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## HYDRODYNAMIC LIMIT FOR A DISORDERED HARMONIC CHAIN

CÉDRIC BERNARDIN, FRANÇOIS HUVENEERS, AND STEFANO OLLA

ABSTRACT. We consider a one-dimensional unpinned chain of harmonic oscillators with random masses. We prove that after hyperbolic scaling of space and time the distributions of the elongation, momentum and energy converge to the solution of the Euler equations. Anderson localization decouples the mechanical modes from the thermal modes, allowing the closure of the energy conservation equation even out of thermal equilibrium. This example shows that the derivation of Euler equations rests primarily on scales separation and not on ergodicity.

## 1. INTRODUCTION

In this paper, we consider a one-dimensional chain of coupled harmonic oscillators with masses  $m_x$  and Hamiltonian

$$H = \sum_{x \in \mathbb{Z}} \left( \frac{p_x^2}{2m_x} + g \frac{(q_{x+1} - q_x)^2}{2} \right).$$

By changing units, one can assume that the stiffness coefficient  $g$  is equal to 1. The dynamics is governed by Hamilton's equations:

$$m_x \dot{q}_x = p_x, \quad \dot{p}_x = (\Delta q)_x,$$

where we have used the notation  $\Delta = \nabla_- \nabla_+ = \nabla_+ \nabla_-$  for the discrete Laplacian, with  $(\nabla_+ f)_x = f_{x+1} - f_x$  and  $(\nabla_- f)_x = f_x - f_{x-1}$ .

This system was first analyzed in finite volume when all masses  $m_x$  are equal. Putting the chain in a non-equilibrium stationary state (NESS) between two heat reservoirs at different temperatures, it was found in [19] that the energy current does not decay with the size of the system, indicating that energy propagates ballistically. The situation changes if the masses are taken to be i.i.d. random variables. This case was first investigated in [20, 9] and subsequently studied in [21, 10, 3]. As it turns out, the disordered harmonic chain is an Anderson insulator in disguise [4]. However, as a consequence of the conservation of momentum, the ground state of the operator  $M^{-1} \Delta$ , featuring in Newton's equation  $\ddot{q} = M^{-1} \Delta q$  with  $M$  the diagonal matrix of the masses, is a "symmetry protected mode" [13], implying a divergent localization length in the lower edge of the spectrum. This leads to a rich and unexpected phenomenology. In particular, if the chain is again in a NESS, the scaling of the energy current with the system size happens to depend on boundary conditions and spectral factors of the reservoirs [10]. This finding reveals also the complete lack of local thermal equilibrium, that results eventually from integrability (see Section 5.2).

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The harmonic chain has three “obvious” conserved quantities: the total energy  $H$ , the total momentum  $P = \sum_x p_x$  and the total stretch or elongation  $R = \sum_x r_x$  with  $r_x = (\nabla_+ q)_x$ . This gives rise to the following microscopic conservation laws:

$$\dot{r}_x = \frac{p_{x+1}}{m_{x+1}} - \frac{p_x}{m_x}, \quad \dot{p}_x = r_x - r_{x-1}, \quad \dot{e}_x = \frac{r_x p_{x+1}}{m_{x+1}} - \frac{r_{x-1} p_x}{m_x}$$

with  $e_x = \frac{1}{2} \left( \frac{p_x^2}{m_x} + r_x^2 \right)$ . After a hyperbolic rescaling of space and time, we ask in this paper whether the empirical densities of these conserved quantities converge to the densities  $\mathbf{r}, \mathbf{p}$  and  $\mathbf{e}$  governed by the macroscopic laws

$$\partial_t \mathbf{r}(y, t) = \frac{1}{\bar{m}} \partial_y \mathbf{p}(y, t), \quad \partial_t \mathbf{p}(y, t) = \partial_y \mathbf{r}(y, t), \quad \partial_t \mathbf{e}(y, t) = \frac{1}{\bar{m}} \partial_y (\mathbf{r}(y, t) \mathbf{p}(y, t)),$$

corresponding to Euler equations in Lagrangian coordinates, with  $\bar{m}$  the average mass.

Instances of rigorous derivation of Euler equations in the smooth regime rest on the ergodicity of the microscopic dynamics. In [8, 18, 16], the Hamiltonian dynamics is perturbed by some stochastic noise acting in such a way that conserved quantities are not destroyed but that the ergodicity of the dynamics can be established rigorously.

The dynamics considered here is purely Hamiltonian, non-ergodic, and possesses actually a full set of invariant quantities (see Section 5.2). In the clean case, i.e. when the masses are all equal, we show in Section 1.1 that Euler equations hold if and only if the temperature profile is constant. Instead, in Section 1.2, we argue that Euler equations hold even out of thermal equilibrium if there is disorder on the masses. We briefly discuss the fate of other conserved quantities in Section 1.3. Theorem 1 in Section 2 constitutes the main result of our paper: We show the convergence to Euler equations for the disordered harmonic chain, almost surely with respect to the masses and on average with respect to an initial local Gibbs state. The rest of the paper is devoted to the proof of this theorem.

**1.1. Clean harmonic chain.** Let us assume that all masses  $m_x$  are equal, say  $m_x = 1$  for simplicity. In this case, the equations of motion read

$$\dot{q}_x = p_x, \quad \dot{p}_x = \Delta q_x. \quad (1.1)$$

Let us first consider the thermal equilibrium case: Assume that the initial configuration of the chain is random and distributed according to a Gaussian law  $\mu_0$  with covariance matrix

$$\langle (\nabla_+ q)_x; (\nabla_+ q)_y \rangle = \langle p_x; p_y \rangle = \beta^{-1} \delta_{x,y}, \quad \langle q_x; p_y \rangle = 0, \quad (1.2)$$

for some inverse temperature  $\beta$ . It is easy to prove that at time  $t > 0$ , the distribution  $\mu_t$  in the phase space is still given by a Gaussian law with the same covariance matrix. To see this, just use Fourier transforms to diagonalize the dynamics:

$$\hat{q}(k, t) = \sum_{x \in \mathbb{Z}} e^{i2\pi kx} q_x(t), \quad \hat{p}(k, t) = \sum_{x \in \mathbb{Z}} e^{i2\pi kx} p_x(t), \quad (1.3)$$

and define the wave function

$$\hat{\psi}(k, t) = \omega(k) \hat{q}(k, t) + i \hat{p}(k, t) \quad (1.4)$$

where  $\omega(k) = |2 \sin(\pi k)|$  is the dispersion relation. Then, the explicit solution of (1.1) is given by

$$\hat{\psi}(k, t) = e^{-i\omega(k)t} \hat{\psi}(k, 0). \quad (1.5)$$

The correlations (1.2) imply that

$$\langle \hat{\psi}(k, 0)^*; \hat{\psi}(k', 0) \rangle = 2\beta^{-1} \delta(k - k'), \quad \langle \hat{\psi}(k, 0); \hat{\psi}(k', 0) \rangle = 0. \quad (1.6)$$

Consequently

$$\begin{aligned} \langle \hat{\psi}(k, t)^*; \hat{\psi}(k', t) \rangle &= e^{i(\omega(k) - \omega(k'))t} \langle \hat{\psi}(k, 0)^*; \hat{\psi}(k', 0) \rangle = 2\beta^{-1} \delta(k - k'), \\ \langle \hat{\psi}(k, t); \hat{\psi}(k', t) \rangle &= e^{-i(\omega(k) + \omega(k'))t} \langle \hat{\psi}(k, 0); \hat{\psi}(k', 0) \rangle = 0. \end{aligned} \quad (1.7)$$

From (1.6) and (1.7), we deduce that the covariances (1.2) are the same at any time  $t$ :

$$\langle (\nabla_+ q)_x(t); (\nabla_+ q)_y(t) \rangle = \langle p_x(t); p_y(t) \rangle = \beta^{-1} \delta_{x,y}, \quad \langle q_x(t); p_y(t) \rangle = 0, \quad (1.8)$$

which implies that the Gaussian distribution  $\mu_t$  differs from  $\mu_0$  only by the averages  $\bar{r}_x(t) = \langle r_x(t) \rangle = \langle (\nabla_+ q)_x(t) \rangle$  and  $\bar{p}_x(t) = \langle p_x(t) \rangle$  that, by linearity of the dynamics, evolve following the same equation (1.1).

Assume now that the initial averages of the momentum  $p_x$  and stretch  $r_x = (\nabla_+ q)_x$  are slowly varying on a macroscopic scale, i.e. given some initial macroscopic profiles  $p(y), r(y)$  where  $y \in \mathbb{R}$ , we have

$$\bar{r}_x(0) = r(x/N), \quad \bar{p}_x(0) = p(x/N). \quad (1.9)$$

Let  $\widehat{p}(\xi)$  and  $\widehat{r}(\xi)$  be the Fourier transforms (in  $\mathbb{R}$ ) of  $p(y)$  and  $r(y)$ . Then, as  $N \rightarrow \infty$ ,  $\frac{1}{N} \widehat{p}(\frac{\xi}{N}) \rightarrow \widehat{p}(\xi)$  and  $\frac{1}{N} \widehat{r}(\frac{\xi}{N}) \rightarrow \widehat{r}(\xi)$ . After a straightforward analysis we have that

$$\frac{1}{N} \widehat{p}(\frac{\xi}{N}, Nt) \rightarrow \widehat{p}(\xi, t), \quad \frac{1}{N} \widehat{r}(\frac{\xi}{N}, Nt) \rightarrow \widehat{r}(\xi, t) \quad (1.10)$$

where

$$\partial_t \widehat{r}(\xi, t) = -\frac{i\xi}{2\pi} \widehat{p}(\xi, t), \quad \partial_t \widehat{p}(\xi, t) = -\frac{i\xi}{2\pi} \widehat{r}(\xi, t). \quad (1.11)$$

