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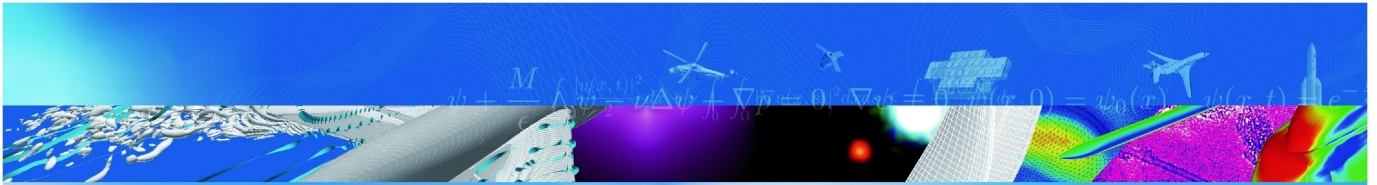
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COMMUNICATION A CONGRES

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Coronagraphic wavefront sensing with COFFEE: high spatial-frequency diversity and other news

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ABSTRACT

The final performance of current and future instruments dedicated to exoplanet detection and characterization is limited by intensity residuals in the scientific image plane, which originate in uncorrected optical aberrations. In order to reach very high contrasts, these aberrations need to be compensated for. We have proposed a focal-plane wave-front sensor called COFFEE (for COronagraphic Focal-plane wave-Front Estimation for Exoplanet detection), which consists in an extension of conventional phase diversity to a coronagraphic system. In this communication, we study the extension of COFFEE to the joint estimation of the phase and the amplitude in the context of space-based coronagraphic instruments: we optimize the diversity phase in order to minimize the reconstruction error; we also propose and optimize a novel low-amplitude high-frequency diversity that should allow the phase-diverse images to still be used for science. Lastly, we perform a first experimental validation of COFFEE in the very high, space-like contrast conditions of the THD bench and show that COFFEE is able to distinguish between phase and amplitude aberrations.

Keywords: Wavefront sensing, phase diversity, phase retrieval, high contrast imaging, inverse problems, exoplanets.

1. INTRODUCTION

The final performance of current and future instruments dedicated to exoplanet detection and characterization (such as SPHERE on the VLT, GPI on Gemini North or future ground or space-based instruments) is limited by intensity residuals in the scientific image plane, which originate in uncorrected optical aberrations. The main contribution to these residuals (after correction of turbulence by Adaptive Optics for ground-based instruments) comes from the quasi-static aberrations introduced upstream of the coronagraph, which create long-lived speckles in the detector plane that can easily be mistaken for a planet. In order to reach very high contrasts such as the ones required to image earth-like planets, these aberrations must be compensated for. We have recently developed a dedicated focal-plane wave-front sensor called COFFEE (for COronagraphic Focal-plane wave-Front Estimation for Exoplanet detection), which consists in an extension of conventional phase diversity to a coronagraphic system: aberrations both upstream and downstream of the coronagraph are estimated using two coronagraphic focal-plane images, recorded from the scientific camera itself, without any differential aberration.^{1,2} Such a system has been successfully validated on the SPHERE instrument, where COFFEE's estimation has been used to compensate for the phase aberration upstream of the coronagraph, leading to a contrast optimization in the whole focal plane area controlled by the AO loop.³ For planet-finder instruments that will aim at detecting and characterizing Earth-like planets from the ground or from space, several orders of magnitude in contrast must still be gained. This requires measuring and compensating not only the phase but also the amplitude aberrations (inhomogeneous intensity in the pupil, due to Fresnel propagation). In this communication, we study the extension of COFFEE to perform a simultaneous estimation of both phase and amplitude aberration from three or more focal plane images. We optimize the diversity phase in order to minimize the reconstruction

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error. We also propose and optimize a novel low-amplitude high-frequency diversity that should allow the phase-diverse images to still be used for science. Lastly, we perform a first experimental validation of COFFEE in the very high, space-like contrast conditions of the THD bench of Observatoire de Paris and show that COFFEE is able to distinguish between phase and amplitude aberrations.

COFFEE has the advantage, with respect to approaches such as speckle nulling⁴ or the Electric Field Conjugation (EFC),⁵ that it does not assume small aberrations and does not perform a Taylor expansion of the complex field. COFFEE is thus likely to lead to a correction that converges faster and needs less image acquisitions. Compared to the SCC approach,⁶ COFFEE needs more images but does not require any modification of the imaging system, and can thus be integrated very easily into any high-contrast imaging system. Compared to ZELDA,⁷ COFFEE has the advantage that it uses coronagraphic images of the scientific detector and is thus immune to Non-Common-Path Aberrations (NCPA).

