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#### LETTER

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# FTIR and Raman spectroscopy characterization of fluorine-bearing titanian clinohumite in antigorite serpentinite and chlorite harzburgite

Kenneth T Koga<sup>1,2,3\*</sup>, Carlos J Garrido<sup>4</sup>, José A Padrón-Navarta<sup>5</sup>, Vicente López Sánchez-Vizcaíno<sup>6</sup> and María T Gómez-Pugnaire<sup>7</sup>

#### Abstract

Titanian clinohumite is an accessory phase found in peridotites, and it can be a key repository of volatile and other trace elements in sub-arc mantle. To characterize spectroscopic variations due to volatile elements in the phase, we have investigated the infrared spectroscopic characteristics of a suite of naturally occurring Ti clinohumite minerals with varying F and Ti concentrations. Samples were ultramafic rocks from Cerro del Almirez, Spain and Cima di Gagnone, Switzerland. The infrared spectra corresponding to OH vibrations showed systematic variation in regard to the abundance of F in Ti clinohumite. In particular, the intensity ratios of infrared absorption spectra at 3,565 and 3,390 cm<sup>-1</sup> correlated linearly with the abundance of F independently of the crystal orientation. Raman spectra did not corroborate the Fourier transform infrared (FTIR) results. This intensity ratio provides an alternative analytical criterion for the characterization of Ti clinohumite.

#### **Findings**

#### Introduction

Titanian clinohumite (TiChu) is a member of the humite group minerals that constitute a polysomatic series with olivine (Thompson 1978; Veblen 1991, 1992). The generalized mineral formula of humite group minerals is  $[n(M_2SiO_4)M_{1 - x}Ti_x(OH, F)_{2 - 2x}O_{2x}]$  where 0 < x < 0.5, M represents divalent cations (mostly Mg with minor Fe, Mn, and Ni), and *n* is one for norbergite, two for chondrodite, three for humite, and four for clinohumite. Clinohumite is the humite group mineral whose crystal structure is most closely related to olivine because it has the highest proportion of filled tetrahedral sites.

TiChu is a common accessory mineral in hydrous ultrabasic rocks including antigorite serpentinite (Damour 1879; Brugnatelli 1904; De Quervain 1938; Bearth 1967; Möckel 1969; Trommsdorff and Evans 1980), metaperidotite from high-pressure metamorphic terrains (Möckel

\* Correspondence: k.koga@opgc.univ-bpclermont.fr

Full list of author information is available at the end of the article



The OH/F ratio of TiChu plays a particularly important role in the stabilization of TiChu at high temperature (Engi and Lindsley 1980; Trommsdorff and Evans 1980; Evans and Trommsdorff 1983), and it is crucial for its stabilization at mantle temperatures beyond the antigorite stability field in hydrous ultrabasic systems (Engi and Lindsley 1980; Trommsdorff and Evans 1980). This is possibly due to the effect of the OHF<sub>-1</sub> substitution that reduces the repulsion of H-H bonds within the brucite



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<sup>&</sup>lt;sup>1</sup>Laboratoire Magmas et Volcans, Clermont Université, Université Blaise Pascal, Clermont-Ferrand 63038, France

<sup>&</sup>lt;sup>2</sup>Laboratoire Magmas et Volcans, CNRS UMR 6524, Clermont-Ferrand 63038, France

layers of the clinohumite crystal structure (Fujino and Takeuchi 1978; Ribbe 1979). Because of its potential role as a repository phase for H, F, and Ti in ultramafic rocks, characterizations of naturally occurring TiChu can provide greater insight into fluid transport in subduction zones.

TiChu is one of the trace minerals in ultramafic lithologies that is often overlooked because its submicroscopic intergrowth with olivine can easily be missed during electron microprobe analysis and its geochemical significance has only recently been discovered. We present a Fourier transform infrared (FTIR) transmittance of TiChu with variable OH/F ratios from high-pressure hydrous ultramafic rocks. Previous spectroscopic studies of natural clinohumite are scarce (Miller et al. 1987; Matsyuk and Langer 2004; Berry et al. 2005; Hermann et al. 2007) and have not explored the potential effect of the compositional variation in this mineral in detail. We found that the FTIR spectrum of natural clinohumite in the range of the OH absorbance bands (3,800 to 3,200  $\text{cm}^{-1}$ ) systematically changes with the OH/F ratio. Combined with wavelength-dispersive X-ray spectroscopy (WDS) electron microprobe data, we propose a calibration that quantitatively correlates the absorbance spectrum with the OH/F ratio for this mineral. This calibration gives an alternative assessment of OH-F substitutions in TiChu by FTIR, which is probably most applicable to inclusions in minerals and fine grains in experimental changes.

