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BLOW-UP PROFILE OF ROTATING 2D FOCUSING BOSE GASES

MATHIEU LEWIN, PHAN THÀNH NAM, AND NICOLAS ROUGERIE

ABSTRACT. We consider the Gross-Pitaevskii equation describing an attractive Bose gas trapped to a quasi 2D layer by means of a purely harmonic potential, and which rotates at a fixed speed of rotation Ω . First we study the behavior of the ground state when the coupling constant approaches a_* , the critical strength of the cubic nonlinearity for the focusing nonlinear Schrödinger equation. We prove that blow-up always happens at the center of the trap, with the blow-up profile given by the Gagliardo-Nirenberg solution. In particular, the blow-up scenario is independent of Ω , to leading order. This generalizes results obtained by Guo and Seiringer (*Lett. Math. Phys.*, 2014, vol. 104, p. 141–156) in the non-rotating case. In a second part we consider the many-particle Hamiltonian for N bosons, interacting with a potential rescaled in the mean-field manner $-a_N N^{2\beta-1} w(N^\beta x)$, with $w \geq 0$ a positive function such that $\int_{\mathbb{R}^2} w(x) dx = 1$. Assuming that $\beta < 1/2$ and that $a_N \rightarrow a_*$ sufficiently slowly, we prove that the many-body system is fully condensed on the Gross-Pitaevskii ground state in the limit $N \rightarrow \infty$.

Dedicated to Herbert Spohn, on the occasion of his 70th birthday

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1. INTRODUCTION

Because of their ability to display quantum effects at the macroscopic scale, Bose-Einstein condensates (BEC) have become an important subject of research, in particular after their first realization in the laboratory in 1995 [3, 15, 18, 35]. Condensates with *attractive* interactions are expected to behave quite differently from the better understood *repulsive* case, and they have generated many experimental, numerical or theoretical works. Some atoms like ${}^7\text{Li}$ indeed have a negative scattering length and were initially believed not to be able to form a condensate, until they were finally realized experimentally in traps [5].

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For attractive interactions, the Gross-Pitaevskii functional, commonly used to describe BECs, predicts a collapse of the system when Na (the number of particles times the scattering length) is too negative [4, 17, 69, 56, 29], an effect which has been observed in some experiments [28, 23]. In addition, attractive Bose-Einstein condensates are believed to respond to rotation in a rather different manner from the repulsive case. In a rotating repulsive Bose gas, a triangular lattice of vortices is formed, with the number of vortices increasing with the speed of rotation [14, 26, 1, 16]. On the contrary, it has been argued [74, 55, 59, 65, 51, 13, 8, 66] that in an attractive rotating Bose gas, vortices should be unstable and it is instead the center of mass of the system which can rotate around the axis.

In this paper, we rigorously establish two results about 2D attractive Bose-Einstein condensates in the critical regime of collapse. We consider a Bose gas trapped to a quasi 2D layer by means of a purely harmonic potential, and which rotates at a fixed speed of rotation Ω . First we look at the Gross-Pitaevskii equation which describes the macroscopic behavior of the condensate. We study its solutions in the regime where the coupling constant approaches the critical blow-up value a_* , given by the best constant in the Gagliardo-Nirenberg inequality. In this case we prove that blow-up always happens at the center of the trap, with the blow-up profile given by the Gagliardo-Nirenberg optimizer. This shows that the rotation does not affect the general blow-up scenario, to leading order. The non-rotating case has been previously studied by Guo and Seiringer in [29]. Other similar mathematical results on the trapped nonlinear Schrödinger equation for non-rotating gases include [52, 21, 30, 61] in the stationary case and [75, 7, 76] in the time-dependent case.

In a second part we consider the many-particle (microscopic) Hamiltonian for N such bosons, interacting with a potential rescaled in the mean-field manner

$$-a_N N^{2\beta-1} w(N^\beta x)$$

with $w \geq 0$ a fixed positive function such that $\int_{\mathbb{R}^2} w(x) dx = 1$. Assuming that $0 < \beta < 1/2$ and that $a_N \rightarrow a_*$ sufficiently slowly, we are able to show that the many-body system is fully condensed on the Gross-Pitaevskii ground state studied in the first step, in the limit $N \rightarrow \infty$. This justifies the validity of the Gross-Pitaevskii equation in this regime of collapse, with complete Bose-Einstein condensation at the point of blow-up. We do not observe fragmented condensation at this order.

The mathematical method used here follows several of our previous papers [41, 43, 44, 46]. Note that some authors have already dealt with the time-dependent equation, in the subcritical regime $a < a_*$, see [10, 46, 34, 57].

The next section contains the precise definition of our model as well as the statement of our main results. The remainder of the paper is then devoted to their proofs. In Appendix A we mention several possible extensions of our findings, without giving the detailed mathematical proofs.

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2. MODELS AND MAIN RESULTS

2.1. Collapse of the rotating Gross-Pitaevskii ground state. The Gross-Pitaevskii functional describing a condensed system of bosons trapped to a 2D plane and rotating

along the third axis at an angular velocity Ω reads

$$\mathcal{E}_{\Omega,a}^{\text{GP}}(u) = \int_{\mathbb{R}^2} |\nabla u(x)|^2 dx + \int_{\mathbb{R}^2} |x|^2 |u(x)|^2 dx - 2\Omega \langle u, Lu \rangle - \frac{a}{2} \int_{\mathbb{R}^2} |u(x)|^4 dx \quad (2.1)$$

where $L = -ix \wedge \nabla = i(x_2 \partial_1 - x_1 \partial_2)$ is the angular momentum. Here we have chosen units such that the trapping potential has a trapping frequency $\Omega_{\text{trap}} = 1$, for simplicity. The system is stable under the assumption that $|\Omega| < 1$ and $a < a_*$, where a_* is the optimal constant of the Gagliardo-Nirenberg inequality [71, 75, 29, 52]

$$\left(\int_{\mathbb{R}^2} |\nabla u|^2 \right) \left(\int_{\mathbb{R}^2} |u|^2 \right) \geq \frac{a_*}{2} \int_{\mathbb{R}^2} |u|^4, \quad \forall u \in H^1(\mathbb{R}^2). \quad (2.2)$$

Equivalently,

$$a_* = \|Q\|_{L^2(\mathbb{R}^2)}^2,$$

where $Q \in H^1(\mathbb{R}^2)$ is the unique positive solution of

$$-\Delta Q + Q - Q^3 = 0 \text{ in } \mathbb{R}^2, \quad (2.3)$$

up to translations. More precisely, Q is symmetric radial decreasing and it is the unique (up to translations and dilations) optimizer for the Gagliardo-Nirenberg inequality (2.2). In the following we therefore always assume that $0 \leq \Omega < 1$ and $0 < a < a_*$. The energy may also be written in the form

$$\mathcal{E}_{\Omega,a}^{\text{GP}}(u) = \int_{\mathbb{R}^2} |\nabla u(x) + i\Omega x_{\perp} u(x)|^2 dx + (1 - \Omega^2) \int_{\mathbb{R}^2} |x|^2 |u(x)|^2 dx - \frac{a}{2} \int_{\mathbb{R}^2} |u(x)|^4 dx \quad (2.4)$$

where $x_{\perp} = (x_2, -x_1)$. We call

$$E_{\Omega}^{\text{GP}}(a) := \min_{u \in H^1(\mathbb{R}^2), \|u\|_{L^2} = 1} \mathcal{E}_{\Omega,a}^{\text{GP}}(u)$$

the ground state energy and look at the limit $a \rightarrow a_*$ at fixed $0 \leq \Omega < 1$. The existence of ground states follows the standard direct method in the calculus of variations. Our first main result is the following.

Theorem 2.1 (Collapse of rotating Gross-Pitaevskii ground states).

Let $0 \leq \Omega < 1$ be any fixed rotation. Then we have

$$E_{\Omega}^{\text{GP}}(a) = E_0^{\text{GP}}(a) + o(E_0^{\text{GP}}(a)) = \sqrt{a_* - a} \left(\frac{2\lambda_*^2}{a_*} + o(1) \right), \quad (2.5)$$

when $a \rightarrow a_*$, where

$$\lambda_* = \left(\int_{\mathbb{R}^2} |x|^2 |Q(x)|^2 dx \right)^{\frac{1}{4}}$$

and Q is the unique radial positive Gagliardo-Nirenberg solution (2.3).

