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1 **Mineral assemblages and temperature associated with Cu enrichment in the Seival**
2 **area (Neoproterozoic Camaquã Basin of Southern Brazil)**

3
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19
20 **Abstract**

21 The Neoproterozoic sequence of volcanic rocks in the Camaquã Basin included in the
22 Lavras do Sul Shoshonitic Association hosts disseminated Cu deposits. The volcanic
23 sequence in the Seival Mine area includes andesitic lava flows, lapilli tuff, volcanic
24 agglomerate and andesitic dikes with pervasive alteration. The hydrothermal alteration is
25 interpreted to be a product of late-magmatic fluids or the mixing of magmatic with
26 meteoric fluids or basinal brines. The late-magmatic hydrothermal system started after
27 volatile oversaturation and degassing in a magmatic system, which was partially segregated
28 into vesicles. The similar REE contents in the andesine-labradorite and albite indicate that
29 the fluid temperatures ranged from 600 to 250 °C. The crystallization of titanite and Fe-
30 clinocllore to Mg-chamosite followed the albitization process. The temperature range
31 associated with the chloritization process was estimated using a chlorite geothermometer

32 and varied from 251 ± 56 to 183 ± 39 °C. The precipitation of Mg–saponite and smectite–
33 rich chlorite/smectite mixed–layers in some andesitic lava flows and the precipitation of
34 quartz, calcite, barite and hematite in fractures suggest fluid circulation with temperatures
35 lower than the chloritization process. The alteration minerals are represented by chlorite
36 and albite associated with pyrite–chalcopyrite, while chlorite/smectite mixed–layers and
37 barite or hematite are associated with bornite–chalcocite–covellite. Thermodynamic
38 calculations confirm the potential coprecipitation of pyrite–chalcopyrite with chlorite and
39 albite. The occurrence of bornite–chalcocite–covellite and barite seems to be favored by
40 low–temperature chlorite/smectite with neutral to mildly acidic water influx. Consequently,
41 albitization and the process of chloritization at high temperatures ($> ca. 251 \pm 56$ °C) can be
42 used as an exploration guide for primary pyrite–chalcopyrite enrichment, and
43 chlorite/smectite at low temperatures (ca. 250 to 50 °C) can be related to bornite–
44 chalcocite–covellite or the process of Cu enrichment.

45 **Keywords:** Late–magmatic and hydrothermal alteration; Albitization; Chlorites.

46

47 **1. Introduction**

48 The Neoproterozoic Camaquã Basin and its associated igneous rocks contain significant
49 Au, Cu and Ag ore deposits, such as the Uruguay–São Luiz mines and the Lavras do Sul
50 Mining District (Remus et al., 2000; Bongioiolo et al., 2011; Renac et al., 2014). In the
51 central–western section of this basin, volcanogenic (Lopes et al., 2014; Fontana et al.,
52 2017) and plutonic rocks (Mexias et al., 2005) show pervasive alteration described as late–
53 magmatic albitization and hydrothermal alteration with the influx of meteoric fluids
54 (Mexias et al., 1990a, 2005; Bongioiolo et al., 2011; Fontana et al., 2017). The largest Cu
55 deposit in this basin is the Camaquã Mine, which is hosted in a Neoproterozoic sedimentary
56 sequence and was exploited from the 18th century until 1996 (Remus et al., 2000; Laux et
57 al., 2005; Renac et al., 2014). This exploited deposit has an ore reserve of ~30.8 Mt with an
58 average Cu grade of ~1.06%, which is manifested as massive veins, stockworks and
59 disseminated chalcopyrite, pyrite, bornite, chalcocite, gold, silver and hematite. The
60 extrusive and intrusive rocks from the Lavras do Sul Mining District (Mexias et al., 2007),
61 which are located on the central–western rim of the Camaquã Basin, host nineteen Au and
62 Au–Cu deposits that were exploited during the same time frame (Bongioiolo et al., 2011).

63 The ore zones consist of quartz and sulfide-rich quartz veins (Mexias et al., 2005)
64 associated with wall-rock alteration characterized by 'sericite', chlorite, pyrite and
65 chalcopyrite (Bongiolo et al., 2011). In the vicinity of the Lavras do Sul area, the
66 volcanogenic sequence in the Seival area hosts Cu-Ag deposits (Reischl, 1978; Lopes et
67 al., 2014). The different mines yielded ca. 0.20 Mt of Cu-ore with an average grade of
68 ~1.40%. The chalcopyrite, pyrite, chalcocite, bornite and covellite sulfides occurred in
69 association with chlorite, hematite, calcite, barite and quartz disseminated in andesitic
70 lapilli tuff and volcanic agglomerates or along fault planes cutting the volcanic sequence. A
71 recent study on the Seival Mine (Fontana et al., 2017) improved current knowledge about
72 the clay minerals in these ore deposits. Fontana et al. (2017) described the zoning and
73 lithotypes with the predominant occurrence of smectite, chlorite/smectite mixed-layers,
74 chlorite and corrensite, and suggested that the base-metal deposition was related to the
75 late-magmatic process of albitization and high-temperature hydrothermal alteration. Later
76 low-temperature hydrothermal alteration represented the mixing of magmatic and meteoric
77 fluids. Both albitization and chloritization processes were observed in the granite-related
78 epithermal systems of the Lavras do Sul area (Mexias et al., 2007; Bongiolo et al., 2011)
79 and the Camaquã Mine (Laux et al., 2005; Renac et al., 2014). A similar alteration
80 mineralogy in the vicinity of the Lavras do Sul Intrusive Complex, located 10 km west of
81 the Seival area, suggests that the Lavras do Sul porphyry-epithermal system contributes to
82 the alteration of the andesitic rocks and ore deposition in the Seival area, as also occurs in
83 the Potrerill Cu-Mo porphyry deposits and El Hueso Au epithermal deposits (Sillitoe,
84 1991; Marsh et al., 1997).

85 Consequently, this study focused on the chemical signatures of albite and chlorite to
86 smectite in the hydrothermal alteration products. The petrographic and chemical analyses of
87 the minerals were used to estimate the temperatures and bracket temperatures of late-
88 magmatic to hydrothermal alteration. Petrographic observations and temperatures estimated
89 by chlorite geothermometers were introduced into a thermodynamic model. This model was
90 used to reconstruct the chemical conditions favorable to ore deposition in the process of
91 base-metal migration in the Seival area.

92

93 **2. Geological Setting**

94 **2.1 Regional geology**

95 The Mantiqueira Province extends from the southeast of Brazil to Uruguay (Almeida
96 and Hasui, 1984; Silva et al., 2005). In the Gondwana supercontinent, the southern portion
97 of collision between the Rio de La Plata (Brazil) and the Kalahari (Africa) cratons is
98 characterized by the Sul–rio–grandense Shield (Fig. 1A) with Neoproterozoic ages. This
99 collision is represented by the Brazilian Orogeny and the opening of the Camaquã Basin
100 related to the post–collisional stage (Lima and Nardi, 1998; Gastal et al., 2005a, b). The
101 Sul–rio–grandense Shield comprises four geotectonic units: the Taquarembó Terrane, São
102 Gabriel Terrene, Tijucas Terrene, and Pelotas Batholith. Recovering these units, the
103 Camaquã Basin was considered as a succession of sub–basins (Paim et al., 2000) that
104 contains volcano–sedimentary sequences. The Maricá Group, which is represented by the
105 volcano–sedimentary rocks of the Maricá Formation, is related to a foreland retro arc–basin
106 setting (630–601 Ma; Paim et al., 2000; Borba et al., 2008). The Bom Jardim Group, in
107 which shoshonitic volcanism was emplaced during the strike–slip movement of the basin
108 (600–590 Ma; Janikian et al., 2008), is represented by the Hilário Formation. The late Santa
109 Bárbara Group (574–544 Ma; Sommer et al., 2006; Bicca et al., 2013; Almeida et al., 2012;
110 Janikian et al., 2012; Matté et al., 2016) is mainly represented by the acidic
111 effusive/explosive episodes and minor basic magmatism of the Acampamento Velho
112 Formation and by the sedimentary sequences of the Santa Fé and Santa Bárbara formations
113 (Fig. 1B). The Guaritas Group (547–535 Ma; Almeida et al., 2012; Oliveira et al., 2014) is
114 represented by volcanic rocks of the Rodeio Velho Formation associated with a sedimentary
115 sequence represented by the Pedra Pintada and Guaritas formations related to transtensional
116 tectonic stages (Paim et al., 2000; Almeida et al., 2012). These effusive rocks have basic to
117 intermediate sodic compositions with mildly alkaline to high–K tholeiitic affinities
118 (Almeida et al., 2012).

119

120 **INSERT Figure 1**

121

122 The volcano–sedimentary units that occur in the Seival area (Fig. 2) have dominant
123 intermediate compositions with basic and acidic end–members. The Bom Jardim Group, as
124 previously described, consists of the effusive units of the Hilário Formation in the study

125 area (Ribeiro and Fantinel, 1978). The deformation of the Hilário Formation along the
126 Lavras do Sul contact suggests that the Hilário Formation is older than the Lavras do Sul
127 Intrusive Complex (Gastal et al., 2015). The Hilário Formation consists of dominantly
128 andesitic rocks represented by andesitic lava flows, lapilli tuff, volcanic agglomerates, and
129 dikes with a shoshonitic geochemical affinity (Lopes et al., 2014; Fontana et al., 2017). The
130 Ar–Ar radiogenic dating of plagioclase suggests an age of 600 to 590 Ma (Janikian et al.,
131 2008) that might be thermally influenced by intrusive magmas such as dikes or the Lavras
132 do Sul granitoids. The other vesicular lava flows and volcanic clastic materials are
133 represented by the polymictic conglomerates of the Santa Fé Formation. The Acampamento
134 Velho Formation occurs as rhyolite dikes that cross-cut the Hilário Formation.

135

136 **INSERT Figure 2**

137

138 The granitoids of the Lavras do Sul Intrusive Complex with U–Pb ages varying from
139 601 to 587 Ma (SHRIMP in zircon, Remus et al., 1999, 2000; Gastal et al., 2006; Liz et al.,
140 2009), comprise a ring structure composed of porphyritic quartz–monzonite granodiorite
141 and syenogranite to monzogranite with monzonitic intrusions in its northeastern margin
142 (Gastal et al., 2006; Liz et al., 2009). In the contact zone between the Lavras do Sul
143 granitoids and the Hilário Formation is a hectometer-thick metamorphic halo (Mexias et al.,
144 1990a, b; Bongiololo et al., 2011) attesting to the thermal influence of granitoids on the
145 volcano–sedimentary pile.

146

147 **2.2 Ore geology**

148 The Seival hosts six inactive Cu–bearing deposits and several prospects. This reserve
149 was estimated at ~0.20 Mt of ore with an average grade of ~1.40% of Cu and 10–70 ppm of
150 Ag (Fig. 3A; Reischl, 1978; Lopes et al., 2014). The ore deposits in the Seival area were
151 associated with the development of NE–SW-oriented shear zones that cross-cut the
152 Camaquã Basin (Fernandes et al., 1992; Philipp et al., 2016). The emplacement of metallic
153 mineralizations are bracketed between 600 and 580 Ma ago based on the ages of lava flows
154 (Janikian et al., 2008), tectonic activity (Gastal et al., 2015) and the Lavras do Sul
155 magmatism (Gastal et al., 2006; Almeida et al., 2012).

