



HAL
open science

Weak and strong solutions of equations of compressible magnetohydrodynamics

Xavier Blanc, Bernard Ducomet

► **To cite this version:**

Xavier Blanc, Bernard Ducomet. Weak and strong solutions of equations of compressible magnetohydrodynamics. Handbook of Mathematical Analysis in Mechanics of Viscous Fluids, 2016, 10.1007/978-3-319-10151-4_72-1 . hal-01201867

HAL Id: hal-01201867

<https://hal.science/hal-01201867>

Submitted on 18 Sep 2015

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Weak and strong solutions of equations of compressible magnetohydrodynamics

Xavier Blanc¹, Bernard Ducomet²

¹ Univ. Paris Diderot, Sorbonne Paris Cité,
Laboratoire Jacques-Louis Lions, UMR 7598,
UPMC, CNRS, F-75205 Paris, France

² CEA, DAM, DIF,
Département de Physique Théorique et Appliquée,
F-91297 Arpajon, France

1 Basic equations of magnetohydrodynamics

A plasma is a mixture of ions, electrons and neutral particles. At the macroscopic level, beyond the kinetic description necessary in rarefied situations, dynamics of dense plasmas is governed by the interaction between the fluid components and the electromagnetic fields. When dealing with collisional plasmas [64] [67], it is often appropriate to use one-fluid approximations which are much simpler than a complete kinetic theory and relevant for a number of applications (astrophysics [20] [89] [100] [104], plasma physics [25] [27], electrometallurgy [45] etc...).

Moreover a further approximation is obtained when one studies the interaction between electromagnetic fields and a electric-conducting fluid with electrical effects negligible compared to magnetic effects. This simplified scheme is called magnetohydrodynamics (MHD). It unifies classical fluid dynamics and magnetism of continuum media in a coupled way: electromagnetic fields induce currents in the moving fluid which produce forces which in turn modify the electromagnetic fields. If one considers a compressible heat-conducting fluid, the equations governing the system will be the compressible Navier-Stokes-Fourier system coupled to the set of Maxwell's equations through suitable momenta and energy sources.

1.1 Modelling of the magnetic field

Recall that the electromagnetic field is governed by the Maxwell system [70] decomposed into

- the *Ampère's law*

$$\partial_t \mathbf{D} + \mathbf{j} = \mathbf{curl}_x \mathbf{H}, \quad (1)$$

- the *Coulomb's law*

$$\operatorname{div}_x \mathbf{D} = \rho_c, \quad (2)$$

- the *Faraday's law*

$$\partial_t \mathbf{B} + \mathbf{curl}_x \mathbf{E} = 0, \quad (3)$$

- the *Gauss's law*

$$\operatorname{div}_x \mathbf{B} = 0, \quad (4)$$

where \mathbf{D} is the electric induction, \mathbf{B} is the magnetic induction, \mathbf{E} is the electric field, \mathbf{H} is the magnetic field, ϱ_c is the density of electric charges and \mathbf{j} the electric current. We suppose that these quantities are linked by the constitutive relations

$$\mathbf{B} = \mu\mathbf{H} \quad \text{and} \quad \mathbf{D} = \varepsilon\mathbf{E}, \quad (5)$$

where μ , the magnetic permeability and ε the dielectric permittivity are positive quantities, generally depending on the electromagnetic field. Furthermore one observes that the electric density charge ϱ_c and the current density \mathbf{j} are interrelated through the *electric charge conservation*

$$\partial_t \varrho_c + \operatorname{div}_x \mathbf{j} = 0, \quad (6)$$

however we assume in the following that electric neutrality holds: then $\varrho_c = 0$.

Finally we suppose that the magnetic induction vector \mathbf{B} is related to the electric field \mathbf{E} and the macroscopic fluid velocity \mathbf{u} via *Ohm's law*

$$\mathbf{j} = \sigma(\mathbf{E} + \mathbf{u} \times \mathbf{B}), \quad (7)$$

where the electrical conductivity $\sigma = \sigma(\varrho, \vartheta, \mathbf{H})$ of the fluid is a positive quantity.

Now the *magnetohydrodynamic approximation* [12] [66] [98] consists in neglecting the term $\partial_t \mathbf{D}$ in (1) and we obtain

$$\mathbf{j} = \operatorname{curl}_x \mathbf{H}. \quad (8)$$

This approximation corresponds to a strongly non-relativistic regime and has been justified in the 2-dimensional case by S. Kawashima and Y. Shizuta in [63], and more recently in 3-dimensional space by S. Jiang and F. Li (see [57] for isentropic flows and [58] for the general case).

Accordingly, equation (3) can be written in the form

$$\partial_t \mathbf{B} + \operatorname{curl}_x (\mathbf{B} \times \mathbf{u}) + \operatorname{curl}_x \left(\frac{1}{\sigma} \operatorname{curl}_x \left(\frac{1}{\mu} \mathbf{B} \right) \right) = 0, \quad (9)$$

where $\mu = \mu(|\mathbf{H}|) > 0$.

Setting

$$\mathcal{M}(s) = \int_0^s \tau \partial_\tau (\tau \mu(\tau)) \, d\tau, \quad (10)$$

equation (3) becomes

$$\partial_t \mathcal{M}(|\mathbf{H}|) + \mathbf{j} \cdot \mathbf{E} = \operatorname{div}_x (\mathbf{H} \times \mathbf{E}). \quad (11)$$

Supposing now that we study the previous system in a bounded region $\Omega \subset \mathbb{R}^3$, we add boundary conditions

$$\mathbf{B} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \mathbf{E} \times \mathbf{n}|_{\partial\Omega} = \mathbf{0}. \quad (12)$$

and initial conditions

$$\mathbf{B}(0, x) = \mathbf{B}_0(x), \quad \mathbf{E}(0, x) = \mathbf{E}_0(x) \quad \text{for any } x \in \Omega. \quad (13)$$

In accordance with the basic principles of continuum mechanics, the magnetofluid dynamics is described (see Cabannes [12]) by the *Navier-Stokes-Fourier system* of equations with magnetic sources. This includes *mass conservation*

$$\partial_t \varrho + \operatorname{div}_x (\varrho \mathbf{u}) = 0, \quad (14)$$

the *momentum balance*

$$\partial_t (\varrho \mathbf{u}) + \operatorname{div}_x (\varrho \mathbf{u} \otimes \mathbf{u}) = \operatorname{div}_x \mathbb{T} + \varrho_c \mathbf{E} + \mathbf{j} \times \mathbf{B} + \mathbf{f}, \quad (15)$$

and the *energy balance*

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e \right) + \operatorname{div}_x \left(\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + p \right) \mathbf{u} - \mathbb{S} \mathbf{u} \right) + \operatorname{div}_x \mathbf{q} = \mathbf{E} \cdot \mathbf{j} + g, \quad (16)$$

where e is the specific internal energy, \mathbf{q} denotes the heat flux, \mathbb{T} is the Cauchy stress tensor, \mathbf{f} is a given body force, g is a given energy source and the Lorentz force $\varrho_c \mathbf{E} + \mathbf{j} \times \mathbf{B}$ is imposed by the electromagnetic field. In the magnetohydrodynamic approximation, it may be shown [12] that this approximation also implies that $\varrho_c \mathbf{E}$ is negligible compared to $\mathbf{j} \times \mathbf{B}$. Therefore, the momentum balance (15) reads

$$\partial_t (\varrho \mathbf{u}) + \operatorname{div}_x (\varrho \mathbf{u} \otimes \mathbf{u}) = \operatorname{div}_x \mathbb{T} + \mathbf{j} \times \mathbf{B} + \mathbf{f}. \quad (17)$$

Furthermore the Cauchy stress tensor \mathbb{T} is given by *Stokes' law*

$$\mathbb{T} = \mathbb{S} - p \mathbb{I}, \quad (18)$$

where p is the pressure, and the symbol \mathbb{S} stands for the viscous stress tensor, given by

$$\mathbb{S} = \nu \left(\nabla_x \mathbf{u} + \nabla_x \mathbf{u}^t - \frac{2}{3} \operatorname{div}_x (\mathbf{u}) \mathbb{I} \right) + \eta \operatorname{div}_x (\mathbf{u}) \mathbb{I}, \quad (19)$$

with the shear viscosity coefficient ν , and the bulk viscosity coefficient η .

Finally we impose a no-slip boundary condition for the velocity

$$\mathbf{u}|_{\partial\Omega} = 0. \quad (20)$$

Using (11) and (16) the energy of the system satisfies

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + \mathcal{M}(|\mathbf{H}|) \right) + \operatorname{div}_x \left(\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + p \right) \mathbf{u} + \mathbf{E} \times \mathbf{B} - \mathbb{S} \mathbf{u} + \mathbf{q} \right) = 0, \quad (21)$$

where e is the specific internal energy, and \mathbf{q} denotes the heat flux.

We prescribe a no-flux boundary condition for \mathbf{q}

$$\mathbf{q} \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (22)$$

Using the previous boundary conditions we deduce that the (conserved) total energy of the system satisfies

$$\frac{d}{dt} \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + \mathcal{M}(|\mathbf{H}|) \right) dx = 0. \quad (23)$$

In the variational formulation it will be more convenient to replace (21) by the *entropy balance*

$$\partial_t (\varrho s) + \operatorname{div}_x (\varrho s \mathbf{u}) + \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) = r, \quad (24)$$

with the specific entropy s , and the entropy production rate r , for which the second law of thermodynamics requires $r \geq 0$.

The specific entropy s , the specific internal energy e , and the pressure are interrelated through the second principle of thermodynamics

$$\vartheta Ds = De + D \left(\frac{1}{\varrho} \right) p, \quad (25)$$

where D stand for the differential with respect to the variable ϱ , ϑ , which implies some compatibility conditions between e and p (Maxwell's relation).

If the motion is smooth, the sum of the kinetic and magnetic energy satisfies

$$\partial_t \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \mathcal{M}(|\mathbf{H}|) \right) + \operatorname{div}_x \left(\left(\frac{1}{2} \varrho |\mathbf{u}|^2 + p \right) \mathbf{u} + \mathbf{E} \times \mathbf{B} - \mathbb{S} \mathbf{u} \right) = p \operatorname{div}_x (\mathbf{u}) - \mathbb{S} : \nabla_x \mathbf{u} - \frac{1}{\mu\sigma} |\operatorname{curl}_x \mathbf{B}|^2. \quad (26)$$

Indeed using (25) one checks that

$$r = \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} + \frac{1}{\sigma} |\mathbf{curl}_x \mathbf{H}|^2 - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right). \quad (27)$$

If the motion is not (known to be) smooth, validity of (16) is no longer guaranteed and the dissipation rate of mechanical energy may exceed the value $\mathbb{S} : \nabla_x \mathbf{u}$.

Accordingly, we replace equation (24) by inequality

$$\partial_t(\varrho s) + \operatorname{div}_x(\varrho s \mathbf{u}) + \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) \geq \frac{1}{\vartheta} \left(\mathbb{S} : \nabla_x \mathbf{u} + \frac{1}{\sigma} |\mathbf{curl}_x \mathbf{H}|^2 - \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right), \quad (28)$$

to be satisfied together with the total energy balance (23).

1.2 Hypotheses on constitutive relations

In order to close the system one needs first to postulate a constitutive equation relating the pressure to the other variables. Several possibilities have been considered. We choose here, in the spirit of [30] (however see [36] [39] [51] for other possible choices) to focus on high-temperature phenomena appearing in various astrophysical contexts (stars, planetar magnetospheres, neighborhood of black holes etc...) [20] [89] [100] [104], or plasma physics applications (tokamaks, inertial confinement fusion devices etc...) [64] [67].

In that case, the so called equilibrium diffusion limit of radiation hydrodynamics involves a modified pressure law where the pressure $p_F(\varrho, \vartheta)$ in the fluid is augmented by a radiation component $p_R(\vartheta)$ related to the absolute temperature ϑ through the Stefan-Boltzmann law

$$p_R(\vartheta) = \frac{a}{3} \vartheta^4, \quad \text{with a constant } a > 0 \quad (29)$$

(see, for instance, Chapter 15 in [33]). Accordingly, the specific internal energy of the fluid must be supplemented by a term

$$e_R(\vartheta) = e_R(\varrho, \vartheta) = \frac{a}{\varrho} \vartheta^4, \quad (30)$$

where ϱ is the fluid density and similarly, the heat conductivity of the fluid is enhanced by a radiation component (see [11] [78])

$$\mathbf{q}_R(\vartheta) = -\kappa_R \vartheta^3 \nabla_x \vartheta, \quad \text{with a constant } \kappa_R > 0 \quad (31)$$

Finally the total pressure p takes the form

$$p(\varrho, \vartheta) = p_F(\varrho, \vartheta) + p_R(\vartheta), \quad (32)$$

where the radiation component is given by (29). Similarly, the specific internal energy reads

$$e(\varrho, \vartheta) = e_F(\varrho, \vartheta) + e_R(\varrho, \vartheta), \quad (33)$$

with e_R determined by (30).

Furthermore we suppose that the gas is perfect and monoatomic

$$p_F(\varrho, \vartheta) = \frac{2}{3} \varrho e_F(\varrho, \vartheta). \quad (34)$$

The following hypothesis expresses the convexity of free energy

$$\frac{\partial p_F(\varrho, \vartheta)}{\partial \varrho} > 0, \quad \frac{\partial e_F(\varrho, \vartheta)}{\partial \vartheta} > 0 \quad \text{for all } \varrho, \vartheta > 0, \quad (35)$$

and finally, we suppose

$$\liminf_{\varrho \rightarrow 0^+} p_F(\varrho, \vartheta) = 0, \quad \liminf_{\varrho \rightarrow 0^+} \frac{\partial p_F(\varrho, \vartheta)}{\partial \varrho} > 0 \text{ for any fixed } \vartheta > 0, \quad (36)$$

and

$$\liminf_{\vartheta \rightarrow 0^+} e_F(\varrho, \vartheta) = 0, \quad \liminf_{\vartheta \rightarrow 0^+} \frac{\partial e_F(\varrho, \vartheta)}{\partial \varrho} > 0 \text{ for any fixed } \varrho > 0. \quad (37)$$

As one can check by solving (25) with (34), there exists a function $P_F \in C^1(0, \infty)$ such that

$$p_F(\varrho, \vartheta) = \vartheta^{\frac{5}{2}} P_F\left(\frac{\varrho}{\vartheta^{\frac{3}{2}}}\right). \quad (38)$$

Consequently

$$\frac{\partial e_F(\varrho, \vartheta)}{\partial \vartheta} = \frac{3}{2} \frac{1}{Y} \left(\frac{5}{3} P_F(Y) - P'_F(Y) Y \right), \quad \text{with } Y = \frac{\varrho}{\vartheta^{\frac{3}{2}}},$$

so according to (35)

$$P'_F(z) > 0, \quad \frac{5}{3} P_F(z) - P'_F(z) z > 0 \text{ for any } z > 0, \quad (39)$$

where the latter inequality yields

$$\left(\frac{P_F(z)}{z^{\frac{5}{3}}} \right)' < 0 \text{ for all } z > 0. \quad (40)$$

In particular

$$\frac{P_F(z)}{z^{\frac{5}{3}}} \rightarrow p_\infty > 0 \text{ for } z \rightarrow \infty. \quad (41)$$

According to (36), (37), we see that $P_F \in C^1[0, \infty)$. Moreover

$$0 < \liminf_{z \rightarrow 0^+} \frac{1}{z} \left(\frac{5}{3} P_F(z) - P'_F(z) z \right) \leq \limsup_{z \rightarrow 0^+} \frac{1}{z} \left(\frac{5}{3} P_F(z) - P'_F(z) z \right) < \infty, \quad (42)$$

$$\limsup_{z \rightarrow \infty} \frac{1}{z} \left(\frac{5}{3} P_F(z) - P'_F(z) z \right) < \infty, \quad (43)$$

and

$$\lim_{z \rightarrow 0^+} P_F(z) = 0, \quad \lim_{z \rightarrow 0^+} P'_F(z) > 0, \quad \lim_{z \rightarrow \infty} \frac{P'_F(z)}{z^{\frac{2}{3}}} = \frac{5}{3} p_\infty > 0. \quad (44)$$

Now it follows from (44) that

$$P'_F(z) \geq cz^{\frac{2}{3}} \text{ for all } z > 0, \text{ and a certain } c > 0.$$

Therefore there exists $p_c > 0$ such that the mapping

$$\varrho \mapsto p_F(\varrho, \vartheta) - p_c \varrho^{\frac{5}{3}}, \quad (45)$$

is a non-decreasing function of ϱ for any fixed $\vartheta > 0$.

Accordingly, using equation (25) we have

$$s(\varrho, \vartheta) = s_F(\varrho, \vartheta) + s_R(\varrho, \vartheta), \quad (46)$$

where

$$s_F(\varrho, \vartheta) = S_F\left(\frac{\varrho}{\vartheta^{\frac{3}{2}}}\right), \quad \text{with } S'_F(z) = -\frac{3}{2} \frac{\frac{5}{3} P_F(z) - P'_F(z) z}{z^2}, \quad (47)$$

and

$$\varrho s_R(\varrho, \vartheta) = \frac{4}{3} a \vartheta^3. \quad (48)$$

Note that after (39) and (42), S_F is a decreasing function such that

$$\lim_{z \rightarrow 0^+} z S'_F(z) = -\frac{2}{5} P'_F(0) < 0,$$

whence, normalizing by $S_F(1) = 0$, we get

$$-c_1 \log(z) \leq S_F(z) \leq -c_2 \log(z), \quad c_1 > 0, \quad \text{for all } 0 < z \leq 1, \quad (49)$$

and

$$0 \geq S_F(z) \geq -c_3 \log(z) \quad \text{for all } z \geq 1. \quad (50)$$

Using the second law of thermodynamics, we see that viscosity coefficients ν and η are non-negative quantities.

In addition, we shall suppose that

$$\nu = \nu(\vartheta, |\mathbf{H}|) > 0, \quad \eta = \eta(\vartheta, |\mathbf{H}|) \geq 0, \quad (51)$$

where the bulk viscosity ν satisfies some technical, but physically relevant, coercivity conditions to be specified below.

Similarly, the heat flux \mathbf{q} obeys Fourier's law

$$\mathbf{q} = \mathbf{q}_R + \mathbf{q}_F, \quad (52)$$

where the radiation heat flux \mathbf{q}_R is given by (31) and

$$\mathbf{q}_F = -\kappa_F(\varrho, \vartheta, |\mathbf{H}|) \nabla_x \vartheta, \quad \kappa > 0. \quad (53)$$

For the sake of simplicity we assume the transport coefficients ν , κ_F , μ and σ to admit a common temperature scaling, namely

$$c_1(1 + \vartheta^\alpha) \leq \kappa_F(\varrho, \vartheta, |\mathbf{H}|), \quad \nu(\vartheta, |\mathbf{H}|), \quad \sigma^{-1}(\varrho, \vartheta, |\mathbf{H}|) \leq c_2(1 + \vartheta^\alpha), \quad c_1, c_2 > 0, \quad (54)$$

with $\alpha \geq 1$ to be specified below. Similarly, we suppose that

$$0 \leq \eta(\vartheta, |\mathbf{H}|) \leq c_3(1 + \vartheta^\alpha), \quad (55)$$

and that the magnetic permeability has the symbol property

$$\underline{c}_k s(1 + s)^{-k} \leq \partial_s^k(s\mu(s)) \leq \bar{c}_k s(1 + s)^{-k}, \quad (56)$$

for any $s \geq 0$ and for $k = 0, 1$, with $\underline{c}_k, \bar{c}_k > 0$.

We finally suppose that external sources are absent: $\mathbf{f} = 0$ and $g = 0$.

Let us conclude this introduction by briefly fixing the limits of our study. As our framework is explicitly viscous and compressible, all the important works about the incompressible model are beyond our scope. However the reader may consult, among standard references: Ladyzhenskaya and Solonnikov [69], Duvaut and Lions [31] [32], Sermange and Temam [88] and Cannone et al. [13] for reviews of the incompressible setting and Germain and Masmoudi [46] (with references therein) for the inviscid case. Finally we refer the reader to Eringen and Maugin [34] for more complex situations coupling electrodynamics to continua.

The article is structured as follows: in Section 2 we consider the main existence result of weak solutions, in Section 3 we present related models (barotropic or polytropic) and in Section 4 we introduce asymptotic models. In Section 5 we present the main results concerning existence and uniqueness of strong solutions including critical aspects and finally results for one dimensional models are given in the last Section 6.

2 Weak solutions

2.1 Variational formulation

As in the pure hydrodynamical model [29] the approach is based on the concept of variational solutions.

1- The mass conservation is expressed through the weak form

$$\int_0^T \int_{\Omega} \left(\varrho \partial_t \varphi + \varrho \mathbf{u} \cdot \nabla_x \varphi \right) dx dt + \int_{\Omega} \varrho_0 \varphi(0, \cdot) dx = 0, \quad (57)$$

to be satisfied for any test function $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$ where ϱ_0 is the initial density.

Assuming that both ϱ and the flux $\varrho \mathbf{u}$ are locally integrable on $[0, T] \times \overline{\Omega}$ equation (57) makes sense. Moreover if ϱ belongs to $L^\infty(0, T; L^\gamma(\Omega))$ for a $\gamma > 1$, then [82] $\varrho \in C([0, T]; L_{weak}^\gamma(\Omega))$ i.e.

$$\varrho(t) \rightarrow \varrho_0 \text{ weakly in } L^\gamma(\Omega) \text{ as } t \rightarrow 0,$$

and the total mass

$$M_0 = \int_{\Omega} \varrho(t) dx = \int_{\Omega} \varrho_0 dx, \quad (58)$$

is a constant of motion.

