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High order exponential integrators for nonlinear Schrödinger equations with application to rotating Bose-Einstein condensates

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Abstract

This article deals with the numerical integration in time of nonlinear Schrödinger equations. The main application is the numerical simulation of rotating Bose-Einstein condensates. The authors perform a change of unknown so that the rotation term disappears and they obtain as a result a nonautonomous nonlinear Schrödinger equation. They consider exponential integrators such as exponential Runge–Kutta methods and Lawson methods. They provide an analysis of the order of convergence and some preservation properties of these methods in a simplified setting and they supplement their results with numerical experiments with realistic physical parameters. Moreover, they compare these methods with the classical split-step methods applied to the same problem.

AMS Classification.

Keywords. Gross-Pitaevskii equation, numerical methods.

1 Introduction

This paper deals with the numerical integration of nonautonomous nonlinear Schrödinger (NLS) equations which read

$$\begin{cases} i\partial_t\psi = -\frac{1}{2}\Delta\psi + V(t, \mathbf{x})\psi + \beta|\psi|^{2\kappa}\psi, & (t, \mathbf{x}) \in \mathbb{R}^+ \times \mathbb{R}^d, \\ \psi(0, \mathbf{x}) = \psi_0(\mathbf{x}), & x \in \mathbb{R}^d. \end{cases} \quad (1.1)$$

The operator Δ denotes the usual Laplace operator on \mathbb{R}^d , $d \in \mathbb{N}^*$. The potential function V is smooth and $\kappa \in \mathbb{N}$. The unknown ψ is a complex-valued wavefunction associated to the given initial datum ψ_0 . The derivation and the analysis of efficient semi-discrete numerical methods for the time integration of these equations have a long history. Some authors are interested in the finite time accuracy of the schemes [28, 23, 10, 41]. With additional hypotheses corresponding to physically relevant situations, equation (1.1) may have several invariants such as energy, momentum, etc, ..., in addition to mass conservation. Several authors show interest in the preservation of the invariants by the numerical schemes [23, 9, 46, 31, 43]. Beyond finite time integration of Eq. (1.1), asymptotic regimes have been considered: long time preservation of invariants [32, 27] and semi-classical regimes [5, 16, 11, 17, 20].

Our goal in this paper is to introduce and analyse two classes of exponential methods for the time integration of (1.1) on a d -dimensional space. The first class is that of exponential Runge–Kutta methods [35, 36, 26]. The second class relies on the Lawson techniques [40, 42]. We focus on situations where the dynamics of equation (1.1) essentially stays in a bounded domain of \mathbb{R}^d . Hence, we replace Eq. (1.1) with the same equation on a large periodic torus \mathbb{T}^d . Our main application is the Gross-Pitaevskii equation modeling rotating Bose-Einstein Condensate (BEC) in \mathbb{R}^2 for which the dynamics of the solutions is much spatially localized. We prove our results for functions in Sobolev spaces $H^\sigma(\mathbb{T}^d)$ for $\sigma > d/2$ and the proofs extend naturally to functions in $H^\sigma(\mathbb{R}^d)$. However, the fully discrete corresponding situations are not similar. Discretizing functions on the whole space requires to deal with carefully chosen specific boundary conditions. Working on a torus has the numerical advantage of avoiding boundary conditions. Moreover the numerical computation of the linear group generated by $i\Delta$ is much simpler in the case of the torus [34].

Before the introduction of the methods (Sections 2 and 3) and our numerical results (Section 4), we introduce below our main application and show how it fits equation (1.1).

1.1 Presentation of the application

A Bose-Einstein Condensate is the state of matter reached by a dilute gas of bosons cooled to very low temperature. A large fraction of bosons occupies the lowest quantum state so that macroscopic quantum phenomena become apparent. This phenomena was theoretically predicted by Bose in 1924 for photons [14] and generalized to atoms by Einstein in 1925 [29]. The first experimental evidence of BEC was obtained in 1995 [1, 22].

At low temperature, a rotating planar BEC can be described by the macroscopic complex-valued wave function $\varphi = \varphi(t, \mathbf{x})$ whose evolution is governed by a Gross-Pitaevskii

Equation (GPE) with an angular momentum rotation term. After a suitable changes of variables [6], the dimensionless GPE in $d = 2$ dimensions satisfied by φ can be written for $\mathbf{x} \in \mathbb{R}^2$:

$$i\partial_t\varphi = -\frac{1}{2}\Delta\varphi + V_c(\mathbf{x})\varphi + \beta|\varphi|^2\varphi - \Omega R\varphi. \quad (1.2)$$

The real-valued function $V_c = V_c(\mathbf{x})$ corresponds to a potential depending only on the space variables denoted by $\mathbf{x} = (x, y)^t$. In our physical context, this potential is confining: this means that $V_c(\mathbf{x})$ tends to $+\infty$ when $\|\mathbf{x}\| = \sqrt{x^2 + y^2}$ tends to $+\infty$. For example, in this paper, we consider potentials of the form

$$V_c(\mathbf{x}) = \frac{1}{2} \left(\gamma_x^2 x^2 + \gamma_y^2 y^2 \right), \quad (1.3)$$

where $\gamma_x, \gamma_y > 0$. This confining potential competes with the rotation operator $-\Omega R = i\Omega(x\partial_y - y\partial_x)$ at angular speed $\Omega \in \mathbb{R}$: the former tends to make bosons stay together at the origin of the plane, while the latter tends to spread the bosons out. The real coefficient β represents the nonlinearity strength, and comes from the averaged effect of the bosons. The evolution equation (1.2) is supplemented with an initial condition

$$\varphi(0, \mathbf{x}) = \varphi_0(\mathbf{x}), \quad \text{for all } \mathbf{x} \in \mathbb{R}^2. \quad (1.4)$$

If one introduces the functional

$$E(\varphi) = \int_{\mathbb{R}^2} \left[\frac{1}{2} |\nabla\varphi|^2 + V_c|\varphi|^2 + \frac{\beta}{2} |\varphi|^4 - \Omega \text{Re}(\varphi^* R\varphi) \right] d\mathbf{x}, \quad (1.5)$$

then (1.2) reads

$$i\partial_t\varphi(t, \cdot) = \nabla_{\varphi(t, \cdot)} E(\varphi(t, \cdot)).$$

The equation happens to be Hamiltonian, and the energy is preserved by the dynamics: along a solution $t \mapsto \varphi(t, \cdot)$ of (1.2)-(1.4), one has $E(\varphi(t, \cdot)) = E(\varphi_0)$. In addition to preserving the energy, the evolution preserves the mass of the wave function : along a solution $t \mapsto \varphi(t, \cdot)$ of (1.2)-(1.4), one has $\int_{\mathbb{R}^2} |\varphi(t, \mathbf{x})|^2 d\mathbf{x} = \int_{\mathbb{R}^2} |\varphi_0(\mathbf{x})|^2 d\mathbf{x}$. Another dynamical feature of this equation is the evolution of the angular momentum expectation: if one denotes by

$$\langle R \rangle (t) = \int_{\mathbb{R}^2} \varphi^*(t, \mathbf{x}) R\varphi(t, \mathbf{x}) d\mathbf{x}, \quad (1.6)$$

then this real-valued function is constant in the special case of a radial harmonic potential ($\gamma_x = \gamma_y$ in (1.3)) and has a more complex dynamics in more general cases (see Lemma 6.2.1 in [45]).

In the last decades this model has been studied a lot [7, 3, 2, 4, 45]. An important issue in the numerical time integration of equations (1.2)-(1.4) comes from the rotation term R . Following [6], we introduce new coordinates that allow to put equation (1.2) in the form of equation (1.1). Let us set for $t \in \mathbb{R}$,

$$A(t) = \begin{pmatrix} \cos(\Omega t) & -\sin(\Omega t) \\ \sin(\Omega t) & \cos(\Omega t) \end{pmatrix}. \quad (1.7)$$

Note that $A(t)$ is orthogonal and hence satisfies $A(t)^{-1} = A(t)^t$. We perform the change of unknown $\varphi \leftrightarrow \psi$ defined by

$$\varphi(t, \mathbf{x}) = \psi(t, A(t)\mathbf{x}). \quad (1.8)$$

This way, φ solves (1.2) if and only if ψ solves

$$\partial_t \psi = \frac{i}{2} \Delta \psi - iV(t, \tilde{\mathbf{x}})\psi - i\beta|\psi|^2\psi, \quad (1.9)$$

where V is a time dependent potential given by

$$V(t, \tilde{\mathbf{x}}) = V_c(A(t)\tilde{\mathbf{x}}),$$

and the initial datum is

$$\psi(0, \tilde{\mathbf{x}}) = \psi_0(\tilde{\mathbf{x}}) = \varphi_0(\tilde{\mathbf{x}}).$$

For convenience, we shall denote explicitly the following time-dependent change of spatial variables:

$$\tilde{\mathbf{x}} = \begin{pmatrix} \tilde{x} \\ \tilde{y} \end{pmatrix} = A(t) \begin{pmatrix} x \\ y \end{pmatrix} = A(t)\mathbf{x}.$$

Note that equation (1.9) satisfied by ψ is a standard nonlinear Schrödinger equation (1.1) with a cubic nonlinearity ($\kappa = 1$) and a space and time dependent potential.

1.2 General setting of the Cauchy problem

The unboundedness of the potential function V makes the numerical analysis of the exponential methods difficult. Therefore, we modify the equation (1.1) by cutting off the potential V smoothly. Let us motivate this modification. In the context of Schrödinger equations with confining potentials, if the mass of the initial datum is essentially concentrated in a bounded set around the origin, then this mass localisation property is to be preserved by the evolution of (1.1), at least for reasonable times. Therefore, modifying the potential V out of a sufficiently large bounded set around the origin will not create huge errors in the solution, at least for not too long times. Let us introduce a smooth function $\chi : \mathbb{R} \rightarrow [0, 1]$ such that

$$\forall x \in [1 - \delta/2, \delta/2 - 1], \quad \chi(x) = 1 \quad \text{and} \quad \forall x \in (-\infty, -\delta/2) \cup (\delta/2, +\infty), \quad \chi(x) = 0,$$

where $\delta \gg 2$ is a given real number, chosen accordingly with the initial datum φ_0 , the other physical parameters and the final computational time $T > 0$. We define the new potential function W for $t \in [0, T]$, $\mathbf{x} = (x_1, \dots, x_d)^t$ by letting

$$W(t, \mathbf{x}) = V(t, \mathbf{x}) \prod_{j=1}^d \chi(x_j). \quad (1.10)$$

Although modifying the potential function as above changes deeply the physical situation, the fact that δ is taken accordingly to φ_0 and the physical parameters and for a finite time interval $[0, T]$ gives some hope in the fact that the evolution of equation (1.1) with the potential function V and that of the same equation with V replaced with W starting from the same initial datum φ_0 will be quite similar.