156 The sulfides are represented by chalcopyrite, bornite, chalcocite and covellite (Fontana
157 et al., 2017) and are part of the main Cu-bearing veins that fill fractures and faults. The
158 alteration halos are related to the distances of mineralized structures (Lopes et al., 2014;
159 Fontana et al., 2017). The mineral association of chlorite, epidote and calcite formed at
160 temperature of ~300 °C (Mexias et al., 1990b; Mexias et al., 2007; Bongiolo et al., 2011) in
161 the volcanogenic rocks of the Hilário Formation close to the Lavras do Sul granitoids.

162 The mineral assemblages allow for the characterization of hydrothermal zones
163 (Seedorff et al., 2005; Sillitoe, 2010). Propylitic alteration was related to chlorite, C/S
164 mixed-layer, hematite, titanite, calcite, barite and quartz associated with pyrite–
165 chalcopyrite and chalcocite–bornite, while argillic alteration was related to smectite and
166 quartz (Fig. 3B; Lopes et al., 2014; Fontana et al., 2017). Cu enrichment was associated
167 with propylitic alteration and the formation of albite (600 to 250 °C; Fontana et al., 2019)
168 and Mg–chamosite to Fe–clinocllore (251 to 180 °C). The late–hydrothermal veins are
169 composed of barite, calcite and minor proportion of veins with calcite–quartz assemblage.
170 These late precipitations represent the lowest temperatures of hydrothermal alteration (150
171 to 80 °C; Fontana et al., 2017). In the preserved parts of these inactive mines,
172 mineralization occurs as a stockwork of veins or disseminated in volcanoclastic rocks and
173 andesitic lava flows (Fontana et al., 2017). The faults, mineralized structures and dikes are
174 considered contemporaneous based on the mineralogy and alteration type. Thus, the
175 hydrothermal processes can be associated with the intrusion of NE–SW-oriented andesite
176 (Lopes et al., 2018) and local hydrothermal cells.

177

178 **INSERT Figure 3**

179

180 **3. Analytical methods**

181 Field work consisted of sample collection and logging of drill cores. The samples
182 collected were representative of lapilli tuff, amygdaloidal to porphyritic andesitic lava
183 flows and dikes. Petrographic observations of thin sections allowed the description of
184 magmatic and hydrothermally altered crystals. The minerals, after of crushing and sieving,
185 were separated using a magnetic separator, heavy liquids and hand-picking under a
186 binocular microscope. Products of alteration were either collected under a binocular

187 microscope in the altered phenocrysts or identified in the clay-size fraction. These
188 petrographic observations, mineral separation and clay-size fraction were characterized by
189 scanning electron microscopy combined with energy-dispersive spectrometry (SEM-EDS,
190 Zeiss EVO LS15, 20 kV, 2.5 nA with X-Max Oxford EDS detectors, IAEA Monaco) and
191 X-ray diffraction (Siemens D5000, $K\alpha$ Cu, 40 kV and 25 mA, Geoscience Institute-CPGq-
192 UFRGS). X-ray patterns of the clay minerals were obtained in $< 0.4 \mu\text{m}$ clay-size fraction.
193 Major element composition as oxides (SiO_2 , TiO_2 , Al_2O_3 , $\text{FeO}_{\text{total}}$, MnO , MgO , CaO , Na_2O ,
194 F and K_2O) in minerals from thin section were quantified by electron microprobe
195 (CAMECA, SXFive model equipped with TAP, PET and LIF crystal detectors) at the
196 Geoscience Institute, Universidade Federal do Rio Grande do Sul, Brazil. The beam
197 diameter is about 10 to 4 μm with an acceleration voltage of 15 kV and 15 nA with
198 counting time per element of 10 sec and background counting times of 5 sec per element.
199 Chemical calibrations were performed on albite, jadeite, diopside, sanidine, rutile,
200 rhodonite and hematite standards mounted in epoxy resin. The mineral composition in thin
201 section were calculated of atom per formula unit (a.p.f.u.) assumed the following number of
202 oxygen atoms for ideal compositions of feldspar (32), pyroxene (12), biotite (22), titanite
203 (5), smectite (22), chlorite/smectite mixed-layer (28) and chlorite (28). Charge balance of
204 structural formula was obtained by the adjustment of Fe^{2+} and Fe^{3+} proportions, including
205 for smectite and biotite, where only Fe^{2+} was considered. Analyzes in ten thin sections were
206 performed to follow changes from preserved magmatic minerals to mineral transformed or
207 precipitated by the hydrothermal alteration. Based on major element compositions,
208 chemical changes will then be related to albitization and chloritization processes. The
209 complete dataset of primary and hydrothermal minerals analyzed are available in
210 Supplementary Table A-G.

211 The major element changes observed in plagioclase were completed by rare earth
212 elements (REE; lanthanides, Y and Sc) analyzes. These REEs have been measured from
213 plagioclase separates preliminary recovered for phenocrysts and the microlithic
214 groundmass subsequently recognized with different major element compositions. These
215 REE concentrations were measured from feldspar separates preliminary acid leached with
216 ultrapure HNO_3 (1N) to dissolve traces carbonates and oxides on the surface of the grains
217 separated then clean in ultrapure pure MQ water. Later, cleaned grains were then acid

218 digested in ultrapure HNO₃ (15N) and HF (20N). Inductively-coupled plasma mass
219 spectrometry analyzed all digested samples diluted in 1% HNO₃ solutions (ICP-MS: Elan
220 DRCII, Perkin Elmer at the University of Nice Sophia-Antipolis). Quantification limits are
221 below 50 ng/L. Element concentration was expressed as ppm and then normalized with
222 respect to chondrite values (McDonough and Sun, 1995). The Eu* the values of the Eu
223 anomaly were calculated using $[Eu^*] = [(Eu_N)/(Sm_N \times Gd_N)^{1/2}]$ from Worrall and Pearson
224 (2001) or $[Eu^*] = [((Eu_N)/(Sm_N^2 \times Tb_N)^{1/3})]$ from Lawrence and Kamber (2006).

225 The different mineralogical and chemical data of magmatic minerals were used with the
226 petrogenetic model (MELTS software modeling; Ghiorso and Sack, 1995) to reconstruct
227 magmatic conditions. The magmatic and hydrothermal minerals were combined in a
228 thermodynamic model (PHREEQC interactive; Parkhurst, 1995) to estimate the favorable
229 conditions for precipitation of hydrothermal phases and sulfides.

230

231 **4. Results**

232 **4.1 Textures and petrographic observations**

233 The host rocks of the Cu deposits in the Seival area are predominantly andesitic lava
234 flows and lapilli tuff (Fig. 4A) to volcanic agglomerates with rounded andesitic pyroclasts.
235 Their aphyric groundmasses are dark to glassy and green with disseminated microlites of
236 plagioclase and fragments of pyroxene. Amygdaloidal and porphyritic andesitic lava flows
237 are predominant in the southern and northeastern parts of the Seival area and are in contact
238 with the Santa Fé Formation (Fig. 2). Subvolcanic rocks have andesitic (Fig. 4B) to
239 trachyandesitic compositions with feldspar phenocrysts, dominant pyroxene and scarce
240 biotite phenocrysts.

241 Most of the rocks are intensively fractured and brecciated, particularly those with
242 higher Cu sulfide contents. The andesitic pyroclasts in the lapilli tuff show vesicles or
243 amygdaloids. The amygdaloidal groundmass in the lapilli tuff consists of fragments of
244 igneous plagioclase and pyroxene. The petrographic observations of pyroxene revealed the
245 replacement by clay minerals and calcite or intergranular spaces filled by clay minerals,
246 quartz, calcite, titanite, pyrite, chalcopyrite, and minor hematite (Fig. 4C).

247 Porphyritic andesitic lava flows have phenocrysts of plagioclase and clinopyroxene
248 in a trachytic groundmass with oriented plagioclase. Andesitic dikes host subhedral grains

249 of feldspars further identified as plagioclase with rims of K-feldspar (Fig. 4D) and
250 phenocrysts of pyroxene. Some trachyandesitic dikes have subhedral phenocrysts of
251 plagioclase, pyroxene and biotite. Some plagioclase and biotite phenocrysts have inclusions
252 of apatite, zircon (50 to 200 μm) and titanite, while the oxidized groundmass contains
253 hematite. The titanite grains occur as euhedral grains associated with biotite and pyroxene
254 and as anhedral grains of smaller granular agglomerates in the altered groundmass.

255 As previously described (Lopes et al., 2014; Fontana et al., 2017), the Hilário
256 volcanic sequence in the studied area is pervasively altered to clay minerals, calcite, barite
257 and quartz. The original igneous mafic minerals and different groundmasses have a light
258 green material recognized as chlorite/smectite mixed-layer (Fig. 4E). Albite, chlorite and
259 sulfide occur in the more altered rocks with traces of calcite and specular hematite, and
260 euhedral or agglomerates of titanite occur with albite and microscopic euhedral quartz (Fig.
261 4F).

262 Less-altered material occurs in the amygdaloidal andesitic lava flows from the south
263 and west. Nevertheless, these lava flows are partially altered to dark clay minerals (Fig. 4G)
264 between crystals or exhibit quartz, pyrite-chalcopyrite and calcite filling vesicles.
265 Petrographic observations indicate the predominance of clay minerals associated with
266 sulfides in the intergranular porosity that are further recognized as chlorite and dominant
267 pyrite-chalcopyrite, or the association of chlorite/smectite mixed-layer (Fig. 4H) with
268 minor pyrite-chalcopyrite and dominant chalcocite-covellite or chalcocite-bornite. The
269 malachite and azurite observed in the different open pits are the product of the recent
270 weathering of Cu-rich sulfides.

271

272 **INSERT Figure 4**

273

274 **4.2 Major element compositions of magmatic to late-magmatic minerals and REE** 275 **compositions of feldspars**

276 Phenocrysts and microlites or crystalloclasts of pyroxene in porphyritic andesitic
277 lava flows and andesitic dikes are augite ($\text{Wo}_{0.40-0.47} \text{En}_{0.32-0.41} \text{Fs}_{0.12-0.28}$; Fig. 5A;
278 Supplementary Table A). According to the structural formula of augite (Table 1), the atomic

279 ratios of $[Mg/(Mg+Fe_{tot})= 0.67-0.85]$ are high and those of $[Na/(Ca+Na)= 0.02-0.05]$ are
280 low.

281 Prismatic biotite phenocrysts recognized in trachyandesitic dikes are classified as
282 Mg-biotite with the structural formula expressed in Supplementary Table B. The less
283 pristine Mg-biotites show slightly different compositions (Table 1), with increases in Al^{VI}
284 (Fig. 5B) and lower contents of Fe, Mg and Mn (Table 1). The euhedral to agglomerate
285 titanite grains (Fig. 5C; Table 1 and Supplementary Table C) are associated with albite and
286 biotite phenocrysts, and the anhedral grains in the groundmass have similar chemical
287 compositions.