Consequently  $r_{[Ny]}(Nt)$  and  $p_{[Ny]}(Nt)$  converge weakly to the solution of the linear wave equation

$$\partial_t \mathbf{r}(y, t) = \partial_y \mathbf{p}(y, t), \quad \partial_t \mathbf{p}(y, t) = \partial_y \mathbf{r}(y, t). \quad (1.12)$$

Let us now consider the energy per particle  $e_x = \frac{1}{2} (p_x^2 + r_x^2)$ . Its average under the distribution  $\mu_t$  is  $\langle e_x(t) \rangle = \beta^{-1} + \frac{1}{2} (\bar{p}_x^2(t) + \bar{r}_x^2(t))$  since by (1.8), the variance of  $p_x$  and  $r_x$ , i.e. the temperature, remains constant in time. In the limit  $N \rightarrow \infty$  we have

$$\langle e_{[Ny]}(Nt) \rangle \rightarrow \mathbf{e}(y, t) = \beta^{-1} + \frac{1}{2} (\mathbf{p}^2(y, t) + \mathbf{r}^2(y, t)),$$

i.e. it solves the equation

$$\partial_t \mathbf{e}(y, t) = \partial_y (\mathbf{p}(y, t) \mathbf{r}(y, t)). \quad (1.13)$$

We recognize that (1.12 - 1.13) are the Euler equations. The above is the simplest example of propagation of local equilibrium and hydrodynamic limit in hyperbolic scaling: in a harmonic chain in thermal equilibrium at temperature  $\beta^{-1}$ , and the mechanical modes not in equilibrium, we will have a persistence of the thermal equilibrium at any time  $t$ , while the mechanical modes evolve independently from the thermal mode following the linear wave equation.

Notice that the argument above does not require the distribution  $\mu_0$  to be the thermal equilibrium measure defined by (1.2), and that it holds for any measure  $\mu_0$  with translation invariant covariance given by

$$\langle \nabla q_x; \nabla q_y \rangle = \langle p_x; p_y \rangle = C(x - y), \quad \langle q_x; p_y \rangle = 0 \quad (1.14)$$

for a positive definite function  $C(x)$ . The only difference is then that in (1.6-1.7) the term  $\beta^{-1}$  has to be replaced by the Fourier transform  $\widehat{C}(k)$ . Actually, the measure  $\mu_t$  is

not even a local equilibrium state [7], underlining that the validity of Euler equations in this example does not require the propagation of local equilibrium.

The above argument rests on the translation invariance of the distribution of the thermal modes and fails if it is space inhomogeneous, for example if the starting distribution is given by a local Gibbs state with a slowly varying temperature  $\beta_{N,x}^{-1} = \beta^{-1}(x/N)$ , i.e. a Gaussian measure with covariances

$$\langle \nabla q_x; \nabla q_y \rangle = \langle p_x; p_y \rangle = \beta_{N,x}^{-1} \delta_{x,y}, \quad \langle q_x; p_y \rangle = 0. \quad (1.15)$$

In this case, even though the wave equation (1.12) still holds, generally the energy equation (1.13) is not valid. In fact the energies of each mode  $k$  evolves autonomously, as we can see studying the limit evolution of the Wigner distribution defined by

$$\begin{aligned} \widehat{W}_N(\xi, k, t) &:= \frac{2}{N} \langle \widehat{\psi}^* \left( k - \frac{\xi}{2N}, Nt \right) \widehat{\psi} \left( k + \frac{\xi}{2N}, Nt \right) \rangle \\ W_N(y, k, t) &:= \int e^{-i2\pi\xi y} \widehat{W}_N(\xi, k, t) d\xi, \end{aligned} \quad (1.16)$$

(the above definitions should be understood as distributions on  $\mathbb{R} \times \Pi$  with  $\Pi = \mathbb{R} \setminus \mathbb{Z}$ ).

In the limit as  $N \rightarrow \infty$  the Wigner distribution converge to a positive distribution with an absolutely continuous part, the local distribution of the thermal modes, and a singular part concentrated on  $k = 0$ , the mechanical modes:

$$\lim_{N \rightarrow \infty} \widehat{W}_N(\xi, k, t) = \widehat{W}_{th}(\xi, k, t) + \widehat{W}_m(\xi, t) \delta_0(dk) \quad (1.17)$$

The mechanical part  $\widehat{W}_m(\xi, t)$  is the Fourier transform of  $\frac{1}{2}(\mathbf{p}^2(y, t) + \mathbf{r}^2(y, t))$ .

A straightforward calculation gives for the thermal part (see [11] or [5] for a rigorous argument):

$$\widehat{W}_{th}(\xi, k, t) = e^{-i\omega'(k)\xi t} \widehat{W}_{th}(\xi, k, 0). \quad (1.18)$$

This implies that the inverse Fourier transform  $W_{th}(y, k, t)$  satisfies the transport equation

$$\partial_t W_{th}(y, k, t) + \frac{\omega'(k)}{2\pi} \partial_y W_{th}(y, k, t) = 0. \quad (1.19)$$

It also follow that

$$\int W_{th}(y, k, t) dk = \tilde{\mathbf{e}}(y, t) \quad (1.20)$$

where  $\tilde{\mathbf{e}}(y, t)$  is the limit profile of thermal energy (or temperature) defined as

$$\frac{1}{2} (\langle r_{[Ny]}(Nt); r_{[Ny]}(Nt) \rangle + \langle p_{[Ny]}(Nt); p_{[Ny]}(Nt) \rangle) \rightarrow \tilde{\mathbf{e}}(y, t). \quad (1.21)$$

Consequently the thermal energy  $\tilde{\mathbf{e}}(y, t)$  evolves non autonomously following the equation

$$\partial_t \tilde{\mathbf{e}}(y, t) + \partial_y J(y, t) = 0, \quad J(y, t) = \int \omega'(k) W_{th}(y, k, t) dk. \quad (1.22)$$

We say that the system is in *local equilibrium* if  $W_{th}(y, k) = \beta^{-1}(y)$  constant in  $k$ . This correspond to the fact that Gibbs measure gives uniform distribution on the modes. Starting in thermal equilibrium means  $W_{th}(y, k, 0) = \beta^{-1}$  and trivially  $W_{th}(y, k, t) = \beta^{-1}$  for any  $t > 0$ . But starting with local equilibrium, i.e.  $W(y, k, 0) = \beta^{-1}(y)$  constant in  $k$ , we have a non autonomous evolution of  $\tilde{\mathbf{e}}(y, t)$ .

**1.2. Disordered harmonic chain.** The situation so far can be summarized as follows. By linearity, the variables  $r_x$  and  $p_x$  admit a macroscopic limit described by (1.12) independently of the initial temperature profile. The macroscopic equation (1.13) predicts that the evolution of the energy is purely mechanical and that the temperature does not evolve with time. As it turns out, the evolution of the mechanical energy is correctly described by Euler equation (see the term  $\mathcal{A}_N(t)$  in our decomposition (4.2) below), but thermal fluctuations do in general evolve with time as well, except if the temperature profile is initially flat.

This picture gets strongly modified if the masses are taken to be random. On the one hand, deriving the macroscopic evolution of the fields  $r_x$  and  $p_x$  becomes less obvious because some homogenization over the masses is required. This difficulty can be solved by the elegant method of the “corrected empirical measure”, see [12, 14, 6] (though we will actually solve it another way). On the other hand, and this is the main point in considering random masses, the evolution of the energy  $e_x$  is now much better approximated by Euler equation. Indeed, at a microscopic level, all thermal fluctuations are frozen thanks to Anderson localization and the evolution of the energy becomes purely mechanical.

To understand this a little bit better, it is good to realize how the disorder modifies the nature of the eigenmodes  $(\psi^k)_{1 \leq k \leq N}$  of the operator  $M^{-1}\Delta$  for a finite chain of size  $N$ . As a consequence of Anderson localization [4], all modes at positive energy are spatially localized. However the localization length  $\xi_k$  diverges as one approaches the ground state:

$$\xi_k^{-1} \sim \omega_k^2 \sim \left(\frac{k}{N}\right)^2,$$

so that only the modes with  $k \gtrsim \sqrt{N}$  are actually localized, while the modes  $k \lesssim \sqrt{N}$  remain comparable to the modes of the clean chain [21, 3]. By imposing a smooth initial profile  $r, p$ , the initial local Gibbs state attributes a weight of order 1 to a few first modes above the ground state, and a weight of order 1 to all other modes together. The first ones are responsible for the transport of mechanical energy; all modes with  $k \gg \sqrt{N}$  are localized and do not transport any thermal energy; all modes with  $1 \ll k \leq o(N)$  have a vanishing weight in the thermodynamic limit and can be neglected in the analysis.

Finally, we would like to mention that, while the disorder considered here and the stochastic velocity exchange noise considered in [8, 16] act in an obviously very different way, e.g. the disorder preserves integrability while the stochastic noise makes the dynamic ergodic, they do produce the same effects in some respect. Indeed the noise has only a very slow effect on  $(r_x, p_x)$  profiles that vary smoothly with  $x$  at the macroscopic scale (because nearly identical momenta are exchanged). This bares some similarity with the fact that the disorder has very little influence on the low modes of the disordered chain. Instead the noise provides an active hopping mechanism among the high modes. However, this produces a sub-ballistic spreading of the energy [15, 16], that is not visible in the hyperbolic scaling. Thus, the noise plays here as well a role analogous to the disorder by freezing the fluctuations (on the scales where we follow the dynamics).

