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We experimentally and numerically investigate the effect of wind forcing on the spectral dynamics of Akhmediev breathers, a wave-type known to model the modulation instability. We develop the wind model to the same order in steepness as the higher order modification of the nonlinear Schrödinger equation, also referred to as the Dysthe equation. This results in an asymmetric wind term in the higher order, in addition to the leading order wind forcing term. The derived model is in good agreement with laboratory experiments within the range of the facility's length. We show that the leading order forcing term amplifies all frequencies equally and therefore induces only a broadening of the spectrum while the asymmetric higher order term in the model enhances higher frequencies more than lower ones. Thus, the latter term induces a permanent upshift of the spectral mean. On the other hand, in contrast to the direct effect of wind forcing, wind can indirectly lead to frequency downshifts, due to dissipative effects such as wave breaking, or through amplification of the intrinsic spectral asymmetry of the Dysthe equation. Furthermore, the definitions of the upand downshift in terms of peak- and mean frequencies, that are critical to relate our work to previous results, are highlighted and discussed.

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11 I. INTRODUCTION

The interaction between wind and ocean waves lies at the heart of ocean dynamics, and ¹³ since the 1950's^{1,2} significant progress has been made in understanding this interplay³⁻⁷. ¹⁴ More recently, a notable number of studies has been devoted to understanding the occurrence ¹⁵ of rogue waves^{8? -10}. Rogue waves are 'monster waves' that have a much higher amplitude ¹⁶ than what can statistically be expected from the current sea-state. The role of wind in ¹⁷ the formation and the evolution of such extreme waves has been investigated numerically ¹⁸ and experimentally¹¹⁻¹³. However, the direct effect of the wind on rogue wave spectra still ¹⁹ needs further investigation. An interesting physical phenomenon related to the wind-wave ²⁰ interaction is the frequency downshift observed in ocean waves^{14,15}. Apart from altering ²¹ a fundamental property of a wave field, namely its frequency, downshift can affect the ²² formation of rogue waves^{16,17}. Yet, a *unified* explanation for the physical origin of downshift ²³ seems to be missing, since proposed mechanisms such as the presence of wind^{18–20}, wave ²⁴ breaking^{21,22} and viscosity^{23,24} are opposite in nature as they respectively force and damp ²⁵ the waves. This study aims to clarify this downshift paradox.

Investigating up- and downshift needs a precise definition. To do so, two approaches can 27 be outlined²⁴⁻²⁶. The first is to look at a shift in position of the mode with the highest 28 amplitude in the spectrum, the *spectral peak* f_p . In doing so, any small asymmetries in the 29 spectrum are disregarded. The second approach is to look at the position of the *spectral* 30 mean f_m , defined as the ratio of the momentum M of the envelope to its norm N. In the 31 time domain, these are defined in Carter and Govan²⁴. In the spectral domain, this can be 32 written as

$$M = \sum_{n=-n_{\rm lim}}^{n=+n_{\rm lim}} f_n |\hat{a}_{f_n}|^2$$
(1)

$$N = \sum_{n=-n_{\rm lim}}^{n=+n_{\rm lim}} |\hat{a}_{f_n}|^2$$
(2)

$$f_m = \frac{M}{N} \tag{3}$$

³³ where \hat{a} is the Fourier mode of the envelope, f the frequency, and n_{lim} the spectral mode ³⁴ that marks the limit of the main carrier wave mode, such that the interval of f excludes the ³⁵ bound waves. The boundary betweeen the main mode an the first harmonic is considered ³⁶ to be at $f_{\text{lim}} = f_0 + f_0/2$. From the spectral definition one can see that M/N is equal to ³⁷ the quadratic weighted average and detects any asymmetry in the spectrum.

In addition, up-/downshifts can be reported as temporary or permanent. In the former 39 case, the mean or the peak shifts to a higher/lower frequency, but eventually shifts back 40 to the original carrier wave frequency. This recovery does not occur in case of permanent 41 up-/downshift.

In this paper, we aim to investigate the effect of the wind on spectra of unidirectional, 42 43 modulationally unstable gravity wave trains, specifically on Akhmediev breathers. Note 44 that when comparing to the ocean, this excludes directionally spread waves. Our main ⁴⁵ finding is to inject the wind terms found by Brunetti and Kasparian²⁷ into existing MNLS²⁸ ⁴⁶ model with viscous dissipation²⁴. This forms a forced-damped MNLS model in which the ⁴⁷ wind, dissipation and nonlinearity are all of the same order. In section II, we present our 48 model for this purely forcing effect of wind on water waves, that is, wind blowing in the ⁴⁹ direction of the wave propagation and consequently adding energy to the system. The wind 50 contribution in the model consists of a leading order forcing term that amplifies all wave ⁵¹ frequency components equally, as well as a higher order asymmetric term that amplifies ⁵² higher frequencies more than lower ones. In sections III and IV, we corroborate the model ⁵³ with wave tank experiments. Subsequently, we perform long range simulations in section V 54 to overcome the limited fetch in the experiment. We demonstrate that the wind forcing by ⁵⁵ itself can only cause an upshift in the spectral mean. Finally, in section VI, we clear up the ⁵⁶ aforementioned downshift paradox, discussing our model in light of previous literature, and 57 taking into account the two downshift definitions.

58 II. WEAKLY NONLINEAR WIND-WAVE MODEL

In a simplified Euler approximation for the water-wave problem, the Coriolis term is neglected with respect to the advective term in the momentum equations. In addition, the water density is considered constant $\rho_{\rm w} = \rho_{\rm w,0}$. Considering unidirectional waves, the transverse coordinate is neglected, and the system of equations that has to be solved is^{29,30}:

 ϕ_z

$$\nabla^2 \phi \equiv \phi_{xx} + \phi_{zz} = 0 \qquad \qquad -h \le z \le \eta(x,t) \tag{4}$$

$$= 0 z = -h (5)$$

$$\eta_t + \phi_x \eta_x - \phi_z = 2\nu \eta_{xx} \qquad \qquad z = \eta(x, t) \tag{6}$$

$$\phi_t + \frac{1}{2}\phi_x^2 + \frac{1}{2}\phi_z^2 + g\eta = -\frac{P}{\rho_w} - 2\nu\phi_{zz} \qquad z = \eta(x,t)$$
(7)

⁶³ where g is the gravitational acceleration, h is the water depth, x is the propagation direc-⁶⁴ tion and z is the upward coordinate. The potential function $\phi = \phi(x, z, t)$ is defined as ⁶⁵ $u = \partial \phi / \partial x = \phi_x$, $w = \partial \phi / \partial z = \phi_z$, where u and w are the velocity components. The ⁶⁶ function $\eta(x, t)$ is the surface elevation with respect to the average level z = 0. The kine-⁶⁷ matic viscosity of the fluid is denoted as ν [m²/s]. The wind is characterized by P = P(x, t), ⁶⁸ the excess pressure at the water surface in the presence of wind, that in the framework of ⁶⁹ the Miles mechanism¹ is given by

$$\frac{P}{\rho_{\rm w}} = \frac{\Gamma}{f_0} \frac{c_p^2}{2\pi} \eta_x = \frac{\omega_0}{k_0^2} \Gamma \eta_x \tag{8}$$

ro where c_p is the phase velocity of the carrier wave, k_0 its wave number, f_0 its frequency, r_1 and $\omega_0 = 2\pi f_0$ its angular frequency. Γ [1/s] is the growth rate of the energy E of the r_2 waves due to the wind blowing in the direction of the wave propagation, that is $\frac{\partial E}{\partial t} = \Gamma E$. r_3 In turn, an expression for Γ as a function of wind speed U and wave frequency f_0 can be r_4 modeled in various ways, as will be discussed in section III. In the approximation that the r_5 envelope varies slowly in comparison to the surface elevation, the Method of Multiple Scales r_6 (MMS) can be used to expand these boundary conditions and obtain a weakly nonlinear r_7 propagation equation for the envelope a(x,t) at each order of interest. In the MMS, the r_8 small order parameter is the steepness of the wave $\epsilon = ak_0$. In addition, deep water waves, $r_9 k_0 h \to \infty$, are considered. The surface elevation to the first order in steepness is given by

$$\eta(x,t) = \operatorname{Re}\{a(x,t)\exp[i(k_0x - \omega_0 t)]\}$$
(9)