288 The chemical compositions of feldspars were analyzed in thin section. The
289 plagioclase phenocrysts have compositions of andesine-labradorite ($An_{0.43-0.53} Ab_{0.43-0.54}$
290 $Or_{0.03-0.05}$) (Table 1). Moreover, some of the andesine-labradorite phenocrysts have a thin
291 rim of K-feldspar ($An_{0-0.01} Ab_{0-0.33} Or_{0.66-1}$; Fig. 4B and Fig. 5D; Table 1), further
292 described as Group I. In contrast, the microlites and phenocrysts of plagioclase
293 (Supplementary Table D) from altered lapilli tuff, porphyritic andesitic lava flows and
294 andesitic dikes consist of dominant albite ($An_{0-0.06} Ab_{0.89-1} Or_{0-0.10}$; Table 1), further
295 described as Group II. The feldspar separates were characterized by X-ray diffraction and
296 observed by SEM-EDS. The separates with traces of illite or quartz were removed from
297 further REE measurements. The total REE content of andesine-labradorite (Group I) and
298 albite (Group II) have similar patterns and are slightly higher in albite. However, the
299 andesine-labradorite (Group I) and albite (Group II) have both positive have positive Eu
300 anomalies (Group I: $[Eu^*]= 0.90-1.22$; Group II: $[Eu^*]= 1.04-4.60$; Table 2).

301

302 **INSERT Figure 5**

303 **INSERT Table 1**

304 **INSERT Table 2**

305

306 **4.3 Mineralogy and major elemental compositions of clay minerals**

307 Petrographic observations show “green clays” replacing pyroxene, biotite or
308 feldspars that fill vesicles and intergranular porosity. Similar to Fontana et al. (2017), using
309 X-ray diffraction techniques on clay-size fractions allowed for the identification of

310 dominant trioctahedral smectite and smectite-rich (C/S) in less altered andesitic lava flows,
311 as well as chlorite-rich chlorite/smectite mixed-layer and chlorite in more altered andesitic
312 lava flows and lapilli tuff. The occurrence of different clay minerals for similar andesitic
313 lithotypes suggests that hydrothermal alteration is not completely related to the primary
314 magmatic mineralogy but to the influx of external fluid. The in-situ electron microprobe
315 analyses allowed for the identification of the composition of the large chlorite flakes (Fig.
316 6) in the cleavage of andesine-labradorite and albite and the replacement of augite
317 phenocrysts or micrometer vugs in the andesitic dikes with a light green color (Table 1).
318 Based on the $\text{Fe}^{2+}/(\text{Fe}^{2+}+\text{Mg})$ ratio, these chlorites have compositions of Fe-clinocllore to
319 Mg-chamosite (Supplementary Table E). The observed $\text{Fe}^{2+}/(\text{Fe}^{2+}+\text{Mg})$ atomic ratios
320 between 0.47 and 0.64 suggest that values < 0.5 are Fe-clinocllore and those of > 0.5 are
321 Mg-chamosite. These two types of chlorite were also recognized in a regional epithermal
322 system (Fig. 7; Mexias et al., 1990b; Bongiolo et al., 2011) and distal-epithermal halo from
323 the Camaquã Mine (Fig. 7; Troian et al., 2010; Renac et al., 2014). Low $\text{Fe}^{2+}/(\text{Fe}^{2+}+\text{Mg})$
324 proportions are structurally compensated by tetrahedral and octahedral aluminum ($\text{Al}^{\text{IV}}=$
325 1.7 to 3.4; $\text{Al}^{\text{VI}}= 2.7$ to 3.11). The Fe-clinocllore to Mg-chamosite in andesitic lava flows
326 and lapilli tuff have lower $\text{Fe}^{2+}+\text{Mg}$ atomic proportions (7.5 to 8.5; atomic ratios of
327 $\text{Fe}^{3+}/\text{Fe}_{\text{tot}}= 0.01$ to 0.06) than in andesitic dikes ($\text{Fe}^{2+}+\text{Mg}= 9$ to 10.1; atomic ratio of
328 $\text{Fe}^{3+}/\text{Fe}_{\text{tot}}= 0.05$). However, the chloritization of augite and filling of vugs reflect increasing
329 contents of Fe+Al and Fe+Mg total contents in lapilli tuff and andesitic dikes, respectively.
330 In addition, the Fe-clinocllore to Mg-chamosite compositions are very similar (Fig. 8A;
331 Table 1) and were described using the same structural formula.

332

333 **INSERT Figure 6**

334 **INSERT Figure 7**

335 **INSERT Figure 8**

336

337 The occurrence of dominant chlorite/smectite, trioctahedral smectite and corrensite
338 recognized by their dark green color and identified using X-ray diffraction (Fig. 8) was
339 confirmed by electron microprobe measurements. The smectite-rich chlorite/smectite,
340 smectite (Supplementary Table F) and corrensite (Fig. 8B and Figs. 9A–B) represent the

341 dominant hydrothermal phase of andesitic lava flows and andesitic dikes. The dominant
342 trioctahedral smectite (Fig. 8C) has a Mg–saponite composition (Table 1). The dominant
343 chlorite–rich chlorite–smectite (Fig., 8D; Supplementary Table G) with micrometer-scale
344 fan–shaped particles (Fig. 4C) occurs in altered porphyritic and amygdaloidal andesitic lava
345 flows. The chlorite–rich chlorite–smectite have Al^{VI} and ^{IV} contents (Al_{tot} 2.21+ 1.25 to
346 2.46+ 1.48 a.p.f.u.) that are higher in the porphyritic andesitic lava flows than the chlorite–
347 rich chlorite–smectite (Al_{tot}= 1.59+ 0.77 to 1.93+ 1.47) in the amygdaloidal andesitic lava
348 flows. The proportion of Fe and its valency balance are represented by changes in
349 tetrahedral and octahedral Al. The chlorite–rich chlorite–smectite structural formulas vary
350 between porphyritic and amygdaloidal andesitic lava flows (Table 1).

351

352 **INSERT Figure 9**

353

354 It is possible to confirm the association of albite and chlorite minerals with Cu
355 enrichment. The magmatic and post–magmatic stages were related to albite and chlorite
356 crystallization, while chlorite, chlorite/smectite mixed–layer and the destabilization of
357 sulfides were related to the precipitation of chalcocite–bornite. These mineral associations
358 suggest a sequence of crystallization and hydrothermal alteration over time.

359

360 **5. Discussion**

361 **5.1 Chemistry of magmatic and late–magmatic minerals**

362 The volcanogenic rocks in the Seival area are interpreted as part of the Lavras do
363 Sul Shoshonitic Association (Lima and Nardi, 1998), which are related to the Lavras do Sul
364 granitoids and lamprophyres. The occurrence of lamprophyres (Lima and Nardi, 1998;
365 Almeida et al., 2012), andesitic dikes and various volcanic events in the region suggests
366 long and intense episodes of magmatism in the Seival area constrained by the dating of
367 zircon (Almeida et al., 2012; Janikian et al., 2012) between 600 ± 2 to 590 ± 6 Ma. The
368 mineralogy of andesitic rocks with andesine–labradorite, augite and Mg–biotite phenocrysts
369 and their whole–rock compositions (Lima and Nardi, 1998; Sommer et al., 2006; Lopes et
370 al., 2014) suggest a petrogenetic link with the Lavras do Sul Shoshonitic Association.

371 Petrogenetic modeling (MELTS Software: Ghiorso and Sack, 1995) used with
372 whole-rock compositions (Lopes et al., 2014) estimates a liquidus temperature of ca. 1080
373 °C. The modeling shows a sequence of crystallization initiated by calcic plagioclase
374 followed by low-Ca-pyroxene, late orthoclase and high-Ca pyroxene or biotite (ca. 800
375 °C). The stages below 800 °C include late-magmatic minerals formed by volatile-rich melt
376 (c.f. Dixon et al., 1995). In this context, the biotites observed with the higher Al^{IV} contents
377 of less pristine biotite compared with the biotite phenocrysts might represent two stages of
378 biotite crystallization or slightly altered crystals. The late-magmatic stage initiated at
379 solidus temperatures (800 to 500 °C) with volatile oversaturation and the degassing of the
380 melt (c.f. Dixon et al., 1995). This interpretation is reinforced by the oxygen isotopic values
381 of andesine-labradorite ($\delta^{18}\text{O}_{\text{andesine-labradorite}}$: 6.6 to 8.2 ‰; Fontana et al., 2019) showing
382 the magmatic signature of fluids in equilibrium with feldspars (Taylor, 1987) evolving
383 toward the mixing of meteoric and magmatic fluids for albite ($\delta^{18}\text{O}_{\text{albite}}$: 14.2 to 16.2 ‰;
384 Fontana et al., 2019). Consequently, presume that observed Mg-biotite formation, euhedral
385 F-rich titanite grains and albite are related to late-magmatic crystallization and albitization
386 processes of andesine-labradorite. According to thermodynamic modelling and the
387 literature (c.f. Higgins and Ribb, 1976; Franz and Spear, 1985; Enami et al., 1993; Tropper
388 et al., 2002; Harlov et al., 2006; Hövelmann et al., 2010; Kaur et al., 2012) this albitization
389 process occurs at temperatures below 800 °C, presumably from ca. 600 to 250 °C. The
390 widely observed albite suggests a magmatic to high-temperature hydrothermal system
391 (epithermal-porphyry: Richards, 2003; Sillitoe, 2010 or Iron Oxide-Copper-Gold; Sillitoe,
392 2003; Williams et al., 2005; Groves et al., 2010). The alkaline and acidic fluids in these
393 systems are related to the albitization process and K-rich alteration (Perez and Boles, 2005;
394 Boulvais et al., 2007) that, to some degree, participated or started the leaching of base
395 metals from the lava (Moody et al., 1985; Perez and Boles, 2005; Hövelmann et al., 2010;
396 Kontonikas-Charos et al., 2014). Non-magmatic basinal brine has also reported in the
397 albitization process (Barton and Johnson, 1996 and 2000), such as in the Nevada porphyry,
398 IOCG deposits (Carten, 1986; Dilles and Einaudi, 1992; Dilles et al., 2000; Sillitoe, 2003;
399 Mark et al., 2006; Kendrick et al., 2007) or Chilean “Manto-type” deposits, such as
400 Michilla, with the mixing of magmatic and basinal brine (Barton, 2014). In the Seival area,
401 the change of andesine-labradorite (An_{0.43-0.53} Ab_{0.43-0.54} Or_{0.03-0.05}) to albite (An_{0-0.06} Ab

402 0.89–1 Or 0–0.10) with rare K–feldspar rims on andesine–labradorite or rare illite attest to the
403 circulation of alkaline fluid. Moreover, the sizes of the quartz and calcite crystals with
404 liquid inclusions did not allow fluid inclusion microthermometry to recognize high–density
405 fluid, as in IOCG-type deposits (Reeve et al., 1990; Davidson and Large, 1998). The REE
406 patterns of albite-rich separates (Group II: $Eu^* = 1.04–4.60$) are similar to the whole–rock
407 (Fig. 10; Lopes et al., 2014). The more positive Eu anomaly of group II and whole–rock
408 (whole–rock: $Eu^* = 6.23$) compared to Group I (Labradorite $Eu^* = 0.90–1.20$). These
409 overall, positive Eu positive anomalies ($Eu^* > 1$; Table 2) higher in the albite (group II;
410 Fig. 10A–B) than andesine–labradorite–rich grains (group I) seem unrealistic assuming,
411 first crystallization of Ca–feldspar then Na–rich feldspar. However, the more positive Eu
412 anomalies of whole–rock and Group II might be related to trace minerals such as apatite
413 inclusions in plagioclase, as mentioned in the petrography section and whole–rock. These
414 inclusions have not been removed totally by preliminary acid–leaching used to clean
415 feldspars of carbonates, Ca–phosphates and Fe oxides (Fontana et al., 2017). The REE
416 signature of Ca–rich feldspar would then represent a chronologically primitive signature in
417 the magmatic chamber, while, the late crystallization include dominant Na–rich feldspar
418 and phosphate traces.

