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Diffusion of energy in chains of oscillators with bulk noise

Cédric Bernardin¹

Abstract: These notes are based on a mini-course given during the conference Particle systems and PDE's - II which held at the Center of Mathematics of the University of Minho in December 2013. We discuss the problem of normal and anomalous diffusion of energy in systems of coupled oscillators perturbed by a stochastic noise conserving energy.

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The goal of statistical mechanics is to elucidate the relation between the microscopic world and the macroscopic world. *Equilibrium statistical mechanics* assume the microscopic systems studied to be in equilibrium. In this course we will be concerned with *non-equilibrium statistical mechanics* where time evolution is taken into account: our interest will not only be in the relation between the microscopic and the macroscopic scale in space but also in time.

By microscopic system we refer to molecules or atoms governed by the classical Newton's equations of motion. The question is then to understand how do these particles manage to organize themselves in such a way as to form a coherent structure on a large scale. The "structure" will be described by few variables (temperature, pressure ...) governed by autonomous equations (Euler's equations, Navier-Stokes's equation, heat equation ...). The microscopic specificities of the system will appear on this scale only through the thermodynamics (equation of state) and through the transport coefficients. Unfortunately, we are very far from understanding how to derive such macroscopic equations for physical relevant interactions.

One of the main ingredients that we need to obtain the macroscopic laws is that the particles, which evolve deterministically, have a behavior that one can consider almost as being random. The reason for this is that the dynamical system considered is expected to have a very sensitive dependence on the initial conditions and therefore is chaotic. This "deterministic chaos" is a poorly understood subject for systems with many degrees of freedom and even a precise consensual formulation is missing.

A first simplification to attack these problems consists in replacing the deterministic evolution of particles *ab initio* by purely stochastic evolutions. Despite this simplification we notice that the derivation of the macroscopic evolution laws is far from being trivial. For example, we do not have any derivation of a system of conservation laws from a stochastic microscopic system after shocks. Nevertheless, since the pioneering work of Varadhan and Yau, important progresses have been performed in several well understood situations by the development of robust probabilistic and analytical methods ([36]).

In this course we will be mainly (but not only) interested in hybrid models for which the time evolution is governed by a combination of deterministic and stochas-

tic dynamics. These systems have the advantage to be mathematically tractable and conserve some aspects of the underlying deterministic evolution. The stochastic noise has to be chosen in order to not destroy the main features of the Hamiltonian system that we perturb.

The central macroscopic equation of these lecture notes is the heat equation:

$$\begin{cases} \partial_t u = \partial_x(D(u)\partial_x u), & x \in \mathring{U}, \quad t > 0, \\ u(0, x) = u_0(x), & x \in U, \\ u(t, x) = b(x), & x \in \partial U, \quad t > 0. \end{cases}$$

Here $u(t, x)$ is a function of the time $t \geq 0$ and the space $x \in U \subset \mathbb{R}^d$, $d \geq 1$, starting from the initial condition u_0 and subject to boundary conditions prescribed by the function b . The advantage of the heat equation with respect to other macroscopic equations such as the Euler or Navier-Stokes equations is that the notion of solution is very well understood. The dream would be to start from a system of $N \gg 1$ particles whose interactions are prescribed by Newton's laws and to show that in the large N limit, the empirical energy converges in the diffusive time scale $t = \tau N^2$ to u (τ is the microscopic time and t the macroscopic time). In fact, this picture is expected to be valid only under suitable conditions and to fail for some low dimensional systems. In the case where the heat equation (or its variants) holds we say that the system has a normal behavior. Otherwise anomalous behavior occurs and the challenging question (even heuristically) is to know by what we shall replace the heat equation and what is the time scale we have to observe the system in order to see this macroscopic behavior.

The course is organized as follows. In Chapter 1 we introduce the models studied. Chapter 2 is concerned with models which have a normal diffusive behavior. In Chapter 3 we are interested in systems producing an anomalous diffusion. An important issue not discussed here is the effect of disorder on diffusion problems. In order to deal with lecture notes of a reasonable size, many of the proofs have been suppressed or only sketchy presented.

Chapter 1

Chains of oscillators

1.1 Chains of oscillators with bulk noise

Chains of coupled oscillators are usual microscopic models of heat conduction in solids. Consider a finite box $\Lambda_N = \{1, \dots, N\}^d \subset \mathbb{Z}^d$, $d \geq 1$, whose boundary $\partial\Lambda_N$ is defined as $\partial\Lambda_N = \{x \notin \Lambda_N; \exists y \in \Lambda_N, |x - y| = 1\}$. Here $|\cdot|$ denotes the Euclidian norm in \mathbb{R}^d and “ \cdot ” the corresponding scalar product. Let us fix a nonnegative symmetric pair interaction potential V and a pinning potential W on \mathbb{R} . The atoms are labeled by $x \in \Lambda_N$. The momentum of atom x is $p_x \in \mathbb{R}$ and its displacement from its equilibrium position ¹ is $q_x \in \mathbb{R}$. The energy \mathcal{E}_x of the atom x is the sum of the kinetic energy, the pinning energy and the interaction energy:

$$\mathcal{E}_x = \frac{|p_x|^2}{2} + W(q_x) + \frac{1}{2} \sum_{\substack{|y-x|=1, \\ y \in \Lambda_N}} V(q_x - q_y). \quad (1.1)$$

The Hamiltonian is given by

$$\mathcal{H}_N = \sum_{x \in \Lambda_N} \mathcal{E}_x + \partial\mathcal{H}_N \quad (1.2)$$

where $\partial\mathcal{H}_N$ is the part of the Hamiltonian corresponding to the boundary conditions which are imposed.

We will consider the following cases:

¹We restrict us to the case where $q_x \in \mathbb{R}^n$ with $n = 1$ because the relevant dimension of the system is the dimension d of the lattice. Most of the results stated in this manuscript can be generalized to the case $n \geq 1$.

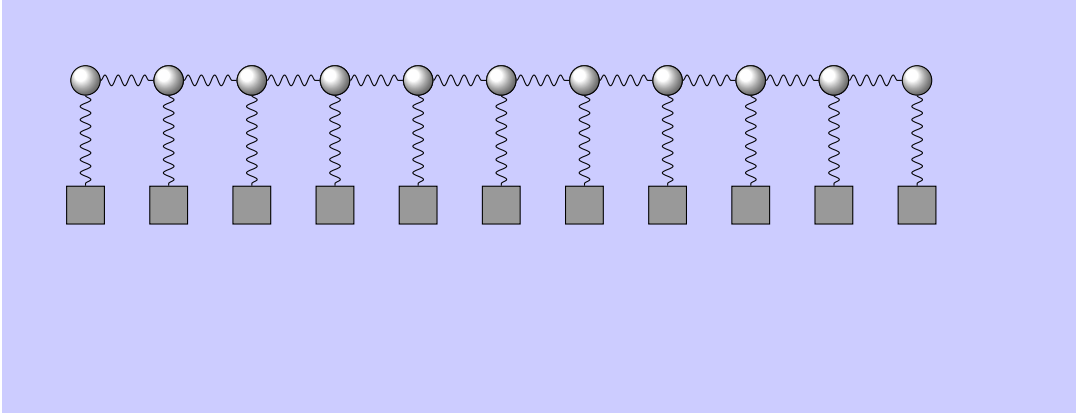


Figure 1.1: A one-dimensional chain of pinned oscillators with free boundary conditions