For $\gamma \geq 2$ one can use the DiPerna-Lions theory [28] to show that the pair (ϱ, \mathbf{u}) satisfies the renormalized equation

$$\int_0^T \int_{\Omega} \left(b(\varrho) \partial_t \varphi + b(\varrho) \mathbf{u} \cdot \nabla_x \varphi + (b(\varrho) - b'(\varrho) \varrho) \operatorname{div}_x(\mathbf{u} \varphi) \right) dx dt + \int_{\Omega} b(\varrho_0) \varphi(0, \cdot) dx = 0, \quad (59)$$

for any $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$, and for any continuously differentiable function b whose derivative vanishes for large arguments. It can also be shown that any renormalized solution ϱ belongs to the space $C([0, T]; L^1(\Omega))$ (see [28]).

2- The variational formulation of the momentum equation reads now

$$\begin{aligned} & \int_0^T \int_{\Omega} \left((\varrho \mathbf{u}) \cdot \partial_t \varphi + (\varrho \mathbf{u} \otimes \mathbf{u}) : \nabla_x \varphi + p \operatorname{div}_x(\varphi) \right) dx dt \\ &= \int_0^T \int_{\Omega} \left(\mathbb{S} : \nabla_x \varphi - (\mathbf{j} \times \mathbf{B}) \cdot \varphi \right) dx dt - \int_{\Omega} (\varrho \mathbf{u})_0 \cdot \varphi(0, \cdot) dx, \end{aligned} \quad (60)$$

which must be satisfied for any vector field $\varphi \in \mathcal{D}([0, T] \times \Omega; \mathbb{R}^3)$. Once more we suppose that the quantities $\varrho \mathbf{u}$, $\varrho \mathbf{u} \otimes \mathbf{u}$, p , \mathbb{S} , $\varrho \nabla_x \Psi$, and $\mathbf{J} \times \mathbf{B}$ are at least locally integrable on the set $[0, T] \times \Omega$.

The initial value of the momentum $(\varrho \mathbf{u})_0$ must satisfy

$$(\varrho \mathbf{u})_0 = 0 \text{ a.e. on the set } \{\varrho_0 = 0\}. \quad (61)$$

Finally using the no-slip boundary condition (20) we get

$$\mathbf{u} \in L^2(0, T; W_0^{1,2}(\Omega; \mathbb{R}^3)). \quad (62)$$

Observe that our choice of the function space has been motivated by the a priori bounds resulting from the entropy balance specified below.

3- According to (28) the weak form of the entropy production reads

$$\begin{aligned} & \int_0^T \int_{\Omega} \left(\varrho s \partial_t \varphi + \varrho s \mathbf{u} \cdot \nabla_x \varphi + \frac{\mathbf{q}}{\vartheta} \cdot \nabla_x \varphi \right) dx dt \leq \\ & \int_0^T \int_{\Omega} \frac{1}{\vartheta} \left(\frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} - \mathbb{S} : \nabla_x \mathbf{u} - \frac{1}{\mu^2 \sigma} |\operatorname{curl}_x B|^2 \right) \varphi dx dt - \int_{\Omega} (\varrho s)_0 \varphi(0, \cdot) dx, \end{aligned} \quad (63)$$

for any test function $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$, $\varphi \geq 0$.

We observe that the presence of ϑ in the denominator forces this quantity to be positive on a set of full measure for (63) to make sense.

4- Finally Maxwell's equation (9) is replaced by

$$\int_0^T \int_{\Omega} \left(\mathbf{B} \cdot \partial_t \varphi - \left(\mathbf{B} \times \mathbf{u} + \frac{1}{\sigma} \mathbf{curl}_x \frac{\mathbf{B}}{\mu} \right) \cdot \mathbf{curl}_x \varphi \right) dx dt + \int_{\Omega} \mathbf{B}_0 \cdot \varphi(0, \cdot) dx, \quad (64)$$

to be satisfied for any vector field $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$.

Hereafter using boundary conditions (12) and (62) one must take

$$\mathbf{B}_0 \in L^2(\Omega), \quad \operatorname{div}_x(\mathbf{B}_0) = 0 \text{ in } \mathcal{D}'(\Omega), \quad \mathbf{B}_0 \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (65)$$

By virtue of Theorem 1.4 in [93], \mathbf{B}_0 belongs to the closure of solenoidal fields on $\mathcal{D}(\Omega)$ with respect to the L^2 -norm.

Applying (23) and (28), we infer that $\mathcal{M}(|\mathbf{H}|) \in L^\infty(0, T; L^1(\Omega; \mathbb{R}^3))$ and $\mathbf{curl}_x \mathbf{H} \in L^2(0, T; L^2(\Omega; \mathbb{R}^3))$. But using (10) and (56) we deduce that $\mathbf{B} \in L^\infty(0, T; L^2(\Omega; \mathbb{R}^3))$, and from (4) and (12) we see that

$$\operatorname{div}_x(\mathbf{B})(t) = 0 \text{ in } \mathcal{D}'(\Omega), \quad \mathbf{B}(t) \cdot \mathbf{n}|_{\partial\Omega} = 0 \text{ for a.e. } t \in (0, T),$$

so using Theorem 6.1 in [32] we conclude that

$$\mathbf{H} \in L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)), \quad \operatorname{div}_x(\mathbf{B})(t) = 0, \quad \mathbf{B} \cdot \mathbf{n}|_{\partial\Omega} = 0 \text{ for a.e. } t \in (0, T). \quad (66)$$

5- According to (21) we shall assume the total energy to be a constant of motion

$$E(t) = \int_{\Omega} \left(\frac{1}{2} \varrho(t) |\mathbf{u}(t)|^2 + (\varrho e)(t) + \mathcal{M}(|\mathbf{H}(t)|) \right) dx = E_0 \text{ for a.e. } t \in (0, T), \quad (67)$$

where

$$E_0 = \int_{\Omega} \left(\frac{1}{2\varrho_0} |(\varrho \mathbf{u})_0|^2 + (\varrho e)_0 + \mathcal{M}(|\mathbf{H}_0|) \right) dx. \quad (68)$$

Of course the initial values of the entropy $(\varrho s)_0$ and the internal energy $(\varrho e)_0$ should be chosen consistently with the constitutive relations. A natural possibility consists in fixing the initial temperature ϑ_0 and setting

$$(\varrho s)_0 = \varrho_0 s(\varrho_0, \vartheta_0), \quad (\varrho e)_0 = \varrho_0 e(\varrho_0, \vartheta_0). \quad (69)$$

2.2 Global existence of weak solutions for large data

The following result holds

Theorem 2.1 *Let $\Omega \subset \mathbb{R}^3$ be a bounded domain with boundary of class $C^{2+\delta}$, $\delta > 0$.*

Suppose that the thermodynamic functions

$$p = p(\varrho, \vartheta), \quad e = e(\varrho, \vartheta), \quad s = s(\varrho, \vartheta)$$

are interrelated through (25), where p , e can be decomposed as in (32), (33), with the components p_F , e_F satisfying (34). Moreover, let $p_F(\varrho, \vartheta)$, $e_F(\varrho, \vartheta)$ be continuously differentiable functions of positive arguments ϱ , ϑ satisfying (35)-(37).

Furthermore, we suppose that the transport coefficients

$$\nu = \nu(\vartheta, |\mathbf{H}|), \quad \eta = \eta(\vartheta, |\mathbf{H}|), \quad \kappa_F = \kappa_F(\varrho, \vartheta, |\mathbf{H}|), \quad \sigma = \sigma(\varrho, \vartheta, |\mathbf{H}|) \text{ and } \mu = \mu(|\mathbf{H}|),$$

are continuously differentiable functions of their arguments obeying (51)-(56), with

$$1 \leq \alpha < \frac{65}{27}. \quad (70)$$

Finally, let the initial data ϱ_0 , $(\varrho \mathbf{u})_0$, ϑ_0 , \mathbf{B}_0 be given so that

$$\varrho_0 \in L^{\frac{5}{3}}(\Omega), \quad (\varrho \mathbf{u})_0 \in L^1(\Omega; \mathbb{R}^3), \quad \vartheta_0 \in L^\infty(\mathbb{R}^3), \quad \mathbf{B}_0 \in L^2(\Omega; \mathbb{R}^3), \quad (71)$$

$$\varrho_0 \geq 0, \quad \vartheta_0 > 0, \quad (72)$$

$$(\varrho s)_0 = \varrho_0 s(\varrho_0, \vartheta_0), \quad \frac{1}{\varrho_0} |(\varrho \mathbf{u})_0|^2, \quad (\varrho e)_0 = \varrho_0 e(\varrho_0, \vartheta_0) \in L^1(\Omega), \quad (73)$$

and

$$\operatorname{div}_x(\mathbf{B}_0) = 0 \text{ in } \mathcal{D}'(\Omega), \quad \mathbf{B}_0 \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (74)$$

Then problem (57)-(65) possesses at least one variational solution ϱ , \mathbf{u} , ϑ , \mathbf{B} in the sense of Section 2.1 on an arbitrary time interval $(0, T)$.

Observe that hypothesis (70) is a technical restriction. However functional dependence on ϑ was rigorously justified in [26] after an asymptotic analysis of some kinetic-fluid models. The same temperature scaling on all transport coefficients was also suggested in [47].

2.2.1 Difficulties and methods

The main difficulty of the problem being the lack of a priori estimates, methods of weak convergence based on the theory of the parametrized (Young) measures represent the main tool [83] to proceed. Accordingly, we denote by $\bar{b}(z)$ a weak limit of any sequence of composed functions $\{b(z_n)\}_{n \geq 1}$. More precisely,

$$\int_{\mathbb{R}^M} \overline{b(y, \mathbf{z})} \varphi(y) \, dy = \int_{\mathbb{R}^M} \varphi(y) \left(\int_{\mathbb{R}^K} b(y, z) \, d\Lambda_y(z) \right) \, dy,$$

where $\Lambda_y(z)$ is a parametrized measure associated to a sequence $\{\mathbf{z}_n\}_{n \geq 1}$ of vector-valued functions, $\mathbf{z}_n : \mathbb{R}^M \rightarrow \mathbb{R}^K$ (see Chapter 1 in [83]).

Assume there is a sequence of approximate solutions resulting from a suitable regularization process. The starting point is to establish the relation

$$\overline{\left(p(\varrho, \vartheta) - \left(\frac{4}{3}\nu + \eta \right) \operatorname{div}_x(\mathbf{u}) \right) b(\varrho)} = \left(\overline{p(\varrho, \vartheta)} \overline{b(\varrho)} - \left(\frac{4}{3}\nu(\vartheta, |\mathbf{B}|) + \eta(\vartheta, |\mathbf{B}|) \right) \operatorname{div}_x(\mathbf{u}) \right) \overline{b(\varrho)}, \quad (75)$$

for any bounded function b .

The quantity $p - (4/3\nu + \eta)\operatorname{div}_x(\mathbf{u})$ is called the effective viscous pressure and relation (75) was proved by P.-L. Lions [72] for the barotropic Navier-Stokes system, with $p = p(\varrho)$, when viscosity coefficients ν and η are constant. The same result for general temperature dependent viscosity coefficients was obtained in [40] with the help of certain commutator estimates in the spirit of [19].

In the present setting, this approach has to be modified in order to accommodate the dependence of ν and η on the magnetic field \mathbf{B} .

The propagation of density oscillations can be described by the renormalized continuity equation [28]

$$\partial_t b(\varrho) + \operatorname{div}_x(b(\varrho)\mathbf{u}) + \left(b'(\varrho)\varrho - b(\varrho) \right) \operatorname{div}_x(\mathbf{u}) = 0, \quad (76)$$

and its “weak” counterpart

$$\partial_t \overline{b(\varrho)} + \operatorname{div}_x(\overline{b(\varrho)\mathbf{u}}) + \overline{\left(b'(\varrho)\varrho - b(\varrho) \right) \operatorname{div}_x(\mathbf{u})} = 0. \quad (77)$$

To prove (76) one must show first boundedness of the oscillations defect measure

$$\mathbf{osc}_{\gamma+1}[\varrho_n \rightarrow \varrho](Q) = \sup_{k \geq 1} \left(\limsup_{n \rightarrow \infty} \int_Q |T_k(\varrho_n) - T_k(\varrho)|^{\gamma+1} dx dt \right) < \infty, \quad (78)$$

where $T_k(\varrho) = \min\{\varrho, k\}$, for any bounded $Q \subset \mathbb{R}^4$ and a $\gamma > 1$.

Due to the poor estimates resulting from (54), (55) and (63) relation (78) has to be replaced by a weaker weighted estimate

$$\sup_{k \geq 1} \left(\limsup_{n \rightarrow \infty} \int_Q (1 + \vartheta_n)^{-\beta} |T_k(\varrho_n) - T_k(\varrho)|^{\gamma+1} dx dt \right) < \infty, \quad (79)$$

for suitable $\beta > 0$, $\gamma > 1$.

Finally one recovers strong convergence of the sequence $\{\vartheta_n\}_{n \geq 1}$ of approximate temperatures, knowing that spatial gradients of ϑ_n are uniformly square integrable and that

$$\overline{\varrho s(\varrho, \vartheta)} = \overline{\varrho s(\varrho, \vartheta)} \vartheta, \quad (80)$$

where (80) can be deduced from the entropy inequality (63) by using Aubin-Lions lemma.

The structure of the proof of Theorem 2.1 goes as follows: introducing a three level approximation scheme adapted from [39] allows to reduce the proof to a weak stability problem. One first proves uniform bounds on the sequence of approximate solutions. Relying on these uniform bounds, one can handle the convective terms in the field equations by means of a compactness argument. The strong convergence of the sequence of approximate temperatures is proved by using entropy inequality together with the theory of parametrized (Young) measures discussed below. Then one finally shows that the sequence of approximate densities converges strongly in the Lebesgue space $L^1((0, T) \times \Omega)$. This requires weighted estimates of the oscillations defect measure in order to show that the limit densities satisfy the renormalized continuity equation.

2.2.2 The approximation scheme

The approximation scheme used in the following is inspired from Chapter 3 in [42].

1- The continuity equation (17) is regularized by a viscous approximation

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = \varepsilon \Delta \varrho, \quad \varepsilon > 0, \quad (81)$$

on $(0, T) \times \Omega$, supplemented by the homogeneous Neumann boundary conditions

$$\nabla_x \varrho \cdot \mathbf{n}|_{\partial \Omega} = 0. \quad (82)$$

The initial density is given by

$$\varrho(0, \cdot) = \varrho_{0, \delta}, \quad (83)$$

where

$$\varrho_{0, \delta} \in C^1(\overline{\Omega}), \quad \nabla_x \varrho_{0, \delta} \cdot \mathbf{n}|_{\partial \Omega} = 0, \quad \inf_{x \in \Omega} \varrho_{0, \delta}(x) > 0, \quad (84)$$

with a positive parameter $\delta > 0$.

The functions $\varrho_{0, \delta}$ are chosen in such a way that

$$\varrho_{0, \delta} \rightarrow \varrho_0 \text{ in } L^{\frac{5}{3}}(\Omega), \quad |\{\varrho_{0, \delta} < \varrho_0\}| \rightarrow 0 \text{ for } \delta \rightarrow 0. \quad (85)$$

Here, of course, the choice of the ‘‘critical’’ exponent $\gamma = 5/3$ is directly related to estimate (44) established above.

2- The regularized momentum equation is

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) + \nabla_x p + \delta \nabla_x \varrho^\Gamma + \varepsilon \nabla_x \mathbf{u} \nabla_x \varrho = \operatorname{div}_x(\mathbb{S}) + \mathbf{j} \times \mathbf{B}, \quad (86)$$

in $(0, T) \times \Omega$ where the approximate velocity field satisfies the no slip boundary condition

$$\mathbf{u}|_{\partial\Omega} = 0. \quad (87)$$

We also prescribe initial conditions

$$(\varrho \mathbf{u})(0, \cdot) = (\varrho \mathbf{u})_{0,\delta}, \quad (88)$$

where

$$(\varrho \mathbf{u})_{0,\delta} = \begin{cases} (\varrho \mathbf{u})_0 & \text{provided } \varrho_{0,\delta} \geq \varrho_0, \\ 0 & \text{otherwise.} \end{cases} \quad (89)$$

The artificial pressure term $\delta \varrho^\Gamma$ in (86) provides additional estimates on the approximate densities in order to facilitate the limit passage $\varepsilon \rightarrow 0$. To this end, one has to take Γ large enough, say, $\Gamma > 8$, and to re-parametrize the initial distribution of the approximate densities so that

$$\delta \int_{\Omega} \varrho_{0,\delta}^\Gamma dx \rightarrow 0 \text{ for } \delta \rightarrow 0. \quad (90)$$

3- The entropy equation (24) is replaced by the regularized internal energy balance

$$\partial_t(\varrho e + \delta \varrho \vartheta) + \operatorname{div}_x((\varrho e + \delta \varrho \vartheta) \mathbf{u}) - \operatorname{div}_x((\kappa_F + \kappa_R \vartheta^3 + \delta \vartheta^\Gamma) \nabla_x \vartheta) \quad (91)$$

$$= \mathbb{S} : \nabla_x \mathbf{u} - p \operatorname{div}_x(\mathbf{u}) + \frac{1}{\sigma} |\operatorname{curl}_x \mathbf{H}|^2 + \varepsilon \Gamma |\nabla_x \varrho|^2 \varrho^{\Gamma-2},$$

to be satisfied in $(0, T) \times \Omega$, together with no-flux boundary conditions

$$\nabla_x \vartheta \cdot \mathbf{n}|_{\partial} = 0. \quad (92)$$

The initial conditions read

$$\varrho(e + \delta \vartheta)(0, \cdot) = \varrho_{0,\delta}(e(\varrho_{0,\delta}, \vartheta_{0,\delta}) + \delta \vartheta_{0,\delta}), \quad (93)$$

where the approximate temperature satisfies

$$\vartheta_{0,\delta} \in C^1(\overline{\Omega}), \quad \nabla_x \vartheta_{0,\delta} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \inf_{x \in \Omega} \vartheta_{0,\delta}(x) > 0, \quad (94)$$

and

$$\left\{ \begin{array}{l} \vartheta_{0,\delta} \rightarrow \vartheta_0 \text{ in } L^p(\Omega) \text{ for any } p \geq 1, \\ \delta \int_{\Omega} \varrho_{0,\delta} \log(\vartheta_{0,\delta}) dx \rightarrow 0, \\ \int_{\Omega} \varrho_{0,\delta} s(\varrho_{0,\delta}, \vartheta_{0,\delta}) dx \rightarrow \int_{\Omega} \varrho_0 s(\varrho_0, \vartheta_0) dx, \end{array} \right. \quad (95)$$

as $\delta \rightarrow 0$, with

$$\int_{\Omega} \varrho_{0,\delta} e(\varrho_{0,\delta}, \vartheta_{0,\delta}) dx < c \text{ uniformly for } \delta > 0. \quad (96)$$

Finally, the magnetic induction vector \mathbf{B} obeys the original Maxwell's equations

$$\partial_t \mathbf{B} + \operatorname{curl}_x(\mathbf{B} \times \mathbf{u}) + \operatorname{curl}_x \left(\frac{1}{\sigma} \operatorname{curl}_x \left(\frac{\mathbf{B}}{\mu} \right) \right) = 0, \quad \operatorname{div}_x \mathbf{B} = 0 \quad (97)$$

in $(0, T) \times \Omega$, supplemented with the initial condition

$$\mathbf{B}(0, \cdot) = \mathbf{B}_{0,\delta}, \quad (98)$$

where, by virtue of Theorem 1.4 in [93], one can take

$$\mathbf{B}_{0,\delta} \in \mathcal{D}(\Omega; \mathbb{R}^3), \quad \operatorname{div}_x \mathbf{B}_{0,\delta} = 0, \quad (99)$$

$$\mathbf{B}_{0,\delta} \rightarrow \mathbf{B}_0 \text{ in } L^2(\Omega; \mathbb{R}^3) \text{ for } \delta \rightarrow 0. \quad (100)$$

For given positive parameters ε, δ , and $\Gamma > 8$, the proof of Theorem 2.1 consists in the following steps:

Step 1 Solving problem (81)-(100) for fixed $\varepsilon > 0, \delta > 0$.

Step 2 Passing to the limit for $\varepsilon \rightarrow 0$.

Step 3 Letting $\delta \rightarrow 0$.

One sees that Step 1 can be achieved by using a simple fixed point argument: equation (86) is solved in terms of the velocity \mathbf{u} with the help of the Faedo-Galerkin method, where ϱ, ϑ , and \mathbf{B} are computed successively from (81), (91), and (97) as functions of \mathbf{u} . It is easy to check that the corresponding approximate solutions satisfy the energy balance

$$\frac{d}{dt} \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + \mathcal{M}(|\mathbf{H}|) + \frac{\delta}{\Gamma - 1} \varrho^\Gamma + \delta \varrho \vartheta \right) dx = 0, \quad (101)$$

in $\mathcal{D}'(0, T)$, and the limit property

$$\lim_{t \rightarrow 0^+} \int_{\Omega} \left(\frac{1}{2} \varrho |\mathbf{u}|^2 + \varrho e + \mathcal{M}(|\mathbf{H}|) + \frac{\delta}{\Gamma - 1} \varrho^\Gamma + \delta \varrho \vartheta \right) dx = \quad (102)$$

$$\int_{\Omega} \left(\frac{1}{2 \varrho_{0,\delta}} |(\varrho \mathbf{u})_{0,\delta}|^2 + \varrho_{0,\delta} e(\varrho_{0,\delta}, \vartheta_{0,\delta}) + \mathcal{M}(|\mathbf{H}_{0,\delta}|) + \frac{\delta}{\Gamma - 1} \varrho_{0,\delta}^\Gamma + \delta \varrho_{0,\delta} \vartheta_{0,\delta} \right) dx.$$

The reason for splitting Steps 2 and 3 into the ε and δ -parts is a refined density estimates based on multipliers $\nabla_x (-\Delta)^{-1} [\varrho^\beta]$, where one must take $\beta = 1$ when the artificial viscosity term $\varepsilon \Delta \varrho$ is present, while uniform estimates require β to be a small positive number. For this reason, we focus only on Step 3: the goal is to establish the weak sequential stability property for the solution set of the approximate problem:

1. The density $\varrho_\delta \geq 0$ and the velocity \mathbf{u}_δ satisfy

$$\int_0^T \int_{\Omega} \left(\varrho_\delta \partial_t \varphi + \varrho_\delta \mathbf{u}_\delta \cdot \nabla_x \varphi \right) dx dt + \int_{\mathbb{R}^3} \varrho_{0,\delta} \varphi(0, \cdot) dx = 0 \quad (103)$$

for any test function $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$. In addition,

$$\mathbf{u}_\delta(t) \in W_0^{1,2}(\Omega; \mathbb{R}^3) \text{ for a.e. } t \in (0, T). \quad (104)$$

2. The momentum equation

$$\int_0^T \int_{\Omega} \left(\varrho_\delta \mathbf{u}_\delta \cdot \partial_t \varphi + (\varrho_\delta \mathbf{u}_\delta \otimes \mathbf{u}_\delta) : \nabla_x \varphi + p_\delta \operatorname{div}_x(\varphi) + \delta \varrho_\delta^\Gamma \operatorname{div}_x(\varphi) \right) dx dt = \quad (105)$$

$$\int_0^T \int_{\Omega} \left(\mathbb{S}_\delta : \nabla_x \varphi - (\mathbf{j}_\delta \times \mathbf{B}_\delta) \cdot \varphi \right) dx dt - \int_{\Omega} (\varrho \mathbf{u})_{0,\delta} \cdot \varphi(0, \cdot) dx$$

holds for any $\varphi \in \mathcal{D}([0, T] \times \Omega; \mathbb{R}^3)$, where $p_\delta = p(\varrho_\delta, \vartheta_\delta)$, and $\mathbb{S}_\delta, \mathbf{J}_\delta$ are determined in terms of $\mathbf{u}_\delta, \vartheta_\delta$, and \mathbf{B}_δ through the constitutive relations.

