# Global existence of a diffusion limit with damping for the compressible radiative Euler system coupled to an electromagnetic field 

X. Blanc, B. Ducomet, S. Nečasová

## - To cite this version:

X. Blanc, B. Ducomet, S. Nečasová. Global existence of a diffusion limit with damping for the compressible radiative Euler system coupled to an electromagnetic field. Topological Methods in Nonlinear Analysis, In press. hal-01695611v1

## HAL Id: hal-01695611

## https://hal.science/hal-01695611v1

Submitted on 29 Jan 2018 (v1), last revised 8 Feb 2018 (v2)

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire HAL, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

# Global existence of a diffusion limit with damping for the compressible radiative Euler system coupled to an electromagnetic field 

X. Blanc, B. Ducomet, S. Nečasová


#### Abstract

We study the Cauchy problem for a system of equations corresponding to a singular limit of radiative hydrodynamics, namely the 3D radiative compressible Euler system coupled to an electromagnetic field through the MHD approximation. Assuming the presence of damping together with suitable smallness hypotheses for the data, we prove that this problem admits a unique global smooth solution.


Keywords: compressible, Euler, magnetohydrodynamics, radiation hydrodynamics.
AMS subject classification: 35Q30, 76N10

## 1 Introduction

In [4], following the study of Buet and Després [5] we considered a singular limit for a compressible inviscid radiative flow. The motion of the fluid was governed by the Euler system with damping for the evolution of the density $\varrho=\varrho(t, x)$, the velocity field $\vec{u}=\vec{u}(t, x)$, and the absolute temperature $\vartheta=\vartheta(t, x)$ as functions of the time $t$ and the Eulerian spatial coordinate $x \in \mathbb{R}^{3}$. A damping term was added to the momentum equation. We proved that the associated Cauchy problem admits a unique global smooth solution, provided that the data are small enough.

In the present work we study the coupling of the previous model with an electromagnetic field through the so called magnetohydrodynamic approximation (MHD) [3].

Let us recall briefly that Maxwell's electromagnetic theory relies on Ampère-Maxwell equation and Faraday's law. The first one reads

$$
\begin{equation*}
\partial_{t} \vec{D}+\vec{J}=\operatorname{curl}_{x} \vec{H} \tag{1.1}
\end{equation*}
$$

where $\vec{D}=\epsilon \vec{E}$ is the electric induction and $\vec{H}$ is the magnetic field. The second one reads

$$
\begin{equation*}
\partial_{t} \vec{B}+\operatorname{curl}_{x} \vec{E}=0 \tag{1.2}
\end{equation*}
$$

where $\vec{B}=\mu \vec{H}$ is the magnetic induction. Here, the constant $\mu>0$ stands for the permeability of free space.
The two last laws are Coulomb's law

$$
\begin{equation*}
\operatorname{div}_{x} \vec{D}=q \tag{1.3}
\end{equation*}
$$

where $q$ is the electric charge density, and Gauss's law

$$
\begin{equation*}
\operatorname{div}_{x} \vec{B}=0 \tag{1.4}
\end{equation*}
$$

We assume that the electric current density $\vec{J}$ is related to the electric field $\vec{E}$ and the macroscopic fluid velocity $\vec{u}$ via Ohm's law

$$
\begin{equation*}
\vec{J}=\sigma(\vec{E}+\vec{u} \times \vec{B}) \tag{1.5}
\end{equation*}
$$

where $\sigma$ is the electrical conductivity of the fluid.
The magnetic force acting on the fluid (Lorentz's force) $\vec{f}_{m}$ and the magnetic energy supply $E_{m}$ are given by

$$
\begin{equation*}
\overrightarrow{f_{m}}=\vec{J} \times \vec{B}, \quad E_{m}:=\vec{J} \cdot \vec{E} \tag{1.6}
\end{equation*}
$$

The MHD approximation consists in neglecting the displacement current $\partial_{t} \vec{D}$ (for the electric induction given by $\vec{D}=\epsilon \vec{E}$ ) in Ampère-Maxwell equation (1.1) and supposing that the charge $q$ is negligible, so we obtain

$$
\begin{equation*}
\mu \vec{J}=\operatorname{curl}_{x} \vec{B}, \mu>0 \tag{1.7}
\end{equation*}
$$

where, as mentioned above, the constant $\mu$ is the permeability of free space.
Moreover, using (1.5), equation (1.2) can be written [6] in the form

$$
\begin{equation*}
\partial_{t} \vec{B}+\operatorname{curl}_{x}(\vec{B} \times \vec{u})+\operatorname{curl}_{x}\left(\lambda \operatorname{curl}_{x} \vec{B}\right)=0 \tag{1.8}
\end{equation*}
$$

where $\lambda=(\mu \sigma)^{-1}$ is the magnetic diffusivity of the fluid.
Finally, from Faraday's law we get

$$
\begin{equation*}
\partial_{t}\left(\frac{1}{2 \mu}|\vec{B}|^{2}\right)+\vec{J} \cdot \vec{E}=\operatorname{div}_{x}\left(\frac{1}{\mu} \vec{B} \times \vec{E}\right) . \tag{1.9}
\end{equation*}
$$

Concerning radiation, we consider the non equilibrium diffusion regime where radiation appears through an extra equation of parabolic type for the radiative temperature which is a priori different from the fluid temperature.

More specifically the system of equations to be studied for the five unknowns $\left(\varrho, \vec{u}, \vartheta, E_{r}, \vec{B}\right)$ reads

$$
\begin{gather*}
\partial_{t} \varrho+\operatorname{div}_{x}(\varrho \vec{u})=0  \tag{1.10}\\
\partial_{t}(\varrho \vec{u})+\operatorname{div}_{x}(\varrho \vec{u} \otimes \vec{u})+\nabla_{x}\left(p+p_{r}\right)+\frac{1}{\mu} \vec{B} \times \operatorname{curl}_{x} \vec{B}+\nu \vec{u}=0,  \tag{1.11}\\
\partial_{t}(\varrho E)+\operatorname{div}_{x}((\varrho E+p) \vec{u})+\vec{u} \cdot \nabla_{x} p_{r}+\frac{1}{\mu}(\vec{u} \times \vec{B}) \cdot \operatorname{curl}_{x} \vec{B}=\operatorname{div}_{x}\left(\kappa \nabla_{x} \vartheta\right)-\sigma_{a}\left(a \vartheta^{4}-E_{r}\right)+\frac{\lambda}{\mu}\left|\operatorname{curl}_{x} \vec{B}\right|^{2},  \tag{1.12}\\
\partial_{t} E_{r}+\operatorname{div}_{x}\left(E_{r} \vec{u}\right)+p_{r} \operatorname{div}_{x} \vec{u}=\operatorname{div}_{x}\left(\frac{1}{3 \sigma_{s}} \nabla_{x} E_{r}\right)-\sigma_{a}\left(E_{r}-a \vartheta^{4}\right),  \tag{1.13}\\
\partial_{t} \vec{B}+\operatorname{curl}_{x}(\vec{B} \times \vec{u})+\operatorname{curl}_{x}\left(\lambda \operatorname{curl}_{x} \vec{B}\right)=0 \tag{1.14}
\end{gather*}
$$

where $\vec{B}$ is a divergence-free vector field, $E=\frac{1}{2}|\vec{u}|^{2}+e(\varrho, \vartheta)$, $E_{r}$ is the radiative energy related to the radiation temperature $T_{r}$ by $E_{r}=a T_{r}^{4}$ and $p_{r}$ is the radiative pressure given by $p_{r}=\frac{1}{3} a T_{r}^{4}=\frac{1}{3} E_{r}$, with $a>0$. We have also supposed for simplicity that $\mu, \sigma_{a}, \sigma_{s}, \sigma$ and $a$ are positive constants. This implies in particular that $\operatorname{curl}_{x}\left(\frac{1}{\sigma} \operatorname{curl}_{x}\left(\frac{1}{\mu} \vec{B}\right)\right)=-\frac{1}{\sigma \mu} \Delta \vec{B}$.

Extending the analysis of [4] and using stability arguments introduced by Beauchard and Zuazua in [1], our goal is to prove global existence of solutions for the system (1.10) - (1.14) when data are sufficiently close to an equilibrium state.

The plan of the paper is as follows: in Section 2 we state our main result (Theorem 2.1) then in Section 3 we study the MHD model and prove Theorem 2.1.

## 2 Main result

Hypotheses imposed on constitutive relations are motivated by the general existence theory for the Euler-Fourier system developed in $[20,21]$. Hypotheses on transport coefficients are physically relevant for the radiative part $[17,19]$. We impose that the pressure $p(\varrho, \vartheta)>0$, the internal energy $e(\varrho, \vartheta)>0$ and the specific entropy $s(\varrho, \vartheta)$ are smooth functions of their arguments. Moreover, we impose the following monotony assumptions:

$$
\begin{equation*}
\forall \vartheta>0, \quad \forall \varrho>0, \quad \frac{\partial p}{\partial \varrho}(\varrho, \vartheta)>0, \quad \frac{\partial e}{\partial \vartheta}(\varrho, \vartheta)>0 . \tag{2.1}
\end{equation*}
$$

In our simplified setting, transport coefficients $\kappa, \sigma_{a}, \sigma_{s}$ and the Planck's coefficient $a$ are supposed to be fixed positive numbers. Finally the damping term with coefficient $\nu>0$ of Darcy type can be interpreted here as a diffusion of a light gas into a heavy one.

We are going to prove that, under the above structural assumptions on the equation of state, system (1.10)-(1.14) has a global smooth solution close to any equilibrium state.

Theorem 2.1. Let $\left(\bar{\varrho}, 0, \bar{\vartheta}, \overline{E_{r}}, \bar{B}\right)$ be a constant state with $\bar{\varrho}>0, \bar{\vartheta}>0$ and $\overline{E_{r}}=a \bar{\vartheta}^{4}>0$. Consider $d>7 / 2$. There exists $\varepsilon>0$ such that, for any initial state $\left(\varrho_{0}, \vec{u}_{0}, \vartheta_{0}, E_{r}^{0}, \vec{B}_{0}\right)$ satisfying

$$
\begin{equation*}
\left\|\left(\varrho_{0}, \vec{u}_{0}, \vartheta_{0}, E_{r}^{0}, \vec{B}_{0}\right)-\left(\bar{\varrho}, 0, \bar{\vartheta}, \overline{E_{r}}, \overline{\vec{B}}\right)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)} \leq \varepsilon \tag{2.2}
\end{equation*}
$$

there exists a unique global solution $\left(\varrho, \vec{u}, \vartheta, E_{r}, \vec{B}\right)$ to (1.10)-(1.11)-(1.12)-(1.13)-(1.14), such that

$$
\left(\varrho-\bar{\varrho}, \vec{u}, \vartheta-\bar{\vartheta}, E_{r}-\overline{E_{r}}, \vec{B}-\overline{\vec{B}}\right) \in C\left([0,+\infty) ; H^{d}\left(\mathbb{R}^{3}\right)\right) \cap C^{1}\left([0,+\infty) ; H^{d-1}\left(\mathbb{R}^{3}\right)\right) .
$$

In addition, this solution satisfies the following energy inequality:

$$
\begin{align*}
& \left\|\left(\varrho(t)-\bar{\varrho}, \vec{u}(t), \vartheta(t)-\bar{\vartheta}, E_{r}(t)-\overline{E_{r}}, \vec{B}(t)-\overline{\vec{B}}\right)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}+\int_{0}^{t}\left\|\nabla_{x}\left(\varrho, \vec{u}, \vartheta, E_{r}, \vec{B}\right)(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2} d s \\
& +\int_{0}^{t}\left(\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
& \leq C\left\|\left(\varrho_{0}-\bar{\varrho}, \vec{u}_{0}, \vartheta_{0}-\bar{\vartheta}, E_{r}^{0}-\overline{E_{r}}, \vec{B}_{0}-\overline{\vec{B}}\right)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2} \tag{2.3}
\end{align*}
$$

for some constant $C>0$ which does not depend on $t$.

