

Convection Effects During Bulk Transparent Alloy Solidification in DECLIC-DSI and Phase-Field Simulations in Diffusive Conditions

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Abstract	conditions, three-dimen on a model transparent Insert in the Device for repeated on Earth under macrosegregation result spacings measured in gr experiments provide un formation under diffusi importance of accuratel interface and the selecti	formation and evolution of cellular and dendritic arrays under diffusive growth asional (3D) directional solidification experiments were conducted in microgravity alloy onboard the International Space Station using the Directional Solidification the study of Critical Liquids and Crystallization. Selected experiments were r gravity-driven fluid flow to evidence convection effects. Radial and axial ting from convection are observed in both ground experiments, and primary round and microgravity experiments are noticeably different. The microgravity tique benchmark data for numerical simulations of spatially extended pattern ve growth conditions. The results of 3D phase-field simulations highlight the ly modeling thermal conditions that strongly influence the front recoil of the ion of the primary spacing. The modeling predictions are in good quantitative
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³ Convection Effects During Bulk Transparent Alloy Solidification ⁴ in DECLIC-DSI and Phase-Field Simulations in Diffusive ⁵ Conditions

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To study the dynamical formation and evolution of cellular and dendritic arrays under diffusive growth conditions, three-dimensional (3D) directional solidification experiments were conducted in microgravity on a model transparent alloy onboard the International Space Station using the Directional Solidification Insert in the Device for the study of Critical Liquids and Crystallization. Selected experiments were repeated on Earth under gravity-driven fluid flow to evidence convection effects. Radial and axial macrosegregation resulting from convection are observed in both ground experiments, and primary spacings measured in ground and microgravity experiments are noticeably different. The microgravity experiments provide unique benchmark data for numerical simulations of spatially extended pattern formation under diffusive growth conditions. The results of 3D phase-field simulations highlight the importance of accurately modeling thermal conditions that strongly influence the front recoil of the interface and the selection of the primary spacing. The modeling predictions are in good quantitative agreements with the microgravity experiments.

INTRODUCTION

33 Solidification microstructures play an important 34 role in the mechanical and physical properties of 35 materials such as metallic alloys. The correlation 36 between solidification microstructure and process-37 ing conditions can be examined by directional 38 solidification experiments in which all parameters 39 can be independently controlled.^{1,2} Most theoretical 40 models have been developed considering diffusive 41 heat and mass transport. These models have been 42 validated by comparison with well-controlled exper-43 iments, most of which have been performed in thin 44 samples to limit the extent of convection through 45 confinement. In bulk samples, natural convection in 46 the melt disturbs the homogeneity of processing conditions and, hence, microstructure formation.³⁻ 47 48 Solidification experiments under low-gravity condi-49 tions provide an exceptional tool to investigate

interfacial pattern formation and solidification50microstructure development in spatially extended51samples with negligible convection and under well-52controlled growth conditions. Those experiments53provide unique benchmark data to validate the54predictions of theoretical and computational models55in a purely diffusive growth regime.56

57 The present study comprises experiments per-58 formed in the Directional Solidification Insert (DSI) dedicated to in situ and real-time characterization 59 of the dynamical selection of the solid-liquid inter-60 face morphology in bulk samples of transparent 61 materials. It was developed by the French Space 62 Agency (CNES) in the framework of the DECLIC 63 project (Device for the study of Critical Liquids and 64 Crystallization). The DECLIC facility was installed 65 onboard the International Space Station (ISS) as 66 part of a joint NASA/CNES research program. A 67 series of microgravity experiments was performed 68

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69 from April 2010 to April 2011, which were followed 70 and controlled remotely in near real-time condi-71 tions. The favorable conditions provided by micro-72 gravity allowed us to observe for the first time the 73 dynamics of extended oscillating cellular patterns 74 that were further investigated by three-dimensional (3D) phase-field simulations.⁸⁻¹⁰ Significant differ-75 ences between the behavior of extended oscillating 76 77 patterns and confined patterns in thin samples 78 were evidenced, stressing the importance of pattern 79 dimensionality. The DSI was then brought back on 80 ground and experiments were performed with sim-81 ilar control parameters as in microgravity.