3. The entropy production inequality reads now

$$\begin{aligned}
& \int_0^T \int_{\Omega} \left((\varrho_{\delta} s_{\delta} + \delta \varrho_{\delta} \log(\vartheta_{\delta})) \partial_t \varphi + (\varrho_{\delta} s_{\delta} \mathbf{u}_{\delta} + \delta \varrho_{\delta} \log(\vartheta_{\delta}) \mathbf{u}_{\delta}) \cdot \nabla_x \varphi \right) dx dt \\
& \quad + \int_0^T \int_{\Omega} \left(\frac{\mathbf{q}_{\delta}}{\vartheta_{\delta}} \cdot \nabla_x \varphi_{\delta} - \delta \vartheta_{\delta}^{\Gamma-1} \nabla_x \vartheta_{\delta} \cdot \nabla_x \varphi \right) dx dt \\
& \leq \int_0^T \int_{\Omega} \frac{1}{\vartheta_{\delta}} \left(\frac{\mathbf{q}_{\delta} \cdot \nabla_x \vartheta_{\delta}}{\vartheta_{\delta}} - \mathbb{S}_{\delta} : \nabla_x \mathbf{u}_{\delta} - \frac{1}{\sigma_{\delta}} |\mathbf{curl}_x \mathbf{H}_{\delta}|^2 - \delta \vartheta_{\delta}^{\Gamma-1} |\nabla_x \vartheta_{\delta}|^2 \right) \varphi dx dt \\
& \quad - \int_{\Omega} \left(\varrho_{0,\delta} s(\varrho_{0,\delta}, \vartheta_{0,\delta}) + \delta \varrho_{0,\delta} \log(\vartheta_{0,\delta}) \right) \varphi(0, \cdot) dx,
\end{aligned} \tag{106}$$

for any $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3)$, $\varphi \geq 0$. Here, ϑ_{δ} is assumed to be positive a.e. on the set $(0, T) \times \Omega$, $s_{\delta} = s(\varrho_{\delta}, \vartheta_{\delta})$, $\sigma_{\delta} = \sigma(\varrho_{\delta}, \vartheta_{\delta})$, and \mathbf{q}_{δ} is a function of ϱ_{δ} , ϑ_{δ} given by (52).

4. The magnetic induction vector \mathbf{B}_{δ} satisfies

$$\int_0^T \int_{\Omega} \left(\mathbf{B}_{\delta} \cdot \partial_t \varphi - (\mathbf{B}_{\delta} \times \mathbf{u}_{\delta} + \frac{1}{\sigma_{\delta}} \mathbf{curl}_x \mathbf{H}_{\delta}) \cdot \mathbf{curl}_x \varphi \right) dx dt + \int_{\Omega_{\delta}} \mathbf{B}_{0,\delta} \cdot \varphi(0, \cdot) dx = 0, \tag{107}$$

for any vector field $\varphi \in \mathcal{D}([0, T] \times \mathbb{R}^3; \mathbb{R}^3)$.

5. Finally, the (total) energy equality

$$\begin{aligned}
& \int_{\Omega} \left(\frac{1}{2} \varrho_{\delta} |\mathbf{u}_{\delta}|^2 + \varrho_{\delta} e + \mathcal{M}(\mathbf{H}_{\delta}) + \frac{\delta}{\Gamma-1} \varrho_{\delta}^{\Gamma} + \delta \varrho_{\delta} \vartheta_{\delta} \right) (t) dx \\
& = \int_{\Omega} \left(\frac{1}{2 \varrho_{0,\delta}} |(\varrho \mathbf{u})_{0,\delta}|^2 + \varrho_{0,\delta} e(\varrho_{0,\delta}, \vartheta_{0,\delta}) + \mathcal{M}(\mathbf{H}_{0,\delta}) \right) dx + \int_{\Omega_{\delta}} \left(\frac{\delta}{\Gamma-1} \varrho_{0,\delta}^{\Gamma} + \delta \varrho_{0,\delta} \vartheta_{0,\delta} \right) dx,
\end{aligned} \tag{108}$$

holds for a.e. $t \in (0, T)$.

The weak sequential stability problem to be addressed below consists now in showing that one can pass to the limit

$$\varrho_{\delta} \rightarrow \varrho, \quad \mathbf{u}_{\delta} \rightarrow \mathbf{u}, \quad \vartheta_{\delta} \rightarrow \vartheta, \quad \mathbf{B}_{\delta} \rightarrow \mathbf{B},$$

as $\delta \rightarrow 0$, in a suitable topology, where the limit quantity $\{\varrho, \mathbf{u}, \vartheta, \mathbf{B}\}$ is a variational solution of problem (57)-(69).

2.2.3 Uniform bounds

One must identify uniform bounds on the sequences $\{\varrho_{\delta}\}_{\delta>0}$, $\{\mathbf{u}_{\delta}\}_{\delta>0}$, $\{\vartheta_{\delta}\}_{\delta>0}$, and $\{\mathbf{B}_{\delta}\}_{\delta>0}$ through total energy balance (108), dissipation inequality (106) and the other relations resulting from (103)-(108).

1. *Total mass conservation.* As ϱ_{δ} , $\varrho_{\delta} \mathbf{u}_{\delta}$ are locally integrable in $[0, T] \times \overline{\Omega}$, it follows from (103) that total mass is a constant of motion, specifically,

$$\int_{\Omega} \varrho_{\delta}(t) dx = \int_{\Omega} \varrho_{0,\delta} dx = M_0 \text{ for a.e. } t \in (0, T). \tag{109}$$

In particular, as $\varrho_{\delta} \geq 0$ and (85) holds, we get

$$\{\varrho_{\delta}\}_{\delta>0} \text{ bounded in } L^{\infty}(0, T; L^1(\Omega)). \tag{110}$$

2. *Energy estimates.* Using (36) and (44) one gets

$$\int_{\Omega} \varrho_{\delta} e(\varrho_{\delta}, \vartheta_{\delta}) dx = \frac{3}{2} \int_{\Omega} p_F(\varrho_{\delta}, \vartheta_{\delta}) dx \geq \frac{3p_{\infty}}{2} \int_{\Omega} \varrho_{\delta}^{\frac{5}{3}} dx.$$

Total energy balance (108) together with bounds on the initial data (85), (89), (90), (96), and (100) yield the energy estimates

$$\{\varrho_\delta\}_{\delta>0} \text{ bounded in } L^\infty(0, T; L^{\frac{5}{3}}(\Omega)), \quad (111)$$

$$\{\vartheta_\delta\}_{\delta>0} \text{ bounded in } L^\infty(0, T; L^4(\Omega)), \quad (112)$$

$$\{\mathbf{H}_\delta\}_{\delta>0} \text{ bounded in } L^\infty(0, T; L^2(\Omega; \mathbb{R}^3)), \quad (113)$$

$$\{\varrho_\delta |\mathbf{u}_\delta|^2\}_{\delta>0}, \{\varrho_\delta e(\varrho_\delta, \vartheta_\delta)\}_{\delta>0} \text{ bounded in } L^\infty(0, T; L^1(\Omega)), \quad (114)$$

and

$$\delta \int_\Omega \varrho_\delta^\Gamma dx \leq c \text{ uniformly with respect to } \delta > 0. \quad (115)$$

In particular, using Hölder's inequality, (111) and (114) imply

$$\{(\varrho \mathbf{u})_\delta\}_{\delta>0} \text{ bounded in } L^\infty(0, T; L^{\frac{5}{4}}(\Omega; \mathbb{R}^3)). \quad (116)$$

3. Dissipation estimates. Taking a spatially homogeneous test function φ such that $\varphi(0, \cdot) = 1$ in the entropy production inequality (106) and using (95) (96), we obtain

$$\int_0^\tau \int_\Omega \frac{1}{\vartheta_\delta} \left(\mathbb{S}_\delta : \nabla_x \mathbf{u}_\delta - \frac{\mathbf{q}_\delta \cdot \nabla_x \vartheta_\delta}{\vartheta_\delta} + \frac{1}{\sigma_\delta} |\mathbf{curl}_x \mathbf{H}_\delta|^2 + \delta \vartheta_\delta^{\Gamma-1} |\nabla_x \vartheta_\delta|^2 \right) dx dt \quad (117)$$

$$\leq c_1 + \int_\Omega \varrho_\delta(\tau) s(\varrho_\delta(\tau), \vartheta_\delta(\tau)) + \delta \varrho_\delta(\tau) \log(\vartheta_\delta(\tau)) dx \leq c_2 + \int_\Omega \varrho_\delta(\tau) s(\varrho_\delta(\tau), \vartheta_\delta(\tau)) dx,$$

for a.e. $\tau \in (0, T)$, where the last inequality follows from (111), (112).

Using (49) and (50), the integral in the right hand side may be estimated

$$\begin{aligned} \int_\Omega \varrho_\delta s(\varrho_\delta, \vartheta_\delta) dx &= \int_\Omega \frac{4}{3} a \vartheta_\delta^3 + \varrho_\delta S_F \left(\frac{\varrho_\delta}{\vartheta_\delta^{\frac{3}{2}}} \right) dx \leq c_1 - c_2 \int_{\{\varrho_\delta \leq \vartheta_\delta^{\frac{3}{2}}\}} \varrho_\delta \left(\log(\varrho_\delta) - \frac{3}{2} \log(\vartheta_\delta) \right) dx \\ &\leq c_3 + c_4 \int_{\{\varrho_\delta \leq \vartheta_\delta^{\frac{3}{2}}\}} \varrho_\delta \log(\vartheta_\delta) dx \leq c_5 + c_6 \int_\Omega \varrho_\delta \vartheta_\delta dx \leq c_7, \end{aligned} \quad (118)$$

where we have taken into account (111) and (112).

Thus the integral in the left hand side of (117) is bounded independently of δ and we conclude after (53)-(56) that

$$\left\{ (1 + \vartheta_\delta)^{\frac{\alpha-1}{2}} \langle \nabla_x \mathbf{u}_\delta \rangle \right\}_{\delta>0} \text{ is bounded in } L^2(0, T; L^2(\Omega; \mathbb{R}^{3 \times 3})), \quad (119)$$

$$\left\{ (1 + \vartheta_\delta)^{\frac{\alpha-1}{2}} \operatorname{div}_x (\mathbf{u}_\delta) \right\}_{\delta>0} \text{ is bounded in } L^2(0, T; L^2(\Omega)), \quad (120)$$

$$\{\nabla_x \log(\vartheta_\delta)\}_{\delta>0}, \left\{ \nabla_x \vartheta_\delta^{\frac{3}{2}} \right\}_{\delta>0} \text{ are bounded in } L^2(0, T; L^2(\Omega; \mathbb{R}^3)), \quad (121)$$

$$\left\{ (1 + \vartheta_\delta)^{\frac{\alpha-1}{2}} \mathbf{curl}_x \mathbf{H}_\delta \right\}_{\delta>0} \text{ is bounded in } L^2(0, T; L^2(\Omega; \mathbb{R}^3)), \quad (122)$$

and

$$\left\{ \sqrt{\delta} \nabla_x \vartheta_\delta^{\frac{\Gamma}{2}} \right\}_{\delta>0} \text{ is bounded in } L^2(0, T; L^2(\Omega; \mathbb{R}^3)), \quad (123)$$

where α satisfies (70), and where we denote by

$$\langle \mathbb{Q} \rangle = \frac{1}{2} (\mathbb{Q} + \mathbb{Q}^T) - \frac{1}{3} \operatorname{trace}(\mathbb{Q}) \mathbb{I},$$

the traceless component of the symmetric part of a tensor \mathbb{Q} .

Since the velocity field satisfies (104) we find

$$\{\mathbf{u}_\delta\}_{\delta>0} \text{ bounded in } L^2(0, T; W_0^{1,2}(\Omega; \mathbb{R}^3)), \quad (124)$$

and (121) and (112) lead to

$$\left\{ \vartheta_\delta^{\frac{3}{2}} \right\}_{\delta>0} \text{ bounded in } L^2(0, T; W^{1,2}(\Omega)). \quad (125)$$

4. *Positivity of the absolute temperature.* According to physics, temperature must be positive a.e. on $(0, T) \times \Omega$. To justify that we use the uniform L^2 - estimates of $\nabla_x \log(\vartheta_\delta)$ given by (124) and the following version of Poincaré's inequality:

Lemma 2.1 *Let $\Omega \subset \mathbb{R}^N$ be a bounded Lipschitz domain, and $\omega \geq 1$ be a given constant. Furthermore, assume $O \subset \Omega$ is a measurable set such that $|O| \geq \underline{\omega} > 0$. Then*

$$\|v\|_{W^{1,2}(\Omega)} \leq c(\underline{\omega}, \Omega, \omega) \left(\|\nabla_x v\|_{L^2(\Omega; \mathbb{R}^3)} + \left(\int_O |v|^{\frac{1}{\omega}} dx \right)^\omega \right).$$

In order to apply Lemma 2.1, we show first the uniform bound

$$0 < \underline{T} \leq \int_\Omega \vartheta_\delta(t) dx \text{ for a.e. } t \in (0, T). \quad (126)$$

As a consequence of (106) we get

$$\int_\Omega \left(\varrho_\delta s(\varrho_\delta, \vartheta_\delta) + \delta \varrho_\delta \log(\vartheta_\delta) \right) (t) dx \geq \int_\Omega \left(\varrho_{0,\delta} s(\varrho_{0,\delta}, \vartheta_{0,\delta}) + \delta \varrho_{0,\delta} \log(\vartheta_{0,\delta}) \right) dx,$$

where, by virtue of (85) and (95)

$$\delta \int_\Omega \varrho_{0,\delta} \log(\vartheta_{0,\delta}) dx \rightarrow 0 \quad \text{and} \quad \int_\Omega \varrho_{0,\delta} s(\varrho_{0,\delta}, \vartheta_{0,\delta}) dx \rightarrow \int_\Omega \varrho_0 s(\varrho_0, \vartheta_0) dx.$$

On the other hand (111) and (112) yield

$$\delta \int_\Omega \varrho_\delta \log(\vartheta_\delta)(t) dx \leq \delta \int_\Omega \varrho_\delta(t) \vartheta_\delta(t) dx \rightarrow 0 \quad \text{for a.e. } t \in (0, T),$$

while, using (46)

$$\int_\Omega \varrho_\delta s(\varrho_\delta, \vartheta_\delta)(t) dx = \int_\Omega \left(\frac{4}{3} a \vartheta_\delta^3 + \varrho_\delta S_F \left(\frac{\varrho_\delta}{\vartheta_\delta^{\frac{3}{2}}} \right) \right) (t) dx.$$

Supposing the opposite of (126), one could extract a sequence $\{\vartheta_\delta(t_\delta)\}_{\delta>0}$ such that

$$\vartheta_\delta(t_\delta) \rightarrow 0 \text{ weakly in } L^4(\Omega), \text{ and strongly in } L^p(\Omega) \text{ for any } 1 \leq p < 4, \quad (127)$$

$$\int_\Omega \left(\frac{4}{3} a \vartheta_0^3 + \varrho_0 S_F \left(\frac{\varrho_0}{\vartheta_0^{\frac{3}{2}}} \right) \right) dx \leq \liminf_{\delta \rightarrow 0^+} \int_\Omega \varrho_\delta(t_\delta) S_F \left(\frac{\varrho_\delta}{\vartheta_\delta^{\frac{3}{2}}} \right) (t_\delta) dx, \quad (128)$$

where

$$\varrho_\delta(t_\delta) \rightarrow \varrho(t) \text{ weakly in } L^{\frac{5}{3}}(\Omega), \quad \int_\Omega \varrho_\delta(t_\delta) dx = \int_\Omega \varrho_{0,\delta} dx.$$

Now, for any fixed $K > 1$, one can write

$$\begin{aligned}
& \int_{\Omega} \varrho_{\delta}(t_{\delta}) S_F\left(\frac{\varrho_{\delta}}{\vartheta_{\delta}^{\frac{3}{2}}}\right)(t_{\delta}) \, dx \tag{129} \\
& \leq \int_{\{\varrho_{\delta}(t_{\delta}) \leq K \vartheta_{\delta}^{\frac{3}{2}}(t_{\delta})\}} \varrho_{\delta}(t_{\delta}) S_F\left(\frac{\varrho_{\delta}}{\vartheta_{\delta}^{\frac{3}{2}}}\right)(t_{\delta}) \, dx + \int_{\{\varrho_{\delta}(t_{\delta}) \geq K \vartheta_{\delta}^{\frac{3}{2}}(t_{\delta})\}} \varrho_{\delta}(t_{\delta}) S_F\left(\frac{\varrho_{\delta}}{\vartheta_{\delta}^{\frac{3}{2}}}\right)(t_{\delta}) \, dx \\
& \leq c \int_{\{\varrho_{\delta}(t_{\delta}) \leq K \vartheta_{\delta}^{\frac{3}{2}}(t_{\delta})\}} \varrho_{\delta}(t_{\delta}) \left(1 + \frac{\vartheta_{\delta}^{\frac{3}{2}}}{\varrho_{\delta}}(t_{\delta})\right) \, dx + S_F(K) \int_{\Omega} \varrho_{0,\delta} \, dx \\
& \leq 2c \int_{\Omega} \vartheta_{\delta}^{\frac{3}{2}}(t_{\delta}) \, dx + S_F(K) \int_{\Omega} \varrho_{0,\delta} \, dx.
\end{aligned}$$

Thus combining (127)-(129) one concludes

$$\int_{\Omega} \frac{4}{3} a \vartheta_0^3 + \varrho_0 S_F\left(\frac{\varrho_0}{\vartheta_0^{\frac{3}{2}}}\right) \, dx \leq S_F(K) \int_{\Omega} \varrho_0 \, dx \text{ for any } K > 1,$$

which is in contradiction with the fact that S_F is strictly decreasing with $\lim_{K \rightarrow \infty} S_F(K) < 0$, and ϑ_0 positive (non-zero) on Ω . This proves the uniform bound (126).

Finally, observing that

$$\underline{T} - \varepsilon |\Omega| \leq \int_{\{\vartheta_{\delta}(t) > \varepsilon\}} \vartheta_{\delta}(t) \, dx \leq |\{\vartheta_{\delta}(t) > \varepsilon\}|^{\frac{3}{4}} \|\vartheta_{\delta}(t)\|_{L^4(\Omega)} \text{ for any } \varepsilon > 0,$$

we deduce, using (112), that there exist $\varepsilon > 0$ and $\underline{\varrho}$ such that

$$|\{\vartheta_{\delta}(t) > \varepsilon\}| > \underline{\varrho} > 0 \text{ for a.e. } t \in (0, T),$$

uniformly for $\delta > 0$. Then using (121) one can apply Lemma 2.1 to $\log(\vartheta_{\delta})$ and we obtain the required estimate

$$\{\log(\vartheta_{\delta})\}_{\delta > 0} \text{ bounded in } L^2(0, T; W^{1,2}(\Omega)). \tag{130}$$

5. *Estimates of the magnetic field.* Using the integral identity (107) we get

$$\operatorname{div}_x(\mathbf{B}_{\delta})(t) = 0 \text{ in } \mathcal{D}'(\Omega), \quad \mathbf{B}_{\delta} \cdot \mathbf{n}|_{\partial\Omega} = 0, \tag{131}$$

which, together with estimates (113), (122) and Theorem 6.1 in [32], yields

$$\{\mathbf{H}_{\delta}\}_{\delta > 0} \text{ bounded in } L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)). \tag{132}$$

6. *Pressure estimates.* We first observe that estimates (111), (124) imply boundedness of the sequences $\{\varrho_{\delta} \mathbf{u}_{\delta}\}_{\delta > 0}$ and $\{\varrho_{\delta} \mathbf{u}_{\delta} \otimes \mathbf{u}_{\delta}\}_{\delta > 0}$ in $L^p((0, T) \times \Omega)$ for a certain $p > 1$.

Now one gets

$$\mathbb{S}_{\delta} = \sqrt{\vartheta_{\delta} \nu(\vartheta_{\delta})} \sqrt{\frac{\nu(\vartheta_{\delta})}{\vartheta_{\delta}}} \langle \nabla_x \mathbf{u}_{\delta} \rangle + \sqrt{\vartheta_{\delta} \eta(\vartheta_{\delta})} \sqrt{\frac{\eta(\vartheta_{\delta})}{\vartheta_{\delta}}} \operatorname{div}_x(\mathbf{u}_{\delta}), \tag{133}$$

where, by virtue of (112), (119), (125), and hypothesis (70), the expression on the right-hand side is bounded in $L^p((0, T) \times \Omega)$ for a certain $p > 1$.

