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Towards Adaptive Grids for Atmospheric Boundary-Layer Simulations

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Abstract We present a proof-of-concept for the adaptive mesh refinement method 9 applied to atmospheric boundary-layer simulations. Such a method may form an 10 attractive alternative to static grids for studies on atmospheric flows that have a 11 12 high degree of scale separation in space and/or time. Examples include the diurnal 13 cycle and a convective boundary layer capped by a strong inversion. For such cases, large-eddy simulations using regular grids often have to rely on a subgrid-scale 14 closure for the most challenging regions in the spatial and/or temporal domain. 15 Here we analyze a flow configuration that describes the growth and subsequent 16 decay of a convective boundary layer using direct numerical simulation (DNS). 17 We validate the obtained results and benchmark the performance of the adaptive 18 solver against two runs using fixed regular grids. It appears that the adaptive-mesh 19 algorithm is able to coarsen and refine the grid dynamically whilst maintaining 20 an accurate solution. In particular, during the initial growth of the convective 21 boundary layer a high resolution is required compared to the subsequent stage 22 of decaying turbulence. More specifically, the number of grid cells varies by two 23 orders of magnitude over the course of the simulation. For this specific (DNS) case, 24 the adaptive solver was not yet more efficient than the more traditional solver that 25 is dedicated to these types of flows. However, the overall analysis shows that the 26 method has a clear potential for numerical investigations of the most challenging 27

²⁸ atmospheric cases.

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 $_{30}$ $\,$ numerical simulations \cdot Large-eddy simulations \cdot Turbulence

31 1 Introduction

The aim of the present work is to introduce adaptive mesh refinement (AMR) as an 32 33 efficient tool for numerical investigations of the atmospheric boundary layer (ABL) using turbulence resolving methods. This refers typically to models that rely on 34 direct numerical simulation (DNS) or large-eddy simulation (LES) techniques. In 35 36 general, AMR solvers aim to distribute the available computational resources efficiently over a domain by dynamically refining and coarsening the computational 37 grid in space and time. AMR techniques have successfully been employed in stud-38 ies concerning flows with a high degree of scale separation throughout the spatial 39 and/or temporal domain. Such studies concern a wide range of topics, e.g. cos-40 mological hydrodynamics (Teyssier 2002), electro hydrodynamics (López-Herrera 41 et al. 2011), multiphase flows (Fuster et al. 2009), flows in complex geometries 42 (Popinet 2003) and turbulence simulations (Schneider and Vasilyev 2010). How-43 ever, to our knowledge, the potential of this technique has not yet been explored 44 for ABL research, and here we aim to do so through an investigation of the con-45 secutive growth and decay of a convective boundary-layer (CBL) system. The flow 46 configuration is modelled after Van Heerwaarden and Mellado (2016) who per-47 formed an in-depth study of this case using a regular grid configuration. As such, 48 the AMR method is tested and benchmarked. 49 Several methods that meet a varying resolution requirement throughout the 50 spatial domain have already been successfully applied in studies on ABL turbu-51 lence. For example, stretching and squeezing of grids (see e.g. Heus et al. 2011, 52 Van Heerwaarden and Mellado 2016, De Roode et al. 2016), nested grids (see e.g. 53 Sullivan et al. 1996, 1998, Moeng et al. 2007, Mirocha et al. 2013, Muñoz-Esparza 54 et al. 2014) and the usage of unstructured anisotropic grids. However, the mesh is 55 always kept fixed during the simulation, whereas dynamical changes in the ABL 56 call for variation of resolution in time. Furthermore, the aforementioned methods 57 of refinement need to be predefined. Consequently, detailed a priori knowledge 58 is needed on the varying resolution requirement throughout the spatial domain. 59 Apart from tailored and well-known cases, this knowledge is usually not avail-60 able beforehand; therefore, we identify three favourable characteristics of an AMR 61

approach for ABL studies. First, the resolution can vary throughout the spatial
domain. Second, the grid can vary in time such that temporal variation in the
local resolution requirement can be met. Third, the grid is generated adaptively
based on the evolution of the numerical solution itself, relaxing the requirement

 $_{\rm 66}$ $\,$ of detailed a priori knowledge on the resolution requirement.

To illustrate our philosophy, we briefly discuss a textbook example of the evolution of the ABL during a diurnal cycle (after Stull 1988). Figure 1 depicts a typical evolution of the ABL during a diurnal cycle. Around sunrise the solar irradiation of the Earth's surface causes a thermal instability that results in the rapid growth of a CBL. The typical size of the largest thermal plumes scales with the boundary-layer height and hence there is a temporal dependency on the resolution requirement to

⁷² resolve these turbulent structures. The growth of the boundary layer slows down

⁷⁴ when the rising thermals reach the inversion layer, which effectively caps turbu-

 $\mathbf{2}$

lent structures at the top of the CBL. The dynamics within an inversion layer are
of pivotal importance for the evolution of the CBL (Garcia and Mellado 2014).

 π Apart from the effective 'lid' on the boundary layer, entrainment processes occur

⁷⁸ here and the formation of stratocumulus clouds is promoted by the large jump in

⁷⁹ temperature with height. Due to the presence of strong stable stratification, tur-

⁷⁵ temperature with height. Due to the presence of strong stable stratification, tur-⁸⁰ bulent length scales are suppressed (De Lozar and Mellado 2015), and in order to ⁸¹ resolve the most prominent turbulent structures here, a much higher resolution is ⁸² necessary compared to the bulk of the CBL (Sullivan and Patton 2011, De Lozar ⁸³ and Mellado 2015). Applying such high resolution everywhere in the domain is ⁸⁴ not feasible given the current status of computational resources, and might not

⁸⁵ be feasible in coming years (Bou-Zeid 2015). For this reason, many LES studies
⁸⁶ have to rely on their subgrid-scale (SGS) parametrizations within the region of
⁸⁷ the inversion layer, partially negating the purpose of a turbulence resolving study.
⁸⁸ Furthermore, the exact height and strength of the inversion layer are not always
⁸⁹ known a priori (except in cases that have been studied before). Fixed nested grids

(Sullivan et al. 1998) are thus not always flexible enough to capture the dynamics
 properly. On the other hand, practically speaking, it should be noted that LES
 results between various studies often tend to converge, signifying that SGS models

have appreciable skill in describing certain characteristics of the inversion layer
(see e.g Nieuwstadt et al. 1993, Siebesma et al. 2003).

