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► **To cite this version:**

Philippe G. LeFloch, Yue Ma. The mathematical validity of the $f(r)$ theory of modified gravity. 2016.
hal-01423831

HAL Id: hal-01423831

<https://hal.science/hal-01423831>

Preprint submitted on 31 Dec 2016

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**THE MATHEMATICAL VALIDITY OF
THE $f(R)$ THEORY OF MODIFIED GRAVITY**

by

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Abstract. — We investigate the Cauchy problem for the $f(R)$ theory of modified gravity, which is a generalization of Einstein’s classical theory of gravitation. The integrand of the Einstein-Hilbert functional is the scalar curvature R of the spacetime, while, in modified gravity, it is a nonlinear function $f(R)$ so that, in turn, the field equations of the modified theory involve *up to fourth-order* derivatives of the unknown spacetime metric. We introduce here a *formulation of the initial value problem in modified gravity* when initial data are prescribed on a spacelike hypersurface. We establish that, in addition to the induced metric and second fundamental form (together with the initial matter content, if any), an initial data set for modified gravity must also provide one with the *spacetime scalar curvature* and its first-order time-derivative. We propose an *augmented conformal formulation* (as we call it), in which the spacetime scalar curvature is regarded as an *independent variable*. In particular, in the so-called wave gauge, we prove that the field equations of modified gravity are equivalent to a coupled system of *nonlinear wave-Klein-Gordon equations* with defocusing potential. We establish the consistency of the proposed formulation, whose main unknowns are the conformally-transformed metric and the scalar curvature (together with the matter fields) and we establish the existence of a maximal globally hyperbolic Cauchy development associated with any initial data set with sufficient Sobolev regularity when, for definiteness, the matter is represented by a massless scalar field. We analyze the so-called *Jordan coupling* and work with the so-called *Einstein metric*, which is conformally equivalent to the physical metric—the conformal factor depending upon the unknown scalar curvature. A main result in this paper is the derivation of quantitative estimates in suitably defined functional spaces, which are uniform in term of the nonlinearity $f(R)$ and show that spacetimes of modified gravity are ‘close’ to *Einstein spacetimes*, when the defining function $f(R)$ is ‘close’ to the Einstein-Hilbert integrand R . We emphasize that this is a highly singular limit problem, since the field equations under consideration are fourth-order in the metric, while the Einstein equations are second-order only. In turn, our analysis provides the first mathematically rigorous validation of the theory of modified gravity.

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1. Introduction

In recent years, new observational data have suggested that alternative theories of gravity, based on extensions of Einstein’s field equations of general relativity, may be relevant in order to explain the accelerated expansion of the Universe as well as certain instabilities observed in galaxies —without explicitly introducing notions such as ‘dark energy’ or ‘dark matter’. Among these theories, the so-called $f(R)$ –**theory of modified gravity** (associated with a prescribed function $f(R)$ of the scalar curvature R) was recognized as a physically viable alternative to Einstein’s theory. Despite the important role played by this theory in physics, the corresponding field equations have not been investigated by mathematicians yet. This is due to the fact that the modified gravity equations are significantly more involved than the Einstein equations: they contain up to *fourth-order derivatives* of the unknown metric, rather than solely second-order derivatives. Extensive works are available in the physical and numerical literature [3, 4, 8, 7, 13, 23, 24, 25]. The study of the well-posedness for this theory was also investigated earlier for instance in [9] by taking advantage of an equivalence with the Brans-Dicke theory. Furthermore, the function f is sometimes taken to be singular (and this leads to a further difficulty [5]), but here we assume this function to be regular.

Our purpose in this article is to initiate a rigorous mathematical study of the modified gravity equations and, specifically,

- to introduce a **notion of initial data set in modified gravity**,
- to describe an **initial value formulation** from an arbitrary spacelike hypersurface,
- to establish the **existence of a globally hyperbolic maximal development** associated with a given initial data set,
- and, importantly, to **provide a rigorous validation** that the modified gravity theory is an ‘approximation’ of Einstein’s theory, in sense that we will make precise with quantitative estimates. For definiteness, we will deal with asymptotically flat solutions, although our arguments are purely local and could be formulated in a domain of dependence of any initial data set. Our setting is appropriate in order to address the global nonlinear stability of Minkowski spacetime which we establish in the series of papers [18]–[22].

As already mentioned, in addition to the (second-order) Ricci curvature terms arising in the Einstein equations, the field equations of the $f(R)$ -theory involve fourth-order derivatives of the metric and, more precisely, second-order derivatives of the scalar curvature. The corresponding system of partial differential equations (after a suitable choice of gauge) consists of a system of *nonlinear wave equations*, which is significantly more involved than

the corresponding system derived from Einstein's equations. Yet, a remarkable mathematical structure is uncovered in the present work, which we refer to as the **augmented conformal formulation**:

- we introduce a *conformally equivalent metric* based on a conformal factor that depends upon the (unknown) scalar curvature,
- we proceed by introducing an *extended system* in which the metric and its scalar curvature are regarded as *independent unknowns*,
- we then establish the *well-posedness* of the initial value problem for this augmented formulation,
- and we finally explain how to recover the solutions to the original system of modified gravity.

Before we present our results in further details, let us first to recall that Einstein's theory is based on **Hilbert-Einstein's action**

$$(1.1) \quad \mathcal{A}_{\text{HE}}[\phi, g] := \int_M \left(\frac{R_g}{16\pi} + L[\phi, g] \right) dV_g$$

associated with a $(3 + 1)$ -dimensional spacetime (M, g) with Lorentzian signature $(-, +, +, +)$ whose canonical volume form is denoted by $dV = dV_g$. Here, and thereafter, we denote by $\text{Rm} = \text{Rm}_g$, $\text{Ric} = \text{Ric}_g$, and $R = R_g$ the Riemann, Ricci, and scalar curvature of the metric g , respectively. Observe that the above functional $\mathcal{A}_{\text{EH}}[g]$ is determined from the scalar curvature R_g and a Lagrangian $L[\phi, g]$, the latter term describing the matter content represented by one or several fields ϕ defined on M .

It is well-known that critical metrics for the action $\mathcal{A}_{\text{EH}}[g]$ (at least formally) satisfy Einstein's equation

$$(1.2) \quad G_g := \text{Ric}_g - \frac{R_g}{2} g = 8\pi T[\phi, g],$$

in which the right-hand side⁽¹⁾

$$(1.3) \quad T_{\alpha\beta}[\phi, g] := -2 \frac{\delta L}{\delta g^{\alpha\beta}}[\phi, g] + g_{\alpha\beta} L[\phi, g]$$

is referred to as the stress-energy tensor of the matter model. In the vacuum, for instance, these equations are equivalent to the Ricci-flat condition

$$(1.4) \quad \text{Ric}_g = 0.$$

The 'higher-order' gravity theory of interest is defined as follows. A smooth function $f : \mathbb{R} \rightarrow \mathbb{R}$ being prescribed, the **action of the $f(R)$ -modified gravity theory** reads⁽²⁾

$$(1.5) \quad \mathcal{A}_{\text{NG}}[\phi, g] =: \int_M \left(\frac{f(R_g)}{16\pi} + L[\phi, g] \right) dV_g,$$

⁽¹⁾Greek indices $\alpha, \beta = 0, \dots, 3$ represent spacetime indices.

⁽²⁾See Buchdahl [6], as well as the earlier proposal by Brans and Dicke [4].

whose critical points satisfy the **field equations of modified gravity**

$$(1.6) \quad \begin{aligned} N_g &:= f'(R_g) G_g - \frac{1}{2} \left(f(R_g) - R_g f'(R_g) \right) g + (g \square_g - \nabla d) (f'(R_g)) \\ &= 8\pi T[\phi, g]. \end{aligned}$$

The modified gravity tensor N_g thus “replaces” Einstein’s tensor G_g , while the right-hand side⁽¹⁾ is still given by the same expression (1.3). Observe that, by taking the trace of (1.6), we deduce the scalar equation

$$(1.7) \quad \operatorname{tr} N_g = f'(R_g) R_g - 2f'(R_g) + 3\square_g f'(R_g) = 8\pi \operatorname{tr} (T),$$

which can be regarded as an *evolution equation for the spacetime curvature* and will play an important role.

Concerning the matter content, we point out (cf. Section 2 below for the derivation) that the modified gravity tensor N_g is *divergence free*, that is,

$$(1.8) \quad \nabla^\alpha N_{\alpha\beta} = 0,$$

so that the matter field satisfies the **matter evolution equation**

$$(1.9) \quad \nabla^\alpha T_{\alpha\beta} = 0.$$

Furthermore, in order for the nonlinear theory to be a formal extension of the classical theory, we assume that $f(R) \simeq R$ in the zero curvature limit $R \rightarrow 0$. Since we will see later that the (positive) sign of the coefficient $\kappa := f''(0) > 0$ is critical for nonlinear stability, it is convenient to set

$$(1.10) \quad f'(R) = 1 + \kappa(R + \kappa\mathcal{O}(R^2)),$$

which after integration yields

$$(1.11) \quad f(R) = R + \kappa \left(\frac{R^2}{2} + \kappa\mathcal{O}(R^3) \right).$$

By definition, the remainder $\mathcal{O}(z^2)/z^2$ remains bounded when $z \rightarrow 0$ (uniformly in κ , if this parameter is taken to vary). In particular, the function f is increasing and strictly convex in a neighborhood of the origin and, therefore, one-to-one. In particular, the term $\kappa\mathcal{O}(R^3)$ in (1.11) could be taken to vanish identically, which corresponds to the quadratic action $\int_M \left(R_g + \frac{\kappa}{2} (R_g)^2 + 16\pi L[\phi, g] \right) dV_g$ often treated in the physical literature.

As we will see, in local coordinates, the field equations (1.6) take the form of a nonlinear system of fourth-order partial differential equations (PDE’s), while the Einstein equation (1.2) leads to only second-order equations. Our challenge in the present work is investigating the role of these fourth-order terms and generalizing the mathematical methods that were originally developed for Einstein’s equations. Furthermore, one formally would expect to recover Einstein’s theory by letting the coefficient κ tend to zero. However, this limit is *very singular*, since this involves analyzing the convergence of a fourth-order system (of no well-defined type) to a system of second-order (hyperbolic-elliptic) PDE’s.

⁽¹⁾further discussed shortly below

Before we can proceed further, we need to make an important observation concerning the modeling of the matter content of the spacetime. In the physics literature, the choice of the frame⁽¹⁾ in which measurements are made is still somewhat controversial as explained in [8, 7, 23]. Yet, this issue is essential in order for properly formulating the coupling between the gravity equations and the matter fields. Two standpoints were proposed by physicists. In the so-called “Jordan frame”, the original metric $g_{\alpha\beta}$ is considered to be the physically relevant metric, while in the “Einstein frame”, the **conformally-transformed metric**

$$(1.12) \quad g^\dagger_{\alpha\beta} := f'(R_g)g_{\alpha\beta}$$

is considered to be the physically relevant metric. In the present work, these two approaches will be referred to as the “Jordan coupling” and “Einstein coupling” for the matter. Hence, the “Jordan coupling” refers to the minimal coupling of the matter field to the geometry of the spacetime (represented by the tensor N_g) described by the “Jordan metric” (i.e. the original metric) $g_{\alpha\beta}$. On the other hand, the “Einstein coupling” refers to the minimal coupling of the matter field to the geometry of the spacetime described by the metric $g^\dagger_{\alpha\beta}$.

It is important to observe that different matter couplings lead to different physical theories, which may or may not be equivalent to each other. Of course, a given physical theory can also be expressed in various choices of metrics, that is, for the problem under consideration, the “Jordan coupling” could also be expressed with the “Einstein metric” $g^\dagger_{\alpha\beta}$, while the “Einstein coupling” could also be stated in the “Jordan metric” $g_{\alpha\beta}$. A coupling which is minimal (in the sense that the action takes the decoupled form (1.5)), in general, will no longer be minimal in another choice of metric. This suggests that the Einstein metric is not the physical metric in the Jordan coupling theory, while the Jordan metric is not the physical metric in the Einstein coupling. This has apparently led to great confusion and controversies in the physical literature, until some clarification was brought in the most recent contributions [8, 7]. Observe that our notion of Einstein coupling is equivalent to the notion of Einstein frame adopted in [23].

In this work, we will treat the Jordan coupling but expressed in the (conformal) Einstein metric g^\dagger . This coupling has the minimal form (1.5), but only in the original metric g . If one would insist on stating the problem in terms of the Einstein metric, then the coupling would not be minimal. This presentation appears to be optimal from the standpoint of establishing a well-posed theory for the initial value problem.

Throughout this article, the matter model of interest is a **massless scalar field**, defined by its standard stress-energy tensor, and we consider the following two possible couplings:

$$(1.13) \quad \begin{aligned} T_{\alpha\beta} &:= \nabla_\alpha\phi\nabla_\beta\phi - \frac{1}{2}g_{\alpha\beta}g^{\delta\lambda}\nabla_\delta\phi\nabla_\lambda\phi, \\ T^\dagger_{\alpha\beta} &:= f'(R_g)\left(\nabla_\alpha\phi\nabla_\beta\phi - \frac{1}{2}g_{\alpha\beta}g^{\delta\lambda}\nabla_\delta\phi\nabla_\lambda\phi\right), \end{aligned}$$

and, for convenience, the Einstein coupling is stated in the Jordan metric. As should be expected from the above discussion, different choices of coupling lead to systems of PDE's

⁽¹⁾From a mathematical standpoint, all frames are of course equivalent.

of rather different nature. In fact, we will show that the Einstein coupling leads to an *ill-defined Cauchy problem*. Therefore, in the rest of this section, we restrict attention to the Jordan coupling.

We are now in a position to state our existence theory, at least in a preliminary form. We recall that the initial value problem for the Einstein equations is classically formulated as follows. (We refer to the textbook by Choquet-Bruhat [10] for the terminology and historical references.) Given a Riemannian 3-manifold (\bar{M}, \bar{g}) together with a 2-covariant tensor field K (plus suitable matter data) satisfying certain constraint equations, one seeks for a (globally hyperbolic) development of this initial data set. By definition, such a development consists of a Lorentzian manifold (M, g) satisfying the Einstein equations such that \bar{M} is embedded in M as a spacelike hypersurface with induced metric \bar{g} and second fundamental form K . The maximal (globally hyperbolic) development, by definition, is the unique development of the initial data set in which any such development can be isometrically embedded.

In short, our formulation of the initial value problem for the theory of modified gravity is as follows. Since the field equations (1.6) are fourth-order in the metric, additional initial data are required, which are denoted by R_0, R_1 and are specified on the initial slice \bar{M} : they represent the scalar curvature and the time derivative of the scalar curvature of the (to-be-constructed) spacetime. They must of course also satisfy certain Gauss-Codazzi-type constraints. In addition, since the matter is modeled by a scalar field, say ϕ , we also prescribe some initial data denoted by ϕ_0, ϕ_1 , and representing the initial values of the scalar field and its time derivative, respectively. The preliminary statement above will be made more precise in the course of our analysis and all necessary terminology will be introduced. For definiteness, the results are stated with asymptotically flat data, but this is unessential.

Theorem 1.1 (Cauchy developments in the theory of modified gravity)

Consider the field equations (1.6) for the theory of modified gravity based on a function $f = f(R)$, satisfying (1.10), and assume that the matter is described by a scalar field with Jordan coupling (1.13). Given an asymptotically flat initial data set⁽¹⁾ $(\bar{M}, \bar{g}, K, R_0, R_1, \phi_0, \phi_1)$, there exists a unique maximal globally hyperbolic development (M, g) of these data, which satisfies⁽²⁾ the modified gravity equations (1.6). Furthermore, if an initial data set $(\bar{M}, \bar{g}, K, R_0, R_1, \phi_0, \phi_1)$ for modified gravity is “close” (in a sense that will be made precise later on) to an initial data set $(\bar{M}, \bar{g}', K', \phi'_0, \phi'_1)$ for the classical Einstein theory, then the corresponding development of modified gravity is also close to the corresponding Einstein development. This statement is uniform in term of the gravity parameter κ and modified gravity developments converge to Einstein developments when $\kappa \rightarrow 0$.

⁽¹⁾in the sense of Definition 6.2, below

⁽²⁾in the sense of Definition 2.6, below

Our results provide the first mathematically rigorous proof that the theory of modified gravity admits a well-posed Cauchy formulation and, furthermore, can be regarded as a “correction” to Einstein’s classical theory, as indicated by physicists.

A key contribution of the present work is a re-formulation of the field equations of modified gravity as a system of second-order hyperbolic equations and, more precisely, a coupled *system of wave-Klein-Gordon equations*. Note that wave-Klein-Gordon systems have brought a lot of attention in mathematical analysis: see, for instance, Bachelot [1, 2], Delort et al. [11, 12], Holzegel and Warnick [15], Katayama [16], Lannes [17], and LeFloch and Ma [18]–[22] and the references therein. For further results on the mathematical aspects of the $f(R)$ theory, we refer to [19]. We advocate here the use of *wave coordinates associated with the Einstein metric* and our formulation in such a gauge leads us to propose the following definition. Importantly, our formulation below contains an *augmented variable* denoted by ρ , which represents the scalar curvature of the spacetime⁽¹⁾.

Definition 1.2. — *The augmented conformal formulation of the field equations of modified gravity (with Jordan coupling and in wave coordinates associated with the Einstein metric) reads:*

$$(1.14) \quad \begin{aligned} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} g^{\dagger}_{\alpha\beta} &= F_{\alpha\beta}(g^{\dagger}; \partial g^{\dagger}, \partial g^{\dagger}) - 12 \partial_{\alpha\rho} \partial_{\beta}\rho + V(\rho) g^{\dagger}_{\alpha\beta} - 16\pi \partial_{\alpha}\phi \partial_{\beta}\phi, \\ g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \phi &= -2g^{\dagger\alpha\beta} \partial_{\alpha}\phi \partial_{\beta}\rho, \\ g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \rho - \frac{\rho}{3\kappa} &= W(\rho) - \frac{4\pi}{3e^{2\rho}} g^{\dagger\alpha\beta} \partial_{\alpha}\phi \partial_{\beta}\phi, \\ g^{\dagger\alpha\beta} \Gamma^{\lambda}_{\alpha\beta} &= 0, \end{aligned}$$

in which $F_{\alpha\beta}(g^{\dagger}; \partial g^{\dagger}, \partial g^{\dagger})$ are quadratic expressions (defined in Section 3 below), ∂g^{\dagger} is determined by the Ricci curvature, and the function $V = V(\rho)$ and $W = W(\rho)$ are of quadratic order as $\rho \rightarrow 0$.

Clearly, we recover Einstein equations by letting $\kappa \rightarrow 0$ and thus $f(R) \rightarrow R$. Namely, we will show that $\rho \rightarrow 0$ so that (1.14) reduces to the standard formulation in wave coordinates [10]. In particular, in this limit, we do recover the expression $R = 8\pi \nabla_{\alpha}\phi \nabla^{\alpha}\phi$ of the scalar curvature in terms of the norm of the scalar field.

An outline of the rest of this article is as follows. In Sections 2 and 3, we formulate the initial value problem first in the Jordan metric and then in the Einstein metric. We find that the second formulation is simpler, since the Hessian of the scalar curvature is eliminated by the conformal transformation. Furthermore, we demonstrate that the Einstein coupling is ill-posed. The conformal formulation is analyzed in Section 4, where the wave gauge is introduced and the wave-Klein-Gordon structure of the field equations is exhibited. Section 5 contains one of our main result and proposes an augmented formulation of the conformal system of modified gravity. The local existence theory with bounds that are

⁽¹⁾specifically $\rho = \frac{1}{2} \ln f'(R_g)$

uniform in κ is developed in Sections 6 to 8 and leads us in Section 9 to our main statement concerning the comparison between the modified and the classical theories.

2. Formulation of the Cauchy problem in the Jordan metric

2.1. The 3 + 1 decomposition of spacetimes

In this section, we formulate the initial value problem for the modified gravity system, by prescribing suitable initial data on a spacelike hypersurface. We follow the presentation in the textbook [10] where the classical gravity is studied. We are interested in a time-oriented spacetime (M, g) endowed with a Lorentzian metric g with signature $(-, +, +, +)$, which is homeomorphic to $[0, t_{\max}) \times M_t$ and admits a global foliation by spacelike hypersurfaces $M_t \simeq \{t\} \times \overline{M}$. The foliation is determined by a time function $t : M \rightarrow [0, t_{\max})$ and a three-dimensional manifold \overline{M} and, throughout, we assume that

$$(2.1) \quad M \text{ is globally hyperbolic and every } M_t \text{ is a Cauchy surface.}$$

This ensures that a wave equation with initial data posed on any such Cauchy surface enjoys the local existence and uniqueness property. (See [10, 14] for the definitions.)

We introduce local coordinates adapted to the above product structure, that is, $(x^\alpha) = (x^0, x^i) = (t, x^i)$, and we call the basis of vectors (∂_i) the **natural frame** defined on each slice M_t . This also provides us with a 'natural frame' (∂_t, ∂_i) on the spacetime M . By definition, the **Cauchy adapted frame** is given by the vectors $e_i = \partial_i$ and $e_0 = \partial_t - \beta^i \partial_i$, where $\beta = \beta^i \partial_i$ a time-dependent field, tangent to M_t , called the **shift vector**. We impose the restriction that e_0 is orthogonal to each M_t .

We also introduce the dual frame (θ^α) of the Cauchy adapted frame (e_α) by setting

$$(2.2) \quad \theta^0 := dt, \quad \theta^i := dx^i + \beta^i dt,$$

so that the spacetime metric reads

$$(2.3) \quad g = -N^2 \theta^0 \theta^0 + g_{ij} \theta^i \theta^j,$$

where the function $N > 0$ is referred to as the **lapse function** of the foliation. The Levi-Civita connection ∇ associated with g is represented by the set of connection coefficients $\omega_{\gamma\alpha}^\beta$, defined by

$$(2.4) \quad \nabla e_\alpha = \omega_{\gamma\alpha}^\beta \theta^\gamma \otimes e_\beta$$

and, consequently,

$$(2.5) \quad \nabla \theta^\alpha = -\omega_{\gamma\beta}^\alpha \theta^\gamma \otimes \theta^\beta.$$

We denote by $\bar{g} = \bar{g}_t$ the induced Riemannian metric associated with the slices M_t and by $\bar{\nabla}$ the Levi-Civita connection of \bar{g} , whose Christoffel symbols (in the natural frame)

are denoted by $\bar{\Gamma}_{ab}^c$. We can also introduce the extrinsic curvature of the slices or **second fundamental form** $K = K_t$, defined by

$$(2.6) \quad K(X, Y) := g(\nabla_X n, Y)$$

for all vectors X, Y tangent to the slices M_t , where n denotes the future-oriented, unit normal to the slices. In the Cauchy adapted frame, it reads

$$(2.7) \quad K_{ij} = -\frac{1}{2N} (\langle e_0, g_{ij} \rangle - g_{lj} \partial_i \beta^l - g_{il} \partial_j \beta^l).$$

Here, and throughout this article, we use the notation $\langle e_0, g_{ij} \rangle$ for the action of the vector field e_0 on the function g_{ij} . Next, we define the **time-operator** $\bar{\partial}_0$ acting on a two-tensor defined on the slice M_t by

$$(2.8) \quad \bar{\partial}_0 T_{ij} = \langle e_0, T_{ij} \rangle - T_{lj} \partial_i \beta^l - T_{il} \partial_j \beta^l,$$

which, by definition, is a two-tensor on M_t . With this notation, we thus have

$$(2.9) \quad K_{ij} = -\frac{1}{2N} \bar{\partial}_0 \bar{g}_{ij}.$$

Elementary calculations (see, for instance, [10, Section VI.3]) yield us the connection coefficients in terms of the (3 + 1)-decomposition:

$$(2.10) \quad \begin{aligned} \omega_{00}^0 &= N^{-1} \langle e_0, N \rangle, \\ \omega_{00}^i &= N g^{ij} \partial_j N, & \omega_{0i}^0 &= \omega_{i0}^0 = N^{-1} \partial_i N, \\ \omega_{ij}^0 &= \frac{1}{2} N^{-2} (\langle e_0, g_{ij} \rangle - g_{hj} \partial_i \beta^h - g_{ih} \partial_j \beta^h) = -N^{-1} K_{ij}, \\ \omega_{0j}^i &= -N K_j^i + \partial_j \beta^i, & \omega_{j0}^i &= -N K_j^i, \\ \omega_{jk}^i &= \bar{\Gamma}_{jk}^i. \end{aligned}$$

Here, $\bar{\Gamma}_{jk}^i$ denotes the Christoffel symbol of the connection $\bar{\nabla}$ in the coordinates $\{x^i\}$.

It is also a standard matter to derive the Gauss–Codazzi equations for each slice:

$$(2.11) \quad \begin{aligned} R_{ij,kl} &= \bar{R}_{ij,kl} + K_{ik} k_{lj} - K_{il} K_{kj}, \\ R_{0i,jk} &= N (\bar{\nabla}_j K_{ki} - \bar{\nabla}_k K_{ji}), \\ R_{0i,0j} &= N (\bar{\partial}_0 K_{ij} + N K_{ik} K_j^k + \bar{\nabla}_i \partial_j N). \end{aligned}$$

In addition by suitable contractions of these identities, we arrive at

$$(2.12a) \quad R_{ij} = \bar{R}_{ij} - \frac{\bar{\partial}_0 K_{ij}}{N} + K_{ij} K_l^l - 2K_{il} K_j^l - \frac{\bar{\nabla}_i \partial_j N}{N},$$

$$(2.12b) \quad R_{0j} = N (\partial_j K_l^l - \bar{\nabla}_l K_j^l),$$

and, for the (0, 0)-component of the Einstein curvature,

$$(2.12c) \quad G_{00} = \frac{N^2}{2} (\bar{R} - K_{ij} K^{ij} + (K_l^l)^2).$$

These equations clarify the relations between the geometric objects of the spacetime M and the ones of the slices M_t .

The equation (2.12a) yields the evolution of the tensor K and together with the definition $\partial_0 \bar{g} = -2NK$, we thus find the following first-order system satisfied by the metric and the second fundamental form:

$$(2.13) \quad \begin{aligned} \bar{\partial}_0 \bar{g}_{ij} &= -2NK_{ij}. \\ \bar{\partial}_0 K_{ij} &= N(\bar{R}_{ij} - R_{ij}) + NK_{ij}K_l^l - 2NK_{il}K_j^l - \bar{\nabla}_i \partial_j N. \end{aligned}$$

2.2. Evolution and constraint equations

Our objective is to combine the equations (2.12) and the field equations (1.6) in order to derive the fundamental equations of modified gravity. We recall first some elementary identities about the Hessian of a function expressed in the Cauchy adapted frame. Given any smooth function $f : M \rightarrow \mathbb{R}$, we can write

$$\nabla df = \nabla(\langle e_\gamma, f \rangle \theta^\gamma) = \langle e_\beta, \langle e_\alpha, f \rangle \rangle \theta^\beta \otimes \theta^\alpha - \langle e_\gamma, f \rangle \omega_{\beta\alpha}^\gamma \theta^\beta \otimes \theta^\alpha.$$

First of all, we compute the components

$$(2.14) \quad \begin{aligned} \nabla_i \nabla_j f &= \partial_i \partial_j f - \langle e_\gamma, f \rangle \omega_{ij}^\gamma = \partial_i \partial_j f - \bar{\Gamma}_{ij}^k \partial_k f - \omega_{ij}^0 \langle e_0, f \rangle \\ &= \bar{\nabla}_i \bar{\nabla}_j f + K_{ij} N^{-1} (\partial_t - \beta^l \partial_l) f \\ &= \bar{\nabla}_i \bar{\nabla}_j f + K_{ij} \mathcal{L}_n f, \end{aligned}$$

where \mathcal{L}_n is the Lie derivative associated with the normal unit vector of the slice M_t . Then, for the other components, we find

$$\begin{aligned} \nabla_j \nabla_0 f &= \nabla_0 \nabla_j f = \langle e_0, \partial_j f \rangle - \langle e_0, f \rangle \omega_{0j}^0 - \partial_i f \omega_{0j}^i \\ &= (\partial_t - \beta^i \partial_i) \partial_j f - (\partial_t - \beta^i \partial_i) f N^{-1} \partial_j N + N \partial_i f K_j^i - \partial_i f \partial_j \beta^i \\ &= (\partial_t - \beta^i \partial_i) \partial_j f - (\partial_t - \beta^i \partial_i) f \partial_j \ln N + N \partial_i f K_j^i - \partial_i f \partial_j \beta^i \\ &= \partial_j (\partial_t - \beta^i \partial_i) f - (\partial_t - \beta^i \partial_i) f \partial_j \ln N + N \partial_i f K_j^i \\ &= N \partial_j (N^{-1} (\partial_t - \beta^i \partial_i) f) + NK_j^i \partial_i f \end{aligned}$$

and

$$\begin{aligned} \nabla_0 \nabla_0 f &= (\partial_t - \beta^i \partial_i) (\partial_t - \beta^i \partial_i) f - (\partial_t - \beta^i \partial_i) f \omega_{00}^0 - \partial_i f \omega_{00}^i \\ &= (\partial_t - \beta^i \partial_i) (\partial_t - \beta^i \partial_i) f - (\partial_t - \beta^i \partial_i) f (N^{-1} \partial_t N) - \partial_i f (N^{-1} (\partial_t - \beta^i \partial_i) N) \\ &= (\partial_t - \beta^i \partial_i) (\partial_t - \beta^i \partial_i) f - (\partial_t - \beta^i \partial_i) f \cdot \partial_t \ln N - \partial_i f (\partial_t - \beta^i \partial_i) \ln N. \end{aligned}$$

In particular, the trace of the Hessian of a function is the so-called wave operator, expressed in the Cauchy adapted frame as

$$\begin{aligned} \square_g f &= g^{\alpha\beta} \nabla_\alpha \nabla_\beta f = -N^2 \nabla_0 \nabla_0 f + g^{ij} \nabla_i \nabla_j f \\ &= -N^2 \nabla_0 \nabla_0 f + \bar{g}^{ij} \bar{\nabla}_i \bar{\nabla}_j f + \bar{g}^{ij} K_{ij} N^{-1} (\partial_t - \beta^l \partial_l) f \\ &= -N^2 \nabla_0 \nabla_0 f + \Delta_{\bar{g}} f + \bar{g}^{ij} K_{ij} \mathcal{L}_n f, \end{aligned}$$

where $\Delta_{\bar{g}}f$ is the Laplace operator associated with the metric \bar{g} .

To proceed with the formulation of the field equation (1.6), we need first to rewrite it in a slightly different form, by defining the tensor:

$$(2.15) \quad E_{g_{\alpha\beta}} := N_{g_{\alpha\beta}} - \frac{1}{2} \text{tr} (N_g) g_{\alpha\beta},$$

where $\text{tr} (\cdot)$ is the trace with respect to the metric g . Then, we have the following relation in terms of the Ricci tensor:

$$(2.16) \quad E_{g_{\alpha\beta}} = f'(R_g)R_{\alpha\beta} + \frac{1}{2}f'(R)W_1(R_g)g_{\alpha\beta} - \left(\frac{1}{2}g_{\alpha\beta}\square_g + \nabla_\alpha \nabla_\beta \right) f'(R_g),$$

where we have introduced the **function** W_1 by

$$(2.17) \quad W_1(r) := \frac{f(r) - rf'(r)}{f'(r)}, \quad r \in \mathbb{R}.$$

In view of (1.6), we know that E_g satisfies the field equations

$$(2.18) \quad E_{g_{\alpha\beta}} = 8\pi(T_{\alpha\beta} - \frac{1}{2}\text{tr} (T)g_{\alpha\beta}) =: 8\pi H_{\alpha\beta},$$

where we have introduced the new **matter tensor** $H_{\alpha\beta}$. More precisely, it will be most convenient to introduce, for different components, a different form of the equations, that is, we write the field equations as:

$$(2.19) \quad \begin{aligned} E_{g_{ij}} &= 8\pi H_{ij}, \\ E_{g_{0j}} &= 8\pi H_{0j}, \\ N_{g_{00}} &= 8\pi T_{00}, \end{aligned}$$

or, equivalently,

$$(2.20a) \quad f'(R_g)R_{ij} + \frac{1}{2}f'(R)W_1(R_g)g_{ij} - \left(\frac{1}{2}g_{ij}\square_g + \nabla_i \nabla_j \right) f'(R_g) = 8\pi(T_{ij} - \frac{1}{2}\text{tr} (T)g_{ij}),$$

$$(2.20b) \quad f'(R_g)R_{0j} - \nabla_0 \nabla_j f'(R_g) = 8\pi T_{0j},$$

$$(2.20c) \quad f'(R_g)G_{g_{00}} - \frac{1}{2}f'(R_g)W_1(R_g)g_{00} + (g_{00}\square_g - \nabla_0 \nabla_0)(f'(R_g)) = 8\pi T_{00}.$$

For completeness we check the following equivalence.

Lemma 2.1. — *If a metric $g_{\alpha\beta}$ and a matter tensor $T_{\alpha\beta}$ satisfy the field equations (1.6), then they also satisfy (2.19). The converse is also true.*

Proof. — The equations (2.19) are clearly equivalent to

$$\begin{aligned} (N_g - 8\pi T)_{ij} &= -\frac{1}{2}g_{ij}\text{tr} (8\pi T - N_g), \\ (N_g - 8\pi T)_{0j} &= 0, \\ N_{g_{00}} - 8\pi T_{00} &= 0. \end{aligned}$$

By taking the trace of the tensor $(N_g - 8\pi T)$, we find $\text{tr}(N_g - 8\pi T) = -\frac{3}{2}\text{tr}(8\pi T - N_g)$ and thus $\text{tr}(8\pi T - N_g) = 0$, which proves the result. \square

Hence, in view of (2.20a) and by using (2.14), we have arrived at the field equations of modified gravity in a preliminary form. First of all, we have

$$\begin{aligned}
(2.21a) \quad R_{ij} &= \frac{1}{f'(R_g)} \left(E_{g_{ij}} - \frac{1}{2} f'(R_g) W_1(R_g) g_{ij} + \frac{1}{2} (g_{ij} \square_g + 2 \nabla_i \nabla_j) f'(R_g) \right) \\
&= \frac{E_{ij}}{f'(R_g)} - \frac{1}{2} W_1(R_g) \bar{g}_{ij} + \frac{(\bar{g}_{ij} \Delta_{\bar{g}} + 2 \bar{\nabla}_i \bar{\nabla}_j) f'(R_g)}{2 f'(R_g)} \\
&\quad - \frac{\bar{g}_{ij} \nabla_0 \nabla_0 f'(R_g)}{2 N^2 f'(R_g)} + (K_{ij} + \frac{1}{2} \bar{g}_{ij} K) \mathcal{L}_n \ln(f'(R_g)),
\end{aligned}$$

where $K := \bar{g}^{ij} K_{ij}$ is the trace of K with respect to \bar{g} . We also have

$$\begin{aligned}
(2.21b) \quad R_{0j} &= \frac{1}{f'(R_g)} \left(N_{g_{0j}} + \nabla_0 \nabla_j f'(R_g) \right) \\
&= \frac{N_{g_{0j}}}{f'(R_g)} + \frac{N \partial_j (N^{-1} (\partial_t - \beta^i \partial_i) f'(R_g))}{f'(R_g)} + N K_j^i \partial_i (\ln(f'(R_g))) \\
&= \frac{N_{g_{0j}}}{f'(R_g)} + \frac{N \partial_j (f''(R_g) \mathcal{L}_n R_g)}{f'(R_g)} + N K_j^i \partial_i (\ln(f'(R_g))),
\end{aligned}$$

and, finally,

$$\begin{aligned}
(2.21c) \quad G_{00} &= \frac{1}{f'(R_g)} \left(N_{g_{00}} + \frac{1}{2} f'(R_g) W_1(R_g) g_{00} - (g_{00} \square_g - \nabla_0 \nabla_0) f'(R_g) \right) \\
&= \frac{N_{g_{00}}}{f'(R_g)} + \frac{1}{2} g_{00} W_1(R_g) - \frac{g_{00}}{f'(R_g)} (\Delta_{\bar{g}} + \bar{g}^{ij} K_{ij} \mathcal{L}_n) f'(R_g).
\end{aligned}$$

Next, by combining (2.21a) with (2.12a), (2.21b) with (2.12b), and (2.21c) with (2.12c), the evolution equations and constraint equations for the system of modified gravity read

$$\begin{aligned}
(2.22) \quad \bar{\partial}_0 K_{ij} &= N \bar{R}_{ij} - N R_{ij} + N K_{ij} K_l^l - 2 N K_{il} K_j^l - \bar{\nabla}_i \partial_j N \\
&= N \bar{R}_{ij} + N K_{ij} K_l^l - 2 N K_{il} K_j^l - \bar{\nabla}_i \partial_j N \\
&\quad - \frac{N E_{ij}}{f'(R_g)} + \frac{N}{2} W_1(R_g) \bar{g}_{ij} - \frac{N (\bar{g}_{ij} \Delta_{\bar{g}} + 2 \bar{\nabla}_i \bar{\nabla}_j) f'(R_g)}{2 f'(R_g)} \\
&\quad + \frac{N \bar{g}_{ij} \nabla_0 \nabla_0 f'(R_g)}{2 N^2 f'(R_g)} - N (K_{ij} + \frac{1}{2} \bar{g}_{ij} K) \mathcal{L}_n \ln(f'(R_g)), \\
\bar{\partial}_0 \bar{g}_{ij} &= -2 N K_{ij},
\end{aligned}$$

and

$$(2.23) \quad \begin{aligned} \bar{R} - K_{ij}K^{ij} + (K_j^j)^2 &= \frac{2N_{g00}}{N^2 f'(R_g)} + \frac{2\Delta_{\bar{g}} f'(R_g)}{f'(R_g)} \\ &\quad + 2\bar{g}^{ij} K_{ij} \mathcal{L}_n(\ln f'(R_g)) - W_1(R_g), \end{aligned}$$

where

$$(2.24) \quad \partial_j K_i^i - \bar{\nabla}_i K_j^i = \frac{N_{g0j}}{N f'(R_g)} + \frac{\partial_j (f''(R_g) \mathcal{L}_n R_g)}{f'(R_g)} + K_j^i \partial_i (\ln(f'(R_g))).$$

It remains to consider the coupling with the matter field, described by the stress-energy tensor $T_{\alpha\beta}$. Recall that the equations read $N_{g\alpha\beta} = 8\pi T_{\alpha\beta}$, where different expressions for $T_{\alpha\beta}$ are provided for the Jordan coupling and for the Einstein coupling. We also define the mass density σ and the momentum vector J (measured by an observer moving orthogonally to the slices) by the relations

$$(2.25) \quad \sigma := N^{-2} T_{00}, \quad J_j := -N^{-1} T_{0j}.$$

We can thus conclude this section and introduce a definition suitable for modified gravity.

Proposition and Definition 2.2. — *The equations for modified gravity in the Cauchy adapted frame $\{e_0, e_1, e_2, e_3\}$ decompose as follows:*

1. **Evolution equations:**

$$(2.26) \quad \begin{aligned} \bar{\partial}_0 K_{ij} &= N \bar{R}_{ij} + N K_{ij} K_l^l - 2N K_{il} K_j^l - \bar{\nabla} \partial_j N \\ &\quad - \frac{8\pi N (T_{ij} - \frac{1}{2} g_{ij} \text{tr}(T))}{f'(R_g)} + \frac{N}{2} W_1(R_g) \bar{g}_{ij} \\ &\quad - \frac{N (\bar{g}_{ij} \Delta_{\bar{g}} + 2\bar{\nabla}_i \bar{\nabla}_j) f'(R_g)}{2f'(R_g)} + \frac{N \bar{g}_{ij} \nabla_0 \nabla_0 f'(R_g)}{2N^2 f'(R_g)} \\ &\quad - N \left(K_{ij} + \frac{1}{2} \bar{g}_{ij} K \right) \mathcal{L}_n \ln(f'(R_g)), \\ \bar{\partial}_0 \bar{g}_{ij} &= -2N K_{ij}. \end{aligned}$$

2. **Hamiltonian constraint:**

$$(2.27) \quad \begin{aligned} \bar{R} - K_{ij}K^{ij} + (K_j^j)^2 &= \frac{16\pi\sigma}{f'(R_g)} + \frac{2\Delta_{\bar{g}} f'(R_g)}{f'(R_g)} \\ &\quad + 2\bar{g}^{ij} K_{ij} \mathcal{L}_n(\ln f'(R_g)) - W_1(R_g), \end{aligned}$$

3. **Momentum constraint:**

$$(2.28) \quad \partial_j K_i^i - \bar{\nabla}_i K_j^i = -\frac{8\pi J_i}{f'(R_g)} + \frac{\partial_j (f''(R_g) \mathcal{L}_n R_g)}{f'(R_g)} + K_j^i \partial_i (\ln(f'(R_g))).$$

Observe that, in the classical gravity theory, the factor $f'(R_g)$ is constant and equal to unit, so that the terms containing $f'(R_g)$ in the right-hand sides of the constraint equations (2.27) and (2.28) vanish identically; consequently, we can recover here the standard equations (2.29) given below.

These new constraint equations are very involved compared with the classical ones: they contain fourth-order derivatives of the metric \bar{g} and, more precisely, second-order derivatives of the scalar curvature R_g . In particular, we can not recognize directly the elliptic nature of the classical constraint equations.

Remark 2.3. — Recall here the constraint equations for the classical theory of general relativity, when the Einstein equations $G_{\alpha\beta} = 8\pi T_{\alpha\beta}$ are imposed: the last two equations in (2.12) yield

$$(2.29) \quad \begin{aligned} \bar{R} + K_{ij} K^{ij} - (K_i^i)^2 &= 16\pi\sigma, \\ \bar{\nabla}^i K_{ij} - \bar{\nabla}_j K_i^l &= 8\pi J_j, \end{aligned} \quad (\text{when } f' \equiv 0).$$

2.3. The divergence identity

As in the classical gravity theory, we expect that the matter should be divergence-free $\nabla^\alpha T_{\alpha\beta} = 0$, which is now proven.

Lemma 2.4 (The divergence identity in modified gravity)

The contracted Bianchi identities

$$(2.30) \quad \nabla^\alpha R_{\alpha\beta} = \frac{1}{2} \nabla_\beta R$$

imply the divergence-free property for the modified gravity tensor

$$(2.31) \quad \nabla^\alpha N_{g_{\alpha\beta}} = 0.$$

Proof. — The following calculation holds in any local frame. We compute the three relevant terms:

$$\begin{aligned} \nabla^\alpha (\nabla_\alpha \nabla_\beta f'(R) - g_{\alpha\beta} \square_g f'(R)) &= (\nabla^\alpha \nabla_\alpha \nabla_\beta - \nabla_\beta \nabla^\lambda \nabla_\lambda) f'(R) \\ &= (\nabla_\alpha \nabla_\beta \nabla^\alpha - \nabla_\beta \nabla_\alpha \nabla^\alpha) (f'(R)) \\ &= [\nabla_\alpha, \nabla_\beta] (\nabla^\alpha (f'(R))) = R_{\alpha\beta} \nabla^\alpha (f'(R)), \end{aligned}$$

then

$$\begin{aligned} \nabla^\alpha (f'(R) R_{\alpha\beta}) &= R_{\alpha\beta} \nabla^\alpha (f'(R)) + f'(R) \nabla^\alpha R_{\alpha\beta} \\ &= R_{\alpha\beta} \nabla^\alpha (f'(R)) + \frac{1}{2} f'(R) \nabla_\beta R, \end{aligned}$$

and, finally,

$$\nabla^\alpha \left(\frac{1}{2} f(R) g_{\alpha\beta} \right) = \frac{1}{2} \nabla_\beta (f(R)) = \frac{1}{2} f'(R) \nabla_\beta R.$$

Combining these three identities together yields us the desired identity. \square

As a first application of Lemma 2.4, we now determine which coupling (formulated in the Jordan metric as far as this section is concerned) is mathematically sound. On one hand, consider first the Jordan coupling, corresponding to

$$(2.32) \quad T_{\alpha\beta} = \partial_\alpha\phi\partial_\beta\phi - \frac{1}{2}g_{\alpha\beta}g^{\alpha'\beta'}\partial_{\alpha'}\phi\partial_{\beta'}\phi.$$

By the field equation $N_{g_{\alpha\beta}} = 8\pi T_{\alpha\beta}$ and (2.30), we find $\nabla^\alpha T_{\alpha\beta} = 0$ and, after some calculations,

$$(2.33) \quad \partial_\beta\phi\Box_g\phi = 0.$$

Consequently, if the scalar field ϕ satisfies the wave equation

$$(2.34) \quad \Box_g\phi = 0,$$

then the tensor $T_{\alpha\beta}$ is divergence-free, as required. Hence, we need to solve a single scalar equation for the evolution of the matter.

On the other hand, if we assume the Einstein coupling

$$(2.35) \quad T_{\alpha\beta} = T_{\alpha\beta}^\dagger = f'(R_g)(\partial_\alpha\phi\partial_\beta\phi - \frac{1}{2}g_{\alpha\beta}g^{\alpha'\beta'}\partial_{\alpha'}\phi\partial_{\beta'}\phi),$$

then the field equation $N_{g_{\alpha\beta}} = 8\pi T_{\alpha\beta}^\dagger$ together with (2.30) lead us to $\nabla^\alpha T_{\alpha\beta}^\dagger = 0$, which now reads

$$f''(R_g)T_{\alpha\beta}\nabla^\alpha R_g + f'(R_g)\partial_\beta\phi\Box_g\phi = 0.$$

This (vectorial) equation can be written as

$$\partial_\beta\phi\Box_g\phi = \frac{f''(R_g)}{f'(R_g)}(\partial_\alpha\phi\partial_\beta\phi - \frac{1}{2}g_{\alpha\beta}g^{\alpha'\beta'}\partial_{\alpha'}\phi\partial_{\beta'}\phi)\nabla^\alpha R_g$$

or, equivalently,

$$(2.36) \quad \left(\frac{f''(R_g)}{f'(R_g)}\partial_\alpha\phi\nabla^\alpha R - \Box_g\phi\right)\nabla_\beta\phi = \frac{f''(R_g)}{2f'(R_g)}\left(g^{\alpha'\beta'}\partial_{\alpha'}\phi\partial_{\beta'}\phi\right)\nabla_\beta R.$$

Now, for general initial data, this is an *over-determined*⁽¹⁾ partial differential system (since the unknown of this vectorial system is a single scalar field): this strongly suggests that the Einstein coupling is not mathematically (nor physically) well-behaved. Consequently, from now on, we *focus our attention on the Jordan coupling*.

2.4. The initial value problem for modified gravity

Before we can formulate the Cauchy problem for the system (1.6), we need to specify the stress-energy tensor. In agreement with our discussion in the previous section, we assume

⁽¹⁾Unless we would impose the very unnatural restriction that $\nabla\phi$ and ∇R be co-linear.

a scalar field and the Jordan coupling (2.32), so that the matter fields read

$$(2.37) \quad \begin{aligned} \sigma &= N^{-2}T_{00} = |\mathcal{L}_n\phi|^2 + \frac{1}{2}|\nabla\phi|_g^2 = \frac{1}{2}(|\mathcal{L}_n\phi|^2 + \bar{g}^{ij}\partial_i\phi\partial_j\phi), \\ J_i &= -\mathcal{L}_n\phi\partial_i\phi. \end{aligned}$$

Definition 2.5. — *An initial data set for the modified gravity theory*

$$(\bar{M}, \bar{g}, K, R_0, R_1, \phi_0, \phi_1)$$

consists of the following data:

- a 3-dimensional manifold \bar{M} endowed with a Riemannian metric \bar{g} and a symmetric $(0, 2)$ -tensor field K ,
- two scalar fields denoted by R_0 and R_1 defined on \bar{M} and representing the (to-be-constructed) spacetime curvature and its time derivative,
- two scalar fields ϕ_0 and ϕ_1 defined on \bar{M} .

