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Performance of the Low-Rank Adaptive Normalized Matched Filter Test Under a Large Dimension Regime

Alice Combernoux, Frédéric Pascal, *Senior Member, IEEE*, Guillaume Ginolhac, *Member, IEEE*

Abstract—When a possible target is embedded in a Low Rank (LR) Gaussian clutter (which is contained in a low dimensional subspace) plus a white Gaussian noise, the detection process can be performed by applying the Low-Rank Adaptive Normalized Matched Filter (LR-ANMF) which is a function of the estimated projector. In a recent work, we derived an approximate distribution of the LR-ANMF under the \mathcal{H}_0 hypothesis by using a restrictive hypothesis (the target has to be orthogonal to the clutter subspace). In this paper, we propose to determine new approximations of the Pfa and the Pd of the LR-ANMF by relaxing this restrictive hypothesis. This new derivation is based on results concerning the convergence in a large dimension regime of quadratic forms. Simulations validate our result, in particular when the tested signal is close to the clutter subspace.

Index Terms—Adaptive detection, Low-Rank Covariance Matrix, Approximate Detector Distribution, Random Matrix Theory, RADAR

I. INTRODUCTION

In order to detect a known signal coming from a target embedded in a Gaussian noise characterized by its unknown covariance matrix, the Generalized Likelihood Ratio Test (GLRT) is derived which leads to the well-known Kelly's detector [1]. This detector allows to choose between two hypothesis \mathcal{H}_1 (target present) and \mathcal{H}_0 (target absent). The theoretical Probability of False Alarm (Pfa) and Probability of Detection (Pd) are derived in order to set a detection threshold for a chosen Pfa. Another approach is to derive the GLRT by assuming to know the covariance matrix and by replacing it in the detector by the Maximum Likelihood Estimator (MLE) of the theoretical covariance matrix [2] (in Gaussian, the MLE is the Sample Covariance Matrix - SCM). The theoretical performance of this detector are also provided even for subspace detectors [3], [4].

Unfortunately, an accurate estimation of the covariance matrix needs a large number of data (denoted secondary data and assumed to be iid and free signal) $K \approx 2m$ (where m is the data size) [5], [6]. To reduce this number K , it is possible to regularize the SCM [7] when any information on the system is available. In array processing applications, the received antenna is often symmetric which allows to take the persymmetry property of the data into account in order to divide by 2 the number of secondary data needed [8]–[10].

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It is also possible to handle with symmetry properties of the disturbance [11], [12] to reduce this number of secondary data. In the case where the disturbance can be decomposed by a sum of two contributions, a White Gaussian Noise (WGN) and a Low-Rank (LR) Gaussian contribution (called clutter in RADAR or SONAR), the LR Adaptive Normalized Matched Filter (LR-ANMF) has been proposed [10], [13]. In this detector, the estimation of the covariance matrix is replaced by the estimation of the projector onto the clutter subspace which is obtained by the Eigenvalue Decomposition (EVD) of the SCM. When the LR subspace r is small compared to the data size m , the numerator of the LR-ANMF has interesting properties: only $K = 2r$ secondary data are needed to obtain equivalent performance compared to the numerator of the ANMF built with the SCM and $K = 2m$ [14], [15]. This theoretical result is difficult to obtain with the LR-ANMF or even for other LR adaptive detectors [16]–[18]. When the LR-ANMF is built with the theoretical projector, its theoretical performance are available in [13]. But to our knowledge, the theoretical performance of the LR-ANMF at fixed dimension does not exist in the literature. Therefore, we have to resort to asymptotic methods to approximate its performance. In [19], we have proposed a result approximating the LR-ANMF distribution under the \mathcal{H}_0 hypothesis. To derive this result, several hypotheses have been used. The two first are mostly common: the norm of the steering vector \mathbf{d} is assumed to be equal to 1 and the Clutter to Noise Ratio (CNR) is assumed to be large. This last assumption is realistic in many applications such as Spatio-Temporal Adaptive Processing (STAP) Radar. For the mathematical derivations, a last hypothesis is needed: the projection of the steering vector onto the theoretical interference subspace is negligible which means that the tested steering vector is not fully embedded in the clutter subspace. This hypothesis is often needed when we derive theoretical performance with a projector [10], [14], [15], [20], [21]. The result derived in [19] is interesting because it first shows the CFAR behavior of the LR-ANMF for a tested vector far from the clutter subspace and allows to quickly evaluate an approximate Pfa for a chosen threshold. But in general, the LR-ANMF is more useful when the tested steering vectors are close to the clutter subspace (such as STAP to detect slow moving targets): it will be more suitable to derive its theoretical performance in these particular conditions. We propose in this paper a new approximation of the LR-ANMF distribution under \mathcal{H}_0 and \mathcal{H}_1 (and not only \mathcal{H}_0 as in [19]) by relaxing the last hypothesis about the orthogonality between

the steering vector and the theoretical interference subspace. The approximations for Pfa and Pd for the LR-ANMF are also deduced from this approximate distribution.

This new derivation is first based on Random Matrix Theory (RMT) results. The LR-ANMF is a function of three Quadratic Forms (QFs), $\mathbf{s}_1^H \hat{\Pi}_c^\perp \mathbf{s}_2$ (where $\hat{\Pi}_c^\perp$ is a projector estimate and \mathbf{s}_1 and \mathbf{s}_2 are two vectors which will be explicitated later) which have been broadly studied since ten years in the signal processing community [22]–[25]. The principle is to examine the spectral behavior of an estimate of the theoretical covariance matrix, like the SCM, by RMT for obtaining its convergence, performance and asymptotic distribution in the desired convergence regime. The spectral behavior is generally studied under two different regimes: when K tends to infinity, for fixed m and when both the data size m and K tend to infinity at the same rate (denoted as *large dimensional regime*), i.e. $c_m \triangleq m/K \rightarrow c \in]0, +\infty[$, for different models of the observed data as in [22], [23], [26], [25] and [24]. These results show that the QFs of the form of $\mathbf{s}_1^H \hat{\Pi}_c^\perp \mathbf{s}_2$ do not converge to the theoretical one in the large dimensional regime. The authors of [26]–[28] have proposed new MUSIC algorithms which ensure the consistency property in this large dimensional regime. In these new algorithms, the estimated projector is replaced by a new one often called pseudo-projector in the literature.

By using these convergence results, we have first proposed in [29] to integrate the new QFs of [26]–[28] in the classical detection framework and we have shown by Monte-Carlo simulations that this new detector leads to better performance in a STAP context in particular when the target is close to the clutter. In [30], we derived a first approximation of the LR-ANMF under the \mathcal{H}_0 hypothesis but the result suffers of a main drawback: the obtained approximate distribution is not a known distribution and is a function of several random variables. Therefore Monte Carlo simulations are still needed to obtain the final result and it is impossible to have a Pfa-threshold relation which easily sets the threshold for a chosen Pfa. In this paper, we propose a new result for the approximation of the LR-ANMF distribution which is related to a known distribution and allows to improve the result of [30]. Moreover, this new derivation allows to have an approximation of the LR-ANMF distribution under the \mathcal{H}_1 hypothesis. This new derivation is based on convergence results in a large dimension regime and the approximate distribution of a ratio of two random QFs proposed in [31]. We validate hereafter in this paper our theoretical result by comparing it with Monte-Carlo simulations in a jamming application which consists in detecting a signal among different strong interferences. These interferences are contained in a LR subspace and therefore this application is well adapted to validate our result. But, our result will remain valid in other applications such as STAP or STAP-MIMO where the LR subspace is composed by different contributions of the ground.

