A Study on Convolution Using Half-Precision Floating-Point Numbers on GPU for Radio Astronomy Deconvolution

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3 Application to radio astronomy



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Deconvolution using FP16 on GPU

Deconvolution

Trying to recover data from an acquisition...



Source: I.M. Steward, Bielefeld University

Helps in the analysis in many fields: from radio astronomy to microscope images.

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Modelisation

A physical measurement can be modeled as:

$$g = Hf + n$$

- $\boldsymbol{g} \in \mathbb{R}^{M}$ is the measured data
- $\boldsymbol{H} \in \mathbb{R}^{M \times N}$ is the linear observation model (here: a convolution)
- $\boldsymbol{f} \in \mathbb{R}^N$ is the ground truth, unknown
- $\boldsymbol{n} \in \mathbb{R}^{M}$ is added noise, usually Gaussian

Regularization

$$oldsymbol{J}(oldsymbol{f}) = \|oldsymbol{g} - oldsymbol{H}oldsymbol{f}\|^2 + \lambda \|oldsymbol{D}oldsymbol{f}\|^2$$

The criterion involves two terms: data fidelity and regularization to compensate the ill-posedness behavior. In the following, we choose $D = I^{M \times M}$ The solution is:

The Hessian matrix's dimensions are too large to inverse it: we fall back on an iterative solver instead.

The gradient descent algorithm

Require: H, λ , g, ϵ , N, c1: Set $\boldsymbol{b} = \boldsymbol{H}^t \boldsymbol{q}$ and $\boldsymbol{Q} = \boldsymbol{H}^t \boldsymbol{H} + \lambda \boldsymbol{I}$ 2: Set $f^{(0)}$ and $n \leftarrow 0$ 3: repeat $\boldsymbol{k} \leftarrow \boldsymbol{H}^{t} \boldsymbol{H} \boldsymbol{f}^{(n)} - \boldsymbol{b} + 2\lambda \boldsymbol{f}^{(n)}$ 4: $\alpha \leftarrow \mathbf{k}^t \mathbf{k} / \mathbf{k}^t \mathbf{Q} \mathbf{k}$ \triangleright or $\alpha \leftarrow c$ with $c \leq \frac{2}{\|Q\|}$ 5: $oldsymbol{f}^{(n+1)} \leftarrow oldsymbol{f}^{(n)} - lpha oldsymbol{k}$ 6: 7: $n \leftarrow n+1$ until Some criterion is met \triangleright e.g.: $n \ge N$ 8. return $f^{(n)}$

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Up to 4 convolutions per loop!

2D convolution



Convolving a $N \times N$ image with a $K \times K$ kernel can mainly be done in two ways:

- In direct space: complexity $O(N^2 K^2)$
- In Fourier space: complexity $O(N^3 \log(N))$

In special cases, optimization exists: separable kernel, overlap-add...

16-bit floats format



- Supported on an increasing number of platforms
- Double the FLOPS on Nvidia GPUs via SIMD instructions
- Decreased memory and bandwidth usage
- Limited range (from 10^{-8} to 65504) and accuracy

From FP32 to FP16



• We suffer from quantization errors in the conversion

• The relative error stays the same

Results on convolutions

Convolution of a 512 \times 512 image with a square kernel of variable size. Only the kernel execution time is measured.



Figure: Left: results with FP32. Right: speed-up with FP16

- Naive and PRCF: own implementations.
- cuFFT looks constant because kernel is padded

Accuracy



Figure: MRE compared to a reference implementation

$$MRE = rac{1}{N} \sum \left\{ ig| rac{x[i] - y[i]}{x[i]}
ight|, ext{ if } x[i]
eq 0$$
0, otherwise

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Application to radio astronomy



PSF (cubic root)

Sky (clipped to 0.001)

1000:1 ratio between star values and background field.

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Errors on "real" data



Figure: Left: convolution with FP32, right: with FP16

• Important loss of precision with a single convolution!

Deconvolution Results



Figure: Reconstructions of f with different precisions (the cubic root is displayed for better contrast)

Minimizing the criterion

$$oldsymbol{J}(oldsymbol{f}) = \|oldsymbol{g} - oldsymbol{H}oldsymbol{f}\|^2 + \lambda \|oldsymbol{f}\|^2$$



- Half-float only implementation does not have a satisfying behavior.
- Mixed strategy performs well.

Summary & Conclusions

- We developed a GPU convolution benchmark
- We made a CUDA/C++ implementation of a deconvolution algorithm

Summary & Conclusions

- We developed a GPU convolution benchmark
- \bullet We made a CUDA/C++ implementation of a deconvolution algorithm
- Half floats can significantly speed-up convolutions (up to $\times 2, \times 3)$
- Loss of precision must be tightly controlled...
- ... but can be mitigated with the use FP32 compute / FP16 storage

Thank you!

L2S, Lab. Lagrange

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Sharan Chetlur et. al. 2015