Note that the sign choice of the argument in the exponential is important for the spectral representation of envelope with respect to the carrier wave. For instance in Carter and Govan²⁴ the opposite sign choice has been made. To avoid confusion, the spectrum of the scarrier wave is plotted throughout this paper. In the absence of both wind and viscosity

⁸⁴ ($\nu = \Gamma = 0$), the MMS yields (i) the nonlinear Schrödinger (NLS) equation when developing ⁸⁵ the boundary conditions to $O(\epsilon^3)^{31-33}$, and (ii) the modified NLS (MNLS) or Dysthe equation ⁸⁶ when developing the boundary conditions to $O(\epsilon^4)^{28}$. Assuming the wind Γ/f_0 and the ⁸⁷ viscosity $\nu k_0^2/f_0$ contribution are both of order $O(\epsilon^2)$, and including these in the MMS ⁸⁸ development up to $O(\epsilon^4)$ using a similar method as Carter and Govan²⁴, yields the following ⁸⁹ damped-forced MNLS:

$$\frac{\partial a}{\partial t} + \frac{\omega_0}{2k_0} \frac{\partial a}{\partial x} = \epsilon \left[-i \frac{\omega_0}{8k_0^2} \frac{\partial^2 a}{\partial x^2} - \frac{1}{2} i k_0^2 \omega_0 a |a|^2 - 2k_0^2 \nu a + \frac{1}{2} \Gamma a \right]
+ \epsilon^2 \left[4i k_0 \nu \frac{\partial a}{\partial x} - \frac{3i}{4k_0} \Gamma \frac{\partial a}{\partial x} - \frac{3i}{4k_0} \Omega \frac{\partial a}{\partial x} - \frac{3i}{4k_0} \Omega \frac{\partial a}{\partial x} + \frac{\omega_0}{16k_0^3} \frac{\partial^3 a}{\partial x^3} - i k_0 a \frac{\partial \bar{\phi}}{\partial x} \right]$$
(10)

⁹⁰ where $\bar{\phi}$ is the potential mean flow. To obtain a propagation in space, in accordance with ⁹¹ the motion in a 1D wave tank, the following coordinate transformation is applied:

$$\begin{cases} \tilde{t} = t - \frac{2k_0}{\omega_0} x \\ \tilde{x} = \epsilon x \end{cases}$$

The potential mean flow term in (10) is replaced by a Hilbert transform term, see ⁹³ Janssen³⁴. The Hilbert transform \mathscr{H} is defined as $\mathscr{F}[\mathscr{H}[u]] = -i \operatorname{sign}(\omega) \mathscr{F}[u]$, where ⁹⁴ \mathscr{F} is the Fourier transform. The damped-forced MNLS becomes:

$$\underbrace{\frac{\partial a}{\partial \tilde{x}} + i\frac{k_0}{\omega_0^2}\frac{\partial^2 a}{\partial \tilde{t}^2} + ik_0^3 a|a|^2}_{\text{NLS}} = \underbrace{+\epsilon\frac{k_0^3}{\omega_0}\left(6|a|^2\frac{\partial a}{\partial \tilde{t}} + 2a\frac{\partial|a|^2}{\partial \tilde{t}} + 2ia\mathscr{H}\left[\frac{\partial|a|^2}{\partial \tilde{t}}\right]\right)}_{\text{MNLS correction}} -\frac{4\frac{k_0^3}{\omega_0}\nu a - \epsilon^2 0i\frac{k_0^3}{\omega_0^2}\nu\frac{\partial a}{\partial \tilde{t}}}_{\text{Viscosity}} + \underbrace{\frac{k_0}{\omega_0}\Gamma a + \epsilon 4i\frac{k_0}{\omega_0^2}\Gamma\frac{\partial a}{\partial \tilde{t}}}_{\text{Wind}}$$
(11)

This full model consists of the corrected NLS, taking into account higher-order dispersion and mean flow (MNLS correction) as well as viscosity and wind effects, where the higher order terms are indicated by ϵ . In the following the tilde's are omitted. The dynamics of the MNLS equation are well known and have been studied numerically by Lo and Mei³⁵? ⁹⁹ and applied to many experiments, see for instance Trulsen and Stansberg³⁶, Tulin and ¹⁰⁰ Waseda³⁷, Slunyaev *et al.*³⁸, Chabchoub *et al.*³⁹.

The viscosity consists of a linear damping term $-4\frac{k_0^3}{\omega_0}\nu a$ in the leading order, as proposed 102 in Dias, Dyachenko, and Zakharov²⁹. In addition, following Carter and Govan²⁴, the higher 103 order viscosity term is given by $-\epsilon 20i\frac{k_0^3}{\omega_0^2}\nu\frac{\partial a}{\partial t}$.

The wind action appears as a leading order linear forcing term $\frac{k_0}{\omega_0}\Gamma a$ which in past work 104 has been included based on either the intuition of a simple forcing⁴⁰, or through more rigorous 105 justification³⁰. There has been no experimental validation of a wind forcing term in an 106 evolution equation for deep water waves. Our addition of a higher order wind term $\epsilon 4i \frac{k_0}{\omega^2} \Gamma \frac{\partial a}{\partial t}$ 107 completes the wind-wave model (11). This higher order wind term restores consistency in 108 the sense that since the nonlinearity is developed up to $O(\epsilon^4)$ in the MNLS framework, the 109 ¹¹⁰ dissipation and forcing also have to be developed up to this order to have a coherent model. The model reported here can be related to the wind-wave model presented in Brunetti 111 and Kasparian²⁷, Brunetti *et al.*⁴¹ by noting that the higher order wind term can be ob-112 tained in two ways. To obtain (11), the wind forcing Γ/f_0 was assumed $O(\epsilon^2)$ in (7), and 113 the kinematic and dynamic boundary conditions were developed up to $O(\epsilon^4)$ in the MMS. 114 Alternatively, following Brunetti *et al.*⁴¹, one can assume that $\Gamma/f_0 = O(\epsilon)$ in (7), thus, 115 strong in comparison to the steepness ϵ . Next, the kinematic and dynamic boundary condi-117 tions only have to be developed up to $O(\epsilon^3)$ in the MMS to obtain the same wind terms as $_{118}$ in (11).

