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## Towards Adaptive Grids for Atmospheric Boundary-Layer Simulations

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1 **Towards Adaptive Grids for Atmospheric Boundary-Layer**  
2 **Simulations**

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9 **Abstract** We present a proof-of-concept for the adaptive mesh refinement method  
10 applied to atmospheric boundary-layer simulations. Such a method may form an  
11 attractive alternative to static grids for studies on atmospheric flows that have a  
12 high degree of scale separation in space and/or time. Examples include the diurnal  
13 cycle and a convective boundary layer capped by a strong inversion. For such cases,  
14 large-eddy simulations using regular grids often have to rely on a subgrid-scale  
15 closure for the most challenging regions in the spatial and/or temporal domain.  
16 Here we analyze a flow configuration that describes the growth and subsequent  
17 decay of a convective boundary layer using direct numerical simulation (DNS).  
18 We validate the obtained results and benchmark the performance of the adaptive  
19 solver against two runs using fixed regular grids. It appears that the adaptive-mesh  
20 algorithm is able to coarsen and refine the grid dynamically whilst maintaining  
21 an accurate solution. In particular, during the initial growth of the convective  
22 boundary layer a high resolution is required compared to the subsequent stage  
23 of decaying turbulence. More specifically, the number of grid cells varies by two  
24 orders of magnitude over the course of the simulation. For this specific (DNS) case,  
25 the adaptive solver was not yet more efficient than the more traditional solver that  
26 is dedicated to these types of flows. However, the overall analysis shows that the  
27 method has a clear potential for numerical investigations of the most challenging  
28 atmospheric cases.

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29 **Keywords** Adaptive mesh refinement · Atmospheric boundary layer · Direct  
30 numerical simulations · Large-eddy simulations · Turbulence

## 31 1 Introduction

32 The aim of the present work is to introduce adaptive mesh refinement (AMR) as an  
33 efficient tool for numerical investigations of the atmospheric boundary layer (ABL)  
34 using turbulence resolving methods. This refers typically to models that rely on  
35 direct numerical simulation (DNS) or large-eddy simulation (LES) techniques. In  
36 general, AMR solvers aim to distribute the available computational resources effi-  
37 ciently over a domain by dynamically refining and coarsening the computational  
38 grid in space and time. AMR techniques have successfully been employed in stud-  
39 ies concerning flows with a high degree of scale separation throughout the spatial  
40 and/or temporal domain. Such studies concern a wide range of topics, e.g. cos-  
41 mological hydrodynamics (Teyssier 2002), electro hydrodynamics (López-Herrera  
42 et al. 2011), multiphase flows (Fuster et al. 2009), flows in complex geometries  
43 (Popinet 2003) and turbulence simulations (Schneider and Vasilyev 2010). How-  
44 ever, to our knowledge, the potential of this technique has not yet been explored  
45 for ABL research, and here we aim to do so through an investigation of the con-  
46 secutive growth and decay of a convective boundary-layer (CBL) system. The flow  
47 configuration is modelled after Van Heerwaarden and Mellado (2016) who per-  
48 formed an in-depth study of this case using a regular grid configuration. As such,  
49 the AMR method is tested and benchmarked.

50 Several methods that meet a varying resolution requirement throughout the  
51 spatial domain have already been successfully applied in studies on ABL turbu-  
52 lence. For example, stretching and squeezing of grids (see e.g. Heus et al. 2011,  
53 Van Heerwaarden and Mellado 2016, De Roode et al. 2016), nested grids (see e.g.  
54 Sullivan et al. 1996, 1998, Moeng et al. 2007, Mirocha et al. 2013, Muñoz-Esparza  
55 et al. 2014) and the usage of unstructured anisotropic grids. However, the mesh is  
56 always kept fixed during the simulation, whereas dynamical changes in the ABL  
57 call for variation of resolution in time. Furthermore, the aforementioned methods  
58 of refinement need to be predefined. Consequently, detailed a priori knowledge  
59 is needed on the varying resolution requirement throughout the spatial domain.  
60 Apart from tailored and well-known cases, this knowledge is usually not avail-  
61 able beforehand; therefore, we identify three favourable characteristics of an AMR  
62 approach for ABL studies. First, the resolution can vary throughout the spatial  
63 domain. Second, the grid can vary in time such that temporal variation in the  
64 local resolution requirement can be met. Third, the grid is generated adaptively  
65 based on the evolution of the numerical solution itself, relaxing the requirement  
66 of detailed a priori knowledge on the resolution requirement.

67 To illustrate our philosophy, we briefly discuss a textbook example of the evolu-  
68 tion of the ABL during a diurnal cycle (after Stull 1988). Figure 1 depicts a typical  
69 evolution of the ABL during a diurnal cycle. Around sunrise the solar irradiation of  
70 the Earth’s surface causes a thermal instability that results in the rapid growth of a  
71 CBL. The typical size of the largest thermal plumes scales with the boundary-layer  
72 height and hence there is a temporal dependency on the resolution requirement to  
73 resolve these turbulent structures. The growth of the boundary layer slows down  
74 when the rising thermals reach the inversion layer, which effectively caps turbu-

75 lent structures at the top of the CBL. The dynamics within an inversion layer are  
76 of pivotal importance for the evolution of the CBL (Garcia and Mellado 2014).  
77 Apart from the effective ‘lid’ on the boundary layer, entrainment processes occur  
78 here and the formation of stratocumulus clouds is promoted by the large jump in  
79 temperature with height. Due to the presence of strong stable stratification, tur-  
80 bulent length scales are suppressed (De Lozar and Mellado 2015), and in order to  
81 resolve the most prominent turbulent structures here, a much higher resolution is  
82 necessary compared to the bulk of the CBL (Sullivan and Patton 2011, De Lozar  
83 and Mellado 2015). Applying such high resolution everywhere in the domain is  
84 not feasible given the current status of computational resources, and might not  
85 be feasible in coming years (Bou-Zeid 2015). For this reason, many LES studies  
86 have to rely on their subgrid-scale (SGS) parametrizations within the region of  
87 the inversion layer, partially negating the purpose of a turbulence resolving study.  
88 Furthermore, the exact height and strength of the inversion layer are not always  
89 known a priori (except in cases that have been studied before). Fixed nested grids  
90 (Sullivan et al. 1998) are thus not always flexible enough to capture the dynamics  
91 properly. On the other hand, practically speaking, it should be noted that LES  
92 results between various studies often tend to converge, signifying that SGS models  
93 have appreciable skill in describing certain characteristics of the inversion layer  
94 (see e.g. Nieuwstadt et al. 1993, Siebesma et al. 2003).

95 At the approach of sunset, thermal plumes gradually decay into so-called resid-  
96 ual turbulence, and due to the radiative cooling of the Earth’s surface, stable  
97 stratification sets in and turbulence is now driven by wind shear only. The stable  
98 boundary layer (SBL) is typically much shallower than the CBL and, furthermore,  
99 the length scales of the turbulent structures that account for the mixing of heat  
100 and momentum within this layer are only a fraction of the size of those associated  
101 with daytime convective turbulence (Basu et al. 2008). Additionally, Anson and  
102 Mellado (2016) argue that the resolution requirement for their simulations of the  
103 intermittently turbulent SBL is dictated by localized dissipative flow structures  
104 that only encompass a fraction of the computational domain.

105 Rather than capturing the cyclic behaviour of the atmosphere as depicted in  
106 Fig. 1, the contrast between daytime and night-time turbulence has resulted in  
107 many numerical studies focusing only on either convective or stable conditions.  
108 The studies that do simulate a diurnal cycle typically struggle to resolve turbu-  
109 lence during the night (Kumar et al. 2006, Basu et al. 2008, Abkar et al. 2015).  
110 Furthermore, the transition period itself (i.e. around sunset) would benefit from  
111 high fidelity numerical studies (Lothon et al. 2014). In summary: the example  
112 shows that the intrinsic dynamic character of the ABL calls for flexible techniques  
113 such as an AMR approach in addition to existing techniques that have successfully  
114 been applied to studies on idealized, steady cases.

115 Apart from our long-term prospects, we focus here on a case corresponding to  
116 the red and grey sections in Fig. 1. This choice is motivated by the fact that as a  
117 first step, we would like to present a proof-of-concept of the AMR approach before  
118 we redirect our attention towards more challenging cases. Therefore, we present  
119 results obtained with DNS, for which all turbulent structures are resolved explicitly  
120 down to the small-scale Kolmogorov length (i.e. the viscous length scale) according  
121 to the Navier-Stokes equations, without any closure for turbulence. Compared to,  
122 for example, LES, the results obtained with DNS should be independent of the  
123 numerical formulations or choice of any SGS model, whereas with LES this is a

topic of discussion (Bretherton et al. 1999, Siebesma et al. 2003, Fedorovich et al. 2004, Beare et al. 2006, De Roode et al. 2017). However, as shown in Sect. 4, the concept of the AMR approach can be easily extended to LES. Since this technique is a popular choice for studies on the ABL, we also briefly discuss results obtained with the AMR technique using a LES formulation.

We realize that we cannot address all questions regarding the AMR technique in relation to ABL simulations. For example, here we focus on a single case whereas we will argue that the performance of an AMR solver varies depending on the particular case specifications (see Appendix 1). Furthermore, we choose a numerical solver called Basilisk (<http://basilisk.fr>) for the adaptive-grid runs and do not assess alternatives.

The paper is organized as follows; in Sect. 2.1 the details of the adaptive-grid solver are described, focusing on the AMR algorithm, and in addition, Sect. 2.2 provides an example analysis of how the algorithm assesses a turbulent signal and adapts the grid accordingly. In Sect. 2.3 the case and the numerical set-up of the different runs are specified. Section 3 presents the obtained results including a performance assessment, while in Sect. 4 we provide an outlook on future plans. We finish with a conclusion combined with a discussion in Sect. 5. Additionally, using a simple flow set-up, Appendix 1 illustrates an important advantage the AMR technique has over a fixed equidistant-grid approach.

