shows in red the IR spectrum recorded after NO deposition (without co-deposition of H). We can see the characteristic, N=O asymmetric stretch ν_5 and the N = O symmetric stretch ν_1 , observed at 1770 and 1860 cm⁻¹, respectively, of NO dimer. 23,51 In blue is displayed the IR spectrum obtained after the co-deposition of NO and H on the gold surface held at 10 K. NO dimer (or monomer) features are no longer detectable, what we observe is a peak at 2233 cm $^{-1}$ which is attributed to N₂O. We note here that the IR features of NH₂OH is composed of broad bands from (900 -1800) cm⁻¹ and (2700 - 3500) cm⁻¹, that are hardly detectable and furthermore unusable for quantitative analysis under our experimental conditions, because of their relative low contrasts compared to narrower peaks such as those of NO or N2O. No evolution of the IR signal is observable during the heating phase of the TPD, except of course the disappearance of the N_2O peak due to its desorption of N_2O .

These results are similar to those previously published^{23,27}. Taking into account the sensitivity of the method, we note that more than 90% of the NO should have reacted during the co-deposition phase so at low temperatures.

We show the TPD curves in figure 2. For pure $\{NO\}$ experiments (in red) we observe a desorption peak at around 50 K. In blue is displayed the desorption after the $\{NO+H\}$ reactive system completion. We can clearly see both N_2O desorption (140-160 K) and NH_2OH desorption (160-200 K). Their ratios are clearly indicating that NH_2OH is the major product, whereas N_2O is the minor product.

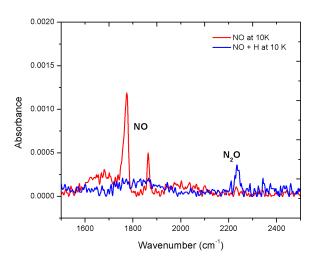


Figure 1: IR spectra obtained after the deposition of 4 ML of NO at 10 K on gold (red line), and after NO (4ML) and H (16 ML) co-deposition (blue line).

Around 50 K, there is still a weak desorption of NO, slightly shifted and broadened compared with the case of pure $\{NO\}$ experiments, indicating some remaining NO. However, the NO consumption is \sim 90%, in agreement with IR spectra. The shift and broadening of the peak is due to the change of the molecular surrounding of the desorbing NO. Actually, in case of $\{NO+H\}$ experiments the NO molecules have to escape from an ice film now composed in majority of NH₂OH and N₂O. These two compounds are desorbing at higher temperatures (\sim 180 K and \sim 150 K, respectively) and therefore are slightly delaying the NO desorption.

The remaining NO could originate from incomplete NO consumption, although we cannot exclude some default in our physical beams overlap on the sample, making a geometrical zone of NO less exposed to H atoms. Longer H expositions time on thinner NO films, however, rule out this possibility. Under these conditions we can achieve a full NO disappearance. More reasonably, we can suppose that the remaining part of NO is due to an insufficient total number of H atoms compared to NO molecules, but it is also possible that a back reaction $HNO + H \longrightarrow NO + H_2$ (reaction 8) impedes a complete consumption of NO by repopulating NO from is product HNO. This possibility will be

discussed later.

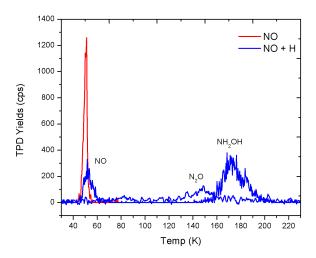


Figure 2: TPD profiles of pure {NO} (red curve) and {NO +H} (blue curves). m/z 30, 33 and 44 are displayed corresponding respectively to NO, N₂O and NH₂OH

3.2.2. Temperature dependency

In previous experiments, reactivity of NO with H has been tested at different temperatures 23 , up to 42 K using porous amorphous ice as a substrate. The aim was to establish that the H + NO reaction is barrier-less. However, only NO consumption was tested, and no products were analyzed. Such product analysis was conducted only for experiments performed on surfaces held at 10 K.

Here we present new results that show the main products of NO hydrogenation, namely hydroxylamine (NH₂OH m/z=33) and nitrous oxide (N₂O m/z=44) for two surface temperatures of 8 K and 40 K, respectively (Figure 3). For both experiments, there is a remaining part of NO. It corresponds to the desorption peak at around 50 K (green curve). The other features of m/z = 30 at higher temperature are due to the cracking patterns of both NH₂OH (peak at 160-200 K) and N₂O (around 140-170 K). More important is the extreme difference in intensity of the NH₂OH and N₂O TPD peaks. At 8 K, NH₂OH is the major product whereas at 40 K N₂O becomes the major product and NH₂OH is almost

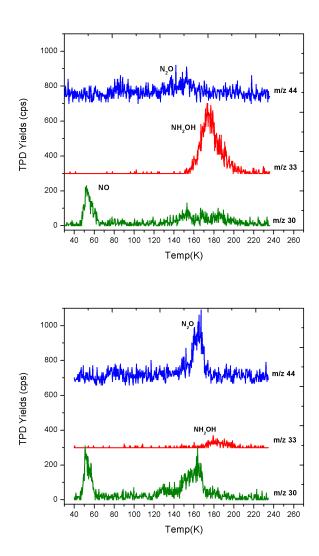


Figure 3: Influence of the surface temperature: TPD traces of m/z=30 (NO, green lines), m/z= 33 (NH₂OH, red lines) and m/z= 44 (N₂O, blue lines) for co-deposition experiments of {NO+H} on gold substrate held at 8 K (upper panel), and 40 K (lower panel). Curves are offset for clarity purpose.

present in trace amounts. We calculated the ratio of the integrated TPD areas between NH_2OH and N_2O , it is 6.7 at 8 K, and 0.01 at 40 K.

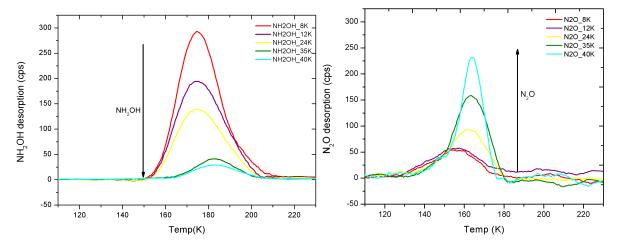


Figure 4: TPD traces of NH₂OH (left panel) and N₂O (right panel) at different surface temperatures. Each temperature (8, 12, 24, 35, 40 K) corresponds to one specific codeposition experiment {NO+H}. Curves are smoothed by adjacent averaging for a better visibility. NH₂OH amount decreases with T_s whereas N₂O increases. The vertical black arrows represent the evolution of the peak with the temperature of the substrate.