*These data are required to satisfy the **Hamiltonian constraint of modified gravity***

$$(2.38) \quad \begin{aligned} \bar{R} - K_{ij}K^{ij} + (K_j^j)^2 &= \frac{8(\phi_1^2 + \bar{g}^{ij}\partial_i\phi_0\partial_j\phi_0)}{f'(R_0)} + \frac{2\Delta_{\bar{g}}f'(R_0)}{f'(R_0)} \\ &\quad + 2\bar{g}^{ij}K_{ij}\frac{f''(R_0)R_1}{f'(R_0)} - W_1(R_0), \end{aligned}$$

*and the **momentum constraint of modified gravity***

$$(2.39) \quad \partial_j K_i^i - \bar{\nabla}_i K_j^i = \frac{8\pi\phi_1\partial_i\phi_0}{f'(R_0)} + \frac{\partial_j(f''(R_0)R_1)}{f'(R_0)} + K_j^i\partial_i(\ln(f'(R_0))).$$

Definition 2.6. — *Given an initial data set $(\bar{M}, \bar{g}, K, R_0, R_1, \phi_0, \phi_1)$ as in Definition 2.5, the **initial value problem for the modified gravity theory** consists of finding a Lorentzian manifold (M, g) and a matter field ϕ defined on M such that the following properties hold:*

1. *The field equations of modified gravity (1.6) are satisfied.*
2. *There exists an embedding $i : \bar{M} \rightarrow M$ with pull-back metric $\bar{g} = i^*g$ and second fundamental form K .*
3. *The field R_0 coincides with the restriction of the spacetime scalar curvature R on \bar{M} , while R_1 coincides with the Lie derivative $\mathcal{L}_n R$ restricted to \bar{M} , where n denotes the normal to \bar{M} .*
4. *The scalar fields ϕ_0, ϕ_1 coincides with the restriction of $\phi, \mathcal{L}_n\phi$ on \bar{M} , respectively.*

*Such a solution to (1.6) is referred to as a **modified gravity development of the initial data set** $(\bar{M}, \bar{g}, K, R_0, R_1, \phi_0, \phi_1)$.*

Similarly as in classical gravity, we can define [10] the notion of **maximal globally hyperbolic** development for the modified gravity theory. Observe that the initial value problem for modified gravity reduces to the classical formulation in the special case of vacuum and vanishing geometric data $\phi_0 = \phi_1 = R_0 = R_1 \equiv 0$. For the modified gravity

theory, we have just shown that, similarly as in classical gravity, these prescribed fields can not be fully arbitrary, and certain constraints (given above) must be assumed.

2.5. Preservation of the constraints

We need to address the following issue: if the evolution equations are satisfied by symmetric two-tensors (\bar{g}, K) , the stress-energy tensor T is divergence-free, and furthermore the constraint equations are satisfied on some initial slice M_0 , then are all of field equations satisfied? In other words, we want to establish the preservation of the constraint equations along a flow of solutions.

This is the first instance where we establish a “preservation property” and, later in this text, other similar situations will occur. The common character of these results is as follows. A differential system being given, the equations therein can be classified into two categories: one is easier to handle (the evolution equations in this example) while the other is more difficult (the constraint equations here). Our strategy is to replace the most difficult equations by some equations which can be deduced from the original system (in this example, the trace-free equation of T) but are also easier to handle. This leads us to a new system to be studied first, and an essential task is to check the equivalence between the original system and the new system.

Before we give a precise statement, we make the following observation. The evolution equations (2.26) are equivalent to (2.20a) and the constraint equations (2.27) and (2.28) are equivalent to (2.20b) and (2.20c). So we suppose that (2.20a) together with the divergence condition $\nabla^\alpha T_{\alpha\beta} = 0$ are satisfied in the spacetime $M = \bigcup_{t \in [0, t_{\max})} M_t$, and the constraint equations (2.20b) and (2.20c) are satisfied on the initial slice. Then we will prove that the equations (2.20b) and (2.20c) are satisfied in the whole spacetime. More precisely, we have the following result.

Proposition 2.7. — *With the notation above, suppose that the equations*

$$(2.40) \quad E_{g_{ij}} - 8\pi H_{ij} = 0 \quad \text{in the spacetime } \bigcup_{t \in [0, t_{\max})} M_t$$

$$(2.41) \quad \nabla^\alpha T_{\alpha\beta} = 0 \quad \text{in the spacetime } \bigcup_{t \in [0, t_{\max})} M_t$$

hold, together with

$$(2.42) \quad E_{g_{0j}} = 8\pi H_{0j}, \quad N_{g_{00}} = 8\pi T_{00} \quad \text{in the initial slice } M_0 = \{t = 0\}.$$

Then, it follows that

$$(2.43) \quad E_{g_{0j}} = 8\pi H_{0j}, \quad N_{g_{00}} = 8\pi T_{00} \quad \text{in the spacetime } \bigcup_{t \in [0, t_{\max})} M_t.$$

Proof. — The calculations are made in the Cauchy adapted frame and, for convenience, we introduce the notation

$$\Sigma_{\alpha\beta} := N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta}.$$

We will prove that $\Sigma_{\alpha\beta} = 0$ which is equivalent to the desired result. First of all, by the condition $\nabla_\alpha T_{\alpha\beta} = 0$ and the identity (2.30), we have

$$(2.44) \quad \nabla^\alpha \Sigma_{\alpha\beta} = 0.$$

By the definition of $E_{g_{\alpha\beta}}$ and $H_{\alpha\beta}$, the following identity holds:

$$(2.45) \quad \Sigma_{\alpha\beta} = E_{g_{\alpha\beta}} - H_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}\text{tr}(E_g - H)$$

and, in particular,

$$(2.46) \quad \Sigma_{00} = E_{g_{00}} - H_{00} - \frac{1}{2}g_{00}\text{tr}(E_g - H).$$

Now, due to the fact that $g^{0i} = g_{0i} = 0$ and (2.40), we have

$$(2.47) \quad \text{tr}(E_g - H) = g^{00}(E_{g_{00}} - H_{00}).$$

Combining (2.46) and (2.47) yields

$$(2.48) \quad \text{tr}(E_g - H) = 2g^{00}\Sigma_{00}$$

and, by substituting this into (2.45),

$$(2.49) \quad \Sigma_{\alpha\beta} = E_{g_{\alpha\beta}} - H_{\alpha\beta} - g_{\alpha\beta}g^{00}\Sigma_{00}.$$

Here, we can compute more precisely the spatial components in view of (2.40)

$$(2.50) \quad \Sigma_{ij} = -g_{ij}g^{00}\Sigma_{00}.$$

Given this material, we are now in a position to calculate $\Sigma_\beta^\alpha = g^{\alpha\alpha'}\Sigma_{\alpha'\beta}$. When $\alpha = \beta = 0$, we find

$$(2.51) \quad \Sigma_0^0 = \Sigma_{0\alpha}g^{\alpha 0} = g^{00}\Sigma_{00}.$$

For $\beta = 0$, $1 \leq \alpha \leq 3$, we set $a = \alpha$ and, by recalling that $g^{a0} = 0$, we obtain

$$\Sigma_0^a = g^{a\alpha'}\Sigma_{\alpha'0} = g^{a\alpha'}\Sigma_{\alpha'0}$$

For $1 \leq b \leq 3$, we have

$$\Sigma_b^0 = g_{b\beta}g^{0\alpha}\Sigma_\alpha^\beta = g^{00}g_{bc}\Sigma_0^c.$$

For $1 \leq b \leq 3$ and $1 \leq a \leq 3$, by applying (2.50) and (2.51), we obtain

$$\Sigma_b^a = g^{a\alpha}\Sigma_{b\alpha} = g^{a\alpha'}\Sigma_{ba'} = -g^{a\alpha'}g_{ba'}g^{00}\Sigma_{00} = -\delta_b^a\Sigma_0^0.$$

Hence, we conclude with

$$(2.52) \quad \Sigma_b^0 = g^{00}g_{bc}\Sigma_0^c, \quad \Sigma_b^a = -\delta_b^a\Sigma_0^0.$$

Now recall that the identity (2.44) can be written as

$$\nabla_\alpha \Sigma_\beta^\alpha = 0,$$

which leads to

$$\langle e_\alpha, \Sigma_\beta^\alpha \rangle - \omega_{\alpha\beta}^\delta \Sigma_\delta^\alpha + \omega_{\alpha\delta}^\alpha \Sigma_\beta^\delta = 0.$$

When $\beta = 0$, we have

$$(2.53) \quad \langle e_0, \Sigma_0^0 \rangle + \partial_a \Sigma_0^a - \omega_{\alpha\beta}^\delta \Sigma_\delta^\alpha + \omega_{\alpha\delta}^\alpha \Sigma_\beta^\delta = 0.$$

For $1 \leq b \leq 3$, we can take the equation (2.52) and write

$$\langle e_0, g^{00} g_{bc} \Sigma_0^c \rangle + \partial_a (-\delta_b^a \Sigma_0^0) - \omega_{\alpha b}^\gamma \Sigma_\gamma^\alpha + \omega_{\alpha\gamma}^\alpha \Sigma_b^\gamma = 0,$$

which leads us to

$$(2.54) \quad \langle e_0, \Sigma_0^c \rangle - g_{00} \bar{g}^{bc} \partial_b \Sigma_0^0 - g_{00} \bar{g}^{bc} (\omega_{\alpha b}^\gamma \Sigma_\gamma^\alpha - \omega_{\alpha\gamma}^\alpha \Sigma_b^\gamma) + g_{00} \bar{g}^{bc} \langle e_0, g^{00} g_{bc'} \rangle \Sigma_0^{c'} = 0.$$

We now consider the equations (2.53) and (2.54) together, and we observe that, in view of (2.52), the lower-order terms are *linear combinations* Σ_0^α with $0 \leq \alpha \leq 3$. Hence, these equations form a first-order differential system with linear source-terms. This system can also be written in a standard symmetric hyperbolic form. Namely, by recalling the notation $\bar{g} = \bar{g}_t$ for the induced Riemannian metric on the slices M_t , we introduce

$$V := (\Sigma_0^0, \Sigma_0^a)^T, \quad \rho^a := g_{00} (\bar{g}^{a1}, \bar{g}^{a2}, \bar{g}^{a3})^T$$

and

$$\sigma_1 = (1, 0, 0), \quad \sigma_2 = (0, 1, 0), \quad \sigma_3 = (0, 0, 1).$$

The principal part of the system defined by (2.53) and (2.54) can be put in the form

$$(2.55) \quad \langle e_0, V \rangle + \sum_a A^a \partial_a V = F,$$

where

$$A^a = \begin{pmatrix} 0 & \sigma_a \\ -g_{00} \rho^a & 0 \end{pmatrix}$$

and F is a linear form on V . By multiplying this equation by the matrix

$$A_0 := \begin{pmatrix} 1 & 0 \\ 0 & -g^{00} g_{ab} \end{pmatrix} = -g_{00} \begin{pmatrix} -g_{00} & 0 \\ 0 & g_{ab} \end{pmatrix},$$

we conclude that (2.55) becomes

$$(2.56) \quad A_0 \langle e_0, V \rangle + \sum_a A_0 A^a \partial_a V = A_0 F.$$

Note that $A_0 A^a$ are symmetric:

$$A_0 A^a = \begin{pmatrix} 0 & \sigma^a \\ (\sigma^a)^T & 0 \end{pmatrix}.$$

and the system (2.55) is thus symmetrizable. Clearly, (2.42) implies that $V = 0$ on the initial slice $\{t = 0\}$. Thanks to our global hyperbolicity assumption (2.1) and by a standard uniqueness argument, we therefore conclude that $V = 0$ in the whole spacetime. \square

3. Formulation of the Cauchy problem in the Einstein metric

3.1. Conformal transformation

In view of the derivation made in Section 2, it is clear that the evolution and the constraint equations of modified gravity are, both, very involved and do not have a standard (hyperbolic, elliptic) type within the general class of PDE's. The main difficulty comes from the fourth-order term

$$\nabla_\alpha \nabla_\beta f'(R_g).$$

As we will now show it, the conformal transformation

$$(3.1) \quad g^\dagger_{\alpha\beta} := e^{2\rho} g_{\alpha\beta}, \quad g^{\dagger\alpha\beta} = e^{-2\rho} g^{\alpha\beta}$$

(which depends upon second-order derivatives of the unknown metric) will overcome some of the difficulties: where the **conformal factor** is defined by

$$(3.2) \quad \rho := \frac{1}{2} \ln f'(R_g)$$

or, equivalently, $f'(R_g) = e^{2\rho}$. We now proceed by deriving several relevant expressions in the **conformal metric** g^\dagger in order to derive a tractable formulation of the field equations.

We begin by deriving an expression for the gravity tensor N_g in terms of the Einstein metric.

Lemma 3.1. — *With the notation above, the following identity holds*

$$(3.3) \quad e^{2\rho} R^\dagger_{\alpha\beta} - 6e^{2\rho} \partial_\alpha \rho \partial_\beta \rho + \frac{g^\dagger_{\alpha\beta}}{2} W_2(\rho) = N_{g_{\alpha\beta}} - \frac{1}{2} g_{\alpha\beta} \text{tr}(N_g),$$

where the **function** $W_2 = W_2(\rho)$ is defined implicitly by

$$(3.4) \quad W_2(s) := \frac{f(r) - f'(r)r}{f'(r)}, \quad e^{2s} = f'(r), \quad r \in \mathbb{R}.$$

We also recall that the function W_1 is defined by (2.17) and it will be also convenient (in the proof below) to introduce the **function** $W_3 = W_3(\rho)$ by

$$(3.5) \quad W_3(s) := f(r), \quad e^{2s} = f'(r), \quad r \in \mathbb{R}.$$

Proof. — We need to analyze the tensor

$$N_{g_{\alpha\beta}} = f'(R_g) R_{\alpha\beta} - \frac{1}{2} f(R_g) g_{\alpha\beta} + (g_{\alpha\beta} \square_g - \nabla_\alpha \nabla_\beta) f'(R_g)$$

and its trace $\text{tr}(N_g) = f'(R_g) R_g - 2f(R_g) + 3\square_g f'(R_g)$. Recall first the identities

$$\begin{aligned} \nabla_\alpha \nabla_\beta e^{2\rho} &= 2e^{2\rho} \nabla_\alpha \nabla_\beta \rho + 4e^{2\rho} \nabla_\alpha \rho \nabla_\beta \rho, \\ \square_g e^{2\rho} &= 2e^{2\rho} \square_g \rho + 4e^{2\rho} g(\nabla \rho, \nabla \rho), \end{aligned}$$

which imply

$$N_{g_{\alpha\beta}} = e^{2\rho} R_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta} W_3(\rho) + 2e^{2\rho} (g_{\alpha\beta} \square_g - \nabla_\alpha \nabla_\beta) \rho + 4e^{2\rho} (g_{\alpha\beta} g(\partial\rho, \partial\rho) - \partial_\alpha \rho \partial_\beta \rho)$$

and

$$\square_g \rho - \frac{W_3(\rho)}{6e^{2\rho}} + 2g(\partial\rho, \partial\rho) - \frac{1}{6} W_2(\rho) = \frac{\text{tr}(N_g)}{6e^{2\rho}}.$$

Moreover, we have the following relation between the Ricci curvature tensors of g and g^\dagger :

$$(3.6) \quad R^\dagger_{\alpha\beta} = R_{\alpha\beta} - 2(\nabla_\alpha \nabla_\beta \rho - \nabla_{\alpha\rho} \nabla_\beta \rho) - (\square_g \rho + 2g(\nabla\rho, \nabla\rho)) g_{\alpha\beta}$$

and, therefore, we see that N_g can be expressed as

$$N_{g_{\alpha\beta}} = e^{2\rho} R^\dagger_{\alpha\beta} - 6e^{2\rho} \nabla_{\alpha\rho} \nabla_\beta \rho + 6e^{2\rho} g_{\alpha\beta} g(\nabla\rho, \nabla\rho) + 3e^{2\rho} g_{\alpha\beta} \square_g \rho - \frac{1}{2} W_3(\rho) g_{\alpha\beta}.$$

It remains to combine this result with the trace equation above. \square

We are now in a position to state the field equations in the conformal metric. At this juncture, it is unclear how the scalar field ρ should be recovered in term of the Einstein metric, and this is an issue that we will address next.

Proposition and Definition 3.2. — *The field equations of modified gravity in the Einstein metric $g^\dagger_{\alpha\beta} = e^{2\rho} g_{\alpha\beta}$ with $\rho = \frac{1}{2} \ln(f'(R_g))$ read*

$$(3.7) \quad e^{2\rho} R^\dagger_{\alpha\beta} - 6e^{2\rho} \partial_\alpha \rho \partial_\beta \rho + \frac{g^\dagger_{\alpha\beta}}{2} W_2(\rho) = 8\pi (T_{\alpha\beta} - \frac{1}{2} g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} T_{\alpha'\beta'}).$$

Remark 3.3. — *For any sufficiently regular function w , one also has*

$$(3.8) \quad \square_{g^\dagger} w = e^{-2\rho} (\square_g w + 2g^{\alpha\beta} \partial_\alpha \rho \partial_\beta w) = e^{-2\rho} \square_g w + 2g^{\dagger\alpha\beta} \partial_\alpha \rho \partial_\beta w,$$

so that the trace equation transforms into

$$(3.9) \quad \square_{g^\dagger} \rho = \frac{W_2(\rho)}{6e^{2\rho}} + \frac{W_3(\rho)}{6e^{4\rho}} + \frac{1}{6e^{4\rho}} \text{tr}(N_g).$$

3.2. Evolution and constraint equations in the conformal metric

As in the previous section, we can formulate the evolution equations and constraint equations associated with the conformal field equation (3.3). To do so, as before, we introduce a foliation of the spacetime $M = [0, +\infty) \times M_t$ and a Cauchy adapted frame $\{e_0, e_1, e_2, e_3\}$ associated with the transformed metric g^\dagger . Then, by similar calculations as above, we find

$$(3.10a) \quad R^\dagger_{ij} = \bar{R}_{ij} - \frac{\bar{\partial}_0 K^\dagger_{ij}}{N} + K^\dagger_{ij} K^{\dagger l}_l - 2K^\dagger_{il} K^{\dagger l}_j - \frac{\bar{\nabla}^\dagger_i \partial_j N^\dagger}{N^\dagger},$$

$$(3.10b) \quad R^\dagger_{0j} = N^\dagger (\partial_j K^{\dagger l}_l - \bar{\nabla}^\dagger_l K^{\dagger l}_j),$$

$$(3.10c) \quad G^\dagger_{00} = \frac{N^{\dagger 2}}{2} (\bar{R}^\dagger - K^\dagger_{ij} K^{\dagger ij} + (K^{\dagger l}_l)^2).$$

Here, $\bar{\nabla}^\dagger$ refers to the covariant derivative on the slice M_t with respect to g^\dagger , and we observe that (3.10a) yields the evolution equations

$$(3.11) \quad \begin{aligned} \bar{\partial}_0 K^\dagger_{ij} &= N^\dagger (\bar{R}^\dagger_{ij} - R^\dagger_{ij}) + N^\dagger K^\dagger_{ij} K^{\dagger l}_l - 2N^\dagger K^\dagger_{il} K^{\dagger l}_j - \bar{\nabla}^\dagger_i \partial_j N^\dagger, \\ \bar{\partial}_0 \bar{g}^\dagger_{ij} &= -2N^\dagger K^\dagger_{ij}. \end{aligned}$$

Moreover, the transformed field equations (3.3) read

$$R^\dagger_{\alpha\beta} = e^{-2\rho} (N_{g_{\alpha\beta}} - \frac{1}{2} g_{\alpha\beta} \text{tr}(N_g)) + 6\partial_\alpha \rho \partial_\beta \rho - \frac{1}{2e^{2\rho}} g^\dagger_{\alpha\beta} W_2(\rho)$$

and, by taking the trace of this equation **with respect to the metric** g^\dagger , we have

$$R^\dagger = -e^{-4\rho} \text{tr}(N_g) + 6g^\dagger(\partial\rho, \partial\rho) g_{\alpha\beta} + e^{-2\rho} g^\dagger_{\alpha\beta} W_2(\rho).$$

This leads us to

$$G^\dagger_{\alpha\beta} = N_{g_{\alpha\beta}} + 6\nabla_\alpha \varrho \nabla_\beta \varrho - 3g^\dagger(\partial\rho, \partial\rho) g^\dagger_{\alpha\beta} + e^{-2\rho} g^\dagger_{\alpha\beta} W_2(\rho),$$

We have thus derived the evolution equations and constraint equations. The evolution equations read

$$\begin{aligned} \bar{\partial}_0 K^\dagger_{ij} &= N^\dagger \bar{R}^\dagger_{ij} + N^\dagger K^\dagger_{ij} K^{\dagger l}_l - 2N^\dagger K^\dagger_{il} K^{\dagger l}_j - \bar{\nabla}^\dagger_i \partial_j N^\dagger \\ &\quad - N^\dagger e^{-2\rho} \left(N_{g_{ij}} - \frac{1}{2e^{2\rho}} g^\dagger_{ij} \text{tr}(N_g) + 6e^{2\rho} \partial_i \rho \partial_j \rho + \frac{1}{2} g^\dagger_{ij} W_2(R_g) \right), \\ \bar{\partial}_0 \bar{g}^\dagger_{ij} &= -2N^\dagger K^\dagger_{ij}, \end{aligned}$$

while the Hamilton constraint equation reads

$$\bar{R}^\dagger - K^\dagger_{ij} K^{\dagger ij} + (K^{\dagger l}_l)^2 = \frac{2N_{g_{00}}}{e^{2\rho} N^{\dagger 2}} + 12|\mathcal{L}_{n^\dagger} \rho|^2 + 6g^\dagger(\nabla^\dagger \rho, \nabla^\dagger \rho) - e^{-2\rho} W_2(\rho)$$

and the momentum constraint equations read

$$\partial_j K^{\dagger l}_l - \bar{\nabla}^\dagger K^{\dagger l}_j = \frac{N_{g_{0j}}}{e^{2\rho} N^{\dagger}} + 6\mathcal{L}_{n^\dagger} \rho \partial_j \rho.$$

Here, n^\dagger denotes the normal unit vector of the slice M_t .

Finally, we consider the Jordan coupling with matter field (this choice of coupling being revisited in the next subsection):

$$(3.12) \quad N_{g_{\alpha\beta}} = 8\pi T_{\alpha\beta}$$

and, furthermore, we define the matter fields

$$(3.13) \quad J^\dagger_j := -\frac{T_{0j}}{N^{\dagger}}, \quad \sigma^\dagger := \frac{T_{00}}{N^{\dagger 2}}.$$

Definition 3.4. — *In the Einstein metric, the equations of modified gravity in a Cauchy adapted frame $\{e_0, e_1, e_2, e_3\}$ can be decomposed as follows:*

1. **Evolution equations:**

$$\begin{aligned}\bar{\partial}_0 K^\dagger_{ij} &= N^\dagger \bar{R}^\dagger_{ij} + N^\dagger K^\dagger_{ij} K^{\dagger l}_l - 2N^\dagger K^\dagger_{il} K^{\dagger l}_j - \bar{\nabla}^\dagger_i \partial_j N^\dagger \\ &\quad - N^\dagger e^{-2\rho} \left(8\pi T_{ij} - 4\pi e^{-2\rho} g^\dagger_{ij} \text{tr} (T) + 6e^{2\rho} \partial_i \rho \partial_j \rho + \frac{1}{2} g^\dagger_{ij} W_1(R_g) \right), \\ \bar{\partial}_0 \bar{g}^\dagger_{ij} &= -2N^\dagger K^\dagger_{ij}.\end{aligned}$$

2. **Hamiltonian constraint:**

$$(3.14) \quad \bar{R}^\dagger - K^\dagger_{ij} K^{\dagger ij} + (K^{\dagger l}_l)^2 = \frac{16\sigma^\dagger}{e^{2\rho}} + 12|\mathcal{L}_{n^\dagger} \rho|^2 + 6g^\dagger(\nabla^\dagger \rho, \nabla^\dagger \rho) - e^{-2\rho} W_2(\rho)$$

3. **Momentum constraints:**

$$(3.15) \quad \partial_j K^{\dagger l}_l - \bar{\nabla}^\dagger K^{\dagger l}_j = -\frac{J^\dagger_j}{e^{2\rho}} + 6\mathcal{L}_{n^\dagger} \rho \partial_j \rho.$$

Let us again emphasize again that the constraint equations are equivalent to $N_{g_{00}} = 8\pi T_{00}$, $N_{g_{0a}} = 8\pi T_{0a}$, while the evolution equations are equivalent to

$$N_{g_{ab}} - \frac{1}{2} g_{ab} \text{tr} (N_g) = 8\pi (T_{ab} - \frac{1}{2} g_{ab} \text{tr} (T)).$$

3.3. The divergence identity

In order to derive an evolution equation for the matter field, we need the divergence of the tensor N_g with respect to the conformal metric g^\dagger .

Lemma 3.5. — *The modified gravity tensor in terms of the conformal metric satisfies the identity*

$$(3.16) \quad \nabla^{\dagger\alpha} N_{g_{\alpha\beta}} = e^{-2\rho} (2g^{\gamma\delta} \partial_\gamma \rho N_{g_{\delta\beta}} - \text{tr} (N_g) \partial_\beta \rho).$$

Proof. — We work in an arbitrary (possibly only locally defined) natural frame. The desired identity follows from

$$\Gamma^{\dagger\gamma}_{\alpha\beta} = \Gamma^\gamma_{\alpha\beta} + g^\gamma_\alpha \partial_\beta \rho + g^\gamma_\beta \partial_\alpha \rho - g_{\alpha\beta} \nabla^\gamma \rho.$$

We have

$$\begin{aligned}\nabla^{\dagger\alpha} N_{g_{\alpha\beta}} &= e^{-2\rho} g^{\alpha\gamma} \nabla^\dagger_\gamma N_{g_{\alpha\beta}} \\ &= e^{-2\rho} g^{\alpha\gamma} (\partial_\gamma N_{g_{\alpha\beta}} - \Gamma^\delta_{\gamma\alpha} N_{g_{\beta\delta}} - \Gamma^\delta_{\gamma\beta} N_{g_{\alpha\delta}}) \\ &\quad - e^{-2\rho} g^{\alpha\gamma} (g^\delta_\gamma \partial_\alpha \rho + g^\delta_\alpha \partial_\gamma \rho - g_{\gamma\alpha} \nabla^\delta \rho) N_{g_{\beta\delta}} \\ &\quad - e^{-2\rho} g^{\alpha\gamma} (g^\delta_\gamma \partial_\beta \rho + g^\delta_\beta \partial_\gamma \rho - g_{\gamma\beta} \nabla^\delta \rho) N_{g_{\alpha\delta}},\end{aligned}$$

thus

$$\begin{aligned}\nabla^{\dagger\alpha} N_{g_{\alpha\beta}} &= e^{-2\rho} \nabla^\alpha N_{g_{\alpha\beta}} \\ &\quad - e^{-2\rho} (\nabla^\delta \rho + \nabla^\delta \rho - 4\nabla^\delta \rho) N_{g_{\beta\delta}} - e^{-2\rho} (\partial_\beta \rho \text{tr} (N_g) + \nabla^\alpha \rho N_{g_{\alpha\beta}} - \nabla^\alpha \rho N_{g_{\alpha\beta}}).\end{aligned}$$

Recalling that $\nabla^\alpha N_{g_{\alpha\beta}} = 0$ by Lemma 2.4, we conclude that

$$\nabla^{\dagger\alpha} N_{g_{\alpha\beta}} = e^{-2\rho} (2g^{\gamma\delta} \partial_{\gamma\rho} N_{g_{\delta\beta}} - \text{tr} (N_g) \partial_{\beta\rho}).$$

□

With the Jordan coupling, the divergence of the stress-energy tensor is thus expressed as

$$(3.17) \quad \nabla^{\dagger\alpha} T_{\alpha\beta} = (2g^{\delta\gamma} \partial_{\delta\rho} T_{\gamma\beta} - \text{tr} (T) \partial_{\beta\rho}) e^{-2\rho},$$

which (together with an equation of state for the matter field) determines the evolution of the matter.

Remark 3.6. — *We conclude this section with a discussion of the Einstein coupling. We rely on (3.16) and now show that the only meaningful choice of coupling (now viewed in the Einstein metric) is the Jordan coupling. Observe first that the Jordan coupling*

$$T_{\alpha\beta} = \partial_\alpha \phi \partial_\beta \phi - \frac{1}{2} g_{\alpha\beta} |\nabla \phi|_g^2$$

implies

$$\begin{aligned} \partial_{\beta\rho} \text{tr} (T) - 2g^{\delta\gamma} \partial_{\delta\rho} T_{\gamma\beta} &= -|\nabla \phi|_g^2 \partial_{\beta\rho} - 2g(\partial\rho, \partial\phi) \partial_\beta \phi + |\nabla \phi|_g^2 \partial_{\beta\rho} \\ &= -2g(\partial\rho, \partial\phi) \partial_\beta \phi, \end{aligned}$$

which leads us to

$$(3.18) \quad (2g^{\delta\gamma} \partial_{\delta\rho} T_{\gamma\beta} - \text{tr} (T) \partial_{\beta\rho}) e^{-2\rho} = 2g^\dagger(\partial\rho, \partial\rho) \partial_\beta \phi.$$

From the identity

$$\nabla^{\dagger\alpha} T_{\alpha\beta} = \partial_\beta \phi \square_{g^\dagger} \phi$$

combined with (3.16), we have

$$\partial_\beta \phi \square_{g^\dagger} \phi = 2e^{-2\rho} g^{\alpha'\beta'} \partial_{\alpha'} \phi \partial_{\beta'} \rho \partial_\beta \phi$$

and this leads us to the **wave equation for the matter field**

$$(3.19) \quad \square_{g^\dagger} \phi = 2g^{\dagger\alpha'\beta'} \partial_{\alpha'} \phi \partial_{\beta'} \rho.$$

On the other hand, let us consider the Einstein coupling:

$$T_{\alpha\beta} = T_{\alpha\beta}^\dagger = e^{2\rho} (\partial_\alpha \phi \partial_\beta \phi - \frac{1}{2} g_{\alpha\beta} |\nabla \phi|_g^2),$$

which gives

$$\begin{aligned} \nabla^{\dagger\alpha} T_{\alpha\beta} &= e^{2\rho} \partial_\beta \phi \square_{g^\dagger} \phi + 2e^{2\rho} g^{\dagger\alpha\alpha'} \partial_{\alpha'} \rho (\partial_\alpha \phi \partial_\beta \phi - \frac{1}{2} g_{\alpha\beta} |\nabla \phi|_g^2) \\ &= e^{2\rho} \partial_\beta \phi (\square_{g^\dagger} \phi + 2g^\dagger(\partial\rho, \partial\phi)) - \partial_{\beta\rho} |\nabla \phi|_g^2. \end{aligned}$$

In combination with (3.23), we find

$$e^{2\rho} \partial_\beta \phi (\square_{g^\dagger} \phi + 2g^\dagger(\partial\rho, \partial\phi)) - \partial_{\beta\rho} |\nabla \phi|_g^2 = 2g^\dagger(\partial\rho, \partial\phi) \partial_\beta \phi,$$

and therefore

$$e^{2\rho}\partial_\beta\phi\left(\square_{g^\dagger}\phi + 2g^\dagger(\partial\rho, \partial\phi) - 2e^{-2\rho}g^\dagger(\nabla^\dagger\rho, \nabla^\dagger\phi)\right) = \partial_\beta\rho|\nabla\phi|_g^2.$$

In agreement with what we noticed with the Jordan metric, the Einstein coupling leads to an over-determined partial differential system. This suggests again that the Einstein coupling cannot lead to a well-posed initial value problem.

3.4. The conformal version of the initial value problem

We are now ready to formulate the notion of initial data set and the notion of Cauchy development in terms of the conformal metric. In agreement to our discussion in the previous section, we work with the Jordan coupling and a massless scalar field:

$$(3.20) \quad T_{\alpha\beta} = \partial_\alpha\phi\partial_\beta\phi - \frac{1}{2}g_{\alpha\beta}|\nabla\phi|_g^2,$$

and we set

$$\begin{aligned} \sigma^\dagger &:= \frac{1}{2}\left(|\mathcal{L}_{n^\dagger}\phi|^2 + \bar{g}^{ij}\partial_i\phi\partial_j\phi\right), \\ J^\dagger_j &:= -\mathcal{L}_{n^\dagger}\phi\partial_j\phi. \end{aligned}$$

Definition 3.7. — *An initial data set for the modified gravity theory in the Einstein metric $(\bar{M}, \bar{g}^\dagger, K^\dagger, \rho_0^\dagger, \rho_1^\dagger, \phi_0^\dagger, \phi_1^\dagger)$ consists of the following data:*

- a 3-dimensional manifold \bar{M} endowed with a Riemannian metric \bar{g}^\dagger and a symmetric $(0, 2)$ -tensor field K^\dagger ,
- two scalar fields denoted by ρ_0^\dagger and ρ_1^\dagger on \bar{M} and representing the (to-be-constructed) conformal factor and its time derivative,
- two scalar field ϕ_0^\dagger and ϕ_1^\dagger defined on \bar{M} .

Furthermore, these data are required to satisfy the **Hamiltonian constraint of modified gravity in the Einstein metric**

$$(3.21) \quad \begin{aligned} \bar{R}^\dagger - K^\dagger_{ij}K^{\dagger ij} + (K^{\dagger j}_j)^2 &= 8e^{-2\rho}\left((\phi_1^\dagger)^2 + \bar{g}^{\dagger ij}\partial_i\phi_0^\dagger\partial_j\phi_0^\dagger\right) + 6(\rho_1^\dagger)^2 \\ &\quad + 6\bar{g}^{\dagger ij}\partial_i\rho_0^\dagger\partial_j\rho_0^\dagger - e^{-2\rho_0^\dagger}W_2(\rho_0^\dagger), \end{aligned}$$

and the **momentum constraint of modified gravity in the Einstein metric**

$$(3.22) \quad \partial_j K^{\dagger i}_i - \bar{\nabla}^\dagger_i K^{\dagger i}_j = \frac{\phi_1^\dagger\partial_j\phi_0^\dagger}{e^{2\rho}} + 6\rho_1^\dagger\partial_j\rho_0^\dagger.$$

Definition 3.8. — *Given an initial data set $(\bar{M}, \bar{g}^\dagger, K^\dagger, \rho_0^\dagger, \rho_1^\dagger, \phi_0^\dagger, \phi_1^\dagger)$ as in Definition 3.7, the initial value problem for the modified gravity theory in the Einstein metric consists of finding a Lorentzian manifold (M, g) and a two-tensor field $T_{\alpha\beta}$ on M*

1. The conformal metric g^\dagger is defined with the relation $g^\dagger_{\alpha\beta} = e^{2\rho}g_{\alpha\beta}$ with the conformal factor $\rho = \frac{1}{2}\ln(f'(R_g))$ where R_g is the scalar curvature of g .
2. The field equations of modified gravity (3.7) are satisfied with $\rho = \frac{1}{2}\ln f'(R_g)$.

3. There exists an embedding $i : \overline{M} \rightarrow M$ with pull-back metric $\overline{g}^\dagger = i^*g^\dagger$ and second fundamental form K^\dagger .
4. The field ρ_0^\dagger coincides with the restriction of the conformal factor ρ on \overline{M} , while ρ_1^\dagger coincides with the Lie derivative $\mathcal{L}_{n^\dagger}\rho$ restricted to \overline{M} , where n^\dagger denotes the normal unit vector of \overline{M} .
5. The scalar fields $\phi_0^\dagger, \phi_1^\dagger$ coincides with the restriction of $\phi, \mathcal{L}_{n^\dagger}\phi$ on \overline{M} .

Such a solution to (3.7) is referred to as a **modified gravity development of the initial data set** $(\overline{M}, \overline{g}^\dagger, K^\dagger, \rho_0^\dagger, \rho_1^\dagger, \phi_0^\dagger, \phi_1^\dagger)$.

The notion of **maximal globally hyperbolic development** is then defined along the same lines as in [10] for the classical gravity. We observe that our formulation of the initial value problem for modified gravity reduces to the classical formulation in the special case of vanishing geometric data $\phi_0^\dagger = \phi_1^\dagger = R_0 = R_1 \equiv 0$. On the other hand, without matter fields and for non-vanishing geometric data R_0 and R_1 , the spacetimes under consideration do not satisfy Einstein vacuum equations. Similarly as in classical gravity, these fields can not be fully arbitrary prescribed but certain constraints (given above) must be assumed.

3.5. Preservation of the constraints

Next, we establish the preservation of the constraints, as follows.

Proposition 3.9. — *Let $(\overline{g}^\dagger, K^\dagger)$ be symmetric two-tensors defined in $M = \cup_{t \in [0, t_{\max})} M_t$. If the following equations hold in M*

$$(3.23) \quad N_{g_{ij}} - \frac{1}{2} \text{tr} (N_g) g_{ij} = 8\pi (T_{ij} - \frac{1}{2} \text{tr} (N_g) g_{ij}),$$

$$(3.24) \quad \nabla^{\dagger\alpha} T_{\alpha\beta} = e^{-2\rho} (g^{\gamma\delta} \partial_\gamma \rho T_{\gamma\beta} - \partial_\beta \rho \text{tr} (T)),$$

and

$$(3.25) \quad N_{g_{0\beta}} = 8\pi T_{0\beta}.$$

holds on the initial slice M_0 , then (3.25) holds throughout the spacetime M .

Proof. — Recalling the notation $\Sigma_{\alpha\beta} = N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta}$, we are going to prove that $\sigma_{00} = 0$. We note that (3.23) can be written as

$$(3.26) \quad \Sigma_{ij} - \frac{1}{2} \text{tr}^\dagger(\Sigma) g^\dagger_{ij} = 0.$$

By taking the trace of the tensor $\Sigma_{\alpha\beta} - \frac{1}{2} \text{tr} \Sigma g_{\alpha\beta}$ **with respect to g^\dagger** , we find

$$\text{tr}^\dagger \Sigma - 2 \text{tr}^\dagger \Sigma = -N^{\dagger-2} \Sigma_{00} + g^{\dagger ij} (\Sigma_{ij} - \frac{1}{2} \text{tr} (\Sigma) g_{ij}).$$

Combining with (3.26), we thus have

$$(3.27) \quad \text{tr}^\dagger \Sigma = -N^{\dagger-2} \Sigma_{00}.$$

Combining (3.27) together with (3.26), we then obtain

$$(3.28) \quad \Sigma_{ij} = -\frac{\Sigma_{00}}{2N^{\dagger 2}} g^{\dagger}_{ij}.$$

Along the same lines as in the proof of Proposition 2.7, we have

$$(3.29) \quad \Sigma_b^0 = g^{\dagger 00} g^{\dagger}_{bc} \Sigma_0^c, \quad \Sigma_b^a = -\delta_b^a \Sigma_0^0.$$

Let us consider the identity (3.16) combined with (3.24), and note the identity

$$(3.30) \quad \nabla^{\dagger \alpha} \Sigma_{\alpha\beta} = e^{-2\rho} (2\partial_{\gamma}\rho \Sigma_{\beta}^{\gamma} - \partial_{\beta}\rho \text{tr}(\Sigma)).$$

We observe that by (3.29), the right-hand-side is a linear form of the function Σ_0^{β} and, by definition,

$$(3.31) \quad \nabla^{\dagger}_{\alpha} \Sigma_{\beta}^{\alpha} = \langle e_{\alpha}, \Sigma_{\beta}^{\alpha} \rangle - \omega^{\dagger \delta}_{\alpha\beta} \Sigma_{\delta}^{\alpha} + \omega^{\dagger \alpha}_{\alpha\delta} \Sigma_{\beta}^{\delta}.$$

By combining (3.30) and (3.31), we arrive at the first-order linear differential system

$$\langle e_{\alpha}, \Sigma_{\beta}^{\alpha} \rangle - \omega^{\dagger \delta}_{\alpha\beta} \Sigma_{\delta}^{\alpha} + \omega^{\dagger \alpha}_{\alpha\delta} \Sigma_{\beta}^{\delta} = e^{-2\rho} (2\partial_{\gamma}\rho \Sigma_{\beta}^{\gamma} - \partial_{\beta}\rho \text{tr}(\Sigma)),$$

whose principal part is

$$(3.32) \quad \begin{aligned} & \langle e_0, \Sigma_0^0 \rangle + \partial_a \Sigma_0^a = \text{lower order terms}, \\ & \langle e_0, g^{\dagger 00} g^{\dagger}_{bc} \Sigma_0^c \rangle - \partial_b \Sigma_0^0 = \text{lower order terms}. \end{aligned}$$

can be symmetrized by the same procedure as we did for the system (2.53) and (2.54). Recall also that by (3.25), this system has vanishing initial data and, therefore, in view of our global hyperbolicity assumption (2.1), the desired result is proven. \square

4. The conformal formulation in wave coordinates

4.1. The wave gauge

We now turn our attention to solving the system (3.7) and our first task is thus to express it in well-chosen coordinates. In view of the expression of the left-hand-side of (3.7), we observe that if we remove the terms in ρ , the principal part (that is, the second-order terms in g^{\dagger}) is determined by $R^{\dagger}_{\alpha\beta}$. In order to investigate its structure, we perform first some basic calculations, which are valid for general Lorentzian manifolds in arbitrary local coordinates. Let (M, g) be a Lorentzian manifold with metric g of signature $(-, +, +, +)$ and consider any local coordinate system $\{x^0, x^1, x^2, x^3\}$. Let $\Gamma^{\gamma}_{\alpha\beta}$ be the associated Christoffel symbols, and consider the wave operator $\square_g = \nabla^{\alpha} \nabla_{\alpha}$ associated with g . The following lemma follows from a straightforward but tedious calculation.

Lemma 4.1 (Ricci curvature in general coordinates). — *With the notation*

$$(4.1) \quad \Gamma^{\lambda} := g^{\alpha\beta} \Gamma^{\lambda}_{\alpha\beta}, \quad \Gamma_{\lambda} := g_{\lambda\beta} \Gamma^{\beta},$$

one has

$$(4.2) \quad R_{\alpha\beta} = -\frac{1}{2}g^{\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}g_{\alpha\beta} + \frac{1}{2}(\partial_{\alpha}\Gamma_{\beta} + \partial_{\beta}\Gamma_{\alpha}) + \frac{1}{2}F_{\alpha\beta}(g; \partial g, \partial g),$$

where $F_{\alpha\beta}(g; \partial g, \partial g)$ are nonlinear functions in the metric coefficients and are quadratic in their first-order derivatives. The wave operator and the reduced wave operator $\tilde{\square}_g u := g^{\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}$ satisfy the relation

$$(4.3) \quad \square_g u = g^{\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}u + \Gamma^{\delta}\partial_{\delta}u = \tilde{\square}_g u + \Gamma^{\delta}\partial_{\delta}u,$$

and, clearly, these two operators coincide if and only if the coefficients Γ^{λ} vanish identically.

Proof. — Recall the definitions

$$\begin{aligned} R_{\alpha\beta} &= \partial_{\lambda}\Gamma_{\alpha\beta}^{\lambda} - \partial_{\alpha}\Gamma_{\beta\lambda}^{\lambda} + \Gamma_{\alpha\beta}^{\lambda}\Gamma_{\lambda\delta}^{\delta} - \Gamma_{\alpha\delta}^{\lambda}\Gamma_{\beta\lambda}^{\delta}, \\ \Gamma_{\alpha\beta}^{\lambda} &= \frac{1}{2}g^{\lambda\lambda'}(\partial_{\alpha}g_{\beta\lambda'} + \partial_{\beta}g_{\alpha\lambda'} - \partial_{\lambda'}g_{\alpha\beta}). \end{aligned}$$

In the expression of the Ricci tensor, we consider the first two terms:

$$\begin{aligned} &\partial_{\lambda}\Gamma_{\alpha\beta}^{\lambda} - \partial_{\alpha}\Gamma_{\beta\lambda}^{\lambda} \\ &= \frac{1}{2}\partial_{\lambda}\left(g^{\lambda\delta}(\partial_{\alpha}g_{\beta\delta} + \partial_{\beta}g_{\alpha\delta} - \partial_{\delta}g_{\alpha\beta})\right) - \frac{1}{2}\partial_{\alpha}\left(g^{\lambda\delta}(\partial_{\beta}g_{\lambda\delta} + \partial_{\lambda}g_{\beta\delta} - \partial_{\delta}g_{\beta\lambda})\right) \\ &= -\frac{1}{2}\partial_{\lambda}\left(g^{\lambda\delta}\partial_{\delta}g_{\alpha\beta}\right) + \frac{1}{2}\partial_{\lambda}\left(g^{\lambda\delta}(\partial_{\alpha}g_{\beta\delta} + \partial_{\beta}g_{\alpha\delta})\right) - \frac{1}{2}\partial_{\alpha}\left(g^{\lambda\delta}\partial_{\beta}g_{\lambda\delta}\right), \end{aligned}$$

so that

$$(4.4) \quad \begin{aligned} &\partial_{\lambda}\Gamma_{\alpha\beta}^{\lambda} - \partial_{\alpha}\Gamma_{\beta\lambda}^{\lambda} \\ &= -\frac{1}{2}g^{\lambda\delta}\partial_{\lambda}\partial_{\delta}g_{\alpha\beta} + \frac{1}{2}g^{\lambda\delta}\partial_{\alpha}\partial_{\lambda}g_{\delta\beta} + \frac{1}{2}g^{\lambda\delta}\partial_{\beta}\partial_{\lambda}g_{\delta\alpha} - \frac{1}{2}g^{\lambda\delta}\partial_{\alpha}\partial_{\beta}g_{\lambda\delta} + \text{l.o.t.}, \end{aligned}$$

where l.o.t. are quadratic terms in the first-order derivatives of the metric.

On the other hand, we can compute the term $\partial_{\alpha}\Gamma_{\beta} + \partial_{\beta}\Gamma_{\alpha}$ and obtain

$$\begin{aligned} \Gamma^{\gamma} &= \Gamma_{\alpha\beta}^{\gamma}g^{\alpha\beta} = \frac{1}{2}g^{\alpha\beta}g^{\gamma\delta}(\partial_{\alpha}g_{\beta\delta} + \partial_{\beta}g_{\alpha\delta} - \partial_{\delta}g_{\alpha\beta}) \\ &= g^{\gamma\delta}g^{\alpha\beta}\partial_{\alpha}g_{\beta\delta} - \frac{1}{2}g^{\alpha\beta}g^{\gamma\delta}\partial_{\delta}g_{\alpha\beta} \end{aligned}$$

and

$$\Gamma_{\lambda} = g_{\lambda\gamma}\Gamma^{\gamma} = g^{\alpha\beta}\partial_{\alpha}g_{\beta\lambda} - \frac{1}{2}g^{\alpha\beta}\partial_{\lambda}g_{\alpha\beta}.$$

So, we have

$$\partial_{\alpha}\Gamma_{\beta} = \partial_{\alpha}(g^{\delta\lambda}\partial_{\delta}g_{\lambda\beta}) - \frac{1}{2}\partial_{\alpha}(g^{\lambda\delta}\partial_{\beta}g_{\lambda\delta}),$$

and, therefore,

$$(4.5) \quad \partial_{\alpha}\Gamma_{\beta} + \partial_{\beta}\Gamma_{\alpha} = g^{\gamma\delta}\partial_{\alpha}\partial_{\lambda}g_{\delta\beta} + g^{\lambda\delta}\partial_{\beta}\partial_{\lambda}g_{\delta\alpha} - g^{\lambda\delta}\partial_{\alpha\beta}g_{\lambda\delta} + \text{l.o.t.}$$

It remains to compare (4.4) and (4.5). □

Observe that the field equation (1.6) or the conformally transformed field equations (3.3), both, contain linear terms in the Ricci curvature. In order to exhibit the hyperbolicity property for the linear part of these systems (at least for the second-order terms in (1.6)), we now introduce local coordinates. Recall that a **wave coordinate system**, by definition, has Christoffel symbols satisfying

$$(4.6) \quad \Gamma^\lambda = g^{\alpha\beta}\Gamma_{\alpha\beta}^\gamma = 0.$$

In view of (3.3), the principal part (after removing the terms in ρ) of

$$N_{g_{\alpha\beta}} - \frac{1}{2}g_{\alpha\beta}\text{tr}(N_g) - \frac{1}{2}e^{2\rho}(\partial_\alpha\Gamma^\dagger + \partial_\beta\Gamma^\dagger)$$

is $\frac{1}{2}e^{2\rho}g^{\dagger\alpha\beta}\partial_{\alpha'}\partial_{\beta'}g^\dagger_{\alpha\beta}$, which is a *quasi-linear wave operator*. From this observation, the equations (3.7) in wave coordinates, with

$$(4.7) \quad \Gamma^\dagger_\lambda = 0,$$

can be reformulated as:

$$(4.8) \quad N_{g_{\alpha\beta}} - \frac{1}{2}g_{\alpha\beta}\text{tr}(N_g) - \frac{1}{2}e^{2\rho}(\partial_\alpha\Gamma^\dagger_\beta + \partial_\beta\Gamma^\dagger_\alpha) = 8\pi(T_{\alpha\beta} - \frac{1}{2}\text{tr}(T)g_{\alpha\beta}),$$

while the trace equation (3.9) becomes

$$(4.9) \quad \square_{g^\dagger}\rho + \Gamma^{\dagger\delta}\partial_\delta\rho = \frac{W_2(\rho)}{6e^{2\rho}} + \frac{W_3(\rho)}{6e^{4\rho}} + \frac{4\pi\text{tr}(T)}{3e^{4\rho}}.$$

Hence, in view of Lemma 4.1, the above system can be written in terms of the metric g^\dagger and its derivatives. We emphasize that the trace equation (4.10c) and the evolution equation of matter field (4.10d) below are coupled to the field equations.