In addition, for any sequence $a_N \rightarrow a_*$ and any sequence $\{u_N\}$ such that $\|u_N\|_{L^2(\mathbb{R}^2)} = 1$ and

$$\mathcal{E}_{\Omega,a_N}^{\text{GP}}(u_N) = E_{\Omega}^{\text{GP}}(a_N) + o(\sqrt{a_* - a_N}), \quad (2.6)$$

(for instance u_N a minimizer of $E_{\Omega}(a_N)$), we have

$$\lim_{N \rightarrow \infty} \left\| u_N - e^{i\theta_N} Q_N \right\|_{L^2(\mathbb{R}^2)} = 0, \quad (2.7)$$

for a properly chosen phase $\theta_N \in [0, 2\pi)$, where

$$Q_N(x) = (a_*)^{-1/2} \lambda_* (a_* - a_N)^{-\frac{1}{4}} Q \left(\lambda_* (a_* - a_N)^{-\frac{1}{4}} x \right) \quad (2.8)$$

is the rescaled Gagliardo-Nirenberg optimizer which blows up at the origin at speed $(a_* - a_N)^{\frac{1}{4}}/\lambda_*$.

This theorem was proved by Guo and Seiringer in [29] in the case $\Omega = 0$, with the convergence of the ground states but not of general “approximate ground states” u_N ’s as in (2.7). Our theorem shows that the results found by Guo and Seiringer remain valid when the system is set in rotation, the blow-up scenario being independent of Ω to leading order. In particular, we do not see a rotation of the center of mass at this order. Our method of proof relies on the non-degeneracy of the minimizer Q , which is known to play a fundamental role in many situations [72, 9]. We expect that the non-degeneracy of Q should provide quantitative estimates for the difference between Q_N and the ground state of $E_\Omega(a_N)$, however we have not investigated this question in details.

Remark 2.2.

(i) It would be interesting to investigate the case where $\Omega = \Omega_N \rightarrow 1$ at the same time as $a_N \rightarrow a_*$. In this case the centrifugal force almost compensates the trapping potential, and this effect could compete with the collapse scenario induced by attractive interactions.

(ii) Our proof covers more general external potentials attaining their minimum at the origin and behaving at least quadratically at zero and at infinity, like for instance the quartic-quadratic potential $V(x) = |x|^2 + k|x|^4$ with $k > 0$. \diamond

2.2. Collapse of the many-body system in the Gross-Pitaevskii ground state.

Next we turn to the N -particle quantum Hamiltonian describing our trapped 2D bosons, which reads

$$H_N = \sum_{j=1}^N (-\Delta_{x_j} + |x_j|^2 - 2\Omega L_{x_j}) - \frac{a}{N-1} \sum_{1 \leq i < j \leq N} w_N(x_i - x_j), \quad (2.9)$$

and acts on $\mathfrak{H}^N = \bigotimes_{\text{sym}}^N L^2(\mathbb{R}^2)$, the Hilbert space of square-integrable symmetric functions. The two-body interaction w_N approaches a Dirac delta and is chosen in the form

$$w_N(x) = N^{2\beta} w(N^\beta x) \quad (2.10)$$

for a fixed parameter $\beta > 0$ and a fixed function w satisfying

$$w(x) = w(-x) \geq 0, \quad (1 + |x|)w, \quad \widehat{w} \in L^1(\mathbb{R}^2), \quad \int_{\mathbb{R}^2} w = 1. \quad (2.11)$$

Finally, $a > 0$ is a parameter which describes the strength of the interaction. We will take

$$a = a_N \rightarrow a_*$$

which is the Gagliardo-Nirenberg critical constant mentioned before.

Hamiltonians of the form (2.9) have generated a huge amount of works in the past decades, in any dimension d . The chosen coupling constant proportional to $1/(N-1)$ ensures that the kinetic and interaction energies are comparable in the limit $N \rightarrow \infty$. Due to the trapping potential, most of the particles will usually accumulate in a bounded region of space, leading to a high density of order N (in our case they will even collapse at one point).

In this paper we are interested in the behavior of the ground state energy per particle of H_N ,

$$E_{\Omega,a}^Q(N) := N^{-1} \inf_{\Psi \in \mathfrak{H}^N, \|\Psi\|=1} \langle \Psi, H_N \Psi \rangle, \quad (2.12)$$

and in the corresponding (non necessarily unique) ground states Ψ_N , when $N \rightarrow \infty$. In the regime considered in this paper, we expect that the particles will essentially become independent (Bose-Einstein condensation), that is, in terms of wave functions:

$$\Psi_N(x_1, \dots, x_N) \approx u^{\otimes N}(x_1, \dots, x_N) := u(x_1)u(x_2)\dots u(x_N). \quad (2.13)$$

Indeed, if $w_N \equiv 0$ then the first eigenfunction Ψ_N of H_N is exactly of this form, with u the first eigenfunction of the one-body operator $-\Delta + |x|^2 - 2\Omega L$. For our interacting Hamiltonian H_N , Ψ_N will *never* be of this form, because there is no reason to believe that *all* the particles ought to be in the state u . Only of the order of N of them would suffice [45]. Nevertheless, the ansatz (2.13) provides the right energy to leading order, as well as the right density matrices, as we will prove in this paper, and as it has already been shown in many other similar situations, see [49, 63, 64] for reviews.

The energy per particle of the fully condensed trial function $u^{\otimes N}$ is given by the Hartree energy functional

$$\begin{aligned} \mathcal{E}_{\Omega,a,N}^H(u) &= \frac{\langle u^{\otimes N}, H_N u^{\otimes N} \rangle}{N} = \int_{\mathbb{R}^2} \left(|\nabla u(x)|^2 + |x|^2 |u(x)|^2 \right) dx - 2\Omega \langle u, Lu \rangle \\ &\quad - \frac{a}{2} \iint_{\mathbb{R}^2 \times \mathbb{R}^2} w_N(x-y) |u(x)|^2 |u(y)|^2 dx dy. \end{aligned} \quad (2.14)$$

The infimum of this functional over the set of all u 's with $\|u\|_{L^2(\mathbb{R}^2)} = 1$,

$$E_{\Omega,a}^H(N) := \inf_{\|u\|_{L^2(\mathbb{R}^2)}=1} \mathcal{E}_{\Omega,a,N}^H(u) \quad (2.15)$$

is thus an upper bound to the many-body energy:

$$E_{\Omega,a}^Q(N) \leq E_{\Omega,a}^H(N).$$

When $N \rightarrow \infty$, since $w_N \rightarrow \delta_0$, the Hartree functional (2.14) *formally* boils down to the trapped nonlinear Gross-Pitaevskii functional $\mathcal{E}_{\Omega,a}^{\text{GP}}$ which we have introduced in (2.1). We can therefore expect that

$$E_{\Omega,a}^H(N) \simeq E_{\Omega}^{\text{GP}}(a),$$

and that their ground states are close. At fixed $a < a_*$ this was shown in [46], but here we need to control the limit $a_N \rightarrow a_*$ at the same time as $N \rightarrow \infty$ and the corresponding estimates will be provided later in the proof of Proposition 4.1.

In [44, 46] we have proved that if $a < a_*$ is fixed and $0 \leq \Omega < 1$, then the many-body ground states of H_N are condensed on the minimizer(s) of the Gross-Pitaevskii functional. In the present paper, we will consider the limit where $a_N \rightarrow a_*$ as $N \rightarrow \infty$. In that case, the Gross-Pitaevskii minimizer blows up at the center $x = 0$ of the trap, as shown in Theorem 2.1 above. We will prove that the many-particle ground state Ψ_N condensates on the exact same function Q_N , hence derive a many-body analogue to the result of Guo and Seiringer [29], at positive rotation.

As usual, the convergence of ground states is formulated using k -particles reduced density matrices, defined for any $\Psi_N \in \mathfrak{H}^N$ by a partial trace

$$\gamma_{\Psi_N}^{(k)} := \text{Tr}_{k+1 \rightarrow N} |\Psi_N\rangle \langle \Psi_N|$$

or, equivalently, $\gamma_{\Psi_N}^{(k)}$ is the trace class operator on \mathfrak{H}^k with kernel

$$\gamma_{\Psi_N}^{(k)}(x_1, \dots, x_k; y_1, \dots, y_k) = \int_{\mathbb{R}^{2(N-k)}} \Psi_N(x_1, \dots, x_k, Z) \overline{\Psi_N(y_1, \dots, y_k, Z)} dZ.$$

Bose-Einstein condensation is properly expressed by the convergence in trace norm

$$\lim_{N \rightarrow \infty} \text{Tr} \left| \gamma_{\Psi_N}^{(k)} - |u^{\otimes k}\rangle\langle u^{\otimes k}| \right| = 0, \quad \forall k \in \mathbb{N}.$$

Our second main result is the following

Theorem 2.3 (Collapse and condensation of the many-body ground state).