To be more specific with the periodization of the problem, let us denote by \mathbb{T}_δ the quotient $\mathbb{R}/(\delta\mathbb{Z})$. We consider the function w defined as δ -periodic in all directions such that for $t \in [0, T]$ and all $\mathbf{x} \in \mathbb{R}^d$ satisfying $|x_j| \leq \delta/2$, $1 \leq j \leq d$, one has $w(t, \mathbf{x}) = W(t, \mathbf{x})$. The mapping $t \mapsto w(t, \cdot)$ is smooth from $[0, T]$ to $H^\sigma(\mathbb{T}_\delta^d)$ as soon as $\sigma \geq 0$. In the following

sections, we replace the continuous problem (1.1) with its periodic counterpart as explained above, and we assume that $\varphi_0 = \psi_0 \in H^\sigma(\mathbb{T}_\delta^d)$ for some $\sigma > d/2$.

We therefore get the following semilinear Cauchy problem in time:

$$\begin{aligned} \partial_t \psi(t, \mathbf{x}) - L\psi(t, \mathbf{x}) &= N_w(t, \psi)(\mathbf{x}), \quad (t, \mathbf{x}) \in [0, T] \times \mathbb{T}_\delta^d, \\ \psi(0, \mathbf{x}) &= \psi_0(\mathbf{x}), \quad \mathbf{x} \in \mathbb{T}_\delta^d, \end{aligned} \tag{1.11}$$

where $T > 0$, $L = \frac{i\Delta}{2}$ and

$$N_w(t, \psi)(\mathbf{x}) = -iw(t, \mathbf{x})\psi(t, \mathbf{x}) - i\beta|\psi|^{2\kappa}\psi(t, \mathbf{x}). \tag{1.12}$$

Remark 1. *The choice of the definition of the linear part L and the nonlinear part N from equation (1.1) is somewhat arbitrary. One could also choose, for example*

$$L = \frac{i\Delta}{2} - iw(t, \mathbf{x}) \quad \text{and} \quad N(t, \psi)(\mathbf{x}) = -i\beta|\psi|^{2\kappa}\psi(t, \mathbf{x}),$$

but this would lead to a nonautonomous linear problem whose spectral properties are not as nice as that of the other case. We will indeed use the nice spectral properties of the operator L later.

One can check that the operators of the form $e^{\alpha L}$ for $\alpha \in \mathbb{R}$ are isometric over $H^\sigma(\mathbb{T}_\delta^d)$. The function $t \mapsto w(t, \cdot)$ belongs to $\mathcal{C}^\infty(\mathbb{R}, H^\sigma(\mathbb{T}_\delta^d))$. Moreover, the nonlinear function N_w satisfies a local Lipschitz condition :

Lemma 2. *For all $T > 0$, for all $r > 0$, there exists a constant $C > 0$ such that for all $t \in [0, T]$, $\varphi_1, \varphi_2 \in H^\sigma(\mathbb{T}_\delta^d)$ such that $\|\varphi_1\|_{H^\sigma(\mathbb{T}_\delta^d)} \leq r$ and $\|\varphi_2\|_{H^\sigma(\mathbb{T}_\delta^d)} \leq r$, we have $\|N_w(t, \varphi_1) - N_w(t, \varphi_2)\|_{H^\sigma(\mathbb{T}_\delta^d)} \leq C\|\varphi_1 - \varphi_2\|_{H^\sigma(\mathbb{T}_\delta^d)}$.*

Remark 3. *Note that this Lipschitz property is also true if one replaces N_w with $N_W(t, \psi) = -iW(t, \cdot)\psi(\cdot) - i\beta|\psi|^{2\kappa}\psi(\cdot)$ and the norms are taken in $H^\sigma(\mathbb{R}^d)$ with $\sigma > d/2$, but is no longer true with $N_V = -iV(t, \cdot)\psi(\cdot) - i\beta|\psi|^{2\kappa}\psi(\cdot)$.*

We recover the fact that the Cauchy problem (1.11) is well-posed in $H^\sigma(\mathbb{T}_\delta^d)$.

1.3 Outline of the paper

The outline of the paper is as follows: Section 2 is devoted to exponential Runge–Kutta methods. Firstly, we briefly describe the construction of these exponential methods, and recall what is the so-called underlying Runge–Kutta method, which is defined using s collocation points. Secondly, we prove for the equation (1.11) that if the s collocation points are distinct, then the exponential Runge–Kutta method applied to Problem (1.11) has order s . Finally, we provide some numerical experiments in the case of Gauss collocation points using physically relevant parameters and initial data. Section 3 is devoted to exponential integrators named Lawson methods. These methods are collocation methods on a new evolution equation, obtained from Problem (1.11) *via* another change of unknown. We show that Lawson methods applied to the simplified equation keep their classical order. In particular, the methods with s stages defined with Gauss points have order $2s$. Moreover, they preserve quadratic invariants up to round off errors. Section 4 is devoted to the comparison of our methods with other methods such as splitting methods as used in [6]. For our numerical experiments, the algorithm in time is supplemented with the fast Fourier transform (FFT) method in space, in a periodic square. Finally we end the paper with conclusion and outlook.

2 Exponential Runge–Kutta methods

The main idea behind exponential integrators is to integrate exactly the linear part of the problem and then to use an appropriate approximation of the nonlinear part. Exponential Runge–Kutta (ERK) methods are particular exponential integrators. These methods have been derived and analysed for semi-linear parabolic Cauchy problems (see for example [36] for collocation methods, [35] for explicit methods and [37] for a survey on exponential integrators). They also have been used in [26] for solving linear and semi-linear Schrödinger Cauchy problems on the d -dimensional torus.

In this section, we introduce, analyse and use ERK methods to solve numerically equation (1.9).

2.1 Notations and description of the method

In order to solve numerically problem (1.11), we consider the following ERK methods of collocation type. We refer to [36] for a derivation of such methods for semi-linear problems based on variation-of-constants formula.

Let $T > 0$ be the final computational time and $(t_n)_{0 \leq n \leq M}$ be a uniform subdivision of $[0, T]$ with $M + 1$ points *i.e.* $t_n = nh$ with $h = T/M$. Let $s \in \mathbb{N}^*$ and $c_1, \dots, c_s \in [0, 1]$ be given such that for all $(i, j) \in \{1, \dots, s\}^2$, $c_i \neq c_j$ if $i \neq j$. For some $n \in \{0, \dots, M - 1\}$, we assume we have an approximation ψ_n of the exact solution $\psi(t_n)$ of problem (1.11) at time t_n . Using an ERK method consists in computing an approximation ψ_{n+1} from ψ_n in the following way. First, we solve the nonlinear system of s equations

$$\psi_{n,k} = e^{c_k h L} \psi_n + h \sum_{\ell=1}^s a_{k,\ell}(hL) N_w(t_n + c_\ell h, \psi_{n,\ell}), \quad 1 \leq k \leq s, \quad (2.1)$$

where the unknowns are $\psi_{n,1}, \dots, \psi_{n,s}$ and the s^2 coefficients $(a_{k,\ell}(hL))_{(k,\ell) \in \{1, \dots, s\}^2}$ are linear continuous operators on $H^\sigma(\mathbb{T}_\delta^d)$ defined by

$$a_{k,\ell}(hL) = \frac{1}{h} \int_0^{c_k h} e^{(c_k h - \sigma)L} \mathcal{L}_\ell(\sigma) d\sigma, \quad (2.2)$$

with $(\mathcal{L}_\ell)_{1 \leq \ell \leq s}$ the Lagrange polynomials defined by

$$\mathcal{L}_\ell(\tau) = \prod_{j=1, j \neq \ell}^s \frac{\tau/h - c_j}{c_\ell - c_j}, \quad 1 \leq \ell \leq s. \quad (2.3)$$

Then we compute ψ_{n+1} using

$$\psi_{n+1} = e^{hL} \psi_n + h \sum_{k=1}^s b_k(hL) N_w(t_n + c_k h, \psi_{n,k}), \quad (2.4)$$

where

$$\forall k \in \{1, \dots, s\}, \quad b_k(hL) = \frac{1}{h} \int_0^h e^{(h-\sigma)L} \mathcal{L}_k(\sigma) d\sigma. \quad (2.5)$$

Remark 4. *If we set $L = 0$ in formulae (2.2) and (2.5), then one recovers classical formulae defining Runge-Kutta collocation methods. The Runge-Kutta collocation method defined by the corresponding coefficients is called the underlying Runge-Kutta method.*

Note that in order to be able to compute ψ_{n+1} from ψ_n , we have to precompute the coefficients $a_{k,\ell}(hL)$ and $b_k(hL)$ for all $k, \ell \in \{1, \dots, s\}$. We present an accurate and efficient way of precomputing these coefficients in the next subsection. Of course, the spatial discretization of the operator L has to be specified.