Lemma 4.2 (Conformal field equations in wave coordinates)

The field equations (3.7) in wave coordinates take the form

$$(4.10a) \quad \begin{aligned} &g^{\dagger\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}g^\dagger_{\alpha\beta} \\ &= F_{\alpha\beta}(g^\dagger; \partial g^\dagger, \partial g^\dagger) - 12\partial_{\alpha\rho}\partial_{\beta\rho} + \frac{W_2(\rho)}{e^{2\rho}}g^\dagger_{\alpha\beta} - \frac{8\pi}{e^{2\rho}}(2T_{\alpha\beta} - \text{tr}(T)g_{\alpha\beta}), \end{aligned}$$

which is supplemented with the algebraic constraint equation (derived from the wave coordinate condition)

$$(4.10b) \quad g^{\dagger\alpha\beta}\Gamma_{\alpha\beta}^{\dagger\gamma} = 0.$$

In addition, this system must be supplemented with the trace equation

$$(4.10c) \quad g^{\dagger\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}\rho = \frac{W_2(\rho)}{6e^{2\rho}} + \frac{W_3(\rho)}{6e^{4\rho}} + \frac{4\pi\text{tr}(T)}{3e^{4\rho}}$$

with $\rho = \frac{1}{2}f'(R_g)$, as well as with the evolution equation for the matter

$$(4.10d) \quad \nabla^{\dagger\alpha}T_{\alpha\beta} = (\text{tr}(T)\partial_{\beta\rho} - 2g^{\delta\gamma}\partial_{\gamma\rho}T_{\gamma\beta})e^{-2\rho}.$$

4.2. A nonlinear wave system for the modified gravity theory

The aim of this subsection is to study the ‘essential system’ consisting of (4.10a) and (4.10b). If we remove the terms in ρ , this is a quasi-linear wave system with constraints (see (4.10b)), whose structure is quite involved. The strategy we propose is to replace these constraints by another differential equation which will turn out to be simpler to handle. We will write a new system which may (a priori) not be equivalent to the original system (3.7). This system is defined as follows.

Definition 4.3. — *The wave-reduced system in geometric form associated with (3.7) is, by definition,*

$$(4.11a) \quad N_{g_{\alpha\beta}} - \frac{1}{2}g_{\alpha\beta} \operatorname{tr}(N_g) - \frac{1}{2}e^{2\rho}(\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha) = 8\pi(T_{\alpha\beta} - \frac{1}{2}\operatorname{tr}(T)g_{\alpha\beta}),$$

$$(4.11b) \quad \nabla^{\dagger\alpha} T_{\alpha\beta} = (\operatorname{tr}(T)\partial_\beta \rho - 2g^{\delta\gamma}\partial_\gamma \rho T_{\gamma\beta})e^{-2\rho},$$

where $N_{g_{\alpha\beta}}$ is defined by (3.3) and $e^{2\rho} = f'(R_g)$.

Thanks to (3.3) and Lemma 4.1, the above system reads also

$$(4.12) \quad \begin{aligned} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} g^\dagger_{\alpha\beta} &= F_{\alpha\beta}(g^\dagger; \partial g^\dagger, \partial g^\dagger) - 12\partial_\alpha \rho \partial_\beta \rho + \frac{W_2(\rho)}{e^{2\rho}} g^\dagger_{\alpha\beta} - \frac{8\pi}{e^{2\rho}} (2T_{\alpha\beta} - \operatorname{tr}(T)g_{\alpha\beta}), \\ \nabla^{\dagger\alpha} T_{\alpha\beta} &= (\operatorname{tr}(T)\partial_\beta \rho - 2g^{\delta\gamma}\partial_\gamma \rho T_{\gamma\beta})e^{-2\rho}, \\ \rho &= \frac{1}{2} \ln f'(R_g), \end{aligned}$$

which is (4.10) without the constraint equations (4.10b) but includes the evolution equation (4.10d). In the following subsection, we will establish the following result.

Proposition 4.4 (Preservation of the wave coordinate conditions in modified gravity)

Consider a globally hyperbolic spacetime $M = \cup_{t \in [0, t_{\max})} M_t$ with metric g^\dagger , together with a matter field T defined on M . Suppose that the wave coordinate conditions

$$(4.13) \quad \Gamma^{\dagger\gamma} := g^{\dagger\alpha\beta} \Gamma^\dagger_{\alpha\beta} = 0 \quad \text{on the initial slice } M_0,$$

together with the constraint equations (3.14) and (3.15). Then, the wave conditions (4.13) are satisfied within the whole of M .

In other words, if one wants to find a solution of (3.7), what we need to do is to find first a solution of (4.12) with the constraint equations (4.13) satisfied on the initial slice, with four additional constraint equations to be required on the initial data set. Recall again that the interest of relying on (4.11) rather than on (3.7) is that the former one has a hyperbolic principal part (after removing the terms in ρ).

4.3. Preservation of the wave coordinate conditions

The key to Proposition 4.4 is the contracted Bianchi identity and, amazingly, the precise form of the modified gravity tensor N_g is not used at this juncture. We begin with some lemmas and the derivations of key identities.

Lemma 4.5. — *For any (pseudo-riemannian) manifold (M, g) , the following identity holds:*

$$(4.14) \quad \begin{aligned} & \nabla^\alpha (\partial_\alpha \Gamma_\beta + \partial_\beta \Gamma_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma_{\beta'}) \\ &= g^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma_\beta - g^{\alpha\alpha'} \Gamma_{\alpha\alpha'}^\gamma \partial_\gamma \Gamma_\beta - g^{\alpha\alpha'} \Gamma_{\alpha'\beta}^\delta \partial_\delta \Gamma_\alpha - \partial_\beta g^{\alpha'\beta'} (\partial_{\alpha'} \Gamma_{\beta'}), \end{aligned}$$

where $\Gamma_{\alpha\beta}^\gamma$ denote the Christoffel symbols and $\Gamma^\gamma := g^{\alpha\beta} \Gamma_{\alpha\beta}^\gamma$ and $\Gamma_\gamma = g_{\gamma'\gamma} \Gamma^{\gamma'}$.

Proof. — From the standard identities

$$\begin{aligned} \nabla^\alpha \partial_\beta u &= g^{\alpha\alpha'} \partial_{\alpha'} \partial_\beta u - \Gamma_{\alpha'\beta}^\gamma \partial_\gamma u, \\ \square_g u &= g^{\alpha\beta} \partial_\alpha \partial_\beta u - g^{\alpha\beta} \Gamma_{\alpha\beta}^\gamma \partial_\gamma u, \end{aligned}$$

we find

$$\begin{aligned} & \nabla^\alpha (\partial_\alpha \Gamma_\beta + \partial_\beta \Gamma_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma_{\beta'}) \\ &= \nabla^\alpha \partial_\alpha \Gamma_\beta + \nabla^\alpha \partial_\beta \Gamma_\alpha - \partial_\beta (g^{\alpha'\beta'} \partial_{\alpha'} \Gamma_{\beta'}) \\ &= \square_g \Gamma_\beta + g^{\alpha\alpha'} \partial_{\alpha'} \partial_\beta \Gamma_\alpha - g^{\alpha\alpha'} \Gamma_{\alpha'\beta}^\delta \partial_\delta \Gamma_\alpha - \partial_\beta g^{\alpha'\beta'} (\partial_{\alpha'} \Gamma_{\beta'}) - g^{\alpha'\beta'} \partial_\beta \partial_{\alpha'} \Gamma_{\beta'} \end{aligned}$$

thus

$$\begin{aligned} & \nabla^\alpha (\partial_\alpha \Gamma_\beta + \partial_\beta \Gamma_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma_{\beta'}) \\ &= \square_g \Gamma_\beta - g^{\alpha\alpha'} \Gamma_{\alpha'\beta}^\delta \partial_\delta \Gamma_\alpha - \partial_\beta g^{\alpha'\beta'} (\partial_{\alpha'} \Gamma_{\beta'}) \\ &= g^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma_\beta - g^{\alpha\alpha'} \Gamma_{\alpha\alpha'}^\gamma \partial_\gamma \Gamma_\beta^\dagger - g^{\alpha\alpha'} \Gamma_{\alpha'\beta}^\delta \partial_\delta \Gamma_\alpha - \partial_\beta g^{\alpha'\beta'} (\partial_{\alpha'} \Gamma_{\beta'}). \end{aligned}$$

□

Our next lemma establishes a relation between the wave condition and the evolution equation of the wave-reduced system (4.12). Recall that sufficiently regular is assumed throughout so that all terms under consideration are continuous functions at least.

Lemma 4.6. — *Consider an arbitrary manifold (M, g^\dagger) and a (matter) tensor $T_{\alpha\beta}$. Then, if in some local coordinate system $\{x^0, x^1, x^2, x^3\}$, g^\dagger , $T_{\alpha\beta}$ satisfy (4.11), then the following equations hold*

$$(4.15) \quad g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta = F_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma),$$

where $F_\beta(\rho, g^\dagger; \cdot, \cdot)$ is a combination of linear and bilinear forms, and one recalls $\Gamma_{\alpha\beta}^{\dagger\gamma}$ are the Christoffel symbols with $\Gamma^\dagger_\beta = g_{\beta\beta'} \Gamma^{\dagger\beta'} = g_{\beta\beta'} g^{\alpha\gamma} \Gamma_{\alpha\gamma}^{\dagger\beta'}$.

Proof. — Taking the trace of (4.11a) with respect to the metric g ,

$$(4.16) \quad \text{tr} (N_g - 8\pi T) = -e^{2\rho} g^{\alpha\beta} \partial_\alpha \Gamma^\dagger_\beta$$

and combining with (4.11a), we obtain

$$(4.17) \quad \begin{aligned} N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta} &= \frac{1}{2} g_{\alpha\beta} \text{tr} (N_g - 8\pi T) + \frac{1}{2} e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha) \\ &= \frac{1}{2} e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \end{aligned}$$

Taking the trace of (4.17) with respect to g^\dagger , we obtain

$$(4.18) \quad \frac{1}{2} \nabla^{\dagger\alpha} (e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'})) = \nabla^{\dagger\alpha} (N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta}),$$

whose left-hand-side is evaluated by using (4.14):

$$\begin{aligned} &\frac{1}{2} \nabla^{\dagger\alpha} (e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'})) \\ &= \frac{1}{2} e^{2\rho} \nabla^{\dagger\alpha} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \\ &\quad + e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \nabla^{\dagger\alpha} \rho. \end{aligned}$$

We thus obtain that

$$\begin{aligned} &\frac{1}{2} \nabla^{\dagger\alpha} (e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'})) \\ &= \frac{1}{2} e^{2\rho} (g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta - g^{\dagger\alpha\alpha'} \Gamma^{\dagger\gamma}_{\alpha\alpha'} \partial_\gamma \Gamma^\dagger_\beta - g^{\dagger\alpha\alpha'} \Gamma^{\dagger\delta}_{\alpha'\beta} \partial_\delta \Gamma^\dagger_\alpha - \partial_\beta g^{\dagger\alpha'\beta'} (\partial_{\alpha'} \Gamma^\dagger_{\beta'})) \\ &\quad + e^{2\rho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \nabla^{\dagger\alpha} \rho \\ &=: \frac{1}{2} e^{2\rho} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta + \frac{1}{2} e^{2\rho} \tilde{F}_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma), \end{aligned}$$

where $F_\beta(\rho, g^\dagger; \cdot)$ is a combination of linear and bilinear forms of the functions Γ^\dagger_γ and $\partial \Gamma^\dagger_\gamma$.

The right-hand side of (4.18) is computed by using the identity (3.16) and (4.11b), that is,

$$\nabla^{\dagger\alpha} (N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta}) = e^{-2\rho} (\text{tr} (N_g - 8\pi T) \partial_\beta \rho - 2g^{\gamma\delta} \partial_\gamma \rho (N_{g_{\delta\beta}} - 8\pi T_{\delta\beta})).$$

Then, by (4.16), we obtain

$$\begin{aligned} &\nabla^{\dagger\alpha} (N_{g_{\alpha\beta}} - 8\pi T_{\alpha\beta}) \\ &= e^{-2\rho} \left(-e^{2\rho} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'} \partial_\beta \rho - e^{2\rho} g^{\gamma\delta} \partial_\gamma \rho (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g^\dagger_{\alpha\beta} g^{\dagger\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \right) \\ &= -g^{\gamma\delta} \partial_\gamma \rho (\partial_\gamma \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\delta) \end{aligned}$$

and, by (4.18),

$$(4.19) \quad g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta = F_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma),$$

where

$$F_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial\Gamma^\dagger_\gamma) = -\tilde{F}_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial\Gamma^\dagger_\gamma) - 2e^{-2\rho}g^{\gamma\delta}\partial_\gamma\rho(\partial_\delta\Gamma^\dagger_\beta + \partial_\beta\Gamma^\dagger_\delta)$$

is a combination of linear and bilinear forms in Γ^\dagger_γ and $\partial\Gamma^\dagger_\gamma$. \square

Lemma 4.7. — *Let (M, g^\dagger) be a globally hyperbolic Lorentzian manifold endowed with a foliation $M = [0, t_{max}) \times M_t$ (and signature $(-, +, +, +)$), together with a tensor field $T_{\alpha\beta}$. Suppose that the equation (4.11) holds on the initial slice M_0 and, furthermore, the wave coordinate conditions and the constraint equations hold on the slice M_0 :*

$$\Gamma^{\dagger\gamma} := g^{\dagger\alpha\beta}\Gamma^{\dagger\gamma}_{\alpha\beta} = 0$$

and (as stated in (3.21)-(3.22))

$$\begin{aligned} \bar{R}^\dagger - K^\dagger_{ij}K^{\dagger ij} + (K^{\dagger l}_l)^2 &= \frac{16\sigma^\dagger}{e^{2\rho}} + \frac{12|\langle e_0, \rho \rangle|^2}{N^{\dagger 2}} + 6g^\dagger(\nabla\rho, \nabla\rho) - e^{-2\rho}W_2(\rho), \\ \partial_j K^{\dagger l}_l - \bar{\nabla}^\dagger K^{\dagger l}_j &= -\frac{J^\dagger_j}{e^{2\rho}} + \frac{6\langle e_0, \rho \rangle \partial_j \rho}{N^\dagger}. \end{aligned}$$

Then, one has

$$\partial_0\Gamma^{\dagger\lambda} = 0 \quad \text{in the spacetime } M.$$

Proof. — We work in a Cauchy adapted frame (e_0, e_1, e_2, e_3) , that is,

$$e_0 = \partial_0 - \beta^i \partial_i, \quad \beta^i = \frac{g^\dagger_{0i}}{g^\dagger_{ii}}, \quad e_i = \partial_i,$$

so that $g^\dagger(e_0, e_i) = 0$. A tensor can be written in, both, the natural frame and the Cauchy adapted frame. We denote by an underlined letter the components in the Cauchy adapted frame. For example, $\underline{T}_{\alpha\beta}$ are the components of T in the Cauchy frame.

Recall that the momentum constraint equations are equivalent to

$$(4.20) \quad \underline{N}_{g_{0j}} = 8\pi \underline{T}_{0j},$$

the Hamiltonian constraint equation is equivalent to

$$(4.21) \quad \underline{N}_{g_{00}} = 8\pi \underline{T}_{00}.$$

Recall also that the Cauchy adapted frame is expressed in the natural frame via $e_\alpha = \Phi_\alpha^\beta \partial_\beta$, where

$$(\Phi_\alpha^\beta)_{\alpha\beta} = \begin{pmatrix} 1 & -\beta^1 & -\beta^2 & -\beta^3 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

Then, we have

$$\underline{N}_{g_{\alpha\beta}} = N_{g_{\alpha'\beta'}} \Phi_\alpha^{\alpha'} \Phi_\beta^{\beta'}, \quad \underline{T}_{\alpha\beta} = T_{\alpha'\beta'} \Phi_\alpha^{\alpha'} \Phi_\beta^{\beta'}.$$

Observe that the wave-reduced field equation (4.11a) can be rewritten in the Cauchy adapted frame as

$$\underline{N}_{g_{\alpha\beta}} - \frac{1}{2} \text{tr } N_g \underline{g}_{\alpha\beta}^\dagger - \frac{1}{2} e^{2\rho} \Phi_\alpha^{\alpha'} \Phi_\beta^{\beta'} (\partial_{\alpha'} \Gamma_{\beta'}^\dagger + \partial_{\beta'} \Gamma_{\alpha'}^\dagger) = 8\pi (\underline{T}_{\alpha\beta} - \frac{1}{2} \text{tr } (T) \underline{g}_{\alpha\beta}^\dagger),$$

which is

$$(4.22) \quad \begin{aligned} & \underline{N}_{g_{\alpha\beta}} - \frac{1}{2} \text{tr } N_g \underline{g}_{\alpha\beta}^\dagger - \frac{1}{2} e^{2\rho} (\Phi_\beta^{\beta'} \langle e_\alpha, \Gamma_{\beta'}^\dagger \rangle + \Phi_\alpha^{\alpha'} \langle e_\beta, \Gamma_{\alpha'}^\dagger \rangle) \\ & = 8\pi (\underline{T}_{\alpha\beta} - \frac{1}{2} \text{tr } (T) \underline{g}_{\alpha\beta}^\dagger). \end{aligned}$$

Next, by combining (4.22) with (4.20) with $\alpha = 0$, $1 \leq \beta = b \leq 3$ in (4.22) and by observing that $\underline{g}_{0j}^\dagger = 0$, we obtain

$$\Phi_b^{\beta'} \langle e_0, \Gamma_{\beta'}^\dagger \rangle + \Phi_0^{\alpha'} \langle e_b, \Gamma_{\alpha'}^\dagger \rangle = 0.$$

We consider this equation on the initial slice M_0 .

Recall that $\Gamma^{\dagger\lambda} = 0$ so that $\partial_b \Gamma_\beta = 0$ for any $1 \leq b \leq 3$ and $0 \leq \beta \leq 3$. Then observe that $\Phi_b^{\beta'} = \delta_b^{\beta'}$ and $\langle e_b, \Gamma_{\alpha'}^\dagger \rangle = \partial_b \Gamma_{\alpha'}^\dagger = 0$, so that

$$(4.23) \quad \langle e_0, \Gamma_b^\dagger \rangle = 0,$$

which leads us to $\partial_t \Gamma_b^\dagger = 0$.

Now, we can combine (4.22) with (4.20) and (4.21) with $\alpha = \beta = 0$ in (4.22):

$$\Phi_0^{\beta'} \langle e_0, \Gamma_{\beta'}^\dagger \rangle + \Phi_0^{\alpha'} \langle e_0, \Gamma_{\alpha'}^\dagger \rangle = -\text{tr } (N_g - 8\pi T) g_{00}^\dagger.$$

We recall (4.20) and (4.21) and the fact that $g_{0j}^\dagger = 0$, so that

$$(4.24) \quad \text{tr } (N_g - 8\pi T) = \underline{g}^{\alpha\beta} (\underline{N}_{g_{\alpha\beta}} - 8\pi \underline{T}_{\alpha\beta}) = \underline{g}^{ij} (\underline{N}_{g_{ij}} - 8\pi \underline{T}_{ij}).$$

We also recall (4.22) with $1 \leq \alpha = i \leq 3$ and $1 \leq \beta = j \leq 3$, and we observe that $\langle e_i, \Gamma_\alpha^\dagger \rangle = 0$. This shows that

$$\underline{N}_{g_{ij}} - 8\pi \underline{T}_{ij} = \frac{1}{2} \text{tr } (N_g - 8\pi T) \underline{g}_{ij},$$

which leads us to

$$\underline{g}^{ij} (\underline{N}_{g_{ij}} - 8\pi \underline{T}_{ij}) = \frac{1}{2} \text{tr } (N_g - 8\pi T) \underline{g}^{ij} \underline{g}_{ij}.$$

Therefore, by (4.24), we have $\text{tr } (N_g - 8\pi T) = \frac{3}{2} \text{tr } (N_g - 8\pi T)$, and thus $\text{tr } (N_g - 8\pi T) = 0$. We substitute this conclusion in (4.23) and obtain

$$\Phi_0^{\beta'} \langle e_0, \Gamma_{\beta'}^\dagger \rangle + \Phi_0^{\alpha'} \langle e_0, \Gamma_{\alpha'}^\dagger \rangle = 0.$$

We finally recall (4.23) and get

$$\Phi_0^0 \langle e_0, \Gamma_0^\dagger \rangle + \Phi_0^0 \langle e_0, \Gamma_0^\dagger \rangle = 0$$

and the desired conclusion is reached. \square

Proof of Proposition 4.4. — In view of Lemma 4.6 and Lemma 4.7, we see that Γ^\dagger_β satisfies the initial value problem

$$g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta = F_\beta(\rho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma)$$

with initial data

$$\Gamma^\dagger_\beta|_{x^0=0} = 0, \quad \partial_0 \Gamma^\dagger_\beta|_{x^0=0} = 0.$$

Since x^0 is the timelike direction and the symmetry of g^\dagger guarantees the hyperbolicity of $g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'}$. We also observe that $\Gamma^\dagger = 0$ is a solution of this initial value problem. Thanks to the global hyperbolicity of g^\dagger , the desired uniqueness result holds within the domain of determinacy of the initial slice, that is, M itself, thanks to our global hyperbolicity assumption. \square

5. The augmented conformal formulation

5.1. A novel formulation

In this section we re-formulate our system and establish a local-in-time existence theory. Since ρ is function of the scalar curvature, the system under consideration now is still *third-order* and is not of a specific PDE type. (The third-order terms are $\partial\rho$, where ρ is a function of the second-order derivatives of $g^\dagger_{\alpha\beta}$.) To bypass this difficulty, we thus introduce still another transformation. In (4.11), we now replace the constraint $e^{2\rho} = f'(R_g)$ by the trace equation leading to the evolution equation for ρ , and we introduce an *augmented formulation*, as we call it. For clarity in the presentation, we switch from the notation ρ to the notation ϱ , in order to emphasize that the relation $e^{2\rho} = f'(R_g)$ is *no longer imposed*.

Let us now define the tensor $N^\dagger_{g_{\alpha\beta}}$ as

$$(5.1) \quad N^\dagger_{g_{\alpha\beta}} - \frac{1}{2} g_{\alpha\beta} \text{tr} (N^\dagger_g) = e^{2\varrho} R^\dagger_{\alpha\beta} - 6e^{2\varrho} \partial_\alpha \varrho \partial_\beta \varrho + \frac{1}{2} g^\dagger_{\alpha\beta} W_2(\varrho).$$

Here, ϱ plays the role of the previous quantity $\rho = \frac{1}{2} \ln f'(R_g)$ (which need not hold), and we work in the **metric of the augmented system**

$$(5.2) \quad g^\dagger_{\alpha\beta} = e^{2\varrho} g_{\alpha\beta}.$$

We also use the notation ∇^\dagger , R^\dagger_α and R^\dagger for the connection, the Ricci curvature, and the scalar curvature of g^\dagger , respectively. Also, we denote by $\Gamma^\dagger_{\alpha\beta}{}^\gamma$ the Christoffel symbols of g^\dagger , and we set $\Gamma^\dagger{}^\gamma := g^{\dagger\alpha\beta} \Gamma^\dagger_{\alpha\beta}{}^\gamma$ and $\Gamma^\dagger_{\gamma'} := g^\dagger_{\gamma\gamma'} \Gamma^\dagger{}^\gamma$.

Definition 5.1. — *The conformal augmented formulation of modified gravity is the following partial differential system:*

$$(5.3a) \quad N^\dagger_{g_{\alpha\beta}} - \frac{1}{2} e^{2\varrho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha) = 8\pi (T_{\alpha\beta} - \frac{1}{2} g_{\alpha\beta} \text{tr} (T)),$$

$$(5.3b) \quad \square_{g^\dagger} \varrho = \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{4\pi \text{tr} (T)}{3e^{4\varrho}} g^\dagger_{\alpha\beta}.$$

In the now proposed standpoint, ϱ is an *independent unknown*, which no longer depends upon the scalar curvature R_g . In this way, the system under consideration is *second-order*. Our first task then is to compute the divergence of N_g^\ddagger .

Lemma 5.2. — *When (5.3b) holds, the following identity also holds:*

$$(5.4) \quad \nabla^{\ddagger\alpha} N_{g_{\alpha\beta}}^\ddagger = e^{-2\varrho} (2g^{\alpha\alpha'} \partial_{\alpha'} \varrho N_{g_{\alpha\beta}}^\ddagger - \text{tr} (8\pi T) \partial_\beta \varrho).$$

Proof. — By (5.3), we have

$$N_{g_{\alpha\beta}}^\ddagger = e^{2\varrho} G_{\alpha\beta}^\ddagger - 6e^{2\varrho} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta} |\nabla \varrho|_{g^\ddagger}^2) - \frac{1}{2} g_{\alpha\beta}^\ddagger W_2(\varrho),$$

where $G_{\alpha\beta}^\ddagger := R_{\alpha\beta}^\ddagger - \frac{1}{2} g_{\alpha\beta}^\ddagger R^\ddagger$ is the Einstein curvature of g^\ddagger . We start from the identities

$$\nabla^{\ddagger\alpha} G_{\alpha\beta}^\ddagger = 0,$$

and

$$\nabla^{\ddagger\alpha} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta}^\ddagger |\nabla \varrho|_{g^\ddagger}^2) = \partial_\beta \varrho \square_{g^\ddagger} \varrho,$$

and we introduce the function

$$(5.5) \quad s(r) = \frac{1}{2} \ln f'(r), \quad r \in \mathbb{R},$$

together with its local inverse near 0, denoted by θ . Then, we have

$$\begin{aligned} \nabla^{\ddagger\alpha\beta} (g_{\alpha\beta}^\ddagger W_2(\varrho)) &= \partial_\beta (W_2(\varrho)) = \partial_\beta (e^{-2\varrho} f(\theta(\varrho)) - \theta(\varrho)) \\ &= \left(-2e^{-2\varrho} f(\theta(\varrho)) + e^{-2\varrho} f'(\theta(\varrho)) \theta'(\varrho) \partial_\beta \varrho - \theta'(\varrho) \right) \partial_\beta \varrho \\ &= -2e^{-2\varrho} f(\theta(\varrho)) \partial_\beta \varrho. \end{aligned}$$

This allows us to compute the divergence of $N_{g_{\alpha\beta}}^\ddagger$:

$$\begin{aligned} \nabla^{\ddagger\alpha} N_{g_{\alpha\beta}}^\ddagger &= 2e^{2\varrho} G_{\alpha\beta}^\ddagger \nabla^{\ddagger\alpha} \varrho - 12e^{2\varrho} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta} |\nabla \varrho|_{g^\ddagger}^2) \nabla^{\ddagger\alpha} \varrho \\ &\quad - 6e^{2\varrho} \partial_\beta \varrho \square_{g^\ddagger} \varrho + e^{2\varrho} \partial_\beta f(\theta(\varrho)) \\ &= 2e^{2\varrho} G_{\alpha\beta}^\ddagger \nabla^{\ddagger\alpha} \varrho - 12e^{2\varrho} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta} |\nabla \varrho|_{g^\ddagger}^2) \nabla^{\ddagger\alpha} \varrho - g_{\alpha\beta}^\ddagger W_2(\varrho) \nabla^{\ddagger\alpha} \varrho \\ &\quad - 6e^{2\varrho} \partial_\beta \varrho \square_{g^\ddagger} \varrho + e^{-2\varrho} \partial_\beta f(\theta(\varrho)) + g_{\alpha\beta}^\ddagger W_2(\varrho) \nabla^{\ddagger\alpha} \varrho \\ &= 2N_{g_{\alpha\beta}}^\ddagger \nabla^{\ddagger\alpha} \varrho - 6e^{2\varrho} \partial_\beta \varrho \left(\square_{g^\ddagger} - \frac{f(\theta(\varrho))}{6e^{4\varrho}} - \frac{W_2(\varrho)}{6e^{2\varrho}} \right). \end{aligned}$$

Then, by (5.9c), we find

$$\square_{g^\ddagger} \varrho = \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{4\pi \text{tr} (T)}{3e^{4\varrho}}$$

and the desired conclusion follows. \square

By Lemma 5.2, we see that the evolution law for T is

$$(5.6) \quad \nabla^{\dagger\alpha} T_{\alpha\beta} = (2g^{\delta\gamma} \partial_\gamma \varrho T_{\gamma\beta} - \text{tr}(T) \partial_\beta \varrho) e^{-2e}.$$

The following question arises at this juncture: Will the relation $e^{2e} = f'(R_g)$ (with $g^\dagger_{\alpha\beta} = e^{2e} g_{\alpha\beta}$) hold if we solely solve the equations (5.3)? The following subsections precisely provide a (positive) answer to this question.

First, in Section 5.2, we will re-formulate the initial value problem for the augmented conformal formulation, by building upon our previous formulations of the field equations in the Jordan and Einstein metrics.

Then, in Section 5.3, in order for the principal part of (5.3) to be hyperbolic, we will write our augmented system in wave coordinates. Finally, in Section 5.4, we will prove that once the wave constraint equations $\Gamma^\dagger_\lambda = 0$ hold on the initial slice, then the augmentation relation $e^{2e} = f'(R_g)$ is guaranteed by (5.10).

5.2. Initial value formulation for the augmented conformal system

In this section, we revisit Definitions 3.7 and 3.8, as follows.

Definition 5.3. — *An initial data set for the augmented conformal formulation of modified gravity $(\overline{M}, \overline{g}^\dagger, K^\dagger, \varrho_0, \varrho_1, \phi_0^\dagger, \phi_1^\dagger)$ consists of the following data:*

- a 3-dimensional manifold \overline{M} endowed with a Riemannian metric \overline{g}^\dagger and a symmetric $(0, 2)$ -tensor field K^\dagger ,
- two scalar fields denoted by ϱ_0 and ϱ_1 on \overline{M} and representing the (to-be-constructed) conformal factor and its time derivative,
- and two scalar field ϕ_0^\dagger and ϕ_1^\dagger defined on \overline{M}

Moreover, these data are required to satisfy the **Hamiltonian constraint of modified gravity in the augmented conformal form**

$$(5.7) \quad \begin{aligned} \overline{R}^\dagger - K^\dagger_{ij} K^{\dagger ij} + (K^{\dagger j}_j)^2 \\ = 8 e^{-2\rho} ((\phi_1^\dagger)^2 + \overline{g}^{\dagger ij} \partial_i \phi_0^\dagger \partial_j \phi_0^\dagger) + 6(\varrho_1)^2 + 6\overline{g}^{\dagger ij} \partial_i \varrho_0 \partial_j \varrho_0 - e^{-2\varrho_0} W_2(\varrho_0) \end{aligned}$$

and the **momentum constraint in the augmented conformal form**

$$(5.8) \quad \partial_j K^{\dagger i}_i - \overline{\nabla}^\dagger_i K^{\dagger i}_j = \frac{\phi_1^\dagger \partial_j \phi_0^\dagger}{e^{2\varrho_0}} + 6\varrho_1 \partial_j \varrho_0.$$

Here, \overline{R}^\dagger and $\overline{\nabla}^\dagger$ is the scalar curvature and the connection of the metric \overline{g}^\dagger , respectively.

Definition 5.4. — *Given an initial data set $(\overline{M}, \overline{g}^\dagger, K^\dagger, \varrho_0, \varrho_1, \phi_0^\dagger, \phi_1^\dagger)$ as in Definition 5.3, the initial value problem in modified gravity in the augmented conformal form consists of finding a Lorentzian manifold (M, g) and a two-tensor field $T_{\alpha\beta}$ on M such that:*

1. The augmented conformal metric $g^\dagger_{\alpha\beta} = e^{2e} g_{\alpha\beta}$ with conformal factor ϱ satisfies the **evolution equation** (5.9c).

2. The augmented conformal field equations (5.9) are satisfied.
3. There exists an embedding $i : \bar{M} \rightarrow M$ with pull-back metric $\bar{g}^\ddagger = i^*g^\ddagger$ and second fundamental form K^\ddagger .
4. The field ϱ_0 coincides with the restriction of the conformal factor ϱ on \bar{M} , while ϱ_1 coincides with the Lie derivative $\mathcal{L}_{n^\ddagger}\varrho$ restricted to \bar{M} , where n^\ddagger denotes the normal unit vector of \bar{M} (with respect to g^\ddagger).
5. The scalar fields $\phi_0^\ddagger, \phi_1^\ddagger$ coincide with the restriction of $\phi, \mathcal{L}_{n^\ddagger}\phi$ on \bar{M} .

Such a solution to (3.7) is referred to as a **modified gravity development of the initial data set** $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$.

Note that, as in Section 3, the geometric form of the constraint equations is the **Hamiltonian constraint** $N_{g_{00}} = 8\pi T_{00}$ and the **momentum constraint** $N_{g_{0\alpha}} = 8\pi T_{0\alpha}$.

5.3. Augmented conformal formulation in wave coordinates

We now reduce the conformal augmented system (5.3) in wave coordinates to a system with hyperbolic principal part. Indeed, we obtain the following formulation, where we replace the wave constraints by the evolution law of T given by (5.4).

Definition 5.5. — *The conformal augmented formulation of modified gravity is, by definition, the following partial differential system:*

$$(5.9a) \quad N_{g_{\alpha\beta}}^\ddagger - \frac{1}{2}g_{\alpha\beta}tr(N_g^\ddagger) - \frac{1}{2}e^{2\varrho}(\partial_\alpha\Gamma_{\beta}^\ddagger + \partial_\beta\Gamma_{\alpha}^\ddagger) = 8\pi(T_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}tr(T)),$$

$$(5.9b) \quad \nabla^{\ddagger\alpha}T_{\alpha\beta} = (2g^{\delta\gamma}\partial_\gamma\varrho T_{\gamma\beta} - tr(T)\partial_\beta\varrho)e^{-2\varrho},$$

$$(5.9c) \quad \square_{g^\ddagger}\varrho + \Gamma^{\ddagger\delta}\partial_\delta\varrho = \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{4\pi tr(T)}{3e^{4\varrho}}g^{\ddagger\alpha\beta}.$$

By Lemma 4.1, we then have the following expressions in coordinates, in which we emphasize that g^\ddagger need not coincide with g^\ddagger .

Lemma 5.6. — *The conformal augmented formulation of modified gravity theory in coordinates reads*

$$(5.10a) \quad g^{\ddagger\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}g_{\alpha\beta}^\ddagger = F_{\alpha\beta}(g^\ddagger; \partial g^\ddagger, \partial g^\ddagger) - 12\partial_\alpha\varrho\partial_\beta\varrho + \frac{W_2(\varrho)}{e^{2\varrho}}g_{\alpha\beta}^\ddagger - 16\pi(T_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}tr(T)),$$

$$(5.10b) \quad \nabla^{\ddagger\alpha}T_{\alpha\beta} = (2g^{\delta\gamma}\partial_\delta\varrho T_{\gamma\beta} - tr(T)\partial_\beta\varrho)e^{-2\varrho},$$

$$(5.10c) \quad g^{\ddagger\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}\varrho = \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{4\pi tr(T)}{3e^{4\varrho}},$$

with

$$(5.11) \quad g_{\alpha\beta}^\ddagger = e^{2\varrho}g_{\alpha\beta}.$$

5.4. Preservation of the constraints

Our first task is to address the problem of the preservation of the constraints.

Proposition 5.7. — *Let (M, g^\dagger) be a globally hyperbolic Lorentzian manifold endowed with foliation $M = [0, t_{max})$ and with signature $(-, +, +, +)$. Let T be a symmetric two-tensor (representing the matter content) and let ϱ be a scalar field defined in $(-\epsilon, \epsilon) \times \mathbb{R}^3$. Furthermore, assume that $(g^\dagger_{\alpha\beta}, \phi, \varrho)$ is a solution to the conformal field equations (5.10). Let $g_{\alpha\beta} := e^{-2\varrho} g^\dagger_{\alpha\beta}$ be the metric conformal to g^\dagger . Then, provided the constraint equations (5.7) and (5.8) together with the constraint equations*

$$(5.12) \quad g^{\dagger\alpha\beta} \Gamma_{\alpha\beta}^{\dagger\gamma} = 0$$

are satisfied on the slice $\{x^0 = 0\}$ (where R_g is the scalar curvature of $g_{\alpha\beta}$ and $\Gamma_{\alpha\beta}^{\dagger\gamma}$ are the Christoffel symbols of g^\dagger), then (5.12) holds in the whole of $(-\epsilon, \epsilon) \times \mathbb{R}^3$. Furthermore, one has

$$(5.13) \quad e^{2\varrho} = e^{2\rho} = \frac{1}{2} \ln(f'(R_g))$$

in $(-\epsilon, \epsilon) \times \mathbb{R}^3$, so that $(g^\dagger_{\alpha\beta}, \phi, \varrho)$ is also a solution to (3.7).

The rest of this section is devoted to the proof of this proposition. Recall that, throughout, we assume that the $g^\dagger_{\alpha\beta}$, ϱ and $T_{\alpha\beta}$ are sufficiently regular, so that all relevant derivatives are continuous at least. For the proof, we need some preliminary material and, first of all, we compute the divergence of N_g^\dagger .

Lemma 5.8. — *When (5.9c) holds, the following identity holds:*

$$(5.14) \quad \nabla^{\dagger\alpha} N_{g_{\alpha\beta}}^\dagger = e^{-2\varrho} (2g^{\alpha\alpha'} \partial_{\alpha'} \varrho N_{g_{\alpha\beta}}^\dagger - \text{tr}(8\pi T) \partial_\beta \varrho) + 6e^{2\varrho} \partial_\beta \varrho \Gamma^{\dagger\delta} \partial_\delta \varrho.$$

Proof. — First of all, in view of (5.3), we have

$$N_{g_{\alpha\beta}}^\dagger = e^{2\varrho} G_{\alpha\beta}^\dagger - 6e^{2\varrho} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta} |\nabla \varrho|_{g^\dagger}^2) - \frac{1}{2} g^\dagger_{\alpha\beta} W_2(\varrho),$$

where $G_{\alpha\beta}^\dagger := R_{\alpha\beta}^\dagger - \frac{1}{2} g^\dagger_{\alpha\beta} R^\dagger$ is the Einstein curvature of g^\dagger . We have the identities

$$\begin{aligned} \nabla^{\dagger\alpha\beta} G_{\alpha\beta}^{(\varrho)} &= 0, \\ \nabla^{\dagger\alpha\beta} (\partial_\alpha \varrho \partial_\beta \varrho - \frac{1}{2} g_{\alpha\beta} |\nabla \varrho|_{g^\dagger}^2) &= \partial_\beta \varrho \square_{g^\dagger} \varrho. \end{aligned}$$

As before, by introducing the function

$$(5.15) \quad s(r) = \frac{1}{2} \ln f'(r), \quad r \in \mathbb{R}$$

with (local) inverse denoted by θ and defined near 0 at least, we can write

$$\begin{aligned}\nabla^{\dagger\alpha\beta}(g_{\alpha\beta}^{\dagger}W_2(\varrho)) &= \partial_{\beta}(W_2(\varrho)) = \partial_{\beta}(e^{-2\varrho}f(\theta(\varrho)) - \theta(\varrho)) \\ &= -2e^{-2\varrho}\partial_{\beta}\varrho f(\theta(\varrho)) + e^{-2\varrho}f'(\theta(\varrho))\theta'(\varrho)\partial_{\beta}\varrho - \theta'(\varrho)\partial_{\beta}\varrho \\ &= -2e^{-2\varrho}\partial_{\beta}\varrho f(\theta(\varrho)) + e^{-2\varrho}e^{2\varrho}\theta'(\varrho)\partial_{\beta}\varrho - \theta'(\varrho)\partial_{\beta}\varrho \\ &= -2e^{-2\varrho}\partial_{\beta}\varrho f(\theta(\varrho)).\end{aligned}$$

We are now in a position to compute the divergence

$$\begin{aligned}\nabla^{\dagger\alpha}N_{g_{\alpha\beta}}^{\dagger} &= 2e^{2\varrho}G_{\alpha\beta}^{\varrho}\nabla^{\dagger\alpha}\varrho - 12e^{2\varrho}(\partial_{\alpha}\varrho\partial_{\beta}\varrho - \frac{1}{2}g_{\alpha\beta}|\nabla\varrho|_{g^{\dagger}}^2)\nabla^{\dagger\alpha}\varrho \\ &\quad - 6e^{2\varrho}\partial_{\beta}\varrho\Box_{g^{\dagger}}\varrho + e^{2\varrho}\partial_{\beta}f(\theta(\varrho)) \\ &= 2e^{2\varrho}G_{\alpha\beta}^{\varrho}\nabla^{\dagger\alpha}\varrho - 12e^{2\varrho}(\partial_{\alpha}\varrho\partial_{\beta}\varrho - \frac{1}{2}g_{\alpha\beta}|\nabla\varrho|_{g^{\dagger}}^2)\nabla^{\dagger\alpha}\varrho - g_{\alpha\beta}^{\dagger}W_2(\varrho)\nabla^{\dagger\alpha}\varrho \\ &\quad - 6e^{2\varrho}\partial_{\beta}\varrho\Box_{g^{\dagger}}\varrho + e^{-2\varrho}\partial_{\beta}f(\theta(\varrho)) + g_{\alpha\beta}^{\dagger}W_2(\varrho)\nabla^{\dagger\alpha}\varrho \\ &= 2N_{g_{\alpha\beta}}^{\dagger}\nabla^{\dagger\alpha}\varrho - 6e^{2\varrho}\partial_{\beta}\varrho\left(\Box_{g^{\dagger}} - \frac{f(\theta(\varrho))}{6e^{4\varrho}} - \frac{W_2(\varrho)}{6e^{2\varrho}}\right).\end{aligned}$$

In view of (5.9c), we find

$$\Box_{g^{\dagger}}\varrho = -\Gamma^{\dagger\delta}\partial_{\delta}\varrho + \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{4\pi\text{tr}(T)}{3e^{4\varrho}}$$

and this yields the desired conclusion. \square

Lemma 5.9. — *The equation (5.10a) leads also to the wave equation for ϱ*

$$(5.16) \quad \Box_{g^{\dagger}}\varrho = \frac{R_g}{6e^{2\varrho}} + \frac{W_2(\varrho)}{3e^{2\varrho}} + \frac{4\pi\text{tr}(T)}{3e^{4\varrho}} - \frac{g^{\alpha\beta}\partial_{\alpha}\Gamma_{\beta}^{\dagger}}{6e^{2\varrho}}.$$

Proof. — We recall that (5.10a) can be written as

$$(5.17) \quad \begin{aligned}e^{2\varrho}R_{\alpha\beta}^{\dagger} - 6e^{2\varrho}\partial_{\alpha}\varrho\partial_{\beta}\varrho + \frac{g_{\alpha\beta}^{\dagger}}{2}W_2(\varrho) - \frac{e^{2\varrho}}{2}(\partial_{\alpha}\Gamma_{\beta}^{\dagger} + \partial_{\beta}\Gamma_{\alpha}^{\dagger}) \\ = 8\pi(T_{\alpha\beta} - \frac{1}{2}g_{\alpha\beta}\text{tr}(T)),\end{aligned}$$

and that (3.6) implies

$$R_{\alpha\beta}^{\dagger} = R_{\alpha\beta} - 2(\nabla_{\alpha}\nabla_{\beta}\varrho - \partial_{\alpha}\varrho\partial_{\beta}\varrho) - (\Box_g\varrho + 2g(\nabla\varrho, \nabla\varrho))g_{\alpha\beta}.$$

We substitute this relation into (5.17) and obtain

$$\begin{aligned}e^{2\varrho}R_{\alpha\beta} - 2e^{2\varrho}\nabla_{\alpha}\nabla_{\beta}\varrho - 4e^{2\varrho}\partial_{\alpha}\varrho\partial_{\beta}\varrho - g_{\alpha\beta}e^{2\varrho}\Box_g\varrho - 2e^{2\varrho}g_{\alpha\beta}g(\nabla\varrho, \nabla\varrho) + \frac{g_{\alpha\beta}^{\dagger}}{2}W_2(\varrho) \\ = 8\pi(T_{\alpha\beta} - \frac{1}{2}\text{tr}(T)g_{\alpha\beta}) + \frac{1}{2}e^{2\varrho}(\partial_{\alpha}\Gamma_{\beta}^{\dagger} + \partial_{\beta}\Gamma_{\alpha}^{\dagger}).\end{aligned}$$

By taking the trace of this equation *with respect to g* , we have

$$e^{2\varrho}R_g - 6e^{2\varrho}\Box_g\varrho - 12e^{2\varrho}g(\nabla\varrho, \nabla\varrho) + 2e^{2\varrho}W_2(\varrho) = -8\pi\text{tr}(T) + e^{2\varrho}g^{\alpha\beta}\partial_{\alpha}\Gamma_{\beta}^{\dagger},$$

which can also be written as

$$\square_g \varrho + 2g(\nabla \varrho, \nabla \varrho) = \frac{R_g}{6} + \frac{W_2(\varrho)}{3} + \frac{4\pi \text{tr}(T)}{3e^{2\varrho}} - \frac{g^{\alpha\beta}}{6} \partial_\alpha \Gamma^\dagger_\beta.$$

In view of the expression (3.8), the desired result is proven. \square

In the next lemma, we identify the geometric form of the system (5.10).

Lemma 5.10. — *Let (M, g^\dagger) be a Lorentzian manifold together with a two-tensor $T_{\alpha\beta}$. If, in some local coordinates $\{x^0, x^1, x^2, x^3\}$, $(g^\dagger, T_{\alpha\beta})$ satisfy (5.10), then the following equation holds (in the domain of the coordinate system):*

$$(5.18) \quad g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta = F_\beta(\varrho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma),$$

where $F_\beta(\varrho, g^\dagger; \cdot, \cdot)$ is a combination of linear and bilinear forms and $\Gamma^{\dagger\gamma}_{\alpha\beta}$ denote the associated Christoffel symbol with

$$\Gamma^\dagger_\beta = g^{\dagger\beta\beta'} \Gamma^{\dagger\beta'} = g^{\dagger\beta\beta'} g^{\dagger\alpha\gamma} \Gamma^{\dagger\beta'}_{\alpha\gamma}.$$

Proof. — By taking the trace of (5.9a) we have

$$(5.19) \quad e^{2\varrho} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'} = -\text{tr} (N_g^\dagger - 8\pi T).$$

Combining this results with (5.9a), we obtain

$$(5.20) \quad \frac{e^{2\varrho}}{2} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) = N_{g_{\alpha\beta}}^\dagger - 8\pi T_{\alpha\beta}.$$

By computing the divergence of this equation (for the metric g^\dagger) and evaluating the left-hand side in (4.14), we get

$$(5.21) \quad \begin{aligned} & \frac{1}{2} \nabla^{\dagger\alpha} \left(e^{2\varrho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \right) \\ &= \frac{1}{2} e^{2\varrho} \left(g^{\dagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta - g^{\dagger\alpha\alpha'} \Gamma^{\dagger\gamma}_{\alpha\alpha'} \partial_\gamma \Gamma^\dagger_\beta - g^{\dagger\alpha\alpha'} \Gamma^{\dagger\delta}_{\alpha'\beta} \partial_\delta \Gamma^\dagger_\alpha - \partial_\beta g^{\dagger\alpha'\beta'} (\partial_{\alpha'} \Gamma^\dagger_{\beta'}) \right) \\ & \quad + e^{2\varrho} \left(e^{2\varrho} (\partial_\alpha \Gamma^\dagger_\beta + \partial_\beta \Gamma^\dagger_\alpha - g_{\alpha\beta} g^{\alpha'\beta'} \partial_{\alpha'} \Gamma^\dagger_{\beta'}) \right) \nabla^{\dagger\alpha} \varrho \\ &=: \frac{1}{2} e^{2\varrho} g^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \Gamma^\dagger_\beta + \tilde{F}_\beta(\varrho, g^\dagger; \Gamma^\dagger_\gamma, \partial \Gamma^\dagger_\gamma). \end{aligned}$$

Here, \tilde{F} is a combination of linear and bilinear forms on Γ^\dagger_γ and $\partial \Gamma^\dagger_\gamma$ depending on ϱ, g^\dagger and their derivatives.