Finally the Lorentz force $\mathbf{j}_{\delta} \times \mathbf{B}_{\delta} = \operatorname{curl}_x \mathbf{H}_{\delta} \times \mu(\mathbf{H}_{\delta}) \mathbf{H}_{\delta}$ is controlled by (113), (132) and (56) which, combined with a simple interpolation argument, yields

$$\{\mathbf{j}_{\delta} \times \mathbf{B}_{\delta}\}_{\delta > 0} \text{ bounded in } L^p((0, T) \times \Omega) \text{ for a certain } p > 1. \tag{134}$$

At this stage, using the main result of [44], we are allowed to use the quantities

$$\varphi(t, x) = \psi(t)\mathcal{B}[\varrho_\delta^\omega], \quad \psi \in \mathcal{D}(0, T) \quad \text{for a sufficiently small parameter } \omega > 0,$$

as test functions in (105), where \mathcal{B} is the Bogovskii operator [9] solving the boundary value problem

$$\operatorname{div}_x \left(\mathcal{B}[v] \right) = v - \frac{1}{|\Omega|} \int_\Omega v \, dx, \quad \mathcal{B}|_{\partial\Omega} = 0. \quad (135)$$

The resulting estimate reads

$$\int_0^T \int_\Omega \left(p(\varrho_\delta, \vartheta_\delta) \varrho_\delta^\omega + \delta \varrho_\delta^{\Gamma+\omega} \right) dx \, dt < c, \quad \text{with } c \text{ independent of } \delta, \quad (136)$$

then

$$\{p(\varrho_\delta, \vartheta_\delta)\}_{\delta>0} \text{ is bounded in } L^p((0, T) \times \Omega) \text{ for a certain } p > 1, \quad (137)$$

and

$$\left\{ \varrho_\delta^{\frac{5}{3}+\omega} \right\}_{\delta>0} \text{ is bounded in } L^1((0, T) \times \Omega). \quad (138)$$

2.2.4 Sequential stability of the equations

1. *Continuity equation.* With the estimates established in the previous section, we can pass to the limit $\delta \rightarrow 0$ in (103).

Indeed, extracting subsequences if necessary, we deduce from (111),(116), (124), and the fact that ϱ_δ and \mathbf{u}_δ satisfy the identity (103)

$$\varrho_\delta \rightarrow \varrho \text{ in } C([0, T]; L^{\frac{5}{3}}_{weak}(\Omega)), \quad (139)$$

$$\mathbf{u}_\delta \rightarrow \mathbf{u} \text{ weakly in } L^2(0, T; W_0^{1,2}(\Omega; \mathbb{R}^3)), \quad (140)$$

$$\varrho_\delta \mathbf{u}_\delta \rightarrow \varrho \mathbf{u} \text{ weakly-}^* \text{ in } L^\infty(0, T; L^{\frac{5}{4}}(\Omega; \mathbb{R}^3)), \quad (141)$$

where the limit quantities satisfy (57).

2. *Momentum equation.* Using the estimates obtained in Section 2.2.3.7 together with (105) we have

$$\varrho_\delta \mathbf{u}_\delta \rightarrow \varrho \mathbf{u} \text{ in } C([0, T]; L^{\frac{5}{4}}(\Omega; \mathbb{R}^3)), \quad (142)$$

$$\varrho_\delta \mathbf{u}_\delta \otimes \mathbf{u}_\delta \rightarrow \varrho \mathbf{u} \otimes \mathbf{u} \text{ weakly in } L^2(0, T; L^{\frac{30}{29}}(\Omega; \mathbb{R}^{3 \times 3})), \quad (143)$$

where we have used the embedding $W_0^{1,2}(\Omega) \hookrightarrow L^6(\Omega)$.

Furthermore, in accordance with (136), (137)

$$p(\varrho_\delta, \vartheta_\delta) \rightarrow \overline{p(\varrho, \vartheta)} \text{ weakly in } L^p((0, T) \times \Omega), \quad (144)$$

and

$$\delta \varrho_\delta^\Gamma \rightarrow 0 \text{ in } L^p((0, T) \times \Omega), \quad (145)$$

for a certain $p > 1$. Here, in agreement with Section 2.2.3,

$$\int_0^T \int_\Omega \overline{p(\varrho, \vartheta)} \varphi \, dx \, dt = \int_0^T \int_\Omega \varphi \left(\int_{\mathbb{R}^2} p(\varrho, \vartheta) \, d\Lambda_{t,x}(\varrho, \vartheta) \right) dx \, dt, \quad \varphi \in \mathcal{D}((0, T) \times \Omega),$$

where $\Lambda_{t,x}(\varrho, \vartheta)$ is a parametrized (Young) measure associated to the (vector valued) sequence $\{\varrho_\delta, \vartheta_\delta\}_{\delta>0}$.

Similarly, one can use (107) together with estimates (113), (132) to deduce

$$\mathbf{H}_\delta \rightarrow \mathbf{H} \text{ weakly in } L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)) \text{ and strongly in } L^2((0, T) \times \Omega; \mathbb{R}^3), \quad (146)$$

and

$$\mu(\mathbf{H}_\delta) \rightarrow \mu(\mathbf{H}) \text{ strongly in } L^2((0, T) \times \Omega; \mathbb{R}^3),$$

then

$$\mathbf{j}_\delta \times \mathbf{B}_\delta \rightarrow \mathbf{j} \times \mathbf{B} \text{ weakly in } L^p((0, T) \times \Omega; \mathbb{R}^3),$$

for a certain $p > 1$. Thus the limit quantities satisfy an averaged momentum equation

$$\begin{aligned} & \int_0^T \int_\Omega \left((\varrho \mathbf{u}) \cdot \partial_t \varphi + (\varrho \mathbf{u} \otimes \mathbf{u}) : \nabla_x \varphi + \overline{p(\varrho, \vartheta)} \operatorname{div}_x(\varphi) \right) dx dt \\ &= \int_0^T \int_\Omega \left(\overline{\mathbb{S}} : \nabla_x \varphi - (\mathbf{j} \times \mathbf{B}) \cdot \varphi \right) dx dt - \int_\Omega (\varrho \mathbf{u})_0 \cdot \varphi(0, \cdot) dx \end{aligned} \quad (147)$$

for any vector field $\varphi \in \mathcal{D}([0, T) \times \Omega; \mathbb{R}^3)$.

The symbol $\overline{\mathbb{S}}$ denotes a weak limit in $L^p(0, T; L^p(\Omega; \mathbb{R}_{sym}^{3 \times 3}))$, $p > 1$ of the approximate viscosity tensors \mathbb{S}_δ specified in (133). Clearly, relation (147) will coincide with the (variational) momentum equation (60) as soon as we show strong (pointwise) convergence of the sequences $\{\varrho_\delta\}_{\delta>0}$ and $\{\vartheta_\delta\}_{\delta>0}$.

2.2.5 Entropy inequality and strong convergence of the temperature

1. *Entropy inequality.* In order to obtain information on the time oscillations of the sequence $\{\vartheta_\delta\}_{\delta>0}$, we use estimates on $\partial_t(\varrho_\delta s(\varrho_\delta, \vartheta_\delta))$ provided by the approximate entropy balance (106).

It first follows from (49), (50) that

$$|\varrho_\delta s(\varrho_\delta, \vartheta_\delta)| \leq c \left(\vartheta_\delta^3 + \varrho_\delta |\log(\varrho_\delta)| + \varrho_\delta |\log(\vartheta_\delta)| \right),$$

therefore, by virtue of the uniform estimates (111), (112), (116), and (130), we can assume

$$\varrho_\delta s(\varrho_\delta, \vartheta_\delta) \rightarrow \overline{\varrho s(\varrho, \vartheta)} \text{ weakly in } L^p((0, T) \times \Omega), \quad (148)$$

$$\varrho_\delta s(\varrho_\delta, \vartheta_\delta) \mathbf{u}_\delta \rightarrow \overline{\varrho s(\varrho, \vartheta) \mathbf{u}} \text{ weakly in } L^p((0, T) \times \Omega; \mathbb{R}^3), \quad (149)$$

for a certain $p > 1$.

Similarly, one estimates the entropy flux

$$\left| \frac{\mathbf{q}}{\vartheta_\delta} + \delta \vartheta_\delta^{\Gamma-1} \nabla_x \vartheta_\delta \right| \leq c \left(|\nabla_x \log(\vartheta_\delta)| + \vartheta_\delta^2 |\nabla_x \vartheta_\delta| + \delta \vartheta_\delta^{\Gamma-1} |\nabla_x \vartheta_\delta| \right), \quad (150)$$

where

$$\vartheta_\delta^2 |\nabla_x \vartheta_\delta| = \frac{2}{3} \vartheta_\delta^{\frac{3}{2}} \left| \nabla_x \vartheta_\delta^{\frac{3}{2}} \right|, \quad \delta \vartheta_\delta^{\Gamma-1} \nabla_x \vartheta_\delta = \frac{2}{\Gamma} \sqrt{\delta} \left| \nabla_x \vartheta_\delta^{\frac{\Gamma}{2}} \right| \sqrt{\delta} \vartheta_\delta^{s \frac{\Gamma}{2}} \vartheta_\delta^{(1-s) \frac{\Gamma}{2}}.$$

Choosing the parameter $0 < s < 1$ small enough so that

$$(1-s) \frac{\Gamma}{2} = 1,$$

we have after Hölder's inequality

$$\left\| \delta \vartheta_\delta^{\Gamma-1} \nabla_x \vartheta_\delta \right\|_{L^p(\Omega; \mathbb{R}^3)} \leq c \left\| \sqrt{\delta} \nabla_x \vartheta_\delta^{\frac{\Gamma}{2}} \right\|_{L^2(\Omega; \mathbb{R}^3)} \left\| \vartheta_\delta \right\|_{L^4(\Omega)} \left\| \sqrt{\delta} \vartheta_\delta^{s \frac{\Gamma}{2}} \right\|_{L^6(\Omega)}.$$

Then using estimates (112), (123) together with the imbedding $W^{1,2}(\Omega) \hookrightarrow L^6(\Omega)$, we get

$$\delta \vartheta_\delta^{\Gamma-1} \nabla_x \vartheta_\delta \rightarrow 0 \text{ in } L^p((0, T) \times \Omega; \mathbb{R}^3) \text{ for a certain } p > 1. \quad (151)$$

Moreover using similar arguments one can also show that

$$\left\{ \frac{\mathbf{q}}{\vartheta_\delta} \right\}_{\delta>0} \text{ is bounded in } L^p((0, T) \times \Omega; \mathbb{R}^3), \text{ for a certain } p > 1. \quad (152)$$

On the other hand according to (121)

$$b(\vartheta_\delta) \rightarrow \overline{b(\vartheta)} \text{ weakly in } L^q((0, T) \times \Omega), \text{ and weakly in } L^2(0, T; W^{1,2}(\Omega)), \quad (153)$$

for any finite $q \geq 1$ and any function b provided both b and b' are uniformly bounded.

Now, as a consequence of the entropy balance (106), we have

$$\begin{aligned} \text{Div}_{t,x} \left[\varrho_\delta s(\varrho_\delta, \vartheta_\delta) + \delta \varrho_\delta \log(\vartheta_\delta), (\varrho_\delta s(\varrho_\delta, \vartheta_\delta) + \delta \varrho_\delta \log(\vartheta_\delta)) \mathbf{u}_\delta + \frac{\mathbf{q}}{\vartheta_\delta} - \delta \vartheta_\delta^{\Gamma-1} \nabla_x \vartheta_\delta \right] \\ \geq 0 \quad \text{in } \mathcal{D}'((0, T) \times \Omega), \end{aligned}$$

while (153) yields

$$\text{Curl}_{t,x} \left[b(\vartheta_\delta), 0, 0, 0 \right] \text{ bounded in } L^2((0, T) \times \Omega; \mathbb{R}^{4 \times 4}).$$

Thus a direct application of the Div-Curl lemma [80, 92] gives the relation

$$\overline{\varrho s(\varrho, \vartheta) b(\vartheta)} = \overline{\varrho s(\varrho, \vartheta)} \overline{b(\vartheta)}, \quad (154)$$

for any b as in (153).

Moreover, seeing that the sequence $\{\varrho_\delta s(\varrho_\delta, \vartheta_\delta) \vartheta_\delta\}_{\delta>0}$ is bounded in $L^p((0, T) \times \Omega)$ we deduce from (155) the formula

$$\overline{\varrho s(\varrho, \vartheta) \vartheta} = \overline{\varrho s(\varrho, \vartheta)} \vartheta. \quad (155)$$

2. Parametrized measures and pointwise convergence of the temperature

To show that (155) implies strong convergence of the sequence $\{\vartheta_\delta\}_{\delta>0}$, we first remark that the approximate solutions solve the renormalized continuity equation

$$\partial_t b(\varrho_\delta) + \text{div}_x (b(\varrho_\delta) \mathbf{u}_\delta) + \left(b'(\varrho_\delta) \varrho_\delta - b(\varrho_\delta) \right) \text{div}_x (\mathbf{u}_\delta) = 0 \quad \text{in } \mathcal{D}'((0, T) \times \mathbb{R}^3), \quad (156)$$

provided that ϱ_δ and \mathbf{u}_δ are extended by zero outside Ω , and for any continuously differentiable function b whose derivative vanishes for large arguments.

Functions ϱ_δ being square-integrable because of the artificial pressure in the energy equality, (156) follows from (103) via the technique of DiPerna and Lions [28].

Now it follows from (156) that

$$b(\varrho_\delta) \rightarrow \overline{b(\varrho)} \text{ in } C([0, T]; L^p_{weak}(\Omega)) \quad \text{for any finite } p > 1 \text{ and bounded } b. \quad (157)$$

Relations (157) and (153) yield

$$\overline{g(\varrho) h(\vartheta)} = \overline{g(\varrho)} \overline{h(\vartheta)},$$

or, in terms of the corresponding parametrized measures (see Section 2.2.1)

$$\Lambda_{t,x}(\varrho, \vartheta) = \Lambda_{t,x}(\varrho) \otimes \Lambda_{t,x}(\vartheta). \quad (158)$$

Relation (158) means that oscillations in the sequences $\{\varrho_\delta\}_{\delta>0}$ and $\{\vartheta_\delta\}_{\delta>0}$ are orthogonal, in the sense that the parametrized measure associated to $\{\varrho_\delta, \vartheta_\delta\}_{\delta>0}$ can be written as a tensor product of the parametrized measures generated by $\{\varrho_\delta\}_{\delta>0}$ and $\{\vartheta_\delta\}_{\delta>0}$.

Let us consider a function $H = H(t, x, r, z)$ defined for $t \in (0, T)$, $x \in \Omega$, and $(r, z) \in \mathbb{R}^2$ through formula

$$H(t, x, r, z) = r \left(s(r, z) - s(r, \vartheta(t, x)) \right) \left(z - \vartheta(t, x) \right).$$

Clearly, H is a Caratheodory function: $H(t, x, \cdot, \cdot)$ is continuous for a.e $(t, x) \in (0, T) \times \Omega$, and $H(\cdot, \cdot, r, z)$ is measurable for any $(r, z) \in \mathbb{R}^2$.

Moreover, as both s_R and s_F are increasing functions of the absolute temperature, we have

$$H(t, x, r, z) \geq \frac{4}{3}a\left(z^3 - \vartheta^3(t, x)\right)\left(z - \vartheta(t, x)\right) \geq 0. \quad (159)$$

At this stage, we use Theorem 6.2 in [83], namely: weak limits of Caratheodory functions can be expressed in terms of the parametrized measure. Accordingly, we obtain

$$\begin{aligned} \lim_{\delta \rightarrow 0} \int_0^T \int_{\Omega} \varphi(t, x) H(t, x, \varrho_{\delta}, \vartheta_{\delta}) \, dx \, dt &= \int_0^T \int_{\Omega} \varphi(t, x) \left(\int_{\mathbb{R}^2} H(t, x, \varrho, \vartheta) \, d\Lambda_{t,x}(\varrho, \vartheta) \right) \, dx \, dt \\ &= \int_0^T \int_{\Omega} \varphi(t, x) \left(\int_{\mathbb{R}^2} \varrho s(\varrho, \vartheta) (\vartheta - \vartheta(t, x)) \, d\Lambda_{t,x}(\varrho, \vartheta) \right) \, dx \, dt \\ &\quad - \int_0^T \int_{\Omega} \varphi(t, x) \left(\int_{\mathbb{R}^2} \varrho s(\varrho, \vartheta(t, x)) (\vartheta - \vartheta(t, x)) \, d\Lambda_{t,x}(\varrho, \vartheta) \right) \, dx \, dt \\ &= \int_0^T \int_{\Omega} \varphi(t, x) \left(\int_{\mathbb{R}^2} \varrho s(\varrho, \vartheta) (\vartheta - \vartheta(t, x)) \, d\Lambda_{t,x}(\varrho, \vartheta) \right) \, dx \, dt \\ &= \int_0^T \int_{\Omega} \varphi \left(\overline{\varrho s(\varrho, \vartheta) \vartheta} - \overline{\varrho s(\varrho, \vartheta) \vartheta} \right) \, dx \, dt = 0 \text{ for any } \varphi \in \mathcal{D}((0, T) \times \Omega), \end{aligned}$$

where we have used (155) to get the last equality together with (158), observing that

$$\int_{\mathbb{R}^2} \varrho s(\varrho, \vartheta(t, x)) (\vartheta - \vartheta(t, x)) \, d\Lambda_{t,x}(\varrho, \vartheta) = \int_{\mathbb{R}} \varrho s(\varrho, \vartheta(t, x)) \, d\Lambda_{t,x}(\varrho) \int_{\mathbb{R}} (\vartheta - \vartheta(t, x)) \, d\Lambda_{t,x}(\vartheta) = 0.$$

In particular, we deduce from (159) that $\overline{\vartheta^3 \vartheta} = \overline{\vartheta^3} \vartheta$, which is equivalent to

$$\vartheta_{\delta} \rightarrow \vartheta \text{ in } L^4((0, T) \times \Omega). \quad (160)$$

2.2.6 Pointwise convergence of densities

1. *The effective viscous pressure.* Let us start with a result of P.-L. Lions [72] on the effective viscous pressure.

In the present setting, it can be stated in terms of the parametrized measures as follows

$$\overline{\psi p(\varrho, \vartheta) b(\varrho)} - \overline{\psi p(\varrho, \vartheta)} \overline{b(\varrho)} = \mathcal{R} : [\psi \mathbb{S}] b(\varrho) - \mathcal{R} : [\psi \mathbb{S}] \overline{b(\varrho)}, \quad (161)$$

for any $\psi \in \mathcal{D}(\Omega)$, and any bounded continuous function b , where $\mathcal{R} = \mathcal{R}_{i,j}$ is the Riesz operator defined by means of the Fourier transform:

$$\mathcal{R}_{i,j}[v] = \partial_{x_i} \Delta_x^{-1} \partial_{x_j} v = \mathcal{F}_{\xi \rightarrow x}^{-1} \left[\frac{\xi_i \xi_j}{|\xi|^2} \mathcal{F}_{x \rightarrow \xi}[v] \right]. \quad (162)$$

Note that (161) does not depend on the specific form of the constitutive relations for p , \mathbb{S} and requires only validity of the momentum equation (105) and the renormalized continuity equation (156) for ϱ_{δ} and \mathbf{u}_{δ} .

2. *Commutator estimates.* If viscosity coefficients ν and η were constant, we would have $\mathcal{R}[\mathbb{S}] = (\frac{4}{3}\nu + \eta) \operatorname{div}_x(\mathbf{u})$, and (161) would give immediately

$$\overline{p(\varrho, \vartheta) b(\varrho)} - \overline{p(\varrho, \vartheta)} \overline{b(\varrho)} = \left(\frac{4}{3}\nu + \eta \right) \overline{\operatorname{div}_x(\mathbf{u}) b(\varrho)} - \left(\frac{4}{3}\nu + \eta \right) \operatorname{div}_x(\mathbf{u}) \overline{b(\varrho)}, \quad (163)$$

where the quantity $p - (\frac{4}{3}\nu + \eta)\text{div}_x(\mathbf{u})$ is called the effective viscous pressure.

In order to establish the same relation for variable viscosity coefficients, we write

$$\mathcal{R} : [\psi \mathbb{S}] = \psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u}) + \left\{ \mathcal{R} : [\psi \mathbb{S}] - \psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u}) \right\}.$$

Applying Lemma 4.2 in [40], the quantity

$$\mathcal{R} : \left[\psi \left(\nu(\vartheta_\delta, \mathbf{B}_\delta) \langle \nabla_x \mathbf{u}_\delta \rangle + \eta(\vartheta_\delta, \mathbf{B}_\delta) \text{div}_x(\mathbf{u}_\delta) \mathbb{I} \right) \right] - \psi \left(\frac{4}{3}\nu(\vartheta_\delta, \mathbf{B}_\delta) + \eta(\vartheta_\delta, \mathbf{B}_\delta) \right) \text{div}_x(\mathbf{u}_\delta),$$

is bounded in the space $L^2(0, T; W^{\omega, p}(\Omega))$ for suitable $0 < \omega < 1$, $p > 1$ in terms of the bounds established in (124), (125) and (132) provided that the viscosity coefficients are globally Lipschitz functions of their arguments.

If this is the case, one can use (157) in order to deduce the desired relation

$$\begin{aligned} \overline{\mathcal{R} : [\psi \mathbb{S}] b(\varrho)} - \overline{\mathcal{R} : [\psi \mathbb{S}]} \overline{b(\varrho)} &= \overline{\psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u}) b(\varrho)} - \overline{\psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u})} \overline{b(\varrho)} = \\ &= \overline{\psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u}) b(\varrho)} - \overline{\psi \left(\frac{4}{3}\nu + \eta \right) \text{div}_x(\mathbf{u})} \overline{b(\varrho)}, \end{aligned}$$

where the last equality is a direct consequence of the pointwise convergence proved in (146) and (160).