At the approach of sunset, thermal plumes gradually decay into so-called resid-95 ual turbulence, and due to the radiative cooling of the Earth's surface, stable 96 stratification sets in and turbulence is now driven by wind shear only. The stable 97 boundary layer (SBL) is typically much shallower than the CBL and, furthermore, 98 the length scales of the turbulent structures that account for the mixing of heat 99 and momentum within this layer are only a fraction of the size of those associated 100 with daytime convective turbulence (Basu et al. 2008). Additionally, Ansorge and 101 Mellado (2016) argue that the resolution requirement for their simulations of the 102 intermittently turbulent SBL is dictated by localized dissipative flow structures 103 that only encompass a fraction of the computational domain. 104

Rather than capturing the cyclic behaviour of the atmosphere as depicted in 105 Fig. 1, the contrast between daytime and night-time turbulence has resulted in 106 many numerical studies focusing only on either convective or stable conditions. 107 The studies that do simulate a diurnal cycle typically struggle to resolve turbu-108 lence during the night (Kumar et al. 2006, Basu et al. 2008, Abkar et al. 2015). 109 Furthermore, the transition period itself (i.e. around sunset) would benefit from 110 high fidelity numerical studies (Lothon et al. 2014). In summary: the example 111 shows that the intrinsic dynamic character of the ABL calls for flexible techniques 112 such as an AMR appoach in addition to existing techniques that have successfully 113 been applied to studies on idealized, steady cases. 114

Apart from our long-term prospects, we focus here on a case corresponding to 115 the red and grey sections in Fig. 1. This choice is motivated by the fact that as a 116 first step, we would like to present a proof-of-concept of the AMR approach before 117 we redirect our attention towards more challenging cases. Therefore, we present 118 results obtained with DNS, for which all turbulent structures are resolved explicitly 119 down to the small-scale Kolmogorov length (i.e. the viscous length scale) according 120 to the Navier-Stokes equations, without any closure for turbulence. Compared to, 121 for example, LES, the results obtained with DNS should be independent of the 122

¹²³ numerical formulations or choice of any SGS model, whereas with LES this is a

¹²⁴ topic of discussion (Bretherton et al. 1999, Siebesma et al. 2003, Fedorovich et al.

¹²⁵ 2004, Beare et al.2006, De Roode et al. 2017). However, as shown in Sect. 4, the

¹²⁶ concept of the AMR approach can be easily extended to LES. Since this technique

¹²⁷ is a popular choice for studies on the ABL, we also briefly discuss results obtained

¹²⁸ with the AMR technique using a LES formulation.

We realize that we cannot address all questions regarding the AMR technique in relation to ABL simulations. For example, here we focus on a single case whereas we will argue that the performance of an AMR solver varies depending on the particular case specifications (see Appendix 1). Furthermore, we choose a numerical solver called Basilisk (http://basilisk.fr) for the adaptive-grid runs and do not assess alternatives. The paper is organized as follows; in Sect. 2.1 the details of the adaptive-grid

solver are described, focusing on the AMR algorithm, and in addition, Sect. 2.2
provides an example analysis of how the algorithm assesses a turbulent signal and
adapts the grid accordingly. In Sect. 2.3 the case and the numerical set-up of the
different runs are specified. Section 3 presents the obtained results including a
performance assessment, while in Sect. 4 we provide an outlook on future plans.
We finish with a conclusion combined with a discussion in Sect. 5. Additionally,
using a simple flow set-up, Appendix 1 illustrates an important advantage the

¹⁴³ AMR technique has over a fixed equidistant-grid approach.

144 2 Methods

¹⁴⁵ 2.1 Basilisk and the Grid Adaptation Algorithm

¹⁴⁶ The AMR runs are performed with the partial-differential-equation solver called

Basilisk, a code that contains a second-order accurate finite-volume solver for the
 Navier-Stokes equations. For a detailed description of the numerical formulations

see Popinet (2003,2009), Lagrée et al. (2011), and references therein.

In order to facilitate local adaptive refinement and coarsening whilst maintain-

ing a Cartesian grid-structure, a so-called tree-based grid is used. To illustrate this 151 mesh structure, Fig. 2 shows the two-dimensional (2D) variant of a tree-based grid 152 (i.e. a quadtree), whose structure introduces a hierarchy between cells at integer 153 levels of refinement. The resolution between the levels of refinement differs by a 154 factor of two and the Basilisk solver allows neighbouring cells to vary up to one 155 level. The formulations of numerical methods (e.g. evaluating spatial derivatives) 156 on equidistant Cartesian grids are relatively straightforward compared to their 157 uneven grid counterparts. Therefore, ghost points are defined, enabling simple 158 Cartesian stencil operations for the cells in the vicinity of a resolution boundary. 159 These points act as virtual cells and are located such that all cells have neighbours 160 that are defined at the same level of refinement, see Fig. 2b. The field values on 161 these ghost cells are defined with interpolation techniques using the original field 162

163 values.

The tree grid facilitates an efficient and convenient structure to perform a multiresolution analysis of a discretized field. During the simulation, such an analysis is used to determine which grid cells require refinement and where in the domain cells can be coarsened. This procedure is discussed next. Consider a 1D signal (f)discretized with an even number (n) of elements f_n , where individual entries of f_n are indexed with *i* such that f_n^i represents the *i*-th entry of f_n . First, we define a downsampling operation (D) that approximates f_n on a coarser level grid with n/2 elements,

$$f_{n/2} = D(f_n). \tag{1}$$

¹⁷² Second, we define an upsampling operator (U) that samples $f_{n/2}$ to a signal that ¹⁷³ is defined with the same element entries as the original signal f_n ,

$$g_n = U(f_{n/2}),\tag{2}$$

noting that in general $f_n \neq g_n$, and the absolute difference χ , defined as,

$$\chi_n^i = \|f_n^i - g_n^i\|,\tag{3}$$

can be interpreted as an estimation of the discretization error. The downsampling 175 operation in the Basilisk solver is defined as local volume averaging of the signal 176 to obtain a value for a corresponding coarser-level grid cell (see Fig. 3 a.). This 177 formulation is exact since in a finite-volume formulation, the grid cell values rep-178 resent volume-averaged quantities. To be in line with the second-order accuracy 179 of the solver, the upsampling operation is chosen to be second-order accurate as 180 well, and entails performing a linear interpolation between the grid points of the 181 coarse level solution (see Fig. 3b). Once these two operations have been applied 182 to the discretized signal, it is possible to evaluate χ_n^i for each of the grid cells. 183 Given an error threshold ζ , the following assessment with regards to a grid-cell's 184 resolution can be made, 185

the *i*-th grid cell is
too fine.
$$\chi_n^i > \zeta$$
,
too fine. $\chi_n^i < \frac{2\zeta}{3}$, (4)
just fine. Otherwise.