Author Proof

82 The aim of this article is to analyze directional 83 solidification experiments realized under micro-84 85 86 87 gravity and under Earth's gravity with similar control parameters to address the effect of natural convection on interface patterns. The article is organized as follows. We first briefly describe the 88 experimental apparatus and procedure as well as 89 the characteristics of convection on Earth during 90 upward solidification experiments and summarize 91 the phase-field simulation methodology. We then 92 present the results for the interface recoil (i.e., the 93 time evolution of the axial interface position) and 94 the axial macrosegregation. The results demon-95 strate that including the effects of thermal diffusion 96 and latent heat rejection at the interface in phase-97 field simulations, so as to model accurately the 98 spatiotemporal evolution of the thermal field in the 99 experimental bulk samples, is necessary in predict-100 ing the front recoil. The comparison of Earth and 101 microgravity front recoils further reveals the exis-102 tence of a typical convecto-diffusive macrosegregation profile on Earth and, more surprisingly, even 103 104 for microgravity experiments at a small growth 105 rate. Next, we present the results pertaining to 106 pattern formation. We first characterize the well-107 known shift of the planar-to-cellular velocity threshold resulting from convection.¹²⁻¹⁴ We then show 108 109 the existence of radial macrosegregation by analyz-110 ing the primary spacing distribution, and we char-111 acterize fluid flow effects on stationary primary 112 spacings. In contrast to previous observations in metallic systems,^{15,16} we find here that suppression 113 of convection in microgravity leads to smaller 114 115 spacings than in ground-based experiments. Finally, we compare spacings observed in micro-116 117 gravity and predicted in phase-field simulations. We 118 find that accurate modeling of the thermal field also 119 improves the quantitative agreement between mea-120 sured and predicted spacings.

121 Experimental Procedure and Numerical122 Simulations

123 The DECLIC-DSI includes a Bridgman furnace 124 and an experimental cartridge.¹⁷ The cartridge 125 comprises a quartz crucible and a system of volume 126 compensation, which is mandatory to accommodate 127 the specimen volume variations associated with phase changes. The cylindrical crucible has an inner128diameter of 10 mm and a length that enables about12910 cm of solidification, allowing the study of the130whole development of extended patterns from their131initial stages up to the steady growth regime.132

Figure 1 shows a representation of the optical 133 apparatus, where the experimental cartridge is 134 reduced to a cylinder with the liquid on top and 135 solid at the bottom. In situ continuous observation 136 led to a large number of images, and bespoke 137 systematic image analysis procedures were devel-138 oped. The main observation mode (axial) takes 139 advantage of the complete transparency of the 140 experimental cartridge: The light coming from 141 LEDs passes through the cartridge from the bottom 142 to the top, crossing the interface whose image is 143 formed on a charge-coupled device (CCD) camera. 144 Examples of cellular and dendritic patterns are 145 given in Fig. 1. The top-view images of the interface 146 are used to study array dynamics and characteris-147 148 tics, such as the primary spacing, i.e., the center-tocenter distance between two first-neighbor cells. On 149 the same axis, a Mach–Zehnder interferometer with 150 a He-Ne laser produces interferometric images 151 (examples in Fig. 2) used for three-dimensional 152 reconstruction of microstructures.¹⁸ Additionally, 153 the transverse observation mode provides side-view 154 interface images (examples in Fig. 3), which allow a 155 real-time characterization of interface motion and 156 macroscopic shape. 157

The studied organic transparent compound was a 158 succinonitrile (SCN)-0.24 wt.% camphor alloy pre-159 pared with SCN purified by NASA by successive 160 distillation and zone melting. Every step of sample 161 preparation was carefully realized under vacuum to 162 avoid humidity contamination. Solidification was 163 performed by pulling the experimental cartridge 164 from the hot zone toward the cold zone at a constant 165 rate V in a range between 0.1 and 30 μ m/s. A 166 thermal gradient G ranging from 10 K/cm to 30 K/ 167 cm was imposed by regulated hot and cold zones, 168 which were located above and below the adiabatic 169 area where the interface was positioned. Further 170 details about the experimental procedure can be 171 found in previous articles.^{19,20} 172

