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| 1 | Changes in the dynamical properties of the North Atlantic atmospheric |
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| 2 | circulation in the past 150 years |
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ABSTRACT

It is of fundamental importance to evaluate the ability of climate models to 13 capture the large-scale atmospheric circulation patterns. In the context of a 14 rapidly changing climate, it is equally crucial to quantify the robustness of the 15 modeled changes in the large-scale atmospheric dynamics. Here we approach 16 this problem from an innovative point of view based on dynamical systems 17 theory. We characterize the atmospheric circulation over the North Atlantic 18 in the CMIP5 historical simulations (1851 to 2000) in terms of two instan-19 taneous metrics: local dimension of the attractor and stability of trajectories. 20 We then use these metrics to compare the models to the 20CR reanalysis over 2 the same historical period. The comparison suggests that: i) all the models 22 capture the mean attractor properties and models with finer grids perform bet-23 ter; ii) extremes of the dynamical systems metrics match the same large-scale 24 patterns in most of the models; iii) changes in the attractor properties observed 25 for the 20CR reanalysis - studied by dividing the 1851-2000 period into 3 sub-26 periods of 50 years each - are not reproduced by the models; iv) some models 27 present significant changes in the dynamical systems metrics over time but 28 there is no agreement on the direction and on the intensity of the shifts. 29

30 1. Introduction

One of the main sources of uncertainty in determining the impact of climate change on ex-31 treme events is the forced response of atmospheric dynamics (Shepherd 2014; Field 2012). While 32 changes in observables such as surface temperature are easily diagnosed, shifts in the mid-latitude 33 atmospheric patterns have proved very difficult to quantify. Some advances have been made by 34 focussing on specific features such as atmospheric blocking (Kay et al. 2015; Cassou and Cattiaux 35 2016; Faranda et al. 2016b), which in turn influence the occurrence of European cold spells and 36 heat waves, but the broader appreciation of circulation changes is still unsatisfactory. Here we 37 address this knowledge gap by using a dynamical systems framework. We illustrate the power of 38 such an approach by considering the well-known Lorenz (1963) system, a conceptual model of 39 atmospheric convection consisting of three differential equations: 40

$$\dot{x} = \sigma(y - x) \quad \dot{y} = rx - y - xz \quad \dot{z} = xy - bz \tag{1}$$

where *x*, *y*, *z* represent respectively the convection strength, the temperature difference between the surface and the top of the troposphere and the asymmetry of the convection cells. The parameters σ , *r* are the Prandtl and the Rayleigh numbers, while *b* is a ratio of critical parameters. A trajectory of the Lorenz (1963) attractor is shown in blue in figure 1. The figure consists of 2000 points obtained by iterating the Lorenz equations with $\Delta t \simeq 0.035$, $\sigma = 28$, r = 10, b = 8/3with a Runge Kutta scheme of order 4.

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To study the effects of an external forcing, we increase σ by 2% with respect to the classical value. Figure 1 shows a trajectory at $\sigma = 28.5$ in magenta. The magenta trajectory favours higher values of the variable *z*, but the changes relative to the original trajectory depend on the point being considered: some points are not displaced, while some others are mapped elsewhere. Assuming no knowledge of the system other than the trajectories' paths, how could we determine whether they both belong to the Lorenz attractors with two different forcings? To answer this question we would need: i) to measure the dynamical properties of an ensemble of trajectories representing the two configurations; ii) to estimate the distance between the obtained trajectories and determine if the shift has changed the properties of the points in a detectable way.

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The atmospheric equivalent of a point on the Lorenz attractor is the ensemble of instantaneous 58 fields describing the atmosphere at a time t. To study the atmospheric circulation over the North 59 Atlantic we will focus on a single field: the sea-level pressure (SLP) over this region. The SLP 60 field reflects the major modes of variability affecting the North Atlantic (Hurrell 1995; Moore 61 et al. 2013) and can further be used to diagnose a wealth of other atmospheric features, ranging 62 from teleconnection patterns to storm track activity to atmospheric blocking e.g. (Rogers 1997; 63 Comas-Bru and McDermott 2014). The trajectories of our dynamical systems are the succession 64 of daily SLP fields from 27 CMIP5 models and the 20CR reanalysis over the period 1851 to 65 2000. In order to measure changes in the systems, one must be able to specify at each point (each 66 day) the local (daily) dynamical properties and track their evolution. Recent contributions to 67 dynamical systems analysis have proven that local properties of the trajectories are characterized 68 by two quantities: the local dimension and stability of the field considered (Lucarini et al. 2016; 69 Faranda et al. 2017). They correspond respectively to the rarity and the typical persistence of the 70 configuration. Faranda et al. (2017) have also shown that these two metrics can be connected to 71 the predictability of a given atmospheric state and that their extremes match climate extremes. 72