**1.3. Other conserved quantities.** Before moving on, let us briefly comment on the issue of the other conserved quantities of the system. These can also be written as a

sum of local terms and lead thus to additional conservation laws. For example,

$$I = \sum_x d_x = \frac{1}{2} \sum_x \left( \frac{(r_x - r_{x-1})^2}{m_x} + \left( \frac{p_{x+1}}{m_{x+1}} - \frac{p_x}{m_x} \right)^2 \right)$$

is conserved (see Sections 3.1 and 5.2) and leads to the microscopic conservation law

$$\dot{d}_x = \left( \frac{p_{x+1}}{m_{x+1}} - \frac{p_x}{m_x} \right) \frac{r_{x+1} - r_x}{m_{x+1}} - \left( \frac{p_x}{m_x} - \frac{p_{x-1}}{m_{x-1}} \right) \frac{r_x - r_{x-1}}{m_x}.$$

It is thus natural to ask whether this relation generates also some macroscopic law. In the cases where we can derive the macroscopic evolution equation (1.13) for the energy, it is easy to argue that the corresponding macroscopic density  $\mathbf{d}(y, t)$  does not evolve with time in the hyperbolic scaling. Indeed, we can decompose  $d_x$  as the sum of a mechanical and a thermal contribution, as we do in (4.2) below for the energy. In this case, contrary to what happens for the energy, the mechanical contribution vanishes in the thermodynamic limit since  $d_x$  depends on  $r$  and  $p$  only through their gradients, while the contribution from the thermal modes does not evolve with time, for the same reasons as it does not for the energy.

All the other conserved quantities in this model that can be written as a sum of local terms are obtained by taking further gradients in the variables  $r$  and  $p$  (see Section 5.2), and have thus no evolution either in the hyperbolic scaling.

## 2. MODEL AND RESULTS

We define the model studied in this paper and we state our main result. For technical reasons, it is easier to work on a finite system of size  $N$  and then let  $N \rightarrow \infty$ .

**2.1. Hamiltonian model.** The Hamiltonian  $H$  on  $\mathbb{R}^{2N}$  is defined by

$$H(q, p) = \frac{1}{2} \sum_{x=1}^N \left( \frac{p_x^2}{m_x} + ((\nabla_+ q)_x)^2 \right)$$

with free boundary conditions:  $q_0 = q_1$  and  $q_{N+1} = q_N$ . The masses  $(m_x)_{1 \leq x \leq N}$  are i.i.d. random variables. In order to avoid any technical difficulty in exploiting known results from the Anderson localization literature, we assume that the law of  $m_x$  admits a smooth density compactly supported in  $[m_-, m_+]$  with  $m_- > 0$ .

The equations of motion read  $M\dot{q} = p$  and  $\dot{p} = \Delta q$  where  $M$  is the square diagonal matrix of size  $N$  with entries defined by  $M_{x,y} = \delta(x-y)m_x$  ( $\delta(z)$  is defined by 1 for  $z = 0$  and 0 otherwise). It is more convenient to express the equations of motion in terms of the displacement variables

$$r_x = (\nabla_+ q)_x \quad (1 \leq x \leq N-1). \quad (2.1)$$

The equations of motion become

$$\dot{r}_x = (\nabla_+ M^{-1} p)_x \quad (1 \leq x \leq N-1), \quad \dot{p}_x = (\nabla_- r)_x \quad (1 \leq x \leq N)$$

where we use fixed boundary conditions for  $r$  in the second equation:  $r_0 = r_N = 0$ .

**2.2. Gibbs and locally Gibbs states.** We consider three locally conserved quantities in the bulk:

$$H = \sum_{x=1}^N e_x = \sum_{x=1}^N \left( \frac{p_x^2}{2m_x} + \frac{r_x^2}{2} \right), \quad P = \sum_{x=1}^N p_x, \quad R = \sum_{x=1}^{N-1} r_x.$$

The energy  $H$  and the momentum  $P$  are actually truly conserved, but the conservation of  $R$  is broken at the boundary:  $\dot{R} = m_N^{-1}p_N - m_1^{-1}p_1$ .

The Gibbs states are characterized by three parameters:  $\beta > 0$  and  $\mathbf{p}, \mathbf{r} \in \mathbb{R}$ . Its probability density writes

$$\rho_G(r, p) = \frac{1}{Z_G} \exp \left\{ -\frac{\beta}{2} \sum_{x=1}^N m_x \left( \frac{p_x}{m_x} - \frac{\mathbf{p}}{\bar{m}} \right)^2 - \frac{\beta}{2} \sum_{x=1}^{N-1} (r_x - \mathbf{r})^2 \right\}.$$

where  $\bar{m}$  denotes the mean mass and  $Z_G := Z_G(\beta, \mathbf{p}, \mathbf{r})$  is a normalizing constant. Local Gibbs states are obtained by replacing the constant parameters  $\beta, \mathbf{p}, \mathbf{r}$  by functions

$$\beta, \mathbf{p}, \mathbf{r} : [0, 1] \rightarrow \mathbb{R},$$

with  $\beta(x) > 0$  for all  $x \in [0, 1]$ , and by considering the measure with density

$$\rho_{\text{loc}}(r, p) = \frac{1}{Z_{\text{loc}}} \exp \left\{ -\frac{1}{2} \sum_{x=1}^N \beta(x/N) m_x \left( \frac{p_x}{m_x} - \frac{\mathbf{p}(x/N)}{\bar{m}} \right)^2 - \frac{1}{2} \sum_{x=1}^{N-1} \beta(x/N) (r_x - \mathbf{r}(x/N))^2 \right\} \quad (2.2)$$

where  $Z_{\text{loc}} := Z_{\text{loc}}(\beta, \mathbf{p}, \mathbf{r})$  is a normalizing constant. We impose the following regularity conditions on  $\beta, \mathbf{p}, \mathbf{r}$ :

$$\beta \in \mathcal{C}^0([0, 1]), \quad \mathbf{r} \in \mathcal{C}^1([0, 1]) \text{ with } r(0) = r(1) = 0, \quad \mathbf{p} \in \mathcal{C}^1([0, 1]). \quad (2.3)$$

We take such a local Gibbs state as initial state. Below, we denote the expectation with respect to it by  $\langle \cdot \rangle$ :

$$\langle F \rangle = \int F(r, p) \rho_{\text{loc}}(r, p) dr dp.$$

Instead, expectation (resp. probability) with respect to the masses is denoted by  $E$  (resp.  $P$ ).

**2.3. Evolution of the locally conserved quantities.** Let us fix some maximal time  $T > 0$ . Let us define the fields  $\mathcal{R}, \mathcal{P}$  and  $\mathcal{E}$  acting on functions  $f \in \mathcal{C}^0([0, 1])$  as

$$\begin{aligned} \mathcal{R}(f, t) &= \int_0^1 \mathbf{r}(y, t) f(y) dy, & \mathcal{P}(f, t) &= \int_0^1 \mathbf{p}(y, t) f(y) dy, \\ \mathcal{E}(f, t) &= \int_0^1 \mathbf{e}(y, t) f(y) dy. \end{aligned} \quad (2.4)$$

for all  $t \in [0, T]$ . The kernels  $\mathbf{r}, \mathbf{p}$  and  $\mathbf{e}$  are defined as follows. First, at  $t = 0$ , we impose

$$\mathbf{r}(y, 0) = r(y), \quad \mathbf{p}(y, 0) = p(y), \quad \mathbf{e}(y, 0) = \frac{1}{\beta(y)} + \frac{p^2(y)}{2\bar{m}} + \frac{r^2(y)}{2}.$$



Next, the evolution at all further time is governed by the following system of conservation laws:

$$\partial_t \mathbf{r}(y, t) = \frac{1}{m} \partial_y \mathbf{p}(y, t), \quad \mathbf{r}(0, t) = \mathbf{r}(1, t) = 0, \quad (2.5)$$

$$\partial_t \mathbf{p}(y, t) = \partial_y \mathbf{r}(y, t), \quad (2.6)$$

$$\partial_t \mathbf{e}(y, t) = \frac{1}{m} \partial_y (\mathbf{r}(y, t) \mathbf{p}(y, t)). \quad (2.7)$$

Thanks to the regularity conditions on  $r, p$  in (2.3), the solutions of these equations are classical. Since  $(\mathbf{r}, \mathbf{p})$  are solution of wave equations with suitable boundary conditions, they can be obtained explicitly by expanding them in Fourier series. Then, by a time integration,  $\mathbf{e}$  may be expressed as a function of  $(\mathbf{r}, \mathbf{p})$ , see (4.1). Later we will use that a classical solution for the system governing  $(\mathbf{r}, \mathbf{p})$  coincides with the (unique) weak solution of this system. Because of the boundary conditions, test functions will have to be chosen appropriately (see (3.8-3.9)).

**Theorem 1.** *Let  $t \in [0, T]$  and  $f \in \mathcal{C}^0([0, 1])$ . Let us assume that the system is initially prepared in a locally Gibbs state such that  $\beta, r$  and  $p$  satisfy (2.3). Then, as  $N \rightarrow \infty$ , almost surely (w.r.t.  $\mathbb{P}$ ),*

$$\mathcal{R}_N(f, t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \langle r_x(Nt) \rangle \quad \rightarrow \quad \mathcal{R}(f, t), \quad (2.8)$$

$$\mathcal{P}_N(f, t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \langle p_x(Nt) \rangle \quad \rightarrow \quad \mathcal{P}(f, t), \quad (2.9)$$

$$\mathcal{E}_N(f, t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \langle e_x(Nt) \rangle \quad \rightarrow \quad \mathcal{E}(f, t). \quad (2.10)$$

**Remark 1.** *As pointed out in the introduction, the situation is much simpler at thermal equilibrium, i.e. for  $\beta$  constant, and these limits hold even for the non-disordered chain. See Section 4.3 for a derivation along the lines used to derive Theorem 1.*

### 3. EVOLUTION OF $\mathcal{R}_N$ AND $\mathcal{P}_N$

In this section, we show the limits (2.8-2.9). Moreover, in order to later deal with the field  $\mathcal{E}_N$ , we show more:

The functions  $\langle r_{[Ny]}(Nt) \rangle$  and  $m_{[Ny]}^{-1} \langle p_{[Ny]}(Nt) \rangle$  are uniformly (in  $N$ ) Hölder regular in  $y \in [0, 1]$ , with exponent at least  $1/2$ . Hence they converge pointwise to  $\mathbf{r}(y, t)$  and  $\mathbf{p}(y, t)$  respectively.

