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YANG-MILLS FIELDS ON THE EXTERIOR OF THE
SCHWARZSCHILD BLACK HOLE**

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INSTABILITY OF INFINITELY-MANY STATIONARY SOLUTIONS OF THE $SU(2)$ YANG-MILLS FIELDS ON THE EXTERIOR OF THE SCHWARZSCHILD BLACK HOLE

DIETRICH HÄFNER, CÉCILE HUNEAU

ABSTRACT. We consider the spherically symmetric $SU(2)$ Yang-Mills fields on the Schwarzschild metric. Within the so called purely magnetic Ansatz we show that there exists a countable number of stationary solutions which are all nonlinearly unstable.

1. INTRODUCTION

1.1. General introduction. We study the $SU(2)$ Yang-Mills equations on the Schwarzschild metric, with spherically symmetric initial data fulfilling the so called purely magnetic Ansatz. This equation has at least a countable number of stationary solutions. Because of energy conservation, the zero curvature solution is stable. In [11] the first author and S. Ghanem show decay estimates for small energy solutions in the exterior of the Schwarzschild black hole within this Ansatz. In this paper we show that the other solutions of the above set of stationary solutions are nonlinearly unstable.

Global existence for Yang-Mills fields on \mathbb{R}^{3+1} was shown by Eardley and Moncrief in a classical result, [6] and [7]. Their result was then generalized by Chruściel and Shatah to general globally hyperbolic curved space-times in [5]. Later, the hypotheses of [5] were weakened in [10].

The purely magnetic Ansatz excludes Coulomb type solutions and reduces the Yang-Mills equations to a nonlinear scalar wave equation:

$$\partial_t^2 W - \partial_x^2 W + \frac{(1 - \frac{2m}{r})}{r^2} W(W^2 - 1) = 0. \quad (1.1)$$

Strong numerical evidence of the existence of a countable number of stationary solutions $(W_n)_{n \in \mathbb{N}}$ in the case of Yang Mills equations coupled with Einstein equations with spherical symmetry was shown in [1] (see also [3]). It was then proved analytically, still in the coupled case, in [18], see also [4]. For sake of completeness, we give an analytical proof of the existence of infinitely many solution to the Yang-Mills equation on Schwarzschild in the appendix of this paper (adapted from [18]). The solution W_n possesses n zeros. The stationary solutions $W_0 = \pm 1$ correspond to the zero curvature solution. Linearizing around a stationary solution W_n leads

to the linear operator

$$\mathcal{A}_n = -\partial_x^2 + \frac{(1 - \frac{2m}{r})}{r^2}(3W_n^2 - 1).$$

In [3] it was numerically observed for the first stationary solutions that \mathcal{A}_n has n negative eigenvalues. In this paper we show analytically that \mathcal{A}_n has at least one negative eigenvalue for $n \geq 1$. Writing the equation as a first order equation one then observes that the spectrum of the linear part meets $\{\text{Re}\lambda > 0\}$. As already observed for example in [16] this leads to nonlinear instability. We will describe in Section 2 a general abstract setting for non linear one dimensional wave equations. Adapting ideas from [14] we find similar results to those obtained by [9] and [16] but in slightly different spaces. This abstract setting is applied in Section 3 to the Yang-Mills equation. This section contains in particular the proof of the existence of a negative eigenvalue for \mathcal{A}_n , $n \geq 1$. The difficulty consists in the fact that the solution W_n and therefore the potential V_n in the operator \mathcal{A}_n is not explicitly known. Nevertheless we are able to show that $V_n \leq F$ for some explicit F with $\int F < 0$. This then gives the existence of at least one negative eigenvalue.

Instability for similar solutions of the Einstein-Yang-Mills system has been investigated in [20], without the purely magnetic ansatz. In the setting of [20], the Yang-Mills field is given by four components instead of one in the purely-magnetic case (see the formula (1.8)). The instability shown analytically in [20] comes from perturbations which are in the "odd parity sector", which, in the notation of our paper corresponds to perturbations for the components W_2, A_0, A_1 of the Yang-Mills field, letting the W_1 component frozen. Consequently it does not imply the instability in the purely magnetic ansatz.

1.2. The exterior of the Schwarzschild black hole. The exterior Schwarzschild spacetime is given by $\mathcal{M} = \mathbb{R}_t \times \mathbb{R}_{r>2m} \times S^2$ equipped with the metric

$$\begin{aligned} g &= -(1 - \frac{2m}{r})dt^2 + \frac{1}{(1 - \frac{2m}{r})}dr^2 + r^2 d\theta^2 + r^2 \sin^2(\theta)d\phi^2 \\ &= N(-dt^2 + dx^2) + r^2 d\sigma^2 \end{aligned}$$

where

$$N = (1 - \frac{2m}{r}) \tag{1.2}$$

and $d\sigma^2$ is the usual volume element on the sphere. The coordinate x is defined by the requirement

$$\frac{dx}{dr} = N^{-1}.$$

The coordinates t, r, θ, ϕ , are called Boyer-Lindquist coordinates. The singularity $r = 2m$ is a coordinate singularity and can be removed by changing coordinates, see [13]. m is the mass of the black hole. We will only be interested in the region outside the black hole, $r > 2m$.

1.3. The spherically symmetric $SU(2)$ Yang-Mills equations on the Schwarzschild metric. Let $G = SU(2)$, the real Lie group of 2×2 unitary matrices of determinant 1. The Lie algebra associated to G is $su(2)$, the antihermitian traceless 2×2 matrices. Let τ_j , $j \in \{1, 2, 3\}$, be the following real basis of $su(2)$:

$$\tau_1 = \frac{i}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \quad \tau_2 = \frac{1}{2} \begin{pmatrix} 0 & -1 \\ 1 & 0 \end{pmatrix}, \quad \tau_3 = \frac{i}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

Note that

$$[\tau_1, \tau_2] = \tau_3, \quad [\tau_3, \tau_1] = \tau_2, \quad [\tau_2, \tau_3] = \tau_1.$$

We are looking for a connection A , that is a one form with values in the Lie algebra $su(2)$ associated to the Lie group $SU(2)$, which satisfies the Yang-Mills equations which are:

$$\mathbf{D}_\alpha^{(A)} F^{\alpha\beta} \equiv \nabla_\alpha F^{\alpha\beta} + [A_\alpha, F^{\alpha\beta}] = 0, \quad (1.3)$$

where $[\cdot, \cdot]$ is the Lie bracket and $F_{\alpha\beta}$ is the Yang-Mills curvature given by

$$F_{\alpha\beta} = \nabla_\alpha A_\beta - \nabla_\beta A_\alpha + [A_\alpha, A_\beta], \quad (1.4)$$

and where we have used the Einstein raising indices convention with respect to the Schwarzschild metric. We also have the Bianchi identities which are always satisfied in view of the symmetries of the Riemann tensor and the Jacobi identity for the Lie bracket:

$$\mathbf{D}_\alpha^{(A)} F_{\mu\nu} + \mathbf{D}_\mu^{(A)} F_{\nu\alpha} + \mathbf{D}_\nu^{(A)} F_{\alpha\mu} = 0. \quad (1.5)$$

The Cauchy problem for the Yang-Mills equations formulates as the following: given a Cauchy hypersurface Σ in M , and a \mathcal{G} -valued one form A_μ on Σ , and a \mathcal{G} -valued one form E_μ on Σ satisfying

$$\left. \begin{aligned} E_t &= 0, \\ \mathbf{D}_\mu^{(A)} E^\mu &= 0 \end{aligned} \right\} \quad (1.6)$$

we are looking for a \mathcal{G} -valued two form $F_{\mu\nu}$ satisfying the Yang-Mills equations such that once $F_{\mu\nu}$ restricted to Σ we have

$$F_{\mu t} = E_\mu \quad (1.7)$$

and such that $F_{\mu\nu}$ corresponds to the curvature derived from the Yang-Mills potential A_μ , i.e. given by (1.4). Equations (1.6) are the Yang-Mills constraints equations on the initial data.

Any spherically symmetric Yang-Mills potential can be written in the following form after applying a gauge transformation, see [8], [12] and [21],

$$\begin{aligned} A &= [-W_1(t, r)\tau_1 - W_2(t, r)\tau_2]d\theta + [W_2(t, r)\sin(\theta)\tau_1 - W_1(t, r)\sin(\theta)\tau_2]d\phi \\ &\quad + \cos(\theta)\tau_3 d\phi + A_0(t, r)\tau_3 dt + A_1(t, r)\tau_3 dr, \end{aligned} \quad (1.8)$$

where $A_0(t, r)$, $A_1(t, r)$, $W_1(t, r)$, $W_2(t, r)$ are arbitrary real functions. We consider here a purely magnetic Ansatz in which we have $A_0 = A_1 = W_2 = 0$, $W_1 =: W$.

The components of the curvature are then

$$\left. \begin{aligned} F_{\theta x} &= W' \tau_1, \\ F_{\theta t} &= \dot{W} \tau_1, \\ F_{\phi x} &= W' \sin(\theta) \tau_2, \\ F_{\phi t} &= \dot{W} \sin(\theta) \tau_2, \\ F_{tx} &= 0, \\ F_{\theta\phi} &= (W^2 - 1) \sin(\theta) \tau_3. \end{aligned} \right\}$$

This kind of Ansatz is preserved by the evolution. Also the principal restriction is $A_0 = A_1 = 0$. The constraint equations then impose that W_1 is proportional to W_2 , a case which can be reduced to $W_2 = 0$. We refer the reader to [11] for details.