Phase-field (PF) modeling is a powerful computa-173 174 tional approach to solving a challenging free-boundary problem and to simulating the dynamics of complex interfaces.²¹⁻²⁴ The quantitative PF model 175 176 used here has been developed over the past decades 177 to explore quantitatively a variety of microstructure 178 dynamics during alloy solidification. This model 179 exploits a thin-interface asymptotic analysis^{25,26} 180 and an anti-trapping current^{27,28} to model accu-181 rately the solid-liquid interface dynamics with a 182 computationally tractable diffuse interface width 183 much larger than the width of the actual physical 184 interface.^{25,26} Using the latest three-dimensional 185 implementation of this model parallelized on graph-186 ics processing units,^{8,10,12} we considered three dif-187 ferent descriptions of the thermal field to assess its 188



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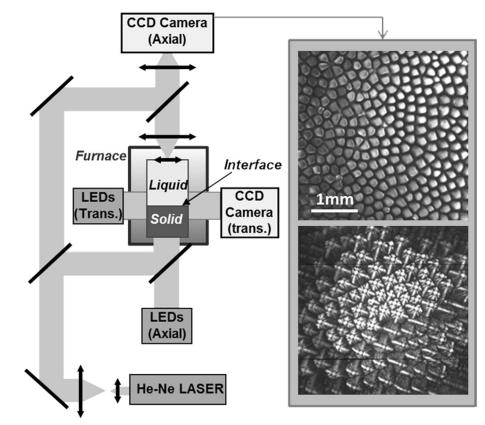


Fig. 1. Representation of the optical apparatus of the DECLIC-DSI (left). Examples of cellular and dendritic patterns obtained under microgravity (right).

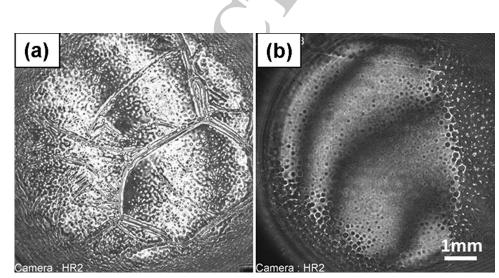


Fig. 2. Axial view images obtained by interferometry of the solid–liquid interface of SCN-0.24 wt.% campbor alloy for V = 1 μ m/s: (a) microgravity and (b) ground experiments (G = 19 K/cm).

influence on interface recoil and microstructure
formation. First, the widely used frozen temperature approximation (FTA) assumes a constant temperature gradient translating uniformly at the
sample pulling velocity. Second, the thermal drift
approximation (TDA) models phenomenologically

the drift of isotherms evidenced experimentally 195 with fit parameters extracted from experimental 196 measurements.²⁰ Third, a time-dependent thermal 197 field calculation (TFC) that is completely free of fit 198 parameters couples the evolution equations of the 199 phase field and solute field to a full time-dependent 200

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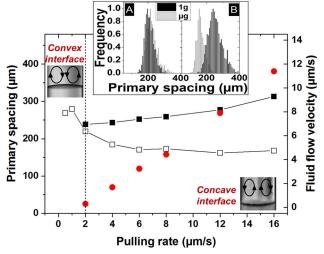


Fig. 3. Primary spacing as a function of pulling rate under microgravity (T) and on Earth (■). Right Y axis: evolution of upward fluid flow velocity (\bullet). Insets: primary spacing histograms for V = 2 (A) and 12 μ m/s (B); examples of macroscopic interface shape with corresponding convection loops (G = 19 K/cm).

201 calculation of axial (1D) thermal diffusion in the 202 adiabatic zone. This calculation assumes that the temperature field is radially uniform and includes latent heat release at the interface. PF simulations use similar parameters as in our recent studies,^{9,10} 205 206 and details of the thermal field calculation will be 207 presented elsewhere.

RESULTS

209 The microstructure formation was studied start-210 ing from the interface at rest until its stationary 211 state, for a thermal gradient of 19 K/cm and pulling 212 rates ranging from 0.5 μ m/s to 16 μ m/s. The first 213 stages of solidification have fast dynamics with the 214 transient development of microstructure, followed 215 by a stationary state growth characterized by a 216 stable spacing even if the patterns keep evolving in 217 of topological order. Description terms of 218 microstructures, primary spacing evolution, and mechanisms of adjustment can be found else-219 where.²⁹ Different effects of convection, found by 220 221 comparative experiments on ground (upward 222 growth, V//-g) and under microgravity, are ana-223 lyzed in the next subsections.