Let $0 \leq \Omega < 1$, and $a_N = a_* - N^{-\alpha}$ with

$$0 < \alpha < \min \left\{ \frac{4}{5}\beta, 2(1 - 2\beta) \right\}.$$

Let Ψ_N be any ground state of H_N . Then we have

$$\boxed{\lim_{N \rightarrow \infty} \text{Tr} \left| \gamma_{\Psi_N}^{(k)} - |Q_N^{\otimes k}\rangle\langle Q_N^{\otimes k}| \right| = 0,} \quad (2.16)$$

for all $k \in \mathbb{N}$, where Q_N is the rescaled Gagliardo-Nirenberg optimizer introduced in (2.8). In addition, we have

$$\boxed{E_{\Omega, a_N}^Q(N) = E_{\Omega}^{\text{GP}}(a_N) + o(E_{\Omega}^{\text{GP}}(a_N)) = \sqrt{a_* - a_N} \left(\frac{2\lambda_*^2}{a_*} + o(1) \right).} \quad (2.17)$$

Remark 2.4.

(i) Note that the condition $\alpha < 2(1 - 2\beta)$ implies that we consider mean-field (by opposition to dilute, see [64, Section 5.1]) interactions, i.e. their range is much larger than the average distance between particles. The latter is set by the length scale of the GP ground state:

$$\text{distance between particles} \propto \sqrt{\frac{(a_* - a_N)^{1/2}}{N}} \propto N^{-1/2 - \alpha/4}.$$

We are in fact somewhat deep in the mean-field regime since the transition to dilute interactions would occur when

$$\text{range of the interaction} \propto N^{-\beta} \propto N^{-1/2 - \alpha/4} \propto \text{distance between particles},$$

i.e. at $\beta \sim 1/2 + \alpha/4$.

The condition $\alpha < 4\beta/5$ is used to ensure that the Hartree and GP ground state problems are close in the limit $N \rightarrow \infty$.

(ii) By using the method of [46], we expect that our result can be extended to a dilute regime as well, i.e. to some (not too large)

$$\beta > \frac{1}{2} + \frac{\alpha}{4}.$$

Here we assume $\beta < 1/2 - \alpha/4$ for simplicity as this ensures the stability of the many-body system immediately [46]. The approach we follow is significantly simpler than that of [46], since we use neither the moments estimates, nor the bootstrap on the energy introduced therein. The proof is less flexible however and deeply relies on the uniqueness of the limit profile for GP ground states.

(ii) When $a_N = a_*$, it is not clear to us what happens in the large N limit. While the existence of the ground state of H_N still holds true, the blow-up phenomenon becomes more complicated. The behavior of the minimizers for the Hartree functional $\mathcal{E}_{\Omega, a_*, N}^H(u)$ in (2.14) when $N \rightarrow \infty$ seems to be open. It seems also difficult to look at the case where Ω depends on N as well and approaches its limit of stability $\Omega_N \rightarrow 1$. \diamond

In [46] and several of our previous works [41, 44, 46, 42, 58], our approach was based on the *quantum de Finetti theorem* [68, 33], a non-commutative version of the de Finetti-Hewitt-Savage theorem for exchangeable random variables in probability theory [19, 20, 22, 32]. More precisely, we used a quantitative version of this theorem in finite-dimensional spaces, which we have proved in [43, Lemmas 3.4, 3.6] and which extends several previous results by different authors [38, 25, 12, 11, 31]. The idea of using de Finetti theorems in the context of mean-field limits is not new. For classical systems, this has been put forward by Spohn [67, 54, 37] and then extended in many directions, see, e.g., [6, 36, 63] and the references therein. For the mean-field limit of quantum systems, the older results in this spirit include [24, 70, 60, 62, 73].

Although one can follow the same strategy here, we give below a different proof of Bose-Einstein condensation, based on a Feynman-Hellman-type argument. This method is much less flexible (it relies on the uniqueness of the limit profile for GP ground states), but it allows to cover a wider range for the parameters β and α .

3. COLLAPSE OF THE ROTATING GP MINIMIZER: PROOF OF THEOREM 2.1

In this section we provide the proof of Theorem 2.1. It is convenient to work at the blow up scale, and thus to rewrite everything in terms of

$$v(x) = \sqrt{\varepsilon} u(\sqrt{\varepsilon}x), \quad \varepsilon = \sqrt{a_* - a}.$$

Since the angular momentum L commutes with dilations about the center of rotation, we get

$$\begin{aligned} \mathcal{E}_{\Omega,a}^{\text{GP}}(u) &= \frac{1}{\varepsilon} \int_{\mathbb{R}^2} |\nabla v|^2 + \varepsilon \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx - 2\Omega \langle v, Lv \rangle - \frac{a}{2\varepsilon} \int_{\mathbb{R}^2} |v(x)|^4 dx \\ &= \frac{\mathcal{F}_{\Omega,\varepsilon}(v)}{\varepsilon} \end{aligned}$$

where

$$\begin{aligned} \mathcal{F}_{\Omega,\varepsilon}(v) &= \int_{\mathbb{R}^2} |\nabla v|^2 - \frac{a_*}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx + \varepsilon^2 \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx \\ &\quad + \frac{\varepsilon^2}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx - 2\varepsilon\Omega \langle v, Lv \rangle. \end{aligned} \quad (3.1)$$

We then introduce

$$\boxed{F_{\Omega}(\varepsilon) = \inf_{\|v\|_{L^2}=1} \mathcal{F}_{\Omega,\varepsilon}(v) = \sqrt{a_* - a} E_{\Omega}^{\text{GP}}(a)}$$

and our goal is to prove that

$$F_{\Omega}(\varepsilon) = F_0(\varepsilon) + o(\varepsilon^2) = \varepsilon^2 \left(\frac{2\lambda_*^2}{a_*} + o(1) \right). \quad (3.2)$$

The behavior of $F_0(\varepsilon)$ and its associated unique ground state is studied in [29]. However, even when $\Omega = 0$ we have to prove the convergence of approximate ground states in the sense of (2.6).

Step 1. Convergence to a Gagliardo-Nirenberg optimizer. By rearrangement inequalities, $F_0(\varepsilon)$ has a radial-decreasing minimizer \tilde{v}_ε . Then $L\tilde{v}_\varepsilon = 0$, hence

$$F_\Omega(\varepsilon) \leq F_0(\varepsilon)$$

for every $0 \leq \Omega < 1$. It is the reverse inequality which is not obvious. From the diamagnetic inequality, we have

$$\int_{\mathbb{R}^2} |\nabla v(x) + i\varepsilon\Omega x_\perp v(x)|^2 dx \geq \int_{\mathbb{R}^2} |\nabla|v|(x)|^2 dx$$

and therefore we obtain

$$\begin{aligned} F_\Omega(\varepsilon) &\geq \min_{\substack{v \in H^1(\mathbb{R}^d) \\ \|v\|_{L^2} = 1}} \left\{ \int_{\mathbb{R}^2} |\nabla v|^2 + (1 - \Omega^2)\varepsilon^2 \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx \right. \\ &\quad \left. - \frac{a_* - \varepsilon^2}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx \right\} \\ &= \sqrt{1 - \Omega^2} F_0(\varepsilon). \end{aligned}$$

This lower bound has the right behavior $O(\varepsilon^2)$ but not the right constant.

We can also write the energy in a different form and obtain the following lower bound

$$\begin{aligned} \mathcal{F}_{\Omega,\varepsilon}(v) &= (1 - \Omega) \int_{\mathbb{R}^2} |\nabla v(x)|^2 dx + \Omega \int_{\mathbb{R}^2} |\nabla v(x) + i\varepsilon x_\perp v(x)|^2 dx \\ &\quad - \frac{a_*}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx + (1 - \Omega)\varepsilon^2 \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx + \frac{\varepsilon^2}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx \\ &\geq \int_{\mathbb{R}^2} |\nabla|v|(x)|^2 dx - \frac{a_*}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx \\ &\quad + (1 - \Omega)\varepsilon^2 \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx + \frac{\varepsilon^2}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx \\ &\geq (1 - \Omega)\varepsilon^2 \int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx + \frac{\varepsilon^2}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx. \end{aligned} \tag{3.3}$$

From these bounds we deduce that any sequence $\{v_\varepsilon\} \subset H^1(\mathbb{R}^2)$ such that $\|v_\varepsilon\|_{L^2} = 1$ and $\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) = O(\varepsilon^2)$ (for instance approximate ground states) is bounded in $H^1(\mathbb{R}^2)$, and that $|x|v_\varepsilon$ is bounded in $L^2(\mathbb{R}^2)$. Such sequences are precompact in $L^p(\mathbb{R}^2)$ for all $2 \leq p < \infty$. Therefore, up to a subsequence, we can pass to the limit $v_\varepsilon \rightarrow v$ and obtain

$$\int_{\mathbb{R}^2} |\nabla v(x)|^2 dx - \frac{a_*}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx = 0 \quad \text{with} \quad \int_{\mathbb{R}^2} |v|^2 = 1. \tag{3.4}$$

This means that v belongs to the set of the Gagliardo-Nirenberg optimizers (up to a phase)

$$\mathcal{Q} := \left\{ Q_{\lambda,X}(x) = \lambda Q_*(\lambda(x - X)), \quad \lambda > 0, \quad X \in \mathbb{R}^2 \right\}. \tag{3.5}$$

Here Q_* is the unique positive radial solution to the equation

$$-\Delta Q_* - a_* Q_*^3 = -Q_*.$$

This solution necessarily satisfies $\int_{\mathbb{R}^2} Q_*^2 = 1$ and it is just given by

$$Q_* = \|Q\|^{-1} Q = (a_*)^{-1/2} Q$$

where $-\Delta Q - Q^3 = -Q$. Note that $Q_{\lambda,X}$ solves the equation $-\Delta Q_{\lambda,X} - a_* Q_{\lambda,X}^3 = -\lambda^2 Q_{\lambda,X}$.