**3.1. A priori estimates.** Given  $d \in \mathbb{N}$ , we denote the standard scalar product on  $\mathbb{R}^d$  by  $\langle \cdot, \cdot \rangle_d$  (we will drop the subscript  $d$  when no confusion seems possible). Let us consider the two following conserved quantities:

$$H(r, p) = \frac{1}{2} (\langle p, M^{-1} p \rangle_N + \langle r, r \rangle_{N-1}), \quad (3.1)$$

$$I(r, p) = \frac{1}{2} (\langle \nabla_- r, M^{-1} \nabla_- r \rangle_N + \langle \nabla_+ M^{-1} p, \nabla_+ M^{-1} p \rangle_{N-1}). \quad (3.2)$$

The conservation of  $I$  follows from the fact that, if  $(r, p)$  solve (2.1), then  $(\nabla_+ M^{-1} p, \nabla_- r)$  solve the same equation, the corresponding Hamiltonian being  $I$  (since we have that

$H(\nabla_+ M^{-1} p, \nabla_- r) = I(r, p)$ ). Notice also that a full set of conserved quantities can be generated by further taking gradients, see Section 5.2.

Thanks to these two conservation laws, and to the smoothness assumptions on  $r$  and  $p$ , we deduce

**Lemma 1.** *There exists  $C$  such that, for any  $t \geq 0$  and any  $N \in \mathbb{N}$ ,*

$$\sum_{x=1}^{N-1} \langle r_x(Nt) \rangle^2 \leq CN, \quad \sum_{x=1}^N \langle p_x(Nt) \rangle^2 \leq CN, \quad (3.3)$$

$$\sum_{x=1}^N \langle (\nabla_- r)_x(Nt) \rangle^2 \leq \frac{C}{N}, \quad \sum_{x=1}^{N-1} \langle (\nabla_+ M^{-1} p)_x(Nt) \rangle^2 \leq \frac{C}{N}. \quad (3.4)$$

*Proof.* By linearity of the equations of motion (2.1),  $(\langle r \rangle, \langle p \rangle)$  solve the same equations as  $(r, p)$ . Therefore, the conservation of  $H(r, p)$  and  $I(r, p)$  implies the conservation of  $H(\langle r \rangle, \langle p \rangle)$  and  $I(\langle r \rangle, \langle p \rangle)$ . Since the quantities to be estimated in (3.3) are bounded by  $H(\langle r \rangle, \langle p \rangle)$  and the quantities to be estimated in (3.4) are bounded by  $I(\langle r \rangle, \langle p \rangle)$ , we conclude that it is enough to establish them respectively for  $H(\langle r \rangle, \langle p \rangle)$  and  $I(\langle r \rangle, \langle p \rangle)$  at  $t = 0$ . This follows from a direct computation, thanks to the product structure of the local Gibbs state (2.2) and to the hypotheses on  $r$  and  $p$  in (2.3) (in particular, this is the place where the boundary condition on  $r$  plays a role).  $\square$

As a corollary, we deduce the existence of a constant  $C \in \mathbb{R}$  such that, for any  $x, y \in \mathbb{Z} \cap [1, N]$ ,

$$\begin{aligned} |\langle r_{x'}(Nt) \rangle - \langle r_x(Nt) \rangle| &\leq C \left| \frac{x'-x}{N} \right|^{1/2}, \\ |m_{x'}^{-1} \langle p_{x'}(Nt) \rangle - m_x^{-1} \langle p_x(Nt) \rangle| &\leq C \left| \frac{x'-x}{N} \right|^{1/2}, \end{aligned} \quad (3.5)$$

and therefore also such that

$$|\langle r_x(Nt) \rangle| \leq C, \quad |\langle p_x(Nt) \rangle| \leq C. \quad (3.6)$$

Indeed, to get e.g. (3.5) for  $r$ , we deduce from (3.4) that

$$\begin{aligned} |\langle r_{x'}(Nt) \rangle - \langle r_x(Nt) \rangle| &= \left| \sum_{z=x+1}^{x'} \langle (\nabla_- r)_z(Nt) \rangle \right| \leq \left( \sum_{z=1}^N \langle (\nabla_- r)_z(Nt) \rangle^2 \right)^{1/2} |x - x'|^{1/2} \\ &\leq C \left| \frac{x'-x}{N} \right|^{1/2}. \end{aligned}$$

**3.2. Averaging lemma for the field  $\mathcal{P}_N$ .** The method of the corrected empirical measure is an elegant method to deal with the randomness on the masses in deriving the hydrodynamic limit for  $\mathcal{R}_N$  and  $\mathcal{P}_N$  [12, 14, 6]. However, in our case, it seems more convenient to use the following lemma:

**Lemma 2.** *Let  $f \in \mathcal{C}^0([0, 1])$  and  $t \geq 0$ . Almost surely (w.r.t. the masses), for  $N \rightarrow \infty$ ,*

$$\frac{1}{N} \sum_{x=1}^N f(x/N) \frac{\langle p_x(Nt) \rangle}{m_x} (m_x - \bar{m}) \rightarrow 0. \quad (3.7)$$

*Proof.* Let  $A_N$  be the quantity in the left hand side of (3.7), let  $\tilde{m}_x = m_x - \bar{m}$ , and let

$$\varphi(x) = f(x/N) \frac{\langle p_x(Nt) \rangle}{m_x}.$$

Let  $0 < \tau < 1$  and let us decompose  $A_N$  as

$$\begin{aligned}
A_N &= \frac{1}{N^{1-\tau}} \sum_{x \in N^{1-\tau} \mathbb{Z} \cap [0, N]} \frac{1}{N^\tau} \sum_{z=1}^{N^\tau} \varphi(x+z) \tilde{m}_{x+z} \\
&= \frac{1}{N^{1-\tau}} \sum_{x \in N^{1-\tau} \mathbb{Z} \cap [0, N]} \frac{\varphi(x)}{N^\tau} \sum_{z=1}^{N^\tau} \tilde{m}_{x+z} \\
&\quad + \frac{1}{N^{1-\tau}} \sum_{x \in N^{1-\tau} \mathbb{Z} \cap [0, N]} \frac{1}{N^\tau} \sum_{z=1}^{N^\tau} (\varphi(x+z) - \varphi(x)) \tilde{m}_{x+z} \\
&=: A_N^{(1)} + A_N^{(2)}.
\end{aligned}$$

To deal with  $A_N^{(1)}$ , we observe that  $\varphi$  is bounded, see (3.6), so that by Jensen's inequality,

$$\mathbb{E}((A_N^{(1)})^4) \leq \frac{C}{N^{1-\tau}} \sum_{x \in N^{1-\tau} \mathbb{Z} \cap [0, N]} \mathbb{E} \left( \frac{1}{N^\tau} \sum_{z=1}^{N^\tau} \tilde{m}_{x+z} \right)^4 \leq \frac{C}{N^{2\tau}}$$

Taking  $\tau > 1/2$ , this shows that  $A_N^{(1)} \rightarrow 0$  almost surely by Borel-Cantelli's lemma. To show that  $A_N^{(2)} \rightarrow 0$  (deterministically), observe that for any  $\varepsilon > 0$ ,  $|\varphi((x+z)/N) - \varphi(x/N)| < \varepsilon$  holds for all  $N$  large enough. This property follows from the continuity of  $f$  and of  $m_x^{-1} \langle p_x(Nt) \rangle$  expressed in (3.5).  $\square$

**3.3. Proof of the convergence to the linear wave equation (2.8-2.9).** For any smooth functions  $f, g : [0, 1] \in \mathbb{R}$  such that  $f(0) = f(1) = 0$ , the limiting fields  $\mathcal{R}$  and  $\mathcal{P}$  defined in (2.4) can be equivalently characterized as follows:

$$\mathcal{R}(f, t) = \mathcal{R}(f, 0) - \frac{1}{\bar{m}} \int_0^t \mathcal{P}(f', s) ds, \quad (3.8)$$

$$\mathcal{P}(g, t) = \mathcal{P}(g, 0) - \int_0^t \mathcal{R}(g', s) ds, \quad (3.9)$$

and

$$\mathcal{R}(f, 0) = \int_0^1 f(x) r(x) dx, \quad \mathcal{P}(g, 0) = \int_0^1 g(x) p(x) dx. \quad (3.10)$$

Let us use this characterization to show that  $\mathcal{R}_N(f, t) \rightarrow \mathcal{R}(f, t)$  and  $\mathcal{P}_N(g, t) \rightarrow \mathcal{P}(g, t)$ .

The convergence at  $t = 0$  follows from the strong law of large numbers:  $\mathcal{R}_N(f, 0)$  and  $\mathcal{P}_N(g, 0)$  converge almost surely to  $\mathcal{R}(f, 0)$  and  $\mathcal{P}(g, 0)$  given by (3.10).