419 Nevertheless, the similar \square REE contents with LREE and HREE variations from
420 albites and andesine–labradorite suggest a similar source of REE (\square REE) that was not
421 affected by fluid dilution, such as that by meteoric fluids. Therefore, the albitization process
422 in the Hilário Formation was related to the neutral to mildly acidic water circulation of Na–
423 rich magmatic fluid, such as late–magmatic glass interactions with outgassing magmatic
424 fluids (CO_2 and H_2O), but these data cannot decipher between a magmatic to late–
425 magmatic crystallization with the mixing of magmatic fluids and brine (e.g. IOCG: Barton
426 and Johnson, 1996 and 2000).

427

428 **INSERT Figure 10**

429

430 No mineralization model has yet been proposed for the Seival Mine. Fontana et al.
431 (2017) indicated mineralization associated with magmatic and meteoric fluids. The aligned
432 Cu–Ag ore bodies and disseminated mineralizations (see Figure 3) with similar mineral

433 assemblages (albite, chlorite, chlorite/smectite mixed-layers, chalcocite-bornite, barite, and
434 minor calcite, hematite and quartz) could be typical of epithermal deposits (Sillitoe, 2010)
435 associated with the geothermal influence of andesitic dikes or the Lavras do Sul granitoids.
436 These observations and interpretations suggest that Seival would be a superficial part of the
437 Lavras do Sul porphyry (Mexias et al., 2007; Bongiolo et al., 2011). We assumed that low-
438 hematite contents in the Seival area associated with calcite in small veins and ore deposits
439 without an iron-oxide caprock are clues discarding the hypotheses of an IOCG type deposit
440 or a metallic enrichment related to weathering profile such as Gossan genetic model of ore
441 deposits.

442

443 **5.2 Clay minerals and chlorite during hydrothermal stages**

444 The hydrothermal stages associated with saponite, chlorite/smectite and chlorite
445 precipitations surrounding albite suggest hydrothermal fluids of below 350 °C. Previous
446 studies (Lopes et al., 2014; Fontana et al., 2017) characterized the occurrence of Fe-Cu-
447 rich sulfides associated with chlorite/smectite. The association of Cu-rich sulfides with Fe
448 -Al -Mg rich chlorite/smectite suggests that the Fe²⁺, Mg and Al contents of the clay
449 minerals are related to ore deposition (Wilkinson et al., 2015; Yoneda et al., 2016).
450 Nevertheless, the origin, chemistry and temperature of the fluids are not well characterized.

451 In the Seival area, crystallography realized by X-ray diffraction and mineral
452 chemistry obtained by electron microprobe analyses identified Mg-saponite
453 (smectite), chlorite/smectite mixed-layer and corrensite infilling the porosity or
454 intergranular spaces. The occurrence of dominant clay minerals with Mg-saponite,
455 chlorite/smectite and corrensite in different lithotypes was interpreted as a lithological
456 control of alteration rather than being derived from local equilibrium (Fontana et al., 2017).
457 The comparison of the major element compositions of magmatic minerals (Part 4.2), clay
458 mineral (part 4.3) and whole-rock samples (Lopes et al., 2014) provided information to
459 support this hypothesis. The predominant Fe- Mg-rich clay minerals and their
460 Fe_{tot}/Fe_{tot}+Mg atomic ratios compared to augite (0.14 to 0.27) and biotite (0.37 to 0.44)
461 indicate that Mg-saponite (0.22 to 0.27) and chlorite/smectite mixed-layers may be derived
462 from the alteration of magmatic minerals and glass. In contrast, the high Fe-clinocllore to
463 Mg-chamosite ratio (0.47 to 0.63; Supplementary Table E) suggests the addition of Fe and

464 Mg during the chloritization process. This interpretation is confirmed by mass–balance
465 calculations performed by Fontana et al. (2019).

466 We estimated the temperatures of chlorites using chemical thermometers (Inoue et
467 al., 2009; Bourdelle and Cathelineau, 2015). These thermometers are the only available
468 temperature constraint considering the small sizes of the fluid inclusions (Fontana et al.,
469 2017). We discarded the chemical thermometer of chlorite related to metamorphic
470 conditions and selected one determined in similar lithotypes and pressure conditions. The
471 chosen chemical thermometers use a structural formula with an uncertainty dominantly
472 controlled by the Fe valency and different octahedral layers. Assuming an error of 20% in
473 the Fe distribution, the uncertainty is approximately 50 °C in both calculations. The
474 calculated temperatures of Fe–clinocllore to Mg–chamosite (Supplementary Table E) vary
475 from 129 to 275 °C in the lapilli tuff–andesite (median= 183 ± 39 °C; n= 18) and 172 to
476 312 °C in andesitic dikes (median= 251 ± 56 °C; n= 8). Moreover, the compositional
477 changes in chlorite in the Seival area compared to those in the Lavras do Sul area (Mexias
478 et al., 1990b; Bongiolo et al., 2011) or the Camaquã Mine (Renac et al., 2014) yield
479 comparable temperatures. These different uncertainties (± 50 °C) and calculated
480 temperatures reflect chloritization below of the minimum temperature of albitization. The
481 higher temperature of chlorite associated with andesitic dikes suggests a temperature
482 related to intrusive bodies or different thermal gradients between lapilli tuff and andesitic
483 dikes.

484

485 **5.3 Hydrothermal evolution and thermodynamic modeling**

486 The chlorite and chlorite/smectite mixed–layers associated with sulfides are typical
487 mineralogical assemblages of hydrothermal deposits (Barnes, 1979; Pirajno, 2009; Sillitoe,
488 2010). In the Seival area, different Fe–Cu –rich sulfide assemblages occur in the
489 intergranular space or are mixed with clay minerals (Fontana et al., 2017).

490 Hydrothermal alteration minerals are associated with a sequence of precipitations
491 rather than a paragenetic assemblage, as occurs in metamorphic reactions. Therefore,
492 thermodynamic modeling can be erroneous when based on the assumption of an output
493 represented by all of the hydrothermal phases observed. Consequently, we tested the
494 different equilibrium conditions of the hydrothermal stage (below 350 °C) with

495 thermodynamic modeling (PHREEQC interactive; Parkhurst, 1995, rather than SUPCRIT
496 or the equivalent CHONZ Rpackage). The PHREEQC thermodynamic models primarily
497 use minerals with an initial water content and propose secondary minerals as outputs with
498 the new composition of the final fluid. In our case, reactions involved observed magmatic
499 minerals observed (feldspars: andesine–labradorite and albite and Ca–pyroxene: augite) and
500 late–magmatic minerals (biotite, titanite, Fe–titanium oxides and silica–rich glass) reacting
501 with initial fluid (low–concentration water similar to rainfall, Salve et al., 2008), and
502 carbonic fluids (CO₂ as recognized in fluid inclusions (Fontana et al., 2017)). The
503 temperature domain of these calculations was bracketed by the albitization process (< 350
504 °C) and the calculations of the chlorite thermometer (120 to 320 °C). Each calculation
505 considers a molar proportion of water a hundred times higher than those of minerals and
506 fluids for varying pH values and temperatures below 350 °C. Consequently, the
507 thermodynamic changes in calculated fluids and the dissolution or precipitation of primary
508 and secondary minerals equilibrated at 350, 300, 250, 200 and 150 °C with pH and Eh were
509 adjusted for charge balance and redox conditions. These preliminary calculations indicate
510 multiple solutions with more products than were identified in the altered rocks. Therefore,
511 we forced the software to uniquely consider results with chlorite, calcite, quartz and Fe–Cu
512 rich sulfide or barite, which are here assumed to be products of hydrothermal reactions,
513 without any constraints related to their precipitation or dissolution.

514 These thermodynamic calculations were obtained for neutral (pH: 7) and acidic (pH:
515 4) initial water conditions. Regardless of the considered temperature, all final waters
516 remained neutral for an initial fluid with a neutral pH (pH: 7 to 7.3), whereas acidic waters
517 changed to a neutral pH (pH 4 to 7.3).

518 The high–temperature (350 °C) calculations indicate the coprecipitation of albite
519 and titanite with the total consumption of the Ca–rich feldspars and pyroxene, but no minor
520 chlorite or smectite. The conservation of some pyroxene and andesine–labradorite observed
521 in the petrography and the lack of clay minerals suggests unrealistic solutions for
522 hydrothermal alteration forming albite with titanite and chlorite. Consequently, we forced
523 the precipitation of chlorites in a second data set.

524 – For temperatures between 350 and 250 °C, the chlorite, quartz and calcite
525 associated with pyrite and chalcopyrite can be formed, whereas silica-rich glass, titanite
526 and Ti-magnetite are consumed anorthite, albite and pyroxene are partially dissolved.
527 – For temperatures between 250 and 150 °C, neutral or acid water dissolved
528 magmatic minerals with predicted precipitations of chlorite, barite, calcite, quartz and
529 hematite with bornite-chalcocite. Mg-saponite (smectite) was precipitated below 200 °C.
530 – Below 150 °C, the lower activities of Fe²⁺ with decreasing HS⁻/(HS⁻+SO₄) molar
531 ratios increased the saturation with respect to chalcocite and sulfate and Fe oxides, such as
532 barite and hematite. These calculations confirm that pyrite and chalcopyrite were probably
533 precipitated from a dominant magmatic fluid at temperatures between 350 to 250 °C. The
534 further mixing of the magmatic fluid with meteoric water or basinal brine at temperatures
535 varying from 250 to 150 °C promoted chalcocite-bornite deposition. Considering the
536 domain of barite saturation and the lack of supergene alteration or gossan profiles in this
537 region, the lower temperature of the Cu enrichment of sulfide ranged from 150 to 50 °C.
538

539 **6. Conclusions**

540 The Seival area contains hydrothermal Cu-sulfide deposits hosted in volcanic rocks
541 with a shoshonitic affinity. Interpretations based on mineralogical and chemical data
542 suggest that volatile-rich magma produced Mg-biotite, followed by late-magmatic
543 alteration with titanite and the albitization of andesine-labradorite (650 to 250 °C), as
544 suggested by the REE contents of albite. Thermodynamic modeling indicates that late
545 albitization (ca. < 350 °C) was associated with the formation of primary pyrite-
546 chalcopyrite, or that high-temperature chloritization (350 to 250 °C) was associated with
547 bornite-chalcocite. Moreover, the Fe_{tot}/Fe_{tot}+Mg atomic ratios of Fe-clinochlore to Mg-
548 chamosite suggest that fluids altered the pyroxene or residual material with the addition of
549 external Fe compared to smectite precipitation. The high Fe_{tot}/Fe_{tot}+Mg atomic ratios and
550 high Fe²⁺/Fe³⁺ atomic ratios of these chlorites suggest an Fe-rich and reduced fluid in
551 agreement with the migration of Fe, Cu and S species producing sulfides. However, the
552 origin of fluid, whether the mixing of magmatic and meteoric fluids or basinal brines
553 (IOCG, Chilean 'Manto type'), was not identified. Nevertheless, the chloritization process
554 occurred from 312 to 120 °C. Despite the errors related to chlorite thermometers, the ranges

555 of temperatures in the lapilli tuff (275 to 129 °C; median= 183 ± 39 °C) and andesitic dikes
556 (312 to 172 °C, median= 251 ± 56 °C) suggest the existence of changing thermal gradients
557 related to the cooling of the lava pile or intrusive dike and fluids with different
558 temperatures at different times.