Note that we know from the above arguments that v_ε converges to $Q_{\lambda,X}$ strongly in $L^2(\mathbb{R}^2) \cap L^4(\mathbb{R}^2)$ and that

$$\int_{\mathbb{R}^2} |\nabla v_\varepsilon(x)|^2 dx - \frac{a_*}{2} \int_{\mathbb{R}^2} |v_\varepsilon(x)|^4 dx \rightarrow 0.$$

It follows that

$$\int_{\mathbb{R}^2} |\nabla v_\varepsilon(x)|^2 dx \rightarrow \frac{a_*}{2} \int_{\mathbb{R}^2} |Q_{\lambda,X}(x)|^4 dx = \int_{\mathbb{R}^2} |\nabla Q_{\lambda,X}(x)|^2 dx$$

and thus that the limit is also strong in $H^1(\mathbb{R}^2)$. For later purposes, we choose the phase of v_ε such that v_ε is the closest to its limit:

$$\|v_\varepsilon - Q_{\lambda,X}\|_{L^2} = \min_{\theta} \|e^{i\theta} v_\varepsilon - Q_{\lambda,X}\|_{L^2}.$$

This gives the orthogonality condition on the imaginary part of v_ε :

$$\int_{\mathbb{R}^2} Q_{\lambda,X} \operatorname{Im}(v_\varepsilon) = 0. \quad (3.6)$$

We have up to now shown that any sequence $\{v_\varepsilon\}$ such that $\mathcal{F}_\Omega(v_\varepsilon) = O(\varepsilon^2)$ converges to an element of \mathcal{Q} , up to a subsequence and a phase. This is optimal for sequences that have an energy of the order $O(\varepsilon^2)$. In order to determine the possible values of X and λ , we have to assume that v_ε is an approximate ground state of $F_\Omega(\varepsilon)$. In the next three steps we actually assume v_ε is a true ground state, so that we can rely on the variational equation and get better estimates. We return to approximate ground states at the end of the proof.

Step 2. Decay of ground states. For v_ε a true ground state, the Euler-Lagrange equation takes the form

$$-\Delta v_\varepsilon - (a_* - \varepsilon^2)|v_\varepsilon|^2 v_\varepsilon + \varepsilon^2|x|^2 v_\varepsilon - 2\Omega\varepsilon L v_\varepsilon + \mu_\varepsilon v_\varepsilon = 0, \quad (3.7)$$

with the Lagrange multiplier given by

$$\mu_\varepsilon = -\mathcal{F}_\Omega(v_\varepsilon) + \frac{a_* - \varepsilon^2}{2} \int_{\mathbb{R}^2} |v_\varepsilon|^4 \rightarrow \frac{a_*}{2} \int_{\mathbb{R}^2} Q_{\lambda,X}^4 = \lambda^2 > 0. \quad (3.8)$$

Using that $\lambda^2 > 0$ we shall obtain uniform decay estimates *à la* Agmon [2] for v_ε . First we need to show that v_ε converges uniformly.

Lemma 3.1 (Uniform convergence).

The sequence $\{v_\varepsilon\}$ is bounded in $H^2(\mathbb{R}^2)$ and converges to $Q_{\lambda,X}$ strongly in $H^1(\mathbb{R}^2)$ and in $L^\infty(\mathbb{R}^2)$.

Proof. We have

$$(-\Delta + \varepsilon^2|x|^2 - 2\Omega\varepsilon L + \lambda^2) v_\varepsilon = (\lambda^2 - \mu_\varepsilon)v_\varepsilon + (a_* - \varepsilon^2)|v_\varepsilon|^2 v_\varepsilon$$

where the right side is bounded in $L^2(\mathbb{R}^2)$. By the Cauchy-Schwarz inequality and the fact that L commutes with $-\Delta + \varepsilon^2|x|^2$, we have

$$2\varepsilon|L| \leq -\Delta + \varepsilon^2|x|^2$$

and

$$(2\Omega\varepsilon L)^2 \leq \Omega^2 (-\Delta + \varepsilon^2|x|^2)^2$$

or, equivalently,

$$\left\| 2\Omega\varepsilon L (-\Delta + \varepsilon^2|x|^2)^{-1} \right\| \leq \Omega < 1. \quad (3.9)$$

By the resolvent formula this proves that

$$\left\| (-\Delta + \varepsilon^2|x|^2) (-\Delta + \varepsilon^2|x|^2 - 2\Omega\varepsilon L)^{-1} \right\| \leq \frac{1}{1-\Omega}$$

and similarly that

$$\left\| (-\Delta + \varepsilon^2|x|^2 + \lambda^2) (-\Delta + \varepsilon^2|x|^2 - 2\Omega\varepsilon L + \lambda^2)^{-1} \right\| \leq \frac{1}{1-\Omega}.$$

Using

$$\|(-\Delta + \varepsilon^2|x|^2)f\|_{L^2}^2 = \|\Delta f\|_{L^2}^2 + \varepsilon^4 \| |x|^2 f \|_{L^2}^2 + \varepsilon^2 \|x \cdot \nabla f\|_{L^2}^2 - 2\varepsilon^2 \|f\|_{L^2}^2$$

we get

$$(-\Delta + \varepsilon^2|x|^2 + \lambda^2)^2 \geq (-\Delta + \lambda^2)^2 - 2\varepsilon^2 \geq (-\Delta + \lambda^2/2)^2$$

for $\varepsilon \ll 1$, and thus

$$\left\| (-\Delta + \lambda^2/2) (-\Delta + \varepsilon^2|x|^2 - 2\Omega\varepsilon L + \lambda^2)^{-1} \right\| \leq \frac{1}{1-\Omega}.$$

Inserting in Equation (3.7), this proves that v_ε is bounded in $H^2(\mathbb{R}^2)$. Since v_ε already converges strongly in $L^2(\mathbb{R}^2)$, it also converges strongly in $H^1(\mathbb{R}^2)$ and in $L^\infty(\mathbb{R}^2)$, by interpolation. \square

Next we can prove an exponential decay estimate. We need it later only to obtain the strong convergence $|x|v_\varepsilon \rightarrow |x|Q_{\lambda,X}$ in $L^2(\mathbb{R}^2)$.

Lemma 3.2 (Exponential decay).

The function v_ε satisfies

$$\int_{\mathbb{R}^2} e^{\lambda|x|} |v_\varepsilon(x)|^2 dx + \int_{\mathbb{R}^2} |\nabla e^{\frac{\lambda}{2}|x|} v_\varepsilon|^2 \leq C,$$

for a constant C independent of ε . In particular, $|x|v_\varepsilon \rightarrow |x|Q_{\lambda,X}$ strongly in $L^2(\mathbb{R}^2)$.

