



HAL
open science

On Post-Minkowskian Hamiltonians in General Relativity

Andrea Cristofoli, N. E. J. Bjerrum-Bohr, Poul H. Damgaard, P. Vanhove

► **To cite this version:**

Andrea Cristofoli, N. E. J. Bjerrum-Bohr, Poul H. Damgaard, P. Vanhove. On Post-Minkowskian Hamiltonians in General Relativity. *Physical Review D*, 2019, 100, pp.084040. 10.1103/PhysRevD.100.084040 . hal-02165253

HAL Id: hal-02165253

<https://hal.science/hal-02165253>

Submitted on 23 Feb 2021

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



Distributed under a Creative Commons Attribution 4.0 International License

On Post-Minkowskian Hamiltonians in General Relativity

Andrea Cristofoli,¹ N.E.J. Bjerrum-Bohr,¹ Poul H. Damgaard,¹ and Pierre Vanhove^{2,3}

¹*Niels Bohr International Academy and Discovery Center*

Niels Bohr Institute, University of Copenhagen

²*Institut de Physique Théorique, Université Paris-Saclay*

CEA, CNRS, F-91191 Gif-sur-Yvette Cedex, France

³*National Research University Higher School of Economics, Russian Federation*

(Dated: October 3, 2019)

Abstract

We describe the computation of post-Minkowskian Hamiltonians in General Relativity from scattering amplitudes. Using a relativistic Lippmann-Schwinger equation, we relate perturbative amplitudes of massive scalars coupled to gravity to the post-Minkowskian Hamiltonians of classical General Relativity to any order in Newton's constant. We illustrate this by deriving a Hamiltonian for binary black holes without spin up to 2nd order in the post-Minkowskian expansion and demonstrate explicitly the equivalence with the recently proposed method based on an effective field theory matching.

I. INTRODUCTION

The detection of gravitational waves by the LIGO/Virgo collaboration has opened up the exciting possibility of testing Einstein’s theory of general relativity at a new and unprecedented level, including the regime of strong gravity as probed by black holes just prior to merging. A combination of Numerical Relativity and analytical methods is needed in order to push theory to the level where it can provide best-fit templates from which physical parameters can be extracted. This has spurred interest in new and innovative ideas that can facilitate computations of the required two-body interaction Hamiltonians to high accuracy.

Conventionally, the calculations of effective interaction Hamiltonians have been carried out in the systematic post-Newtonian expansion of General Relativity. The problem can, however, be attacked from an entirely different angle, that of relativistic scattering amplitudes as computed by standard quantum field theory methods in a quantum field theory of gravity coupled to matter [1]. Modern methods of amplitude computations greatly facilitate this program [2–9]. Incoming and outgoing particles in the scattering process are taken to past and future infinity where the metric by definition is flat Minkowskian, and the full metric is treated perturbatively around that Minkowskian background. The classical piece of the scattering amplitude solves the scattering problem of two black holes to the given order in Newton’s constant G_N . When expanding to the appropriate post-Newtonian order and defining the interaction potential with the inclusion of the required lower-order Born subtractions as explained in detail in the next section, the amplitude also contains all the information of the bound state problem of two massive objects to the given order in the expansion in Newton’s constant. For the bound-state regime one has, on account of the virial theorem, a double expansion in both Newton’s constant and momentum. However, a more daring angle of attack is to treat the bound state problem as not expanded in momentum while still expanding to fixed order in Newton’s constant. Such an approach has recently been proposed by Cheung, Rothstein and Solon [8], and it has already been pushed one order higher in the expansion in Newton’s constant [9] (and compared to the post-Newtonian expansions in [10]). Here the method of effective field theory is used to extract the interaction Hamiltonian: the underlying Einstein-Hilbert action coupled to matter produces the classical part of the scattering amplitude while an effective theory of two massive objects define the interaction Hamiltonian. The correct matching between the two theories is performed

by insisting that the two theories produce the same scattering amplitude to the given order in Newton's constant.

The post-Newtonian expansion (see, *e.g.*, refs. [11–13] for recent comprehensive reviews) of General Relativity dates back to the founding days of the theory. Its perturbation theory is ideal for the low-velocity situations of planetary orbits, satellites, and large-distance effects of General Relativity that occur at velocities far below the speed of light. In contrast to this, the computation of observables in General Relativity based on relativistic scattering amplitudes is valid for all velocities and in particular this is the proper framework for high energy scattering where kinetic energies can exceed potential energies by arbitrarily large amounts. This leads naturally to what has become known in the theory of General Relativity as the post-Minkowskian expansion [14–20].

Extracting the interaction energy from the relativistic scattering amplitude, for consistency with the virial theorem in the bound-state problem one would perform a double expansion where velocity v and G_N are both kept to the appropriate order. To any given order in G_N this would imply a truncation of a Taylor-expanded amplitude in powers of momenta. There is no general argument for whether keeping higher powers of only one of the expansion parameters in the regime where they are of comparable magnitude will increase the accuracy. Considering its potential impact, it is nevertheless of much interest to explore the consequences of keeping higher-order terms of momenta even in the bound state regime where they would not ordinarily have been included [8–10]. We will here show how that post-Minkowskian Hamiltonian also follows directly from the full relativistic amplitude without having to perform the amplitude matching to the effective field theory, thereby explicitly showing equivalence between the two approaches [7, 8].