2. MEASURING AMPLITUDE AND PHASE WITH COFFEE

2.1 Image formation model

We consider a coronagraphic imaging system observing a point source. Let $\exp i \cdot \phi_u + \xi_u$ be the complex field to be estimated in a pupil plane upstream of the coronagraph, and ϕ_d the phase aberrations in a pupil plane downstream of the coronagraph. We denote by $h_c(\phi_u, \xi_u, \phi_d)$ the model of the coronagraphic PSF as a function of ϕ_u, ξ_u, ϕ_d . We assume that we record three coronagraphic images in the focal plane: one focused image i_c^{foc} and two diversity images $\phi_{\text{div}1}$ and $\phi_{\text{div}2}$, differing from the focused image by a known diversity, $\phi_{\text{div}1}$ and $\phi_{\text{div}2}$ respectively. These images can be written as:

$$\begin{aligned} i_c^{\text{foc}} &= \alpha_{\text{foc}} h_{\text{det}} \star h_c(\phi_u, \xi_u, \phi_d) + n_{\text{foc}} + \beta_{\text{foc}} \\ i_c^{\text{div}1} &= \alpha_{\text{div}1} h_{\text{det}} \star h_c(\phi_u + \phi_{\text{div}1}, \xi_u, \phi_d) + n_{\text{div}1} + \beta_{\text{div}1} \\ i_c^{\text{div}2} &= \alpha_{\text{div}2} h_{\text{det}} \star h_c(\phi_u + \phi_{\text{div}2}, \xi_u, \phi_d) + n_{\text{div}2} + \beta_{\text{div}2}, \end{aligned} \quad (1)$$

where α_p is the flux of the star ($p \in \{\text{foc}, \text{div}1, \text{div}2\}$), β_p represents a constant residual background in image i_c^p , h_{det} is the detector PSF, and n_p denotes the measurement noise. This noise is usually a combination of photon and detector noises, which can be well approximated by a white Gaussian inhomogeneous noise for the flux levels at hand in high contrast imaging. Its variance is then the sum of the variances of the photon noise $\sigma_{\text{det}p}^2$ and of the detector noise $\sigma_{\text{phot}p}^2$ ⁸:

$$\sigma_{n_p}^2[k] = \sigma_{\text{det}p}^2 + \sigma_{\text{phot}p}^2[k], \quad (2)$$

where k is the current pixel index. In this equation, $\sigma_{\text{det}p}$ is a scalar determined by the calibration of the detector, and the variance map $\sigma_{\text{phot}p}^2[k]$ can be estimated simply from the image i_c^p itself⁸:

$$\sigma_{\text{phot}p}^2[k] = \max(i_c^p[k], 0) \quad (3)$$

2.2 MAP metric for phase and amplitude reconstruction

The COFFEE method is based on the optimization of a Maximum *A Posteriori* (MAP) metric. It consists in searching for the upstream phase and amplitude aberrations $\hat{\phi}_u$ and $\hat{\xi}_u$, the downstream phase aberrations $\hat{\phi}_d$, as well as the fluxes $\hat{\alpha} = [\hat{\alpha}_{\text{foc}}, \hat{\alpha}_{\text{div}1}, \hat{\alpha}_{\text{div}2}]$ and the background values $\hat{\beta} = [\hat{\beta}_{\text{foc}}, \hat{\beta}_{\text{div}1}, \hat{\beta}_{\text{div}2}]$ that maximize the posterior likelihood $p(\alpha, \beta, \phi_u, \xi_u, \phi_d | i_c^{\text{foc}}, i_c^{\text{div}1}, i_c^{\text{div}2})$. In practice, the solution is obtained by the minimization of the neg-log-likelihood metric J defined as the opposite of the logarithm of this posterior likelihood, *i.e.*:

$$(\hat{\alpha}, \hat{\beta}, \hat{\phi}_u, \hat{\xi}_u, \hat{\phi}_d) = \arg \min_{\alpha, \beta, \phi_u, \xi_u, \phi_d} J(\alpha, \beta, \phi_u, \xi_u, \phi_d), \quad (4)$$

with:

$$\begin{aligned}
J(\alpha, \beta, \phi_u, \xi_u, \phi_d) = & \frac{1}{2} \left\| \frac{\mathbf{i}_c^{\text{foc}} - [\alpha_{\text{foc}} \mathbf{h}_{\text{det}} \star \mathbf{h}_c(\phi_u, \xi_u, \phi_d) + \beta_{\text{foc}}]}{\sigma_{n_{\text{foc}}}} \right\|^2 \\
& + \frac{1}{2} \left\| \frac{\mathbf{i}_c^{\text{div1}} - [\alpha_{\text{div1}} \mathbf{h}_{\text{det}} \star \mathbf{h}_c(\phi_u + \phi_{\text{div1}}, \xi_u, \phi_d) + \beta_{\text{div1}}]}{\sigma_{n_{\text{div1}}}} \right\|^2 \\
& + \frac{1}{2} \left\| \frac{\mathbf{i}_c^{\text{div2}} - [\alpha_{\text{div2}} \mathbf{h}_{\text{det}} \star \mathbf{h}_c(\phi_u + \phi_{\text{div2}}, \xi_u, \phi_d) + \beta_{\text{div2}}]}{\sigma_{n_{\text{div2}}}} \right\|^2 \\
& + \mathcal{R}(\phi_u) + \mathcal{R}(\xi_u) + \mathcal{R}(\phi_d),
\end{aligned} \tag{5}$$

where $\mathcal{R}(\phi_q)$ represents the regularization term on the phase aberration ϕ_q , with $q \in [u, d]$, and $\mathcal{R}(\xi_u)$ represents the regularization term on the amplitude aberrations ξ_u . The minimization of metric J (Equation 5) is performed by means of the VMLM-B method,⁹ and relies on the analytical expressions of gradients $\partial J/\partial \phi_u$, $\partial J/\partial \xi_u$, $\partial J/\partial \phi_d$, $\partial J/\partial \alpha$ and $\partial J/\partial \beta$.

2.3 Choice of appropriate basis and regularization for the aberrations

The use of a pixel basis for the phase reconstruction is required for COFFEE to estimate high order aberrations. However, this leads to a large number of unknowns, which in turn calls for a regularization metric in order to reduce the noise sensitivity. We chose a regularization metric that is based on the available a priori knowledge on the quasi-static aberrations. Indeed, they can be reasonably assumed to be Gaussian, homogeneous and thus endowed with a power spectral density (PSD). In this communication, we use the regularizations in $1/f^2$ and $1/f^4$ whose implementations are described in Ref. 2, both for the phase and for the amplitude regularizations.

The choice of the diversity phases ϕ_{div1} and ϕ_{div2} is a degree of freedom of the method, and is studied in the following sections.

3. OPTIMIZING THE DIVERSITY FOR COFFEE

In this section, we assess the quality of the reconstruction of phase and amplitude aberrations with COFFEE. We consider a set of three coronagraphic images, one at the best focus, and two additional images with known diversity phases ϕ_d^1 and ϕ_d^2 . The imaging wavelength is 1600 nm, representative of a science camera in H band. Additionally, because we consider a calibration context, we consider that the images are high flux (10^9 photons/image) and that the detector is low noise (3 electrons/pixel). The aberrations to be measured have a RMS value of about 0.15 rad (40 nm).

We first illustrate the use of COFFEE with simulations that use the conventional defocus diversity, with 125 nm RMS and 500 nm RMS diversity. The coronagraphic images and the images reconstructed by COFFEE are shown on Fig. 1. These images lead to nanometric reconstruction errors shown on Fig. 2.

Next, we wish to optimize the diversity phases. We do this in two steps: we begin by optimizing the focus diversity in the simplified setting where we have only two images and search only for the upstream and downstream phases. Figure 3 shows the reconstruction error as a function of the amplitude of the diversity. We see that a diversity of at least 10 nm is required to estimate the upstream phase with a nanometric error, and that a much larger diversity, of more than 125 nm, is required if one wants a nanometric error both on the upstream and downstream phases, which is consistent with the results obtained in Ref. 2. We then set the first diversity to 10 nm, and explore the influence of the second diversity on the reconstruction error in the setting where we have three images and search for the upstream phase and amplitude aberrations as well as the downstream aberrations. The results are shown on Figure 4.

(log-scale)	Simulated image	Reconstructed image	Error criterion
Focused plane			
Diversity 1 plane			
Diversity 2 plane			

Figure 1. Simulated images (left column), images reconstructed by COFFEE (middle column), and difference (right column), for a focus diversity. All images are shown in log scale.