Proof. It is well known that v_ε is analytic with all its derivatives decaying fast at infinity. We seek here for an explicit bound, independent of ε . We use that

$$\begin{aligned} -\operatorname{Re} \langle v_\varepsilon, e^{\alpha|x|} \Delta v_\varepsilon \rangle &= -\frac{1}{2} \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\overline{v_\varepsilon}(x) \Delta v_\varepsilon(x) + v_\varepsilon(x) \Delta \overline{v_\varepsilon}(x) \right) dx \\ &= -\frac{1}{2} \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\Delta |v_\varepsilon|^2(x) - 2|\nabla v_\varepsilon(x)|^2 \right) dx \\ &= -\frac{1}{2} \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\left(\frac{\alpha}{|x|} + \alpha^2 \right) |v_\varepsilon(x)|^2 - 2|\nabla v_\varepsilon(x)|^2 \right) dx \\ &= \int_{\mathbb{R}^2} e^{\alpha|x|} |\nabla v_\varepsilon|^2 - \frac{1}{2} \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\frac{\alpha}{|x|} + \alpha^2 \right) |v_\varepsilon(x)|^2 dx \\ &= \int_{\mathbb{R}^2} |\nabla e^{\frac{\alpha}{2}|x|} v_\varepsilon|^2 - \frac{\alpha^2}{4} \int_{\mathbb{R}^2} e^{\alpha|x|} |v_\varepsilon(x)|^2 dx. \end{aligned}$$

Then we integrate the Euler-Lagrange equation (3.7) against $e^{\alpha|x|}\bar{v}_\varepsilon$ and obtain

$$\begin{aligned} 0 &= \int_{\mathbb{R}^2} |\nabla e^{\frac{\alpha}{2}|x}|v_\varepsilon|^2 + \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\varepsilon^2|x|^2 - (a_* - \varepsilon^2)|v_\varepsilon|^2 + \mu_\varepsilon - \frac{\alpha^2}{4} \right) |v_\varepsilon(x)|^2 dx \\ &\quad - 2\Omega\varepsilon \left\langle e^{\frac{\alpha}{2}|x}|v_\varepsilon, Le^{\frac{\alpha}{2}|x}|v_\varepsilon \right\rangle \\ &\geq (1 - \Omega) \int_{\mathbb{R}^2} |\nabla e^{\frac{\alpha}{2}|x}|v(x)|^2 dx + (1 - \Omega)\varepsilon^2 \int_{\mathbb{R}^2} e^{\alpha|x|}|x|^2|v(x)|^2 dx \\ &\quad + \int_{\mathbb{R}^2} e^{\alpha|x|} \left(\mu_\varepsilon - (a_* - \varepsilon^2)|v_\varepsilon|^2 - \frac{\alpha^2}{4} \right) |v_\varepsilon(x)|^2 dx. \end{aligned}$$

Choosing $\alpha = \lambda$ and using the uniform convergence of v_ε towards $Q_{\lambda,X}$, we can find a radius R independent of ε such that

$$\mu_\varepsilon - (a_* - \varepsilon^2)|v_\varepsilon|^2 - \frac{\alpha^2}{4} \geq \frac{\lambda^2}{2}, \quad \forall |x| \geq R,$$

and then

$$\frac{\lambda^2}{2} \int_{\mathbb{R}^2 \setminus B_R} e^{\lambda|x|}|v_\varepsilon(x)|^2 dx + (1 - \Omega) \int_{\mathbb{R}^2} |\nabla e^{\frac{\lambda}{2}|x}|v_\varepsilon|^2 \leq e^{\lambda R} \left(\mu_\varepsilon + \frac{\lambda^2}{4} + a_* \|v_\varepsilon\|_{L^\infty(B_R)}^2 \right)$$

for all $\varepsilon > 0$ small enough. This proves the desired exponential decay estimate. The strong convergence of $|x|v_\varepsilon$ in $L^2(\mathbb{R}^2)$ then follows by interpolation. \square

Step 3. The imaginary part is (very) small. We split v_ε into real and imaginary parts

$$v_\varepsilon = q_\varepsilon + ir_\varepsilon$$

and get bounds on r_ε using energy estimates (we could similarly use the equation (3.7)). Observe that,

$$\langle v_\varepsilon, Lv_\varepsilon \rangle = - \int_{\mathbb{R}^2} x^\perp \cdot \text{Im}(v_\varepsilon \nabla \bar{v}_\varepsilon) = - \int_{\mathbb{R}^2} x^\perp \cdot (r_\varepsilon \nabla q_\varepsilon - q_\varepsilon \nabla r_\varepsilon) = 2 \int_{\mathbb{R}^2} x^\perp \cdot q_\varepsilon \nabla r_\varepsilon \quad (3.10)$$

where we have integrated by parts and used that $\text{div } x^\perp = 0$. Thus

$$|\langle v_\varepsilon, Lv_\varepsilon \rangle| \leq C \|\nabla r_\varepsilon\|_{L^2}.$$

Here we have used the fact that $|x|q_\varepsilon$ is bounded in $L^2(\mathbb{R}^2)$. Then the energy reads

$$\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) \geq \int_{\mathbb{R}^2} |\nabla q_\varepsilon|^2 + \int_{\mathbb{R}^2} |\nabla r_\varepsilon|^2 - \frac{a_*}{2} \int_{\mathbb{R}^2} (q_\varepsilon^4 + r_\varepsilon^4 + 2q_\varepsilon^2 r_\varepsilon^2) - C\varepsilon \|\nabla r_\varepsilon\|_{L^2}.$$

Since $q_\varepsilon \rightarrow Q_{\lambda,X}$ and $r_\varepsilon \rightarrow 0$ uniformly by Lemma 3.1, we obtain

$$\int_{\mathbb{R}^2} |q_\varepsilon^2 - Q_{\lambda,X}^2| r_\varepsilon^2 + \int_{\mathbb{R}^2} r_\varepsilon^4 = o(\|r_\varepsilon\|_{L^2}^2).$$

Moreover, using the Gagliardo-Nirenberg inequality (2.2) for the real part q_ε , we have

$$\int_{\mathbb{R}^2} |\nabla q_\varepsilon|^2 - \frac{a_*}{2} \int_{\mathbb{R}^2} |q_\varepsilon|^2 \geq \left(\int_{\mathbb{R}^2} |\nabla q_\varepsilon|^2 \right) \left(1 - \int_{\mathbb{R}^2} |q_\varepsilon|^2 \right) = (\lambda^2 + o(1)) \int_{\mathbb{R}^2} |r_\varepsilon|^2.$$

Here in the second estimate we have used the facts that $\|q_\varepsilon\|_{L^2}^2 + \|r_\varepsilon\|_{L^2}^2 = 1$ and that $q_\varepsilon \rightarrow Q_{\lambda,X}$ strongly in $H^1(\mathbb{R}^2)$ by Lemma 3.1. Thus we can bound the energy from below as

$$\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) \geq \int_{\mathbb{R}^2} |\nabla r_\varepsilon|^2 - a_* \int_{\mathbb{R}^2} Q_{\lambda,X}^2 r_\varepsilon^2 + (\lambda^2 + o(1)) \int_{\mathbb{R}^2} |r_\varepsilon|^2 - C\varepsilon \|\nabla r_\varepsilon\|_{L^2}. \quad (3.11)$$

Now we use some non-degeneracy property of $Q_{\lambda,X}$ [72, 53, 9, 27]. Since $Q_{\lambda,X}$ is positive, it must be the first eigenfunction of the operator

$$\mathcal{L}_- := -\Delta - a_* Q_{\lambda,X}^2 + \lambda^2 \quad (3.12)$$

and the corresponding eigenvalue 0 is non-degenerate [47, Cor. 11.9]. In particular, we get

$$\langle f, \mathcal{L}_- f \rangle_{L^2(\mathbb{R}^2)} \geq \lambda_2 \|f\|_{L^2(\mathbb{R}^2)}^2 \quad (3.13)$$

for all f orthogonal to $Q_{\lambda,X}$ where $\lambda_2 > 0$ is the second eigenvalue of \mathcal{L}_- . Since on the other hand

$$\langle f, \mathcal{L}_- f \rangle_{L^2(\mathbb{R}^2)} \geq \|\nabla f\|_{L^2(\mathbb{R}^2)}^2 - a_* \|Q_{\lambda,X}\|_{L^\infty(\mathbb{R}^2)}^2 \|f\|_{L^2(\mathbb{R}^2)}^2, \quad (3.14)$$

we may combine (3.13) with (3.14) (add a large constant times the second inequality to the first one) and obtain the well known estimate [72]

$$\langle f, \mathcal{L}_- f \rangle_{L^2(\mathbb{R}^2)} \geq c \|f\|_{H^1(\mathbb{R}^2)}^2 \quad (3.15)$$

for a constant $c > 0$ and all f orthogonal to $Q_{\lambda,X}$. Inserting (3.15) in (3.11) using the fact that r_ε is orthogonal to $Q_{\lambda,X}$ as we have seen in (3.6), we obtain

$$\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) \geq c_1 \|r_\varepsilon\|_{H^1}^2 - C\varepsilon \|\nabla r_\varepsilon\|_{L^2}$$

for a constant $c_1 > 0$. Combining with the energy upper bound $\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) = O(\varepsilon^2)$ we conclude that

$$\|r_\varepsilon\|_{H^1(\mathbb{R}^2)} \leq C\varepsilon. \quad (3.16)$$

Step 4. Change of gauge and convergence of ground states. Since $|x|q_\varepsilon$ converges to $|x|Q_{\lambda,X}$ strongly in L^2 by Lemma 3.2, we deduce from (3.10) and (3.16) that

$$\langle v_\varepsilon, Lv_\varepsilon \rangle = 2 \int_{\mathbb{R}^2} x^\perp \cdot Q_{\lambda,X} \nabla r_\varepsilon + o(\varepsilon) = -2 \int_{\mathbb{R}^2} x^\perp \cdot \nabla Q_{\lambda,X} r_\varepsilon + o(\varepsilon).$$