On the other hand, the right-hand side is computed from (5.9b) and (5.14), as follows:

$$\begin{aligned} & \nabla^{\dagger\alpha} (N_{g_{\alpha\beta}}^\dagger - 8\pi T_{\alpha\beta}) \\ &= e^{-2\varrho} (2g^{\alpha\alpha'} \partial_{\alpha'} \varrho N_{g_{\alpha\beta}}^\dagger - \text{tr} (8\pi T) \partial_\beta \varrho) + 6e^{2\varrho} \partial_\beta \varrho \Gamma^{\dagger\delta} \partial_\delta \varrho \\ & \quad - 8\pi e^{-2\varrho} (2g^{\delta\gamma} \partial_\gamma \varrho T_{\gamma\beta} - \text{tr} (T) \partial_\beta \varrho) \\ &= e^{-2\varrho} \left(2g^{\gamma\alpha} \partial_\gamma \varrho (N_{g_{\alpha\beta}}^\dagger - 8\pi T_{\alpha\beta}) - \text{tr} (N_g^\dagger - 8\pi T) \partial_\beta \varrho \right) + 6\partial_\beta \varrho \Gamma^{\dagger\delta} \delta \partial_\delta \varrho. \end{aligned}$$

Then, in view of (5.19) and (5.20), we obtain

$$(5.22) \quad \nabla^{\dagger\alpha} (N_{g_{\alpha\beta}}^{\dagger} - 8\pi T_{\alpha\beta}) = g^{\gamma\alpha} \partial_{\gamma} \varrho (\partial_{\alpha} \Gamma_{\beta}^{\dagger} + \partial_{\beta} \Gamma_{\alpha}^{\dagger}) + 6\partial_{\beta} \rho \Gamma^{\dagger\delta} \partial_{\delta} \rho,$$

which is a linear form in the functions $\Gamma_{\alpha}^{\dagger}$ and $\partial_{\gamma} \Gamma_{\alpha}^{\dagger}$. Finally, we arrive at the desired conclusion by combining (5.21) and (5.22) with (5.20) together. \square

Lemma 5.11. — *Provided the equations (5.10) hold on the initial slice $\{t = 0\}$ and the condition*

$$\Gamma_{\lambda}^{\dagger} = 0$$

and the constraint equations (5.7) and (5.8) also hold on the initial slice $\{t = 0\}$, then it follows that, on this initial slice,

$$\partial_t \Gamma_{\lambda}^{\dagger} = 0.$$

Proof. — We observe that the constraint equations (5.7) and (5.8) are equivalent to $N_{g_{0j}}^{\dagger} = 8\pi T_{00}$ and $N_{g_{00}}^{\dagger} = 8\pi T_{00}$, respectively. Consequently, the proof of the lemma follows from the same calculation which was performed in the proof of Lemma 4.7. Therefore, we omit the details. \square

Proof of Proposition 5.7. — By Lemmas 5.10 and 5.11 and by the global hyperbolicity of the metric g^{\dagger} , we see that, in the spacetime M ,

$$(5.23) \quad \Gamma_{\beta}^{\dagger} = 0.$$

Recalling Lemma 5.9 and combining (5.16) with (5.23), we obtain

$$(5.24) \quad \square_{g^{\dagger}} \varrho = \frac{R_g}{6e^{2\varrho}} + \frac{W_2(\varrho)}{3e^{2\varrho}} + \frac{4\pi \operatorname{tr}(T)}{3e^{4\varrho}}.$$

In a similar way, (5.9c) becomes

$$(5.25) \quad \square_{g^{\dagger}} \varrho = \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{f(\theta(\varrho))}{6e^{4\varrho}} + \frac{4\pi \operatorname{tr}(T)}{3e^{4\varrho}}.$$

By comparing these two equations, we thus get

$$R_g + W_2(\varrho) - f(\theta(\varrho)) = 0,$$

which (by the definition of W_2) leads us to $\theta(\varrho) = R_g$, and it remains to recall the notation $e^{2\varrho} = R_g$. \square

5.5. The local existence theory

The standard theory of local-in-time existence for the initial value problem associated with hyperbolic problems can now be applied to an arbitrary initial set in these sense of Definition 3.7. Let us sketch the strategy of proof. For simplicity in this discussion and without genuine restriction, we can consider that the initial data set and, therefore, the solutions are close to data in Minkowski space.

First of all, we need to construct the (local-in-time) solution of the problem (6.1) whose initial data set must be expressed in wave coordinates, say $(\overline{M}, g_{0\alpha\beta}^{\dagger}, g_{1\alpha\beta}^{\dagger}, \varrho_0, \varrho_1, \phi_0, \phi_1)$.

This PDE initial data set is determined from the geometric initial data set denoted by $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$. Without restriction, smallness (and regularity) assumptions are here made on the initial data set.

Second, we need to check that this (local-in-time) solution leads to a globally hyperbolic spacetime. Then, according to Proposition 5.7, we can conclude that this solution preserves the constraints (5.12) and (5.13) and, consequently, is also a solution to the field equations (3.7) with $g^\ddagger = g^\dagger$ and $\varrho = \rho = \frac{1}{2} \ln f'(R_g)$. We thus conclude that this solution solves the initial value problem in Definition 3.7.

Third, we need to observe that the solution (g, ϕ) constructed from $g_{\alpha\beta} := e^{-2\rho} g^\dagger_{\alpha\beta}$ is a solution to the original initial value problem, stated in Definition 2.6 with the corresponding initial data set determined by $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$ (via the conformal transformation).

We omit the details and refer to Choquet-Bruhat [10] for the existence and uniqueness statements in classical gravity, which based on our reformulation and discussion above can be exactly restated for the field equations of modified gravity. For the rest of this work, our objective is to revisit such a theory and, while re-proving this existence result, to establish that *modified gravity developments* remain close to *classical Einstein developments*.

6. Local existence theory. Formulation and main statement

6.1. Construction of the PDE initial data set

Our objective is thus to establish an existence theory for the Cauchy problem associated with the modified gravity field equations (5.10), when the initial data are assumed to be asymptotic flat. For the sake of simplicity and without genuine loss of generality as far as our method of proof is concerned, we focus our presentation on quadratic functions $f(r) = r + \frac{\kappa}{2}r^2$. It is straightforward to modify our argument to cover more general functions f .

We need first to introduce several notations, before we can state one of our main results in Theorem 6.3 below. Recall that the matter model we are considering is the massless scalar field with Jordan coupling and that, in agreement with Section 3.3 (see (3.19)), the system (5.10) under consideration reads

$$(6.1a) \quad \begin{aligned} & g^{\ddagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} g^\ddagger_{\alpha\beta} \\ &= F_{\alpha\beta}(g^\ddagger; \partial g^\ddagger, \partial g^\ddagger) - 12 \partial_\alpha \varrho \partial_\beta \varrho + \frac{W_2(\varrho)}{e^{2\varrho}} g^\ddagger_{\alpha\beta} - 16\pi \partial_\alpha \phi \partial_\beta \phi, \end{aligned}$$

$$(6.1b) \quad g^{\ddagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \phi = 2g^{\ddagger\alpha'\beta'} \partial_{\alpha'} \phi \partial_{\beta'} \varrho,$$

$$(6.1c) \quad g^{\ddagger\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} \varrho = \frac{W_2(\varrho)}{6e^{2\varrho}} + \frac{f(\theta(\varrho))}{6e^{4\varrho}} - \frac{4\pi}{3e^{2\varrho}} g^{\ddagger\alpha'\beta'} \partial_{\alpha'} \phi \partial_{\beta'} \phi.$$

Clearly, this is quasi-linear system of wave equations in diagonalized form and, in order to formulate a well-posed problem, the initial data set should include the functions

$$(6.2) \quad \begin{aligned} g^\dagger_{\alpha\beta}(0, x) &= g^\dagger_{0\alpha\beta}, & \partial_t g^\dagger_{\alpha\beta}(0, x) &= g^\dagger_{1\alpha\beta}, \\ \varrho(0, x) &= \varrho_0, & \partial_t \varrho(0, x) &= \varrho_1, \\ \phi(0, x) &= \phi_0 = \phi_0^\dagger, & \partial_t \phi(0, x) &= \phi_1 = \phi_1^\dagger. \end{aligned}$$

There are 24 functions to be prescribed, but the geometrical initial data set

$$(\overline{M}, \overline{g}^\dagger, K^\dagger, \varrho_0, \varrho_1, \phi_0^\dagger, \phi_1^\dagger)$$

contains 16 functions only. In fact, in order to construct a solution of (6.1) that also satisfies (3.7), we see that, by Proposition 5.7, the conditions (5.12), (5.7) and (5.8) must hold on the initial hypersurface. These conditions form a nonlinear PDE's system of eight equations, and it is expected that the 8 remaining initial data components could be determined from these 8 constraint equations. This task, however, is not a trivial one and further investigation would be needed to fully clarify the set of initial data.

From now on, we assume that this system of 8 constraint equations together with the 16 functions prescribed by the geometric initial data set uniquely determine our PDE's initial data set. Throughout, we denote by $(\overline{M}, g^\dagger_{0\alpha\beta}, g^\dagger_{1\alpha\beta}, \varrho_0, \varrho_1, \phi_0, \phi_1)$ the PDE initial data determined by $(\overline{M}, \overline{g}^\dagger, K^\dagger, \varrho_0, \varrho_1, \phi_0^\dagger, \phi_1^\dagger)$ and the constraint equations (5.7), (5.8) and (5.12).

6.2. Simplifying the field equations of $f(R)$ gravity

For definiteness, we focus on the role of second-order terms in f and assume that

$$(6.3) \quad f(r) := r + \frac{\kappa}{2}r^2, \quad r \in \mathbb{R}$$

for some $\kappa \geq 0$. We recall $e^{2\rho} = f'(R_g)$, so that

$$e^{2\rho} = 1 + \kappa R_g, \quad R_g = \frac{e^{2\rho} - 1}{\kappa},$$

and

$$\begin{aligned} W_2(s) &= \frac{f(r) - r f'(r)}{f'(r)} \\ &= \frac{r + \frac{\kappa}{2}r^2 - r(1 + \kappa r)}{1 + \kappa r} = -\frac{\kappa r^2}{2(1 + \kappa r)}, \end{aligned}$$

with $e^{2s} = f'(r) = 1 + \kappa r$ so that

$$(6.4) \quad W_2(s) = -\frac{(e^{2s} - 1)^2}{2\kappa e^{2s}}.$$

The spacetime (M, g^\dagger) under consideration is endowed with a foliation

$$M = [0, T] \times M_t, \quad t \in [0, T]$$

and we assume that, for each t , M_t is diffeomorphic to \mathbb{R}^3 . The spacetime metric is supposed to be sufficiently close to Minkowski metric and, especially, is asymptotically flat, so that the following notation is convenient:

$$(6.5) \quad h_{\alpha\beta} := g^\dagger_{\alpha\beta} - m_{\alpha\beta}$$

and we thus seek for unknowns triples $(h_{\alpha\beta}, \varrho, \phi)$. Sometimes, we will write $(h_{\alpha\beta}, \varrho, \phi) = (h_{\alpha\beta}^\kappa, \varrho^\kappa, \phi^\kappa)$ in order to emphasize that the solutions of (5.10) depend on the coefficient κ .

With these notation the system (5.10) take in the alternative form:

$$(6.6a) \quad \begin{aligned} & (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^\kappa \\ & = F_{\alpha\beta}(h^\kappa; \partial h^\kappa, \partial h^\kappa) - 16\pi \partial_\alpha \phi^\kappa \partial_\beta \phi^\kappa - 12 \partial_\alpha \varrho^\kappa \partial_\beta \varrho^\kappa \\ & \quad - \frac{(e^{2\varrho^\kappa} - 1)^2}{2\kappa e^{4\varrho^\kappa}} (m_{\alpha\beta} + h_{\alpha\beta}^\kappa), \end{aligned}$$

$$(6.6b) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \partial_{\beta'} \phi^\kappa = 2(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \phi^\kappa \partial_{\beta'} \varrho^\kappa,$$

$$(6.6c) \quad \begin{aligned} & (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \partial_{\beta'} \varrho^\kappa - \frac{e^{2\varrho^\kappa} - 1}{6\kappa e^{4\varrho^\kappa}} \\ & = -\frac{4\pi}{3e^{2\varrho^\kappa}} (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \phi^\kappa \partial_{\beta'} \phi^\kappa, \end{aligned}$$

where, from $h^\kappa = (h_{\alpha\beta}^\kappa)$, we have determined

$$(6.7) \quad (m^{\alpha\beta} + H^{\alpha\beta}(h^\kappa)) \text{ as the inverse of } (m_{\alpha\beta} + h_{\alpha\beta}^\kappa).$$

With this notation, the PDE initial data denoted by $(\overline{M}, h_0, h_1, \varrho_0, \varrho_1, \phi_0^\dagger, \phi_1^\dagger)$ is thus rewritten in terms of h , with

$$h_{\alpha\beta}|_{t=0} = h_{0\alpha\beta} := g^\dagger_{0\alpha\beta} - m_{\alpha\beta}, \quad \partial_t h_{\alpha\beta} = h_{1\alpha\beta} := g^\dagger_{1\alpha\beta}.$$

The system under consideration is composed of 12 quasi-linear wave equations: 11 of them (those on $h_{\alpha\beta}^\kappa$ and ϕ^κ) are quasi-linear wave equations, while the equation on ϱ^κ is a quasi-linear Klein-Gordon equation with defocusing potential.

Remark 6.1. — *The coefficients $H^{\alpha\beta}$ are clearly smooth functions of h , in a sufficiently small neighborhood of the origin, at least. Hence, we can find a positive constant ϵ_0 such that if $|h| \leq \epsilon_0$, then for any integer k , the k -th order derivatives of $H^{\alpha\beta}$ with respect to h , say $D^k H^{\alpha\beta}$, are well-defined and*

$$\sup_{|h| \leq \epsilon'} |D^k H^{\alpha\beta}(h)| \leq C(k, \epsilon_0)$$

Standard linear algebra arguments show that when $|h| \leq \epsilon_0$ (with ϵ_0 sufficiently small), then

$$(6.8) \quad H^{\alpha\beta}(h) = -h_{\alpha\beta} + Q^{\alpha\beta}(h, h)(1 + R^{\alpha\beta}(h)),$$

where $Q^{\alpha\beta}$ is a quadratic form in its argument and $|R^{\alpha\beta}(h)| \leq C(\epsilon_0) |h|$.

6.3. Vector fields and notation

We introduce the three generators of the spatial rotations

$$(6.9) \quad \Omega_{ij} := x_i \partial_j - x_j \partial_i = x^i \partial_j - x^j \partial_i,$$

which are known to commute with the wave operator, as well as with the Klein-Gordon operator. Here, the coordinate indices are raised and lowered with the Minkowski metric. We also write

$$\Omega_1 := \Omega_{12}, \quad \Omega_2 := \Omega_{23}, \quad \Omega_3 := \Omega_{13}.$$

Note that

$$(6.10) \quad [\Omega_a, \square] = 0, \quad [\Omega_a, \square + 1] = 0.$$

The following notation about multi-indices will be used. Given a finite set $\mathcal{I} = \{\alpha_1, \alpha_2, \dots, \alpha_n\}$, we call n the order of \mathcal{I} , denoted by $|\mathcal{I}| = n$. We introduce an ordering relation denoted by \preceq on \mathcal{I} , defined by

$$\alpha_i \preceq \alpha_j \quad \text{if and only if} \quad i \leq j.$$

The pair (\mathcal{I}, \preceq) is called an abstract multi-index. Obviously, a subset of \mathcal{I} can also be regarded as a multi-index endowed with the same (restricted) order. The order \preceq describes the location of each differential operator in a product.

A partition of an abstract index \mathcal{I} is defined as follows. Let \mathcal{J}_k be family of subsets of \mathcal{I} , with

$$\bigcup_{k=1}^m \mathcal{J}_k = \mathcal{I}, \quad \mathcal{J}_k \cap \mathcal{J}_{k'} = \emptyset.$$

Then, we say that $\{\mathcal{J}_k\}$ is an m -partition of \mathcal{I} and we write $\sum_{k=1}^m \mathcal{J}_k = \mathcal{I}$. We observe that each \mathcal{J}_k can be regarded as a multi-index and $\sum_{k=1}^m |\mathcal{J}_k| = |\mathcal{I}|$.

If for all $k = 1, 2, \dots, m$, we have $\mathcal{J}_k \neq \emptyset$, then we say $\{\mathcal{J}_k\}$ is a proper m -partition of \mathcal{I} and we write $\sum_{k=1}^m \mathcal{J}_k = \mathcal{I}, |\mathcal{J}_k| \geq 1$.

Now, let us return to the case of multi-indices in the context of differential operators. Let Z be a family of order one differential operator, say $Z = \{Z_1, Z_2, \dots, Z_p\}$. A n -multi-index on the family Z is a map

$$I : \mathcal{I} \rightarrow \{1, 2, \dots, p\}, \\ \alpha_i \mapsto I(\alpha_i) \in \{1, 2, \dots, p\},$$

and we write $Z^I := Z_{I(\alpha_1)} \circ Z_{I(\alpha_2)} \circ \dots \circ Z_{I(\alpha_n)}$. With some abuse of notation, we often write $I = (\alpha_n, \alpha_{n-1}, \dots, \alpha_1)$ with $\alpha_i \in \{1, 2, \dots, p\}$, where each α_i is replaced by $I(\alpha_i)$.

An m -partition of index I is defined as follows. Let $\{\mathcal{J}_k\}$ be a (proper) m -partition of an abstract index \mathcal{I} . Then we restrict the map I on each ordered set \mathcal{J}_k , and this yields us an m -multi-index, denoted by $\{J_k\} = I(\{\mathcal{J}_k\})$. Then, we call J_k a (proper) m -partition of I and we denote it by $I = \sum_{k=1}^m J_k, (|J_k| \geq 1)$.

We often consider the set $\mathcal{P}(\mathcal{I}, m)$ composed by all possible m -partitions of \mathcal{I} . Then, each partition in $\mathcal{P}(\mathcal{I}, m)$ can be associated with a partition of I . We observe that if

$\prod_{k=1}^m u_k$ is a product of m functions, then

$$Z^I \left(\prod_{k=1}^m u_k \right) = \sum_{\{\mathcal{J}_k\} \in \mathcal{P}(\mathcal{J}, m)} \prod_{i=1}^m Z^{J_k} u_k$$

with $\{J_k\} = I(\{\mathcal{J}_k\})$. However, for the sake of simplicity in the notation and whenever there is no risk of confusion, we will often write

$$Z^I \left(\prod_{k=1}^m h_k \right) = \sum_{\sum_{k=1}^m J_k = I} \prod_{k=1}^m Z^{J_k} u_k.$$

6.4. Functional spaces of interest

Recall that, in classical gravity, an initial data set for the Cauchy problem must satisfy the constraint equations (2.29). As first proven by Lichnerowicz, solutions can be constructed by solving a nonlinear elliptic system and, by the positive mass theorem, the non-trivial part of the metric $\bar{g} - \bar{m}$ decreases precisely at the rate r^{-1} at spatial infinity. Indeed, if $\bar{g} - \bar{m}$ would decrease faster than r^{-1} , then $\bar{g} = \bar{m}$.

In modified gravity, the constraint equations (5.7) and (5.8) are much more involved, an analogue result related to the positive mass theorem is not known. Yet, since we focus our attention on the (local) existence and ($\kappa \rightarrow 0$) convergence theory, we can consider an ‘‘Einsteinian initial data set’’ having $\varrho_0 = \varrho_1 \equiv 0$ satisfying the classical constraint equations. We need to handle the quasi-linear wave system with initial data decreasing like r^{-1} at spatial infinity. These functions (in general) do not lie in $L^2(\mathbb{R}^3)$, and we need to construct our local solution in suitably weighted functional spaces.

We need to introduce some norms about $C_c^\infty(\mathbb{R}^3)$ functions, that is, smooth functions with compact support. A first norm to be introduced is

$$\|\varrho\|_{X^d} := \sum_{|I_1|+|I_2|\leq d} \|\partial_x^{I_1} \Omega^{I_2} \varrho\|_{L^2(\mathbb{R}^3)}.$$

The L^2 norm is with respect to the standard volume form, i.e. the Lebesgue measure.

The second norm defined for the $C_c^\infty(\mathbb{R}^3)$ functions is

$$\|\varrho\|_{X_P^{d+1}} := \sum_{|I_1|+|I_2|\leq d} \|\partial_x \partial_x^{I_1} \Omega^{I_2} \varrho\|_{L^2(\mathbb{R}^3)},$$

where $\partial_x \phi$ refers to the spatial gradient of ϕ .

The first functional space to be used in our analysis, the space X_P^{d+1} , is defined as the completion of the norm $\|\cdot\|_{X_P^{d+1}}$ on the C_c^∞ functions. We denote by

$$(6.11) \quad \|\cdot\|_{X_R^{d+1}} := \|\cdot\|_{X^d} + \|\cdot\|_{X_P^{d+1}}.$$

The second functional space, X_R^{d+1} , to be used in our analysis, is defined as the completion of the norm $\|\cdot\|_{X_R^{d+1}}$ on the C_c^∞ functions.

We also define the weighted sup-norm

$$\|f\|_{\mathcal{E}_{-1}} := \sup_{r \geq 0} \{(1+r)|f|\}$$

and

$$\|f\|_{X_H^{d+1}} := \|f\|_{\mathcal{E}_{-1}} + \|f\|_{X_P^{d+1}},$$

so that the functional spaces \mathcal{E}_{-1} and X_H^{d+1} are obtained by completion from the $C_c(\mathbb{R}^3)$ functions with respect of the norm under consideration.

The relations among these functional spaces are as follows:

$$(6.12) \quad X^{d+1} \subset X_R^{d+1} \subset X^d, \quad X_R^d \subset X_P^d, \quad X_H^d \subset X_P^d.$$

In the next section, when $d \geq 2$, by (7.8), we will also see that

$$(6.13) \quad X^d \subset \mathcal{E}_{-1}, \quad X_R^d \subset X_H^d \subset X_P^d.$$

Finally, we define the norm of a triple $S_0 := (h_{0\alpha\beta}, \varrho_0, \phi_0)$:

$$\|S\|_{X_0^{d+1}} := \sum_{\alpha, \beta} \|h_{0\alpha\beta}\|_{X_H^{d+1}} + \|\varrho_0\|_{X_R^{d+1}} + \|\phi_0\|_{X_P^{d+1}}$$

and we set

$$(6.14) \quad X_0^{d+1} = X_H^{d+1} \times X_R^{d+1} \times X_P^{d+1}.$$

Similarly, for triples $S_1 = (h_{1\alpha\beta}, \phi_1, \varrho_1)$, we define

$$\|S_1\|_{X_1^d} := \sum_{\alpha, \beta} \|h_{1\alpha\beta}\|_{X^d} + \|\varrho_1\|_{X^d} + \|\phi_1\|_{X^d}$$

We are now ready to discuss the notion of asymptotically flat PDE initial data.

Definition 6.2. — *A PDE initial data set in wave coordinates for the initial value problem stated in Definition 3.7, say $(\bar{M}, h_{0\alpha\beta}, h_{1\alpha\beta}, \varrho_0, \varrho_1, \phi_0, \phi_1)$, is said to be **asymptotically flat** if*

- *the initial slice M_0 is diffeomorphic to \mathbb{R}^3 and in its canonical coordinate system, the initial data set satisfies the wave constraint equations.*
- *in the canonical coordinate system (x^1, x^2, x^3) ,*

$$\|h_{0\alpha\beta}\|_{\mathcal{E}_{-1}} \leq \epsilon_0,$$

where ϵ_0 represents the ADM mass.

- *the $\mathcal{E}_{-1}(\mathbb{R}^3)$ norm of $\varrho_0, \varrho_1, \phi_1$ are finite.*

Hence, the initial data behaves like r^{-1} at spatial infinity. A geometrical initial data $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$ is called asymptotically flat, if it gives a asymptotically flat PDE initial data.

We recall that the components of the solution to the system (6.1) are functions defined in \mathbb{R}^4 with three spatial variables and one time variable. To study these functions, say $u = u(t, \cdot)$, we need to the following norms and corresponding spaces:

$$\begin{aligned} \|u(t, \cdot)\|_{E^d} &:= \sum_{|I_1|+|I_2|\leq d} \|\partial^{I_1}\Omega^{I_2}u(t, \cdot)\|_{L^2(\mathbb{R}^3)} \\ &= \sum_{k+|J_1|+|J_2|\leq d} \|\partial_t^k \partial_x^{J_1} \Omega^{J_2} u(t, \cdot)\| = \sum_{k+l\leq d} \|\partial_t^k u(t, \cdot)\|_{X^l}, \\ \|u(t, \cdot)\|_{E_P^{d+1}} &:= \sum_{\substack{\alpha \\ |I_1|+|I_2|\leq d}} \|\partial_\alpha \partial^{I_1} \Omega^{I_2} u(t, \cdot)\|_{L^2(\mathbb{R}^3)} \\ &= \sum_{\alpha} \sum_{k+|J_1|+|J_2|\leq d} \|\partial_\alpha \partial_t^k \partial_x^{J_1} \Omega^{J_2} u(t, \cdot)\|_{L^2(\mathbb{R}^3)} = \sum_{k+l\leq d} \|\partial_t^k u(t, \cdot)\|_{X_P^l}, \end{aligned}$$

and then

$$\begin{aligned} \|u(t, \cdot)\|_{E_R^{d+1}} &:= \|u(t, \cdot)\|_{E^d} + \|u(t, \cdot)\|_{E_P^{d+1}}, \\ \|u(t, \cdot)\|_{E_H^{d+1}} &:= \|u(t, \cdot)\|_{\varepsilon_{-1}} + \|u(t, \cdot)\|_{E_P^{d+1}}. \end{aligned}$$

We also define several norms on the time interval $[0, T]$ for $p \in [1, \infty]$:

$$\begin{aligned} \|u\|_{L^p([0, T]; E^d)} &:= \left\| \|u(t, \cdot)\|_{E^d} \right\|_{L^p([0, T])}, \quad \|u\|_{L^p([0, T]; E_P^d)} := \left\| \|u(t, \cdot)\|_{E_P^d} \right\|_{L^p([0, T])}, \\ \|u\|_{L^p([0, T]; E_R^d)} &:= \left\| \|u(t, \cdot)\|_{E_R^d} \right\|_{L^p([0, T])}, \quad \|u\|_{L^p([0, T]; E_H^d)} := \left\| \|u(t, \cdot)\|_{E_H^d} \right\|_{L^p([0, T])}. \end{aligned}$$

Finally, for $p < \infty$, we have the functional spaces of interest

$$L^p([0, T]; E^d), \quad L^p([0, T]; E_P^d), \quad L^p([0, T]; E_R^d),$$

and $L^p([0, T]; E_H^d)$, defined by completion with respect to the corresponding norms in the space of $C_c^\infty(\mathbb{R}^4)$ functions. This leads us to the definition of $C([0, T]; E^d)$, $C([0, T]; E_P^d)$ and $C([0, T]; E_H^d)$ by completion from the space C_c^∞ functions with respect to the following norms:

$$L^\infty([0, T]; E^d), \quad L^\infty([0, T]; E_P^d), \quad L^\infty([0, T]; E_R^d),$$

6.5. Existence theorem for the nonlinear field equations

We introduce the following norm for the initial data $S_0 := (h_0, h_1, \phi_0, \phi_1, \varrho_0, \varrho_1)$:

$$\begin{aligned} \|S_0\|_{X_\kappa^{d+1}} &:= \max\{\|h_{0\alpha\beta}\|_{X_H^{d+1}}, \|h_{1\alpha\beta}\|_{X^d}, \kappa^{-(1/2)[d/2]+1/4}\|\phi_0\|_{X_P^{d+1}}, \\ &\quad \kappa^{-(1/2)[d/2]+1/4}\|\phi_1\|_{X^d}, \kappa^{-[d/2]-1/2}\|\varrho_0\|_{X_R^{d+1}}, \kappa^{-[d/2]-1/2}\|\varrho_1\|_{X^d}\}. \end{aligned}$$

We are ready to state one of our main results.

Theorem 6.3 (Local existence with uniform bounds). — *Given any integer $d \geq 4$, assume that*

$$(h_{0\alpha\beta}, h_{1\alpha\beta}) \in X_H^{d+1} \times X^d, \quad (\phi_0, \phi_1) \in X_P^{d+1} \times X^d, \quad (\varrho_0, \varrho_1) \in X_R^{d+1} \times X^d$$

and suppose that $S_0 := (h_0, h_1, \phi_0, \phi_1, \varrho_0, \varrho_1)$ satisfies

$$(6.15) \quad \|S_0\|_{X_\kappa^{d+1}} \leq \epsilon \leq \epsilon_0$$

for some sufficiently small $\epsilon_0 > 0$. Then, there exist constants $A, T^* > 0$ which are **independent** of κ and such that, within the time interval $[0, T^*]$, the Cauchy problem (6.6) (with $0 < \kappa \leq 1$) has a unique solution $(h_{\alpha\beta}^\kappa, \phi^\kappa, \varrho^\kappa)$ in the following functional space (with $0 \leq k \leq d-1$):

$$\begin{aligned} \partial_t^k h_{\alpha\beta}^\kappa &\in C([0, T], X_H^{d-k}) \cap C^1([0, T], X_H^{d-k-1}), \\ \partial_t^k \phi^\kappa &\in C([0, T], X_P^{d-k}) \cap C^1([0, T], X_P^{d-k-1}), \\ \partial_t^k \varrho^\kappa &\in C([0, T], X_R^{d-k}) \cap C^1([0, T], X_R^{d-k-1}), \end{aligned}$$

Furthermore, the following estimates hold with constant **independent** of κ :

$$(6.16a) \quad \|h_{\alpha\beta}^\kappa(t, \cdot)\|_{E_H^d} \leq A\epsilon,$$

$$(6.16b) \quad \|\phi^\kappa\|_{E_P^d} \leq A\epsilon,$$

$$(6.16c) \quad \|\varrho_n^\kappa\|_{E_P^d} + \kappa^{-1/2} \|\varrho_n^\kappa\|_{E^{d-1}} \leq A\epsilon.$$

Equipped with the above theorem, we are thus able to build the local solution of the original Cauchy problem stated in Definition 5.4.

Theorem 6.4 (Existence of modified gravity developments)

Consider an initial data set $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$ for the Cauchy problem in Definition 5.3 and assume that its associated PDE initial data $S_0 = (h_0, h_1, \phi_0, \phi_1, \varrho_0, \varrho_1)$ is asymptotically flat and satisfies the conditions in Theorem 6.3. Let $(h_{\alpha\beta}^\kappa, \phi^\kappa, \varrho^\kappa)$ be the corresponding solution of (5.10) associated with S_0 . Then the spacetime $([0, T] \times \mathbb{R}^3, g^\ddagger)$ is a modified gravity development of the initial data $(\bar{M}, \bar{g}^\ddagger, K^\ddagger, \varrho_0, \varrho_1, \phi_0^\ddagger, \phi_1^\ddagger)$.

Proof. — We simply note that the local solution $(h_{\alpha\beta}^\kappa, \phi^\kappa, \varrho^\kappa)$ is sufficiently regular and that $h = g^\ddagger - m$ sufficiently small, which guarantees that the metric g^\ddagger is globally hyperbolic on $[0, T] \times \mathbb{R}^3$. We can apply the result about the preservation of constraints in Proposition 5.7. Once the constraints

$$e^{2\varrho} = \frac{1}{2} \ln f'(R_g), \quad \Gamma^\ddagger{}^\lambda = 0$$

hold, we see that the pair (g^\ddagger, ϕ) satisfies the conformal field equations (3.7). \square

7. Technical tools for the local existence theory

7.1. Estimates on commutators

From this subsection we will make some preparations in order to prove Theorem 6.3. In this subsection we derive commutator estimates. First we point out the following commutation

relations:

$$(7.1) \quad [\partial_a, \Omega_{bc}] = \delta_{ab}\partial_c - \delta_{ac}\partial_b, \quad [\partial_t, \Omega_{ab}] = 0.$$

Lemma 7.1. — *If u is a smooth function defined on $[0, T] \times \mathbb{R}^3$, then for any multi-index I_1, I_2 with $|I_1| + |I_2| = n$, the following estimates hold:*

$$(7.2) \quad [\Omega^J, \partial_\alpha]u = \sum_{|J| < |I|} \Theta_{\alpha J}^{Ic} \partial_c \Omega^J u,$$

where $\Theta_{\alpha J}^{Ic}$ are constants. When $\alpha = 0$, one has $\Theta_{\alpha J}^{Ic} = 0$.

Proof. — First, we observe that when $\alpha = 0$, $\partial_0 = \partial_t$ commutes with Ω^J . When $\alpha > 0$, this is proven by induction on the order of $|J|$. When $|J| = 1$, the result is proven by (7.1). We denote by

$$[\Omega_a, \partial_\beta] = \theta_{a\beta}^c \partial_c.$$

Now assume that for $|J| \leq k$, (7.2) is valid. For $|J| = k$, we find

$$\Omega^J = \Omega_{a_1} \Omega^{J_1},$$

where $|J_1| = k - 1$. Then, we have

$$\begin{aligned} [\Omega^J, \partial_a]u &= \Omega_{a_1} \Omega^{J_1} \partial_a u - \partial_a \Omega_{a_1} \Omega^{J_1} u = \Omega_{a_1} \Omega^{J_1} \partial_a u - \Omega_{a_1} \partial_a \Omega^{J_1} u + \Omega_{a_1} \partial_a \Omega^{J_1} u - \partial_a \Omega_{a_1} \Omega^{J_1} u \\ &= \Omega_{a_1} ([\Omega^{J_1}, \partial_a]u) + [\Omega_{a_1}, \partial_a] \Omega^{J_1} u \\ &= \Omega_{a_1} \left(\sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} \partial_b \Omega^{K_1} u \right) + \theta_{a_1 a}^b \partial_b \Omega^{J_1} u \\ &= \sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} \Omega_{a_1} \partial_b \Omega^{K_1} u + \theta_{a_1 a}^b \partial_b \Omega^{J_1} u \\ &= \sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} \partial_b \Omega_{a_1} \Omega^{K_1} u + \theta_{a_1 a}^b \partial_b \Omega^{J_1} u + \sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} [\Omega_{a_1}, \partial_b] \Omega^{K_1} u \\ &= \sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} \partial_b \Omega_{a_1} \Omega^{K_1} u + \theta_{a_1 a}^b \partial_b \Omega^{J_1} u + \sum_{|K_1| < |J_1|} \Theta_{\alpha K_1}^{J_1 b} \theta_{a_1 b}^c \partial_c \Omega^{K_1} u. \end{aligned}$$

□

Lemma 7.2. — *Let u be a smooth function defined in $[0, T] \times \mathbb{R}^3$. Then the following estimates hold for $|I_1| + |I_2| = d$:*

$$(7.3a) \quad |\partial_x^{I_1} \Omega^{I_2} \partial_x u| \leq |\partial_x \partial_x^{I_1} \Omega^{I_2} u| + C(d) \sum_{|I'_2| < |I_2|} |\partial_x \partial_x^{I'_2} \Omega^{I_2} u|,$$

$$(7.3b) \quad |\partial_x \partial_x^{I_1} \Omega^{I_2} u| \leq |\partial_x^{I_1} \Omega^{I_2} \partial_x u| + C(d) \sum_{|I'_1| + |I'_2| < d} |\partial_x^{I'_1} \Omega^{I'_2} \partial_x u|,$$

$$(7.3c) \quad |[\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha \partial_\beta]u| \leq C(d) \sum_{\substack{a, \alpha \\ |J_2| < |I_2|}} |\partial_a \partial_\alpha \partial_x^{I_1} \Omega^{J_2} u|,$$

$$(7.3d) \quad |\partial_r^k \Omega^{I_1} u(t, x)| \leq C(k) \sum_{|J_1| \leq k} |\partial_x^{J_1} \Omega^{I_1} u(t, x)|, \quad \text{for } |x| \geq 1,$$

$$(7.3e) \quad \begin{aligned} |[\partial_x^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u| &\leq \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} |\partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta}| |\partial_\alpha \partial_\beta \partial_x^{J_1} \Omega^{J_2} u| \\ &+ C(d) \sum_{\substack{J_1+J'_1=I_1 \\ |J_2|+|J'_2|<|I_2| \\ \alpha, \beta, \alpha', a}} |\partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta}| |\partial_{\alpha'} \partial_a \partial_x^{J_1} \Omega^{J_2} u|, \end{aligned}$$

where $H^{\alpha\beta}$ are smooth functions.

Proof. — In view of (7.2), the following identity is immediate:

$$[\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha] = \sum_{|K| < |I_2|} \Theta_{\alpha K}^{I_2 b} \partial_b \partial_x^{I_1} \Omega^K u.$$

To derive (7.3a), we observe that

$$[\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha] u = \partial_x^{I_1} ([\Omega^{I_1}, \partial_\alpha] u) = \sum_{c, |J_2| \leq |I_2|} \partial_x^{I_1} (\Theta_{\alpha J_2}^{I_1 c} \partial_c \Omega^{J_2} u) = \sum_{c, |J_2| \leq |I_2|} \Theta_{\alpha J_2}^{I_1 c} \partial_c \partial_x^{I_1} \Omega^{J_2} u$$

and

$$(7.4) \quad \sum_{c, |J_2| \leq |I_2|} |\Theta_{\alpha J_2}^{I_1 c} \partial_c \partial_x^{I_1} \Omega^{J_2} u| \leq C(d) \sum_{c, |I'_2| < |I_2|} |\partial_c \partial^{I_1} \Omega^{I'_2} u|.$$

This establishes (7.3a).

The equation (7.3b) is derived by induction on d . Clearly, (7.3b) holds for $d = 0$. If it holds for $d \leq k$, let us prove the case $d = k + 1$:

$$|\partial_\alpha \partial^{I_1} \Omega^{I_2} u| \leq |\partial^{I_1} \Omega^{I_2} \partial_\alpha u| + |[\partial_\alpha, \partial^{I_1} \Omega^{I_2}] u|.$$

Then by (7.4),

$$|\partial_\alpha \partial^{I_1} \Omega^{I_2} u| \leq |\partial^{I_1} \Omega^{I_2} \partial_\alpha u| + C(d) \sum_{c, |I'_2| < |I_2|} |\partial_c \partial^{I_1} \Omega^{I'_2} u|.$$

Note that $|I'_2| \leq |I_2| - 1 \leq k$. We apply the induction assumption on the second term in the right-hand side and obtain

$$|\partial_c \partial^{I_1} \Omega^{I'_2} u| \leq |\partial^{I_1} \Omega^{I'_2} \partial_c u| + C(d-1) \sum_{c', |I''_2| > |I'_2|} |\partial^{I_1} \Omega^{I''_2} \partial_{c'} u|,$$

which proves (7.3b).

The proof of (7.3c) is a direct application of (7.2).

$$\begin{aligned}
(7.5) \quad [\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha \partial_\beta] u &= [\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha] \partial_\beta u + \partial_\alpha ([\partial_x^{I_1} \Omega^{I_2}, \partial_\beta] u) \\
&= \sum_{|J_2| < |I_2|} \Theta_{\alpha J_2}^{I_2 c} \partial_c \partial_x^{I_1} \Omega^{J_2} \partial_\beta u + \sum_{|J_2| < |I_2|} \Theta_{\beta J_2}^{I_2 c} \partial_\alpha \partial_c \partial_x^{I_1} \Omega^{J_2} u \\
&= \sum_{|J_2| < |I_2|} \Theta_{\alpha J_2}^{I_2 c} \partial_c \partial_x^{I_1} \partial_\beta \Omega^{J_2} u + \sum_{|J_2| < |I_2|} \Theta_{\alpha J_2}^{I_2 c} \partial_c \partial_x^{I_1} [\Omega^{J_2}, \partial_\beta] u \\
&+ \sum_{|J_2| < |I_2|} \Theta_{\beta J_2}^{I_2 c} \partial_\alpha \partial_c \partial_x^{I_1} \Omega^{J_2} u \\
&= \sum_{\substack{|J_2| < |I_2| \\ |K_2| < |I_2|}} \Theta_{\alpha J_2}^{I_2 c} \Theta_{\beta K_2}^{J_2 c'} \partial_c \partial_{c'} \partial_x^{I_1} \Omega^{K_2} u \\
&+ \sum_{|J_2| < |I_2|} \Theta_{\beta J_2}^{I_2 c} \partial_\alpha \partial_c \partial_x^{I_1} \Omega^{J_2} u + \sum_{|J_2| < |I_2|} \Theta_{\alpha J_2}^{I_2 c} \partial_c \partial_x^{I_1} \partial_\beta \Omega^{J_2} u
\end{aligned}$$

which leads to (7.3c).

The proof of (7.3d) needs the notion of homogeneous functions. A smooth function f defined in the pointed region $\mathbb{R}^3 \setminus \{0\}$ is said to be **homogeneous of degree i** if

$$f(rx) = r^i f(x), \text{ for any } r > 0 \text{ and } x \in \mathbb{R}^3 \setminus \{0\}.$$

It is well-known that the partial derivatives of a homogeneous function of degree i are also homogeneous and of degree $i - 1$.

We denote by $\omega^a := \frac{x^a}{r}$ and we note that they are homogeneous functions of degree 0. And recall the definition of radial derivative $\partial_r = \omega^a \partial_a$. We will prove that

$$(7.6) \quad \partial_r^k = \sum_{\substack{a, \\ |I| \leq k}} A_I^k \partial_x^I,$$

where A_I^k is a homogeneous function of degree $-k + |I|$. For $k = 1$, this is guaranteed by the expression of ∂_r . Assume that this holds for the integer less than or equal to k , we will prove the case of $k + 1$:

$$\begin{aligned}
\partial_r^{k+1} u &= \partial_r \partial_r^k u = \partial_r \left(\sum_{I,a} A_I^k \partial_x^I \right) \\
&= \sum_{I,a} \omega^b \partial_b (A_I^k \partial_x^I) = \sum_{I,a} \omega^b \partial_b A_I^k \partial_x^I + \sum_{I,a} \omega^b A_I^k \partial_b \partial_x^I
\end{aligned}$$

We observe that $\omega^b \partial_b A_I^k$ is homogeneous of degree $-k - 1 + |I|$ and $\omega^b A_I^k$ is homogeneous of degree $-(k+1) + |I| + 1$. This concludes (7.6). Next, we see that (7.3d) follows immediately from (7.6).

To prove (7.3e), we perform the following calculation:

$$\begin{aligned}
& [\partial_x^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta u] \\
&= \sum_{\substack{J_1+J'_1=I_1, J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_x^{J_1} \Omega^{J_2} \partial_\alpha \partial_\beta u + H^{\alpha\beta} [\partial_x^{I_1} \Omega^{I_2}, \partial_\alpha \partial_\beta] u \\
&= \sum_{\substack{J_1+J'_1=I_1, J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_\alpha \partial_\beta \partial_x^{J_1} \Omega^{J_2} u + \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2}} \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta} [\partial_x^{J_1} \Omega^{J_2}, \partial_\alpha \partial_\beta] u \\
&= \sum_{\substack{J_1+J'_1=I_1, J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_\alpha \partial_\beta \partial_x^{J_1} \Omega^{J_2} u \\
&+ \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2}} \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \left(\sum_{\substack{|K_2|<|J_2| \\ |K'_2|<|K_2|}} \Theta_{\alpha K_2}^{J_2 c} \Theta_{\beta K'_2}^{c' K'_2} \partial_c \partial_{c'} \partial_x^{J_1} \Omega^{K'_2} u \right. \\
&\quad \left. + \sum_{|K_2|<|J_2|} \Theta_{\beta K_2}^{J_2 c} \partial_\alpha \partial_c \partial_x^{J_1} \Omega^{K_2} u \right),
\end{aligned}$$

where (7.5) is applied. Then (7.3e) follows from this identity. \square

We also need the commutator estimates on the product in the form $\partial^{I_1} \Omega^{I_2}$.

Lemma 7.3. — *Let u be a smooth function defined in $[0, T] \times \mathbb{R}^3$. Then the following estimates hold for $|I_1| + |I_2| = d$:*

$$(7.7a) \quad |\partial^{I_1} \Omega^{I_2} \partial_\alpha u| \leq |\partial_\alpha \partial^{I_1} \Omega^{I_2} u| + C(d) \sum_{a, |I'_2| < |I_2|} |\partial_a \partial^{I_1} \Omega^{I'_2} u|,$$

$$(7.7b) \quad |\partial_\alpha \partial^{I_1} \Omega^{I_2} u| \leq |\partial^{I_1} \Omega^{I_2} \partial_\alpha u| + C(d) \sum_{\substack{\beta \\ |I'_1|+|I'_2|<d}} |\partial^{I'_1} \Omega^{I'_2} \partial_\beta u|,$$

$$(7.7c) \quad |[\partial^{I_1} \Omega^{I_2}, \partial_\alpha \partial_\beta] u| \leq C(d) \sum_{\substack{a, \alpha \\ |J_2| < |I_2|}} |\partial_a \partial_\alpha \partial^{I_1} \Omega^{J_2} u|,$$

$$\begin{aligned}
(7.7d) \quad |[\partial^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u| &\leq \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} |\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta}| |\partial_\alpha \partial_\beta \partial^{J_1} \Omega^{J_2} u| \\
&+ C(d) \sum_{\substack{J_1+J'_1=I_1 \\ |J_2|+|J'_2|<|I_2| \\ \alpha, \beta, \alpha', a}} |\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta}| |\partial_{\alpha'} \partial_a \partial^{J_1} \Omega^{J_2} u|.
\end{aligned}$$

Proof. — From (7.2), the following identity is immediate:

$$[\partial^{I_1} \Omega^{I_2}, \partial_\alpha] = \sum_{|K| < |I_2|} \Theta_{\alpha K}^{I_2 b} \partial_b \partial_x^{I_1} \Omega^K u.$$

Then we perform exactly the same calculation as in the proof of Lemma 7.2 with $\partial_x^{I_1}$ replaced by ∂^{I_1} . \square

7.2. Global Sobolev inequalities and embedding properties

For completeness, we re-derive a classical estimate (due to Klainerman).

Lemma 7.4. — *For all $u \in X^d$ with $d \geq 2$, one has*

$$(7.8) \quad \|u\|_{\varepsilon_{-1}} \leq C \|u\|_{X^d}.$$

Proof. — We only prove this inequality for smooth functions since, by regularization, it then extends to the whole X^d . We consider \mathbb{R}^3 equipped with the polar coordinates, i.e.:

$$x^1 = r \sin \theta \cos \varphi, \quad x^2 = r \sin \theta \sin \varphi, \quad x^3 = r \cos \theta.$$

Note that (7.8) is equivalent to the following inequality for all $x_0 \in \mathbb{R}^3$:

$$|u(x_0)|(1 + r_0) \leq C \|u\|_{X^d}$$

with $x_0 = (x_0^1, x_0^2, x_0^3) = (r_0 \sin \theta_0 \cos \varphi_0, r_0 \sin \theta_0 \sin \varphi_0, r_0 \cos \theta_0)$.

The case $r_0 \leq 1$ is direct by classical Sobolev inequality. We thus focus on the case $r_0 > 1$ and we consider the estimate on the the following open subset of \mathbb{R}^3 defined by

$$V = (r_0 - 1/2, r_0 + 1/2) \times (\theta_0 - 1/2, \theta_0 + 1/2) \times (\varphi_0 - 1/2, \varphi_0 + 1/2)$$

Now let u be a smooth function and denote also by u its restriction on V , with

$$v(r, \theta, \varphi) := u(r \sin \theta \cos \varphi, r \sin \theta \sin \varphi, r \cos \theta)$$

with $r_0 > 1$. Then by the classical Sobolev inequality, we have

$$|u(x_0)|^2 = |v(r_0, \theta_0, \varphi_0)|^2 \leq C \sum_{k_0+k_1+k_2 \geq 2} \int_V |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 dr d\theta d\varphi.$$

Note that in V , $r_0 - 1/2 < r < r_0 + 1/2$, which leads to $1 - 1/(2r_0) < r/r_0 < 1 + 1/(2r_0)$. Recall that $r_0 > 1$, then

$$\frac{1}{2} < \frac{r}{r_0} < \frac{3}{2}.$$

Thus, we have

$$\begin{aligned}
\frac{1}{2} \int_V |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 dr d\theta d\varphi &\leq \int_V |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 \frac{r^2}{r_0^2} dr d\theta d\varphi \\
&= r_0^{-2} \int_V |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 r^2 dr d\theta d\varphi \\
&= r_0^{-2} \int_V |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 dx \\
&\leq r_0^{-2} \int_{\mathbb{R}^3} |\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v|^2 dx,
\end{aligned}$$

where dx is the standard volume form of \mathbb{R}^3 .

Here, we observe that

$$\partial_\theta v = \cos \varphi \Omega_{31} u + \sin \varphi \Omega_{32} u = \frac{x^1}{((x^1)^2 + (x^2)^2)^{1/2}} \Omega_{31} u + \frac{x^2}{((x^1)^2 + (x^2)^2)^{1/2}} \Omega_{32} u$$

and

$$\partial_\theta v = \Omega_{12} u, \quad \partial_r v = \partial_r u.$$

Note that $\Omega^J \cos \varphi$, $\Omega^J \sin \varphi$ are homogeneous of degree 0. So, by homogeneity, that for $r \geq 1/2$, we have

$$|\partial_r^{k_0} \partial_\theta^{k_1} \partial_\varphi^{k_2} v| = |\partial_r^{k_0} (\cos \varphi \Omega_{31} + \sin \varphi \Omega_{33})^{k_1} \Omega_{12}^{k_2} u| \leq C \sum_{|I| \leq k_1 + k_2} |\partial_r^{k_0} \Omega^I u|,$$

which leads to

$$(7.9) \quad |u(x_0)|^2 \leq C r_0^{-2} \sum_{|I| + k_0 \leq 2} \int_{\mathbb{R}^3} |\partial_r^{k_0} \Omega^I u|^2 dx.$$

Then by (7.3d), the desired result is proven. \square

We will also need the following embedding result.

Lemma 7.5. — *Let u be a function in X_H^{d+2} and $v \in L^\infty([0, T]; E_H^{d+2})$. Then for all pair of multi-indices (I_1, I_2) with $|I_1| + |I_2| \leq d$, the following estimate holds:*

$$(7.10a) \quad \left\| \partial_x^{I_1} \Omega^{I_2} u \right\|_{L^\infty(\mathbb{R}^3)} \leq C \|u\|_{X_H^{d+2}}.$$

$$(7.10b) \quad \left\| \partial^{I_1} \Omega^{I_2} v(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \leq C \|v(t, \cdot)\|_{E_H^{d+2}} \quad \text{for } t \in [0, T].$$

Proof. — We only prove this inequality when $u \in C_c^\infty(\mathbb{R}^3)$ and $v \in C_c^\infty(\mathbb{R}^4)$. Then by regularization, it extends on X_H^{d+2} and $L^\infty([0, T]; E_H^{d+2})$.