The threshold on the estimated error for refinement ζ is called the refinement 186 criterion, with ζ having the same physical units as f. Note that the described 187 method is formally linked to wavelet thresholding that has already been employed 188 for fluid dynamical simulations (Schneider and Vasilyev 2010). The grid can be 189 refined and coarsened according to Eq. 4 and field values for the new refined and 190 coarsened cells can be defined using an identical formulation as is used for the U191 and D operator, respectively. However, the Basilisk solver allows the formulations 192 for upsampling and downsampling during the grid-resolution assessment and the 193 actual refinement and coarsening of cells to differ. 194

In general, the tree grid that results from applying the adaptation algorithm 195 results in the presence of the aforementioned resolution boundaries and accompa-196 nying ghost cells within the domain (see Fig. 2). To define the field values of ghost 197 points, the Basilisk solver uses the downsampling and upsampling operations. The 198 implementation is visually represented for a 1D scenario in Fig. 4. First, down-199 sampling is used to define the field values of ghost points on the high-resolution 200 side of a resolution boundary. Second, an upsampling method is used to define the 201 field values of the ghost points on the coarse side of the resolution boundary. By 202 using this method, the estimation error in the ghost cells' field values scales with 203 ζ. 204

The formulations used for downsampling and upsampling as exemplified in 205 Figs. 3 and 4 can be easily extended to two and three dimensions, for so-called 206 quadtree and octree grids, respectively. In order to demonstrate the algorithm and 207 the effect of different ζ values on the representation of a turbulent field, the next 208 section shows the results of the algorithm applied to a slice of a 3D turbulent field. 209 The Basilisk solver can run in parallel on many processors by applying a domain 210 decomposition using the Message Passing Interface (MPI). As the grid structure 211 may change during a simulation run, an important issue is load-balancing; the de-212 composition of the domain between processors must then be modified as the grid 213 is locally refined or coarsened. This is achieved in the Basilisk solver using the nat-214 ural decomposition of a Z-ordering space-filling curve applied to the quad/octree 215 structure (Griebel and Zumbusch 2001). 216

217 2.2 An Example of the Adaptation Algorithm

This section aims to exemplify how the adaption algorithm assesses a discretized 218 signal and adapts the grid according to a refinement criterion ζ . For this pur-219 pose, we apply the algorithm to a subset of the data from the simulation of forced 220 isotropic turbulence in Li et al. (2008). The simulation is run using a fixed equidis-221 tant grid with 1024^3 nodes; in terms of the Kolmogorov length scale (η) , the grid 222 spacing (Δ_i) is $\Delta_i = 2.2\eta$. For the analysis we assume the data to be resolved well 223 enough, and the results are kindly made available via the Johns Hopkins turbu-224 lence databases (http://turbulence.pha.jhu.edu/). We analyze a 2D slice of the 225 data (i.e. 1024^2 cells) and for simplicity, we only consider the velocity component 226 perpendicular to the sliced plane (u_{\perp}) . The data are presented in Fig. 5a; using 227 the algorithm described in Sect. 2.1, we can evaluate the χ field corresponding to 228 the original u_{\perp} field. A section of the resulting field, indicated by the black box in 229 Fig. 5a, is shown in Fig. 5b, where we can clearly see that the estimated discretiza-230 tion error is not distributed uniformly by the equidistant-grid approach that was 231 used in the simulation. Rather, it appears that there are anisotropic structures 232 present, visualized by relatively high χ values (in yellow). These structures appear 233 to correspond to vortex filaments that characterize the dissipative structures of 234 high-Reynolds-number turbulence (Frisch, 1995). This result motivates the appli-235 cation of the grid refinement algorithm to the data sample shown. Note that we 236 cannot 'add' new information by refinement and at this point we do not make any 237 claims regarding what χ values are reasonable for a turbulence-resolving simula-238 tion (this will depend on the numerical formulations and is the topic of a future 239 study). As such, we only allow the algorithm to coarsen the field with a maximum 240 error threshold ζ (as defined in Eq. 4). The number of grid cells resulting from the 241 application of the adaptation algorithm for a range of ζ values is shown in Fig. 5c; 242 as expected, the number of grid cells decreases with an increasing ζ value. Note 243 that the plot also shows that even for the high ζ values, the grid still contains cells 244 at the maximum resolution. 245

The main concept of employing the described grid-adaption algorithm is visualized in Fig.5d. Here histograms of the number of grid cells within 512 equallyspaced χ bins are presented for the original data and the data obtained from applying the grid adaptation technique with three different refinement criteria. It appears that for the original dataset, the histogram is monotonically decreasing

 $\mathbf{6}$

with increasing χ . This shows that many grid cells exist where the numerical solution is relatively smooth compared to cells in the tail of the histogram. Hence, if

the grid is chosen such that the discretization errors in the latter region do not affect the relevant statistics of the flow evolution, then the grid must be over-refined elsewhere. The histograms of the adapted grids show that the algorithm is able to lower the number of grid cells with low χ values, such that fewer grid cells are employed. Note that the grid coarsening does not introduce new grid cells with $\chi > 2\zeta/3$, as this part of the histogram remains unaltered.

When grid cells with a small but finite χ value are coarsened, some of the data 259 are lost and in general cannot be exactly reconstructed by interpolation techniques 260 (see Sect. 2.4). In order to assess how the data from the adapted grids compare 261 with the original data, Fig. 5e presents the corresponding power spectra. It ap-262 pears that none of the adapted grid data are able to exactly reproduce the original 263 power spectrum; more specifically, with increasing ζ values, the wavenumbers (k)264 that show a significant deviation in E(k) from the original appear to decrease. 265 We would like to point out that in order to evaluate the spectrum we have lin-266 early interpolated the data from the non-uniform grids to an equidistant grid with 267 1024×1024 data points. The choice of the interpolation technique is arbitrary and 268 will pollute the diagnosed spectrum in a non-trivial manner. As such, we directly 269 compare all $1024^2 u_{\perp}(x,y)$ samples in Fig. 5f, where we see that the deviation of 270 the data from the 1:1 line is a function of ζ . 271

The example presented in Fig. 5 is meant to demonstrate the used adaptation algorithm. The following sections are dedicated to assessing its application to timedependent numerical simulations of a turbulent field for an atmospheric case.