## 2 Methods

### 2.1 Basilisk and the Grid Adaptation Algorithm

The AMR runs are performed with the partial-differential-equation solver called Basilisk, a code that contains a second-order accurate finite-volume solver for the Navier-Stokes equations. For a detailed description of the numerical formulations see Popinet (2003,2009), Lagr e et al. (2011), and references therein.

In order to facilitate local adaptive refinement and coarsening whilst maintaining a Cartesian grid-structure, a so-called tree-based grid is used. To illustrate this mesh structure, Fig. 2 shows the two-dimensional (2D) variant of a tree-based grid (i.e. a quadtree), whose structure introduces a hierarchy between cells at integer levels of refinement. The resolution between the levels of refinement differs by a factor of two and the Basilisk solver allows neighbouring cells to vary up to one level. The formulations of numerical methods (e.g. evaluating spatial derivatives) on equidistant Cartesian grids are relatively straightforward compared to their uneven grid counterparts. Therefore, ghost points are defined, enabling simple Cartesian stencil operations for the cells in the vicinity of a resolution boundary. These points act as virtual cells and are located such that all cells have neighbours that are defined at the same level of refinement, see Fig. 2b. The field values on these ghost cells are defined with interpolation techniques using the original field values.

The tree grid facilitates an efficient and convenient structure to perform a multi-resolution analysis of a discretized field. During the simulation, such an analysis is used to determine which grid cells require refinement and where in the domain cells can be coarsened. This procedure is discussed next. Consider a 1D signal ( $f$ ) discretized with an even number ( $n$ ) of elements  $f_n$ , where individual entries of

169  $f_n$  are indexed with  $i$  such that  $f_n^i$  represents the  $i$ -th entry of  $f_n$ . First, we define  
 170 a downsampling operation ( $D$ ) that approximates  $f_n$  on a coarser level grid with  
 171  $n/2$  elements,

$$f_{n/2} = D(f_n). \quad (1)$$

172 Second, we define an upsampling operator ( $U$ ) that samples  $f_{n/2}$  to a signal that  
 173 is defined with the same element entries as the original signal  $f_n$ ,

$$g_n = U(f_{n/2}), \quad (2)$$

174 noting that in general  $f_n \neq g_n$ , and the absolute difference  $\chi$ , defined as,

$$\chi_n^i = \|f_n^i - g_n^i\|, \quad (3)$$

175 can be interpreted as an estimation of the discretization error. The downsampling  
 176 operation in the Basilisk solver is defined as local volume averaging of the signal  
 177 to obtain a value for a corresponding coarser-level grid cell (see Fig. 3 a.). This  
 178 formulation is exact since in a finite-volume formulation, the grid cell values rep-  
 179 resent volume-averaged quantities. To be in line with the second-order accuracy  
 180 of the solver, the upsampling operation is chosen to be second-order accurate as  
 181 well, and entails performing a linear interpolation between the grid points of the  
 182 coarse level solution (see Fig. 3b). Once these two operations have been applied  
 183 to the discretized signal, it is possible to evaluate  $\chi_n^i$  for each of the grid cells.  
 184 Given an error threshold  $\zeta$ , the following assessment with regards to a grid-cell's  
 185 resolution can be made,

$$\text{the } i\text{-th grid cell is } \begin{cases} \text{too coarse.} & \chi_n^i > \zeta, \\ \text{too fine.} & \chi_n^i < \frac{2\zeta}{3}, \\ \text{just fine.} & \text{Otherwise.} \end{cases} \quad (4)$$

186 The threshold on the estimated error for refinement  $\zeta$  is called the refinement  
 187 criterion, with  $\zeta$  having the same physical units as  $f$ . Note that the described  
 188 method is formally linked to wavelet thresholding that has already been employed  
 189 for fluid dynamical simulations (Schneider and Vasilyev 2010). The grid can be  
 190 refined and coarsened according to Eq. 4 and field values for the new refined and  
 191 coarsened cells can be defined using an identical formulation as is used for the  $U$   
 192 and  $D$  operator, respectively. However, the Basilisk solver allows the formulations  
 193 for upsampling and downsampling during the grid-resolution assessment and the  
 194 actual refinement and coarsening of cells to differ.

195 In general, the tree grid that results from applying the adaptation algorithm  
 196 results in the presence of the aforementioned resolution boundaries and accompa-  
 197 nyng ghost cells within the domain (see Fig. 2). To define the field values of ghost  
 198 points, the Basilisk solver uses the downsampling and upsampling operations. The  
 199 implementation is visually represented for a 1D scenario in Fig. 4. First, down-  
 200 sampling is used to define the field values of ghost points on the high-resolution  
 201 side of a resolution boundary. Second, an upsampling method is used to define the  
 202 field values of the ghost points on the coarse side of the resolution boundary. By  
 203 using this method, the estimation error in the ghost cells' field values scales with  
 204  $\zeta$ .

The formulations used for downsampling and upsampling as exemplified in Figs. 3 and 4 can be easily extended to two and three dimensions, for so-called quadtree and octree grids, respectively. In order to demonstrate the algorithm and the effect of different  $\zeta$  values on the representation of a turbulent field, the next section shows the results of the algorithm applied to a slice of a 3D turbulent field.

The Basilisk solver can run in parallel on many processors by applying a domain decomposition using the Message Passing Interface (MPI). As the grid structure may change during a simulation run, an important issue is load-balancing; the decomposition of the domain between processors must then be modified as the grid is locally refined or coarsened. This is achieved in the Basilisk solver using the natural decomposition of a Z-ordering space-filling curve applied to the quad/octree structure (Griebel and Zumbusch 2001).

## 2.2 An Example of the Adaptation Algorithm

This section aims to exemplify how the adaption algorithm assesses a discretized signal and adapts the grid according to a refinement criterion  $\zeta$ . For this purpose, we apply the algorithm to a subset of the data from the simulation of forced isotropic turbulence in Li et al. (2008). The simulation is run using a fixed equidistant grid with  $1024^3$  nodes; in terms of the Kolmogorov length scale ( $\eta$ ), the grid spacing ( $\Delta_i$ ) is  $\Delta_i = 2.2\eta$ . For the analysis we assume the data to be resolved well enough, and the results are kindly made available via the Johns Hopkins turbulence databases (<http://turbulence.pha.jhu.edu/>). We analyze a 2D slice of the data (i.e.  $1024^2$  cells) and for simplicity, we only consider the velocity component perpendicular to the sliced plane ( $u_\perp$ ). The data are presented in Fig. 5a; using the algorithm described in Sect. 2.1, we can evaluate the  $\chi$  field corresponding to the original  $u_\perp$  field. A section of the resulting field, indicated by the black box in Fig. 5a, is shown in Fig. 5b, where we can clearly see that the estimated discretization error is not distributed uniformly by the equidistant-grid approach that was used in the simulation. Rather, it appears that there are anisotropic structures present, visualized by relatively high  $\chi$  values (in yellow). These structures appear to correspond to vortex filaments that characterize the dissipative structures of high-Reynolds-number turbulence (Frisch, 1995). This result motivates the application of the grid refinement algorithm to the data sample shown. Note that we cannot ‘add’ new information by refinement and at this point we do not make any claims regarding what  $\chi$  values are reasonable for a turbulence-resolving simulation (this will depend on the numerical formulations and is the topic of a future study). As such, we only allow the algorithm to coarsen the field with a maximum error threshold  $\zeta$  (as defined in Eq. 4). The number of grid cells resulting from the application of the adaptation algorithm for a range of  $\zeta$  values is shown in Fig. 5c; as expected, the number of grid cells decreases with an increasing  $\zeta$  value. Note that the plot also shows that even for the high  $\zeta$  values, the grid still contains cells at the maximum resolution.

The main concept of employing the described grid-adaption algorithm is visualized in Fig. 5d. Here histograms of the number of grid cells within 512 equally-spaced  $\chi$  bins are presented for the original data and the data obtained from applying the grid adaptation technique with three different refinement criteria. It appears that for the original dataset, the histogram is monotonically decreasing

with increasing  $\chi$ . This shows that many grid cells exist where the numerical solution is relatively smooth compared to cells in the tail of the histogram. Hence, if the grid is chosen such that the discretization errors in the latter region do not affect the relevant statistics of the flow evolution, then the grid must be over-refined elsewhere. The histograms of the adapted grids show that the algorithm is able to lower the number of grid cells with low  $\chi$  values, such that fewer grid cells are employed. Note that the grid coarsening does not introduce new grid cells with  $\chi > 2\zeta/3$ , as this part of the histogram remains unaltered.

When grid cells with a small but finite  $\chi$  value are coarsened, some of the data are lost and in general cannot be exactly reconstructed by interpolation techniques (see Sect. 2.4). In order to assess how the data from the adapted grids compare with the original data, Fig. 5e presents the corresponding power spectra. It appears that none of the adapted grid data are able to exactly reproduce the original power spectrum; more specifically, with increasing  $\zeta$  values, the wavenumbers ( $k$ ) that show a significant deviation in  $E(k)$  from the original appear to decrease. We would like to point out that in order to evaluate the spectrum we have linearly interpolated the data from the non-uniform grids to an equidistant grid with  $1024 \times 1024$  data points. The choice of the interpolation technique is arbitrary and will pollute the diagnosed spectrum in a non-trivial manner. As such, we directly compare all  $1024^2$   $u_{\perp}(x, y)$  samples in Fig. 5f, where we see that the deviation of the data from the 1 : 1 line is a function of  $\zeta$ .