Figure 4 shows the TPD series of NH_2OH (left panel) and N_2O (right panel) at different surface temperatures. By looking at NH_2OH and N_2O peaks, we see a clear decrease in NH_2OH , whereas N_2O yield is increasing with the surface temperature. We have calculated the integrated area of NH_2OH and N_2O desorption peak for each temperature, and plotted the result in Figure 5. We can see an exchange of the chemical output of the $\{NO+H\}$ reactive system. There is a decrease in NH_2OH , which can be fitted by an exponential decay, while the N_2O yield may exhibit a double increment, one at the lowest temperature and the other one at around 30 K.

The main learning that comes from this new set of experiment is that the ratio between the different reactions vary with surface temperature. The yields of all kinetic experiments, as those presented here, are always the result of a competition between the different reactions. Every single reaction rate evolves differently with temperature. In the given set of equations we can distinguish and classify some of them. Reaction 1 (H + H \longrightarrow H₂), is central to this reactive system, since it regulates the H surface density, so the efficiency

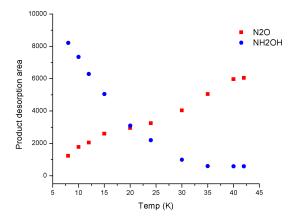


Figure 5: Integrated areas of NH_2OH (blue circles) and N_2O (red squares) desorptions as a function of the surface temperature, obtained after the same co-deposition of NO and H.

of the hydrogenation. In absence of any other reactants, the H surface density reduces with surface temperature. Between 8 and 14 K it is due to the increase in the H mobility that limits the reaction. At higher temperature, desorption of H start to be significant and opens a new competitive exit channel. However, even if the residence time of H is very short, reaction 2 (NO+H \longrightarrow HNO) turns out to be faster since the chemical network is fed by this reaction and also NO is still highly consumed at 42 K.

To calculate a rate such as k_1 or k_2 , it is usually assumed that it can be decomposed in different parts, one which is intrinsic to the reaction, and corresponds to the probability per attempt to cross the barrier, and one which depends on the conditions which are determining the number of attempts per second, so the diffusion and the surface density (or concentration). As an example we write k_1 and k_2 .

$$k_1 = k_{diff(H)} p_{R1}[H][H]$$

$$k_2 = k_{diff(NO,H)} p_{R2}[NO][H]$$

where [H] is the surface density of H.

In first approximation we can consider that the diffusion is dominated by H diffusion

so the diffusive terms $k_{diff(NO,H)}$ $k_{diff(H)}$ are equal. Moreover, the probability to overcome the barrier of reaction p_{R1} and p_{R2} are equal and close to unity because the two reactions are associating two radicals. Therefore, we can see that once the [NO] surface concentration is larger than the [H] surface concentration, the rate k_2 will dominate over k_1 .

Therefore a decrease of [H] does not necessarily reduce the efficiency of reaction 2, because it affects firstly reaction 1 due to the square function of [H]. Having verified that reaction 4 ($H_2NO + H \longrightarrow NH_2OH$) on the 4 water cluster is also barrier-less, it should proceed promptly as soon as some H_2NO is appearing on the surface, and should be kinetically behaving like reaction 2. Reactions 5 and 7, which lead to N_2O and H_2O , are not hydrogenation reactions. Therefore, the limiting factor compared to hydrogenation reactions for these reactions is certainly the diffusive part of the reaction because the diffusion of H is believed to dominate over all other species diffusion. At low temperatures, only the increase of the surface density of HNO, so the relative inefficiency of its destruction, is able to explain why at 10 K we can observe the formation of N_2O .

Now we have to consider other reactions, especially reaction 3 (HNO + H \longrightarrow H₂NO), which enables the formation of hydroxylamine. We have seen before that the environmental part (diffusion and surface concentration) of the rate k_3 is very favourable. It is hydrogen addition so the diffusion is about the same level as other hydrogenation, and follow the same temperature trend, and the surface concentration of HNO is likely to be high enough to accelerate reaction 5. Therefore, the limiting factor should be the probability p_{R3} which is lower than other hydrogenation reactions, implying a barrier to the reaction. From our theoretical investigations we have located a barrier of 734 K for the addition of a H atom on the nitrogen of HNO adsorbed on the water cluster. We have also calculated for the HNO + H \longrightarrow NHOH reaction a barrier above 10000 K (Table 3 in the supporting information). Since the ice surface is not expected to lower such a high barrier at a level where tunneling could become efficient, this reaction should be a minor one on the ice. Therefore, as we observe a relative decrease in hydroxylamine with temperature,

likely due to reaction 3, and since we have located a low energy barrier for $H+HNO-4H_2O$ reaction, it is fair to suppose that this addition proceed via tunneling.

Figure 6 shows a sketch of the {NO+H} reactive system, including the main change with temperature of the chemical pathways, which is probably induced by the different dependency on the temperature of reaction 3, if compared to other hydrogenation reactions.

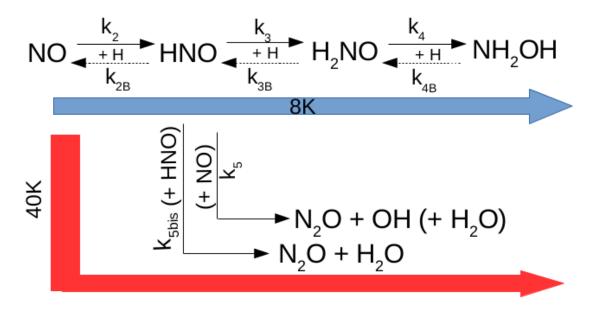


Figure 6: Sketch of the {NO+H} reactive system as it proceeds at 8 K (blue arrow, and and 40 K (red arrow).

3.2.3. The $\{NO + D\}$ reactive system at various temperatures

It is possible to test the hypothesis of the reaction proceeding through quantum tunneling by substituting H atoms by D atoms. Indeed, the mass variation is supposed to lower considerably the reaction rate in case of quantum tunneling.

