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▶ To cite this version:

Hubert Luce, Richard Wilson, Francis Dalaudier, Fanny Truchi, Hiroyuki Hashiguchi, et al.. Estimating length scales for tropospheric turbulence from MU radar and balloon data. 14th International Workshop on Technical and Scientific Aspects of MST14/IMST1, May 2014, Sao José dos Campos, Brazil. hal-01108891

HAL Id: hal-01108891 https://hal.science/hal-01108891v1

Submitted on 23 Jan 2015 $\,$

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Questions or comments? Corresponding authors: hubert.luce@univ-tln.fr richard.wilson@upmc.fr

京都大学生存圈研究所 Research Institute for Sustainable e Humanosohere

Estimating length scales for tropospheric turbulence from MU radar and balloon data

Hubert Luce, Université de Toulon, La Garde, France Richard Wilson, F. Dalaudier, F. Truchi LATMOS-IPSL, UPMC Univ Paris 06, Univ. Versailles St-Ouentin, CNRS/INSU, UMR 8190, Paris, France Hirdyuki Hashiguchi, Masayuki K. Yamamoto, Mamoru Yamamoto Research Institute for Sustainathe Humanosphere, Kyoto University, Uji, Japan. Lakshmi Kantha Aerospace Engineering Sciences, University of Colorado, Boulder, USA

UPMC

This work is dedicated to the memory of Prof. Shoichiro Fukao,

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Estimating atmospheric turbulence parameters from ST radar measurements is an important issue. The methods used in the Iterature are based on the hypothesis that radar exchose at oblique incidence result from isotropic turbulence within the inertial subrange (e.g. Atsatróm and Eaton, 2005). In such a case, the width of Doppier spectrum can be related to the turbulent energy dissipation rate c.¹ Whom the outer scale of turbulence is smaller than the dimensions of the radar volume, e depends on the background stability *NP* (e.g., Hocking, EPS, 1999). The latter parameter is usually settimated (rom standard halloom exeruments).

Introduction

than the dimensions of the radar volume, c depends on the background stability AP (e.g., Hocking, EPS, 1999). The latter parameter is usually estimated from standard balloon measurements. Our studies aim at estimating truthulence parameters by taking advantage of the high time resolution of radiosondes (1 sec) and concurrent radar observations. Wilson et al. (JAOT, 2011) showed that detecting turbulence from PTU measurements is possible by using Thorpe analysis (Thorpe, DSR, 1977) despite the instrumental measurement noise. The method was applied by Clayson and Kantan (JAOT, 2005) but which considering the noise effects. Turbulent layers can be directly identified through the detection of private produced by mixing and billows in or most potential temperature profiles. The instrument noise can lead to the detection of private sites specially in weakly strailider deposition. Wilson et al. (JAOT, 2010) proposed objective selection criteria and procedures for minimizing noise effects. Wilson et al. (AMT, 2013) considered the air saturation effects.

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Instrumental set-up

The Middle and Upper Atmosphere radar (MUR) was operated in range imaging (FII) mode (5 frequencies) from 01 November 2013 16:00 LT to 09 November 2013 06:00 LT at a time resolution of 6.14 s (e.g. Luce et al. 2006). The radar beam was steered into 3 directions (vertical, and 10° off zenith toward North and East). The range sampling was performed from 1.32 km to 20.37 km. The FII mode was used for imaging the turbulent and stable layers. The Doppler spectra and moments were estimated at the range resolution of 150 m.





Meisei RS-06 GPS Vaisala RS92G

Among the other instruments operated at the same time, two types of radiosondes (Vaisala R592G and Maisel R596 GPS) were used. The positionation of the Maisir andiosondes are given at http://www.mbisico.ip/signithyproductsmeteor/motion_goc_motionode.html Beh radiocondes provide PTU and wind) profiles at a vertical sampling of 1 sec. A total of 36 radiocondes were launched auring pint predicts every-01 hour 30 min (eccept on 01-02 No during which Melesi and Vaisala radiocondes were launched aurinanously three mines). Launching times are indicated by the red dots in Figures 1a and b. The PTU profiles collected from both radiosondes were processed for applying the Thorpe analysis



07/11 09:00 Time (LT) is 1(a)-1(b) Heigh-time cross-section of vertical MUR echo power in range imaging mode for all the campaign. The passage of frontal zo early visible. The smoothed appearance of the radar echoes in clouds and precipitations strongly differ with the appearance of the ed layers in dry air. The detailed analysis of the synoptic conditions will be performed in a subsequent work.

06/11 03:00



Figure 2: Close up of Figure (1a) between 01-NOV 20:00 LT and 02-NOV 05:20 LT. The nearly straight lines show the balloon altitude vs time (net: Vaisala, Black Meise), VQ2M03, VVMM05 and V06M07 were launched almost simultaneously but separately. The groffles are the most potential temperature profiles used for Thorps sorting. They were estimated from the numerical integration of dry N^a when relative humidity was below the saturation thresholds defined by Zhang et al. (2010) for Vaisala sondes (and also used for Mese's sondes here) or most N (Lais and Einaud, JAS, 1974) when above. Also show are the vertical profiles of N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a algo effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to N^a effet teclotion of 50 m. (The vertical dashed line correspond to teclotion of 50 m. (The vertical dashed li