224 The Origin of Curvature and Resulting 225 **Macroscopic Effects**

226 At rest, the planar solid-liquid interface is located 227 close to the *liquidus* temperature of the alloy. Its 228 macroscopic interface shape is determined by the 229 isotherm shape, usually curved in 3D. The interface 230 is convex (inset in the top-left of Fig. 3) for both 231 thermal gradients, meaning that the crucible 232 receives heat from the outside. Consequently, the 233 liquid at the border of the crucible is hotter than in 234 the center.

Once a pulling rate is applied, the interface moves 235 236 closer to the cold zone and isotherms become concave. Moreover, as a result of the low thermal 237 conductivity of the alloy compared with the quartz 238 crucible, evacuation of latent heat generated during 239 growth mainly occurs through the crucible, thus, 240 inducing a thermal radial gradient from the border 241 (colder) to the center (hotter) of the crucible. The 242 interface then progressively becomes more concave 243 (inset in the bottom-right of Fig. 3) as velocity 244 increases. Lastly, isotherms are shifted downward 245 by heat transport in the cartridge, also leading to an 246 increase of concavity. Both effects are approxi-247 mately proportional to the pulling rate as detailed 248 in a previous study.²⁰ 249

On the ground, the radial thermal gradient 250 induces thermal convection.³⁰ For convex interfaces, 251 the hotter liquid with lower density at the border is 252 driven upward by buoyancy, so that a toric convec-253 254 tion loop, ascending at the border and descending in the center, is formed. For concave interfaces, the 255 liquid in the center is hotter, so that fluid flow rises 256 in the center of the crucible. The fluid flow direction 257 is schematically represented in the top-left and 258 bottom-right insets of Fig. 3, respectively, for con-259 vex and concave interfaces. 260

Interface Motion and Axial Macrosegregation 261

When pulling of the sample at a constant velocity 262 V is initiated, the interface recoils within the 263 temperature frame, and the solute concentration 264 builds up on the liquid side of the interface. This 265 266 interface motion from rest to its steady-state position is referred to as the front recoil. The planar 267 front undergoes the Mullins-Sekerka instability^{31,32} 268 when the solute concentration gradient in front of 269 the interface $G_{\rm c}$ exceeds the critical value ${\rm G}/m_{\rm L}$, 270 with $m_{\rm L}$ the alloy *liquidus* slope. This initial 271 interface dynamics before the destabilization of the 272 planar front can usually be well predicted by the 273 274 Warren–Langer model.³

275 Figure 4a gives an example of the experimental interface recoil for $V = 4 \ \mu m/s$ (μg experiment) 276 obtained from transverse observation, focused on 277 the initial solidification stage. The three different 278 thermal descriptions, namely, FTA, TDA, and TFC 279 are compared with the experimental measurements. 280 The FTA calculation, representative of the classic 281 Warren-Langer model,³³ underestimates both the 282 recoil rate and its final extent compared with 283 present experiments. Other phase-field simulations 284 using the TDA or the TFC provide a much better 285 agreement with experiments. Not only is the time 286 evolution of the interface position better estimated, 287 but the time for the occurrence of the morphological 288 instability ($L \sim 1.88$ mm for the TDA and 1.76 mm 289 for the TFC) is also in much better quantitative 290 agreement with experimental observations 291 (1.55 mm < L < 1.96 mm). Steady-state cell/den-292 293 drite tip positions after morphological instability



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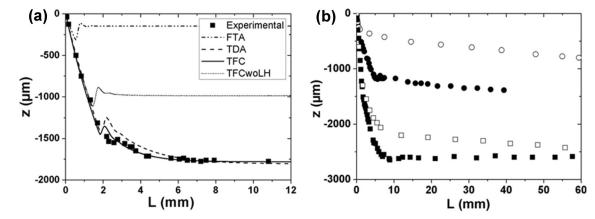


Fig. 4. Interface position (z) as a function of solidified length (L = V t). (a) The experimental interface dynamics at V = 4 μ m/s is compared with PF simulations with different thermal conditions. (b) Comparison between results obtained under microgravity (full symbols) and on the ground (empty symbols) for V = 8 (squares) and 0.5 μ m/s (circles) (G = 19 K/cm).