## 3 The Euler-MHD system

### 3.1 The linearized Euler-MHD system

Multiplying (1.11) by $\vec{u}$ and using (1.10) we get

$$
\partial_{t}\left(\frac{1}{2} \varrho|\vec{u}|^{2}\right)+\operatorname{div}_{x}\left(\frac{1}{2} \varrho|\vec{u}|^{2} \vec{u}\right)+\nabla_{x}\left(p+p_{r}\right) \cdot \vec{u}+\nu|\vec{u}|^{2}=\vec{f}_{m} \cdot \vec{u} .
$$

Subtracting this relation to (1.12), using the definition $C_{v}=\partial_{\vartheta} e$ and the thermodynamical identity $\partial_{\varrho} e=\frac{1}{\varrho^{2}}\left(p-\vartheta \partial_{\vartheta} p\right)$ (Maxwell's relation), equation (1.12) can be replaced by an equation for temperature

$$
\begin{equation*}
\varrho C_{v}\left(\partial_{t} \vartheta+\vec{u} \cdot \nabla_{x} \vartheta\right)+\vartheta p_{\vartheta} \operatorname{div}_{x} \vec{u}-\nu \vec{u}^{2}=\operatorname{div}_{x}\left(\kappa \nabla_{x} \vartheta\right)-\sigma_{a}\left(a \vartheta^{4}-E_{r}\right)+E_{m}-\vec{f}_{m} \cdot \vec{u} . \tag{3.1}
\end{equation*}
$$

Linearizing the system $(1.10)(1.11)(3.1)(1.13)(1.14)$ around the constant state $\left(\bar{\varrho}, 0, \bar{\vartheta}, \bar{E}_{r}, \overline{\vec{B}}\right)$ with the compatibility condition $\bar{E}_{r}=a \bar{\vartheta}^{4}$ and setting $\varrho=r+\bar{\varrho}, \vartheta=T+\bar{\vartheta}, E_{r}=e_{r}+\bar{E}_{r}$ and $\vec{B}=\vec{b}+\overline{\vec{B}}$ we get

$$
\begin{gather*}
\partial_{t} r+\bar{\varrho} \operatorname{div}_{x} \vec{u}=0  \tag{3.2}\\
\partial_{t} \vec{u}+\frac{\bar{p}_{\varrho}}{\bar{\varrho}} \nabla_{x} r+\frac{\bar{p}_{\vartheta}}{\bar{\varrho}} \nabla_{x} T+\frac{1}{3 \bar{\varrho}} \nabla_{x} e_{r}+\frac{1}{\mu} \overline{\vec{B}} \times\left(\operatorname{curl}_{x} \vec{b}\right)+\nu \vec{u}=0  \tag{3.3}\\
\partial_{t} T+\frac{\bar{\vartheta} \bar{p}_{\vartheta}}{\bar{\varrho} \bar{C}_{v}} \operatorname{div}_{x} \vec{u}=\operatorname{div}_{x}\left(\frac{\kappa}{\bar{\varrho} \bar{C}_{v}} \nabla_{x} T\right)-\frac{\sigma_{a}}{\bar{\varrho} \bar{C}_{v}}\left(4 a \bar{\vartheta}^{3} T-e_{r}\right), \tag{3.4}
\end{gather*}
$$

$$
\begin{gather*}
\partial_{t} e_{r}+\frac{4}{3} \bar{E}_{r} \operatorname{div}_{x} \vec{u}=\operatorname{div}_{x}\left(\frac{1}{3 \sigma_{s}} \nabla_{x} e_{r}\right)-\sigma_{a}\left(e_{r}-4 a \bar{\vartheta}^{3} T\right),  \tag{3.5}\\
\partial_{t} \vec{b}+\overline{\vec{B}} \operatorname{div}_{x} \vec{u}-\left(\overline{\vec{B}} \cdot \nabla_{x}\right) \vec{u}=\lambda \Delta \vec{b} \tag{3.6}
\end{gather*}
$$

Using the vector notation $U:=\left(\begin{array}{c}r \\ u_{1} \\ u_{2} \\ u_{3} \\ T \\ e_{r} \\ b_{1} \\ b_{2} \\ b_{3}\end{array}\right)$, the linearized system (3.2) - (3.6) rewrites

$$
\begin{equation*}
\partial_{t} U+\sum_{j=1}^{3} \mathcal{A}_{j} \partial_{j} U=\mathcal{D} \Delta U-\mathcal{B} U \tag{3.7}
\end{equation*}
$$

with

$$
\begin{aligned}
& \mathcal{A}_{1}=\left(\begin{array}{ccccccccc}
0 & \bar{\varrho} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
\alpha^{\prime} & 0 & 0 & 0 & \beta^{\prime} & \beta^{\prime \prime} & 0 & \bar{B}_{2} / \mu & \bar{B}_{3} / \mu \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{1} / \mu & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{1} / \mu \\
0 & \gamma^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \gamma^{\prime \prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{2} & -\bar{B}_{1} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{3} & 0 & -\bar{B}_{1} & 0 & 0 & 0 & 0 & 0
\end{array}\right), \\
& \mathcal{A}_{2}=\left(\begin{array}{ccccccccc}
0 & 0 & \bar{\varrho} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{2} / \mu & 0 & 0 \\
\alpha^{\prime} & 0 & 0 & 0 & \beta^{\prime} & \beta^{\prime \prime} & \bar{B}_{1} / \mu & 0 & \bar{B}_{3} / \mu \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{2} / \mu \\
0 & 0 & \gamma^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \gamma^{\prime \prime} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & -\bar{B}_{2} & \bar{B}_{1} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \bar{B}_{3} & -\bar{B}_{2} & 0 & 0 & 0 & 0 & 0
\end{array}\right) \\
& \mathcal{A}_{3}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & \bar{\varrho} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{3} / \mu & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{3} / \mu & 0 \\
\alpha^{\prime} & 0 & 0 & 0 & \beta^{\prime} & \beta^{\prime \prime} & \bar{B}_{1} / \mu & \bar{B}_{2} / \mu & 0 \\
0 & 0 & 0 & \gamma^{\prime} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \gamma^{\prime \prime} & 0 & 0 & 0 & 0 & 0 \\
0 & -\bar{B}_{3} & 0 & \bar{B}_{1} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & -\bar{B}_{3} & \bar{B}_{2} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0
\end{array}\right),
\end{aligned}
$$

and

$$
\mathcal{D}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \delta^{\prime} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \delta^{\prime \prime} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda
\end{array}\right) \quad \mathcal{B}:=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \nu & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \nu & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \nu & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \zeta & -\eta & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & -\pi & \sigma_{a} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0
\end{array}\right),
$$

where

$$
\begin{gathered}
\alpha^{\prime}=\frac{\bar{p}_{\varrho}}{\bar{\varrho}}, \quad \beta^{\prime}=\frac{\bar{p}_{\vartheta}}{\bar{\varrho}}, \quad \beta^{\prime \prime}=\frac{1}{3 \bar{\varrho}}, \quad \gamma^{\prime}=\frac{\bar{p}_{\vartheta}}{\bar{C}_{v}}, \quad \delta^{\prime}=\frac{\kappa}{\bar{\varrho} \bar{C}_{v}} \\
\gamma^{\prime \prime}=\frac{4}{3} \bar{E}_{r}, \quad \delta^{\prime \prime}=\frac{1}{3 \sigma_{s}}, \quad \zeta=\frac{4 a \sigma_{a} \bar{\vartheta}^{3}}{\bar{\varrho}_{v}}, \quad \eta=\frac{\sigma_{a}}{\bar{\varrho} \bar{C}_{v}}, \quad \pi=4 a \sigma_{a} \bar{\vartheta}^{3} .
\end{gathered}
$$

In order to apply the Kreiss theory we have to put the system (3.7) in a symmetric form [2]. For that purpose it is sufficient to consider a diagonal symmetrizer

$$
\widetilde{\mathcal{A}}_{0}=\left(\begin{array}{ccccccccc}
\mu \frac{\alpha^{\prime}}{\bar{\varrho}} & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0  \tag{3.8}\\
0 & \mu & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \mu & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \mu & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \mu \frac{\beta^{\prime}}{\gamma^{\prime}} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \mu \frac{\beta^{\prime \prime}}{\gamma^{\prime \prime}} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1
\end{array}\right) .
$$

Multiplying the first equation (3.7) by $\widetilde{\mathcal{A}}_{0}$ on the left, we get

$$
\begin{equation*}
\widetilde{\mathcal{A}}_{0} \partial_{t} U+\sum_{j=1}^{3} \widetilde{\mathcal{A}}_{j} \partial_{j} U=\widetilde{\mathcal{D}} \Delta U-\widetilde{\mathcal{B}} U \tag{3.9}
\end{equation*}
$$

where the matrices $\widetilde{\mathcal{A}}_{j}=\widetilde{\mathcal{A}}_{0} \mathcal{A}_{j}$ are symmetric, for all $j=1,2,3$. More specifically,

$$
\widetilde{\mathcal{A}}_{1}=\left(\begin{array}{ccccccccc}
0 & \mu \alpha^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
\mu \alpha^{\prime} & 0 & 0 & 0 & \mu \beta^{\prime} & \mu \beta^{\prime \prime} & 0 & \bar{B}_{2} & \bar{B}_{3} \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{1} & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{1} \\
0 & \mu \beta^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \mu \beta^{\prime \prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{2} & -\bar{B}_{1} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{3} & 0 & -\bar{B}_{1} & 0 & 0 & 0 & 0 & 0
\end{array}\right)
$$

$$
\begin{aligned}
& \widetilde{\mathcal{A}}_{2}=\left(\begin{array}{ccccccccc}
0 & 0 & \mu \alpha^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{2} & 0 & 0 \\
\mu \alpha^{\prime} & 0 & 0 & 0 & \mu \beta^{\prime} & \mu \beta^{\prime \prime} & \bar{B}_{1} & 0 & \bar{B}_{3} \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{2} \\
0 & 0 & \mu \beta^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \mu \beta^{\prime \prime} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & -\bar{B}_{2} & \bar{B}_{1} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \bar{B}_{3} & -\bar{B}_{2} & 0 & 0 & 0 & 0 & 0
\end{array}\right), \\
& \widetilde{\mathcal{A}}_{3}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & \mu \alpha^{\prime} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{3} & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -\bar{B}_{3} & 0 \\
\mu \alpha^{\prime} & 0 & 0 & 0 & \mu \beta^{\prime} & \mu \beta^{\prime \prime} & \bar{B}_{1} & \bar{B}_{2} & 0 \\
0 & 0 & 0 & \mu \beta^{\prime} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \mu \beta^{\prime \prime} & 0 & 0 & 0 & 0 & 0 \\
0 & -\bar{B}_{3} & 0 & \bar{B}_{1} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & -\bar{B}_{3} & \bar{B}_{2} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0
\end{array}\right) .
\end{aligned}
$$