We begin with (7.10a), and the proof is decomposed as follows.

Case 1: $I_1 = I_2 = 0$. The left-hand side of (7.10a) is controlled by its \mathcal{E}_{-1} norm so (7.10a) holds.

Case 2: $|I_1| > 0$. In this case we suppose that $I_1 = (a_1, a_2, \dots, a_n)$ and denote by $I'_1 = (a_2, \dots, a_n)$. Then I'_1 is of order $n - 1 \geq 0$ and, by classical Sobolev's inequality,

$$\|\partial_x^{I_1} \Omega^{I_2} u\|_{L^\infty(\mathbb{R}^3)} = \|\partial_{a_1} \partial_x^{I'_1} \Omega^{I_2} u\|_{L^\infty(\mathbb{R}^3)} \leq C \sum_{|J_1| \leq 2} \|\partial_x^{J_1} \partial_{a_1} \partial_x^{I'_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} \leq C \|u\|_{X_P^{d+2}}.$$

Case 3: $|I_2| = 0, |I_2| > 0$. In this case we suppose that $I_2 = (b_1, b_2, \dots, b_n)$ and denote by $I'_2 = (b_2, \dots, b_n)$. Then I'_2 is of order $n - 1 \geq 0$, and

$$\|\partial_x^{I_1} \Omega^{I_2} u\|_{L^\infty(\mathbb{R}^3)} = \|\Omega_{b_1} \Omega^{I'_2} u\|_{L^\infty(\mathbb{R}^3)} \leq C \sum_b \|(1+r) \partial_b \Omega^{I'_2} u\|_{L^\infty(\mathbb{R}^3)} = C \sum_b \|\partial_b \Omega^{I'_2} u\|_{\mathcal{E}_{-1}}.$$

Then by (7.8), we have

$$\begin{aligned} \|\partial_x^{I_1} \Omega^{I_2} u\|_{L^\infty(\mathbb{R}^3)} &\leq C \sum_b \|\partial_b \partial_x^{I_1} \Omega^{I_2} u\|_{X^2} = C \sum_{|J_1|+|J_2| \leq 2} \sum_b \|\partial_x^{J_1} \Omega^{J_2} \partial_b \partial_x^{I_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} \\ &\leq C \sum_{\substack{|J_1|+|J'_2| \leq 2 \\ b'}} \|\partial_{b'} \partial_x^{I'_1} \partial_x^{J_1} \Omega^{J'_2} \Omega^{I'_2} u\|_{L^2(\mathbb{R}^3)} \leq C \|u\|_{X_P^{d+2}}, \end{aligned}$$

where the commutator estimate (7.3a) is used.

By combining these three cases together, (7.10a) is proven.

We then prove (7.10b). The proof is similar and we also discuss three different cases.

Case 1: $I_1 = I_2 = 0$. The left-hand side of (7.10b) is controlled by its \mathcal{E}_{-1} norm so (7.10b) holds.

Case 2: $I_1 > 0$. In this case we also suppose that $I_1 = (\alpha_1, \alpha_2, \dots, \alpha_n)$ and denote by $I'_1 = (\alpha_2, \dots, \alpha_n)$. Then $|I'_1| \geq 0$, and also by the classical Sobolev's inequality,

$$\begin{aligned} \|\partial^{I_1} \Omega^{I_2} v(t, \cdot)\|_{L^\infty} &\leq C \sum_{|J_1| \leq 2} \|\partial_x^{J_1} \partial_{\alpha_1} \partial^{I'_1} \Omega^{I_2} v(t, \cdot)\|_{L^2(\mathbb{R}^3)} \\ &= C \sum_{|J_1| \leq 2} \|\partial_{\alpha_1} \partial_x^{J_1} \partial^{I'_1} \Omega^{I_2} v(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C \|v(t, \cdot)\|_{E_P^{d+2}}. \end{aligned}$$

Case 3: $|I_1| = 0, |I_2| > 0$. We suppose that $I_2 = (b_1, b_2, \dots, a_n)$ and denote by $I'_2 = (b_2, \dots, b_n)$. Then, we have $|I'_2| \geq 0$, and

$$\begin{aligned} \|\partial^{I_1} \Omega^{I_2} v(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} &= \|\Omega_{b_1} \Omega^{I'_2} v(t, \cdot)\|_{L^\infty} \\ &\leq C \sum_b \|(1+r) \partial_b \Omega^{I'_2} v(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} = C \sum_b \|\Omega^{I'_2} v(t, \cdot)\|_{\varepsilon_{-1}} \\ &\leq C \sum_b \|\partial_b \Omega^{I'_2} v(t, \cdot)\|_{X^2} = C \sum_{\substack{b \\ |J_1|+|J_2| \leq 2}} \|\partial_x^{J_1} \Omega^{J_2} \partial_b \Omega^{I'_2} v(t, \cdot)\|_{X^2} \\ &\leq C \sum_{\substack{b' \\ |J_2|+|J'_2| \leq 2}} \|\partial_x^{J_1} \partial_{b'} \Omega^{J_2} \Omega^{I'_2} v(t, \cdot)\|_{X^2} \leq C \|v(t, \cdot)\|_{E_P^{d+2}}. \end{aligned}$$

By combining these three cases, (7.10b) is established. \square

Lemma 7.6. — *Let $u \in X_H^d$ and $v \in L^\infty([0, T]; E_P^d)$, then the following estimate holds for all pair of multi-index (I_1, I_2) with $1 \leq |I_1| + |I_2| \leq d$:*

$$(7.11a) \quad \|(1+r)^{-1} \partial_x^{I_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} \leq C \|u\|_{X_P^d}.$$

$$(7.11b) \quad \|(1+r)^{-1} \partial^{I_1} \Omega^{I_2} v(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C \|v(t, \cdot)\|_{E_P^d}, \quad \text{for } t \in [0, T].$$

Proof. — The proof of (7.11a) is decomposed into several cases, as follows.

Case 1: $|I_1| > 0$. In this case we suppose that $I_1 = (a_1, a_2, \dots, a_n)$ and denote by $I'_1 = (a_2, \dots, a_n)$. Then I'_1 is of order $n-1 \geq 0$, and we obtain

$$\|\partial_x^{I_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} = \|\partial_{a_1} \partial_x^{I'_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} \leq \|u\|_{X_P^d}.$$

Case 2: $|I_1| = 0, |I_2| > 0$. In this case we suppose that $I_2 = (b_1, b_2, \dots, b_n)$ and denote by $I'_b = (b_2, \dots, b_n)$. Then I'_b is of order $n-1 \geq 0$ and we have

$$\begin{aligned} \|(1+r)^{-1} \partial_x^{I_1} \Omega^{I_2} u\|_{L^2(\mathbb{R}^3)} &= \|(1+r)^{-1} \Omega_{b_1} \Omega^{I'_b} u\|_{L^2(\mathbb{R}^3)} \\ &\leq C \sum_b \|\partial_b \Omega^{I'_b} u\|_{L^2(\mathbb{R}^3)} \leq C \|u\|_{X_P^d}. \end{aligned}$$

By combining these two cases, (7.11a) is proven.

The proof of (7.11b) is exactly the same if we replace ∂_x by ∂ in the above proof. \square

Lemma 7.7. — *For all function u of class X^2 , one has*

$$(7.12) \quad \|u\|_{L^\infty(\mathbb{R}^3)} \leq C \sum_{|I_1|+|I_2| \leq 2} \|(1+|x|)^{-1} \partial_x^{I_1} \Omega^{I_2} u\|_{L^2}.$$

Proof. — This inequality is equivalent to

$$(7.13) \quad |u(x)| \leq C \sum_{|I_1|+|I_2| \leq 2} \|(1+|x|)^{-1} \partial_x^{I_1} \Omega^{I_2} u\|_{L^2}, \quad x \in \mathbb{R}^3$$

for all $u \in C_c^\infty(\mathbb{R}^3)$. Then by regularization, this inequality is hold for all $u \in X^2$. This is proven by distinguish different x . Let $\chi(\cdot)$ a C^∞ function defined on $[0, \infty)$ with

$$\chi(s) = \begin{cases} 1, & s \leq 1/2, \\ 0, & 1 \leq s < \infty. \end{cases}$$

Then $\chi(|\cdot|)$ is a $C_c^\infty(\mathbb{R}^3)$ function.

Now we consider the case $|x| \leq 1/2$. We consider the function

$$f(x) := \chi(|x|)u(x)$$

which is in class $C_c^\infty(\{|x| < 1\})$. Then we apply the classical Sobolev's inequality:

$$\begin{aligned} |u(x)| = |f(x)| &\leq C \sum_{|I| \leq 2} \|\partial_x^I f\|_{L^2(\{|x| < 1\})} = C \sum_{|I| \leq 2} \|\partial_x^I (\chi(|x|)u(x))\|_{L^2(\{|x| < 1\})} \\ &= C \sum_{\substack{|I_1|+|I_2| \leq |I|, \\ |I| \leq 2}} \|\partial_x^{I_1} \chi(|x|) \partial_x^{I_2} u(x)\|_{L^2(\{|x| < 1\})} \leq C \sum_{|I_2| \leq 2} \|\partial_x^{I_2} u(x)\|_{L^2(\{|x| < 1\})} \\ &\leq C \sum_{|I_2| \leq 2} \|(1+|x|)^{-1} \partial_x^{I_2} u(x)\|_{L^2(\{|x| < 1\})} \leq C \sum_{|I_2| \leq 2} \|(1+|x|)^{-1} \partial_x^{I_2} u(x)\|_{L^2(\mathbb{R}^3)}, \end{aligned}$$

where we used that $\partial_x^{I_1} \chi(|x|)$ are bounded.

When $|x| \geq 1/2$, we have

$$\begin{aligned} |u(x)| &= |(1 - \chi(2|x|))u(x)| = \|(1+|x|) \cdot (1+|x|)^{-1} (1 - \chi(2|x|))u(x)\|_{L^\infty(\mathbb{R}^3)} \\ &\leq \|(1+|x|)^{-1} (1 - \chi(2|x|))u(x)\|_{\varepsilon_{-1}} \end{aligned}$$

and the by (7.8),

$$\begin{aligned} |u(x)| &\leq \|(1+|x|)^{-1} (1 - \chi(2|x|))u(x)\|_{\varepsilon_{-1}} \\ &\leq C \sum_{|I_1|+|I_2| \leq 2} \|\partial_x^{I_1} \Omega^{I_2} ((1+|x|)^{-1} (1 - \chi(2|x|))u(x))\|_{L^2(\mathbb{R}^3)} \\ &= C \sum_{\substack{|J_1|+|J'_1| = |I_1| \\ |I_1|+|I_2| \leq 2}} \|\partial_x^{J'_1} ((1+|x|)^{-1} (1 - \chi(2|x|))) \\ &\quad \partial_x^{J_1} \Omega^{I_2} ((1+|x|)^{-1} (1 - \chi(2|x|))u(x))\|_{L^2(\mathbb{R}^3)}, \end{aligned}$$

where for the last equality we have used the fact that

$$\Omega_a((1+|x|)^{-1} (1 - \chi(2|x|))) = 0,$$

since both factors are radial symmetric. Then we will prove that

$$\partial_x^{J'_1} ((1+|x|)^{-1} (1 - \chi(2|x|))) \leq C(1+|x|)^{-1}$$

for $|J'_1| \leq 2$. This is check directly by calculating and the fact that $(1 - \chi(2|x|))$ and its derivatives are supported out of the ball $\{|x| < 1/4\}$. Then the desired result is established. \square

7.3. Linear estimates

We begin with the linear theory of wave equation with the initial data given in $X_H^{d+1}(\mathbb{R}^3) \times X^d(\mathbb{R}^3)$. For the simplicity of proof, we introduce the energy functional with respect to a metric $g^{\alpha\beta}$ defined in \mathbb{R}^4 as follows:

$$(7.14) \quad E_g(t, u) := \left(\int_{\mathbb{R}^3} (-g^{00}|\partial_t u|^2 + g^{ab}\partial_a u \partial_b u)(t, \cdot) dx \right)^{1/2}$$

and

$$(7.15) \quad E_{g,c}(t, u) := \left(\int_{\mathbb{R}^3} (-g^{00}|\partial_t u|^2 + g^{ab}\partial_a u \partial_b u + c^2 v^2)(t, \cdot) dx \right)^{1/2}.$$

A metric $g^{\alpha\beta}$ is said to be **coercive with constant** $C > 0$ if

$$(7.16) \quad C^{-1} \|\nabla u\|_{L^2(\mathbb{R}^3)} \leq E_g(t, u) \leq C \|\nabla u\|_{L^2(\mathbb{R}^3)},$$

where ∇u refers to the spacetime gradient of u . At this juncture, let us introduce a notation for the C_c^∞ functions defined in the region $[0, T] \times \mathbb{R}^3$:

$$\|u(t, \cdot)\|_{X_E^{d+1}} := \sum_{|I_1|+|I_2|\leq d} \|\nabla \partial_x^{I_1} \Omega^{I_2} u(t, \cdot)\|_{L^2(\mathbb{R}^3)}.$$

By (7.3a) and (7.3b), the norm $\|\cdot\|_{X_E^{d+1}}$ and $\|\nabla(\cdot)\|_{X^d}$ are equivalent.

We also introduce

$$E_g^d(s, u) := \sum_{|I_1|+|I_2|\leq d} E_g(s, \partial^{I_1} \Omega^{I_2} u), \quad E_{g,c}^d(s, u) := \sum_{|I_1|+|I_2|\leq d} E_{g,c}(s, \partial^{I_1} \Omega^{I_2} u).$$

Then when the coercivity condition (7.16) is assumed,

$$(7.17) \quad C^{-1} \|u(t, \cdot)\|_{E_P^{d+1}} \leq E_g^d(s, u) \leq C \|u(t, \cdot)\|_{E_P^{d+1}}.$$

The existence result in the next section is based on the following linear estimate.

Lemma 7.8 (*L^∞ type estimate for wave equation*). — *Let u be a smooth function defined in the region $[0, T]$ and let $F = -\square u$, then for any $0 \leq t \leq T$*

$$(7.18) \quad \|u(t, \cdot)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} \leq Ct \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} ds + C(\|u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} + t\|\nabla u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)}),$$

where ∇u refers to the spacetime gradient of u .

Proof. — This estimate based on the explicit expression of the linear wave equation. We consider the Cauchy problem:

$$(7.19) \quad \begin{aligned} \square u &= -F, \\ u(0, x) &= -f(x), \quad \partial_t u(0, x) = -g(x). \end{aligned}$$

Then, u can be expressed by

$$\begin{aligned} u(t, x) &= \frac{1}{4\pi} \int_0^t \frac{1}{t-s} \int_{|y|=t-s} F(s, x-y) d\sigma(y) ds \\ &\quad + \frac{1}{4\pi t} \int_{|y|=t} g(x-y) d\sigma(y) + \frac{1}{4\pi t^2} \int_{|y|=t} (f(x-y) - \langle \partial_x f(x-y), y \rangle) d\sigma(y) \\ &:= u_1(t, x) + u_2(t, x) + u_3(t, x). \end{aligned}$$

Here, $d\sigma(y)$ refers to the standard Lebesgue measure on the sphere $|y| = t-s$ or $|y| = t$, and $\langle \cdot, \cdot \rangle$ refers to the standard scalar product on \mathbb{R}^3 . The notation $\partial_x f$ stands for the (spacial) gradient of f . The remaining work is to estimate u_i with $i = 1, 2, 3$.

When $|x| \leq 1$, we make use of the fact that $\|\cdot\|_{L^\infty(\mathbb{R}^3)} \leq \|\cdot\|_{\varepsilon_{-1}}$. Then, we have

$$\begin{aligned} |u_1(t, x)| &\leq \frac{1}{4\pi} \int_0^t \frac{\|F(s, \cdot)\|_{\varepsilon_{-1}}}{t-s} \int_{|y|=t-s} d\sigma(y) ds \\ &\leq \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} (t-s) ds \leq t \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds. \end{aligned}$$

The other terms u_2 and u_3 are estimated similarly. Then (7.18) is proven in this case.

When $|x| > 1$, we need to establish the decay estimate of the solution at spatial infinity. We begin with u_1 .

$$\begin{aligned} |u_1(t, x)| &\leq \frac{1}{4\pi} \int_0^t \frac{1}{t-s} \int_{|y|=t-s} \frac{\|F(s, \cdot)\|_{\varepsilon_{-1}}}{1+|x-y|} d\sigma(y) ds \\ &= \frac{1}{4\pi} \int_0^t \frac{\|F(s, \cdot)\|_{\varepsilon_{-1}}}{t-s} \int_{|y|=t-s} (1+|x-y|)^{-1} d\sigma(y) ds. \end{aligned}$$

We focus on the expression $\int_{|y|=t-s} (1+|x-y|)^{-1} d\sigma(y)$. We make the following parametrization of the sphere $\{|y| = t-s\}$. Let θ be the angle from the vector $-x$ to the vector $-y$, and φ refers the angle from the plan determined by pair of vector (x, y) to a fixed plan containing x (for example the plan determined by x and $(0, 0, 1)$). With this parametrization, the volume form $\sigma(y)$ has the following expression:

$$d\sigma(y) = (t-s)^2 \sin \theta d\theta d\varphi.$$

Also, by the classical trigonometrical theorem ‘‘Law of cosines’’,

$$|x-y|^2 = |x|^2 + |t-s|^2 - 2|x|(t-s) \cos \theta.$$

Then, we obtain

$$\begin{aligned} \int_{|y|=t-s} (1+|x-y|)^{-1} d\sigma(y) &= \int_0^{2\pi} \int_0^\pi \frac{(t-s)^2 \sin \theta d\theta d\varphi}{1 + (|x|^2 + |t-s|^2 - 2|x|(t-s) \cos \theta)^{1/2}} \\ &= 2\pi \int_{-1}^1 \frac{(t-s)^2 d\gamma}{1 + (|x|^2 + |t-s|^2 - 2|x|(t-s)\gamma)^{1/2}}, \end{aligned}$$

where $\gamma := \cos \theta$. Then, we have

$$\begin{aligned} \int_{|y|=t-s} (1 + |x - y|)^{-1} d\sigma(y) &= \frac{2\pi(t-s)}{|x|} \int_{-1}^1 \frac{|x|(t-s)d\gamma}{1 + (|x|^2 + |t-s|^2 - 2|x|(t-s)\gamma)^{1/2}} \\ &= \frac{2\pi(t-s)}{|x|} \int_{||x|-(t-s)|}^{|x|+(t-s)} \frac{\tau d\tau}{1 + \tau}, \end{aligned}$$

where $\tau := (|x|^2 + |t-s|^2 - 2|x|(t-s)\gamma)^{1/2}$. Then, we obtain

$$\begin{aligned} |u_1(t, x)| &\leq \frac{1}{4\pi} \int_0^t \frac{\|F(s, \cdot)\|_{\mathbb{B}^{-1}}}{t-s} \int_{|y|=t-s} (1 + |x - y|)^{-1} d\sigma(y) ds \\ &\leq \frac{1}{2|x|} \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} \int_{||x|-(t-s)|}^{|x|+(t-s)} \frac{\tau d\tau}{1 + \tau} ds. \end{aligned}$$

Now the discussion should distinguish between two cases.

Case 1: $|x| > t$. In this case $|x| > t - s$ always holds, and we find

$$\begin{aligned} |u_1(t, x)| &\leq \frac{1}{2|x|} \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} \int_{||x|-(t-s)|}^{|x|+(t-s)} \frac{\tau d\tau}{1 + \tau} ds \\ &= \frac{1}{2|x|} \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} \int_{|x|-(t-s)}^{|x|+(t-s)} \frac{\tau d\tau}{1 + \tau} ds \\ &\leq \frac{1}{2|x|} \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} \int_{|x|-(t-s)}^{|x|+(t-s)} d\tau ds \\ &= \frac{1}{|x|} \int_0^t (t-s) \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds \leq \frac{t}{|x|} \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds, \end{aligned}$$

which leads to

$$|x||u(t, x)| \leq t \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds.$$

Case 2: $|x| \leq t$. In this case we need a more precise calculation:

$$\begin{aligned}
|u_1(t, x)| &\leq \frac{1}{2|x|} \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} \int_{||x|-(t-s)|}^{|x|+(t-s)} \frac{\tau d\tau}{1+\tau} ds \\
&= \frac{1}{2|x|} \int_0^{t-|x|} \|F(s, \cdot)\|_{\varepsilon_{-1}} \int_{(t-s)-|x|}^{|x|+(t-s)} \frac{\tau d\tau}{1+\tau} ds \\
&\quad + \frac{1}{2|x|} \int_{t-|x|}^t \|F(s, \cdot)\|_{\varepsilon_{-1}} \int_{|x|-(t-s)}^{|x|+(t-s)} \frac{\tau d\tau}{1+\tau} ds \\
&\leq \frac{1}{2|x|} \int_0^{t-|x|} 2|x| \|F(s, \cdot)\|_{\varepsilon_{-1}} ds + \frac{1}{2|x|} \int_{t-|x|}^t 2(t-s) \|F(s, \cdot)\|_{\varepsilon_{-1}} ds \\
&\leq \int_0^{t-|x|} \|F(s, \cdot)\|_{\varepsilon_{-1}} ds + \int_{t-|x|}^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds = \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds.
\end{aligned}$$

Recalling that $|x| \leq t \leq T$, we find

$$|u_1(t, x)| \leq \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds \leq \frac{t}{|x|} \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds,$$

so that

$$|x| |u_1(t, x)| \leq t \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds.$$

The estimate of u_2 and u_3 are similar:

$$\begin{aligned}
|u_2(t, x)| &\leq \frac{1}{4\pi t} \int_{|y|=t} |g(x-y)| d\sigma(y) \\
&\leq \frac{\|g\|_{\varepsilon_{-1}}}{4\pi t} \int_{|y|=t} (|x-y|+1)^{-1} d\sigma(y).
\end{aligned}$$

By the same parametrization made in the estimate of u_1 and similar calculation, we can conclude with

$$|x| |u_2(t, x)| \leq t \|g\|_{\varepsilon_{-1}}.$$

In the same way, we have

$$\begin{aligned}
|u_3(t, x)| &\leq \frac{1}{4\pi t^2} \int_{|y|=t} \frac{\|f\|_{\varepsilon_{-1}} d\sigma(y)}{|x-y|+1} + \frac{1}{4\pi t} \int_{|y|=t} \frac{\|\partial_x f\|_{\varepsilon_{-1}} d\sigma(y)}{|x-y|+1} \\
&\leq \frac{1}{|x|} \|f\|_{\varepsilon_{-1}} + \frac{t}{|x|} \|\partial_x f\|_{\varepsilon_{-1}}.
\end{aligned}$$

By combining the estimates made in $|x| \leq 1$ and $|x| > 1$, the desired result is established. \square

Lemma 7.9 (*L^2 type estimate for wave equation*). — Let u be a smooth function defined in the region $[0, T] \times \mathbb{R}^3$ and $g^{\alpha\beta} = m^{\alpha\beta} + H^{\alpha\beta}$ be a smooth metric, where $m^{\alpha\beta}$

is the Minkowski metric with signature $(-, +, +, +)$. Assume that $g^{\alpha\beta}$ is coercive with constant C . Let $f = -g^{\alpha\beta}\partial_\alpha\partial_\beta u$, then for any $0 \leq t \leq T$,

$$(7.20) \quad \frac{d}{dt}E_g(t, u) \leq C\|f(t, \cdot)\|_{L^2(\mathbb{R}^3)} + C \sum_{\alpha, \beta} \|\nabla g^{\alpha\beta}(t, \cdot)\|_{L^\infty} E_g(t, u),$$

where ∇u refers to the spacetime gradient of u .

Proof. — This is a standard calculation and we write

$$\begin{aligned} \partial_t u g^{\alpha\beta} \partial_\alpha \partial_\beta u &= \frac{1}{2} \partial_0 (g^{00} (\partial_0 u)^2 - g^{ab} \partial_a u \partial_b u) + \partial_a (g^{a\beta} \partial_t u \partial_\beta u) \\ &\quad + \frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha u \partial_\beta u - \partial_\alpha g^{\alpha\beta} \partial_t u \partial_\beta u. \end{aligned}$$

Integrating on the slice $\{t = \tau\}$ and applying Stokes' formula, we find

$$\begin{aligned} &\frac{1}{2} \int_{\mathbb{R}^3} \partial_0 (-g^{00} (\partial_t u)^2 + g^{ab} \partial_a u \partial_b u) dx \\ &= \int_{\mathbb{R}^3} \partial_t u f dx + \frac{1}{2} \int_{\mathbb{R}^3} (\partial_t g^{\alpha\beta} \partial_\alpha u \partial_\beta u - 2 \partial_\alpha g^{\alpha\beta} \partial_t u \partial_\beta u) dx, \end{aligned}$$

which leads to (by the coercivity condition)

$$\frac{1}{2} \frac{d}{dt} (E_g(t, u))^2 = \int_{\mathbb{R}^3} \partial_t u f dx + \frac{1}{2} \int_{\mathbb{R}^3} (\partial_t g^{\alpha\beta} \partial_\alpha u \partial_\beta u - 2 \partial_\alpha g^{\alpha\beta} \partial_t u \partial_\beta u) dx,$$

which is

$$E_g(t, u) \frac{d}{dt} E_g(t, u) = \int_{\mathbb{R}^3} \partial_t u f dx + \frac{1}{2} \int_{\mathbb{R}^3} (\partial_t g^{\alpha\beta} \partial_\alpha u \partial_\beta u - 2 \partial_\alpha g^{\alpha\beta} \partial_t u \partial_\beta u) dx.$$

So we have

$$E_g(t, u) \frac{d}{dt} E_g(t, u) \leq \|\partial_t u\|_{L^2(\mathbb{R}^3)} \|f\|_{L^2(\mathbb{R}^3)} + C \|\nabla g^{\alpha\beta}\|_{L^\infty(\mathbb{R}^3)} \|\nabla u\|_{L^2(\mathbb{R}^3)}^2$$

and by recalling (7.16)

$$\frac{d}{dt} E_g(t, u) \leq C \|f\|_{L^2(\mathbb{R}^3)} + C \|\nabla g^{\alpha\beta}\|_{L^\infty(\mathbb{R}^3)} E_g(t, u).$$

□

Now we combine Lemmas 7.8 and 7.9 with the Sobolev estimate (7.8) in order to get a estimate on the E_H^{d+1} norm.

Lemma 7.10 (E_H^d norm estimate on wave equation). — *Let u be a smooth function defined in the region $[0, T] \times \mathbb{R}^3$. Let $g^{\alpha\beta}$ be a smooth metric, $g^{\alpha\beta} = m^{\alpha\beta} + H^{\alpha\beta}$ with $m^{\alpha\beta}$ the standard Minkowski metric. Assume that g satisfies the following coercivity condition with a constant $C > 0$ and $|H^{00}| \leq 1/2$. Let $g^{\alpha\beta} \partial_{\alpha\beta} u = -F$. Then the following estimates hold for $d \geq 3$:*

$$(7.21a) \quad \|u(t, \cdot)\|_{E_P^{d+1}} \leq C \|u(0, \cdot)\|_{E_P^{d+1}} e^{C \int_0^t D_d(\tau) d\tau} + C \int_0^t \|F(\tau, \cdot)\|_{E^d} e^{C \int_\tau^t D_d(s) ds} d\tau,$$

(7.22)

$$\begin{aligned} \|u(t, \cdot)\|_{\varepsilon_{-1}} &\leq Ct \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds + Ct E_g^3(0, u) \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} e^{C \int_0^s D_3(\tau) d\tau} ds \\ &\quad + Ct \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} \int_0^s \|F(\tau, \cdot)\|_{E^3} e^{C \int_\tau^s D_3(\lambda) d\lambda} ds + Ct E_g^3(0, u), \end{aligned}$$

where

$$D_k(t) := \max_{\alpha, \beta} \|H^{\alpha\beta}(t, \cdot)\|_{E_H^{k+1}}.$$

Proof. — We begin with the estimate on E^d norm of ∇u . We differentiate the equation with respect to $\partial^{I_1} \Omega^{I_2}$, where $|I_1| + |I_2| = l$. Remember that this product of operator commute with the linear wave operator, Then

$$-\partial^{I_1} \Omega^{I_2} F = \square \partial^{I_1} \Omega^{I_2} u + H^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{I_1} \Omega^{I_2} u + [\partial^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u$$

which is

$$g^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{I_1} \Omega^{I_2} u = -F - [\partial^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u.$$

Apply (7.20),

$$(7.23) \quad \begin{aligned} \frac{d}{dt} E_g(t, \partial^{I_1} \Omega^{I_2} u) &\leq C \|\partial^{I_1} \Omega^{I_2} F(t, \cdot)\|_{L^2(\mathbb{R}^3)} + C \|[\partial^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u(t, \cdot)\|_{L^2(\mathbb{R}^3)} \\ &\quad + C \sum_{\alpha, \beta} \|\nabla H^{\alpha\beta}(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} E_g(t, \partial^{I_1} \Omega^{I_2} u). \end{aligned}$$

We should estimate the second term in the right-hand side. By (7.3e),

$$(7.24) \quad \begin{aligned} \|[\partial^{I_1} \Omega^{I_2}, H^{\alpha\beta} \partial_\alpha \partial_\beta] u\|_{L^2(\mathbb{R}^3)} &\leq \sum_{\substack{J_1 + J'_1 = I_1 \\ J_2 + J'_2 = I_2 \\ |J'_1| + |J'_2| > 0}} \|\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{J_1} \Omega^{J_2} u\|_{L^2(\mathbb{R}^3)} \\ &\quad + \sum_{\substack{J_1 + J'_1 = I_1 \\ |J_2| + |J'_2| < |I_2| \\ \alpha, \beta, \alpha', a}} \|\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_{\alpha'} \partial_a \partial^{J_1} \Omega^{J_2} u\|_{L^2(\mathbb{R}^3)} \\ &=: T_1 + T_2. \end{aligned}$$

Here, we make the convention that when $l \leq 0$, $[\partial^{I_1} \Omega^{I_1}, \partial_\alpha \partial_\beta] = 0$. We see that both terms can be bounded by $CD_d E^d(s, u)$: For T_1 , when $|J'_1| + |J'_2| \leq d - 1$, Then by (7.10b),

$$\begin{aligned} \|\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{J_1} \Omega^{J_2} u\|_{L^2(\mathbb{R}^3)} &\leq \|\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta}\|_{L^\infty(\mathbb{R}^3)} \|\partial_\alpha \partial_\beta \partial^{J_1} \Omega^{J_2} u\|_{L^2(\mathbb{R}^3)} \\ &\leq C \|H^{\alpha\beta}\|_{E_H^{d+1}} \|u\|_{E_P^{d+1}}. \end{aligned}$$

When $|J'_1| + |J'_2| \geq d$, then $J_1 = J_2 = 0$. Then by (7.11b) and (7.8),

$$\begin{aligned} \|\partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{J_1} \Omega^{J_2} u\|_{L^2(\mathbb{R}^3)} &\leq \|(1+r)^{-1} \partial^{J'_1} \Omega^{J'_2} H^{\alpha\beta}\|_{L^2(\mathbb{R}^3)} \|(1+r) \partial_\alpha \partial_\beta u\|_{L^\infty(\mathbb{R}^3)} \\ &\leq C \|H^{\alpha\beta}\|_{E_H^{d+1}} \|\partial_\alpha \partial_\beta u\|_{E^2} \leq C \|H^{\alpha\beta}\|_{E_H^{d+1}} \|u\|_{E_P^4}. \end{aligned}$$

The term T_2 is bounded in the same manner and we omit the details.

Combining this with (7.23), we find

$$(7.25) \quad \begin{aligned} \frac{d}{dt} E_g(t, \partial^{I_1} \Omega^{I_2} u) &\leq C \|\partial^{I_1} \Omega^{I_2} F(t, \cdot)\|_{L^2(\mathbb{R}^3)} + C D_d E_g^d(t, u) \\ &+ C \sum_{\alpha, \beta} \|\nabla H^{\alpha\beta}(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} E_g(t, \partial^{I_1} \Omega^{I_2} u) \end{aligned}$$

Taking the sum with respect to (I_1, I_2) for $|I_1| + |I_2| \leq d$, then

$$(7.26) \quad \frac{d}{dt} E_g^d(t, u) \leq C \|F(t, \cdot)\|_{E^d} + C D_d E_g^d(t, u).$$

Integrating (7.26), we obtain

$$(7.27) \quad E_g^d(t, u) \leq E_g^d(0, u) e^{C \int_0^t D_d(\tau) d\tau} + C \int_0^t \|F(\tau, \cdot)\|_{E^d} e^{C \int_\tau^t D_d(s) ds} d\tau,$$

which leads to (by (7.17)) (7.21a).

We now turns to estimate the \mathcal{E}_{-1} norm. We can easily deduce that

$$(7.28) \quad \square u = -(1 - H^{00})^{-1} (H^{ab} + H^{00} m^{ab}) \partial_a \partial_b u - 2(1 - H^{00})^{-1} H^{a0} \partial_t \partial_a u - (1 - H^{00})^{-1} F.$$

Here, we use the assumption that $H^{00} \leq 1/2$ to make sure that $(1 - H^{00})^{-1}$ is well defined. By Lemma 7.8, we have

$$\begin{aligned} \|u\|_{\mathcal{E}_{-1}} &\leq Ct \int_0^t \|(1 - H^{00})^{-1} F(s, \cdot)\|_{\mathcal{E}_{-1}} ds \\ &+ Ct \int_0^t \|(1 - H^{00})^{-1} H^{ab} \partial_a \partial_b u\|_{\mathcal{E}_{-1}} ds + Ct \int_0^t \|(1 - H^{00})^{-1} H^{a0} \partial_a \partial_t u\|_{\mathcal{E}_{-1}} ds \\ &+ C(\|u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} + t \|\nabla u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)}) \\ &\leq Ct \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds + Ct \int_0^t \|H^{ab} \partial_a \partial_b u\|_{\mathcal{E}_{-1}} ds + Ct \int_0^t \|H^{a0} \partial_a \partial_t u\|_{\mathcal{E}_{-1}} ds \\ &+ C(\|u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} + t \|\nabla u(0, x)\|_{\mathcal{E}_{-1}(\mathbb{R}^3)}). \end{aligned}$$

We observe that the following estimates are guaranteed by (7.8):

$$\begin{aligned} \|H^{a\beta} \partial_a \partial_\beta u\|_{\mathcal{E}_{-1}} &\leq \|H^{a\beta}\|_{L^\infty} \|\partial_a \partial_\beta u\|_{\mathcal{E}_{-1}} \\ &\leq \|H^{\alpha\beta}\|_{L^\infty} \|\partial_a \partial_\beta u\|_{E^2} \leq C \sum_{\alpha, \beta} \|H^{a\beta}\|_{L^\infty} E_g^3(s, u) \end{aligned}$$

and

$$\|\nabla u(0, \cdot)\|_{\mathcal{E}_{-1}} \leq C \|\nabla u(0, \cdot)\|_{E^2}.$$

By combining these two estimates, we get

$$(7.29) \quad \begin{aligned} &\|u(t, \cdot)\|_{\mathcal{E}_{-1}} \\ &\leq Ct \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds + Ct \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} E_g^3(s, u) ds + Ct E_g^2(0, u). \end{aligned}$$

By combining (7.27) and (7.29), we obtain

$$(7.30) \quad \begin{aligned} \|u(t, \cdot)\|_{\varepsilon_{-1}} &\leq Ct \int_0^t \|F(s, \cdot)\|_{\varepsilon_{-1}} ds + Ct E_g^3(0, u) \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} e^{C \int_0^s D_3(\tau) d\tau} ds \\ &\quad + Ct \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} \int_0^s \|F(\tau, \cdot)\|_{E_R^3} e^{C \int_\tau^s D_3(\lambda) d\lambda} ds + Ct E_g^3(0, u). \end{aligned}$$

□

Furthermore, the following L^2 estimate for Klein-Gordon equations is essential in our analysis.

Lemma 7.11 (*L^2 -type estimate for KG equations*). — *Let v be in a smooth function defined $[0, T] \times \mathbb{R}^3$ and let $F = c^2 v - g^{\alpha\beta} \partial_\alpha \partial_\beta v$, $c > 0$. Suppose that $g^{\alpha\beta} = m^{\alpha\beta} + H^{\alpha\beta}$ satisfies the coercivity condition with a constant C , i.e. (7.16). Then, the following estimate holds for $0 \leq t < T$:*

$$(7.31a) \quad \begin{aligned} c\|v(t, \cdot)\|_{E^d} + \|v(t, \cdot)\|_{E_P^{d+1}} &\leq C(\|v(0, \cdot)\|_{E_P^{d+1}} + c\|v(0, \cdot)\|_{E^d}) e^{C \int_0^t D_d(\tau) d\tau} \\ &\quad + C \int_0^t \|F(s, \cdot)\|_{E^d} e^{C \int_s^t D_d(\tau) d\tau} ds. \end{aligned}$$

Proof. — The proof is essentially the same to that of Lemma 7.8. The only difference comes from the potential term:

$$\begin{aligned} &\partial_t v (g^{\alpha\beta} \partial_\alpha \partial_\beta v - c^2 v) \\ &= \frac{1}{2} \partial_0 (g^{00} (\partial_0 v)^2 - g^{ab} \partial_a v \partial_b v) - \frac{1}{2} (c \partial_0 v)^2 + \partial_a (g^{\alpha\beta} \partial_t v \partial_\beta v) + \frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha v \partial_\beta v - \partial_\alpha g^{\alpha\beta} \partial_t v \partial_\beta v. \end{aligned}$$

Then the same calculation of the proof in (7.8) leads to

$$(7.32) \quad \frac{d}{dt} E_{g,c}(t, v) \leq C \|F(t, \cdot)\|_{L^2(\mathbb{R}^3)} + C \sum_{\alpha, \beta} \|\nabla H^{\alpha\beta}(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} E_{g,c}(t, v).$$

Now we derive the equation with respect to $\partial^{I_1} \Omega^{I_2}$, and perform the same calculation as we done in the proof of Lemma 7.10, then we arrive at:

$$(7.33) \quad E_{g,c}^d(t, u) \leq E_{g,c}^d(0, u) e^{C \int_0^t D_d(\tau) d\tau} + C \int_0^t \|F(\tau, \cdot)\|_{E^d} e^{C \int_\tau^t D_d(s) ds} d\tau.$$

Then combined with the expression of $E_{g,c}$ and the coercivity condition (7.16), the desired result is proven. □

At the end of this subsection, we establish the following estimate on second order time derivative of the solution.

Lemma 7.12. — *Let u be a smooth function defined in \mathbb{R}^4 and suppose that u satisfies the following wave/Klein-Gordon equation:*

$$g^{\alpha\beta} \partial_\alpha \partial_\beta u - c^2 u = -F,$$

where $c \geq 0$. Suppose that $g^{\alpha\beta} = m^{\alpha\beta} + H^{\alpha\beta}$ with $|H^{00}| \leq 1/2$. Then the following estimate hold for all pair of multi-index (I_1, I_2)

$$(7.34) \quad \begin{aligned} & \left\| \partial_x^{I_1} \Omega^{I_2} \partial_t \partial_t u \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| \partial_x^{I_1} \Omega^{I_2} \left((1 - H^{00})^{-1} (m^{ab} + H^{ab}) \partial_a \partial_b u \right) \right\|_{L^2(\mathbb{R}^3)} \\ & \quad + c^2 \left\| \partial_x^{I_1} \Omega^{I_2} \left((1 - H^{00})^{-1} u \right) \right\|_{L^2(\mathbb{R}^3)} \\ & \quad + 2 \left\| \partial_x^{I_1} \Omega^{I_2} \left((1 - H^{00})^{-1} H^{0a} \partial_t \partial_a u \right) \right\|_{L^2(\mathbb{R}^3)} + \left\| \partial_x^{I_1} \Omega^{I_2} \left((1 - H^{00})^{-1} F \right) \right\|_{L^2(\mathbb{R}^3)}. \end{aligned}$$

Proof. — By decomposing the wave operator, we have

$$g^{\alpha\beta} \partial_\alpha \partial_\beta u = (-1 + H^{00}) \partial_t \partial_t u + 2H^{0a} \partial_t \partial_a u + (m^{ab} + H^{ab}) \partial_a \partial_b u$$

and thanks to the equation

$$(-1 + H^{00}) \partial_t \partial_t u + 2H^{0a} \partial_t \partial_a u + (m^{ab} + H^{ab}) \partial_a \partial_b u - c^2 u = -F,$$

we have

$$\partial_t \partial_t u = -\frac{c^2 u}{1 - H^{00}} + \frac{2H^{0a} \partial_t \partial_a u}{1 - H^{00}} + \frac{(m^{ab} + H^{ab}) \partial_a \partial_b u}{1 - H^{00}} + \frac{F}{1 - H^{00}}.$$

□

7.4. Existence results for linear equations

We now establish the existence theory for linear wave and Klein-Gordon equations with initial data in the corresponding functional spaces defined in subsection 6.4. We begin with the wave equation.

Proposition 7.13 (Existence of linear wave equation in E_H^{d+1})

Let $d \geq 3$ be an integer. Assume that $F \in L^1([0, T]; E^d)$, $(u_0, u_1) \in E_H^{d+1} \times E^d$. Assume that $g^{\alpha\beta}$ is a smooth metric defined on $[0, T] \times \mathbb{R}^3$ and coercive with constant $C > 0$, and $H^{\alpha\beta} = g^{\alpha\beta} - m^{\alpha\beta}$ is in the class $C([0, T]; E_H^{d+1})$ with $|H^{00}| < 1/2$. Then the following Cauchy problem

$$(7.35) \quad \begin{aligned} & g^{\alpha\beta} \partial_\alpha \partial_\beta u = -F, \\ & u(0, x) = u_0(x), \quad \partial_t u(0, x) = u_1(x) \end{aligned}$$

has a unique solution in class $C([0, T]; E_H^{d+1})$ with $\partial_t^k u \in C([0, T]; E^{d+1-k})$, $1 \leq k \leq d$. Furthermore, this solution satisfies the following estimate:

$$(7.36a) \quad \|u(t, \cdot)\|_{E_P^{d+1}} \leq C \|\nabla u(0, \cdot)\|_{E^d} e^{C \int_0^t D_d(\tau) d\tau} + C \int_0^t \|F(\tau, \cdot)\|_{E^d} e^{C \int_\tau^t D_d(s) ds} d\tau,$$

$$\begin{aligned}
(7.37) \quad \|u(t, \cdot)\|_{\mathcal{E}_{-1}} &\leq Ct \int_0^t \|F(s, \cdot)\|_{\mathcal{E}_{-1}} ds + Ct E_g^3(0, u) \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} e^{\int_0^s D_3(\tau) d\tau} ds \\
&+ Ct \sum_{\alpha, \beta} \int_0^t \|H^{\alpha\beta}(s, \cdot)\|_{L^\infty} \int_0^s \|F(\tau, \cdot)\|_{E^3} e^{\int_\tau^s D_3(\lambda) d\lambda} ds + Ct E_g^3(0, u),
\end{aligned}$$

where

$$D_k(t) := \max_{\alpha, \beta} \|H^{\alpha\beta}(t, \cdot)\|_{E_H^{k+1}}.$$

Proof of Proposition 7.13. — The uniqueness is direct by applying Lemma 7.9.

The existence is based on the regularization and the estimate proved in Lemma 7.10. We proceed by make a series of triple (u_0^n, u_1^n, F^n) which converges to (u_0, u_1, F) in the following sense:

$$\begin{aligned}
\lim_{n \rightarrow \infty} \|u_0^n - u_0\|_{E_H^{d+1}} &= 0, \quad \lim_{n \rightarrow \infty} \|u_1^n - u_0\|_{E^d} = 0 \\
\lim_{n \rightarrow \infty} \|F^n - F\|_{L^1([0, T]; E^d)} &= 0, \quad \lim_{n \rightarrow \infty} \|H_n^{\alpha\beta} - H^{\alpha\beta}\|_{L^\infty([0, T]; E^d)} = 0,
\end{aligned}$$

where u_0^n, u_1^n are $C_c^\infty(\mathbb{R}^3)$ functions and for all $t \in [0, T]$, $F^n(t, \cdot) \in C_c^\infty(\mathbb{R}^3)$.

By classical existence theorem of linear wave equation (see for example [26]), fix the time interval $[0, T]$, each triple (u_0^n, u_1^n, F^n) determines a unique smooth solution by (7.35). These solution, denoted by u^n , formes a series.

Now we take the difference of the equation satisfied by u^n and u^{n-1} :

$$g_n^{\alpha\beta} \partial_\alpha \partial_\beta (u^n - u^{n-1}) = (g_{n-1}^{\alpha\beta} - g_n^{\alpha\beta}) \partial_\alpha \partial_\beta u^{n-1} + (F^n - F^{n-1}).$$

The apply to this equation the estimate (7.22), we see that the sequence $\{u^n\}$ converges with respect the norm $L^\infty([0, T]; \mathcal{E}_{-1})$.

By estimate (7.21a), $\{u^n\}$ is bounded in $L^\infty([0, T], E_P^{d+1})$. We recall the estimate (7.27) and apply this on the time interval $[t', t''] \subset [0, T]$, we get

$$E_g^d(t'', u^n) - E_g^d(t', u^n) \leq E_g^d(e^{\int_{t'}^{t''} D_d^{\alpha\beta}(\tau) d\tau} - 1) + C \int_{t'}^{t''} \|F^n(\tau, \cdot)\|_{E^d} e^{\int_\tau^{t''} D_d^{\alpha\beta}(s) ds} d\tau,$$

where

$$D_k(t) := \max_{\alpha, \beta} \|H^{\alpha\beta}(t, \cdot)\|_{E_H^{k+1}}.$$

Recall that $D_d^n(\tau)$ and $\|F(\tau, \cdot)\|_{E^d}$ are uniformly (with respect to n) bound. This implies that $\{u^n\}$ is equicontinuous with respect to the norm $L^\infty([0, T], E_P^{d+1})$. Then there is a sub-sequence of $\{u^n\}$ converges in the sense of $L^\infty([0, T], E_P^{d+1})$. We denote it again by u^n . Then we see that $\{u^n\}$ converges in $L^\infty([0, T], E_H^{d+1})$. We denote by u its limit.

When $d \geq 3$, (7.22) shows that $\{u^n\}$ is a Cauchy sequence in $L^\infty([0, T], \mathcal{E}_{-1})$. So that $\{u^n\}$ converges in $C([0, T], E_H^{d+1})$ (u^n are C_c^∞ functions so $u^n \in C([0, T]; \mathcal{E}_{-1})$). Furthermore, since u^n are C_c^∞ functions so they are in $C([0, T], E_H^{d+1})$ which is a closed subspace of $L^\infty([0, T], E_H^{d+1})$. Then $\{u^n\}$ converges in $C([0, T], E_H^{d+1})$. We denote by u the $C([0, T], E_H^{d+1})$ -limit of $\{u^n\}$. Then we see that $u \in C([0, T]; E_H^d)$

We apply the same argument on $\{\partial_t^k u\}$ and get the desired regularity. The estimate on u is gained by taking limit of the estimate on u^n . \square

If we analyse carefully the proof of Proposition 7.13, we can conclude that if the triple (F, u_0, u_1) is only supposed to be in $L^1([0, T], E^d) \times E_P^{d+1} \times E^d$, the Cauchy problem (7.35) determines also a unique solution in $C([0, T], E_P^{d+1})$. We prefer to state this result separately in the following proposition:

Proposition 7.14. — *Let $d \geq 3$ and assume that the triple (F, u_0, u_1) is only supposed to be in $L^1([0, T], E^d) \times E_P^{d+1} \times E^d$. And assume that $g^{\alpha\beta}$ is a C^∞ metric defined on $[0, T] \times \mathbb{R}^3$ and coercive with constant $C > 0$, and $H^{\alpha\beta} = g^{\alpha\beta} - m^{\alpha\beta}$ is in the class $C([0, T]; E_H^{d+1})$. Then the Cauchy problem (7.35) has a unique solution u in $C([0, T], E_P^{d+1})$ and $\partial_t^k u \in C([0, T]; E^{d+1-k})$ for $0 \leq k \leq d$. Furthermore, it satisfies the following estimate:*

$$(7.38a) \quad \|u(t, \cdot)\|_{E_P^{d+1}} \leq C \|\nabla u(0, \cdot)\|_{E^d} e^{\int_0^t CD_d(\tau) d\tau} + C \int_0^t \|F(\tau, \cdot)\|_{E^d} e^{\int_\tau^t D_d(s) ds} d\tau,$$

where

$$D_k(t) := \max_{\alpha, \beta} \|H^{\alpha\beta}(t, \cdot)\|_{E_H^{k+1}}.$$

Apply Lemma 7.11 and taking the same regularization argument as in the proof of Proposition 7.13, the following existence result for linear Klein-Gordon equation holds.