559 The lowest-temperature stage was related to the widespread formation of Mg–
560 saponite and smectite–rich C/S (200 to 50 °C). The formations of chlorites,
561 chlorite/smectite mixed–layers and Mg–saponite were related to the influx of fluids
562 (meteoric water or basinal brine) at lower thermal fluxes associated with the sulfide
563 assemblage of bornite–chalcocite, chalcocite, and barite with the increasing concentration
564 of sulfate. Consequently, the petrographic association of albite and chlorite with pyrite and
565 chalcopyrite suggests that the albitization to chloritization stage and structural control may
566 be prospective guides for Cu fertile magmatic systems. The chlorite/smectite mixed–layers
567 emphasized favorable Cu–enriched zones, such as fluid circulation in faults or along dikes
568 and percolation in the porosity of the lapilli tuff.

569

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581

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848

849 **Figures and Tables Legends**

850 **Figure 1.** Geological setting of the study area. A) Schematic Sul–rio–grandense Shield map
851 (modified from Hartmann et al., 2007); B) Central–western Camaquã Basin geological map
852 with Seival Mine area and Cu occurrences localization (modified from Toniolo et al., 2007;
853 Gastal et al., 2015). Keys: 1= Seival Mine (Cu and Ag); 2= Silveira Mine (Cu); 3= Crespos
854 Mine (Cu); 4= Volta Grande and Merita Mine (Cu and Au).

855

856 **Figure 2.** Geological map of the Seival Mine. Sequence of lapilli tuff, andesitic lava flows
857 and volcanic agglomerate with andesitic and trachyandesitic dike (NE direction) from
858 Hilário Formation (HF), polymictic conglomerate and vesicular lava flows from Santa Fé
859 Formation (SFF), rhyolite dikes of Acampamento Velho Formation (AVF), faults,
860 mineralization projections, inactive open pit mines and bore–hole localization (modified
861 from Reischl, 1978; Lopes et al., 2014). Keys: 1= Barita Mine; 2= João Dahne Mine; 3=
862 Morcego Mine; 4= Meio Mine; 5= Cruzeta Mine; 6= Alcides Mine.

863

864 **Figure 3.** Mineralization in the Seival Mine area. A) Cu soil anomaly (modified from
865 Reischl, 1978) and inactive open pit mines, faults and Cu–mineralization; B) Schematic
866 section of alteration zones in Hilário Formation. Keys: Chl= Chlorite; C/S=
867 Chlorite/smectite mixed–layer; Hem= Hematite; Tit= Titanite; Cal= Calcite; Qtz= Quartz;
868 Cc= Chalcocite; Bn= Bornite; Smec= Smectite.

869

870 **Figure 4.** Macroscopic photograph (A and B), photomicrograph in natural (C and F) and
871 polarized light (D, E, G and H). A) Hand sample containing altered pyroclasts of lapilli tuff;
872 B) Hand sample with plagioclase and altered groundmass in andesite dike; C) Pervasive
873 clay minerals alteration in lapilli tuff; D) Zoned phenocryst of plagioclase with indication
874 of unaltered andesine–labradorite and altered plagioclase with albite composition in
875 andesitic dike; E) Clinopyroxene and vesicle fulfilled by smectite–rich chlorite/smectite in
876 less altered amygdaloidal andesitic lava flows; F) Amygdaloidal andesitic lava flows
877 pyroclast of lapilli tuff fulfilled by chlorite, calcite, sulfide, titanite and quartz; G) Smectite
878 in less altered andesitic lava flows; H) Vesicle fulfilled by chlorite/smectite mixed–layers in
879 andesitic lava flows. Keys: Pl= Plagioclase; Sul= Sulfide; Cal= Calcite; Tit= Titanite;

880 Hem= Hematite; Qtz= Quartz; K-fds= K-feldspar; And-lab= Andesine-labradorite; Ab=
881 Albite; C/S= Chlorite/smectite mixed-layer; Px= Pyroxene; Chl= Chlorite, Smec=
882 Smectite.

883

884 **Figure 5.** Plots of mineral compositions. A) Ca–Mg–Fe clinopyroxenes showing augite
885 compositions in andesitic lava flows and andesitic dike; B) Biotite (white point) and less
886 pristine biotite (gray point) in the Al_{tot} vs Al^{VI} diagram; C) $Al+Fe^{3+}$ vs Ti and Al vs Fe binary
887 diagram of the hydrothermal titanite compositions (diagram extracted from Cao et al.,
888 2015); D) An–Ab–Or ternary feldspar diagram showing albitization in andesine–labradorite
889 and rims with K–feldspar in some zoned plagioclase (An= Anorthite; Ab= Albite; Or=
890 Orthoclase).

891

892 **Figure 6.** Chemical compositions of a series of saponite, chlorite and chlorite/smectite
893 mixed-layers in projection M^+ , 4Si and $3R^{2+}$ (Velde, 1985; adapted from Inoue et al., 1984)
894 with detail showing the smectite, chlorite-rich chlorite–smectite mixed-layer (C/S) and
895 chlorite compositions from Seival Mine, as well as comparison between chlorites analyzes
896 from epithermal Lavras do Sul (Mexias et al., 1990b, c; Bongiolo et al., 2011) and
897 Camaquã Mine (Renac et al., 2014).

898

899 **Figure 7.** Atomic proportions of chlorites in vein and groundmass of andesitic dikes
900 presented in Wiewiora and Weiss diagram (1990), of Si and divalent transition metals
901 normalized to chlorite atom per formula (28 oxygen total anions atoms) in comparison with
902 epithermal system from Lavras do Sul (Mexias et al., 1990b; Bongiolo et al., 2011) and
903 Camaquã Mine (Renac et al., 2014) and graphical representation of low-temperature
904 chlorite geothermometry (Bourdelle and Cathelineau, 2015).

905

906 **Figure 8.** X-ray diffraction patterns of < 0.4 μ m clay-size fraction of chlorite (A), corrensite
907 (B), smectite (C) and chlorite/smectite (C/S) mixed-layer (D). Keys: Air dried (black line);
908 EG= Ethylene glycol solvation (gray line); Heating treatment (light gray line); Chl=
909 Chlorite; Smec= Smectite.

910

911 **Figure 9.** S.E.M. images obtained from drill hole fresh rock sample fragment of
912 amygdaloidal andesitic lava flows described by Fontana et al. (2017). A) Albite fragment
913 associated with corrensite clay minerals of the andesitic lava flows groundmass; B) Detail
914 of corrensite morphology of the altered andesitic lava flows groundmass. Keys: Ab= Albite;
915 Cor= Corrensite.

916

917 **Figure 10.** REE normalized patterns by the chondrite values (McDonough and Sun, 1995).
918 A) Andesine–labradorite with slight albitization of group I; B) Albite–rich plagioclases of
919 group II.

920

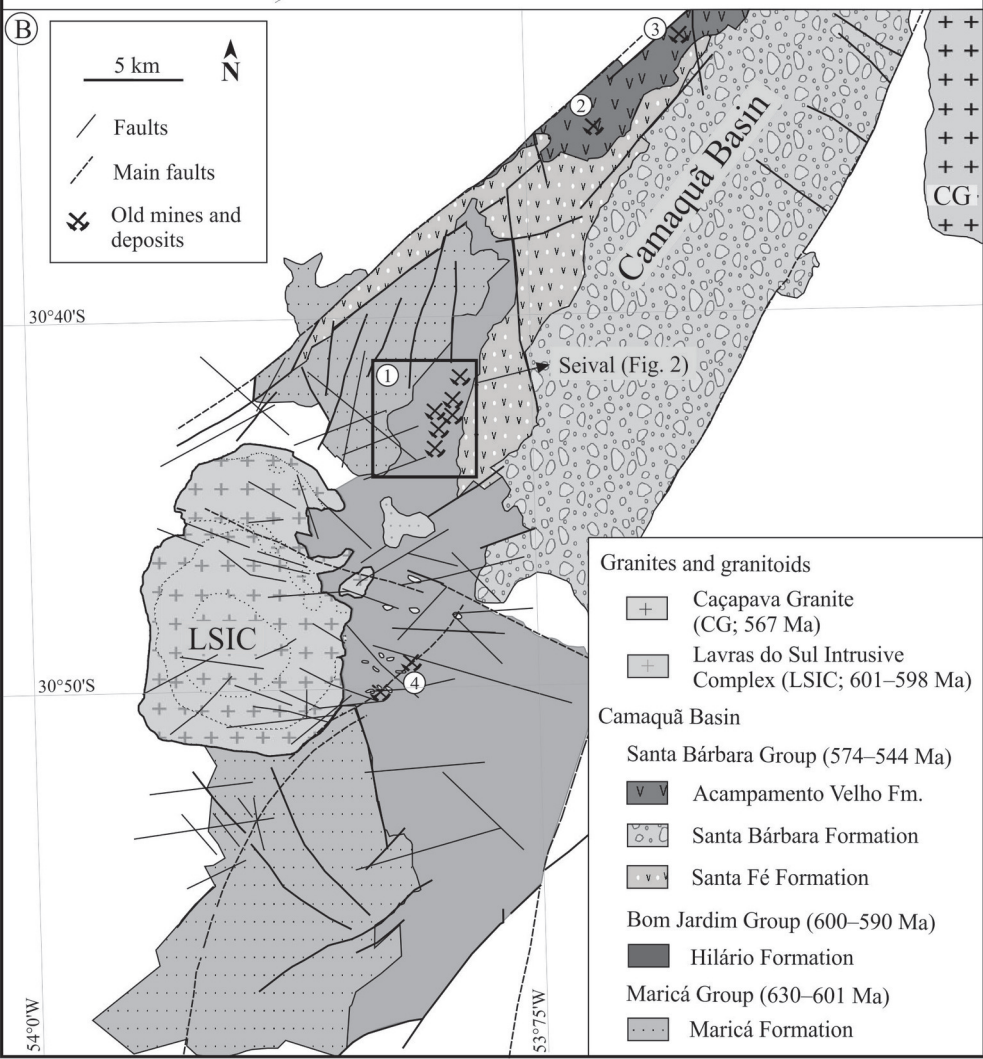
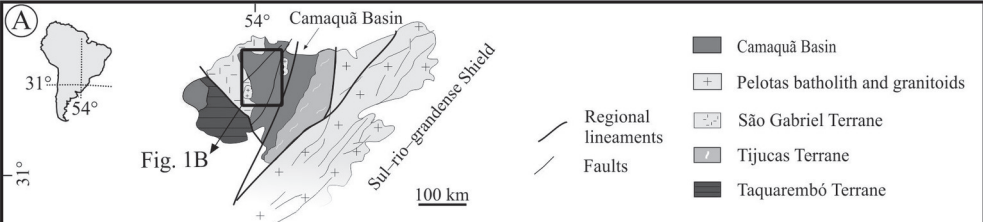
921 **Table 1.** Summary of mineral, host rock and structural formula of magmatic, magmatic to
922 late–magmatic and hydrothermal phases. Nomenclature of chlorites after Bayliss (1975).