The hyperbolic part of system (3.9) is now symmetric while its symmetric dissipative part is given by

$$
\widetilde{\mathcal{D}}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0  \tag{3.10}\\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \frac{\mu \beta^{\prime} \delta^{\prime}}{\gamma^{\prime}} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \frac{\mu \beta^{\prime \prime} \delta^{\prime \prime}}{\gamma^{\prime \prime}} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda
\end{array}\right) \widetilde{\mathcal{B}}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \mu \nu & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \mu \nu & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \mu \nu & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \frac{\mu \beta^{\prime} \zeta}{\gamma^{\prime}} & -\frac{\mu \beta^{\prime} \eta}{\gamma^{\prime} \eta} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & -\frac{\mu \beta^{\prime \prime} \pi}{\gamma^{\prime \prime}} & \frac{\mu \sigma_{a} \beta^{\prime \prime}}{\gamma^{\prime \prime}} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0
\end{array}\right) .
$$

A simple computation proves that $\widetilde{\mathcal{B}}$ is a semi-definite positive matrix, that is,

$$
{ }^{t} X \widetilde{\mathcal{B}} X \geq 0, \quad \text { for any vector } X \in \mathbb{R}^{9} .
$$

Applying the Fourier transform in $x$ to (3.9) we get

$$
\begin{equation*}
\widetilde{\mathcal{A}}_{0} \partial_{t} \widehat{U}+i \sum_{j=1}^{3} \xi_{j} \widetilde{\mathcal{A}}_{j} \widehat{U}=-|\xi|^{2} \widetilde{\mathcal{D}} \widehat{U}-\widetilde{\mathcal{B}} \widehat{U} \tag{3.11}
\end{equation*}
$$

or

$$
\begin{equation*}
\widetilde{\mathcal{A}}_{0} \partial_{t} \widehat{U}=E(\xi) \widehat{U} \tag{3.12}
\end{equation*}
$$

with

$$
E(\xi)=-\mathcal{B}(\xi)-i \mathcal{A}(\xi)
$$

where

$$
\mathcal{A}(\xi)=\sum_{j=1}^{3} \xi_{j} \tilde{\mathcal{A}}_{j}=
$$

$$
\left(\begin{array}{ccccccccc}
0 & \mu \alpha^{\prime} \xi_{1} & \mu \alpha^{\prime} \xi_{2} & \mu \alpha^{\prime} \xi_{3} & 0 & 0 & 0 & 0 & 0 \\
\mu \alpha^{\prime} \xi_{1} & 0 & 0 & 0 & \mu \beta^{\prime} \xi_{1} & \mu \beta^{\prime \prime} \xi_{1} & -\bar{B}_{2} \xi_{2}-\bar{B}_{3} \xi_{3} & \bar{B}_{2} \xi_{1} & \bar{B}_{3} \xi_{1} \\
\mu \alpha^{\prime} \xi_{2} & 0 & 0 & 0 & \mu \beta^{\prime} \xi_{2} & \mu \beta^{\prime \prime} \xi_{2} & \bar{B}_{1} \xi_{2} & -\bar{B}_{1} \xi_{1}-\bar{B}_{3} \xi_{3} & \bar{B}_{3} \xi_{2} \\
\mu \alpha^{\prime} \xi_{3} & 0 & 0 & 0 & \mu \beta^{\prime} \xi_{3} & \mu \beta^{\prime \prime} \xi_{3} & \bar{B}_{1} \xi_{3} & \bar{B}_{2} \xi_{3} & -\bar{B}_{1} \xi_{1}-\bar{B} \\
0 & \mu \beta^{\prime} \xi_{1} & \mu \beta^{\prime} \xi_{2} & \mu \beta^{\prime} \xi_{3} & 0 & 0 & 0 & 0 & 0 \\
0 & \mu \beta^{\prime \prime} \xi_{1} & \mu \bar{B}^{\prime \prime} \xi_{2} & \bar{\beta}^{\prime \prime} \xi_{3} & 0 & 0 & 0 & 0 & 0 \\
0 & -\bar{B}_{2} \xi_{2}-\bar{B}_{3} \xi_{3} & \bar{B}_{1} \xi_{2} & \bar{B}_{1} \xi_{3} & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{2} \xi_{1} & -\bar{B}_{1} \xi_{1}-\bar{B}_{3} \xi_{3} & \bar{B}_{2} \xi_{3} & 0 & 0 & 0 & 0 & 0 \\
0 & \bar{B}_{3} \xi_{1} & -\bar{B}_{1} \xi_{1}-\bar{B}_{2} \xi_{2} & 0 & 0 & 0 & 0 & 0 \\
0
\end{array}\right.
$$

and

$$
\mathcal{B}(\xi):=\widetilde{\mathcal{B}}+|\xi|^{2} \widetilde{\mathcal{D}}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0  \tag{3.14}\\
0 & \nu & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \nu & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \nu & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \frac{\mu \beta^{\prime} \zeta \zeta}{\gamma^{\prime}}+\frac{\mu \beta^{\prime} \delta^{\prime}}{\gamma^{\prime}}|\xi|^{2} & -\frac{\mu \beta^{\prime} \eta}{\gamma^{\prime} \eta} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & -\frac{\mu \beta^{\prime \prime} \pi}{\gamma^{\prime \prime \prime}} & \frac{\mu \beta^{\prime \prime} \sigma_{a}}{\gamma^{\prime \prime}}+\frac{\mu \beta^{\prime \prime} \delta^{\prime \prime}}{0}|\xi|^{2} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & \lambda|\xi|^{2} & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda|\xi|^{2} & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda|\xi|^{2}
\end{array}\right) .
$$

Solving this equation with initial condition $\widehat{U}_{0}(\xi)$ we get

$$
\begin{equation*}
\widehat{U}(t, \xi)=\exp \left[t \widetilde{\mathcal{A}}_{0}^{-1} E(\xi)\right] \widehat{U}_{0}(\xi) . \tag{3.15}
\end{equation*}
$$

In the strictly hyperbolic case $\widetilde{\mathcal{D}}=0$, under the Kalman rank condition [12] for the pair $(\mathcal{A}(\xi), \mathcal{B})$, it can be proved [1] that

$$
\exists C>0, \quad \lambda(\xi)>0 \quad: \quad \exp \left[t \widetilde{\mathcal{A}}_{0}^{-1} E(\xi)\right] \leq C e^{-\lambda(\xi) t} .
$$

Due to the partially parabolic nature of the system, one can expect a similar result when $\widetilde{\mathcal{D}} \neq 0$ with a parabolic smoothing effect at low frequencies and an extra damping in the high frequency regime.

Taking benefit of the damping, we can use the Shizuta-Kawashima condition (SK) [22] which applies to the previous system. Following the arguments of Beauchard and Zuazua [1], we have

Lemma 1. For any $\xi \in S^{2}$, a necessary and sufficient condition for the matrices $\mathcal{B}(\xi)$ and $\mathcal{A}(\xi)$ defined by (3.13) and (3.14) satisfy the Shizuta-Kawashima condition (SK):

$$
\begin{equation*}
\left\{\text { eigenvectors of }\left(\widetilde{\mathcal{A}}_{0}\right)^{-1} \mathcal{A}(\xi)\right\} \cap \text { ker } \mathcal{B}(\xi)=\{0\} \text {, } \tag{3.16}
\end{equation*}
$$

is that $\nu>0$.

## Proof:

1. We first consider the case $\nu \neq 0$. One checks that $\operatorname{ker} \mathcal{B}(\xi)$ is the 1 -dimensional subspace spanned by the vector $(1,0,0,0,0,0,0,0,0)$. Therefore, if $X \in \operatorname{ker} \mathcal{B}(\xi) \backslash\{0\}$ is an eigenvector of $\left(\widetilde{\mathcal{A}}_{0}\right)^{-1} \mathcal{A}(\xi)$, we have $X=$ $\left(x_{1}, 0,0,0,0,0,0,0,0\right), x_{1} \neq 0$, and

$$
\mathcal{A}(\xi) X=\lambda \widetilde{\mathcal{A}}_{0} X,
$$

for some $\lambda \in \mathbb{R}$. This, together with the definition of $\widetilde{\mathcal{A}}_{0}$ and $\mathcal{A}(\xi)$, implies that $\lambda=0, \xi_{1}=\xi_{2}=\xi_{3}=0$, which is in contradiction with the hypothesis $\xi \in S^{2}$.
2. Next, we assume that $\nu=0$. One checks that $\operatorname{ker} \mathcal{B}(\xi)$ is the 4 -dimensional subspace spanned by the vectors $\left(x_{1}, x_{2}, x_{3}, x_{4}, 0,0,0,0,0\right)$.
Let us denote by $(\lambda, X)$ an eigenpair of $\mathcal{A}(\xi)$, with non zero eigenvector $X \in \operatorname{ker} \mathcal{B}(\xi)$. $X$ satisfies the system

$$
\begin{gathered}
\mu \alpha^{\prime} \xi_{1} x_{2}+\mu \alpha^{\prime} \xi_{2} x_{3}+\mu \alpha^{\prime} \xi_{3} x_{4}=\lambda x_{1} \\
\mu \alpha^{\prime} \xi_{1} x_{1}=\lambda x_{2} \\
\mu \alpha^{\prime} \xi_{2} x_{1}=\lambda x_{3} \\
\mu \alpha^{\prime} \xi_{3} x_{1}=\lambda x_{4} \\
\mu \beta^{\prime} \xi_{1} x_{2}+\mu \beta^{\prime} \xi_{2} x_{3}+\mu \beta^{\prime} \xi_{3} x_{4}=0 \\
\mu \beta^{\prime \prime} \xi_{1} x_{2}+\mu \beta^{\prime \prime} \xi_{2} x_{3}+\mu \beta^{\prime \prime} \xi_{3} x_{4}=0 \\
-\left(\bar{B}_{2} \xi_{2}+\bar{B}_{3} \xi_{3}\right) x_{2}+\bar{B}_{1} \xi_{2} x_{3}+\bar{B}_{1} \xi_{3} x_{4}=0 \\
\bar{B}_{2} \xi_{1} x_{2}-\left(\bar{B}_{1} \xi_{1}+\bar{B}_{3} \xi_{3}\right) x_{3}+\bar{B}_{2} \xi_{3} x_{4}=0 \\
\bar{B}_{3} \xi_{1} x_{2}+\bar{B}_{3} \xi_{2} x_{3}-\left(\bar{B}_{1} \xi_{1}+\bar{B}_{2} \xi_{2}\right) x_{4}=0
\end{gathered}
$$

Denoting $\vec{B}=\left(\bar{B}_{1}, \bar{B}_{2}, \bar{B}_{3}\right), \vec{\xi}=\left(\xi_{1}, \xi_{2}, \xi_{3}\right)$, and $\vec{x}=\left(x_{2}, x_{3}, x_{4}\right)$, the system rewrites

$$
\begin{gathered}
\mu \alpha^{\prime} \vec{x} \cdot \vec{\xi}=\lambda x_{1} \\
\mu \alpha^{\prime} x_{1} \vec{\xi}=\lambda \vec{x} \\
\mu \beta^{\prime} \vec{x} \cdot \vec{\xi}=0 \\
\mu \beta^{\prime \prime} \vec{x} \cdot \vec{\xi}=0 \\
-(\vec{B} \cdot \vec{\xi}) \vec{x}+(\vec{\xi} \cdot \vec{x}) \vec{B}=0
\end{gathered}
$$

In particular, this implies $\vec{x} \cdot \vec{\xi}=0$, which in turn implies $\lambda x_{1}=0$. As a consequence, we have

$$
\begin{gathered}
\lambda x_{1}=0 \\
\vec{x} \cdot \vec{\xi}=0 \\
\mu \alpha^{\prime} x_{1} \vec{\xi}=\lambda \vec{x} \\
(\vec{\xi} \cdot \vec{B}) \vec{x}=0
\end{gathered}
$$

Choosing $\lambda=0$, we see that any $\vec{x} \in \vec{\xi}^{\perp}$ for $\xi \in \vec{B}^{\perp}$ gives a nontrivial eigenpair $(\lambda, X)$ with $\lambda=0$ and $X=(0, \vec{x}, 0,0,0,0,0)$. Hence the SK condition is not satisfied.