Let us next consider  $t \geq 0$ , and let us first deal with  $\mathcal{R}_N$ . Integrating the equations of motion yields

$$\begin{aligned}
\mathcal{R}_N(f, t) &= \mathcal{R}_N(f, 0) + \int_0^{Nt} \frac{1}{N} \sum_{x=1}^N f(x/N) \langle (\nabla_+ M^{-1} p)_x(s) \rangle ds \\
&= \mathcal{R}_N(f, 0) - \int_0^{Nt} \frac{1}{N} \sum_{x=1}^N \nabla_- f(x/N) m_x^{-1} \langle p_x(s) \rangle ds
\end{aligned}$$

where we used the boundary condition  $f(0) = f(1) = 0$  to perform the integration by part. Since  $\nabla_- f(x/N) = N^{-1}f'(x/N) + \mathcal{O}(N^{-2})$ , we obtain

$$\mathcal{R}_N(f, t) = \mathcal{R}_N(f, 0) - \int_0^t \frac{1}{N} \sum_{x=1}^N f'(x/N) m_x^{-1} \langle p_x(Ns) \rangle ds + \mathcal{O}\left(\frac{1}{N}\right).$$

Now we use Lemma 2 to replace  $m_x^{-1}$  by  $(\bar{m})^{-1}$  up to an error that vanishes almost surely in the limit  $N \rightarrow \infty$  (strictly speaking, due to the presence of the integral over time, we cannot use this lemma as such, but one checks that the proof is basically not affected). Thus

$$\mathcal{R}_N(f, t) = \mathcal{R}_N(f, 0) - \frac{1}{\bar{m}} \int_0^t \mathcal{P}_N(f', s) ds + \varepsilon_N, \quad (3.11)$$

where  $\varepsilon_N \rightarrow 0$  almost surely as  $N \rightarrow \infty$ . Let us next deal with  $\mathcal{P}_N$ . This case is simpler since no homogenization over the masses is needed. Proceeding similarly, we find

$$\begin{aligned} \mathcal{P}_N(g, t) &= \mathcal{P}_N(g, 0) + \int_0^{Nt} \frac{1}{N} \sum_{x=1}^N g(x/N) \langle (\nabla_- r)_x(s) \rangle ds \\ &= \mathcal{P}_N(g, 0) - \int_0^{Nt} \frac{1}{N} \sum_{x=1}^N \nabla_+ g(x/N) \langle r_x(s) \rangle ds \\ &= \mathcal{P}_N(g, 0) - \int_0^t \mathcal{R}_N(g', s) ds + \tilde{\varepsilon}_N \end{aligned} \quad (3.12)$$

where we used the boundary condition  $r_0(s) = r_N(s) = 0$  for all time  $s \geq 0$  to perform the integration by part, and where  $\tilde{\varepsilon}_N \rightarrow 0$  deterministically as  $N \rightarrow \infty$ .

The families  $(\mathcal{R}_N(f, \cdot))_N$  and  $(\mathcal{P}_N(g, \cdot))_N$  are equicontinuous since a uniform bound on the time derivative of  $\mathcal{R}_N(f, \cdot)$  and  $\mathcal{P}_N(g, \cdot)$  holds. Hence, the relations (3.11) and (3.12) implies that any limiting point must satisfy (3.8-3.9).

**3.4. Pointwise convergence.** Thanks to the Hölder regularity of both  $\langle r_x(Nt) \rangle$  and  $m_x^{-1} \langle p_x(Nt) \rangle$  expressed by (3.5), we deduce a stronger result:

**Proposition 1.** *Let  $y \in ]0, 1[$  and let  $t \in ]0, 1[$ . As  $N \rightarrow \infty$ , almost surely (w.r.t. P),*

$$\langle r_{[Ny]}(Nt) \rangle \rightarrow \mathbf{r}(y, t), \quad \frac{\langle p_{[Ny]}(Nt) \rangle}{m_{[Ny]}} \rightarrow \frac{\mathbf{p}(y, t)}{\bar{m}}.$$

*Proof.* Let us first deal with  $\langle r_{[Ny]}(Nt) \rangle$ . Let  $(\rho_\epsilon)_{\epsilon > 0}$  be a regularizing family:  $\rho_\epsilon \in \mathcal{C}^\infty(\mathbb{R})$ ,  $\text{supp}(\rho_\epsilon) = [-\epsilon, \epsilon]$ ,  $\rho_\epsilon \geq 0$  and  $\int \rho_\epsilon(y) dy = 1$ . For  $y \in ]\epsilon, 1 - \epsilon[$ , we decompose

$$\begin{aligned} \langle r_{[Ny]}(Nt) \rangle &= \int \rho_\epsilon(y - y') \langle r_{[Ny]}(Nt) \rangle dy' \\ &= \int \rho_\epsilon(y - y') \langle r_{[Ny']}(Nt) \rangle dy' + \int \rho_\epsilon(y - y') (\langle r_{[Ny]}(Nt) \rangle - \langle r_{[Ny']}(Nt) \rangle) dy'. \end{aligned}$$

By (3.5) the second term is bounded in absolute value by

$$\int \rho_\epsilon(y - y') |\langle r_{[Ny]}(Nt) \rangle - \langle r_{[Ny']}(Nt) \rangle| dy' \leq C\sqrt{\epsilon}, \quad (3.13)$$

while the first term is approximated uniformly in  $N$  by

$$\frac{1}{N} \sum_{x=1}^N \rho_\epsilon \left( \frac{x}{N} - y \right) \langle r_x(Nt) \rangle$$

that converges, by the result shown in the previous paragraph, to

$$\int_0^1 \rho_\epsilon(y - y') \mathbf{r}(y', t) dy' \quad (3.14)$$

as  $N \rightarrow \infty$ . Letting next  $\epsilon \rightarrow 0$ , the continuity of  $\mathbf{r}(\cdot, t)$  implies that (3.14) converges to  $\mathbf{r}(y, t)$  while (3.13) converges to 0 as  $N \rightarrow \infty$ . To deal with  $m_{[Ny]}^{-1} \langle p_{[Ny]}(Nt) \rangle$ , we proceed similarly, using Lemma 2, to get the analog of (3.14).  $\square$

Finally, thanks to the bound (3.6) and the pointwise convergence result in Proposition 1, and thanks to the averaging Lemma 2 for the field  $\mathcal{P}_N$ , we derive (2.8-2.9) by applying the dominated convergence theorem.

#### 4. EVOLUTION OF THE ENERGY $\mathcal{E}_N$

In this section we show the limit (2.10). We will assume that  $f \in \mathcal{C}^1([0, 1])$ . We can then recover the result (2.10) for  $f \in \mathcal{C}^0([0, 1])$  by density, and using the a priori estimate  $\sum_x \langle e_x(t) \rangle \leq CN$  at all time  $t \geq 0$ .

**4.1. Main decomposition of the energy.** In order to derive the limit of  $\mathcal{E}_N$ , we separate the contribution to the total energy from the temperature (that does not evolve with time) and from mechanical energy, i.e. the average kinetic and potential energy (that does evolve due to the transport of momentum and displacement).

At the macroscopic level, we deduce from (2.5-2.7) that

$$\begin{aligned} \mathbf{e}(y, t) &= \mathbf{e}(y, 0) + \frac{1}{m} \int_0^t \partial_y (\mathbf{r}(y, s) \mathbf{p}(y, s)) ds \\ &= \frac{1}{\beta(y)} + \frac{\mathbf{p}(y, 0)}{2\bar{m}} + \frac{\mathbf{r}(y, 0)}{2} + \frac{1}{m} \int_0^t \partial_s (\mathbf{p}^2(y, s)/2 + \bar{m} \mathbf{r}(y, s)/2) ds \\ &= \frac{\mathbf{p}^2(y, t)}{2\bar{m}} + \frac{\mathbf{r}^2(y, t)}{2} + \frac{1}{\beta(y)}. \end{aligned} \quad (4.1)$$

At the microscopic level, we decompose

$$\begin{aligned} \mathcal{E}_N(f, t) &= \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle p_x^2 \rangle}{2m_x} + \frac{\langle r_x^2 \rangle}{2} \right) (Nt) \\ &= \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle p_x \rangle^2}{2m_x} + \frac{\langle r_x \rangle^2}{2} \right) (Nt) + \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle \tilde{p}_x^2 \rangle}{2m_x} + \frac{\langle \tilde{r}_x^2 \rangle}{2} \right) (Nt) \\ &=: \mathcal{A}_N(t) + \mathcal{F}_N(t), \end{aligned} \quad (4.2)$$

with

$$\tilde{p}_x = p_x - \langle p_x \rangle, \quad \tilde{r}_x = r_x - \langle r_x \rangle,$$

and where  $\mathcal{A}$  and  $\mathcal{F}$  stands respectively for “average” and “fluctuations”. Comparing (4.1) and (4.2), we conclude that it is enough to show that, P almost surely, as  $N \rightarrow \infty$ ,

$$\mathcal{A}_N(t) \rightarrow \int_0^1 f(y) \left( \frac{\mathbf{p}^2(y, t)}{2\bar{m}} + \frac{\mathbf{r}^2(y, t)}{2} \right) dy, \quad (4.3)$$

$$\mathcal{F}_N(t) - \mathcal{F}_N(0) \rightarrow 0, \quad \mathcal{F}_N(0) \rightarrow \int_0^1 \frac{f(y)}{\beta(y)} dy. \quad (4.4)$$

The limit (4.3) is deduced in the same ways as (2.8-2.9): To deal with the term involving  $\langle p_x \rangle^2 / 2m_x$ , we use an averaging result similar to Lemma 2 (and which proof follows exactly the same lines): given  $f \in \mathcal{C}^1([0, 1])$ , almost surely (w.r.t. the masses), as  $N \rightarrow \infty$ ,

$$\frac{1}{N} \sum_{x=1}^N f(x/N) \frac{\langle p_x(Nt) \rangle^2}{m_x^2} (m_x - \bar{m}) \rightarrow 0.$$

Next, thanks to the bound (3.6) and the pointwise convergence result in Proposition 1, we derive (4.3) by applying the dominated convergence theorem.

The limit (4.4) will be established thanks to the localization of the high modes of the chain; this is the only place where localization is used. Moreover, we will show in Section 4.3 that in thermal equilibrium, the equality  $\mathcal{F}_N(t) = \mathcal{F}_N(0)$  holds without any assumption on the distribution of the masses (besides positivity). This shows thus that Theorem 1 holds actually also for a clean chain if  $\beta$  is constant.

