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REHH 2.0: a reimplementation of the R package REHH to detect positive selection from haplotype structure.

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keywords: footprints of selection, EHH, iHS, Rsb, XP-EHH

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Abstract

Identifying genomic regions with unusually high local haplotype homozygosity represents a powerful strategy to characterize candidate genes responding to natural or artificial positive selection. To that end, statistics measuring the extent of haplotype homozygosity within (e.g., EHH, iHS) and between (Rsb or XP-EHH) populations have been proposed in the literature. The REHH package for R was previously developed to facilitate genome-wide scans of selection, based on the analysis of long-range haplotypes. However, its performance wasn't sufficient to cope with the growing size of available data sets. Here we propose a major upgrade of the REHH package, which includes an improved processing of the input files, a faster algorithm to enumerate haplotypes, as well as multi-threading. As illustrated with the analysis of large human haplotype data sets, these improvements decrease the computation time by more than an order of magnitude. This new version of REHH will thus allow performing iHS-, Rsb- or XP-EHH-based scans on large data sets. The package REHH 2.0 is available from the CRAN repository (http://cran.r-project.org/web/packages/rehh/index.html) together with help files and a detailed manual.

Introduction

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Next-generation sequencing (NGS) technologies have deeply transformed the nature of polymorphism data. Although population geneticists were, until recently, limited by the amount of available data in a handful of presumably independent markers, they now have access to dense single nucleotide polymorphism (SNP) data in both model and non-model species (Davey et al, 2011). In those species where 19 genome assemblies are available, the analysis of haplotype structure in a population has proved useful to detect recent positive selection (Sabeti et al., 2002). Consider neutral mutations appearing in a pop-21 ulation: if, by chance, any of these increases in frequency after some time, then recombination should tend to break down linkage disequilibrium (LD) around it, thereby decreasing the length of haplotypes 23 on which this mutation stands. Common variants are therefore expected to be old and standing on short haplotypes. If a mutation is selected for, however, it should expand in the population before recombination has time to break down the haplotype on which it occurred. A powerful strategy to characterize candidate genes responding to natural or artificial positive selection thus consists in identifying genomic regions with unusually high local haplotype homozygosity, relatively to neutral expectation (Sabeti et al, 2002). For that purpose, Sabeti et al (2002) introduced a new metric, referred to as the extended haplotype homozygosity (EHH), which measures the decay of identity by descent, as function of distance, between randomly sampled chromosomes carrying a focal SNP. Tests of departure of EHH from neutral expec-32 tation were proposed, based on coalescent simulations of demographic history. Voight et al (2006) later introduced a test statistic (iHS) based on the standardized log-ratio of the integrals of the observed decay of EHH computed for the ancestral and the derived alleles at the focal SNP. Finally, cross-population statistics were proposed, to contrast EHH profiles between populations: XP-EHH (Sabeti et al, 2007) and Rsb (Tang et al, 2007). These haplotype-based methods of detecting selection have largely been applied on human data (Vitti et al, 2013), a wide range of livestock (see, e.g. Flori et al, 2014; Bosse et al, 2015; Barson et al, 2015) and plant species (see, e.g. Wang et al, 2014; Jin et al, 2016), and also

A few years ago, we developed REHH (Gautier & Vitalis, 2012), a package for the statistical software package R (R Development Core Team, 2008), to detect recent positive selection from the analysis of long-range haplotypes. Since then, two alternative programs were released: selscan (Szpiech & Hernandez,

non-model species (see, e.g. Roesti et al, 2015; Mueller et al, 2016).

44 2014), which introduces multithreading to improve computational efficiency and happin (Maclean et al,

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s 2015), which in addition to multithreading offers considerable gain in computation time thanks to a new

computational approach based on a bitwise algorithm.

Here we propose a major upgrade of the REHH package (Gautier & Vitalis, 2012), which includes an improved algorithm to enumerate haplotypes, as well as multi-threading. These improvements decrease the computation time by more than an order of magnitude, as compared to the previous REHH version (1.13), which eases the analysis of big datasets.

Below we provide a brief overview of the statistics and tests available in REHH 2.0, and give a detailed worked example of the analysis of chromosome 2 in humans (HSA2), from HapMap samples CEU (Utah residents with Northern and Western European ancestry from the CEPH collection) and JPT+CHB (Japanese in Tokyo, Japan and Chinese from Beijing, China). We use this example as a guideline to use REHH2.0. We further show how REHH was improved since the previous version, and how it compares to the alternative programs selscan (Szpiech & Hernandez, 2014) and hapbin (Maclean et al, 2015).