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In this work we will first assess whether the models and reanalysis present similar attractor prop erties over the full time period considered. To do this, we compute daily values for the dimension

⁷⁶ and stability of the SLP fields and study their average and extreme properties. Next, we study the ⁷⁷ attractors for three sub-periods of 50 years each and quantify their changes across the periods. We ⁷⁸ then compare the changes seen in the models to those observed in the reanalysis. Finally, we use ⁷⁹ statistical mechanics arguments to attribute the changes observed in the reanalysis to greenhouse ⁸⁰ forcing.

81 2. Data & Methods

We use daily model output from the historical simulations of 27 CMIP5 models (see Table 1). The data is publicly available from the CMIP5 archive (Taylor et al. 2012). We then compare these to the 20th Century Reanalysis (20CRv2c) ensemble mean dataset (Compo et al. 2011). The analysis focuses on the region $22.5^{\circ}N - 70^{\circ}N$ and $80^{\circ}W - 50^{\circ}E$.

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In order to compute the dynamical systems metrics, we combine the statistical tools of extreme value theory with the results obtained by Freitas et al. (2010) for Poincaré recurrences. The parameters mentioned in the introduction (local dimension *d* and stability θ) are computed for the points ζ on the attractor obtained as sequence of states of the system. The dynamical indicators are linked to the probability \mathscr{P} that a trajectory x(t) of the distances of a trajectory emerging from a sphere of center ζ and diameter 2ε , i.e. the recurrence rate of the configuration ζ . We briefly outline the physical meaning of these quantities and the way they are computed below.

⁹⁴ (*i*) Local Dimensions: The Freitas et al. (2010) theorem and its modification in Lucarini et al. ⁹⁵ (2012) states that the probability \mathscr{P} for chaotic attractors is a generalized Pareto distribution ⁹⁶ (Pickands III 1975). We first compute the distance δ between the SLP field ζ and all other obser-⁹⁷ vations along the trajectory. We then weight the time series of the distance:

$$g(x(t)) = -\log(\delta(x(t), \zeta)).$$

The reason for taking the logarithm is explained by Collet and Eckmann (2009): in the dynamical system set-up the logarithm increases the discrimination of small values of $\delta(x, y)$ which correspond to large values of g(x(t)). The generalized Pareto distribution then reduces to:

$$\mathscr{P}(g(x(t)) > q, \zeta) \simeq \exp(-[x - \mu(\zeta)]/\beta)$$

¹⁰¹ namely an exponential law whose parameters μ and σ depend on the point ζ chosen on the ¹⁰² attractor. Remarkably, $\beta(\zeta) = 1/d(\zeta)$, where $d(\zeta)$ is the local dimension around the point ζ . ¹⁰³ This result has recently been applied to SLP fields in Faranda et al. (2017).

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(*ii*) Local Stability: θ , the inverse of the residence time within a neighborhood of the config-105 uration, is exactly the extremal index introduced in extreme value theory to measure clustering 106 (Freitas et al. 2012; Faranda et al. 2016a). As in the extreme value theory, θ varies between 0 and 107 1. The value $\theta = 0$ corresponds to a stable fixed point of the dynamics where the observation ζ is 108 repeated infinite times (as for a pendulum left in its equilibrium position). This is of course never 109 observed in the atmospheric dynamics. A value of $\theta = 1$ indicates a point immediately leaving 110 the neighborhood of ζ . Since θ is the inverse of a persistence time, it depends on the Δt used. If 111 Δt is too large, the time dependence structure is hidden and θ will to be close to 1. If Δt is too 112 small, θ is close to zero. In Faranda et al. (2017) it has been observed that θ for SLP fields over 113 the North Altantic is between 0.3 and 0.5, when $\Delta t = 1$ day. In this work we use the same Δt . 114

Figure 2 illustrates of the meaning of the indicators: the local dimension *d* is the number of degrees of freedom needed to describe the dynamics of the system linearized around the state ζ and it is therefore proportional to the number of possible states resulting from ζ . The inverse of the persistence time θ is linked to the probability that the trajectory follows a path where each field resembles those of the previous and subsequent days.