If ν and η are only continuously differentiable as required by hypotheses of Theorem 2.1, one can write

$$\begin{cases} \nu(\vartheta, \mathbf{B}) = Y(\vartheta, \mathbf{B})\nu(\vartheta, \mathbf{B}) + (1 - Y(\vartheta, \mathbf{B}))\nu(\vartheta, \mathbf{B}), \\ \eta(\vartheta, \mathbf{B}) = Y(\vartheta, \mathbf{B})\eta(\vartheta, \mathbf{B}) + (1 - Y(\vartheta, \mathbf{B}))\eta(\vartheta, \mathbf{B}), \end{cases} \quad (164)$$

where $Y \in \mathcal{D}(\mathbb{R}^2)$ is a suitable function.

Now we have

$$\begin{aligned} &\nu(\vartheta_\delta, \mathbf{B}_\delta) \langle \mathbf{u}_\delta \rangle + \eta(\vartheta_\delta, \mathbf{B}_\delta) \text{div}_x(\mathbf{u}_\delta) \mathbb{I} = \\ &Y(\vartheta_\delta, \mathbf{B}_\delta) \nu(\vartheta_\delta, \mathbf{B}_\delta) \langle \mathbf{u}_\delta \rangle + Y(\vartheta_\delta, \mathbf{B}_\delta) \eta(\vartheta_\delta, \mathbf{B}_\delta) \text{div}_x(\mathbf{u}_\delta) \mathbb{I} + \\ &\left\{ (1 - Y(\vartheta_\delta, \mathbf{B}_\delta)) \nu(\vartheta_\delta, \mathbf{B}_\delta) \langle \mathbf{u}_\delta \rangle + (1 - Y(\vartheta_\delta, \mathbf{B}_\delta)) \eta(\vartheta_\delta, \mathbf{B}_\delta) \text{div}_x(\mathbf{u}_\delta) \mathbb{I} \right\}, \end{aligned}$$

where the expression in the curl brackets can be made arbitrarily small in the norm of $L^p((0, T) \times \Omega)$, with a certain $p > 1$, by a suitable choice of Y (see estimates (112), (119), (125) and formula (133)).

Thus relation (163) holds for any ν and η satisfying hypotheses of Theorem 2.1

3. The oscillations defect measure. In order to describe oscillations in the sequence $\{\varrho_\delta\}_{\delta>0}$, we use the renormalized continuity equation (59) together with its counterpart resulting from letting $\delta \rightarrow 0$ in (156).

Although we have already shown that the limit quantities ϱ and \mathbf{u} satisfy the momentum equation (57), the validity of its renormalization (59) is not clear as the sequence $\{\varrho_\delta\}_{\delta>0}$ is not known to be uniformly square integrable and the Di Perna-Lions machinery [28] does not work a priori.

In order to solve this problem, a concept of *oscillations defect measure* was introduced in [38].

To be more specific we set

$$\text{osc}_q[\varrho_\delta \rightarrow \varrho](Q) = \sup_{k \geq 1} \left(\limsup_{\delta \rightarrow 0} \int_Q |T_k(\varrho_\delta) - T_k(\varrho)|^q \, dx \, dt \right), \quad (165)$$

where T_k are the cut-off functions

$$T_k(z) = \text{sgn}(z) \min\{|z|, k\}.$$

As shown in [38], the limit functions ϱ , \mathbf{u} solve the renormalized equation (59) provided

- $\varrho_\delta, \mathbf{u}_\delta$ satisfy (156),
- $\{\mathbf{u}_\delta\}_{\delta>0}$ is bounded in $L^2(0, T; W^{1,2}(\Omega))$,
-

$$\mathbf{osc}_q[\varrho_\delta \rightarrow \varrho]((0, T) \times \Omega) < \infty \text{ for a certain } q > 2. \quad (166)$$

Then, in order to establish (59) it is clearly enough to show (166).

4. *Weighted estimates of the oscillations defect measure.* In order to prove (166) we use weighted estimates, where the corresponding weight function depends on the absolute temperature.

Taking $b = T_k$ in (163) we get

$$\begin{aligned} & \overline{p(\varrho, \vartheta) T_k(\varrho)} - \overline{p(\varrho, \vartheta)} \overline{T_k(\varrho)} \\ &= \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \overline{\operatorname{div}_x(\mathbf{u}) T_k(\varrho)} - \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \operatorname{div}_x(\mathbf{u}) \overline{T_k(\varrho)}. \end{aligned} \quad (167)$$

As observed in (45), there is a positive constant p_c such that $p_F(\varrho, \vartheta) - p_c \varrho^{\frac{5}{3}}$ is a non-decreasing function of ϱ for any ϑ .

Accordingly we have

$$\overline{p(\varrho, \vartheta) T_k(\varrho)} - \overline{p(\varrho, \vartheta)} \overline{T_k(\varrho)} \geq p_c \left(\overline{\varrho^{\frac{5}{3}} T_k(\varrho)} - \overline{\varrho^{\frac{5}{3}}} \overline{T_k(\varrho)} \right). \quad (168)$$

Now let us choose a weight function $w \in C^1[0, \infty)$

$$w(\vartheta) > 0 \text{ for } \vartheta \geq 0, \quad w(\vartheta) = \vartheta^{-\frac{1+\alpha}{2}} \text{ for } \vartheta \geq 1, \quad (169)$$

where α is the exponent appearing in hypothesis (54).

Multiplying (167) by $w(\vartheta)$ and using (168) we obtain

$$\begin{aligned} & w(\vartheta) \left(\overline{\varrho^{\frac{5}{3}} T_k(\varrho)} - \overline{\varrho^{\frac{5}{3}}} \overline{T_k(\varrho)} \right) \\ & \leq \frac{w(\vartheta)}{p_c} \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \overline{\operatorname{div}_x(\mathbf{u}) T_k(\varrho)} - \frac{w(\vartheta)}{p_c} \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \operatorname{div}_x(\mathbf{u}) \overline{T_k(\varrho)}, \end{aligned} \quad (170)$$

where the right-hand side is a weak limit of

$$\frac{w(\vartheta_\delta)}{p_c} \left(\frac{4}{3} \nu(\vartheta_\delta, \mathbf{B}_\delta) + \eta(\vartheta_\delta, \mathbf{B}_\delta) \right) \operatorname{div}_x(\mathbf{u}_\delta) \left(T_k(\varrho_\delta) - T_k(\varrho) \right) + \frac{w(\vartheta)}{p_c} \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \operatorname{div}_x(\mathbf{u}) \left(T_k(\varrho) - \overline{T_k(\varrho)} \right).$$

Consequently, using the uniform weighted estimates (119) together with hypothesis (19) and the growth restriction on w specified in (169), we conclude that

$$\begin{aligned} & \int_0^T \int_\Omega \left[\frac{w(\vartheta)}{p_c} \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \overline{\operatorname{div}_x(\mathbf{u}) T_k(\varrho)} - \frac{w(\vartheta)}{p_c} \left(\frac{4}{3} \nu(\vartheta, \mathbf{B}) + \eta(\vartheta, \mathbf{B}) \right) \operatorname{div}_x(\mathbf{u}) \overline{T_k(\varrho)} \right] dx dt \\ & \leq c \liminf_{\delta \rightarrow 0} \|T_k(\varrho_\delta) - T_k(\varrho)\|_{L^2((0, T) \times \Omega)}, \text{ with } c \text{ independent of } k. \end{aligned} \quad (171)$$

On the other hand, it was shown in Chapter 6 in [39] that the left-hand side of (170) could be bounded below as

$$\int_0^T \int_\Omega w(\vartheta) \left(\overline{\varrho^{\frac{5}{3}} T_k(\varrho)} - \overline{\varrho^{\frac{5}{3}}} \overline{T_k(\varrho)} \right) dx dt \geq \limsup_{\delta \rightarrow 0} \int_0^T \int_\Omega w(\vartheta) |T_k(\varrho_\delta) - T_k(\varrho)|^{\frac{8}{3}} dx dt. \quad (172)$$

Then (170) together with (171), (172), give the weighted estimate

$$\limsup_{\delta \rightarrow 0} \int_0^T \int_\Omega w(\vartheta) |T_k(\varrho_\delta) - T_k(\varrho)|^{\frac{8}{3}} dx dt \leq c \liminf_{\delta \rightarrow 0} \|T_k(\varrho_\delta) - T_k(\varrho)\|_{L^2((0, T) \times \Omega)}. \quad (173)$$

Now taking $q > 2$ and $\frac{8}{3q}\omega = \frac{1+\alpha}{2}$, we can use Hölder's inequality to obtain

$$\begin{aligned} \int_0^T \int_{\Omega} |T_k(\varrho_\delta) - T_k(\varrho)|^q dx dt &= \int_0^T \int_{\Omega} |T_k(\varrho_\delta) - T_k(\varrho)|^q (1+\vartheta)^{-\omega} (1+\vartheta)^\omega dx dt \\ &\leq c \left(\int_0^T \int_{\Omega} |T_k(\varrho_\delta) - T_k(\varrho)|^{\frac{8}{3}} (1+\vartheta)^{-\frac{1+\alpha}{2}} dx dt + \int_0^T \int_{\Omega} (1+\vartheta)^{\frac{3q(1+\alpha)}{2(8-3q)}} dx dt \right). \end{aligned} \quad (174)$$

On the other hand, in accordance with (112), (125) $\vartheta \in L^\infty(0, T; L^4(\Omega)) \cap L^3(0, T; L^9(\Omega))$, and a simple interpolation argument yields $\vartheta \in L^r((0, T) \times \Omega)$, with $r = \frac{46}{9}$. Consequently, one can find $q > 2$ such that

$$\frac{3q(1+\alpha)}{2(8-3q)} = r = \frac{46}{9},$$

provided α complies with hypothesis (70).

Then finally relations (173), (174) give rise to the desired estimate

$$\limsup_{\delta \rightarrow 0} \int_0^T \int_{\Omega} w(\vartheta) |T_k(\varrho_\delta) - T_k(\varrho)|^q dx dt \leq c \left(1 + \liminf_{\delta \rightarrow 0} \|T_k(\varrho_\delta) - T_k(\varrho)\|_{L^2((0, T) \times \Omega)} \right),$$

yielding (166).

5. Propagation of oscillations and strong convergence. According to (166) we know that the renormalized continuity equation (59) is satisfied by the limit functions ϱ and \mathbf{u} . In particular

$$\partial_t L_k(\varrho) + \operatorname{div}_x(L_k(\varrho)\mathbf{u}) + T_k(\varrho) \operatorname{div}_x(\mathbf{u}) = 0 \text{ in } \mathcal{D}'((0, T) \times \mathbb{R}^3), \quad (175)$$

provided ϱ, \mathbf{u} have been extended to be zero outside Ω , where $L_k(\varrho) = \varrho \int_1^{\varrho} \frac{T_k(z)}{z^2} dz$.

On the other hand, one can let $\delta \rightarrow 0$ in (156) and one obtains

$$\partial_t \overline{L_k(\varrho)} + \operatorname{div}_x(\overline{L_k(\varrho)\mathbf{u}}) + \overline{T_k(\varrho)} \operatorname{div}_x(\mathbf{u}) = 0 \text{ in } \mathcal{D}'((0, T) \times \mathbb{R}^3). \quad (176)$$

Now, following step by step Chapter 3 in [42] we take the difference of (175) and (176) and use (167) and (168) to deduce

$$\int_{\Omega} \overline{L_k(\varrho)}(\tau) - L_k(\varrho)(\tau) dx \leq \int_0^\tau \int_{\Omega} (\operatorname{div}_x(\mathbf{u})) \left(T_k(\varrho) - \overline{T_k(\varrho)} \right) dx dt \text{ for any } \tau \in [0, T],$$

where the right-hand sides vanishes for $k \rightarrow \infty$ due to (166).

Consequently

$$\overline{\varrho \log(\varrho)} = \varrho \log(\varrho) \text{ in } (0, T) \times \Omega,$$

a relation equivalent to strong convergence of $\{\varrho_\delta\}_{\delta>0}$, which means in fact that

$$\varrho_\delta \rightarrow \varrho \text{ in } L^1((0, T) \times \Omega). \quad (177)$$

2.2.7 End of the proof

Note that we have already shown that ϱ, \mathbf{u} satisfy the continuity equation (57) as well as its renormalized version (59). Moreover, it is easy to see that the ‘‘averaged’’ momentum equation (147) coincides in fact with (60) in view of the strong convergence results established in (146), (160) and (177). By the same token, one can pass to the limit in the energy equality (108) in order to obtain (67). Note that the regularizing δ -dependent terms on the left-hand side disappear by virtue of the estimates (111), (112), and (136).

Similarly, one can handle Maxwell's equation (107). Here, the only thing to observe is that the terms

$$\frac{1}{\sigma(\varrho_\delta, \vartheta_\delta, \mathbf{H}_\delta)} \mathbf{curl}_x \mathbf{H}_\delta \text{ are bounded in } L^p((0, T) \times \Omega), \text{ for a certain } p > 1,$$

uniformly with respect to δ (such a bound can be obtained exactly as in (133)).

To conclude, we have to deal with the entropy inequality (106). First of all, the extra terms on the left-hand side tend to zero because of (111), (112), (130) and (151). Furthermore, it is standard to pass to the limit in the entropy production rate keeping the correct sense of the inequality as all terms are convex with respect to the spatial gradients of \mathbf{u} , ϑ and \mathbf{B} .

Finally, the limit in the logarithmic terms can be carried over thanks to estimate (130) and the following result (see Lemma 5.4 in [29]) holds

Lemma 2.2 *Let $\Omega \subset \mathbb{R}^N$ be a bounded Lipschitz domain. Assume that*

$$\vartheta_\delta \rightarrow \vartheta \text{ in } L^2((0, T) \times \Omega) \text{ and } \log(\vartheta_\delta) \rightarrow \overline{\log(\vartheta)} \text{ weakly in } L^2((0, T) \times \Omega).$$

Then ϑ is positive a.e. on $(0, T) \times \Omega$ and $\log(\vartheta) = \overline{\log(\vartheta)}$.

3 Other models

3.1 Polytropic models

The magnetic Navier-Stokes-Poisson case where magnetic permeability μ is a pure positive constant was studied by Ducomet and Feireisl in [30].

In the absence of magnetic field a model of pressure law more adapted to cold plasmas of the type

$$p(\varrho, \vartheta) = p_e(\varrho) + \vartheta p_\vartheta(\varrho), \tag{178}$$

with associated internal energy $e(\varrho, \vartheta)$ such that

$$e_\rho = \frac{1}{\rho^2}(p - \theta p_\theta), \quad e_\vartheta = C_V(\vartheta), \tag{179}$$

introduced by Feireisl [39] in the case of constant viscosities and generalized to temperature dependent viscosities and conductivity in [40], has been extended to the MHD case (with Dirichlet boundary condition $\mathbf{B}|_{\partial\Omega} = 0$ and constant resistivity) by Hu and Wang in [51]:

Theorem 3.1 *Let $\Omega \subset \mathbb{R}^3$ be a bounded domain with boundary of class $C^{2+\delta}$, $\delta > 0$.*

Suppose that the pressure $p = p(\varrho, \vartheta)$ is given by (178), that p_e, p_ϑ are C^1 functions of their arguments on $[0, \infty)$ such that

$$\begin{aligned} p_e(0) = p_\vartheta(0) &= 0, \\ p'_e(\rho) &\geq a_1 \rho^{\gamma-1}, \quad p_\vartheta(\rho) \geq 0 \text{ for all } \rho > 0, \\ p_e(\rho) &\leq a_2 \rho^\gamma, \quad p_\vartheta(\rho) \leq a_3(1 + \rho^{\gamma/3}) \text{ for all } \rho \geq 0. \end{aligned}$$

Furthermore, we suppose that the magnetic permeability μ and the electrical conductivity σ are positive constants and that the transport coefficients

$$\kappa = \kappa(\vartheta), \quad \nu = \nu(\vartheta), \quad \frac{1}{\sigma\mu} := \lambda = \lambda(\vartheta) := \eta(\vartheta) - 2/3, \quad c_V(\theta),$$

are continuously differentiable functions of their argument satisfying

$$c_1(1 + \vartheta^\alpha) \leq \kappa(\vartheta) \leq c_2(1 + \vartheta^\alpha), \quad (180)$$

$$0 \leq \lambda(\vartheta) \leq \bar{\lambda}, \quad (181)$$

$$0 < \underline{\nu} \leq \nu(\vartheta) \leq \bar{\nu}, \quad (182)$$

and

$$0 < \underline{c_V} \leq c_V(\vartheta) \leq \bar{c_V}, \quad (183)$$

for $c_1, c_2, \bar{\lambda}, \underline{\nu}, \bar{\nu}, \underline{c_V}, \bar{c_V} > 0$ and $\alpha > 2$.

Finally assume the homogeneous boundary conditions

$$\mathbf{H}|_{\partial\Omega} = 0, \quad \mathbf{u}|_{\partial\Omega} = 0, \quad \mathbf{q} \cdot \mathbf{n}|_{\partial\Omega} = 0,$$

and let initial data $\varrho_0, (\varrho\mathbf{u})_0, \vartheta_0, \mathbf{B}_0$ be given so that

$$\varrho_0 \in L^{\frac{5}{3}}(\Omega), \quad (\varrho\mathbf{u})_0 \in L^1(\Omega; \mathbb{R}^3), \quad \vartheta_0 \in L^\infty(\mathbb{R}^3), \quad \mathbf{B}_0 \in L^2(\Omega; \mathbb{R}^3), \quad (184)$$

$$\varrho_0 \geq 0, \quad \vartheta_0 > 0, \quad (185)$$

$$(\varrho s)_0 = \varrho_0 s(\varrho_0, \vartheta_0), \quad \frac{1}{\varrho_0} |(\varrho\mathbf{u})_0|^2, \quad (\varrho e)_0 = \varrho_0 e(\varrho_0, \vartheta_0) \in L^1(\Omega), \quad (186)$$

and

$$\operatorname{div}_x \mathbf{B}_0 = 0 \text{ in } \mathcal{D}'(\Omega), \quad \mathbf{B}_0 \cdot \mathbf{n}|_{\partial\Omega} = 0. \quad (187)$$

Then problem (57)-(65) possesses at least one variational solution $\varrho, \mathbf{u}, \vartheta, \mathbf{B}$ on an arbitrary time interval $(0, T)$.

This model, with boundary conditions

$$\mathbf{B} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \operatorname{curl}_x \mathbf{B} \times \mathbf{n}|_{\partial\Omega} = 0,$$

has been extended to the variable conductivity case

$$0 < \underline{\sigma} \leq \sigma(\varrho, \vartheta) \leq \bar{\sigma},$$

for constants $\underline{\sigma}, \bar{\sigma} > 0$ and for $\alpha \geq 2$, but under additional constraints on viscosities, by Fan and Yu [36].

3.2 Barotropic models

When the temperature fluctuations can be neglected, one obtains the barotropic-MHD system for the density ϱ , the velocity \mathbf{u} and the magnetic field \mathbf{H} in a bounded region Ω

$$\partial_t \varrho + \operatorname{div}_x(\varrho\mathbf{u}) = 0, \quad (188)$$

$$\partial_t(\varrho\mathbf{u}) + \operatorname{div}_x(\varrho\mathbf{u} \otimes \mathbf{u}) = \operatorname{div}_x \mathbb{T} + \mathbf{j} \times \mathbf{B}, \quad (189)$$

$$\partial_t \mathbf{H} + \operatorname{curl}_x(\mathbf{H} \times \mathbf{u}) + \operatorname{curl}_x(\lambda \operatorname{curl}_x \mathbf{H}) = 0, \quad \operatorname{div}_x \mathbf{H} = 0, \quad (190)$$

$$\mathbf{u}|_{\partial\Omega} = 0, \quad \mathbf{H}|_{\partial\Omega} = 0. \quad (191)$$

where the pressure is $p(\varrho) = A\varrho^\gamma$ and the transport coefficients ν, η, λ are positive constants.

Hu and Wang proved the following existence result [51]

Theorem 3.2 Let $\Omega \subset \mathbb{R}^3$ be a bounded domain with boundary of class $C^{2+\delta}$, $\delta > 0$ and $\gamma > 3/2$. Assume that the initial data ϱ_0 , $(\varrho \mathbf{u})_0$ and \mathbf{B}_0 be given so that

$$\varrho_0 \in L^\gamma(\Omega), (\varrho \mathbf{u})_0 \in L^1(\Omega; \mathbb{R}^3), \mathbf{H}_0 \in L^2(\Omega; \mathbb{R}^3), \quad (192)$$

$$\varrho_0 \geq 0, \quad (193)$$

and

$$\operatorname{div}_x \mathbf{H}_0 = 0 \text{ in } \mathcal{D}'(\Omega), \mathbf{H}_0|_{\partial\Omega} = 0. \quad (194)$$

Assume finally that $f, g \in L^\infty(0, T \times \Omega)$.

Then problem (188)-(191) possesses at least one variational solution ϱ , \mathbf{u} , \mathbf{H} on an arbitrary time interval $(0, T)$.

Let us mention that the same result holds (R. Sart [86]) with different boundary conditions on \mathbf{H} .

Let us quote for completeness some other models recently introduced.

Y. Amirat and K. Hamdache [1] studied a barotropic ferrofluid model with a magnetization dependent pressure $p = p(\varrho, \mathbf{M})$ including in the fluid description an extra equation for the angular momentum.

Finally X. Hu and D. Wang [53] considered density dependent viscosities, extending to MHD the result of D. Bresch and B. Desjardins [10] relying on analysis of the so-called BD-entropy. In this last case, presence of the radiative Stefan-Boltzmann contribution ϑ^4 is no-more requested but additional constraints on the pressure and transport coefficients (including magnetic permeability) are needed.

4 Singular limits

As usual in hydrodynamical problems, when putting the system in non dimensional form, various non dimensional numbers [101] (Mach, Reynolds, Peclet etc...) appear, which may be small (or large) in certain regimes of motion. Studying the corresponding limits may simplify drastically the description. Typically low Mach number limits of compressible flows are expected to be close to their incompressible limit. In MHD flows, asymptotic regimes appear naturally when the so called Mach number, Alfvén number, Froude number or Péclet number are small.