- Periodic boundary conditions: we identify the site 1 to the site N and denote the corresponding box by \mathbb{T}_N , the discrete torus of length N (then $\partial\mathcal{H}_N = 0$).
- Free boundary conditions: this corresponds to the absence of boundary conditions, i.e. to $\partial\mathcal{H}_N = 0$.
- Fixed boundary conditions: introduce the positions $q_y = 0$, $y \in \partial\Lambda_N$, of some fictive walls. We add to the Hamiltonian \mathcal{H}_N a boundary term $\partial\mathcal{H}_N = \partial^f \mathcal{H}_N$ given by

$$\partial^f \mathcal{H}_N = \sum_{\substack{|y-x|=1, \\ x \in \Lambda_N, y \in \partial\Lambda_N}} V(q_x - q_y).$$

- Forced boundary conditions: site $\mathbf{1} = (1, \dots, 1)$ is in contact with a wall at position $q_0 = 0$ and each site $y \in \partial\Lambda_N \setminus \{0\}$ is driven by a constant force τ_y . This results in a boundary term $\partial\mathcal{H}_N = \partial^\tau \mathcal{H}_N$ given by

$$\partial^\tau \mathcal{H}_N = \sum_{\substack{|y-x|=1, \\ x \in \Lambda_N, y \in \partial\Lambda_N}} V(q_x - q_y) - \sum_{y \in \partial\Lambda_N \setminus \{0\}} \tau_y q_y. \quad (1.3)$$

The equations of motion of the atoms are

$$\dot{q}_x = \partial_{p_x} \mathcal{H}_N, \quad \dot{p}_x = -\partial_{q_x} \mathcal{H}_N \quad (1.4)$$

and the generator \mathcal{A}_N of the system is given by the Liouville operator

$$\mathcal{A}_N = \sum_{x \in \Lambda_N} \{ \partial_{p_x} \mathcal{H}_N \partial_{q_x} - \partial_{q_x} \mathcal{H}_N \partial_{p_x} \}.$$

It will be also useful to consider the chain of oscillators in infinite volume, i.e. replacing Λ_N by \mathbb{Z}^d , $d \geq 1$, in the definitions above. The formal generator \mathcal{A}_N is then denoted by \mathcal{A} . The dynamics can be defined for a large set of initial conditions if V and W do not behave too badly ([39], [43], [10]). We define the set Ω as the subset of $\mathbb{R}^{\mathbb{Z}}$ given by

$$\Omega = \bigcap_{\alpha > 0} \left\{ \eta \in \mathbb{R}^{\mathbb{Z}}; \sum_{x \in \mathbb{Z}} e^{-\alpha|x|} \eta_x^2 < +\infty \right\} \quad (1.5)$$

and $\tilde{\Omega} = \Omega \times \Omega$. We equip Ω with its natural product topology and its Borel σ -field and $\tilde{\Omega}$ by the corresponding product topology. For $X = \Omega$ or $X = \tilde{\Omega}$, the set of Borel probability measures on X is denoted by $\mathcal{P}(X)$. A function $f : X \rightarrow \mathbb{R}$ is said to be *local* if it depends of η only through the coordinates $\{\eta_x; x \in \Lambda_f\}$, Λ_f being a finite box of \mathbb{Z} . We also introduce the sets $C_0^k(X)$, $k \geq 1$ (resp. $k = 0$), of bounded local functions on X which are differentiable up to order k with bounded partial derivatives (resp. continuous and bounded).

In the rest of the manuscript, apart from specific cases, we will assume that the following conditions hold:

- The potentials V and W have bounded second derivatives and $W'' \geq c > 0$ for a fixed positive constant c . Then the infinite dynamics $(\omega(t))_{t \geq 0}$ can be defined for any initial condition $\omega^0 = (\mathbf{q}^0, \mathbf{p}^0) \in \tilde{\Omega}$. Moreover $\tilde{\Omega}$ is invariant by the dynamics. This defines a semigroup $(P_t)_{t \geq 0}$ on $C_0^0(\tilde{\Omega})$ and the Chapman-Kolmogorov equations

$$(P_t f)(\omega) - f(\omega) = \int_0^t (P_s \mathcal{A} f)(\omega) ds = \int_0^t (\mathcal{A} P_s f)(\omega) ds \quad (1.6)$$

are valid for any $f \in C_0^1(\tilde{\Omega})$.

- The potential $W = 0$ and the interaction potential V has a second derivative uniformly bounded by above and by bellow. It is more convenient to go over the deformation field $\eta_{(x,y)} = q_y - q_x$, $|x - y| = 1$, which by construction is constrained to have zero curl. In $d = 1$ we will denote $\eta_{(x-1,x)} = q_x - q_{x-1}$ by r_x . The dynamics (1.4) can be read as a dynamics for the deformation field and the momenta. Given say q_0 , the scalar field $\mathbf{q} = \{q_x\}_{x \in \mathbb{Z}^d}$ can be reconstructed from η . In the sequel, when $W = 0$, we will use these coordinates

without further mentions. The dynamics for the coordinates $\omega = (\eta, \mathbf{p}) = (\eta_{(x, x+\mathbf{e})}, p_x)_{|\mathbf{e}|=1, x \in \mathbb{Z}}$ can be defined if the initial condition satisfies $\omega^0 \in \tilde{\Omega}$. Moreover the set $\tilde{\Omega}$ is invariant by the dynamics. This defines a semigroup $(P_t)_{t \geq 0}$ on $C_0^0(\tilde{\Omega})$ and the Chapman-Kolmogorov equations

$$(P_t f)(\omega) - f(\omega) = \int_0^t (P_s \mathcal{A} f)(\omega) ds = \int_0^t (\mathcal{A} P_s f)(\omega) ds \quad (1.7)$$

are valid for any $f \in C_0^1(\tilde{\Omega})$.²

Let us first consider the problem related to the characterization of equilibrium states. For simplicity we take the finite volume dynamics with periodic boundary conditions. Then it is easy to see that the system conserves one or two physical quantities depending if the chain is pinned or not. The total energy \mathcal{H}_N is always conserved. If $W = 0$ the system is translation invariant and the total momentum $\sum_x p_x$ is also conserved. Notice that because of the periodic boundary conditions the sum of the deformation field $\sum_x \eta_{(x, x+\mathbf{e}_i)}$ is automatically fixed equal to 0 for any $i = 1, \dots, d$.