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Before beginning the analysis, it is necessary to outline how the method of recurrence deals with 122 changes in the attractor. There are few theoretical results on non-stationary statistics of dynamical 123 systems, as well as on non-stationary extreme value theory. Luckily, the recurrence technique 124 also allows to bypass most of the technical difficulties linked to non-stationarity because the 125 dynamical properties are measured with respect to each single state of the attractor. If the change 126 affects the neighbourhood of a state, it will change its dynamical properties d and θ . If most of 127 the states are affected by the changes in the dynamics, then the average dimension of the attractor 128 and the average persistence will change accordingly. 129

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In order to test this idea, we again consider the two Lorenz (1963) systems discussed in the Introduction and perform two sets of 30 realizations (trajectories) at $\sigma = 28$ and $\sigma = 28.5$, each of them consisting of 50000 points. These values correspond respectively to the size of the CMIP5 ensemble (about 30) and to the number of days in the period 1850-2000 (about 50000). If the method of recurrence is capable to distinguish between the $\sigma = 28$ and $\sigma = 28.5$ Lorenz attractors, then the *d* and θ distributions of the two sets of realizations should be significantly different.

First, we have to define a metric to compare (d, θ) distributions. The simplest idea is to compute 139 the median of the d, θ distribution for each realization and verify that the clouds of median 140 centroids for $\sigma = 28$ and for $\sigma = 28.5$ are separated. This is shown in the top panel of Figure 141 3. As a further test, we can compute the empirical probability density functions $\Lambda(\sigma_{28})$ and 142 $\Lambda(\sigma_{28.5})$ of the pairwise distances between the d, θ median-based centroids. These distributions 143 should be significantly different from the $\Lambda(\sigma_{28})/\Lambda(\sigma_{28.5})$ distribution obtained by mixing 144 together the two sets of realizations. The visual inspection of Λ distributions (Figure 3-bottom 145 panel) suggests that distances computed when mixing the $\sigma = 28$ and $\sigma = 28.5$ realizations are 146 generally higher than those computed by realizations of the same attractor. This claim is sta-147 tistically supported by the Kolmogorov-Smirnov test (Von Storch 1999) results reported in Table 2. 148

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Table 2 also contains the results of the Kolmogorov-Smirnov test obtained when considering the 150 Wasserstein distances \mathscr{W}_2 (Villani 2008) between the full d, θ distributions. As described by Robin 151 et al. (2017), the Wasserstein distance is the proper tool to measure distance in multivariate set-ups 152 so that our median-based indicator should be tested against \mathcal{W}_2 . The test results suggest that the 153 median-based centroids are good proxies of the Wasserstein distances. We will see in the next 154 section that this result also hold for the CMIP5 ensemble. Since in the climate system we do not 155 have distinct trajectories (one before and one after climate change) we will divide the dataset in 156 three different periods imagining three separate trajectories obtained under different (greenhouse) 157 forcings. 158

3. Aggregate analysis of model and reanalysis attractors

We begin the analysis of the daily SLP fields from 1851 to 2000 by presenting the scatterplot of *d* versus θ for the 20CR reanalysis (Figure 4). The average of *d* is proportional to the

number of degrees of freedom needed to represent the systems' dynamics while the average of 162 θ is the inverse of the mean persistence time. Maxima (minima) of d correspond to the most 163 complex (simple) trajectories of the system. Maxima (minima) of θ correspond to the most 164 unstable (stable) trajectories (Messori et al. 2017). Such extremes are associated to specific 165 weather patterns that closely resemble the canonical North-Atlantic weather regimes. The 166 top panels in Figure 4 show the composite SLP anomalies for days beyond the 0.98 and 0.02 167 quantiles of the d and θ distributions. Maxima of θ reproduce an Atlantic Ridge pattern, 168 while minima of θ correspond to a negative North Atlantic Oscillation (NAO) phase. Similarly, 169 maxima of d correspond to a Blocking pattern and minima to a positive NAO (Faranda et al. 2017). 170 171