#### Sample location and petrography

We analyzed a wide variety of petrographic types of TiChu from the Cerro del Almirez ultramafic body, Spain. López Sánchez-Vizcaíno et al. (2005) provided a detailed account of the petrology of TiChu in this massif. Our samples can be divided into: TiChu associated with antigorite serpentinite (atg-serpentinite), chlorite harzburgite (chl-harzburgite), and enstatite-rich veins within chl-harzburgite. TiChu associated with atg-serpentinite is almost fluorine free and occurs as dispersed grains embedded in a matrix of flaky antigorite (type serp-1; Al96-20), as veins with millimeter- to centimeter-sized agglomerations of TiChu and olivine (type serp-2; Al96-15B), and as rare small grains formed together with chlorite after relic Ti-rich clinopyroxene grains (type serp-3; Al96-20). TiChu associated with chl-harzburgite is exclusively F-OH-TiChu and occurs as single grains in the matrix of chlharzburgite minerals (type hz-1; Al96-10), as rare patches, intergrowths, and lamellae within olivine (type hz-2), and as individual grains within enstatite-chlorite veins (type hz-3; Al01-7B). We also studied a third textural type occurring in carbonate-rich lenses near the contact with the country rock. These samples display a 'sunny-side up' texture characterized by yellow cores of F-rich TiChu sequentially rimmed by light yellow and then clear olivine (type 4; Al98-29A).

In addition, we analyzed three fluorine-bearing TiChu samples from a garnet metaperidotite in an ultramafic lens enclosed in kyanite schist from the Cima di Gagnone (Central Alps, Ticino, Switzerland). These samples have already been described by Evans and Trommsdorff (1983) (all are from locality 304; samples MG-304-B). TiChu grains occur here in brown patches within a matrix of olivine, orthopyroxene, and spherical aggregates of chlorite interpreted as breakdown products from garnet. These TiChu samples expand the compositional range of TiChu investigated here to low OH/F and X<sub>Ti</sub> compositions.

#### Analytical methods

#### Electron microprobe analysis

Major elemental compositions of TiChu were measured on the same spots from which FTIR spectra were collected; this was done using a SX-100 CAMECA electron microprobe (EMP) at the Laboratoire Magmas et Volcans (Clermont-Ferrand, France) with a 15-kV accelerating voltage, a 30-nA sample current, and a counting time ranging from 10 to 30 s using thallium acid phthalate (TAP), and pentaerythritol (PET) refraction crystals. Fluorine analyses were performed following the procedure of Moune et al. (2007) but modified to a 40-nA sample current and 30-s acquisition time using three TAP crystals. Fluorine concentrations were determined using CaF<sub>2</sub> calibration standards. The estimated precision of F depended on the concentrations and ranged from 2% to 32% relative deviation in repeated analyses. The detection limit was 300 ppm or lower.

#### FTIR sample preparation and measurements

Micro-FTIR spectroscopy analyses were made on doublesided polished sample slabs prepared on glass slides with Crystalbond<sup>TM</sup> that were subsequently dissolved off with acetone and then rinsed thoroughly with ethanol. For spectral calibration, we measured the slab thickness with a micrometer, and the value had a precision of  $\pm 5 \ \mu m$ .

We collected in situ FTIR spectra using a Bruker™ HYPERION microscope at the Laboratoire Magmas et Volcans, Université Blaise-Pascal (Clermont-Ferrand, France). The FTIR spectra were acquired using the Video Assisted Measure option of the Bruker<sup>™</sup> OPUS software (v. 5.0). For each spot, we acquired one spectrum in unpolarized infrared (IR) light and two spectra in polarized IR light taken at orthogonal orientations corresponding to maximum and minimum pleochroism under polarized light. The IR polarizer was gold-coated thallium bromide iodide, and it had 87% polarization efficiency. The IR source was a mid-IR lamp, and a KBr beam splitter was used. Spectra were collected with a × 15 IR objective and a nitrogen-cooled mercury cadmium telluride (MCT) detector, and the values were calculated over the range of 4,000 to 700 cm<sup>-1</sup> at 4 cm<sup>-1</sup> resolution