Liouville's Theorem implies that the uniform measure λ^N on the manifold Σ^N composed of the configurations with a fixed total energy (and eventually a fixed total momentum) is invariant for the dynamics. The micro canonical ensemble is defined as the probability measure λ^N . The dynamics restricted to Σ^N is not necessarily ergodic. Two examples for which one can show it is not the case are the harmonic lattice (V and W quadratic) and the Toda lattice ($d = 1$, $W = 0$, $V(r) = e^{-r} - 1 + r$) which is a completely integrable system ([51]). In fact what is really needed for our purpose is not the ergodicity of the finite dynamics but of the infinite dynamics. We expect that even if the finite dynamics are never ergodic the fraction of Σ^N corresponding to non ergodic behavior decreases as N increases, and probably disappears as $N = \infty$ (apart from very peculiar cases). Therefore a good notion of ergodicity has to be stated for infinite dynamics. The definition of a conserved quantity is not straightforward in infinite volume (the total energy of the infinite chain is usually equal to $+\infty$). To give a precise definition we will use the notion of space-time invariant probability measures for the infinite dynamics defined above.

The infinite volume Gibbs grand canonical ensembles are such probability measures. They form a set of probability measures indexed by one (pinned chains) or $d+2$ (unpinned chains) parameters and are defined by the so-called Dobrushin-Landford-Ruelle's equations. To avoid a long discussion we just give a formal definition (see e.g. [31] for a detailed study).

²The generator \mathcal{A} has to be written in terms of the deformation field.

- Pinned chains ($W \neq 0$): the infinite volume Gibbs grand canonical ensemble μ_β with inverse temperature $\beta > 0$ is the probability measure on $\tilde{\Omega}$ whose density with respect to the Lebesgue measure is

$$Z^{-1}(\beta) \exp \left(-\beta \sum_{x \in \mathbb{Z}^d} \mathcal{E}_x \right).$$

- Unpinned chains ($W = 0$): the infinite volume Gibbs grand canonical ensemble ³ $\mu_{\beta, \bar{p}, \tau}$ with inverse temperature $\beta > 0$, average momentum $\bar{p} \in \mathbb{R}$ and tension $\tau = \beta^{-1} \lambda \in \mathbb{R}^d$ is the probability measure on $\tilde{\Omega}$ whose density with respect to the Lebesgue measure is

$$Z^{-1}(\beta, \bar{p}, \tau) \exp \left(- \sum_{x \in \mathbb{Z}^d} \{ \beta \mathcal{E}_x + \bar{p} p_x + \sum_{i=1}^d \lambda_i \eta_{(x, x+e_i)} \} \right).$$

Observe that in the one dimensional unpinned case we have simply product measures.

Fix an arbitrary Gibbs grand canonical ensemble μ . A probability measure ν is said to be μ -regular if for any finite box $\Lambda \subset \mathbb{Z}^d$ whose cardinal is denoted by $|\Lambda|$, the relative entropy of $\nu|_\Lambda$ w.r.t. $\mu|_\Lambda$ is bounded above by $C|\Lambda|$ for a constant C independent of Λ . We recall that the relative entropy $H(\nu|\mu)$ of $\nu \in \mathcal{P}(X)$ with respect to $\mu \in \mathcal{P}(X)$, X being a probability space, is defined as

$$H(\nu|\mu) = \sup_{\phi} \left\{ \int \phi d\nu - \log \left(\int e^{\phi} d\mu \right) \right\}, \quad (1.8)$$

where the supremum is carried over all bounded measurable functions ϕ on X .

If for any arbitrary Gibbs grand canonical ensembles μ and μ' , μ is μ' -regular and μ' is μ -regular then ν μ -regular is equivalent to ν μ' -regular and we simply say that ν is regular.

We believe that a good notion of ergodicity for infinite dynamics is the following.

Definition 1.1.1 (Macro-Ergodicity). ⁴ *Assume that all the Gibbs grand canonical ensembles are mutually regular. We say that the dynamics generated by \mathcal{A} is macro-ergodic if and only if the only time-space invariant regular measures ν for \mathcal{A} are mixtures of Gibbs grand canonical ensembles.*

³They are defined with respect to the gradient fields $\eta_{(x,y)}$. It would be more coherent to call them *gradient Gibbs measures*.

⁴The name has been proposed by S. Goldstein.

If the microscopic dynamics is macro-ergodic, then, by using the relative entropy method developed in [44], we can derive the hydrodynamic equations ⁵ in the Euler time scale of the chain before the appearance of the shocks, at least in $d = 1$ ([10]). These limits form a triplet of compressible Euler equations (for energy \mathbf{e} , momentum \mathbf{p} and deformation \mathbf{r}) in the form

$$\begin{cases} \partial_t \mathbf{r} = \partial_q \mathbf{p} \\ \partial_t \mathbf{p} = \partial_q \tau \\ \partial_t \mathbf{e} = \partial_q (\mathbf{p} \tau) \end{cases} \quad (1.9)$$

where the pressure $\tau := \tau(\mathbf{r}, \mathbf{e} - \frac{\mathbf{p}^2}{2})$ is a suitable thermodynamic function depending on the potential V . A highly challenging open question is to extend these results after the shocks. The proof can be adapted to take into account the presence of mechanical boundary conditions ([19]).

We do not claim that the macro-ergodicity is a necessary condition to get Euler equations for purely Hamiltonian systems. We could imagine that weaker or different conditions are sufficient but in the actual state of the art the macro-ergodicity is a clear and simple mathematical statement of what we could require to deterministic systems in order to derive rigorously Euler equations. We refer the interested reader to [20] and [49] for general discussions about related problems in statistical mechanics.

1.1.1 Conserving noises

In [29], Fritz, Funaki and Lebowitz prove a weak form of macro-ergodicity for a chain of oscillators.

Theorem 1.1.2 ([29]). ⁶ *Consider the pinned chain $W \neq 0$ generated by \mathcal{A} or an unpinned chain $W = 0$ in $d = 1$. The only regular time and space invariant measures for \mathcal{A} which are such that conditionally to the positions configuration $\mathbf{q} := \{q_x; x \in \mathbb{Z}^d\}$ the law of the momenta $\mathbf{p} := \{p_x; x \in \mathbb{Z}^d\}$ is exchangeable are given by mixtures of Gibbs grand canonical ensembles.*

They also proposed to perturb the dynamics by a stochastic noise consisting to exchange at random exponential times, independently for each pair of nearest neighbors site $x, y \in \mathbb{Z}^d$, $|x - y| = 1$, the momenta p_x and p_y . The formal generator \mathcal{L} of this dynamics, that we will call the *stochastic energy-momentum conserving*

⁵The notion of hydrodynamic limits is detailed in Section 2.2.1 and Section 3.2.2.