II. PERTURBATIVE GRAVITY AS A FIELD THEORY

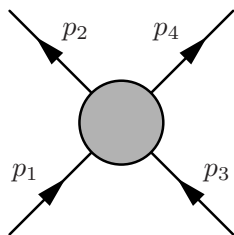
We start by introducing the Einstein-Hilbert action minimally coupled to massive scalar fields ϕ_a

$$\mathcal{S} = \int d^4x \sqrt{-g} \left[\frac{R}{16\pi G_N} + \frac{1}{2} \sum_a \left(g^{\mu\nu} \partial_\mu \phi_a \partial_\nu \phi_a - m_a^2 \phi_a^2 \right) \right], \quad (1)$$

where R defines the Ricci scalar and $g \equiv \det(g_{\mu\nu})$. Perturbatively, we expand the metric around a Minkowski background: $g_{\mu\nu}(x) = \eta_{\mu\nu} + \sqrt{32\pi G_N} h_{\mu\nu}(x)$. At large distances we can

treat the scattering of two massive objects m_a and m_b as that of two point-like particles with the same masses. This has all been well elucidated in the literature (see, *e.g.*, refs. [21, 22]), although most focus until now seems to have been on considering the quantum mechanical effects. The way classical terms appear from the quantum mechanical loop expansion is subtle [1, 23]; see ref. [24] for a very nice and clear discussion of this issue. Instead of expanding the action (1) in terms of ordinary Feynman rules, it pays to use modern amplitude methods to extract the needed non-analytic pieces in momentum transfer \vec{q} through the appropriate cuts at loop level [2–4].

The scattering $m_a + m_b \rightarrow m_a + m_b$ mediated by gravitons at an arbitrary loop order is described by



$$\mathcal{M} = \text{diagram} = \sum_{L=0}^{+\infty} \mathcal{M}^{L\text{-loop}} \quad , \quad \mathcal{M}^{L\text{-loop}} \sim O(G_N^{L+1}), \quad (2)$$

We choose the center-of-mass frame and parametrize the momenta as follows:

$$\begin{aligned} p_1^\mu &= (E_a, \vec{p}), & p_2^\mu &= (E_a, \vec{p}'), \\ p_3^\mu &= (E_b, -\vec{p}), & p_4^\mu &= (E_b, -\vec{p}'), \end{aligned} \quad (3)$$

and $|\vec{p}| = |\vec{p}'|$. We also define

$$q^\mu = p_1^\mu - p_2^\mu = p_4^\mu - p_3^\mu \equiv (0, -\vec{q}), \quad \vec{q} \equiv \vec{p}' - \vec{p}, \quad (4)$$

and the total energy $E_p = E_a + E_b$.

III. THE LIPPMANN-SCHWINGER EQUATION

It is a classical problem in perturbative scattering theory to relate the scattering amplitude \mathcal{M} to an interaction potential V . This is typically phrased in terms of non-relativistic quantum mechanics, but it is readily generalized to the relativistic case. Crucial in this respect is the fact that we shall consider particle solutions to the relativistic equations only.

There will thus be, in the language of old-fashioned (time-ordered) perturbation theory, no back-tracking diagrams corresponding to multiparticle intermediate states. This is trivially so since we neither wish to treat the macroscopic classical objects such as heavy neutron stars as indistinguishable particles with their corresponding antiparticles, nor do we wish to probe the scattering process in any potential annihilation channel. The classical objects that scatter will always be restricted to classical distance scales.

We now briefly outline a systematic procedure for connecting the scattering amplitude in perturbative gravity with post-Minkowskian potentials in classical General Relativity. We start by introducing a bit of notation. First, we assume the existence of a relativistic one-particle Hamiltonian of only particle states describing what in bound-state problems is known as the Salpeter equation,

$$\hat{\mathcal{H}} = \hat{\mathcal{H}}_0 + \hat{V}, \quad \hat{\mathcal{H}}_0 = \sqrt{\hat{k}^2 + m_a^2} + \sqrt{\hat{k}^2 + m_b^2} \quad (5)$$

where \hat{V} is a so far unspecified potential describing our post-Minkowskian system. We also define, on a proper subset of the complex plane, the following \mathbb{C} -valued operators

$$\hat{G}_0(z) \equiv (z - \hat{\mathcal{H}}_0)^{-1}, \quad \hat{G}(z) \equiv (z - \hat{\mathcal{H}})^{-1}, \quad (6)$$

$$\hat{T}(z) \equiv \hat{V} + \hat{V}\hat{G}(z)\hat{V} \quad (7)$$

Here \hat{G}_0 and \hat{G} are the Green's operator for the free and interacting case, while \hat{T} is the off-shell scattering matrix whose on-shell matrix elements provide the non-trivial S -matrix elements. We can relate the two Green's operator by means of the following operator identity

$$A^{-1} = B^{-1} + B^{-1}(B - A)A^{-1} \quad \Rightarrow \quad \hat{G} = \hat{G}_0 + \hat{G}_0\hat{V}\hat{G} \quad (8)$$

Multiplying both sides of (7) by \hat{G}_0 , combined with (8), one has

$$\hat{G}_0\hat{T} = \hat{G}_0\hat{V} + \hat{G}_0\hat{V}\hat{G}\hat{V} = \hat{G}\hat{V}, \quad (9)$$

$$\hat{T}(z) = \hat{V} + \hat{V}\hat{G}_0(z)\hat{T}(z) \quad (10)$$

which is the basis for a perturbative knowledge of \hat{T} and it usually known as *Lippmann-*

Schwinger equation.