(Linear scale)	Simulated phase	Reconstructed phase	Reconstruction error
Upstream aberration phase map			
WFE [nm]	50	50	0.3
Downstream aberration phase map			
WFE [nm]	20	20	0.5

Figure 2. Simulated aberrations (left column), aberrations reconstructed by COFFEE (middle column), and difference (right column), for the upstream aberrations (top) and the downstream ones (bottom). The simulated aberrations are reconstructed with a nanometric RMS precision.

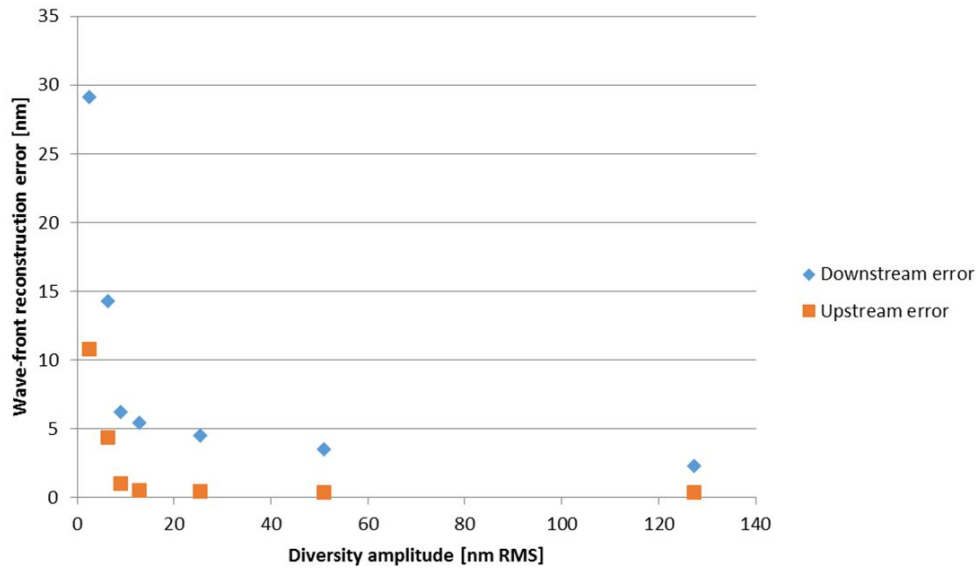


Figure 3. Optimization of the diversity phase. The diversity is a defocus, and only the upstream and downstream phases are reconstructed, using the focused image and one diversity image.

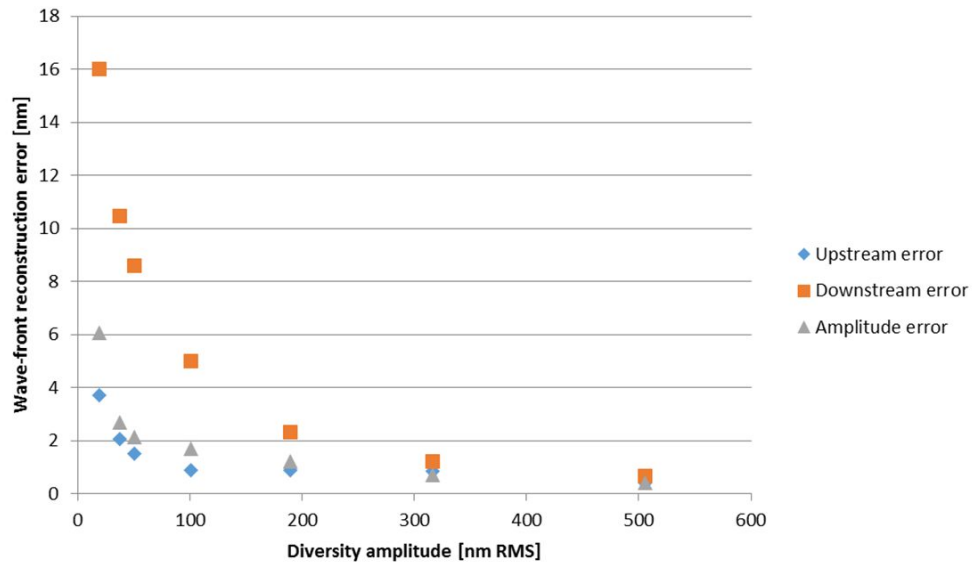


Figure 4. Optimization of the diversity phase for the second diversity image. The diversity of the first diversity image is a defocus, of amplitude 0.05 rad (12 nm), that of the second diversity image is a defocus of increasing amplitude. Upstream and downstream phases as well as upstream amplitude are reconstructed, using the focused image and two diversity images.