Paper is organized as follows: section II presents the problem statement, section III contains the main result of the paper which is the approximate distribution of the LR-ANMF and section IV shows the results of numerical simulations.

The following convention is adopted: italic indicates a

scalar quantity, lower case boldface indicates a vector quantity and upper case boldface a matrix. T denotes the transpose operator and H the transpose conjugate. $E[\cdot]$ is the expected value operator. $\mathcal{CN}(\mathbf{a}, \mathbf{M})$ (resp. $\mathcal{N}(\mathbf{a}, \mathbf{M})$) is a complex (resp. real) Gaussian vector of mean \mathbf{a} and of covariance matrix \mathbf{M} . \mathbf{I}_m is the $m \times m$ -identity matrix. $\chi^2(n)$ is a Chi-square random variable with n degrees of freedom. $\mathcal{N}_c(a, \sigma^2)$ is complex Gaussian random variable of mean a and variance σ^2 . \sim means "distributed as". $\Re(a)$ and $\Im(a)$ designs the real and imaginary part of a complex number a .

II. PROBLEM STATEMENT

A. Derivation of the LR-ANMF

The stated problem is to infer if the received signal $\mathbf{x} \in \mathbb{C}^{m \times 1}$, corrupted by an additive disturbance, also contains a complex signal $\alpha \mathbf{d}(\Theta)$. One also has a set of K secondary data $\{\mathbf{x}_k\}$ which are signal free realizations of the disturbance. The two hypotheses are then:

$$\begin{cases} \mathcal{H}_0 : \mathbf{x} = \mathbf{c} + \mathbf{n} & \mathbf{x}_k = \mathbf{c}_k + \mathbf{n}_k, \quad k \in \llbracket 1, K \rrbracket \\ \mathcal{H}_1 : \mathbf{x} = \alpha \mathbf{d}(\Theta) + \mathbf{c} + \mathbf{n} & \mathbf{x}_k = \mathbf{c}_k + \mathbf{n}_k, \quad k \in \llbracket 1, K \rrbracket \end{cases}, \quad (1)$$

where \mathbf{d} is the steering vector, α is an unknown deterministic parameter and Θ is an unknown deterministic vector (in the following, Θ will be omitted). $\mathbf{n} \in \mathbb{C}^{m \times 1}$ (or \mathbf{n}_k) $\sim \mathcal{CN}(0, \gamma \mathbf{I}_m)$ is the additive WGN complex vector. $\mathbf{c} \in \mathbb{C}^{m \times 1}$ (or \mathbf{c}_k) $\sim \mathcal{CN}(0, \mathbf{C})$ is the Gaussian clutter. Consequently, the covariance matrix of the secondary data can be written as $\mathbf{R} = \mathbf{C} + \gamma \mathbf{I}_m \in \mathbb{C}^{m \times m}$. Moreover, the clutter is considered low-rank r^1 . Hence, $\text{rank}(\mathbf{C}) = r \ll m$ and one could write the eigendecomposition of \mathbf{C} :

$$\mathbf{C} = \sum_{i=1}^r \gamma_i \mathbf{u}_i \mathbf{u}_i^H, \quad (2)$$

where γ_i and \mathbf{u}_i , $i \in \llbracket 1; r \rrbracket$ are respectively the non zero eigenvalues and the associated eigenvectors of \mathbf{C} . We define the projector onto the clutter subspace Π_c and the projector onto the orthogonal of the clutter subspace $\Pi_c^\perp = \mathbf{I}_m - \Pi_c$ [14], [15]:

$$\begin{aligned} \Pi_c &= \sum_{i=1}^r \mathbf{u}_i \mathbf{u}_i^H \\ \Pi_c^\perp &= \mathbf{I}_m - \sum_{i=1}^r \mathbf{u}_i \mathbf{u}_i^H. \end{aligned} \quad (3)$$

Since Π_c^\perp is not known in practice, we have to estimate it from the secondary data $\{\mathbf{x}_k\}$. The classical estimation is based on the EVD of the SCM:

$$\hat{\mathbf{R}} = \frac{1}{K} \sum_{k=1}^K \mathbf{x}_k \mathbf{x}_k^H = \sum_{i=1}^r \hat{\lambda}_i \hat{\mathbf{u}}_i \hat{\mathbf{u}}_i^H + \sum_{i=r+1}^m \hat{\lambda}_i \hat{\mathbf{u}}_i \hat{\mathbf{u}}_i^H, \quad (4)$$

where $\hat{\lambda}_i$ and $\hat{\mathbf{u}}_i$ are the estimated eigenvalues and eigenvectors. Finally, the estimated projectors are:

$$\begin{aligned} \hat{\Pi}_c &= \sum_{i=1}^r \hat{\mathbf{u}}_i \hat{\mathbf{u}}_i^H \\ \hat{\Pi}_c^\perp &= \mathbf{I}_m - \sum_{i=1}^r \hat{\mathbf{u}}_i \hat{\mathbf{u}}_i^H. \end{aligned} \quad (5)$$

¹The rank is assumed to be known in this paper such as STAP applications according to Brennan's formula [32], [33] $r = N + (M - 1) * \beta$ (N is the number of sensors, M the number of pulses and β depends on radar parameters). If not, it is possible to estimate it for example with new methods based on Random Matrix Theory tools [34].

The conventional adaptive test to the problem of detection (1) has been proposed in [13] and is called the LR-Adaptive NMF (LR-ANMF):

$$\hat{\Lambda}_{LR} = \frac{|\mathbf{d}^H \hat{\mathbf{\Pi}}_c^\perp \mathbf{x}|^2}{(\mathbf{d}^H \hat{\mathbf{\Pi}}_c^\perp \mathbf{d})(\mathbf{x}^H \hat{\mathbf{\Pi}}_c^\perp \mathbf{x})} \underset{\mathcal{H}_0}{\overset{\mathcal{H}_1}{\gtrless}} \eta, \quad (6)$$

where $\underset{\mathcal{H}_0}{\overset{\mathcal{H}_1}{\gtrless}} \eta$ means that the \mathcal{H}_1 hypothesis (respectively \mathcal{H}_0) is decided if the test $\hat{\Lambda}_{LR}$ is over (respectively under) the threshold η .

B. Previous result on theoretical LR-ANMF distribution

In [19], we have obtained the following result for an approximate distribution under the \mathcal{H}_0 hypothesis of $\hat{\Lambda}_{LR}$.