The simplest case has been studied by P. Kukučka [65], for which one expects convergence toward an incompressible limit.

More precisely for a suitable scaling (see [65]) one considers the following non dimensional form of the MHD system (9)-(14)-(17)-(24) where magnetic permeability μ is supposed to be a positive constant (and we note $\lambda(\varrho, \vartheta, \mathbf{H}) := \frac{1}{\mu\sigma(\varrho, \vartheta, \mathbf{H})}$ the magnetic diffusivity), to simplify the exposition.

$$\partial_t \mathbf{B} + \operatorname{curl}_x(\mathbf{B} \times \mathbf{u}) + \operatorname{curl}_x(\lambda \operatorname{curl}_x \mathbf{B}) = 0, \quad (195)$$

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (196)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) \frac{1}{Ma^2} = \operatorname{div}_x \mathbb{S} + \frac{1}{Al^2} \mathbf{j} \times \mathbf{B} + \frac{1}{Fr^2} \varrho \nabla_x \Phi, \quad (197)$$

$$\frac{d}{dt} \left(\frac{Ma^2}{2} \varrho |\mathbf{u}|^2 + \varrho e + \frac{Ma^2}{Al^2} \frac{1}{2\mu} |\mathbf{B}|^2 \right) dx = 0, \quad (198)$$

$$\partial_t(\varrho s) + \operatorname{div}_x(\varrho s \mathbf{u}) + \frac{1}{Pe} \operatorname{div}_x \left(\frac{\mathbf{q}}{\vartheta} \right) = \sigma, \quad (199)$$

with

$$\sigma \geq \frac{1}{\vartheta} \left(Ma^2 \mathbb{S} : \nabla_x \mathbf{u} + \frac{Ma^2}{Al^2} |\operatorname{curl}_x \mathbf{H}|^2 - \frac{1}{Pe} \frac{\mathbf{q} \cdot \nabla_x \vartheta}{\vartheta} \right), \quad (200)$$

where Φ is an external potential (gravitation for example).

In order to simplify the exposition, following [65], we suppose that Mach number Ma and Alfvén number Al are both small $Ma = Al = \varepsilon$ with $0 < \varepsilon \ll 1$ while $Pe = 1$. We also discard the external potential force ($\Phi = 0$).

We assume boundary conditions

$$\mathbf{B} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \mathbf{E} \times \mathbf{n}|_{\partial\Omega} = \mathbf{0}, \quad \mathbf{u} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \mathbb{S}\mathbf{n} \times \mathbf{n}|_{\partial\Omega} = \mathbf{0}, \quad \nabla_x \vartheta \cdot \mathbf{n}|_{\partial\Omega} = \mathbf{0}, \quad (201)$$

and initial conditions

$$(\mathbf{B}, \varrho, \mathbf{u}, \vartheta)(0, x) = (\mathbf{B}_0, \varrho_0, \mathbf{u}_0, \vartheta_0)(x) \quad \text{for any } x \in \Omega, \quad (202)$$

“ill-prepared” in the sense that

$$\mathbf{B}_0(x) = \mathbf{B}_{0,\varepsilon}^{(1)}(x), \quad \varrho_0(x) = \bar{\varrho} + \varepsilon_{0,\varepsilon}^{(1)}(x), \quad \mathbf{u}_0(x) = \mathbf{u}_{0,\varepsilon}^{(1)}(x), \quad \vartheta_0(x) = \bar{\vartheta} + \vartheta_{0,\varepsilon}^{(1)}(x), \quad (203)$$

where $\bar{\varrho}$ and $\bar{\vartheta}$ are two positive constants.

Formally expanding (195)-(199) with respect to ε we actually find at lowest order the incompressible MHD system with temperature

$$\partial_t \bar{\mathbf{B}} + \mathbf{curl}_x(\bar{\mathbf{B}} \times \mathbf{U}) + \mathbf{curl}_x(\lambda(\bar{\varrho}, \bar{\vartheta}, \bar{\mathbf{B}}) \mathbf{curl}_x \bar{\mathbf{B}}) = 0, \quad (204)$$

$$\operatorname{div}_x(\mathbf{U}) = 0, \quad (205)$$

$$\bar{\varrho}(\partial_t \mathbf{U} + \mathbf{U} \cdot \nabla_x \mathbf{U}) + \nabla_x \Pi = \operatorname{div}_x(\mathbb{S}(\mathbf{U}, \bar{\vartheta})) + \mathbf{curl}_x \bar{\mathbf{B}} \times \bar{\mathbf{B}}, \quad (206)$$

$$\bar{\varrho} c_P(\bar{\varrho}, \bar{\vartheta})(\partial_t \Theta + \mathbf{U} \cdot \nabla_x \Theta) = \operatorname{div}_x \left(\left(\kappa_F(\bar{\varrho}, \bar{\vartheta}, \bar{\mathbf{B}}) + \kappa_r \bar{\vartheta}^2 \right) \nabla_x \Theta \right), \quad (207)$$

where $c_P(\varrho, \vartheta) = e_\vartheta + \frac{\vartheta}{\varrho^2} \frac{p_\vartheta^2}{p_\varrho}$, with boundary conditions

$$\mathbf{B} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \mathbf{U} \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad \nabla_x \Theta \cdot \mathbf{n}|_{\partial\Omega} = 0, \quad (208)$$

and initial conditions

$$\mathbf{B}(0, x) = \mathbf{B}_0(x), \quad \mathbf{U}(0, x) = \mathbf{u}_0(x), \quad \Theta(0, x) = \vartheta_0(x) \quad \text{for any } x \in \Omega. \quad (209)$$

then one has the convergence result [65]

Theorem 4.1 *Let $\Omega \subset \mathbb{R}^3$ be a bounded domain with boundary of class $C^{2+\delta}$, $\delta > 0$ and assume that properties of the state functions and transport coefficients given in Theorem 2.1 are valid.*

Let $\{\varrho_\varepsilon, \mathbf{u}_\varepsilon, \vartheta_\varepsilon, \mathbf{B}_\varepsilon\}_{\varepsilon > 0}$ be a family of weak solutions to the scaled MHD system (195)-(199) on $(0, T) \times \Omega$ supplemented with boundary conditions (201) and initial conditions (203) such that

$$\begin{aligned} \varrho_{0,\varepsilon} &\rightarrow \varrho_0^{(1)} \quad \text{weakly* in } L^\infty(\Omega), \\ \mathbf{u}_{0,\varepsilon} &\rightarrow \mathbf{u}_0^{(1)} \quad \text{weakly* in } L^\infty(\Omega; \mathbb{R}^3), \\ \vartheta_{0,\varepsilon} &\rightarrow \vartheta_0^{(1)} \quad \text{weakly* in } L^\infty(\Omega), \\ \mathbf{B}_{0,\varepsilon} &\rightarrow \mathbf{B}_0^{(1)} \quad \text{weakly* in } L^\infty(\Omega; \mathbb{R}^3), \end{aligned}$$

as $\varepsilon \rightarrow 0$. Then

$$\begin{aligned} \operatorname{ess\,sup}_{t \in (0, T)} \|\varrho_\varepsilon(t) - \bar{\varrho}\|_{L^{\frac{5}{3}}(\Omega)} &\leq C\varepsilon, \\ \mathbf{u}_\varepsilon &\rightarrow \mathbf{U} \quad \text{weakly in } L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)), \\ \frac{\mathbf{B}_\varepsilon}{\varepsilon} &\rightarrow \mathbf{B} \quad \text{weakly in } L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)), \end{aligned}$$

$$\frac{\vartheta_\varepsilon - \bar{\vartheta}}{\varepsilon} \rightarrow \Theta \text{ weakly in } L^2(0, T; W^{1,2}(\Omega; \mathbb{R}^3)),$$

where $(\mathbf{U}, \Theta, \mathbf{B})$ is a weak solution of the incompressible MHD system with temperature (204)-(207) with boundary conditions (209), initial temperature

$$\Theta_0 = \frac{\bar{\vartheta}}{c_P(\bar{\varrho}, \bar{\vartheta})} \left(s_\varrho(\bar{\varrho}, \bar{\vartheta}) \varrho_0^{(1)} + s_\vartheta(\bar{\varrho}, \bar{\vartheta}) \vartheta_0^{(1)} \right),$$

and initial velocity given by

$$\bar{\varrho} \mathbf{U}(0) = \lim_{\varepsilon \rightarrow 0} H[\varrho_\varepsilon \mathbf{u}_\varepsilon](0),$$

where H is Helmholtz's projector on the space of solenoidal fields.

The proof follows the strategy of [42]:

1. One first proves uniform estimates for the sequence $\{\varrho_\varepsilon, \mathbf{u}_\varepsilon, \vartheta_\varepsilon, \mathbf{B}_\varepsilon\}_{\varepsilon > 0}$ and the entropy rate σ_ε .
2. One pass to the limit $\varepsilon \rightarrow 0$ in the system, the most delicate term being the convective term $\varrho_\varepsilon \mathbf{u}_\varepsilon \otimes \mathbf{u}_\varepsilon$, for which one uses the Helmholtz decomposition together with the associated acoustic equation.

More complex situations have been investigated: A. Novotný, M. Růžička and G. Thäter [81] have considered the case $Ma = \varepsilon$, $Al = \varepsilon$, $Fr = \varepsilon$, $Pe = \varepsilon^2$, Y.S. Kwon and K. Trivisa [68] have studied the case $Ma = \varepsilon$, $Al = \sqrt{\varepsilon}$, $Fr = \sqrt{\varepsilon}$, $Pe = 1$, E. Feireisl, A. Novotný and Y. Sun [43] have considered $Ma = \varepsilon$, $Al = Fr = Pe = 1$, $Re = 1/\varepsilon^n$, for a $n > 0$ (large Reynolds number regime) and S. Jiang, Q. Ju, F. Li and Z. Xin [56] have studied more complex ε^k regimes. In all of these situations the target system is the incompressible MHD system.

5 Strong solutions

Previous Sections dealt with weak solutions, belonging to Sobolev spaces, for system (9)-(14)-(17)-(21), together with (7) and (8) for the definition of the electric field and the current density. It is not known in general if these solutions are unique.

Another approach is to look directly for strong solutions, which are assumed to be regular. In order to do so, the general method is to prove local existence of smooth solutions. This is done by applying general local existence results on hyperbolic and hyperbolic-parabolic systems (Subsection 5.1 below). Then, it is possible, assuming that the initial data are sufficiently close to an equilibrium state, to prove that the parabolic nature of the equation dominates. This implies a priori bounds which allow to prove that the solution remains in the neighbourhood of the equilibrium as long as it exists. This in turn allows to prove global existence. This method, due to Kawashima and his co-authors, is described in Subsection 5.2. Subsection 5.3 is devoted to existence and uniqueness results in critical (Besov) spaces.

5.1 Local existence results

In [95], it is proved that a symmetric hyperbolic-parabolic system has a unique local solution. Therefore, its application to the present system only amounts to proving that the system is symmetrizable. The use of entropy variables is necessary to give a symmetric system for the fluid part (equations (14)-(17)-(21)). See for instance [21] for the details. The part with the magnetic field is dealt with using the method of [87], so the results of [95] apply. The result may be summarized as follows:

Theorem 5.1 *Let $\Omega = \mathbb{R}^3$, and assume that the thermodynamic functions*

$$p = p(\varrho, \vartheta), \quad e = e(\varrho, \vartheta), \quad s = s(\varrho, \vartheta)$$

are interrelated through (25), where p, e are continuously differentiable functions of positive arguments ϱ, ϑ satisfying (34)-(35)-(36)-(37).

Furthermore, we suppose that each of the transport coefficients

$$\nu = \nu(\vartheta, |\mathbf{H}|), \quad \eta = \eta(\vartheta, |\mathbf{H}|), \quad \kappa_F = \kappa_F(\varrho, \vartheta, |\mathbf{H}|), \quad \sigma = \sigma(\varrho, \vartheta, |\mathbf{H}|) \quad \text{and} \quad \mu = \mu(|\mathbf{H}|),$$

are positive smooth functions of their arguments. Consider an equilibrium state $\bar{\varrho}, \bar{\mathbf{u}} \equiv 0, \bar{\vartheta}, \bar{\mathbf{B}} \equiv 0$ of the system (9)-(14)-(17)-(21), and let the initial data $\varrho_0, \mathbf{u}_0, \vartheta_0, \mathbf{B}_0$ be given so that

$$(\varrho_0 - \bar{\varrho}, \mathbf{u}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{B}_0) \in H^3(\mathbb{R}^3). \quad (210)$$

Then there exists $T > 0$ such that the system (9)-(14)-(17)-(21), with the initial condition $\varrho_0, (\varrho \mathbf{u})_0, \vartheta_0, \mathbf{B}_0$, has a unique strong solution on $[0, T]$. This solution satisfies $\varrho, \vartheta > 0$ in $\mathbb{R}^3 \times [0, T]$, and T depends only on $\|(\varrho_0 - \bar{\varrho}, \mathbf{u}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{B}_0)\|_{H^3}, \inf_{\mathbb{R}^3} \vartheta_0$ and $\inf_{\mathbb{R}^3} \varrho_0$.

Some remarks are in order.

First, the above result is stated in [61], for electro-magneto-fluids (which, in addition to the above systems, includes equations for the charge ϱ_e and the electric field \mathbf{E}) with two-dimensional symmetries. However, the aim of [61] is to prove global existence, and this is why their study is restricted to two-dimensional flows. The local existence result is still valid for the 3D case. As a matter of fact, it is an application of the results of [95], which are valid in any dimension.

Second, the same result can be stated with H^s instead of H^3 , for any $s \geq 3$ (see [95, Section 4]).

Third, Theorem 5.1 is proved in [84, Theorem 2.1] in the special case of constant coefficients $\nu, \eta, \kappa_F, \sigma$ and μ . The proof is not given, and is claimed to be an easy adaptation of that of [77]. Looking at the proof of [77], it is restricted to the case $\mathbf{B} = 0$, that is, the compressible Navier-Stokes system with thermal conductivity. Their proof readily applies to the present case, and allows in fact for variable coefficients satisfying the above assumptions. A complete proof is also given in [37], in the case of a bounded domain (the method is easily adapted to $\Omega = \mathbb{R}^3$), with possible occurrence of vacuum, that is, $\varrho \geq 0$ instead of $\varrho > 0$.

Note also that, in Theorem 5.1, we have assumed (34), that is, the fluid is a perfect gas. Now, the result is still valid for a general equation of state satisfying (35)-(36)-(37) only (see [61]).

Finally, the assumptions $\nu, \eta, \kappa_F, \sigma, \mu > 0$ may be weakened in the following way: it is in fact possible to treat cases in which some of these coefficients are identically 0. For instance, using the same method as in [61], it is possible to deal with the case $\nu, \eta, \sigma, \mu > 0, \kappa_F \equiv 0$, the case $\nu, \kappa_F, \sigma, \mu > 0, \eta \equiv 0$, and the case $\nu, \sigma, \mu > 0, \kappa_F, \eta \equiv 0$.

5.2 Global existence results for small data

Theorem 5.1 being proved, one can then turn to the proof of global existence of a smooth solution for the system under consideration. For this purpose, the key ingredient is an a priori estimate which we now detail. For this purpose, we define, for any $0 \leq t_1 \leq t_2$, the quantity

$$N_s(t_1, t_2) = \sup_{t_1 \leq t \leq t_2} \|(\varrho - \bar{\varrho}, \mathbf{u}, \vartheta - \bar{\vartheta}, \mathbf{B})(t)\|_{H^s(\mathbb{R}^3)} + \left[\int_{t_1}^{t_2} \left(\|(\nabla \varrho, \nabla \mathbf{B})\|_{H^{s-1}(\mathbb{R}^3)}^2(t) + \|(\nabla \mathbf{u}, \nabla \vartheta)\|_{H^s(\mathbb{R}^3)}^2(t) \right) dt \right]^{1/2}. \quad (211)$$

Then, differentiating the system if necessary, and multiplying it by some well chosen test functions, it is possible to prove the following kind of estimate: there exists $\delta > 0$ such that, if

$$N_s(0, t)^2 \leq C (N_s(0, 0)^2 + N(0, t)^3),$$

for some constant C which depends on δ , but not on $t > 0$. This in turn allows to prove that there exists $\delta > 0$ and $C(\delta) > 0$, both independent of t , such that

$$N_3(0, 0) \leq \delta \quad \Rightarrow \quad N_3(0, t) \leq C(\delta)N_3(0, 0). \quad (212)$$

Using (212), one proves a bound on $N_3(0, t)$ for any time t of existence of the local in time solution. Applying the local existence result successively any each small time interval, one finally proves global existence:

Theorem 5.2 *Let $\Omega = \mathbb{R}^3$, and assume that the thermodynamic functions*

$$p = p(\varrho, \vartheta), \quad e = e(\varrho, \vartheta), \quad s = s(\varrho, \vartheta)$$

are interrelated through (25), where p, e are continuously differentiable functions of positive arguments ϱ, ϑ satisfying (35)-(37).

Furthermore, we suppose that each of the transport coefficients $\nu, \eta, \kappa_F, \sigma, \mu$ are positive constants.

Consider an equilibrium state $\bar{\varrho}, \bar{\mathbf{u}} \equiv 0, \bar{\vartheta}, \bar{\mathbf{B}} \equiv 0$ of the system (9)-(14)-(17)-(21), and let the initial data $\varrho_0, \mathbf{u}_0, \vartheta_0, \mathbf{B}_0$ be given so that

$$(\varrho_0 - \bar{\varrho}, \mathbf{u}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{B}_0) \in H^3(\mathbb{R}^3). \quad (213)$$

Then there exists $\delta > 0$ such that, if $N_3(0, 0) < \delta$ (N_3 is defined by (211), then problem (9)-(14)-(17)-(21), with the initial condition $\varrho_0, \mathbf{u}_0, \vartheta_0, \mathbf{B}_0$, has a unique global strong solution $(\varrho, \mathbf{u}, \vartheta, \mathbf{B})$ such that

$$\begin{aligned} (\varrho - \bar{\varrho}, \mathbf{B}) \text{ in } C^0(\mathbb{R}^+, H^3) \cap C^1(\mathbb{R}^+, H^2), \quad (\mathbf{u}, \vartheta - \bar{\vartheta}) \in C^0(\mathbb{R}^+, H^3) \cap C^1(\mathbb{R}^+, H^1), \\ \nabla(\mathbf{u}, \vartheta - \bar{\vartheta}) \in L^2(\mathbb{R}^+, H^3). \end{aligned} \quad (214)$$

This solution satisfies the estimate

$$\forall t \in \mathbb{R}^+, \quad N_3(0, t) \leq CN_3(0, 0),$$

where C depends only on $\delta > 0$, and N_3 is defined by (211).

This result is proved in [84, Theorem 2.5]. Considering the strategy of proof, it seems that the fact that the coefficients are constant is not important, although we did not check all the details.

In [61], the same result is proved for a more general system and with non-constant coefficients, but only with two-dimensional symmetries. Note that, due to the structure of their system, which includes an equation for the electric field and the density of charge, the equilibrium state does not necessarily satisfies $\bar{\mathbf{B}} \equiv 0$.

In essence, the above result is proved by studying the linearized equation near the equilibrium state $(\bar{\varrho}, \bar{\mathbf{u}} \equiv 0, \bar{\vartheta}, \bar{\mathbf{B}} \equiv 0)$. This is a linear hyperbolic-parabolic system, for which energy-type estimates allow to prove the above result. Hence, a natural consequence is that the solution converges, as time tends to infinity, to the equilibrium state $(\bar{\varrho}, \bar{\mathbf{u}} \equiv 0, \bar{\vartheta}, \bar{\mathbf{B}} \equiv 0)$. Such a convergence has been studied in [84], with the above assumptions. The convergence following convergence is proved in this paper:

$$\forall k = 1, 2, 3, \quad \|\nabla^k (\varrho - \bar{\varrho}, \mathbf{u}, \vartheta - \bar{\vartheta}, \mathbf{B})\|_{L^2(\mathbb{R}^3)} \leq \frac{C}{(1+t)^{\frac{3}{2}(\frac{1}{q}-\frac{1}{2})+\frac{k}{2}}} \quad (215)$$

$$\forall p > q, \quad \|(\varrho - \bar{\varrho}, \mathbf{u}, \vartheta - \bar{\vartheta}, \mathbf{B})\|_{L^p(\mathbb{R}^3)} \leq \frac{C}{(1+t)^{\frac{3}{2}(\frac{1}{q}-\frac{1}{p})}} \quad (216)$$

Here, $q \in [1, \frac{6}{5})$ is such that $(\varrho_0 - \bar{\varrho}, \mathbf{u}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{B}_0) \in L^q(\mathbb{R}^3)$.

As in the case of variational solutions, barotropic fluids have been extensively studied. In such a case, the system becomes (188)-(189)-(190)-(191), with a pressure law $p(\varrho) = \varrho^\gamma$, $\gamma > 1$. All the coefficients

are supposed to be constant. In [73, 97], local existence is proved in the case $\lambda = 0$, that is, when the parabolic term in (190) vanishes. The main idea of the method is to fix the velocity field u , and then solve the problem with respect to all other variables. Then the Schauder fixed-point theorem is applied. Note that in [97], $\varrho \mapsto p(\varrho)$ is only assumed to be strictly convex and increasing. Global existence is proved in [14] under the assumption of small initial data, using a method similar to the one explained above. [91] proposes a generalization to the case of a system with Coulomb interaction, and [99] to the case of a bounded domain. The result in [75] is similar, but allows for vacuum in the system of equations. Finally, [99] gives a similar kind of result, but with an H^2 norm to estimate the initial data. Note finally that the asymptotic behavior as $t \rightarrow +\infty$ is also proved in [14, 91, 99] with the same kind of results as above.

5.3 Critical regularity

After the works of R. Danchin [22] [23] [24] on compressible Navier-Stokes system in a critical functional framework, some recent works have been devoted to the extension of his results to the compressible MHD system.