1.4. The initial value problem for the purely magnetic Ansatz. We look at initial data prescribed on $t = 0$ where there exists a gauge transformation such that once applied on the initial data, the potential A can be written in this gauge as

$$\left. \begin{aligned} A_t(t=0) &= 0, \\ A_r(t=0) &= 0, \\ A_\theta(t=0) &= -W_0(r) \tau_1, \\ A_\phi(t=0) &= -W_0(r) \sin(\theta) \tau_2 + \cos(\theta) \tau_3, \end{aligned} \right\} \quad (1.9)$$

and, we are given in this gauge the following one form E_μ on $t = 0$:

$$\left. \begin{aligned} E_\theta(t=0) &= F_{\theta t}(0) = W_1(r) \tau_1, \\ E_\phi(t=0) &= F_{\phi t}(0) = W_1(r) \sin(\theta) \tau_2, \\ E_r(t=0) &= F_{rt}(0) = 0, \\ E_t(t=0) &= F_{tt}(t=0) = 0. \end{aligned} \right\} \quad (1.10)$$

Notice that with this Ansatz the constraint equations (1.6) are automatically fulfilled

$$(\mathbf{D}^{(A)\theta} E_\theta + \mathbf{D}^{(A)\phi} E_\phi + \mathbf{D}^{(A)r} E_r)(t=0) = 0.$$

The Yang-Mills equations now reduce to

$$\left. \begin{aligned} \ddot{W} - W'' + PW(W^2 - 1) &= 0, \\ W(0) &= W_0, \\ \partial_t W(0) &= W_1, \end{aligned} \right\} \quad (1.11)$$

where

$$P = \frac{(1 - \frac{2m}{r})}{r^2}.$$

It is easy to check that the following energy is conserved, see also [11],

$$\mathcal{E}(W, \dot{W}) = \int \dot{W}^2 + (W')^2 + \frac{P}{2}(W^2 - 1)^2 dx.$$

We note by $\dot{H}^k = \dot{H}^k(\mathbb{R}, dx)$ and $H^k = H^k(\mathbb{R}, dx)$, the homogeneous and inhomogeneous Sobolev spaces of order k , respectively.

Definition 1.1. (1) We define the spaces L_P^4 , resp. L_P^2 , as the completion of $C_0^\infty(\mathbb{R})$ for the norm

$$\|v\|_{L_P^4}^4 := \int P|v|^4 dx \quad \text{resp.} \quad \|v\|_{L_P^2}^2 := \int P|v|^2 dx. \quad (1.12)$$

(2) We also define for $1 \leq k \leq 2$ the space \mathcal{H}^k as the completion of $C_0^\infty(\mathbb{R})$ for the norm

$$\|u\|_{\mathcal{H}^k}^2 = \|u\|_{H^k}^2 + \|u\|_{L_P^4}^2. \quad (1.13)$$

We note that \mathcal{H}^k is a Banach space which contains all constant functions. It turns out that $\mathcal{E} := \mathcal{H}^1 \times L^2$ is exactly the space of finite energy solutions, see [11] for details. We then have [11, Theorem 1]

Theorem 1. *Let $(W_0, W_1) \in \mathcal{H}^2 \times H^1$. Then there exists a unique strong solution of (1.11) with*

$$\begin{aligned} W &\in C^1([0, \infty); \mathcal{H}^1) \cap C([0, \infty); \mathcal{H}^2), \\ \partial_t W &\in C^1([0, \infty); L^2) \cap C([0, \infty); H^1), \\ \sqrt{P}(W^2 - 1) &\in C^1([0, \infty); L^2) \cap C([0, \infty); H^1). \end{aligned}$$

We can reformulate the above theorem in the following way

Corollary 1.1. *We suppose that the initial data for the Yang-Mills equations is given after suitable gauge transformation by*

$$\left. \begin{aligned} A_t(0) &= A_r(0) = 0, \\ A_\theta(0) &= -W_0 \tau_1, \\ A_\phi(0) &= -W_0 \sin \theta \tau_2 + \cos \theta \tau_3, \\ E_\theta(0) &= W_1 \tau_1, \\ E_\phi(0) &= W_1 \sin \theta \tau_2, \\ E_r(0) &= E_t(0) = 0 \end{aligned} \right\}$$

with $(W_0, W_1) \in \mathcal{H}^2 \times H^1$. Then, the Yang-Mills equation (1.3) admits a unique solution F with

$$F_{\theta x}, \frac{1}{\sin \theta} F_{\phi x}, F_{\theta t}, \frac{1}{\sin \theta} F_{\phi t}, \sqrt{P} \frac{1}{\sin \theta} F_{\theta \phi} \in C^1([0, \infty); L^2) \cap C([0, \infty); H^1).$$

1.5. Energies. We now introduce the Yang-Mills energy momentum tensor

$$T_{\mu\nu} = \langle F_{\mu\beta}, F_\nu^\beta \rangle - \frac{1}{4} g_{\mu\nu} \langle F_{\alpha\beta}, F^{\alpha\beta} \rangle.$$

Here $\langle \cdot, \cdot \rangle$ is an Ad-invariant scalar product on the Lie algebra $su(2)$. We have

$$\nabla^\nu T_{\mu\nu} = 0.$$

For a vector field X^ν we define

$$J_\mu(X) = X^\nu T_{\mu\nu}$$

and the energy on the spacelike slice Σ_t ($\Sigma_{t_0} = \{t = t_0\}$) by

$$E^{(X)}(F(t)) = \int_{\Sigma_t} J_\mu(X) n^\nu d\Sigma_t.$$

By the divergence theorem this energy is conserved if X is Killing. In particular

$$E^{(\partial_t)}(F(t)) = \int_{\Sigma_t} J_\mu(\partial_t)n^\mu d\Sigma_t$$

is conserved. If F is the curvature associated to (W, \dot{W}) , then

$$E^{(\partial_t)}(F(t)) = \mathcal{E}(W, \dot{W}),$$

see [11] for details.

1.6. Main result. We first recall the following result which is implicit in the paper [3] of P. Bizoń, A. Rostworowski and A. Zenginoglu.

Theorem 2. *There exists a decreasing sequence $\{a_n\}_{n \in \mathbb{N} \geq 1}$, $0 < \dots < a_n < a_{n-1} < \dots < a_1 = \frac{1+\sqrt{3}}{3\sqrt{3}+5}$ and smooth stationary solutions W_n of (1.11) with*

$$-1 \leq W_n \leq 1, \quad \lim_{x \rightarrow -\infty} W_n(x) = a_n, \quad \lim_{x \rightarrow \infty} W_n(x) = (-1)^n.$$

The solution W_n has exactly n zeros.

Remark 1.1. There is an explicit formula for the first stationary solution (see [2])

$$W_1 = \frac{c - \frac{r}{2m}}{\frac{r}{2m} + 3(c-1)}, \quad c = \frac{3 + \sqrt{3}}{2}.$$

This solution corresponds to $\lim_{x \rightarrow -\infty} W_1(x) = a_1 = \frac{1+\sqrt{3}}{3\sqrt{3}+5}$.

We give a detailed proof of this result in the appendix, where we follow arguments of Smoller, Wasserman, Yau and McLeod. The above solutions are all nonlinearly unstable :

Theorem 3 (Main Theorem). *For all $n \geq 1$ the solution W_n of (1.11) is unstable. More precisely there exists $\epsilon_0 > 0$ and a sequence $(W_{0,n}^m, W_{1,n}^m)$ with $\|(W_{0,n}^m, W_{1,n}^m) - (W_n, 0)\|_{\mathcal{E}} \rightarrow 0$, $m \rightarrow \infty$, but for all m*

$$\sup_{t \geq 0} \|(W_n^m(t), \partial_t W_n^m(t)) - (W_n, 0)\|_{\mathcal{E}} \geq \epsilon_0 > 0.$$

Remark 1.2. We don't show in this paper that there is no stationary solution with $W(2m) > a_1$. We do not exclude either the fact that there may exist solutions with an infinite number of zeros which tend to zero at infinity. Our main theorem does not apply to this two categories of hypothetical stationary solutions.

For n given we construct initial data from W_n as in Section 1.4. Let F_n be the corresponding curvature at time $t = 0$. We obtain

Corollary 1.2. *For all $n \geq 1$ the solution F_n of (1.3) is unstable. More precisely there exists $\epsilon_0 > 0$ and a sequence of initial data giving rise to the curvature $F_{0,n}^m$ with*

$$E^{(\partial_t)}(F_{0,n}^m - F_n) \rightarrow 0, \quad m \rightarrow \infty,$$

but for all m

$$\sup_{t \geq 0} E^{(\partial_t)}(F_n^m(t) - F_n) \geq \epsilon_0,$$

where $F_n^m(t)$ is the solution associated to the initial data corresponding to the curvature $F_{0,n}^m$.

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2. ABSTRACT SETTING

In this section we describe the general abstract framework. We first write the abstract wave equation as a first order equation. The properties of the potential then give rise to spectral properties of the first order operator, see Sections 2.2-2.3. In Section 2.4 we then adapt arguments of Henry [14] to our situation. This gives results close to those obtained in [9], [16]. These results then have to be adapted to slightly different spaces in Section 3.2, see Remark 2.5.

2.1. Abstract result. We consider the one dimensional wave equation

$$\begin{cases} \ddot{u} - u'' + Vu &= F(u), \\ u|_{t=0} &= u_0, \\ \partial_t u|_{t=0} &= u_1 \end{cases} \quad (2.1)$$

with $\dot{\cdot} = \partial_t$, $' = \partial_x$ and

$$V \in C(\mathbb{R}) \cap L^1(\mathbb{R}), \quad \lim_{|x| \rightarrow \infty} V(x) = 0, \quad \int_{\mathbb{R}} V(x) dx < 0. \quad (\text{HV})$$

We also suppose that

$$\|F(u) - F(v)\|_{L^2} \leq M_F(\|u\|_{H^1} + \|v\|_{H^1})\|u - v\|_{H^1} \quad (\text{HF})$$

for $\|u\|_{H^1} \leq 1, \|v\|_{H^1} \leq 1$. Let $X = H^1 \times L^2$. We then have the following

Theorem 4. *The zero solution of (2.1) is unstable. More precisely there exists $\epsilon_0 > 0$ and a sequence (u_0^m, u_1^m) with $\|(u_0^m, u_1^m)\|_X \rightarrow 0, m \rightarrow \infty$, but for all m*

$$\sup_{t \geq 0} \|(u^m(t), \partial_t u^m(t))\|_X \geq \epsilon_0 > 0.$$

Here $u^m(t)$ is the solution of (2.1) with initial data (u_0^m, u_1^m) and the supremum is taken over the maximal interval of existence of $u^m(t)$.

Let

$$\mathcal{A} = -\partial_x^2 + V, \quad D(\mathcal{A}) = H^2(\mathbb{R}).$$

We note that \mathcal{A} is a selfadjoint operator.

2.2. Spectral analysis of \mathcal{A} .

Proposition 2.1. *We have*

$$\sigma(\mathcal{A}) = \{-\lambda_n^2\}_{n \in \mathcal{N}} \cup [0, \infty),$$

where $-\lambda_n^2$, $\lambda_0 > \lambda_1 > \dots > \lambda_n > \dots > 0$ is a finite ($\mathcal{N} = \{0, \dots, N\}$) or infinite ($\mathcal{N} = \mathbb{N}$) sequence of negative eigenvalues with only possible accumulation point 0.