Proposition II.1 *The LR-ANMF of Eq. (6), $\hat{\Lambda}_{LR}$, can be expressed as a function of 6 independent random variables:*

$$\hat{\Lambda}_{LR} \sim \hat{\Lambda}_{LR_{th}} = \frac{|\alpha|^2}{\beta}, \quad (7)$$

with

$$\begin{aligned} \alpha &= b_1 - \frac{1}{K} X_1 X_3 s_1 \\ \beta &= |b_1|^2 + X_2^2 - \frac{2}{K} X_1 X_3 \Re(b_1 s_1 + X_2 s_2), \end{aligned} \quad (8)$$

where $X_1 \sim \sqrt{\frac{1}{2}\chi^2(2r)}$, $X_2 \sim \sqrt{\frac{1}{2}\chi^2(2(m-r-1))}$, $X_3 \sim \sqrt{\frac{1}{2}\chi^2(2K)}$, $s_1 \sim \mathcal{N}_c(0, 1)$, $s_2 \sim \mathcal{N}_c(0, 1)$ and $b_1 \sim \mathcal{N}_c(0, 1)$ are independent random variables.

To derive this result, several assumptions have been used. First, we assumed that the norm of steering vector \mathbf{d} is equal to 1 and that the CNR is large which leads to: $\gamma_1, \dots, \gamma_r \gg \gamma$. This assumption is realistic in many applications such as STAP. The last assumption is the most restrictive but it is needed for the mathematical derivations: the projection of the steering vector onto the theoretical interference subspace is negligible, i.e. $\mathbf{u}_i^H \mathbf{d} \approx 0$ for $i = 1, \dots, r$. This means that the tested steering vector is not fully embedded in the clutter ridge. This assumption is often needed when we derive theoretical performance with projector [10], [14], [15], [20], [21].

The result of Prop. II.1 shows that the LR-ANMF is Constant False Alarm Rate (CFAR) when the tested target is far from the LR subspace. But in practice and by simulation, it is well known that this property is not guaranteed when the target is close to the LR subspace. Actually, the Pfa then depends on the distance between the tested vector and the clutter subspace. Therefore, we propose in this paper a new approximation of the LR-ANMF distribution by relaxing the last assumption about the orthogonality between the steering vector and the theoretical interference subspace which allows to obtain a valid result for any tested target. Moreover, this new derivation will allow us to obtain an approximate distribution (and not a function of several random variables) under \mathcal{H}_0 and \mathcal{H}_1 hypothesis leading to obtain the approximations for Pfa and Pd. The approach used in this paper is based on RMT tools and in particular on the convergences of the three QFs of Eq. (6).

C. RMT model and convergence results

In this subsection, we present the standard convergence results in a large dimension regime of the QFs.

1) *Covariance matrix models:* These convergence results are based on the behavior of the eigenvalues of the SCM when $m, K \rightarrow \infty$ at the same rate, i.e. $c_m \triangleq m/K \rightarrow c \in]0, +\infty[$. The general assumptions that are useful to the different convergences will be first presented. Next, the covariance model used in this paper, the *G-MUSIC* model [22] [23], will be provided. The studied data models are characterized by their formulation and in a more important way by the asymptotic behavior of the covariance matrix eigenvalues when $m, K \rightarrow \infty$ at the same rate. Then, to prove convergences, we make the following standard assumptions:

- (As1) \mathbf{R} has uniformly bounded spectral norm $\forall m \in \mathbb{N}^*$, i.e. $\forall i \in \llbracket 1, m \rrbracket$, $|\lambda_i| < \infty$. More precisely, let \mathbf{R}_m be the covariance matrix indexed by its size m and $\lambda_{1,m}, \dots, \lambda_{m,m}$ its eigenvalues. Then, we assume that $\sup_m \max_{1 \leq i \leq m} |\lambda_{i,m}| < \infty$.
- (As2) The vectors $\mathbf{s}_1, \mathbf{s}_2 \in \mathbb{C}^{m \times 1}$ used in the QFs (here $\mathbf{d}(\Theta)$ or $\mathbf{x}/\|\mathbf{x}\|_2$) have uniformly bounded Euclidean norm $\forall m \in \mathbb{N}^*$.
- (As3) The elements of the Random Matrix have to be Gaussian².

Under the previous assumptions, the behavior of the SCM eigenvalues can be studied in a large dimension regime. By building the matrix $\mathbf{X} = [\mathbf{x}_1, \dots, \mathbf{x}_K]$, the SCM is rewritten as $\hat{\mathbf{R}} = \frac{1}{K} \mathbf{X} \mathbf{X}^H$. The matrix \mathbf{X} is a function of the theoretical covariance matrix:

$$\mathbf{X} = \mathbf{R}^{1/2} \mathbf{Y} = (\mathbf{I}_m + \mathbf{C})^{1/2} \mathbf{Y} \quad (9)$$

where $\mathbf{R}^{1/2}$ is the $m \times m$ Hermitian positive definite square root of the covariance matrix and the matrix $\mathbf{Y} \in \mathbb{C}^{m \times K}$ has iid entries $y_{ij} \sim \mathcal{CN}(0, 1)$. Under the previous assumptions, the following properties are satisfied for this matrix: the probability law of \mathbf{Y} is invariant by left multiplication with a deterministic unitary matrix. The eigenvalues empirical PDF of $\frac{1}{K} \mathbf{Y} \mathbf{Y}^H$ a.s. converges to the Marčenko-Pastur distribution [35] with support $[(1-\sqrt{c})^2, (1+\sqrt{c})^2]$ and the maximum (resp. minimum) eigenvalue of $\frac{1}{K} \mathbf{Y} \mathbf{Y}^H$ a.s. tends to $(1+\sqrt{c})^2$ (resp. to $(1-\sqrt{c})^2$). From this last result, the convergence of the eigenvalues of the SCM are addressed through the convergence of the Stieltjes transform of the eigenvalues CDF and are given for the G-MUSIC model in [36].

Finally, a final assumption for the G-MUSIC model is needed to obtain the convergences of quadratic forms that are functions of estimated projectors. This will provide the value of c which leads to separated clusters in $f(x)$ for all the theoretical eigenvalues of \mathbf{R} . This will be referred as the *separation condition*.

- (As4) The ratio $c = m/K$ is chosen such that $1/c$ is strictly greater than $\{\psi(\bar{f}_i) : i \in \llbracket 1, t \rrbracket\}$ which depends on m , where $\psi(f)$ is defined in [22].

²Although the paper only considers Gaussian random variables, the results are still valid with random variables which have finite eighth order moments.

Finally, the useful convergences of simple QFs (as $\mathbf{s}_1^H \hat{\Pi} \mathbf{s}_2$) will be given in the next subsection.