As in the equilibrium case, (3.16) is equivalent to the existence of a compensating matrix:
Proposition 3.1. For any $\xi \in S^{2}$, the matrices $\widetilde{\mathcal{A}}_{0}, \mathcal{B}(\xi)$ and $\mathcal{A}(\xi)$ being defined by (3.8), (3.13) and (3.14), there exists a matrix-valued function

$$
\begin{aligned}
K: \quad S^{2} & \longrightarrow \mathbb{R}^{6 \times 6} \\
\omega & \longmapsto K(\omega)
\end{aligned}
$$

such that

1. $\omega \mapsto K(\omega)$ is a $C^{\infty}$ function, and satisfies $K(-\omega)=-K(\omega)$ for any $\omega \in S^{2}$.
2. $K(\omega) \tilde{\mathcal{A}}_{0}$ is a skew-symmetric matrix for any $\omega \in S^{2}$.
3. Denoting by $[A]=\frac{1}{2}\left(A+A^{T}\right)$ the symmetric part of $A$, the matrix $[K(\omega) \mathcal{A}(\omega)]+\mathcal{B}(\omega)$ is symmetric positive definite for any $\omega \in S^{2}$.

### 3.2 Entropy properties

Adding equations (1.9), (1.12) and (1.13) we get

$$
\begin{align*}
\partial_{t}\left(\frac{1}{2} \varrho|\vec{u}|^{2}+\varrho e+E_{r}+\frac{1}{2 \mu}|\vec{B}|^{2}\right)+\operatorname{div}_{x}\left(\left(\varrho E+E_{r}\right) \vec{u}+\left(p+p_{r}\right) \vec{u}\right. & \left.+\frac{1}{\mu} \vec{E} \times \vec{B}\right) \\
& =\operatorname{div}_{x}\left(\kappa \nabla_{x} \vartheta\right)+\operatorname{div}_{x}\left(\frac{1}{3 \sigma_{s}} \nabla_{x} E_{r}\right) . \tag{3.17}
\end{align*}
$$

We define $S_{r}:=\frac{4}{3} a T_{r}^{3}$ the radiative entropy. With these definitions, equation (1.13) rewrites

$$
\partial_{t} S_{r}+\operatorname{div}_{x}\left(S_{r} \vec{u}\right)=\frac{1}{T_{r}} \operatorname{div}_{x}\left(\frac{1}{3 \sigma_{s}} \nabla_{x} E_{r}\right)-\sigma_{a} \frac{E_{r}-a \vartheta^{4}}{T_{r}}
$$

that is,

$$
\begin{equation*}
\partial_{t} S_{r}+\operatorname{div}_{x}\left(S_{r} \vec{u}\right)=\operatorname{div}_{x}\left(\frac{1}{3 \sigma_{s} T_{r}} \nabla_{x} E_{r}\right)+\frac{4 a}{3 \sigma_{s}} T_{r}\left|\nabla_{x} T_{r}\right|^{2}-\sigma_{a} \frac{E_{r}-a \vartheta^{4}}{T_{r}} \tag{3.18}
\end{equation*}
$$

Replacing equation (1.12) by the internal energy equation

$$
\begin{equation*}
\partial_{t}(\varrho e)+\operatorname{div}_{x}(\varrho e \vec{u})+p \operatorname{div}_{x} \vec{u}-\nu|\vec{u}|^{2}=\operatorname{div}_{x}\left(\kappa \nabla_{x} \vartheta\right)-\sigma_{a}\left(a \vartheta^{4}-E_{r}\right)+\frac{1}{\sigma \mu^{2}}\left|\operatorname{curl}_{x} \vec{B}\right|^{2} . \tag{3.19}
\end{equation*}
$$

The entropy $s$ of the fluid is defined by the Gibbs law $\vartheta d s=d e+p d\left(\frac{1}{\varrho}\right)$. Hence, dividing (3.19) by $\vartheta$, we find

$$
\begin{equation*}
\partial_{t}(\varrho s)+\operatorname{div}_{x}(\varrho s \vec{u})-\frac{\nu}{\vartheta}|\vec{u}|^{2}=\operatorname{div}_{x}\left(\frac{\kappa \nabla_{x} \vartheta}{\vartheta}\right)+\frac{\kappa\left|\nabla_{x} \vartheta\right|^{2}}{\vartheta^{2}}-\sigma_{a} \frac{a \vartheta^{4}-E_{r}}{\vartheta}+\frac{1}{\sigma \mu^{2} \vartheta}\left|\operatorname{curl}_{x} \vec{B}\right|^{2} \tag{3.20}
\end{equation*}
$$

So adding (3.20) and (3.18) we obtain

$$
\begin{gather*}
\partial_{t}\left(\varrho s+S_{r}\right)+\operatorname{div}_{x}\left(\left(\varrho s+S_{r}\right) \vec{u}\right)-\operatorname{div}_{x}\left(\frac{\kappa \nabla_{x} \vartheta}{\vartheta}+\frac{1}{3 \sigma_{s} T_{r}} \nabla_{x} E_{r}\right) \\
=\frac{\kappa\left|\nabla_{x} \vartheta\right|^{2}}{\vartheta^{2}}+\frac{4 a}{3 \sigma_{s}} T_{r}\left|\nabla_{x} E_{r}\right|^{2}+\frac{a \sigma_{a}}{\vartheta T_{r}}\left(\vartheta-T_{r}\right)^{2}\left(\vartheta+T_{r}\right)\left(\vartheta^{2}+T_{r}^{2}\right)+\frac{1}{\sigma \mu^{2} \vartheta}\left|\operatorname{curl}_{x} \vec{B}\right|^{2}+\frac{\nu}{\vartheta}|\vec{u}|^{2} \tag{3.21}
\end{gather*}
$$

Introducing the Helmholtz functions $H_{\bar{\vartheta}}(\varrho, \vartheta):=\varrho(e-\bar{\vartheta} s)$ and $H_{r, \bar{\vartheta}}\left(T_{r}\right):=E_{r}-\bar{\vartheta} S_{r}$, we check that the quantities $H_{\bar{\vartheta}}(\varrho, \vartheta)-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})$ and $H_{r, \bar{\vartheta}}\left(T_{r}\right)-H_{r, \bar{\vartheta}}\left(\bar{T}_{r}\right)$ are non-negative and strictly coercive functions reaching zero minima at the equilibrium state $\left(\bar{\varrho}, \bar{\vartheta}, \bar{E}_{r}\right)$.
Lemma 2. Let $\bar{\varrho}$ and $\bar{\vartheta}=\bar{T}_{r}$ be given positive constants. Let $\mathcal{O}_{1}$ and $\mathcal{O}_{2}$ be the sets defined by

$$
\begin{align*}
\mathcal{O}_{1}:= & \left\{(\varrho, \vartheta) \in \mathbb{R}^{2}: \frac{\bar{\varrho}}{2}<\varrho<2 \bar{\varrho}, \frac{\bar{\vartheta}}{2}<\vartheta<2 \bar{\vartheta},\right\} .  \tag{3.22}\\
& \mathcal{O}_{2}:=\left\{T_{r} \in \mathbb{R}: \frac{\bar{T}_{r}}{2}<T_{r}<2 \bar{T}_{r},\right\} . \tag{3.23}
\end{align*}
$$

There exist positive constants $C_{1,2}(\bar{\varrho}, \bar{\vartheta})$ and $C_{3,4}\left(\bar{T}_{r}\right)$ such that
1.

$$
\begin{gather*}
C_{1}\left(|\varrho-\bar{\varrho}|^{2}+|\vartheta-\bar{\vartheta}|^{2}\right) \leq H_{\bar{\vartheta}}(\varrho, \vartheta)-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta}) \\
\leq C_{2}\left(|\varrho-\bar{\varrho}|^{2}+|\vartheta-\bar{\vartheta}|^{2}\right), \tag{3.24}
\end{gather*}
$$

for all $(\varrho, \vartheta) \in \mathcal{O}_{1}$,
2.

$$
\begin{equation*}
C_{3}\left|T_{r}-\bar{T}_{r}\right|^{2} \leq H_{r, \bar{\vartheta}}\left(T_{r}\right)-H_{r, \bar{\vartheta}}\left(\bar{T}_{r}\right) \leq C_{4}\left|T_{r}-\bar{T}_{r}\right|^{2} \tag{3.25}
\end{equation*}
$$

for all $T_{r} \in \mathcal{O}_{2}$.