We next compare the (d, θ) bivariate histograms obtained for the 20CR with those computed 172 for the CMIP5 models (Figure 5-right). Two different behaviors emerge: some of the models 173 (e.g. CMCC-CMS reported in Figure 5-centre) yeld a single mode distribution resembling 174 that obtained for the 20CR; other models show bimodal distributions (e.g. the IPCC-CM5A) 175 histograms reported in Figure 5-bottom. We find the different behaviors to be related to the 176 seasonal cycle: in Figure 5-left, we plot the (d, θ) diagrams for the same models by coloring 177 each point according to the month of the year. In the 20CR and the CMCC-CMS model, the 178 different seasons are spread across the cloud, although maxima of θ mostly occur in winter. 179 The IPSL-CM5A displays a much stronger seasonal discrimination, with two distinct (d, θ) 180 clouds for the winter and for the summer seasons corresponding to the different modes of the 181 bivariate histograms. This implies that both the bulk statistics and the extremes are modified by 182 the seasonal cycle. 183

Given the variety of the possible behaviors, we will analyze separately the mean and the 185 extreme behavior of the dynamical properties. We report the aggregate analysis in Table 3 and 186 in Figure 7. The dots correspond to the centres of the ellipses and represent the median values 187 for each model; the semiaxes correspond to the standard deviation of the mean. Models are 188 numbered as in Tables 1, 3 and are ordered by increasing horizontal resolution. In Table 3 we 189 also provide the distances $\delta(d)$ and $\delta(\theta)$ from the median of the 20CR, the relative distances 190 $R(d) = \delta(d) / \max \delta(d)$ and $R(\theta) = \delta(\theta) / \max \delta(\theta)$ with respect to the farthest model and a 191 global score $R_{tot} = (R(d) + R(\theta))/2$. To check the validity of this global score, we compare 192 R_{tot} with the Wasserstein distance \mathscr{W}_2 between 20CR and the CMIP5 ensemble, computed as 193 described in Robin et al. (2017). The results are displayed in Figure 6. The two indicators are 194 so similar (Pearson coefficient: $r_{pear} = 0.90$ and Spearman coefficient $r_{spear} = 0.85$ (Von Storch 195 (1999)) that we will use the simpler R_{tot} when discussing our results. R_{tot} further indicates the 196 direction of the changes (larger or smaller d, θ) while \mathcal{W}_2 only provides this information if the 197 transport plan is computed (Villani 2008). the latter would be particularly complex to compute 198 for the dataset analysed here. 199

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²⁰¹ By using the R_{tot} metrics, we find that all the models are within one standard deviation of the ²⁰² 20CR ensemble mean. At the same time, most models display median values in *d* and θ which ²⁰³ are statistically different from those found in 20CR. The results of a Wilcoxon ranksum test ²⁰⁴ (Von Storch 1999) are reported in Table S1. We remark that both Figure 7 and Table S1 indicate ²⁰⁵ that models with a higher horizontal resolution have median values generally closer to those of ²⁰⁶ the 20CR.

Different models have different spreads in d and θ , making it worth to investigate the extremes 208 of these quantities and their relation with the weather regimes found in 20CR. Figures 8 and 9 209 display the composite SLP anomalies for the d and θ extremes - computed as in Figure 4 - for 210 the three closest and the three farthest models (in terms of R_{tot}) relative to the 20CR. The models 211 display similar composites independently their R_{tot} score. A quantitative analysis is reported in 212 Table S2 using the Root Mean Square Error (RMSE) between 20CR and CMIP5 SLP composite 213 anomalies. In general, we find the NAO- and NAO+ patterns to have higher RMSE whereas 214 Atlantic Ridge and Blocking pattern are better represented. 215

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The aggregate analysis shows that many CMIP5 models provide a dynamical picture coherent with that of the 20CR reanalysis. At the same time, most models reproduce statistically different medians of the two dynamical systems metrics relative to what found in the reanalysis.