2.2 Precomputation of the coefficients

It is well known that the operator $L = i\Delta/2$ on $H^\sigma(\mathbb{T}_\delta^d)$ is self-adjoint with eigenvalues

$$\omega_p = -\left(\frac{2\pi}{\delta}\right)^2 (p_1^2 + \dots + p_d^2), \quad p = (p_1, \dots, p_d) \in \mathbb{Z}^d.$$

We discretize the problem in space using a uniform grid with K^2 points and rely on FFT techniques. The eigenvalues of the discretized version of the periodic Laplace operator are $i\omega_p/2$ for $0 \leq |p|_\infty \leq K-1$, where $|p|_\infty = \max\{|p_1|, \dots, |p_d|\}$. Since this operator is also self-adjoint, the computation of the discretized versions of the operators $a_{k,\ell}$ and b_k amounts to the computation of the values of

$$a_{k,\ell}(ih\omega_p/2) = \frac{1}{h} \int_0^{c_k h} e^{(c_k h - \sigma)i\omega_p/2} \mathcal{L}_\ell(\sigma) d\sigma, \quad (2.6)$$

and

$$b_k(ih\omega_p/2) = \frac{1}{h} \int_0^h e^{(h-\sigma)i\omega_p/2} \mathcal{L}_k(\sigma) d\sigma. \quad (2.7)$$

We explain how one can compute these coefficients. These functions $a_{k,\ell}$ and b_k are holomorphic functions over \mathbb{C} . After integration (*e.g.* using integration by parts) over $(0, h)$ of the integrands defining these operators, we obtain holomorphic functions (say, of hL) with a removable singularity at the origin of the complex plane. For example, when the method has $s = 2$ stages, the coefficient $a_{1,1}$ reads

$$a_{1,1}(ih\omega_p/2) = c_1^2 \frac{e^{ic_1 h \frac{\omega_p}{2}} (1 - ic_2 h \frac{\omega_p}{2}) - 1 + ih \frac{\omega_p}{2} (c_2 - c_1)}{(c_1 - c_2)((ic_1 h \omega_p/2)^2)}. \quad (2.8)$$

Therefore, applying directly a formula such as (2.8) yields numerical instabilities in the computation of the ratio when $h|\omega_p| \ll 1$. To circumvent this problem, we use a method for computing the coefficients that distinguishes between those such that $h|\omega_p| \leq 1/2$ and those such that $h|\omega_p| > 1/2$ following in that way the method presented in [39, 44]. Essentially, we compute the former coefficients using a discretized version of the Cauchy representation formula for a holomorphic function f

$$f(z) = \frac{1}{2i\pi} \int_{\mathcal{C}} \frac{f(\omega)}{\omega - z} d\omega, \quad (2.9)$$

(where \mathcal{C} is the positively oriented unit circle and $|z| < 1$). For the latter coefficients, we simply evaluate directly formulas of the form (2.8).

Remark 5. For the evaluation of $f(ih\omega_p)$ when $h|\omega_p| \leq 1/2$, we use the trapezoidal quadrature rule to compute an approximation of the integral in (2.9). Since the unit circle is a smooth path in the complex plane, the function $t \mapsto ie^{it}f(e^{it})/(e^{it} - ih\omega_p)$ is a smooth 2π -periodic function and hence the trapezoidal rule has infinite order.

One can in fact compute these $s \times (s+1) \times K^2$ coefficients independently and use parallel computing. Once these coefficients are computed, we can begin the time-stepping method and no additional computation is required provided the discretization parameters h and K are fixed.

2.3 Result

For the solution of the Cauchy problem (1.11), using an ERK method defined by the s points $0 \leq c_1 < \dots < c_s \leq 1$, we have the following

Theorem 6. *For all $\psi_0 \in H^\sigma(\mathbb{T}_\delta^d)$ and all $T > 0$ such that the exact solution of (1.11) is smooth over $[0, T]$, there exists $C, h_0 > 0$ such that for all $h \in (0, h_0)$, the ERK method defined by (2.1)-(2.4) starting from ψ_0 is well-defined. Moreover, we have for all $h \in (0, h_0)$ and $n \in \mathbb{N}$ such that $nh \leq T$,*

$$\|\psi(t_n) - \psi_n\|_{H^\sigma(\mathbb{T}_\delta^d)} \leq Ch^s$$

Proof. This is a consequence of Theorem 3.6 in [26]. Let us check the hypotheses of Theorem 3.6. Hypothesis 3.1 is straightforward from (1.12). Hypothesis 3.2 follows from Lemma 2 and the fact that $H^\sigma(\mathbb{T}_\delta^d)$ is an algebra for $\sigma > d/2$. Hypotheses 3.3 and 3.4 follow from our assumption on the temporal smoothness of the exact solution. \square

Remark 7. *Note that our proof extends to general nonlinear Schrödinger equations with smooth time-dependant potential on the whole space \mathbb{R}^d as soon as $\sigma > d/2$.*

Remark 8. *Even if we are not able to prove it, we observe in the numerical experiments that if the s collocation points of the ERK method are Gauss points, the method is of numerical order $2s$ (see section 4.1). This indeed may be true when the methods are applied to the problem (1.11) on the torus \mathbb{R}^d because of the periodic boundary conditions while it may not be true when the methods are applied to the problem (1.9) set on the whole space \mathbb{R}^d . This is the case for example when one considers such methods applied to semilinear parabolic problems [36]. This would indeed be another limitation of the reduction of the problem from \mathbb{R}^d to \mathbb{T}^d (see Section 1.2). In the following we call Gauss-ERK method an ERK method using Gauss collocation points.*

3 Lawson methods

In [40], Lawson considers the problem of designing some Runge–Kutta type methods for stiff ordinary differential equations. The idea is to perform a change of unknowns to transform the stiff system into a related nonstiff one. Then some basic Runge–Kutta method is applied to the related problem. The combination is termed a generalized Runge–Kutta method in [40] and often called Lawson method.

The goal of this section is to describe the implementation of such methods on the problem (1.11) seen as an ordinary differential equation in time, to perform an analysis of the order of convergence of these methods as well as some of their preservation properties and to illustrate our results numerically.

3.1 Notations and description of the method

Let us consider the equation (1.11), and set the following change of unknowns (also called Lawson transformation),

$$u(t, \mathbf{x}) = e^{-Lt}\psi(t, \mathbf{x}). \tag{3.1}$$

Then ψ solves (1.11) if and only if u solves

$$\begin{aligned} \partial_t u(t, \mathbf{x}) &= e^{-Lt} N_w(t, e^{Lt} u(t, \mathbf{x})), \quad (t, \mathbf{x}) \in [0, T] \times \mathbb{T}_\delta^d \\ u(0, \mathbf{x}) &= \psi(0, \mathbf{x}) = \psi_0(\mathbf{x}), \quad \mathbf{x} \in \mathbb{T}_\delta^d. \end{aligned} \quad (3.2)$$

Now one can apply a classical Runge–Kutta method to (3.2) seen as an ordinary differential equation in time. Assume $(a_{k,\ell})_{1 \leq k, \ell \leq s}$ and $(b_k)_{1 \leq k \leq s}$ are given real numbers. Set

$$\forall k \in \{1, \dots, s\}, \quad c_k := \sum_{\ell=1}^s a_{k,\ell}, \quad (3.3)$$

and consider the s -stages classical Runge–Kutta method with Butcher tableau given by

$$\begin{array}{c|ccc} c_1 & a_{1,1} & \cdots & a_{1,s} \\ \vdots & \vdots & & \vdots \\ c_s & a_{s,1} & \cdots & a_{s,s} \\ \hline & b_1 & \cdots & b_s \end{array}$$

Assume that this Runge–Kutta method is of order at least 1. This means that the following condition is fulfilled:

$$\sum_{k=1}^s b_k = 1. \quad (3.4)$$

Applying this Runge–Kutta method to the problem (3.2) defines a Lawson method : we compute an approximation u_{n+1} at time t_{n+1} of the exact solution from an approximation u_n at time t_n by solving the system of s nonlinear equations (the unknowns being the $(u_{n,k})_{1 \leq k \leq s}$):

$$u_{n,k} = u_n + h \sum_{\ell=1}^s a_{k,\ell} e^{-(t_n + c_\ell h)L} N_w \left(t_n + c_\ell h, e^{(t_n + c_\ell h)L} u_{n,\ell} \right), \quad (3.5)$$

and then we compute u_{n+1} through the formula

$$u_{n+1} = u_n + h \sum_{k=1}^s b_k e^{-(t_n + c_k h)L} N_w \left(t_n + c_k h, e^{(t_n + c_k h)L} u_{n,k} \right). \quad (3.6)$$

Equivalently, the Lawson method for the unknowns

$$\psi_{n,i} := e^{(t_n + c_i h)L} u_{n,i} \quad \text{and} \quad \psi_n := e^{t_n L} u_n, \quad (3.7)$$

consists in solving the s nonlinear equations

$$\psi_{n,k} = e^{c_k h L} \psi_n + h \sum_{\ell=1}^s a_{k,\ell} e^{(c_k - c_\ell) h L} N_w(t_n + c_\ell h, \psi_{n,\ell}), \quad (3.8)$$

and then computing ψ_{n+1} through the formula

$$\psi_{n+1} = e^{hL} \psi_n + h \sum_{k=1}^s b_k e^{(1-c_k)hL} N_w(t_n + c_k h, \psi_{n,k}). \quad (3.9)$$

We simply denote these relations (3.8)-(3.9) by $\psi_{n+1} = \Phi_{t_n \rightarrow t_{n+1}}(\psi_n)$. As for the ERK methods of Section 2, the Runge–Kutta method defined by $a_{k,\ell}, b_k$ is referred to as the underlying Runge–Kutta method.

Note that, in view of Lemma 2, the Lawson method (3.7)-(3.9) is well defined in $H^\sigma(\mathbb{T}_\delta^d)$ provided $h > 0$ is small enough.

3.2 Results

In this section, we present some results on the Lawson method given by (3.8)-(3.9). First of all, since equation (1.11) is time reversible, we give a sufficient condition for the Lawson method to be symmetric. We follow ideas developed in [19], where the authors are solving an autonomous NLS equation, and we show that although equation (1.11) is non autonomous, the sufficient condition is quite similar.

Theorem 9. *Assume that the s -stages Runge-Kutta method defined by $(a_{k,\ell})_{1 \leq k, \ell \leq s}$ and $(b_k)_{1 \leq k \leq s}$ satisfies (3.4) so that it is of order at least 1. Assume that this method satisfies*

$$\forall(k, \ell) \in \{1, \dots, s\}^2, \quad a_{s+1-k, s+1-\ell} + a_{k,\ell} = b_\ell, \quad (3.10)$$

so that it is symmetric (see Theorem 2.3 in [33]). Then the Lawson method defined by (3.8)-(3.9) is also symmetric.