Figure 7 shows the TPD profiles of the $\{NO+D\}$ co-deposition experiment for a surface held at 10 K. ND_2OD desorbs between 170 K and 190 K a,nd N_2O between 140 K and 180 K. The presence of the two main products indicates that the reactive network is similar but if we consider the height of the peak of hydroxylamine, we can see by comparison

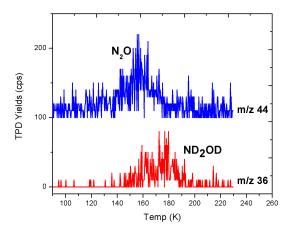


Figure 7: TPD profiles of masses m/z = 36 a.m.u. (in red) and m/z = 44 (in blue) corresponding to ND₂OD and N₂O, respectively, recorded after a co-deposition experiment of {NO+D} at 10 K.

with Figure 3 that it is relatively smaller, whereas N_2O is about the same size.

Table 1: Ratios of the TPD areas of N_2O , NH_2OH (or ND_2OD) obtained after co-deposition experiments and the TPD area of NO in the pure NO experiment (indicated by NO*). {NO + D} and {NO + H} are performed on a gold substrate hold at 10 K whereas {NO + H + H_2O } is made on a H_2O substrate.

Experiment	N ₂ O/NO*	NX ₂ OX/NO*	NO/NO*
{NO*}	< 2%	<1%	100%
${NO + D}$	16%	8%	25%
${NO + H}$	14%	53%	13%
$\{NO + H + H_2O\}$	12%	82%	2%

Table 1 presents the relative production of N_2O and NH_2OH (or ND_2OD), at the end of co-deposition experiments performed at 10 K. We divide the TPD areas of the products by the TPD area of NO molecules obtained during the $\{NO\}$ experiment. We have added a star in the NO^* case to emphasize that the normalization is not done with respect to the remaining NO in each experiments, but according to the amount of NO obtained in the pure NO experiment. Therefore, the ratio NO/NO^* is 100% for the first experiment. We first point out that the efficiency of detection of the QMS is not exactly equal from one molecule to the other. It can vary within a few tens of % for these molecules. It is due to

the different ionization cross sections (which could be calculated), but also to the different spatial and kinetic distributions of the different desorbing species (which is unknown). We added TPD areas of all the cracking patterns of a given molecule. This can be done even for masses showing several peaks such as mass 30, which is attributed to 3 parent molecules (NO, N_2O , NH_2OH), because they these species not desorb in the same window of temperature. As a conclusion, the ratios presented in the table do not exactly represent the ratios of the species, but they give a good indication. In the same column, the relative error is similar, so if we underestimate or overestimate a species, we do the same for all the experiments. Therefore, if the absolute values can be debated, the relative values are reliable.

In Table 1 we first notice that the total yield for the {NO+H} experiment is 80% of the deposited dose. It is less than 100% which could be due to the inaccuracy of the method. But this is certainly not an explanation for the {NO+D} experiment whose total yield is 49%. So there are missing products in the deuterium experiment.

Our hypothesis is that some of the products are lost because of the chemical desorption (or reactive desorption). The chemical desorption is the direct return to the gas phase of a newly formed molecule on a surface due to the excess energy released in case of an exothermic reaction. Following our previous findings, this is very likely to occur when the excess of chemical energy is large and when the products have a small number of atoms, such as in the following reactions for example:

$$NO + H \longrightarrow HNO_{gas}$$
 (11)

$$HNO + H \longrightarrow NO_{gas} + H_2$$
 (12)

$$HNO + H \longrightarrow NO + H_{2gas}$$
 (13)

However, the chemical desorption efficiency is not believed to change dramatically by substitution of isotope, and remains a relatively low probability effect (<10-20% per

reaction) most of the time. So we conclude that to increase the loss of products (i.e., species that go to the gas phase), the efficiency of some reactions should increase, and that probably reactions 8 and 12, which are abstraction reactions, are enhanced in the case of the {NO+D} experiment.

Indeed, as the abstraction reaction 8 and 12 have much smaller barriers than addition reactions, they are enhanced for D atoms as tunneling is less efficient for D than it is for H. Thus the reaction

$$DNO + D \longrightarrow D_2NO$$
 (14)

is unfavored with respect to

$$DNO + D \longrightarrow NO + D_2 \tag{15}$$

From Table 1 we can see that the production of N_2O may be higher, but it is at the limit of our experimental uncertainties. The increase of DNO on the surface, that is induced by the reduction of the second step of deuteration favors the N_2O production. However the absence of mobility of the reactants at low temperature, keep the yield relatively low, because N_2O does not come directly from an hydrogenation reaction. So, the only remaining possibility for DNO is to react via the back reaction 8, which favors the chemical desorption, hence the prompt release of molecules into the gas phase.

Figure 8 compares the yields of NH_2OH and ND_2OD after $\{NO+H\}$ and $\{NO+D\}$ experiments carried out at different surface temperatures. We can see that the trend is similar. Hydroxylamine is less efficiently produced when the temperature is increased. But the hydrogenation is always more efficient than the deuteration, especially at low temperatures. On the other hand, the formation yield of N_2O is slightly favored in the deuteration experiments, but the most striking point is the lost of a large fraction of the initial reactants, probably expelled from the surface due to the chemical desorption. This

non-thermal desorption efficiency is enhanced in the case of reactions that proceed in both directions. We can conclude that DNO + D \rightarrow D₂NO is much slower than HNO + H \rightarrow H₂NO. This is an indication that quantum tunneling should be at play for this specific reaction. This is consistent with our calculated low energy barrier for this reaction. Of course, it could also be an effect of zero-point energy shift, D atoms lying usually deeper in adsorption wells, and therefore may have a higher barrier to cross. However, in this peculiar case one has to explain why the abstraction reaction DNO + D \rightarrow NO + D₂ is relatively more efficient than the reaction HNO + H \rightarrow NO + H₂. Our best explanation remains that reaction 3 has a higher barrier than reaction 8, and that reaction 3 should be overcome via quantum tunneling.