57 Overview of the EHH-based tests

58 Within population tests

The allele-specific extended haplotype homozygosity: EHH (Sabeti et al., 2002)

At a focal SNP and for a given core allele (the ancestral or derived), the allele-specific extended haplotype homozygosity (EHH) is defined as the probability that two randomly chosen chromosomes (carrying the core allele considered) are identical by descent (IBD). IBD is assayed by computing homozygosity at all SNPs within an interval surrounding the core region (Sabeti *et al*, 2002). The EHH thus aims at measuring to which extent an extended haplotype is transmitted without recombination. In practice, the EHH (EHH $_{a_s,t}$) of a tested core allele a_s ($a_s = 1$ or $a_s = 2$) for a focal SNP s over the chromosome interval comprised between the core allele a_s and the SNP t is computed as:

$$EHH_{a_s,t} = \frac{1}{n_{a_s}(n_{a_s} - 1)} \sum_{k=1}^{K_{a_s,t}} n_k(n_k - 1)$$
(1)

where $K_{a_s,t}$ represents the number of different extended haplotypes (from SNP s to SNP t) carrying
the core allele a_s , n_k is the number of the kth haplotype, and n_{a_s} represents the number of haplotypes
carrying the core allele a_s , i.e., $n_{a_s} = \sum_{k=1}^{K_{a_s,t}} n_k$.

The integrated (allele-specific) EHH: iHH (Voight et al, 2006)

By definition, irrespective of the allele considered, EHH starts at 1, and decays monotonically to 0 as one 71 moves away from the focal SNP. For a given core allele, the integrated EHH (iHH) (Voight et al, 2006) is defined as the area under the EHH curve with respect to map position. In REHH (Gautier & Vitalis, 2012), 73 this definite integral is computed using the trapezoidal rule. In practice, the integral is only computed for the regions of the curve above an arbitrarily small EHH value (e.g., EHH>0.05). In their seminal 75 paper, Voight et al (2006) considered genetic distances and apply a penalty (proportional to physical distances) for successive SNPs separated by more than 20 kb. In addition, they did not compute iHH 77 if any physical distance between a pair of neighboring SNPs was above 200 kb. We did not implement such an approach in REHH although this might easily be done by modifying the positions of the markers in SNP information input file. In addition, large gaps between successive SNPs (e.g., centromeres) might 80 also be treated by splitting the chromosomes. For instance, when analyzing metacentric chromosomomes (e.g., HSA2), each chromosome arm may be considered separately in the analyses by assigning a different chromosome name (e.g., 2a and 2b) to the underlying SNPs.

- The standardized ratio of core alleles iHH: iHS (Voight et al, 2006)
- Let UniHS represent the log-ratio of the iHH for its ancestral (iHH $_a$) and derived (iHH $_d$) alleles:

$$UniHS = \log \left(\frac{iHH_a}{iHH_d} \right) \tag{2}$$

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The iHS of a given focal SNP s (iHS(s)) is then defined following (Voight et al, 2006) as:

$$iHS(s) = \frac{UniHS(s) - \mu_{UniHS}^{p_s}}{\sigma_{UniHS}^{p_s}}$$
(3)

where $\mu_{\text{UniHS}}^{p_s}$ and $\sigma_{\text{UniHS}}^{p_s}$ represent, respectively, the average and the standard deviation of the UniHS computed over all the SNPs with a derived allele frequency p_s similar to that of the core SNP s. In practice, the derived allele frequencies are generally binned so that each bin is large enough (e.g., >10

SNPs) to obtain reliable estimates of $\mu_{\text{UniHS}}^{p_s}$ and $\sigma_{\text{UniHS}}^{p_s}$. The iHS is constructed to have an approximately standard Gaussian distribution and to be comparable across SNPs regardless of their underlying allele frequencies. Hence, one may further transform iHS into p_{iHS} (Gautier & Naves, 2011):

$$p_{\text{iHS}} = -\log_{10} \left(1 - 2|\Phi \left(\text{iHS} \right) - 0.5| \right) \tag{4}$$

6

where $\Phi(x)$ represents the Gaussian cumulative distribution function. Assuming most of the genotyped SNPs behave neutrally (i.e., that the genome-wide empirical iHS distribution is a fair approximation of the neutral distribution), p_{iHS} may thus be interpreted as a two-sided p-value (in a $-\log_{10}$ scale) associated with the null hypothesis of selective neutrality.

97 Pairwise-population tests

The site-specific extended haplotype homozygosity: EHHS (Tang et al, 2007; Sabeti et al, 2007)

At a focal SNP, the site-specific extended haplotype homozygosity (EHHS) is defined as the probability

that two randomly chosen chromosomes are IBD at all SNPs within an interval surrounding the core

region. EHHS might roughly be viewed as linear combination of the EHH's for the two alternative alleles,

with some weights depending on the corresponding allele frequencies. Two different EHHS estimators

further referred to as EHHS^{Sabeti} and EHHS^{Tang} have been proposed by Sabeti et al (2007) and Tang

et al (2007), respectively. For a focal SNP s over a chromosome interval extending to SNP t, these are

computed as (using the same notation as above):