We now take the inner product on scattering states $|p\rangle, |p'\rangle$

$$\langle p|\hat{T}(z)|p'\rangle = \langle p|\hat{V}|p'\rangle + \int \frac{d^3k}{(2\pi)^3} \frac{\langle p|\hat{V}|k\rangle \langle k|\hat{T}(z)|p'\rangle}{z - E_k} \quad (11)$$

and use the crucial relation

$$\lim_{\epsilon \rightarrow 0} \langle p|\hat{T}(E_p + i\epsilon)|p'\rangle = \mathcal{M}(p, p') \quad (12)$$

which provides the link to the conventionally defined scattering amplitude \mathcal{M} in quantum field theory restricted to the particle sector. Substituting (12) into (11) we have a recursive relation between the amplitude and the post-Minkowskian potential

$$\mathcal{M}(p, p') = \langle p|V|p'\rangle + \int \frac{d^3k}{(2\pi)^3} \frac{\langle p|V|k\rangle \mathcal{M}(k, p')}{E_p - E_k + i\epsilon} \quad (13)$$

Solving this equation iteratively, we can invert it in order to arrive at a relativistic equation for the potential V

$$\langle p|V|p'\rangle = \mathcal{M}(p, p') - \int \frac{d^3k}{(2\pi)^3} \frac{\mathcal{M}(p, k) \mathcal{M}(k, p')}{E_p - E_k + i\epsilon} + \dots \quad (14)$$

or, in position space,

$$V(p, r) = \int \frac{d^3q}{(2\pi)^3} e^{iq \cdot r} V(p, q), \quad (15)$$

with

$$V(p, q) \equiv \langle p|V|p'\rangle \quad (16)$$

At this stage there has not been any restriction to a non-relativistic limit. The anti-particle sector has been eliminated by hand, as dictated by the physical scattering process. We can thus regard (15) as defining a post-Minkowskian potential.

B. The post-Minkowskian potential to second order

In order to consider a post-Minkowskian potential at second order in G_N^2 , we will need to consider a contribution coming from the iterated tree-level amplitude, as dictated by (14)

$$V_{2\text{PM}}(p, q) = \mathcal{M}^{1\text{-loop}}(p, p') + \mathcal{M}^{\text{Iterated}}(p, p'), \quad (22)$$

$$\mathcal{M}^{\text{Iterated}}(p, p') \equiv - \int \frac{d^d k}{(2\pi)^d} \frac{\mathcal{M}^{\text{tree}}(p, k) \mathcal{M}^{\text{tree}}(k, p')}{E_p - E_k + i\epsilon}. \quad (23)$$

Infrared divergences are regularized by temporarily switching to $d+1$ space-time dimensions. The classical terms of the one-loop amplitude have been given elsewhere [6–8, 25–28]. They can be decomposed in terms of scalar integrals with coefficients that are independent of the exchanged three-momentum \vec{q} ,

$$\mathcal{M}^{1\text{-loop}} = \frac{i16\pi^2 G_N^2}{E_a E_b} \left(c_{\square} \mathcal{I}_{\square} + c_{\boxtimes} \mathcal{I}_{\boxtimes} + c_{\triangleright} \mathcal{I}_{\triangleright} + c_{\triangleleft} \mathcal{I}_{\triangleleft} + \dots \right) \quad (24)$$

where the symbol of each coefficient refers to the topology of the contributions involved while the ellipses denote quantum mechanical contributions that we neglect.

In detail, the scalar box and crossed-box integrals are given by

$$\mathcal{I}_{\square} = \int \frac{d^{d+1} \ell}{(2\pi)^{d+1}} \frac{1}{((\ell + p_1)^2 - m_a^2 + i\epsilon)((\ell - p_3)^2 - m_b^2 + i\epsilon)(\ell^2 + i\epsilon)((\ell + q)^2 + i\epsilon)}, \quad (25)$$

$$\mathcal{I}_{\boxtimes} = \int \frac{d^{d+1} \ell}{(2\pi)^{d+1}} \frac{1}{((\ell + p_1)^2 - m_a^2 + i\epsilon)((\ell + p_4)^2 - m_b^2 + i\epsilon)(\ell^2 + i\epsilon)((\ell + q)^2 + i\epsilon)}, \quad (26)$$