#### Raman spectroscopy

Laser Raman measurements on the same samples were acquired at the Ecole Normale Supérieur (ENS), in Lyon, France using a Jobin-Yvon LabRam HR800 spectrometer equipped with a charge-coupled device (CCD) detector. An argon ion laser from Spectra Physics, which was set at a wavelength of 632.8 nm, was used to excite Raman scattering. Final spectra were obtained by averaging three spectra with an acquisition time of 30 s for shift ranges of 500 to 2,000 cm<sup>-1</sup> (aluminosilicate framework domain, TOT) and 3,200 to 3,800 cm<sup>-1</sup> (OH stretching domain) relative to the laser light wavelength.

#### Results

#### Compositional variations of titanian clinohumite

Table 1 shows the composition of TiChu samples used in this study. An averaged value for each sample is reported although not every crystal in a sample is necessarily homogenous. The degree of heterogeneity was inferred from the standard deviations. For example, because the analytical uncertainty of FeO is approximately  $\pm 0.5$  wt.%, the standard deviation of a homogeneous sample should give roughly the same value. Thus, larger variations of MgO, FeO, and TiO<sub>2</sub> in samples Al96-20 and Al98-29A were due to the compositional heterogeneity. The analytical spots for FTIR, Raman, and the electron microprobe were all at the same coordinates and identified by the sample textures. The correlations of FTIR spectra and TiChu compositions were established from the analyses for each point rather than averaged sample compositions.

Figure 1 shows notable compositional variation due to exchange vectors. The mole fractions of F and Ti in their respective crystallographic sites formed a negative correlation. This has been described elsewhere as being the result of the exchange vectors of OH-F and  $TiO_2$ -M(OH)<sub>2</sub> (Evans and Trommsdorff 1983). Our samples showed a coherent variation that corresponds to these exchange vectors. Consequently, F-rich TiChu is OH and Ti poor, and vice versa.

#### FTIR spectra

Figure 2 shows a series of unpolarized Ti-clinohumite spectra. Here, we focus on the spectrum wavenumbers

Table 1 Titanian clinohumite from Cerro del Almirez, Spain (Al) and Cima di Gagnone, Switzerland (MG)

	Sample: rock type													
	Al96-20: Ant serp		Al96-15B: TiCl vein Serp		Al96-10: Px Iamella		Al96-1C: Chl harz		Al01-7B: En vein		Al98-29A: Marble		MG-304B: Chl perid	
SiO <sub>2</sub>	35.8	(9)	36.4	(5)	36.8	(1)	36.2	(2)	36.9	(4)	36.8	(4)	36.2	(4)
TiO <sub>2</sub>	5.4	(1)	5.3	(2)	5.6	(2)	4.9	(1)	4.8	(1)	2.4	(1.1)	1.9	(6)
FeO	7.1	(2)	8.7	(5)	8.1	(6)	8.3	(1)	8.9	(2)	9.1	(8)	8.4	(3)
MnO	0.36	(4)	0.32	(3)	0.39	(3)	0.13	(2)	0.10	(5)	0.10	(3)	0.08	(4)
MgO	47.8	(1.2)	47.3	(5)	47.6	(4)	47.7	(2)	48.1	(3)	49.4	(1.0)	49.8	(4)
NiO	0.14	(3)	0.16	(1)	0.15	(2)	0.25	(1)	0.19	(1)	0.19	(1)	0.14	(2)
F	0.02	(2) <sup>c</sup>	0.02	(2) <sup>c</sup>	0.014	(4) <sup>d</sup>	0.49	(5)	0.62	(4)	1.58	(50)	1.78	(21)
$H_2O^a$	2.71	(26)	2.73	(2)	2.74	(2)	2.50	(2)	2.47	(3)	2.02	(23)	1.91	(10)
Total	99.6	(8)	101.0	(5)	101.5	(3)	100.8	(4)	102.1	(6)	101.6	(3)	100.4	(5)
n	4		16		5		5		4		8		5	
Ti	0.45	(19)	0.43	(2)	0.46	(1)	0.41	(1)	0.39	(1)	0.20	(9)	0.16	(5)
Fe	0.66	(7)	0.80	(4)	0.74	(6)	0.76	(1)	0.81	(1)	0.82	(7)	0.76	(3)
Mn	0.03	(2)	0.03	(0)	0.04	(0)	0.01	(0)	0.01	(0)	0.01	(0)	0.01	(0)
Mg	7.87	(64)	7.73	(6)	7.74	(5)	7.80	(1)	7.77	(1)	7.97	(12)	8.10	(5)
Ni	0.01	(2)	0.01	(0)	0.01	(0)	0.02	(0)	0.02	(0)	0.02	(0)	0.01	(0)
F	0.005	(10)	0.006	(10)	0.009	(10)	0.17	(2)	0.21	(1)	0.54	(17)	0.61	(7)
$OH^b$	1.10	(32)	1.12	(3)	1.07	(3)	1.02	(3)	1.01	(2)	1.07	(5)	1.08	(3)