Proof. First of all, let us mention that the symmetry condition (3.10) gives

$$b_\ell = b_{s+1-\ell}, \quad 1 \leq \ell \leq s. \quad (3.11)$$

Moreover, summing (3.10) over ℓ and using (3.4) we have

$$1 - c_k = c_{s+1-k}, \quad 1 \leq k \leq s. \quad (3.12)$$

The adjoint $\Phi_{t_n \rightarrow t_{n+1}}^*$ of the method $\Phi_{t_n \rightarrow t_{n+1}}$ is by definition $\Phi_{t_{n+1} \rightarrow t_n}^{-1}$. The relation $\hat{\psi}_{n+1} = \Phi_{t_n \rightarrow t_{n+1}}^*(\hat{\psi}_n)$ is equivalent to $\hat{\psi}_n = \Phi_{t_{n+1} \rightarrow t_n}(\hat{\psi}_{n+1})$. This corresponds to exchanging t_n with t_{n+1} and h with $-h$ in (3.8)-(3.9). This leads to

$$\hat{\psi}_{n,k} = e^{-c_k h L} \hat{\psi}_{n+1} - h \sum_{\ell=1}^s a_{k,\ell} e^{-(c_k - c_\ell) h L} N_w(t_{n+1} - c_\ell h, \hat{\psi}_{n,\ell}), \quad (3.13)$$

$$\hat{\psi}_n = e^{-h L} \hat{\psi}_{n+1} - h \sum_{k=1}^s b_k e^{-(1-c_k) h L} N_w(t_{n+1} - c_k h, \hat{\psi}_{n,k}). \quad (3.14)$$

Extracting $\hat{\psi}_{n+1}$ from (3.14) gives

$$\hat{\psi}_{n+1} = e^{h L} \hat{\psi}_n + h \sum_{k=1}^s b_k e^{c_k h L} N_w(t_{n+1} - c_k h, \hat{\psi}_{n,k}). \quad (3.15)$$

Plugging this expression into (3.13), we get

$$\hat{\psi}_{n,k} = e^{(1-c_k) h L} \hat{\psi}_n + h \sum_{\ell=1}^s (b_\ell - a_{k,\ell}) e^{-(c_k - c_\ell) h L} N_w(t_{n+1} - c_\ell h, \hat{\psi}_{n,\ell}).$$

Using (3.10), we infer

$$\hat{\psi}_{n,k} = e^{(1-c_k) h L} \hat{\psi}_n + h \sum_{\ell=1}^s a_{s+1-k, s+1-\ell} e^{-(c_k - c_\ell) h L} N_w(t_{n+1} - c_\ell h, \hat{\psi}_{n,\ell}). \quad (3.16)$$

Setting $\tilde{\psi}_{n,k} = \hat{\psi}_{n, s+1-k}$ and reordering the sums in (3.15)-(3.16), we obtain

$$\tilde{\psi}_{n,k} = e^{(1-c_{s+1-k}) h L} \hat{\psi}_n + h \sum_{\ell=1}^s a_{k,\ell} e^{-(c_{s+1-k} - c_{s+1-\ell}) h L} N_w(t_n + (1 - c_{s+1-\ell}) h, \tilde{\psi}_{n,\ell}),$$

$$\hat{\psi}_{n+1} = e^{hL}\hat{\psi}_n + h \sum_{k=1}^s b_{s+1-k} e^{c_{s+1-k}hL} N_w \left(t_n + (1 - c_{s+1-k})h, \tilde{\psi}_{n,k} \right).$$

Using (3.11)-(3.12) we conclude that

$$\begin{aligned} \tilde{\psi}_{n,k} &= e^{c_k hL} \hat{\psi}_n + h \sum_{\ell=1}^s a_{k,\ell} e^{(c_k - c_\ell)hL} N_w \left(t_n + c_\ell h, \tilde{\psi}_{n,\ell} \right), \\ \hat{\psi}_{n+1} &= e^{hL} \hat{\psi}_n + h \sum_{k=1}^s b_k e^{(1-c_k)hL} N_w \left(t_n + c_k h, \tilde{\psi}_{n,k} \right). \end{aligned}$$

This proves that the Lawson method is symmetric. \square

If the underlying Runge–Kutta method preserves quadratic invariants in the sense that it satisfies the Cooper condition, then so does the associated Lawson method.

Theorem 10. *Assume that the underlying Runge–Kutta method satisfies (3.4) so that it is of order at least 1. Assume that it satisfies the Cooper condition,*

$$b_k a_{k,\ell} + b_\ell a_{\ell,k} = b_k b_\ell, \quad \forall 1 \leq k, \ell \leq s, \quad (3.17)$$

so that it preserves quadratic invariants. Then the Lawson method (3.8)-(3.9) preserves the L^2 -norm:

$$\|\psi_n\|_{L^2(\mathbb{T}_\delta^d)} = \|\psi_0\|_{L^2(\mathbb{T}_\delta^d)}, \quad \forall n \geq 0.$$

Proof. Since the evolution equation (1.11) preserves the L^2 -norm, and so does the change of variables (3.1), the evolution equation (3.2) also preserves the L^2 -norm. Therefore $t \mapsto \int_{\mathbb{T}_\delta^d} |\psi(t, \mathbf{x})|^2 d\mathbf{x}$ is a quadratic invariant of equation (3.2). The Cooper condition guarantees that the Runge–Kutta method (3.5)-(3.6) preserves quadratic invariants. Hence for all n , $\int_{\mathbb{T}_\delta^d} |v_n(\mathbf{x})|^2 d\mathbf{x} = \int_{\mathbb{T}_\delta^d} |v_0(\mathbf{x})|^2 d\mathbf{x}$ and then $\int_{\mathbb{T}_\delta^d} |\psi_n(\mathbf{x})|^2 d\mathbf{x} = \int_{\mathbb{T}_\delta^d} |\psi_0(\mathbf{x})|^2 d\mathbf{x}$. \square

Definition 11. *We call Gauss–Lawson method any Lawson method of the form (3.8)-(3.9) such that the underlying Runge–Kutta method is a Gauss collocation method (see section II.1.3 in [33]).*

Corollary 12. *A Gauss–Lawson method applied to (1.11) is symmetric and preserves the L^2 -norm.*

Proof. The underlying Runge–Kutta method is a collocation method at Gauss points and therefore it is symmetric (see Corollary 2.2, Chapter V in [33]). Hence the Gauss–Lawson method applied to (1.11) is symmetric by theorem 9. Moreover the underlying Runge–Kutta method satisfies the Cooper condition (3.17) (see exercise 5 in chapter IV in [33]). Therefore the Gauss–Lawson method preserves the L^2 -norm by theorem 10. \square

We now want to prove that the Gauss–Lawson method with s stages applied to (1.11) has order $2s$ in $H^\sigma(\mathbb{T}_\delta^d)$ for $\sigma > d/2$ (see Theorem 14). Our strategy is the following. First, we consider an equivalent autonomous form of (1.11). Second, we show that applying a Lawson method to this autonomous form is essentially the same as applying the method to (1.11) directly. Third, we rely on an Alekseev–Gröbner lemma for autonomous systems

which provides a representation of the error that allows to conclude that, since Gauss-Lawson methods are collocation methods, they have order $2s$.

Let us set

$$U(t) = \begin{pmatrix} t \\ u(t) \end{pmatrix} \in \mathbb{R} \times H^\sigma(\mathbb{T}_\delta^d),$$

so that (1.11) reads

$$\frac{d}{dt}U(t) = F(U(t)), \quad (3.18)$$

with

$$F \begin{pmatrix} t \\ u \end{pmatrix} = \begin{pmatrix} 1 \\ e^{-tL}N_w(t, e^{tL}u) \end{pmatrix}. \quad (3.19)$$

We have the following useful

Lemma 13. *Let us fix $u_0 \in H^\sigma(\mathbb{T}_\delta^d)$. For all $h > 0$ sufficiently small, we apply the Lawson method (3.5)-(3.6) and denote by u_n the corresponding numerical values. Similarly, we start with $U_0 = (0, u_0)^t$, we apply the same method to the Cauchy problem (3.18) with initial datum $U(0) = U_0$ and we denote by U_n the corresponding numerical values. We have for all $n \in \mathbb{N}$ such that $nh \leq T$,*

$$U_n = \begin{pmatrix} nh \\ u_n \end{pmatrix}.$$

Proof. Since the function N_w satisfies a local Lipschitz condition (see Lemma 2), so does F (see (3.19)). Hence, the Lawson methods are well defined locally for $h > 0$ small enough. We perform the proof by induction. The relation is true for $n = 0$. Assume it holds for some $n \in \mathbb{N}$ such that $(n+1)h \leq T$. The definition of the coefficients $(c_k)_{1 \leq k \leq s}$ (see (3.3)) and the first component of the function F ensure that for all $k \in \{1, \dots, s\}$, the first coefficient of $U_{n,k}$ is $t_{n,k} = nh + c_k h$. Therefore, in view of (3.5), the $(U_{n,k})_{1 \leq k \leq s} = (nh + c_k h, u_{n,k})^t$ are the unique local solutions for the Lawson inner problem in U . Similarly, the consistency relation (3.4) ensures that the first component of U_{n+1} is $(n+1)h$. Hence, in view of (3.6), we have that the second component of U_{n+1} is actually u_{n+1} . \square

We are now able to state

Theorem 14. *Assume $u_0 \in H^\sigma(\mathbb{T}_\delta^d)$ and $T > 0$ are given. There exists constants $C > 0$ and $h_0 > 0$ such that the exact solution $t \mapsto u(t)$ of the Cauchy problem (1.11) and the corresponding numerical approximations provided by the Gauss-Lawson method with s stages u_n satisfy*

$$\forall h \in (0, h_0), \forall n \in \mathbb{N} \text{ s.t. } 0 \leq nh \leq T, \quad \|u(t_n) - u_n\|_{H^\sigma(\mathbb{T}_\delta^d)} \leq Ch^{2s}.$$

Proof. Thanks to Lemma 13, it is sufficient to prove that the Gauss-Lawson method applied to the Cauchy problem (3.18) with initial datum $U_0 = (0, u_0)$ is of order $2s$. We chose $h_0 > 0$ sufficiently small to ensure that the method is well-defined for all $h \in (0, h_0)$. We follow the proof of [38] for collocation methods applied to ODEs. First, we recall that the mapping $U_n \mapsto U_{n+1}$ can be defined through a collocation problem. The same problem

allows to express the consistency error: Find P a polynomial of degree at most s (with coefficients in $\mathbb{R} \times H^\sigma(\mathbb{T}_\delta^d)$) such that

$$\begin{cases} P(t_n) = U(t_n), \\ P'(t_n + c_k h) = F(P(t_n + c_k h)), \end{cases} \quad \text{for } 1 \leq k \leq s,$$

and set $\tilde{U}_{n+1} = P(t_{n+1})$, so that the consistency error reads $\tilde{U}_{n+1} - U(t_{n+1})$. Second, define the default function d as

$$d(t, P(t)) = P'(t) - F(P(t)).$$

Then, we use the Alekseev-Gröbner lemma (Theorem 2 of [24]), which ensures that there exists a smooth function g (corresponding to $\partial_2 \mathcal{E}_G$ in [24]) from $[0, T] \times (\mathbb{R} \times H^\sigma(\mathbb{T}_\delta^d))$ to the linear continuous operators of $\mathbb{R} \times H^\sigma(\mathbb{T}_\delta^d)$ such that the following identity holds in $H^\sigma(\mathbb{T}_\delta^d)$: for all $t \in [0, h]$,

$$P(t_n + t) - U(t_n + t) = \int_{t_n}^{t_n+t} g(t_n + t - \tau, P(\tau)) d(\tau, P(\tau)) d\tau.$$

Setting $t = h$ in this formula and using the fact that g is smooth and all the derivatives of $t \mapsto d(t, P(t))$ are bounded on $[0, h]$ independantly of h (see Lemma 1.6 of Chapter 2 in [33]), we infer that the consistency error in $H^\sigma(\mathbb{T}_\delta^d)$ is of order h^{2s+1} . Hence the global error is of order h^{2s} via a classical discrete Gronwall lemma. \square

4 Numerical experiments

In this section, we make some numerical experiments and show the numerical efficiency of the methods. First, we consider one dimensional problems and compare some ERK methods and Gauss–Lawson methods with the classical splitting methods. Second, we compare the same kinds of methods applied to an actual 2D problem with a rotating BEC.