¹¹⁹ Surprisingly, the outcome of the MMS is that the viscosity and wind contributions in ¹²⁰ both the leading and higher order have the same form in (11), despite of their non-similar ¹²¹ appearance in the kinematic and dynamic boundary conditions (6 and 7), respectively. As ¹²² these terms have opposite signs, the viscosity can cancel out the wind forcing, and vice versa. ¹²³ However, note that their balance is different for different orders. This balance occurs as

$$\delta_0 \equiv \Gamma - 4k_0^2 \nu = 0 \tag{12}$$

 $_{124}$ for the leading order, as studied by Kharif *et al.*³⁰. In our model (11), the higher order $_{125}$ terms are balanced if

$$\delta_1 \equiv \Gamma - 5k_0^2 \nu = 0 \tag{13}$$

The NLS is a spectrally symmetric evolution equation, that is, as the envelope propagates, 127 the frequencies below f_0 evolve in the same way as their counterparts above f_0 . An odd

¹²⁸ derivative in time is needed to have an asymmetric evolution of the spectrum³⁵. In the full ¹²⁹ model (11), the odd terms are the MNLS correction, the higher order wind and the higher ¹³⁰ order viscosity term. Indeed, the leading order term δ_0 has the effect of either damping or ¹³¹ amplifying all frequencies in the spectrum, depending on its sign. For the higher order, if ¹³² viscosity is dominant, $\delta_1 < 0$, the higher frequencies $(f > f_0)$ are damped more than lower ¹³³ ones $(f < f_0)$ causing a permanent downshift of the spectral mean, as evidenced by Carter ¹³⁴ and Govan²⁴. Our derivation of the higher order wind term shows that if the wind action ¹³⁵ is dominant, $\delta_1 > 0$, the higher order effect is opposite and the spectral mean is upshifted.

136 III. EXPERIMENTAL AND COMPUTATIONAL DETAILS

137 A. Initial condition

To validate our model, we are interested in cases in which the system experiences a ¹³⁹ spectral broadening. To model the modulation instability, the Akhmediev Breather $(AB)^{42}$ ¹⁴⁰ has been used to generate our initial conditions for the experiment. Starting the dynamics ¹⁴¹ from an exact AB expression is useful for experimental investigations, since it allows to ¹⁴² trigger the modulation instability dynamics in relation to the length of the facility. The AB ¹⁴³ gives an approximate prediction of the growth and subsequent decay cycle that is to occur, ¹⁴⁴ rather than just a prediction of a linear growth rate as calculated from the Benjamin-Feir ¹⁴⁵ instability analysis⁴³. This theoretical expectation allows us to identify deviations from this ¹⁴⁶ growth-decay pattern. The AB is a solution to the *NLS* part of (11) and reads⁴⁴

$$a_{\rm AB}(x^*, t) = a_0 \frac{\sqrt{2A}\cos(\Omega\frac{t}{T_0}) + (1 - 4A)\cosh(R\frac{x^*}{L_0}) + iR\sinh(R\frac{x^*}{L_0})}{\sqrt{2A}\cos(\Omega\frac{t}{T_0}) - \cosh(R\frac{x^*}{L_0})}\exp(i\frac{x^*}{L_0}), \qquad (14)$$

¹⁴⁷ where x^* is the distance to the focal point, $L_0 = (k_0^3 a_0^2)^{-1}$ and $T_0 = \sqrt{2}(k_0 \omega_0 a_0)^{-1}$ are ¹⁴⁸ the rescaling coordinates, $\Omega = 2\sqrt{1-2A}$ is the dimensionless modulation frequency, ¹⁴⁹ $R = \sqrt{8A(1-2A)}$ the growth rate, and 0 < A < 0.5 is the Akhmediev parameter. ¹⁵⁰ The case A = 0.25 corresponds to a maximal growth rate in the Benjamin-Feir theory. Note ¹⁵¹ that since the AB is a solution of the NLS only, the maximally unstable mode in the NLS ¹⁵² framework does not exactly coincide with that of our full system. When $x \to \pm \infty$ the AB ¹⁵³ tends to a regular wave train, while at the focal point x = 0, the breather reaches maximal

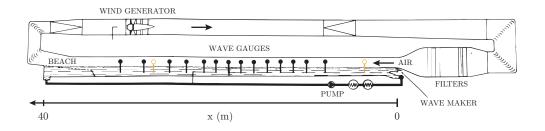


FIG. 1: IRPHE wind-wave tank facility: 40 m long, 3 m wide, water depth 0.80 m, air channel height 1.5 m. Wave gauges have been placed approximately evenly along the tank.

Capacitive wave gauges are indicated with open circles, others are resistive gauges

¹⁵⁴ temporal compression and consequently maximal amplitude. In the spectral domain, the
¹⁵⁵ focusing of the breather corresponds to a broadening of the spectrum, see Wetzel *et al.*⁴⁵.
¹⁵⁶ Thus, the focal point is marked by maximal spectral width, and by minimal amplitude of
¹⁵⁷ the carrier wave Fourier component.

¹⁵⁸ Choosing an initial condition at a given distance from the focal point allows us to control ¹⁵⁹ the number of developed sidebands. The signal given to the wave maker is the surface ¹⁶⁰ elevation $\eta(x^*, t)$ (9), based on the dimensional AB envelope $a_{AB}(x^*, t)$ (14).

¹⁶¹ B. Experimental Setup

Experiments have been performed in the closed wind-wave facility at IRPHE (Luminy)-162 ¹⁶³ Aix Marseille University. A schematic depiction of the facility is shown in figure 1. As detailed in Coantic et al.⁴⁶, the tank is 3 m wide, has a water depth of 80 cm, and a 164 length of 40 m. At the end of the wave tank, an 8 m sloping beach was installed to prevent 165 wave reflection. At the beginning of the tank, a 1.5 m long plastic sheet floating on the 166 water surface allowed the incoming wind to be tangential to the water surface, and damped 167 possible high-frequency mechanical wave modes. Mechanical waves have been generated 168 by an underwater piston-like wavemaker controlled by a computer. The system was able 169 to produce arbitrary surface gravity waves in the frequency range of 0.5-1.8 Hz. The air-170 channel above the tank is 1.5 m high. The wind was generated by a closed-loop air flow 171 $_{172}$ system, up to a maximum velocity of U = 15 m/s in the direction of the wave propagation. ¹⁷³ A total of 16 wave gauges have been placed at fixed positions along the tank to measure the 174 surface elevation. The gauges were positioned approximately evenly between x = 3 m and $_{175} x = 32$ m. The first and 14^{th} gauge were capacitive-, the others were resistive-type wave $_{176}$ probes. All gauges had a sampling rate of 400 Hz. The wind speed was measured by a pitot $_{177}$ tube at different positions in the tank to verify a constant and homogeneous air flow.

The parameter space of the experiment is limited in several directions. Firstly, the 179 steepness of the background plane wave of the AB (when $x^* \to -\infty$) is constrained to a 180 range of $0.08 \leq \epsilon \leq 0.10$. The lower the steepness, the larger the propagation distance 181 required for the modulation instability to develop in the small fetch, restricted by the size of 182 the facility (figure 1). Conversely, if the carrier waves are initially too steep, wave breaking 183 is inevitable as a consequence of focusing. Secondly, all measurements have been performed 184 with $f_0 = 1.67$ Hz, thus $\omega_0 = 10.5$ rad Hz, and $k_0 = 11.2$ rad/m. This yields a Bond number 185 of ~1000, confirming these waves are in the gravity wave regime⁴⁷. Higher frequencies waves 186 were not possible to generate with the installed wavemaker. Thirdly, wind speeds were 187 limited, as for $U \gtrsim 4$ m/s breaking would occur for waves generated with initial steepness 188 in the described range.

Besides the breather-type waves generated by the wavemaker, the wind naturally gener-190 ates additional waves. The frequency range of these wind-waves shifts down as a function 191 of wind speed and fetch. At the last wave gauge, for U = 4 m/s, the wind-waves have been 192 measured to be in the range of approximately 2.2 - 3.2 Hz. Therefore, these waves were 193 considered not to overlap with the breather-type waves. The wind-generated waves showed 194 micro-breaking at wind speeds of $U \sim 4$ m/s towards the end of the tank.