4. Changes in the attractor properties

We next investigate whether the SLP's dynamical indicators have changed as an effect of 221 past greenhouse forcing by separating the results into three periods: 1851-1900, 1901-1950 and 222 1951-2000. To provide a visually immediate depiction of the changes, we compute the joint 223 histograms of (d, θ) . The left-hand side and middle columns in Figure 10 display these for 224 CCSM4 (the model with the highest horizontal resolution among those analysed) and 20CR. 225 The median values are highlighted by magenta lines. The right-hand side panels show the 226 differences between 20CR and CCSM4. Results for all the other models are shown in Figures 227 S1-S26. In the reanalysis, both the median of the local dimension d and the median of θ 228 increase from the first to the second and the third period. These increases, although small 229 relative to the metrics' spread, are significant at the 95% level under a Wilcoxon ranksum test 230

(Table S1). For the models, the sign and significance of the changes depend on the period
 considered and are often not significant (Table S1). As a general feature, models tend to repro duce the dynamical properties better for the period 1951-2000 (see Figure 10 and Figures S1-S26).

We also perform a simple analysis of extremes in d and θ for the three different periods. We 235 first compute the quantiles 0.02 and 0.98 for the d and θ distributions of the whole dataset. For the 236 20 CR these quantiles correspond to the black lines of Figure 4. Then we compute the number of 237 days falling beyond the 0.02 and 0.98 quantiles in each of the three different periods. The results 238 are shown in Figure 11. From top to bottom, we report the four extremes corresponding to Atlantic 239 Ridge (maxima of θ), NAO- (minima of θ), Blocking (maxima of d) and NAO+ (minima of d). 240 Again, the changes observed for the 20CR are generally larger than those observed for the models. 241 None of the models displays changes coherent with those observed in the reanalysis. The results 242 are stable with respect to reasonable changes in the quantiles. 243

5. Discussion and Conclusions

²⁴⁵ We have computed the instantaneous dynamical properties of the SLP fields for the 20CR and ²⁴⁶ the CMIP5 historical runs, over the period 1851-2000. The goal of our analysis was to assess ²⁴⁷ whether different models with different physics and resolutions quantitatively represent the *same* ²⁴⁸ dynamical system and therefore possess attractors with similar characteristics. The metrics we ²⁴⁹ used are the local dimension *d* and the inverse of the persistence time θ . As described in Faranda ²⁵⁰ et al. (2017), these two quantities give a complete characterization of the attractor of the system.

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²⁵² When the whole analysis period is considered, we find that the models successfully capture ²⁵³ some of the dynamical systems features identified in the reanalysis. For example, the range and variability of the dynamical metrics are consistent across the datasets. At the same time, some models exagerate the effects of the seasonal cycle on the dynamical indicators, and the statistical agreement in the median values of the metrics is generally poor. Models with higher horizontal resolutions tend to perform better. The SLP fields corresponding to extremes in *d* and θ are mostly similar across the models and reanalysis. The main differences are found for the minima of θ and *d*, which in 20CR correspond to the NAO- and NAO+ patterns.

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To detect the changes in the attractor properties with time, we have then divided the results into 261 three periods: 1851-1900, 1901-1950 and 1951-2000. We have analysed the joint histograms 262 of the d, θ variables and compared them to those obtained for the 20CR. The reanalysis shows 263 significant increases in d and θ throughout the time period analysed. These changes also reflect 264 in the number of days having extreme d and/or θ values. Days with Blocking and Atlantic Ridge 265 patterns increase in frequency with time, while days with NAO patterns are decreasing. This is 266 coherent with the results obtained by Alvarez-Castro et al. (2017). We note that the decrease in 267 the frequency of NAO patterns does not imply a more negative NAO index. Indeed, the frequency 268 of the NAO- pattern decreases more than that of the NAO+ pattern suggesting that – if anything 269 - the positive NAO phase becomes more dominant towards the end of the analysis period, in 270 agreement with SLP measurements (Hurrell et al. 2001). None of the models show comparable 271 changes: changes in d are mostly not significant and the shifts in θ are significantly smaller 272 than in the reanalysis. This is reflected in the fact that the median values of the two variables 273 in the models do not show a clear upward trend trhoughout the three sub-periods considered. 274 As a caveat we note that our analysis does not attempt to separate the forced variability from 275 natural low-frequency oscillations and that, especially during the first two periods analysed, 276 it is unclear whether the greenhouse forcing can be clearly discerned above the background 277

²⁷⁸ "climate noise" (Paeth et al. 1999; Lyu et al. 2015). We must therefore take into account the possi²⁷⁹ bility that the model's internal variability dominates over the trends for the time period considered.