Proof. First note that $\sigma(\mathcal{A}) \cap \mathbb{R}^- \neq \emptyset$. Indeed let $\chi \in C_0^\infty(\mathbb{R})$, $\chi(0) = 1$, $\chi \geq 0$, $\chi_R(\cdot) = \chi(\frac{\cdot}{R})$. Then

$$\langle \mathcal{A}\chi_R, \chi_R \rangle = \frac{1}{R} \int |\chi'(x)|^2 dx + \int V(x)\chi_R^2 dx \rightarrow \int V(x) dx < 0, \quad R \rightarrow \infty.$$

We now introduce the comparison operator

$$\mathcal{B} = -\partial_x^2.$$

We compute

$$(\mathcal{B} - z^2)^{-1} - (\mathcal{A} - z^2)^{-1} = (\mathcal{A} - z^2)^{-1} V (\mathcal{B} - z^2)^{-1}.$$

Using that $\lim_{x \rightarrow \pm\infty} V(x) = 0$ we see that this is a compact operator. By the Weyl criterion

$$\sigma_{ess}(\mathcal{A}) = \sigma_{ess}(\mathcal{B}) = [0, \infty).$$

On the other hand we already know that \mathcal{A} has negative spectrum. It therefore has at least one negative eigenvalue. \mathcal{A} being bounded from below the proposition follows. \square

2.3. The wave equation as a first order equation.

2.3.1. *The linear equation.* The equation

$$\ddot{v} + \mathcal{A}v = 0$$

is equivalent to

$$\partial_t \psi = L\psi, \quad L = \begin{pmatrix} 0 & i \\ i\mathcal{A} & 0 \end{pmatrix}, \quad \psi = \begin{pmatrix} v \\ \frac{1}{i}\partial_t v \end{pmatrix}.$$

Remark 2.1. Let

$$\mathcal{A}\phi_0 = -\lambda^2\phi_0.$$

Then we have

(1) $\phi_0 \in H^2$.

(2) Let $\psi_0^\pm = \begin{pmatrix} \phi_0 \\ \pm \frac{1}{i}\lambda\phi_0 \end{pmatrix}$. Then

$$L\psi_0^\pm = \pm\lambda\psi_0^\pm.$$

Let V_- be the negative part of the potential. For $\mu^2 > \|V_-\|_\infty (\geq \lambda_0^2)$ we introduce the scalar product

$$\langle u, v \rangle_\mu = \langle (\mathcal{A} + \mu^2)u_0, v_0 \rangle + \langle u_1, v_1 \rangle$$

where $\langle \cdot, \cdot \rangle$ is the usual scalar product on $\mathcal{H} = L^2(\mathbb{R})$. We note $\|\cdot\|_\mu$ the corresponding norm. It is easy to check that the norms $\|\cdot\|_\mu$ and $\|\cdot\|_X$ are equivalent.

Proposition 2.2. *L is the generator of a C^0 - semigroup e^{tL} on X .*

Proof. Let $\mu^2 > \|V_-\|$ and

$$L_\mu = \begin{pmatrix} 0 & i \\ i(\mathcal{A} + \mu^2) & 0 \end{pmatrix}, \quad B_\mu = \begin{pmatrix} 0 & 0 \\ -i\mu^2 & 0 \end{pmatrix}.$$

iL_μ is a selfadjoint operator on $(X, \langle \cdot, \cdot \rangle_\mu)$ and in particular the generator of a C^0 - semigroup $e^{L_\mu t}$. We have $L = L_\mu + B_\mu$. B_μ being bounded, we can apply [15, Theorem 3.1.1] to see that L is the generator of a C^0 - semigroup on $(X, \|\cdot\|_\mu)$ and thus on $(X, \|\cdot\|_X)$. \square

Let now

$$M_i = \begin{pmatrix} \mathbb{1} & \mathbb{1} \\ \frac{\lambda_i}{i} & -\frac{\lambda_i}{i} \end{pmatrix}.$$

Note that $\det M_i = 2i\lambda_i \neq 0$ and that M_i is thus invertible. We define $P_i = \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})M_i$ and $X_i = P_i X$. We also define $X_\infty = \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\mathbb{1}_2 X$. Here $\mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})$ and $\mathbb{1}_{\mathbb{R}^+}(\mathcal{A})$ are defined by the spectral theorem. In particular $\mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})$ is the projection on the eigenspace of \mathcal{A} associated to the eigenvalue $-\lambda_i^2$.

Lemma 2.1.

$$X = (\oplus_{i \in \mathcal{N}} X_i) \oplus X_\infty.$$

Remark 2.2. Note that the sum is orthogonal with respect to the scalar product $\langle \cdot, \cdot \rangle_\mu$.

Proof. Let $(\phi, \psi) \in X$. We put

$$\phi_i = \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})\phi, \quad \psi_i = \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})\psi, \quad \begin{pmatrix} \tilde{\phi}_i \\ \tilde{\psi}_i \end{pmatrix} = M_i^{-1} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix}.$$

Since \mathcal{A} is self-adjoint, we can write

$$\phi = \sum_{i \in \mathcal{N}} \phi_i + \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\phi, \quad \psi = \sum_{i \in \mathcal{N}} \psi_i + \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\psi.$$

Then

$$\begin{pmatrix} \phi \\ \psi \end{pmatrix} = \sum_{i \in \mathcal{N}} M_i \begin{pmatrix} \tilde{\phi}_i \\ \tilde{\psi}_i \end{pmatrix} + \begin{pmatrix} \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\phi \\ \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\psi \end{pmatrix}$$

gives the required decomposition. For uniqueness let

$$\begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} = \sum_{i \in \mathcal{N}} M_i \begin{pmatrix} \tilde{\phi}_i \\ \tilde{\psi}_i \end{pmatrix} + \begin{pmatrix} \phi_\infty \\ \psi_\infty \end{pmatrix}$$

Applying $\mathbb{1}_{\mathbb{R}^+}(\mathcal{A})$, $\mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})$ to each line immediately gives

$$\phi_\infty = \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\phi, \quad \psi_\infty = \mathbb{1}_{\mathbb{R}^+}(\mathcal{A})\psi, \quad \begin{pmatrix} \tilde{\phi}_i \\ \tilde{\psi}_i \end{pmatrix} = M_i^{-1} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix},$$

where $\psi_i = \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})\psi$, $\phi_i = \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A})\phi$. \square

Let

$$X_i^\pm = M_i \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A}) P_\pm X,$$

where $P_+(\phi, \psi) = (\phi, 0)$, $P_-(\phi, \psi) = (0, \psi)$. Clearly $X_i = X_i^+ \oplus X_i^-$ and thus

$$X = \left(\bigoplus_{i \in \mathcal{N}} (X_i^+ \oplus X_i^-) \right) \oplus X_\infty.$$

Remark 2.3. Let $(\phi_i, \psi_i) \in X_i^\pm$. Then $L(\phi_i, \psi_i) = \pm \lambda_i(\phi_i, \psi_i)$.

Remark 2.4. On X_i the norm $\|\cdot\|_{\sqrt{2}\lambda_i}$ is equivalent to the norm $\|\cdot\|_X$ and X_i^+ , X_i^- are orthogonal with respect to this scalar product. Indeed :

$$\left\langle \begin{pmatrix} \phi \\ \frac{\lambda_i}{i} \phi \end{pmatrix}, \begin{pmatrix} \psi \\ -\frac{\lambda_i}{i} \psi \end{pmatrix} \right\rangle_{\sqrt{2}\lambda_i} = \lambda_i^2 \langle \phi, \psi \rangle - \lambda_i^2 \langle \phi, \psi \rangle = 0.$$

Proposition 2.3. (1) *The spaces X_i , X_∞ are e^{tL} invariant.*

(2) *For all $\epsilon > 0$ there exists $C_\epsilon > 0$ such that for all $i \in \mathcal{N}$ and for all $t \in \mathbb{R}$*

$$\|e^{tL}|_{X_i}\|_{X \rightarrow X} \leq C_\epsilon e^{(\lambda_i + \epsilon)|t|}.$$

(3) *For all $\epsilon > 0$ there exists $C_\epsilon > 0$ such that for all $t \in \mathbb{R}$*

$$\|e^{tL}|_{X_\infty}\|_{X \rightarrow X} \leq C_\epsilon e^{\epsilon|t|}.$$

Proof. (1) We have

$$e^{tL} M_i \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A}) \mathbb{1}_2 \begin{pmatrix} \phi \\ \psi \end{pmatrix} = M_i \mathbb{1}_{\{-\lambda_i^2\}}(\mathcal{A}) \mathbb{1}_2 \begin{pmatrix} e^{t\lambda_i} \phi \\ e^{-t\lambda_i} \psi \end{pmatrix}$$

and thus X_i is invariant under the evolution. The fact that X_∞ is invariant follows from the fact that $\mathbb{1}_{\mathbb{R}^+}(\mathcal{A})$ commutes with L .

(2) Because of the equivalence of the norms it is sufficient to estimate the $\|\cdot\|_\mu$ norm. Let

$$\begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} \in X_i.$$

We compute

$$\begin{aligned} \left\| e^{tL} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} \right\|_\mu^2 &= \left\| \begin{pmatrix} (\mu^2 - \lambda_i^2)^{1/2} & 0 \\ 0 & \mathbb{1} \end{pmatrix} M_i \begin{pmatrix} e^{t\lambda_i} & 0 \\ 0 & e^{-\lambda_i t} \end{pmatrix} M_i^{-1} \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} \right\|_{\mathcal{H} \times \mathcal{H}}^2 \\ &\leq \|N_i\|_{\mathbb{R}^2 \rightarrow \mathbb{R}^2}^2 \left\| \begin{pmatrix} \phi_i \\ \psi_i \end{pmatrix} \right\|_\mu^2, \end{aligned}$$

where

$$N_i = \begin{pmatrix} (\mu^2 - \lambda_i^2)^{1/2} & 0 \\ 0 & \mathbb{1} \end{pmatrix} M_i \begin{pmatrix} e^{t\lambda_i} & 0 \\ 0 & e^{-t\lambda_i} \end{pmatrix} M_i^{-1} \begin{pmatrix} (\mu^2 - \lambda_i^2)^{-1/2} & 0 \\ 0 & \mathbb{1} \end{pmatrix}.$$