But $Q_{\lambda,X} = \lambda Q_*(\lambda(x - X))$ with Q_* a radial function, hence

$$(x^\perp - X^\perp) \cdot \nabla Q_{\lambda,X} = 0.$$

Inserting in the above, using (3.16) and strong L^2 -convergence of $|x|q_\varepsilon$ again we obtain

$$\langle v_\varepsilon, Lv_\varepsilon \rangle = 2 \int_{\mathbb{R}^2} X^\perp \cdot q_\varepsilon \nabla r_\varepsilon + o(\varepsilon).$$

Inserting this in the energy gives

$$\begin{aligned} \mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) &= \int_{\mathbb{R}^2} |\nabla v_\varepsilon|^2 - \frac{a_*}{2} \int_{\mathbb{R}^2} |v_\varepsilon|^4 + 2\varepsilon \Omega \int_{\mathbb{R}^2} X^\perp \cdot \text{Im}(v_\varepsilon \nabla \bar{v}_\varepsilon) \\ &\quad + \varepsilon^2 \left(\frac{1}{2} \int_{\mathbb{R}^2} |v_\varepsilon|^4 + \int_{\mathbb{R}^2} |x|^2 |v_\varepsilon|^2 \right) + o(\varepsilon^2). \end{aligned}$$

Now, define a new function f_ε by setting

$$v_\varepsilon(x) = f_\varepsilon(x) e^{i\varepsilon \Omega X^\perp \cdot x}$$

and observe that

$$\int_{\mathbb{R}^2} |\nabla v_\varepsilon|^2 + 2\varepsilon \Omega \int_{\mathbb{R}^2} X^\perp \cdot \text{Im}(v_\varepsilon \nabla \bar{v}_\varepsilon) = \int_{\mathbb{R}^2} |\nabla f_\varepsilon|^2 - \varepsilon^2 \Omega^2 |X|^2 \int_{\mathbb{R}^2} |f_\varepsilon|^2.$$

Using the optimal Gagliardo-Nirenberg inequality again and the convergence $f_\varepsilon \rightarrow Q_{\lambda,X}$, we obtain

$$\liminf_{\varepsilon \rightarrow 0} \frac{\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon)}{\varepsilon^2} \geq \frac{1}{2} \int_{\mathbb{R}^2} Q_{\lambda,X}^4 + \int_{\mathbb{R}^2} (|x|^2 - \Omega^2 |X|^2) Q_{\lambda,X}^2.$$

Recalling that $Q_{\lambda,X} = \lambda Q_*(\lambda(x - X))$ for a radial function Q_* this finally yields

$$\liminf_{\varepsilon \rightarrow 0} \frac{\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon)}{\varepsilon^2} \geq (1 - \Omega^2)|X|^2 + \frac{1}{\lambda^2} \int_{\mathbb{R}^2} |x|^2 |Q_*(x)|^2 dx + \frac{\lambda^2}{2} \int_{\mathbb{R}^2} |Q_*(x)|^4 dx.$$

Since $\Omega^2 < 1$, the minimum of the right side is attained for $X = 0$ and $\lambda = \lambda_*$ as in [29]. This also concludes the proof of (3.2), hence of (2.5). Also, this shows that any sequence of minimizers must, modulo rescaling, choice of a constant phase (in (3.6)) and passing to a subsequence, converge strongly in H^1 to $Q_{\lambda^*,0}$. By uniqueness of the limit we conclude that passing to a subsequence is unnecessary, which concludes the proof of (2.7) for true ground states.

Step 5. Convergence of approximate ground states for $\Omega = 0$. Here we assume $\Omega = 0$ and show that any sequence $\{v_\varepsilon\}$ such that

$$\mathcal{F}_{0,\varepsilon}(v_\varepsilon) = F_0(\varepsilon) + o(\varepsilon^2)$$

must converge to $Q_{\lambda^*,0}$, as we have proved before for the exact minimizers, using the Euler-Lagrange equation. Indeed, from Step 1 we already know that $v_\varepsilon \rightarrow Q_{\lambda,X}$ after extraction of a subsequence and choice of a good phase. Hence we have

$$\begin{aligned} \mathcal{F}_{0,\varepsilon}(v_\varepsilon) &\geq \varepsilon^2 \left(\int_{\mathbb{R}^2} |x|^2 |v(x)|^2 dx + \frac{1}{2} \int_{\mathbb{R}^2} |v(x)|^4 dx \right) \\ &\geq \varepsilon^2 \left(\int_{\mathbb{R}^2} |x|^2 Q_{\lambda,X}(x)^2 dx + \frac{1}{2} \int_{\mathbb{R}^2} Q_{\lambda,X}(x)^4 dx \right) + o(\varepsilon^2) \\ &= \varepsilon^2 \left(|X|^2 + \frac{1}{\lambda^2} \int_{\mathbb{R}^2} |x|^2 Q_*(x)^2 dx + \frac{\lambda^2}{2} \int_{\mathbb{R}^2} Q_*(x)^4 dx \right) + o(\varepsilon^2). \end{aligned}$$

Again, the minimum of the term in the parenthesis is attained uniquely for $X = 0$ and $\lambda = \lambda_*$. Since

$$\mathcal{F}_{0,\varepsilon}(v_\varepsilon) = F_0(\varepsilon) + o(\varepsilon^2) = \varepsilon^2 \left(\frac{1}{\lambda_*^2} \int_{\mathbb{R}^2} |x|^2 Q_*(x)^2 dx + \frac{\lambda_*^2}{2} \int_{\mathbb{R}^2} Q_*(x)^4 dx \right)$$

we must have $X = 0$ and $\lambda = \lambda_*$. The limit being unique, the whole sequence must converge to $Q_{\lambda^*,0}$. By usual arguments the limit must be strong in $H^1(\mathbb{R}^2)$.

Step 6. Convergence of approximate ground states for $0 < \Omega < 1$. Next we turn to the rotating case $\Omega < 1$. As before we take a sequence $\{v_\varepsilon\}$ such that $\mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) = F_\Omega(\varepsilon) + o(\varepsilon^2) = F_0(\varepsilon) + o(\varepsilon^2)$. We artificially increase the rotation speed by choosing a $\eta < 1$ be such that $\Omega/\eta < 1$ and we remark that

$$\begin{aligned} F_\Omega(\varepsilon) + o(\varepsilon^2) &= \mathcal{F}_{\Omega,\varepsilon}(v_\varepsilon) = \eta \mathcal{F}_{\Omega/\eta,\varepsilon}(v_\varepsilon) + (1 - \eta) \mathcal{F}_{0,\varepsilon}(v_\varepsilon) \\ &\geq \eta F_{\Omega/\eta}(\varepsilon) + (1 - \eta) F_0(\varepsilon). \end{aligned}$$

Since the terms $\varepsilon^{-2} F_\Omega(\varepsilon)$, $\varepsilon^{-2} F_{\Omega/\eta}(\varepsilon)$ and $\varepsilon^{-2} F_0(\varepsilon)$ all have the same limit, this proves that $\mathcal{F}_{\Omega/\eta,\varepsilon}(v_\varepsilon) = F_{\Omega/\eta}(\varepsilon) + o(\varepsilon^2)$ and $\mathcal{F}_{0,\varepsilon}(v_\varepsilon) = F_0(\varepsilon) + o(\varepsilon^2)$. But then $\{v_\varepsilon\}$ is a sequence of approximate ground states in the case $\Omega = 0$ and we can apply the previous step.

This concludes the proof of Theorem 2.1. □

4. COLLAPSE OF THE MANY-BODY GROUND STATE: PROOF OF THEOREM 2.3

This section is devoted to the proof of Theorem 2.3. We start with the convergence of the ground state energy, and then settle some energy estimates for the ground state.

Step 1. Convergence of the many-body ground state energy. We provide the proof of the convergence of $E_{\Omega, a_N}^{\text{Q}}(N)$, using a method described in [40, Section 3]. The precise statement is the following.

Proposition 4.1 (Convergence of the many-body ground state energy).

Let $\beta \in (0, 1/2)$ and $a_* - a_N = N^{-\alpha}$ with

$$0 < \alpha < \min \left\{ \frac{4}{5}\beta, 2(1 - 2\beta) \right\}. \quad (4.1)$$

Then

$$E_{\Omega, a_N}^{\text{Q}}(N) = E_{\Omega}^{\text{GP}}(a_N) + o(E_{\Omega}^{\text{GP}}(a_N)) = \sqrt{a_* - a_N} \left(\frac{2\lambda^2}{a_*} + o(1) \right). \quad (4.2)$$

Proof. We start with the lower bound. From the arguments in [40, Section 3] and the fact that the Fourier transform of w_N satisfies $|\widehat{w}_N| \leq CN^{2\beta}$ we have

$$E_{\Omega, a}^{\text{Q}}(N) \geq \min_{\substack{\gamma = \gamma^* \geq 0 \\ \text{Tr } \gamma = 1}} \left\{ \text{Tr} [(-\Delta + |x|^2 - 2\Omega L)\gamma] - \frac{a}{2} \int_{\mathbb{R}^d} \int_{\mathbb{R}^d} w_N(x-y) \rho_{\gamma}(x) \rho_{\gamma}(y) dx dy \right\} - CN^{2\beta-1}. \quad (4.3)$$

Here the one-body density is defined by writing

$$\rho_{\gamma} = \sum_j n_j |u_j|^2$$

with the spectral decomposition $\gamma = \sum_j n_j |u_j\rangle\langle u_j|$ ($0 \leq n_j \leq 1$ and $\sum_j n_j = 1$).