⁶The proof given in [29] assume $W \neq 0$ but it can be adapted to the unpinned one dimensional case (see [10]). It would be interesting to extend this theorem to the general unpinned case.

dynamics, is given by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$, $\gamma > 0$, where \mathcal{A} is the Liouville operator and \mathcal{S} is defined for any local function $f : \Omega \rightarrow \mathbb{R}$ by

$$(\mathcal{S}f)(\mathbf{q}, \mathbf{p}) = \sum_{\substack{x, y \in \mathbb{Z}^d \\ |x-y|=1}} [f(\mathbf{q}, \mathbf{p}^{x,y}) - f(\mathbf{q}, \mathbf{p})]. \quad (1.10)$$

Here the momenta configuration $\mathbf{p}^{x,y}$ is the configuration obtained from \mathbf{p} by exchanging p_x with p_y . The previous discussion about existence of the dynamics on $\tilde{\Omega}$ for the deterministic case and its relation with its formal generator is also valid for this dynamics and the other dynamics defined in this section.

With some non-trivial entropy estimates we get the following result.

Theorem 1.1.3 ([29]). *Consider the pinned ($W \neq 0$) or the one-dimensional unpinned ($W = 0$) stochastic energy-momentum conserving dynamics. The only regular time and space invariant measures for these dynamics are given by mixtures of Gibbs grand canonical ensembles, i.e. the stochastic energy-momentum conserving dynamics is macro-ergodic.*

Consequently the stochastic energy-momentum conserving dynamics has in the Euler time scale and before the appearance of the shocks the same hydrodynamics (1.9) as the deterministic model. This is because the noise has some macroscopic effects only in the diffusive time scale.

We consider now a different stochastic perturbation. Let us define the flipping operator $\sigma_x : \mathbf{p} \in \Omega \rightarrow \mathbf{p}^x \in \Omega$ where \mathbf{p}^x is the configuration such that $(\mathbf{p}^x)_z = p_z$ for $z \neq x$ and $(\mathbf{p}^x)_x = -p_x$. In [29] is also proved that the only time-space regular stationary measures for the Liouville operator \mathcal{A} such that conditionally to the positions the momenta distribution is invariant by any flipping operator σ_x are mixtures of Gibbs grand canonical ensembles with zero momentum average. Then we consider the dynamics on $\tilde{\Omega}$ generated by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$, $\gamma > 0$, with \mathcal{S} the noise defined by

$$(\mathcal{S}f)(\mathbf{q}, \mathbf{p}) = \frac{1}{2} \sum_{x \in \mathbb{Z}^d} [f(\mathbf{q}, \mathbf{p}^x) - f(\mathbf{q}, \mathbf{p})] \quad (1.11)$$

for any local function $f : \tilde{\Omega} \rightarrow \mathbb{R}$. This dynamics conserves the energy and the deformation of the lattice but destroys all the other conserved quantities. We call this system the *velocity-flip model* (sometimes the *stochastic energy conserving model*).

Theorem 1.1.4 ([29]). *Consider the pinned d -dimensional velocity-flip model or the one-dimensional unpinned velocity-flip model. The only regular time and space invariant measures are given by mixtures of Gibbs grand canonical ensembles. In other words the velocity-flip model is macro-ergodic.*

Notice that the velocity flip-model has trivial hydrodynamics in the Euler time scale: initial profiles do not evolve. This is only in the diffusive scale that an evolution takes place.

1.1.2 NESS of chains of oscillators perturbed by an energy conserving noise

These models can also be considered in a non-equilibrium stationary state (NESS) by letting them in contact with thermal baths at different temperatures and imposing various mechanical boundary conditions. Let us only give some details for the NESS of the one-dimensional velocity-flip model.

Consider a chain of N unpinned oscillators where the particle 1 (resp. N) is subject to a constant force τ_ℓ (resp. τ_r). Moreover we assume that the particle 1 (reps. N) is in contact with a Langevin thermal bath at temperature T_ℓ (resp. T_r). The generator \mathcal{L}_N of the dynamics on the phase space $\Omega_N = \mathbb{R}^{N-1} \times \mathbb{R}^N$ is given by

$$\mathcal{L}_N = \mathcal{A}_N^{\tau_\ell, \tau_r} + \gamma \mathcal{S}_N + \mathcal{B}_{1, T_\ell} + \mathcal{B}_{N, T_r}, \quad \gamma > 0, \quad (1.12)$$

where $\mathcal{A}_N^{\tau_\ell, \tau_r}$ is the Liouville operator, $\mathcal{B}_{j, T}$ the generator of the Langevin bath at temperature T acting on the j -th particle and \mathcal{S}_N the generator of the noise. The strength of noise and thermostats are regulated by γ , γ_ℓ and γ_r . The Liouville operator is defined by

$$\begin{aligned} \mathcal{A}_N^{\tau_\ell, \tau_r} = & \sum_{x=2}^N (p_x - p_{x-1}) \partial_{r_x} + \sum_{x=2}^{N-1} (V'(r_{x+1}) - V'(r_x)) \partial_{p_x} \\ & - (\tau_\ell - V'(r_2)) \partial_{p_1} + (\tau_r - V'(r_N)) \partial_{p_N}. \end{aligned} \quad (1.13)$$

The generators of the thermostats are given by

$$\mathcal{B}_{j, T} = T \partial_{p_j}^2 - p_j \partial_{p_j}. \quad (1.14)$$

The noise corresponds to independent velocity change of sign, i.e.

$$(\mathcal{S}_N f)(\mathbf{r}, \mathbf{p}) = \frac{1}{2} \sum_{x=2}^{N-1} (f(\mathbf{r}, \mathbf{p}^x) - f(\mathbf{r}, \mathbf{p})), \quad f : \Omega_N \rightarrow \mathbb{R}. \quad (1.15)$$

We will also consider the case where the chain has fixed boundary conditions.

Proposition 1.1.5 ([11], [12],[21]). *Consider a finite chain of pinned or unpinned oscillators with fix, free or forced boundary conditions in contact with two thermal baths at different temperatures and perturbed by one of the energy conserving noises defined above. Then, there exists a unique non-equilibrium stationary state for this dynamics which is absolutely continuous w.r.t. Lebesgue measure.*

Proof. The proof of the existence of the invariant state can be obtained from the knowledge of a suitable Liapounov function and uniqueness from a coupling argument. \square

1.2 Simplified perturbed Hamiltonian systems

Introducing a noise into the deterministic dynamics help us to solve some ergodicity problems. Nevertheless, as we will see, several open challenging problems remain open for chains of oscillators perturbed by a conservative noise. In [13] we proposed to simplify still these models and the main message addressed there is that the models introduced in [13] have qualitatively the same behaviors as the unpinned chains. For simplicity we define only the dynamics in infinite volume.