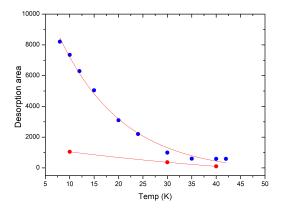


Figure 8: Blue circles: NH₂OH integrated areas after the co-deposition of NO molecules and H atoms. Red circles: ND₂OD integrated areas after the co-deposition of NO molecules and D atoms. ND₂OD yield is less than that of NH₂OH at the same surface temperatures.

3.2.4. Catalytic role of water

In dark molecular clouds, the major component of molecular mantles covering dust grains is solid water.⁵⁵ A water substrate should have an effect on the kinetic of surface reactions because it changes at least the environmental factors of the hydrogenation. Indeed, the H diffusion is dependent on the type of substrate⁴⁶ and is known to affect the hydrogenation of CO.⁵⁶ The comparison of the efficiency of formation of hydroxylamine is presented in

Figure 9. The blue curve, obtained in the $\{NO+H\}$ experiment on a porous ice substrate, has the same shape but a higher amplitude than the one obtained after the hydrogenation of NO on a gold substrate (in red). Therefore, we can conclude that the presence of water is helping the formation of hydroxylamine. The last line of Table 1 displays the amount of the different products. The consumption of NO is higher, the formation of NH₂OH is also enhanced while the production of N₂O is stable or slightly reduced. The higher consumption of NO is a precious indication. Actually we know that reaction 2 is barrier-less and therefore that NO should be continuously consumed. The remaining amount of NO at the end of the experiment is probably due to the presence of the (reverse) abstraction reaction 8. So we can conclude that the presence of water is limiting reaction 8, probably by imposing geometrical constraints on the orientation of HNO.

Our DFT calculations (see Table 4 in the supporting information) indicate that the most stable geometry for HNO bound to water is via the H_2O ...H-NO interaction. Then the H atom abstraction is more difficult due to geometrical constraints, and so limiting the abstraction reaction. If reaction 8 is less efficient, there are less chemical loops between HNO and NO and therefore the chemical desorption total efficiency is lesser. Moreover, water ice substrates typically impede the chemical desorption. 52,54 We note that the total amount of NO-bearing products in presence of water (Table 1, $\{NO + H + H_2O\}$ experiment) is the highest of all experiments. There is no increase in N_2O since its chemical pathway is not much changed. There is not more HNO formed, and the water does not help or even reduce the probability of having NO close to HNO, because of its larger surface area.

3.2.5 The possibility of back reaction NH₂OH + H \longrightarrow H₂NO + H₂

To test this possibility, we first synthesize a ND_2OD film in *situ* and later expose it to H atoms. The ND_2OD is obtained by co-depositing $\{NO+D\}$ at 10 K and thus we heat the sample to 160 K. At this temperature ND_2OD does not desorb, but other species do (i.e D_2 , NO, and N_2O). Then, the surface temperature is cooled to 10 K and H atoms (half the dose

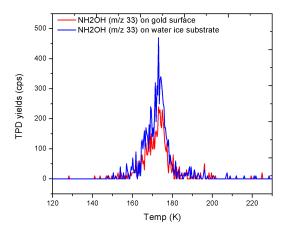


Figure 9: TPD profiles of NH₂OH desorbing after {NO+H} codeposition experiment on gold surface (red) and water ice substrate (blue), at the same sample temperature of 10 K.

of the target molecule) are sent.

The top panel of Figure 10 shows that ND₂OD is reduced by the addition of H atoms. The bottom panel shows a large peak of mass 35. This mass can be attributed to NDHOD, or its isomer ND₂OH. This substitution reaction is unlikely to proceed directly, actually D bonds are usually slightly stronger than H bonds due to the shift of zero point energy. So the most probable case is that a loop of abstraction and addition occurs, making the substitution efficient. Our DFT calculations show that, in the gas phase, the H atom abstraction by H on the NH₂ and OH group of NH₂OH involve barriers around 1500 K. They also show that the most stable geometry for NH₂OH bound on our water cluster involves an H₂O..H-ON interaction (see Table 4 in the Supporting information). Therefore, the H atom abstraction from the OH group of NH₂OH should be inhibited because of geometrical constraints and the back reaction toward the formation of NHOH favoured if tunneling can take place.

The abstraction- addition steps are the following:

$$ND_2OD + H \longrightarrow NDOD + HD$$
 (16)

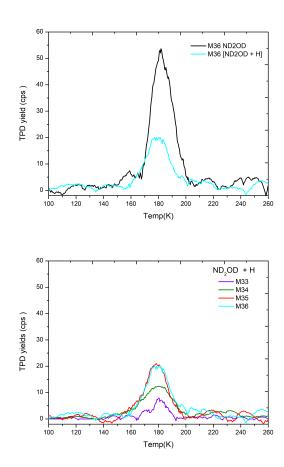


Figure 10: Top panel: TPD profiles of initial ND₂OD (black curve) and remaining ND₂OD (cyan curve) after reacting with H atoms at 10 K. Bottom panel: TPD profiles of products detected after exposure of H atoms onto a ND₂OD film. m/z 35 (red curve), m/z 34 (green curve), m/z 33 (violet curve), and m/z 36 (cyan curve)

$$NDOD + H \longrightarrow NDHOD$$
 (17)

The presence of mass 34 could only be due to the cracking pattern of ND_2OD . But in this case, it should be around a tenth of the mass 36. Therefore a large fraction of the mass 34 is probably the doubly hydrogenated hydroxylamine. This would be the indication that the abstraction - addition mechanism is probably occurring on the ND_2 group more than on the OD group (leading to simple hydrogenation). Even if this is speculative because we know that proton exchange is occurring during the TPD^{57-59} it is in agreement with our theoretical calculations showing that the back hydrogen abstraction reaction should favour the NHOH formation.

3.2.6 Binding energies

The TPDs performed in this work allow us to determine the binding energies of NO, NH_2OH and N_2O . These values, shown on Table 1 in the Supporting information, are compared to the values calculated in the present study and to those given by the semi-empirical model developed by Wakelam et al.. ⁴⁶ The agreement is good.