¹⁹⁵ C. Simulation parameters

The envelope *a*, which is periodic in time, was integrated forward in space according ¹⁹⁷ to (11) by means of a split-step Fourier scheme⁴⁸. As described in Agrawal⁴⁹, the linear ¹⁹⁸ and nonlinear part of (11) can be solved in separate steps. The linear part is an ordinary ¹⁹⁹ differential equation (ODE) which was solved in Fourier space, and to integrate the nonlinear ²⁰⁰ part, the fourth order Runge-Kutta method has been used.

This initial condition for the simulations is the envelope based on the surface elevation 202 measured by the first wave gauge. To ensure periodic boundary conditions, we selected a 203 time interval equal to a multiple of the envelope period (see figure 2). As such, the spectral 204 resolution of the simulation is the same as that in the experiment. At an acquisition rate of

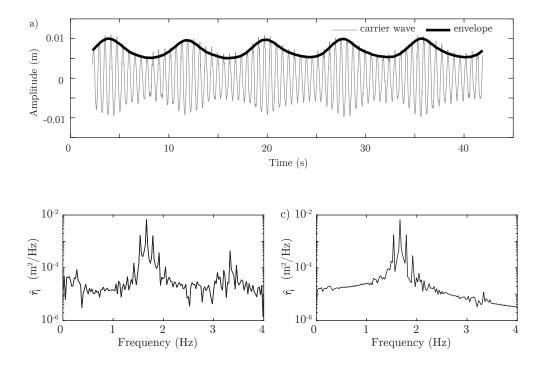


FIG. 2: a) Periodic initial condition used for the simulations, as measured by the first wave gauge at x = 3.7 m at an acquisition rate of 400 Hz. b) Log-linear plot of the spectrum of

the surface elevation as measured by the first wave gauge. c) Log-linear plot of the spectrum of the carrier wave to first order as calculated from the envelope. Carrier wave parameters $\epsilon = 0.08$, T = 0.6 s; AB parameters $x^* = -30$ m and A = 0.25. The envelope was calculated using the Hilbert transform.

²⁰⁵ 400 Hz, depending on the exact conditions, the time series were ~ 40 s long, corresponding ²⁰⁶ to ~ 60 periods of the carrier wave and spectral resolution of ~ 0.025 Hz. The complex ²⁰⁷ envelope was extracted from the real valued envelope η as follows

$$a(x_0, t) = \left[\eta(x_0, t) + i\tilde{\eta}(x_0, t)\right] e^{-i(k_0 x - w_0 t)}$$
(15)

where $\tilde{\eta}$ is the Hilbert transform of the surface elevation, as described in Osborne⁵⁰ and $x_0 =$ 209 3.1 m is the position of the first wave gauge. Starting the simulations from the experimental 210 signal accounts for the possible imperfections of the wavemaker and reduces the number 211 of free parameters in the model. The only free parameters to be determined are the wind 212 growth rate Γ and the viscosity parameter ν , both of which are fitted from the experiment. 213 Following the method of Segur *et al.*²³ and Carter and Govan²⁴, the viscosity parameter

Wind speed	$\Gamma_{ m Miles}$	$\Gamma_{ m Simulation}$	leading order	higher order
4 m/s	$8.3 \times 10^{-3} {\rm s}^{-1}$	$8.5 \times 10^{-3} \mathrm{s}^{-1}$	$\delta_0 > 0$	$\delta_1 > 0$
$2 \mathrm{m/s}$	$1.6 \times 10^{-3} {\rm s}^{-1}$	$2.0 \times 10^{-3} {\rm s}^{-1}$	$\delta_0 < 0$	$\delta_1 < 0$

TABLE I: Predicted values and values used in the simulations for the wind growth rate Γ

²¹⁴ ν in (11) is in fact an effective term that includes not only viscosity, but all sources of ²¹⁵ dissipation in the experiment such as the side-wall effects and surface contamination. It ²¹⁶ has been determined experimentally by fitting the exponential decay of a the wave train ²¹⁷ propagating down the tank without wind forcing: $E = E_0 e^{-\frac{1}{2}\zeta x}$, where $\zeta = (4k_0^3/\omega_0)\nu$. All ²¹⁸ waves used in this work are of the same frequency and steepness, for which we found a ²¹⁹ measured value of the viscosity of $\nu = (1.18 \pm 0.35) \times 10^{-5} \text{ m}^2/\text{s}$. Consequently, we used ²²⁰ $\nu = 1 \times 10^{-5} \text{ m}^2/\text{s}$ in all simulations.

To determine Γ we rely on predictions by the Miles mechanism for wind growth⁵¹

$$\Gamma = \omega_0 \alpha \frac{\rho_{air}}{\rho_w} \frac{u_*^2}{c_p^2} \tag{16}$$

where u_* is the friction velocity, c_p is the phase velocity, ρ_{air} the density of air, ρ_w the density of water, and α an empirical parameter of 32.5. Assuming a logarithmic profile of U as a function of z, the measured wind speed U is related to the friction velocity u_* by¹

$$U(z) = \frac{u_*}{K} \log\left(\frac{z}{z_0}\right),\tag{17}$$

²²⁵ where z is the height where U is measured, K = 0.41 the Von Karman constant and z_0 is ²²⁶ the effective roughness length

$$z_0 = \kappa u_*^2/g,\tag{18}$$

²²⁷ with $\kappa = 0.0144$ the Charnock constant.

Since estimations for Γ can have deviations up to a factor two with field and wave tank measurements (figure 1 in Banner and Song⁵¹), we adjusted the latter value to match the spectral widening and growth in the experiment for each wind speed, see table I. Values obtained for u_* through (17) are accurate within a range of 10 % compared to those measured in an elaborate study by Caulliez, Makin, and Kudryavtsev⁵² in the same facility.

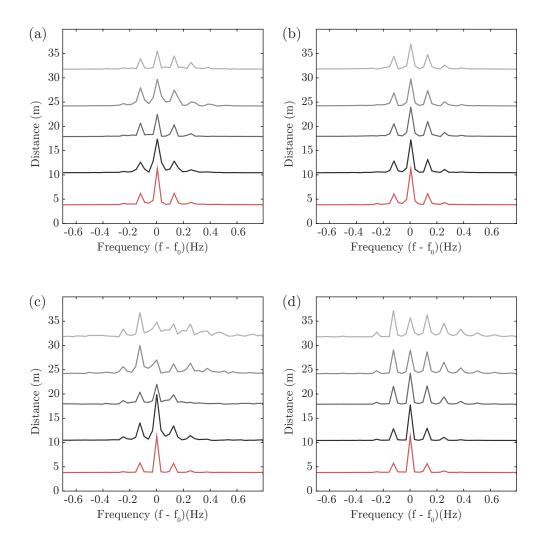


FIG. 3: Spectral amplitude $|\hat{\eta}|$ (normalized) offset by distance, for carrier wave parameters $\epsilon = 0.08$, $T_0 = 0.6$ s, AB parameters $x^* = -30$ m and A = 0.25, and simulation parameter $\nu = 1 \times 10^{-5} \,\mathrm{m}^2/\mathrm{s}$. (a,c) Experiments, (b,d) corresponding simulations. (a,b) U = 0 m/s, (c,d) U = 4 m/s. Initial condition (simulation) and measurement (experiment) at x = 3.1 m are indicated by the red line.

233 IV. EXPERIMENTAL RESULTS AND MODEL VALIDATION

Our numerical wind-wave model was compared to laboratory experiments on AB-type 235 waves. All results presented here are based on carrier wave parameters $T = \frac{2\pi}{\omega_0} = 0.6$ s, 236 $\epsilon = 0.08$ and AB parameter A = 0.25, the case for maximal modulation growth rate. For 237 the simulations, we set $\nu = 1 \times 10^{-5}$ m²/s and Γ as presented in table I.