Let U and V be two potentials on \mathbb{R} and consider the Hamiltonian system $(\omega(t))_{t \geq 0} = (\mathbf{r}(t), \mathbf{p}(t))_{t \geq 0}$ described by the equations of motion

$$\frac{dp_x}{dt} = V'(r_{x+1}) - V'(r_x), \quad \frac{dr_x}{dt} = U'(p_x) - U'(p_{x-1}), \quad x \in \mathbb{Z}, \quad (1.16)$$

where p_x is the momentum of particle x , q_x its position and $r_x = q_x - q_{x-1}$ the “deformation”. Standard chains of oscillators are recovered for a quadratic kinetic energy $U(p) = p^2/2$. The dynamics conserves (at least) three physical quantities: the total momentum $\sum_x p_x$, the total deformation $\sum_x r_x$ and the total energy $\sum_x \mathcal{E}_x$ with $\mathcal{E}_x = V(r_x) + U(p_x)$. Consequently, every Gibbs grand canonical ensemble $\nu_{\beta, \lambda, \lambda'}$ defined by

$$d\nu_{\beta, \lambda, \lambda'}(\eta) = \prod_{x \in \mathbb{Z}} \mathcal{Z}(\beta, \lambda, \lambda')^{-1} \exp \{-\beta \mathcal{E}_x - \lambda p_x - \lambda' r_x\} dr_x dp_x \quad (1.17)$$

is invariant under the evolution. To simplify we assume that the potentials U and V are smooth potentials with bounded by below and by above second derivatives.

To overcome our ignorance about macro-ergodicity of the dynamics, as before, we add a stochastic conserving perturbation. In the general case $U \neq V$, the Hamiltonian dynamics can be perturbed by the energy-momentum conserving noise acting on the velocities (as proposed in [29]) but conserving the three physical invariants mentioned above. Then the infinite volume dynamics can be defined on the state space $\tilde{\Omega}$. Its generator \mathcal{L} is given by $\mathcal{L} = \mathcal{A} + \gamma \mathcal{S}$, $\gamma > 0$, where

$$\begin{aligned} (\mathcal{A}f)(\mathbf{r}, \mathbf{p}) &= \sum_{x \in \mathbb{Z}} \{ (V'(r_{x+1}) - V'(r_x)) \partial_{p_x} f + (U'(p_x) - U'(p_{x+1})) \partial_{r_x} f \} (\mathbf{r}, \mathbf{p}) \\ (\mathcal{S}f) &= \sum_{x \in \mathbb{Z}} [f(\mathbf{r}, \mathbf{p}^{x, x+1}) - f(\mathbf{r}, \mathbf{p})] \end{aligned} \quad (1.18)$$

for any $f \in C_0^1(\tilde{\Omega})$.

Theorem 1.2.1 ([13]). *The dynamics generated by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$ with $\gamma > 0$ and \mathcal{A}, \mathcal{S} given by (1.18) is macro-ergodic. Consequently, before the appearance of the shocks, in the Euler time scale, the hydrodynamic limits are given by a triplet of compressible Euler equations.*

Our motivation being to simplify as much as possible the dynamics considered in [1, 2] without destroying the anomalous behavior of the energy diffusion, we mainly focus on the symmetric case $U = V$. Then the \mathbf{p} 's and \mathbf{r} 's play a symmetric role so there is no reason that momentum conservation is more important than deformation conservation. We propose thus to add a noise conserving only the energy and $\sum_x [r_x + p_x]$. It is more convenient to use the variables $\{\eta_x; x \in \mathbb{Z}\} \in \mathbb{R}^{\mathbb{Z}}$ defined by $\eta_{2x} = p_x$ and $\eta_{2x-1} = r_x$ so that (1.16) becomes

$$d\eta_x = [V'(\eta_{x+1}) - V'(\eta_{x-1})] dt, \quad x \in \mathbb{Z}. \quad (1.19)$$

In order to give names to things we interpret this dynamics for the η 's as the dynamics of an interface whose height (resp. energy) at site x is η_x (resp. $V(\eta_x)$). The quantity $\sum_x \eta_x$ is called the volume.

Hence, we introduce the so-called *stochastic energy-volume conserving dynamics*, which is still described by (1.19) between random exponential times where two nearest neighbors heights η_x and η_{x+1} are exchanged. Observe that in the momenta-deformation picture this noise is less degenerate than the momenta exchange noise since exchange between momenta and positions is now allowed. The generator \mathcal{L} of the infinite volume dynamics, well defined on the state space Ω , is given by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$, $\gamma > 0$, where for any $f \in C_0^1(\Omega)$,

$$\begin{aligned} (\mathcal{A}f)(\eta) &= \sum_{x \in \mathbb{Z}} [V'(\eta_{x+1}) - V'(\eta_{x-1})] (\partial_{\eta_x} f)(\eta), \\ (\mathcal{S}f)(\eta) &= \sum_{x \in \mathbb{Z}} [f(\eta^{x,x+1}) - f(\eta)]. \end{aligned} \quad (1.20)$$

The noise still conserves the total energy and the total volume but destroys the conservation of momentum and deformation. Therefore, only two quantities are conserved and the invariant Gibbs grand canonical measures of the stochastic dynamics correspond to the choice $\lambda = \lambda'$ in (1.17). We denote $\nu_{\beta, \lambda, \lambda}$ (resp. $\mathcal{Z}(\beta, \lambda, \lambda)$) by $\mu_{\beta, \lambda}$ (resp. $Z(\beta, \lambda)$).

Chapter 2

Normal diffusion

Normal diffusion of energy in purely deterministic homogeneous chains of oscillators is expected to hold in high dimension ($d \geq 3$) or if momentum is not conserved, i.e. in the presence of a pinning potential. The problem of anomalous diffusion will be discussed in the next chapter. In this chapter we consider the case of normal diffusion.

The first step to show such normal behavior is to prove that the transport coefficient, the thermal conductivity, is well defined. Once it has been achieved, the following non-equilibrium problems can be considered:

- Hydrodynamic limits in the diffusive time scale $t\varepsilon^{-2}$, ε being the scaling parameter: if the system has trivial hydrodynamics in the time scale $t\varepsilon^{-1}$, i.e. if momentum is not conserved, we would like to show that in the diffusive time scale, the macroscopic energy profile evolves according to a diffusion equation. If the system has non-trivial hydrodynamics given by the Euler equations in the hyperbolic scaling (i.e. if momentum is conserved), in the diffusive time scale, we would like to derive the incompressible Navier-Stokes equations. These would be obtained by starting with some initial momentum macroscopic profile of order $\mathcal{O}(\varepsilon)$, energy profile of order $\mathcal{O}(1)$.
- Validity of Fourier's law: we consider the NESS of the system in contact at the boundaries with thermal baths at different temperatures. Fourier's law expresses that the average of the energy current in the NESS is proportional to the gradient of the local temperature. The proportionality coefficient is called the thermal conductivity.