As a further caveat we note that several studies (Krueger et al. 2013; Ferguson and Villarini 281 2012, 2014) have questioned the consistency of the 20CR dataset. So, how much can we trust 282 the results obtained when investigating the three separate periods? Most of the observations used 283 to constrain 20CR in the first part of the dataset are located in Europe or eastern North America 284 (Cram et al. 2015); the North Atlantic sector can therefore be expected to perform better than 285 elsewhere. The dataset has a sufficiently high horizontal resolution to obtain a good estimate of 286 the local dimension distribution (Faranda et al. 2017). However, the fact that the 20CR data is 287 increasingly constrained to follow the SLP observations as the time approaches the present day 288 causes a decrease of the ensemble spread with time, since the system is more closely pinned to 289 a specific manifold (the observations), without the possibility of exploring the full phase space. 290 This may explain the changes in the 20CR local dimension with time. 291

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If the results obtained for 20CR do not depend on the quality of the dataset but have a real phyis-293 cal meaning, the increase in dimension with time could be explained by using the results obtained 294 in Faranda et al. (2013) for simple dynamical systems. In these conceptual models, the dimension 295 increases with the *temperature* of the system defined – following the Einstein model for Brownian 296 motion – as the variance of a stochastic noise term added to the deterministic dynamics. Schubert 297 and Lucarini (2015) have demonstrated that this modeling approach is relevant for climate and, in 298 general, for any multiscale system. Faranda et al. (2013) have further shown that $\theta \to 1$ when the 299 temperature is increased, coherently with what observed in the 20CR. 300

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| 402 | TABLE 1. List of CMIP5 Models Analysed and 20CR reanalysis from 1851 to 2001. Models are ordered by |
|-----|---|
| 403 | increasing in resolution. |

| No. ¹ | Model Institution/ID | | Country | Resolution ² |
|------------------|----------------------|---|-----------|-------------------------|
| 1 | 20CRv2c | NOAA-CIRES | USA | 2 x 2 |
| 2 | CMCC-CESM | Centro Euro-Mediterraneo sui Cambiamenti Climatici | Italy | 3,75x3,75 |
| 3 | CanESM2 | Canadian Centre for Climate Modelling and Analysis, CCCMa | Canada | 2.81x 2.79 |
| 4 | MIROC-ESM-CHEM | MIROC^4 | Japan | 2,81x 2,79 |
| 5 | MIROC-ESM | $MIROC^4$ | Japan | 2,81x 2,79 |
| 6 | BCC-CSM1-1 | Beijing Climate Center | China | 2,81x 2,79 |
| 7 | IPSL-CM5B-LR | Institute Pierre Simon Laplace, IPSL | France | 3,75x1,89 |
| 8 | NorESM1-M | Norwegian Climate Center | Norway | 2,5x1,89 |
| 9 | FGOALS-2 | Institute of Atmospheric Physics, Chinese Academy of Sciences | China | 2,81x2,81 |
| 10 | MPI-ESM-P | Max Planck Institute for Meteorology, MPI | Germany | 1,87x1,87 |
| 11 | MPI-ESM-LR | Max Planck Institute for Meteorology, MPI | Germany | 1,87x1,87 |
| 12 | CSIRO-MK3-6-0 | CSIRO-BOM ⁵ | Australia | 1,87x1,87 |
| 13 | CMCC-CMS | Centro Euro-Mediterraneo sui Cambiamenti Climatici | Italy | 1,87x1,87 |
| 14 | MPI-ESM-MR | Max Planck Institute for Meteorology, MPI | Germany | 1,87x1,87 |
| 15 | IPSL-CM5A-MR | Institute Pierre Simon Laplace, IPSL | France | 2,5x1,26 |
| 16 | INM-CM4 | Institute for Numerical Mathematics, INM | Russia | 2x1,5 |
| 17 | ACCESS 1-0 | CSIRO-BOM ⁵ | Australia | 1,87x1,25 |
| 18 | MIROC5 | MIROC^4 | Japan | 1,40x1,40 |
| 19 | CNRM-CM5 | CNRM-CERFACS ³ | France | 1,40x1,40 |
| 20 | MRI-ESM1 | Meteorological Research Institute, MRI | Japan | 1,125x1,125 |
| 21 | BCC-CSM1-M | Beijing Climate Center | China | 1,125x1,125 |
| 22 | MRI-CGCM3 | Meteorological Research Institute, MRI | Japan | 1,125x1,125 |
| 23 | EC-EARTH | Danish Meteorological Institute, DMI | Denmark | 1,125x1,125 |
| 24 | CESM1-FASTCHEM | Community Earth System Model Contributors, NCAR | USA | 1,25x0,94 |
| 25 | CESM1-CAM5 | Community Earth System Model Contributors, NCAR | USA | 1,25x0,94 |
| 26 | CESM1-BGC | Community Earth System Model Contributors, NCAR | USA | 1,25x0,94 |
| 27 | CCSM4 | National Center for Atmospheric Research, NCAR | USA | 1,25x0,94 |