At leading order in the momentum transfer \vec{q} the coefficients of these integrals are finite at $d = 3$ and given by [3, 22]

$$c_{\square} = c_{\boxtimes} = 4(m_a^2 m_b^2 - 2(p_1 \cdot p_3)^2)^2. \quad (27)$$

The scalar triangle integrals are given by

$$\mathcal{I}_{\triangleright} = \int \frac{d^{d+1}\ell}{(2\pi)^{d+1}} \frac{1}{((\ell + q)^2 + i\varepsilon)(\ell^2 + i\varepsilon)((\ell + p_1)^2 - m_a^2 + i\varepsilon)} \quad (28)$$

$$\mathcal{I}_{\triangleleft} = \int \frac{d^{d+1}\ell}{(2\pi)^{d+1}} \frac{1}{((\ell - q)^2 + i\varepsilon)(\ell^2 + i\varepsilon)((\ell - p_3)^2 - m_b^2 + i\varepsilon)} \quad (29)$$

with coefficients, at the leading order in $|\vec{q}|$ and around $d = 3$, given by

$$c_{\triangleright} = 3m_a^2(m_a^2m_b^2 - 5(p_1 \cdot p_3)^2) \quad , \quad c_{\triangleleft} = 3m_b^2(m_a^2m_b^2 - 5(p_1 \cdot p_3)^2) \quad (30)$$

These scalar integrals are conveniently evaluated by performing proper contour integrals in ℓ^0 as explained in [7]. Doing so, we see that the box, crossed-box, and triangle contributions are given by [22, 29]

$$\mathcal{I}_{\square} = -\frac{i}{16\pi^2|\vec{q}|^2} \left(-\frac{1}{m_a m_b} + \frac{m_a(m_a - m_b)}{3m_a^2 m_b^2} + \frac{i\pi}{|p|E_p} \right) \left(\frac{2}{3-d} - \log|\vec{q}|^2 \right) + \dots \quad , \quad (31)$$

$$\mathcal{I}_{\boxtimes} = -\frac{i}{16\pi^2|\vec{q}|^2} \left(\frac{1}{m_a m_b} - \frac{m_a(m_a - m_b)}{3m_a^2 m_b^2} \right) \left(\frac{2}{3-d} - \log|\vec{q}|^2 \right) + \dots \quad , \quad (32)$$

$$\mathcal{I}_{\triangleright} = -\frac{i}{32m_a} \frac{1}{|\vec{q}|} + \dots \quad , \quad (33)$$

$$\mathcal{I}_{\triangleleft} = -\frac{i}{32m_b} \frac{1}{|\vec{q}|} + \dots \quad , \quad (34)$$

at leading order in the $|\vec{q}|^2$ expansion and around $d = 3$. We thus arrive at the one-loop amplitude to leading order in $|\vec{q}|^2$,

$$\mathcal{M}^{1\text{-loop}} = \frac{\pi^2 G_N^2}{E_p^2 \xi} \left[\frac{1}{2|\vec{q}|} \left(\frac{c_{\triangleright}}{m_a} + \frac{c_{\triangleleft}}{m_b} \right) + \frac{i}{E_p} \frac{c_{\square}}{|\vec{p}|} \frac{(\frac{2}{3-d} - \log|\vec{q}|^2)}{\pi|\vec{q}|^2} \right] \quad (35)$$

The imaginary part of this which arises from the box and crossed-box integrals is the infrared divergent Weinberg phase [30]. By restoring the \hbar -counting, one sees that it scales as \hbar^{-1} , a behavior dubbed super-classical in [24]. We will show below that it cancels in the properly defined potential, a fact already noted in the post-Newtonian expansion [25].

We next evaluate the iterated tree-level contribution given by

$$\mathcal{M}^{\text{Iterated}} = -\frac{16\pi^2 G_N^2}{E_a(p^2)E_b(p^2)} \int \frac{d^d k}{(2\pi)^d} \frac{A(\vec{p}, \vec{k}) A(\vec{k}, \vec{p}')}{|\vec{p} - \vec{k}|^2 |\vec{p}' - \vec{k}|^2} \frac{\mathcal{G}(p^2, k^2)}{E_a(k^2)E_b(k^2)} \quad (36)$$

where we have introduced the Green function

$$\mathcal{G}(p^2, k^2) = \frac{1}{E_p - E_k + i\epsilon} \quad (37)$$

The function A is the numerator of the tree-level amplitude (18) with the k -legs satisfying 3-momentum (but not energy) conservation. We notice that $A(\vec{p}, \vec{k})$ and $A(\vec{k}, \vec{p}')$ can be written as

$$A(\vec{p}, \vec{k}) = \tilde{A}(p^2, k^2) + B(\vec{p}, \vec{k}), \quad (38)$$

$$A(\vec{k}, \vec{p}') = \tilde{A}(p^2, k^2) + B(\vec{p}', \vec{k}) \quad (39)$$

where \tilde{A} is \vec{q} -independent and function of $|\vec{p}| = p$ and $|\vec{k}| = k$. The classical contribution from the iterated Born amplitude is hence