Based on the analytical solution to the NLS, it is expected for the AB to reach its 238 239 maximal focusing after propagating 30 m from the wave maker ($x^* = -30$ m). Figure ²⁴⁰ 3 compares the evolution of the spectra retrieved from the experiment to the numerical $_{241}$ simulations, without (a,b) and with the presence of wind U = 4 m/s (c,d). Note that the ²⁴² initial spectrum is still symmetric, both with and without wind, i.e. the lower sideband $\hat{\eta}_{-1}$ ²⁴³ and the upper sideband $\hat{\eta}_{+1}$ have a similar amplitude. In the experiment without wind, see figure 3a, due to the MNLS correction^{36,38}, and taking into account the effect of viscosity, 244 the focal point is expected to be around x = 36 m. Note that the higher modes $\hat{\eta}_{+2}$ and $\hat{\eta}_{+3}$ 245 slightly grow towards the end of the tank. Indeed, our simulations in figure 3b reproduce 246 these features. In the presence of wind, U = 4 m/s, both the experimental measurements 247 in figure 3c and the simulations in figure 3d show an upstream shifting of the focal point, 248 ²⁴⁹ that is, $\hat{\eta}_0$ reaches a minimum around x = 28 m. In addition, under the wind action, the ²⁵⁰ spectrum broadens, albeit in an asymmetric manner as a broad range of higher frequencies $_{251} \hat{\eta} > \hat{\eta_0}$ grow, while only a narrow range of the lower frequencies $\hat{\eta} < \hat{\eta_0}$ do.

In order to characterize the appearance of the spectral asymmetry, and see the effect of 252 the wind on the shifting of the focal point, figure 4 displays the evolution of the carrier 253 wave Fourier component $\hat{\eta}_0$ as well as the first upper and lower side bands as a function of 254 propagation distance. With increasing wind speed the decay rate for $\hat{\eta}_0$ increases, equally, 255 the growth rate of $\hat{\eta}_{-1}$ and $\hat{\eta}_{+1}$ increases. Consequently, the crossing point, at which the 256 amplitude of the sidebands overtake $\hat{\eta}_0$, moves upstream with increasing wind speed. For 257 U = 4 m/s, since $\hat{\eta}_{-1}$ has a higher growth rate than $\hat{\eta}_{+1}$, a downshift of the spectral peak 258 ²⁵⁹ originates, as measured at the last wave gauge. The numerical simulations of our developed ²⁶⁰ model reproduce this shifting behavior. Similarly, the simulations matches the trend for the ²⁶¹ evolution of the total energy, in spite of the experimental fluctuations due to the inherent variability related to wind. Note that due to the width of the tank, a certain amount of this 262 fluctuation in total energy measured by each wave gauge can also vary due to for instance 263 reflected waves crossing the central line of the wave tank were the gauges were lined up, 264 causing fluctuations superimposed on the trend set by the wind and viscosity input. 265

Figure 5 shows the evolution of spectral peak f_p (dots) and the spectral mean f_m (solid 267 line) as a function of distance. The spectral peak remains equal to f_0 for U = 0 m/s, while 268 for U = 4 m/s a downshift of the peak occurs. Due to the length limitations of the tank 269 and the short considered fetch, we are not able to experimentally assess and state whether

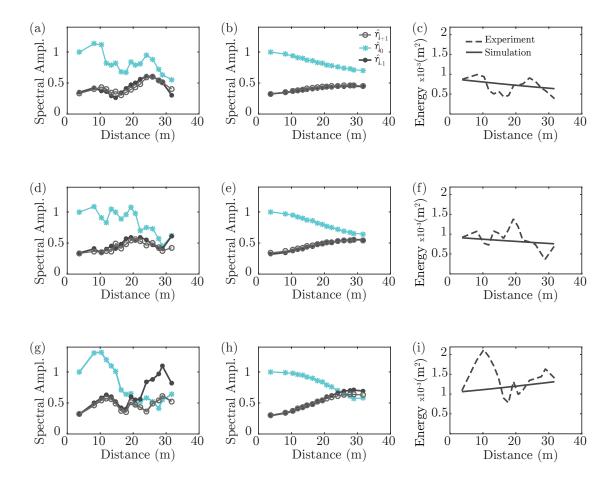


FIG. 4: Evolution of the three central Fourier components $\hat{\eta}_0$ (*), $\hat{\eta}_{+1}$ (\circ) and $\hat{\eta}_{-1}$ (\bullet), for the same parameters as in figure 3. (a,d,g) Experiments, (b,e,h) simulations. Values are normalized to the amplitude of $\hat{\eta}_0$ at the first wave gauge. (c,f,i) Energy evolution. (a,b,c) U = 0 m/s, (d,e,f) U = 2 m/s, (g,h,i) U = 4 m/s.

²⁷⁰ this shift is permanent or temporary. In contrast to the spectral peak, the spectral mean ²⁷¹ in figure 5 demonstrates a clear upshift. This dissimilar behavior illustrates the need for a ²⁷² careful definition of up- and downshift in order to allow for an accurate description of the ²⁷³ physics and dynamics at play.

Performing another experimental investigation for U = 2 m/s and for a shorter expected 275 focal distance of about 20 m from the wave maker ($x^* = -20$ m), as shown in figure 276 6, allows to quantify the spectral dynamics after the focal point. The spectral evolution 277 indeed shows the same downshift behavior in the spectral peak, and a broad growth of the 278 higher frequencies ($f > f_0$). However, for both of these features a reversion towards the

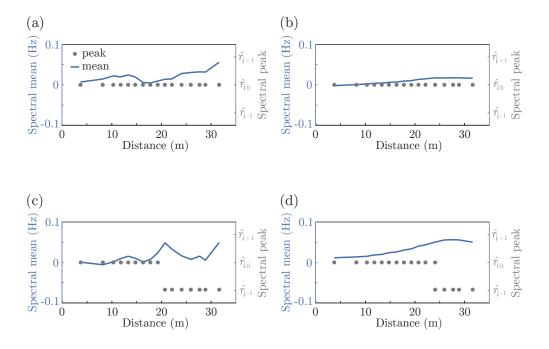


FIG. 5: Evolution of the spectral mean f_m (blue solid line, left vertical axis) and spectral peak f_p (gray dots, right vertical axis), for same parameters as in figure 3. (a,c) Experiments, (b,d) simulations. (a,b) U = 0 m/s, (c,d) U = 4 m/s. Two dots at the same distance indicates the spectral heights are within 1 percent range.

279 initial condition occurs at the end of the tank. The same is observed in figure 9 in Tulin ²⁸⁰ and Waseda³⁷ or figure 4 in Chabchoub *et al.*³⁹. Indeed, for the homogeneous NLS, the ²⁸¹ Fermi-Pasta-Ulam recurrence⁵³ predicts a cyclic pattern after which the initial condition ²⁸² is completely recovered and then repeats several times. However, the MNLS framework ²⁸³ predicts only a near-recovery^{35,37}, that is, a quasi-recurrence. With respect to the spectral mean, this quasi-recurrence causes an increase of the value of $f_{\rm m}$ near the focal point and a decrease to the original value when the cycle is finished. As a consequence, the spectral mean 'oscillates' if multiple quasi-recurrence cycles occur. The same oscillatory behavior is 286 observed for the spectral peak as it shifts down and recovers. Together with the long range 287 simulations in section V, this reveals that the spectral peak downshift downshift due to wind 288 in the previous case, when the focal point was expected to be 30 m from the wave generator, 289 ²⁹⁰ is temporary too.