Assume for simplicity that $d = 1$ and that the energy is the only conserved quantity. The corresponding microscopic current, denoted by $j_{x,x+1}^e$, is defined by

the local energy conservation law

$$\mathcal{L}\mathcal{E}_x = -\nabla j_{x-1,x}^e \quad (2.1)$$

where \mathcal{L} is the generator of the infinite dynamics under investigation. In the actual state of the art, in all the problems mentioned above, the usual approach consists to prove that there exist functions $\varphi_x = \theta_x \varphi_0$ and $h_x = \theta_x h_0$ (actually only approximate solutions are needed) such that the following decomposition

$$j_{x,x+1}^e = \nabla \varphi_x + \mathcal{L}h_x \quad (2.2)$$

holds. Here θ_x denotes the shift by $x \in \mathbb{Z}^d$. Equation (2.2) is called a *microscopic fluctuation-dissipation equation*. Then, taking arbitrary large integer $\ell \geq 1$, by using a multi-scale analysis we replace the block averaged function $\ell^{-1} \sum_{|y-x| \leq \ell} \varphi_y$ by $D(\mathcal{E}_x^\ell) \nabla \mathcal{E}_x^\ell$ where the function D is identified to a diffusion coefficient which depends on the empirical energy $\mathcal{E}_x^\ell = \ell^{-1} \sum_{|y-x| \leq \ell} \mathcal{E}_y$ in the mesoscopic box of length $(2\ell + 1)$ centered around x . Intuitively, $\mathcal{L}h_x$ represents rapid fluctuation (integrated in time, it is a martingale) and the term $\nabla \varphi_x$ represents the dissipation. Gradient models are systems for which the current is equal to the gradient of a function ($h_x = 0$ with the previous notations).

There are at least two reasons for which the problems listed above are difficult:

- The existence of a microscopic fluctuation-dissipation equation has been given for the first time for reversible systems. It has been extended to asymmetric systems satisfying a *sector condition*. Roughly speaking this last condition means that the antisymmetric part of the generator is a bounded perturbation of the symmetric part of the generator¹. Later this condition has been relaxed into the so-called *graded sector condition*: there exists a gradation of the space where the generator is defined and the asymmetric part is bounded by the symmetric part on each graded part (see [37], [32] and references therein). The Hamiltonian systems perturbed by a noise are non-reversible and since the noise (the symmetric part of the generator) is very degenerate, none of these conditions hold.
- The system evolves in a non compact space and one needs to show that energy can not concentrate on a site. This technical problem turns out to be difficult since no general techniques are available. For deterministic nonlinear chains the bounds on the average energy moments are usually polynomial in the size N of the system. Typically we need bounds of order one with respect to N .

¹The antisymmetric (resp. symmetric) part of the generator \mathcal{L} is given by $\frac{\mathcal{L}-\mathcal{L}^*}{2}$ (resp. $\frac{\mathcal{L}+\mathcal{L}^*}{2}$) where \mathcal{L}^* is the adjoint of \mathcal{L} in $L^2(\mu)$, μ being any Gibbs grand canonical measure. For the models considered in this course, the antisymmetric part is \mathcal{A} and due to the deterministic dynamics, the symmetric part is \mathcal{S} and due to the noise.

2.1 Anharmonic chain with velocity-flip noise

2.1.1 Linear response theory: Green-Kubo formula

Green-Kubo formula is one of the most important formulas of non-equilibrium statistical mechanics. In the two problems mentioned in the introduction of the chapter (hydrodynamic limits and Fourier's law) the limiting objects are defined via some macroscopic coefficients which can be expressed by a Green-Kubo formula. The latter is a formal expression and showing it is well defined is a difficult problem. It is usually introduced in the context of the linear response theory that we describe below.

Consider a one dimensional unpinned chain of N harmonic oscillators with forced boundary conditions and perturbed by the velocity-flip noise. The two external constant forces are denoted by τ_ℓ and τ_r . Furthermore on the boundary particles 1 and N there are acting Langevin thermostats at different temperature $T_\ell = \beta_\ell^{-1}$ and $T_r = \beta_r^{-1}$. The generator \mathcal{L}_N of the dynamics is given by (1.12) and we denote the unique non-equilibrium stationary state by μ_{ss} . The expectation w.r.t. μ_{ss} is denoted by $\langle \cdot \rangle_{ss}$.

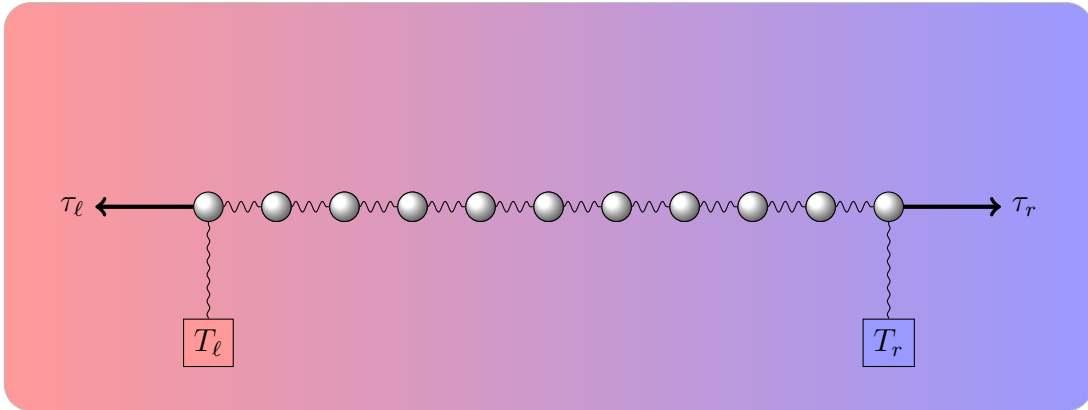


Figure 2.1: The unpinned chain with boundary thermal reservoirs and forced boundary conditions.

The energy ² of atom x is defined by

$$\mathcal{E}_1 = \frac{p_1^2}{2}, \quad \mathcal{E}_x = \frac{p_x^2}{2} + V(r_x), \quad x = 2, \dots, N.$$

²The definition of the energy is slightly modified w.r.t. (1.1). It is more convenient since the energies are then independent random variables in the Gibbs grand canonical ensemble.

The local conservation of energy is expressed by the microscopic continuity equation

$$\mathcal{L}_N(\mathcal{E}_x) = -\nabla j_{x-1,x}^e, \quad x = 1, \dots, N,$$

where the energy current $j_{x,x+1}^e$ from site x to site $x+1$ is given by

$$\begin{aligned} j_{0,1}^e &= -\tau_\ell p_1 + \gamma_\ell(T_\ell - p_1^2), \\ j_{N,N+1}^e &= -\tau_r p_N - \gamma_r(T_r - p_N^2), \\ j_{x,x+1}^e &= -p_x V'(r_{x+1}), \quad x = 1, \dots, N-1. \end{aligned} \tag{2.3}$$

The energy current $j_{0,1}^e$ (and similarly for $j_{N,N+1}^e$) is composed of two terms: the term $-\tau_\ell p_1$ corresponds to the work done on the first particle by the linear force and the term $\gamma_\ell(T_\ell - p_1^2)$ is the heat current due to the left reservoir.