$$\mathcal{M}^{\text{Iterated}} = -\frac{16\pi^2 G_N^2}{E_a(p^2)E_b(p^2)} \int \frac{d^d k}{(2\pi)^d} \frac{\mathcal{G}(p^2, k^2)Q(p^2, k^2)}{|\vec{p} - \vec{k}|^2 |\vec{p}' - \vec{k}|^2} \quad (40)$$

where

$$Q(p^2, k^2) = \frac{\tilde{A}^2(p^2, k^2)}{E_a(k^2)E_b(k^2)} \quad (41)$$

We now expand Q around p^2 ,

$$Q(p^2, k^2) = Q_{k=p} + (k^2 - p^2)\partial_{k^2}Q_{k^2=p^2} + \dots, \quad (42)$$

$$Q_{k^2=p^2} = \frac{\tilde{A}_{k^2=p^2}^2}{E_a(p^2)E_b(p^2)} = \frac{c_1^2}{E_p^2 \xi}, \quad (43)$$

$$\partial_{k^2}Q_{k^2=p^2} = -\frac{1}{E_p^2 \xi^2} \left(2c_1 p_1 \cdot p_3 + \frac{c_1^2}{2E_p^2 \xi} (1 - 2\xi) \right) \quad (44)$$

The Green function \mathcal{G} likewise admits a Laurent expansion in k^2

$$\mathcal{G}(p^2, k^2) = \frac{2E_p \xi}{p^2 - k^2} + \frac{3\xi - 1}{2E_p \xi} + \dots \quad (45)$$

Combining terms, the Born subtraction can hence be expressed as

$$\begin{aligned} \mathcal{M}^{\text{Iterated}} = & \frac{32\pi^2 G_N^2}{E_p^3 \xi} c_1^2 \int \frac{d^d k}{(2\pi)^d} \frac{1}{|\vec{p}' - \vec{k}|^2 |\vec{p} - \vec{k}|^2 (k^2 - p^2)} \\ & - \frac{16\pi^2 G_N^2}{E_p^3 \xi^2} \left(\frac{c_1^2 (1 - \xi)}{2E_p^2 \xi} + 4c_1 p_1 \cdot p_3 \right) \int \frac{d^d k}{(2\pi)^d} \frac{1}{|\vec{p}' - \vec{k}|^2 |\vec{p} - \vec{k}|^2} + \dots \end{aligned} \quad (46)$$

Evaluating the remaining three-dimensional integrals, we find

$$\mathcal{M}^{\text{Iterated}} = \frac{i\pi G_N^2}{E_p^3 \xi} \frac{4c_1^2}{|\vec{p}'|} \frac{(\log |\vec{q}'|^2 - \frac{2}{3-d})}{|\vec{q}'|^2} + \frac{2\pi^2 G_N^2}{E_p^3 \xi^2 |\vec{q}'|} \left(\frac{c_1^2 (\xi - 1)}{2E_p^2 \xi} - 4c_1 p_1 \cdot p_3 \right) \quad (47)$$

The second-order post-Minkowskian potential in momentum space is thus given by

$$V_{2\text{PM}}(p, q) = \mathcal{M}^{1\text{-loop}} + \mathcal{M}^{\text{Iterated}} \quad (48)$$

leading to

$$V_{2\text{PM}}(p, q) = \frac{\pi^2 G_N^2}{E_p^2 \xi |\vec{q}'|} \left[\frac{1}{2} \left(\frac{c_{\triangleright}}{m_a} + \frac{c_{\triangleleft}}{m_b} \right) + \frac{2}{E_p \xi} \left(\frac{c_1^2 (\xi - 1)}{2E_p^2 \xi} - 4c_1 p_1 \cdot p_3 \right) \right] \quad (49)$$

or, in coordinate space,

$$V_{2\text{PM}}(p, r) = \frac{G_N^2}{r^2} \frac{1}{E_p^2 \xi} \left[\frac{1}{4} \left(\frac{c_{\triangleright}}{m_a} + \frac{c_{\triangleleft}}{m_b} \right) + \left(\frac{c_1^2 (\xi - 1)}{2E_p^3 \xi^2} - \frac{4c_1 p_1 \cdot p_3}{E_p \xi} \right) \right]. \quad (50)$$

This agrees with what has been previously obtained in ref. [8] (taking into account that c_1 here is $E^2 \xi$ times c_1 in [8]). As expected on physical grounds, the imaginary part which is composed of super-classical and infrared divergent pieces has cancelled, leaving a finite and well-defined post-Minkowskian potential at $d = 3$. That such cancellation had to occur was expected on physical ground, since the imaginary part clearly cannot affect classical motion. Interestingly, the evaluation of the same potential in $\mathcal{N} = 8$ supergravity has shown no contributions coming from triangle topologies [31].