## Proof:

1. Point 1 is proved in [9] and we only sketch the proof for convenience.

We have the decomposition

$$
\varrho \mapsto H_{\bar{\vartheta}}(\varrho, \vartheta)-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})=\mathcal{F}(\varrho)+\mathcal{G}(\varrho),
$$

where $\mathcal{F}(\varrho)=H_{\bar{\vartheta}}(\varrho, \bar{\vartheta})-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})$ and $\mathcal{G}(\varrho)=H_{\bar{\vartheta}}(\varrho, \vartheta)-H_{\bar{\vartheta}}(\varrho, \bar{\vartheta})$. Using Gibbs law $\vartheta d s=d e+p d\left(\frac{1}{\varrho}\right)$, one easily proves that $\partial_{\varrho}^{2} H_{\bar{\vartheta}}(\varrho, \bar{\vartheta})=\frac{1}{\varrho \bar{\vartheta}} \partial_{\varrho} p(\varrho, \bar{\vartheta})$, which is positive according to (2.1). Hence, $\mathcal{F}$ is strictly convex and reaches a zero minimum at $\bar{\varrho}$. Turning to $\mathcal{G}$, we have, still using Gibbs law, $\partial_{\vartheta} H_{\bar{\vartheta}}(\varrho, \bar{\vartheta})=\varrho \frac{\vartheta-\bar{\vartheta}}{\vartheta} \partial_{\vartheta} e(\varrho, \bar{\vartheta})$. Thus, using (2.1) again, we infer that $\mathcal{G}$ is strictly decreasing for $\vartheta<\bar{\vartheta}$ and strictly increasing for $\vartheta>\bar{\vartheta}$. Computing the derivatives of $H_{\bar{\vartheta}}$ leads directly to the estimate (3.24).
2. Point 2 follows from the properties of the function $x \mapsto H_{r, \bar{\vartheta}}(x)-H_{r, \bar{\vartheta}}\left(\bar{T}_{r}\right)=a x^{3}\left(x-\frac{4}{3} \bar{\vartheta}\right)+\frac{a}{3} \bar{\vartheta}^{4}$

From this simple result, we can obtain an identity leading to energy estimates. Multiplying (3.21) by $\bar{\vartheta}$, subtracting the result to (3.17) and using the conservation of mass, we get

$$
\begin{gather*}
\partial_{t}\left(\frac{1}{2} \varrho|\vec{u}|^{2}+H_{\bar{\vartheta}}(\varrho, \vartheta)-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})+H_{r, \bar{\vartheta}}\left(T_{r}\right)-H_{r, \bar{\vartheta}}\left(\bar{T}_{r}\right)+\frac{1}{2 \mu}|\vec{B}|^{2}\right) \\
+\operatorname{div}_{x}\left(\left(\varrho(E-\bar{e})+E_{r}\right) \vec{u}+\left(p+p_{r}\right) \vec{u}-\bar{\vartheta}\left(\varrho(s-\bar{s})+S_{r}\right) \vec{u}+\frac{1}{\mu} \vec{E} \times \vec{B}\right) \\
=\operatorname{div}_{x}\left(\kappa \nabla_{x} \vartheta+\frac{1}{3 \sigma_{s}} \nabla_{x} E_{r}\right)-\bar{\vartheta} \operatorname{div}_{x}\left(\frac{\kappa \nabla_{x} \vartheta}{\vartheta}+\frac{1}{3 \sigma_{s} T_{r}} \nabla_{x} E_{r}\right) \\
-\bar{\vartheta} \frac{\kappa\left|\nabla_{x} \vartheta\right|^{2}}{\vartheta^{2}}-\bar{\vartheta} \frac{4 a}{3 \sigma_{s}} T_{r}\left|\nabla_{x} E_{r}\right|^{2}-\bar{\vartheta} \frac{a \sigma_{a}}{\vartheta T_{r}}\left(\vartheta-T_{r}\right)^{2}\left(\vartheta+T_{r}\right)\left(\vartheta^{2}+T_{r}^{2}\right)-\frac{\nu}{\vartheta}|\vec{u}|^{2}-\frac{1}{\sigma \mu^{2} \vartheta}\left|\operatorname{curl}_{x} \vec{B}\right|^{2} . \tag{3.26}
\end{gather*}
$$

In the sequel, we define $V=\left(\rho, \vec{u}, \vartheta, E_{r}, \vec{B}\right)^{T}, \bar{V}=\left(\bar{\rho}, 0, \bar{\vartheta}, \overline{E_{r}}, \overline{\vec{B}}\right)^{T}$, and

$$
\begin{align*}
N(t)^{2}=\sup _{0 \leq s \leq t} & \|V(s)-\bar{V}\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2} \\
& +\int_{0}^{t}\left(\left\|\nabla_{x} V(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
& +\int_{0}^{t}\left(\left\|\vartheta(s)-T_{r}(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \tag{3.27}
\end{align*}
$$

Recall that $T_{r}=E_{r}^{1 / 4} a^{-1 / 4}$. Note also that, since $\operatorname{div}_{x}(\vec{B})=0$, we have

$$
\begin{equation*}
\int_{\mathbb{R}^{3}}\left|\operatorname{curl}_{x} \vec{B}\right|^{2}=\int_{\mathbb{R}^{3}}\left|\nabla_{x} \vec{B}\right|^{2}, \tag{3.28}
\end{equation*}
$$

as far as $\vec{B} \in H^{1}\left(\mathbb{R}^{3}\right)$, and similarly for any $H^{s}$ norm. This allows, in the sequel, to replace $\operatorname{curl}_{x} \vec{B}$ by $\nabla_{x} \vec{B}$ in all bounds.

### 3.2.1 $\quad L^{\infty}\left(H^{d}\right)$ estimates

Using these entropy properties, we are going to prove the following result:
Proposition 3.2. Let the assumptions of Theorem 2.1 be satisfied. Consider a solution ( $\varrho, \vec{u}, \vartheta, E_{r}$ ) of system (1.10)-(1.11)-(1.12)-(1.13)-(1.14) on $[0, t]$, for some $t>0$. Then, the energy defined by (3.27) satisfies

$$
\begin{align*}
& \|V(t)-\bar{V}\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} \\
& \quad+\int_{0}^{t}\left(\left\|\nabla_{x} \vartheta(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\vartheta(s)-T_{r}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
& \leq C(N(t)) N(0)^{2} \tag{3.29}
\end{align*}
$$

where the function $C$ is non-decreasing.
Proof: Following the proof of [13, Lemma 3.1] we define

$$
\begin{equation*}
\eta(t, x)=H_{\bar{\vartheta}}(\varrho, \vartheta)-(\varrho-\bar{\varrho}) \partial_{\varrho} H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})-H_{\bar{\vartheta}}(\bar{\varrho}, \bar{\vartheta})+H_{r, \bar{\vartheta}}\left(T_{r}\right)-H_{r, \bar{\vartheta}}\left(\bar{T}_{r}\right) . . \tag{3.30}
\end{equation*}
$$

We multiply (3.21) by $\bar{\vartheta}$, and subtract the result to (3.17). Integrating over $[0, t] \times \mathbb{R}^{3}$, we find

$$
\begin{aligned}
\int_{\mathbb{R}^{3}}\left(\frac{1}{2} \varrho(t)|\vec{u}|^{2}(t)+\right. & \left.\eta(t, x)+\frac{1}{2 \mu}|\vec{B}|^{2}\right) d x+\int_{0}^{t} \int_{\mathbb{R}^{3}} \kappa \frac{\bar{\vartheta}}{\vartheta^{2}}\left|\nabla_{x} \vartheta\right|^{2}+\frac{4 a}{3 \sigma_{s}} T_{r}\left|\nabla_{x} E_{r}\right|^{2} \bar{\vartheta} \\
& +\int_{0}^{t} \int_{\mathbb{R}^{3}} \bar{\vartheta} \frac{a \sigma_{a}}{\vartheta T_{r}}\left(\vartheta+T_{r}\right)\left(\vartheta-T_{r}\right)^{2}\left(\vartheta^{2}+T_{r}^{2}\right)+\frac{\bar{\vartheta}}{\vartheta} \nu|\vec{u}|^{2}+\frac{\bar{\vartheta}}{\sigma \mu^{2} \vartheta}\left|\operatorname{curl}_{x} \vec{B}\right|^{2} \\
& \leq \int_{\mathbb{R}^{3}} \eta(0, x) d x+\int_{\mathbb{R}^{3}} \varrho_{0}\left|\vec{u}_{0}\right|^{2}+\frac{1}{2 \mu} \int_{\mathbb{R}^{3}}\left|\vec{B}_{0}\right|^{2}
\end{aligned}
$$

Defining

$$
\begin{equation*}
M(t)=\sup _{0 \leq s \leq t} \sup _{x \in \mathbb{R}^{3}}\left[\max \left(|\varrho(s, x)-\bar{\varrho}|,|\vec{u}(s, x)|,|\vartheta(s, x)-\bar{\vartheta}|,\left|E_{r}(s, x)-\overline{E_{r}}\right|,|\vec{B}-\overline{\vec{B}}|\right)\right], \tag{3.31}
\end{equation*}
$$

and applying Lemma 2, we find that

$$
\begin{aligned}
& \|V(t)-\bar{V}\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} \\
& \quad+\int_{0}^{t}\left(\left\|\nabla_{x} \vartheta(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\vartheta(s)-T_{r}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\operatorname{curl}_{x} \vec{B}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
& \leq C(M(t)) N(0)
\end{aligned}
$$

where $C: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+}$is non-decreasing. Equation (3.28) allows to replace $\operatorname{curl}_{x} \vec{B}$ by $\nabla_{x} \vec{B}$ in the above estimate. Finally, we point out that, since $d>7 / 2>3 / 2$, due to Sobolev embeddings, there exists a universal constante $C_{0}$ such that $M(t) \leq C_{0} N(t)$. Since $C$ is non-decreasing, this proves (3.29).
Proposition 3.3. Setting $V=\left(\varrho, \vec{u}, \vartheta, E_{r}, \vec{B}\right)^{T}$, under the same assumptions as in Theorem 2.1, we have the following estimate:

$$
\begin{align*}
&\left\|\partial_{t} V(t)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C(N(t))\left(\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\nabla_{x} \vartheta\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}\right. \\
&\left.+\left\|\nabla_{x} E_{r}\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}+\left\|\vartheta-T_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\|\vec{u}\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\nabla_{x} \vec{B}\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}\right) \tag{3.32}
\end{align*}
$$

Proof: The system satisfied by $V$ may be written formally

$$
\begin{equation*}
\partial_{t} V+\sum_{j=1}^{3} \widehat{\mathcal{A}}_{j}(V) \partial_{x_{j}} V=\widehat{\mathcal{D}}(V) \Delta V-\widehat{\mathcal{B}}(V)=0 \tag{3.33}
\end{equation*}
$$

where

$$
\begin{aligned}
& \widehat{\mathcal{A}}_{1}=\left(\begin{array}{ccccccccc}
0 & \varrho & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
\widehat{\alpha}^{\prime} & 0 & 0 & 0 & \widehat{\beta}^{\prime} & \widehat{\beta}^{\prime \prime} & 0 & B_{2} / \varrho \mu & B_{3} / \varrho \mu \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -B_{1} / \varrho \mu & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & -B_{1} / \varrho \mu \\
0 & \widehat{\gamma}^{\prime} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & \widehat{\gamma}^{\prime \prime} & 0 & 0 & 0 & u_{1} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & u_{1} & 0 & 0 \\
0 & B_{2} & -B_{1} & 0 & 0 & 0 & 0 & u_{1} & 0 \\
0 & B_{3} & 0 & -B_{1} & 0 & 0 & 0 & 0 & u_{1}
\end{array}\right), \\
& \widehat{\mathcal{A}}_{2}
\end{aligned}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \varrho & 0 & 0 & 0 & 0 & 0 \\
\widehat{\alpha}^{\prime} & 0 & 0 & 0 & 0 & 0 & -B_{2} / \varrho \mu & 0 & 0 \\
0 & 0 & 0 & 0 & \widehat{\beta}^{\prime} & \widehat{\beta}^{\prime \prime} & B_{1} / \varrho \mu & 0 & B_{3} / \varrho \mu \\
0 & 0 & \widehat{\gamma}^{\prime} & 0 & 0 & 0 & 0 & 0 & -B_{2} / \varrho \mu \\
0 & 0 & \widehat{\gamma}^{\prime \prime} & 0 & 0 & u_{2} & 0 & 0 & 0 \\
0 & -B_{2} & B_{1} & 0 & 0 & 0 & u_{2} & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & u_{2} & 0 \\
0 & 0 & B_{3} & -B_{2} & 0 & 0 & 0 & 0 & u_{2} \\
0 & 0 & 0 & & & 0 & & 0 & 0 \\
0 & 0 & 0 & \varrho & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & -B_{3} / \varrho \mu & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & -B_{3} / \varrho \mu & 0 \\
\widehat{\alpha}^{\prime} & 0 & 0 & 0 & \widehat{\beta}^{\prime} & \widehat{\beta}^{\prime \prime} & B_{1} / \varrho \mu & B_{2} / \varrho \mu & 0 \\
0 & 0 & 0 & \widehat{\gamma}^{\prime} & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \widehat{\gamma}^{\prime \prime} & 0 & u_{3} & 0 & 0 & 0 \\
0 & -B_{3} & 0 & B_{1} & 0 & 0 & u_{3} & 0 & 0 \\
0 & 0 & -B_{3} & B_{2} & 0 & 0 & 0 & u_{3} & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & u_{3}
\end{array}\right),
$$