The upper limit of the model, with respect to the steepness, is reached when wave breaking 292 occurs. Figure 7 shows the spectral evolution for the expected focal point at 30 m from

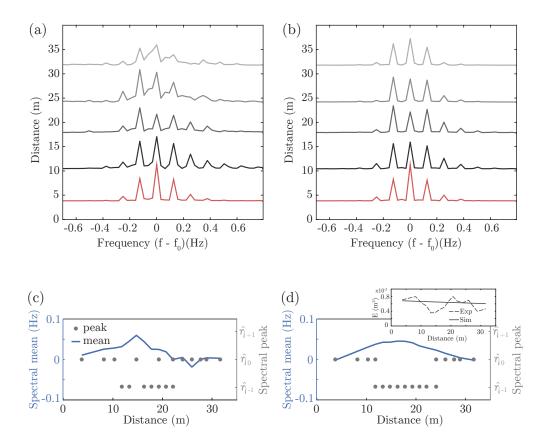


FIG. 6: (a) Spectral amplitude $|\hat{\eta}|$ (normalized) offset by distance at U = 2 m/s for (a) experiment and (b) simulations. Evolution of $f_{\rm m}$ (blue solid line) and $f_{\rm p}$ (gray dots) for experiment (c) and simulation (d). For the same parameters as in figure 3, except shorter focal distance $x^* = -20$ m. The inset shows the Energy evolution for the simulations and experiment

²⁹³ the wave maker $(x^* = -30 \text{ m})$ and a stronger wind of U = 6 m/s. Clearly, a downshift ²⁹⁴ is observed in the experiments, from which $\hat{\eta}_0$ does not recover. In the simulations, this ²⁹⁵ permanent downshift cannot be reproduced, instead, a quasi-recurrence occurred. This ²⁹⁶ disagreement can be attributed to the wave breaking that has in fact been observed to occur ²⁹⁷ around $x \approx 15$ m, and is not considered in the model.

This observation underlines the importance of dissipation associated with wave breaking to induce a permanent downshift for both the spectral peak and mean. To further exemplify this fact, figure 8a shows the experimentally measured spectral peak shifts down several nodes, up to $\hat{\eta}_{-3}$, for increasing wind speeds. A similar downward trend is observed for the

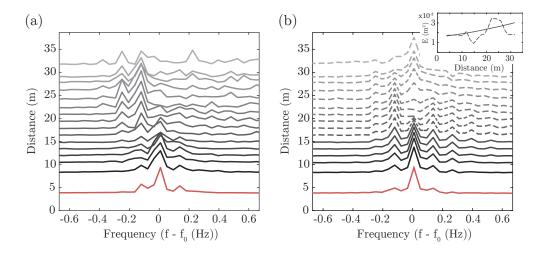


FIG. 7: Spectral amplitude $|\hat{\eta}|$ (normalized) offset by distance at U = 6 m/s for (a) experiment and (b) simulation, for the same parameters as in figure 3. The simulation is no longer accurate after the wave breaking event at $x \approx 15$ m, indicated by the dashed

lines. The inset shows the Energy evolution for the simulations and experiment.

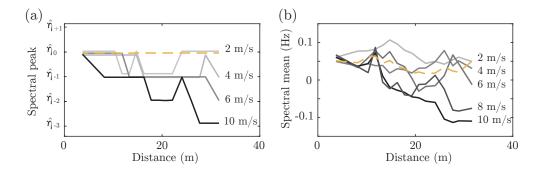


FIG. 8: Experimental evolution of (a) the spectral peak f_p and (b) the spectral mean f_m as a function of distance for increasing wind speeds, for same parameters as in figure 3, except shorter focal distance $x^* = -20$ m. A systematic downshift occurs for strong winds of (a) the spectral peak up to the position of $\hat{\eta}_{-3}$ and (b) the spectral mean. Dashed line indicates U = 0 m/s. Wavebreaking occurs for $U \ge 4$ m/s.

³⁰² spectral mean in figure 8b. For our carrier wave parameters, breaking occurs for U > 4 m/s, ³⁰³ and therefore cannot be compared to simulations.

³⁰⁴ While the focus of this work lies on the forcing aspect of the wind, that is, up to the

³⁰⁵ point of wave breaking, it is worth mentioning efforts to model the wave evolution after ³⁰⁶ the point of wave breaking. Kato and Oikawa ²² propose an ad-hoc higher order term that ³⁰⁷ activates at high steepness and leads to a permanent downshift of the peak. In a more ³⁰⁸ theoretically structured approach, Trulsen and Dysthe ^{21,40} add a symmetric source term ³⁰⁹ to the MNLS equation and observe a permanent downshift to the most unstable mode. ³¹⁰ However, applying this model to our data did not yield the permanent downshift observed ³¹¹ in for instance figure 7a. A rigorously derived model for the symmetric NLS is proposed ³¹² and validated by experimental data in Tulin and Li⁵⁴, Hwung *et al.* ⁵⁵, where a downshift ³¹³ to the most unstable mode is also observed. Here, the asymmetry is caused by an integral ³¹⁴ over the envelope. This model seems a promising candidate to simulate wave breaking in ³¹⁵ the framework of an MNLS equation.

Since the NLS inherently only applies to narrow banded spectra, it seems an open question whether a wave breaking term in an NLS-like evolution equation can account for a downshift multiple modes, as observed in figure 8a, since this means the spectrum is broadened significantly. For fully nonlinear simulations, efforts to simulate wave breaking have been successfully conducted by Tian, Perlin, and Choi^{56,57}.

In summary, we observe a temporary downshift of the spectral peak towards the lower szz satellite, and a temporary upshift of the mean. Our model (11) reproduces the qualitative features of the experimental results well. These include the correct length scale and magnitude of the spectral mean and peak shift, the crossing of the sidebands, and the broadening of the spectrum. Considering the variability inherent to wind experiments, a qualitative agreement on the spectral dynamics is the best one can expect. To overcome the length limit of wave tank and investigate multiple semi-recurrence cycles, long range simulations have been performed as described in the next section

329 V. LONG RANGE SIMULATIONS

Numerical simulations have been performed over a length of 100 m on AB with parameters $x^* = -20 \text{ m}, A = 0.25 \text{ and carrier wave parameters } T_0 = 0.6 \text{ s}, \epsilon = 0.1, \nu = 1 \times 10^{-5} \text{ m}^2/\text{s},$ $T_0 = 8.0 \times 10^{-3} \text{ s}^{-1}$ under wind forcing, in which case $\Gamma > 5\nu k_0^2$. To demonstrate the effect

of our higher order wind term, three simulation cases are compared:

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1. Absence of wind: Equation (11) including the MNLS correction and both leading and higher order viscosity terms, however without wind. In this case, $\delta_0 < 0$ and $\delta_1 < 0$.

2. Leading order wind: Equation (11) including the MNLS correction and both order viscosity terms, with only the leading order wind term, as posed in Kharif *et al.*³⁰, Trulsen and Dysthe⁴⁰, Onorato and Proment⁵⁸. In this case, $\delta_0 > 0$ and $\delta_1 < 0$.

340 3. Full model: Equation (11) including the MNLS correction and both order viscosity 341 terms, with both the leading order and the higher order wind term. In this case, $\delta_0 > 0$ 342 and $\delta_1 > 0$.