EHHS^{Sabeti}_{s,t} =
$$\frac{1}{n_s(n_s - 1)} \sum_{a_s = 1}^{a_s = 2} \left(\sum_{k=1}^{K_{a_s,t}} n_k(n_k - 1) \right)$$
 (5)

where $n_s = \sum\limits_{a_s=1}^{a_s=2} n_{a_s}$ and

$$EHHS_{s,t}^{Tang} = \frac{1 - h_{hap}^{(s,t)}}{1 - h_{all}^{(s)}}$$
(6)

107 where:

• $h_{all}^{(s)} = \frac{n_s}{n_s - 1} \left(1 - \frac{1}{n_s^2} \sum_{a_s = 1}^{a_s = 2} n_{a_s}^2 \right)$ is an estimator of the focal SNP heterozygosity

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$$h_{hap}^{(s,t)} = \frac{n_s}{n_s-1} \left(1 - \frac{1}{n_s^2} \sum_{a_s=1}^{a_s=2} \left(\sum_{k=1}^{K_{a_s,t}} n_k^2 \right) \right)$$
 is an estimator of haplotype heterozygosity over the chromosome region interval extending from SNP s to SNP t .

111 The integrated EHHS: iES

As for the EHH (see above), EHHS starts at 1 and decays monotonically to 0 with increasing distance from the focal SNP. At a focal SNP, and in a similar fashion as the iHH, iES is defined as the integrated EHHS (Tang *et al*, 2007). Depending on the EHHS estimator considered, EHHS^{Sabeti} or EHHS^{Tang}, two different iES estimators, that we further refer to as iES^{Sabeti} and iES^{Tang} can be computed.

The standardized ratios of pairwise population iES: XP-EHH (Sabeti *et al*, 2007) and Rsb (Tang *et al*, 2007)

For a given SNP s, let LRiES^{Sabeti}(s) (respectively LRiES^{Tang}(s)) represent the (unstandardized) log-ratio of the iES^{Sabeti}(s) and iES^{Sabeti}(s) (respectively iES^{Tang}(s) and iES^{Tang}(s)) computed in two different populations:

$$LRiES^{Sabeti}(s) = \log \left(\frac{iES_{pop1}^{Sabeti}(s)}{iES_{pop2}^{Sabeti}(s)} \right) \quad and \quad LRiES^{Tang}(s) = \log \left(\frac{iES_{pop1}^{Tang}(s)}{iES_{pop2}^{Tang}(s)} \right)$$
 (7)

The XP-EHH (Sabeti et al, 2007) and the Rsb (Tang et al, 2007) for a given focal SNP are then standardized, as:

$$xpEHH(s) = \frac{LRiES^{Sabeti}(s) - med_{LRiES^{Sabeti}}}{\sigma_{LRiES^{Sabeti}}} \quad and \quad rSB(s) = \frac{LRiES^{Tang}(s) - med_{LRiES^{Tang}}}{\sigma_{LRiES^{Tang}}} \quad (8)$$

where $med_{LRiESSab}$ (respectively $med_{LRiESTang}$) and $\sigma_{LRiESSab}$ (respectively $\sigma_{LRiESTang}$) represent the median and standard deviation of the $LRiES^{Sabeti}(s)$ (respectively $LRiES^{Tang}(s)$) computed over all the analyzed SNPs. As recommended by Tang *et al* (2007), the median is used instead of the mean because it is less sensitive to extreme data points. As for the iHS (see above), XP-EHH and Rsb are constructed to have an approximately standard Gaussian distribution. They may further be transformed into p_{xpEHH} or p_{rsb} :

$$p_{\text{xpEHH}} = -\log_{10} (1 - 2|\Phi(\text{xpEHH}) - 0.5|)$$
 and $p_{\text{rSB}} = -\log_{10} (1 - 2|\Phi(\text{rSB}) - 0.5|)$ (9)

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where $\Phi(x)$ represents the Gaussian cumulative distribution function. Assuming most of the genotyped SNPs behave neutrally (i.e., the genome-wide empirical distributions of XP-EHH and Rsb are fair approximations of their corresponding neutral distributions), p_{xpEHH} and p_{rSB} may thus be interpreted as a two-sided p-values (in a $-\log_{10}$ scale) associated with a null hypothesis of selective neutrality. Alternatively, one may also compute p'_{xpEHH} or p'_{rSB} as:

$$p_{\text{xpEHH}} = -\log_{10} \left(\Phi \left(\text{xpEHH} \right) \right) \quad \text{and} \quad p_{\text{rSB}} = -\log_{10} \left(|\Phi \left(\text{rSB} \right)| \right)$$
 (10)

(see Gautier & Naves, 2011); p'_{xpEHH} and p'_{rSB} may then be interpreted as a one-sided p-values (in a $-\log_{10}$ scale) allowing the identification of those sites displaying outstandingly high EHHS in population pop2 (represented in the denominator of the corresponding LRiES) relatively to the reference population (pop1).