and

$$
\widehat{\mathcal{D}}=\left(\begin{array}{ccccccccc}
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \widehat{\delta}^{\prime} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \hat{\delta}^{\prime \prime} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & \lambda
\end{array}\right) \widehat{\mathcal{B}}:=\left(\begin{array}{c}
0 \\
-\nu \frac{\vec{u}}{\varrho} \\
-\frac{\sigma_{a}}{\varrho C_{v}}\left(a \vartheta^{4}-E_{r}\right)+\frac{1}{\varrho C_{v}} \frac{\lambda}{\mu}\left|\operatorname{curl}_{x} \vec{B}\right|^{2}+\frac{\nu \mid \overrightarrow{u^{2}}}{\varrho C_{v}} \\
\sigma_{a}\left(a \vartheta^{4}-E_{r}\right) \\
0 \\
0 \\
0
\end{array}\right),
$$

where

$$
\widehat{\alpha}^{\prime}=\frac{p_{\varrho}}{\varrho}, \quad \widehat{\beta}^{\prime}=\frac{p_{\vartheta}}{\varrho}, \quad \widehat{\beta}^{\prime \prime}=\frac{1}{3 \varrho}, \quad \widehat{\gamma}^{\prime}=\frac{3 \varrho p_{\vartheta}}{3 \varrho C_{v}}, \quad \widehat{\delta}^{\prime}=\frac{\kappa}{\varrho C_{v}}, \quad \widehat{\gamma}^{\prime \prime}=\frac{4}{3} E_{r}, \quad \widehat{\delta}^{\prime \prime}=\frac{1}{3 \sigma_{s}} .
$$

It is possible to symmetrize this nonlinear system in the same spirit as what we have done for the linearized system (3.7). However, we do not need to do so here. So we write

$$
\partial_{t} V=-\sum_{j=1}^{3}\left[\widehat{\mathcal{A}}_{j}(V)-\widehat{\mathcal{A}}_{j}(\bar{V})\right] \partial_{x_{j}} V-\sum_{j=1}^{n} \widehat{\mathcal{A}}_{j}(\bar{V}) \partial_{x_{j}} V+[\widehat{\mathcal{D}}(V)-\widehat{\mathcal{D}}(\bar{V})] \Delta V+\widehat{\mathcal{D}}(\bar{V}) \Delta V-\widehat{\mathcal{B}}(V) .
$$

We first observe that these matrices are Lipschitz continuous with respect to $V$, away from $\varrho=0$ and $\vartheta=0$ and also that the matrices $\widehat{\mathcal{B}}$ and $\widehat{\mathcal{D}}$ have, respectively, the same structure as those defined in (3.10). Note also that, since
$d-1>5 / 2=3 / 2+1$, Sobolev embeddings imply that $H^{d-1}\left(\mathbb{R}^{3}\right)$ is an algebra. Therefore, we have

$$
\begin{aligned}
& \left\|\partial_{t} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C_{0}\left(1+\sum_{j=1}^{3}\left\|\widehat{\mathcal{A}}_{j}(V)-\widehat{\mathcal{A}}_{j}(\bar{V})\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \\
& +C_{0}\left(1+\|\widehat{\mathcal{D}}(V)-\widehat{\mathcal{D}}(\bar{V})\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)\left(\|\Delta \vartheta\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\Delta E_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\operatorname{curl}_{x}\left(\operatorname{curl}_{x} \vec{B}\right)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right) \\
& \\
& +C_{0}\left(1+\|\widehat{\mathcal{B}}(V)-\widehat{\mathcal{B}}(\bar{V})\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)\left(\left\|\vartheta-T_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\|\vec{u}\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)
\end{aligned}
$$

whence,

$$
\begin{aligned}
\left\|\partial_{t} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C_{0}\left(1+\|V-\bar{V}\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)\left(\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\|\Delta \vartheta\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\Delta E_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right. \\
\left.+\|\Delta \vec{B}\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\vartheta-T_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\|\vec{u}\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\right)
\end{aligned}
$$

which proves (3.32).
Next, we bound the spatial derivatives as follows:
Proposition 3.4. Assume that the hypotheses of Theorem 2.1 are satisfied. Let $k \in \mathbb{N}^{3}$ be such that $1 \leq|k| \leq d$, where $d>7 / 2$. Then, we have

$$
\begin{align*}
&\left\|\partial_{x}^{k} V(t)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\int_{0}^{t}\left(\left\|\partial_{x}^{k} \nabla_{x} \vartheta(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\right.\left\|\partial_{x}^{k} \nabla_{x} E_{r}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\partial_{x}^{k}\left(\vartheta-T_{r}\right)(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} \\
&\left.+\left\|\partial_{x}^{k} \nabla_{x} \vec{B}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\partial_{x}^{k} \vec{u}(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
& \leq C_{0} N(0)^{2}+C_{0} N(t) \int_{0}^{t}\left(\left\|\nabla_{x} V(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\right. \\
&\left.\left\|\vartheta(s)-T_{r}(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \tag{3.34}
\end{align*}
$$

Proof: Here, we need to symmetrize the nonlinear system. For this purpose, we multiply (3.33) on the right by the matrix

$$
\widehat{\mathcal{A}}_{0}(V)=\left(\begin{array}{ccccccccc}
\mu \frac{\widehat{\alpha}^{\prime}}{\varrho} & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0  \tag{3.35}\\
0 & \mu & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \mu & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & \mu & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & \mu \frac{\widehat{\beta}^{\prime}}{\widehat{\gamma}^{\prime}} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \mu \frac{\widehat{\beta}^{\prime \prime}}{\gamma^{\prime \prime}} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 1 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 1
\end{array}\right) .
$$

This gives

$$
\begin{equation*}
\widehat{\mathcal{A}}_{0}(V) \partial_{t} V=-\sum_{j=1}^{3} \check{\mathcal{A}}_{j}(V) \partial_{x_{j}} V+\check{\mathcal{D}}(V) \Delta V-\check{\mathcal{B}}(V)=0 \tag{3.36}
\end{equation*}
$$

where $\breve{\mathcal{A}}_{j}(V)=\widehat{\mathcal{A}}_{0}(V) \widehat{\mathcal{A}}_{j}(V), \breve{\mathcal{B}}(V)=\widehat{\mathcal{A}}_{0}(V) \widehat{\mathcal{B}}(V)$, and $\check{\mathcal{D}}(V)=\widehat{\mathcal{A}}_{0}(V) \widehat{\mathcal{D}}(V)$ are all symmetric matrices. Applying $\partial_{x}^{k}$ to (3.36) then taking the scalar product with the vector $\partial_{x}^{k} V$, and integrating over $[0, t] \times \mathbb{R}^{3}$, we find

$$
\begin{aligned}
\frac{1}{2} \int_{\mathbb{R}^{3}}\left[\left(\widehat{\mathcal{A}}_{0} \partial_{x}^{k} V\right) \cdot \partial_{x}^{k} V\right]_{0}^{t} d x+\int_{0}^{t} \int_{\mathbb{R}^{3}}\left(\check{\mathcal{D}} \nabla_{x}\left(\partial_{x}^{k} V\right)\right) \cdot \nabla_{x}\left(\partial_{x}^{k} V\right) d x d t & +\int_{0}^{t} \int_{\mathbb{R}^{3}}\left(\check{\mathcal{B}}(V) \partial_{x}^{k} V\right) \cdot\left(\partial_{x}^{k} V\right) d x d t \\
& =\int_{0}^{t} \int_{\mathbb{R}^{3}}\left(\frac{1}{2}\left(I_{1}+I_{2}\right)-I_{3}-I_{4}-I_{5}\right) d x d t
\end{aligned}
$$

where

$$
\begin{array}{ll}
I_{1}=\partial_{t}\left(\widehat{\mathcal{A}}_{0}(V)\right) \partial_{x}^{k} V \cdot \partial_{x}^{k} V, & I_{2}=\sum_{j=1}^{3} \partial_{x_{j}}\left(\check{\mathcal{A}}_{j}(V)\right) \partial_{x}^{k} V \cdot \partial_{x}^{k} V, \quad I_{3}=\left[\partial_{x}^{k}, \widehat{\mathcal{A}}_{0}(V)\right] \partial_{t} V \cdot \partial_{x}^{k} V, \\
I_{4} & =\sum_{j=1}^{3}\left[\partial_{x}^{k}, \breve{\mathcal{A}}_{j}(V)\right] \partial_{x_{j}} V \cdot \partial_{x}^{k} V,
\end{array} I_{5}=\partial_{x}^{k}(\check{\mathcal{B}}(V)) \cdot \partial_{x}^{k} V .
$$

We estimate separately each term of the right-hand side.
First, we have

$$
\begin{gathered}
\int_{0}^{t} \int_{\mathbb{R}^{3}}\left|I_{1}\right| \leq C \int_{0}^{t} \int_{\mathbb{R}^{3}}\left|\partial_{x}^{k} V\right|^{2}\left|\partial_{t} V\right| \leq C \int_{0}^{t} \int_{\mathbb{R}^{3}}\left|\partial_{x}^{k} V\right|^{2}\left(\left|\nabla_{x} V\right|+|\mathcal{B}(V)|+|\mathcal{D} \Delta V|\right) \\
\leq C N(t) \int_{0}^{t}\left\|\partial_{x}^{k} V(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} d s
\end{gathered}
$$

where we have used Sobolev embeddings and the fact that $d>7 / 2$. A similar computation gives

$$
\int_{0}^{t} \int_{\mathbb{R}^{3}}\left|I_{2}\right| \leq C N(t) \int_{0}^{t}\left\|\partial_{x}^{k} V(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} d s
$$

We estimate $I_{3}$ by applying Cauchy-Schwarz inequality:

$$
\int_{0}^{t} \int_{\mathbb{R}^{3}}\left|I_{3}\right| \leq \int_{0}^{t}\left\|\partial_{x}^{k} V\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}\left\|\left[\partial_{x}^{k}, \widehat{\mathcal{A}}_{0}(V)\right] \partial_{t} V\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}
$$