4.1 One dimensional examples

We provide in this subsection some numerical experiments to show the efficiency of the ERK and the Gauss–Lawson methods compared with the traditional splitting methods. We use here splitting methods respectively of order 1, 2, 4 and 6 [12]. In order to present the usual splitting schemes, we have to define the operators S_L and S_{N_w} associated respectively to the evolution of the pdes

$$\partial_t u(t, \mathbf{x}) = Lu(t, \mathbf{x}) \quad \partial_t v(t, \mathbf{x}) = N_w v(t, \mathbf{x}),$$

where L and N_w are defined in (1.11). The operators satisfy by definition the following relations involving the exact solutions of the associated PDEs:

$$u(t + h, \mathbf{x}) = S_L(h)u(t, \mathbf{x}) \quad \text{and} \quad v(t + h, \mathbf{x}) = S_{N_w}(h)v(t, \mathbf{x}).$$

As explained for example in [12], a splitting idea for building a splitting scheme of even order consists in approximating the continuous flow associated to $\partial_t \zeta(t, \mathbf{x}) = L\zeta(t, \mathbf{x}) + N_w(t, \zeta(t, \mathbf{x}))$ by a composition of operators S_L and S_{N_w} of the form

$$S_{N_w}(a_1 h) S_L(b_1 h) \cdots S_{N_w}(a_r h) S_L(b_r h) S_{N_w}(a_{r+1} h),$$

with $r \geq 1$ and for all ℓ , $a_{r+2-\ell} = a_\ell$ and $b_{r+1-\ell} = b_\ell$. The coefficients of the splitting methods we implemented are given in table 1. Let us mention that alternative ways of

Order 2	Order 4	Order 6
$a_1 = 1/2$ $b_1 = 1$	$a_1 = \theta$ $a_2 = 1/2 - \theta$ $b_1 = 2\theta$ $b_2 = 1 - 4\theta$ $\theta = (2 + 2^{1/3} + 2^{-1/3})/6$	$a_1 = 0.0502627644003922$ $a_2 = 0.413514300428344$ $a_3 = 0.0450798897943977$ $a_4 = -0.188054853819569$ $a_5 = 0.541960678450780$ $a_6 = 1 - 2(a_1 + a_2 + a_3 + a_4 + a_5)$ $b_1 = 0.148816447901042$ $b_2 = -0.132385865767784$ $b_3 = 0.067307604692185$ $b_4 = 0.432666402578175$ $b_5 = 0.5 - (b_1 + b_2 + b_3 + b_4)$

Table 1: Coefficient of splitting methods.

deriving splitting methods for such problems exist, some of which authorize the use of complex coefficients (see for example [15],[13]).

We compute the numerical solution to the one-dimensional cubic NLS equation

$$\partial_t \psi = i \partial_x^2 \psi + iq |\psi|^2 \psi, \quad (t, x) \in [0, T] \times \mathbb{R}. \quad (4.1)$$

An exact solution for $(t, x) \in [0, T] \times \mathbb{R}$ is given by the soliton formula

$$\psi_{ex}(t, x) = \frac{2a}{q} \operatorname{sech}((x - x_0) - ct) \exp\left(ic \frac{(x - x_0) - ct}{2}\right) \exp\left(i\left(a + \frac{c^2}{4}\right)t\right).$$

For the numerical approximation of this solution, as explained in section 1.2, we take a periodic finite interval (x_ℓ, x_r) of big enough length, and we discretize the space operators using Fourier spectral approximation. We choose the spatial mesh size $k = \Delta x > 0$ with $k = (x_r - x_\ell)/M$ with $M = 2^P$, $P \in \mathbb{N}^*$. The time step is denoted by $h = \Delta t$ and $h = T/N_T$ for some $N_T \in \mathbb{N}^*$. The grid points and the discrete times are

$$x_j := x_\ell + jk, \quad t_n := nh, \quad j = 0, 1, \dots, M, \quad n = 0, 1, \dots, N_T.$$

Let ψ_j^n be the approximation of $\psi(t_n, x_j)$. Since we discretize (4.1) by the Fourier spectral method, ψ_j^n and its Fourier transform satisfy the following relations:

$$\psi_j^n = \frac{1}{M} \sum_{m=-M/2}^{M/2-1} \hat{\psi}_m^n e^{i\mu_m(x_j - x_\ell)}, \quad j = 0, \dots, M-1,$$

and

$$\hat{\psi}_m^n = \sum_{j=0}^{M-1} \psi_j^n e^{-i\mu_m(x_j - x_\ell)}, \quad m = -\frac{M}{2}, \dots, \frac{M}{2} - 1,$$

where $\mu_m = \frac{2\pi m}{x_r - x_\ell}$ for all $m = -\frac{M}{2}, \dots, \frac{M}{2} - 1$. Let us define the discrete gradient operator ∇_k

$$\widehat{(\nabla_k v)}_m = i\mu_m \hat{v}_m, \quad v \in \mathbb{C}^M.$$

Let us denote by Π_k the projection operator

$$\begin{aligned} \Pi_k : \mathcal{C}^0([x_\ell, x_r], \mathbb{C}) &\rightarrow \mathbb{C}^M \\ \varphi &\mapsto (\varphi(x_j))_{0 \leq j \leq M-1} . \end{aligned}$$

We define the discrete ℓ^r norm on \mathbb{C}^M as

$$\|v\|_{\ell^r} = \left(k \sum_{j=0}^{M-1} |v_j|^r \right)^{1/r}, \quad v \in \mathbb{C}^M, \quad r \geq 1.$$

Using these definitions, we consider the following errors:

$$\mathcal{E}_{P,h} = \sup_{n \in \{0, \dots, N\}} \left\| \Pi_k(\psi_{ex}(t_n, \cdot)) - (\psi_j^n)_j \right\|_{\ell^2}, \quad (4.2)$$

$$\mathcal{E}_{M,h} = \sup_{n \in \{0, \dots, N\}} \left(\left\| \Pi_k(\psi_{ex}(t_n, \cdot)) \right\|_{\ell^2} - \left\| (\psi_j^n)_j \right\|_{\ell^2} \right) / \left\| \Pi_k(\psi_{ex}(0, \cdot)) \right\|_{\ell^2}. \quad (4.3)$$

We also define the discrete energy:

$$E_k(v) = \frac{1}{2} \|\nabla_k v\|_{\ell^2}^2 - \frac{q}{4} \|v\|_{\ell^4}^4,$$

the energy conservation is seen through the following relative error

$$\mathcal{E}_{E,h} = \sup_{n \in \{0, \dots, N\}} \left(E_k(\Pi_k(\psi_{ex}(t_n, \cdot))) - E_k((\psi_j^n)_j) \right) / E_k(\Pi_k(\psi_{ex}(0, \cdot))). \quad (4.4)$$

For all the following simulations, we consider the computational domain $(x_\ell, x_r) = (-15, 15)$ and the final time $T = 5$. The experiments are performed with the discretization parameters $P = 10$ and various time steps k and the chosen physical parameters are $q = 8$, $a = q^2/16$ and $c = 0.5$. We provide for all methods the evolution of $\mathcal{E}_{P,h}$, $\mathcal{E}_{M,h}$ and $\mathcal{E}_{E,h}$ for various time steps h . All of our methods being implicit, it is necessary to solve a nonlinear problem at each time step for both ERK and Lawson method. This is performed here through a fixed point algorithm. It is therefore important to show the efficiency of the schemes to plot the evolution of the Cputime with respect to the error $\mathcal{E}_{P,h}$.

The legends of the figures are respectively given on Fig. 1(a) and Fig. 1(b) for ERK and Lawson methods (we recall that s denotes the number of collocation points) and splitting schemes. We showed in the previous sections that if one uses arbitrary distinct collocation points, the order of ERK and Lawson methods are h^s . Therefore, in all figures, curves associated to $s = 1, 2$ and 4 respectively correspond to splitting of order 1, 2 and 4. If we would like to compare the splitting of order 6 to ERK and Lawson methods, we would have to consider $s = 6$ collocation points.

We gather the evolution of $\mathcal{E}_{P,h}$ for various schemes on Fig. 2. Note that since we use Gauss collocation points, we clearly obtain numerically an order $2s$ for both Gauss-ERK and Gauss-Lawson methods with s stages. It is not surprising for Gauss-Lawson schemes thanks to theorem 14. The same result seems to be valid for ERK schemes. Let us remark that there is a saturation phenomenon with higher s . Let us remark that there is a saturation process that appears in $\mathcal{E}_{P,h}$ when h gets small, due to the boundary conditions (see the transformation of the problem from \mathbb{R}^d to \mathbb{T}_δ^d in section 1.2). Actually, we can

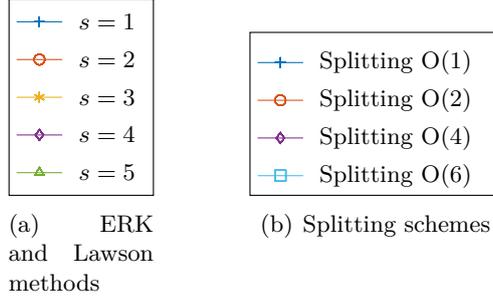


Figure 1: Legends

reduce this limiting error either by increasing the size δ of the computational domain or by changing the periodic boundary conditions to, for example, homogeneous Dirichlet ones (in this case, one can replace FFT with the discrete sine transform). If we compare figures 2(a) and 2(b), it is noticeable that the constants of error are better for Gauss-ERK method.