4. DISCUSSION AND CONCLUSIONS

4.1 Summary of the results

Our experimental re-investigation of the NO+H reactive system has demonstrated the following points: (i) it exists an activation barrier to the H + HNO reaction, (ii) this barrier is likely to be crossed at low temperature through quantum tunneling, (iii) the direct evidence of the back reaction $NH_2OH + H$ has been shown, which reminds us of the importance of taking into account this process. Indirect evidence of the H + HNO back reaction has also been provided. (iv) Water has a catalytic role. It increases the efficiency

of H additions probably by reducing the efficiency of the back reaction HNO + H \rightarrow NO + H₂. Alternatively, it is the water propensity to retain the products after reactions have occurred or, in other words, its property of reducing the chemical desorption can also partly explain why more products are obtained on a water ice substrate.

Our theoretical calculations corroborate these experimental findings and quantum chemistry help us to understand our results. More precisely we calculated at the MO6-2X level within the cluster approach:

 $H + NO \longrightarrow HNO$: no barrier

 $H + HNO \longrightarrow NO + H_2$: geometrically unfavored on water

 $H + HNO \longrightarrow NH_2O$: barrier of 734 K

 $H + NH_2O \longrightarrow NH_2OH$: no barrier

At the M06-2X level in the gaze phase:

 $H + HNO \longrightarrow NHOH$: barrier of 8167 K

 $H + NHOH \longrightarrow NH_2OH$: no barrier

 $H + NH_2OH \longrightarrow NH_2O + H_2$: barrier of 1528 K

 $H + NH_2OH \longrightarrow NHOH + H_2$: barrier of 1540 K

4.2 New insights into the catalytic role of dust and water ice

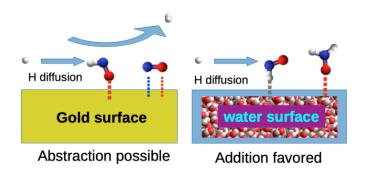


Figure 11: Schematic view of the role of water ice, compared with a gold surface, in the abstraction reaction H+ HNO. Due to H-bonding in a water environment, the HNO orientation does not favor H abstraction, despite its very low barrier in the gas phase

Faced with the difficulty of finding effective reaction pathways leading to hydrogenation of molecules in the gas phase, interstellar dust grains have long been identified as a probable source of H additions. 60 Cold surfaces are said to be catalytic centres thanks to four main properties. i) The grains collect species coming from the gas phase and promote encounters. ii) The grains act as third bodies and are able to dissipate excess energy. For this, the case of H_2 is emblematic. The $H+H \longrightarrow H_2 + hv$ reaction (where hv is a photon to remind that it is a radioactive recombination), is highly inefficient in the gas phase for such a small system. Here, the dust grains act as a third body and dissipate the excess energy of the reaction. iii) In addition to accretion, property i), the grains considerably increase the interaction time of the reactive partners. In the gas phase, once the collision is over, reactants move away from each other (if the reaction does not proceed). On the contrary, in the solid phase, reagents can easily form long-life complexes simply by being in two adjacent adsorption sites, and improve the overall efficiency of the reaction. Also in the case of reactions with an important barrier, solid-state chemistry allows hydrogenation to proceed by quantum tunnelling.

iv) Dust grains modify the reaction barriers, and especially can lower them, or as we can see in our study, the particular environment may impose a geometric constraint on the adsorbate and hinder certain reaction pathways that could have been fast in the gas phase. Figure 11 illustrates this aspect. Calculations have shown that HNO forms a H bond with an oxygen atom in the water cluster, limiting the access of H atoms that would diffuse at the entrance of the abstraction channel. Therefore, in this case, even if the hydrogenation reaction of HNO has a moderate barrier (1200K) and is certainly higher than that of the abstraction reaction channel (in the gas phase), the addition reaction is favoured (right sketch in Fig. 11). On the gold substrate, there is less geometric stress on the entry channel of the abstraction reaction, and thanks to its low barrier, it becomes a competitive reactive pathway. Therefore, we would like to add an important piece of evidence resulting from this study: v) the molecular environment (such as H₂O ice), can change the balance of the

different competitive reactive pathways, and most importantly, the H bond can prevent an abstraction reaction. Thus, our studies seem to show that cosmic grains not only have a catalytic role in amplifying reactivity, but also a catalytic role in guiding reactions, which can reduce the diversity of reactive pathways on grains.

4.3 Astrophysical implications and conclusions

Our present study does not fundamentally change the conclusions of the previous experimental works. At low temperatures, hydroxylamine is the major product of the direct hydrogenation of NO. However, if the grains temperature increases, hydrogenation becomes less efficient. Back reactions (abstraction reactions) have to be taken into account. In this reactive scheme, and under the assumption that H atoms are the most abundant reaction partner on interstellar dust grains, their accretion can produce in situ radicals, even from stable molecules such as NH₂OH that would already be accreted on the grain. In our case this radical is NHOH. These radicals, issued from hydrogenation of saturated molecules, could be the carrier of a new generation of COMs. In Dulieu et al., 61 the formation of formamide (NH₂CHO), is thought to be the result of the reaction of such radicals, and therefore, it is possible that the absence of detection of hydoxylamine in space is not due to low formation efficiency, but more likely to its further chemical evolution, in particular towards formamide if some H₂CO is also adsorbed on grains. The presence of such *in situ* radicals produced by hydrogenation has already been used to form COMs. ⁶² In the present case, because we have measured a fast inclusion of D in NH₂OH (Fig. 10), and found a low barrier to abstraction, we can identify the NHOH radical as a potential important missing link of the solid phase chemistry. In this way, NH₂OH is not the end of the NO hydrogenation, but a step toward inclusion of N in the skeleton of COMs.

Our final conclusion is more general and concerns the method. Conducted independently, both quantum chemistry and laboratory studies lead to conclusions that may be open to discussion, while when they are combined, they can be very conclusive. We believe

that it is now possible to disentangle complex chemical networks at play on cold surfaces through quantitative and various measurements, and to perform advanced quantum calculations that lift shadows and bring decisive arguments. Thanks to this synergy, the COMs synthesis pathways on interstellar grains are beginning to reveal their secrets.

Acknowledgement

TN thanks the Labex Michem program for her PHD grant. This work was supported by the Programme National "Physique et Chimie du Milieu Interstellaire" (PCMI) of CNRS/INSU with INC/INP co-funded by CEA and CNES, and by the DIM-ACAV+, a funding program of the Region Ile de France. This work was granted access to the HPC resources of [TGCC/CINES/IDRIS] under the allocation 2018 [A0040805116] made by GENCI.