Next, we apply the same estimate for commutators and composition of functions (see [15, Proposition 2.1]), and $|k| \leq d:$

$$
\begin{aligned}
& \left\|\left[\partial_{x}^{k}, \widehat{\mathcal{A}}_{0}(V)\right] \partial_{t} V\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}=\left\|\left[\partial_{x}^{k}, \widehat{\mathcal{A}}_{0}(V)-\widehat{\mathcal{A}}_{0}(\bar{V})\right] \partial_{t} V\right\|_{L^{2}\left(\mathbb{R}^{3}\right)} \\
& \quad \leq C\left(\left\|\partial_{t} V\right\|_{L^{\infty}\left(\mathbb{R}^{3}\right)}\left\|\nabla_{x} \widehat{\mathcal{A}}_{0}(V)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}+\left\|\partial_{t} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}\left\|\nabla_{x} \widehat{\mathcal{A}}_{0}(V)\right\|_{L^{\infty}\left(\mathbb{R}^{3}\right)}\right) .
\end{aligned}
$$

Moreover, we have

$$
\left\|\nabla_{x} \widehat{\mathcal{A}}_{0}(V)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C\|V-\bar{V}\|_{H^{d}\left(\mathbb{R}^{3}\right)} \leq C N(t),
$$

and

$$
\left\|\nabla_{x} \widehat{\mathcal{A}}_{0}(V)\right\|_{L^{\infty}\left(\mathbb{R}^{3}\right)} \leq C\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C N(t)
$$

Hence, $I_{3}$ satisfies

$$
\begin{aligned}
& \int_{0}^{t} \int_{\mathbb{R}^{3}}\left|I_{3}\right| \leq C(N(t)) N(t) \int_{0}^{t}\left(\left\|\nabla_{x} V(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right. \\
&\left.+\left\|\vartheta(s)-T_{r}(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s .
\end{aligned}
$$

Here, we have used (3.32).
The integral of $I_{4}$ is dealt with using similar computations.
Turning to $I_{5}$, we use the particular form of $\partial_{x}^{k} \breve{\mathcal{B}}(V)$. More precisely, we have

$$
\begin{gathered}
\partial_{x}^{k}(\check{\mathcal{B}}(V)) \cdot \partial_{x}^{k} V=\partial_{x}^{k}\left(\frac{\vec{f}_{m}}{\varrho}-\nu \frac{\vec{u}}{\varrho}\right) \cdot \partial_{x}^{k} \vec{u}-\partial_{x}^{k}\left(\frac{\sigma_{a}}{\varrho}\left(a \vartheta^{4}-E_{r}\right)+\frac{E_{m}-\vec{f}_{m} \cdot \vec{u}}{\varrho}\right) \cdot \partial_{x}^{k} \vartheta \\
+\partial_{x}^{k}\left(\sigma_{a}\left(a \vartheta^{4}-E_{r}\right)\right) \cdot \partial_{x}^{k} E_{r}
\end{gathered}
$$

from which, using estimates for composition of functions (see Proposition 2.1 in [15]) we infer

$$
\int_{0}^{t} \int_{\mathbb{R}^{3}}\left|I_{5}\right| \leq C N(t) \int_{0}^{t}\left\|\partial_{x}^{k} V(s)\right\|_{L^{2}\left(\mathbb{R}^{3}\right)}^{2} d s
$$

Collecting the estimates on $I_{1}, I_{2}, I_{3}, I_{4}$ and $I_{5}$, we have proved (3.34).
The above results allow to derive the following bound:
Proposition 3.5. Assume that the assumptions of Theorem 2.1 are satisfied. Then, there exists a non-decreasing function $C: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+}$such that

$$
\begin{align*}
\|V-\bar{V}\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\int_{0}^{t}\left(\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right. & +\left\|\nabla_{x} E_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\vartheta(s)-T_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2} \\
& \left.+\|\vec{u}(s)\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s \\
\leq C(N(t))\left[N(0)^{2}+N(t) \int_{0}^{t}\right. & \left(\left\|\nabla_{x} V(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vartheta(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\right. \\
& \left.\left.\left\|\vartheta(s)-T_{r}(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\|\vec{u}(s)\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}(s)\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right) d s\right] . \tag{3.37}
\end{align*}
$$

Proof: We sum up estimates (3.34) over all multi-indices $k$ such that $|k| \leq d$, and add this to (3.29). This leads to (3.37).

### 3.2.2 $L^{2}\left(H^{d-1}\right)$ estimates

In this section, we derive bounds on the right-hand side of (3.37). For this purpose, we adapt the strategy of [22], which was further developed in [10]. We apply the Fourier transform to the linearized system and use the compensating matrix $K$ to prove estimates on the space derivatives of $V$.
Proposition 3.6. Assume that the assumptions of Theorem 2.1 are satisfied. Then there exists a non-decreasing function $C: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+}$such that

$$
\begin{equation*}
\int_{0}^{t}\left\|\nabla_{x} V(s)\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} d s \leq C(N(t))\left(N(t)+\left\|V_{0}-\bar{V}\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}\right) \tag{3.38}
\end{equation*}
$$

Proof: As a first step, we apply the symmetrizer of the linearized system (3.7) (which leads to (3.9)) to the nonlinear system (1.10)-(1.11)-(1.12), which then reads

$$
\widetilde{\mathcal{A}}_{0}(V) \partial_{t} V+\sum_{j=1}^{3} \widetilde{\mathcal{A}}_{j}(V) \partial_{x_{j}} V=\widetilde{\mathcal{D}} \Delta V-\widetilde{\mathcal{B}}(V) V
$$

Of course, this system is not symmetric. However, the corresponding linearized system (3.9) is symmetric. Next, we rewrite the nonlinear system by setting $U=V-\bar{V}$ :

$$
\widetilde{\mathcal{A}}_{0}(V) \partial_{t} U+\sum_{j=1}^{3} \widetilde{\mathcal{A}}_{j}(V) \partial_{x_{j}} U=\widetilde{\mathcal{D}} \Delta U-\widetilde{\mathcal{B}}(V) U-\widetilde{\mathcal{B}}(V) \bar{V}
$$

Therefore, multiplying this system on the left by $\widetilde{\mathcal{A}}_{0}(\bar{V})\left(\widetilde{\mathcal{A}}_{0}(V)\right)^{-1}$, we find

$$
\begin{equation*}
\widetilde{\mathcal{A}}_{0}(\bar{V}) \partial_{t} U+\sum_{j=1}^{3} \widetilde{\mathcal{A}}_{j}(\bar{V}) \partial_{x_{j}} U=H \tag{3.39}
\end{equation*}
$$

where

$$
\begin{aligned}
& H=-\widetilde{\mathcal{A}}_{0}(\bar{V}) \sum_{j=1}^{3}\left[\left(\widetilde{\mathcal{A}}_{0}(V)\right)^{-1} \widetilde{\mathcal{A}}_{j}(V)-\left(\widetilde{\mathcal{A}}_{0}(\bar{V})\right)^{-1} \widetilde{\mathcal{A}}_{j}(\bar{V})\right] \partial_{x_{j}} V \\
&+\widetilde{\mathcal{A}}_{0}(\bar{V})\left(\widetilde{\mathcal{A}}_{0}(V)\right)^{-1} \widetilde{\mathcal{D}} \Delta U-\widetilde{\mathcal{A}}_{0}(\bar{V})\left(\widetilde{\mathcal{A}}_{0}(V)\right)^{-1} \widetilde{\mathcal{B}}(V) U-\widetilde{\mathcal{A}}_{0}(\bar{V})\left(\widetilde{\mathcal{A}}_{0}(V)\right)^{-1} \widetilde{\mathcal{B}}(V) \bar{V}
\end{aligned}
$$

We apply the Fourier transform to (3.39), and then multiply on the left by $-i(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right)$, where * denotes the transpose of the complex conjugate, and $K$ is the compensating matrix (see Proposition 3.1). Taking the real part of the result, we infer

$$
\begin{equation*}
\operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \partial_{t} \widehat{U}\right)+|\xi|(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}\left(\frac{\xi}{|\xi|}\right) \widehat{U}=\operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \widehat{H}\right) \tag{3.40}
\end{equation*}
$$

where the matrix $\mathcal{A}\left(\frac{\xi}{|\xi|}\right)$ is defined by (3.13). According to Proposition 3.1, $K \mathcal{A}_{0}(\bar{V})$ is skew-symmetric, hence

$$
\operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \partial_{t} \widehat{U}\right)=\frac{1}{2} \frac{d}{d t} \operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \widehat{U}\right)
$$

Next, we also have

$$
\begin{equation*}
|\xi|(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}\left(\frac{\xi}{|\xi|}\right) \widehat{U}=|\xi|(\widehat{U})^{*}\left[K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}\left(\frac{\xi}{|\xi|}\right)+\mathcal{B}\left(\frac{\xi}{|\xi|}\right)\right] \widehat{U}-|\xi|(\widehat{U})^{*} \widetilde{\mathcal{B}} \widehat{U}-|\xi|(\widehat{U})^{*} \widetilde{\mathcal{D}} \widehat{U} \tag{3.41}
\end{equation*}
$$

Hence, still applying Proposition 3.1, there exists $\alpha_{1}>0$ and $\alpha_{2}>0$ such that

$$
\begin{align*}
&|\xi|(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}\left(\frac{\xi}{|\xi|}\right) \widehat{U} \geq \alpha_{1}|\xi||\widehat{U}|^{2} \\
&-\alpha_{2} \frac{1}{|\xi|}\left(|\xi(\widehat{\vartheta-\bar{\vartheta}})|^{2}+\left|\xi\left(\widehat{E_{r}-\overline{E_{r}}}\right)\right|^{2}+|\xi|^{2}|\widehat{\vec{B}-\overline{\vec{B}}}|^{2}+|\xi|^{2}|\widehat{\vec{u}}|^{2}+|\xi|^{2}\left|\widehat{\vartheta-T_{r}}\right|^{2}\right) \tag{3.42}
\end{align*}
$$

Finally, we estimate the right-hand side of (3.40) using Cauchy-Schwarz inequality and Young inequality:

$$
\begin{equation*}
\left|\operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \widehat{H}\right)\right| \leq \varepsilon|\xi||\widehat{U}|^{2}+C_{\varepsilon} \frac{1}{|\xi|}|\widehat{H}|^{2} \tag{3.43}
\end{equation*}
$$

for any $\varepsilon>0$. We choose $\varepsilon$ small enough, insert (3.41)-(3.42)-(3.43) into (3.40), and find

$$
\begin{aligned}
|\xi||\widehat{U}|^{2} \leq C\left[\frac { 1 } { | \xi | } \left(|\xi(\widehat{\vartheta-\bar{\vartheta}})|^{2}+\left|\xi\left(\widehat{E_{r}-\overline{E_{r}}}\right)\right|^{2}+|\xi|^{2}|\widehat{\vec{B}-\vec{B}}|^{2}\right.\right. & \left.+|\xi|^{2}|\widehat{\vec{u}}|^{2}+|\xi|^{2}\left|\widehat{\vartheta-T_{r}}\right|^{2}\right) \\
& \left.+\frac{1}{|\xi|}|\widehat{H}|^{2}-\frac{d}{d t} \operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \widehat{U}\right)\right]
\end{aligned}
$$

We multiply this inequality by $|\xi|^{2 l-1}$, for some $1 \leq l \leq d$, and get

$$
\begin{array}{r}
|\xi|^{2 l}|\widehat{U}|^{2} \leq C\left[|\xi|^{2 l-2}\left(|\xi(\widehat{\vartheta-\bar{\vartheta}})|^{2}+\left|\xi\left(\widehat{E_{r}-\overline{E_{r}}}\right)\right|^{2}+|\xi|^{2}|\widehat{\vec{B}-\overline{\vec{B}}}|^{2}+|\xi|^{2}|\widehat{\vec{u}}|^{2}+|\xi|^{2}\left|\widehat{\vartheta-T_{r}}\right|^{2}\right)+|\xi|^{2 l-2}|\widehat{H}|^{2}\right. \\
\left.-|\xi|^{2 l-1} \frac{d}{d t} \operatorname{Im}\left((\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \widehat{U}\right)\right] \tag{3.44}
\end{array}
$$