We then estimate uniformly in $i \in \mathcal{N}$:

$$\|N_i\|_{\mathbb{R}^2 \rightarrow \mathbb{R}^2}^2 \lesssim \left\| \frac{1}{2} \begin{pmatrix} e^{t\lambda_i} + e^{-t\lambda_i} & \frac{1}{i\lambda_i}(e^{-t\lambda_i} - e^{t\lambda_i}) \\ \frac{\lambda_i}{i}(e^{t\lambda_i} - e^{-t\lambda_i}) & e^{t\lambda_i} + e^{-t\lambda_i} \end{pmatrix} \right\|_2^2.$$

We have for $t \geq 0$

$$\begin{aligned} \frac{1}{\lambda_i}(e^{t\lambda_i} - e^{-t\lambda_i}) &= 2 \sum_{i=1}^{\infty} \frac{(t\lambda_i)^{2i+1}}{\lambda_i(2i+1)!} \\ &\leq 2t \sum_{i=1}^{\infty} \frac{(t\lambda_i)^{2i}}{(2i)!} \leq t(e^{t\lambda_i} + e^{-t\lambda_i}) \leq \tilde{C}_\epsilon e^{(\lambda_i + \epsilon)t}. \end{aligned}$$

Using that $\lambda_i \leq \lambda_0$ we find uniformly in $i \in \mathcal{N}$:

$$\|N_i\|_{\mathbb{R}^2 \rightarrow \mathbb{R}^2} \lesssim e^{(\lambda_i + \epsilon)t}.$$

(3) We consider the case $t \geq 0$. First note that

$$\|u\|_{X_\epsilon}^2 = \langle \mathcal{A}u_0, u_0 \rangle + \|u_1\|^2 + \epsilon^2 \|u_0\|^2$$

defines a norm on X_∞ . We estimate for $u(t) = e^{tL}u$

$$\begin{aligned} \frac{d}{dt} \|u\|_{X_\epsilon}^2 &= 2\operatorname{Re} (\langle \mathcal{A}u_0, \dot{u}_0 \rangle + \langle u_1, \dot{u}_1 \rangle + \epsilon^2 \langle u_0, \dot{u}_0 \rangle) \\ &= 2\operatorname{Re} \epsilon^2 \langle u_0, iu_1 \rangle \\ &\leq 2\epsilon^2 \|u_0\| \|u_1\| \leq \epsilon^3 \|u_0\|^2 + \epsilon \|u_1\|^2 \leq \epsilon \|u\|_{X_\epsilon}^2. \end{aligned}$$

By the Gronwall lemma we obtain:

$$\|u(t)\|_{X_\epsilon}^2 \leq \tilde{C}_\epsilon e^{\epsilon t} \|u\|_{X_\epsilon}^2.$$

We now claim that on X_∞ the X and the X_ϵ norms are equivalent. Indeed

$$\langle \mathcal{A}u_0, u_0 \rangle + \|u_1\|^2 + \epsilon^2 \|u_0\|^2 \lesssim \|u_0\|_{H^1}^2 + \|u_1\|^2.$$

Also,

$$\begin{aligned} \|u_0\|_{H^1}^2 + \|u_1\|^2 &= \langle (-\partial_x^2 + V)u_0, u_0 \rangle - \langle Vu_0, u_0 \rangle + \|u_0\|^2 + \|u_1\|^2 \\ &\lesssim \langle \mathcal{A}u_0, u_0 \rangle + \|u_0\|^2 + \|u_1\|^2 \lesssim \|u\|_{X_\epsilon}^2. \end{aligned}$$

Then we can estimate

$$\|u(t)\|_X \lesssim \|u(t)\|_{X_\epsilon} \lesssim e^{\epsilon t} \|u\|_{X_\epsilon} \lesssim e^{\epsilon t} \|u\|_X.$$

□

Let $Y = X_0^- \oplus \left(\bigoplus_{i=1}^N X_i \right) \oplus X_\infty$. We have $X = X_0^+ \oplus Y$ and both spaces are invariant under e^{tL} .

Corollary 2.1. *For all $\epsilon > 0$ there exists $M_{L,\epsilon} > 0$ such that for all $t \geq 0$ we have*

$$\|e^{tL}|_Y\|_{X \rightarrow X} \leq M_{L,\epsilon} e^{(\lambda_1 + \epsilon)t}.$$

Proof. Because of the equivalence of the norms $\|\cdot\|_X$ and $\|\cdot\|_\mu$ ($\mu^2 > \|V_-\|_\infty$) it is sufficient to show the estimate with respect to the norm $\|\cdot\|_\mu$. We choose $\epsilon < \lambda_1$ and apply Proposition 2.3. Let

$$\phi = \phi_0^- + \sum_{i=1}^N \phi_i + \phi_\infty$$

with $\phi_0^- \in X_0^-$, $\phi_i \in X_i$, $\phi_\infty \in X_\infty$. We have

$$\begin{aligned} \|e^{tL}\phi\|_\mu^2 &= e^{-\lambda_0 t} \|\phi_0^-\|_\mu^2 + \sum_{i=1}^N \|e^{tL}\phi_i\|_\mu^2 + \|\phi_\infty\|_\mu^2 \\ &\lesssim e^{2(\lambda_1+\epsilon)t} (\|\phi_0^-\|_\mu^2 + \sum_{i=0}^N \|\phi_i\|_\mu^2 + \|\phi_\infty\|_\mu^2) = e^{2(\lambda_1+\epsilon)t} \|\phi\|_\mu^2. \end{aligned}$$

□

Let

$$E_0 = \mathbb{1}_{\{-\lambda_0^2\}}(\mathcal{A})M_0P_+M_0^{-1}, \quad E_1 = \mathbb{1} - E_0.$$

We easily check that

$$\forall \psi \in X, E_0\psi \in X_0^+; \quad \forall \psi \in X, E_1\psi \in Y; \quad E_0 + E_1 = \mathbb{1}.$$

2.3.2. The nonlinear equation. The nonlinear equation can be written now as a first order equation

$$\begin{cases} \partial_t \psi &= L\psi + G(\psi), \\ \psi(0) &= \psi_0 \end{cases} \quad (2.2)$$

with

$$G(\psi) = \begin{pmatrix} 0 \\ F(P_+(\psi)) \end{pmatrix}.$$

From hypothesis (HF) we directly obtain

$$\|G(\psi) - G(\phi)\|_X \leq M_F(\|\psi\|_X + \|\phi\|_X)\|\psi - \phi\|_X \quad (2.3)$$

for $\|\psi\|_X \leq 1$, $\|\phi\|_X \leq 1$. The abstract theorem then reads

Theorem 5. *The zero solution of (2.2) is unstable. More precisely there exists $\epsilon_0 > 0$ and a sequence ψ_0^m with $\|\psi_0^m\|_X \rightarrow 0$, $m \rightarrow \infty$, but for all m*

$$\sup_{t \geq 0} \|\psi^m(t)\|_X \geq \epsilon_0 > 0.$$

Here $\psi^m(t)$ is the solution of (2.2) with initial data ψ_0^m and the supremum is taken over the maximal interval of existence of ψ^m .

Remark 2.5. We could in principle apply [9, Theorem 2.1] or [16, Theorem 1], Proposition 2.3 and Corollary 2.1 establish the necessary spectral information of e^{tL} . However the energy space we are working with is not exactly the space which is used for the spectral analysis. We therefore have to adapt the proof to the present situation, see Section 3.2. For the convenience of the reader we repeat in the following two pages the principal arguments in the proofs of the instability theorems. Our proof is an adaption of the proof of [14, Theorem 5.1.3]. Note however that this last theorem cannot be applied directly, because it requires the linear part to be sectorial, which is not the case here.

2.4. Proof of the abstract theorem. We note L_0 the restriction of L to X_0^+ and L_1 the restriction of L to Y . For $\psi_0 \in X_0^+$ with small norm we consider for a certain parameter $\tau > 0$ the integral equation

$$\psi(t) = e^{L_0(t-\tau)}\psi_0 + \int_{\tau}^t e^{L_0(t-s)}E_0G(\psi)ds + \int_{-\infty}^t e^{L_1(t-s)}E_1G(\psi)ds =: \mathcal{I}(\psi). \quad (2.4)$$

We fix $\epsilon > 0$ in Corollary 2.1 small enough such that $\tilde{\lambda}_1 := \lambda_1 + \epsilon < \lambda_0$. We will drop in the following the index ϵ ($M_L = M_{L,\epsilon}$). We fix $\beta > 0$ such that $\lambda_0 > 2\beta > \tilde{\lambda}_1$. Let

$$Z = \{\psi \in C([0, \tau]; X); \|\psi\|_X \leq e^{\beta(t-\tau)}\rho\}.$$

We equip Z with the norm

$$\|\psi\|_Z = \sup_{0 \leq t \leq \tau} \|e^{-\beta(t-\tau)}\psi(t)\|_X.$$

Let ψ_0 such that $\|\psi_0\|_X = \frac{\rho}{3}$. We claim that for ρ small enough

$$\mathcal{I} : \overline{B}_Z(0, \rho) \rightarrow \overline{B}_Z(0, \rho)$$

and that it is a contraction on that space. First note that

$$\mathcal{I}(\psi) = \mathcal{I}_0(\psi) + \mathcal{I}_1(\psi) + \mathcal{I}_2(\psi)$$

with

$$\begin{aligned} \mathcal{I}_0(\psi) &= e^{L_0(t-\tau)}\psi_0, \\ \mathcal{I}_1(\psi) &= - \int_t^\tau e^{L_0(u-\tau)}E_0G(\psi(t+\tau-u))du, \\ \mathcal{I}_2(\psi) &= \int_{-\infty}^t e^{L_1(t-s)}E_1G(\psi(s))ds. \end{aligned}$$

We first estimate for $t \leq \tau$

$$\|\mathcal{I}_0(\psi)\|_X = e^{\lambda_0(t-\tau)}\|\psi_0\|_X \leq 1/3e^{\beta(t-\tau)}\rho.$$