Material and Methods

A new efficient algorithm to explore haplotype variability

In the previous version of REHH (1.13) the distribution of haplotype counts for the entire interval from the core SNP to the distance x was computed for each x independently, entailing repeatedly the same calculations. In the new version of REHH (2.0), the distribution of haplotype counts for the interval from the core SNP to the distance x is updated consecutively from the distribution of haplotype counts corresponding to the interval between the core SNP and x-1. The new algorithm doesn't affect the output, in particular, as in the previous version, all haplotypes carrying missing data are discarded from the computation of long-range homozygosity.

47 Human haplotype data

Two HSA2 haplotype data sets were downloaded from the HAPMAP project (phase III) (The Interna-148 tional HapMap3 Consortium, 2010) website (ftp://ftp.ncbi.nlm.nih.gov/hapmap/). They consisted of 236 haplotypes of 116,430 SNPs from the CEU and 342 haplotypes from the JPT+CHB popula-150 tions, respectively. Further details about these data (including the phasing procedure) can be found 151 on the HAPMAP website. For each SNP, the ancestral (and derived) allele was determined accord-152 ing to the Chimpanzee genome reference (using the dbsnp_chimp_B36.qff annotation file available at 153 ftp://ftp.ncbi.nlm.nih.gov/hapmap/gbrowse/2010-08_phaseII+III/gff/). Such ancestral information is indeed required to carry out iHS-based tests (see above). As a result, 6,230 SNPs (5.35%) 155 for which ancestral/derived states could not be unambiguously determined were discarded from further analyses leading to a total of 110,200 SNPs per analyzed haplotype. 157

158 Computation

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For comparison purposes, the different haplotype data sets were analyzed using the software packages 159 REHH (both the previous version 1.13 and the new version 2.0), selscan (version 1.1.0b) (Szpiech & Hernandez, 2014) and happin (version 1.0.0) (Maclean et al, 2015). Default options were generally used 161 except for the minimal threshold on the minor allele frequency (MAF) that was set to 0.01 for all programs. In addition, for the selscan program, both the window size around the core SNPs (--ehh-win 163 option) and the maximum allowed gap in bp between two consecutive SNPs (--max-gap option) were set 164 to 10⁹ (this was made to disallow these options that are not considered in other programs). Similarly, 165 the --max-extent option was inactivated by setting --max-extent=-1. For the hapbin programs (i.e., 166 ihsbin and xpehhbin), the EHH and EHHS cut-off values (defined to stop the calculation of unstandard-167 ized iHS and iES) were set to 0.05 (i.e., the default value in selscan and REHH). For all programs, the 168 standardization of iHS was performed with allele frequency bins of 0.01, as controlled by the frequency argument in the ihh2ihs() function of the REHH package, and the bins argument for the program norm 170 of the selscan package and the program ihsbin of the hapbin package. The command lines used for the different programs, together with the corresponding input data files are provided in the Supplementary 172 Materials. 173

Finally, for each analysis and parameter set, the estimation of computation times was averaged over

ten independent runs. All analyses were run on a standard computer running under Linux Debian 8.5
and equipped with an Intel[®] Xeon[®] 6-core processor W3690 (3.46 GHz, 12M cache). Note that the Unix
command taskset was used to control the number of working threads for the analyses with the hapbin
programs (since neither the ihsbin nor the xpehhbin programs allow to chose the number of threads to
be used).

Results and Discussion

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Analysis of the human chromosome 2 data sets

For illustration purpose, we used REHH 2.0 to analyze two human data sets consisting of 236 and 342 182 haplotypes of 110,200 SNPs mapping to HSA2 that were sampled in the CEU and JPT+CHB popu-183 lations respectively. The chromosome-wide scans of iHS for the CEU and the JPT+CHB populations, 184 respectively, are plotted in Figure 1A. The most significant SNP map at position 136.503,121 bp for the CEU population (iHS=-5.35) and at position 111,506,728 bp for the JPT+CHB population (iHS=-4.92). 186 The chromosome-wide scans of XP-EHH and Rsb, which contrast EHH profiles between the CEU and the JPT+CHB populations, are plotted in Figure 1B. The most significant SNP mapped at position 188 136,533,558 bp for Rsb-based test (Rsb=6.13) and at position 136,523,244 bp for the XP-EHH-based test (XP-EHH=5.59). For this latter SNP (mapping to region #7 as defined below), the haplotype bifurca-190 tion diagrams for the ancestral and derived alleles within the CEU population are plotted in (Figure 1C) and (Figure 1D), respectively, using the bifurcation diagram function from the REHH package. Note 192 the extent of haplotype homozygosity associated with the derived allele (Figure 1C), relatively to that associated with the ancestral allele (Figure 1D), which is consistent with the negative iHS measure at this SNP (iHS=-3.24). 195

[Figure 1 about here.]