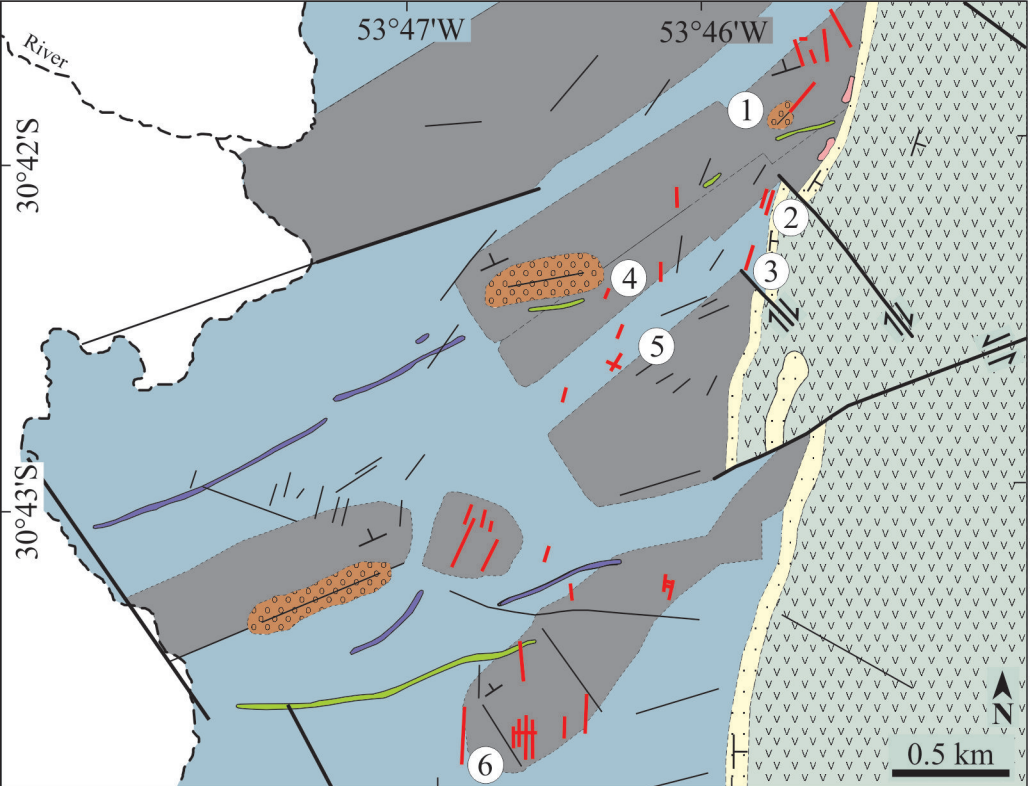
923

924 **Table 2.** Measured REE analyzes and normalized values to McDonough and Sun (1995) in
925 separate plagioclase grains of the Seival Mine, andesitic lava flows located outside of the
926 area (black line) and whole–rock analyzes (Lopes et al., 2014). Keys: And= Andesitic lava
927 flows; AD= Andesitic dike; Outs. area= Andesitic lava flows located of outside studied
928 area; Light REE= La to Sm; Heavy REE= Eu to Lu; Σ REE= Sum of total concentration of
929 not normalized rare earth elements; Eu^* = The value of the Eu anomaly calculated by the
930 formula $[Eu^*] = [(Eu_N)/(Sm_N \times Gd_N)^{1/2}]$ from Worrall and Pearson (2001) or $[Eu^*] =$
931 $[(Eu_N)/(Sm_N^2 \times Tb_N)^{1/3}]$ from Lawrence and Kamber (2006).

932

933 **Supplementary Table.** Representative compositions of pyroxene (A), biotite (B), titanite
934 (C), feldspar (D) and clay minerals as chlorite (E), smectite (F) and chlorite/smectite
935 mixed–layer (G) analyzes and structural formula. Chlorite temperatures values are showed
936 in Supplementary Table E (Inoue et al., 2009; Bourdelle and Cathelineau, 2015).





CAMAQUÃ BASIN (630–540 Ma)

Santa Bárbara Group (574–544 Ma)

Acampamento Velho Formation

Rhyolite dike

Santa Fé Formation

Polimitic conglomerate

Vesicular lava flows

Bom Jardim Group (600–590 Ma)

Hilário Formation

Trachyandesite dike

Andesite dike

Volcanic agglomerate

Andesitic lava flow

Lapilli tuff

Legend

Faults

Main faults

Mineralization

River

Dipping (S0)

Inactive mine

① Barita

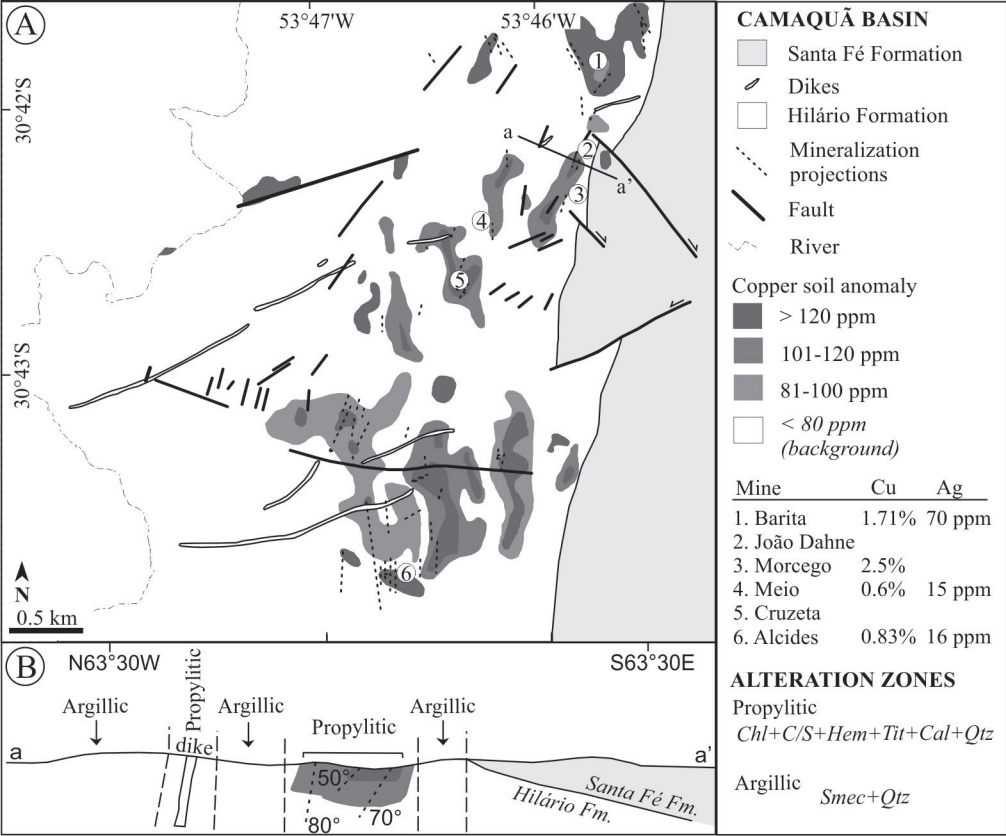
② João Dahne

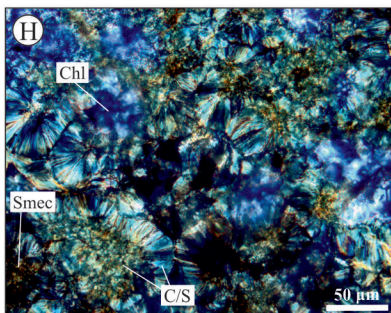
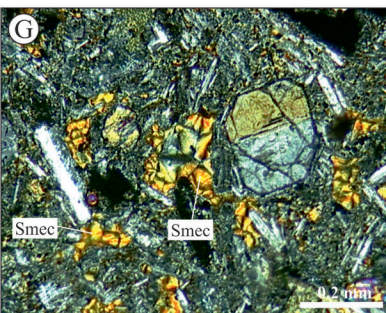
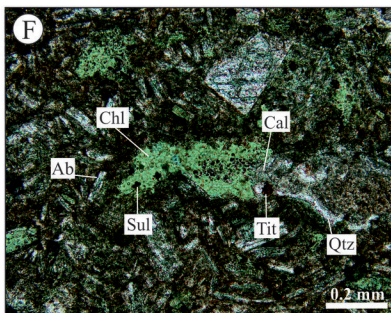
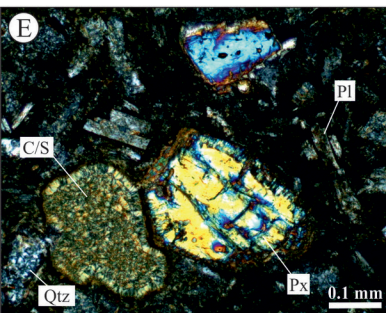
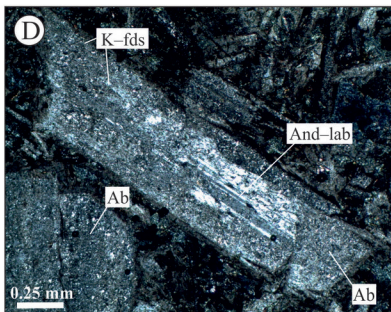
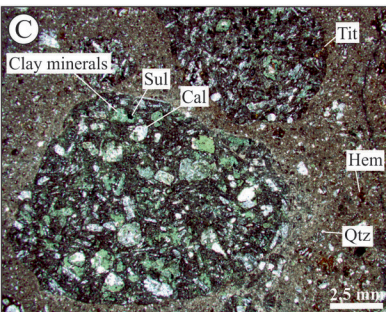
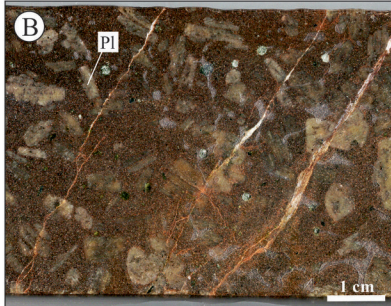
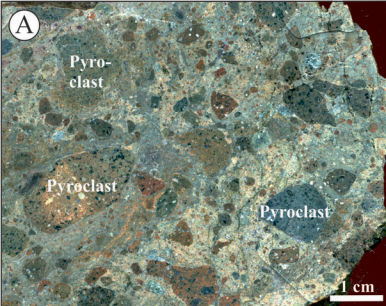
③ Morcego