Order by horizontal resolution (Decreasing)
 ²Longitude x Latitude (°)
 ³Centre National de Recherches Meteorologiques - Centre Europeen de Recherche et de Formation Avance en Calcul Scientifique
 ⁴Atmosphere and Ocean Research Institute (University of Tokyo), National Institute for Environmental Studies, and Japan Agency for Marine-Earth Science and Technology
 ⁵Commonwealth Scientific and Industrial Research Organisation(CSIRO), Bureau of Meteorology(BOM)

TABLE 2. Two sample Kolmogorov Smirnov (KS) tests results and p-values for the distribution of distances between realizations of the Lorenz system for two different metrics: (\mathscr{W}_2) Wasserstein distance between full d, θ distributions, (*m*) distances between the medians of the d, θ distributions. The null hypothesis is that data in $\Lambda(1)$ and $\Lambda(2)$ are from the same continuous distribution, with a significance level of 95%. The null hypothesis is always rejected.

| | $\Lambda(\sigma_{28})/\Lambda(\sigma_{28/28.5})$ | $\Lambda(\sigma_{28.5})/\Lambda(\sigma_{28/28.5})$ |
|-------------------------------|--|--|
| KS(₩2) | 0.16 | 0.22 |
| p -value(\mathcal{W}_2) | $5 \cdot 10^{-10}$ | $2 \cdot 10^{-20}$ |
| KS(m) | 0.29 | 0.4 |
| p-value(<i>m</i>) | $3 \cdot 10^{-33}$ | $1 \cdot 10^{-63}$ |

TABLE 3. List of median values for dimension *d* and inverse of persistence time θ , distances $\delta(d)$ and $\delta(\theta)$ from the median of the 20CR, relative distances $R(d) = \delta(d) / \max \delta(d)$ and $R(\theta) = \delta(\theta) / \max \delta(\theta)$ with respect to the farthest model and global score $R_{tot} = (R(d) + R(\theta))/2$.