To further identify regions displaying strong footprints of selection, we split the HSA2 chromosome into 950 consecutive 500 kb-windows (with a 250 kb overlap). Windows with at least 2 SNPs displaying a statistic > 4 (in absolute value that roughly corresponds to a two-sided p-value $< 10^{-4}$, see above) for at least one of the four test statistics were deemed significant. Significant overlapping windows were then merged, leading to a total of 11 regions harboring strong signals of selection, which characteristics are

detailed in Table 1 (see also Figure 1). As expected, most of the regions identified in previously published genome-scans for samples with the same origin (Sabeti et al, 2007; Tang et al, 2007; Voight et al, 2006) overlap with the regions identified here (Table 1). For instance, regions #6 and #7 that are in the vicinity of the EDAR gene (under selection in Asian populations) and the LCT gene (under selection in European populations), respectively, have been extensively characterized in the literature (e.g., Peter et al, 2012). Interestingly, we detected more regions than previously reported in the aforementioned studies, most probably because our analyses are based on a larger dataset and different cut-off values. A more detailed description of the newly identified regions is however beyond the scope of the present article.

[Table 1 about here.]

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Note finally that XP-EHH- and Rsb-based scans gave consistent results, with the exception of the region in the vicinity of the LCT gene (#7 in Table 1 and Figure 1) where a double peak was observed with Rsb (consistent with the iHS profile within CEU) and a single peak with XP-EHH. The overall correlation between these statistics was equal to 0.843, which illustrates the close similarity of these two metrics.

Comparing the performances of REHH 2.0 relatively to REHH 1.13, selscan and hapbin packages

The two CEU and JPT+CHB human data sets were further analyzed with REHH 1.13 to evaluate the gain in computation time resulting from the modifications introduced in version 2.0. Note that extensive tests were done during the development of version 2.0, to ensure that the same estimates were obtained with both versions. Only very marginal differences were however sometimes observed in the estimates of iES^{Tang}. For instance, the correlation between the resulting Rsb computed across the CEU and JPT+CHB populations with version REHH 1.13 and REHH 2.0 was found equal to 0.999992 (instead of 1.0). This is actually due to the introduction of the computation of iES^{Sabeti} in version 2.0 to estimate XP-EHH. Indeed, we chose to define the same cut-off value for both statistics during the computation of the component variable EHHS (controlled with the option limehhs, set to 0.05 by default).

An improved processing of the input file

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The first major modification introduced in REHH version 2.0 deals with the processing of input files 228 (haplotype and SNP information files) using the function data2haplohh. Indeed our own experience with earlier versions of the package together with feedback from several users prompted us to optimize data 230 import and to improve allele recoding, which was inefficient in version 1.x. Considering standard input 231 haplotype file format (which is common to both versions), and with alleles encoded in the appropriate 232 format ({0,1,2} for missing data, ancestral and derived alleles respectively), the new data2haplohh 233 function is about 2.5 times faster than the previous one (see Table 2). In addition, the allele recoding 234 option results in slightly better processing performances, and is no more prone to errors as in version 1.x. 235 Finally, the new haplotype format (with haplotypes in columns), corresponding to the output file of the 236 SHAPEIT phasing program (O'Connell et al, 2014), was found to be the most efficient to process (see 237 Table 2).

[Table 2 about here.]

With datasets of increasing complexity and size, such improvement in the processing of input files is critical to REHH users. Processing a data set as large as the JPT+CHB one (consisting of 342 haplotype with 110,200 SNPs) now takes less than 12 seconds. Note however that for this file a maximum of about 1 Gb RAM was used, for a net memory size change of 240 Mb. For larger data sets, RAM requirements may therefore be limiting for some computers.

²⁴⁵ A faster and parallel algorithm to explore haplotype variability

The second major modification introduced in REHH version 2.0 concerns the core algorithm that computes 246 the distribution of haplotype counts, which underlies the calculation of all the metrics of interest (iHS, Rsb and XP-EHH). As shown in Table 3, this new algorithm allows to decrease the computation times 248 by more than one order of magnitude, as compared to the algorithm implemented in REHH version 1.13. 249 Hence, for the computation of iHS in the CEU population (respectively, the JPT+CHB population) on 250 a single thread, computation times were 13.7 (respectively 21.8) times smaller on average. Interestingly, 251 the computation time for the JPT+CHB dataset (which is approximately 1.34 times larger than the CEU 252 one in terms of number of SNPs × number of haplotype) was only 1.09 times slower than for the latter. 253 Conversely, the computation time was 1.73 times slower for JPT+CHB relatively to CEU with REHH version 1.13. Although a more detailed profiling of the algorithm would be required, these results suggest
that computational burden is approximately linearly related to the data set complexity.

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To further improve computational speed, haplotype structure is now performed using OpenMP parallelization across SNPs in genome-wide scans. Using four threads then lead to an additional decrease of about 3.5 times in computation times (see Table 3).

Overall, the whole analysis of the HSA2 haplotype files used in this study took about 1.5 minutes (including the processing of input files) with REHH 2.0, and more than 1.3 hours with REHH 1.3. This corresponds to the computation of iHS within the CEU and within the JPT+CHB populations, as well as the computation of Rsb and XP-EHH.