For the reader's convenience, let us first recall a few basic definitions concerning Besov spaces (a much more detailed presentation can be found in [5]).

Let us first introduce the homogeneous Littlewood-Paley decomposition. Let $\chi : \mathbb{R}^n \rightarrow [0, 1]$ be a smooth nonincreasing radial function supported in $B(0, 1)$ and such that $\chi \equiv 1$ on $B(0, 1/2)$, and let

$$\varphi(\xi) := \chi(\xi/2) - \chi(\xi).$$

Then we define the homogeneous Littlewood-Paley decomposition $(\dot{\Delta}_k)_{k \in \mathbb{Z}}$ over \mathbb{R}^n by

$$\dot{\Delta}_k u := \varphi(2^{-k} D)u = \mathcal{F}^{-1}(\varphi(2^{-k} \cdot) \mathcal{F}u),$$

where \mathcal{F} stands for the Fourier transform on \mathbb{R}^n .

For any $s \in \mathbb{R}$ and $(p, r) \in [1, +\infty]^2$, the homogeneous Besov space $\dot{B}_{p,r}^s(\mathbb{R}^n)$ stands for the set of tempered distributions f such that

$$\|f\|_{\dot{B}_{p,r}^s(\mathbb{R}^n)} := \left\| 2^{sk} \|\dot{\Delta}_k f\|_{L^p(\mathbb{R}^n)} \right\|_{\ell^r(\mathbb{Z})} < \infty,$$

and we also note $B^s := B_{2,2}^s$.

One has the formal decomposition

$$u = \sum_{k \in \mathbb{Z}} \dot{\Delta}_k u.$$

Owing to the fact that the solutions that we consider here do not have the same regularity in low and high frequencies, it is useful to use hybrid Besov spaces by splitting any distribution u into its low frequency part and its high frequency part [5] as follows:

$$u^\ell = \sum_{k \leq 0} \dot{\Delta}_k u \quad \text{and} \quad u^h = \sum_{k > 0} \dot{\Delta}_k u. \quad (217)$$

Then we define for $s, t \in \mathbb{R}$

$$\|f\|_{B^{s,t}} := \sum_{k \leq 0} 2^{sk} \|\dot{\Delta}_k f\|_{L^2} + \sum_{k > 0} 2^{tk} \|\dot{\Delta}_k f\|_{L^2},$$

and the hybrid Besov space $B^{s,t}$ is defined for $m = -[d/2 + 1 - s]$ by

$$B^{s,t} := \{f \in \mathcal{S}'(\mathbb{R}^d) : \|f\|_{B^{s,t}} < \infty\},$$

if $m < 0$, or

$$B^{s,t} := \{f \in \mathcal{S}'(\mathbb{R}^d)/\mathcal{P}_m : \|f\|_{B^{s,t}} < \infty\},$$

if $m \geq 0$, where \mathcal{P}_m is the set of polynomials of degree less or equal to m (of course in any case $B^s := B^{s,s}$).

Let us first rewrite the barotropic system

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (218)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) - \nabla_x p(\varrho) = \mu \Delta \mathbf{u} + (\mu + \lambda) \nabla_x \operatorname{div}_x(\mathbf{u}) + \mathbf{H} \cdot \nabla_x \mathbf{H} - \frac{1}{2} \nabla_x(|\mathbf{H}|^2), \quad (219)$$

$$\partial_t \mathbf{H} + \operatorname{div}_x(\mathbf{u} \mathbf{H} + \mathbf{u} \cdot \nabla_x \mathbf{H}) = \nu \Delta \mathbf{H}, \quad \operatorname{div}_x \mathbf{H} = 0, \quad (220)$$

$$(\varrho, \mathbf{u}, \mathbf{H})_{t=0} = (\varrho_0, \mathbf{u}_0, \mathbf{H}_0), \quad \operatorname{div}_x \mathbf{H}_0 = 0. \quad (221)$$

where the pressure is $p(\varrho) = A\varrho^\gamma$ and the transport coefficients ν, η, λ are positive constants.

Hao proved the following existence result [48]

Theorem 5.3 *Let $d \geq 3$, $\bar{\rho} > 0$, $2\mu + d\lambda > 0$, $\nu > 0$ and \mathbf{h} an arbitrary constant vector.*

Assume that the initial data satisfy $\varrho_0 - \bar{\rho} \in B^{d/2-1, d/2}$, $\mathbf{u}_0 \in B^{d/2-1}$ and $\mathbf{H}_0 \in B^{d/2-1}$.

Then there exists a small number $\varepsilon > 0$ and a constant M such that

$$\|\varrho_0 - \bar{\rho}\|_{B^{d/2-1, d/2}} + \|\mathbf{u}_0\|_{B^{d/2-1}} + \|\mathbf{H}_0\|_{B^{d/2-1}} \leq \varepsilon.$$

then system (218)-(221) has a unique global solution on an arbitrary time interval $(0, T)$ such that $(\varrho - \bar{\rho}, \mathbf{u}, \mathbf{H} - \mathbf{h}) \in E$, with

$$E := C\left(\mathbb{R}_+; B^{d/2-1, d/2} \times (B^{d/2})^{d+d}\right) \cap L^1\left(\mathbb{R}_+; B^{d/2+1, d/2+2} \times (B^{d/2+1})^{d+d}\right),$$

and satisfies

$$\|(\varrho - \bar{\rho}, \mathbf{u}, \mathbf{H} - \mathbf{h})\|_E \leq M(\|\varrho_0 - \bar{\rho}\|_{B^{d/2-1, d/2}} + \|\mathbf{u}_0\|_{B^{d/2-1}} + \|\mathbf{H}_0\|_{B^{d/2-1}}).$$

In the case of a polytropic fluid with constant viscosities the system rewrites

$$\partial_t \varrho + \operatorname{div}_x(\varrho \mathbf{u}) = 0, \quad (222)$$

$$\partial_t(\varrho \mathbf{u}) + \operatorname{div}_x(\varrho \mathbf{u} \otimes \mathbf{u}) - \nabla_x p(\varrho) = \mu \Delta \mathbf{u} + (\mu + \lambda) \nabla_x \operatorname{div}_x(\mathbf{u}) + \mathbf{H} \cdot \nabla_x \mathbf{H} - \frac{1}{2} \nabla_x(|\mathbf{H}|^2) + \varrho \mathbf{f}, \quad (223)$$

$$\partial_t(\varrho \vartheta) + \operatorname{div}_x(\varrho \mathbf{u} \vartheta) - \kappa \Delta \vartheta + p \operatorname{div}_x(\mathbf{u}) = \frac{\mu}{2} |\nabla_x \mathbf{u} + \nabla_x^t \mathbf{u}| = \mu (\operatorname{div}_x \mathbf{u})^2 + \sigma (\operatorname{curl}_x \mathbf{H})^2 + \mathbf{H} \cdot \nabla_x \mathbf{H} - \frac{1}{2} \nabla_x(|\mathbf{H}|^2) + \varrho \mathbf{f}, \quad (224)$$

$$\partial_t \mathbf{H} + \operatorname{div}_x(\mathbf{u} \mathbf{H} + \mathbf{u} \cdot \nabla_x \mathbf{H}) = \nu \Delta \mathbf{H}, \quad \operatorname{div}_x \mathbf{H} = 0, \quad (225)$$

$$(\varrho, \mathbf{u}, \mathbf{H})_{t=0} = (\varrho_0, \mathbf{u}_0, \mathbf{H}_0), \quad \operatorname{div}_x \mathbf{H}_0 = 0. \quad (226)$$

Bian and Guo proved the following local existence result [7] (see also [6] [8])

Theorem 5.4 *Let $d \geq 2$, $\bar{\rho} > 0$, $2\mu + d\lambda > 0$, $\nu > 0$.*

Assume that the initial data satisfy $\varrho_0 - \bar{\rho} \in B_{p,1}^{d/p}$, $\mathbf{u}_0 \in (B_{p,1}^{d/p})^d$, $\vartheta_0 \in B_{p,1}^{d/p}$ and $\mathbf{H}_0 \in (B_{p,1}^{d/p})^d$, with $\varrho_0 - \bar{\rho}$ bounded away from 0.

Then there exists a positive number T such that the system (222)-(225) has a unique solution $(\varrho, \mathbf{u}, \vartheta, \mathbf{H})$ on the time interval $(0, T)$ such that

$$\varrho - \bar{\rho} \in C([0, T]; B_{p,1}^{d/p}),$$

$$\begin{aligned}\mathbf{u} &\in \left(C([0, T]; B_{p,1}^{d/p-1}) \cap L^1([0, T]; B_{p,1}^{d/p+1}) \right)^d, \\ \vartheta &\in C([0, T]; B_{p,1}^{d/p-2}) \cap L^1([0, T]; B_{p,1}^{d/p}), \\ \mathbf{H} &\in \left(C([0, T]; B_{p,1}^{d/p-1}) \cap L^1([0, T]; B_{p,1}^{d/p+1}) \right)^d.\end{aligned}$$

Moreover if $p \in [2, d]$ the solution is unique.

The first step of the proofs of both of these critical results is to use a smooth approximation of the data. For this regularized problem, previous results of R. Danchin [23] allow to prove local existence using a linearization argument. Then, uniform bounds are proved on the solution, allowing to pass to the limit as the regularization parameter vanishes.

6 One-dimensional models

In this section, we give an account of results in the one-dimensional case, which allows for drastic simplification. In such a case, the use of Lagrangian variables gives a much simpler system which is amenable by standard techniques.

In dimension one, it is commonly assumed that the unknowns depend only on the first coordinate of x_1 of x . The velocity of the fluid is still a three-dimensional vector, but its first component plays a special role. We denote it by u , while the transverse velocity is \mathbf{w} :

$$\mathbf{u} = \begin{pmatrix} u \\ w_1 \\ w_2 \end{pmatrix} = \begin{pmatrix} u \\ \mathbf{w} \end{pmatrix}.$$

Hence, denoting $\mathbf{B} = (b_1, b_2, b_3)^T$, the constraint $\operatorname{div}(\mathbf{B}) = 0$ gives $\partial_1 b_1 = 1$, so that b_1 is independent of x_1 . Further, if f is a vector field which is independent of x_1 , then the first component of $\operatorname{curl}_x(f)$ vanishes. Hence, (9) implies that $\partial_t b_1 = 1$. As a consequence, we may assume without loss of generality that $b_1 = 1$. Hence, denoting by $\mathbf{b} = (b_2, b_3)^T$ the transverse magnetic field, that is,

$$\mathbf{B} = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = \begin{pmatrix} 1 \\ \mathbf{b} \end{pmatrix},$$

we find the following system

$$\begin{cases} \partial_t \varrho + \partial_{x_1}(\varrho u) = 0, \\ \partial_t(\varrho u) + \partial_{x_1}(\varrho u^2 + p) = \partial_{x_1} \left[\left(\frac{4}{3} \nu + \eta \right) \partial_{x_1} u \right] - \mathbf{b} \cdot \partial_{x_1} \left(\frac{\mathbf{b}}{\mu} \right), \\ \partial_t(\varrho \mathbf{w}) + \partial_{x_1} \left(\varrho u \mathbf{w} - \frac{\mathbf{b}}{\mu} \right) = \partial_{x_1}(\nu \partial_{x_1} \mathbf{w}), \\ \partial_t \mathcal{E} + \partial_{x_1} \left[\left(\mathcal{E} + \frac{|\mathbf{b}|^2}{2\mu} + p \right) u - \mathbf{w} \cdot \mathbf{b} \right] = \partial_{x_1} \left[\left(\frac{4}{3} \nu + \eta \right) u \partial_{x_1} u \right] + \partial_{x_1}(\kappa_F \partial_{x_1} \vartheta) \\ \quad + \partial_{x_1} \left(\frac{\mathbf{b}}{\sigma} \cdot \partial_{x_1} \left(\frac{\mathbf{b}}{\mu} \right) \right) + \partial_{x_1}(\nu \mathbf{w} \cdot \partial_{x_1} \mathbf{w}), \\ \partial_t \mathbf{b} + \partial_{x_1}(u \mathbf{b} - \mathbf{w}) = \partial_{x_1} \left(\frac{1}{\sigma} \partial_{x_1} \frac{\mathbf{b}}{\mu} \right), \end{cases} \quad (227)$$

where the total energy \mathcal{E} is given by

$$\mathcal{E} = \frac{1}{2}\varrho(u^2 + |\mathbf{w}|^2) + \varrho e + \mathcal{M}\left(\frac{|\mathbf{b}|}{\mu}\right).$$

Recall that \mathcal{M} is defined by (10), and that, if μ is a constant, we have

$$\mathcal{M}\left(\frac{|\mathbf{b}|}{\mu}\right) = \frac{|\mathbf{b}|^2}{2\mu}.$$

In what follows, we will always assume that μ is a constant.

In the special case of a barotropic fluid, that is, without the energy equation (fourth equation of (227)) and with a pressure law given by $p(\varrho) = A\varrho^\gamma$, $\gamma \geq 1$, it is proved in [102] that the above system has a unique global smooth solution in the spatial domain $\Omega = (0, 1)$, for initial data $(\varrho_0, u_0, \mathbf{w}_0, \mathbf{b}_0)$ such that $\varrho_0 \geq 0$, $\varrho_0 \in H^1$, $(u_0, \mathbf{w}_0) \in H_0^1 \cap H^2$, $\mathbf{b}_0 \in H_0^1$. Further, a compatibility condition is needed, in the spirit of those used in the case of Navier-Stokes equations [17, 18, 74]. All the coefficients ν, η, σ, μ are assumed to be constant. The method of proof is here again to prove local existence using Banach fixed-point theorem. Then a priori estimates are proved, which allow to apply local existence again and extend the time interval to infinity. These a priori estimates give, in addition, a stability of the solution in Sobolev spaces.

Next, in the paper [103], Zhang and Xie study the full system (227), with additional gravitation terms:

$$\left\{ \begin{array}{l} \partial_t \varrho + \partial_{x_1}(\varrho u) = 0, \\ \partial_t(\varrho u) + \partial_{x_1}(\varrho u^2 + p) = \partial_{x_1} \left[\left(\frac{4}{3}\nu + \eta \right) \partial_{x_1} u \right] - \mathbf{b} \cdot \partial_{x_1} \left(\frac{\mathbf{b}}{\mu} \right) + \varrho \partial_{x_1} \psi, \\ \partial_t(\varrho \mathbf{w}) + \partial_{x_1} \left(\varrho u \mathbf{w} - \frac{\mathbf{b}}{\mu} \right) = \partial_{x_1}(\nu \partial_{x_1} \mathbf{w}), \\ \partial_t \mathcal{E} + \partial_{x_1} \left[\left(\mathcal{E} + \frac{|\mathbf{b}|^2}{2\mu} + p \right) u - \mathbf{w} \cdot \mathbf{b} \right] = \partial_{x_1} \left[\left(\frac{4}{3}\nu + \eta \right) u \partial_{x_1} u \right] + \partial_{x_1}(\kappa_F \partial_{x_1} \vartheta) \\ \quad + \partial_{x_1} \left(\frac{\mathbf{b}}{\sigma} \cdot \partial_{x_1} \left(\frac{\mathbf{b}}{\mu} \right) \right) + \partial_{x_1}(\nu \mathbf{w} \cdot \partial_{x_1} \mathbf{w}) + \varrho u \partial_{x_1} \psi, \\ \partial_t \mathbf{b} + \partial_{x_1}(u \mathbf{b} - \mathbf{w}) = \partial_{x_1} \left(\frac{1}{\sigma} \partial_{x_1} \frac{\mathbf{b}}{\mu} \right), \end{array} \right. \quad (228)$$

In this system, ψ is the gravitation field, defined by Poisson's equation:

$$\left\{ \begin{array}{l} -\partial_{x_1}^2 \psi = G\rho \text{ in } \Omega, \\ \psi = 0 \text{ on } \partial\Omega. \end{array} \right. \quad (229)$$

where $G > 0$ is the gravitational constant. The conductivity $\kappa_F = \kappa_F(\varrho, \vartheta)$ is not assumed to be a constant, but satisfies the growth assumptions

$$\frac{1}{C}(1 + \vartheta^q) \leq \kappa_F, \quad \left| \frac{\partial \kappa_F}{\partial \varrho} \right| \leq 1 + \vartheta^q, \quad \forall q > \frac{5}{2}, \quad (230)$$

for some constant $C > 0$. Further, the equation of state includes the radiative term ϑ^4 , that is

$$p = R\rho\vartheta + \frac{a}{3}\vartheta^4, \quad e = C_V\vartheta + \frac{a}{\varrho}\vartheta^4. \quad (231)$$

The method is similar to [102]. This result has been improved by Qin and Hu [85], in which $q > \frac{2+\sqrt{211}}{9} \approx 0.3$ is allowed.

Another approach in this one-dimensional setting is to use Lagrangian coordinates. Here, we assume that we use a bounded domain, which, without loss of generality, we may assume to be $(0, 1)$. We also assume that $\int_0^1 \rho_0(x_1) dx_1 = 1$. Setting

$$y = y(t, x_1) = \int_0^{x_1} \varrho(t, \xi) d\xi, \quad (232)$$

we then have $0 \leq y \leq 1$, and, due to mass conservation

$$\int_0^1 \varrho(t, x_1) dx_1 = \int_0^1 \varrho_0(t, x_1) dx_1 = 1.$$

Thus, defining $v = 1/\rho$ the specific volume of the fluid, we immediately have, using Dirichlet Boundary conditions

$$\partial_t y + \varrho u = 0.$$

Hence, system (227) becomes

$$\begin{cases} \partial_t v - \partial_y u = 0, \\ \partial_t u + \partial_y p = \partial_y \left(\frac{\frac{4}{3}\nu + \eta}{v} \partial_y u \right) - \mathbf{b} \cdot \partial_y \left(\frac{\mathbf{b}}{\mu} \right), \\ \partial_t \mathbf{w} - \partial_y \left(\frac{\mathbf{b}}{\mu} \right) = \partial_y \left(\frac{\nu}{v} \partial_y \mathbf{w} \right), \\ \partial_t (v\mathcal{E}) + \partial_y \left(pu + \frac{|\mathbf{b}|^2}{2\mu} u - \mathbf{w} \cdot \mathbf{b} \right) = \partial_y \left[\left(\frac{4}{3}\nu + \eta \right) u \frac{\partial_y u}{v} + \kappa_F \frac{\partial_y \vartheta}{v} + \frac{\mathbf{b}}{\sigma v} \partial_y \left(\frac{\mathbf{b}}{\mu} \right) + \frac{\nu}{v} \mathbf{w} \cdot \partial_y \mathbf{w} \right], \\ \partial_t (v\mathbf{b}) - \partial_y \mathbf{w} = \partial_y \left(\frac{1}{\sigma v} \partial_y \left(\frac{\mathbf{b}}{\mu} \right) \right). \end{cases} \quad (233)$$

In principle, this system is set on the interval $\Omega = (0, 1)$ with homogeneous Dirichlet boundary conditions:

$$(u, \mathbf{w}, \mathbf{b}, \partial_y \vartheta) = 0 \text{ on } \partial\Omega. \quad (234)$$

However, the corresponding Cauchy problem ($\Omega = \mathbb{R}$) is also studied by some authors, with the same kind of boundary conditions. System (233)-(234) is completed by initial conditions $(v_0, u_0, \mathbf{w}_0, \vartheta_0, \mathbf{b}_0)$.

The first existence result concerning the Cauchy problem (233) is the global existence of Kawashima and Okada [62]. The equation of state is assumed to satisfy natural conditions similar to (34). It is proved in [62] that, given a constant state $(\bar{v}, \bar{u} = 0, \bar{\mathbf{w}} = 0, \bar{\vartheta}, \bar{\mathbf{b}})$, (with $\bar{v} > 0$ and $\bar{\vartheta} > 0$) if the initial condition is close to this state in $H^2(\mathbb{R})$, then system (233)-(234) has a unique global smooth solution $(v, u, \mathbf{w}, \vartheta, \mathbf{b})$ such that, for any $t \geq 0$, $(v(t) - \bar{v}, u(t), \mathbf{w}(t) - \bar{\mathbf{w}}, \vartheta(t) - \bar{\vartheta}, \mathbf{b}(t) - \bar{\mathbf{b}}) \in H^2(\mathbb{R})$. Here again, local existence is proved using classical results, and global energy estimates are proved for these smooth solutions. This allows to extend the existence result to any positive time. The result may be summarized as follows:

Theorem 6.1 (Kawashima-Okada, 1982 [62]) *Consider system (233) set in $\Omega = \mathbb{R}$. Assume that the equation of state satisfies*

$$\frac{\partial p}{\partial \varrho} > 0, \quad \frac{\partial e}{\partial \vartheta} > 0,$$

and that $de = \vartheta dS - pd\left(\frac{1}{\varrho}\right)$, where S is the entropy of the fluid. Assume that μ is a positive constant, and that η, ν, κ_F are positive smooth functions of ϱ and ϑ . Let

$$(\bar{v}, \bar{u}, \bar{\mathbf{w}}, \bar{\vartheta}, \bar{\mathbf{b}}) = (\bar{v}, 0, 0, \bar{\vartheta}, 0)$$

be a constant equilibrium state of (233). Let $(v_0, u_0, \mathbf{w}_0, \vartheta_0, \mathbf{b}_0)$ be an initial data such that

$$(v_0 - \bar{v}, u_0, \mathbf{w}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{b}_0) \in H^2(\mathbb{R}).$$

If $\|v_0 - \bar{v}, u_0, \mathbf{w}_0, \vartheta_0 - \bar{\vartheta}, \mathbf{b}_0\|_{H^2(\mathbb{R})}$ is sufficiently small, then (233) has a unique global smooth solution.

Actually, coefficients are allowed to vanish, provided they are constant: the situation allowed are (a) $\nu = \eta = 0, \mu, \kappa_F > 0$, (b) $\nu, \eta, \mu > 0, \kappa_F = 0$, (c) $\eta = \nu = \kappa_F = 0, \mu > 0$.