C. The post-Minkowskian scattering angle

In [7] a one-loop formula for the gravitational eikonal limit [32, 33] generalized to the scattering of two objects of different masses m_a and m_b was used to deduce the classical scattering angle to second post-Minkowskian order directly from the scattering amplitude. An alternative method based on the Hamiltonian [34] has recently been revived in connection with the third post-Minkowskian scattering amplitude calculation [9, 10] and we here briefly summarize the method at second order in G_N . Since the motion lies on a plane, we can introduce the following coordinates on the phase space (r, ϕ, p_r, p_ϕ) so as to express the momentum in the center of mass frame as

$$p^2 = p_r^2 + \frac{L^2}{r^2} \quad (51)$$

being L the conserved angular momentum of our binary system, with constant energy E

$$\sqrt{p^2 + m_a^2} + \sqrt{p^2 + m_b^2} + V_{1\text{PM}}(p, r) + V_{2\text{PM}}(p, r) = E \quad (52)$$

This equation can be solved perturbatively in G_N for $p^2 = p^2(E, L, r)$

$$p^2 = p_0^2 + \frac{G_N f_1}{r} + \frac{G_N^2 f_2}{r^2} + \dots \quad (53)$$

Using $s = (p_1 + p_3)^2$

$$p_0^2 = \frac{(p_1 \cdot p_3)^2 - m_1^2 m_2^2}{s}, \quad f_1 = -\frac{2c_1}{\sqrt{s}}, \quad f_2 = -\frac{1}{2\sqrt{s}} \left(\frac{c_\triangleright}{m_a} + \frac{c_\triangleleft}{m_b} \right) \quad (54)$$

It is straightforward to derive the following expression for the change in the angular variable ϕ during scattering (see for instance [16, 34])

$$\Delta\phi = \pi + \chi(E, L), \quad (55)$$

where the scattering angle is given by

$$\chi(E, L) = -2 \int_{r_{\min}}^{+\infty} dr \frac{\partial p_r}{\partial L} - \pi \quad (56)$$

Here r_{min} is the positive root for the condition of turning point at $p_r = 0$ with

$$p_r = \sqrt{p_0^2 - \frac{L^2}{r^2} + \frac{G_N f_1}{r} + \frac{G_N^2 f_2}{r^2}} \quad (57)$$

Introducing $r_0 \equiv L/p_0$ we note that p_r can be rewritten as

$$p_r = \frac{p_0}{r} \sqrt{r^2 + r \frac{G_N f_1}{p_0^2} + \frac{G_N^2 f_2}{p_0^2} - r_0^2} = \frac{p_0}{r} \sqrt{r - r^+} \sqrt{r - r^-}, \quad (58)$$

$$r^\pm = -\frac{G_N f_1}{2p_0^2} \pm \sqrt{\frac{G_N^2 f_1^2}{4p_0^4} - \frac{G_N^2 f_2}{p_0^2} + r_0^2} \quad (59)$$

Since $r_{min} = r^+$, the scattering angle becomes

$$\chi(E, L) = 2 \int_{r^+}^{+\infty} \frac{dr}{r} \frac{r_0}{\sqrt{(r - r^+)(r - r^-)}} - \pi \quad (60)$$

The integral so expressed can be performed analytically without the need of regularization.

We get

$$\chi(E, L) = \frac{4r_0}{\sqrt{-r^+ r^-}} \arccos \sqrt{\frac{r^+}{r^+ - r^-}} - \pi \quad (61)$$

Taylor-expanding the scattering angle to second post-Minkowskian order we arrive at the final result

$$\chi(E, L) = \frac{G_N f_1}{p_0 L} + \frac{G_N^2 f_2 \pi}{2L^2} + \dots \quad (62)$$

In terms of $\hat{\mathcal{M}}^2 \equiv s - m_a^2 - m_b^2$ and the impact parameter b , where $L = pb$, we have

$$\chi(E, b) = \frac{4G_N s}{b} \left[\frac{\hat{\mathcal{M}}^4 - 2m_a^2 m_b^2}{\hat{\mathcal{M}}^4 - 4m_a^2 m_b^2} + \frac{3\pi}{16} \frac{G_N (m_a + m_b)}{b} \frac{5\hat{\mathcal{M}}^4 - 4m_a^2 m_b^2}{\hat{\mathcal{M}}^4 - 4m_a^2 m_b^2} \right] \quad (63)$$

which agrees with the result of [14] at second post-Minkowskian order. In particular, since f_1 and f_2 do not depend on box topologies (54), also the scattering angle (62) receives no contributions from these, a known fact from the eikonal approach in four dimensions. The details of the calculation based on the Hamiltonian is, on the surface, quite different from the eikonal approach. It would be interesting to establish the precise link between the two, first identifying the precise exponentiation formula for the eikonal limit beyond second

post-Minkowskian order.

V. CONCLUSION

Using the conventional approach to determining the interaction potential in perturbative gravity we have demonstrated that it can be extended to the relativistic setting by means of a one-particle Hamiltonian and associated Salpeter equation. We have used the Lippmann-Schwinger equation to derive straightforwardly the needed Born subtractions at arbitrary loop order. The resulting Fourier-transformed post-Minkowskian Hamiltonian

$$\mathcal{H}_{2\text{PM}}(p, r) = \sqrt{p^2 + m_a^2} + \sqrt{p^2 + m_b^2} + V_{1\text{PM}}(p, r) + V_{2\text{PM}}(p, r), \quad (64)$$

agrees with the one derived in ref. [8] based on an effective field theory expansion in operators that can contribute to the given order, supplemented with the matching condition that the scattering amplitude as computed in the effective theory agrees with the one computed from the full one-loop expression of the Einstein-Hilbert action (plus scalars).