Supporting Information Available

- The comparison of the binding energies of NO, NH₂OH, and N₂O between experimental data and models
- Energy barriers on the H + HNO reactions paths: Energies relative to H + HNO are in kJ/mol and are corrected for the unscaled ZPE energies. Basis sets are: ¹aug-cc-pVTZ, ²cc-pVQZ-F12, ³6-311++G(d,p), ⁴CBS, ⁵cc-pVQZ
- Energy barries on the H + H_2NOH reaction paths: M062X/aug-cc-pVTZ energies relative energies to H + H_2NOH are in kJ/mol and are corrected for the unscaled ZPE energies
- Structure and binding energies of NO, HNO, H_2NO , and NH_2OH on the 4 water clusters: Binding energies (Δ E) in kJ/mol are calculated at the M06X/6-311++G(d,p)

level as the difference between the molecule bond to the cluster and the sum of the energies of the molecule and the energy of the cluster. They are corrected for unscaled zero point energies calculated at the same level

• Structure for the H-HNO- $4H_2O$ transition state and energy barrier with respect to H + HNO- $4H_2O$: The energy barrier in kJ/mol is calculated at the M062X/6-311++G(d,p) level. It is corrected for the unscaled zero point energies calculated at the same level

References

- (1) Ando, R. et al. Diverse Nuclear Star-forming Activities in the Heart of NGC 253 Resolved with 10-pc-scale ALMA Images. *Astrophys. J.* **2017**, *849*, 81.
- (2) Lopez-Sepulcre, A.; Jaber, A. A.; Mendoza, E.; Lefloch, B.; Ceccarelli, C.; Vastel, C.; Bachiller, R.; Cernicharo, J.; Codella, C.; Kahane, C.; Kama, M.; Tafalla, M. Shedding light on the formation of the pre-biotic molecule formamide with ASAI. *Mon. Not. R. Astron. Soc.* **2015**, *449*, 2438–2458.
- (3) Pagani, L.; Favre, C.; Goldsmith, P. F.; Bergin, E. A.; Snell, R.; Melnick, G. The complexity of Orion: an ALMA view. I. Data and first results. *Astron. Astrophys.* **2017**, 604, A32.
- (4) Cernicharo, J.; Marcelino, N.; Roueff, E.; Gerin, M.; Jiménez-Escobar, A.; Caro, G. M. M. Discover of the methoxy radical, CH₃O, toward B1: Dust grain and gas phase chemistry in cold dark clouds. *Astrophys. J.* 2012, 759, L43.
- (5) Ceccarelli, C. et al. Seeds Of Life In Space (SOLIS): The Organic Composition Diversity at 300–1000 au Scale in Solar-type Star-forming Regions. *Astrophys. J.* **2017**, *850*, 176.
- (6) Shannon, R. J.; Caravan, R. L.; Blitz, M. A.; Heard, D. E. A combined experimental and theoretical study of reactions between the hydroxyl radical and oxygenated

- hydrocarbons relevant to astrochemical environments. *Phys. Chem. Chem. Phys.: PCCP* **2014**, *16*, 3466–78.
- (7) Dulieu, F.; Amiaud, L.; Congiu, E.; Fillion, J.-H.; Matar, E.; Momeni, A.; Pirronello, V.; Lemaire, J. L. Experimental evidence for water formation on interstellar dust grains by hydrogen and oxygen atoms. *Astron. Astrophys.* **2010**, *512*, A30.
- (8) Jing, D.; He, J.; Brucato, J.; De Sio, A.; Tozzetti, L.; Vidali, G. On water formation in the interstellar medium: Laboratory study of the O + D reaction on surfaces. *Astrophys. J.* **2011**, 741, L9.
- (9) Oba, Y.; Watanabe, N.; Kouchi, A.; Hama, T.; Pirronello, V. Experimental Study of CO₂ Formation by Surface Reactions of Non-energetic OH Radicals with CO Molecules. *Astrophys. J. Lett.* **2010**, *712*, L174–L178.
- (10) Ioppolo, S.; van Boheemen, Y.; Cuppen, H. M.; van Dishoeck, E. F.; Linnartz, H. Surface formation of CO2 ice at low temperatures. *Mon. Not. R. Astron. Soc.* **2011**, 413, 2281–2287.
- (11) Noble, J. A.; Dulieu, F.; Congiu, E.; Fraser, H. J. CO₂ formation in quiescent clouds: An experimental study of the CO + OH pathway. *Astrophys. J.* **2011**, *735*, 121.
- (12) Minissale, M.; Congiu, E.; Manicò, G.; Pirronello, V.; Dulieu, F. CO₂ formation on interstellar dust grains: a detailed study of the barrier of the CO + O channel. *Astron. Astrophys.* **2013**, *559*, A49.
- (13) Minissale, M.; Loison, J.-C.; Baouche, S.; Chaabouni, H.; Congiu, E.; Dulieu, F. Solidstate formation of CO₂ via the H₂CO + O reaction. *Astron. Astrophys.* **2015**, *577*, A2.
- (14) Hiraoka, K.; Sato, T.; Sato, S.; Sogoshi, N.; Yokoyama, T.; Takashima, H.; Kitagawa, S. Formation of Formaldehyde by the Tunneling Reaction of H with Solid CO at 10 K Revisited. *Astrophys. J.* **2002**, *577*, 265–270.