275 2.3 Physical Case Set-up

As indicated in the Introduction, we ran a DNS case from the referenced literature 276 to validate, benchmark and exemplify the adaptive-grid approach. The cases from 277 virtually all atmospheric-turbulence-resolving studies prescribe the periodicity of 278 the solution in the horizontal directions. Unfortunately, at the time of writing, the 279 Basilisk solver cannot yet handle an adaptive grid in combination with periodic 280 boundaries. To circumvent this limitation, we limit ourselves to a case where there 281 is no mean horizontal forcing such that we can apply a no-penetration bound-282 ary condition for the normal-velocity component at the lateral boundaries. This 283 is supplemented with a Neumann-boundary condition for the tangential velocity 284 components, pressure and buoyancy fields. We realize that this choice might affect 285 the solution and therefore its impact is assessed by re-running the case using a 286 fixed and regular grid with both sets of lateral boundary conditions (not shown). 287 It appears that for the chosen set-up of the case, the simulation results are in-288 sensitive to the choice of the horizontal boundary conditions. Note that in future 289 work, we will update the adaptive solver such that periodic boundary conditions 290 can be combined with the AMR technique. 291

We study a case introduced by Van Heerwaarden and Mellado (2016) that was designed to investigate the growth and decay of a CBL. In Fig. 6 a schematic overview of the physical system is presented, and in their physical model a linearly stratified fluid at rest with kinematic viscosity (ν) and thermal diffusivity (κ) is heated from below by a surface with a constant temperature. For generality, ²⁹⁷ buoyancy (b) is used as the thermodynamic variable. The buoyancy is related to ²⁹⁸ the potential temperature (θ) according to;

$$b = \frac{g}{\theta_{\rm ref}} (\theta - \theta_{\rm ref}), \tag{5}$$

where θ_{ref} is a reference potential temperature and g the acceleration due to gravity. The initial linear stratification is expressed as $b(z) = N^2 z$, where N^2 is the Brunt-Väisälä frequency associated with the initial stratification and z is the height above the surface. We assign a surface buoyancy b_0 larger than zero. Van Heerwaarden and Mellado (2016) identified relevant length, time, velocity fluctuation and buoyancy flux scales, \mathcal{L}, T, U and B, respectively, according to;

$$\mathcal{L} = \frac{b_0}{N^2},\tag{6a}$$

$$T = \frac{b_0^{2/3}}{N^2 \kappa^{1/3}},\tag{6b}$$

$$U = \frac{b_0^{7/9} \kappa^{1/9}}{N^{2/3}},\tag{6c}$$

$$B = b_0^{4/3} \kappa^{1/3}, \tag{6d}$$

- and are used to analyze the results in a non-dimensional framework. Two dimensionless groups can be identified that describe the system for any given set of
- sionless groups c $\{\nu, \kappa, N^2, b_0\},\$

$$Pr = \frac{\nu}{\kappa},\tag{7a}$$

$$Re = \left(\frac{b_0^{4/3}}{\nu^{2/3}N^2}\right)^{4/3},\tag{7b}$$

where Pr is the Prandtl number and Re is the Reynolds number. Note that for Pr = 1, the definition of the Reynolds number is consistent with $Re = U\mathcal{L}/\nu$.

310 2.4 Numerical Set-up and Formulation

For the evolution of the three velocity components (u_i) , modified pressure (p) and buoyancy (b), the Navier-Stokes equations for an incompressible fluid are solved under the Boussinesq approximation, according to,

$$\frac{\partial u_i}{\partial t} + \frac{\partial u_j u_i}{\partial x_j} = -\frac{\partial p}{\partial x_i} + \nu \frac{\partial^2 u_i}{\partial x_i^2} + b\delta_{i3},\tag{8}$$

$$\frac{\partial b}{\partial t} + \frac{\partial u_j b}{\partial x_j} = \kappa \frac{\partial^2 b}{\partial x_j^2},\tag{9}$$

$$\frac{\partial u_j}{\partial x_j} = 0,\tag{10}$$

and with respect to no-slip and a fixed buoyancy (b_0) condition at the bottom boundary. At the top boundary, no-penetration with a free-slip condition is used

8

and for the buoyancy, a fixed vertical gradient (N^2) is prescribed. Furthermore, a damping layer in the top 25% of the domain is active that damps buoyancy and velocity fluctuations to prevent the artificial reflection of gravity waves at the top boundary. The adaptive-grid runs are initialized with a grid at the minimum reso-

³²⁰ lution that is locally refined to the maximum resolution near the bottom boundary ³²¹ (i.e. $z < \mathcal{L}/10$) before a random perturbation is added to the velocity components

(i.e. $z < \mathcal{L}/10$) before a random pertu and buoyancy field in each grid cell.

Each integration timestep, grid adaptation is based on the estimated error 323 (see Sect. 2.1) of the three velocity components, and the buoyancy field. For each 324 field a refinement criterion (ζ) is specified (ζ_{u_i}, ζ_b) , where we non-dimensionalize the refinement criteria according to $\xi_b = \zeta_b b_0^{-1}$ and $\xi_{u_i} = \zeta_{u_i} U^{-1}$. In order to validate the results and assess the performance of the adaptive solver, we iteratively 325 326 327 decrease the refinement criterion between runs whilst we limit the minimum grid-328 box size. This maximum resolution is inspired by Van Heerwaarden and Mellado 329 (2016), and to limit the degrees of freedom, we choose; $\xi_{u_1} = \xi_{u_2} = \xi_{u_3} = 2.7 \times \xi_b$. 330 We realize that this choice (based on trial and error) is rather arbitrary, as currently 331 a solid framework of how the refinement criteria should be chosen is still lacking. 332 The results are validated by a comparison with runs using a regular and fixed 333 grid at the maximum resolution, performed with the Basilisk and MicroHH flow 334 solvers: MicroHH is the numerical code used by Van Heerwaarden and Mellado 335 (2016) to obtain their results. This code represents a state-of-the-art flow solver 336 that is dedicated to studying atmospheric systems (Van Heerwaarden and Mellado 337 2016, Shapiro et al. 2016); for a detailed description of the MicroHH code see Van 338 Heerwaarden et al. (2017). In addition, the fixed grid results of the Basilisk and 339 MicroHH flow solvers are compared to each other. 340

We choose Pr = 1 and Re = 3000 with a domain size of $3\mathcal{L} \times 3\mathcal{L} \times 3\mathcal{L}$ and simulate the evolution of the system until the physical time t = 45T. In order to limit the computational costs, the evolution of the Basilisk-based run with a fixed regular grid is only computed until t = 10T. To illustrate the physical size of such a numerical experiment in reality; for a domain size of $0.5 \text{ m} \times 0.5 \text{ m} \times 0.5 \text{ m}$ and $\theta_{\text{ref}} = 21 \ ^{\circ}\text{C}$, the corresponding parameters are: $\mathcal{L} = 0.16 \text{ m}$, $\theta_{\text{bottom}} = 36 \ ^{\circ}\text{C}$ and T = 153 s. This could be interpreted as a modest laboratory experiment.

The simulations are performed with Surfsara's supercomputer Cartesius located in Amsterdam, The Netherlands (SURFsara 2017). An overview of the different runs, including the number of cores used, integration timesteps and total run time is listed in Table 1. Additional information on the case set-up for both models can be found at:

353 Basilisk:

355 MicroHH:

³⁵⁴ http://basilisk.fr/sandbox/Antoonvh/freeconv.c¹

³⁵⁶ https://github.com/microhh/microhh/tree/master/cases/vanheerwaarden2016

¹ From a users' perspective, the case definition for the adaptive-grid runs is (subjectively) not more involved than the fixed-grid counterpart. The more complex adaptation-specific formulations are addressed by a low-level part of the Basilisk toolbox that does not require explicit attention from the users' side.