294 are also well predicted. Note that the coupled TFC 295 provides a better estimation than the phenomeno-296 logical TDA, while involving no adjustable fit 297 parameter. Furthermore, the TFC simulation 298 neglecting latent heat rejection contribution (TFC-299 woLH in Fig. 4a) underestimates the steady-state 300 interface position, thereby clearly demonstrating 301 the critical influence of latent heat rejection on the 302 recoil of the interface.

303 Since the interface grows in local thermodynamic 304 equilibrium, solute interfacial composition varia-305 tions in the liquid induce similar variations of 306 composition in the solid. The solute build-up during 307 planar front recoil, and the subsequent transient 308 development of patterns of cells or dendrites before a 309 steady-state growth regime is reached, is a first cause 310 of axial macrosegregation. If the interface tempera-311 ture keeps evolving, meaning that steady state is not 312 reached, an additional source of axial macrosegrega-313 tion, such as convection, must be considered.

314 In Fig. 4b, the fast motion characterizing the 315 initial transient is similar on the ground and in 316 microgravity, but the interface decelerates earlier 317 on the ground. Depending on processing conditions, it may then stabilize at a steady-state position or 318 keep drifting slowly with time. Stabilization (i.e., 319 320 $\mu g - V = 8 \mu m/s$ in Fig. 4b) corresponds to a *plateau* 321 of concentration reached after the initial transient: 322 It is typical of a purely diffusive mode, and most 323 microgravity results follow this tendency. The 324 absence of stabilization corresponds to a character-325 istic "S-shape" macrosegregation profile, typical of a 326 convecto-diffusive growth mode.³⁴ This evolution is 327 observed for ground experiments (i.e., 1 g - V = 0.5328 and 8 μ m/s in Fig. 4b) but also surprisingly for microgravity experiments performed at the lowest 329 pulling rate (i.e., $\mu g - V = 0.5 \mu m/s$ in Fig. 4b). For 330 331 these conditions, an order of magnitude analysis of fluid flow^{35,36} revealed that diffusive conditions 332 would be obtained if the gravity in ISS was actually $10^{-6}g$ (Fig. 13 of Mota et al.²⁰), but as it is much 333 334

closer to $10^{-3}g$ or $10^{-4}g$, experiments performed at low velocities still appear to be affected by residual convection. 336

338

Pattern Development and Spacing Selection

For the lowest studied pulling rate, i.e., 0.5 μ m/s, 339 significant differences were found between ISS and 340 ground experiments. Cells are fully developed under 341 microgravity, whereas on the ground, the interface 342 is still smooth, indicating that the critical velocity 343 from planar to cellular interface is shifted toward 344 higher velocity on the ground. This shift of planar 345 stability at higher velocities (for k < 1) resulting 346 from liquid convection was already evidenced the-347 oretically^{12,37} and experimentally.¹ 348

A classic approach to assessing the effect of liquid 349 convection consists of considering a diffusive bound-350 ary layer in the liquid close to the solid-liquid 351 352 interface, as initially introduced by Burton et al.³⁸. Fluid flow reduces the thickness δ of the boundary 353 layer where transport is diffusive in front of the 354 355 interface; the concentration ahead of this boundary layer is considered homogeneous and well mixed at 356 the nominal concentration. Solute conservation 357 then implies the decrease of the liquid solute 358 concentration at the interface compared with pure 359 diffusive conditions. The solute concentration in the 360 solid also decreases because it is proportional to the 361 concentration in the liquid by the partition coeffi-362 cient k. The driving force for the interface instability 363 is thus reduced inducing a stabilization of the 364 interface and an increase of the critical velocity 365 (see, e.g., Clarke et al.¹² and its supplementary 366 material for a discussion and an analytical deriva-367 tion of the constitutional supercooling criterion in 368 the presence of a diffusive boundary layer). 369

The interface instability driving force on ground $m_{\rm L} \cdot G_{\rm c,conv}$ can be written as $k_{\rm eff} \cdot m_{\rm L} \cdot G_{\rm c,diff}$ where $G_{\rm c}$ is the concentration gradient at the interface and $k_{\rm eff}$ is the effective partition coefficient defined as the ratio between the solute concentration in the solid at the 374