The example presented in Fig. 5 is meant to demonstrate the used adaptation algorithm. The following sections are dedicated to assessing its application to time-dependent numerical simulations of a turbulent field for an atmospheric case.

### 2.3 Physical Case Set-up

As indicated in the Introduction, we ran a DNS case from the referenced literature to validate, benchmark and exemplify the adaptive-grid approach. The cases from virtually all atmospheric-turbulence-resolving studies prescribe the periodicity of the solution in the horizontal directions. Unfortunately, at the time of writing, the Basilisk solver cannot yet handle an adaptive grid in combination with periodic boundaries. To circumvent this limitation, we limit ourselves to a case where there is no mean horizontal forcing such that we can apply a no-penetration boundary condition for the normal-velocity component at the lateral boundaries. This is supplemented with a Neumann-boundary condition for the tangential velocity components, pressure and buoyancy fields. We realize that this choice might affect the solution and therefore its impact is assessed by re-running the case using a fixed and regular grid with both sets of lateral boundary conditions (not shown). It appears that for the chosen set-up of the case, the simulation results are insensitive to the choice of the horizontal boundary conditions. Note that in future work, we will update the adaptive solver such that periodic boundary conditions can be combined with the AMR technique.

We study a case introduced by Van Heerwaarden and Mellado (2016) that was designed to investigate the growth and decay of a CBL. In Fig. 6 a schematic overview of the physical system is presented, and in their physical model a linearly stratified fluid at rest with kinematic viscosity ( $\nu$ ) and thermal diffusivity ( $\kappa$ ) is heated from below by a surface with a constant temperature. For generality,



297 buoyancy ( $b$ ) is used as the thermodynamic variable. The buoyancy is related to  
 298 the potential temperature ( $\theta$ ) according to;

$$b = \frac{g}{\theta_{\text{ref}}}(\theta - \theta_{\text{ref}}), \quad (5)$$

299 where  $\theta_{\text{ref}}$  is a reference potential temperature and  $g$  the acceleration due to  
 300 gravity. The initial linear stratification is expressed as  $b(z) = N^2 z$ , where  $N^2$   
 301 is the Brunt-Väisälä frequency associated with the initial stratification and  $z$  is  
 302 the height above the surface. We assign a surface buoyancy  $b_0$  larger than zero.  
 303 Van Heerwaarden and Mellado (2016) identified relevant length, time, velocity  
 304 fluctuation and buoyancy flux scales,  $\mathcal{L}, T, U$  and  $B$ , respectively, according to;

$$\mathcal{L} = \frac{b_0}{N^2}, \quad (6a)$$

$$T = \frac{b_0^{2/3}}{N^2 \kappa^{1/3}}, \quad (6b)$$

$$U = \frac{b_0^{7/9} \kappa^{1/9}}{N^{2/3}}, \quad (6c)$$

$$B = b_0^{4/3} \kappa^{1/3}, \quad (6d)$$

305 and are used to analyze the results in a non-dimensional framework. Two dimen-  
 306 sionless groups can be identified that describe the system for any given set of  
 307  $\{\nu, \kappa, N^2, b_0\}$ ,

$$Pr = \frac{\nu}{\kappa}, \quad (7a)$$

$$Re = \left( \frac{b_0^{4/3}}{\nu^{2/3} N^2} \right)^{4/3}, \quad (7b)$$

308 where  $Pr$  is the Prandtl number and  $Re$  is the Reynolds number. Note that for  
 309  $Pr = 1$ , the definition of the Reynolds number is consistent with  $Re = U\mathcal{L}/\nu$ .

## 310 2.4 Numerical Set-up and Formulation

311 For the evolution of the three velocity components ( $u_i$ ), modified pressure ( $p$ ) and  
 312 buoyancy ( $b$ ), the Navier-Stokes equations for an incompressible fluid are solved  
 313 under the Boussinesq approximation, according to,

$$\frac{\partial u_i}{\partial t} + \frac{\partial u_j u_i}{\partial x_j} = -\frac{\partial p}{\partial x_i} + \nu \frac{\partial^2 u_i}{\partial x_i^2} + b \delta_{i3}, \quad (8)$$

$$\frac{\partial b}{\partial t} + \frac{\partial u_j b}{\partial x_j} = \kappa \frac{\partial^2 b}{\partial x_j^2}, \quad (9)$$

$$\frac{\partial u_j}{\partial x_j} = 0, \quad (10)$$

314 and with respect to no-slip and a fixed buoyancy ( $b_0$ ) condition at the bottom  
 315 boundary. At the top boundary, no-penetration with a free-slip condition is used

316 and for the buoyancy, a fixed vertical gradient ( $N^2$ ) is prescribed. Furthermore, a  
 317 damping layer in the top 25% of the domain is active that damps buoyancy and  
 318 velocity fluctuations to prevent the artificial reflection of gravity waves at the top  
 319 boundary. The adaptive-grid runs are initialized with a grid at the minimum reso-  
 320 lution that is locally refined to the maximum resolution near the bottom boundary  
 321 (i.e.  $z < \mathcal{L}/10$ ) before a random perturbation is added to the velocity components  
 322 and buoyancy field in each grid cell.

323 Each integration timestep, grid adaptation is based on the estimated error  
 324 (see Sect. 2.1) of the three velocity components, and the buoyancy field. For each  
 325 field a refinement criterion ( $\zeta$ ) is specified ( $\zeta_{u_i}, \zeta_b$ ), where we non-dimensionalize  
 326 the refinement criteria according to  $\xi_b = \zeta_b b_0^{-1}$  and  $\xi_{u_i} = \zeta_{u_i} U^{-1}$ . In order to  
 327 validate the results and assess the performance of the adaptive solver, we iteratively  
 328 decrease the refinement criterion between runs whilst we limit the minimum grid-  
 329 box size. This maximum resolution is inspired by Van Heerwaarden and Mellado  
 330 (2016), and to limit the degrees of freedom, we choose;  $\xi_{u_1} = \xi_{u_2} = \xi_{u_3} = 2.7 \times \xi_b$ .  
 331 We realize that this choice (based on trial and error) is rather arbitrary, as currently  
 332 a solid framework of how the refinement criteria should be chosen is still lacking.  
 333 The results are validated by a comparison with runs using a regular and fixed  
 334 grid at the maximum resolution, performed with the Basilisk and MicroHH flow  
 335 solvers: MicroHH is the numerical code used by Van Heerwaarden and Mellado  
 336 (2016) to obtain their results. This code represents a state-of-the-art flow solver  
 337 that is dedicated to studying atmospheric systems (Van Heerwaarden and Mellado  
 338 2016, Shapiro et al. 2016); for a detailed description of the MicroHH code see Van  
 339 Heerwaarden et al. (2017). In addition, the fixed grid results of the Basilisk and  
 340 MicroHH flow solvers are compared to each other.

341 We choose  $Pr = 1$  and  $Re = 3000$  with a domain size of  $3\mathcal{L} \times 3\mathcal{L} \times 3\mathcal{L}$  and  
 342 simulate the evolution of the system until the physical time  $t = 45T$ . In order to  
 343 limit the computational costs, the evolution of the Basilisk-based run with a fixed  
 344 regular grid is only computed until  $t = 10T$ . To illustrate the physical size of such  
 345 a numerical experiment in reality; for a domain size of  $0.5 \text{ m} \times 0.5 \text{ m} \times 0.5 \text{ m}$  and  
 346  $\theta_{\text{ref}} = 21 \text{ }^\circ\text{C}$ , the corresponding parameters are:  $\mathcal{L} = 0.16 \text{ m}$ ,  $\theta_{\text{bottom}} = 36 \text{ }^\circ\text{C}$  and  
 347  $T = 153 \text{ s}$ . This could be interpreted as a modest laboratory experiment.

348 The simulations are performed with Surfsara's supercomputer Cartesius lo-  
 349 cated in Amsterdam, The Netherlands (SURFsara 2017). An overview of the dif-  
 350 ferent runs, including the number of cores used, integration timesteps and total  
 351 run time is listed in Table 1. Additional information on the case set-up for both  
 352 models can be found at:

353 Basilisk:

354 <http://basilisk.fr/sandbox/Antoonvh/freeconv.c><sup>1</sup>

355 MicroHH:

356 <https://github.com/microhh/microhh/tree/master/cases/vanheerwaarden2016>

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<sup>1</sup> From a users' perspective, the case definition for the adaptive-grid runs is (subjectively) not more involved than the fixed-grid counterpart. The more complex adaptation-specific formulations are addressed by a low-level part of the Basilisk toolbox that does not require explicit attention from the users' side.

### 3 Results

#### 3.1 Grid Structure

First, we study the evolution of the solution and grid structure qualitatively. Vertical slices of the magnitude of the gradient of the buoyancy field ( $\|\nabla b\|$ ) and the used grid at  $t = \{2, 10, 20\}T$  for run BA-0.0025 are presented in Fig. 7. At  $t = 2T$  a complex grid structure is generated by the AMR algorithm, and within the ABL, the grid is refined at locations where vigorous turbulent structures are present. Above the ABL (i.e.  $z/\mathcal{L} > 1$ ), turbulence is absent and the grid is coarse. Both effects are appealing from a physical perspective as the computations are focused on the regions where the activity is present. As the physical time progresses, the boundary layer becomes more neutrally stratified and the turbulence intensity decreases. And again, in response, the adaptive-grid algorithm has coarsened the grid at  $t = 10T$ . This remarkable effect is even more pronounced at  $t = 20T$ , where the coarsened regions have grown in size, indicating that the number of grid cells is decreasing over time. Physically speaking, this is facilitated by the fact that the size of the smallest eddies increases as turbulence decays.