2) *Convergence of simple QFs:* Under (As1-As4) and when $m, K \rightarrow \infty$ at the same rate ($m/K \rightarrow c \in \mathbb{R}^{+*}$), then [22]:

$$\mathbf{s}_1^H \hat{\Pi}_c^\perp \mathbf{s}_2 \xrightarrow[m/K \rightarrow c < \infty]{m, K \rightarrow \infty \text{ a.s.}} \mathbf{s}_1^H \bar{\Pi}_c^\perp \mathbf{s}_2 \quad (10)$$

with, $\bar{\Pi}_c^\perp = \sum_{i=1}^m \psi_i \mathbf{u}_i \mathbf{u}_i^H$ [22], and:

$$\psi_i = \begin{cases} 1 - \frac{1}{m-r-1} \sum_{n=1}^r \left(\frac{1}{\lambda_n-1} - \frac{\mu_m}{\lambda_n-\mu_m} \right), & \text{if } i > r \\ \frac{1}{\lambda_i-1} - \frac{\mu_m}{\lambda_i-\mu_m}, & \text{else} \end{cases}, \quad (11)$$

where $\mu_1 \geq \dots \geq \mu_m$ are the eigenvalues of $\text{diag}(\boldsymbol{\lambda}) - \frac{1}{K} \sqrt{\boldsymbol{\lambda}} \sqrt{\boldsymbol{\lambda}}^T$ and $\boldsymbol{\lambda} = [\lambda_1, \dots, \lambda_m]^T$.

In the following, we denote:

$$\bar{\Pi}_c^\perp = \mathbf{U} \boldsymbol{\Psi} \mathbf{U}^H, \quad (12)$$

where $\mathbf{U} = [\mathbf{u}_1 \dots \mathbf{u}_m]$ and $\boldsymbol{\Psi} = \text{diag}(\psi_i)_{i \in [1, m]}$

III. APPROXIMATE PFA AND PD OF THE LR-ANMF

In this section, we derive an approximate Pfa and Pd for the LR-ANMF. The derivation is decomposed in these steps:

- first, we compute the convergence of the LR-ANMF in the large dimension regime in order to obtain an equivalent detector without a random projector. After some algebraic manipulations, we will obtain a ratio of two QFs which are function of real Gaussian random variables;
- from this ratio, we will compute an approximate distribution and we will present the algorithm to determine the parameters of this new distribution;
- finally the Pfa and the Pd will be evaluated from this approximate distribution.

A. Approximate distribution of the LR-ANMF

Let \mathbf{x} be the signal under test. From Eq. (1), $\mathbf{x} \sim \mathcal{CN}(\boldsymbol{\mu}_x, \mathbf{R})$ where $\boldsymbol{\mu}_x = \mathbf{0}$ under \mathcal{H}_0 and $\boldsymbol{\mu}_x = \alpha \mathbf{d}$ under \mathcal{H}_1 .

Proposition III.1 Equivalent detector LR-ANMF $\hat{\Lambda}_{LR}$ in a large dimension regime

Under (As1-As4) and when $m, K \rightarrow \infty$ at the same rate ($m/K \rightarrow c \in \mathbb{R}^{+*}$), we have:

$$\hat{\Lambda}_{LR} \xrightarrow[m/K \rightarrow c < \infty]{m, K \rightarrow \infty \text{ a.s.}} \bar{\Lambda} = \frac{\tilde{\mathbf{z}}^T \tilde{\mathbf{A}} \tilde{\mathbf{z}}}{\tilde{\mathbf{z}}^T \tilde{\mathbf{H}} \tilde{\mathbf{z}}}, \quad (13)$$

with $\tilde{\mathbf{z}} \in \mathbb{R}^{2m \times 1}$ and $\tilde{\mathbf{z}} \sim \mathcal{N}(\boldsymbol{\mu}_{\tilde{\mathbf{z}}} = \sqrt{2} [\Re(\boldsymbol{\Lambda}^{-1/2} \mathbf{U}^H \boldsymbol{\mu}_x)^T, \Im(\boldsymbol{\Lambda}^{-1/2} \mathbf{U}^H \boldsymbol{\mu}_x)^T]^T, \mathbf{I}_{2m})$, $\boldsymbol{\Lambda} = \text{diag}([\lambda_1, \dots, \lambda_m])$ and

$$\begin{cases} \tilde{\mathbf{A}} = \begin{bmatrix} \Re(\mathbf{A}) & -\Im(\mathbf{A}) \\ \Im(\mathbf{A}) & \Re(\mathbf{A}) \end{bmatrix} \in \mathbb{R}^{2m \times 2m} \\ \tilde{\mathbf{H}} = \begin{bmatrix} \mathbf{H} & \mathbf{O}_m \\ \mathbf{O}_m & \mathbf{H} \end{bmatrix} \in \mathbb{R}^{2m \times 2m} \end{cases}, \quad (14)$$

where $\mathbf{A} = \mathbf{H}^{1/2} \mathbf{f} \mathbf{f}^H \mathbf{H}^{1/2}$, $\mathbf{H} = \text{diag}([\eta_1, \dots, \eta_m])$, $\eta_i = \psi_i \lambda_i$ and $\mathbf{f} = \boldsymbol{\Psi}^{1/2} \mathbf{U}^H \mathbf{d} / \|\boldsymbol{\Psi}^{1/2} \mathbf{U}^H \mathbf{d}\|_2$.

Proof: See appendix VI-A. •

The expression $\bar{\Lambda}$ is now a ratio of two simple QFs, which are functions of real Gaussian random variables. The approximation by polynomial adjustment given in [31] is then fully adapted to this new form. The distribution of $\bar{\Lambda}$ under hypothesis \mathcal{H}_k , $k \in \{0, 1\}$ and its approximation will be respectively denoted $f_{\bar{\Lambda}}(x|\mathcal{H}_k)$ and $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$. Next, we have the following proposition:

Proposition III.2 Approximate distribution of $\hat{\Lambda}_{LR}$ in a large dimension regime

$f_{\bar{\Lambda}}(x|\mathcal{H}_k)$ is approximated by the following distribution:

$$f_{\bar{\Lambda}}(x|\mathcal{H}_k) \approx \tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k) = \varphi(x|\mathcal{H}_k) \sum_{i=0}^d \nu_i |_{\mathcal{H}_k} x^i, \quad (15)$$

where $\varphi(x|\mathcal{H}_k)$ is a basic distribution (for example the exponential or the Beta distribution) of same support than $f_{\bar{\Lambda}}(x|\mathcal{H}_k)$ and $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$, $\nu_i |_{\mathcal{H}_k}$ are the coefficients of the polynomial of adjustment and d its degree.

Proof: The proof is straightforward by following the result of [31]. •

B. Evaluation of $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$

The choice of d is important and is linked to the quality of the approximation but also to the computational cost (especially for \mathcal{H}_1). Next, the evaluation of $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$ is performed in several steps:

- computation of the moments of $\bar{\Lambda}$;
- computation of the moments of $\varphi(x|\mathcal{H}_k)$;
- equalization of the moments of $\bar{\Lambda}$ and $\varphi(x|\mathcal{H}_k)$ in order to compute the elements $\{\nu_i |_{\mathcal{H}_k}\}_{(1, d)}$.

It is important to notice that the support of $\varphi(x|\mathcal{H}_k)$ has to match with that of $f_{\bar{\Lambda}}(x|\mathcal{H}_k)$ and $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$.