Note that the arguments in [40, Section 3] contain two ingredients. One is to get a lower bound involving the mean-field interaction following an idea from [39, 50], using auxiliary classical particles that repel each other in order to model the attractive part of the interaction. The other ingredient is to apply the Hoffmann-Ostenhof inequality¹ to the kinetic energy in order to get the Hartree energy. Here we cannot use the second part when $\Omega \neq 0$. We thus bypass it and only bound the interaction from below. The price to pay is that we end up with the mixed state type Hartree energy on the right side of (4.3).

By the Cauchy-Schwarz inequality, we have

$$\begin{aligned} \iint |u(x)|^2 |u(y)|^2 w_N(x-y) dx dy &\leq \frac{1}{2} \iint (|u(x)|^4 + |u(y)|^4) w_N(x-y) dx dy \\ &= \int |u(x)|^4 dx \end{aligned} \quad (4.4)$$

for any $u \in H^1(\mathbb{R}^2)$. Applying this to $u = \sqrt{\rho_{\gamma}}$, we therefore get

$$E_{\Omega, a}^{\text{Q}}(N) \geq \min_{\substack{\gamma = \gamma^* \geq 0 \\ \text{Tr } \gamma = 1}} \left\{ \text{Tr} [(-\Delta + |x|^2 - 2\Omega L)\gamma] - \frac{a}{2} \int_{\mathbb{R}^d} \rho_{\gamma}(x)^2 dx \right\} - CN^{2\beta-1}. \quad (4.5)$$

Here $\rho_{\gamma}(x) := \gamma(x, x)$ for every $\gamma = \gamma^* \geq 0$ with $\text{Tr}(\gamma) = 1$. The usual Gross-Pitaevskii energy (2.1) corresponds to $\gamma = |u\rangle\langle u|$, a rank one projection. Now we remark that the minimum on the right of (4.5) is nothing but the original Gross-Pitaevskii minimum.

¹Bounding the full kinetic energy from below by that of the one-body density $\sqrt{\rho_{\Psi_N}}$.

Lemma 4.2 (Mixed Gross-Pitaevskii energy).

For every $0 \leq \Omega < 1$ and every $0 \leq a < a_*$, we have

$$\min_{\substack{\gamma = \gamma_* \geq 0 \\ \text{Tr } \gamma = 1}} \left\{ \text{Tr} [(-\Delta + |x|^2 - 2\Omega L)\gamma] - \frac{a}{2} \int_{\mathbb{R}^d} \rho_\gamma(x)^2 dx \right\} = E_\Omega^{\text{GP}}(a). \quad (4.6)$$

Proof of Lemma 4.2. This follows from the concavity of the energy with respect to γ . Indeed, writing $\gamma = \sum_j n_j |u_j\rangle\langle u_j|$ with $0 \leq n_j \leq 1$ and $\sum_j n_j = 1$, we have $\rho_\gamma = \sum_j n_j |u_j|^2$, hence

$$\text{Tr} [(-\Delta + |x|^2 - 2\Omega L)\gamma] - \frac{a}{2} \int_{\mathbb{R}^d} \rho_\gamma(x)^2 dx \geq \sum_j n_j \mathcal{E}_{\Omega,a}^{\text{GP}}(u_j) \geq E_\Omega^{\text{GP}}(a)$$

as we wanted. \square

Inserting in (4.5) and using the behavior of $E_\Omega^{\text{GP}}(a_N)$ proved in Theorem 2.1, we get the simple lower bound

$$\boxed{E_{\Omega,a_N}^{\text{Q}}(N) \geq E_\Omega^{\text{GP}}(a_N) \left(1 - CN^{2\beta-1}(a_* - a_N)^{-1/2}\right)}.$$

The error term $N^{2\beta-1}(a_* - a_N)^{-1/2}$ goes to zero when $a_* - a_N = N^{-\alpha}$ and $\alpha < 2(1 - 2\beta)$.

Now we turn to the upper bound. By the variational principle

$$E_{\Omega,a}^{\text{Q}}(N) = \inf_{\Psi \in \mathfrak{H}^N, \|\Psi\|=1} \frac{\langle \Psi, H_N \Psi \rangle}{N} \leq \inf_{\|u\|_{L^2}=1} \frac{\langle u^{\otimes N}, H_N u^{\otimes N} \rangle}{N} = E_{\Omega,a}^{\text{H}}(N).$$

We therefore need to bound the Hartree functional from above by the Gross-Pitaevskii functional. Introducing the variable $z = N^\beta(x - y)$, we may write

$$\begin{aligned} & \iint_{\mathbb{R}^2 \times \mathbb{R}^2} |u(x)|^2 N^{2\beta} w(N^\beta(x - y)) |u(y)|^2 dx dy - \left(\int_{\mathbb{R}^2} w \right) \int_{\mathbb{R}^2} |u(x)|^4 dx \\ &= \iint_{\mathbb{R}^2 \times \mathbb{R}^2} |u(x)|^2 w(z) \left(|u(x - N^{-\beta}z)|^2 - |u(x)|^2 \right) dx dz \\ &= \iint_{\mathbb{R}^2 \times \mathbb{R}^2} |u(x)|^2 w(z) \left(\int_0^1 (\nabla |u|^2)(x - tN^{-\beta}z) \cdot (N^{-\beta}z) dt \right) dx dz. \end{aligned}$$

Using $|\nabla |u|^2| \leq 2|\nabla u| \cdot |u|$ and Hölder's inequality, we find

$$\begin{aligned} & \int_{\mathbb{R}^2} |u(x)|^2 |(\nabla |u|^2)(x - tN^{-\beta}z)| dx \\ & \leq 2 \left(\int_{\mathbb{R}^2} |u(x)|^6 dx \right)^{1/3} \left(\int_{\mathbb{R}^2} |\nabla u(x - tN^{-\beta}z)|^2 dx \right)^{1/2} \left(\int_{\mathbb{R}^2} |u(x - tN^{-\beta}z)|^6 dx \right)^{1/6} \\ & = 2 \|\nabla u\|_{L^2(\mathbb{R}^2)} \|u\|_{L^6(\mathbb{R}^2)}^3. \end{aligned}$$

Thus

$$\begin{aligned} |\mathcal{E}_{\Omega,a_N,N}^{\text{H}}(u) - \mathcal{E}_{\Omega,a_N}^{\text{GP}}(u)| &= \frac{a_N}{2} \left| \iint_{\mathbb{R}^2 \times \mathbb{R}^2} |u(x)|^2 w_N(x - y) - \int_{\mathbb{R}^2} |u(x)|^4 dx \right| \\ &= \frac{a_N}{2} \left| \iint_{\mathbb{R}^2 \times \mathbb{R}^2} |u(x)|^2 w(z) \left(\int_0^1 (\nabla |u|^2)(x - tN^{-\beta}z) \cdot (N^{-\beta}z) dt \right) dx dz \right| \\ &\leq a_N N^{-\beta} \left(\int_{\mathbb{R}^2} |w(z)| |z| dz \right) \|\nabla u\|_{L^2(\mathbb{R}^2)} \|u\|_{L^6(\mathbb{R}^2)}^3 \\ &\leq CN^{-\beta} \|\nabla u\|_{L^2(\mathbb{R}^2)} \|u\|_{L^6(\mathbb{R}^2)}^3. \end{aligned}$$

Now choosing the trial state Q_N as in (2.8) we find that

$$\begin{aligned} \mathcal{E}_{\Omega, a_N, N}^H(Q_N) &\leq \mathcal{E}_{\Omega, a_N}^{\text{GP}}(Q_N) + CN^{-\beta} \|\nabla Q_N\|_{L^2} \|Q_N\|_{L^6}^3 \\ &= \sqrt{a_* - a_N} \left(\frac{2\lambda^2}{a_*} + CN^{-\beta} (a_* - a_N)^{-\frac{5}{4}} \right). \end{aligned}$$