4. EXPLORING A NOVEL LOW-AMPLITUDE AND HIGH-FREQUENCY DIVERSITY

The usual diversity phases used in classical (*i.e.*, non coronagraphic) phase diversity,¹⁰ in classical phase retrieval (see, *e.g.*, Ref. 11), and in COFFEE^{1,3,12} are of low order and typically consist of defocus and/or astigmatism. For coronagraphic imaging, this has the unfortunate effect that it renders the diversity images unusable for science, as it brings a substantial amount of additional light onto the detector.

That is why we explore below the possibility to use a high-spatial-frequency and low-amplitude diversity, with the aim of keeping most of the corrected zone of the focal plane useable for science. The first diversity we choose is a waffle mode, made of the product of two orthogonal cosine waves (one along x and one along y), having 20 cycles in the aperture. The second diversity is the same as the first one but rotated by 45° . The amplitude aberrations to be estimated are deduced from a pupil intensity measurement made on the THD bench at Observatoire de Meudon.¹³ The phase aberrations are drawn with a $1/f^{2.4}$ power law which is also representative of aberrations on this bench.

Figure 5 shows an example of simulated and reconstructed images, and the difference between the two, with the proposed waffle diversity. The good fit between the two sets of images suggests that the optimization went well, and indeed the reconstruction errors are quite small, as shown on Fig. 6.

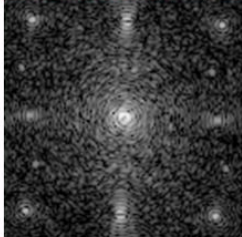
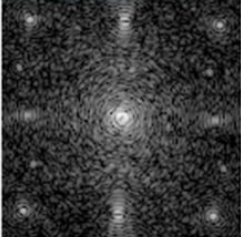
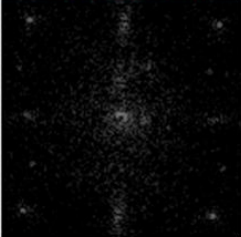
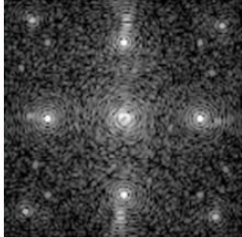
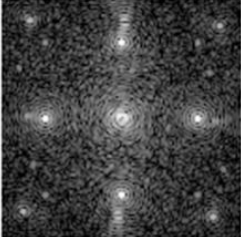
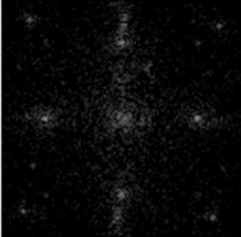
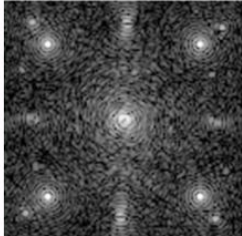
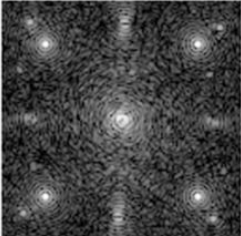
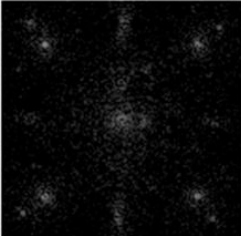
(log-scale)	Simulated image	Reconstructed image	Error criterion
Focused plane			
Diversity 1 plane			
Diversity 2 plane			

Figure 5. Simulated images (left column), images reconstructed by COFFEE (middle column), and difference (right column), for a waffle diversity. The effect of the waffle is to introduce four lateral peaks in the images. All images are shown in log scale.

Then, we study the influence of this waffle diversity on the reconstruction error. Figure 7 shows the reconstruction errors on the upstream phase, the upstream amplitude and the downstream phase as a function of the waffle amplitude. As with the usual low-order diversity, the errors go down when the diversity increases, but here a diversity of 50 nm is sufficient in order to obtain a nanometric accuracy both on the upstream and the downstream phases.