#### 3.2 Validation

Next we compare the results obtained with the AMR and fixed-uniform-grid runs. Following Van Heerwaarden and Mellado (2016), we compare the domain integrated quantities: a boundary-layer height  $z_i$  that is based on the buoyancy profile, kinetic energy  $I_e$ , buoyancy flux  $I_b$  and dissipation  $I_\epsilon$  according to,

$$z_i = \frac{2}{N^2} \int_0^\infty (\langle b \rangle - N^2 z) dz, \quad (11)$$

$$I_\alpha = \int_0^\infty \langle \alpha \rangle dz, \quad (12)$$

where  $\alpha$  is a dummy variable for  $\{e, b, \epsilon\}$  and  $\langle \alpha \rangle$  denotes the horizontally-averaged value of the quantity  $\alpha$ . Figure 8a shows the evolution of the boundary-layer height, where good agreement between all simulations is found. The boundary-layer height is an integral measure for the amount of buoyancy (i.e. analogous to heat) in the system, though due to the case set-up, this integral quantity is not a very sensitive measure to assess the accuracy of the resolved turbulent motions. Therefore, we focus on higher-order statistics. In general, the evolution of the total kinetic energy shows similar behaviour between all runs (see Fig. 8). Nevertheless small discrepancies on the order of 5% are present, particularly between the runs with the adaptive grid and the fixed uniform grids, and as expected, this discrepancy decreases when the refinement criterion is more strict. In order to analyze the evolution of kinetic energy in further detail, Fig. 8c presents the evolution of the domain-integrated buoyancy flux, which represents the energy-production rate for this system. The buoyancy flux agrees well for all different runs and the observed differences between the runs are a result of turbulent fluctuations within the chaotic system rather than systematic discrepancies. This indicates that the overall structure and characteristics of the energy-producing motions are resolved accurately for all

395 runs, and for free convection, these motions are associated with the large thermal  
396 plumes. In order to assess the representation of the small-scale structures in these  
397 simulations, Fig. 8d presents the evolution of the resolved energy-dissipation rate.  
398 Compared to the fixed-grid runs, the AMR-based runs slightly underestimate the  
399 resolved absolute dissipation, an aspect that is present throughout the simulation.  
400 Again, the discrepancy appears to be controlled by the refinement criterion, for  
401 which using stricter (i.e. smaller) criteria, the results seem to converge towards  
402 the values found with the fixed-grid runs. The fact that the runs diagnosed with a  
403 lower dissipation rate are also associated with lower kinetic energy indicates that  
404 a small part of the dissipation has a numerical/non-physical origin.

405 Figure 9 shows the vertical profiles of the kinetic energy at  $t/T = \{2, 4, 25\}$ , and  
406 shows discrepancies at  $t/T = 2$  between all runs. The highly chaotic flow structure  
407 at this early stage of the simulation could explain some of the differences. However,  
408 consistent with Fig. 8b, the adaptive-grid runs show a systematically lower kinetic  
409 energy content over the entire domain. At  $t/T = 4$ , the profiles of the fixed-grid  
410 runs agree well, and furthermore, the energy found in the adaptive-grid run BA-  
411 0.0025 also compares well. It can be seen from the time series in Fig. 8b that for  
412  $t/T < 5$ , the evolution of kinetic energy shows large fluctuations. Therefore, we also  
413 compare the energy profiles at  $t/T = 25$ , where we see again that the fixed-grid run  
414 still contains more energy than the adaptive-grid runs. Again, the adaptive run  
415 with the smallest refinement criterion is closest to the fixed-grid profile compared  
416 to the other adaptive-grid runs.

417 Although it appears that the adaptive-grid algorithm is able to refine the grid  
418 at locations of the turbulent structures, discrepancies in the simulations results  
419 remain present. This can be explained by the fact that the process of refining and  
420 coarsening the mesh relies on a linear interpolation strategy for defining values  
421 on new grid cells. This interpolation introduces additional errors compared to a  
422 simulation that employs a static grid, and these errors are similar to the truncation  
423 errors of fixed grid advection schemes and thus lead to similar additional numerical  
424 dissipation of energy. More accurate interpolation techniques could be tested to  
425 limit the error due to interpolation. Therefore, this relevant aspect will be studied  
426 in more detail in the future.

### 427 3.3 Performance

428 As discussed in the introduction, for highly dynamic flow configurations such as a  
429 diurnal cycle, model performance may benefit from the AMR approach. Although  
430 the present case of decaying convection is less dynamic than a full diurnal cycle,  
431 it is tempting to compare the simulation performance of the AMR-based run to  
432 its counterparts using a fixed and regular grid. Thereupon, several performance  
433 characteristics are presented in Fig. 10. Figure 10a shows, for the AMR-based  
434 runs, the evolution of the number of grid cells, that appear to be controlled by the  
435 refinement criterion, in which a smaller value causes the algorithm to use a more  
436 refined grid. As illustrated in the snapshots of Fig. 7, the number of grid cells varies  
437 significantly over the course of the simulation. Supposedly, the computational re-  
438 sources are distributed more efficiently over time. Furthermore, even in the run  
439 with the most strict refinement criterion, the number of grid cells does not exceed  
440 21% of the maximum-resolution value. Figure 10b shows how the computational

441 speed (i.e. defined here as wall clock time per integration timestep) is correlated  
 442 with the number of grid cells. It appears that there are several regimes in how the  
 443 performance is affected by the number of grid cells. For a large number of grid  
 444 cells (i.e.  $> 10^6$ ) the amount of integration timesteps per second increases with a  
 445 decreasing number of grid cells, indicating that the solver does indeed speed up  
 446 when the grid is coarsened. Note that the simulations apply many grid cells in the  
 447 early stage of the runs (i.e. at the right-hand side of Fig. 10b and uses fewer cells  
 448 as time progresses (towards the left-hand side of Fig. 10b). However, as denoted  
 449 by the  $x^{0.6}$ -scaling line, in this regime the simulation speed is not linearly depen-  
 450 dent on the amount of grid cells. Furthermore, for lower number of grid points  
 451 (i.e.  $< 10^6$ ) the simulation speed appears to slow down when the simulation runs  
 452 with fewer grid cells, i.e. there is a performance penalty for coarser grids! Possible  
 453 causes for these performance characteristics are listed below:

- 454 1. For this case, the grid structure of the coarsened grids at later stages in the  
 455 simulation contains a relatively larger fraction of resolution boundaries (see Fig.  
 456 7). These boundaries are associated with additional overhead as they require  
 457 special attention by the solver (see Sect. 2.1).
- 458 2. The number of used processors (linked to domain decomposition for paralleliza-  
 459 tion) is fixed throughout the simulations. Therefore, the relative overhead of  
 460 MPI-domain communication routines compared to actual calculations increases  
 461 as the number of grid cells decreases.
- 462 3. For coarse grids, the physical timestep taken per integration timestep increases  
 463 (Courant-Friedrichs-Lewy criterion). Since diagnostic analysis of the solution  
 464 is performed with a regular interval in the physical time, i.e.  $\Delta t = T$  for profiles  
 465 and slices and  $\Delta t = T/20$  for the domain-integrated quantities. The frequency  
 466 of calls to diagnostic routines increases (i.e. say, calls per 100 integration steps)  
 467 on average resulting in an increased effort per integration step.

468 In Fig. 10c the amount of system billing units (i.e. the used *number of cores* $\times$ *hours*)  
 469 spending for the different runs is presented. Before an interpretation of the results  
 470 is made, it is important to realize that the performance of a simulation run is a  
 471 function of many aspects that ranges from the details of the hardware configuration  
 472 to the exact case set-up. Therefore, the results presented here are intended as  
 473 an illustration rather than as absolute values. Nevertheless, it is clear that the  
 474 MicroHH run is notably cheaper compared to the runs performed with the Basilisk  
 475 solver. This can be explained by the different numerical schemes that are employed.  
 476 Most notably, for obtaining the pressure field, the Basilisk code uses a multigrid  
 477 strategy for solving the corresponding Poisson equation whereas the isotropic-fixed  
 478 grid in MicroHH facilitates the usage of a spectral Poisson solver. Although the  
 479 spectral method requires more MPI communication for parallelization when using  
 480 a large number of processors, it is known to be more efficient (Fornberg 1998). If  
 481 we compare the adaptive and non-adaptive simulation runs performed with the  
 482 Basilisk solver, we do see a considerable decrease in costs for the adaptive method  
 483 runs.

484 In Fig. 10d the memory used for the different simulation runs is presented.  
 485 Compared to the fixed-grid runs, the adaptive-grid runs require less memory. This  
 486 is due to the fact that the maximum number of grid cells is considerably lower  
 487 than the number of grid cells in the fixed-grid runs (see Fig. 10a). From this per-  
 488 spective, the adaptive-grid approach can also be attractive for applications where

489 the available memory is limited. However, even though the run with MicroHH  
490 employs many more grid cells, the required memory is comparable to that of run  
491 BA-0.0025, meaning that per grid cell, the MicroHH code is more efficient in terms  
492 of memory.