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375 interface and in the liquid far from the interface. $k_{\rm eff}$ 376 tends toward 1 in purely diffusive conditions and 377 toward the partition coefficient k in convectiondominated conditions. The effective segregation coef-378 379 ficient can then be estimated as the ratio between the 380 critical velocities under microgravity and on the 381 ground. An estimation of critical velocities is per-382 formed using the instantaneous interface velocity 383 obtained following the interface position by side-view 384 observation, combined to top-view observation to 385 detect instability triggering. Onboard the ISS, the 386 critical velocity is 0.48 μ m/s and on the ground 387 0.95 μ m/s, which leads to a $k_{\rm eff}$ of 0.51. This is a sign 388 of partial mixing of the bulk liquid by convection 89 90 91 92 93 94 95 during solidification. As mentioned, the critical velocity measured onboard the ISS at $V = 0.5 \ \mu m/s$ may be affected by residual convection: the value of $k_{\rm eff} = 0.51$ is then slightly overestimated as $k_{\rm eff}$ should be calculated using a critical velocity determined in a purely diffusive growth regime.

The onset of morphological instability is shown on 396 Fig. 2 by interferometry top-view observations. 397 Under microgravity, destabilization occurs homoge-398 neously over the whole interface; for upward growth 399 on ground, instability starts near the border and 400 progressively propagates toward the center. A sim-401 ilar effect of inhomogeneous triggering was previ-402 ously observed for succinonitrile-acetone alloys but for a macroscopically concave interface.³⁹ In this 403 404 case, microstructures first developed in a small 405 region near the center of the interface, and the 406 inward radial transport of solute by fluid flow 407 eventually led to a solute accumulation in the 408 center where instability first developed. In the 409 present study, from rest up to morphological insta-410 bility, regardless of experimental conditions, the 411 interface is macroscopically convex (i.e., with con-412 vection descending in the center). The radial con-413 centration gradient induced by fluid flow is opposite compared with the case of a concave interface. 414 415 Morphological instability thus triggers at the border 416 and propagates toward the center.

417 Morphological instability eventually leads to the formation and growth of cells, which compete until 418 they reach a stable spacing. The radial concentra-419 420 tion gradient induced by fluid flow also affects the 421 final microstructure. Under microgravity, patterns 422 are more homogeneous than on the ground where a 423 radial size gradient from large cells in the center to small cells at the crucible border is observed.^{14,30,39} 424 425 The size distribution histograms presented in Fig. 3 426 highlight those differences. At 12 μ m/s, the concave 427 interface is associated with a noticeable convection 428 on the ground that increases the average primary 429 spacing but also enlarges the distribution. Con-430 versely, at 2 μ m/s, the interface is macroscopically 431 flat and the radial thermal gradient almost van-432 ishes, so that even on ground conditions, thermal 433 convection is very weak: the resulting peak and the 434 width of size histograms are thus similar on the 435 ground and under microgravity.

Comparative analyses of microgravity and ground 436 experiments pertaining to primary spacing depen-437 dence on the pulling rate, presented in Fig. 3, also 438 reveal major differences. Similar primary spacing 439 values are found for a macroscopically flat interface 440 $(2 \ \mu m/s)$ as convection is weak. For higher pulling 441 rates, spacing under microgravity is lower than on 442 the ground. Such results may seem surprising 443 because the results of previous studies on metallic 444 systems pointed out spacing in microgravity larger 445 than on the ground with a convergence when the pulling rate increases. 15,16 The difference may come 446 447 from the different origin of convection. As previ-448 ously explained, in the studied transparent system, 449 the radial thermal gradient, and associated concav-450 ity, is induced by the low conductivity of the sample 451 compared with the crucible: The radial thermal 452 gradient associated with concavity then increases 453 with latent heat release, so fluid flow and its effect 454 increase with pulling rate. In metallic systems, 455 thermal convection stems from the difference of 456 conductivities between the solid $(K_{\rm S})$ and the liquid 457 $(K_{\rm L})$, with $K_{\rm S} > K_{\rm L}$, which promotes interface con-458 vexity. The latent heat diffusion effects are negligi-459 ble because it is fast in the sample; when the pulling 460 rate becomes large enough, the effect of fluid flow is 461 thus negligible⁴⁰ and the primary spacing is no 462 longer affected by fluid flow. The current opposite 463 effect of fluid flow on primary spacing can be 464 attributed to the flow direction, resulting from the 465 concavity or convexity of the interface: In metallic 466 systems, the liquid sinks in the center and rises on 467 the border, whereas it is the opposite in the 468 469 currently studied transparent system.