Computation of the moments of $\bar{\Lambda}$:

In this paragraph, we compute the moments $\mu_i^{(\bar{\Lambda})} |_{\mathcal{H}_k}$ of $\bar{\Lambda}$. They are determined by the following relation [31], [37] from the ratio of Eq (13) in the proposition III.1:

$$\begin{aligned} \mu_i^{(\bar{\Lambda})} |_{\mathcal{H}_k} &= \frac{1}{\Gamma(i)} \exp\left(-\frac{\|\boldsymbol{\mu}_{\tilde{\mathbf{z}}}\|_2^2}{2}\right) \int_0^{+\infty} y^{i-1} \frac{\partial^i}{\partial s^i} \left\{ |\mathbf{I}_{2m} - 2s\tilde{\mathbf{A}} + 2y\tilde{\mathbf{H}}|^{-1/2} \right. \\ &\quad \left. \times \exp\left(\frac{1}{2} \boldsymbol{\mu}_{\tilde{\mathbf{z}}}^T (\mathbf{I}_{2m} - 2s\tilde{\mathbf{A}} + 2y\tilde{\mathbf{H}})^{-1} \boldsymbol{\mu}_{\tilde{\mathbf{z}}}\right) \right\} \Big|_{s=0} dy \end{aligned} \quad (16)$$

where $\Gamma(\cdot)$ is the Gamma function.

Choice of the distribution of $\varphi(x|\mathcal{H}_k)$ and computation of its moments:

Since the detector $\hat{\Lambda}_{LR}$ takes values between $[0, 1]$, it seems to be logical that its limit in the large dimension regime $\bar{\Lambda}$ takes the same values. In this case, the choice for $\varphi(x|\mathcal{H}_k)$ of a Beta distribution is interesting because its support is also $[0, 1]$ and the expressions of its moments are simple. The non central moments linked to the Beta distribution are iteratively computed:

$$\mu_i^{(\varphi)} |_{\mathcal{H}_k} = \frac{\alpha_k + i - 1}{\alpha_k + \beta_k + i - 1} \mu_{i-1}^{(\varphi)} |_{\mathcal{H}_k}, \quad (17)$$

with $\mu_{1|\mathcal{H}_k}^{(\varphi)} = \frac{\alpha_k}{\alpha_k + \beta_k}$ and where α_k and β_k are the parameters of the Beta distribution:

$$\begin{aligned}\alpha_k &= -\mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})} + \frac{(1 - \mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})})(\mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})})^2}{\mu_{2|\mathcal{H}_k}^{(\bar{\Lambda})} - (\mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})})^2} \\ \beta_k &= -1 - \alpha_k + \frac{(1 - \mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})})\mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})}}{\mu_{2|\mathcal{H}_k}^{(\bar{\Lambda})} - (\mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})})^2}.\end{aligned}\quad (18)$$

Equalization of the moments of $\bar{\Lambda}$ and $\varphi(x|\mathcal{H}_k)$:

This last step allows to compute the elements $\{\nu_i|\mathcal{H}_k\}_{\{0,d\}}$. For this computation, the non central moments of $\bar{\Lambda}$ are used to obtain those of $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$ from the following relation [31]:

$$\mu_{h|\mathcal{H}_k}^{(\bar{\Lambda})} = \sum_{i=0}^d \nu_i|\mathcal{H}_k \mu_{h+i|\mathcal{H}_k}^{(\varphi)}, \quad \forall h \in \llbracket 0, d \rrbracket, \quad (19)$$

where $\mu_{h|\mathcal{H}_k}^{(\bar{\Lambda})}$ (resp. $\mu_{h|\mathcal{H}_k}^{(\varphi)}$), $h \in \llbracket 1, d \rrbracket$ are the non central moments of order h of $\bar{\Lambda}$ (resp. associated to $\varphi(x|\mathcal{H}_k)$) and $\mu_{0|\mathcal{H}_k}^{(\bar{\Lambda})} = \mu_{0|\mathcal{H}_k}^{(\varphi)} = 1$. By setting $\boldsymbol{\nu}_k = [\nu_{0|\mathcal{H}_k} \ \nu_{1|\mathcal{H}_k} \ \dots \ \nu_{d|\mathcal{H}_k}]^T$, $\boldsymbol{\mu}_k^{(\bar{\Lambda})} = [\mu_{0|\mathcal{H}_k}^{(\bar{\Lambda})} \ \mu_{1|\mathcal{H}_k}^{(\bar{\Lambda})} \ \dots \ \mu_{d|\mathcal{H}_k}^{(\bar{\Lambda})}]^T$ and

$$\mathbf{M}_k = \begin{bmatrix} \mu_{0|\mathcal{H}_k}^{(\varphi)} & \mu_{1|\mathcal{H}_k}^{(\varphi)} & \dots & \mu_{d|\mathcal{H}_k}^{(\varphi)} \\ \mu_{1|\mathcal{H}_k}^{(\varphi)} & \mu_{2|\mathcal{H}_k}^{(\varphi)} & \dots & \mu_{d+1|\mathcal{H}_k}^{(\varphi)} \\ \vdots & \vdots & \ddots & \vdots \\ \mu_{d|\mathcal{H}_k}^{(\varphi)} & \mu_{d+1|\mathcal{H}_k}^{(\varphi)} & \dots & \mu_{2d|\mathcal{H}_k}^{(\varphi)} \end{bmatrix}$$

we have

$$\boldsymbol{\nu}_k = \mathbf{M}_k^{-1} \boldsymbol{\mu}_k^{(\bar{\Lambda})}. \quad (20)$$

We give the summary of the different steps in the following algorithm.

Summary of the computation of the approximate distribution

- Choice of initial parameters: d .
- 1. Preliminary calculus: \mathbf{f} , $\bar{\mathbf{A}}$ and $\tilde{\mathbf{H}}$ according to **Proposition III.1**.
- 2. Computation of the d non central moments $\mu_{i|\mathcal{H}_k}^{(\bar{\Lambda})}$ of $\bar{\Lambda}$ thanks to Eq (16).
- 3. Computation of the parameters of the base distribution, α_k and β_k from Eq (18).
- 4. Computation of the $2d$ non central moments $\mu_{i|\mathcal{H}_k}^{(\varphi)}$ from Eqs. (17).
- 5. Computation of the vector of $d + 1$ coefficients $\boldsymbol{\nu}_k$ from Eq (20).
- Computation of the approximate density of probability $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k)$, its cumulative distribution function, etc.

C. Approximate Pfa and Pd of the LR-ANMF

We now characterize the performance of the LR-ANMF detector by deriving the relations Pfa-threshold and Pd-threshold. This derivation is based on the approximate distribution $\tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_1)$ computed under \mathcal{H}_0 and \mathcal{H}_1 . These relations are given in the following proposition.