**4.2. Convergence of  $\mathcal{F}_N(t)$ .** To deal with the limit (4.4), we will use the fact that any mode of the chain at positive energy is spatially localized in the thermodynamic limit. Hence, we will expand the solutions to the equations of motion into the eigenmodes of the chain. In Section 5 below, we carry this expansion in details and we deduce the needed localization estimates. For our present purposes, it suffices to know the following: There exists a basis  $\{\psi^k\}_{0 \leq k \leq N-1}$  of  $\mathbb{R}^N$ , the basis of the eigenmodes of the chain, so that the solutions to the equations of motion read

$$\tilde{r}_x(t) = \sum_{k=1}^{N-1} \left( \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \cos \omega_k t + \langle \psi^k, \tilde{p}(0) \rangle \sin \omega_k t \right) \frac{(\nabla_+ \psi^k)_x}{\omega_k}, \quad (4.5)$$

$$\tilde{p}_x(t) = \sum_{k=0}^{N-1} \left( \langle \psi^k, \tilde{p}(0) \rangle \cos \omega_k t - \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \sin \omega_k t \right) (M\psi^k)_x, \quad (4.6)$$

where  $\omega_0 = 0$  and  $\omega_k > 0$  for  $1 \leq k \leq N$  are the corresponding eigenfrequencies of the chain. Observe that the first term starts from  $k = 1$  while the second one starts from  $k = 0$ . Moreover, the orthogonality relation  $\langle \psi^k, M\psi^j \rangle = \delta(k - j)$  holds and  $\{\omega_k^{-1} \nabla_+ \psi^k\}_{1 \leq k \leq N-1}$  forms an orthonormal basis of  $(\mathbb{R}^{N-1}, \langle \cdot, \cdot \rangle_{N-1})$ . See Section 5.1 for more details. This representation is useful to exploit localization: all modes with  $k \gtrsim \sqrt{N}$  are spatially localized. See Section 5.3 for more quantitative estimates. However, low modes with  $k \lesssim \sqrt{N}$  remain extended, and we will have to show that the contribution of these modes vanish since their proportion  $N^{1/2}/N \rightarrow 0$  in the thermodynamic limit. Below, for technical reasons, we will replace  $1/2$  by  $1 - \alpha$ , for some  $\alpha > 0$  that we will need to choose small enough.

Let  $0 < \alpha \ll 1$ , let

$$F_1 = \mathbb{Z} \cap [0, N^{1-\alpha}], \quad F_2 = \mathbb{Z} \cap [N^{1-\alpha}, N - 1],$$

and let us decompose  $\tilde{r}(t) = \tilde{r}^{(1)}(t) + \tilde{r}^{(2)}(t)$  and  $\tilde{p}(t) = \tilde{p}^{(1)}(t) + \tilde{p}^{(2)}(t)$  with

$$\tilde{r}^{(i)}(t) = \sum_{k \in F_i \setminus \{0\}} (\dots), \quad \tilde{p}^{(i)}(t) = \sum_{k \in F_i} (\dots),$$

for  $i = 1, 2$  and with (...) the summand featuring in (4.5) or (4.6). We insert this decomposition in  $\mathcal{F}_N$ :

$$\mathcal{F}_N(t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle (\tilde{p}_x^{(1)} + \tilde{p}_x^{(2)})^2 \rangle}{2m_x} + \frac{\langle (\tilde{r}_x^{(1)} + \tilde{r}_x^{(2)})^2 \rangle}{2} \right) (Nt).$$

Let us show the two following limits:

$$\mathcal{F}_N^{(1)}(t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle (\tilde{p}_x^{(1)})^2 \rangle}{2m_x} + \frac{\langle (\tilde{r}_x^{(1)})^2 \rangle}{2} \right) (Nt) \rightarrow 0, \quad (4.7)$$

$$\mathcal{F}_N^{(2)}(t) = \frac{1}{N} \sum_{x=1}^N f(x/N) \left( \frac{\langle (\tilde{p}_x^{(2)})^2 \rangle}{2m_x} + \frac{\langle (\tilde{r}_x^{(2)})^2 \rangle}{2} \right) (Nt) \rightarrow \int \frac{f(y)}{\beta(y)} dy, \quad (4.8)$$

which, by Cauchy-Schwarz inequality, implies (4.4).

Let us show (4.7). Let us bound  $|f(x/N)| \leq C$ , and use the explicit solution (4.5-4.6):

$$\begin{aligned} & |\mathcal{F}_N^{(1)}(t)| \\ & \leq \frac{C}{2N} \sum_{x=1}^N \frac{1}{m_x} \left\langle \left( \sum_{k \in F_1} (\langle \psi^k, \tilde{p}(0) \rangle \cos(\omega_k Nt) - \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \sin(\omega_k Nt)) m_x \psi_x^k \right)^2 \right\rangle \\ & + \frac{C}{2N} \sum_{x=1}^N \left\langle \left( \sum_{k \in F_1 \setminus \{0\}} \left( \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \cos(\omega_k Nt) + \langle \psi^k, \tilde{p}(0) \rangle \sin(\omega_k Nt) \right) \frac{(\nabla_+ \psi^k)_x}{\omega_k} \right)^2 \right\rangle. \end{aligned}$$

In both terms, one may expand the square so as to get a double sum over  $k, j \in F_1$  or  $k, j \in F_1 \setminus \{0\}$ , and insert the sum over  $x$  inside the sum over  $k, j$ . This yields

$$\sum_{x=1}^N \frac{m_x^2}{m_x} \psi_x^j \psi_x^k = \langle \psi^j, M \psi^k \rangle = \delta(k-j), \quad \sum_{x=1}^N \frac{(\nabla_+ \psi^j)_x (\nabla_+ \psi^k)_x}{\omega_j \omega_k} = \delta(k-j).$$

Thus

$$\begin{aligned} |\mathcal{F}_N^{(1)}(t)| & \leq \frac{C}{2N} \sum_{k \in F_1} \left\langle \left( \langle \psi^k, \tilde{p}(0) \rangle \cos(\omega_k Nt) - \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \sin(\omega_k Nt) \right)^2 \right\rangle \\ & + \frac{C}{2N} \sum_{k \in F_1 \setminus \{0\}} \left\langle \left( \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \cos(\omega_k Nt) + \langle \psi^k, \tilde{p}(0) \rangle \sin(\omega_k Nt) \right)^2 \right\rangle. \end{aligned}$$

At this point, it suffices to show that there exists a constant  $C$  such that, for all  $k \in F_1$ ,

$$\langle \langle \psi^k, \tilde{p}(0) \rangle^2 \rangle \leq C, \quad \frac{\langle \langle \nabla_+ \psi^k, \tilde{r}(0) \rangle^2 \rangle}{\omega_k^2} \leq C, \quad (4.9)$$

since, bounding sin and cos by 1, and using Cauchy-Schwarz, we obtain

$$|\mathcal{F}_N^{(1)}(t)| \leq \frac{C}{N} \sum_{k \in F_1} 1 = \frac{CN^{1-\alpha}}{N} \rightarrow 0.$$

Let us deal with  $\langle \langle \psi^k, \tilde{p}(0) \rangle \rangle$  (the other case is analogous):

$$\begin{aligned} \langle \langle \psi^k, \tilde{p}(0) \rangle \rangle &= \left\langle \left( \sum_x \psi_x^k \tilde{p}_x(0) \right)^2 \right\rangle = \left\langle \sum_{x,y} \psi_x^k \psi_y^k \tilde{p}_x(0) \tilde{p}_y(0) \right\rangle \\ &= \sum_x (\psi_x^k)^2 \langle \langle \tilde{p}_x(0) \rangle \rangle \end{aligned}$$

where we have used the fact that  $\langle \cdot \rangle$  is a product measure and that  $\langle \tilde{p}_x(0) \rangle = 0$  for all  $x \in \{1, \dots, N\}$ . We compute  $\langle \langle \tilde{p}_x(0) \rangle \rangle = \frac{m_x}{\beta(x/N)}$ . Since  $\beta$  is positive and continuous on  $[0, 1]$ , there exists  $\beta_- > 0$  such that  $\beta(x/N) \geq \beta_-$  for all  $x \in \{1, \dots, N\}$ . Hence

$$\langle \langle \psi^k, \tilde{p}(0) \rangle \rangle \leq \frac{1}{\beta_-} \sum_{x=1}^N m_x (\psi_x^k)^2 = \frac{1}{\beta_-} \langle \psi^k, M \psi^k \rangle = \frac{1}{\beta_-}. \quad (4.10)$$

Let us now show (4.8). A computation using the initial measure shows that  $\mathcal{F}(0) \rightarrow \int \frac{f(y)}{\beta(y)} dy$  as  $N \rightarrow \infty$ . Hence, thanks to (4.7), it holds that  $\mathcal{F}_N^{(2)}(0) \rightarrow \int \frac{u(y)}{\beta(y)} dy$  as  $N \rightarrow \infty$ . Thus it suffices to show that  $\mathcal{F}_N^{(2)}(t) - \mathcal{F}_N^{(2)}(0) \rightarrow 0$  as  $N \rightarrow \infty$ . Let us write  $\mathcal{F}_N^{(2)}(t)$  as a scalar product and expand it in the eigenmodes basis:

$$\begin{aligned} \mathcal{F}_N^{(2)}(t) &= \frac{1}{2N} \left\langle \langle (f \cdot \tilde{p}^{(2)})(Nt), M^{-1} \tilde{p}^{(2)}(Nt) \rangle + \langle (f \cdot \tilde{r}^{(2)})(Nt), \tilde{r}^{(2)}(Nt) \rangle \right\rangle \\ &= \frac{1}{2N} \sum_{k \in F_2} \left\langle \langle (f \cdot \tilde{p}^{(2)})(Nt), \psi^k \rangle \langle \psi^k, \tilde{p}(Nt) \rangle \right. \\ &\quad \left. + \frac{1}{\omega_k^2} \langle (f \cdot \tilde{r}^{(2)})(Nt), \nabla_+ \psi^k \rangle \langle \nabla_+ \psi^k, \tilde{r}(Nt) \rangle \right\rangle. \end{aligned}$$