TiCl compositions are an average of several electron probe analyses on a rock sample shown as weight percent, and the number of points used for the average is indicated as 'n' large standard deviations of the average represent the extent of sample heterogeneity, as indicated by italics. Numbers in parentheses indicate the uncertainty expressed as the last digits of reported values, as in '1.07 (5)' corresponds to 1.07  $\pm$  0.05. <sup>a</sup>H<sub>2</sub>O weight percent was calculated from the stoichiometric balance. <sup>c</sup>Three low F abundance measurements were below the detection limit and the relative standard deviations were more than 100% hence, these values should be considered as maximum possible values. <sup>c</sup>The value verified by a secondary ion mass spectrometer.





between 3,700 and 3,250 cm<sup>-1</sup> that correspond to the absorbance bands of hydroxyl groups. Spectra were baseline-corrected and normalized to 1-cm thickness. The shape of the spectra and the relative strength of the bands varied significantly for F-rich and F-poor TiChu (Figure 2). The spectra of all investigated F-poor TiChu are characterized by two weak bands at 3,565 and 3,528 cm<sup>-1</sup> of similar strength (typically between 2 to 7 absorbance units), and a stronger, wider band between 3,415 and 3,394 cm<sup>-1</sup> with a shoulder at lower wavenumbers. For example, the strong band at 3,390 cm<sup>-1</sup> was present in all spectra from the F-rich to F-poor clinohumites. This spectrum is similar to those described



elsewhere (Kitamura et al. 1987; Miller et al. 1987; Kurosawa et al. 1997; Berry et al. 2005; Matsyuk and Langer 2004). We did not observe the broad 3,840 cm<sup>-1</sup> band reportedly present in a high F clinohumite (Satish-Kumar and Niimi 1998).

The shape of the spectra was largely orientation independent (Figure 3). The similarity of the shape implies constant relative peak heights, although there may be significant changes in absorption. However, closer inspection of the relative heights of the stronger bands at 3,415 and 3,394 cm<sup>-1</sup> showed that they were dependent on crystal orientation as there were more prominent shoulders in stronger spectra. On the contrary, the wavenumbers for the weaker bands at 3,565 and 3,528 cm<sup>-1</sup> remained constant, but their relative height varied slightly with orientation.

The spectrum of F-rich TiChu was characterized by the 3,565 cm<sup>-1</sup> band being more intense than the 3,528 cm<sup>-1</sup> band. The 3,528 cm<sup>-1</sup> band was weaker than that of the 3,565 cm<sup>-1</sup> band for F-poor specimens, and the exact peak position was obscured by the larger 3,565 cm<sup>-1</sup> band (Figure 2, sample MG304B). In this case, deconvolution of the spectra from 3,570 to 3,520 cm<sup>-1</sup> identified the presence of the 3,528 cm<sup>-1</sup> band. These results indicate that the shape of TiChu spectra depends on its composition, and the intensity ratio of the 3,565 to 3,528 cm<sup>-1</sup> peaks appears to be sensitive to the F/OH ratio (Figure 2).

#### Laser micro-Raman spectroscopy

Figure 4 shows the laser micro-Raman spectra for TiChu with variable F content. The spectra were similar for all samples and displayed Raman bands characteristic of TiChu in the aluminosilicate framework domain (500 to  $1,100 \text{ cm}^{-1}$ , Figure 4a; Frost et al. 2007). However, we