Another result can be found in [15]. In this contribution, system (227) is considered on the domain $\Omega = \Omega(t) = (0, x(t))$, where the free boundary $x(t)$ is defined by $x'(t) = u(t, x(t))$. The boundary conditions are given by

$$(\mathbf{w}, \mathbf{b}, \partial_{x_1} \vartheta)_{\partial\Omega(t)} = 0, \quad u(t, 0) = 0, \quad \left[p - \left(\frac{4}{3} \nu + \eta \right) \partial_{x_1} u \right]_{x_1=x(t)} = 1. \quad (235)$$

Then, the Lagrangian coordinates are used to transform the system into (233), in the domain $\Omega = (0, 1)$, with the boundary conditions

$$(\mathbf{w}, \mathbf{b}, \partial_y \vartheta)_{\partial\Omega} = 0, \quad u(t, 0) = 0, \quad \left[p - \frac{1}{v} \left(\frac{4}{3} \nu + \eta \right) \partial_y u \right]_{y=1} = 1. \quad (236)$$

We are back to a system which is similar to (233)-(234), to which the strategy of Kawashima-Okada [62] is applied, thereby proving global existence if the initial data is close to a given constant state, in Sobolev spaces. Moreover, in the special case of a perfect gas equation of state, a global existence of weak solutions is proved. For this purpose, the initial data, which may be discontinuous, is approximated by smooth functions. This allows to apply the global smooth existence result. The estimates established in the smooth case are then used again to prove that one can pass to the limit as the regularization parameter tends to 0. This is based on Aubin-Lions Lemma (see [4, 71]). All these results are proved using the assumption that $\sigma = 1, \eta, \nu$ are bounded smooth functions of v , that the energy and pressure of the fluid satisfy some polynomial growth property as functions of ϑ , and the fact that

$$\kappa_F(v, \vartheta) + |\partial_v \kappa_F(v, \vartheta)| + |\partial_v^2 \kappa_F(v, \vartheta)| \leq C(1 + \vartheta^q),$$

for some constant $C > 0$ and some power $q \geq 2$. This is in sharp contrast with [62], in which κ_F is assumed to be constant.

Using a similar method, Wang [96] proved that system (233)-(234) has a global weak solution in $H^1(\Omega)$. The same kind of assumptions are used on the equation of state and the diffusion coefficient κ_F .

In a similar setting, it is proved in [16] that the solutions depend continuously on the initial data.

In the case of constant coefficients, with an equation of state satisfying

$$\left| \frac{\partial p}{\partial e}(v, e) \right| + e(\vartheta) + \left| \frac{\partial e}{\partial \vartheta}(v, \vartheta) \right| \leq C,$$

the Cauchy problem (233) set on $\Omega = \mathbb{R}$ is studied by Hoff and Tsyganov in [49, 50]. It is proved that weak solutions exist and are unique, provided the total energy is small. Under an additional integrability assumption on $\partial_t \mathbf{b}$, continuity with respect to the initial data is proved (in $L^2([0, T] \times \mathbb{R})$, for any $T > 0$).

In [35], for system (233)-(234), the limit of vanishing shear viscosity $\nu \rightarrow 0$ is studied, under the same assumptions as in [15]. It is proved that the solution converges to that of the limit problem strongly in $L^2([0, T] \times \Omega)$ for $\rho, \mathbf{w}, \vartheta$, and strongly in $L^2([0, T], H^1(\Omega))$ for u, \mathbf{b} .

Finally, still in Lagrangian coordinates, but in the special case $\mathbf{w} = 0$, system (233)-(234) has been studied by Iskenderova [55]. Assuming constant coefficients and a perfect gas equation of state, Iskenderova proves that the system has a unique weak solution. Note that the initial density ρ_0 is allowed to vanish, provided that $\partial_y \left(\rho_0^{1/4} \right) \in L^2$, and that

$$\rho_0(y) \leq C \inf_{0 \leq z \leq y} \rho_0(z),$$

for some constant $C > 0$. Here again, the method of proof is a regularization of the system: existence and uniqueness is easily proved for a system in which $\rho_0 \geq \varepsilon > 0$. Then, the limit $\varepsilon \rightarrow 0$ is studied using a priori estimates.

References

- [1] Y. Amirat, K. Hamdache. Weak solutions to the equations of motion for compressible magnetic fluids. *J. Math. Pures Appl.*, **91**:433–467, 2009.
- [2] A.A. Amosov, A.A. Zlotnik. A family of difference schemes for the equations of one-dimensional magnetogasdynamics: properties and global error estimates. *Soviet Math. Dokl.*, **37**:545–550, 1988.
- [3] A.A. Amosov, A.A. Zlotnik. A family of difference schemes for one-dimensional magnetic gas dynamics equations (viscous barotropic case). *Sov. J. Anal. Math. Modelling*, **3**:431–451, 1988.
- [4] Aubin, J.-P. Un théorème de compacité. *C. R. Acad. Sci. Paris* **256**:5042–5044, 1963.
- [5] H. Bahouri, J.-Y. Chemin, R. Danchin. *Fourier Analysis and Nonlinear Partial Differential Equations*. Grundlehren der mathematischen Wissenschaften, **343**, Springer (2011).
- [6] D. Bian, B. Guo. Global existence and large time behavior of solutions to the equations of electric-magnetohydrodynamic equations. *Kinetic and Related Models*, **6**:481–503, 2013.
- [7] D. Bian, B. Guo. Well-posedness in critical spaces for the full compressible MHD equations. *Acta Math. Scient.*, **33**:1153–1176, 2013.
- [8] D. Bian, B. Yuan. Well-posedness in super critical Besov spaces for the compressible MHD equations. *Int. J. Dyn. Syst. Differ. Equ.* **3** (3), 383–399, 2011.
- [9] M. E. Bogovskii. Solution of some vector analysis problems connected with operators div and grad (in Russian). *Trudy Sem. S.L. Sobolev*, **80**(1):5–40, 1980.
- [10] D. Bresch, B. Desjardins. On the existence of global weak solutions to the Navier-Stokes equations for viscous compressible and heat conducting fluids. *J. Math. Pures Appl.*, **87**:57–90, 2007.
- [11] C. Buet, B. Després. Asymptotic analysis of fluid models for the coupling of radiation and hydrodynamics. *J. Quant. Spect. and Rad. Trans.*, **85**:385–418, 2003.
- [12] H. Cabannes. *Theoretical magnetohydrodynamics*. Academic Press, New York, 1970.
- [13] M. Cannone, C. Miao, N. Prioux, B. Yuan. The Cauchy problem for the magnetohydrodynamic system. *Banach Center Publications*, **74**:59–93, 2006.
- [14] Q. Chen, Z. Tan. Global existence and convergence rates of smooth solutions for the compressible magnetohydrodynamic equations. *Nonlinear Anal.*, **72**:4438–4451, 2010.
- [15] G-Q. Chen, D. Wang. Global solutions of nonlinear magnetohydrodynamics with large initial data. *J. of Diff. Eq.*, **182**:344–376, 2002.
- [16] G-Q. Chen, D. Wang. Existence and continuous dependence of large solutions for the magnetohydrodynamic equations. *Z. angew. Math. Phys.*, **54**:608–632, 2003.
- [17] H. J. Choe, H. Kim, Global existence of the radially symmetric solutions of the Navier-Stokes equations for the isentropic compressible fluids. *Math. Methods Appl. Sci.* **28**(1):1–28, 2005.
- [18] H. J. Choe, H. Kim, Strong solutions of the Navier-Stokes equations for isentropic compressible fluids. *J. Differential Equations* **190**(2):504–523, 2003.

- [19] R. Coifman, Y. Meyer. On commutators of singular integrals and bilinear singular integrals. *Trans. Amer. Math. Soc.*, **212**:315–331, 1975.
- [20] J.P. Cox, R.T. Giuli. *Principles of stellar structure, I,II*. Gordon and Breach, New-York, 1968.
- [21] C. M. Dafermos. *Hyperbolic conservation laws in continuum physics*. Grundlehren der Mathematischen Wissenschaften **325**, Springer-Verlag, Berlin, 2000.
- [22] R. Danchin. Global existence in critical spaces for compressible Navier-Stokes equations. *Inventiones Math.*, **141**, 579–614, 2000.
- [23] R. Danchin. Local theory in critical spaces for compressible viscous and heat-conductive gases. *Comm. P.D.E*, **26**, 1183–1233, 2001.
- [24] R. Danchin. Global existence in critical spaces for flows of compressible viscous and heat-conductive gases. *Arch. Rational Mech. Anal.*, **160**, 1–39, 2001.
- [25] A. Decoster, P.A. Markowich, B. Perthame. *Models of collisions*. Elsevier, 1998.
- [26] P. Degond, M. Lemou. On the viscosity and thermal conduction of fluids with multivalued internal energy. *Euro. J. Mech. B- Fluids*, **20**:303–327, 2001.
- [27] J.L. Delcroix, A. Bers, *Physique des plasmas*. Savoirs actuels, InterÉditions/CNRS Editions, 1994.
- [28] R.J. DiPerna, P.-L. Lions. Ordinary differential equations, transport theory and Sobolev spaces. *Invent. Math.*, **98**:511–547, 1989.
- [29] B. Ducomet, E. Feireisl. On the dynamics of gaseous stars. *Arch. Rational Mech. Anal.*, **174**:221–266, 2004.
- [30] B. Ducomet, E. Feireisl. The equations of magnetohydrodynamics: on the interaction between matter and radiation in the evolution of gaseous stars. *Commun. Math. Phys.*, **266**:595–629, 2006.
- [31] G. Duvaut, J.-L. Lions. Inequations en thermoélasticité et magnétohydrodynamique. *Arch. Rational Mech. Anal.*, **46**:241–279, 1972.
- [32] G. Duvaut, J.-L. Lions. *Inequalities in mechanics and physics*. Springer-Verlag, Heidelberg, 1976.
- [33] S. Eliezer, A. Ghatak, H. Hora. *An introduction to equations of states, theory and applications*. Cambridge University Press, Cambridge, 1986.
- [34] A.C. Eringen, G.A. Maugin. *Electrodynamics of continua*. Springer Verlag, New-York, Berlin, Heidelberg, 1990.
- [35] J. Fan, S. Jiang, G. Nakamura. Vanishing shear viscosity limit in the magnetohydrodynamic equations. *Commun. Math. Phys.*, **270**:691–708, 2007.
- [36] J. Fan, W. Yu. Global variational solutions to the compressible magnetohydrodynamic equations. *Nonlinear Analysis*, **66**:3637–3660, 2008.
- [37] J. Fan, W. Yu. Strong solution to the compressible magnetohydrodynamic equations with vacuum. *Nonlinear Anal. Real World Appl.*, **10**:392–409, 2008.
- [38] E. Feireisl. On compactness of solutions to the compressible isentropic Navier-Stokes equations when the density is not square integrable. *Comment. Math. Univ. Carolinae*, **42**(1):83–98, 2001.
- [39] E. Feireisl. *Dynamics of viscous compressible fluids*. Oxford University Press, Oxford, 2003.

- [40] E. Feireisl. On the motion of a viscous, compressible, and heat conducting fluid. *Indiana Univ. Math. J.*, **53**:1707–1740, 2004.
- [41] E. Feireisl. Stability of flows of real monoatomic gases. *Commun. Partial Differential Equations*, **31**:325–348, 2006.
- [42] E. Feireisl, A. Novotný. *Singular limits in thermodynamics of viscous fluids*. Birkhauser, Basel, 2009.
- [43] E. Feireisl, A. Novotný, Y. Sun. Dissipative solutions and the the incompressible inviscid limits of the compressible magnetohydrodynamic system in unbounded domains. *Discrete and Continuous Dynamical Systems*, **34**:121–143, 2014.
- [44] E. Feireisl, H. Petzeltová. On integrability up to the boundary of the weak solutions of the Navier-Stokes equations of compressible flow. *Commun. Partial Differential Equations*, **25**(3-4):755–767, 2000.
- [45] J.F. Gerbeau, C. Le Bris, T. Lelièvre. *Mathematical methods for the magnetohydrodynamics of liquid metal*. Oxford University Press, Oxford, 2006.
- [46] P. Germain, N. Masmoudi. Global existence for the Euler-Maxwell system. ArXiv:1107.1595v1, 2011.
- [47] V. Giovangigli. *Multicomponent flow modeling*. Birkhäuser, Basel, 1999.
- [48] C. Hao. Well-posedness to the compressible viscous magnetohydrodynamic system. *Nonlinear Analysis: Real World Applications*, **12**:2962–2972, 2011.
- [49] D. Hoff, E. Tsyganov. Uniqueness and continuous dependence of weak solutions in compressible magnetohydrodynamics. *Z. angew. Math. Phys.*, **56**:791–804, 2005.
- [50] D. Hoff, E. Tsyganov. Systems of partial differential equations of mixed hyperbolic-parabolic type. *J. Differential Equations*, **204** (1), 163–201, 2004.
- [51] X. Hu, D. Wang. Global solutions to the three-dimensional full compressible magnetohydrodynamic flows. *Commun. Math. Phys.*, **283**:255–284, 2008.
- [52] X. Hu, D. Wang. Compactness of weak solutions to the three-dimensional compressible magnetohydrodynamic equations. *J. Differential Equations*, **245**:2176–2198, 2008.
- [53] X. Hu, D. Wang. Low Mach number limit of viscous compressible magnetohydrodynamic flows. *SIAM. J. Math. Anal.*, **41**:1272–1294, 2009.
- [54] X. Hu, D. Wang. Global existence and large time behavior of solutions to the three-dimensional equations of compressible magnetohydrodynamic flows. *Arch. Rational Mech. Anal.*, **197**:203–238, 2010.
- [55] D.A. Iskenderova. An initial-boundary value problem for Magnetogasdynamic equations with degenerate viscosity. *Differential Equations*, **36**:847–856, 2000.
- [56] S. Jiang, Q. Ju, F. Li, Z. Xin. Low Mach number limit for the full compressible magnetohydrodynamics equations with general initial data. *Advances in Mathematics*, **259**:384–420, 2015.
- [57] S. Jiang, F. Li. Rigorous derivation of the compressible magnetohydrodynamics equations from the electromagnetic fluid system. *Nonlinearity*, **25**:1735–1752, 2012.
- [58] S. Jiang, F. Li. Convergence of the complete electromagnetic fluid system to the full compressible magnetohydrodynamics equations. arXiv:1309.3668v2, 2015.

- [59] T. Kato. The Cauchy problem for quasi-linear symmetric hyperbolic systems. *Arch. Rational Mech. Anal.* **58**(3): 181–205, 1975.
- [60] S. Kawashima. Systems of a hyperbolic-parabolic composite type, with applications to the equations in magnetohydrodynamics. *Doctoral Thesis*, Kyoto University, 1983.
- [61] S. Kawashima. Smooth global solutions for two-dimensional equations of electro-magneto-fluid dynamics. *Japan J. Appl. Math.*, **1**:207–222, 1984.
- [62] S. Kawashima, M. Okada. Smooth global solutions for the one-dimensional equations in magnetohydrodynamics. *Proc Japan Acad.*, **58**:384–387, 1982.
- [63] S. Kawashima, Y. Shizuta. Magnetohydrodynamic approximation of the complete equations for an electromagnetic fluid. *Tsukuba J. Math.*, **10**:131–149, 1986.
- [64] N.A. Krall, A.W. Trivelpiece. *Principle of plasma physics*. McGraw Hill Book Company, New York, London, Sydney 1973.
- [65] P. Kukučka. Singular limits of the equations of magnetohydrodynamics. *J. Math. Fluid Mech.*, **13**:173–189, 2011.
- [66] A.G. Kulikovskiy, G.A. Lyubimov. *Magnetohydrodynamics*. Addison-Wesley, Reading, Massachusetts, 1965.
- [67] R.M. Kulsrud. *Plasma physics for astrophysics*. Princeton University Press, Princeton, 2005.
- [68] Y-S. Kwon, K. Trivisa. On the incompressible limits for the full magnetohydrodynamics flows. *J. Differential Equations*, **251**:1990–2023, 2011.
- [69] O.A. Ladyzhenskaya, V.A. Solonnikov. Solution of some non-stationary problems of magnetohydrodynamics for a viscous incompressible fluid. *Trudy Mat. Inst. Steklov*, **59**:115–173, 1960.
- [70] L.D. Landau, E.M. Lifschitz. *Electrodynamics of continuous media*. Pergamon Press, Oxford, 1960.
- [71] J.-L. Lions. *Quelque methodes de résolution des problèmes aux limites non linéaires*. Paris: Dunod, Gauthier-Villars, 1969.
- [72] P.-L. Lions. *Mathematical topics in fluid dynamics, Vol.2, Compressible models*. Oxford Science Publication, Oxford, 1998.
- [73] X. Li, N. Su, D. Wang. Local strong solution to the compressible magnetohydrodynamic flow with large data. *J. of Hyperbolic Differ. Equ.*, **8**:415–436, 2011.
- [74] T. Luo, Z. Xin, T. Yang. Interface behavior of compressible Navier-Stokes equations with vacuum. *SIAM J. Math. Anal.* **31**(6):1175–1191, 2000.
- [75] H.-L. Li, X. Xu, J. Zhang. Global classical solutions to 3D compressible magnetohydrodynamics equations with large oscillations and vacuum. *SIAM J. Math. Anal.*, **45**(3):1356–1387, 2013.
- [76] S. Liu, H. Yu, J. Zhang. Global weak solutions of 3D compressible MHD with discontinuous initial data and vacuum. *J. Differential Equations*, **254**:229–255, 2013.
- [77] A. Matsumura, T. Nishida. The initial value problem for the equations of motion of viscous and heat-conductive gases. *J. Math. Kyoto Univ.* **20**(1):67–104, 1980.
- [78] B. Mihalas, B. Weibel-Mihalas. *Foundations of radiation hydrodynamics*. Dover Publications, Dover, 1984.

- [79] I. Müller, T. Ruggeri. *Rational extended thermodynamics*. Springer Tracts in Natural Philosophy 37, Springer-Verlag, Heidelberg, 1998.
- [80] F. Murat, Compacité par compensation, *Ann. Scuola Norm. Sup. Pisa Cl. Sci.* **5** (4):485–507, 1978.
- [81] A. Novotný, M. Růžička, G. Thäter. Singular limit of the equations of magnetohydrodynamics in the presence of strong stratification. *Math. Models and Methods in the Applied Sciences*, **21**:115–147, 2011.
- [82] A. Novotný, I. Straškraba. *Introduction to the theory of compressible flow*. Oxford University Press, Oxford, 2004.
- [83] P. Pedregal. *Parametrized measures and variational principles*. Birkhäuser, Basel, 1997.
- [84] X. Pu, B. Guo. Global existence and convergence rates of smooth solutions for the full compressible MHD equations. *Z. angew. Math. Phys.*, **64**, 519–538, 2013.
- [85] Y. Qin, G. Hu. Global smooth solutions for 1D thermally radiative magnetohydrodynamics. *J. of Mathematical Physics*, **52**, 023102, 2011.
- [86] R. Sart. Existence of finite energy weak solutions for the equations MHD of compressible fluids. *Applicable Analysis*, **88**:357–379, 2009.
- [87] P. Secchi. Well-posedness for a mixed problem for the equations of ideal magneto-hydrodynamics. *Arch. Math. (Basel)* **64**(3):237–245, 1995.
- [88] M. Sermange, R. Temam. Some mathematical questions related to the MHD equations. *Comm. Pure Appl. Math.*, **36**:635–664, 1983.
- [89] S. N. Shore. *An introduction to atrophysical hydrodynamics*. Academic Press, New York, 1992.
- [90] Ch.S. Smagulov, A.A. Durmagambetov, D.A. Iskenderova. Cauchy problem for the equations of magnetogasdynamics (In Russian). *Differential Equations*, **29**:337–348, 1993.
- [91] Z. Tan, L. Tong, Y. Wang. Large time behavior of the compressible magnetohydrodynamic equations with Coulomb force. *J. Math. Anal. Appl.*, **427**:600–617, 2015.
- [92] L. Tartar. Compensated compactness and applications to partial differential equations. Nonlinear Analysis and Mechanics, Herriot-Watt Symposium, *L.J. Knops editor, Pitman*, 1984.
- [93] R. Temam. *Navier-Stokes equations*. North-Holland, Amsterdam, 1977.
- [94] T. Umeda, S. Kawashima, Y. Shizuta. On the decay of solutions to the linearized equations of electro-magneto-fluid dynamics. *Japan J. Appl. Math.*, **1**:435–457, 1984.
- [95] A.I. Vol’pert, S.I. Hudjaev. The Cauchy problem for composite systems of nonlinear differential equations. (Russian) *Mat. Sbornik* **87**:504–528, 1972.
- [96] D. Wang. Large solutions to the initial-boundary value problem for planar magnetohydrodynamics. *SIAM. J. Appl. Math.*, **63**:1424–1441, 2003.
- [97] Y.-Z. Wang, Y.-X. Wang. Local existence of strong solutions to the three dimensional compressible MHD equation with partial viscosity. *Comm. Pure Appl. Anal.*, **12**:851–866, 2013.
- [98] L.C. Woods. *Principles of magnetoplasma dynamics*. Clarendon Press, Oxford, 1987.
- [99] Y.-F. Yang, X. Gu, G. Dou. Global well-posedness of strong solutions to the magnetohydrodynamic equations of compressible flows. *Nonlinear Anal.*, **95**:23–37, 2014.

- [100] J.P. Zahn, J. Zinn-Justin. *Astrophysical fluid dynamics, les houches, XLVII*. Elsevier, Amsterdam, 1993.
- [101] R.K. Zeytounian. *Mécanique des fluides fondamentale*. Springer Verlag, Berlin, Heidelberg, New York, 1991.
- [102] J. Zhang. Global existence and uniqueness of strong for the magnetohydrodynamic equations. *Boundary Value Problems*, **2008**:735846 (27 November 2007).
- [103] J. Zhang, F. Xie. Global solutions for a one-dimensional model problem in thermally radiative magnetohydrodynamics. *J. Differential Equations*, **245**:1853–1882, 2008.
- [104] H. Zirin. *Astrophysics of the sun*. Cambridge University Press, Cambridge, 1988.