Figure 3: Close up of Figure 1(a) between 02-NOV 20:00 LT and 03-NOV 05:05 LT. The black curves show the mole 1 M² and the vertical dashed lines show N⁻²⁰. Despregions of weak stability can be seen in the height range 06:15 L5 M and concide with a smooth appearament of the range are chose are not aspect sensitive (see Figure 6). Their structure contracts with the thin layering balow 6.0 km and above 11.5 km where profiles of M² show a succession of narrow peaks. A frontal auralos (strong pearare) thurburs the thin layering balow 6.0 km and above 11.5 km where profiles of M² show a succession of narrow peaks. A frontal auralos (strong pearare) the thin layering balow 6.0 km and above 11.5 km indicated by the dotted blue line (and coincide with the strong peak of N². Despregions of the sense no both sides of the cloud base and within the cloud. The nearly vertical straintors of the exchanse underneath cloud can be the signature of KH billows or convective rolls as already reported from MUR observations in FII mode for such atmospheric conditions (e.g. Luce et al., MWR, 2010).



Figure 4: Time-height cross-section of variance of Doppler spectra measured at vertical incidence corrected from the non turbulent effects between 02-NOV 20-20. LT and 02-NOV 05:00 LT. For the purpose of the present work, each Doppler spectrum was inspected and any spurious or supplicious peak (dutter, rain, interferences..., value, as for as possible, manually removed. The black curves show the most Richardson number RV. The deep echologi layers identified in Figure 3 are clearly associated with enhancements of variance on both sites of the frontile surface and reverywhere within the cloud (with vinable levels). Except at some locations in the cloud, RV. Rvisible layers layer to a clearly associated the radiosondes and MUR observed the same turbulent events.





ted by eddies in a turbulent layer (Smyth et al. JPO, 2001) $N^2 = \frac{g}{\overline{\theta} * g} \frac{\theta_{max}}{L_T}$



Figure 6 (a) Preliminary comparisons between buoyancy and Thorpe lengths for 18 selected cases from balloon and MUR data collected in 2011 (black dots) and 2012 (red dots) using similar radar configurations. (b) Systematic comparison for the layers execeeding 100 m in depth and detected above 5 k m() just blevan within the cloud). The blue dots corresponds to the deepset layers shown in Figure 6 and described in Table 1. The ratio $L_{T}L_{a}$ is less than a factor 2 to 3 and correlation coefficient is 0.56. A linear regression indicates LT –LB. (p-1)



 $\varepsilon_k \propto \frac{\sigma_{turb}^3}{I} = a \sigma_{turb}^2 N$ (e.g. Fukao et al. JGR, 1994)

 $L_{o} = \sqrt{\varepsilon_{k}/N^{3}}$

Layer	z (m)	TPE (mJ.kg ⁻¹)	TKE (mJ.kg ⁻¹)	(mW.kg^{-1})	$(mW.kg^{-1})$	L _T (m)	L _B (m)	(n
LB1	6.65 ±0.3	238-344	971	1.34	0.70	230	296	17
LAI	8.14 ± 0.29	604-772	618	0.47	1.15	259	184	9
LTI	10.67 ± 0.69	3302-3229	947	0.57	5.19	383	128	7
L82	6.45 ± 0.47	527-615	931	1.33	1.03	419	359	21
LA2	7.98±0.46	551-551	301	0.17	0.56	397	248	11
LT2	9.72 ± 0.21	118-180	454	0.22	0.28	125	184	9
183	6.22 ± 0.24	73-94	800	1.05	0.36	93	204	11
LAD	7.29 ± 0.42	925-931	725	0.72	2.24	228	137	7
184	6.82 ± 0.24	175-234	341	1.02	0.62	153	221	12
LM	7.65 ± 0.36	271-428	341	0.22	0.46	212	195	9
LAS	7.76±0.3	791-820	323	0.2	1.23	242	127	5
L06	6.51 ± 0.3	200-400	618	0.69	0.481	224	280	15
LAS	7.78 ± 0.4	714-1168	549	0.46	1.09	360	232	12
1.16	10.35 ± 0.62	710-912	939	0.65	1.05	450	330	- 19

Table 1: Energetic parameters estimated for the deepest layers (labeled in Figure 6) below cloud (LB), above cloud (LA) and inside cloud or near cloud top (LT). The values for cloudy, ar are emphasized by gray rectangles. Note that kinetic energy dissipation rates inside clouds are always smaller than values underneath clouds.

Conclusions

The purpose of our investigations is to infer key parameters of atmospheric turbulence from the combination of (MU) radar and balloon data. Our approach is first based on the direct identification of turbulent layers in potential temperature profiles from Thorpe analysis with careful selection procedures. Results from two types of radiosendes (Vaisala and Neise) were obtained with similar performances. The present studies extend the results described by Wilson et al. (LASTP 2014) obtained for a few cases only. -For the first time, a statistical comparison between the Thorpe length and the buoyancy scale is presented for turbulent events observed from six consecutive balloons near an upper level front and within cloud could be made and for selected cases from previous campaigns. A clear relationship between L₁ and L₄ was found (i.e. L₆ -1, despite an important dispersion to be more thoroughly interpreted, see Figure 6b). It is consistent with studies conducted for oceanic turbulence. This fundamental result may have dramatic impacts on the characterization of turbulence in the troposphere and may lustly the estimation of energetic parameters from standard balloon data alone (Clayson and Kantha, JAOT 2006). -Estimation of energy dissipation rates in clear air and upper level clouds could also be obtained. Preliminary results show an overall consistency of the inferred turbulence parameters.