Proposition III.3 Approximate relations Pfa-threshold and Pd-threshold for the LR-ANMF detector

For a threshold denoted $\bar{\eta} \in [0, 1]$ we have the following theoretical relations Pfa-threshold and Pd-threshold:

$$Pfa \simeq \sum_{i=0}^d \nu_i|\mathcal{H}_0 \frac{\Gamma(\alpha_0 + \beta_0)\Gamma(i + \alpha_0)}{\Gamma(\alpha_0)\Gamma(i + \alpha_0 + \beta_0)} (1 - I_{\bar{\eta}}(i + \alpha_0, \beta_0)) \quad (21)$$

$$Pd \simeq \sum_{i=0}^d \nu_i|\mathcal{H}_1 \frac{\Gamma(\alpha_1 + \beta_1)\Gamma(i + \alpha_1)}{\Gamma(\alpha_1)\Gamma(i + \alpha_1 + \beta_1)} (1 - I_{\bar{\eta}}(i + \alpha_1, \beta_1)), \quad (22)$$

where α_k and β_k are obtained from Eq (18).

Proof: See appendix VI-B. •

Finally, we have the approximate relations Pfa-threshold and Pd-threshold of the LR-ANMF as a function of the radar parameters.

IV. VALIDATIONS

In this section, we validate the different propositions given in this paper. Since these results of theoretical performance are obtained whatever the tested target, we choose a steering vector \mathbf{d} close to the clutter subspace. Therefore, we do not present the results of [19] to obtain the clearest figures because they are not valid in this configuration.

A. Application presentation and parameters

Since the disturbance has to be composed of a LR subspace plus a WGN, the jamming application is chosen. Indeed, the main purpose of this application is to detect a target with an uniform linear antenna composed of $m = 20$ sensors despite the presence of r jammers. Their response \mathbf{c} is composed of $r = 3$ synthetic targets, similar to $\mathbf{d}(\theta)$, of Angles of Arrival (AoA) θ_c equal to -20° , 0° and 20° . Therefore, the rank r of the covariance matrix of the low rank contribution, \mathbf{C} , is equal to 3 and the eigenvalues are chosen equal to $\gamma_1 = 6$, $\gamma_2 = 2$ and $\gamma_3 = 1$. The covariance matrix of the total noise is then defined as:

$$\mathbf{R} = \frac{\alpha_{JNR}}{\text{tr}(\mathbf{\Gamma})} \mathbf{C} + \gamma \mathbf{I}_m, \quad (23)$$

with $\mathbf{\Gamma} = \text{diag}([\gamma_1, \gamma_2, \gamma_3])$, $\gamma = 1$ and where α_{JNR} is used to set the Jamming to Noise Ratio (JNR). The JNR and the SNR (signal \mathbf{d} to noise ratio) are fixed to 10dB and 18dB respectively, $K = r + 1$ which results in $c_m = 5$ and the AoA of the data \mathbf{x} under test is $\theta_d = 20.1^\circ$. We notice that the tested data is very close to the jammer subspace. The wavelength is fixed at 0.667m.

B. Validation of the approximate distribution of the LR-ANMF

In this subsection we validate the asymptotic equivalence of the LR-ANMF test given in the propositions III.1 and III.2. The goal is first to measure the impact of the asymptotic approximation in the large dimension regime on the distribution and next to validate the approximation by polynomial adjustment in particular with a small value of d .

We visualize in Fig. 1 the ECDF (*Empirical Cumulative Distribution Function*) of $\bar{\Lambda}$ of Eq. (13) and the CDF of the approximate distribution, $\tilde{f}_{\bar{\Lambda}}$ of Eq. (15), with a minimal

approximation degree $d = 2$. We also compare them to the ECDFs of the LR-ANMF, $\hat{\Lambda}_{LR}$ of Eq. (6), and the LR-NMF detector (built with the theoretical projector and denoted Λ_{LR}). All these quantities are first evaluated under the \mathcal{H}_0 hypothesis in Fig. 1(a) and under the \mathcal{H}_1 hypothesis in Fig. 1(b). To quantify the distance between two ECDFs or one ECDF and one CDF, we use the Cramer-Von Mises distance defined by:

$$d_{cvm} = E_x [|F_1(x) - F_2(x)|^2], \quad (24)$$

where the expectation is taken over x and $F_1(x)$ and $F_2(x)$ are two ECDF (or CDF). In the following, the Cramer-Von Mises distances will be evaluated with 10^5 trials.

Under \mathcal{H}_0 , d_{cvm} between the ecdfs of the LR-ANMF $\hat{\Lambda}_{LR}$ and that of the LR-NMF Λ_{LR} is equal to 0.06 and d_{cvm} between the ecdfs of the LR-ANMF $\hat{\Lambda}_{LR}$ and that of $\bar{\Lambda}$ is equal to 0.007: we first notice that the approximation in a large dimension regime provides an ECDF much closer to the LR-ANMF ecdf than the LR-NMF one. Moreover, d_{cvm} between the ecdf of $\hat{\Lambda}_{LR}$ and the cdf derived from the distribution $\tilde{f}_{\bar{\Lambda}}$ is equal to around 4.10^{-5} : we conclude that even for a small degree of approximation, the theoretical approximation proposed in this paper is very accurate.

Under \mathcal{H}_1 , d_{cvm} between the ecdfs of the LR-ANMF $\hat{\Lambda}_{LR}$ and that of the LR-NMF Λ_{LR} is equal to 0.224 and d_{cvm} between the ecdfs of the LR-ANMF $\hat{\Lambda}_{LR}$ and that of $\bar{\Lambda}$ is equal to 0.002: as for \mathcal{H}_0 , the approximation in a large dimension regime provides an ECDF closer to the LR-ANMF ecdf than the LR-NMF one. Under \mathcal{H}_1 , d_{cvm} between the ecdf of $\hat{\Lambda}_{LR}$ and the cdf derived from the distribution $\tilde{f}_{\bar{\Lambda}}$ is approximately $3.7 \cdot 10^{-4}$: the approximation $\tilde{f}_{\bar{\Lambda}}$ is less accurate than under \mathcal{H}_0 but the result allows to conclude that the theoretical result proposed in this paper provides a good approximation of the behavior of $\bar{\Lambda}$.

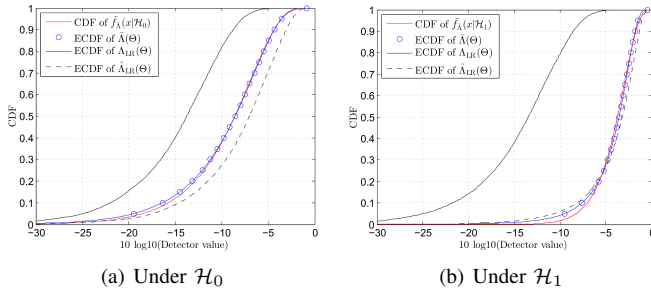


Fig. 1. CDF of the approximate distribution $\tilde{f}_{\bar{\Lambda}}$ with $d = 2$, ECDF of $\bar{\Lambda}$, of the LR-ANMF $\hat{\Lambda}_{LR}$ and of the LR-NMF Λ_{LR} . $m = 20$, $K = r + 1$ and $\theta = 20, 1^\circ$. The JNR is fixed at 10dB. 10^5 trials have been used to evaluate the ECDFs.