357 3 Results

358 3.1 Grid Structure

First, we study the evolution of the solution and grid structure qualitatively. Ver-359 tical slices of the magnitude of the gradient of the buoyancy field $(\|\nabla b\|)$ and the 360 used grid at $t = \{2, 10, 20\}T$ for run BA-0.0025 are presented in Fig. 7. At t = 2T a 361 complex grid structure is generated by the AMR algorithm, and within the ABL, 362 the grid is refined at locations where vigorous turbulent structures are present. 363 Above the ABL (i.e. $z/\mathcal{L} > 1$), turbulence is absent and the grid is coarse. Both 364 effects are appealing from a physical perspective as the computations are focused 365 on the regions where the activity is present. As the physical time progresses, the 366 boundary layer becomes more neutrally stratified and the turbulence intensity de-367 creases. And again, in response, the adaptive-grid algorithm has coarsened the 368 grid at t = 10T. This remarkable effect is even more pronounced at t = 20T, where 369 the coarsened regions have grown in size, indicating that the number of grid cells 370 is decreasing over time. Physically speaking, this is facilitated by the fact that the 371 372 size of the smallest eddies increases as turbulence decays.

373 3.2 Validation

³⁷⁴ Next we compare the results obtained with the AMR and fixed-uniform-grid runs.

³⁷⁵ Following Van Heerwaarden and Mellado (2016), we compare the domain inte-

grated quantities: a boundary-layer height z_i that is based on the buoyancy profile, kinetic energy L buoyancy flux L and discipation L according to

 $_{377}$ kinetic energy I_e , buoyancy flux I_b and dissipation I_ϵ according to,

$$z_i = \frac{2}{N^2} \int^{\infty} (\langle b \rangle - N^2 z) \mathrm{d}z, \qquad (11)$$

$$I_{\alpha} = \int^{\infty} \langle \alpha \rangle \mathrm{d}z, \tag{12}$$

where α is a dummy variable for $\{e, b, \epsilon\}$ and $\langle \alpha \rangle$ denotes the horizontally-averaged 378 value of the quantity α . Figure 8a shows the evolution of the boundary-layer height, 379 where good agreement between all simulations is found. The boundary-layer height 380 is an integral measure for the amount of buoyancy (i.e. analogous to heat) in the 381 system, though due to the case set-up, this integral quantity is not a very sensitive 382 measure to assess the accuracy of the resolved turbulent motions. Therefore, we 383 focus on higher-order statistics. In general, the evolution of the total kinetic energy 384 shows similar behaviour between all runs (see Fig. 8). Nevertheless small discrepan-385 cies on the order of 5% are present, particularly between the runs with the adaptive 386 grid and the fixed uniform grids, and as expected, this discrepancy decreases when 387 the refinement criterion is more strict. In order to analyze the evolution of kinetic 388 energy in further detail, Fig. 8c presents the evolution of the domain-integrated 389 buoyancy flux, which represents the energy-production rate for this system. The 390 buoyancy flux agrees well for all different runs and the observed differences be-391 tween the runs are a result of turbulent fluctuations within the chaotic system 392 rather than systematic discrepancies. This indicates that the overall structure 393 and characteristics of the energy-producing motions are resolved accurately for all 394

³⁹⁵ runs, and for free convection, these motions are associated with the large thermal ³⁹⁶ plumes. In order to assess the representation of the small-scale structures in these

³⁹⁷ simulations, Fig. 8d presents the evolution of the resolved energy-dissipation rate.

³⁹⁸ Compared to the fixed-grid runs, the AMR-based runs slightly underestimate the ³⁹⁹ resolved absolute dissipation, an aspect that is present throughout the simulation.

³⁹⁹ resolved absolute dissipation, an aspect that is present throughout the simulation.
⁴⁰⁰ Again, the discrepancy appears to be controlled by the refinement criterion, for
⁴⁰¹ which using stricter (i.e. smaller) criteria, the results seem to converge towards
⁴⁰² the universe found with the found with the found with the results.

the values found with the fixed-grid runs. The fact that the runs diagnosed with a
lower dissipation rate are also associated with lower kinetic energy indicates that
a small part of the dissipation has a numerical/non-physical origin.

Figure 9 shows the vertical profiles of the kinetic energy at $t/T = \{2, 4, 25\}$, and 405 shows discrepancies at t/T = 2 between all runs. The highly chaotic flow structure 406 at this early stage of the simulation could explain some of the differences. However, 407 consistent with Fig. 8b, the adaptive-grid runs show a systematically lower kinetic 408 energy content over the entire domain. At t/T = 4, the profiles of the fixed-grid 409 runs agree well, and furthermore, the energy found in the adaptive-grid run BA-410 0.0025 also compares well. It can be seen from the time series in Fig. 8b that for 411 t/T < 5, the evolution of kinetic energy shows large fluctuations. Therefore, we also 412 compare the energy profiles at t/T = 25, where we see again that the fixed-grid run 413 still contains more energy than the adaptive-grid runs. Again, the adaptive run 414 with the smallest refinement criterion is closest to the fixed-grid profile compared 415 to the other adaptive-grid runs. 416

Although it appears that the adaptive-grid algorithm is able to refine the grid 417 at locations of the turbulent structures, discrepancies in the simulations results 418 remain present. This can be explained by the fact that the process of refining and 419 coarsening the mesh relies on a linear interpolation strategy for defining values 420 on new grid cells. This interpolation introduces additional errors compared to a 421 simulation that employs a static grid, and these errors are similar to the truncation 422 errors of fixed grid advection schemes and thus lead to similar additional numerical 423 dissipation of energy. More accurate interpolation techniques could be tested to 424 limit the error due to interpolation. Therefore, this relevant aspect will be studied 425 in more detail in the future. 426