- (15) Watanabe, N.; Kouchi, A. Efficient Formation of Formaldehyde and Methanol by the Addition of Hydrogen Atoms to CO in H₂O-CO Ice at 10 K. *Astrophys. J.* **2002**, *571*, L173–L176.
- (16) Fuchs, G. W.,; Cuppen, H. M.,; Ioppolo, S.,; Romanzin, C.,; Bisschop, S. E.,; Andersson, S.,; van Dishoeck, E. F.,; Linnartz, H., Hydrogenation reactions in interstellar CO ice analogues A combined experimental/theoretical approach. *Astron. Astrophys.* **2009**, 505, 629–639.
- (17) Minissale, M.; Fedoseev, G.; Congiu, E.; Ioppolo, S.; Dulieu, F.; Linnartz, H. Solid state chemistry of nitrogen oxides–Part I: surface consumption of NO. *Phys. Chem. Chem. Phys.* **2014**, *16*, 8257–8269.
- (18) Ioppolo, S.; Fedoseev, G.; Minissale, M.; Congiu, E.; Dulieu, F.; Linnartz, H. Solid state chemistry of nitrogen oxides–Part II: surface consumption of NO₂. *Phys. Chem. Chem. Phys.* **2014**, *16*, 8270–8282.
- (19) Minissale, M.; Moudens, A.; Chaabouni, H.; ; Baouche, S.; Dulieu, F. Hydrogenation of CO-bearing species on grains: unexpected chemical desorption of CO. *Mon. Not. R. Astron. Soc.* **2016**, *458*, 2953–2961.
- (20) Velilla Prieto, L.; Sánchez Contreras, C.; Cernicharo, J.; Agúndez, M.; Quintana-Lacaci, G.; Alcolea, J.; Bujarrabal, V.; Herpin, F.; Menten, K. M.; Wyrowski, F. New N-bearing species towards OH 231.8+4.2. HNCO, HNCS, HC₃N, and NO. *Astron. Astrophys.* **2015**, *575*, A84.
- (21) Akyilmaz, M.,; Flower, D. R.,; Hily-Blant, P.,; Pineau des Forêts, G.,; Walmsley, C. M., The depletion of NO in pre-protostellar cores *. *Astron. Astrophys.* **2007**, *462*, 221–230.
- (22) Liszt, H. S.; Turner, B. E. Microwave detection of interstellar NO. *Astrophys. J. Lett.* **1978**, 224, L73–L76.

- (23) Congiu, E.; Chaabouni, H.; Laffon, C.; Parent, P.; Baouche, S.; Dulieu, F. Efficient surface formation route of interstellar hydroxylamine through NO hydrogenation. I. The submonolayer regime on interstellar relevant substrates. *J. Chem. Phys.* **2012**, *137*, 054713.
- (24) He, J.; Vidali, G.; Lemaire, J.-L.; Garrod, R. T. Formation of Hydroxylamine on dust grains via ammonia oxidation. *Astrophys. J.* **2015**, *799*, 49.
- (25) Fedoseev, G.; Chuang, K.-J.; van Dishoeck, E. F.; Ioppolo, S.; Linnartz, H. Simultaneous hydrogenation and UV-photolysis experiments of NO in CO-rich interstellar ice analogues; linking HNCO, OCN⁻, NH₂CHO, and NH₂OH. *Mon. Not. R. Astron. Soc.* **2016**, *460*, 4297–4309.
- (26) Congiu, E.; Fedoseev, G.; Ioppolo, S.; Dulieu, F.; Chaabouni, H.; Baouche, S.; Lemaire, J. L.; Laffon, C.; Parent, P.; Lamberts, T.; Cuppen, H. M.; Linnartz, H. NO ice hydrogenation: A soild pathway to NH₂OH formation in space. *Astrophysical. J.* **2012**, 750, L12.
- (27) Fedoseev, G.; Ioppolo, S.; Lamberts, T.; Zhen, J. F.; Cuppen, H. M.; Linnartz, H. Efficient surface formation route of interstellar hydroxylamine through NO hydrogenation. II. The multilayer regime in interstellar relevant ices. *J. Chem. Phys.* **2012**, 137, 054714.
- (28) Oba, Y.; Chigai, T.; Osamura, Y.; Watanabe, N.; Kouchi, A. Hydrogen isotopic substitution of solid methylamine through atomic surface reactions at low temperatures: A potential contribution to the D/H ratio of methylamine in molecular clouds. *Meteorit. Planet. Sci.* **2014**, *49*, 117–132.
- (29) Kobayashi, H.; Hidaka, H.; Lamberts, T.; Hama, T.; Kawakita, H.; Kästner, J.; Watanabe, N. Hydrogenation and Deuteration of C₂H₂ and C₂H₄ on Cold Grains: A Clue to

- the Formation Mechanism of C_2H_6 with Astronomical Interest. *Astrophys. J.* **2017**, *837*, 155.
- (30) Oba, Y.; Tomaru, T.; Lamberts, T.; Kouchi, A.; Watanabe, N. An infrared measurement of chemical desorption from interstellar ice analogues. *Nat. Astron.* **2018**, 2, 228–232.
- (31) Glarborg, P.; stberg, M.; Alzueta, M. U.; Dam-Johansen, K.; Miller, J. A. The recombination of hydrogen atoms with nitric oxide at high temperatures. *Symposium (International) on Combustion* **1998**, 27, 219 226, Twenty-Seventh Sysposium (International) on Combustion Volume One.
- (32) Soto, M. R.; Page, M. Ab initio variational transition—state—theory reactionrate calculations for the gas—phase reaction H+HNO→H₂+NO. *J. Chem. Phys.* **1992**, *97*, 7287–7296.
- (33) Nguyen, H. M. T.; Zhang, S.; Peeters, J.; Truong, T. N.; Nguyen, M. T. Direct ab initio dynamics studies of the reactions of HNO with H and OH radicals. *Chem. Phys. Lett.* **2004**, *388*, 94 99.
- (34) Vasyunin, A. I.; Herbst, E. Reactive desorption and radiative association as possible drivers of complex molecule formation in the cold interstellar medium. *Astrophys. J.* **2013**, 769, 34.
- (35) Ruaud, M.; Wakelam, V.; Hersant, F. Gas and grain chemical composition in cold cores as predicted by the Nautilus three-phase model. *Mon. Not. R. Astron. Soc.* **2016**, 459, 3756–3767.
- (36) Cuppen, H. M.; Walsh, C.; Lamberts, T.; Semenov, D.; Garrod, R. T.; Penteado, E. M.; Ioppolo, S. Grain Surface Models and Data for Astrochemistry. *Space Sci. Rev.* **2017**, 212, 1–58.