Figure 9 compares the envelope amplitudes for these three cases. In absence of wind, 343 ³⁴⁴ simulation (i), the viscosity attenuates the amplitude. In addition, as described by Kim-³⁴⁵ moun *et al.*⁵⁹, it induces a shift in the quasi-recurrence pattern of the envelope, as indicated ³⁴⁶ in figure 9a. Here, the dashed line is perpendicular on the gradient lines of the envelope amplitude. In simulation (ii), the wind amplifies the amplitude of the envelope with in-347 creasing fetch, as shown in figure 9b. In addition, the forcing cancels the shift caused by the 348 viscosity term as now only maxima occur on the dashed line in figure 9b. In line with results 349 reported in Kharif *et al.*³⁰, due to wind forcing and the increasing steepness, the position 350 of the focal point is moved upstream to x = 25 m, and every subsequent quasi-recurrence 351 cycle is more compressed in space compared to the previous one. The result of simulation 352 (iii) is similar to that of simulation (ii). The consideration of the additional higher order 353 wind term causes a slight further increase of the envelope's amplitude, and shortens the 354 length of the quasi-recurrence cycle: figure 9c now fits 3 maxima of the envelope. Figure 9d 355 displays the corresponding evolution of the normalized amplitude spectrum. Several MNLS 356 quasi-recurrence cycles can be observed in which the spectrum broadens asymmetrically and 357 narrows again. This quasi-recurrence pattern is superimposed on the general broadening of 358 the spectrum due to wind action. 359

³⁶⁰ Comparing the spectral evolution of the simulation cases, however, the influence of the ³⁶¹ higher order wind term can be clearly observed. Figure 10a displays the spectral mean of the ³⁶² three simulations. Simulation (i) confirms the result of Carter and Govan²⁴, namely that ³⁶³ the higher order viscosity term causes a downshift in the spectral mean. Since $\Gamma = 0$, $\delta_1 < 0$,

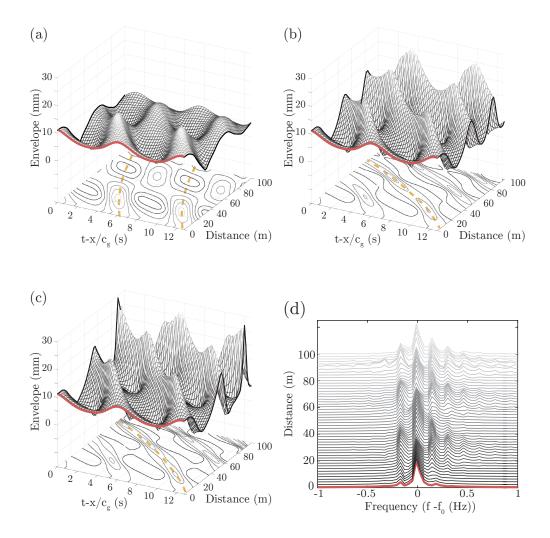


FIG. 9: Evolution of envelope in space and time for the (a) no wind, (b) leading order wind, (c) full model simulations. The orange dashed line shows the space-time evolution of the wave packets. (d) Simulated spectral amplitude $|\hat{\eta}|$ (normalized), for the full model simulation. Each spectrum is offset according to the distance Red line indicates the initial condition based on the theoretical Akhmediev solution at $x^* = -20$ m and A = 0.25 that is subsequently propagated in space by numerical integration. For carrier wave parameters $\epsilon = 0.1, T_0 = 0.6$ s, and simulation parameters $\nu = 1 \times 10^{-5} \text{ m}^2/\text{s}$ and $\Gamma = 8.0 \times 10^{-3} \text{ s}^{-1}$ when wind is active.

³⁶⁴ and the light dashed line, indicating the trend of the spectral mean, has a negative slope. ³⁶⁵ In simulation (ii), the addition of the leading order wind term accelerates the oscillation of ³⁶⁶ the spectral mean, without affecting the trend: the oscillations follow the same downward

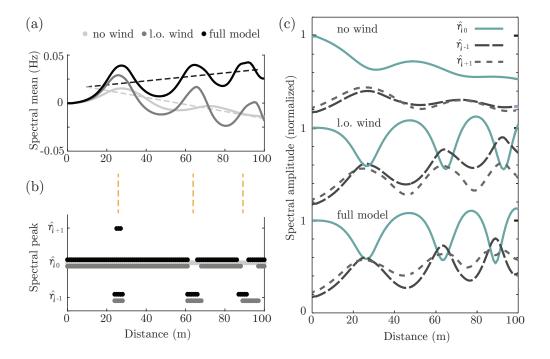


FIG. 10: (a) Simulated evolution of the spectral mean f_m. Dark dashed line is the trend set by the higher order wind term, light dashed line the trend set by the higher order viscosity term. (b) Evolution of the spectral peak f_p. Dots are offset for clarity. The dashed lines between (a) and (b) indicate the focal points of the quasi-recurrence cycles.
(c) Simulated evolution of the Fourier amplitudes for of the three central modes \$\hat{\eta}_0\$ (green solid line), \$\hat{\eta}_{-1}\$ (dashed dark gray line) and \$\hat{\eta}_{+1}\$ (dashed light gray line). Comparing the no wind, leading order (l.o.) wind and full model simulations. Simulations based on the same parameters as figure 9.

³⁶⁷ slope as set by the higher order viscosity term ($\delta_1 < 0$). In contrast, in the full model ³⁶⁸ simulation (iii), there is indeed a clear tendency towards a permanent upshift in the spectral ³⁶⁹ mean, indicated by the dark dashed line, as now there is a forcing effect in the higher order ³⁷⁰ ($\Gamma > 5\nu k_0^2, \delta_1 > 0$). In brief, due to the MNLS modification terms the mean oscillates around ³⁷¹ the slope induced by a balance between the higher order viscosity term and the higher order ³⁷² wind term.

Figure 10b shows the position of spectral peak f_p as a function of distance. As observed in the experimental data in Section IV, without wind forcing f_p remains equal to f_0 . In simulation (ii) a temporary downshift is observed. The origin of this downshift can be

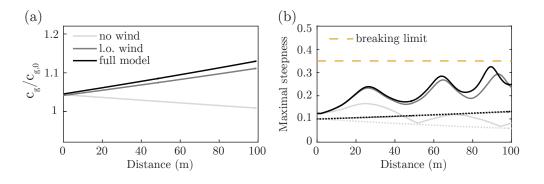


FIG. 11: Simulated evolution of (a) the group velocity and (b) the maximal steepness. Comparing the no wind, leading order (l.o.) wind and full model simulations. The dashed line indicates the critical steepness for wave breaking Simulations based on the same parameters as figure 9. The dotted lines indicate the characteristic steepness

 $\epsilon_{\rm char} = {
m RMS}(\eta^2)k_0$

³⁷⁶ revealed by analyzing figure 10c. Without wind action, the amplitudes of the modes oscillate, but the sidebands do not overtake the main mode. Considering the leading order wind effect, 377 the wave steepness is amplified and consequently the growth and decay rates of all modes, 378 increasing the frequency and amplitude of their oscillation. In figure 10c the sidebands do 379 overtake the main peak within a quasi-recurrence cycle. Due to the initially slightly different growth rates, $\hat{\eta}_{-1}$ reaches a slightly higher amplitude than $\hat{\eta}_{+1}$ and a temporary downshift 381 in the spectral peak sense occurs. The observations of a similar spectral peak downshift pattern by Tulin and Waseda³⁷ without wind forcing but at higher steepness, confirm that 383 ³⁸⁴ the spectral peak downshift is not a direct effect of wind forcing. Rather, it is an consequence due to the wind's influence on the steepness. That is, the spectral asymmetry inherent to 385 the MNLS, where the lower sideband has a slightly higher growth rate than the upper 386 sideband, is amplified by the wind as wave steepness is naturally increased. In simulation 387 (iii) this downshift pattern is not significantly altered, although the quasi-recurrence cycles 388 are slightly shorter. We can notice in figure 10c that instead of $\hat{\eta}_{-1}$ continuously being the 389 dominant sideband, $\hat{\eta}_{-1}$ and $\hat{\eta}_{+1}$ alternate. 390

The addition of the forcing terms affects the group velocity. The homogeneous NLS implies a linear group velocity $c_{g,0}$. Taking into account the MNLS correction terms increases the wave packets' speed⁶⁰. The higher the steepness, the higher the importance of the MNLS ³⁹⁴ correction and the higher the increase in c_g . For the *no wind* simulation the dashed line ³⁹⁵ in figure 9a shows a curvature of the direction of the wave packet propagation in the x - t³⁹⁶ plane that is attenuated towards the end of the tank due to the decrease in steepness caused ³⁹⁷ by viscosity and dissipation. In contrast, for the wind simulations, the dashed lines indicate ³⁹⁸ that this induced variation of group velocity increases with the increase of the steepness. ³⁹⁹ Figure 11a quantifies the deviation from the linear group velocity for the three simulation ⁴⁰⁰ cases (i), (ii) and (iii). In the experiment, a similar increase in the group velocity is indeed ⁴⁰¹ observed.