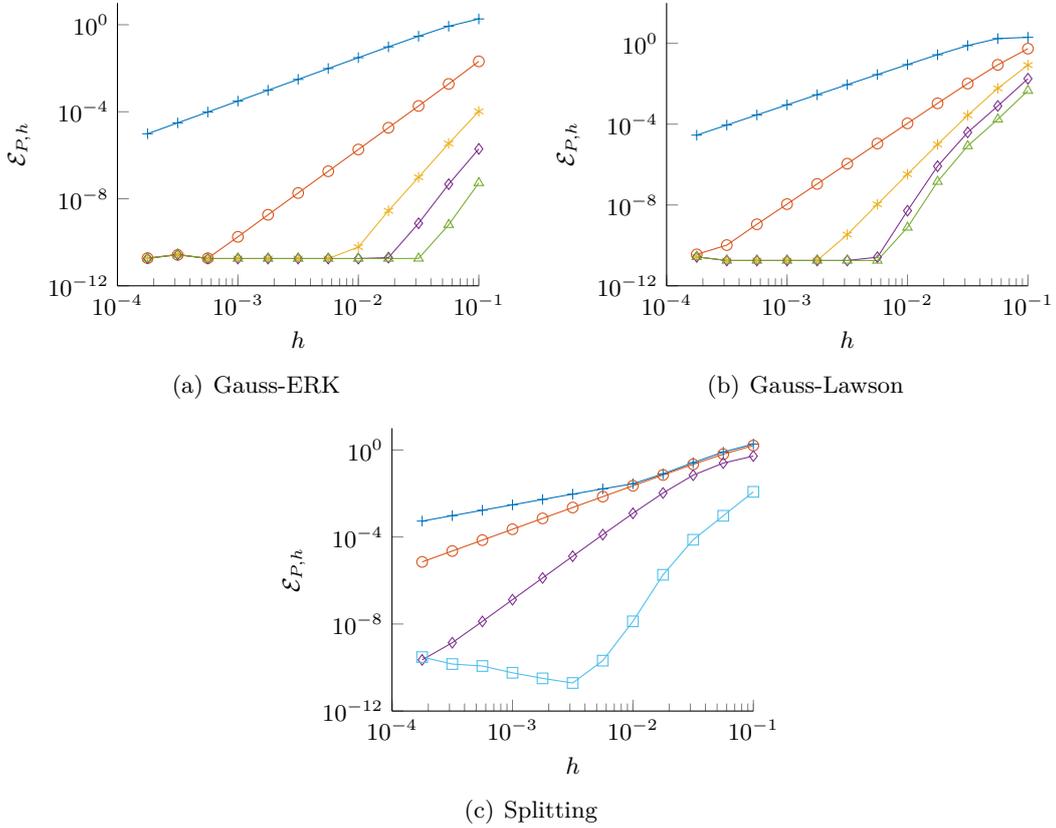


Figure 2: Evolution of $\mathcal{E}_{P,h}$ error with respect to the time step

The preservation of mass and energy are fundamental when one deals with dispersive equations. The figures presented on Fig. 3 show the evolution of $\mathcal{E}_{M,h}$ with respect to the time step for all schemes. As predicted by theorem 10, the L^2 norm is very well preserved by Gauss-Lawson schemes as for splitting schemes. The same preservation does not hold exactly with ERK methods, as one can check numerically. However, if we consider two

collocation Gauss points or more, the mass is preserved at an equivalent level for $h < 10^{-2}$.

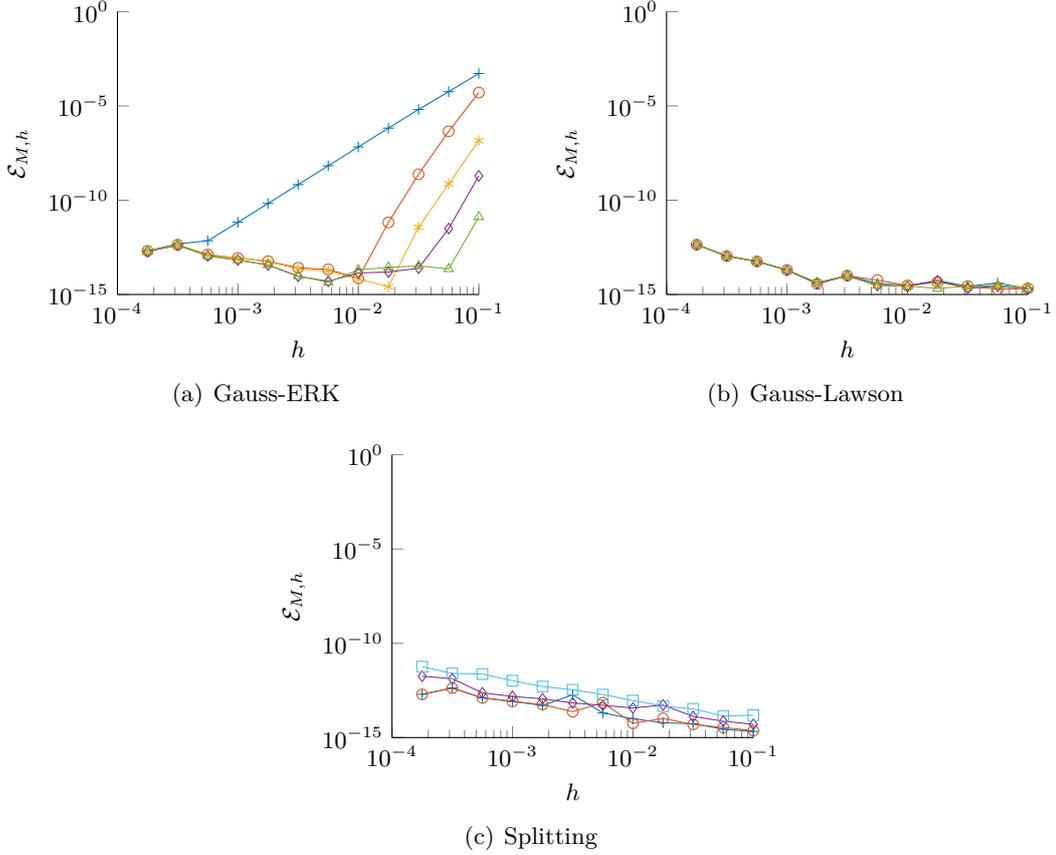


Figure 3: Evolution of $\mathcal{E}_{M,h}$ error with respect to the time step

Concerning the energy conservation, none of the presented methods is able to handle it properly. However, it is clear on Fig. 4, which displays the evolution of $\mathcal{E}_{E,h}$ as a function of the time step for all schemes, that, with respect to energy preservation, Gauss-ERK methods are of better quality than Gauss-Lawson ones which are themselves better than splitting schemes.

Finally, Fig. 5 shows the evolution of the Cputime with respect to the $\mathcal{E}_{P,h}$ error. These figures are very interesting since one could think that implicit methods are more costly compared to splitting schemes. For a given error $\mathcal{E}_{P,h}$, the Cpu time is clearly lower for Gauss-ERK schemes. For an “big” phase error $\mathcal{E}_{P,h} \geq 10^{-2}$, the splitting schemes are less costly than Gauss-Lawson ones. The situation is reversed when one is interested in “small” phase errors $\mathcal{E}_{P,h} \leq 10^{-2}$ and Gauss-Lawson schemes have to be recommended. For both ERK and Lawsons schemes, it is not necessarily interesting to use a high number of collocation points.

In conclusion, even if the ERK schemes do not preserve theoretically the L^2 norm, they allow to preserve it numerically for reasonable time step h and clearly are the better methods for one dimensional simulations.

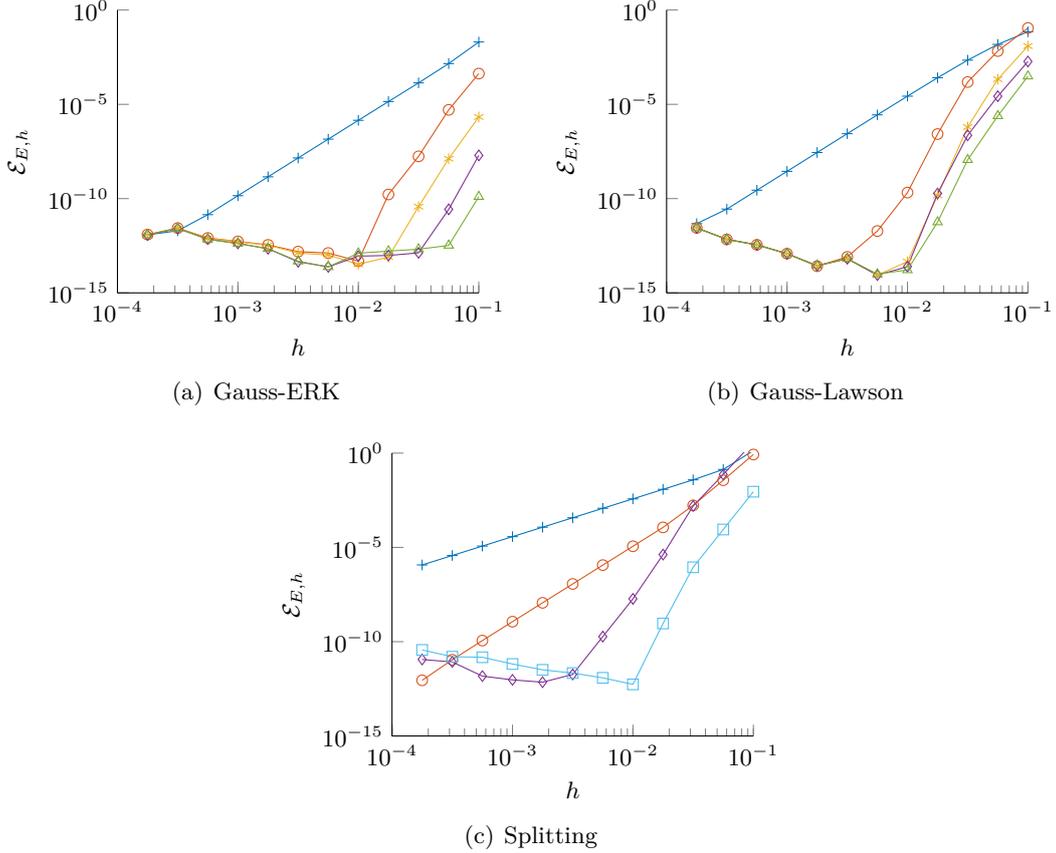


Figure 4: Evolution of $\mathcal{E}_{E,h}$ error with respect to the time step

4.2 Two dimensional simulations

We propose in this subsection to illustrate the efficiency of our methods on the simulation of a soliton for the two-dimensional equation and the simulation of the evolution of a Bose–Einstein condensate.