[Table 3 about here.]

Comparing REHH 2.0 to the selscan and hapbin programs

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Finally, we compared REHH 2.0 with selscan (Szpiech & Hernandez, 2014) and hapbin (Maclean et al, 266 2015), which were recently published. Both programs are written in C++ language and include paral-267 lelisation. Computation times for the different analyses, either on a single or four threads, are provided 268 in Table 3. The new version of REHH outperforms selscan by about one order of magnitude. Moreover, running REHH on a single thread is still more than twice as fast as running selscan on four threads. It 270 should also be noticed that running a full analysis consisting of the estimation of iHS within and XP-EHH between the CEU and JPT+CHB populations result in a significant additional burden with selscan (Ta-272 ble 3). Conversely, hapbin was found to be more than five times faster than REHH 2.0, most likely as a result of its more efficient algorithm to explore haplotype variability. Yet, given the small computation 274 times achieved by both programs, REHH 2.0 remains competitive relative to happin for most practical applications. 276 Correlation between the estimated iHS and XP-EHH obtained with the different programs are given in 277 Table 4. Estimates of XP-EHH were in almost perfect agreement among the different software packages. 278 Similarly, estimates for iHS were almost the same between REHH 2.0 and selscan but slightly depart from 279 those obtained with hapbin. Although we did not further investigate the origin of these discrepancies, 280 this might probably be related to a different definition of haplotype homozygosity in hapbin, as compared 281 to Sabeti $et \ al \ (2007)$.

[Table 4 about here.]

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84 Conclusion

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Although the R package REHH (Gautier & Vitalis, 2012) has been widely used since its first release, the increasing dimension of haplotype datasets typically available in most species led to serious limitations. 286 This stimulated the development of alternative R-free solutions (Szpiech & Hernandez, 2014; Maclean et al, 2015). In this study, we introduced substantial changes in the REHH package to improve its 288 computational efficiency by one to several orders of magnitude. This was achieved by modifying the processing of the input files and, most importantly, by improving and parallelizing the core algorithm 290 that computes the distribution of haplotype counts. As a result, REHH 2.0 clearly outperforms the selscan (Szpiech & Hernandez, 2014) package and competes with hapbin (Maclean et al, 2015), the 292 fastest program to date. A decisive advantage of REHH 2.0 over these programs is that it allows working within the multi-platform R environment. As such, it benefits from several graphical tools that facilitate 294 visual interpretation of the results. 295 REHH 2.0 is available from the CRAN repository (http://cran.r-project.org/web/packages/ 296 rehh/index.html). A help file together with a detailed vignette manual (the current version is pro-297 vided as a Supplementary File S2) are included in the package.

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5 Supplementary Material

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- File S1: compressed archive named FileS1.tar.gz containing example input haplotype data and
 SNP information files in the REHH, selscan and hapbin format. The archive also contains command
 lines that were used to run the different programs
 - File S2: Detailed user manual (vignette) for the REHH 2.0.

List of Figures

 FIGURES 19

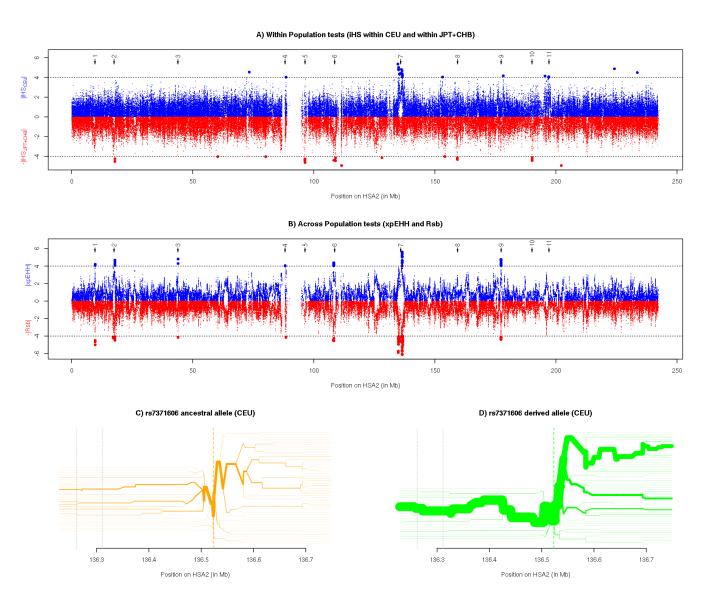


Figure 1. Analysis of the human chromosome 2 haplotype data sets (hg18 human genome assembly) for the CEU and JPT+CHB populations with REHH 2.0. A) Plot of iHS against physical distance, in the CEU (|iHS| in blue) and the JPT+CHB (- |iHS| in red) populations. B) Plot of XP-EHH (|XP-EHH| in blue) and Rsb (- |Rsb| in red) between the CEU and JPT+CHB populations. In A) and B), the horizontal dotted lines indicate the |iHS| significance threshold of 4 that was used to identify significant regions (see Table 1) and the arrows at the top of the graph indicate the mid-position of the significant regions described in Table 1). C) and D) Haplotype bifurcation diagrams drawn for the ancestral and derived allele, respectively, of the rs7377606 SNP in the CEU population (XP-EHH peak position of region #7 described in Table 1 and containing the LCT gene). In C) and D), the two grey vertical dotted lines delimit the LCT gene.