We then estimate for $\psi \in \overline{B}_Z(0, \rho)$

$$\begin{aligned} \|\mathcal{I}_1(\psi)\|_X &\leq M_F\|E_0\| \int_t^\tau e^{\lambda_0(u-\tau)}\|\psi\|_X^2(t+\tau-u)du \\ &\leq M_F\|E_0\| \int_t^\tau e^{\lambda_0(u-\tau)}\rho^2e^{2\beta(t-u)}du \\ &\leq M_F\|E_0\|e^{2\beta t}e^{-\lambda_0\tau}\rho^2 \int_t^\tau e^{(\lambda_0-2\beta)u}du \\ &\leq M_F\|E_0\|\rho^2e^{2\beta t}e^{-\lambda_0\tau} \frac{1}{\lambda_0-2\beta}e^{(\lambda_0-2\beta)\tau} \\ &= \frac{M_F\|E_0\|\rho^2}{\lambda_0-2\beta}e^{2\beta(t-\tau)} \\ &\leq \frac{M_F\|E_0\|\rho^2}{\lambda_0-2\beta}e^{\beta(t-\tau)} \leq 1/3\rho e^{\beta(t-\tau)} \end{aligned}$$

for ρ small enough. We then estimate for $\psi \in \overline{B}_Z(0, \rho)$:

$$\begin{aligned} \|\mathcal{I}_2(\psi(t))\|_X &\leq M_L M_F \|E_1\| \int_{-\infty}^t e^{\tilde{\lambda}_1(t-s)} \rho^2 e^{2\beta(s-\tau)} ds \\ &\leq \frac{M_L M_F \|E_1\| \rho^2}{2\beta - \tilde{\lambda}_1} e^{\tilde{\lambda}_1 t} e^{-2\beta\tau} e^{(2\beta - \tilde{\lambda}_1)t} \\ &= \frac{M_L M_F \|E_1\| \rho^2}{2\beta - \tilde{\lambda}_1} e^{2\beta(t-\tau)} \leq 1/3 \rho^{\beta(t-\tau)} \end{aligned}$$

for ρ small enough. We have just proven $\mathcal{I}(\psi) \in \overline{B}_Z(0, \rho)$. Let us now show that \mathcal{I} is a contraction. We estimate

$$\begin{aligned} \|\mathcal{I}_1(\psi) - \mathcal{I}_1(\phi)\|_X &\leq 2M_F \|E_0\| \int_t^\tau e^{\lambda_0(u-\tau)} \rho e^{\beta(t-u)} \|\psi - \phi\|_X(t + \tau - u) du \\ &\leq 2M_F \|E_0\| \rho \|\psi - \phi\|_Z \int_t^\tau e^{\lambda_0(u-\tau)} e^{2\beta(t-u)} du \\ &= 2M_F \|E_0\| \rho \|\psi - \phi\|_Z e^{2\beta t} e^{-\lambda_0\tau} \int_t^\tau e^{(\lambda_0 - 2\beta)u} du \\ &\leq \frac{2M_F \|E_0\| \rho}{\lambda_0 - 2\beta} e^{2\beta(t-\tau)} \leq 1/4 e^{\beta(t-\tau)} \end{aligned}$$

for ρ sufficiently small. We then estimate

$$\begin{aligned} \|\mathcal{I}_2(\psi) - \mathcal{I}_2(\phi)\|_X &\leq \int_{-\infty}^t 2M_L M_F \|E_1\| \rho e^{\tilde{\lambda}_1(t-s)} e^{\beta(s-\tau)} \|\psi - \phi\|_X ds \\ &\leq 2M_L M_F \|E_1\| \rho \|\psi - \phi\|_Z \int_{-\infty}^t e^{\tilde{\lambda}_1(t-s)} e^{2\beta(s-\tau)} ds \\ &= 2M_L M_F \|E_1\| \rho \|\psi - \phi\|_Z e^{\tilde{\lambda}_1 t} e^{-2\beta\tau} \int_{-\infty}^t e^{(2\beta - \tilde{\lambda}_1)s} ds \\ &\leq \frac{2M_L M_F \|E_1\| \rho}{2\beta - \tilde{\lambda}_1} e^{2\beta(t-\tau)} \leq 1/4 e^{\beta(t-\tau)} \end{aligned}$$

for ρ sufficiently small.

It follows that for ρ sufficiently small there exists a solution of (2.4) in $\overline{B}_Z(0, \rho)$. We note this solution $\psi(t, \tau)$. We easily check that $\psi(t, \tau)$ is also solution of (2.2) with initial data satisfying

$$\|\psi(0, \tau)\|_X \leq \rho e^{-\beta\tau} \rightarrow 0, \tau \rightarrow \infty.$$

We also estimate

$$\begin{aligned} \|\psi(\tau)\|_X &\geq \|\psi_0\|_X - M_L M_F \|E_1\| \int_{-\infty}^\tau e^{\tilde{\lambda}_1(\tau-s)} \rho^2 e^{2\beta(s-\tau)} ds \\ &= \rho/3 - M_L M_F \|E_1\| \rho^2 e^{(\tilde{\lambda}_1 - 2\beta)\tau} \int_{-\infty}^\tau e^{(2\beta - \tilde{\lambda}_1)s} ds \\ &\geq \rho/3 - \frac{M_L M_F \|E_1\| \rho^2}{2\beta - \tilde{\lambda}_1} \geq \rho/6 \end{aligned}$$

for ρ small enough. It follows that $\psi^m(t) = \psi(t, m)$ does the job. \square

3. APPLICATION OF THE ABSTRACT RESULT TO THE YANG-MILLS EQUATION

First note that if $W(t, r)$ is a solution of the Yang-Mills equation (1.11) (written in the r variable), then $W(2mt, 2mr)$ is a solution of the same equation with $m = 1/2$ and vice versa. We can therefore suppose in the following $m = 1/2$. We linearize around $W = W_n$ and obtain for $v = W - W_n$:

$$\ddot{v} - v'' + P(3W_n^2 - 1)v + Pv^2(v + 3W_n) = 0.$$

The linear operator

$$\mathcal{A}_n = -\partial_x^2 + P(3W_n^2 - 1)$$

depends on the stationary solution which we don't know explicitly. We put

$$V_n = P(3W_n^2 - 1).$$

We first want to apply our abstract result on $X = H^1 \times L^2$. It is easy to see that the nonlinear part fulfills the hypotheses of the abstract theorem. Indeed we have

Proposition 3.1. *We have for $\|v\|_{H^1} \leq 1$, $\|u\|_{H^1} \leq 1$:*

$$\|F(v) - F(u)\|_{L^2} \lesssim (\|v\|_{H^1} + \|u\|_{H^1})\|u - v\|_{H^1}.$$

Proof. We compute

$$F(v) - F(u) = P(v^2 + u^2 + uv + 3(W_nv + W_nu))(u - v).$$

Thus

$$\begin{aligned} \|F(v) - F(u)\|_{L^2} &\lesssim (\|v^2\|_{L^2} + \|u^2\|_{L^2})\|u - v\|_{L^\infty} + (\|v\|_{L^\infty} + \|v\|_{L^\infty}\|u\|_{L^\infty} + \|u\|_{L^\infty})\|u - v\|_{L^2} \\ &\lesssim (\|v\|_{L^4}^2 + \|u\|_{L^4}^2)\|u - v\|_{H^1} + (\|v\|_{H^1} + \|v\|_{H^1}\|u\|_{H^1} + \|u\|_{H^1})\|v - u\|_{H^1} \\ &\lesssim (\|v\|_{H^1}^2 + \|u\|_{H^1}^2 + \|v\|_{H^1} + \|u\|_{H^1})\|u - v\|_{H^1} \\ &\lesssim (\|v\|_{H^1} + \|u\|_{H^1})\|u - v\|_{H^1} \end{aligned}$$

for $\|u\|_{H^1} \leq 1$, $\|v\|_{H^1} \leq 1$. Here we have used the Gagliardo Nirenberg inequality and the Sobolev embedding $H^1 \hookrightarrow L^\infty$. \square

In the next subsection we will show that

$$\int_{\mathbb{R}} V_n(x) dx < 0.$$

3.1. Study of the potential V_n . Going back to the r variable we see that the potential W_n fulfills the following equation

$$\left(1 - \frac{1}{r}\right) \partial_r^2 W_n + \frac{1}{r^2} \partial_r W_n + \frac{1}{r^2} W_n (1 - W_n^2) = 0 \quad (3.1)$$

with initial data (or boundary condition) $W_n(1) = a_n$, for $0 < a_n \leq \frac{1+\sqrt{3}}{5+3\sqrt{3}}$. We also have $\lim_{r \rightarrow \infty} W_n(r) = (-1)^n$. We will drop the index n in the rest of this subsection.

3.1.1. *A bound on W .*

Lemma 3.1. *We have $-a \leq W \leq a$ for $1 \leq r \leq 3$.*

Proof. Since the initial data for W are $W(1) = a$ and $W'(1) = -a(1 - a^2) < 0$, there exists $r_0 > 1$ such that for $1 \leq r \leq r_0$ we have

$$-a \leq W(r) \leq a$$

Then Lemma A.1 implies that on this interval we have

$$-a \leq \partial_r W(r) \leq a.$$

W is initially decreasing and can not have a local minimum in the region $W > 0$ (this is a consequence of the maximum principle, see Lemma A.2). Consequently there exists $r_1 > 1$ such that $0 \leq W \leq a$ on $[1, r_1]$ and $W(r_1) = 0$. Because of the bound of the derivative we have $r_1 \geq 2$. By the same bound we have $-a \leq W \leq a$ on $[r_1, r_1 + 1]$. \square

Let $Q(r) = 1 - \frac{1}{r} - \frac{1}{2r^2}$.