Proposition 7.15 (Existence for KG equation). — *Let $d \geq 3$ and the triple (v_0, v_1, F) be in class $E_R^{d+1} \times E_R^d \times L^1([0, T], E_R^d)$. Assume that $g^{\alpha\beta}$ is a C^∞ metric defined on $[0, T] \times \mathbb{R}^3$ and coercive with constant $C > 0$, and $H^{\alpha\beta} = g^{\alpha\beta} - m^{\alpha\beta}$ is in the class $C[0, T]; E_H^{d+1}$. Then the following Cauchy problem*

$$(7.39) \quad \begin{aligned} g^{\alpha\beta} \partial_\alpha \partial_\beta v - c^2 v &= F, \quad c > 0, \\ v(0, x) &= v_0(x), \quad \partial_t v(0, x) = v_1(x) \end{aligned}$$

has a unique solution in class $C([0, T]; E_R^{d+1}) \cap C^1([0, T]; E_R^d)$. Furthermore, it satisfies the following estimate

$$(7.40a) \quad \begin{aligned} &\|v(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t v(t, \cdot)\|_{E^d} + c \|v(t, \cdot)\|_{E^d} \\ &\leq C (\|\nabla u(0, \cdot)\|_{E^d} + c \|v(0, \cdot)\|_{E^d}) e^{C \int_0^t D_d(\tau, \cdot) d\tau} \\ &\quad + C \int_0^t \|F(s, \cdot)\|_{E^d} e^{C \int_s^t D_d(\tau, \cdot) d\tau} ds, \end{aligned}$$

$$(7.40b) \quad \begin{aligned} \|v(t, \cdot)\|_{E_P^{d+1}} + c \|v(t, \cdot)\|_{E^d} &\leq C (\|\nabla u(0, \cdot)\|_{E^d} + c \|v(0, \cdot)\|_{E^d}) e^{C \int_0^t D_d(\tau, \cdot) d\tau} \\ &\quad + C \int_0^t \|F(s, \cdot)\|_{E^d} e^{C \int_s^t D_d(\tau, \cdot) d\tau} ds, \end{aligned}$$

where

$$D_k(t) := \max_{\substack{|I_1|+|I_2|\leq k \\ \alpha, \beta}} \|\partial^{I_1}\Omega^{I_2}H^{\alpha\beta}(t, \cdot)\|_{L^\infty(\mathbb{R}^3)}.$$

7.5. Nonlinear estimates

To estimate the solution of quasi-linear system, we will need the following estimates on nonlinear terms.

Lemma 7.16. — *Let F be a C^∞ function from \mathbb{R}^m to \mathbb{R} and u a C^∞ application from \mathbb{R}^4 to \mathbb{R}^m with components denoted by $u = (u_1, u_2, \dots, u_m)$. Let Z be a family of one order linear differential operator $Z = \{Z_\alpha\}$ with $\alpha \in \Lambda$, where Λ is a subset of \mathbb{N}^* . Then the following identity holds for all multi-index $I = (\alpha_1, \alpha_2, \dots, \alpha_{|I|})$ with $|I| \geq 1$:*

$$(7.41) \quad Z^I(F(u)) = \sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\sum_{ji} K_{ji} = I} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j.$$

Here

$$P^L = \prod_{i=1}^m \partial_i^{l_i}$$

is a product of partial differential operator with $L = (l_1, \dots, l_m)$ and the convention:

$$Z^I u = 1, \quad \text{if } |I| = 0$$

is applied. Furthermore, in a product if the set of index is empty, this product is regarded as 1. For example,

$$\prod_{i=1}^{l_j} Z^{K_{ji}} u_j = 1, \quad \text{if } l_j = 0.$$

Proof. — We observe that, in the right-hand side and for a fixed L , the sum is taken over all the proper $|L|$ -partition of index I . That is, over all the proper L -partition of abstract index \mathcal{I} with $|\mathcal{I}| = I$. We denote by

$$\mathcal{K}_{ji} = \mathcal{I} \quad \text{with} \quad \{K_{ji}\} = I(\{\mathcal{K}_{ji}\})$$

and we denote by $\mathcal{P}_p(\mathcal{I}, |L|)$ the set of all proper $|L|$ -partition of \mathcal{I} . Then (7.41) can be written as

$$Z^I(F(u)) = \sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j, \quad \{K_{ji}\} = I(\{\mathcal{K}_{ji}\}).$$

Now we associate each term in the right-hand side to a pair $(L, \{\mathcal{K}_{ij}\})$: in the sum, each term $P^L F(u) \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j$ corresponds to an operator P^L . The quantity $\{K_{ji}\}$ is a partition of I which is a restriction of I on a abstract partition $\{\mathcal{K}_{ji}\}$. Note that for fixed

L the sum is taken on $\mathcal{P}_p(\mathcal{I}, |L|)$, so we have constructed a bijection from the terms in the right-hand side to the following set

$$\mathcal{Q}(\mathcal{I}) = \{(L, \{\mathcal{K}_{ji}\}) | \mathcal{K}_{ji} \in \mathcal{P}_p(\mathcal{I}, |L|), 1 \leq |L| \leq |\mathcal{I}|\}$$

We will prove (7.16) by induction on the order of $|I|$ with associated abstract multi-index $\mathcal{I} = \{\alpha_1, \alpha_2, \dots, \alpha_n\}$. We check by direct calculation that this identity is valid for $|I| = 1$. Suppose that it holds for $|I| \leq n$, we consider $|I'| = n + 1$. Let \mathcal{I}' be the associated $n + 1$ order abstract multi-index composed by $\{\alpha_1, \alpha_2, \dots, \alpha_n, \alpha_{n+1}\}$ and the restriction of I' on \mathcal{I} is coincide with I .

Then

$$\begin{aligned} & Z^{I'}(F(u)) \\ &= Z_{I'(\alpha_{n+1})} \left(\sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \right) \\ &= \sum_{1 \leq |L| \leq |I|} Z_{I'(\alpha_{n+1})}(P^L F(u)) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &+ \sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} Z_{I'(\alpha_{n+1})} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &=: T_1 + T_2. \end{aligned}$$

For T_1 , we observe that

$$\begin{aligned} T_1 &= \sum_{1 \leq |L| \leq |I|} \sum_{k=1}^m P^{L'_k} F(u) Z_{I'(\alpha_{n+1})} u_k \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &= \sum_{1 \leq |L| \leq |I|} \sum_{k=1}^m P^{L'_k} F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l'_j} Z^{K'_{ji}} u_j \end{aligned}$$

with $L'_k = (l'_1, l'_2, \dots, l'_k, \dots, l'_m)$ with $l'_j = l_j$ for $j \neq k$ and $l'_k = l_k + 1$ and $\mathcal{K}'_{ji} = \mathcal{K}_{ji}$ with $(j, i) \neq (k, l'_k)$ and $\mathcal{K}'_{kl'_k} = \alpha_{n+1}$. Here, K_{ij} is the restriction of I on \mathcal{K}_{ji} while K'_{ji} is the restriction of I' on \mathcal{K}'_{ji} .

For T_2 ,

$$\begin{aligned} T_2 &= \sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} Z_{I'(\alpha_{n+1})} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &= \sum_{1 \leq |L| \leq |I|} P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_p(\mathcal{I}, |L|)} \sum_{\substack{1 \leq j_0 \leq m \\ 1 \leq i_0 \leq l_{j_0}}} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K'_{ji}} u_j, \end{aligned}$$

where $\mathcal{K}'_{ji} = \mathcal{K}_{ji}$ when $(j, i) \neq (j_0, i_0)$ and $\mathcal{K}'_{j_0 i_0} = \mathcal{K}_{j_0 i_0} \cup \{\alpha_{n+1}\}$.

Now, we associate to each term in T_1 and T_2 a pair $o(L', \{\mathcal{K}_{ji}\})$ in the same manner. This defines an injection from the terms contained in T_1 and T_2 to the set

$$\mathcal{Q}(\mathcal{I}') = \{(L', \{\mathcal{K}_{ji}\}) \mid 1 \leq |L'| \leq |\mathcal{I}'|, \{\mathcal{K}_{ji}\} \in \mathcal{P}_P(|L|, \mathcal{I}')\}.$$

The injectivity property is by checked from the fact that for two terms if $L' = \tilde{L}'$, the different terms correspond to a different partition (by our definition of sum over partitions).

Denote by \mathcal{A} the image of the terms in T_1 and T_2 under this injection. This is a subset of $\mathcal{Q}(\mathcal{I}')$. We will prove that $\mathcal{A} = \mathcal{Q}(\mathcal{I}')$ which leads to the equality in the case $|I'| = n+1$. To do so, let

$$(L', \{\mathcal{K}'_{ji}\}) \in \mathcal{Q}(\mathcal{I}'), \quad L' = (l'_1, l'_2, \dots, l'_m)$$

Then we see that as $\alpha_{n+1} \in \mathcal{I}'$ there is one and only one (j_0, i_0) such that $\alpha_{n+1} \in \mathcal{K}'_{j_0 i_0}$ since $\bigcup_{(j,i)} \mathcal{K}'_{ji} = \mathcal{I}'$ and they are disjoint to each other. We will prove that $(L', \{\mathcal{K}'_{ji}\}) \in \mathcal{A}$.

We define $L = (l_1, l_2, \dots, l_m)$ with $l_j = l'_j$ for $j \neq j_0$ and $l_{j_0} = l'_{j_0} - 1$. We construct $\{\mathcal{K}_{ji}\}$ as follows: $\mathcal{K}_{ji} = \mathcal{K}'_{ji}$ if $(j, i) \neq (j_0, i_0)$ and $\mathcal{K}_{j_0 i_0} = \mathcal{K}'_{j_0 i_0} \cap \{\alpha_{n+1}\}^c$. Such constructed pair (L, \mathcal{K}_{ji}) to be in $\mathcal{Q}(\mathcal{I})$.

When $\mathcal{K}_{j_0 i_0} = \emptyset$, we see that $(L', \{\mathcal{K}'_{ji}\})$ corresponds to a term in T_1 . More precisely in the following term:

$$\begin{aligned} & Z_{I'(\alpha_{n+1})}(P^L F(u)) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_P(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &= P^{L'_k} F(u) \sum_{k=1}^m Z_{\alpha_{n+1}} u_k \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_P(\mathcal{I}, |L|)} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j, \end{aligned}$$

where we can see it by fixing $k = j_0$ and $\{\mathcal{K}_{ji}\}$ in the sum.

When $\mathcal{K}_{j_0 i_0} \neq \emptyset$, $(L', \{\mathcal{K}'_{ji}\})$ corresponds to a term in T_2 :

$$\begin{aligned} & P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_P(\mathcal{I}, |L|)} Z_{I'(\alpha_{n+1})} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{ji}} u_j \\ &= P^L F(u) \sum_{\{\mathcal{K}_{ji}\} \in \mathcal{P}_P(\mathcal{I}, |L|)} \sum_{\substack{1 \leq j_0 \leq m \\ 1 \leq i_0 \leq l_{j_0}}} \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K'_{ji}} u_j \end{aligned}$$

we see it by fixing (i_i, j_0) and $\{\mathcal{K}_{ji}\}$ in the sum. \square

Lemma 7.17. — *Let F be a C^∞ function from \mathbb{R}^m to \mathbb{R} and u a C^∞ application from \mathbb{R}^4 to \mathbb{R}^m . Then the following identity holds for any multi-index I_1, I_2 with $|I_1| + |I_2| \geq 1$:*

$$(7.42) \quad \partial^{I_1} \Omega^{I_2} (F(u)) = \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_{j=1}^m l_j=l} P^L F(u) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j,$$

where $L = (l_1, l_2, \dots, l_m)$ is a m -dimensional vector with its components taking value in \mathbb{N} and P^L the partial differential operator

$$P^L = \prod_{i=1}^m \partial_i^{l_i}$$

and the convention:

$$\partial^{I_1} \Omega^{I_2} u = 1, \quad \text{if } |I_1| = |I_2| = 0$$

is applied.

Proof. — The proof is an application of (7.41). Let $D^I = \partial^{I_1} \Omega^{I_2}$ with $D = \{\partial_\alpha, \Omega_\alpha\}$. We denote by

$$D_\alpha = \partial_\alpha \text{ for } \alpha = 0, 1, 2, 3, \quad D_\alpha = \Omega_{\alpha-3}, \text{ for } \alpha = 4, 5, 6.$$

We denote the components of I_1 and I_2 by

$$I_1 = (\beta_1, \beta_2, \dots, \beta_{n_1}), \quad I_2 = (\gamma_1, \gamma_2, \dots, \gamma_{n_2})$$

Then I is determined by

$$I = (\alpha_1, \alpha_2, \dots, \alpha_{n_1}, \alpha_{n_1+1}, \dots, \alpha_{n_1+n_2})$$

with $\alpha_i = \beta_i$ for $i = 1, 2, n_1$ and $\alpha_{i+n_1} = \gamma_i + 3$ for $i = 1, 2, n_2$.

Remake that D is a family of first-order linear differential operator. Then by (7.41)

$$(7.43) \quad \partial^{I_1} \Omega^{I_2} (F(u)) = \sum_{1 \leq |L| \leq |I|} \sum_{\sum_{j_i} K_{j_i} = I} P^L F(u) \prod_{j=1}^m \prod_{i=1}^{l_j} Z^{K_{j_i}} u_j.$$

Then, since $\sum_{j_i} K_{j_i} = I$ is a partition of I . Then

$$D^{K_{j_i}} = \partial^{K_{1j_i}} \Omega^{K_{2j_i}}$$

with $\sum_{j_i} K_{1j_i} = I_1$ a partition of I_1 and $\sum_{j_i} K_{2j_i} = I_2$ a partition of I_2 . This gives

$$\partial^{I_1} \Omega^{I_2} (F(u)) = \sum_{1 \leq |L| \leq |I|} \sum_{\substack{\sum_{j_i} K_{1j_i} = I_1 \\ \sum_{j_i} K_{2j_i} = I_2}} P^L F(u) \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1j_i}} \Omega^{K_{2j_i}} u_j.$$

Then the desired result is proven. \square

The result of Lemma 7.17 will be applied in the following case where $F(\cdot)$ is supposed to vanish at 0 in second order.

Lemma 7.18. — *Let F be a C^∞ function defined in a compact neighborhood V of 0 in \mathbb{R}^m and $F(0) = \nabla F(0) = 0$. Let $d \geq 3$ and suppose that u map from \mathbb{R}^4 to V with its components u_j in $L^\infty([0, T]; E^d)$. Then the following estimates hold for any couple of index (I_1, I_2) with $|I_1| + |I_2| \leq d$:*

$$(7.44) \quad \left\| \partial^{I_1} \Omega^{I_2} (F(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \leq C(F, V, d) \sum_{k=2}^{|I_1|+|I_2|} \|u(t, \cdot)\|_{E^d}^k, \quad \text{for } t \in [0, T],$$

where $C(F, V, d)$ is a constant determined by F , V and d , and $\|u(t, \cdot)\|_{E^d} := \max_j \|u_j(t, \cdot)\|_{E^d}$.

Proof. — When $|I_1| + |I_2| = 0$, by the condition $F(0) = \nabla F(0) = 0$,

$$|F(u)| \leq C(F, V) \sum_{j=1}^m |u_j|^2$$

which leads to the desired result.

For $|I_1| + |I_2| \geq 1$, proof is based on Lemma 7.17. We take the expression:

$$\partial^{I_1} \Omega^{I_2} (F(u)) = \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_{j=1}^m l_j=l} P^L F(u) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j$$

and observe that for $|L| = \sum_{j=1}^m l_j = 1$, $P^L F(0) = 0$. Then we have, in the compact neighborhood V of 0,

$$|\partial_j F(u)| \leq C(F, V) |u|,$$

where $C(F, V)$ is determined by V and F and $|u| := \max_j |u_j|$. Then,

$$\begin{aligned} (7.45) \quad & |\partial^{I_1} \Omega^{I_2} (F(u))| \leq \sum_{j=1}^m |\partial_j F(u)| |\partial^{I_1} \Omega^{I_2} u_j| \\ & + \sum_{l=2}^{|I_1|+|I_2|} \sum_{\sum_{j=1}^m l_j=l} |P^L F(u)| \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} |\partial^{K_{1ji}} \Omega^{K_{2ji}} u_j| \\ & \leq C(F, V) |u| |\partial^{I_1} \Omega^{I_2} u| \\ & + \sum_{l=2}^{|I_1|+|I_2|} \sum_{\sum_{j=1}^m l_j=l} |P^L F(u)| \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} |\partial^{K_{1ji}} \Omega^{K_{2ji}} u_j|. \end{aligned}$$

The first term in the right-hand side is estimated as follows:

$$\begin{aligned} \|u(t, \cdot) \partial^{I_1} \Omega^{I_2} u(t, \cdot)\|_{L^2(\mathbb{R}^3)} & \leq \|u(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} \|\partial^{I_1} \Omega^{I_2} u(t, \cdot)\|_{L^2(\mathbb{R}^3)} \\ & \leq C \|u(t, \cdot)\|_{E^2} \|u(t, \cdot)\|_{X^d} \leq C \|u(t, \cdot)\|_{E^d}^2, \end{aligned}$$

where the Sobolev's inequality is applied.

The second term in the right-hand side of (7.45) is estimated as follows: we observe the term:

$$\sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} |\partial^{K_{1ji}} \Omega^{K_{2ji}} u_j|$$

Recall that $|L| = \sum_j l_j \geq 2$, then we see that in the product there are at least two factors and

$$\sum_{ji} K_{1ji} = I_1, \quad \sum_{ji} K_{2ji} = I_2$$

is a partition of (I_1, I_2) in $|L| = \sum_j l_j$ pieces with $|L| \geq 2$. Note that

$$\sum_{ji} |K_{1ji}| = |I_1|, \quad \sum_{ji} |K_{2ji}| = |I_2|.$$

We observe that there among the index K_{1ji} there is at most one, denoted by $K_{1i_0j_0}$, is of order higher than $|I_1|/2$. In an other world,

$$|K_{1ji}| \leq \lceil |I_1|/2 \rceil, \quad \text{if } i \neq i_0, j \neq j_0,$$

where $\lceil x \rceil$ denotes the biggest integer less than or equal to x . The same result holds for the index K_{2ji} . Then we conclude that in the decomposition of (I_1, I_2) there is at most one pair of index, denoted by $(K_{1i_0j_0}, K_{2i_0j_0})$, is of order higher than $d/2$. In an other world,

$$|K_{1ji}| + |K_{2ji}| \leq \lceil d/2 \rceil, \quad \text{if } i \neq i_0, j \neq j_0.$$

Then if we take the L^2 norm, we will find that

$$\begin{aligned} & \left\| \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \prod_{\substack{j=1 \\ j \neq j_0}}^m \prod_{\substack{i=1 \\ i \neq i_0}}^{l_j} \left\| \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \cdot \left\| \partial^{K_{1i_0j_0}} \Omega^{K_{2i_0j_0}} u_{j_0}(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \prod_{\substack{j=1 \\ j \neq j_0}}^m \prod_{\substack{i=1 \\ i \neq i_0}}^{l_j} \left\| \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{E^2} \cdot \|u_{j_0}(t, \cdot)\|_{E^d} \\ & \leq \prod_{\substack{j=1 \\ j \neq j_0}}^m \prod_{\substack{i=1 \\ i \neq i_0}}^{l_j} \|u(t, \cdot)\|_{E^{2+[d/2]}} \|u(t, \cdot)\|_{E^d}. \end{aligned}$$

Here, we have applied (7.3a). Now recall $d \geq 3$, then $2 + \lceil d/2 \rceil \leq d$. Then, we get

$$\left\| \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \leq \|u\|_{E^d}^{|L|}.$$

Also, we observe that in the compact neighborhood V , $\sup_{x \in V} |P^L F(u)| \leq C(F, V)$ with $C(F, V, |L|)$ a constant determined by F and V and $|L|$. Then the desired result is proven. \square

Now, we combine Lemma 7.18 with the global Sobolev inequality (7.8).

Lemma 7.19. — *Based on the same assumptions on F and u as in Lemma 7.18, for $|I_1| + |I_2| \leq d - 2$, the following estimate holds:*

$$(7.46) \quad \|\partial^{I_1} \Omega^{I_2} (F(u))(t, \cdot)\|_{\varepsilon_{-1}} \leq C(F, V, d) \sum_{k=2}^{|I_1|+|I_2|+2} \|u(t, \cdot)\|_{E^d}^k.$$

We also need the following estimate in the following discussion.

Lemma 7.20. — *Let $d \geq 3$ and assume that $H(\cdot)$ be a C^∞ function defined in a compact neighborhood V of 0 in \mathbb{R}^m and assume that u is a map from \mathbb{R}^3 to V with its components u_j in class $L^\infty([0, T]; E_H^d)$. Then the following estimate holds for $d \geq |I_1| + |I_2| \geq 1$:*

$$(7.47) \quad \|(1+r)^{-1} \partial^{I_1} \Omega^{I_2} (H(u))(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C(H, V, d) \sum_{k=1}^{|I_1|+|I_2|} \|u(t, \cdot)\|_{E_H^d}^k, \quad t \in [0, T].$$

Proof. — We apply the expression (7.42):

$$\partial^{I_1} \Omega^{I_2} (H(u)) = \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_{j=1}^m l_j=l} P^L H(u) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j$$

Note that $P^L H(u)$ is bounded by a constant $C(H, V, |L|)$ determined by V . The estimate of $\|\partial^{I_1} \Omega^{I_2} (H(u))(t, \cdot)\|_{L^2(\mathbb{R}^3)}$ reduced into the estimate of

$$\left\| (1+r)^{-1} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{L^2(\mathbb{R}^3)},$$

where K_{1ji} and K_{2ji} is a possible partition of (I_1, I_2) . We take the same argument to that the proof of Lemma 7.18. Suppose that $|K_{1j_0 i_0}| + |K_{2j_0 i_0}| \geq |K_{1ji}| + |K_{2ji}|$ for all pares (j, i) , i.e. $(K_{1j_0 i_0}, K_{2j_0 i_0})$ is the pair of index with highest order. Then, we find

$$|K_{1ji}| + |K_{2ji}| \leq [d/2], \quad \text{if } i \neq i_0, j \neq j_0$$

and

$$|K_{1j_0 i_0}| + |K_{2j_0 i_0}| \geq 1,$$

and we also have

$$\begin{aligned} & \left\| (1+r)^{-1} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \prod_{\substack{j=1, \\ j \neq j_0}}^m \prod_{\substack{i=1, \\ i \neq i_0}}^{l_j} \left\| \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \left\| (1+r)^{-1} \partial^{K_{1j_0 i_0}} \Omega^{K_{2j_0 i_0}} u_{j_0}(t, \cdot) \right\|_{L^2(\mathbb{R}^3)}. \end{aligned}$$

In view of (7.10b) and (7.11b), the desired result is proven. \square

Lemma 7.21. — *By taking the same assumption on H and u as in Lemma 7.20, and assuming furthermore that $H(0) = 0$. Then for all $|I_1| + |I_2| \leq d - 2$, the following estimate holds:*

$$(7.48) \quad \left\| \partial^{I_1} \Omega^{I_2} (H(u))(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \leq C(H, V, d) \sum_{k=1}^{|I_1|+|I_2|+2} \|u(t, \cdot)\|_{E_H^d}^k.$$

Proof. — When $I_1 = I_2 = 0$, we recall that the condition $H(0) = 0$ implies

$$\|H(u)\|_{L^\infty(\mathbb{R}^3)} \leq C(H, V) \|u\|_{L^\infty(\mathbb{R}^3)} \leq C(H, V) \|u\|_{\varepsilon_{-1}}.$$

When $|I_1| + |I_2| \geq 1$, we apply (7.47) combined with Lemma 7.7:

$$\begin{aligned} \left\| \partial^{I_1} \Omega^{I_2} (H(u))(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} &\leq C \sum_{|J_1|+|J_2| \leq 2} \left\| \partial_x^{J_1} \Omega^{J_2} \partial^{I_1} \Omega^{I_2} (H(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ &\leq C \sum_{\substack{|J_1|+|J_2'| \leq 2 \\ |I_1'| \leq |I_1|}} \left\| \partial_x^{J_1} \partial^{I_1'} \Omega^{J_2'} \Omega^{I_2} (H(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)}. \end{aligned}$$

Here, we observe that when $|I_1| \geq 1$, by applying (7.7a) successively ($|I_1|$ times), we see that $|I_1'| \geq 1$, this leads to the fact that $|J_1| + |I_1'| + |J_2| + |I_2'| \geq 1$. Then we apply (7.47).

When $|I_1| = 0$, then $|I_2| \geq 1$ then we can also apply (7.47). \square

Lemma 7.22. — *Let F be a C^∞ function defined in a compact neighborhood V_1 of 0 in \mathbb{R}^m and H be a C^∞ function defined in a compact neighborhood V_2 of 0 in \mathbb{R}^n . Assume that $F(0) = 0$ and $\nabla(F) = 0$. Let u be a map from \mathbb{R}^4 to V_1 with its components u_j in class $L^\infty([0, T]; E_H^d)$ and v be a map from \mathbb{R}^4 to V_2 with its components v_i in class $L^\infty([0, T]; E^d)$. Then the following estimate holds for $|I_1| + |I_2| \leq d$ with $d \geq 3$:*

$$(7.49) \quad \begin{aligned} &\left\| \partial^{I_1} \Omega^{I_2} (H(u)F(v))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ &\leq C(H, F, V, d) \left(\sum_{k=0}^d \|u(t, \cdot)\|_{E_H^d}^k \right) \left(\sum_{k=2}^d \|v(t, \cdot)\|_{E^d}^k \right), \quad t \in [0, T]. \end{aligned}$$

Proof. — The proof is based on Lemmas 7.18, 7.19, 7.20, and 7.21:

$$\partial^{I_1} \Omega^{I_2} (H(u)F(v)) = \sum_{\substack{J_1+J_1'=I_1 \\ J_2+J_2'=I_2}} \partial^{J_1} \Omega^{J_2} (H(u) - H(0)) \partial^{J_1'} \Omega^{J_2'} (F(v)) + H(0) \partial^{I_1} \Omega^{I_2} (F(v)).$$

When $d \geq |J_1'| + |J_2'| \geq d/2$, $|J_1| + |J_2| \leq [d/2] \leq d - 2$:

$$\begin{aligned} &\left\| \partial^{J_1} \Omega^{J_2} (H(u) - H(0))(t, \cdot) \partial^{J_1'} \Omega^{J_2'} (F(v))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ &\leq \left\| \partial^{J_1} \Omega^{J_2} (H(u) - H(0))(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial^{J_1'} \Omega^{J_2'} (F(v))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ &\leq C(H, F, V, d) \sum_{k=1}^{|J_1|+|J_2|+2} \|u(t, \cdot)\|_{E_H^d}^k \sum_{k=2}^{|J_1'|+|J_2'|} \|v(t, \cdot)\|_{E^d}^k, \end{aligned}$$

where Lemma 7.21 and Lemma 7.19 are applied (on function $H(u) - H(0)$ and $F(v)$). The term $H(0)\partial^{I_1}\Omega^{I_2}(F(v))$ is estimated by (7.44).

When $|J_1| + |J_2| \geq d/2 > 1$, $|J'_1| + |J'_2| \leq [d/2] \leq d - 2$:

$$\begin{aligned} & \left\| \partial^{J_1}\Omega^{J_2}(H(u))(t, \cdot) \partial^{J'_1}\Omega^{J'_2}(F(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| (1+r)^{-1} \partial^{J_1}\Omega^{J_2}(H(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \left\| (1+r) \partial^{J'_1}\Omega^{J'_2}(F(u))(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \\ & = \left\| (1+r)^{-1} \partial^{J_1}\Omega^{J_2}(H(u))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \left\| \partial^{J'_1}\Omega^{J'_2}(F(u))(t, \cdot) \right\|_{\varepsilon_{-1}} \\ & \leq C(H, F, V, d) \left(\sum_{k=1}^{|J_1|+|J_2|} \|u(t, \cdot)\|_{E_H^d}^k \right) \left(\sum_{k=2}^{|J'_1|+|J'_2|+2} \|v(t, \cdot)\|_{E^d}^k \right), \end{aligned}$$

where Lemma 7.20 and Lemma 7.19 are applied. \square

We will need the following estimate on multi-linear functions.

Lemma 7.23. — *Let $d \geq 3$ be an integer and u_i, v_i , $i = 1, 2, \dots, m$ be functions in class $L^\infty([0, T]; E^d)$. Then the following estimates hold for $m \geq 2$ and $|I_1| + |I_2| \leq d$:*

$$(7.50) \quad \begin{aligned} & \left\| \partial^{I_1}\Omega^{I_2} \left(\prod_{i=1}^m u_i - \prod_{i=1}^m v_i \right) (t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(V, d) \|u(t, \cdot) - v(t, \cdot)\|_{E^d} \sum_{k=1}^{m-1} \sum_{k_1+k_2=k} \|u(t, \cdot)\|_{E^d}^{k_1} \|v(t, \cdot)\|_{E^d}^{k_2}, \end{aligned}$$

where $\|u\|_{X^d} := \max_i \|u_i\|_{X^d}$ and $\|v\|_{X^d} := \max_i \|v_i\|_{X^d}$.

Proof. — We observe the following identity

$$\prod_{i=1}^m u_i - \prod_{i=1}^m v_i = \sum_{k=1}^m (u_k - v_k) \prod_{j < k} u_j \prod_{j > k} v_j.$$

Then, for each term, we have

$$\begin{aligned} & \partial^{I_1}\Omega^{I_2} \left((u_k - v_k) \prod_{j < k} u_j \prod_{j > k} v_j \right) \\ & = \sum_{\substack{\sum_i K_{1i} = I_1 \\ \sum_i K_{2i} = I_2}} \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j. \end{aligned}$$

To estimate the product

$$\partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j,$$

we apply the same reasoning as in the proof of Lemma 7.18, there is at most one pair of multi index of order bigger than $d/2$. Then (7.50) is proven by applying the classical Sobolev's inequality. \square

Lemma 7.24. — Let $d \geq 3$ be an integer and $u_i, v_i, i = 1, 2, \dots, m$ be functions in class $L^\infty([0, T]; E_H^d)$. Then the following estimates hold for $m \geq 2$ for $1 \leq |I_1| + |I_2| \leq d$:

$$(7.51) \quad \begin{aligned} & \left\| (1+r)^{-1} \partial^{I_1} \Omega^{I_2} \left(\prod_{i=1}^m u_i - \prod_{i=1}^m v_i \right) (t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(V, d) \|u(t, \cdot) - v(t, \cdot)\|_{E_H^d} \sum_{k=1}^{m-1} \sum_{k_1+k_2=k} \|u(t, \cdot)\|_{E_H^d}^{k_1} \|v(t, \cdot)\|_{E_H^d}^{k_2}, \quad t \in [0, T], \end{aligned}$$

where $\|u\|_{E_H^d} := \max_i \|u_i\|_{E_H^d}$ and $\|v\|_{X^d} := \max_i \|v_i\|_{E_H^d}$.

Proof. — We observe the following identity

$$\prod_{i=1}^m u_i - \prod_{i=1}^m v_i = \sum_{k=1}^m (u_k - v_k) \prod_{j < k} u_j \prod_{j > k} v_j.$$

Then, for each term, we have

$$\begin{aligned} & \partial^{I_1} \Omega^{I_2} \left((u_k - v_k) \prod_{j < k} u_j \prod_{j > k} v_j \right) \\ & = \sum_{\substack{\sum_i K_{1i} = I_1 \\ \sum_i K_{2i} = I_2}} \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j. \end{aligned}$$

To estimate the product

$$\partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j,$$

we also apply the same reasoning as in the proof of Lemma 7.18, there is at most one pair of multi index of order bigger than $d/2$. We discuss three cases.

Case 1: $|K_{1k}| + |K_{2k}| \geq |K_{1j}| + |K_{2j}|$ for $j = 1, 2, \dots, m$. Hence, (K_{1k}, K_{2k}) takes the highest order. For $j \neq k$, $|K_{1j}| + |K_{2j}| \leq [d/2] \leq d-2$ and $|K_{1k}| + |K_{2k}| \geq 1$, and we have

$$\begin{aligned} & \left\| (1+r)^{-1} \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| (1+r)^{-1} \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \right\|_{L^2(\mathbb{R}^3)} \left\| \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j \right\|_{L^\infty(\mathbb{R}^3)} \\ & \leq C \|u_k - v_k\|_{E_P^d} \left\| \prod_{j > k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j < k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j \right\|_{L^\infty(\mathbb{R}^3)}. \end{aligned}$$

Here, we used (7.10a) on the first factor. Note that $m \geq 2$ then there exists a $j_0 \neq k$. So the second factor is bounded by applying (7.48) and the fact that $|K_{1j}| + |K_{2j}| \leq d-2$.

Case 2: $|K_{1k}| + |K_{2k}| \leq [d/2] \leq d-2$ and there exists $j_0 < k$ such that $|K_{1j_0}| + |K_{2j_0}| \geq |K_{1j}| + |K_{2j}|$. Then, we have

$$\begin{aligned} & \left\| (1+r)^{-1} \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \prod_{j>k} \partial^{K_{1j}} \Omega^{K_{2j}} u_j \prod_{j<k} \partial^{K_{1j}} \Omega^{K_{2j}} v_j \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| \partial^{K_{1k}} \Omega^{K_{2k}} (u_k - v_k) \right\|_{L^\infty(\mathbb{R}^3)} \prod_{j>k} \left\| \partial^{K_{1j}} \Omega^{K_{2j}} u_j \right\|_{L^\infty(\mathbb{R}^3)} \prod_{\substack{j<k \\ j \neq j_0}} \left\| \partial^{K_{1j}} \Omega^{K_{2j}} v_j \right\|_{L^\infty(\mathbb{R}^3)} \\ & \times \left\| (1+r)^{-1} \partial^{K_{1j_0}} \Omega^{K_{2j_0}} u^{j_0} \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C \|u - v\|_{E_H^d} \prod_{j>k} \|u_j\|_{E_H^d} \prod_{\substack{j<k \\ j \neq j_0}} \|v_j\|_{E_H^d} \|u^{j_0}\|_{E_H^d}, \end{aligned}$$

where (7.10b) and (7.11b) are applied.

Case 3: $|K_{1k}| + |K_{2k}| \leq [d/2] \leq d-2$ and there exists $j_0 > k$ such that $|K_{1j_0}| + |K_{2j_0}| \geq |K_{1j}| + |K_{2j}|$. The proof is exactly the same as in the last case provided we exchange the roles of u_j and v_j . \square

Finally, we are able to estimate the difference of two quadratic functions.

Lemma 7.25. — *Let $d \geq 3$ be an integer. Let F be a C^∞ function defined in a compact neighborhood V of 0 in \mathbb{R}^m with $F(0) = \nabla F(0) = 0$. Assume that u and v are maps from \mathbb{R}^4 to V with their components in $L^\infty([0, T]; E^d)$. Then the following estimate hold for all pair of index (I_1, I_2) with $|I_1| + |I_2| \leq d$:*

$$(7.52) \quad \left\| \partial_x^{I_1} \Omega^{I_2} (F(u) - F(v)) \right\|_{L^2(\mathbb{R}^3)} \leq C(F, V, d) \|u - v\|_{X^d} \sum_{k=1}^{|I_1|+|I_2|-1} \sum_{k_1+k_2=k} \|u\|_{X^d}^{k_1} \|v\|_{X^d}^{k_2}.$$

Proof. — When $I_1 = I_2 = 0$, we apply the mean value theorem: there is a $\theta \in [0, 1]$ such that

$$|F(u) - F(v)| \leq |\nabla F(\theta u + (1 - \theta)v)| |u - v|.$$

Then by the condition $\nabla F(0) = 0$,

$$|F(u) - F(v)| \leq |\nabla F(\theta u + (1 - \theta)v)| |u - v| \leq C(F, V) |\theta u + (1 - \theta)v| |u - v|.$$

Then

$$\begin{aligned} \|F(u)(t, \cdot) - F(v)(t, \cdot)\|_{L^2(\mathbb{R}^3)} & \leq C(F, V) \|u - v\|_{L^2(\mathbb{R}^3)} \|\theta u + (1 - \theta)v\|_{L^\infty(\mathbb{R}^3)} \\ & \leq C(V) \|u(t, \cdot) - v(t, \cdot)\|_{E^d} (\|u(t, \cdot)\|_{E^d} + \|v(t, \cdot)\|_{E^d}). \end{aligned}$$

For the case for $|I_1| + |I_2| \geq 1$, we recall the expression (7.42) :

$$\begin{aligned}
& \partial^{I_1} \Omega^{I_2} (F(u) - F(v)) \\
&= \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} \left(P^L F(u) \sum_{\substack{\sum_{j,i} i K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j \right. \\
&\quad \left. - P^L F(v) \sum_{\substack{\sum_{j,i} i K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \right) \\
&= \sum_j \partial_j F(u) (\partial^{I_1} \Omega^{I_2} u_j - \partial^{I_1} \Omega^{I_2} v_j) \\
&+ \sum_{l=2}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} P^L F(u) \sum_{\substack{\sum_{j,i} i K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \left(\prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j - \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \right) \\
&+ \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} (P^L F(u) - P^L F(v)) \sum_{\substack{\sum_{j,i} i K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \\
&=: T_1 + T_2 + T_3.
\end{aligned}$$

To estimate T_1 , we take into consideration of the condition $\nabla F(0) = 0$ with leads to

$$|\partial_j F(u)(t, \cdot)| \leq C(V) |u|.$$

So $\|\partial_j F(u)(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} \leq C(V) \|u(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} \leq C(F, V) \|u(t, \cdot)\|_{X^d}$. Then

$$\|T_1(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C(F, V) \|u(t, \cdot) - v(t, \cdot)\|_{X^d} \|u(t, \cdot)\|_{X^d}.$$

To estimate T_2 , we need to apply Lemma 7.23. To do so, we observe the following relation:

$$(7.53) \quad \partial^{I_1} \Omega^{I_2} \left(\prod_{i=1}^m u_i^{l_i} - \prod_{i=1}^m v_i^{l_i} \right) = \sum_{\substack{\sum_{j,i} i K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \left(\prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j - \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \right)$$

Then

$$T_2 = \sum_{l=2}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} P^L F(u) \partial^{I_1} \Omega^{I_2} \left(\prod_{j=1}^m u_j^{l_j} - \prod_{j=1}^m v_j^{l_j} \right)$$

Recall that $|P^L F(u)|$ is bounded by a constant $C(F, V, |L|)$ determined by the neighborhood V , the function F and the order $|L|$. Then we apply (7.50).

The estimate of T_3 is as follows:

$$(7.54) \quad \|T_3\|_{L^2(\mathbb{R}^3)} \leq \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} \left\| P^L F(u) - P^L F(v) \right\|_{L^\infty(\mathbb{R}^3)} \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \left\| \prod_{j=1}^m \prod_{i=1}^{l_j} \partial_x^{K_{1ji}} \Omega^{K_{2ji}} v_j \right\|_{L^2(\mathbb{R}^3)}$$

As in the estimate on T_2 , we see that

$$\sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \left\| \prod_{j=1}^m \prod_{i=1}^{l_j} \partial_x^{K_{1ji}} \Omega^{K_{2ji}} v(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \leq \sum_{l=1}^{|I_1|+|I_2|} \|v(t, \cdot)\|_{X^d}^l.$$

and

$$\begin{aligned} \|P^L F(u)(t, \cdot) - P^L F(v)(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} &\leq C(F, V, |L|) \|u(t, \cdot) - v(t, \cdot)\|_{L^\infty} \\ &\leq C(V, |L|) \|u(t, \cdot) - v(t, \cdot)\|_{E^d} \end{aligned}$$

and then

$$\|T_3(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C(V) \|u(t, \cdot) - v(t, \cdot)\|_{E^d} \sum_{k=1}^{d-1} \|v\|_{E^d}^k$$

which leads us to the conclusion (7.52). \square

The following L^∞ type estimate is a direct application of Lemma 7.25 and the global Sobolev inequality (7.8):

Lemma 7.26. — *Let F, G, u and v take the assumption as in Lemma 7.25, then for any $|I_1| + |I_2| \leq d - 2$, the following estimate holds:*

$$(7.55) \quad \begin{aligned} &\left\| \partial^{I_1} \Omega^{I_2} (F(u)(t, \cdot) - F(v)(t, \cdot))(t, \cdot) \right\|_{\mathcal{E}_{-1}(\mathbb{R}^3)} \\ &\leq C(F, V, d) \|u(t, \cdot) - v(t, \cdot)\|_{E^d} \sum_{k=1}^{d-1} \sum_{k_1+k_2=k} \|u(t, \cdot)\|_{E^d}^{k_1} \|v(t, \cdot)\|_{E^d}^{k_2}. \end{aligned}$$

Lemma 7.27. — *Let $d \geq 3$ be an integer. Let H be a C^∞ function which is defined in a compact neighborhood V of 0 in \mathbb{R}^m and let u, v be maps from \mathbb{R}^4 to V with their components in class $L^\infty([0, T], E_H^d)$. Then the following estimate holds for $1 \leq |I_1| + |I_2| \leq d$:*

$$(7.56) \quad \begin{aligned} &\left\| (1+r)^{-1} \partial^{I_1} \Omega^{I_2} (H(u) - H(v))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ &\leq C(V) \|u(t, \cdot) - v(t, \cdot)\|_{E_H^d} \sum_{k=0}^{|I_1|+|I_2|} \sum_{k_1+k_2=k} \|u(t, \cdot)\|_{E_H^d}^{k_1} \|v(t, \cdot)\|_{E_H^d}^{k_2}. \end{aligned}$$

Proof. — The proof is quite similar to that of Lemma 7.25.

When $|I_1| + |I_2| \geq 1$, we apply the same calculation:

$$\begin{aligned}
& \partial^{I_1} \Omega^{I_2} (H(u) - H(v)) \\
&= \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} \left(P^L H(u) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j \right. \\
&\quad \left. - P^L H(v) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \right) \\
&= \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} P^L H(u) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \left(\prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} u_j - \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \right) \\
&+ \sum_{l=1}^{|I_1|+|I_2|} \sum_{\sum_j l_j=l} (P^L H(u) - P^L H(v)) \sum_{\substack{\sum_{j,i} K_{1ji}=I_1 \\ \sum_{j,i} K_{2ji}=I_2}} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial^{K_{1ji}} \Omega^{K_{2ji}} v_j \\
&=: T_1 + T_2.
\end{aligned}$$

The estimate of T_1 and T_2 are similar to the estimate made in the proof of Lemma 7.25. T_1 is estimated by apply Lemma 7.24 and T_2 is by applying the following estimate

$$\|P^L H(u) - P^L H(v)\|_{L^\infty(\mathbb{R}^3)} \leq C(H, V, |L|) \|u - v\|_{L^\infty} \leq C(V, |L|) \|u - v\|_{X^d}$$

and the fact that (by applying (7.47))

$$\left\| (1+r)^{-1} \prod_{j=1}^m \prod_{i=1}^{l_j} \partial_x^{K_{1ji}} \Omega^{K_{2ji}} v_j \right\|_{L^2(\mathbb{R}^3)} \leq C(V) \|v\|_{X_H^d}^{|L|}.$$

□

We also need a L^∞ estimate on $H(u) - H(v)$:

Lemma 7.28. — *Let $d \geq 3$. Assume that H be a C^∞ function defined in a compact neighborhood V of 0 in \mathbb{R}^m and assume that u, v are maps from \mathbb{R}^4 to V with their components in the class $L^\infty([0, T]; E_H^d)$. Then the following estimate holds for $|I_1| + |I_2| \leq d - 2$:*

$$\begin{aligned}
(7.57) \quad & \|\partial^{I_1} \Omega^{I_2} (H(u) - H(v))(t, \cdot)\|_{L^\infty(\mathbb{R}^3)} \\
& \leq C(H, V, d) \|u(t, \cdot) - v(t, \cdot)\|_{E_H^d} \sum_{k=0}^{|I_1|+|I_2|+1} \sum_{k_1+k_2=k} \|u(t, \cdot)\|_{E_H^d}^{k_1} \|v(t, \cdot)\|_{E_H^d}^{k_2}
\end{aligned}$$

Proof. — When $I_1 = I_2 = 0$, then there exists $\theta \in [0, 1]$ such that

$$|H(u) - H(v)| \leq |\nabla H(\theta u + (1 - \theta)v)| |u - v| \leq C(H, V) |u - v|.$$

which proves (7.57).

For $|I_1| + |I_2| \geq 1$, we apply Lemma 7.7 combined with (7.56). \square

Lemma 7.29. — *Let F be a C^∞ function defined in a compact neighborhood V_1 of 0 in \mathbb{R}^m and H be a C^∞ function defined in a neighborhood V_2 of 0 in \mathbb{R}^n . Let d be an integer and $d \geq 3$. Assume that $F(0) = 0, \nabla F(0) = 0$. Let u_1 and u_2 be maps from \mathbb{R}^4 to V_1 with their components in $L^\infty([0, T]; E_H^d)$ and v_1, v_2 be maps from \mathbb{R}^4 to V_2 with their components in class $L^\infty([0, T]; E^d)$. Then the following estimate holds for $|I_1| + |I_2| \leq d$ and $d \geq 3$. Then, we have*

$$(7.58) \quad \begin{aligned} & \left\| \partial^{I_1} \Omega^{I_2} (H(u_1)F(v_1) - H(u_2)F(v_2))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(V) (\|u_1(t, \cdot) - u_2(t, \cdot)\|_{E_H^d} + \|v_1(t, \cdot) - v_2(t, \cdot)\|_{E^d}) \sum_{k=1}^{d-1} M^k(t), \end{aligned}$$

where

$$M(t) := \max\{\|u_{1j}(t, \cdot)\|_{E_H^d}, \|u_{2j}(t, \cdot)\|_{E_H^d}, \|v_{1j}(t, \cdot)\|_{E^d}, \|v_{2j}(t, \cdot)\|_{E^d}\}.$$

Proof. — We have

$$\begin{aligned} & \partial^{I_1} \Omega^{I_2} (H(u_1)F(v_1) - H(u_2)F(v_2)) \\ & = \partial^{I_1} \Omega^{I_2} (H(u_1)F(v_1) - H(u_1)F(v_2) + H(u_1)F(v_2) - H(u_2)F(v_2)) \\ & = \partial^{I_1} \Omega^{I_2} (H(u_1)(F(v_1) - F(v_2))) + \partial^{I_1} \Omega^{I_2} (F(v_2)(H(u_1) - H(u_2))) \\ & =: T_1 + T_2. \end{aligned}$$

The term T_1 is estimated as follows:

$$T_1 = \partial^{I_1} \Omega^{I_2} (H(u_1)(F(v_1) - F(v_2))) = \sum_{\substack{J_1+J_2=I_1 \\ J_1'+J_2'=I_2}} \partial^{J_1} \Omega^{J_2} (H(u_1)) \partial^{J_1'} \Omega^{J_2'} (F(v_1) - F(v_2)).$$

When $|J_1| + |J_2| \leq [d/2] \leq d - 2$, by applying (7.48) and (7.52)

$$\begin{aligned} & \left\| \partial^{J_1} \Omega^{J_2} (H(u_1)) \partial^{J_1'} \Omega^{J_2'} (F(v_1) - F(v_2))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & = \left\| \partial^{J_1} \Omega^{J_2} (H(u_1))(t, \cdot) \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial^{J_1'} \Omega^{J_2'} (F(v_1) - F(v_2))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(H, F, V, d) \left(\sum_{k=0}^{|J_1|+|J_2|+2} \|u_1(t, \cdot)\|_{E_H^d} \right) \\ & \quad \left(\|v_1(t, \cdot) - v_2(t, \cdot)\|_{E^d} \sum_{k=1}^{|J_1'|+|J_2'|-1} \sum_{k_1+k_2=k} \|v_1(t, \cdot)\|_{E^d}^{k_1} \|v_2(t, \cdot)\|_{E^d}^{k_2} \right) \end{aligned}$$

When $|J_1| + |J_2| \geq d/2 > 1$, $|J'_1| + |J'_2| \leq [d/2]$, we apply (7.47):

$$\begin{aligned}
& \left\| \partial^{J_1} \Omega^{J_2} (H(u_1)) \partial^{J'_1} \Omega^{J'_2} (F(v_1) - F(v_2))(t, \cdot) \right\|_{L^2(\mathbb{R}^3)} \\
& \leq \|(1+r)^{-1} \partial^{J_1} \Omega^{J_2} (H(u_1))(t, \cdot)\|_{L^2(\mathbb{R}^3)} \|\partial^{J'_1} \Omega^{J'_2} (F(v_1) - F(v_2))(t, \cdot)\|_{\varepsilon_{-1}} \\
& \leq C(V) \sum_{k=1}^{|J_1|+|J_2|} \|u\|_{E_H^d}^k \cdot C(H, F, V, d) \|v_1(t, \cdot) - v_2(t, \cdot)\|_{E^d} \sum_{k=1}^{|J'_1|+|J'_2|+1} \sum_{k_1+k_2=k} \|v_1(t, \cdot)\|_{E^d}^{k_1} \|v_2(t, \cdot)\|_{E^d}^{k_2} \\
& \leq C(H, F, V, d) \|v_1(t, \cdot) - v_2(t, \cdot)\|_{E^d} \sum_{k=1}^{|I_1|+|I_2|+2} \sum_{k_1+k_2+k_3=k} \|u_1(t, \cdot)\|_{E_H^d}^{k_1} \|v_1(t, \cdot)\|_{E^d}^{k_2} \|v_2(t, \cdot)\|_{E^d}^{k_3}
\end{aligned}$$

The term T_2 is estimated similarly by applying (7.44), (7.56), (7.46) and (7.57). We omit the details, but we write out the estimate

$$\|T_2(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq \|u_1(t, \cdot) - u_2(t, \cdot)\|_{E_H^d} \sum_{k=2}^{|I_1|+|I_2|+1} \sum_{k_1+k_2+k_3=k} \|u_1\|_{E_H^d}^{k_1} \|u_2\|_{E_H^d}^{k_2} \|v_2\|_{E^d}^{k_3}.$$

□

8. Proof of the local existence

8.1. Iteration and uniform bound

In this section we will begin the proof of Theorem 6.3. The proof of this theorem will occupy the following two subsections and follows a classical iteration procedure:

$$\begin{aligned}
(8.1a) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1} &= F_{\alpha\beta}(h_n^\kappa; \partial h_n^\kappa, \partial h_n^\kappa) - 16\pi \partial_\alpha \phi_n^\kappa \partial_\beta \phi_n^\kappa \\
&\quad - 12 \partial_\alpha \varrho_n^\kappa \partial_\beta \varrho_n^\kappa - \kappa^{-1} V_h(\varrho_n^\kappa) (m_{\alpha\beta} + h_{\alpha\beta}^{\kappa, n}),
\end{aligned}$$

$$(8.1b) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \partial_{\beta'} \phi_{n+1}^\kappa = 2(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \phi_n^\kappa \partial_{\beta'} \varrho_n^\kappa,$$

$$\begin{aligned}
(8.1c) \quad & (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \partial_{\beta'} \varrho_{n+1}^\kappa - \frac{\varrho_{n+1}^\kappa}{3\kappa} \\
& = \kappa^{-1} V_\rho(\varrho_n^\kappa) - \frac{4\pi}{3e^{2\varrho_n^\kappa}} (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \phi_n^\kappa \partial_{\beta'} \phi_n^\kappa,
\end{aligned}$$

where

$$\kappa^{-1} V_h(s) := \frac{(e^{2s} - 1)^2}{2\kappa e^{4s}}, \quad \kappa^{-1} V_\rho(s) := \frac{e^{2s} - 1}{6\kappa e^{4s}} - \frac{s}{3\kappa}$$

and with initial data

$$\begin{aligned}
h_{\alpha\beta}^{\kappa, n}(0, x) &= h_{0\alpha\beta}, & \phi_n^\kappa(0, x) &= \phi_0(x), & \varrho_n^\kappa(0, x) &= \varrho_0(x), \\
\partial_t h_{\alpha\beta}^{\kappa, n}(0, x) &= h_{1\alpha\beta}, & \partial_t \phi_n^\kappa(0, x) &= \phi_1(x) & \partial_t \varrho_n^\kappa(0, x) &= \varrho_1(x).
\end{aligned}$$

Recall the function $H^{\alpha\beta}(h)$ are defined in (6.7) and the associated estimates are in (6.8). We take $(h_{\alpha\beta}^{\kappa,0})$ and set

$$S_n^\kappa := (h_{\alpha\beta}^{\kappa,n}, \phi_n^\kappa, \varrho_n^\kappa).$$

We also denote by $F_H(S_n^\kappa)$, $F_P(S_n^\kappa)$ and $F_R(S_n^\kappa)$ the terms in the right-hand side of (8.1a), (8.1b) and (8.1c).