#### 493 **4 Outlook: Towards Adaptive Mesh Refinement in Atmospheric LES**

494 We have based our test and performance benchmark on an idealized flow config-  
495 uration of a CBL using DNS, providing a ground truth for our intercomparison.  
496 In the future, we plan to study more practically-oriented cases by using an LES  
497 formulation. For many atmospheric cases, LES is preferred over DNS, because it  
498 provides an efficient tool for studying high-Reynolds-number flows. Therefore, the  
499 next step is to test the AMR approach in combination with an LES formulation.  
500 In this section, we briefly discuss some preliminary results on this topic. Because  
501 this is part of ongoing research, we do not perform a quantitative discussion of the  
502 test runs, the results and performance characteristics. The presented results aim  
503 to exemplify the AMR method for a different case and show the flexibility of the  
504 AMR approach. The example is based on the LES intercomparison study case by  
505 Bretherton et al. (1999), in which a boundary layer is filled with a smoke cloud  
506 that cools from the top due to longwave emission. The boundary layer is initially  
507 capped by a strong temperature inversion (i.e. 7 K over 50 m) at  $z \approx 700$  m and  
508 rises over the course of the simulation due to entrainment. The inversion layer  
509 is identified as a region where turbulent length scales are suppressed and turbu-  
510 lent motions are anisotropic due to the stable stratification. As such, this region  
511 requires a high resolution to capture the predominant turbulent structures accu-  
512 rately. In contrast, the convective turbulence in the boundary layer itself can be  
513 captured on a relatively coarse grid (Sullivan and Patton 2011). Accordingly, we  
514 decide not to base the grid adaptation upon the estimated discretization error in  
515 the representation of the velocity-component fields, but only on the estimated er-  
516 ror in the smoke-cloud fraction and temperature fields. With such an approach the  
517 AMR algorithm does not refine the mesh in order to resolve the small turbulent  
518 structures in the near-neutral boundary layer, but allows the LES to employ the  
519 SGS model effectively in this region. In this run, the numerical grid varies by three  
520 levels of refinement, i.e. between 25 m and 3.125 m. Figure 11 presents snapshots  
521 of the temperature and numerical grid taken at  $t = 3$  h after initialization. It is  
522 clear from Fig. 11a that an inversion layer is present, while Fig. 11b shows that  
523 the numerical grid has a high resolution in the region of the inversion layer and  
524 remains coarse in the boundary layer itself. Furthermore, we see the subsiding  
525 shells in the boundary layer that are qualitatively similar to those observed in the  
526 laboratory experiment performed by Jonker and Jiménez (2014).

527 For this case, the AMR algorithm dynamically adapts to the flow by redirecting  
528 the grid refinement to those regions of the spatial domain where it is required.  
529 Hence in this case, adaptation is predominantly spatially focussed, whereas in the  
530 DNS case the refinement was most prominent in the temporal domain (see Fig.  
531 10a). As such, both examples in this study are complementary and both effects  
532 (spatial and temporal adaptive grid refinement) are expected to play an important  
533 role in future simulations of full diurnal cycles (cf. Fig. 1).

534 Finally, we note the following; the present cases were restricted to spatially  
 535 homogeneous set-ups, where ‘scale separation’ naturally occurs through the inter-  
 536 nal variability of turbulence, originating from the non-linearity of the governing  
 537 equations. In reality, heterogeneity in the *surface boundary conditions* also becomes  
 538 important and provides an additional cause of scale separation that may call for  
 539 adaptive grid refinement. For example, refinement may be preferred at sharp tran-  
 540 sitions between different types of land use, such as land–sea interfaces.

## 541 5 Discussion and Conclusions

542 We have introduced and tested an adaptive mesh refinement (AMR) method for  
 543 studies of the atmospheric boundary layer (ABL). This work is motivated by a  
 544 desire to numerically study highly dynamic cases. Such cases are characterized by  
 545 a high degree of scale separation throughout the spatial and temporal domain.  
 546 This work should be viewed as the first step in our AMR-based research that  
 547 assesses the usage of an AMR method for studies of the ABL. We have based our  
 548 adaptive-grid simulations on the flow solver implemented in the Basilisk code.

549 The method is tested using DNS based on a case introduced by Van Heerwaar-  
 550 den and Mellado (2016), describing the growth and subsequent decay of a CBL.  
 551 The AMR algorithm was able to identify the time-varying turbulent regions in the  
 552 domain and refined/coarsened the grid accordingly. The AMR-based simulations  
 553 can reproduce the simulation results of their fixed grid counterparts with minor  
 554 discrepancies. Furthermore, the AMR algorithm can be tuned to apply more grid  
 555 cells such that these discrepancies are suppressed. For all AMR runs, the number  
 556 of grid cells varies significantly over time, resulting in more efficient simulations  
 557 compared to using a regular fixed grid with identical numerical formulations. This  
 558 provides a proof of principle for the AMR method regarding ABL systems.

559 For this case, a numerical solver dedicated to ABL systems (MicroHH) outper-  
 560 formed all other runs in terms of computational efficiency, indicating that there is  
 561 an overhead associated with the usage of the adaptive solver. In general, the exact  
 562 impact of this overhead depends on the details of the studied case. The most chal-  
 563 lenging ABL systems typically owe their complexity to the dynamical interplay  
 564 between various processes at different length and time scales. Hence, the AMR  
 565 technique is likely to be more favourable as complexity increases. More specifi-  
 566 cally, as discussed in Popinet (2011), the cost of an adaptive simulation, relative  
 567 to a constant resolution simulation ( $G$ ) is expected to scale as

$$G = \frac{C_a \Delta^{-D}}{C_c \Delta^{-3}} = \frac{C_a}{C_c} \Delta^{3-D}, \quad (13)$$

568 where  $C_a$  and  $C_c$  are constants related to the absolute speed of the computation for  
 569 the adaptive- and constant-resolution simulations, respectively ;  $\Delta$  is the ratio of  
 570 the minimum to the maximum scale of the physical system (i.e. a measure of scale  
 571 separation) and  $D$  is the effective (or fractal) dimension of the physical process  
 572 (which is necessarily  $\leq 3$ ). In the present study,  $\Delta$  is relatively large (i.e. of order  
 573  $10^{-2}$ ) and the computational gain using the adaptive method is correspondingly  
 574 small, whereas for challenging cases  $\Delta$  can be several orders of magnitude smaller,  
 575 with a correspondingly larger potential gain in efficiency of the adaptive method

576 relative to constant-resolution methods. This important aspect of the overall scal-  
 577 ing behaviour is illustrated in Appendix 1 for a canonical flow set-up. The results  
 578 shown herein thus motivate our continued research using the AMR technique.

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## 584 Appendix 1 The Lid-Driven Cavity in Two Dimensions

585 We study the relation between the computational costs and the scale separation  
 586 for a simulation of a fluid in a lid-driven cavity in two dimensions, and compare the  
 587 results from a regular fixed grid and the adaptive-grid-refinement approach as is  
 588 presented herein. The chosen physical set-up consists of a no-slip box (size  $L \times L$ ),  
 589 in which an incompressible fluid with kinematic viscosity  $\nu$  is set in motion by the  
 590 top lid that moves with a constant velocity ( $U_{lid}$ ) in the left-to-right direction. It  
 591 is well known that this configuration results in a large circulation cell within the  
 592 domain. With system parameters  $L, \nu$  and  $U_{lid}$  we can identify a Reynolds number  
 593 ( $Re_{lid}$ ) according to

$$Re_{lid} = \frac{U_{lid}L}{\nu}. \quad (14)$$

594 In order to study the effect of varying scale separation on the performance statis-  
 595 tics, the simulations cover a range of different Reynolds numbers. Following the  
 596 analysis of Clercx and Van Heijst (2017) on vortex-wall interactions in two dimen-  
 597 sions, we take the (minimum) grid-box size inversely proportional to the Reynolds  
 598 number. As such, the Reynolds number represents the separation of scales in our  
 599 simulations (i.e.  $\Delta$  in Eq. 13). The runs are initialized with the fluid at rest and the  
 600 flow evolution is simulated for a physical time  $t_{end} = 20L/U_{lid}$ . For the adaptive  
 601 grid simulations, a refinement criterion for the velocity components  $\zeta = 0.005U_{lid}$   
 602 is chosen. All runs are performed using a single processor core. A snapshot of the  
 603 vorticity field and the corresponding grid structure at  $t = t_{end}$  for  $Re_{lid} = 500$  are  
 604 presented in Fig. 12. The maximum resolution of this simulation corresponds to a  
 605  $256 \times 256$  grid. First, the solution is validated against the results obtained with the  
 606 fixed equidistant-grid runs in Fig. 13. Here the vorticity fields ( $\omega(x, y)$ ) obtained  
 607 from the fixed-grid and adaptive-grid simulations are directly compared for the  
 608 runs with  $Re_{lid} = \{250, 500\}$ . We conclude that the chosen refinement criterion is  
 609 sufficiently small to accurately reproduce the results obtained with the equidis-  
 610 tant grid. Second, Fig. 14 presents the scaling of the computational costs with the  
 611 Reynolds number. The simulation costs when employing the fixed and equidistant  
 612 grid appear to scale with the third power of the Reynolds number. This exponent  
 613 can be understood from the fact that the total number of grid cells scales with the  
 614 Reynolds number to the second power (i.e. in 2D, doubling the resolution requires  
 615 four times as many grid cells). Furthermore, the well-known numerical-stability



616 criterion of Courant-Friedrichs-Lewy limits the timestep and scales inversely  
617 proportional to the grid-box size, meaning that the total number of timesteps is pro-  
618 portional to the Reynolds number. Combined, the computational costs scale with  
619 the Reynolds number to the power of (2 space+1 time =)3. This analysis holds for  
620 all equidistant-grid approaches, and as such, we can anticipate the computational  
621 costs when using an equidistant-grid code that is an order of magnitude faster  
622 than the solver we have chosen for our fixed-grid approach (i.e. the Basilisk solver  
623 running in fixed-grid mode). Interestingly, with an increasing Reynolds number,  
624 the observed scaling of the adaptive grid simulations is favourable compared to the  
625 equidistant grid counterpart. The observed scaling (i.e.  $\propto Re^{1.9}$ ) reflects that the  
626 resolution requirement is not space filling. Although that for the lower Reynolds  
627 numbers (i.e.  $Re \lesssim 1000$ ), the (theorized) fast equidistant-grid solver is more ef-  
628 ficient than the adaptive grid approach, there exists a crossing point where the  
629 latter technique becomes the more effective option. This feature is indicative to all  
630 processes in nature. in which, with an increasing scale separation, the space-filling  
631 approach of an equidistant grid represents the worst-case scenario, neglecting the  
632 so-called fractal dimension of the problem. Note that this concept also applies to  
633 three-dimensional turbulence (see Chap. 8 in Frisch, 1995). However, the corre-  
634 sponding scaling behaviour for ABL cases is not obvious.