Lehmann et al. Ref.⁴¹ explain the variation of 470 primary spacing with fluid flow by the modification 471 of the concentration of the interdendritic liquid. On 472 this basis, they propose a relationship between the 473 component of fluid flow velocity U| parallel to the 474 pulling velocity (here taken negative for a down-475 ward component) and the primary spacing λ and λ_0 , 476 respectively. with and without influence of 477 convection: 478

$$U_{||} = V \left[1 - \left(\frac{\lambda}{\lambda_0} \right)^2
ight]$$
 (1)

For a convex interface, flow is descending in the 480 center so U| < 0 and $\lambda_0 > \lambda$ as observed in experiments in metallic systems.^{15,40} Equation 1 is pri-482 483 marily designed for downward flows, but its use can 484 be extended to upward flows as long as the fluid 485 velocity remains lower than the pulling rate. Using 486 Eq. 1, we calculated fluid flow velocities U| as a 487 function of V, as plotted in Fig. 3 (right y axis). At 488 $2 \mu m/s$, U is almost zero, corresponding to the 489 experimental conditions on the ground closer to 490 pure diffusive transport; U| increases with pulling 491 492 rate (increase of concavity) leading to a progressively larger spacing on the ground compared with 493



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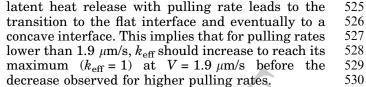
494 that in microgravity. The fluid flow velocity can 495 then be taken into account in the mass transfer 496 equation, enabling expression of the convecto-diffusive parameter $\Delta = \frac{\delta V}{D}$, where *D* is the solute diffusion coefficient in the liquid, as a function of fluid flow velocity $U \downarrow$.^{40,42} In this approach, $\Delta = 1$ for a 497 498 499 500 purely diffusive regime and tends to zero for highly 501 convective regimes. The characteristic equation for 502 the solutal boundary layer is then 40,42

$$\frac{D}{\delta V} = 1 + \frac{U_{||}}{V} \tag{2}$$

504 and the effective partition coefficient k_{eff} initially 505 expressed as a function of Δ then becomes, for an 506 upward component of fluid flow:

$$k_{eff} = \frac{k}{1 - \frac{1 - k}{1 + \frac{U_{||}}{1 + \frac{U_{||}}{$$

The evolution of $k_{\rm eff}$ with pulling rate based on 510 Fig. 3 data is presented in Fig. 5. Using the fluid 511 flow velocities from Fig. 3 (right y axis), $k_{\text{eff}} = 1$ 512 corresponds to a component U| = 0, obtained for a 513 pulling velocity of $1.9 \,\mu\text{m/s}$, which yields similar 514 average primary spacing on the ground and in 515 microgravity. For higher pulling rates, $k_{\rm eff}$ 516 decreases rapidly with V in the vicinity of $k_{\text{eff}} = 1$ (i.e., $V \approx 1.9 \ \mu m/s$) before a very slow decrease 517 around $k_{\text{eff}} \cong 0.2$, which is still far from a complete 518 mixing that would correspond to $k_{\text{eff}} = k = 0.07$, as assessed experimentally.⁴³ The k_{eff} determined 519 520 through the ratio of critical velocities is additionally 521 522 represented in Fig. 5. It should be remembered that 523 it corresponds to a range of pulling velocities where 524 the interface is convex. The progressive increase of