Here  $g \cdot h$  denotes a function on  $\mathbb{Z} \cap [1, N]$  obtained by the usual multiplication in real space between a function  $g$  on  $[0, 1]$  and  $h$  on  $\mathbb{Z} \cap [1, N]$ , i.e.  $(g \cdot h)_x = g(x/N)h_x$ . By Lemma 3 stated in Section 5 below, one may associate a localization center  $x_0(k)$  to each mode  $\psi^k$  with  $k \in F_2$ :  $x_0(k)$  is the center of the interval  $J(k)$  featuring there (assuming that  $\alpha$  is small enough so that the hypotheses of Lemma 3 are satisfied). For each  $k \in F_2$ , let us decompose  $f$  as

$$f = f_0(k) + \tilde{f}_k \quad \text{with} \quad f_0(k) = f(x_0(k)/N)$$

(thus  $f_0(k)$  is a constant and  $\tilde{f}_k$  vanishes at  $x_0(k)/N$ ). We insert this decomposition in the above expression for  $\mathcal{F}_N^{(2)}(t)$ :

$$\mathcal{F}_N^{(2)}(t) = \frac{1}{2N} \sum_{k \in F_2} f_0(k) \left\langle \langle \tilde{p}(Nt), \psi^k \rangle^2 + \frac{\langle \tilde{r}(Nt), \nabla_+ \psi^k \rangle^2}{\omega_k^2} \right\rangle \quad (4.11)$$

$$\begin{aligned} &+ \frac{1}{2N} \sum_{k \in F_2} \left\langle \langle (\tilde{f}_k \cdot \tilde{p}^{(2)})(Nt), \psi^k \rangle \langle \psi^k, \tilde{p}(Nt) \rangle \right. \\ &\quad \left. + \frac{1}{\omega_k^2} \langle (\tilde{f}_k \cdot \tilde{r}^{(2)})(Nt), \nabla_+ \psi^k \rangle \langle \nabla_+ \psi^k, \tilde{r}(Nt) \rangle \right\rangle. \quad (4.12) \end{aligned}$$

Each expression between  $\langle \dots \rangle$  in the sum in (4.11) represents the energy of the mode  $\psi^k$  and does not evolve with time, see (5.4) in Section 5 below. Hence, to show  $\mathcal{F}_N^{(2)}(t) - \mathcal{F}_N^{(2)}(0) \rightarrow 0$ , we only need to show that the sum in (4.12) converges to 0 as  $N \rightarrow \infty$ .



Let us consider a single term in the sum (4.12), and let us focus on the term involving  $\tilde{p}$  (the one involving  $\tilde{r}$  is treated the same way). First, by Cauchy-Schwarz,

$$\langle \langle (\tilde{f}_k \cdot \tilde{p}^{(2)})(Nt), \psi^k \rangle \langle \psi^k, \tilde{p}(Nt) \rangle \rangle \leq \langle \langle (\tilde{f}_k \cdot \tilde{p}^{(2)})(Nt), \psi^k \rangle^2 \rangle^{1/2} \langle \langle \psi^k, \tilde{p}(Nt) \rangle^2 \rangle^{1/2}. \quad (4.13)$$

The second factor in (4.13) is bounded by a constant:

$$\begin{aligned} \langle \langle \psi^k, \tilde{p}(Nt) \rangle^2 \rangle &= \left\langle \left( \langle \psi^k, \tilde{p}(0) \rangle \cos \omega_k Nt - \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \sin \omega_k Nt \right)^2 \right\rangle \\ &\leq 2 \left\langle \langle \psi^k, \tilde{p}(0) \rangle^2 + \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle^2}{\omega_k^2} \right\rangle \leq C, \end{aligned} \quad (4.14)$$

see (4.9). For the first factor in (4.13), we use again Cauchy-Schwarz to get

$$\begin{aligned} \langle \langle (\tilde{f}_k \cdot \tilde{p}^{(2)})(Nt), \psi^k \rangle^2 \rangle &= \left\langle \left( \sum_x \tilde{f}_k(x/N) \tilde{p}_x^{(2)}(Nt) \psi_x^k \right)^2 \right\rangle \\ &\leq \left( \sum_x \tilde{f}_k^2(x/N) (\psi_x^k)^2 \right) \left\langle \sum_x (\tilde{p}_x^{(2)})^2(Nt) \right\rangle. \end{aligned} \quad (4.15)$$

For  $\tilde{f}_k$ , we have the bound

$$|\tilde{f}_k(x/N)| = |f_k(x/N) - f_k(x_0(k)/N)| \leq C \frac{|x - x_0(k)|}{N}$$

(this is the only place where we use  $f \in \mathcal{C}^1([0, 1])$ ). Hence, from Lemma 3 below, we deduce that for any  $\epsilon > 0$ , the first factor in the right hand side of (4.15) can be bounded by  $1/N^{2-\epsilon}$  by taking  $\alpha > 0$  small enough. From the conservation of energy (see (5.4) and the bound (4.9)), the second factor in (4.15) is  $\mathcal{O}(N)$ . Hence, for  $\alpha > 0$  small enough, (4.15) goes to zero as  $N \rightarrow \infty$ . Combining this with (4.14), we find that (4.13) goes to zero as  $N \rightarrow \infty$ , and hence that (4.12) converges to 0 as  $N \rightarrow \infty$ .

**4.3. Thermal equilibrium case.** Assume here that there exists some  $\bar{\beta} > 0$  such that  $\beta(y) = \bar{\beta}$  for all  $y \in [0, 1]$ . Then, we may relax the assumptions on the masses: requiring only that they are all strictly positive, let us show that  $\mathcal{F}_N(t) = F_N(0)$  for all  $t \geq 0$ . This results from an exact computation.

Since  $f$  is arbitrary, it is necessary and sufficient to prove that, for any  $x$ ,

$$\frac{d}{dt} \left( \frac{\langle \tilde{p}_x^2(t) \rangle}{2m_x} + \frac{\langle \tilde{r}_x^2(t) \rangle}{2} \right) = 0.$$

We compute

$$\begin{aligned} \frac{\tilde{p}_x^2(t)}{2m_x} &= \frac{1}{2} \sum_{j,k} \left( \langle \psi^j, \tilde{p}(0) \rangle \cos \omega_j t - \frac{\langle \nabla_+ \psi^j, \tilde{r}(0) \rangle}{\omega_j} \sin \omega_j t \right) \\ &\quad \left( \langle \psi^k, \tilde{p}(0) \rangle \cos \omega_k t - \frac{\langle \nabla_+ \psi^k, \tilde{r}(0) \rangle}{\omega_k} \sin \omega_k t \right) m_x \psi_x^j \psi_x^k. \end{aligned}$$

A similar expression holds for  $\tilde{r}^2(t)/2$ . In order to obtain the expectation with respect to  $\langle \cdot \rangle$ , we compute

$$\langle \langle \psi^k, \tilde{p}(0) \rangle \langle \psi^j, \tilde{p}(0) \rangle \rangle = \frac{\delta(k-j)}{\beta}, \quad (4.16)$$

$$\langle \langle \psi^k, \tilde{p}(0) \rangle \langle \nabla_+ \psi^j, \tilde{r}(0) \rangle \rangle = 0, \quad (4.17)$$

$$\langle \langle \nabla_+ \psi^k, \tilde{r}(0) \rangle \langle \nabla_+ \psi^j, \tilde{r}(0) \rangle \rangle = \frac{\omega_k^2 \delta(k-j)}{\beta}. \quad (4.18)$$

These three properties result from the fact the product structure of  $\rho_{\text{loc}}$ , from the fact that the variables  $\tilde{p}$  and  $\tilde{r}$  are centered, and from the fact that  $\beta$  is constant for (4.16) and (4.18). E.g. (4.16) is obtained by

$$\begin{aligned} \langle \langle \psi^k, \tilde{p}(0) \rangle \langle \psi^j, \tilde{p}(0) \rangle \rangle &= \sum_{x,y} \psi_x^k \psi_y^j \langle \tilde{p}_x(0) \tilde{p}_y(0) \rangle = \frac{1}{\beta} \sum_x m_x \psi_x^k \psi_x^j \\ &= \frac{\delta(k-j)}{\beta}. \end{aligned}$$

Hence we have that

$$\frac{\langle \tilde{p}_x^2(t) \rangle}{2m_x} = \frac{1}{2\beta} \sum_k (\cos^2 \omega_k t + \sin^2 \omega_k t) m_x (\psi_x^k)^2 = \frac{1}{2\beta}$$

and similarly  $\langle \tilde{r}_x^2(t) \rangle / 2 = \frac{1}{2\beta}$ .

## 5. EIGENMODES EXPANSION: INTEGRABILITY, LOCALIZATION

We describe an explicit solution to the equations of motion (2.1) in terms of the eigenmodes of the system. This representation is useful to establish the integrability of the system and to exploit the localization at all energies above the ground states (in the thermodynamic limit).

**5.1. Solution to the equations of motion.** From (2.1), one can deduce second order equations for  $r$  and  $p$  separately:

$$\ddot{r}_x = (\nabla_+ M^{-1} \nabla_- r)_x \quad (1 \leq x \leq N-1), \quad \ddot{p}_x = (\Delta M^{-1} p)_x \quad (1 \leq x \leq N),$$

where, besides the boundary conditions  $r_0 = r_N = 0$ , we have assumed free boundary conditions for  $M^{-1}p$ , i.e.  $m_0^{-1}p_0 = m_1^{-1}p_1$  and  $m_{N+1}^{-1}p_{N+1} = m_N^{-1}p_N$ . Notice that there are two different vector spaces: a  $(N-1)$ -dimensional space for  $r$  with fixed boundary conditions, and a  $N$ -dimensional space for  $M^{-1}p$  with free boundary conditions. Moreover, we observe that  $\nabla_+ = -(\nabla_-)^\dagger$  with fixed boundary conditions, and that  $\Delta = \Delta^\dagger$  with free boundary conditions.