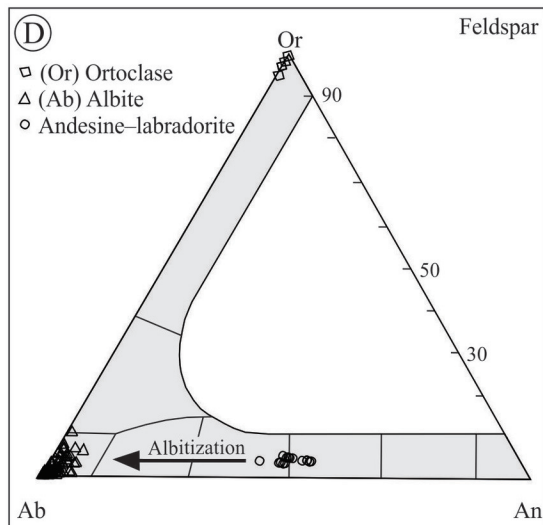
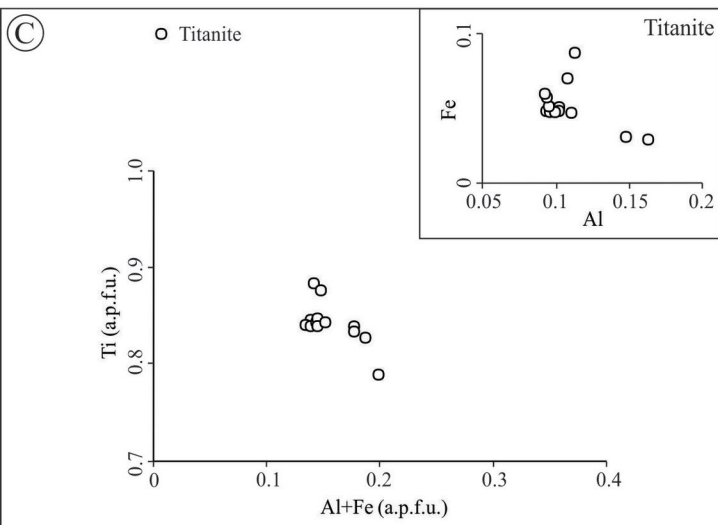
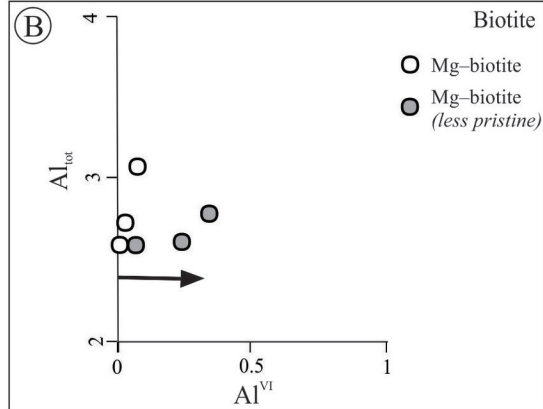
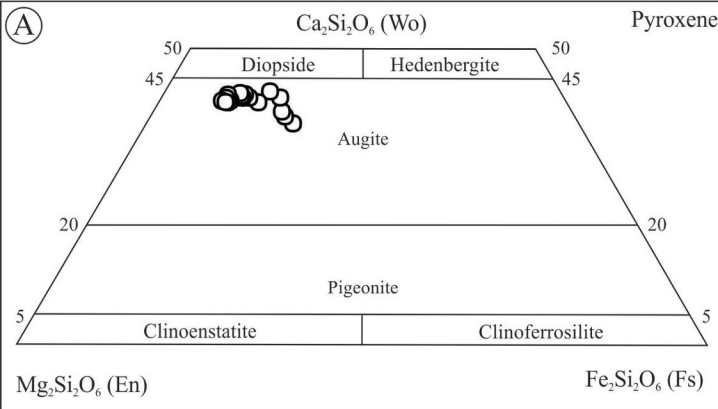
④ Meio

⑤ Cruzeta

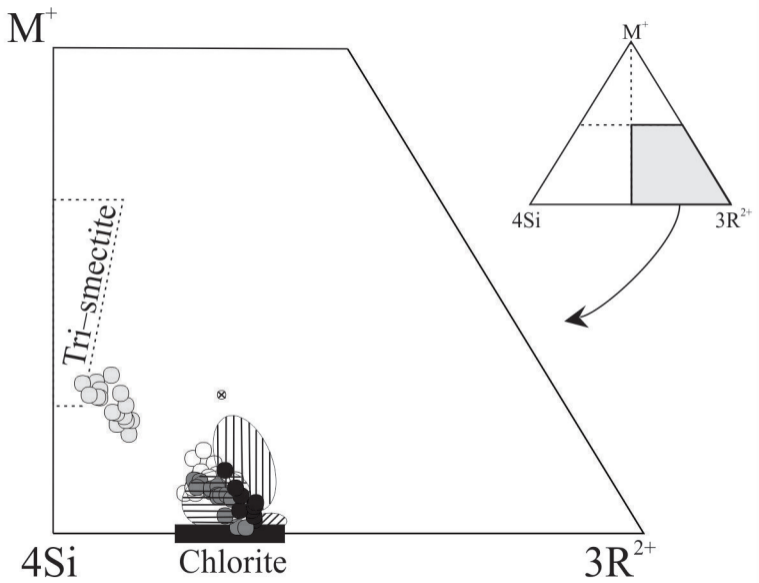
⑥ Alcides

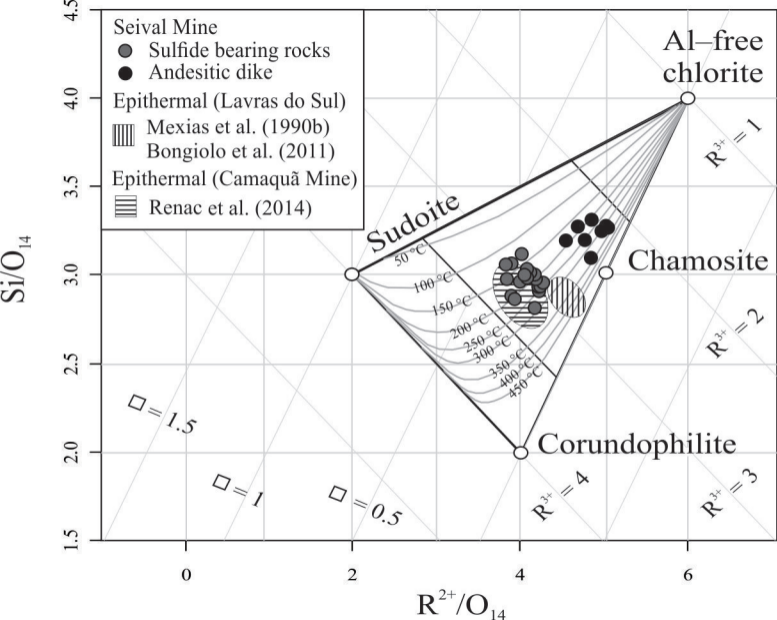


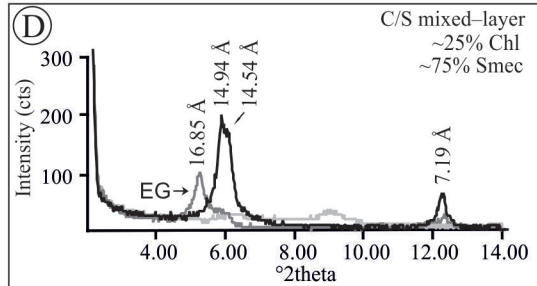
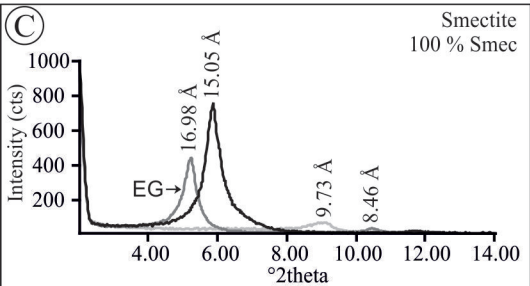
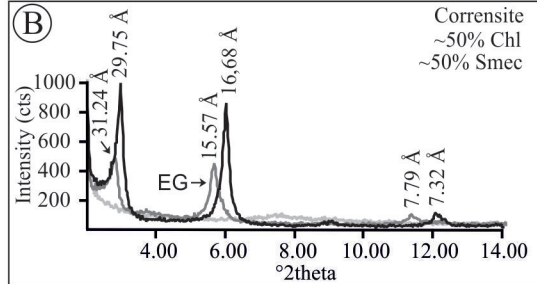
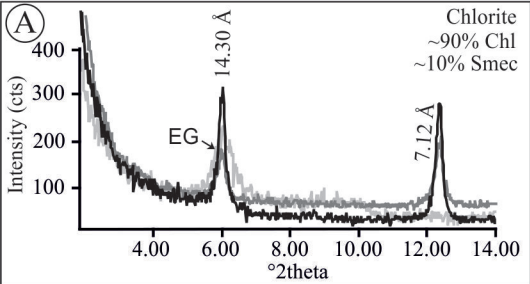


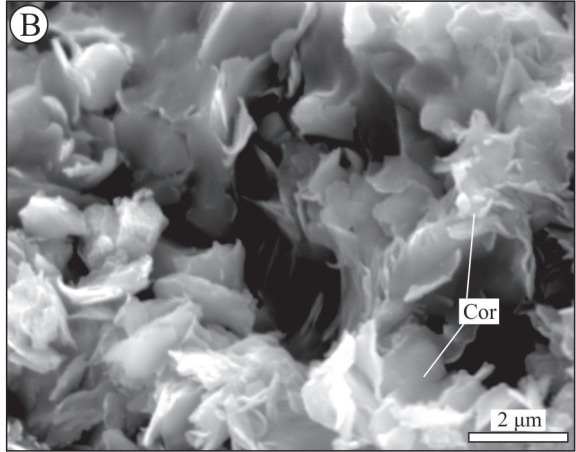
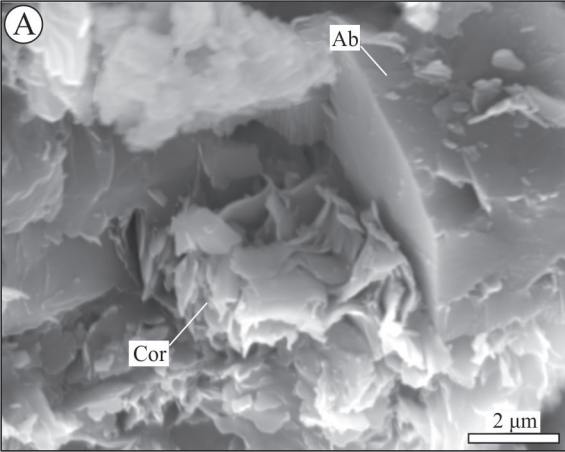