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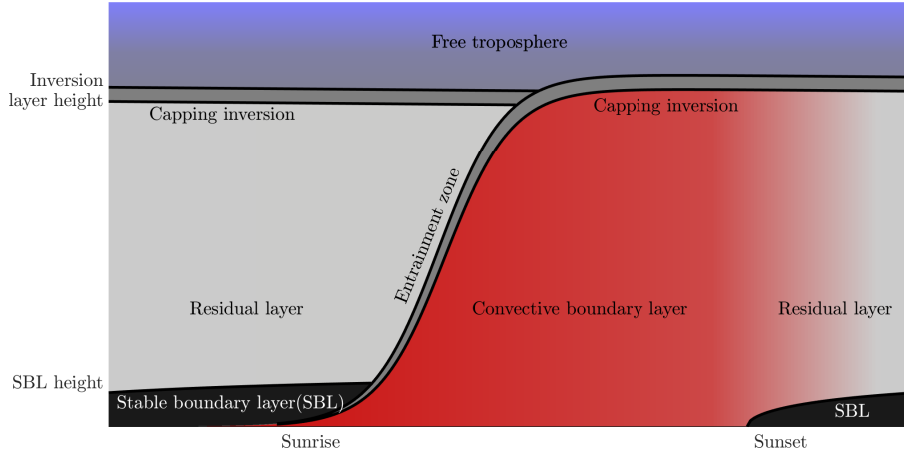
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**Table 1** Overview of the different simulation run details. In the top section a reference name, the used solver, grid type, the (maximal) numerical grid resolution, lateral boundary conditions and refinement criterion ( $\xi_b$ , if applicable) are listed for each run. In the bottom section the used number of cores, the total amount of integration steps taken at  $t/T = \{10, 45\}$  and the total wall clock time of each run are presented.

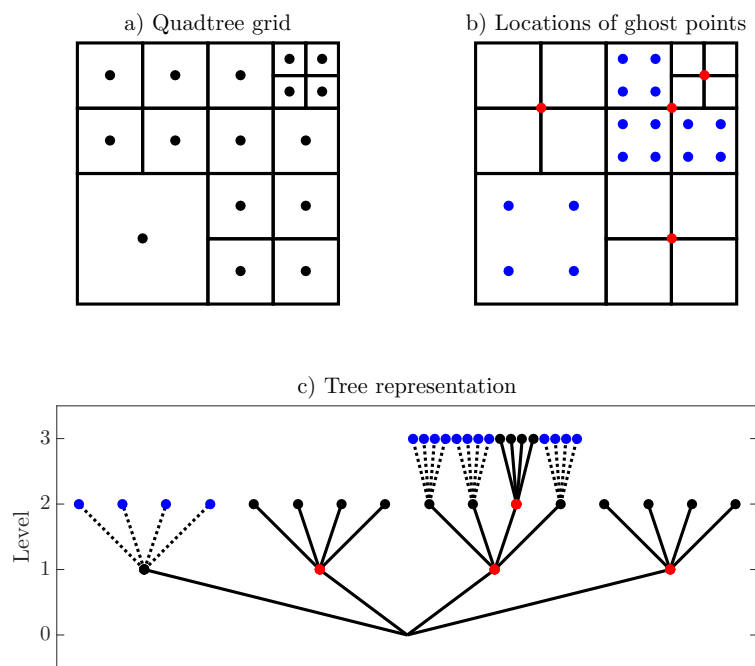
Run name	Code	Grid	$n_x \times n_y \times n_z$ (Maximal)	Lateral BCs	$\xi_b$
MicroHH	MicroHH	Fixed & stretched	$512^2 \times 387$	Periodic	-
BA-512 <sup>3</sup>	Basilisk	Fixed	$512^3$	Neumann & No-pen.	-
BA-0.0025	Basilisk	AMR	$512^3$	Neumann & No-pen.	0.0025
BA-0.005	Basilisk	AMR	$512^3$	Neumann & No-pen.	0.005
BA-0.01	Basilisk	AMR	$512^3$	Neumann & No-pen.	0.01

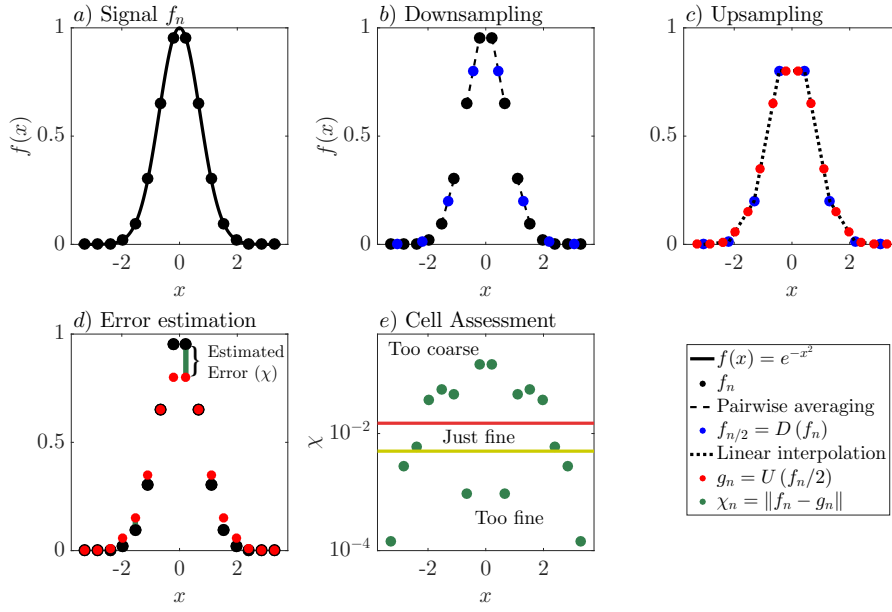
Run name	Number of cores	Integration steps at $t/T = \{10, 45\}$	total wall clock time (D:HH:MM)
MicroHH	64	{13920, 35670}	0:12:22
BA-512 <sup>3</sup>	64	{14073, (35670)} ( <i>estimated</i> )	2:16:12 ( $t/T = 10$ )
BA-0.0025	96	{14095, 30144}	2:10:30
BA-0.005	96	{14061, 28704}	1:18:19
BA-0.01	96	{14167, 25544}	1:02:16



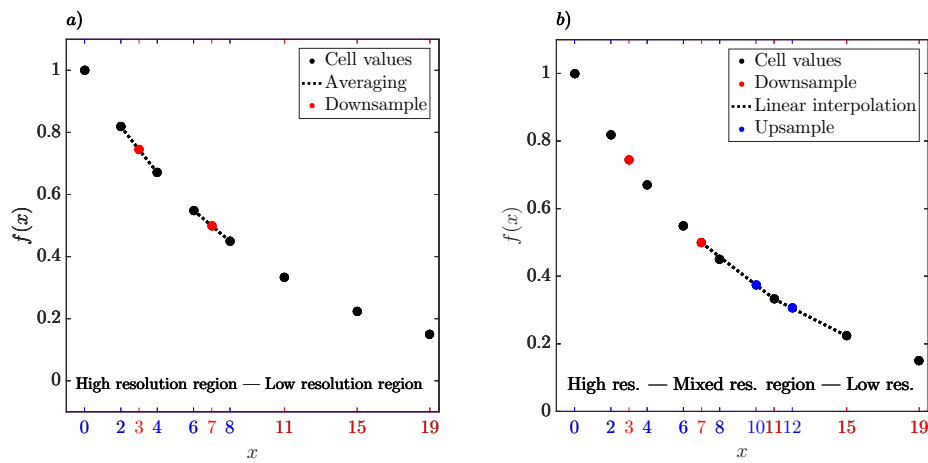
**Fig. 1** Sketch of a prototypical diurnal cycle evolution. Adapted from Stull (1988).



**Fig. 2** Example of a tree-grid structure. The top row presents the spatial structure of the grid cells with varying levels of refinement (a) and the locations of two types of ghost points whose field values are defined by the downsampling (red dots) and upsampling (blue dots) operations (b, see text). The plot on the bottom row presents a corresponding tree representation of the various grid cells and ghost points at different levels (c).