The first 2D example concerns the equation

$$\partial_t \psi = i\Delta \psi + i|\psi|^2 \psi, \quad (t, \mathbf{x}) \in [0, T] \times \mathbb{R}^2, \quad (4.5)$$

which is a nonlinear Schrödinger equation without confinement or rotation. We look for a solution of the form

$$\psi(t, \mathbf{x}) = e^{it} \Theta(\mathbf{x}),$$

where Θ solves

$$\Delta \Theta + |\Theta|^2 \Theta = \Theta,$$

with $\lim_{\|\mathbf{x}\| \rightarrow \infty} \Theta(\mathbf{x}) = 0$. Since we do not have access to an exact solution Θ to this problem, we generate it through a classical shooting point method [25]. Since the decay of this kind of solutions is slow, we consider $(x, y) \in [-38, 38]^2$ with $P = 9$.

As for one dimensional simulations, we evaluate several errors for the different methods. We generalize the error functions $\mathcal{E}_{P,h}$, $\mathcal{E}_{M,h}$ and $\mathcal{E}_{E,h}$ defined in (4.2), (4.3) and (4.4) to two dimensional cases. We only present here simulations for ERK and Lawson methods

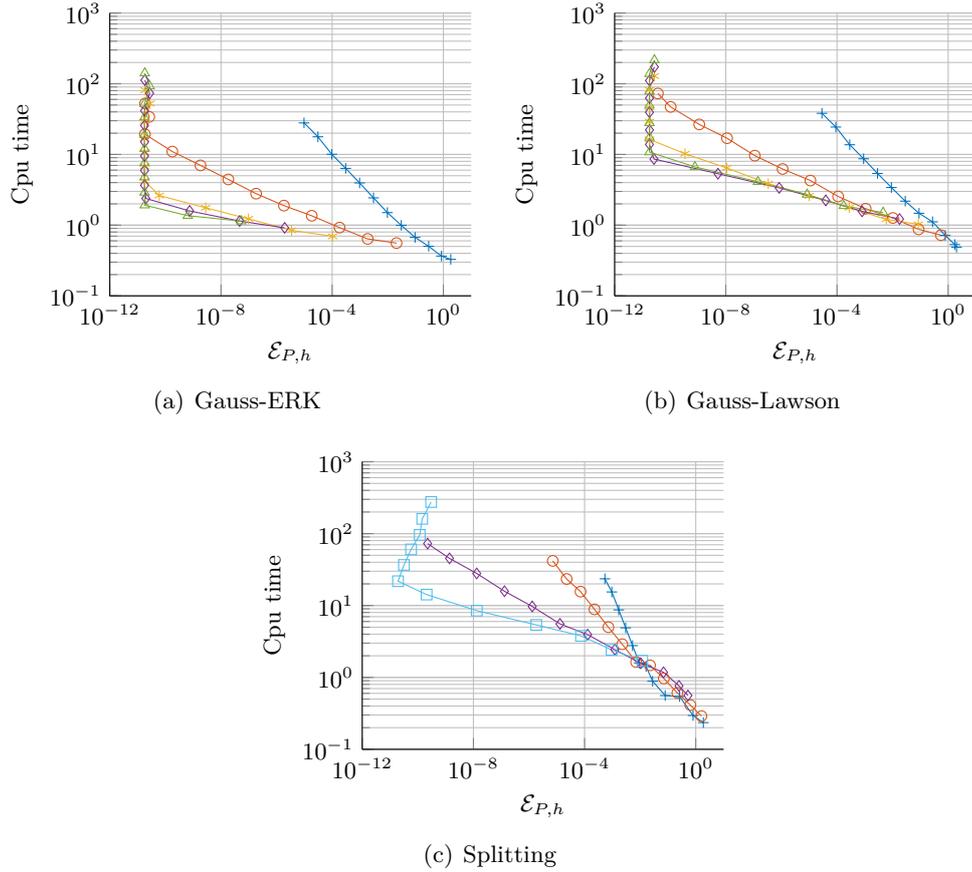


Figure 5: Evolution of the Cputime with respect to $\mathcal{E}_{P,h}$ error

with $s = 1, 2$ and 3 since we have seen on one dimensional experiments that we do not gain much more precision for $s > 3$. The numerical results are displayed on figures 6, 7, 8 and 9.

Our second 2D numerical example corresponds to a rotating BEC modelled by equation (1.2). We reproduce here with the Gauss-ERK method of order 6 the simulations realized in [6] for (1.2), see figure 10. The parameters are $\beta = 1000$, $\Omega = 0.9$, and the potential is (1.3) with $\gamma_x = 1.05$ and $\gamma_y = 0.95$. The computational domain is $(-16, 16)^2$ with 2^9 Fourier modes in each direction. The time step is $h = 10^{-3}$ with a final time $T = 7$. The initial datum is the ground state of the stationary equation and was generated using Matlab toolbox GPESLab¹. We recover the same behaviour and we get good conservation properties.

Conclusion

We presented ERK and Lawson methods that allow to compute numerical solutions of NLS equations modelling a rotating BEC in a fairly general setting. This procedure allows to derive neatly high order methods (in contrast to finding coefficients for high

¹<http://gpelab.math.cnrs.fr>

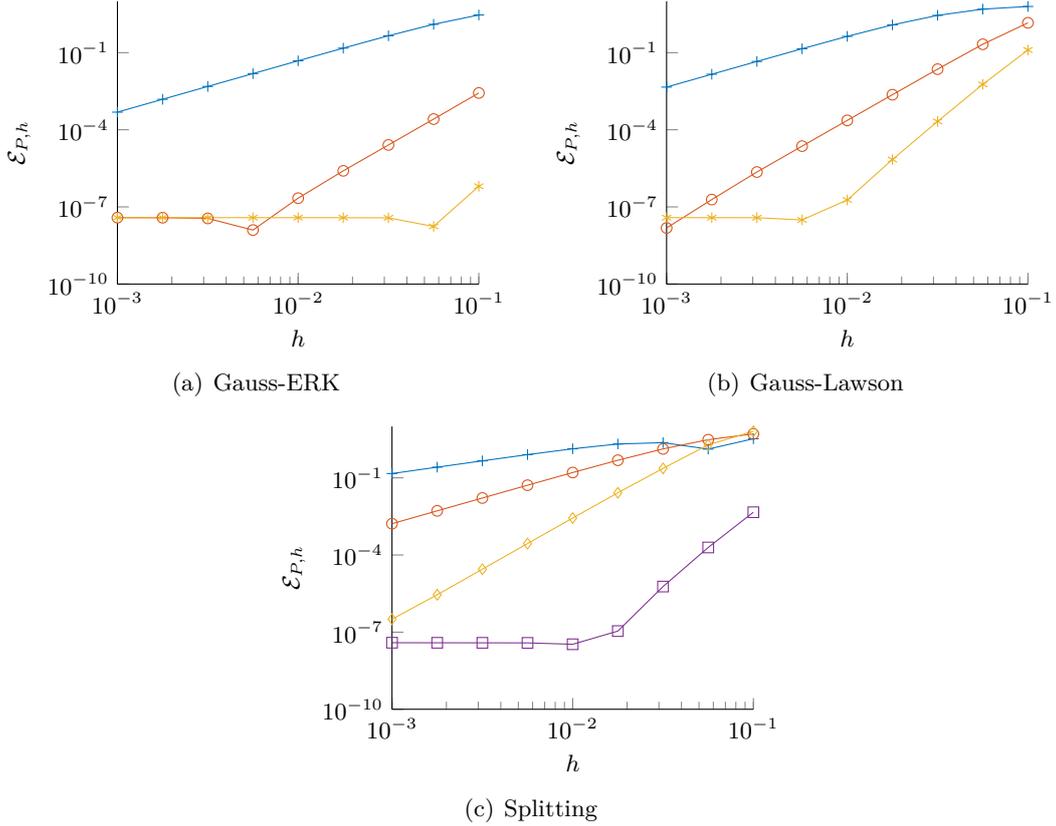


Figure 6: Evolution of $\mathcal{E}_{P,h}$ error with respect to the time step

order splitting methods with small error constants, for example). We proved finite-time convergence in Sobolev norms for these methods in a simplified framework. We compared the numerical results provided by these methods to that obtained via classical splitting methods in several configurations : 1D problems, 2D problems without confinement or rotation, 2D realistic problems with rotation and confinement. It is worth noticing that all the methods presented in this paper allow to deal with non-autonomous problems (no matter whether the lack of autonomy comes from the physical situation or a change of unknown).

When it comes to finite time accuracy, Gauss-ERK methods outperformed Lawson methods and splitting methods in all cases, since they have very low error constants numerically. Moreover, even if they do not preserve mass exactly (up to roundoff errors), their relative error on the mass is comparable to that of Gauss-Lawson and splitting methods (which preserve the mass up to roundoff errors) for reasonably small time steps (see Fig. 3 for time steps of order 10^{-4}). This can be explained by the accumulation of roundoff errors when h gets smaller for the methods preserving mass exactly.

When it comes to computational times, Gauss-ERK methods, though implicit, also outperformed Lawson and splitting methods (see Fig 5) on the 1D example. To be completely fair, one must say that ERK methods (and, to a lesser extent, Lawson methods) require the precomputation of some coefficients. This computation only needs to be carried out once, before one starts the time stepping procedure, and can be parallelized if necessary.