Let V_s be the velocity of the center of mass of the system and J_s be the average energy current, which are defined by

$$V_s = \langle p_x \rangle_{ss} \quad \text{and} \quad J_s = \langle j_{x,x+1}^e \rangle_{ss}.$$

We have the simple relation between these two quantities

$$J_s = -\tau_\ell V_s + \gamma_\ell(T_\ell - \langle p_1^2 \rangle_{ss}), \quad J_s = -\tau_r V_s - \gamma_r(T_r - \langle p_N^2 \rangle_{ss}). \tag{2.4}$$

The value of V_s can be determined exactly and is independent of the nonlinearities present in the system. It follows that the tension profile, defined by $\tau_x = \langle V'(r_x) \rangle_{ss}$, is linear. This is the content of the following straightforward lemma.

Lemma 2.1.1 ([12]). *The velocity V_s of the center of mass is given by*

$$V_s = \frac{\tau_r - \tau_\ell}{\gamma(N-2) + \gamma_\ell + \gamma_r} \tag{2.5}$$

and the tension profile is linear:

$$\tau_x = \frac{\gamma(x-1) + \gamma_\ell}{\gamma(N-2) + \gamma_\ell + \gamma_r} (\tau_r - \tau_\ell) + \tau_\ell, \tag{2.6}$$

that implies

$$\lim_{n \rightarrow \infty} \tau_{[Nu]} = \tau_\ell + (\tau_r - \tau_\ell)u, \quad u \in [0, 1]. \tag{2.7}$$

For purely deterministic chain ($\gamma = 0$), the velocity V_s is of order 1, while the tension profile is flat at the value $(\gamma_\ell + \gamma_r)^{-1} [\gamma_\ell \tau_r + \gamma_r \tau_\ell]$. The first effect of the noise is to make V_s of order N^{-1} and to give a normal tension profile.

It is expected there exists a positive constant C independent of the size N such that $\langle \mathcal{E}_x \rangle_{ss} \leq C$ for any $x = 1, \dots, N$. Apart from the harmonic case we do not know how to prove such a bound.

We shall denote by \tilde{f}_{ss} the derivative of the stationary state μ_{ss} with respect to the local Gibbs equilibrium state μ_{lg} defined by $\mu_{lg}(d\mathbf{r}, d\mathbf{p}) = g(\mathbf{r}, \mathbf{p}) d\mathbf{r} d\mathbf{p}$ with

$$g(\mathbf{r}, \mathbf{p}) = \prod_{x=1}^N \frac{e^{-\beta_x(\mathcal{E}_x - \tau_x r_x)}}{Z(\tau_x \beta_x, \beta_x)}, \quad \beta_x = \beta_\ell + \frac{x}{N}(\beta_r - \beta_\ell). \quad (2.8)$$

In the formula above we have introduced $r_1 = 0$ to avoid annoying notations.

The function \tilde{f}_{ss} is solution, in the sense of distributions, of the equation

$$\tilde{\mathcal{L}}_N^* \tilde{f}_{ss} = 0 \quad (2.9)$$

where $\tilde{\mathcal{L}}_N^*$ is the adjoint of \mathcal{L}_N in $\mathbb{L}^2(\mu_{lg})$. We assume that $T_r = T + \delta T, T_\ell = T$ and $\tau_r = \tau - \delta\tau, \tau_\ell = \tau$ with $\delta T, \delta\tau$ small. At first order in δT and $\delta\tau$, we have

$$\tilde{\mathcal{L}}_N^* = \mathcal{L}_{N,eq}^* + \gamma_r \delta T \partial_{p_N}^2 - \delta\tau \partial_{p_N} - \frac{\delta T}{T^2 N} \sum_{x=1}^{N-1} (j_{x,x+1}^e + \tau p_x) + \frac{\delta\tau}{NT} \sum_{x=1}^{N-1} p_x + o(\delta T, \delta\tau)$$

where $\mathcal{L}_{N,eq}^* = -\mathcal{A}_N^{\tau, \tau} + \gamma \mathcal{S}_N + \gamma_\ell \mathcal{B}_{1,T} + \gamma_r \mathcal{B}_{N,T}$ is the adjoint in $\mathbb{L}^2(\mu_{\tau,T}^N)$ of

$$\mathcal{L}_{N,eq} = \mathcal{A}_N^{\tau, \tau} + \gamma \mathcal{S}_N + \gamma_\ell \mathcal{B}_{1,T} + \gamma_r \mathcal{B}_{N,T} \quad (2.10)$$

and $\mu_{\tau,T}^N$ is the finite volume Gibbs grand canonical ensemble with tension τ and temperature T . We now expand \tilde{f}_{ss} at the linear order in δT and $\delta\tau$:

$$\tilde{f}_{ss} = 1 + \tilde{u} \delta T + \tilde{v} \delta\tau + o(\delta T, \delta\tau) \quad (2.11)$$

and we get that \tilde{u} and \tilde{v} are solution of

$$\begin{aligned} \mathcal{L}_{N,eq}^* \tilde{u} &= \frac{1}{T^2 N} \sum_{x=1}^{N-1} (j_{x,x+1}^e + \tau p_x), \\ \mathcal{L}_{N,eq}^* \tilde{v} &= \frac{1}{NT} \sum_{x=1}^{N-1} p_x. \end{aligned} \quad (2.12)$$

It is clear that the function h_x appearing in the microscopic fluctuation-dissipation equation (2.2) is closely related (up to a time reversal) to the function \tilde{u}, \tilde{v} , i.e. to the *first order correction to local equilibrium*.

We can now compute the average energy current at the first order in δT and $\delta\tau$ as $N \rightarrow \infty$ but we need to introduce few notations. We recall that the generator of the infinite dynamics is given by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$ where, for any $f \in C_0^1(\tilde{\Omega})$,

$$\begin{aligned} (\mathcal{A}f)(\mathbf{r}, \mathbf{p}) &= \sum_{x \in \mathbb{Z}} [(p_x - p_{x-1}) \partial_{r_x} f + (V'(r_{x+1}) - V'(r_x)) \partial_{p_x} f](\mathbf{r}, \mathbf{p}), \\ (\mathcal{S}f)(\mathbf{r}, \mathbf{p}) &= \frac{1}{2} \sum_{x \in \mathbb{Z}} [f(\mathbf{r}, \mathbf{p}^x) - f(\mathbf{r}, \mathbf{p})]. \end{aligned}$$

Let $\mathbb{H} := \mathbb{H}_{\tau, T}$ be the completion w.r.t. the semi-inner product $\ll \cdot, \cdot \gg$ defined for bounded local functions $f, g : \tilde{\Omega} \rightarrow \mathbb{R}$, by

$$\ll f, g \gg = \sum_{x \in \mathbb{Z}} \{\mu_{\tau, T}(f\theta_x g) - \mu_{\tau, T}(f)\mu_{\tau, T}(g)\}. \quad (2.13)$$

Observe that in \mathbb{H} every constant and discrete gradient $\theta_1 f - f$ is equal to zero.