⁴²⁷ 3.3 Performance

As discussed in the introduction, for highly dynamic flow configurations such as a 428 diurnal cycle, model performance may benefit from the AMR approach. Although 429 the present case of decaying convection is less dynamic than a full diurnal cycle, 430 it is tempting to compare the simulation performance of the AMR-based run to 431 its counterparts using a fixed and regular grid. Thereupon, several performance 432 characteristics are presented in Fig. 10. Figure 10a shows, for the AMR-based 433 runs, the evolution of the number of grid cells, that appear to be controlled by the 434 refinement criterion, in which a smaller value causes the algorithm to use a more 435 refined grid. As illustrated in the snapshots of Fig. 7, the number of grid cells varies 436 significantly over the course of the simulation. Supposedly, the computational re-437 sources are distributed more efficiently over time. Furthermore, even in the run 438 with the most strict refinement criterion, the number of grid cells does not exceed 439 21% of the maximum-resolution value. Figure 10b shows how the computational 440

speed (i.e. defined here as wall clock time per integration timestep) is correlated 441 with the number of grid cells. It appears that there are several regimes in how the 442 performance is affected by the number of grid cells. For a large number of grid 443 cells (i.e. $> 10^6$) the amount of integration timesteps per second increases with a 444 decreasing number of grid cells, indicating that the solver does indeed speed up 445 when the grid is coarsened. Note that the simulations apply many grid cells in the 446 early stage of the runs (i.e. at the right-hand side of Fig. 10b and uses fewer cells 447 as time progresses (towards the left-hand side of Fig. 10b). However, as denoted 448 by the $x^{0.6}$ -scaling line, in this regime the simulation speed is not linearly depen-449 dent on the amount of grid cells. Furthermore, for lower number of grid points 450 (i.e. $< 10^6$) the simulation speed appears to slow down when the simulation runs 451 with fewer grid cells, i.e. there is a performance penalty for coarser grids! Possible 452 causes for these performance characteristics are listed below: 453

 For this case, the grid structure of the coarsened grids at later stages in the simulation contains a relatively larger fraction of resolution boundaries (see Fig.
 These boundaries are associated with additional overhead as they require special attention by the solver (see Sect. 2.1).

2. The number of used processors (linked to domain decomposition for parallelization) is fixed throughout the simulations. Therefore, the relative overhead of
MPI-domain communication routines compared to actual calculations increases
as the number of grid cells decreases.

462 3. For coarse grids, the physical timestep taken per integration timestep increases 463 (Courant-Friedrichs-Lewy criterion). Since diagnostic analysis of the solution 464 is performed with a regular interval in the physical time, i.e. $\Delta t = T$ for profiles 465 and slices and $\Delta t = T/20$ for the domain-integrated quantities. The frequency 466 of calls to diagnostic routines increases (i.e. say, calls per 100 integration steps) 467 on average resulting in an increased effort per integration step.

In Fig. 10c the amount of system billing units (i.e. the used number of cores \times hours) 468 spending for the different runs is presented. Before an interpretation of the results 469 is made, it is important to realize that the performance of a simulation run is a 470 function of many aspects that ranges from the details of the hardware configuration 471 to the exact case set-up. Therefore, the results presented here are intended as 472 an illustration rather than as absolute values. Nevertheless, it is clear that the 473 MicroHH run is notably cheaper compared to the runs performed with the Basilisk 474 solver. This can be explained by the different numerical schemes that are employed. 475 Most notably, for obtaining the pressure field, the Basilisk code uses a multigrid 476 strategy for solving the corresponding Poisson equation whereas the isotropic-fixed 477 grid in MicroHH facilitates the usage of a spectral Poisson solver. Although the 478 spectral method requires more MPI communication for parallelization when using 479 a large number of processors, it is known to be more efficient (Fornberg 1998). If 480 we compare the adaptive and non-adaptive simulation runs performed with the 481 Basilisk solver, we do see a considerable decrease in costs for the adaptive method 482 runs. 483

In Fig. 10d the memory used for the different simulation runs is presented. Compared to the fixed-grid runs, the adaptive-grid runs require less memory. This is due to the fact that the maximum number of grid cells is considerably lower than the number of grid cells in the fixed-grid runs (see Fig. 10a). From this perspective, the adaptive-grid approach can also be attractive for applications where 489 the available memory is limited. However, even though the run with MicroHH

 $_{490}$ $\,$ employs many more grid cells, the required memory is comparable to that of run

⁴⁹¹ BA-0.0025, meaning that per grid cell, the MicroHH code is more efficient in terms

492 of memory.

493 4 Outlook: Towards Adaptive Mesh Refinement in Atmospheric LES

We have based our test and performance benchmark on an idealized flow config-494 495 uration of a CBL using DNS, providing a ground truth for our intercomparison. In the future, we plan to study more practically-oriented cases by using an LES 496 formulation. For many atmospheric cases, LES is preferred over DNS, because it 497 provides an efficient tool for studying high-Reynolds-number flows. Therefore, the 498 next step is to test the AMR approach in combination with an LES formulation. 499 In this section, we briefly discuss some preliminary results on this topic. Because 500 this is part of ongoing research, we do not perform a quantitative discussion of the 501 test runs, the results and performance characteristics. The presented results aim 502 to exemplify the AMR method for a different case and show the flexibility of the 503 AMR approach. The example is based on the LES intercomparison study case by 504 Bretherton et al. (1999), in which a boundary layer is filled with a smoke cloud 505 that cools from the top due to longwave emission. The boundary layer is initially 506 capped by a strong temperature inversion (i.e. 7 K over 50 m) at $z \approx 700$ m and 507 rises over the course of the simulation due to entrainment. The inversion layer 508 is identified as a region where turbulent length scales are suppressed and turbu-509 lent motions are anisotropic due to the stable stratification. As such, this region 510 requires a high resolution to capture the predominant turbulent structures accu-511 rately. In constrast, the convective turbulence in the boundary layer itself can be 512 captured on a relatively coarse grid (Sullivan and Patton 2011). Accordingly, we 513 decide not to base the grid adaptation upon the estimated discretization error in 514 the representation of the velocity-component fields, but only on the estimated er-515 ror in the smoke-cloud fraction and temperature fields. With such an approach the 516 AMR algorithm does not refine the mesh in order to resolve the small turbulent 517 structures in the near-neutral boundary layer, but allows the LES to employ the 518 SGS model effectively in this region. In this run, the numerical grid varies by three 519 levels of refinement, i.e. between 25 m and 3.125 m. Figure 11 presents snapshots 520 of the temperature and numerical grid taken at t = 3 h after initialization. It is 521 clear from Fig. 11a that an inversion layer is present, while Fig. 11b shows that 522 the numerical grid has a high resolution in the region of the inversion layer and 523 remains coarse in the boundary layer itself. Furthermore, we see the subsiding 524 shells in the boundary layer that are qualitatively similar to those observed in the 525 laboratory experiment performed by Jonker and Jiménez (2014). 526

For this case, the AMR algorithm dynamically adapts to the flow by redirecting the grid refinement to those regions of the spatial domain where it is required. Hence in this case, adaptation is predominantly spatially focussed, whereas in the DNS case the refinement was most prominent in the temporal domain (see Fig. 10a). As such, both examples in this study are complementary and both effects (spatial and temporal adaptive grid refinement) are expected to play an important role in future simulations of full diurnal cycles (cf. Fig. 1). Finally, we note the following; the present cases where restricted to spatially homogeneous set-ups, where 'scale separation' naturally occurs through the internal variability of turbulence, originating from the non-linearity of the governing equations. In reality, heterogeneity in the *surface boundary conditions* also becomes important and provides an additional cause of scale separation that may call for adaptive grid refinement. For example, refinement may be preferred at sharp transitions between different types of land use, such as land-sea interfaces.