With respect to average or long time behaviour, in contrast to splitting methods, ERK

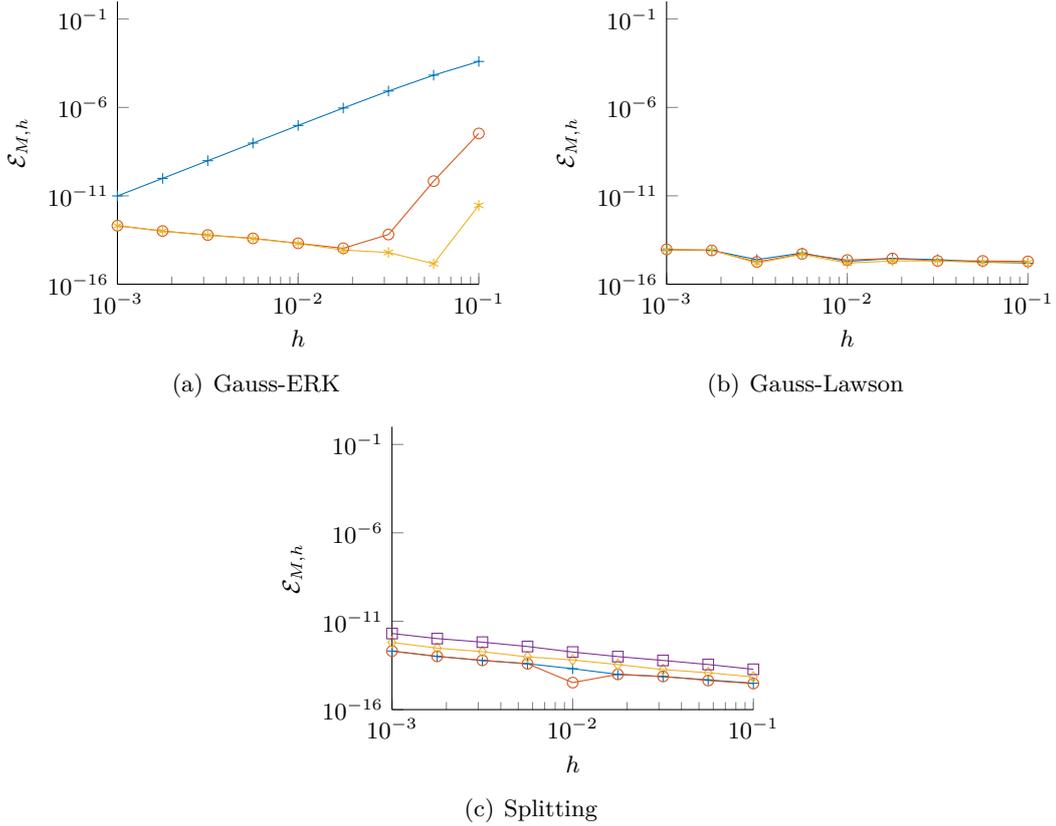


Figure 7: Evolution of $\mathcal{E}_{M,h}$ error with respect to the time step

and Lawson methods do not show resonances phenomena [27, 30].

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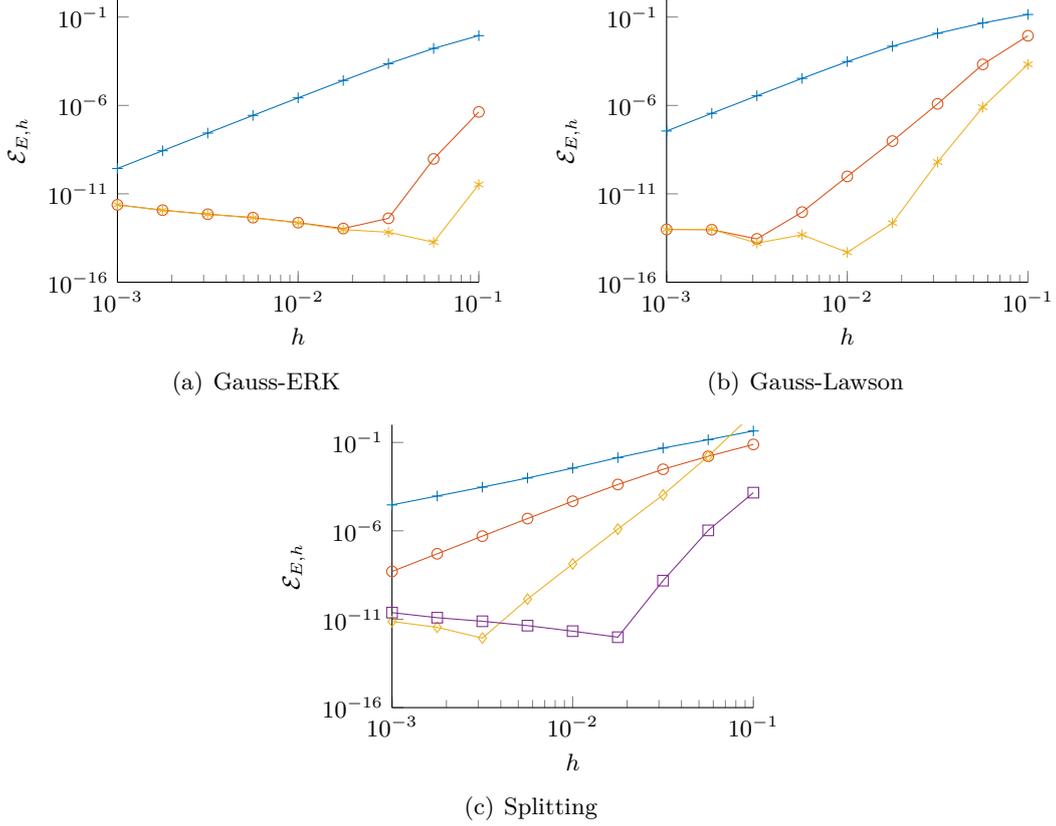


Figure 8: Evolution of $\mathcal{E}_{E,h}$ error with respect to the time step

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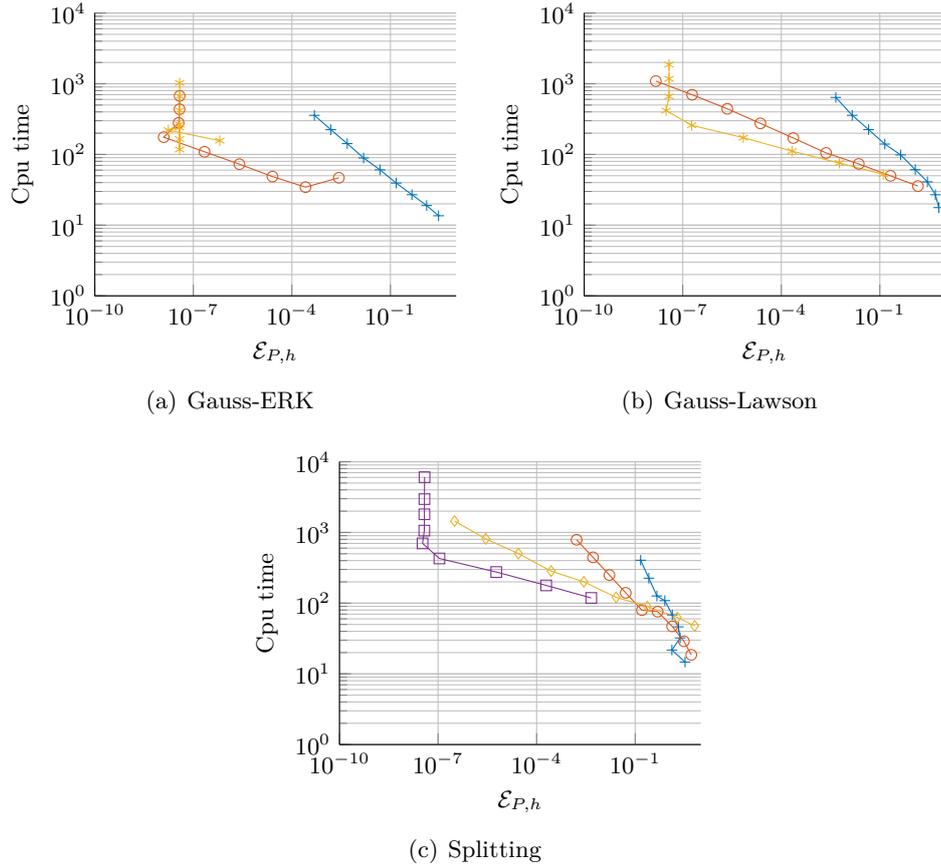


Figure 9: Evolution of the Cputime with respect to $\mathcal{E}_{P,h}$ error

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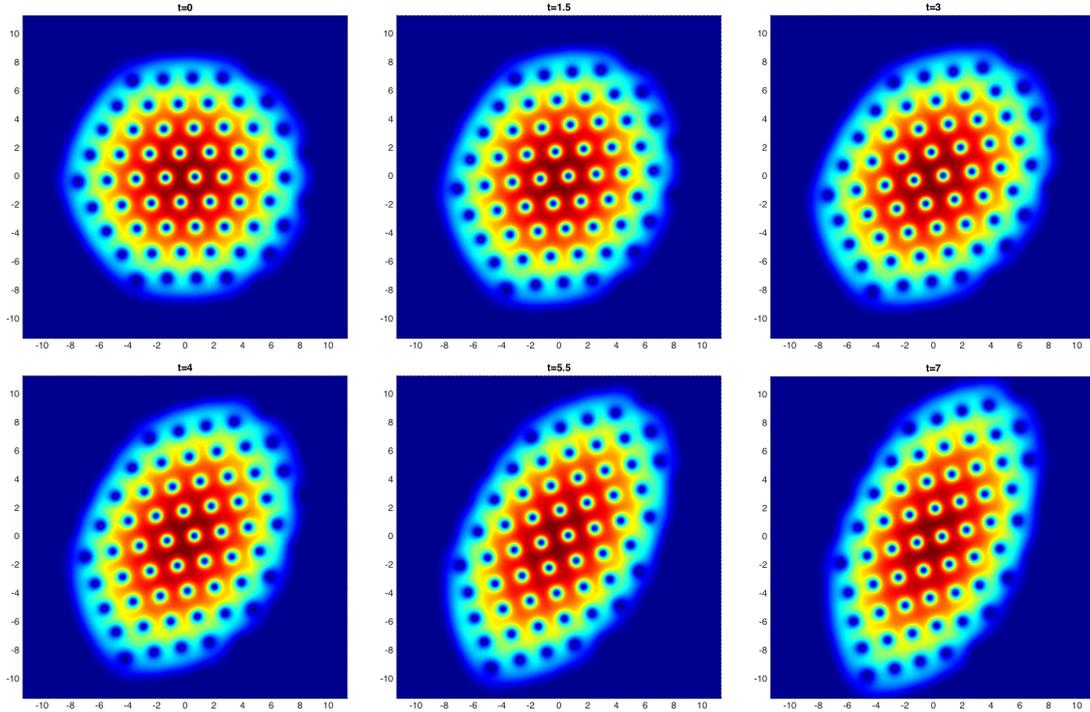


Figure 10: Contour plots of the density function $|\varphi(t, \mathbf{x})|^2$ in a rotating BEC.

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