Proposition 3.2. *We have for $r \geq 3$*

$$-Q(r) \leq W(r) \leq Q(r)$$

Let

$$L(u, r) = \left(1 - \frac{1}{r}\right) \partial_r^2 u + \frac{1}{r^2} \partial_r u + \frac{1}{r^2} u(1 - u^2).$$

Before proving Proposition 3.2, we need the following lemma

Lemma 3.2. *For $r \geq 3$ we have $L(Q, r) < 0$ and $L(-Q, r) > 0$.*

Proof. Since L is odd in u , it is sufficient to prove $L(Q, r) < 0$. We calculate

$$\begin{aligned} L(Q, r) &= \left(1 - \frac{1}{r}\right) \left(-\frac{2}{r^3} - \frac{3}{r^4}\right) + \frac{1}{r^2} \left(\frac{1}{r^2} + \frac{1}{r^3}\right) + \frac{1}{r^2} \left(1 - \frac{1}{r} - \frac{1}{2r^2}\right) \left(1 - \left(1 - \frac{1}{r} - \frac{1}{2r^2}\right)^2\right) \\ &= -\frac{2}{r^4} + \frac{2}{r^5} + \frac{3}{4r^6} + \frac{3}{4r^7} + \frac{1}{8r^8}. \end{aligned}$$

Consequently, for $r \geq 3$ we have

$$L(Q, r) \leq \frac{1}{r^4} \left(-2 + \frac{2}{3} + \frac{3}{4 * 3^2} + \frac{3}{4 * 3^3} + \frac{1}{8 * 3^4}\right) \leq -\frac{1}{r^4} < 0.$$

\square

Proof of Proposition 3.2. We have $-a \leq W(3) \leq a$ and

$$a < \frac{11}{18} = 1 - \frac{1}{3} - \frac{1}{2 * 9} = Q(3).$$

If the inequality of Proposition 3.2 is false, there exists $r_1 < r_2$ with r_2 which can be infinite such that

$$W(r_1) = Q(r_1), \quad W(r_2) = Q(r_2)$$

and $W > Q$ on $]r_1, r_2[$ (The case $W < -Q$ is treated in a similar way). Consider r_0 such that $W - Q$ is maximum at r_0 . Note that such a maximum always exists independently if $\lim_{r \rightarrow \infty} W(r) = -1$ (in which case $r_2 < \infty$) or $\lim_{r \rightarrow \infty} W(r) = 1 = \lim_{r \rightarrow \infty} Q(r)$. Then we have

$$L(W, r_0) - L(Q, r_0) = -L(Q, r_0) > 0$$

so

$$\left(1 - \frac{1}{r_0}\right) (\partial_r^2 W - \partial_r^2 Q)(r_0) + \frac{1}{r_0^2} (W(1 - W^2) - Q(1 - Q^2)) > 0$$

Since

$$W(r_0) > Q(r_0) \geq Q(3) = \frac{11}{18} \geq \frac{1}{\sqrt{3}}$$

and the function $x \mapsto x(1 - x^2)$ is decreasing for $x \geq \frac{1}{\sqrt{3}}$ we have

$$(W(1 - W^2) - Q(1 - Q^2)) \leq 0$$

and consequently

$$\left(1 - \frac{1}{r_0}\right) (\partial_r^2 W - \partial_r^2 Q)(r_0) > 0$$

which is a contradiction with the fact that $W - Q$ is maximum at r_0 . \square

3.1.2. *A bound on the potential.* We now come back to the potential

$$V = P(3W^2 - 1)$$

Proposition 3.3. *We have*

$$\int_{\mathbb{R}} V(x) dx < 0.$$

Proof. First note that

$$\int_{\mathbb{R}} V(x) dx = \int_1^\infty \frac{3W^2 - 1}{r^2} dr.$$

We estimate

$$\int_1^3 \frac{3W^2 - 1}{r^2} \leq \int_1^3 \frac{3a^2 - 1}{r^2} = \frac{2(3a^2 - 1)}{3}$$

and

$$\begin{aligned} \int_3^\infty \frac{3W^2 - 1}{r^2} &\leq \int_3^\infty \frac{1}{r^2} \left(3 \left(1 - \frac{1}{r}\right)^2 - 1\right) \leq \int_3^\infty \frac{1}{r^2} \left(2 - \frac{6}{r} + \frac{3}{r^2}\right) = \left[-\frac{2}{r} + \frac{3}{r^2} - \frac{1}{r^3}\right]_3^\infty \\ &= \frac{1}{3} + \frac{1}{27} = \frac{10}{27} \end{aligned}$$

Note that

$$\frac{2(3a^2 - 1)}{3} + \frac{10}{27} < 0$$

because $a \leq \frac{1+\sqrt{3}}{5+3\sqrt{3}} < \frac{2}{3\sqrt{3}}$. Therefore we have

$$\int_{\mathbb{R}} V(x) dx < 0.$$

\square

3.2. Proof of Theorem 3. The main theorem with \mathcal{E} replaced by X now follows from the abstract result. In order to be able to replace X by \mathcal{E} we need the following lemma. We will drop the index n .

Lemma 3.3. *Let ϕ_0 be an eigenfunction of \mathcal{A} with eigenvalue $-\lambda^2$. Then we have*

$$\int_{\mathbb{R}} P|\phi_0|^2 \geq \lambda^2 \int_{\mathbb{R}} |\phi_0|^2. \quad (3.2)$$

$$- \int V|\phi_0|^2 \geq 0. \quad (3.3)$$

Proof. Let us first show (3.2). We have

$$(-\partial_x^2 + V)\phi_0 = -\lambda^2\phi_0.$$

Multiplication by ϕ_0 and integration by parts gives

$$\int |\phi_0'|^2 + \int V|\phi_0|^2 + \lambda^2 \int |\phi_0|^2 = 0. \quad (3.4)$$

Now recall that $V = P(3W^2 - 1)$, thus

$$\int P|\phi_0|^2 \geq \lambda^2 \int |\phi_0|^2.$$

We now show (3.3). From (3.4) we obtain :

$$- \int V|\phi_0|^2 = \int |\phi_0'|^2 + \lambda^2 \int |\phi_0|^2 \geq 0.$$

□

Let $\tilde{\mathcal{H}}^1$ the completion of C_0^∞ for the norm

$$\|u\|_{\tilde{\mathcal{H}}^1}^2 = \|u\|_{H^1}^2 + \|u\|_{L^2_P}^2$$

We put $\tilde{\mathcal{E}} = \tilde{\mathcal{H}}^1 \times L^2$.

Proof of Theorem 3

We continue using the notations of the abstract setting. We claim that it is sufficient to show the following :

$$\begin{aligned} \text{There exists } \epsilon_0 > 0 \text{ and a sequence } \psi_0^m \text{ with } \|\psi_0^m\|_X \rightarrow 0, m \rightarrow \infty, \\ \text{but for all } m \quad \sup_{t \geq 0} \|\psi^m(t)\|_{\tilde{\mathcal{E}}} \geq \epsilon_0 > 0. \end{aligned} \quad (\text{IM})$$

To see this we first note that

$$\|\psi_0^m\|_{\mathcal{E}} \leq \|\psi_0^m\|_X$$

because

$$\left(\int P|u|^4 \right)^{1/4} \lesssim \|u\|_\infty^{1/2} \|u\|_{L^2}^{1/2} \leq \|u\|_{H^1}$$

by the Sobolev embedding $H^1 \hookrightarrow L^\infty$. On the other hand

$$\|u\|_{L^2_P} = \left(\int P|u|^2 \right)^{1/2} \leq \left(\int P \right)^{1/4} \left(\int P|u|^4 \right)^{1/4} \lesssim \|u\|_{L^2_P}$$

and thus

$$\|\psi^m(t)\|_{\mathcal{E}} \gtrsim \|\psi^m(t)\|_{\tilde{\mathcal{E}}}.$$

Let us now show (IM). We follow the proof of the main theorem. We choose

$$\psi_0 = \begin{pmatrix} \phi_0 \\ \frac{\lambda_0}{i} \phi_0 \end{pmatrix}, \phi_0 \in \mathbb{1}_{\{-\lambda_0^2\}}(\mathcal{A})\mathcal{H}, \|\phi_0\| = \frac{1}{3(1 + \|V_-\|_{\infty})^{1/2}} \rho.$$

We estimate

$$\begin{aligned} \|\psi_0\|_X^2 &= \langle (-\partial_x^2 + V)\phi_0, \phi_0 \rangle - \langle V\phi_0, \phi_0 \rangle + \|\phi_0\|^2 + \lambda_0^2 \|\phi_0\|^2 \\ &\leq (\|V_-\|_{\infty} + 1) \|\phi_0\|^2 = 1/9 \rho^2. \end{aligned}$$

Thus the first part of the proof goes through without any changes. We then have to estimate $\|\psi(\tau)\|_{\tilde{\mathcal{E}}}$. We estimate

$$\begin{aligned} \|\psi_0\|_{\tilde{\mathcal{E}}}^2 &= \langle \mathcal{A}\phi_0, \phi_0 \rangle - \langle V\phi_0, \phi_0 \rangle + \int P|\phi_0|^2 + \lambda_0^2 |\phi_0|^2 \\ &= -\langle V\phi_0, \phi_0 \rangle + \int P|\phi_0|^2 \\ &\geq \int P|\phi_0|^2 \geq \lambda_0^2 \int |\phi_0|^2 \\ &= \lambda_0^2 \frac{1}{9(1 + \|V_-\|_{\infty})} \rho^2. \end{aligned}$$

Here we have used Lemma 3.3. Using

$$\|u\|_{\tilde{\mathcal{E}}} \leq C_1 \|u\|_X$$

we find

$$\|\psi(\tau)\|_{\tilde{\mathcal{E}}} \geq \frac{\lambda_0}{3(1 + \|V_-\|_{\infty})^{1/2}} \rho - \frac{2C_1 M_L M_F \|E_1\|}{2\beta - \lambda_1} \rho^2 \geq \frac{\lambda_0}{6(1 + \|V_-\|_{\infty})} \rho$$

for ρ small enough. \square

3.3. Proof of Corollary 1.2. We recall

$$E^{(\partial_t)}(F(t)) = \mathcal{E}(W, \dot{W}).$$

We take the same sequence of data $W_{0,n}^m$ as in Theorem 3. We first have to show that

$$\int P((W_{0,n}^m)^2 - W_n^2)^2 \rightarrow 0, \quad m \rightarrow \infty.$$

This follows from

$$\begin{aligned} \int P((W_{0,n}^m)^2 - W_n^2)^2 &\lesssim \int P(W_{0,n}^m - W_n)^4 + \int P W_n^2 (W_{0,n}^m - W_n)^2 \\ &\lesssim \int P(W_{0,n}^m - W_n)^4 + \left(\int P(W_{0,n}^m - W_n)^4 \right)^{1/2} \rightarrow 0, \quad m \rightarrow \infty \end{aligned}$$

by Theorem 3. In the first inequality we have used the estimate

$$(A^2 - B^2)^2 = (A - B)^2 (A + B)^2 = (A - B)^2 (A - B + 2B)^2 \leq 2(A - B)^4 + 8B^2 (A - B)^2,$$

and the fact that $\|W_n\|_{L^\infty} \leq 1$. Now we have to show that

$$\sup_{t \geq 0} \int (\dot{W}_n^m)^2 + ((W_n^m)' - W_n')^2 + P((W_n^m)^2 - W_n^2)^2 \geq \epsilon_1 > 0. \quad (3.5)$$

We know by Theorem 3 that

$$\sup_{t \geq 0} \int (\dot{W}_n^m)^2 + ((W_n^m)' - W_n')^2 + P(W_n^m - W_n)^4 \geq \epsilon_0 > 0. \quad (3.6)$$