| N. | Model | median(d) | $\delta(d)$ | R(d) | $median(\theta)$ | $\boldsymbol{\delta}(\boldsymbol{	heta})$ | $R(\theta)$ | R _{tot} |
|----|----------------|-----------|-------------|------|------------------|---|-------------|------------------|
| 1 | 20CRv2c | 11,56 | - | - | 0.5 | - | - | _ |
| 2 | CMCC-CESM | 12,22 | 0,67 | 0,54 | 0,51 | 0,01 | 0,13 | 0,33 |
| 3 | CanESM2 | 11,99 | 0,43 | 0,35 | 0,51 | 0 | 00.2 | 0,19 |
| 4 | MIROC-ESM-CHEM | 12,54 | 0,98 | 0,8 | 0,47 | 0,04 | 0,75 | 0,77 |
| 5 | MIROC-ESM | 12,48 | 0,92 | 0,75 | 0,47 | 0,04 | 0,76 | 0,76 |
| 6 | BCC-CSM1 | 12,12 | 0,57 | 0,46 | 0,51 | 0,01 | 0,12 | 0,29 |
| 7 | IPSL-CM5B | 12,73 | 1,17 | 0,95 | 0,46 | 0,05 | 0,93 | 0,94 |
| 8 | NorESM1-M | 12,12 | 0,56 | 0,46 | 0,48 | 0,02 | 0,44 | 0,45 |
| 9 | FGOALS-S2 | 11,63 | 0,07 | 0,06 | 0,45 | 0,05 | 1,00 | 0,53 |
| 10 | MPI-ESM-P | 12,17 | 0,61 | 0,5 | 0,51 | 0 | 0,06 | 0,28 |
| 11 | MPI-ESM-LR | 12,13 | 0,58 | 0,47 | 0,51 | 0,01 | 0,14 | 0,3 |
| 12 | CSIRO-MK3-6-0 | 12,66 | 1,11 | 0,9 | 0,5 | 0 | 0,02 | 0,46 |
| 13 | CMCC-CMS | 11,95 | 0,39 | 0,32 | 0,52 | 0,01 | 0,22 | 0,27 |
| 14 | MPI-ESM-MR | 12,09 | 0,53 | 0,43 | 0,51 | 0 | 0,09 | 0,26 |
| 15 | IPSL-CM5A | 11,86 | 0,31 | 0,25 | 0,48 | 0,03 | 0,51 | 0,38 |
| 16 | INM-CM4 | 12,79 | 1,23 | 1 | 0,47 | 0,04 | 0,70 | 0,85 |
| 17 | ACCESS-1-0 | 11,74 | 0,19 | 0,15 | 0,49 | 0,02 | 0,31 | 0,23 |
| 18 | MIROC5 | 12,58 | 1,02 | 0,83 | 0,49 | 0,02 | 0,33 | 0,58 |
| 19 | CNRM-CM5 | 12,36 | 0,8 | 0,65 | 0,47 | 0,03 | 0,59 | 0,62 |
| 20 | MRI-ESM1 | 11,72 | 0,16 | 0,13 | 0,51 | 0,01 | 0,12 | 0,13 |
| 21 | BCC-CSM1-M | 11,45 | 0,11 | 0,09 | 0,55 | 0,05 | 0,91 | 0,5 |
| 22 | MRI-CGCM3 | 11,74 | 0,18 | 0,15 | 0,51 | 0 | 0,07 | 0,11 |
| 23 | EC-EARTH | 11,87 | 0,32 | 0,26 | 0,5 | 0 | 0,01 | 0,13 |
| 24 | CESM1-FASTCHEM | 11,56 | 0 | 0 | 0,51 | 0,01 | 0,18 | 0,09 |
| 25 | CESM1-CAM5 | 11,88 | 0,32 | 0,26 | 0,51 | 0,01 | 0,13 | 0,2 |
| 26 | CESM1-BGC | 11,53 | 0,02 | 0,02 | 0,51 | 0,01 | 0,11 | 0,06 |
| 27 | CCSM4 | 11,57 | 0,02 | 0,01 | 0,51 | 0,01 | 0,12 | 0,07 |

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FIG. 1. Two realizations of the Lorenz attractor. Blue: classic attractor $\Delta t \simeq 0.035$, $\sigma = 28$, r = 10, b = 8/3; Violet: $\sigma = 28.5$



FIG. 2. Schematic representation of the dynamical indicators: the local dimension *d* is proportional to the number of possible configurations originating that of the day analyzed and resulting from it. θ , is the inverse of the persistence time of a certain configuration.





FIG. 3. Top: Medians of *d* and θ for 30 realizations of the Lorenz attractor with $\sigma = 28$ (blue) and $\sigma = 28.5$ (red); the crosses display the median of the ensemble of realizations. Bottom: Empirical probability density functions (pdf) of the pairwise distances among *d* and θ medians in $\sigma = 28$ realizations ($\Lambda(\sigma_{28})$), $\sigma = 28.5$ realizations ($\Lambda(\sigma_{28})$) and mixed realizations ($\Lambda(\sigma_{28}/\sigma_{28.5})$).



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FIG. 11. Change in the number of extreme days, computed using the 0.98 and 0.02 quantiles of the full time series for each model, for three different sub-periods. From top to bottom different regimes: Atlantic Ridge (maxima of θ), NAO- (minima of θ), Blocking (maxima of d), NAO+ (minima of d)