Primary spacing selection mechanisms have been 531 extensively studied analytically and experimen-532 tally, mostly in thin sample geometries.⁴⁴ In the 533 specific case of the current DECLIC-DSI experi-534 ments, previous 3D PF simulations based on the 535 FTA had so far fallen short of accurately predicting 536 primary spacing.⁸⁻¹⁰ The region between the upper 537 and lower bars in Fig. 6 shows the stable spacing 538 ranges predicted by PF simulations using a similar 539 procedure as in previous studies^{9,10,49} Stable spacing 540 ranges are determined by the FTA as the steady-541 state thermal gradient is only weakly affected by 542 the different thermal representations. All experi-543 mentally measured average spacings (squares) fall 544 within the calculated stable spacing ranges. 545 Although stability limits are similar for different 546 thermal representations, the dynamically selected 547 spacings within these limits in spatially extended 548 simulations using the TFC (circles) are larger than 549 with the FTA (diamonds). Hence, the more accurate 550 TFC significantly improves the agreement between predicted and measured spacings. Nevertheless, experimental spacings remain slightly larger than in spatially extended PF simulations. This could be linked to additional uncertainties on exact alloy or processing parameters, as well as to a size effect in the PF simulations, for which only about seven cells 557 558 in total were present in spatially extended TFC simulations for $V = 2 \mu m/s$. 559

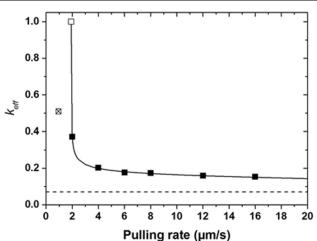


Fig. 5. Evolution of the effective partition coefficient keff as a function of pulling rate: x, estimated from the ratio of critical velocity under microgravity and on the ground; I, estimated using experimental data on fluid flow velocity, where the solid line represents Eq. 3 for $V \geq$ 1.9 $\mu m/s,$ with the boundary condition k_{eff} = 1 for V = 1.9 $\mu m/s$ (T) when fluid flow is zero. The dashed line corresponds to k = 0.07assessed experimentally⁴³ (G = 19 K/cm).

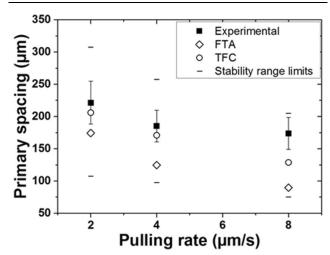
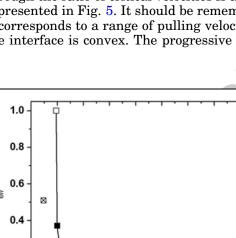


Fig. 6. Experimental average primary spacings with standard deviations (error bars) measured in microgravity experiments compared with selected average primary spacings obtained in PF simulations with a large system and different thermal conditions. Limits of calculated stable spacing ranges are plotted as short horizontal lines. All observed spacings in experiments and simulations fall inside the stable spacing ranges (G = 19 K/cm).



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CONCLUSION

561 In this article, we presented results obtained with the Directional Solidification Insert of the DECLIC 562 563 facility, dedicated to in situ and real-time observa-564 tion of the solid-liquid interface during directional 565 solidification of bulk transparent alloys. Solidifica-566 tion experiments were performed onboard the ISS to 567 image the whole dynamics of microstructure forma-568 tion and evolution under diffusive growth condi-569 tions. After the return of the DSI on the ground, 570 comparative experiments were performed using the 571 same control parameters to assess the influence of 572 convection.

73 The comparison of microgravity and ground 74 experiments highlighted the major effects of ther-75 76 77 mal convection at the macroscopic scale (both radial and axial macrosegregation) and at the microscopic scale (modification of the average primary spacing). 78 In the bulk transparent system studied, the convec-79 tion originates from a radial thermal gradient, 580 which is highly sensitive to latent heat release. We 581 demonstrated that the atypical effect of convection 582 on primary spacing (continuous increase of spacing 583 with pulling rate) compared with previous studies conducted in metallic systems^{15,35} stems from con-584 585 vection characteristics and, especially, from its 586 sensitivity to latent heat release.

587 The reduced gravity environment of space pro-588 vides unique conditions to obtain quantitative 589 benchmark data with homogeneous patterns. By 590 performing a quantitative comparison of those 591 unique data and 3D phase-field simulations, we 592 have demonstrated that accurate modeling of the 593 thermal field within the adiabatic zone is essential 594 to predicting quantitatively both the transient recoil 595 of the planar interface and subsequent dynamical 596 selection of the primary spacing. This is especially 597 relevant to transparent organic compounds in bulk 598 configuration.

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Convection Effects During Bulk Transparent Alloy Solidification in DECLIC-DSI and Phase-Field Simulations in Diffusive Conditions

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