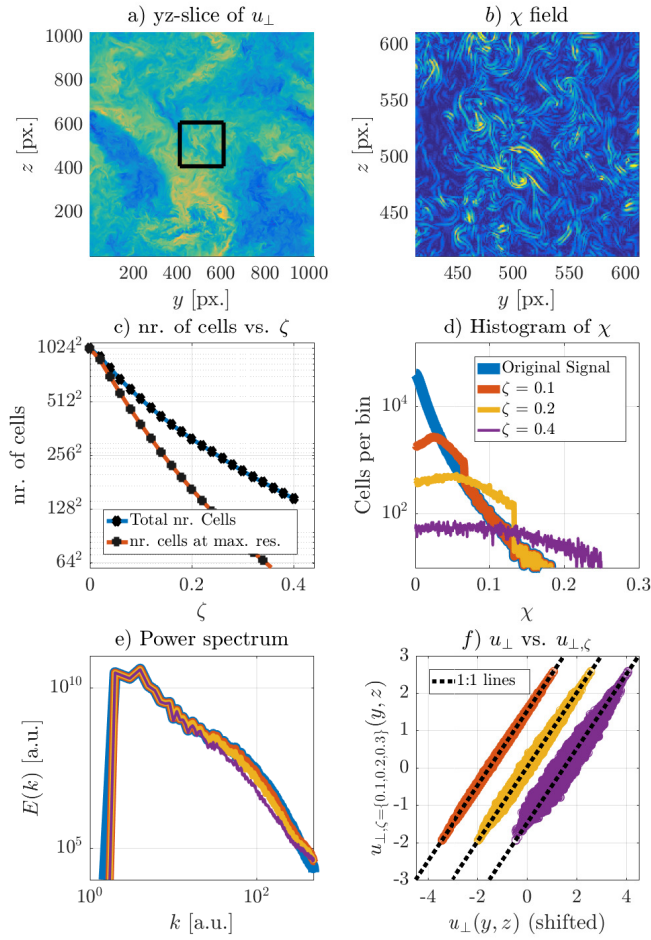


**Fig. 3** A one-dimensional, visual representation of how the adaptation algorithm assesses the discretization of a curved field  $f(x)$ : a) A coarser level estimate of the discretized solution is obtained using the downsampling operation. b) Using these coarse level values, the original discretized solution can be estimated using the upsampling operation. c) The difference between the estimated and original values is interpreted as an error estimator ( $\chi$ ) and can be compared against fixed thresholds (e.g.  $\zeta$ ). d) and e) If the refinement criterion is exceeded, new cells at one level higher are initialized locally by applying a linear interpolation technique using the initial cell values. Alternatively, if the estimated error is smaller than the coarsening criterion for multiple cells, these cells can be merged if that does not violate the general grid-structure requirements (see text and Fig. 2).

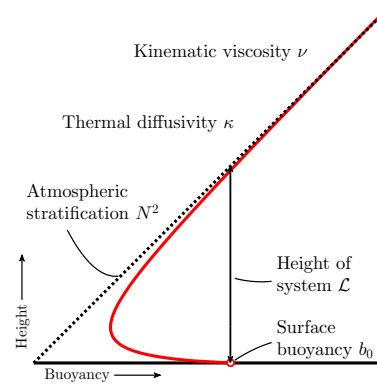


**Fig. 4** Example of the treatment of a resolution boundary in a one-dimensional scenario. First, the high level region near the resolution boundary is downsampled to obtain values for the coarse-level ghost points in red (a). Second, linear interpolation of the coarse level solution is used to define the field values of high level ghost points in blue (b).

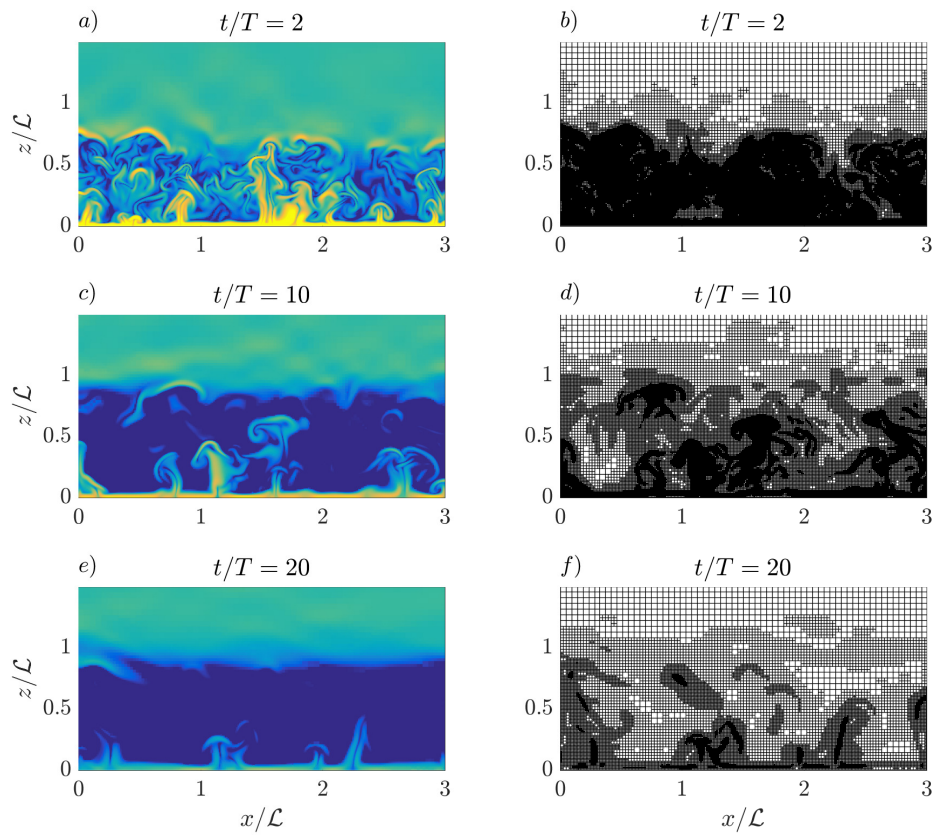




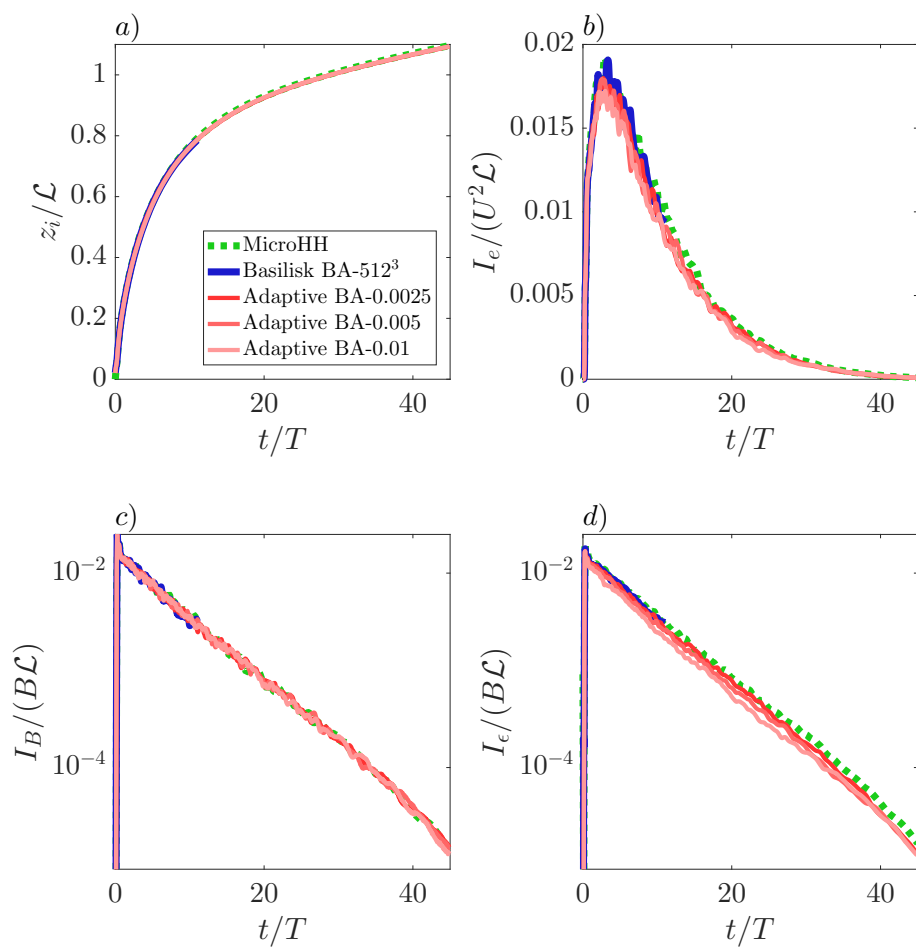
**Fig. 5** Example of the adaption algorithm applied to a (2D) slice of a 3D turbulent field. a) Shows the data slice of the velocity component in the plane-perpendicular direction ( $u_{\perp}$ , obtained from Li et al. (2008)). b) Presents the  $\chi$  field, evaluated using the method described in Sect. 2.4. Only the centre part of the slice, indicated by the black box in (a), is shown to reveal the small scale details in this simulation. c) shows the grid cell number dependence on the chosen refinement criterion ( $\zeta$ ), note the logarithmic vertical axis. A histogram of the  $\chi$  field with 512 bins for the original data, and the data corresponding to three  $\zeta$  values are presented in d). Using the same colour coding as in d), power spectra and a direct comparison of the  $u_{\perp}(y, z)$  field are shown in e) and f), respectively.



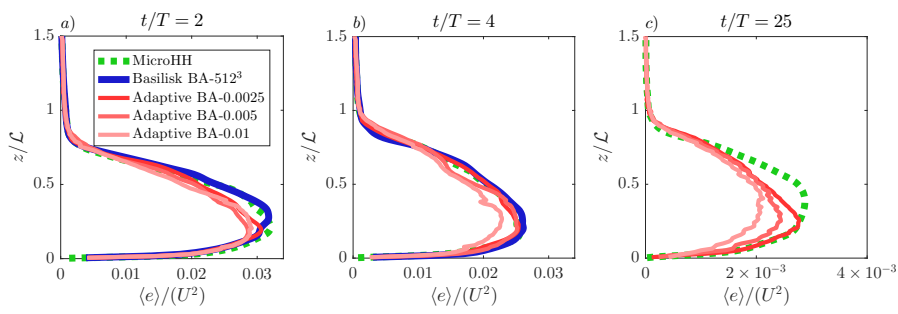
**Fig. 6** Sketch of the system and its parameters. The red line illustrates a typical buoyancy profile within the CBL during the initial development. Adapted from Van Heerwaarden and Mellado (2016).