We also know from the proof of Theorem 3 that this supremum is achieved on the interval $[0, m]$ and that on this interval

$$\|W_n^m - W_n\|_{L^2} \leq \rho$$

for some $\rho > 0$. Now observe that for $u \in H^1$ we have

$$\int Pu^2 \leq 2 \left(\int \frac{1}{r^2} u^2 \right)^{1/2} \left(\int (u')^2 \right)^{1/2}. \quad (3.7)$$

Indeed by density we can suppose $u \in C_0^\infty(\mathbb{R})$ and then compute

$$\int Pu^2 = \int \partial_x \left(-\frac{1}{r} \right) u^2 = 2 \int \frac{1}{r} u u' \leq 2 \left(\int \frac{1}{r^2} u^2 \right)^{1/2} \left(\int (u')^2 \right)^{1/2}.$$

Let us now show (3.5). We can suppose that

$$\sup_{t \geq 0} \int ((W_n^m)' - W_n')^2 \leq \frac{\epsilon_0^2}{4^4 \rho^2}$$

because otherwise there is nothing to show. Then we estimate

$$\begin{aligned} & \int \dot{W}_n^m{}^2 + ((W_n^m)' - W_n')^2 + P((W_n^m)^2 - W_n^2)^2 \\ & \geq \int \dot{W}_n^m{}^2 + ((W_n^m)' - W_n')^2 + \frac{1}{2} P(W_n^m - W_n)^4 - 4PW_n^2(W_n^m - W_n)^2 \\ & \geq \int \frac{1}{2} (\dot{W}_n^m{}^2 + ((W_n^m)' - W_n')^2 + P(W_n^m - W_n)^4) \\ & \quad - 4 \left(\int ((W_n^m)' - W_n')^2 \right)^{1/2} \left(\int \frac{(W_n^m - W_n)^2}{r^2} \right)^{1/2} \\ & \geq \frac{1}{2} \int \dot{W}_n^m{}^2 + ((W_n^m)' - W_n')^2 + P(W_n^m - W_n)^4 - \epsilon_0/4, \end{aligned}$$

where in the first inequality we have used the estimate

$$\begin{aligned} (A^2 - B^2)^2 &= (A - B)^2(A - B + 2B)^2 = (A - B)^2((A - B)^2 + 4B(A - B) + 4B^2) \\ &\geq (A - B)^2((A - B)^2 - \frac{1}{2}(A - B)^2 - 8B^2 + 4B^2) = \frac{1}{2}(A - B)^4 - 4B^2(A - B)^2, \end{aligned}$$

and in the second inequality we have used (3.7) and the fact that $\|W_n\|_{L^\infty} \leq 1$.

The supremum over $t \geq 0$ of this expression is $\geq \epsilon_0/4$ by (3.6). \square

APPENDIX A. PROOF OF THEOREM 2

In this appendix we give an explicit proof of theorem 2. We adapt in the simpler uncoupled case the arguments of Smoller Wasserman Yau, McLeod [19]; Smoller, Wasserman, Yau [18] and Smoller, Wasserman [17] to show the existence of infinitely

many solutions. In this appendix we work with the r variable and we note $' = \partial_r$. Again we can suppose that $m = 1/2$. Recall that the stationary equation reads

$$\left(1 - \frac{1}{r}\right) W'' + \frac{1}{r^2} W' + \frac{1}{r^2} W(1 - W^2) = 0. \quad (\text{A.1})$$

A.1. Local solutions.

Proposition A.1. *Let $0 < \alpha < 1$ and $0 \leq a \leq 1$. There exists $r_a > 1$ and a unique solution $W \in C^{2,\alpha}([1, r_a])$ with boundary condition*

$$W(1) = a, \quad W'(1) = b, \quad W''(1) = c$$

where

$$b = -a(1 - a^2), \quad 2c = -b(1 - 3a^2)$$

Proof. We set $z = W'$ to write the equation as a first order system. We consider

$$X = \{(w, z) \in C^{(2,\alpha)}([1, 1+\epsilon]) \times C^{(1,\alpha)}([1, 1+\epsilon]), w(1) = a, w'(1) = z(1) = b, w''(1) = z'(1) = c\}$$

and the map $T : (w, z) \in X \mapsto (\tilde{w}, \tilde{z})$ with

$$\begin{aligned} \tilde{w} &= a + \int_1^r z, \\ \tilde{z} &= b - \int_1^r \frac{1}{\rho(\rho-1)} (z + w(1 - w^2)). \end{aligned}$$

We first show that T preserves the boundary conditions. We calculate

$$\begin{aligned} \tilde{z}' &= -\frac{1}{r(r-1)} (z + w(1 - w^2)) \\ &= -\frac{1}{r(r-1)} \left(z + a(1 - a^2) + \int_1^r w'(1 - 3w^2) \right) \\ &= -\frac{1}{r(r-1)} (z - b) - \frac{1}{r(r-1)} \int_1^r w'(1 - 3w^2) \end{aligned}$$

so $\tilde{z}'(r) \rightarrow -z'(1) - w'(1)(1 - 3w^2(1)) = -c + 2c = c$ when $r \rightarrow 1$. We now show that T is a contraction in $B_X(0, A)$ for ϵ small enough. For this the only difficulty

is to estimate

$$\begin{aligned}
\frac{|\tilde{z}'(r) - \tilde{z}'(1)|}{|r-1|^\alpha} &\leq \frac{1}{|r-1|^\alpha} \left| -\frac{1}{r(r-1)}(z + w(1-w^2)) - c \right| \\
&\leq \frac{1}{|r-1|^\alpha} \left| -\frac{1}{r(r-1)} \left(b + \int_1^r (z'(\rho) - z'(1))d\rho + c(r-1) \right. \right. \\
&\quad \left. \left. + a(1-a^2) + \int_1^r (w(1-w^2))'(\rho) - (w(1-w^2))'(1)d\rho + b(1-3a^2)(r-1) \right) - c \right| \\
&\leq \frac{1}{|r-1|^\alpha} \left| -\frac{1}{r(r-1)} \int_1^r (z'(\rho) - z'(1))d\rho \right| \\
&\quad + \frac{1}{|r-1|^\alpha} \left| -\frac{1}{r(r-1)} \left(-(r-1)c + \int_1^r (w(1-w^2))'(\rho) - (w(1-w^2))'(1)d\rho \right) - c \right| \\
&\leq \frac{1}{r(r-1)^{1+\alpha}} \int_1^r |z'(\rho) - z'(1)| + c \frac{1}{(r-1)^\alpha} \left(1 - \frac{1}{r} \right) \\
&\quad + \frac{1}{(r-1)^{1+\alpha}} \|(w(1-w^2))'\|_{C^1} (r-1)^2 \\
&\leq \frac{1}{r(r-1)^{1+\alpha}} \|z'\|_{C^{0,\alpha}} \int_1^r |\rho-1|^\alpha + c\epsilon^{1-\alpha} + C(\|w\|_{C^2})\epsilon^{1-\alpha} \\
&\leq \frac{1}{1+\alpha} \|z'\|_{C^{0,\alpha}} + c\epsilon^{1-\alpha} + C(\|w\|_{C^2})\epsilon^{1-\alpha}.
\end{aligned}$$

Consequently we can show that for ϵ small enough, T is a contraction, with contracting constant $\frac{1}{1+\alpha} + C(A)\epsilon^{1-\alpha}$, and consequently it has a unique fixed point. \square

As a corollary of the proof of local existence we obtain the continuity of the family of solutions W_a with respect to the initial data a .

Corollary A.1. *Let $\delta > 0$. If W_a is a solution on $[1, R]$ with $-1 \leq W_a \leq 1$ and a' is sufficiently close to a , then $W_{a'}$ is defined on $[1, R]$ we have*

$$\|W_a - W_{a'}\|_{C^{2,\alpha}([1,R])} \leq \delta.$$

A.2. Basic facts.

Lemma A.1. *Let $0 < B \leq 1$. As long as W is a C^2 solution with $|W| \leq B$ we have $|W'| \leq B$.*

Proof. Assume that in $[1, r_0]$ we have $|W| \leq B$. Then

$$-\frac{B}{r^2} \leq \left(1 - \frac{1}{r}\right)W'' + \frac{1}{r^2}W' \leq \frac{B}{r^2}$$

and consequently

$$\left[\frac{B}{r}\right]_1^r \leq \left[\left(1 - \frac{1}{r}\right)W'\right]_1^r \leq \left[-\frac{B}{r}\right]_1^r$$

so

$$-B \leq W'(r) \leq B.$$

\square

Corollary A.2. *The solution W exists and is $C^{2,\alpha}$ as long as $|W| \leq 1$.*

We now consider the solution W on $[0, r_a[$ where r_a is the smallest r such that $|W| = 1$ if it exists, and $r_a = \infty$ otherwise.

Lemma A.2. *The solution W cannot have a local minimum with $W > 0$ nor a local maximum with $W < 0$.*

Proof. If W has a positive local minimum at r_0 then

$$\left(1 - \frac{1}{r_0}\right) W''(r_0) + \frac{1}{r_0^2} W(r_0)(1 - W^2(r_0)) = 0$$

but $\frac{1}{r_0^2} W(r_0)(1 - W^2(r_0)) > 0$ (the local minimum cannot be 1), and $W''(r_0) \geq 0$, which is a contradiction. \square

Lemma A.3. *The solution W can not have a limit $l \neq -1, 0, 1$.*

Proof. Assume that $W \rightarrow l$ with $0 < l < 1$. We can write for r_1 big enough and $r_n \geq r_1$

$$\left[\frac{l(1-l^2) + \epsilon}{r}\right]_{r_1}^{r_n} \leq \left[\left(1 - \frac{1}{r}\right)W'\right]_{r_1}^{r_n} \leq \left[\frac{l(1-l^2) - \epsilon}{r}\right]_{r_1}^{r_n}.$$

Since $W \rightarrow l$ there exists a sequence $r_n \rightarrow \infty$ such that $W'(r_n) \rightarrow 0$. Letting $n \rightarrow \infty$ we obtain

$$\frac{l(1-l^2) - \epsilon}{r_1} \leq W'(r_1) \left(1 - \frac{1}{r_1}\right) \leq \frac{l(1-l^2) + \epsilon}{r_1}$$

so

$$\frac{l(1-l^2) - \epsilon}{r_1 - 1} \leq W'(r_1) \leq \frac{l(1-l^2) + \epsilon}{r_1 - 1}$$

and there exists a constant C such that for r big enough

$$(l(1-l^2) - \epsilon) \ln(r-1) \leq W(r) - C$$

which is a contradiction. \square

A.3. More technical facts.