The resulting post-Minkowskian Salpeter equation is not an effective low-energy theory (momentum is not limited), but rather a small $|\vec{q}|/m$ approximation where small momentum is exchanged and only particle states are summed over. It is encouraging that preliminary results indicate that the corresponding two-loop Hamiltonian [9] may improve the computation of two-body dynamics as compared to the conventional post-Newtonian expansion for bound states [10]. The post-Minkowskian Hamiltonian also appears to provide a short-cut towards computing the scattering angle without first demonstrating exponentiation (and potential correction terms) as in the eikonal approach. It would be interesting to demonstrate the equivalence between those two scattering angle computations in all generality.

Acknowledgments

This work has been based partly on funding from the European Union’s Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant agreement No. 764850 (“SAGEX”). The work of N.E.J.B.-B. and P.H.D. was supported in part by the Danish National Research Foundation (DNRF91). N.E.J.B.-B. in addition acknowledges partial support from the Carlsberg Foundation. The research of P.V. has received funding the

ANR grant “Amplitudes” ANR-17- CE31-0001-01, and is partially supported by Laboratory of Mirror Symmetry NRU HSE, RF Government grant, ag. N° 14.641.31.0001. P.V. also thanks the Galileo Galilei Institute for Theoretical Physics and INFN for hospitality and partial support during the workshop "String Theory from a worldsheet perspective" where part of this work has been done.

-
- [1] Y. Iwasaki, “Quantum theory of gravitation vs. classical theory. - fourth-order potential,” *Prog. Theor. Phys.* **46**, 1587 (1971). doi:10.1143/PTP.46.1587
 - [2] D. Neill and I. Z. Rothstein, “Classical Space-Times from the S Matrix,” *Nucl. Phys. B* **877**, 177 (2013) doi:10.1016/j.nuclphysb.2013.09.007 [[arXiv:1304.7263](#) [hep-th]].
 - [3] N. E. J. Bjerrum-Bohr, J. F. Donoghue and P. Vanhove, “On-shell Techniques and Universal Results in Quantum Gravity,” *JHEP* **1402**, 111 (2014) doi:10.1007/JHEP02(2014)111 [[arXiv:1309.0804](#) [hep-th]].
 - [4] V. Vaidya, “Gravitational spin Hamiltonians from the S matrix,” *Phys. Rev. D* **91**, no. 2, 024017 (2015) doi:10.1103/PhysRevD.91.024017 [[arXiv:1410.5348](#) [hep-th]].
 - [5] F. Cachazo and A. Guevara, [[arXiv:1705.10262](#) [hep-th]].
 - [6] A. Guevara, *JHEP* **1904**, 033 (2019) doi:10.1007/JHEP04(2019)033 [[arXiv:1706.02314](#) [hep-th]].
 - [7] N. E. J. Bjerrum-Bohr, P. H. Damgaard, G. Festuccia, L. Planté and P. Vanhove, “General Relativity from Scattering Amplitudes,” *Phys. Rev. Lett.* **121**, no. 17, 171601 (2018) doi:10.1103/PhysRevLett.121.171601 [[arXiv:1806.04920](#) [hep-th]].
 - [8] C. Cheung, I. Z. Rothstein and M. P. Solon, *Phys. Rev. Lett.* **121**, no. 25, 251101 (2018) doi:10.1103/PhysRevLett.121.251101 [[arXiv:1808.02489](#) [hep-th]].
 - [9] Z. Bern, C. Cheung, R. Roiban, C. H. Shen, M. P. Solon and M. Zeng, *Phys. Rev. Lett.* **122**, no. 20, 201603 (2019) doi:10.1103/PhysRevLett.122.201603 [[arXiv:1901.04424](#) [hep-th]].
 - [10] A. Antonelli, A. Buonanno, J. Steinhoff, M. van de Meent and J. Vines, “Energetics of Two-Body Hamiltonians in Post-Minkowskian Gravity,” *Phys. Rev. D* **99** (2019) no.10, 104004 doi:10.1103/PhysRevD.99.104004 [[arXiv:1901.07102](#) [gr-qc]].
 - [11] L. Blanchet, “Gravitational Radiation from Post-Newtonian Sources and Inspiralling Compact Binaries,” *Living Rev. Rel.* **17**, 2 (2014) doi:10.12942/lrr-2014-2 [[arXiv:1310.1528](#) [gr-qc]].