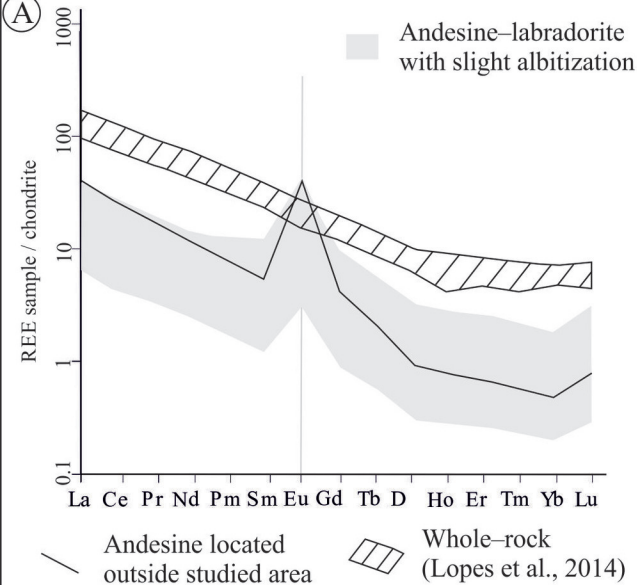
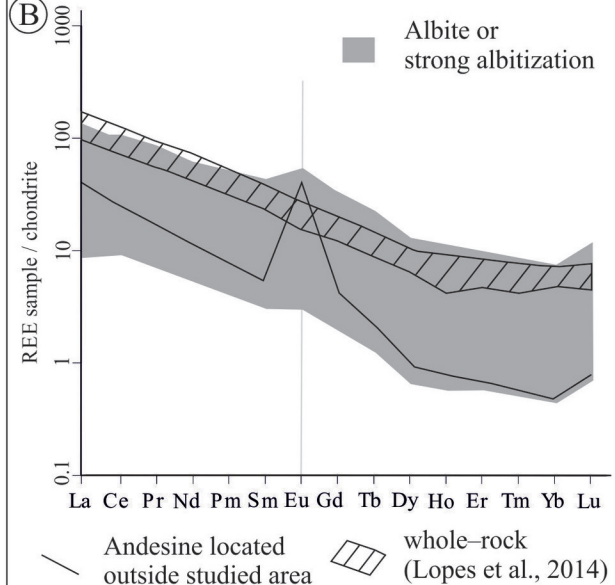
- Seival Mine
- Less altered andesitic lava flow
 - Smectite
- More altered andesitic lava flow
 - Chlorite-rich C/S
- Sulfide bearing rocks
 - Chlorite
- Andesitic dike
 - Chlorite
- Epithermal (Lavras do Sul)
 - ▨ Mexias et al. (1990b)
 - ▩ Bongioiolo et al. (2011)
- Corrensite
 - ⊗ Mexias et al. (1990c)
- Epithermal (Camaquã Mine)
 - ▨ Renac et al. (2014)









(A)**(B)**

Host rock	Mineral	Structural formula
Magmatic phase		
Andesitic lava flow and dike: less altered	Augite	$[(Ca_{0.70-0.81} Na_{0.01-0.03}) (Fe^{+2}_{0.15-0.32} Mg_{0.86-0.97}) (Al_{0.06-0.19} Ti_{0-0.02} Fe^{+3}_{0-0.06}) (Si_{1.88-1.99}) O_{12}]$
Trachyandesitic dike: magmatic or pristine	Mg-biotite	$[(Ca_{0.01-0.03} Na_{0.09-0.25} K_{0.61-1.41}) (Fe^{+2}_{1.92-2.3} Mg_{2.95-3.81} Mn_{0.01-0.03} Al_{0-0.07} Ti_{0.44-0.76})^{VI} (Al_{2.61-2.98} Si_{5.02-5.39})^{IV} O_{20} OH_{3.75-3.96} F_{0.03-0.25}]$
Trachyandesitic dike: less pristine	Mg-biotite	$[(Ca_{0.01-0.02} Na_{0.20-0.24} K_{1.22-1.49}) (Fe^{+2}_{1.79-2.02} Mg_{2.72-2.89} Mn_{0.01-0.02} Al_{0.07-0.33} Ti_{0.61-0.75})^{VI} (Al_{2.39-2.54} Si_{5.46-5.61})^{IV} O_{20} OH_{3.73-3.85} F_{0.15-0.27}]$
Magmatic to late-magmatic phase		
Andesitic dike: euhedral to agglomerates	Titanite	$[(Ca_{0.94-1.02} \square_{0.02} Ti_{0.79-0.88} Al_{0.09-0.16} Fe^{3+}_{0.03-0.09} Mg_{0-0.03} Si_{1.01-1.07} O_5)]$
Andesitic dike and lava flow: less altered	Andesine-labradorite	$[(An_{0.43-0.53} Ab_{0.43-0.54} Or_{0.03-0.05})]$
Andesitic dike	K-feldspar	$[(An_{0.43-0.53} Ab_{0.43-0.54} Or_{0.03-0.05})]$
Andesitic lava flow and dike: more altered	Albite	$[(An_{0-0.06} Ab_{0.89-1} Or_{0-0.10})]$
Hydrothermal phase		
* Porphyritic andesitic lava flow: augite replacement	Mg-chamosite to Fe-clinocllore	$[(Ca_{0.04-0.06} Na_{0-0.04} K_{0.02-0.11}) (Al_{2.70-3.03} Ti_{0-0.01} Fe^{3+}_{0.37-0.57} Fe^{2+}_{3.61-5.23} Mg_{3.07-3.97} Mn_{0.01-0.04})^{VI} (Si_{5.94-6.24} Al_{1.76-2.06})^{IV} O_{20} (OH)_8]$
** Lapilli tuff: filled vugs	Mg-chamosite	$[(Ca_{0-0.03} Na_{0-0.04} K_{0-0.11}) (Al_{2.70-3.11} Ti_{0-0.01} Fe^{3+}_{0.31-0.46} Fe^{2+}_{4.38-5.14} Mg_{3.17-3.50} Mn_{0.01-0.02})^{VI} (Si_{5.62-6.01} Al_{1.99-2.38})^{IV} O_{20} (OH)_8]$
*** Andesitic dikes: matrix	Mg-chamosite to Fe-clinocllore	$[(Ca_{0.03-0.12} Na_{0-0.03} K_{0.01-0.05}) (Al_{1.63-2.13} Ti_{0-0.01} Fe^{3+}_{0.1-0.28} Fe^{2+}_{4.26-5.49} Mg_{4.4-4.74} Mn_{0.03-0.08})^{VI} (OH)_4 (Si_{6.20-6.62} Al_{1.38-1.80})^{IV} O_{20} (OH)_8]$
Amygdaloidal andesitic lava flow: less altered rocks	Mg-saponite	$[(Ca_{0.18-0.31} Na_{0.02-0.05} K_{0.02-0.07}) (Al_{0.06-0.23} Ti_{0-0.01} Fe^{+2}_{1.35-1.78} Mg_{4.50-4.81} Mn_{0.02-0.03})^{VI} (Si_{6.41-6.98} Al_{1.02-1.59})^{IV} O_{20} (OH)_2]$
Amygdaloidal andesitic lava flow: less altered rocks	Chl-rich C/S mixed-layer	$[(Ca_{0.05-0.16} Na_{0-0.09} K_{0.01-0.05}) (Al_{1.59-1.93} Ti_{0-0.01} Fe^{3+}_{0.18-0.49} Fe^{2+}_{2.68-3.77} Mg_{5.70-6.37} Mn_{0.03-0.05})^{VI} (Si_{6.53-7.23} Al_{0.77-1.47})^{IV} O_{20} (OH)_8]$
Porphyritic andesitic lava flow: more altered rock	Chl-rich C/S mixed-layer	$[(Ca_{0.05-0.07} Na_{0.02-0.06} K_{0.01-0.03}) (Al_{2.21-2.46} Ti_{0-0.02} Fe^{3+}_{0.46-0.61} Fe^{2+}_{2.19-2.74} Mg_{5.59-5.86} Mn_{0.06-0.07})^{VI} (Si_{6.52-6.75} Al_{1.25-1.48})^{IV} O_{20} (OH)_8]$

Sample (mg/kg)	FM 13	DD 62	RFM 66	RFM 55	WL 13	WL 19	FM 33	FM 32	DD 76	RFM 68
	Group I			Group II						Outs. Area
	AD	AD	AD	And	And	And	AD	AD	AD	And
Y	0.23	1.61	1.63	9.18	3.15	2.26	0.41	1.36	11.03	1.24
La	1.66	6.18	9.95	28.39	7.70	5.47	2.16	4.30	31.66	8.09
Ce	2.68	15.41	17.54	55.31	19.25	16.08	5.57	14.19	60.18	13.61
Pr	0.32	1.65	1.90	8.15	2.29	2.06	0.66	1.89	7.79	1.46
Nd	1.14	6.10	6.54	32.46	8.71	8.07	2.51	7.31	29.43	5.04
Sm	0.18	1.77	1.51	5.85	2.12	2.25	0.45	2.61	6.47	0.82
Eu	0.17	1.79	2.36	2.28	0.60	0.67	0.17	0.82	3.05	1.75
Gd	0.18	1.89	1.62	5.55	1.94	1.92	0.37	2.19	6.50	0.90
Tb	0.02	0.21	0.17	0.74	0.26	0.26	0.04	0.28	0.81	0.09
Dy	0.07	0.78	0.58	3.07	1.11	1.05	0.16	1.04	3.20	0.30
Ho	0.02	0.15	0.11	0.65	0.23	0.21	0.03	0.20	0.63	0.06
Er	0.04	0.40	0.31	1.80	0.59	0.56	0.09	0.54	1.64	0.16
Tm	0.01	0.06	0.04	0.28	0.09	0.09	0.01	0.08	0.24	0.02
Yb	0.03	0.29	0.22	1.46	0.45	0.45	0.07	0.39	1.21	0.11
Lu	0.01	0.08	0.05	0.35	0.12	0.12	0.02	0.12	0.29	0.03
Normalization standard CI chondrite (McDonough and Sun, 1995)										
La _N	6.99	26.07	42.00	119.77	32.50	23.10	9.10	18.16	133.6	34.12
Ce _N	4.37	25.14	28.62	90.23	31.40	26.23	9.09	23.15	98.18	22.21
Pr _N	3.41	17.80	20.48	87.81	24.68	22.19	7.12	20.42	83.91	15.70
Nd _N	2.49	13.35	14.32	71.04	19.06	17.66	5.48	16.01	64.4	11.03
Sm _N	1.20	11.98	10.22	39.55	14.32	15.20	3.02	17.66	43.73	5.53
Eu _N	3.04	31.85	41.95	40.48	10.57	11.87	2.98	14.55	54.09	31.07
Gd _N	0.89	9.50	8.14	27.90	9.74	9.66	1.88	10.98	32.67	4.50
Tb _N	0.55	5.73	4.63	20.44	7.31	7.15	1.22	7.73	22.49	2.41
Dy _N	0.30	3.16	2.35	12.48	4.50	4.26	0.64	4.24	13.02	1.21
Ho _N	0.27	2.71	2.03	11.83	4.12	3.83	0.57	3.70	11.45	1.04
Er _N	0.26	2.51	1.94	11.27	3.69	3.53	0.56	3.39	10.26	0.98
Yb _N	0.20	1.82	1.37	9.09	2.80	2.81	0.43	2.43	7.54	0.69
Lu _N	0.28	3.09	2.15	14.27	4.72	4.88	0.69	4.67	11.91	1.06
ΣLREE	5.98	31.11	37.44	130.16	40.07	33.93	11.35	30.30	135.53	29.02
ΣHREE	0.55	5.65	5.46	16.18	5.39	5.33	0.96	5.66	17.57	3.42
ΣREE	6.76	38.37	44.53	155.52	48.61	41.52	12.72	37.32	164.13	33.68
[Eu*]= [(Eu _N)/(Sm _N x Gd _N ^{1/2})]	1.22	0.90	0.98	2.94	2.99	4.60	1.25	1.04	1.43	6.23
[Eu*]= [(Eu _N)/(Sm _N ² x Tb _N ^{1/3})]	1.28	0.92	1	3.29	3.40	5.35	1.33	1.09	1.54	7.41