C. Validation of the approximate Pfa and Pd of the LR-ANMF

We are here interested in the validation of the Pfa-threshold and Pd-threshold relations given in proposition III.3.

First, Fig.2(a) compares the Pfa-threshold relations of the LR-ANMF, of the LR-NMF and of Eq. (21) derived from the distribution $\tilde{f}_{\bar{\Lambda}}$. Similarly to the previous subsection, we conclude that the Pfa-threshold relation (21) is a correct approximation in particular for small Pfa. Next, Fig.2(b) shows

the Pd-threshold relations of the LR-ANMF, of the LR-NMF and of Eq. (22) derived from the distribution $\tilde{f}_{\bar{\Lambda}}$. Like for the Pfa-threshold, we notice that the relation (22) is a good approximation of the Pd of the LR-ANMF.

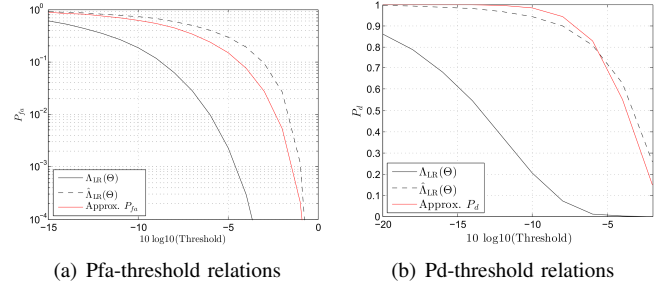


Fig. 2. Pfa-threshold and Pd-threshold relations of LR-NMF Λ_{LR} , of LR-ANMF $\hat{\Lambda}_{LR}$ and of Eq. (22) derived from the distribution $\tilde{f}_{\bar{\Lambda}}$. $m = 20$, $K = r + 1$ and $\theta = 20, 1^\circ$. For the LR-NMF and LR-ANMF, 10^5 trials have been used to evaluate the Pfa and the Pd. The JNR and the SNR are fixed at 10dB and 18dB.

In Fig. 3, we show the Pd as a function of the SNR for a fixed $Pfa = 10^{-2}$. We notice that the LR-NMF is not able to predict the performance of the LR-ANMF. Even if we have an error between the theoretical result and the Monte Carlo simulations, we conclude that it provides a correct approximation of the detection performance when the target under test is near the LR subspace.

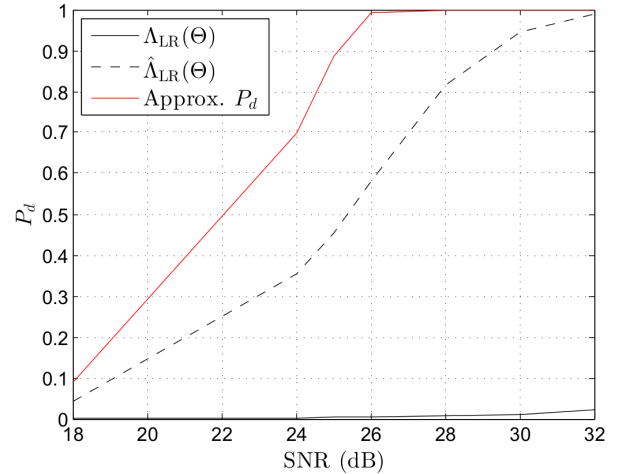


Fig. 3. Pd-SNR relations of LR-NMF Λ_{LR} , of LR-ANMF $\hat{\Lambda}_{LR}$ and of Eq. (22) derived from the distribution $\tilde{f}_{\bar{\Lambda}}$. $m = 20$, $K = r + 1$ and $\theta = 20, 1^\circ$. For the LR-NMF and LR-ANMF, 10^5 trials have been used to evaluate the Pd. The JNR is fixed at 10dB and the Pfa is fixed at 10^{-3} .

Finally, the table I presents the average calculation time for one point of the Pfa-threshold relation with 10^4 iterations when the moments are already calculated³. We notice that the approximation (21) given in this paper allows to strongly reduce the computational cost compared to a classic Monte Carlo simulation. But, this equation requires computing the moment generator function of Eq. (16). If this step has a computational cost of one hour under \mathcal{H}_0 , it becomes prohibitive under \mathcal{H}_1

³The computer used has an Intel(R) Core(TM) i7-3667U CPU @ 2,00GHz processor and RAM memory of 8Go.

	$K = r + 1$	$K = 2r$
With $\hat{\Lambda}_{\text{LR}}$	3,0342 s	3,0188 s
With the theoretical relation (Eq.(21))	0,0252 s	0,0240 s

TABLE I

AVERAGE CALCULATION TIME FOR ONE POINT OF THE PFA-THRESHOLD RELATION WITH 10^4 ITERATIONS AND WHEN THE PFA IS SET AT 10^{-4} .

(around 2 days). Therefore, the approximation (21) allows to obtain a threshold for a given Pfa with a small computation time. On the other hand, Monte Carlo trials are more suitable to evaluate the Pd.

V. CONCLUSION

In this paper, we derived the approximate Pfa-threshold and Pd-threshold relations of the LR-ANMF in a large dimension regime. Compared to the previous work [19], the tested target is assumed to be close to the LR-subspace. In this configuration, we showed that the LR-ANMF lost its CFAR property and its performance depended on the distance between the target and the LR subspace. More generally, this paper illustrated that the tools of RMT are useful for deriving an approximation of a complex statistic form like the LR-ANMF under conditions where the classical tools of asymptotic analysis are unable to provide an interesting result.

We validated this theoretical result in a jamming application where the data under test is close to the LR subspace. Finally, we illustrated that the computational cost is low under \mathcal{H}_0 , which allows a threshold to be easily obtained for a chosen Pfa without the use of Monte Carlo simulations. Unfortunately the computational cost under \mathcal{H}_1 is too high for a practical use.

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VI. APPENDIX

A. Proof of proposition III.1

From the convergence results for the QFs given in subsection II-C2 (under hypotheses (AS1-AS4)) and the *continuous mapping theorem* we directly have:

$$\hat{\Lambda}_{\text{LR}} \xrightarrow[m/K \rightarrow c < \infty]{p.s.} \bar{\Lambda} = \frac{|\mathbf{d}^H \bar{\Pi}_c^\perp \mathbf{x}|^2}{(\mathbf{d}^H \bar{\Pi}_c^\perp \mathbf{d})(\mathbf{x}^H \bar{\Pi}_c^\perp \mathbf{x})}. \quad (25)$$

We notice that the LR-ANMF test is not consistent in a large dimension regime.