The error term goes to zero when $a_* - a_N \geq N^{-\alpha}$ with $\alpha < 4\beta/5$. We have proved the upper bound

$$\boxed{E_{\Omega, a_N}^Q(N) \leq E_{\Omega}^{\text{GP}}(a_N) \left(1 + CN^{-\beta} (a_* - a_N)^{-\frac{5}{4}} \right)}$$

which concludes the proof of Proposition 4.1. \square

Step 2. Convergence of reduced density matrices. This is a Feynman-Hellmann-type argument. Let $\eta > 0$ be a small parameter to be fixed later on and A be a bounded self-adjoint operator on $L^2(\mathbb{R}^2)$. Consider the perturbed Hamiltonian

$$H_{N, \eta} = \sum_{j=1}^N (-\Delta_{x_j} + |x_j|^2 - 2\Omega L_{x_j} + \eta A_j) - \frac{a}{N-1} \sum_{1 \leq i < j \leq N} w_N(x_i - x_j) \quad (4.7)$$

with ground-state energy per particle denoted E_{η}^Q hereafter. Introduce the associated Gross-Pitaevskii energy functional (we drop some Ω, a subscripts during this proof, for lightness of notation)

$$\mathcal{E}_{\eta}^{\text{GP}}(u) = \int_{\mathbb{R}^2} |\nabla u(x)|^2 dx + \int_{\mathbb{R}^2} |x|^2 |u(x)|^2 dx - 2\Omega \langle u, Lu \rangle + \eta \langle u, Au \rangle - \frac{a}{2} \int_{\mathbb{R}^2} |u(x)|^4 dx. \quad (4.8)$$

In what follows we denote by u_{η} a ground state for the latter and $E_{\eta}^{\text{GP}} = \mathcal{E}_{\eta}^{\text{GP}}(u_{\eta})$ the corresponding ground-state energy.

Let Ψ_N be a ground state for $H_N = H_{N,0}$ and $\gamma_{\Psi_N}^{(1)}$ its one-body reduced density matrix. We write

$$\begin{aligned} \eta \text{Tr} \left[A \gamma_{\Psi_N}^{(1)} \right] &= N^{-1} \left(\langle \Psi_N | H_{N, \eta} | \Psi_N \rangle - \langle \Psi_N | H_{N, 0} | \Psi_N \rangle \right) \\ &\geq E_{\eta}^Q - E_0^Q \\ &\geq E_{\eta}^{\text{GP}} - E_0^{\text{GP}} + O(N^{2\beta-1}) + O(N^{3\alpha/4-\beta}) \\ &\geq \mathcal{E}_{\eta}^{\text{GP}}(u_{\eta}) - \mathcal{E}_0^{\text{GP}}(u_{\eta}) + O(N^{2\beta-1}) + O(N^{3\alpha/4-\beta}) \\ &= \eta \langle u_{\eta} | A | u_{\eta} \rangle + O(N^{2\beta-1}) + O(N^{3\alpha/4-\beta}). \end{aligned} \quad (4.9)$$

The first inequality is the variational principle, the second uses the estimates of the previous step. In that regard, observe that the energy lower bound applies *mutatis mutandis* to the problem perturbed by ηA . The error term in (4.3) solely comes from applying the Lévy-Leblond method to the interaction as in [40, Section 3]. It is therefore independent of the one-body term (in particular, of η and A). Lemma 4.2 generalizes to the perturbed functional, for the only property of the one-body energy used in its proof is its linearity in the density matrix. The third inequality in (4.9) is the variational principle again.

Under the assumption that

$$\alpha < \min \left(\frac{4\beta}{5}, 2(1 - 2\beta) \right)$$

one can pick some $\eta = \eta_N \rightarrow 0$ as $N \rightarrow \infty$, such that

$$\eta^{-1}N^{2\beta-1} + \eta^{-1}N^{3\alpha/4-\beta} \xrightarrow{N \rightarrow \infty} 0$$

and also

$$\eta = o(E^{\text{GP}}) = o(\sqrt{a_* - a_N}) = o(N^{-\alpha/2}).$$

Then, dividing (4.9) by η and repeating the argument with A changed to $-A$ yields

$$\langle u_\eta | A | u_\eta \rangle + o(1) \leq \text{Tr} \left[A \gamma_{\Psi_N}^{(1)} \right] \leq \langle u_{-\eta} | A | u_{-\eta} \rangle + o(1). \quad (4.10)$$

On the other hand, with the above choice of η , since

$$\mathcal{E}_0^{\text{GP}}(u_\eta) = \mathcal{E}_\eta^{\text{GP}}(u_\eta) + O(\eta \|A\|) \leq \mathcal{E}_\eta^{\text{GP}}(u_0) + O(\eta \|A\|) = E_0^{\text{GP}} + O(\eta \|A\|),$$

it follows that (u_η) and $(u_{-\eta})$ are sequences of quasi-minimizers for $\mathcal{E}_0^{\text{GP}}$. We may apply Theorem 2.1 to them, and thus both sequences satisfy (2.7). Combining with (4.10), we get, after a dilation of space variables, trace-class weak- \star convergence of $\gamma_{\Psi_N}^{(1)}$ to $|Q_N\rangle\langle Q_N|$. Since no mass is lost in the limit, the convergence must hold in trace-class norm, which gives (2.16) for $k = 1$.

To obtain (2.16) for $k > 1$, observe that, after dilation, $\gamma_{\Psi_N}^{(1)}$ converges in trace-class norm to a rank-one operator. It is well-known that this implies convergence of higher order density matrices to tensor powers of the limiting operator (see e.g. the discussion following [49, Theorem 7.1] or [41, Cor. 2.4]). \square

Remark 4.3.

Note that we have been able to obtain the convergence of density matrices from that of the energy by a rather soft argument. What makes this possible is Lemma 4.2 and the uniqueness of (the limit of) the GP minimizer. Lemma 4.2 relies strongly on the fact that the interaction is attractive. For repulsive interactions², the argument can be adapted provided the one-body part is positivity-preserving (hence, without rotation), using that the bosonic and boltzonic minimization problems then coincide [48, Theorem 3.3]. \diamond

APPENDIX A. EXTENSION TO ANHARMONIC POTENTIALS

The arguments given in this paper can be extended in various directions. One possibility is to consider anharmonic potentials. For completeness we state here the corresponding result when the external potential is chosen in the form

$$V(x) = c_0|x|^s$$

but we expect similar results when V has a unique minimizer and behaves like this in a neighborhood of this point, similarly to what was done in [29]. When $s \neq 2$ the limit $a \rightarrow a_*$ requires to have $\Omega = 0$. Although a stronger confinement $s > 2$ can control the rotating gas at infinity, it is not sufficient to control rotating effects near the blow up point. So we do not consider any rotation here. The many-particle Hamiltonian then takes the form

$$\tilde{H}_N = \sum_{j=1}^N (-\Delta_{x_j} + c_0|x_j|^s) - \frac{a}{N-1} \sum_{1 \leq i < j \leq N} w_N(x_i - x_j). \quad (\text{A.1})$$

The following can be proved by arguing exactly as we did for $s = 2$.

²There is no blow-up then, but one might want to adapt the method to obtain convergence of density matrices to the stable GP minimizers.

Theorem A.1 (Collapse and condensation of the many-body ground state for anharmonic potentials).

Let $s > 0$, $0 < \beta < 1/2$ and $a_N = a_* - N^{-\alpha}$ with

$$0 < \alpha < \min \left\{ \frac{\beta(s+2)}{s+3}, \frac{(1-2\beta)(s+2)}{s} \right\}.$$

Let Ψ_N be the unique ground state of \tilde{H}_N . Then we have

$$\lim_{N \rightarrow \infty} \text{Tr} \left| \gamma_{\Psi_N}^{(k)} - |\tilde{Q}_N^{\otimes k}\rangle \langle \tilde{Q}_N^{\otimes k}| \right| = 0, \quad (\text{A.2})$$

for all $k \in \mathbb{N}$, where \tilde{Q}_N is the rescaled Gagliardo-Nirenberg optimizer given by

$$\tilde{Q}_N(x) = (a_*)^{-1/2} \tilde{\lambda} (a_* - a_N)^{-\frac{1}{2+s}} Q \left(\tilde{\lambda} (a_* - a_N)^{-\frac{1}{2+s}} x \right),$$

with

$$\tilde{\lambda} = \left(\frac{s}{2} c_0 \int_{\mathbb{R}^2} |x|^s |Q(x)|^2 dx \right)^{\frac{1}{2+s}}.$$

In addition, we have

$$\frac{\min \sigma(\tilde{H}_N)}{N} = (a_* - a_N)^{\frac{s}{s+2}} \left(\frac{\tilde{\lambda}^2 s + 2}{a_* s} + o(1) \right). \quad (\text{A.3})$$

The right side of (A.3) is of course the expansion of the Gross-Pitaevskii energy, which has already been derived in [29].

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