FIGURES 20

List of Tables

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1 21 2 Comparison of the computation times (in seconds) required to process input data files with the data2haplohh function for the versions 1.13 and 2.0 of the REHH package. Two data sets consisting, respectively, of 236 and 342 haplotypes of 110,200 SNPs for the CEU and JPT+CHB populations were considered (see the main text). For each of these datasets, the table gives the average computation times (\pm standard deviation) across ten independent runs, either with or without (in parentheses) allele 22 3 Comparison of the time (in seconds) required to compute the different EHHbased statistics for the versions 1.13 and 2.0 of the REHH, the selscan and the hapbin packages. For each analysis, the table gives the average computation time (\pm standard deviation) across ten independent runs. For each program, analyses were run either on a single thread or on four threads (except for REHH 1.13 version, which is not parallelised). 23 Correlation between the estimated iHS and XP-EHH statistics across the programs REHH (version 2.0), selscan and hapbin. Pairwise correlation for the iHS computed in the CEU and the JPT+CHB (in parenthesis) populations are given in the upper diagonal. Pairwise correlation for the XP-EHH computed across the CEU and JPT+CHB populations are given in the lower diagonal. 24

ID	Position a	Candidate Gene	Test	Peak Position ^b	Selected Population	
	(Size)	(position)	1680	1 Cak 1 Osition	(Overlap with other studies c)	
1	9.250-10.00	YWHAQ	ХР-ЕНН	9.700 (-4.21; 3)	JPT+CHB	
1	(0.75)	(9.641-9.688)	Rsb	9.701 (-5.00; 3)	31 1 CHB	
	(0.10)	(3.041-3.000)	iHS _{CEU}	9.700 (2.18; 0)		
			iHS _{JPT+CHB}	9.732 (3.52; 0)		
2	16.75-18.25	MSGN1	ХР-ЕНН	17.871 (-4.68; 17)	JPT+CHB	
-	(1.50)	(17.861-17.862)	Rsb	17.890 (-4.49 ; 8)		
	(1.50)	(17.001 17.002)	iHS _{CEU}	18.150 (3.68; 0)		
			iHS _{JPT+CHB}	17.856 (4.51; 2)		
3	43.50-44.25	ABCG8	XP-EHH	43.955 (-4.80 ; 2)	JPT+CHB	
	(1.75)	(43.919-43.959)	Rsb	43.957 (-4.15; 1)	J = 1 , 3 = =	
	(' ' ')		iHS_{CEU}	44.177 (-3.13; 0)		
			$iHS_{JPT+CHB}$	43.783 (3.86; 0)		
4	87.75-88.50	SMYD1	ХР-ЕНН	88.173 (-4.04; 2)	JPT+CHB	
	(0.75)	(88.148-88.194)	Rsb	88.187 (-3.53 ; 0)		
	,		iHS_{CEU}	88.173 (-2.80 ; 0)		
			$iHS_{JPT+CHB}$	88.198 (-3.01; 0)		
5	96.00-96.75	NCAPH	ХР-ЕНН	96.281 (2.32; 0)	JPT+CHB	
	(0.75)	(96.365-96.405)	Rsb	96.281 (3.07; 0)		
	, ,		iHS_{CEU}	96.609 (-3.88; 0)		
			$iHS_{JPT+CHB}$	96.403 (-4.61; 4)		
6	108.00-109.25	SULT1C2	XP-EHH	108.273 (-4.38 ; 18)	JPT+CHB	
	(1.25)	(108.271-108.292)	Rsb	108.253 (-4.55; 3)	(Vo., Ta., Sa.)	
		EDAR	iHS_{CEU}	109.016 (2.46; 0)		
		(108.877 - 108.972)	$iHS_{JPT+CHB}$	108.982 (4.44; 4)		
7	134.50-137.25	LCT	XP-EHH	136.523 (5.60; 16)	CEU	
	(2.75)	(136.262-136.311)	Rsb	136.533 (6.13; 71)	(Vo., Ta., Sa.)	
		MCM6	iHS_{CEU}	134.706 (-5.35; 19)		
		(136.314-136.335)	$iHS_{JPT+CHB}$	134.727 (-3.70; 0)		
8	159.00-159.75	PKP4	XP-EHH	159.381 (-2.96; 0)	JPT+CHB	
	(0.75)	(159.021 - 159.246)	Rsb	159.380 (-2.86; 0)		
			iHS_{CEU}	159.745 (2.86; 0)		
			$iHS_{JPT+CHB}$	159.293 (4.31; 2)		
9	177.00-177.75	n.a.	XP-EHH	177.338 (-4.77; 16)	JPT+CHB	
	(0.75)		Rsb	177.337 (-4.40; 7)	(Sa.)	
			iHS_{CEU}	177.336 (-2.57; 0)		
			$iHS_{JPT+CHB}$	177.108 (3.43; 0)		
10	189.75-190.50	SLC40A1	XP-EHH	190.040 (0.58; 0)	JPT+CHB	
	(0.75)	(190.133-190.154)	Rsb	189.919 (0.99; 0)		
			iHS_{CEU}	190.326 (2.92; 0)		
			$iHS_{JPT+CHB}$	190.177 (4.41; 3)		
11	196.75-197.50	HECW2	XP-EHH	196.794 (2.11; 0)	CEU	
	(0.75)	(196.772-197.166)	Rsb	196.755 (2.08; 0)	(Ta.)	
			iHS_{CEU}	197.030 (4.06; 2)		
			$iHS_{JPT+CHB}$	197.332 (2.32; 0)		