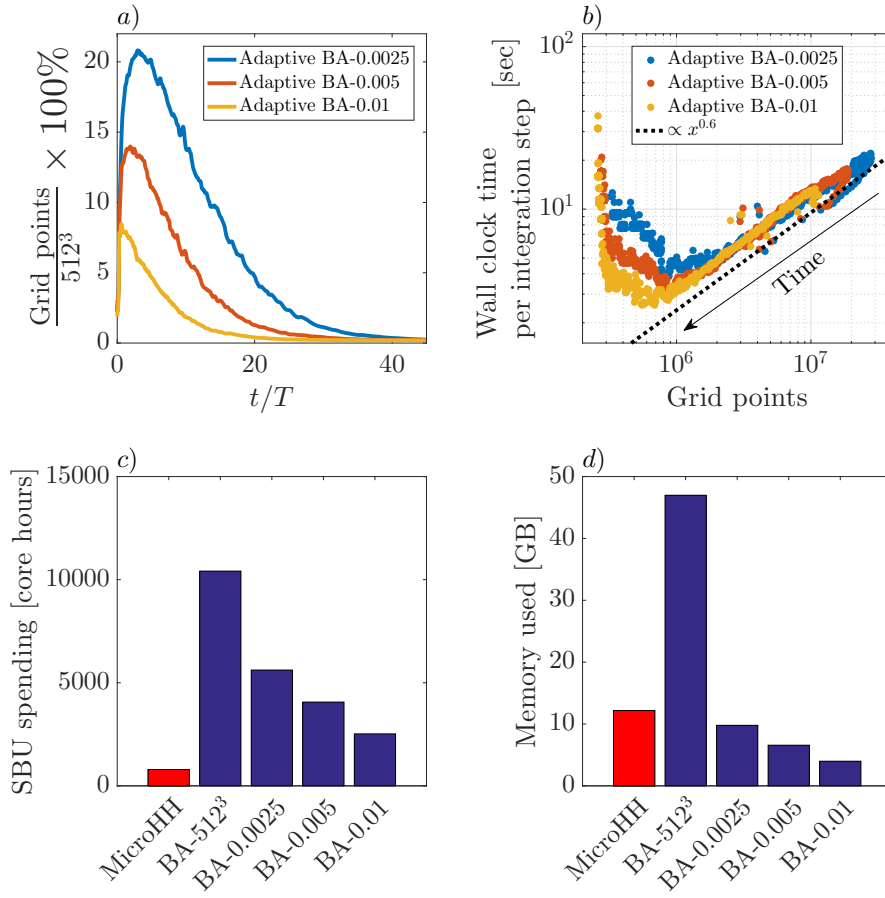
**Fig. 7** Vertical slices of the  $\|\nabla b\|$  field (left column) and the corresponding numerical grid (right column) in the lowest half of the domain. The top, middle and bottom rows represent snapshots taken at  $t/T = \{2, 10, 20\}$ , respectively. These snapshots are taken from the adaptive-grid run BA-0.0025.



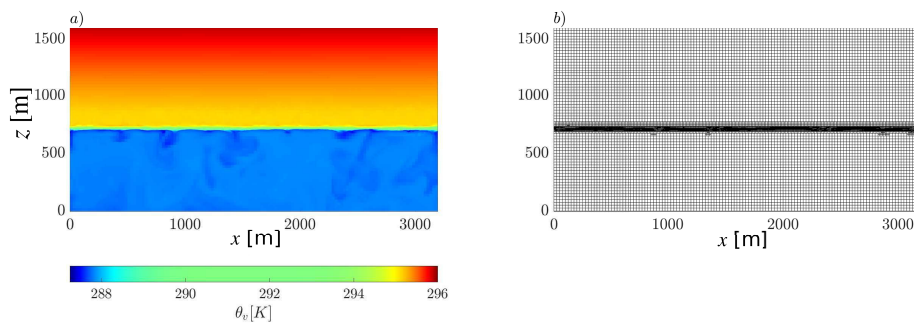
**Fig. 8** Time series of the domain integrated quantities, a) boundary-layer height ( $z_i$ ), b) kinetic energy ( $I_e$ ), c) buoyancy flux ( $I_B$ ) and d) dissipation rate ( $I_\epsilon$ ) according to Eq. 11. The results are obtained with both Basilisk and MicroHH using fixed grids and Basilisk using the adaptive mesh refinement algorithm. Note that plots c) and d) use a logarithmic scale.



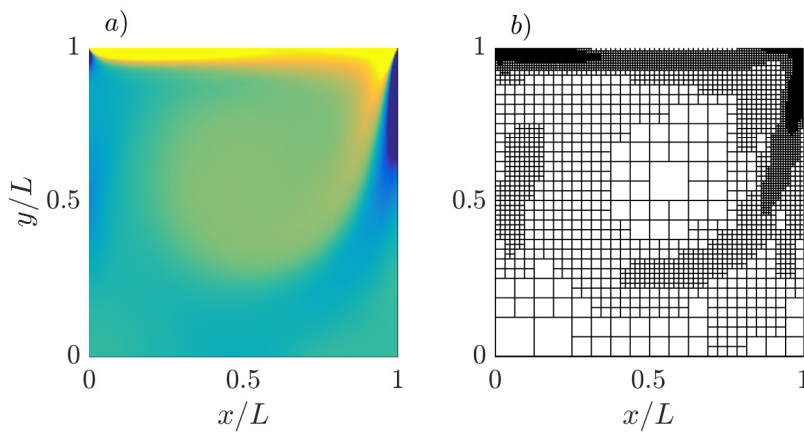
**Fig. 9** Vertical profiles of the horizontally-averaged kinetic energy  $\langle e \rangle$  at  $t/T = \{2, 4, 25\}$  in left, middle and right plot, respectively. The results are obtained with both Basilisk and MicroHH using fixed grids and Basilisk using the adaptive mesh refinement algorithm. Note that in panel c) the horizontal axis is rescaled and that regular-grid computations with Basilisk are not available (see text, Sect. 2.4).



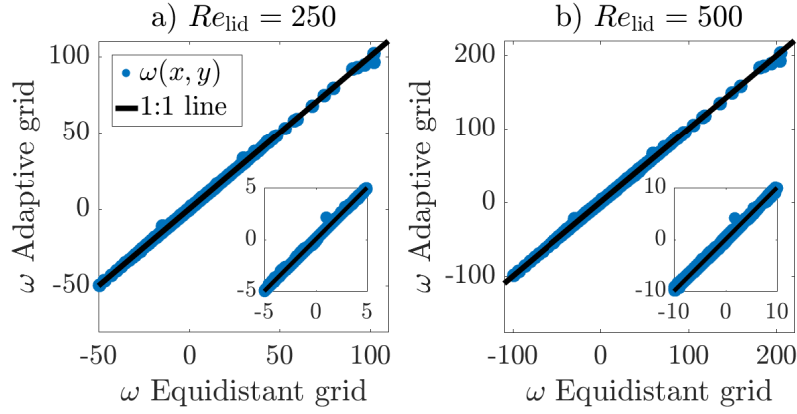
**Fig. 10** Overview of the performance characteristics of the adaptive and fixed-grid simulation runs. a) Time series of the number of grid points for the adaptive runs normalized by the maximum-resolution value (i.e.  $512^3$ ). b) Scatter plot of the wall clock time per integration step versus the used number of grid cells in the adaptive-grid runs. c) The total amount of System Billing Units (SBU, i.e. *number of cores*  $\times$  *hours*) spending on each simulation run. Note that the value for BA-512<sup>3</sup> is estimated as if it were run until  $t/T = 45$ . d) The total RAM memory used in each simulation run in gigabytes (GB).



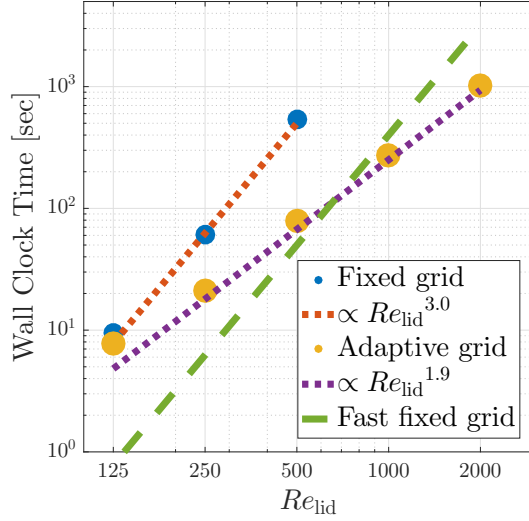
**Fig. 11** Snapshots of a) the vertical slices of the virtual potential temperature field and b) the numerical grid at  $t = 3$  h. The case is based on the work of Bretherton et al. (1999).



**Fig. 12** Snapshots of a) The vorticity field and b) the numerical grid at  $t = t_{end}$  for the lid-driven cavity simulation with  $Re_{lid} = 500$ .



**Fig. 13** Validation of the vorticity field ( $\omega$ ) from the adaptive grid simulation against the results obtained with a fixed equidistant grid. For a)  $Re_{lid} = 250$  and b)  $Re_{lid} = 500$ . The inserts show a zoom-in (i.e. rescaled axes), containing  $\approx 95\%$  of the total number of data points.



**Fig. 14** The correlation of the computational costs and the Reynolds number ( $Re_{lid}$ ) for different approaches. The green line represents the theorized results from a solver that is an order of magnitude faster than the fixed-grid approach that we have used (blue dots).