We take $S_0^\kappa = (h_{\alpha\beta}^{\kappa,0}, \phi_0^\kappa, \varrho_0^\kappa)$ as the solution of the following homogeneous Cauchy problem:

$$\begin{aligned} \square h_{\alpha\beta} &= 0, & \square \phi &= 0, & \square \varrho - \frac{\varrho}{3\kappa} &= 0, \\ h_{\alpha\beta}^{\kappa,0}(0, x) &= h_{0\alpha\beta}, & \phi_0^\kappa(0, x) &= \phi_0(x), & \varrho_0^\kappa(0, x) &= \varrho_0(x), \\ \partial_t h_{\alpha\beta}^{\kappa,0}(0, x) &= h_{1\alpha\beta}, & \partial_t \phi_0^\kappa(0, x) &= \phi_1(x) & \partial_t \varrho_0^\kappa(0, x) &= \varrho_1(x). \end{aligned}$$

We see that the source terms and metric coefficients in (8.1) are sufficiently regular and the initial data are in the corresponding class required in Propositions 7.13, 7.14 and 7.15. Then, by the theory of local existence for linear equations, this iteration procedure is well defined in a fixed time interval $[0, T]$, where the metric coefficients and source terms are in the corresponding class and $|H(h_n^\kappa)^{00}| \leq 1/2$. We see that this iteration defines a sequence of triple $S_n^\kappa := (h_{\alpha\beta}^{\kappa,n}, \phi_n^\kappa, \varrho_n^\kappa)$. In order to get the local existence, we will prove that S_n^κ converges in the following norm:

$$\begin{aligned} &\|S(t, \cdot)_n^\kappa\|_{E_\kappa^{d+1}} \\ &:= \max\{\|h_{\alpha\beta}^{\kappa,n}(t, \cdot)\|_{E_H^{d+1}}, \|\phi_n^\kappa(t, \cdot)\|_{E_P^{d+1}}, \|\varrho_n^\kappa(t, \cdot)\|_{E_P^{d+1}}, \kappa^{-1/2}\|\varrho_n^\kappa(t, \cdot)\|_{E^d}\}, \end{aligned}$$

for all $t \in [0, T^*]$, in which $T^* > 0$ will be defined.

We suppose that for certain $d \geq 3$, for all $k \leq n$ and all $t \in [0, T]$,

$$(8.2) \quad \|S_k^\kappa(t, \cdot)\|_{E_\kappa^{d+1}} \leq A\epsilon,$$

we will show that $\|S_{n+1}^\kappa(t, \cdot)\|_{X_\kappa^{d+1}} \leq A\epsilon$ with T and A well chosen.

First, we observe that when (8.2) holds with $A\epsilon \leq \epsilon_0$ with ϵ_0 small enough, the metric $g_{\alpha\beta}^{\kappa,k} := m^{\alpha\beta} + h_{\alpha\beta}^{\kappa,k}$ are coercive with constant $C(\epsilon_0)$, where $C(\epsilon_0)$ is determined by ϵ_0 .

Now, by combining (8.2) with Lemma 7.18, the following estimates on the source terms follow.

Lemma 8.1. — *Assume that (8.2) holds with $d \geq 3$. Suppose that $0 \leq A\epsilon \leq \epsilon_0$, where ϵ_0 is a constant sufficiently small. Then the following estimates holds for $k \leq n$:*

$$(8.3a) \quad \|F_{\alpha\beta}(h_k^\kappa, \partial h_k^\kappa, \partial h_k^\kappa)(t, \cdot)\|_{E^d} + \|\partial_\alpha \phi_k^\kappa \partial_\beta \phi_k^\kappa(t, \cdot)\|_{E^d} + \|\partial_\alpha \varrho_k^\kappa \partial_\beta \varrho_k^\kappa(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d)(A\epsilon)^2,$$

$$(8.3b) \quad \|(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_k^\kappa))\partial_{\alpha'} \phi_k^\kappa \partial_{\beta'} \phi_k^\kappa(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d)(A\epsilon)^2,$$

$$(8.3c) \quad \|e^{-2\varrho_k^\kappa}(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'} \phi_k^\kappa \partial_{\beta'} \phi_k^\kappa(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d)(A\epsilon)^2,$$

$$(8.3d) \quad \|\kappa^{-1}V_\rho(\varrho_k^\kappa)(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d)(A\epsilon)^2, \quad \|\kappa^{-1}V_h(\varrho_k^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^{\kappa,k})(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d)(A\epsilon)^2.$$

These are classical estimate when we establish the local theory of existence for quasi-linear wave equations with quadratic nonlinearity. The only thing important is (8.3d), where the κ appears in the left-hand side in order to get estimates independent of κ .

Proof. — These estimates are applications of (7.44) and (7.49) combined with (8.2). The only thing we need to be pay attention is that to guarantee the C^∞ regularity of the function $H^{\alpha\beta}(\cdot)$, we need to restrict its defined in a compact neighborhood V of 0 in \mathbb{R}^{10} with

$$V = \{ \max_{1 \leq i \leq 10} |x^i| \leq \epsilon_0 \}$$

with ϵ_0 sufficiently small. This can be guaranteed by taking $\epsilon \leq \epsilon_0$ in (8.2).

We observe that $F_{\alpha\beta}(h_n^\kappa, \partial h_n^\kappa, \partial h_n^\kappa)$ is quadratic with respect to ∂h_n^κ and C^∞ with respect to h_n^κ . Then we apply (7.49). The estimate on the term

$$(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \phi_k^\kappa \partial_{\beta'} \phi_k^\kappa$$

is established in the same manner.

The estimate of (8.3d) is checked by using the estimate

$$\kappa^{-1/2} \|\varrho_k^\kappa\|_{E^d} \leq A\epsilon,$$

which follows from (8.2). □

Now we begin the discussion of the commutators such as $[\partial_x^{I_1} \Omega^{I_2}, H^{\alpha'\beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}] h_{\alpha\beta}^{\kappa,n}$ which appears in the estimates of $\|h_{\alpha\beta}^{\kappa,n+1}\|_{X_E^d}$.

Lemma 8.2. — *There exists a positive constant ϵ_0 such that if (8.2) holds with $d \geq 3$, $A\epsilon \leq \epsilon_0 \leq 1$, then the following estimates hold for all couple of index (I_1, I_2) with $1 \leq |I_1| + |I_2| \leq d$:*

$$(8.4a) \quad \left\| [\partial^{I_1} \Omega^{I_2}, H^{\alpha'\beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}] h_{\alpha\beta}^{\kappa,n+1}(t, \cdot) \right\|_{E^d} \leq C(\epsilon_0, d) A\epsilon \|h_{\alpha\beta}^{\kappa,n+1}(t, \cdot)\|_{E_P^{d+1}},$$

$$(8.4b) \quad \left\| [\partial^{I_1} \Omega^{I_2}, H^{\alpha'\beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}] \phi_{n+1}^\kappa(t, \cdot) \right\|_{E^d} \leq C(\epsilon_0, d) A\epsilon \|\phi_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}},$$

$$(8.4c) \quad \left\| [\partial^{I_1} \Omega^{I_2}, H^{\alpha'\beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}] \varrho_{n+1}^\kappa(t, \cdot) \right\|_{E^d} \leq C(\epsilon_0, d) A\epsilon \|\varrho_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}}.$$

Proof. — The estimate of these three commutators are similar, and we only prove the first statement. Let (I_1, I_2) be a pair of multi-indices, $|I_1| + |I_2| \leq d$. Recall the estimate of commutator (7.3e):

$$\begin{aligned} |[\partial^{I_1} \Omega^{I_2}, H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'}] h_{\alpha'\beta'}^{\kappa,n+1}| &\leq \sum_{\substack{J_1+J'_1=I_1 \\ |J_2|+|J'_2| \leq |I_2| \\ \alpha', \beta', \alpha'', a}} |\partial^{J'_1} \Omega^{J'_2} H^{\alpha'\beta'}| |\partial_{\alpha''} \partial_a \partial^{J_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa,n+1}| \\ &+ \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2 \\ |J'_1|+|J'_2| > 0}} |\partial^{J'_1} \Omega^{J'_2} H^{\alpha'\beta'}| |\partial_{\alpha'} \partial_{\beta'} \partial^{J_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa,n+1}| \\ &=: T_1 + T_2. \end{aligned}$$

We begin with T_2 and distinguish between two cases.

Case 1: $1 \leq |J'_1| + |J'_2| \leq d - 1$, $1 \leq |J_1| + |J_2| \leq d - 1$. In this case, we have

$$\begin{aligned} & \left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} (h_n^\kappa) \partial_{\alpha'} \partial_{\beta'} \partial_x^{J_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa, n+1} \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} (h_n^\kappa) \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial_{\alpha'} \partial_{\beta'} \partial_x^{J_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa, n+1} \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(\epsilon_0, d) A \epsilon \|h_n^\kappa(t, \cdot)\|_{E_P^{d+1}}, \end{aligned}$$

where we used $A \epsilon \leq 1$ and (7.48).

Case 2: $J'_1 = I_1$, $J'_2 = I_2$, $J_1 = J_2 = 0$. Recall that $d \geq 3$ then $|J_1| + |J_2| \leq d - 3$:

$$\begin{aligned} & \left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} (h_n^\kappa) \partial_{\alpha'} \partial_{\beta'} \partial_x^{J_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa, n+1} \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \left\| (1+r)^{-1} \partial_x^{I_1} \Omega^{I_2} H^{\alpha' \beta'} (h_n^\kappa) \right\|_{L^2(\mathbb{R}^3)} \left\| \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1} \right\|_{\mathcal{E}_{-1}} \\ & \leq C(\epsilon_0, d) A \epsilon \|h_{\alpha\beta}^{\kappa, n+1}\|_{E_P^{d+1}}, \end{aligned}$$

where we used $A \epsilon \leq 1$ and (7.47).

The estimate of term T_1 is quite simpler. Recall that in the expression of T_1 , the sum is taken over the index satisfying the following conditions:

$$J_1 + J'_1 = I_1, \quad |J_2| + |J'_2| \leq |I_2| - 1$$

So

$$(|J_1| + |J'_1|) + (|J_2| + |J'_2|) \leq d - 1$$

which leads to

$$|J'_1| + |J'_2| \leq d - 1.$$

So

$$\left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} \partial_{\alpha''} \partial_a \partial_x^{I_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa, n+1} \right\|_{L^2(\mathbb{R}^3)} \leq \left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} \right\|_{L^\infty} \left\| \partial_{\alpha''} \partial_a \partial_x^{I_1} \Omega^{J_2} h_{\alpha\beta}^{\kappa, n+1} \right\|_{L^2(\mathbb{R}^3)}$$

As in the estimate of T_2 , we see that with $|J'_1| + |J'_2| \leq d - 1$,

$$\left\| \partial_x^{J'_1} \Omega^{J'_2} H^{\alpha' \beta'} \right\|_{L^\infty(\mathbb{R}^3)} \leq C(\epsilon_0, d) A \epsilon.$$

The second factor on $h_{\alpha\beta}^{\kappa, n+1}$ is bounded directly by $\|h_{\alpha\beta}^{\kappa, n+1}\|_{X_E^{d+1}}$. So we conclude with (8.4a). \square

Now we need to discuss the bound of the initial data $E_{g_n}^d(0, h_{\alpha\beta}^{\kappa, n+1})$, $E_{g_n}^d(0, \phi_{n+1}^\kappa)$ and $E_{g_n, c}^d(0, \varrho_{n+1}^\kappa)$. We will see that these norms are controlled by $\|S_0\|_{X_\kappa^{d+1}}$:

Lemma 8.3. — *When $\|S_0\|_{X_\kappa^{d+1}}$ is supposed to be bounded by $\epsilon \leq \epsilon_0$ for ϵ_0 sufficiently small, there exists a positive constant determined by ϵ_0 and d such that*

$$E_{g_n}^d(0, h_{n+1}^\kappa) + E_{g_{n+1}}^d(0, \phi^\kappa) + E_{g_{n+1}, \kappa^{-1/2}}^d(0, \varrho^\kappa) \leq C(\epsilon_0, d) \epsilon.$$

Proof. — We recall that

$$\|S_0\|_{X_\kappa^{d+1}} := \max\{\|h_{0\alpha\beta}\|_{X_H^{d+1}}, \|h_{1\alpha\beta}\|_{X^d}, \kappa^{-(1/2)[d/2]+1/4}\|\phi_0\|_{X_P^{d+1}}, \\ \kappa^{-(1/2)[d/2]+1/4}\|\phi_1\|_{X^d}, \kappa^{-[d/2]-1/2}\|\varrho_0\|_{X_P^{d+1}}, \kappa^{-[d/2]-1/2}\|\varrho_1\|_{X^d}\}.$$

We observe that when $0 \leq k \leq 1$ the norm $\|\partial_t^k \partial_x^{I_1} \Omega^{I_2} \varrho^\kappa(0, \cdot)\|_{L^2(\mathbb{R}^3)}$ is determined directly by ϱ_0 and ϱ_1 thus, bounded by $\kappa^{-[d/2]-1/2} \leq \kappa^{-1/2}$.

When $2 \leq k \leq d$, we need to use the equation.

We will prove that for $0 \leq |k| \leq d-2$ and $|I_1| + |I_2| \leq d-1-k$

$$(8.5) \quad \begin{aligned} \|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} h_{\alpha\beta}^\kappa(0, \cdot)\|_{L^2(\mathbb{R}^3)} &\leq C(\epsilon_0, d)\epsilon, \\ \|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} \varrho^\kappa(0, \cdot)\|_{L^2(\mathbb{R}^3)} &\leq C(\epsilon_0, d)\kappa^{[d/2]+1/2-[k/2]}\epsilon, \\ \|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} \phi^\kappa(0, \cdot)\|_{L^2(\mathbb{R}^3)} &\leq C(\epsilon_0, d)\kappa^{1/2[d/2]-1/4}\epsilon \end{aligned}$$

This is proven by induction on k . We see that for $k = 0, 1$, the estimates hold by direct verification. Suppose that (8.5) holds for $(k-1, k)$ we will prove the case $k+1$.

The estimate of $\|\partial_t \partial_t \partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} \phi^\kappa\|_{L^2(\mathbb{R}^3)}$, is a bit complicated. We see that by Lemma 7.12,

$$\begin{aligned} \|\partial_t \partial_t \partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} \varrho^\kappa\|_{L^2(\mathbb{R}^3)} &\leq \|\partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})^{-1}(m^{ab} + H^{ab})\partial_a \partial_b \varrho^\kappa)\|_{L^2(\mathbb{R}^3)} \\ &\quad + 2\|\partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})^{-1}H^{0a}\partial_t \partial_a \varrho^\kappa)\|_{L^2(\mathbb{R}^3)} \\ &\quad + \frac{3}{\kappa}\|\partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} \varrho^\kappa\|_{L^2(\mathbb{R}^3)} \\ &\quad + \|\partial_t^{k+1} \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})F_R)\|_{L^2(\mathbb{R}^3)} \end{aligned}$$

Then we see that

$$\|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} \varrho^\kappa\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)\kappa^{[d/2]+1/2-[d/2]}\epsilon.$$

We observe that by Lemma 7.12,

$$\begin{aligned} \|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} h_{\alpha\beta}^\kappa\|_{L^2(\mathbb{R}^3)} &\leq \|\partial_t^k \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})^{-1}(m^{ab} + H^{ab})\partial_a \partial_b h_{\alpha\beta}^\kappa)\|_{L^2(\mathbb{R}^3)} \\ &\quad + 2\|\partial_t^k \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})^{-1}H^{0a}\partial_t \partial_a h_{\alpha\beta}^\kappa)\|_{L^2(\mathbb{R}^3)} \\ &\quad + \|\partial_t^k \partial_x^{I_1} \Omega^{I_2} ((1-H^{00})F_H)\|_{L^2(\mathbb{R}^3)} \end{aligned}$$

Then by the bounds prescribed by $\|S_0\|_\kappa^{d+1}$ and (8.5), we see that $\|\partial_t \partial_t \partial_t^{I_1} \Omega^{I_2} h_{\alpha\beta}^\kappa\|_{L^2(\mathbb{R}^3)}$ is bounded by $C(\epsilon_0, d)\epsilon$.

In the same manner, we see that for $0 \leq k \leq d-2$

$$\|\partial_t \partial_t \partial_t^k \partial_x^{I_1} \Omega^{I_2} \phi^\kappa\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)\kappa^{-(1/2)[d/2]+1/4}\epsilon.$$

□

Now we are ready to estimate the L^2 type norm of S_n^κ .

Lemma 8.4. — *There exists a positive constant ϵ_0 such that if (8.2) holds for $A\epsilon \leq \epsilon_0 \leq 1$ and $d \geq 3$, then*

$$(8.6a) \quad \|h_{\alpha\beta}^{\kappa,n+1}(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t h_{\alpha\beta}^{\kappa,n+1}(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d) (\epsilon e^{C(\epsilon_0, d)A\epsilon t} + A\epsilon (e^{C(\epsilon_0, d)A\epsilon t} - 1))$$

$$(8.6b) \quad \|\phi_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t \phi_{n+1}^\kappa(t, \cdot)\|_{E^d} \leq C(\epsilon_0, d) (\epsilon e^{C(\epsilon_0, d)A\epsilon t} + A\epsilon (e^{C(\epsilon_0, d)A\epsilon t} - 1))$$

$$(8.6c) \quad \begin{aligned} \|\varrho_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t \varrho_{n+1}^\kappa(t, \cdot)\|_{E^d} + \kappa^{-1/2} \|\varrho_{n+1}^\kappa(t, \cdot)\|_{E^d} \\ \leq C(\epsilon_0, d) (\epsilon e^{C(\epsilon_0, d)A\epsilon t} + A\epsilon (e^{C(\epsilon_0, d)A\epsilon t} - 1)), \end{aligned}$$

where $C(\epsilon_0, d)$ is a positive constant determined by ϵ_0 and d .

Proof. — This is an application of the L^2 estimate (7.20). We consider the estimate for $h_{\alpha\beta}^{\kappa,n+1}$. To do so, we derive the equation (8.1a) with respect to a product $\partial^{I_1}\Omega^{I_2}$ with $|I_1| + |I_2| \leq d$:

$$(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa)) \partial_{\alpha'} \partial_{\beta'} \partial^{I_1}\Omega^{I_2} h_{\alpha\beta}^{\kappa,n+1} = \partial^{I_1}\Omega^{I_2} F_H(h_n^\kappa, \phi_n^\kappa, \varrho_n^\kappa) - [\partial^{I_1}\Omega^{I_2}, H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'}] h_{\alpha\beta}^{\kappa,n+1},$$

where F_H denotes the terms in the right-hand side of (8.1a), which is

$$F_H(h_n^\kappa, \phi_n^\kappa, \varrho_n^\kappa) = F_{\alpha\beta}(h_n^\kappa; \partial h_n^\kappa, \partial h_n^\kappa) - 16\pi \partial_\alpha \phi_n^\kappa \partial_\beta \phi_n^\kappa - 12\partial_\alpha \varrho_n^\kappa \partial_\beta \varrho_n^\kappa - \kappa^{-1} V_h(\varrho_n^\kappa) (m_{\alpha\beta} + h_{\alpha\beta}^{\kappa,n}).$$

Then by (7.20),

$$\begin{aligned} \frac{d}{dt} E_{g_n}(t, \partial^{I_1}\Omega^{I_2} h_{\alpha\beta}^{\kappa,n+1}) &\leq C \|\partial^{I_1}\Omega^{I_2} F_H\|_{L^3(\mathbb{R}^3)} + C \|[\partial^{I_1}\Omega^{I_2}, H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'}] h_{\alpha\beta}^{\kappa,n+1}\|_{L^2(\mathbb{R}^3)} \\ &\quad + C \sum_{\alpha\beta} \|\nabla H^{\alpha'\beta'}(h_n^\kappa)\|_{L^\infty(\mathbb{R}^3)} E_{g_n}(t, \partial^{I_1}\Omega^{I_2} h_{\alpha\beta}^{\kappa,n+1}). \end{aligned}$$

By Lemma 8.1 and (8.2), and the (equi-)coercivity of g_n guaranteed by $A\epsilon \leq \epsilon_0$,

$$\begin{aligned} \frac{d}{dt} E_{g_n}(t, \partial^{I_1}\Omega^{I_2} h_{\alpha\beta}^{\kappa,n+1}) &\leq C(\epsilon_0, d) (A\epsilon)^2 + C(\epsilon_0, d) A\epsilon \sum_{\alpha\beta} E_g^d(t, h_{\alpha\beta}^{\kappa,n+1}) \\ &\quad + C(\epsilon_0, d) A\epsilon E_{g_n}(t, \partial^{I_1}\Omega^{I_2} h_{\alpha\beta}^{\kappa,n+1}). \end{aligned}$$

Taking the sum over the index (I_1, I_2) with $|I_1| + |I_2| \leq d$ and α, β :

$$\frac{d}{dt} \sum_{\alpha\beta} E_{g_n}^d(t, h_{\alpha\beta}^{\kappa,n+1}) \leq C(\epsilon_0, d) (A\epsilon)^2 + C(\epsilon_0, d) A\epsilon \sum_{\alpha\beta} E_{g_n}^d(t, h_{\alpha\beta}^{\kappa,n+1}),$$

which leads to

$$(8.7) \quad \sum_{\alpha\beta} E_{g_n}^d(t, h_{\alpha\beta}^{\kappa,n+1}) \leq \sum_{\alpha\beta} E_{g_n}^d(0, h_{\alpha\beta}^{n+1}) e^{C(\epsilon_0, d)A\epsilon t} + A\epsilon (e^{C(\epsilon_0, d)A\epsilon t} - 1)$$

Note that by Lemma 8.3, $E_{g_n}^d(0, h_{\alpha\beta}^{\kappa,n+1})$ is controlled by $\|h_0\|_{X_P^{d+1}}$ and $\|h_1\|_{X^d}$, so it can be controlled by $C(\epsilon_0, d)\epsilon$, where $C(\epsilon_0, d)$ is a constant depending only on d and ϵ_0 .

The estimate of $E_{g_n}^d(t, \phi_{n+1}^\kappa)$ is exactly the same, and we omit the details.

The estimate of on ϱ_{n+1}^κ is similar. By the same energy estimate, we arrive at the following estimate:

$$(8.8) \quad \begin{aligned} \frac{d}{dt} E_{g_n, \kappa^{-1/2}}(t, \partial^{I_1} \Omega^{I_2} \varrho_{n+1}^\kappa) &\leq C \|\partial^{I_1} \Omega^{I_2} F_R\|_{L^3(\mathbb{R}^3)} + C \left\| [\partial^{I_1} \Omega^{I_2}, H^{\alpha' \beta'} \partial_{\alpha'} \partial_{\beta'}] \varrho_{n+1}^\kappa \right\|_{L^2(\mathbb{R}^3)} \\ &\quad + C \sum_{\alpha \beta} \left\| \nabla H^{\alpha' \beta'}(h_n^\kappa) \right\|_{L^\infty(\mathbb{R}^3)} E_{g_n, \kappa^{-1/2}}(t, \partial^{I_1} \Omega^{I_2} \varrho_{n+1}^\kappa), \end{aligned}$$

where F_R denotes the right-hand side of (8.1c). Then, also by Lemma 8.1 and (8.4c) and the same calculation in the estimate of $h_{\alpha\beta}^{\kappa, n+1}$,

$$E_{g_n, \kappa^{-1/2}}^d(t, \varrho_{n+1}^\kappa) \leq E_{g_n, \kappa^{-1/2}}^d(0, \varrho_{n+1}^\kappa) e^{C(\epsilon_0, d) A \epsilon t} + A \epsilon (e^{C(\epsilon_0, d) A \epsilon t} - 1).$$

□

Now we begin to make the choice of the couple (A_0, T_0) such that when $A \leq A_0$, $T \leq T_0$, (8.6) implies

$$(8.9) \quad \begin{aligned} &\|h_{\alpha\beta}^{\kappa, n+1}(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t h_{\alpha\beta}^{\kappa, n+1}(t, \cdot)\|_{E^d} \leq A \epsilon, \\ &\|\phi_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t \phi_{n+1}^\kappa(t, \cdot)\|_{E^d} \leq A \epsilon, \\ &\|\varrho_{n+1}^\kappa(t, \cdot)\|_{E_P^{d+1}} + \|\partial_t \varrho_{n+1}^\kappa(t, \cdot)\|_{E^d} + \kappa^{-1/2} \|\varrho_{n+1}^\kappa(t, \cdot)\|_{E^d} \leq A \epsilon \end{aligned}$$

on the time interval $[0, T]$.

Lemma 8.5. — *There exists a couple of positive constants $(\epsilon_0, A_0(\epsilon_0, d))$, where ϵ_0 is a universal constant and $A_0(\epsilon_0, d)$ is determined by d and ϵ_0 such that when (8.2) is valid with $A \geq A_0$ and $A \epsilon \leq \epsilon_0 \leq 1$ on the time interval $[0, T]$ with*

$$T \leq T_0 := \frac{\ln(1 + (2C(\epsilon_0, d))^{-1})}{C(\epsilon_0, d) A \epsilon},$$

where $C(\epsilon_0, d)$ is a constant determined by ϵ_0 and d . Therefore, (8.9) hold.

Proof. — By Lemma 8.4, we chose $A_0(\epsilon_0, d)$ and T such that when $A \geq A_0(\epsilon_0, d)$ and $t \leq T$

$$C(\epsilon_0, d) (\epsilon e^{C(\epsilon_0, d) A \epsilon t} + A \epsilon (e^{C(\epsilon_0, d) A \epsilon t} - 1)) \leq A \epsilon.$$

This can be guaranteed by

$$\begin{cases} e^{C(\epsilon_0, d) A \epsilon T_0} - 1 \leq \frac{1}{2C(\epsilon_0, d)}, \\ e^{C(\epsilon_0, d) A \epsilon T_0} \leq \frac{A}{2C(\epsilon_0, d)} \end{cases}$$

which is equivalent to

$$T \leq \frac{\ln(1 + (2C(\epsilon_0, d))^{-1})}{C(\epsilon_0, d) A \epsilon} \quad A \geq 2C(\epsilon_0, d) + 1.$$

Then we can take $A_0 = 2C(\epsilon_0, d) + 1$ and $T_0 = \frac{\ln(1+(2C(\epsilon_0, d))^{-1})}{C(\epsilon_0, d)A\epsilon}$. \square

Then we are about to estimate the \mathcal{E}_{-1} norm of $h_{\alpha\beta}^{\kappa, n+1}$.

Lemma 8.6. — *There exists a positive constant ϵ_0 such that if (8.2) holds with $d \geq 3$, $A \geq A_0$, $A\epsilon \leq \epsilon_0$ on the time interval $[0, T]$, $T \leq T_0$, where (A_0, T_0) are constants determined in Lemma 8.5. Then the following estimate holds:*

$$(8.10) \quad \|h_{\alpha\beta}^{\kappa, n+1}(t, \cdot)\|_{\mathcal{E}_{-1}} \leq C(\epsilon_0)T^2(A\epsilon)^2 + C(T+1)\epsilon.$$

Proof. — We will apply (7.18), and we note that

$$\square h_{\alpha\beta}^{\kappa, n+1} = F_H - H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1}.$$

Then by (7.18), we have

$$\begin{aligned} \|h_{\alpha\beta}^{\kappa, n+1}(t, \cdot)\|_{\mathcal{E}_{-1}} &\leq Ct \int_0^t \|F_H(\tau, \cdot)\|_{\mathcal{E}_{-1}} d\tau + Ct \sum_{\alpha, \beta} \int_0^t \|H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1}(\tau, \cdot)\|_{\mathcal{E}_{-1}} d\tau \\ &\quad + C(\|h_{\alpha\beta}^{\kappa, n+1}(0, \cdot)\|_{\mathcal{E}_{-1}} + t\|\nabla h_{\alpha\beta}^{\kappa, n+1}(0, \cdot)\|_{\mathcal{E}_{-1}}) \end{aligned}$$

Then we can apply on each term the global Sobolev inequality (7.8) to get estimates on the \mathcal{E}_{-1} norms:

$$\|F_H\|_{\mathcal{E}_{-1}} \leq C\|F_H\|_{X^2} \leq C\|F_H\|_{E^2} \leq C(\epsilon_0)(A\epsilon)^2,$$

where we used (8.3a) and (7.8). We have

$$\begin{aligned} \|H^{\alpha'\beta'} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1}(t, \cdot)\|_{\mathcal{E}_{-1}} &\leq \|H^{\alpha'\beta'}\|_{L^\infty(\mathbb{R}^3)} \|\partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n+1}\|_{X^2} \\ &\leq C\|H^{\alpha\beta}\|_{E_H^2} \|h_{\alpha\beta}^{\kappa, n+1}\|_{E_P^4} \\ &\leq C(\epsilon_0)(A\epsilon)^2. \end{aligned}$$

Here, we used Lemma 8.5.

The initial terms $\|h_{\alpha\beta}^{\kappa, n+1}(0, x)\|_{\mathcal{E}_{-1}}$ and $\|\nabla h_{\alpha\beta}^{\kappa, n+1}\|_{\mathcal{B}_{-1}}$ are determined by the initial data h_0 and h_1 , hence, can be controlled by $C\epsilon$, where C is a universal constant. So we conclude with the desired result. \square

Now we can conclude that, with suitable choice of A and T and sufficient small ϵ , the sequence $\{S_n\}$ is bounded with respect to the norm $\|\cdot\|_{X_S^{d+1}}$. More rigorously, the following proposition.

Proposition 8.7. — *There exists a couple of positive constant (A, T) depends only on ϵ_0 , ϵ and d such that if (8.2) holds on $[0, T]$, then*

$$(8.11) \quad \|S_n^\kappa(t, \cdot)\|_{E_\kappa^{d+1}} \leq A\epsilon \leq \epsilon_0 \leq 1,$$

which means that the sequence of triple $\{S_n^\kappa\}$ is bounded in the Banach space E_κ^{d+1} . Furthermore, if $\epsilon \rightarrow 0^+$,

$$T \rightarrow \infty.$$

Note that the choice of (A, T) are **independent** of κ .

Proof. — By Lemma 8.5, we take already $A \geq A_0(\epsilon_0, d)$ and $T \leq T_0$ such that (8.9) holds. In order to prove (8.11), we need only to guarantee, by (8.10), the following inequality:

$$C(\epsilon_0)T^2(A\epsilon)^2 + C(T+1)\epsilon \leq A\epsilon.$$

This can be guaranteed by

$$\begin{cases} T \leq \frac{A}{2C} - 1, \\ T^2 \leq \frac{1}{2C(\epsilon_0)A\epsilon}. \end{cases}$$

So we require that $\frac{A}{2C} - 1 > 0$. Taking into consideration of the conditions in Lemma 8.5:

$$\begin{cases} A \geq A_0(\epsilon_0, d) = 2C(\epsilon_0, d) + 1 \\ T \leq T_0 = \frac{\ln(1 + (2C(\epsilon_0, d))^{-1})}{C(\epsilon_0, d)A\epsilon} \end{cases}$$

together with the condition $A\epsilon \leq \epsilon_0$. So we see that when ϵ sufficiently small such that

$$A_0(\epsilon_0, d) \leq \epsilon_0\epsilon^{-1/3},$$

we can take $A = \epsilon_0\epsilon^{-1/3}$ and $T = \min\{A(2C)^{-1} - 1, (2C(\epsilon_0)A\epsilon)^{-1/2}, T_0\}$. Then there exist a constant $C'(\epsilon_0, d)$ such that $T \geq C'(\epsilon_0, d)\epsilon^{-1/3}$. This proves the desired result. \square

8.2. Contraction property and local existence

To establish theorem 6.3, we need to prove that the sequence $\{S_n^\kappa\}$ is contracting.

Proposition 8.8. — *Let (8.11) holds with (A, T) determined in by Proposition 8.7. Assume that $d \geq 4$. Then there exist a time interval $[0, T^*]$ determined by ϵ_0, ϵ and d such that the sequence $\{S_n\}$ is contracting in the following sense:*

$$(8.12) \quad \|S_{n+1}^\kappa - S_n^\kappa\|_{L^\infty([0, T^*]; X_\kappa^d)} \leq \lambda \|S_n^\kappa(t, \cdot) - S_{n-1}^\kappa(t, \cdot)\|_{L^\infty([0, T^*]; X_\kappa^d)},$$

with a fixed $0 < \lambda < 1$. Furthermore, when $\epsilon \rightarrow 0^+$, we can take T^* such that $T^*(\epsilon) \rightarrow +\infty$.

We emphasize that here the lower bound of the life-span-time T^* given here **does not** depend on the coefficient κ .

The rest of this section is mainly devoted to the proof of this proposition. To do so, we start by taking the difference of between the iteration relation for the pair $(S_{n+1}^\kappa, S_n^\kappa)$ and that of $(S_n^\kappa, S_{n-1}^\kappa)$. This leads to the following differential system

$$(8.13a) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n}) = (H^{\alpha'\beta'}(h_{n-1}^\kappa) - H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^{\kappa, n} + F_H(S_n^\kappa) - F_H(S_{n-1}^\kappa),$$

$$(8.13b) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}(\phi_{n+1}^\kappa - \phi_n^\kappa) = (H^{\alpha'\beta'}(h_{n-1}^\kappa) - H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}\phi_n^\kappa + F_P(S_n^\kappa) - F_P(S_{n-1}^\kappa),$$

$$(8.13c) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}(\varrho_{n+1}^\kappa - \varrho_n^\kappa) = (H^{\alpha'\beta'}(h_{n-1}^\kappa) - H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}\varrho_n^\kappa + F_R(S_n^\kappa) - F_R(S_{n-1}^\kappa)$$

with zero initial data

$$\begin{aligned} h_{\alpha\beta}^{\kappa,n+1}(0, x) - h_{\alpha\beta}^{\kappa,n}(0, x) &= 0, & \partial_t(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})(0, x) &= 0, \\ \phi_{n+1}^\kappa(0, x) - \phi_n^\kappa(0, x) &= 0, & \partial_t(\phi_{n+1}^\kappa - \phi_n^\kappa) &= 0, \\ \varrho_{n+1}^\kappa(0, x) - \varrho_n^\kappa(0, x) &= 0, & \partial_t(\varrho_{n+1}^\kappa - \varrho_n^\kappa)(0, x) &= 0 \end{aligned}$$

For simplicity of expression, we denote by $D_H(S_n, S_{n-1})$ the right-hand side of (8.13a), by $D_P(S_n, S_{n-1})$ the right-hand side of (8.13b) and by $D_R(S_n, S_{n-1})$ the right-hand side of (8.13c). We need to estimate

$$(8.14) \quad \begin{aligned} &\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{X_E^d}, \quad \|\phi_{n+1}^\kappa - \phi_n^\kappa\|_{X_E^d}, \quad \|\varrho_{n+1}^\kappa - \varrho_n^\kappa\|_{X_E^d}, \\ &\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{\varepsilon_{-1}}. \end{aligned}$$

First we recall the uniform bound of the sequence constructed in the last subsection:

$$(8.15) \quad \|S_n^\kappa(t, \cdot)\|_{E_\kappa^{d+1}} \leq A\epsilon \leq \epsilon_0 \leq 1$$

with $d \geq 3$ for $0 \leq t \leq T$. We observe that this condition is equivalent to (8.9) for all $n \in \mathbb{N}^*$.

Now we will make a series of estimates to bound the norm listed in (8.14).

Lemma 8.9. — *Let $\{S_n^\kappa\}$ be the sequence constructed by (8.1) which satisfies the uniform bound condition (8.15) with $d \geq 4$. Then the following estimate holds for $|I_1| + |I_2| \leq d - 1$:*

$$(8.16) \quad \|\partial^{I_1}\Omega^{I_2}D_H(S_{n+1}^\kappa, S_n^\kappa)(t, \cdot)\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)A\epsilon\|S_n^\kappa(t, \cdot) - S_{n-1}^\kappa(t, \cdot)\|_{E_\kappa^d}.$$

Proof. — This is guaranteed by (7.58) and (8.2). Recall that

$$\begin{aligned} D_H(S_n, S_{n-1}) &= -(H^{\alpha'\beta'}(h_n^\kappa) - H^{\alpha'\beta'}(h_{n-1}^\kappa))\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^{\kappa,n} \\ &\quad + (F_{\alpha\beta}(h_n^\kappa, \partial h_n^\kappa, \partial h_n^\kappa) - F_{\alpha\beta}(h_{n-1}^\kappa, \partial h_{n-1}^\kappa, \partial h_{n-1}^\kappa)) \\ &\quad - 16\pi(\partial_\alpha\phi_n^\kappa\partial_\beta\phi_n^\kappa - \partial_\alpha\phi_{n-1}^\kappa\partial_\beta\phi_{n-1}^\kappa) - 12(\partial_\alpha\varrho_n^\kappa\partial_\beta\varrho_n^\kappa - \partial_\alpha\varrho_{n-1}^\kappa\partial_\beta\varrho_{n-1}^\kappa) \\ &\quad - \kappa^{-1}\left(V_h(\varrho_n^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^{\kappa,n}) - V_h(\varrho_{n-1}^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^{\kappa,n-1})\right) \\ &=: T_1 + T_2 + T_3 + T_4. \end{aligned}$$

We observe that $\|T_1\|_{E^d}$ is bounded by $C(\epsilon_0, d)A\epsilon\|h_n^\kappa - h_{n-1}^\kappa\|_{X_H^d}$.

$$\|\partial^{I_1}\Omega^{I_2}T_1\|_{L^2(\mathbb{R}^3)} \leq \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2}} \|\partial^{J_1}\Omega^{J_2}(H^{\alpha'\beta'}(h_n^\kappa) - H^{\alpha'\beta'}(h_{n-1}^\kappa))\partial^{J'_1}\Omega^{J'_2}\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^{\kappa,n}\|_{L^2(\mathbb{R}^3)}.$$

When $|J_1| + |J_2| \leq d - 3$, $d - 1 \geq |J'_1| + |J'_2| \geq 2$,

$$\begin{aligned}
& \left\| \partial^{J_1} \Omega^{J_2} (H^{\alpha' \beta'}(h_n^\kappa) - H^{\alpha' \beta'}(h_{n-1}^\kappa)) \partial^{J'_1} \Omega^{J'_2} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n} \right\|_{L^2(\mathbb{R}^3)} \\
& \leq \left\| \partial^{J_1} \Omega^{J_2} (H^{\alpha' \beta'}(h_n^\kappa) - H^{\alpha' \beta'}(h_{n-1}^\kappa)) \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial^{J'_1} \Omega^{J'_2} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n} \right\|_{L^2(\mathbb{R}^3)} \\
& \leq C(\epsilon_0, d) \|h_n^\kappa - h_{n-1}^\kappa\|_{E_H^d} \|h_{\alpha\beta}^{\kappa, n}\|_{E_P^{d+1}} \\
& \leq C(\epsilon_0, d) A\epsilon \|h_n^\kappa - h_{n-1}^\kappa\|_{E_H^d} \\
& \leq C(\epsilon_0, d) A\epsilon \|S_n^\kappa - S_{n-1}^\kappa\|_{E_k^d},
\end{aligned}$$

Where (7.57) is applied.

When $d - 1 \geq |J_1| + |J_2| \geq d - 2$, we have $0 \leq |J'_1| + |J'_2| \leq 1$. Then recall that $d \geq 4$, $|J'_1| + |J'_2| + 1 \leq d - 2$. Then

$$\begin{aligned}
& \left\| \partial^{J_1} \Omega^{J_2} (H^{\alpha' \beta'}(h_n^\kappa) - H^{\alpha' \beta'}(h_{n-1}^\kappa)) \partial^{J'_1} \Omega^{J'_2} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n} \right\|_{L^2(\mathbb{R}^3)} \\
& \leq \left\| (1+r)^{-1} \partial^{J_1} \Omega^{J_2} (H^{\alpha' \beta'}(h_n^\kappa) - H^{\alpha' \beta'}(h_{n-1}^\kappa)) \right\|_{L^2(\mathbb{R}^3)} \left\| \partial^{J'_1} \Omega^{J'_2} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n} \right\|_{\mathcal{E}_{-1}} \\
& \leq C(\epsilon_0, d) \|h_n^\kappa - h_{n-1}^\kappa\|_{E_H^d} \left\| \partial^{J'_1} \Omega^{J'_2} \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n} \right\|_{X^2} \\
& \leq C(\epsilon_0, d) \|S_n^\kappa - S_{n-1}^\kappa\|_{E_k^d} \|h_{\alpha\beta}^{\kappa, n}\|_{E_P^d} \\
& \leq C(\epsilon_0, d) A\epsilon \|S_n^\kappa - S_{n-1}^\kappa\|_{E_k^d},
\end{aligned}$$

where (7.8), (7.56) and (7.7a) are applied. Note that since of the term with second order derivative $\partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^{\kappa, n}$, we can only bound the E^d norm, i.e. one order of regularity is lost.

The E^d term T_2 and T_3 are bounded by (7.52) and (8.15).

We should pay additional attention to the term T_4 :

$$\kappa^{-1} \left(V_h(\varrho_n^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^{\kappa, n}) - V_h(\varrho_{n-1}^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^{\kappa, n-1}) \right)$$

The E^d norm of this term can be bounded by $C(\epsilon_0, d)(A\epsilon)^2$. This is guaranteed by (7.44) and the assumption

$$\kappa^{-1/2} \|\varrho_k^\kappa\|_{X^d} \leq C(\epsilon_0, d) A\epsilon$$

deduced from (8.15). □

Lemma 8.10. — *Let $\{S_n^\kappa\}$ be the sequence constructed by (8.1) which satisfies the uniform bound condition (8.15). Then the following estimate holds for $|I_1| + |I_2| \leq d - 1$:*

$$(8.17) \quad \left\| [\partial^{I_1} \Omega^{I_2}, H^{\alpha' \beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}] (h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n}) \right\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0) A\epsilon \|h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n}\|_{E_P^d}.$$

Proof. — We perform the same calculation as in the proof of Lemma 8.2:

$$\begin{aligned}
& |[\partial^{I_1}\Omega^{I_2}, H^{\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}](h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})| \\
& \leq \sum_{\substack{J_1+J'_1=I_1 \\ |J_2|+|J'_2|\leq|I_2| \\ \alpha',\beta',\alpha'',a}} |\partial^{J'_1}\Omega^{J'_2}H^{\alpha'\beta'}| |\partial_{\alpha''}\partial_a\partial^{J_1}\Omega^{J_2}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})| \\
& + \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2 \\ |J'_1|+|J'_2|>0}} |\partial^{J'_1}\Omega^{J'_2}H^{\alpha'\beta'}| |\partial_{\alpha'}\partial_{\beta'}\partial^{J_1}\Omega^{J_2}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})| \\
& =: T_1 + T_2.
\end{aligned}$$

To estimate T_1 , we observe that since $|J'_1| + |J'_2| + |J_2| + |J'_2| \leq d - 2$, by (7.48) and (8.15),

$$\|\partial^{J'_1}\Omega^{J'_2}H^{\alpha'\beta'}(h_{\alpha\beta}^{\kappa,n})\|_{L^\infty(\mathbb{R}^3)} \leq C(\epsilon_0, d)A\epsilon$$

and also since $|J_1| + |J_2| \leq d - 2$:

$$\|\partial_{\alpha'}\partial_a\partial^{I_1}\Omega^{J_2}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})\|_{L^2(\mathbb{R}^3)} \leq C\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_P^d}$$

Then we see that

$$\|T_1\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0)A\epsilon\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_P^d}.$$

The estimate on T_2 is established in a bit complicated. We see that in T_2 , $|J_1| + |J_2| \leq d - 2$ so

$$\|\partial_{\alpha'}\partial_a\partial^{I_1}\Omega^{J_2}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})\|_{L^2(\mathbb{R}^3)} \leq C\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_P^d}$$

When $1 \leq |J'_1| + |J'_2| \leq d - 1$

$$\|\partial^{J'_1}\Omega^{J'_2}H^{\alpha'\beta'}(h_{\alpha\beta}^{\kappa,n})\|_{L^\infty(\mathbb{R}^3)} \leq C(\epsilon_0, d)A\epsilon,$$

where we used (7.48) combined with (8.15). \square

Now we are ready to estimate the term $\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_P^d}$.