Next, we transform the previous equation in a ratio of two QFs, which are functions of complex Gaussian random variables:

$$\bar{\Lambda} = \frac{|\mathbf{d}^H \mathbf{U} \Psi \mathbf{U}^H \mathbf{x}|^2}{(\mathbf{d}^H \mathbf{U} \Psi \mathbf{U}^H \mathbf{d})(\mathbf{x}^H \mathbf{U} \Psi \mathbf{U}^H \mathbf{x})} = \frac{|\mathbf{e}^H \mathbf{y}|^2}{(\mathbf{e}^H \mathbf{e})(\mathbf{y}^H \mathbf{y})}, \quad (26)$$

with $\mathbf{e} = \Psi^{1/2} \mathbf{U}^H \mathbf{d}$, $\mathbf{y} = \Psi^{1/2} \mathbf{U}^H \mathbf{x} \sim \mathcal{CN}(\boldsymbol{\mu}_y = \Psi^{1/2} \mathbf{U}^H \boldsymbol{\mu}_x, \mathbf{H})$, $\mathbf{H} = \Lambda \Psi$ and $\Lambda = \text{diag}([\lambda_1, \dots, \lambda_m])$.

Therefore,

$$\begin{aligned} \bar{\Lambda} &= \frac{|\mathbf{e}^H \mathbf{H}^{1/2} \mathbf{z}|^2}{(\mathbf{e}^H \mathbf{e})(\mathbf{z}^H \mathbf{H} \mathbf{z})} \\ &= \frac{\mathbf{z}^H \mathbf{H}^{1/2} \mathbf{e} \mathbf{e}^H \mathbf{H}^{1/2} \mathbf{z}}{\mathbf{z}^H \mathbf{H}^{1/2} \mathbf{f} \mathbf{f}^H \mathbf{H}^{1/2} \mathbf{z}} \\ &= \frac{\mathbf{z}^H \mathbf{H} \mathbf{z}}{\mathbf{z}^H \mathbf{A} \mathbf{z}}, \end{aligned}$$

with $\mathbf{z} = \mathbf{H}^{-1/2} \mathbf{y} \sim \mathcal{CN}(\boldsymbol{\mu}_z = \Lambda^{-1/2} \mathbf{U}^H \boldsymbol{\mu}_x, \mathbf{I}_m)$, $\mathbf{f} = \mathbf{e}/\|\mathbf{e}\|_2$ and $\mathbf{A} = \mathbf{H}^{1/2} \mathbf{f} \mathbf{f}^H \mathbf{H}^{1/2}$.

Finally, it is easy to transform the previous relation which is in a complex form into an equivalent real form:

$$\bar{\Lambda} = \frac{\mathbf{z}^H \mathbf{A} \mathbf{z}}{\mathbf{z}^H \mathbf{H} \mathbf{z}} = \frac{\tilde{\mathbf{z}}^T \tilde{\mathbf{A}} \tilde{\mathbf{z}}}{\tilde{\mathbf{z}}^T \tilde{\mathbf{H}} \tilde{\mathbf{z}}} \quad (27)$$

with

$$\begin{cases} \tilde{\mathbf{z}} = \sqrt{2}[\Re(\mathbf{z})^T, \Im(\mathbf{z})^T]^T \in \mathbb{R}^{2m \times 1} \\ \tilde{\mathbf{A}} = \begin{bmatrix} \Re(\mathbf{A}) & -\Im(\mathbf{A}) \\ \Im(\mathbf{A}) & \Re(\mathbf{A}) \end{bmatrix} \in \mathbb{R}^{2m \times 2m} \\ \tilde{\mathbf{H}} = \begin{bmatrix} \mathbf{H} & \mathbb{O}_{2m} \\ \mathbb{O}_{2m} & \mathbf{H} \end{bmatrix} \in \mathbb{R}^{2m \times 2m} \end{cases}, \quad (28)$$

with $\tilde{\mathbf{z}} \sim \mathcal{N}(\boldsymbol{\mu}_{\tilde{\mathbf{z}}} = \sqrt{2}[\Re(\boldsymbol{\mu}_z)^T, \Im(\boldsymbol{\mu}_z)^T]^T, \mathbf{I}_{2m})$.

B. Proof of proposition II.1

We have by definition of the Pfa and the Pd:

$$\begin{aligned} P_{fa}, P_d &\simeq \int_{\bar{\eta}}^{+\infty} \tilde{f}_{\bar{\Lambda}}(x|\mathcal{H}_k) dx \\ &= \int_{\bar{\eta}}^{+\infty} \varphi(x|\mathcal{H}_k) \sum_{i=0}^d \nu_i|\mathcal{H}_k x^i dx, \quad (29) \\ &= \sum_{i=0}^d \nu_i|\mathcal{H}_k P_i(\bar{\eta}|\mathcal{H}_k) dx \end{aligned}$$

with $P_i(\bar{\eta}|\mathcal{H}_k) = \int_{\bar{\eta}}^{+\infty} x^i \varphi(x|\mathcal{H}_k) dx$ and $k \in \{0, 1\}$. Since $\varphi(x|\mathcal{H}_k)$ is a Beta distribution, then

$$\begin{aligned} P_i(\bar{\eta}|\mathcal{H}_k) &= \int_{\bar{\eta}}^{+\infty} x^i \frac{\Gamma(\alpha_k + \beta_k)}{\Gamma(\alpha_k)\Gamma(\beta_k)} x^{\alpha_k - 1} (1-x)^{\beta_k - 1} \mathbb{1}_{[0,1]} dx \\ &= \frac{\Gamma(\alpha_k + \beta_k)}{\Gamma(\alpha_k)\Gamma(\beta_k)} \int_{\bar{\eta}}^{+\infty} x^{i+\alpha_k - 1} (1-x)^{\beta_k - 1} \mathbb{1}_{[0,1]} dx \\ &= \frac{\Gamma(\alpha_k + \beta_k)}{\Gamma(\alpha_k)\Gamma(\beta_k)} \int_{\bar{\eta}}^{+\infty} \frac{\Gamma(i + \alpha_k)\Gamma(\beta_k)}{\Gamma(i + \alpha_k + \beta_k)} \mathcal{B}(i + \alpha_k, \beta_k) dx \\ &= \frac{\Gamma(\alpha_k)\Gamma(i + \alpha_k)}{\Gamma(\alpha_k + \beta_k)\Gamma(i + \alpha_k)} \int_{\bar{\eta}}^{+\infty} \mathcal{B}(i + \alpha_k, \beta_k) dx \\ &= \frac{\Gamma(\alpha_k)\Gamma(i + \alpha_k + \beta_k)}{\Gamma(\alpha_k + \beta_k)\Gamma(i + \alpha_k)} \left(1 - \int_{-\infty}^{\bar{\eta}} \mathcal{B}(i + \alpha_k, \beta_k) dx\right) \\ &= \frac{\Gamma(\alpha_k)\Gamma(i + \alpha_k + \beta_k)}{\Gamma(\alpha_k)\Gamma(i + \alpha_k)} \left(1 - I_{\bar{\eta}}(i + \alpha_k, \beta_k)\right) \end{aligned} \quad (30)$$

where $I_{\bar{\eta}}$ (resp. \mathcal{B}) is the regularized incomplete beta function (resp. the beta function). Therefore, the proof is finished by using the properties of $I_{\bar{\eta}}$.

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