302 755 780 328 345

963

361

536

(a)



did not detect many of the reported weak bands, which was probably a result of the weak excitation laser used in this study. In particular, our Raman observation in the OH band region differed significantly from that of Frost et al. (2007), who described several peaks of Raman shift in the region of 3,527 to 3,578 cm<sup>-1</sup>, whereas we detected a weak band at 3,565 cm<sup>-1</sup>. The IR absorption peaks corresponded to all the reported OH Raman bands. We did not detect a clear variation in the locations of low-frequency Raman bands among clinohumites with variable OH/F ratios (Figure 4a). All clinohumites showed a band at 3,412 to 3,398 cm<sup>-1</sup> with variable strength in the OH stretching domain (3,200 to 3,750 cm<sup>-1</sup>, Figure 4b). The strength of the band

is probably related to the OH/F ratio and orientation of the sample. The peak was well developed in the F-poor samples; F-rich samples lost their 3,200 to 3,750 cm<sup>-1</sup> Raman signal, which is consistent with the lower OH abundance.

#### Discussion

#### Orientation-independent spectral shapes

For crystalline material (except for that with cubic symmetry), polarized spectra normally vary in intensity depending on the orientation of the crystal with respect to the photon vibration vector. This type of behavior is also expected for TiChu, which is orthorhombic. However, the polarized IR spectra of orthogonal orientations, which were collected on randomly oriented Ti-clinohumite crystals, showed a general similarity in shape. Specifically, it should be noted that relative heights of some peaks were constant while absolute intensities of spectra varied significantly from one to the other. Polarized and unpolarized spectra also showed similar shapes (Figure 3). However, there were minor deviations from this generalization; for example, there were significant variations in shoulder peak intensities at 3,415 and 3,400 cm<sup>-1</sup> depending on the crystal orientation.

We examined the similarity of the spectra shapes by plotting the ratios of peak intensities at 3,565, 3,528, and 3,390 cm<sup>-1</sup>. If the spectra obtained for two orthogonal orientations are symmetric, then the plot of the intensity ratios for one orientation against the other should result in a perfect correlation. Figure 5 shows the intensity ratios of (a) the 3,565 and 3,390 cm<sup>-1</sup> bands and (b) the 3,565 and 3,628 cm<sup>-1</sup> bands. The good correlation (Figure 5a) demonstrates that the intensities of the 3,565 and 3,390 cm<sup>-1</sup> bands vary proportionally with respect to the orientation. The correlation improved for an unpolarized spectra against one of the polarized orientations, which was probably due to the averaging effect by the unpolarized photon vector.

Infrared light absorption in the range of 3,200 to 3,700 normally represents the interaction of photons with OH stretching vibrations. Therefore, the variation in IR absorption depends on the angle between the photon vibration vector and the stretching vibration vector. If the orientations of the crystallographically distinct OH bonds corresponding to the 3,565 and 3,390 cm<sup>-1</sup> bands are parallel in the TiChu crystal, there should be a perfect correlation between the intensity ratios. However, Figure 5a shows moderate scatter around a general correlation, which suggests that these OH bond orientations deviate from the parallel. This deviation from the parallel was not significant enough to invalidate the correlation of Figure 5a. It should also be noted that the correlations of peak intensity ratios for other vibration bands were weak or nonexistent (Figure 5b).



#### Correlation of IR spectra and chemical composition

Because each IR absorption band shown here represents an OH bond in a certain crystallographic site, inherent relationships can be expected between the shape of the IR spectra and the chemical composition, especially in the case of OH-F exchange in TiChu. A detailed study of oriented clinohumites should enable the exact nature of OH-F substitutions to be identified, as well as the site preferences for these elements. The linear relationship between the intensity ratio of 3,565/3,390 and F abundance (Figure 6) implies a near parallel OH bond orientation. Furthermore, because the intensity ratio of 3,565/3,390 was more or less independent of the orientation, the ratio can be used to quantitatively calibrate the correlation between the shape of the IR spectrum and the F abundance.

The spectroscopy results show that the shape and strength of the FTIR spectrum of TiChu is sensitive to the  $X_F$  value ( $X_F = F/(F + OH)$  in atoms per formula unit). A robust correlation was present between F wt.% and the unpolarized intensity ratio of 3,565/3,390 bands



 $(R_{3,565/3,390})$ ; Figure 6). Using this correlation, it is possible to estimate the approximate composition of clinohumite from FTIR spectra. The weighted regression in Figure 6 gives the following formula:

$$F[wt.\%] = -0.36 + 2.3 R_{3.565/3.390}$$
, where  $r^2 = 0.938$ .