Lemma 8.11. — *Let $\{S_n^\kappa\}$ be the sequence constructed by (8.1) which satisfies the uniform bound condition (8.15) with $d \geq 4$. Then the following estimate holds:*

$$\begin{aligned}
(8.18) \quad \frac{d}{dt}E_{g_n}^{d-1}(t, (h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})) & \leq C(\epsilon_0, d)A\epsilon\|S_n(t, \cdot) - S_{n-1}(t, \cdot)\|_{E_\kappa^d} \\
& + C(\epsilon_0, d)A\epsilon E_{g_n}^{d-1}(t, (h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})).
\end{aligned}$$

Proof. — We differentiate the equation (8.13a) with respect to a product $\partial^{I_1}\Omega^{I_2}$ with $|I_1| + |I_2| \leq d - 1$. Recall the relation of commutation, we get

$$\begin{aligned}
(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}\partial^{I_1}\Omega^{I_2}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}) & = \partial^{I_2}\Omega^{I_2}\left((H^{\alpha'\beta'}(h_{n-1}^\kappa) - H^{\alpha'\beta'}(h_n^\kappa))\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^{\kappa,n}\right) \\
& + \partial^{I_1}\Omega^{I_2}(F_H(S_n^\kappa) - F_H(S_{n-1}^\kappa)) \\
& - [\partial^{I_1}\Omega^{I_2}, H^{\alpha'\beta'}\partial_{\alpha'}\partial_{\beta'}](h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}).
\end{aligned}$$

Then we apply (7.20),

$$\begin{aligned} & \frac{d}{dt} E_{g_n}(t, \partial^{I_1} \Omega^{I_2} (h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n})) \\ & \leq C \|\partial^{I_1} \Omega^{I_2} D_H(S_n^\kappa, S_{n-1}^\kappa)\|_{L^2(\mathbb{R}^3)} + \|[\partial^{I_1} \Omega^{I_2}, H^{\alpha'\beta'}(h_n^\kappa) \partial_{\alpha'} \partial_{\beta'}](h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n})\|_{L^2(\mathbb{R}^3)} \\ & + \sum_{\alpha', \beta'} \|\nabla H^{\alpha'\beta'}(h_n^\kappa)\|_{L^\infty(\mathbb{R}^3)} E_{g_n}(t, \partial^{I_1} \Omega^{I_2} (h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n})). \end{aligned}$$

Note that by (7.48) and (8.15),

$$\|\nabla H^{\alpha'\beta'}(h_n^\kappa)\|_{L^\infty(\mathbb{R}^3)} \leq C(\epsilon_0, d) A \epsilon.$$

Then by Lemma 8.9 and 8.10,

$$\begin{aligned} \frac{d}{dt} E_{g_n}(t, \partial^{I_1} \Omega^{I_2} (h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n})) & \leq C(\epsilon_0, d) A \epsilon \|S_n(t, \cdot) - S_{n-1}(t, \cdot)\|_{E_\kappa^d} \\ & + C(\epsilon_0, d) A \epsilon \|h_{\alpha, \beta}^{\kappa, n+1} - h_{\alpha, \beta}^{\kappa, n}\|_{E_P^d} \\ & + C(\epsilon_0) A \epsilon E_{g_n}(t, \partial_x^{I_1} \Omega^{I_2} (h_{\alpha, \beta}^{\kappa, n+1} - h_{\alpha, \beta}^{\kappa, n})). \end{aligned}$$

Then by taking the sum over all the pair of multi-index (I_1, I_2) with $|I_1| + |I_2| \leq d - 1$, and observe that (by (8.15)):

$$\|h_{\alpha, \beta}^{\kappa, n+1} - h_{\alpha, \beta}^{\kappa, n}\|_{E_P^d} \leq C(\epsilon_0, d) E_{g_n}^{d-1}(t, (h_{\alpha, \beta}^{\kappa, n+1} - h_{\alpha, \beta}^{\kappa, n})),$$

and the desired result is proven. \square

The estimates on $(\phi_{n+1}^\kappa - \phi_n^\kappa)$ and $(\varrho_{n+1}^\kappa - \varrho_n^\kappa)$ are established in the same manner.

Lemma 8.12. — *Let $\{S_n^\kappa\}$ be the sequence constructed by (8.1) which satisfies the uniform bound condition (8.15) with $d \geq 4$. Then the following estimates hold:*

$$(8.19a) \quad \begin{aligned} & \frac{d}{dt} E_{g_n}^{d-1}(t, (\phi_{n+1}^\kappa - \phi_n^\kappa)) \\ & \leq C(\epsilon_0, d) A \epsilon \|S_n(t, \cdot) - S_{n-1}(t, \cdot)\|_{E_\kappa^d} + C(\epsilon_0, d) A \epsilon E_{g_n}^{d-1}(t, (\phi_{n+1}^\kappa - \phi_n^\kappa)), \end{aligned}$$

$$(8.19b) \quad \begin{aligned} & \frac{d}{dt} E_{g_n, \kappa^{-1/2}}^{d-1}(t, (\varrho_{n+1}^\kappa - \varrho_n^\kappa)) \\ & \leq C(\epsilon_0, d) A \epsilon \|S_n(t, \cdot) - S_{n-1}(t, \cdot)\|_{E_\kappa^d} + C(\epsilon_0, d) A \epsilon E_{g_n, \kappa^{-1/2}}^{d-1}(t, (\varrho_{n+1}^\kappa - \varrho_n^\kappa)). \end{aligned}$$

At this juncture, we can finally estimate the \mathcal{E}_{-1} norm of $(h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n})$.

Lemma 8.13. — *Let $\{S_n^\kappa\}$ be the sequence constructed by (8.1) which satisfies the uniform bound condition (8.15) with $d \geq 4$. Then the following estimate holds:*

$$(8.20) \quad \|h_{\alpha\beta}^{\kappa, n+1} - h_{\alpha\beta}^{\kappa, n}\|_{\mathcal{E}_{-1}} \leq C(\epsilon_0) t^2 A \epsilon \|S_n^\kappa - S_{n-1}^\kappa\|_{E_\kappa^d}.$$

Proof. — We are going to apply (7.18). To do so we need to establish the following estimates:

$$(8.21) \quad \|H^{\alpha'\beta'}(h_n^\kappa)\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})\|_{\mathcal{E}_{-1}} \leq C(\epsilon_0)A\epsilon\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_H^d},$$

$$(8.22) \quad \|D_H(S_n^\kappa, S_{n-1}^\kappa)\|_{\mathcal{E}_{-1}} \leq C(\epsilon_0)A\epsilon\|S_n^\kappa - S_{n-1}^\kappa\|_{E_\kappa^d}.$$

We see that (8.21) follows from (8.16) and (7.8). To establish (8.21), we see that

$$\begin{aligned} \|H^{\alpha'\beta'}(h_n^\kappa)\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})\|_{\mathcal{E}_{-1}} &\leq \|H^{\alpha'\beta'}(h_n^\kappa)\|_{\mathcal{E}_{-1}}\|\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n})\|_{L^\infty(\mathbb{R}^3)} \\ &\leq C(\epsilon_0)A\epsilon\|h_{\alpha\beta}^{\kappa,n+1} - h_{\alpha\beta}^{\kappa,n}\|_{E_P^4}, \end{aligned}$$

where (7.7) is applied. \square

Proof of Proposition 8.8. — We integrate (8.18), (8.19a) and (8.19b) and get the following estimates:

$$(8.23a) \quad E_{g_n}^{d-1}(t, (h_{\alpha,\beta}^{\kappa,n+1} - h_{\alpha,\beta}^{\kappa,n})) \leq (e^{C(\epsilon_0,d)A\epsilon t} - 1)\|S_n^\kappa - S_{n-1}^\kappa\|_{L^\infty([0,T^*];E_\kappa^d)}$$

$$(8.23b) \quad E_{g_n}^{d-1}(t, (\phi_{n+1}^\kappa - \phi_n^\kappa)) \leq (e^{C(\epsilon_0,d)A\epsilon t} - 1)\|S_n^\kappa - S_{n-1}^\kappa\|_{L^\infty([0,T^*];E_\kappa^d)}$$

$$(8.23c) \quad E_{g_n, \kappa^{-1/2}}^{d-1}(t, (\varrho_{n+1}^\kappa - \varrho_n^\kappa)) \leq (e^{C(\epsilon_0,d)A\epsilon t} - 1)\|S_n^\kappa - S_{n-1}^\kappa\|_{L^\infty([0,T^*];E_\kappa^d)}$$

Recall that the metric g_n is coercive with constant $C(\epsilon_0)$. We have

$$\begin{aligned} &\|S_{n+1}^\kappa(t, \cdot) - S_n^\kappa(t, \cdot)\|_{E_\kappa^d} \\ &\leq C(\epsilon_0, d) \max\{E_{g_n}^{d-1}(t, (h_{\alpha,\beta}^{\kappa,n+1} - h_{\alpha,\beta}^{\kappa,n})), E_{g_n}^{d-1}(t, (\phi_{n+1}^\kappa - \phi_n^\kappa)), E_{g_n, \kappa^{-1/2}}^{d-1}(t, (\varrho_{n+1}^\kappa - \varrho_n^\kappa))\}. \end{aligned}$$

Then we conclude with

$$(8.24) \quad \|S_{n+1}^\kappa - S_n^\kappa\|_{L^\infty([0,T^*];E_\kappa^d)} \leq C(\epsilon_0, d)(e^{C(\epsilon_0,d)A\epsilon T^*} - 1)\|S_n^\kappa - S_{n-1}^\kappa\|_{L^\infty([0,T^*];E_\kappa^d)}.$$

Then if we choose

$$T^* = \frac{\ln(1 + (2C(\epsilon_0))^{-1})}{C(\epsilon_0)A\epsilon},$$

then

$$\lambda := e^{C(\epsilon_0)A\epsilon t} - 1 = 1/2 < 1,$$

which satisfies the contraction condition. Furthermore, recall that in Proposition 8.7 we can take $A = \epsilon_0\epsilon^{-1/3}$ and $T \geq C'(\epsilon_0, d)\epsilon^{-1/3}$ when ϵ sufficiently small. So here we can also take $T^* = C''(\epsilon_0, d)\epsilon^{-1/3}$ for ϵ sufficiently small. This leads to the limit of $T^*(\epsilon)$ when $\epsilon \rightarrow 0^+$. \square

Now we apply Banach's fixed point theorem and see that $\{S_n^\kappa\}$ converges to a triple $S^\kappa := (h^\kappa, \phi^\kappa, \varrho^\kappa)$ in the sense of $L^\infty([0, T^*], E_\kappa^d)$. Then we will prove that S^κ is a solution of (8.1).

Proposition 8.14. — *When $d \geq 4$, the function S^κ constructed above is a solution of (8.1) and, furthermore, for $k \leq d$*

$$\begin{aligned} h_{\alpha\beta}^\kappa &\in C([0, T_0], E_H^d), & \partial_t^k h_{\alpha\beta}^\kappa &\in C([0, T_0], E_P^{d-k}), \\ \phi^\kappa &\in C([0, T_0], E_P^d), & \partial_t^k h_{\alpha\beta}^\kappa &\in C([0, T_0], E_P^{d-k}), \\ \varrho^\kappa &\in C([0, T_0], E_R^d), & \partial_t^k h_{\alpha\beta}^\kappa &\in C([0, T_0], E_R^{d-k}), \end{aligned}$$

with, in addition, $\|S^\kappa(t, \cdot)\|_{E_\kappa^d} \leq A\epsilon$ with $0 \leq t \leq T^*$.

Proof. — The proof is based on taking the limit in both side of (8.1). The convergence of $\{S_n^\kappa\}$ in sense of $L^\infty([0, T^*]; E_\kappa^d)$ can guarantee the convergence of both sides of (8.1). Recall that the sequence $\{\partial_t \partial_t h_{\alpha\beta}^{\kappa, n}\}$ also converges in the sense $L^\infty([0, T^*]; E^{d-2})$ and so does $\{\partial_t \partial_t \phi_n^\kappa\}$ and $\{\partial_t \partial_t \varrho_n^\kappa\}$.

The convergence of $\{S_n^\kappa\}$ in E_κ^d guarantees the following convergence (remark that $\kappa \leq 1$):

$$(8.25) \quad \begin{aligned} h_{\alpha\beta}^{\kappa, n} &\rightarrow h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E_H^d), & \nabla h_{\alpha\beta}^{\kappa, n} &\rightarrow \nabla h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E^{d-1}), \\ \phi_n^\kappa &\rightarrow \phi^\kappa \text{ in } L^\infty([0, T^*]; E_P^d), & \nabla \phi_n^\kappa &\rightarrow \nabla \phi^\kappa \text{ in } L^\infty([0, T^*]; E^{d-1}), \\ \varrho_n^\kappa &\rightarrow \varrho^\kappa \text{ in } L^\infty([0, T^*]; E_P^d) \cap L^\infty([0, T^*]; E^{d-1}), & \nabla \varrho_n^\kappa &\rightarrow \nabla \varrho^\kappa \text{ in } L^\infty([0, T^*]; E^{d-1}). \end{aligned}$$

Here, ∇ denotes the spacetime divergence. By Sobolev embedding ($d - 1 \geq 2$), $\{h_n^\kappa\}$, $\{\nabla \phi_n^\kappa\}$ and $\{\varrho_n^\kappa\}$ converges in $L^\infty([0, T^*] \times \mathbb{R}^3)$. Furthermore, we have

$$(8.26) \quad \begin{aligned} \partial_t \partial_x h_{\alpha\beta}^{\kappa, n} &\rightarrow \partial_t \partial_x h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}), & \partial_x \partial_x h_{\alpha\beta}^{\kappa, n} &\rightarrow \partial_x \partial_x h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}) \\ \partial_t \partial_x \phi_n^\kappa &\rightarrow \partial_t \partial_x \phi^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}), & \partial_x \partial_x \phi_n^\kappa &\rightarrow \partial_x \partial_x \phi^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}) \\ \partial_t \partial_x h_{\alpha\beta}^{\kappa, n} &\rightarrow \partial_t \partial_x h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}), & \partial_x \partial_x h_{\alpha\beta}^{\kappa, n} &\rightarrow \partial_x \partial_x h_{\alpha\beta}^\kappa \text{ in } L^\infty([0, T^*]; E^{d-2}). \end{aligned}$$

These convergence properties are sufficient to guarantee the convergence of both side of (8.1) since both side depend linearly the terms with second order derivatives. And the lower order terms converge in L^∞ sense. \square

Proof of Theorem 6.3. — We have checked that the triple S^κ is a local solution of (6.6). Furthermore, we notice that the lower bound of life-span-time T^* constructed in Proposition 8.14 does not depend on κ . The estimates are established by taking the limit of the (8.15). \square

9. Comparing the $f(R)$ theory to the classical theory

9.1. Statement of the main estimate

In this section, we compare the solutions given by the $f(R)$ theory with the solutions of the classical Einstein theory. We denote by $S^0 := (h_{\alpha\beta}, \phi)$ the triple determined by the following Cauchy problem:

$$(9.1a) \quad (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h)) \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta} = F_{\alpha\beta}(h, \partial h, \partial h) - 16\pi \partial_{\alpha'} \phi \partial_{\beta'} \phi,$$

$$(9.1b) \quad (m^{\alpha\beta'} + H^{\alpha\beta'}(h))\partial_{\alpha'}\partial_{\beta'}\phi = 0$$

with initial data

$$(9.2) \quad \begin{aligned} h_{\alpha\beta}(0, x) &= h_{0\alpha\beta}, & \partial_t h_{\alpha\beta}(0, x) &= h_{1\alpha\beta}, \\ \phi(0, x) &= \phi_0, & \partial_t \phi(0, x) &= \phi_1. \end{aligned}$$

This limiting problem is defined by replacing ρ^κ by 0 in our formulation (6.3). As before, if the initial data satisfies the corresponding constraint conditions, then $g_{\alpha\beta} = m_{\alpha\beta} + h_{\alpha\beta}$ and ϕ satisfy the classical Einstein's field equation coupled with the massless scalar field ϕ . For the convenience of discussion we introduce the norm

$$\|S_0\|_{X_0^{d+1}} := \max \left(\|h_{0\alpha\beta}\|_{X_H^{d+1}}, \|h_{1\alpha\beta}\|_{X^d}, \|\phi_0\|_{X_P^{d+1}}, \|\phi_1\|_{X^d} \right).$$

Proposition 9.1 (Local existence theory for the classical gravity system)

Suppose that $(h_{0\alpha\beta}, h_{1\alpha\beta}) \in X_H^{d+1} \times X^d$ and $(\phi_0, \phi_1) \in X_P^{d+1} \times X^d$ and with $d \geq 4$, and denote by $S_0 = (h_0, h_1, \phi_0, \phi_1)$ and assume that for a sufficiently small ϵ_0

$$\|S_0\|_{X_0^{d+1}} \leq \epsilon \leq \epsilon_0 \leq 1.$$

Then there exist positive constants A, T^* determined from ϵ_0, ϵ and d such that the Cauchy problem (9.1) with initial data S_0 has a unique solution $(h_{\alpha\beta}, \phi)$ in the time interval $[0, T^*]$. Here

$$h_{\alpha\beta} \in C([0, T^*]; E_H^d), \quad \phi \in C([0, T^*]; E_P^d).$$

When $\epsilon \rightarrow 0^+$, we can take

$$\lim_{\epsilon \rightarrow 0^+} T^* = +\infty.$$

Furthermore, the local solution satisfies the following estimates in the time interval $[0, T^*]$:

$$(9.3) \quad \|h_{\alpha\beta}\|_{E_H^d} + \|\phi\|_{E_P^d} \leq A\epsilon.$$

The proof is similar to that of Theorem 6.3: we make an iteration and estimate the sequence constructed by this iteration and we prove that with suitable choice of (A, T^*) , this sequence is contracting. The details of the argument are omitted.

Let $S_0 = (h_{0\alpha\beta}^0, h_{1\alpha\beta}^0, \phi_0^0, \phi_1^0)$ be an initial data which satisfies the Einstein's constraint equation (2.29) and $S_1 = (h_{0\alpha\beta}, h_{1\alpha\beta}, \phi_0, \phi_1, \varrho_0, \varrho_1)$ be an initial data which satisfies the nonlinear constraint equations (5.7) and (5.8). Define the following function $\mathcal{D}_\kappa(S_0, S_1)$:

$$\begin{aligned} \mathcal{D}_\kappa^d(S_0, S_1) := \max \{ & \|h_{0\alpha\beta}^0 - h_{0\alpha\beta}\|_{X_H^{d+1}}, \|h_{1\alpha\beta}^0 - h_{1\alpha\beta}\|_{X^d}, \|\phi_0^0 - \phi_0\|_{X_P^{d+1}}, \\ & \|\phi_1^0 - \phi_1\|_{X_P^{d+1}}, \|\varrho_0\|_{X_P^{d+1}}, \|\varrho_1\|_{X^d}, \kappa^{-1/2} \|\varrho_0\|_{X^d} \}. \end{aligned}$$

Denote by

$$S^0(t) = (h_{\alpha\beta}^0(t), \phi^0(t)) \in C([0, T^*]; X_H^d) \cap C^1([0, T]; X_H^{d-1}) \times C([0, T^*]; X_P^d) \cap C^1([0, T]; X_P^{d-1}).$$

the local solution of Cauchy problem (9.1) with initial data $S^0(0) = S_0$, and

$$\begin{aligned} S^\kappa(t) = (h_{\alpha\beta}^\kappa, \phi^\kappa, \varrho^\kappa) \in & C([0, T^*]; X_H^d) \cap C^1([0, T]; X_H^{d-1}) \times C([0, T^*]; X_P^d) \cap C^1([0, T]; X_P^{d-1}) \\ & \times C([0, T^*]; X_P^d) \cap C^1([0, T]; X_P^{d-1}) \cap C([0, T^*]; X_R^{d-1}). \end{aligned}$$

We introduce the “distance” from S^0 to S^κ :

$$\mathcal{D}^d(S^0, S^\kappa)(t) := \sum_{\alpha\beta} \|h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa\|_{E_H^d} + \|\phi^0 - \phi^\kappa\|_{E_P^d}$$

and we are ready to state the key estimate derived in the present work.

Theorem 9.2 (Comparison estimate). — *There exists a positive constant ϵ_0 such that if*

$$\max\{\|S_1\|_{X_\kappa^{d+1}}, \|S_0\|_{X_0^{d+1}}\} \leq \epsilon \leq \epsilon_0 \leq 1$$

(with $d \geq 4$), then in the common interval of existence $[0, T^*]$ (which depends only on ϵ, ϵ_0 and d), the following estimates hold:

$$(9.4) \quad \mathcal{D}^{d-1}(S^0, S^\kappa)(t) \leq C(\epsilon_0, d) (\mathcal{D}^{d-1}(S^0, S^\kappa)(0) + (E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2 + \kappa^{1/2}(A\epsilon)^3).$$

9.2. Derivation of the comparison estimate

The proof of Theorem 9.2 requires better estimates on $\|\varrho_{n+1}^\kappa\|_{X^{d-1}}$. First we establish an improved bound on the L^2 norm of $\partial_t \partial^{I_1} \Omega^{I_2} \varrho^\kappa$. The following lemma is immediate from (6.16c).

Lemma 9.3. — *Let $S^\kappa = (h_{\alpha\beta}^\kappa, \phi^\kappa, \varrho^\kappa)$ be the solution of Cauchy problem (6.6) with $d \geq 4$. Then the following estimate holds for all $|I_1| + |I_2| \leq d - 2$:*

$$(9.5) \quad \kappa^{-1/2} \|\partial_t \partial^{I_1} \Omega^{I_2} \varrho^\kappa\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d) A\epsilon.$$

Lemma 9.4. — *There exists a positive constant ϵ_0 such that if (8.2) holds for $d \geq 4$ and $A\epsilon \leq \epsilon_0$, then*

$$(9.6) \quad \|\varrho^\kappa(t, \cdot)\|_{E^{d-2}}^2 \leq \kappa^{3/2} C(\epsilon_0, d) t (A\epsilon)^3 + C(\epsilon_0, d) \kappa (E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2.$$

Proof. — We derive an energy-type estimate by differentiating (6.6c) with respect to $\partial^{I_1} \Omega^{I_2}$ with $|I_1| + |I_2| \leq d - 2$. Similarly as in the proof of Lemma 7.11, we have

$$\begin{aligned} & \partial_t \partial^{I_1} \Omega^{I_2} \varrho^\kappa (g^{\alpha\beta} \partial_\alpha \partial_\beta \partial^{I_1} \Omega^{I_2} \varrho^\kappa - 3\kappa^{-1} \partial^{I_1} \Omega^{I_2} \varrho^\kappa) \\ &= \frac{1}{2} \partial_0 (g^{00} (\partial_0 \partial^{I_1} \Omega^{I_2} \varrho^\kappa)^2 - g^{ab} \partial_a \partial^{I_1} \Omega^{I_2} \varrho^\kappa \partial_b \partial^{I_1} \Omega^{I_2} \varrho^\kappa) - \frac{1}{2} \partial_0 ((3\kappa)^{-1/2} \partial^{I_1} \Omega^{I_2} \varrho^\kappa)^2 \\ & \quad + \partial_a (g^{a\beta} \partial^{I_1} \Omega^{I_2} \varrho^\kappa \partial_\beta \partial^{I_1} \Omega^{I_2} \varrho^\kappa) + \frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha \partial^{I_1} \Omega^{I_2} \varrho^\kappa \partial_\beta \partial^{I_1} \Omega^{I_2} \varrho^\kappa \\ & \quad - \partial_\alpha g^{\alpha\beta} \partial_0 \partial^{I_1} \Omega^{I_2} \varrho^\kappa \partial_\beta \partial^{I_1} \Omega^{I_2} \varrho^\kappa. \end{aligned}$$

For simplicity, we set $v = \partial^{I_1} \Omega^{I_2} \varrho^\kappa$ and obtain

$$\begin{aligned} & \partial_t v (\partial^{I_1} \Omega^{I_2} F_R(h^\kappa, \phi^\kappa, \varrho^\kappa)) \\ &= \frac{1}{2} \partial_t (g^{00} (\partial_t v)^2 - g^{ab} \partial_a v \partial_b v) - \frac{1}{2} \partial_t ((3\kappa)^{-1/2} v)^2 + \partial_a (g^{a\beta} \partial_t v \partial_\beta v) \\ & \quad + \frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha v \partial_\beta v - \partial_\alpha g^{\alpha\beta} \partial_t v \partial_\beta v. \end{aligned}$$

Integrating this equation in the region $[0, t] \times \mathbb{R}^3$ and using Stokes' formula, we obtain

$$(9.7) \quad \begin{aligned} & \int_0^t \int_{\mathbb{R}^3} \partial_t v (\partial^{I_1} \Omega^{I_2} F_R(h^\kappa, \phi^\kappa, \varrho^\kappa)) dx dt \\ &= \frac{1}{2} (E_{g, \kappa^{-1/2}}(t, \partial^{I_1} \Omega^{I_2} \varrho^\kappa))^2 - \frac{1}{2} (E_{g, \kappa^{-1/2}}(0, \partial^{I_1} \Omega^{I_2} \varrho^\kappa))^2 \\ & \quad + \int_0^t \int_{\mathbb{R}^3} \left(\frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha v \partial_\beta v - \partial_\alpha g^{\alpha\beta} \partial_t v \partial_\beta v \right) dx dt \end{aligned}$$

Then by (7.44), (7.49), (9.5) and (6.16), we see that

$$\left| \int_0^t \int_{\mathbb{R}^3} \partial_t v (\partial^{I_1} \Omega^{I_2} F_R(h^\kappa, \phi^\kappa, \varrho^\kappa)) dx dt \right| \leq C(\epsilon_0, d) t (A\epsilon)^3 \kappa^{1/2}$$

then, by (6.16) and since $d \geq 4$,

$$\left| \partial_\alpha g^{\alpha\beta} \right|_{L^\infty(\mathbb{R}^3)} \leq C(\epsilon_0) \|h_{\alpha\beta}\|_{E_P^{d-1}} \leq C(\epsilon_0) A\epsilon,$$

and thus

$$\left| \int_0^t \int_{\mathbb{R}^3} \left(\frac{1}{2} \partial_t g^{\alpha\beta} \partial_\alpha v \partial_\beta v - \partial_\alpha g^{\alpha\beta} \partial_t v \partial_\beta v \right) dx dt \right| \leq C(\epsilon_0, d) t (A\epsilon)^3 \kappa.$$

Combining the above two estimates together with (9.7), we find (with $0 < \kappa \leq 1$)

$$(E_{g, \kappa^{-1/2}}(t, \partial^{I_1} \Omega^{I_2} \varrho^\kappa))^2 \leq (E_{g, \kappa^{-1/2}}(0, \partial^{I_1} \Omega^{I_2} \varrho^\kappa))^2 + C(\epsilon_0, d) t (A\epsilon)^3 \kappa^{1/2}.$$

In view of the definition of $E_{g, \kappa^{-1/2}}(t, u)$, this leads us to

$$\kappa^{-1} \|\partial^{I_1} \Omega^{I_2} \varrho^\kappa(t, \cdot)\|_{L^2(\mathbb{R}^3)}^2 \leq C(\epsilon_0, d) (E_{g, \kappa^{-1/2}}(0, \partial^{I_1} \Omega^{I_2} \varrho^\kappa))^2 + C(\epsilon_0, d) t (A\epsilon)^3 \kappa^{1/2}.$$

□

Proof of Theorem 9.2. — By taking the difference of (6.6a) and (9.1a), and the difference of (6.6b) and (9.1b), we obtain

$$(9.8a) \quad \begin{aligned} & (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^0)) \partial_{\alpha'} \partial_{\beta'} ((h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)) \\ &= -(H^{\alpha'\beta'}(h^0) - H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \partial_{\beta'} h_{\alpha\beta}^\kappa + (F_{\alpha\beta}(h^0, \partial h^0, \partial h^0) - F_{\alpha\beta}(h^\kappa, \partial h^\kappa, \partial h^\kappa)) \\ & \quad - 16\pi (\partial_\alpha \phi^0 \partial_\beta \phi^0 - \partial_\alpha \phi^\kappa \partial_\beta \phi^\kappa) + 12 \partial_\alpha \varrho^\kappa \partial_\beta \varrho^\kappa + \kappa^{-1} V_h(\varrho^\kappa) (m_{\alpha\beta} + h_{\alpha\beta}^\kappa) \end{aligned}$$

and

$$(9.8b) \quad \begin{aligned} & (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^0)) \partial_{\alpha'} \partial_{\beta'} (\phi^0 - \phi^\kappa) \\ &= -(H^{\alpha'\beta'}(h^0) - H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \partial_{\beta'} \phi^\kappa - 2(m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^\kappa)) \partial_{\alpha'} \phi^\kappa \partial_{\beta'} \varrho^\kappa. \end{aligned}$$

We are going to derive (9.4) from these two equations. The proof is quite similar to that of Lemmas 8.11 and 8.12.

Step I. L^2 -Type estimates. Let us begin with the E_P^{d-2} norm of $(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)$. Let (I_1, I_2) be a pair of multi-indices with $|I_1| + |I_2| \leq d - 2$. We differentiate (9.8) with respect to $\partial^{I_1}\Omega^{I_2}$:

$$\begin{aligned}
& (m^{\alpha'\beta'} + H^{\alpha'\beta'}(h^0))\partial_{\alpha'}\partial_{\beta'}\partial^{I_1}\Omega^{I_2}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) \\
&= -\partial^{I_1}\Omega^{I_2}\left((H^{\alpha'\beta'}(h^0) - H^{\alpha'\beta'}(h^\kappa))\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^\kappa\right) + \partial^{I_1}\Omega^{I_2}(F_{\alpha\beta}(h^0, \partial h^0, \partial h^0) - F_{\alpha\beta}(h^\kappa, \partial h^\kappa, \partial h^\kappa)) \\
&\quad - 16\pi\partial^{I_1}\Omega^{I_2}(\partial_\alpha\phi^0\partial_\beta\phi^0 - \partial_\alpha\phi^\kappa\partial_\beta\phi^\kappa) - [\partial^{I_1}\Omega^{I_2}, H^{\alpha'\beta'}(h^0)\partial_{\alpha'}\partial_{\beta'}](h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) \\
&\quad + 12\partial^{I_2}\Omega^{I_2}(\partial_\alpha\varrho^\kappa\partial_\beta\varrho^\kappa) + \kappa^{-1}\partial_x^{I_1}\Omega^{I_2}\left(V_h(\varrho^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^\kappa)\right) \\
&=: T_1 + T_2 + T_3 + T_4 + T_5 + T_6.
\end{aligned}$$

Then combined with (7.20),

$$\begin{aligned}
& \frac{d}{dt}E_g(t, \partial^{I_1}\Omega^{I_2}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)) \\
(9.9) \quad & \leq C \sum_{i=1}^6 \|T_i\|_{L^2} + C \sum_{\alpha, \beta} \|H^{\alpha'\beta'}(h^0)\|_{L^\infty(\mathbb{R}^3)} E_g(t, \partial^{I_1}\Omega^{I_2}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)).
\end{aligned}$$

We will need to control the L^2 norm of these T_i for $i = 1, \dots, 6$. The term T_i for $i = 1, 2, 3, 4$ can be bounded as follows:

$$\begin{aligned}
(9.10) \quad & \|T_i\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)A\epsilon(\|h^0 - h^\kappa\|_{E_H^{d-1}} + \|h^0 - h^\kappa\|_{E_P^{d-1}}\|\phi^0 - \phi^\kappa\|_{E_P^{d-1}}) \\
& \leq C(\epsilon_0, d)A\epsilon\mathcal{D}^{d-1}(S^0, S^\kappa)(t).
\end{aligned}$$

The proof is exactly the same to the one of (8.16) and (8.17) and we omit the details.

The *key terms* T_5 and T_6 are bounded as follows:

$$\begin{aligned}
(9.11) \quad & \|T_5\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)\kappa(A\epsilon)^2, \\
& \|T_6\|_{L^2(\mathbb{R}^3)} \leq C(\epsilon_0, d)\kappa^{1/2}(A\epsilon)^3 + C(\epsilon_0, d)(E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2.
\end{aligned}$$

The estimates on T_5 and T_6 is related to the refined estimates (9.5) and (9.6). More precisely, T_5 is estimated by (7.44), (6.16b) and (9.5). The term T_6 is estimated by (7.44) and (9.6).

Next, we combine together the above estimates on T_i and observe that

$$\|H^{\alpha'\beta'}(h^0)\|_{L^\infty} \leq C(\epsilon_0)A\epsilon$$

and

$$E_g(t, \partial^{I_1}\Omega^{I_2}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)) \leq C(\epsilon_0, d)\mathcal{D}^{d-1}(S^0, S^\kappa)(t).$$

We can thus deduce from (9.9) that

$$\begin{aligned}
(9.12) \quad & \frac{d}{dt}E_g^{d-2}(t, (h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)) \leq C(\epsilon_0, d)A\epsilon\mathcal{D}^{d-1}(S^0, S^\kappa)(t) + C(\epsilon_0, d)\kappa^{1/2}(A\epsilon)^3 \\
& \quad + C(\epsilon_0, d)(E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2,
\end{aligned}$$

where we recall the definition

$$\mathcal{D}^d(S^0, S^\kappa)(t) := \sum_{\alpha\beta} \|h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa\|_{E_H^d} + \|\phi^0 - \phi^\kappa\|_{E_P^d}.$$

The estimate on the norm $\|\phi^0 - \phi^1\|_{X_E^d}$ is similar to that of $h^0 - h^\kappa$ (even simpler). We claim that the following estimate on the right-hand side of (9.8b):

$$(9.13a) \quad \|(H^{\alpha'\beta'}(h^0) - H^{\alpha'\beta'}(h^\kappa))\partial_{\alpha'}\partial_{\beta'}\phi^\kappa\|_{E^{d-2}} \leq C(\epsilon_0)A\epsilon\|h^0 - h^\kappa\|_{E_H^{d-1}},$$

$$(9.13b) \quad \|(m^{\alpha\beta} + H^{\alpha\beta}(h^\kappa))\partial_\alpha\phi^\kappa\partial_\beta\varrho^\kappa\|_{E^{d-2}} \leq C(\epsilon_0)\kappa^{1/2}(A\epsilon)^2.$$

The first can be proved exactly as in the proof of (9.12). The second one is proven as follows: for any (I_1, I_2) with $|I_1| + |I_2| \leq d - 2$,

$$\begin{aligned} & \left\| \partial^{I_2} \Omega^{I_2} ((m^{\alpha\beta} + H^{\alpha\beta}(h^\kappa)) \partial_\alpha \phi^\kappa \partial_\beta \varrho^\kappa) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq \sum_{\substack{J_1 + J'_1 = I_1 \\ J_2 + J'_2 = I_2}} \left\| \partial^{J_1} \Omega^{J_2} (m^{\alpha\beta} + H^{\alpha\beta}(h^\kappa)) \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial^{J'_1} \Omega^{J'_2} (\partial_\alpha \phi^\kappa \partial_\beta \varrho^\kappa) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(\epsilon_0, d) \sum_{\substack{\alpha, \beta \\ |J'_1| + |J'_2| \leq d-2}} \left\| \partial^{J'_1} \Omega^{J'_2} (\partial_\alpha \phi^\kappa \partial_\beta \varrho^\kappa) \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(\epsilon_0, d) \sum_{\substack{\alpha, \beta \\ |J'_1| + |J'_2| \leq d-2}} \sum_{\substack{K_1 + K'_1 = J'_1 \\ K_2 + K'_2 = J'_2}} \left\| \partial^{K_1} \Omega^{K_2} \partial_\alpha \phi^\kappa \partial_x^{K'_1} \Omega^{K'_2} \partial_\beta \varrho^\kappa \right\|_{L^2(\mathbb{R}^3)} \end{aligned}$$

Then, when $|K_1| + |K_2| \leq d - 3$,

$$\begin{aligned} \left\| \partial^{K_1} \Omega^{K_2} \partial_\alpha \phi^\kappa \partial_x^{K'_1} \Omega^{K'_2} \partial_\beta \varrho^\kappa \right\|_{L^2(\mathbb{R}^3)} & \leq \left\| \partial^{K_1} \Omega^{K_2} \partial_\alpha \phi^\kappa \right\|_{L^\infty(\mathbb{R}^3)} \left\| \partial^{K'_1} \Omega^{K'_2} \partial_\beta \varrho^\kappa \right\|_{L^2(\mathbb{R}^3)} \\ & \leq C(\epsilon_0, d) A \epsilon \|\varrho^\kappa\|_{E^{d-1}} \\ & \leq C(\epsilon_0, d) \kappa^{1/2} (A \epsilon)^2. \end{aligned}$$

When $|K_1| + |K_2| = d - 2$ and $K'_1 = K'_2 = 0$, recall that $d \geq 4$:

$$\begin{aligned} \left\| \partial^{K_1} \Omega^{K_2} \partial_\alpha \phi^\kappa \partial_x^{K'_1} \Omega^{K'_2} \partial_\beta \varrho^\kappa \right\|_{L^2(\mathbb{R}^3)} & \leq \left\| \partial^{K_1} \Omega^{K_2} \partial_\alpha \phi^\kappa \right\|_{L^2(\mathbb{R}^3)} \left\| \partial_\beta \varrho^\kappa \right\|_{L^\infty(\mathbb{R}^3)} \\ & \leq C(\epsilon_0, d) A \epsilon \|\varrho^\kappa\|_{E^{d-1}} \leq C(\epsilon_0, d) \kappa^{1/2} (A \epsilon)^2 \end{aligned}$$

So we conclude with (9.13b), and combined with (7.20),

$$(9.14) \quad \frac{d}{dt} E_g^{d-2}(t, (\phi^0 - \phi^\kappa)) \leq C(\epsilon_0, d) A \epsilon \mathcal{D}^{d-1}(S^0, S^\kappa)(t) + C(\epsilon_0, d) \kappa^{1/2} (A \epsilon)^2.$$

Step II. Estimate of the \mathcal{E}_{-1} norm. To do so we rewrite the equation (9.8a) into the following form:

$$\begin{aligned}
(9.15) \quad \square((h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)) &= -H^{\alpha'\beta'}(h^0)\partial_{\alpha'}\partial_{\beta'}((h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) - (H^{\alpha'\beta'}(h^0) - H^{\alpha'\beta'}(h^\kappa))\partial_{\alpha'}\partial_{\beta'}h_{\alpha\beta}^\kappa \\
&\quad + (F_{\alpha\beta}(h^0, \partial h^0, \partial h^0) - F_{\alpha\beta}(h^\kappa, \partial h^\kappa, \partial h^\kappa)) - 16\pi(\partial_\alpha\phi^0\partial_\beta\phi^0 - \partial_\alpha\phi^\kappa\partial_\beta\phi^\kappa) \\
&\quad + 12\partial_\alpha\varrho^\kappa\partial_\beta\varrho^\kappa + \kappa^{-1}V_h(\varrho^\kappa)(m_{\alpha\beta} + h_{\alpha\beta}^\kappa) \\
&=: T_0 + T_1 + T_2 + T_3 + T_4 + T_5 + T_6.
\end{aligned}$$

By (7.18) we need to control the \mathcal{E}_{-1} norm of the terms T_i for $i = 1, \dots, 6$. By (7.8), we need only to control the X^2 norm of these terms. Recall the condition $d \geq 4$, then $d - 2 \geq 2$. So we only need to control the X^{d-2} norm of these terms. Note that in the **Step I** we have already controlled this norm for the terms T_i with $i \geq 1$. Now we only need to control the X^2 norm of T_0 . Let (I_1, I_2) be a pair of multi-indices with $|I_1| + |I_2| \leq 2$. Then, we have

$$\begin{aligned}
&\|\partial_x^{I_1}\Omega^{I_2}(H^{\alpha'\beta'}(h^0)\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa))\|_{L^2(\mathbb{R}^3)} \\
&\leq \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2}} \|\partial_x^{J_1}\Omega^{J_2}(H^{\alpha'\beta'}(h^0))\partial_x^{J'_1}\Omega^{J'_2}\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)\|_{L^2(\mathbb{R}^3)} \\
&\leq \sum_{\substack{J_1+J'_1=I_1 \\ J_2+J'_2=I_2}} \|\partial_x^{J_1}\Omega^{J_2}(H^{\alpha'\beta'}(h^0))\|_{L^\infty(\mathbb{R}^3)} \|\partial_x^{J'_1}\Omega^{J'_2}\partial_{\alpha'}\partial_{\beta'}(h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa)\|_{L^2(\mathbb{R}^3)} \\
&\leq C(\epsilon_0)\|h^0\|_{X^4}\|h^0 - h^\kappa\|_{E_P^3} \leq C(\epsilon_0)A\epsilon\|h^0 - h^\kappa\|_{E_P^{d-1}}.
\end{aligned}$$

Then by (7.18), the following estimate on \mathcal{E}_{-1} norm holds:

$$\begin{aligned}
(9.16) \quad &\|h_{\alpha\beta}^0(t, \cdot) - h_{\alpha\beta}^\kappa(t, \cdot)\|_{\mathcal{E}_{-1}} \\
&\leq C(\epsilon_0)tA\epsilon \int_0^t \mathcal{D}^{d-1}(S^0, S^\kappa)(\tau)d\tau + C(1+t)(\|h_{0\alpha\beta}^0 - h_{0\alpha\beta}^\kappa\|_{X_H^{d-1}}).
\end{aligned}$$

Step III: Conclusion. Now by integrating (9.12) and (9.15), we get the following estimate:

$$\begin{aligned}
(9.17a) \quad E_g^{d-2}(t, h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) &\leq E_g^{d-2}(0, h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) + C(\epsilon_0, d)A\epsilon \int_0^t \mathcal{D}^{d-1}(S^0, S^\kappa)(\tau)d\tau \\
&\quad + C(\epsilon_0, d)\kappa^{1/2}(A\epsilon)^3t + C(\epsilon_0, d)t(E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2,
\end{aligned}$$

$$\begin{aligned}
(9.17b) \quad E_g^{d-2}(t, \phi^0 - \phi^\kappa) &\leq E_g^{d-2}(0, (\phi^0 - \phi^\kappa)) + C(\epsilon_0, d)A\epsilon \int_0^t \mathcal{D}^{d-1}(S^0, S^\kappa)(\tau)d\tau \\
&\quad + C(\epsilon_0, d)\kappa^{1/2}(A\epsilon)^2t.
\end{aligned}$$

Recall that g is coercive with constant $C(\epsilon_0)$ when ϵ_0 is sufficiently small. Then, we have

$$(9.18) \quad \begin{aligned} & \mathcal{D}^{d-1}(S^0 - S^\kappa)(t) \\ & \leq C(\epsilon_0, d) \sum_{\alpha, \beta} E_g^{d-2}(t, h_{\alpha\beta}^0 - h_{\alpha\beta}^\kappa) + C(\epsilon_0, d) \sum_{\alpha, \beta} \|h_{\alpha\beta}^0(t, \cdot) - h_{\alpha\beta}^\kappa(t, \cdot)\|_{\mathcal{E}_{-1}} \\ & \quad + C(\epsilon_0, d) E_g^{d-2}(t, \phi^0 - \phi^\kappa) \leq C(\epsilon_0, d)^2 \mathcal{D}^{d-1}(S^0 - S^\kappa)(t). \end{aligned}$$

Then by combining (9.16), (9.17a), (9.17b) and (9.18), the following estimate holds:

$$(9.19) \quad \begin{aligned} \mathcal{D}^{d-1}(S^0, S^\kappa)(t) & \leq C(\epsilon_0, d)(1 + T^*) \mathcal{D}^{d-1}(S^0, S^\kappa)(0) + C(\epsilon_0, d) T^* (E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2 \\ & \quad + C(\epsilon_0, d) T^* \kappa^{1/2} (A\epsilon)^3 \\ & \quad + C(\epsilon_0, d)(1 + T^*) A\epsilon \int_0^t \mathcal{D}^{d-1}(S^0, S^\kappa)(\tau) d\tau, \end{aligned}$$

which yields

$$\begin{aligned} & \mathcal{D}^{d-1}(S^0, S^\kappa)(t) \\ & \leq C(\epsilon_0, d)(1 + T^*) \left(\mathcal{D}^{d-1}(S^0, S^\kappa)(0) + (E_{g, \kappa^{-1/2}}^{d-2}(0, \varrho^\kappa))^2 + \kappa^{1/2} (A\epsilon)^3 \right) e^{C(\epsilon_0, d) A\epsilon (1 + T^*) t}. \end{aligned}$$

□

Acknowledgments

The authors were partially supported by the Agence Nationale de la Recherche (ANR) through the grant 06-2-134423 and ANR SIMI-1-003-01. Part of this research was done when the first author (PLF) was a visiting professor for the Fall Semester 2013 at the Mathematical Sciences Research Institute (Berkeley) and was supported by the National Science Foundation under Grant No. 0932078 000.

References

- [1] A. BACHELOT, Problème de Cauchy global pour des systèmes de Dirac-Klein-Gordon, *Ann. Inst. Henri Poincaré* 48 (1988), 387–422.
- [2] A. BACHELOT, Asymptotic completeness for the Klein-Gordon equation on the Schwarzschild metric, *Ann. Inst. Henri Poincaré: Phys. Théor.* 61 (1994), 411–441.
- [3] N. BEDJAOU, P.G. LEFLOCH, J. MARTIN-GARCIA, AND J. NOVAK, Existence of naked singularities for the Brans-Dicke theory. An analytical and numerical study, *Class. Quantum Gravity* 27 (2010), 245010.
- [4] C. BRANS AND R.H. DICKE, Mach principle and a relativistic theory of gravitation, *Phys. Rev.* 124 (1961), 925–935.
- [5] F. BRISCESE, E. ELIZALDE, S. NOJIRI, AND S.D. ODINTSOV, Phantom scalar dark energy as modified gravity: Understanding the origin of the Big Rip singularity, *Phys. Lett. B* 646 (2007), 105–111.

- [6] H.A. BUCHDAHL, Nonlinear Lagrangians and cosmological theory, *Monthly Notices Royal Astr. Soc.* 150 (1970), 1–8.
- [7] S. CAPOZZIELLO, S. NOJIRI, S.D. ODINTSOV, AND A. TROISI, Cosmological viability of $f(R)$ -gravity as an ideal fluid and its compatibility with a matter dominated phase, *Phys. Lett. B* 639 (2006), 135–143.
- [8] S. CAPOZZIELLO, S. NOJIRI, AND S.D. ODINTSOV, Dark Energy: the equation of state description versus scalar-tensor or modified gravity, *Phys. Lett. B* 634 (2006), 93–100.
- [9] S. CAPOZZIELLO AND S. VIGNOLO, The Cauchy problem for metric-affine $f(R)$ -gravity in the presence of perfect-fluid matter, *Class. Quantum Grav.* 26, 175013.
- [10] Y. CHOQUET-BRUHAT, *General relativity and the Einstein equations*, Oxford Math. Monograph, Oxford University Press, 2009
- [11] J.-M. DELORT, Existence globale et comportement asymptotique pour l'équation de Klein-Gordon quasi-linéaire à données petites en dimension 1, *Ann. Sci. École Norm. Sup.* 34 (2001), 1–61.
- [12] J.-M. DELORT, D. FANG, AND R. XUE, Global existence of small solutions for quadratic quasilinear Klein-Gordon systems in two space dimensions, *J. Funct. Anal.* 211 (2004), 288–323.
- [13] R. GARCÍA-SALCEDO, R. GONZALEZ, C. MORENO, Y. NAPOLES, Y. LEYVA, AND I. QUIROS, Asymptotic properties of a supposedly regular (Dirac-Born-Infeld) modification of general relativity, *J. Cosm. Astr. Physics* 2 (2010), 027.
- [14] S.W. HAWKING AND G.F.R. ELLIS, *The large scale structure of spacetime*, Cambridge Monographs on Math. Phys., Vol. 1, Cambridge Univ. Press, Cambridge, 1973.
- [15] G. HOLZEGEL AND C.M. WARNICK, The Einstein-Klein-Gordon-AdS system for general boundary conditions, *J. Hyperbolic Differ. Equ.* 12 (2015), 293–342.
- [16] S. KATAYAMA, Global existence for coupled systems of nonlinear wave and Klein-Gordon equations in three space dimensions, *Math. Z.* 270 (2012), 487–513.
- [17] D. LANNES, Space-time resonances [after Germain, Masmoudi, Shatah], *Séminaire Bourbaki*, Vol. 2011/2012, Astérisque No. 352 (2013), 355–388.
- [18] P.G. LEFLOCH AND Y. MA, *The hyperboloidal foliation method*, World Scientific, 2014.
- [19] P.G. LEFLOCH AND Y. MA, The global nonlinear stability of Minkowski spacetime for the Einstein equations in presence of massive fields, *Note C.R. Acad. Sc. Paris* 354 (2016), 948–953.
- [20] P.G. LEFLOCH AND Y. MA, The global nonlinear stability of Minkowski space for self-gravitating massive fields. The wave-Klein-Gordon model, *Comm. Math. Phys.* 346 (2016), 603–665.
- [21] P.G. LEFLOCH AND Y. MA, The global nonlinear stability of Minkowski space for self-gravitating massive fields. Compact Schwarzschild perturbations. Preprint ArXiv:1511.03324.
- [22] P.G. LEFLOCH AND Y. MA, The global nonlinear stability of Minkowski space for self-gravitating massive fields, in preparation.
- [23] G. MAGNANO AND L.M. SOKOLOWSKI, On physical equivalence between modified gravity theories and a general relativistic self-gravitating scalar field, *Phys. Rev. D* 50 (1994), 5039–5059.
- [24] S. NOJIRI AND S.D. ODINTSOV, Unified cosmic history in modified gravity: from $f(R)$ -theory to Lorentz non-invariant models, *Phys. Rept.* 505 (2011), 59–144.

- [25] M. SALGADO AND D. MARTINEZ-DEL RIO, The initial value problem of scalar-tensor theories of gravity, in: VII Mexican School on Gravitation and Mathematical Physics, J. Phys. Conf. Ser. 91, (2007), 012004.
- [26] M.E. TAYLOR, *Partial differential equations III. Nonlinear equations*, Appl. Math. Sc., Vol. 117, Springer Verlag, New York, 2011.

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