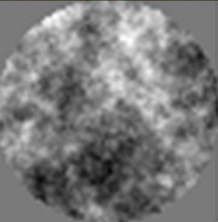
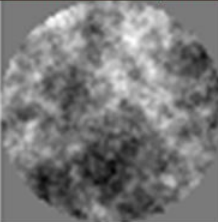

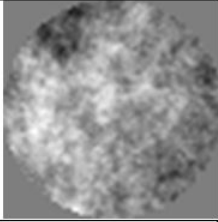
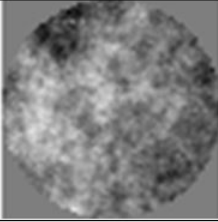
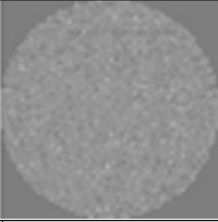
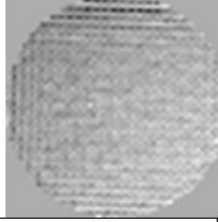
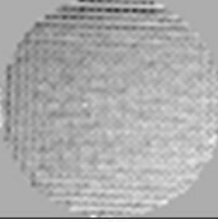

(Linear scale)	Simulated phase	Reconstructed phase	Difference
Upstream phase			
WFE [nm]	50		1
Downstream phase			
WFE [nm]	20		1
Upstream amplitude			
WFE [nm]	15	15	4

Figure 6. Simulated aberrations (left column), aberrations reconstructed by COFFEE (middle column), and difference (right column), for the upstream phase (top), the downstream phase (center) and the upstream amplitude (bottom). These reconstructions validate the use of waffle as a diversity phase.

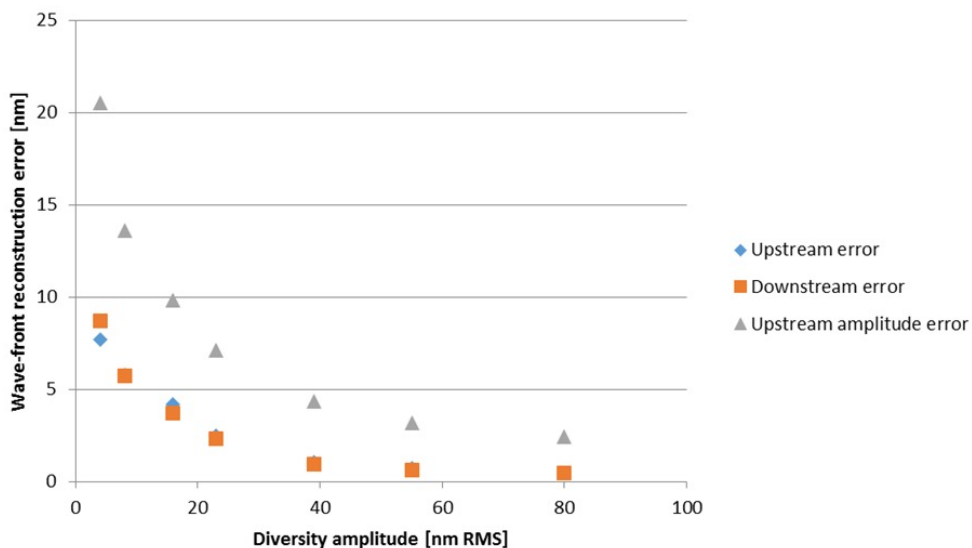


Figure 7. Optimization of the diversity phase for a waffle diversity: the RMS reconstruction errors of the upstream and downstream phases as well as of the upstream amplitude are plotted as a function of the amplitude of the waffle used for the two diversity images.