- [12] M. Levi, “Effective Field Theories of Post-Newtonian Gravity: A comprehensive review,” [arXiv:1807.01699](#) [hep-th].
- [13] A. Cristofoli, “An Effective Field Theory Approach to the Two-Body Problem in General Relativity,” Master thesis Padua university 2018
- [14] K. Westpfahl, “High-Speed Scattering of Charged and Uncharged Particles in General Relativity,” *Fortsch.Phys.*,33,417 (1985).
- [15] T. Ledvinka, G. Schafer and J. Bicak, “Relativistic Closed-Form Hamiltonian for Many-Body Gravitating Systems in the Post-Minkowskian Approximation,” *Phys. Rev. Lett.* **100**, 251101 (2008) doi:10.1103/PhysRevLett.100.251101 [[arXiv:0807.0214](#) [gr-qc]].
- [16] T. Damour, “Gravitational scattering, post-Minkowskian approximation and Effective One-Body theory,” *Phys. Rev. D* **94**, no. 10, 104015 (2016) doi:10.1103/PhysRevD.94.104015 [[arXiv:1609.00354](#) [gr-qc]].
- [17] D. Bini and T. Damour, “Gravitational spin-orbit coupling in binary systems at the second post-Minkowskian approximation,” *Phys. Rev. D* **98**, no. 4, 044036 (2018) doi:10.1103/PhysRevD.98.044036 [[arXiv:1805.10809](#) [gr-qc]].
- [18] L. Blanchet and A. S. Fokas, “Equations of motion of self-gravitating N -body systems in the first post-Minkowskian approximation,” *Phys. Rev. D* **98**, no. 8, 084005 (2018) doi:10.1103/PhysRevD.98.084005 [[arXiv:1806.08347](#) [gr-qc]].
- [19] T. Damour, “High-energy gravitational scattering and the general relativistic two-body problem,” *Phys. Rev. D* **97**, no. 4, 044038 (2018) doi:10.1103/PhysRevD.97.044038 [[arXiv:1710.10599](#) [gr-qc]].
- [20] A. Guevara, A. Ochirov and J. Vines, “Scattering of Spinning Black Holes from Exponentiated Soft Factors,” [arXiv:1812.06895](#) [hep-th].
- [21] J. F. Donoghue, “General relativity as an effective field theory: The leading quantum corrections,” *Phys. Rev. D* **50**, 3874 (1994) doi:10.1103/PhysRevD.50.3874 [[gr-qc/9405057](#)].
- [22] N. E. J. Bjerrum-Bohr, J. F. Donoghue and B. R. Holstein, “Quantum gravitational corrections to the nonrelativistic scattering potential of two masses,” *Phys. Rev. D* **67**, 084033 (2003) Erratum: [*Phys. Rev. D* **71**, 069903 (2005)] doi:10.1103/PhysRevD.71.069903, 10.1103/PhysRevD.67.084033 [[hep-th/0211072](#)].
- [23] B. R. Holstein and J. F. Donoghue, “Classical physics and quantum loops,” *Phys. Rev. Lett.* **93**, 201602 (2004) doi:10.1103/PhysRevLett.93.201602 [[hep-th/0405239](#)].

- [24] D. A. Kosower, B. Maybee and D. O’Connell, “Amplitudes, Observables, and Classical Scattering,” *JHEP* **1902**, 137 (2019) doi:10.1007/JHEP02(2019)137 [[arXiv:1811.10950](#) [hep-th]].
- [25] B. R. Holstein and A. Ross, “Spin Effects in Long Range Gravitational Scattering,” [arXiv:0802.0716](#) [hep-ph]; “Spin Effects in Long Range Electromagnetic Scattering,” [arXiv:0802.0715](#) [hep-ph].
- [26] A. Koemans Collado, P. Di Vecchia and R. Russo, “Revisiting the 2PM eikonal and the dynamics of binary black holes,” [arXiv:1904.02667](#) [hep-th].
- [27] A. Brandhuber and G. Travaglini, “On higher-derivative effects on the gravitational potential and particle bending,” [arXiv:1905.05657](#) [hep-th].
- [28] W. T. Emond and N. Moynihan, “Scattering Amplitudes, Black Holes and Leading Singularities in Cubic Theories of Gravity,” [arXiv:1905.08213](#) [hep-th].
- [29] J. F. Donoghue and T. Torma, “On the Power Counting of Loop Diagrams in General Relativity,” *Phys. Rev. D* **54** (1996) 4963 doi:10.1103/PhysRevD.54.4963 [[hep-th/9602121](#)].
- [30] S. Weinberg, “Infrared photons and gravitons,” *Phys. Rev.* **140**, B516 (1965). doi:10.1103/PhysRev.140.B516
- [31] S. Caron-Huot and Z. Zahraee, “Integrability of Black Hole Orbits in Maximal Supergravity,” [arXiv:1810.04694](#) [hep-th].
- [32] D. N. Kabat and M. Ortiz, “Eikonal quantum gravity and Planckian scattering,” *Nucl. Phys. B* **388**, 570 (1992) doi:10.1016/0550-3213(92)90627-N [[hep-th/9203082](#)].
- [33] R. Akhouri, R. Saotome and G. Sterman, “High Energy Scattering in Perturbative Quantum Gravity at Next to Leading Power,” [arXiv:1308.5204](#) [hep-th].
- [34] T. Damour and G. Schaefer, “Higher Order Relativistic Periastron Advances and Binary Pulsars,” *Nuovo Cim. B* **101**, 127 (1988). doi:10.1007/BF02828697