We integrate this inequality over $[0, t] \times \mathbb{R}^{3}$, and use Plancherel's theorem:

$$
\begin{array}{r}
\int_{0}^{t} \int_{\mathbb{R}^{3}} \sum_{|k|=l-1}\left|\partial_{x}^{k} \nabla_{x} V\right|^{2} \leq C \int_{0}^{t} \int_{\mathbb{R}^{3}} \sum_{|k|=l-1}\left(\left|\partial_{x}^{k} \nabla_{x} \vartheta\right|^{2}+\left|\partial_{x}^{k} \nabla_{x} E_{r}\right|^{2}+\left|\partial_{x}^{k} \nabla_{x} \vec{B}\right|^{2}+\left|\partial_{x}^{k} \nabla_{x} \vec{u}\right|^{2}+\left|\partial_{x}^{k} H\right|^{2}\right) \\
+C \operatorname{Im} \int_{\mathbb{R}^{3}}|\xi|^{2 l-1}\left[(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \widehat{U}\right]_{0}^{t} \tag{3.45}
\end{array}
$$

The matrix $K\left(\frac{\xi}{|\xi|}\right)$ is uniformly bounded for $\xi \in \mathbb{R}^{3} \backslash\{0\}$, so we have

$$
\begin{aligned}
\operatorname{Im} \int_{\mathbb{R}^{3}}|\xi|^{2 l-1}\left[(\widehat{U})^{*} K\left(\frac{\xi}{|\xi|}\right) \mathcal{A}_{0}(\bar{V}) \widehat{U}\right]_{0}^{t} \leq C\left(\int_{\mathbb{R}^{3}}\left(1+|\xi|^{2}\right)^{l}|\widehat{U}(t)|^{2}\right. & \left.+\int_{\mathbb{R}^{3}}\left(1+|\xi|^{2}\right)^{l}\left|\widehat{U}_{0}\right|^{2}\right) \\
& \leq C\left(\|V-\bar{V}\|_{H^{l}\left(\mathbb{R}^{3}\right)}^{2}+\left\|V_{0}-\bar{V}\right\|_{H^{l}\left(\mathbb{R}^{3}\right)}^{2}\right)
\end{aligned}
$$

We insert this estimate into (3.45), sum the result over $1 \leq l \leq d$, which leads to

$$
\begin{align*}
& \int_{0}^{t}\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)} \leq C\left(\|V-\bar{V}\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}+\left\|V_{0}-\bar{V}\right\|_{H^{d}\left(\mathbb{R}^{3}\right)}^{2}\right. \\
& \left.\quad+\int_{0}^{t}\left(\left\|\nabla_{x} \vartheta\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} E_{r}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{u}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\left\|\nabla_{x} \vec{B}\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}+\|H\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2}\right)\right) \tag{3.46}
\end{align*}
$$

In order to conclude, we need to estimate the perturbation $H$. For this purpose, we use that $H^{d-1}\left(\mathbb{R}^{3}\right)$ is an algebra: for any $s \leq t$,

$$
\|H(s)\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}^{2} \leq C N(t)\left\|\nabla_{x} V\right\|_{H^{d-1}\left(\mathbb{R}^{3}\right)}
$$

Inserting this into (3.46), we prove (3.38).
We are now in position to conclude with the
Proof of Theorem 2.1: We first point out that local existence for system (1.10)-(1.11)-(1.12) may be proved using standard fix-point methods. We refer to [15] for the proof. The existence is proved in the following functional space:

$$
X(0, T)=\left\{V, \quad V-\bar{V} \in C\left([0, T] ; H^{d}\left(\mathbb{R}^{3}\right)\right), \quad \nabla_{x} V \in L^{2}\left([0, T] ; H^{d-1}\left(\mathbb{R}^{3}\right)\right), ~ 子 \quad \nabla_{x} \vartheta, \nabla_{x} E_{r}, \nabla_{x} \vec{B} \in L^{2}\left([0, T] ; H^{d}\left(\mathbb{R}^{3}\right)\right)\right\} .
$$

In order to prove global existence, we argue by contradiction, and assume that $T_{c}>0$ is the maximum time existence. Then, we necessarily have

$$
\begin{equation*}
\lim _{t \rightarrow T_{c}} N(t)=+\infty \tag{3.47}
\end{equation*}
$$

where $N(t)$ is defined by (3.27). We are thus reduced to prove that $N$ is bounded. For this purpose, we use the method of [13], which was also used in [18]. First note that, due to Proposition 3.5 on the one hand, and to Proposition 3.6 on the other hand, we know that there exists a non-decreasing continuous function $C: \mathbb{R}^{+} \rightarrow \mathbb{R}^{+}$such that

$$
\begin{equation*}
\forall T \in\left[0, T_{c}\right], \quad N(t)^{2} \leq C(N(t))\left(N(0)^{2}+N(t)^{3}\right) \tag{3.48}
\end{equation*}
$$

Hence, setting $N(0)=\varepsilon$, we have

$$
\begin{equation*}
\frac{N(t)^{2}}{\varepsilon^{2}+N(t)^{3}} \leq C(N(t)) \tag{3.49}
\end{equation*}
$$

Studying the variation of $\phi(N)=N^{2} /\left(\varepsilon^{2}+N^{3}\right)$, we see that $\phi^{\prime}(0)=0$, that $\phi$ is increasing on the interval $\left[0,\left(2 \varepsilon^{2}\right)^{1 / 3}\right]$ and decreasing on the interval $\left[\left(2 \varepsilon^{2}\right)^{1 / 3},+\infty\right)$. Hence,

$$
\max \phi=\phi\left(\left(2 \varepsilon^{2}\right)^{1 / 3}\right)=\frac{1}{3}\left(\frac{2}{\varepsilon}\right)^{2 / 3}
$$

Hence, the function $C$ being independent of $\varepsilon$, we can choose $\varepsilon$ small enough to have $\phi(N) \leq C(N)$ for all $N \in\left[0, N^{*}\right]$, where $N^{*}>0$. Since $C$ is continuous, (3.49) implies that $N \leq N^{*}$. This is clearly in contradiction with (3.47).

Acknowlegment: Šárka Nečasová acknowledges the support of the GAČR (Czech Science Foundation) project P201-13-00522S in the framework of RVO: 67985840.

Bernard Ducomet is partially supported by the ANR project INFAMIE (ANR-15-CE40-0011).

## References

[1] C. Beauchard, E. Zuazua. Large time asymptotics for partially dissipative hyperbolic systems. Arch. Rational Mech. Anal., 199:177-227, 2011.
[2] S. Benzoni-Gavage, D. Serre. Multi-dimensional hyperbolic partial differential equations. Oxford Science Publications, 2007.
[3] X. Blanc, B. Ducomet. Weak and strong solutions of equations of compressible magnetohydrodynamics. arXiv:150905573v1 [math.AP] 18 Sep 2015.
[4] X. Blanc, B. Ducomet, Š. Nečasová. On some singular limits in damped radiation hydrodynamics. Journal of Hyperbolic Differential Equations, 13:249-271, 2016.
[5] C. Buet, B. Després. On some singular limits in damped radiation hydrodynamicsAsymptotic analysis of fluid models for the coupling of radiation and hydrodynamics. Journal of Qantitative Spectroscopy and Radiative Transfer, 85:385-418, 2004.
[6] H. Cabannes. Theoretical magnetohydrodynamics. Academic Press, New York, 1970.
[7] A.R. Choudhuri The physics of fluids and plasmas: an introduction for astrophysicists. Cambridge University Press, 1998.
[8] G.S. Dulikravich, S.R. Lynn. Unified electro-magneto-fluid dynamics: a survey of mathematical models. Int. J. Non-Linear Mechanics, 32 (1997) 923-932.
[9] E. Feireisl and A. Novotný, Singular limits in thermodynamics of viscous fluids. Birkhauser, Basel, 2009.
[10] B. Hanouzet, R. Natalini. Global existence of smooth solutions for partially dissipative hyperbolic systems with a convex entropy. Arch. Ration. Mech. Anal., 69(2):89-117, 2003.
[11] I. Imai. General principles of magneto-fluid dynamics. Suppl. of the Progress of Theor. Phys. 24 1-34, 1962.
[12] R. E. Kalman, P. L. Falb, M. A. Arbib. Topics in mathematical system theory McGraw-Hill Book Co., New York-Toronto, Ont.-London, 1969.
[13] C. Lin, T. Goudon. Global existence of the equilibrium diffusion model in radiative hydrodynamics. Chin. Ann. Math., 32B:549-568, 2011.
[14] R.B. Lowrie, J.E. Morel, J.A.. Hittinger. The coupling of radiation and hydrodynamics. The Astrophysical Journal, 521:432-450, 1999.
[15] A. Majda. Compressible fluid flow and systems of conservation laws in several variables. Springer-Verlag, NewYork, Berlin, Heidelberg, Tokyo, 1984.
[16] P.A. Markowich, C. Ringhofer, C. Schmeiser. Semiconductor equations. Springer-Verlag, New-York, Berlin, Heidelberg, Tokyo, 1990.
[17] D. Mihalas, B. Weibel-Mihalas. Foundations of radiation hydrodynamics. Oxford University Press, New York, 1984.
[18] T. Nishida, Nonlinear hyperbolic equations and related topics in fluid dynamics, Publications Mathématiques d'Orsay, No. 78-02, Département de Mathématique, Université de Paris-Sud, Orsay, 1978.
[19] G.C. Pomraning. Radiation hydrodynamics. Dover Publications, Inc., Mineola, New York, 2005.
[20] D. Serre. Systèmes de lois de conservation I, II. Diderot Editeur, Arts et Sciences, Paris, New-York, Amsterdam, 1996.
[21] D. Serre. Systems of conservation laws with dissipation. Lecture Notes SISSA, 2007.
[22] Y. Shizuta, S. Kawashima. Systems of equation of hyperbolic-parabolic type with application to the discrete Boltzmann equation. Hokkaido Math. J., 14:249-275, 1985.
[23] Y. Ueda, S. Wang, S. Kawashima, Dissipative structure of the regularity-loss type and time asynptotic decay of solutions for the Euler-Maxwell system. SIAM J. Math. Anal., 44 2002-2017, 2012.

Xavier Blanc<br>Univ. Paris Diderot, Sorbonne Paris Cité,<br>Laboratoire Jacques-Louis Lions,<br>UMR 7598, UPMC, CNRS, F-75205 Paris, France<br>E-mail: blanc@ann.jussieu.fr<br>Bernard Ducomet<br>CEA, DAM, DIF<br>F-91297 Arpajon, France<br>E-mail: bernard.ducomet@cea.fr<br>Šárka Nečasová<br>Institute of Mathematics of the Academy of Sciences of the Czech Republic<br>Žitná 25, 11567 Praha 1, Czech Republic<br>E-mail: matus@math.cas.cz