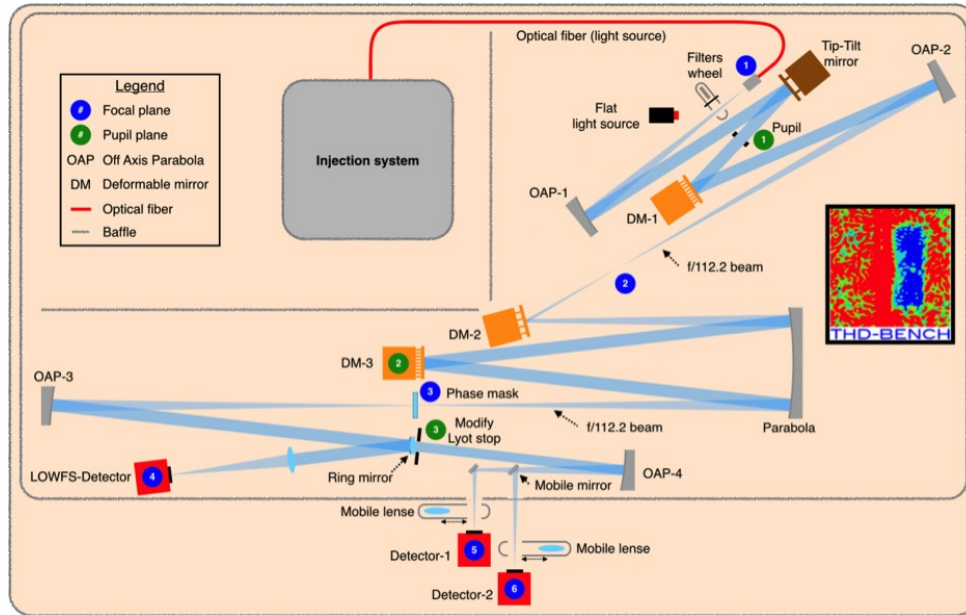


Figure 8. Schematics of the THD bench, including optics, components and beams.

5. EXPERIMENTS OF COFFEE ON THE THD BENCH

In this section, we validate COFFEE on the THD bench, which is a very high contrast imaging bench designed and built by Observatoire de Paris à Meudon,^{6,14} and shown on Fig. 8. Deformable mirror DM3 is a Boston Micro-Machines with 32×32 actuators in a pupil plane used to produce phase aberrations. DM1 and DM2 are placed away from a pupil plane in order to introduce amplitude aberrations. In the experiments described in this paper, DM1 is flattened and DM2 is replaced by a flat mirror. A preliminary work consisted in carefully calibrating the sampling of the detector and the diversity produced by the DM3. Indeed, such a mirror does not have a linear response, and phase diversity needs a precise calibration of the introduced diversity. In order to perform this calibration, we used classical phase diversity with the FQPM coronagraph removed, the following way: with DM3 flat, we recorded a high-SNR focal image of the point-source at $\lambda = 783.25$ nm, then moved the camera by 16.5 mm, which for an F# of 60.6 translates into a peak-to-valley defocus of 4.51 radians, and recorded a diversity image. From these two images we used our phase diversity code to estimate the phase of the setup. Then we sent voltages to DM3 to produce, in theory, a focus, and we repeated the image measurements as well as the phase diversity estimation. The difference of the two estimated phases was then projected onto the first 210 Zernike polynomials in order to filter out noise. The results are shown on Fig. 9 and illustrate the diversity effectively produced by the DM and used in our subsequent, coronagraphic, reconstructions.

Starting from a nominal state of the bench with a contrast of the order of $2 \cdot 10^{-8}$, we added a sine wave at 45° with DM3, and we took 4 coronagraphic images, one focused and three with different levels of diversity. We then ran COFFEE to estimate the phase and the amplitude, using the diversity estimated in (non-coronagraphic) phase diversity as described above. Fig. 10 shows the recorded images and the images that COFFEE reconstructed at convergence. The spots at 45° that are visible in the two first images are the focal-plane traces of the sine wave that we added on the DM. One notices that the images reconstructed by COFFEE are visually close to the recorded ones, which suggests that the minimization performed by COFFEE was effective.

Figure 11 shows the phase and the log-amplitude estimated by COFFEE. Although preliminary, this result shows that COFFEE has been able to rightfully assess that the sine wave was in phase and not in amplitude. The RMS value of the estimated phase is 4 nm RMS. The pattern in the log-amplitude can most likely be attributed to the DM itself.

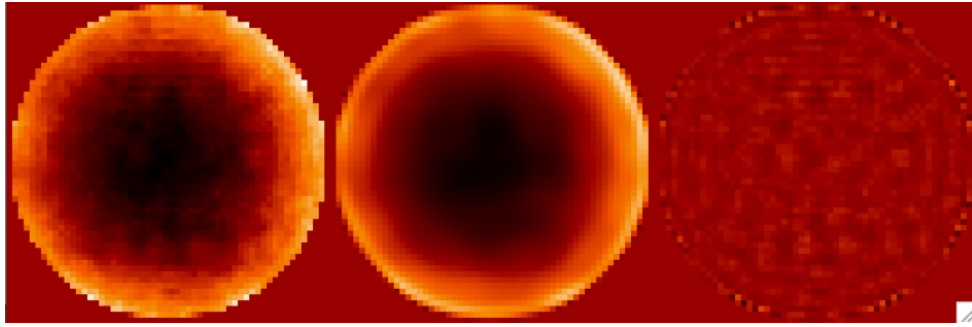


Figure 9. Left: diversity produced by the DM3, obtained as the difference of two classical phase diversity measurements. Middle: diversity on the left projected onto the first 210 Zernike polynomials to filter out noise. Right: difference of the two first images.