541 5 Discussion and Conclusions

We have introduced and tested an adaptive mesh refinement (AMR) method for studies of the atmospheric boundary layer (ABL). This work is motivated by a desire to numerically study highly dynamic cases. Such cases are characterized by a high degree of scale separation throughout the spatial and temporal domain. This work should be viewed as the first step in our AMR-based research that assesses the usage of an AMR method for studies of the ABL. We have based our adaptive-grid simulations on the flow solver implemented in the Basilisk code.

The method is tested using DNS based on a case introduced by Van Heerwaar-549 den and Mellado (2016), describing the growth and subsequent decay of a CBL. 550 The AMR algorithm was able to identify the time-varying turbulent regions in the 551 domain and refined/coarsened the grid accordingly. The AMR-based simulations 552 can reproduce the simulation results of their fixed grid counterparts with minor 553 discrepancies. Furthermore, the AMR algorithm can be tuned to apply more grid 554 cells such that these discrepancies are suppressed. For all AMR runs, the number 555 of grid cells varies significantly over time, resulting in more efficient simulations 556 compared to using a regular fixed grid with identical numerical formulations. This 557 provides a proof of principle for the AMR method regarding ABL systems. 558

For this case, a numerical solver dedicated to ABL systems (MicroHH) outper-559 formed all other runs in terms of computational efficiency, indicating that there is 560 an overhead associated with the usage of the adaptive solver. In general, the exact 561 impact of this overhead depends on the details of the studied case. The most chal-562 lenging ABL systems typically owe their complexity to the dynamical interplay 563 between various processes at different length and time scales. Hence, the AMR 564 technique is likely to be more favourable as complexity increases. More specifi-565 cally, as discussed in Popinet (2011), the cost of an adaptive simulation, relative 566 to a constant resolution simulation (G) is expected to scale as 567

$$G = \frac{C_a \Delta^{-D}}{C_c \Delta^{-3}} = \frac{C_a}{C_c} \Delta^{3-D},$$
(13)

where C_a and C_c are constants related to the absolute speed of the computation for 568 the adaptive- and constant-resolution simulations, respectively; Δ is the ratio of 569 the minimum to the maximum scale of the physical system (i.e. a measure of scale 570 separation) and D is the effective (or fractal) dimension of the physical process 571 (which is necessarily ≤ 3). In the present study, Δ is relatively large (i.e of order 572 10^{-2}) and the computational gain using the adaptive method is correspondingly 573 small, whereas for challenging cases Δ can be several orders of magnitude smaller, 574 with a correspondingly larger potential gain in efficiency of the adaptive method 575

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⁵⁷⁶ relative to constant-resolution methods. This important aspect of the overall scal⁵⁷⁷ ing behaviour is illustrated in Appendix 1 for a canonical flow set-up. The results
⁵⁷⁸ shown herein thus motivate our continued research using the AMR technique.

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584 Appendix 1 The Lid-Driven Cavity in Two Dimensions

We study the relation between the computational costs and the scale separation 585 for a simulation of a fluid in a lid-driven cavity in two dimensions, and compare the 586 results from a regular fixed grid and the adaptive-grid-refinement approach as is 587 presented herein. The chosen physical set-up consists of a no-slip box (size $L \times L$), 588 in which an incompressible fluid with kinematic viscosity ν is set in motion by the 589 top lid that moves with a constant velocity (U_{lid}) in the left-to-right direction. It 590 is well known that this configuration results in a large circulation cell within the 591 domain. With system parameters L, ν and U_{lid} we can identify a Reynolds number 592 (Re_{lid}) according to 593

$$Re_{\rm lid} = \frac{U_{lid}L}{\nu}.$$
 (14)

In order to study the effect of varying scale separation on the performance statis-594 tics, the simulations cover a range of different Reynolds numbers. Following the 595 analysis of Clercx and Van Heijst (2017) on vortex-wall interactions in two dimen-596 sions, we take the (minimum) grid-box size inversely proportional to the Reynolds 597 number. As such, the Reynolds number represents the separation of scales in our 598 simulations (i.e. Δ in Eq. 13). The runs are initialized with the fluid at rest and the 599 flow evolution is simulated for a physical time $t_{end} = 20L/U_{lid}$. For the adaptive 600 grid simulations, a refinement criterion for the velocity components $\zeta = 0.005 U_{\text{lid}}$ 601 is chosen. All runs are performed using a single processor core. A snapshot of the 602 vorticity field and the corresponding grid structure at $t = t_{end}$ for $Re_{lid} = 500$ are 603 presented in Fig. 12. The maximum resolution of this simulation corresponds to a 604 256×256 grid. First, the solution is validated against the results obtained with the 605 fixed equidistant-grid runs in Fig. 13. Here the vorticity fields $(\omega(x, y))$ obtained 606 from the fixed-grid and adaptive-grid simulations are directly compared for the 607 runs with $Re_{lid} = \{250, 500\}$. We conclude that the chosen refinement criterion in 608 sufficiently small to accurately reproduce the results obtained with the equidis-609 tant grid. Second, Fig. 14 presents the scaling of the computational costs with the 610 Reynolds number. The simulation costs when employing the fixed and equidistant 611 grid appear to scale with the third power of the Reynolds number. This exponent 612 can be understood from the fact that the total number of grid cells scales with the 613 Reynolds number to the second power (i.e. in 2D, doubling the resolution requires 614

⁶¹⁵ four times as many grid cells). Furthermore, the well-known numerical-stability

criterion of Courant-Friedrichs-Lewy limits the timestep and scales inversely pro-616 portional to the grid-box size, meaning that the total number of timesteps is pro-617 portional to the Reynolds number. Combined, the computational costs scale with 618 the Reynolds number to the power of (2 space+1 time =)3. This analysis holds for 619 all equidistant-grid approaches, and as such, we can anticipate the computational 620 costs when using an equidistant-grid code that is an order of magnitude faster 621 than the solver we have chosen for our fixed-grid approach (i.e. the Basilisk solver 622 running in fixed-grid mode). Interestingly, with an increasing Reynolds number, 623 the observed scaling of the adaptive grid simulations is favourable compared to the 624 equidistant grid counterpart. The observed scaling (i.e $\propto Re^{1.9}$) reflects that the 625 resolution requirement is not space filling. Although that for the lower Reynolds 626 numbers (i.e. $Re \leq 1000$), the (theorized) fast equidistant-grid solver is more ef-627 ficient than the adaptive grid approach, there exists a crossing point where the 628 latter technique becomes the more effective option. This feature is indicative to all 629 processes in nature. in which, with an increasing scale separation, the space-filling 630 approach of an equidistant grid represents the worst-case scenario, neglecting the 631 so-called fractal dimension of the problem. Note that this concept also applies to 632 three-dimensional turbulence (see Chap. 8 in Frisch, 1995). However, the corre-633 sponding scaling behaviour for ABL cases is not obvious. 634

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Table 1 Overview of the different simulation run details. In the top section a reference name, the used solver, grid type, the (maximal) numerical grid resolution, lateral boundary conditions and refinement criterion (ξ_b , if applicable) are listed for each run. In the bottom section the used number of cores, the total amount of integration steps taken at $t/T = \{10, 45\}$ and the total wall clock time of each run are presented.