The uncertainties of the regression are not reported here because the intention of this study was to demonstrate the presence of a composition-spectra correlation. This formula is useful in estimating F content upon detection of clinohumite by spectroscopic methods. It is clear that a calibration based on Raman spectra would be more versatile in natural sample applications.

Application of the correlation to complex natural situations Fluorine abundance in solid materials can often be measured using an electron microprobe, as in the case of the large TiChu crystals in this study, but accurate F measurements by an electron microprobe are not routine and require some developments (Ottolini et al. 2000). Estimation of the chemical composition from IR spectra is probably a useful alternative for cases such as complex intergrowths of TiChu with olivine. Clinohumite crystals in olivine can occur as lamellae at scales ranging from several hundred nanometers to several tens of micrometers (Wirth et al. 2001; Mosenfelder et al. 2006b; Berry et al. 2007). Because a higher F abundance stabilizes TiChu at higher temperatures (Evans and Trommsdorff 1983), the identification of such lamellae in olivines can provide new information about the volatile element sink that is transported deep into the mantle.

Published FTIR spectra of F-rich TiChu from Dabie Shan, China (Hermann et al. 2007) and those obtained here for F-rich TiChu from Cima di Gagnone (Figure 2) are similar. Accordingly, FTIR spectra may allow us to infer that relatively F-poor TiChu from the Almirez chl-harzburgite formed at lower temperatures than F-rich ones from Dabie Shan or Cima di Gagnone. This grossly simplified conclusion can be made despite ignoring initial F abundances and activity in the respective proto-liths, and it is supported by independent P-T estimations made in each of these localities.

Among recent reports on the hydrogen incorporation mechanisms in olivine, the presence of clinohumite-type OH defects has been proposed as one of the critical natural incorporation mechanisms (Miller et al. 1987; Berry et al. 2005; Mosenfelder et al. 2006b; Berry et al. 2007; Walker et al. 2007). Our IR characterization should not be applied to the investigation of olivine point defect structures because the reported IR spectra in olivine are suggestive of dissimilarities between TiChu structures and TiChu-like defects. For instance, forsterites synthesized under an MgO and TiO<sub>2</sub> buffer (Berry et al. 2005) show stronger IR absorption peaks at 3,572 and 3,525 than around 3,415 to 3,390, whereas our results consistently showed stronger peaks around the 3,415 to 3,390 region than either of the peaks at 3,572 or 3,525. In Berry et al. (2005), the 3,572 and 3,525 peaks are attributed to TiChu-like OH-incorporating defects. We note that the TiChu OH stretching peaks are at 3,565 and 3,528, offsets of -7 and +3 cm<sup>-1</sup> from that of the OH defect in olivine. Looking at the spectra qualitatively (Figure 1 in Berry et al. 2005), the defect peaks are similar to the clinohumite peaks, but these offsets are significantly beyond the uncertainty of the instrument. These disparities point to the fact that the casually termed 'titanian clinohumite-like defect' is structurally dissimilar to TiChu minerals apart from the fact that both have Ti in the sixfold coordinate site. Mosenfelder et al. (2006b) presents a similar conclusion based on the inconsistency of peak intensities found in experimental and natural olivine crystals with TiChu-like defects.

#### **Competing interests**

The authors declare that they have no competing interests.

#### Authors' contributions

The manuscript was written by KTK and verified for accuracy by all authors. CJG, JAPN, and VLSV collected the samples and provided all petrological information. Data acquisitions were done by KTK and CJG. MTGP provided crystallographic information of the sample. All authors read and approved the final manuscript.

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#### Author details

<sup>1</sup>Laboratoire Magmas et Volcans, Clermont Université, Université Blaise Pascal, Clermont-Ferrand 63038, France. <sup>2</sup>Laboratoire Magmas et Volcans, CNRS UMR 6524, Clermont-Ferrand 63038, France. <sup>3</sup>Laboratoire Magmas et Volcans, IRD M163, Clermont-Ferrand 63038, France. <sup>4</sup>Instituto Andaluz de Ciencias de la Tierra (IACT), CSIC-UGR, Armilla, Granada 18100, Spain. <sup>5</sup>Géosciences Montpellier, Université Montpellier 2 & CNRS, Montpellier 34095, France. <sup>6</sup>Department Geología, Universidad de Jaén, Escuela Politécnica Superior, Linares 23700, Spain. <sup>7</sup>Department Mineralogia y Petrología, Facultad de Ciencias, Universidad de Granada, Granada 18002, Spain.

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