 $[^]a\mathrm{All}$ the position are given in Mb with respect to the hg18 human genome assembly

Table 1. Regions of HSA2 harboring strong signals of selection.

 $[^]b$ In parentheses: the value of the test statistics at the peak position; the number of SNPs in the window that have a test statistic (in absolute value) above the threshold of 4

^cSignificant tests of selection found in other studies for the same regions are indicated: Vo. stands for Voight et al (2006); Ta. stands for Tang et al (2007) and Sa. stands for Sabeti et al (2007)

	haplotype format	CEU haplotypes	CHB+JPT haplotypes
пенн 1.13	standard	$>36000^a (29.97 \pm 0.29)$	$>36000^a (34.62 \pm 0.60)$
REHH 2.0	standard	$9.858 \pm 0.39 \; (10.73 \pm 0.16)$	$14.56 \pm 0.17 \ (15.61 \pm 0.26)$
REHH 2.0	$transposed^b$	$7.882 \pm 0.10 \; (8.832 \pm 0.50)$	$11.80 \pm 0.20 \ (12.91 \pm 0.14)$

^aAs mentioned in the manual, REHH version 1.x is quite inefficient in allele recoding. Versions 1.x are also prone to error (e.g., if some alleles are coded as "T").

Table 2. Comparison of the computation times (in seconds) required to process input data files with the *data2haplohh* function for the versions 1.13 and 2.0 of the REHH package. Two data sets consisting, respectively, of 236 and 342 haplotypes of 110,200 SNPs for the CEU and JPT+CHB populations were considered (see the main text). For each of these datasets, the table gives the average computation times (± standard deviation) across ten independent runs, either with or without (in parentheses) allele recoding (using the option allele.recode).

busing the new option haplotype.in.columns=T

program	#threads	iHS_{ceu}	$iHS_{chb+jpt}$	ХР-ЕНН	Rsb	$Total^a$
пенн 1.13	1	1759 ± 29	3045 ± 31	n.a.	4803 ± 58	4805 ± 58
REHH 2.0	1	128 ± 1.0	140 ± 2.1	268 ± 1.8	268 ± 1.8	269 ± 1.8
	4	37.8 ± 0.3	40.2 ± 0.3	77.1 ± 0.5	77.1 ± 0.5	78.5 ± 0.5
selscan	1	1237 ± 17	1503 ± 29	3833 ± 100	n.a.	6573 ± 86
	4	324 ± 6.5	391 ± 6.5	969 ± 5.6	n.a.	1684 ± 9.3
hapbin	1	17.6 ± 0.2	20.0 ± 0.1	47.4 ± 0.2	n.a.	85.0 ± 0.3
	4	5.68 ± 0.7	7.42 ± 0.1	13.2 ± 0.0	n.a.	26.2 ± 0.7

^aIn REHH the function scanhh computes iHH and iES simultaneously. It therefore needs to be run only once per haplotype data set. As a result, computing XP-EHH (and/or Rsb) requires almost no extra time, once iHS for the two populations has been computed.

Table 3. Comparison of the time (in seconds) required to compute the different EHH-based statistics for the versions 1.13 and 2.0 of the REHH, the selscan and the hapbin packages. For each analysis, the table gives the average computation time (± standard deviation) across ten independent runs. For each program, analyses were run either on a single thread or on four threads (except for REHH 1.13 version, which is not parallelised)

	REHH	selscan	hapbin	
REHH	na	$0.991 \ (0.993)$	0.907 (0.945)	
selscan	0.985	na	0.907 (0.945)	
hapbin	0.986	0.994	na	

Table 4. Correlation between the estimated iHS and XP-EHH statistics across the programs REHH (version 2.0), selscan and hapbin. Pairwise correlation for the iHS computed in the CEU and the JPT+CHB (in parenthesis) populations are given in the upper diagonal. Pairwise correlation for the XP-EHH computed across the CEU and JPT+CHB populations are given in the lower diagonal.