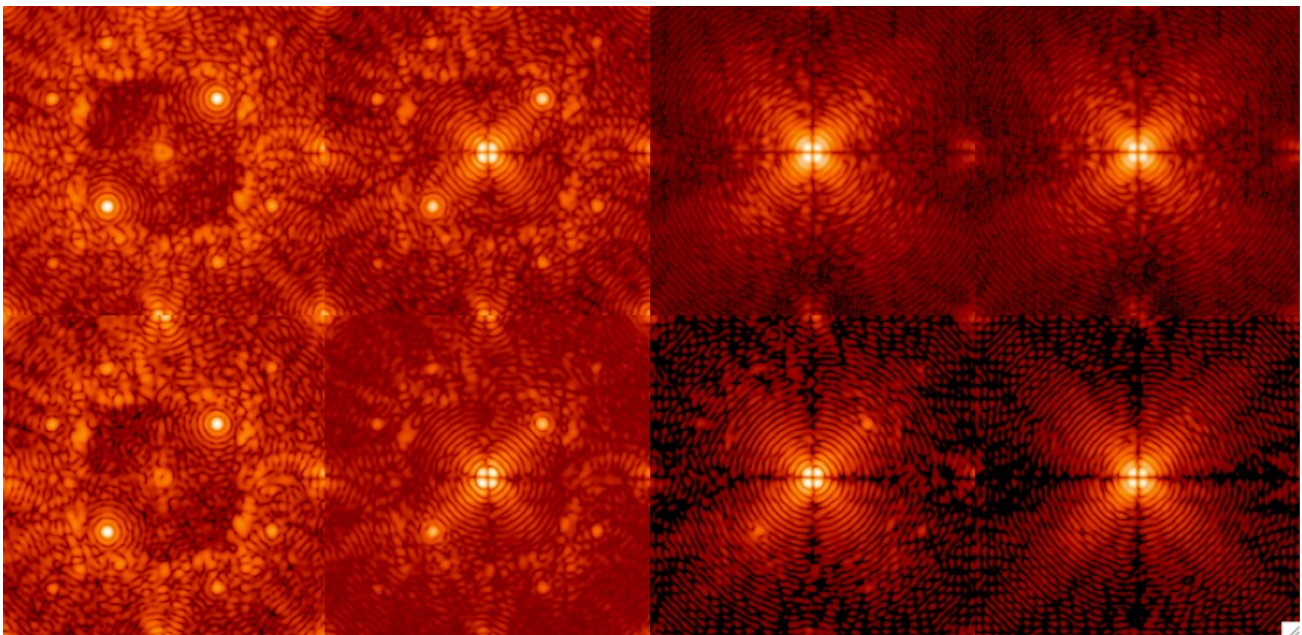


Figure 10. Top: focused image and diversity images recorded on the THD bench. Bottom: images reconstructed by COFFEE at convergence.

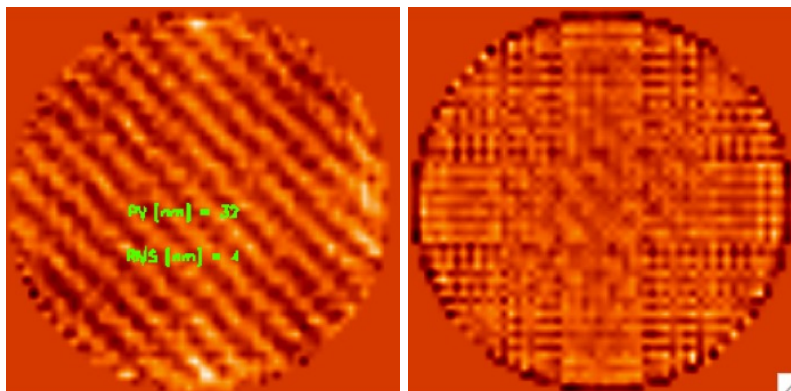


Figure 11. Field reconstructed by COFFEE: phase (left) and log-amplitude (right).

6. CONCLUSION

In this communication, we have studied the extension of COFFEE to perform a simultaneous estimation of both phase and amplitude aberrations from three or more focal plane images. We have optimized the diversity phase in order to minimize the reconstruction error. We have also proposed and optimized a novel low-amplitude high-frequency diversity that should allow the phase-diverse images to still be used for science. Lastly, we have performed a preliminary experimental validation of COFFEE in the very high, space-like contrast conditions of the THD bench of Observatoire de Paris and showed that COFFEE is able to distinguish between phase and amplitude aberrations.

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