Run name	Code	Grid	$\begin{array}{c} n_x \times n_y \times n_z \\ \text{(Maximal)} \end{array}$	Lateral BCs	ξ_b
MicroHH	MicroHH	Fixed &	$512^2 \times 387$	Periodic	-
$BA-512^{3}$	Basilisk	Fixed	512^{3}	Neumann & No-pen.	-
BA-0.0025	Basilisk	AMR	512^{3}	Neumann & No-pen.	0.0025
BA-0.005	Basilisk	AMR	512^{3}	Neumann & No-pen.	0.005
BA-0.01	Basilisk	AMR	512^{3}	Neumann & No-pen.	0.01
Run name	Number of cores	Integration at $t/T = \{1$	steps 0, 45}	total wall clock time (D:HH:MM)	
$\begin{array}{c} \text{MicroHH} \\ \text{BA-512}^3 \\ \text{BA-0.0025} \\ \text{BA-0.005} \\ \text{BA-0.01} \end{array}$	HH 64 $\{13920, 35670\}$ 2^3 64 $\{14073, (35670)\}$ (estimated) 0025 96 $\{14095, 30144\}$ 005 96 $\{14061, 28704\}$ 01 96 $\{14167, 25544\}$			$\begin{array}{c} 0:12:22\\ ed) & 2:16:12 \ (t/T=10)\\ & 2:10:30\\ & 1:18:19\\ & 1:02:16\end{array}$	•



Fig. 1 Sketch of a prototypical diurnal cycle evolution. Adapted from Stull (1988).





Fig. 2 Example of a tree-grid structure. The top row presents the spatial structure of the grid cells with varying levels of refinement (a) and the locations of two types of ghost points whose field values are defined by the downsampling (red dots) and upsampling (blue dots) operations (b, see text). The plot on the bottom row presents a corresponding tree representation of the various grid cells and ghost points at different levels (c).



Fig. 3 A one-dimensional, visual representation of how the adaptation algorithm assesses the discretization of a curved field f(x): a) A coarser level estimate of the discretized solution is obtained using the downsampling operation. b) Using these coarse level values, the original discretized solution can be estimated using the upsampling operation. c) The difference between the estimated and original values is interpreted as an error estimator (χ) and can be compared against fixed thresholds (e.g. ζ). d) and e) If the refinement criterion is exceeded, new cells at one level higher are initialized locally by applying a linear interpolation technique using the initial cell values. Alternatively, if the estimated error is smaller than the coarsening criterion for multiple cells, these cells can be merged if that does not violate the general grid-structure requirements (see text and Fig. 2).



Fig. 4 Example of the treatment of a resolution boundary in a one-dimensional scenario. First, the high level region near the resolution boundary is downsampled to obtain values for the coarse-level ghost points in red (a). Second, linear interpolation of the coarse level solution is used to define the field values of high level ghost points in blue (b).



Fig. 5 Example of the adaption algorithm applied to a (2D) slice of a 3D turbulent field. a) Shows the data slice of the velocity component in the plane-perpendicular direction $(u_{\perp},$ obtained from Li et al. (2008). b) Presents the χ field, evaluated using the method described in Sect. 2.4. Only the centre part of the slice, indicated by the black box in (a), is shown to reveal the small scale details in this simulation. c) shows the grid cell number dependence on the chosen refinement criterion (ζ), note the logarithmic vertical axis. A histogram of the χ field with 512 bins for the original data, and the data corresponding to three ζ values are presented in d). Using the same colour coding as in d), power spectra and a direct comparison of the $u_{\perp}(y, z)$ field are shown in e) and f), respectively.



Fig. 6 Sketch of the system and its parameters. The red line illustrates a typical buoyancy profile within the CBL during the initial development. Adapted from Van Heerwaarden and Mellado (2016).



Fig. 7 Vertical slices of the $||\nabla b||$ field (left column) and the corresponding numerical grid (right column) in the lowest half of the domain. The top, middle and bottom rows represent snapshots taken at $t/T = \{2, 10, 20\}$, respectively. These snapshots are taken from the adaptive-grid run BA-0.0025.



Fig. 8 Time series of the domain integrated quantities, a) boundary-layer height (z_i) , b) kinetic energy (I_e) , c) buoyancy flux (I_B) and d) dissipation rate (I_e) according to Eq. 11. The results are obtained with both Basilisk and MicroHH using fixed grids and Basilisk using the adaptive mesh refinement algorithm. Note that plots c) and d) use a logarithmic scale.



Fig. 9 Vertical profiles of the horizontally-averaged kinetic energy $(\langle e \rangle)$ at $t/T = \{2, 4, 25\}$ in left, middle and right plot, respectively. The results are obtained with both Basilisk and MicroHH using fixed grids and Basilisk using the adaptive mesh refinement algorithm. Note that in panel c) the horizontal axis is rescaled and that regular-grid computations with Basilisk are not available (see text, Sect. 2.4).



Fig. 10 Overview of the performance characteristics of the adaptive and fixed-grid simulation runs. a) Time series of the number of grid points for the adaptive runs normalized by the maximum-resolution value (i.e. 512^3). b) Scatter plot of the wall clock time per integration step versus the used number of grid cells in the adaptive-grid runs. c) The total amount of System Billing Units (SBU, i.e. *number of cores* × *hours*) spending on each simulation run. Note that the value for BA-512³ is estimated as if it were run until t/T = 45. d) The total RAM memory used in each simulation run in gigabytes (GB).



Fig. 11 Snapshots of a) the vertical slices of the virtual potential temperature field and b) the numerical grid at t = 3 h. The case is based on the work of Bretherton et al. (1999).



Fig. 12 Snapshots of a) The vorticity field and b) the numerical grid at $t = t_{end}$ for the lid-driven cavity simulation with $Re_{lid} = 500$.



Fig. 13 Validation of the vorticity field (ω) from the adaptive grid simulation against the results obtained with a fixed equidistant grid. For a) $Re_{lid} = 250$ and b) $Re_{lid} = 500$. The inserts show a zoom-in (i.e. rescaled axes), containing $\approx 95\%$ of the total number of data points.



Fig. 14 The correlation of the computational costs and the Reynolds number (Re_{lid}) for different approaches. The green line represents the theorized results from a solver that is an order of magnitude faster than the fixed-grid approach that we have used (blue dots).