Proposition A.2. *Let $0 \leq a \leq 1$. There exists $\epsilon > 0$ and $R > 0$ such that if there exists $R < r_0 < r_a$ such that W has a local extremum at r_0 with $1 - \epsilon \leq |W(r_0)| < 1$ then $r_a < \infty$ and W has one and only one zero in $[r_0, r_a]$.*

Proof. We consider the case $W(r_0) > 0$. The other case can be treated similarly. We consider

$$H = r^2 \frac{(W')^2}{2} + \frac{W^2}{2} - \frac{W^4}{4}.$$

We calculate

$$\begin{aligned} H'(r) &= r(W')^2 + r^2 \frac{1}{r(1-r)} (W' + W(1-W^2))W' + WW' - W^3W' \\ &= (W')^2 \left(r + \frac{r^2}{r(1-r)} \right) + WW'(1-W^2) \left(1 + \frac{r^2}{r(1-r)} \right) \\ &= (W')^2 \left(r + \frac{r^2}{r(1-r)} \right) + WW'(1-W^2) \frac{1}{1-r}. \end{aligned}$$

Let R be such that for $r \geq R$ we have

$$\left(r + \frac{r^2}{r(1-r)} \right) > 0$$

then for $r \geq R$ and $WW' \leq 0$ we have $H'(r) > 0$. With our assumption on r_0 we can estimate

$$H(r_0) \geq \frac{(1-\epsilon)^2}{2} - \frac{1}{4} \geq \frac{1}{4\delta}$$

for a suitable δ which will be precised later, and ϵ small enough. Since W has a local maximum at r_0 , there are two possibilities

- We have $r_a = \infty$, W is decreasing on $[r_0, +\infty[$ and $W \rightarrow 0$ at ∞ .
- There exists $r_1 < r_a$ such that $W(r_1) = 0$ and W is decreasing on $[r_0, r_1]$.

In the first case we obtain for all $r \geq r_a$, $H'(r) > 0$ so

$$H(r) \geq \frac{1}{4\delta}$$

and since $W \rightarrow 0$ the expression of H yields the existence of r_2 such that for $r \geq r_2$

$$W'(r)^2 \geq \frac{1}{(2\delta+1)r^2}$$

so $W'(r) \leq -\frac{1}{\sqrt{2\delta+1}r}$ and $W(r) \leq W(r_2) - \frac{1}{\sqrt{2\delta+1}} \ln(r)$ which is a contradiction. Consequently we are in the second case. We have $H(r_1) \geq H(r_0)$ so we can estimate $W'(r_1)$

$$W'(r_1) \leq -\frac{1}{\sqrt{2\delta}r_1}$$

Moreover, when $-1 \leq W \leq 1$ we have $W(1-W^2) \leq \frac{2}{3\sqrt{3}}$ and consequently we can write for $r_a > r_2 > r_1$

$$\left[W'(r) \left(1 - \frac{1}{r} \right) \right]_{r_1}^{r_2} \leq \left[-\frac{2}{3\sqrt{3}r} \right]_{r_1}^{r_2}$$

and consequently

$$\left(1 - \frac{1}{r_2} \right) W'(r_2) \leq - \left(1 - \frac{1}{r_1} \right) \frac{1}{\sqrt{2\delta}r_1} + \frac{2}{3\sqrt{3}r_1} - \frac{2}{3\sqrt{3}r_2}.$$

For r_1 big enough (which is possible by choosing R big enough) and δ close enough to 1 (which is possible by choosing ϵ small enough) we have

$$- \left(1 - \frac{1}{r_1} \right) \frac{1}{\sqrt{2\delta}r_1} + \frac{2}{3\sqrt{3}r_1} \leq 0,$$

since $\frac{2}{3\sqrt{3}} < \frac{1}{\sqrt{2}}$. and consequently

$$W'(r_2) \leq -\frac{2}{3\sqrt{3}(r_2 - 1)},$$

and therefore $r_a < \infty$, $W(r_a) = -1$ and W' is decreasing on $[r_1, r_a]$. This concludes the proof of Proposition A.2 \square

Proposition A.3. *Let $N > 0$. Then for a small enough, the solution W has more than N zeros on $[1, r_a]$*

Proof. To count the number of zeros of a solution W we can introduce the function θ which is the continuous function such that

$$\tan(\theta) = \frac{W'}{W}$$

and $-\frac{\pi}{2} < \theta(1) < \frac{\pi}{2}$. Then W has N zero between 1 and r_0 if and only if

$$-\frac{\pi}{2} - N\pi < \theta(r_0) < \frac{\pi}{2} - N\pi.$$

It is totally similar to count the number of zero thanks to the function ψ defined by

$$\tan(\psi) = \frac{rW'}{W}.$$

We estimate ψ'

$$\begin{aligned} \psi'(r) &= \frac{1}{1 + \left(\frac{rW'}{W}\right)^2} \frac{W(W' + rW'') - r(W')^2}{W^2} \\ &= \frac{WW' - W\frac{1}{r-1}(W' + W(1 - W^2)) - (rW')^2}{W^2 + (rW')^2} \\ &= \frac{WW'\frac{r-2}{r-1} - \frac{1}{r-1}W^2(1 - W^2) - (rW')^2}{W^2 + (rW')^2}. \end{aligned}$$

We first estimate

$$\frac{WW'\frac{r-2}{r-1}}{W^2 + (rW')^2} \leq \frac{1}{r} \frac{|rWW'|}{W^2 + (rW')^2} \leq \frac{1}{2r}.$$

We assume that $|W| \leq \delta$. To estimate the other terms we consider three cases

- $2|W|^2 \leq |rW'|^2$. Then we have

$$\psi'(r) \leq \frac{1}{2r} - \frac{r(W')^2}{W^2 + (rW')^2} \leq \frac{1}{2r} - \frac{2}{3r} = -\frac{1}{6r}.$$

- $2|rW'|^2 \leq |W|^2$. Then we have

$$\psi'(r) \leq \frac{1}{2r} - \frac{W^2(1 - W^2)}{(r-1)(W^2 + (rW')^2)} \leq \frac{1}{2r} - \frac{2(1 - \delta^2)}{3(r-1)} \leq \frac{3 - 4(1 - \delta^2)}{6r}.$$

- $\frac{1}{2}|W|^2 \leq |rW'|^2 \leq 2|W|^2$. Then we have

$$\psi'(r) \leq \frac{1}{2r} - \frac{r(W')^2}{W^2 + (rW')^2} - \frac{W^2(1 - W^2)}{(r-1)(W^2 + (rW')^2)} \leq \frac{1}{2r} - \frac{1}{3r} - \frac{(1 - \delta^2)}{3r} \leq \frac{3 - 4 + 2\delta^2}{6r}.$$

If we take δ small enough we then have

$$\psi'(r) \leq -\frac{1}{12r}.$$

Let now R be such that $-\frac{1}{12} \ln(R) \leq -N\pi$. Thanks to Corollary A.1, we can find a_0 small enough such that for $0 \leq a \leq a_0$ the solution exists on $[1, R]$ and satisfies $|W| \leq \delta$ on this interval. Then $\psi(R) - \psi(1) \leq -N\pi$ so W has at least N zero on $[1, R]$. This concludes the proof of Proposition A.3. \square

Corollary A.3. *Let W be a solution with $r_a = \infty$ and a finite number of zeros. Then $W \rightarrow \pm 1$.*

Proof. Because of Lemma A.2 a solution with a finite number of zeros has a finite limit. Because of Lemma A.3 this limit must be 0 or ± 1 . If it was 0 we could find R_0 such that $|W| \leq \delta$ for $r \geq R_0$, with δ defined in the proof of Proposition A.3. then for $r \geq R_0$ we have

$$\psi'(r) \leq -\frac{1}{12r},$$

consequently ψ is unbounded from above, so W has an infinite number of zeros. \square

A.4. Proof of the Theorem.

Lemma A.4. *Let X_n be the set of initial data a such that the corresponding solution has n zeros and satisfies $r_a < \infty$. Then X_n is open and if α is a limit point of X_n the corresponding solution satisfies $r_\alpha = \infty$ has m zeros, with $m = n$ or $m = n - 1$ and tends to $(-1)^m$ at infinity.*

Proof. The fact that X_n is open is a direct consequence of Corollary A.1. Let α be a limit point and let $a_i \in X_n$ be such that $a_i \rightarrow \alpha$.

Assume first that W_α is such that $r_a < \infty$. Then we can compare the solution on the fixed interval $[1, r_\alpha]$ so Corollary A.1 implies that W_α has exactly n zeros, so $\alpha \in X_n$ which is a contradiction. Consequently $r_a = \infty$. This also implies that the sequence of r_{a_i} is not bounded.

Assume now that there exists R such that W_α has strictly more than n zeros before R . Once again Corollary A.1 yields a contradiction.

Assume now that W_α has m zeros with $m < n$. Then thanks to Corollary A.3 W_α tend to $(-1)^m$ and Proposition A.2 implies that the W_{a_i} have m or $m + 1$ zeros. \square

Proof of Theorem 2. Let \tilde{X}_n be the set of initial data a such that the corresponding solution has less than n zeros. Let $\alpha = \min(\tilde{X}_n)$. Proposition A.3 implies that $\alpha > 0$. There are two case

- If $\alpha \in \tilde{X}_n$, then W_α is a solution with $m \leq n$ zeros with $r_\alpha = \infty$. Then Corollary A.1 and Proposition A.2 imply that for a close to α either the solutions have m zeros, either have $m + 1$ zeros and $r_a < \infty$. Consequently

we have $m = n$ and considering a sequence $a_i < \alpha$ converging to α we have shown that X_{n+1} is non empty.

- If $\alpha \notin \widetilde{X}_n$ then W_α must be a solution with $r_\alpha = \infty$ and $k > n$ zeros. But we have shown in the previous point that in a neighborhood of such a solution we can only have solutions with k or $k + 1$ zeros, so this case can not occur.

We start the iteration with the function

$$W_1 = \frac{c - r}{r + 3(c - 1)}, \quad c = \frac{3 + \sqrt{3}}{2},$$

which is a special solution of (3.1), with only one 0 (see [2]). Note also that

$$W_1(1) = \frac{1 + \sqrt{3}}{5 + 3\sqrt{3}} = a_1.$$

We then obtain at least one solution $-1 \leq W_{a_n} \leq 1$ for each number of zeros n . This concludes the proof of Theorem 2. □

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