- (37) Vasyunin, A. I.; Caselli, P.; Dulieu, F.; Jiménez-Serra, I. Formation of Complex Molecules in Prestellar Cores: A Multilayer Approach. *Astrophys. J.* **2017**, *842*, 33.
- (38) Minissale, M.; Nguyen, T.; Dulieu, F. Experimental study of the penetration of oxygen and deuterium atoms into porous water ice. *Astron. Astrophys.* **2019**, 622, A148.
- (39) Zhao, Y.; Truhlar, D. G. The M06 suite of density functionals for main group thermochemistry, thermochemical kinetics, noncovalent interactions, excited states, and transition elements: two new functionals and systematic testing of four M06-class functionals and 12 other functionals. *Theor. Chem. Acc.* **2008**, *120*, 215–241.
- (40) Korth, M.; Grimme, S. Mindless DFT Benchmarking. *J. Chem. Theory Comput.* **2009**, *5*, 993–1003, PMID: 26609608.
- (41) Raghavachari, K.; Trucks, G. W.; Pople, J. A.; Head-Gordon, M. A fifth-order perturbation comparison of electron correlation theories. *Chem. Phys. Lett.* **1989**, *157*, 479 483.
- (42) Knizia, G.; Adler, T. B.; Werner, H.-J. Simplified CCSD(T)-F12 methods: Theory and benchmarks. *J. Chem. Phys.* **2009**, *130*, 054104.
- (43) Peterson, K. A.; Adler, T. B.; Werner, H.-J. Systematically convergent basis sets for explicitly correlated wavefunctions: The atoms H, He, B–Ne, and Al–Ar. *J. Chem. Phys.* **2008**, *128*, 084102.
- (44) Sylvetsky, N.; Peterson, K. A.; Karton, A.; Martin, J. M. L. Toward a W4-F12 approach: Can explicitly correlated and orbital-based ab initio CCSD(T) limits be reconciled? *J. Chem. Phys.* **2016**, *144*, 214101.
- (45) Karton, A.; Rabinovich, E.; Martin, J. M. L.; Ruscic, B. W4 theory for computational thermochemistry: In pursuit of confident sub-kJ/mol predictions. *J. Chem. Phys.* **2006**, 125, 144108.

- (46) Wakelam, V.; Bron, E.; Cazaux, S.; Dulieu, F.; Gry, C.; Guillard, P.; Habart, E.; Hornekær, L.; Morisset, S.; Nyman, G.; Pirronello, V.; Price, S. D.; Valdivia, V.; Vidali, G.; Watanabe, N. H₂ formation on interstellar dust grains: The viewpoints of theory, experiments, models and observations. *Mol. Astrophys.* **2017**, *9*, 1–36.
- (47) M.J. Frisch,; G.W. Trucks,; H.B. Schlegel,; G.E. Scuseria,; M.A. Robb,; J.R. Cheeseman,; G. Scalmani,; V. Barone,; B. Mennucci,; G.A. Petersson, Gaussian09, Revision A. *Inc.*, *Wallingford CT.* **2009**, *121*, 150–166.
- (48) Werner, H.-J.; Knowles, P. J.; Knizia, G.; Manby, F. R.; Schtz, M. Molpro: a general-purpose quantum chemistry program package. *Wiley Interdisciplinary Reviews: Computational Molecular Science*. **2012**, 2, 242–253.
- (49) Washida, N.; Akimoto, H.; Okuda, M. Nitrosyl hydride formed in the H + NO + M reaction system. *J. Phys. Chem.* **1978**, *82*, 2293–2299.
- (50) Karton, A.; Sylvetsky, N.; Martin, J. M. L. W4-17: A diverse and high-confidence dataset of atomization energies for benchmarking high-level electronic structure methods. J. Comput. Chem. 2017, 38, 2063–2075.
- (51) Fateley, W. G.; Bent, H. A.; Crawford, B. Infrared Spectra of the Frozen Oxides of Nitrogen. *J. Chem. Phys.* **1959**, *31*, 204–217.
- (52) Dulieu, F.; Congiu, E.; Noble, J.; Baouche, S.; Chaabouni, H.; Moudens, A.; Minissale, M.; Cazaux, S. How micron-sized dust particles determine the chemistry of our Universe. *Sci. Rep.* **2013**, *3*, 1338.
- (53) Minissale, M.; Fedoseev, G.; Congiu, E.; Ioppolo, S.; Dulieu, F.; Linnartz, H. Solid state chemistry of nitrogen oxides —Part I: surface consumption of NO. *Phys. Chem. Chem. Phys.* **2014**, *16*, 8257–8269.

- (54) Minissale, M.,; Dulieu, F.,; Cazaux, S.,; Hocuk, S., Dust as interstellar catalyst. I. Quantifying the chemical desorption process. *Astron. Astrophys.* **2016**, *585*, A24.
- (55) Boogert, A. A.; Gerakines, P. A.; Whittet, D. C. Observations of the Icy Universe. *Annu. Rev. Astron. Astrophys.* **2015**, *53*, 541–581.
- (56) Watanabe, N.; Nagaoka, A.; Shiraki, T.; Kouchi, A. Hydrogenation of CO on Pure Solid CO and CO–H₂O Mixed Ice. *Astrophys. J.* **2004**, *616*, 638–642.
- (57) R. Scott Smith,; C. Huang,; Kay*, B. D. Evidence for Molecular Translational Diffusion during the Crystallization of Amorphous Solid Water. *J. Phys. Chem. B* **1997**, 101, 6123–6126.
- (58) Mokrane, H.; Chaabouni, H.; Accolla, M.; Congiu, E.; Dulieu, F.; Chehrouri, M.; Lemaire, J. L. Experimental evidence for water formation via ozone hydrogenation on dust grains at 10 K. *Astrophys. J.* **2009**, 705, L195–L198.
- (59) Dulieu, F.; Minissale, M.; Bockelée-Morvan, D. Production of O₂ through dismutation of H₂O₂ during water ice desorption: a key to understanding comet O₂ abundances. *Astron. Astrophys.* **2017**, 597, A56.
- (60) Gould, R. J.; Salpeter, E. E. The Interstellar Abundance of the Hydrogen Molecule. I. Basic Processes. *Astrophys. J.* **1963**, *138*, 393.
- (61) Dulieu, F.; Nguyen, T.; Congiu, E.; Baouche, S.; Taquet, V. Efficient formation route of the prebiotic molecule formamide on interstellar dust grains. *Mon. Not. R. Astron. Soc. Lett.* 2019, 484, L119–L123.
- (62) Chuang, K.-J.; Fedoseev, G.; Ioppolo, S.; van Dishoeck, E. F.; Linnartz, H. H-atom addition and abstraction reactions in mixed CO, H₂CO and CH₃OH ices an extended view on complex organic molecule formation. *Mon. Not. R. Astron. Soc.* **2015**, 455, 1702–1712.

Graphical TOC Entry