In order to solve the equations of motion, we need to diagonalize two matrices:  $(\nabla_+ M^{-1} \nabla_-)^\dagger = \nabla_+ M^{-1} \nabla_-$  (of size  $N-1$ ) and  $(\Delta M^{-1})^\dagger = M^{-1} \Delta$  (of size  $N$ ). First, the matrix  $\nabla_+ M^{-1} \nabla_-$  is symmetric and non-negative. Hence there exists an orthonormal set of eigenvectors, that we denote by  $|\tilde{\psi}^k\rangle$  for  $1 \leq k \leq N-1$  and corresponding eigenvalues that we denote by  $\tilde{\omega}_k^2$  (below we will connect this basis to another one and drop the notation with tilde most of the time). Second, the matrix  $M^{-1} \Delta$  is not symmetric but the matrix  $M^{-1/2} (-\Delta) M^{-1/2}$  is symmetric and non-negative. It admits thus an orthonormal set of eigenvectors,  $\{\varphi^k\}_{0 \leq k \leq N-1}$  and corresponding eigenvalues  $\omega_k^2$ . Moreover, the spectrum is P-almost surely non-degenerate (see e.g. Proposition

II.1 in [17], considering here a perturbation around the non-degenerate equal masses case). Therefore the vectors  $\psi^k = M^{-1/2}\varphi^k$  are such that

$$M^{-1}(-\Delta)\psi^k = \omega_k^2\psi^k, \quad \langle \psi^j, M\psi^k \rangle = \delta(j-k). \quad (5.1)$$

Because of free boundary conditions,  $\omega_0^2 = 0$ , and

$$\psi_0 = \left( \sum_x m_x \right)^{-1/2} (1, \dots, 1)^\dagger.$$

We can now connect  $\tilde{\omega}_k, \tilde{\psi}^k$  to  $\omega_k, \psi^k$ :

$$\tilde{\psi}^k = \frac{1}{\omega_k} \nabla_+ \psi^k, \quad \tilde{\omega}_k = \omega_k$$

for  $1 \leq k \leq N-1$ . We observe that, by the free boundary conditions on  $\psi^k$ ,  $\tilde{\psi}^k(0) = \tilde{\psi}^k(N) = 0$ .

Given initial conditions  $r(0), p(0)$ , we can write an explicit solution for  $r(t), p(t)$ :

$$\langle \tilde{\psi}^k, \ddot{r} \rangle = -\omega_k^2 \langle \tilde{\psi}^k, r \rangle \quad (1 \leq k \leq N-1), \quad \langle \psi^k, \ddot{p} \rangle = -\omega_k^2 \langle \psi^k, p \rangle \quad (0 \leq k \leq N-1).$$

Thus

$$\begin{aligned} \langle \tilde{\psi}^k, r(t) \rangle &= \langle \tilde{\psi}^k, r(0) \rangle \cos \omega_k t + \frac{\langle \tilde{\psi}^k, \nabla_+ M^{-1} p(0) \rangle}{\omega_k} \sin \omega_k t \quad (1 \leq k \leq N-1), \\ \langle \psi^k, p(t) \rangle &= \langle \psi^k, p(0) \rangle \cos \omega_k t + \frac{\langle \nabla_- r(0), \psi^k \rangle}{\omega_k} \sin \omega_k t \quad (0 \leq k \leq N-1) \end{aligned}$$

with the convention  $\frac{\sin 0}{0} = 1$  in the second expression at  $k=0$  (notice that  $\langle \nabla_- r(0), \psi^k \rangle = -\langle r(0), \nabla_+ \psi^k \rangle = 0$  for  $k=0$ ). This yields therefore

$$r(t) = \sum_{k=1}^{N-1} \left( \frac{\langle \nabla_+ \psi^k, r(0) \rangle}{\omega_k} \cos \omega_k t + \langle \psi^k, p(0) \rangle \sin \omega_k t \right) \frac{\nabla_+ \psi^k}{\omega_k}, \quad (5.2)$$

$$p(t) = \sum_{k=0}^{N-1} \left( \langle \psi^k, p(0) \rangle \cos \omega_k t - \frac{\langle \nabla_+ \psi^k, r(0) \rangle}{\omega_k} \sin \omega_k t \right) M \psi^k. \quad (5.3)$$

**5.2. Full set of invariant quantities.** We observe that the dynamics has  $N$  invariant quantities, corresponding to the energy of each mode. It is thus an integrable system. Indeed, let us write the full energy as

$$\begin{aligned} H &= \frac{1}{2} (\langle p, M^{-1} p \rangle + \langle r, r \rangle) = \frac{1}{2} \sum_{k=0}^{N-1} \langle p, \psi^k \rangle \langle M \psi^k, M^{-1} p \rangle + \frac{1}{2} \sum_{k=1}^{N-1} \langle r, \tilde{\psi}^k \rangle \langle \tilde{\psi}^k, r \rangle \\ &= \frac{1}{2} \sum_{k=0}^{N-1} \left( \langle p, \psi^k \rangle^2 + \frac{\langle r, \nabla_+ \psi^k \rangle^2}{\omega_k^2} \right) \end{aligned}$$

with the convention that the second term in the last sum is 0 at  $k=0$ . From the evolution equation of the dynamics, one gets that actually

$$\frac{d}{dt} \left( \langle p, \psi^k \rangle^2 + \frac{\langle r, \nabla_+ \psi^k \rangle^2}{\omega_k^2} \right) = 0 \quad \text{for all } 0 \leq k \leq N-1. \quad (5.4)$$

Moreover, by taking specific linear combinations of these conserved quantities, one can obtain conserved quantities that can be written as a sum of local terms. This is for instance the case of the quantity  $I$  defined in (3.2), that reads also

$$I(r, p) = \frac{1}{2} \sum_{k=1}^{N-1} \omega_k^2 \left( \langle p, \psi^k \rangle^2 + \frac{\langle r, \nabla_+ \psi^k \rangle^2}{\omega_k^2} \right).$$

**5.3. Localization.** Localization can be expressed mathematically in the following strong sense, see [17, 1, 2] for the general theory and [21, 3] for precise estimates on the localization length as one approaches the ground state. Let  $0 < \alpha < \frac{1}{2}$  and let  $I(\alpha) = ]N^{(1-\alpha)}, N-1] \cap \mathbb{Z}$ . There exist constants  $C, c > 0$  such that

$$\mathbb{E} \left( \sum_{k \in I(\alpha)} |\psi_x^k \psi_y^k| \right) \leq C e^{-c|x-y|/\xi(\alpha)} \quad \text{with} \quad \xi(\gamma) = N^{2\alpha}.$$

We will use this estimate to show that every mode in  $k \in I(\alpha)$  is supported in a small subset of  $[1, N] \cap \mathbb{Z}$  up to a small error:

**Lemma 3.** *Let  $\alpha, \gamma > 0$  be such that  $2\alpha < \gamma < 1$ . There exists almost surely  $N_0 \in \mathbb{N}$  so that for all  $N \geq N_0$ , and for all  $k \in I(\alpha)$ , there exists an interval  $J(k)$  with  $|J(k)| \leq 2N^\gamma$  such that  $|\psi_x^k| \leq N^{-1/\gamma}$  for all  $x \notin J(k) \cap \mathbb{Z}$ .*

*Proof.* Let us first show that

$$\begin{aligned} P &:= \mathbb{P}(\exists k \in I(\alpha), \exists x, y \in [1, N] \cap \mathbb{Z} : |x-y| \geq N^\gamma, |\psi_x^k| \geq N^{-1/\gamma}, |\psi_y^k| \geq N^{-1/\gamma}) \\ &\leq C(\alpha, \gamma) e^{-N^{(\gamma-2\alpha)/2}} \end{aligned} \quad (5.5)$$

Indeed we compute

$$\begin{aligned} P &\leq \sum_{k \in I(\alpha)} \sum_{x, y : |x-y| \geq N^\gamma} \mathbb{P}(|\psi_x^k| \geq N^{-1/\gamma}, |\psi_y^k| \geq N^{-1/\gamma}) \\ &\leq \sum_{k \in I(\alpha)} \sum_{x, y : |x-y| \geq N^\gamma} \mathbb{P}(|\psi_x^k \psi_y^k| \geq N^{-2/\gamma}) \\ &\leq \sum_{k \in I(\alpha)} \sum_{x, y : |x-y| \geq N^\gamma} N^{2/\gamma} \mathbb{E}(|\psi_x^k \psi_y^k|) \\ &\leq CN^{2/\gamma} \sum_{x, y : |x-y| \geq N^\gamma} e^{-c|x-y|/\xi(\alpha)} \leq CN^{2/\gamma} N^{2\gamma} e^{-N^{\gamma-2\alpha}} \leq C(\alpha, \gamma) e^{-N^{(\gamma-2\alpha)/2}}. \end{aligned}$$

Therefore, there exists almost surely  $N_0$  so that for for all  $N \geq N_0$ , the event featuring in (5.5) does not occur. Hence in this case, for all  $k \in I(\alpha)$  and all  $|x-y| > N^\gamma$ , we must have either  $|\psi_x^k| \leq 1/N^{1/\gamma}$  or  $|\psi_y^k| \leq 1/N^{1/\gamma}$ . This means thus that for any  $k \in I(\alpha)$  there exists an interval  $J(k)$  with  $|J(k)| \leq 2N^\gamma$  such that  $|\psi_x^k| \leq N^{-1/\gamma}$  for all  $x \notin J(k) \cap \mathbb{Z}$ .  $\square$

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