It should be noted that the regime in which the higher order wind term becomes relevant 402 403 is hard to reach experimentally due to its proximity to the wave breaking threshold. An in-404 creased steepness makes the higher order wind term of greater influence and hereby increases the deviation from simulation (ii). However, at the same time, high steepness brings waves 405 closer to the breaking threshold, beyond which our model (11) is incomplete. Furthermore, 406 the significance of the higher order wind term increases with the wind strength and fetch, 407 see figure 10a. Indeed, the steepness is also increased as a consequence of wind strength and 408 fetch, as displayed in figure 11b. Thus it is important to monitor the steepness in the simula-409 ⁴¹⁰ tions to signal possible wave breaking. In our long range simulations, the maximal steepness 411 of the wave $\epsilon_{\text{max}} = k_0 a_{\text{max}}$, remains below the breaking limit of $\epsilon = 0.35$ as considered by Trulsen and Dysthe⁴⁰. While this value for the threshold number was calculated for Stokes 412 waves in the absence of wind, Saket *et al.*⁶¹ show the breaking threshold is very similar for 413 ⁴¹⁴ wind driven waves. Note that other studies suggest an even higher critical steepness^{62–64}.

415 VI. DISCUSSION

For the spectral peak, data and simulations alike show that a forcing wind can induce 417 a downshift. However, the underlying cause for the faster growth of the lower sideband is 418 the asymmetry introduced by the MNLS correction terms, which is amplified by the wind. 419 Moreover, this downshift is only temporary. Considering the spectral mean, our long range 420 simulations show that the higher order wind term creates a permanent upward trend, while 421 the higher order viscosity term causes a permanent downward trend. As both terms have 422 the same form in (11), the balance between these two, the sign of δ_1 , determines whether an 423 upshift or a downshift in the spectral mean will be observed. Finally, when the wind action

TABLE II: Our observations of the influence on the spectrum of the MNLS correction, and the leading order (l.o.) and higher order (h.o.) wind and viscosity terms in (11). US = upshift, DS =downshift.

	Effect		
Terms	Mean	Peak	
MNLS correction	temporary DS	temporary US	
l.o. viscosity \rightarrow lower ϵ	slower MNLS dynamics	slower, damped oscillation	
h.o. viscosity	-	permanent DS	
l.o. wind \rightarrow higher ϵ	faster MNLS dynamics	faster, amplified oscillation	
h.o. wind	-	permanent US	

⁴²⁴ is sufficiently strong, wave breaking is a natural result. We experimentally confirm the well ⁴²⁵ known notion that wave breaking induces a permanent downshift in both the spectral peak ⁴²⁶ and spectral mean. Our observations on the effect of the wind, viscosity, and the MNLS ⁴²⁷ modification are summarized in table II.

These findings might seem contradictory with respect to existing literature in which wind 429 is often associated with spectral downshift of gravity waves^{18–20,22,40}, as discussed. However, 430 by taking into account the different downshift interpretations and making the distinction 431 between wind forcing, in which energy is added to the system and forms of energy dissipation 432 that can be triggered by wind, our results extend the existing framework, as upon closer 433 inspection instances of permanent downshift are associated with dissipation. This distinction 434 between the direct and potential indirects effects of wind solves the downshift paradox. This 435 idea is confirmed for unidirectional waves in the review of the frequency downshift by Dias 436 and Kharif⁴⁷.

The most obvious dissipative phenomenon that can occur as a result of wind forcing is that waves reach a critical steepness and break. Wave breaking shifts the spectral peak are to a lower frequency. This has indeed been already experimentally observed by Tulin and Waseda³⁷, Melville⁶², Lake *et al.*⁶⁵ and is explained along the general line of energy being dissipated from the higher modes into the lower modes. Efforts have been made to model wave breaking theoretically^{21,22,40,54,55,57}, as discussed. Another instance in which wind can ⁴⁴³ have a dissipative effect is when it blows in opposite direction of the wave propagation and ⁴⁴⁴ as such damps the waves. This configuration has been modeled by Schober and Strawn²⁰, $_{445}$ a study in which the dissipation term as defined by Kato and Oikawa 22 has been taken into account and a permanent downshift is modeled as well. Finally, even when the direct forcing 446 447 effect of wind is included in a study, the dominant regime for spectral movement can still be dissipative when the viscosity is strong. For example, Touboul and Kharif¹⁹ observed 448 a permanent downshift in the spectral peak due to the effect of wind. However, they are 449 exactly on the balance of forcing and dissipation in the leading order, $\Gamma \approx 4k_0^2\nu$, and thus, 450 in the higher order regime the dissipation is slightly dominating, $\delta_1 < 0$. A similar argument holds for Hara and Mei¹⁸. While our study applies to surface gravity waves, it is interesting 452 ⁴⁵³ to note that for capillary-gravity waves Hara and Mei⁶⁶ numerically found a frequency ⁴⁵⁴ upshift of the spectral peak due to wind, and Skandrani, Kharif, and Poitevin⁶⁷ have shown ⁴⁵⁵ numerically that the shift to lower frequencies is promoted by a damping mechanism.

456 VII. CONCLUSION

We derive a higher order $O(\epsilon^4)$ wind forcing contribution to the MNLS framework, re-457 ⁴⁵⁸ sulting in a forced-damped MNLS equation (11). The direct effect of this term, when it ⁴⁵⁹ exceeds the viscosity at the same order, is an upshift of the spectral mean. This trend is superimposed on the oscillation caused by the MNLS correction. For significant wind action, the higher order wind term cancels the downshifting effect of the higher order viscosity 461 term and moves towards an upward trend of the spectral mean. The leading order wind 462 term is symmetric but can amplify the asymmetric growth initiated by the MNLS correction 463 terms, resulting in a temporary downshift of the spectral peak. Finally, we confirm that the 464 permanent downshift of the spectral mean and of the spectral peak often observed in wind 465 ⁴⁶⁶ experiments is an indirect effect associated with dissipation, including wave breaking.

This is the first time a propagation equation for deep water waves including wind forcing 468 is validated by laboratory experiments. The evolution of an Akhmediev breather in the 469 presence of wind shows good agreement with the model in the limited fetch, dictated by the 470 length of the facility. A natural continuation of this work is an experimental exploration of 471 the full damping-forcing range that we have now modeled in the leading and higher order in 472 (2.8), in particular the upshift in the spectral mean predicted by the higher order wind term.

⁴⁷³ In addition, increasing the fetch in the laboratory environment would improve the validation ⁴⁷⁴ analysis of the numerical results. However, it is a delicate balance between observing the ⁴⁷⁵ effect of the higher order term, and driving the model out of its validity range due to wave ⁴⁷⁶ breaking, since the elements that increase the effect of the higher order term, *i.e.* fetch, ⁴⁷⁷ wind speed and steepness, at the same time drive waves toward breaking.

478 ACKNOWLEDGMENTS

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