Let \tilde{J}_s and \hat{V}_s be the limiting average energy current and velocity:

$$\tilde{J}_s = \lim_{N \rightarrow \infty} N \langle j_{0,1}^e \rangle_{ss}, \quad \hat{V}_s = \lim_{N \rightarrow \infty} N \langle p_0 \rangle_{ss}, \quad (2.14)$$

and define $\hat{J}_s = \tilde{J}_s + \tau \hat{V}_s$. We expect that as N goes to infinity and, at first order in δT and $\delta\tau$,

$$\begin{pmatrix} \hat{J}_s \\ \hat{V}_s \end{pmatrix} = -\kappa(T, \tau) \begin{pmatrix} \delta T \\ \delta\tau \end{pmatrix}$$

with

$$\kappa(T, \tau) = \begin{pmatrix} \kappa^e & \kappa^{e,r} \\ \kappa^{r,e} & \kappa^r \end{pmatrix} \quad (2.15)$$

the *thermal conductivity* matrix. By (2.12), we get that in the thermodynamic limit $N \rightarrow \infty$, the transport coefficients are given by the Green-Kubo formulas

$$\begin{aligned} \kappa^e &= T^{-2} \ll j_{0,1}^e + \tau p_0, (-\mathcal{L})^{-1} (j_{0,1}^e + \tau p_0) \gg, \\ \kappa^{e,r} &= -T^{-1} \ll p_0, (-\mathcal{L})^{-1} (j_{0,1}^e + \tau p_0) \gg, \end{aligned} \quad (2.16)$$

and

$$\begin{aligned} \kappa^r &= T^{-1} \ll p_0, (-\mathcal{L})^{-1} (p_0) \gg, \\ \kappa^{r,e} &= -T^{-2} \ll j_{0,1}^e + \tau p_0, (-\mathcal{L})^{-1} (p_0) \gg. \end{aligned} \quad (2.17)$$

The argument above is formal. In fact even proving the existence of the transport coefficients defined by (2.16), (2.17) is a non-trivial task. It can be made

rigorous for \hat{V}_s since we have the exact expression of V_s . From Lemma 2.1.1, we have, even for $\delta\tau, \delta T$ that are not small,

$$\hat{V}_s = -\frac{\delta\tau}{\gamma}$$

and the quantities $\kappa^r, \kappa^{r,e}$, defined by (2.17), shall satisfy

$$\kappa^r = \gamma^{-1}, \quad \kappa^{r,e} = 0. \quad (2.18)$$

2.1.2 Existence of the Green-Kubo formula

One of the main results of [12] is the existence of the Green-Kubo formula for the conductivity matrix. Let \mathbb{H}^a (resp. \mathbb{H}^s) be the set of functions $f : \tilde{\Omega} \rightarrow \mathbb{R}$ antisymmetric (resp. symmetric) in \mathbf{p} , i.e. $f(\mathbf{r}, \mathbf{p}) = -f(\mathbf{r}, -\mathbf{p})$ (resp. $f(\mathbf{r}, \mathbf{p}) = f(\mathbf{r}, -\mathbf{p})$) for every configuration $(\mathbf{r}, \mathbf{p}) \in \tilde{\Omega}$. For example, the functions $j_{0,1}^e, p_0$ and every linear combination of them are antisymmetric in \mathbf{p} .

Theorem 2.1.2 ([12], [9]). *Let $f, g \in \mathbb{H}^a$. The limit*

$$\sigma(f, g) = \lim_{z \rightarrow 0} \ll f, (z - \mathcal{L})^{-1} g \gg$$

exists and $\sigma(f, g) = \sigma(g, f)$. Then, the conductivity matrix $\kappa(T, \tau)$ is well defined in the following sense: the limits

$$\begin{aligned} \kappa^e &= \lim_{z \rightarrow 0} T^{-2} \ll j_{0,1}^e + \tau p_0, (z - \mathcal{L})^{-1} (j_{0,1}^e + \tau p_0) \gg, \\ \kappa^{e,r} &= -\lim_{z \rightarrow 0} T^{-1} \ll p_0, (z - \mathcal{L})^{-1} (j_{0,1}^e + \tau p_0) \gg, \\ \kappa^r &= \lim_{z \rightarrow 0} T^{-1} \ll p_0, (z - \mathcal{L})^{-1} (p_0) \gg = (2\gamma)^{-1}, \\ \kappa^{r,e} &= -\lim_{z \rightarrow 0} T^{-2} \ll j_{0,1}^e + \tau p_0, (z - \mathcal{L})^{-1} (p_0) \gg \end{aligned} \quad (2.19)$$

exist and are finite. Moreover Onsager's relation $\kappa^{e,r} = \kappa^{r,e} (= 0)$ holds.

We have a nice thermodynamical consequence of the previous result. If δT and $\delta\tau$ are small and of the same order, the system can not be used as a refrigerator or a boiler: at the first order, a gradient of tension does not contribute to the energy current. The argument above says nothing about the possibility to realize a heater or a refrigerator if $\delta\tau$ is not of the same order as δT . For the harmonic chain, we will see that it is possible to get a heater if $\delta\tau$ is of order $\sqrt{\delta T}$.

Remark 2.1.3. *1. The existence of the Green-Kubo formula is also valid for a pinned or unpinned chain in any dimension.*

2. Observe that with respect to the establishment of a microscopic fluctuation-dissipation equation (2.2) the computation of the Green-Kubo formula is less demanding since only the knowledge of $\sum_x h_x$ is necessary.

The proof of Theorem 2.1.2 is based on functional analysis arguments. The first main observation is that there exists of a spectral gap for the operator \mathcal{S} restricted to the space \mathbb{H}^a .

Lemma 2.1.4. *The noise operator \mathcal{S} lets \mathbb{H}^a and \mathbb{H}^s invariant. For any local function $f \in \mathbb{H}^a$ we have that*

$$\ll f, f \gg \leq \ll f, -\mathcal{S}f \gg. \quad (2.20)$$

Moreover, for any local function $f \in \mathbb{H}^a$, there exists a local function $h \in \mathbb{H}^a$ such that

$$\mathcal{S}h = f.$$

Proof. Since the Gibbs states are Gaussian states in the p_x 's it is convenient to decompose the operator \mathcal{S} (which acts only on the p_x 's) in the orthogonal basis of Hermite polynomials. The the lemma follows easily. \square

Proof of Theorem 2.1.2. We observe first that \mathbb{H}^a and \mathbb{H}^s are orthogonal Hilbert spaces such that $\mathbb{H} = \mathbb{H}^a \oplus \mathbb{H}^s$. It is also convenient to define the following semi-inner product

$$\ll f, g \gg_1 = \ll f, (-\mathcal{S})g \gg.$$

Let \mathbb{H}^1 be the associated Hilbert space. We also define the Hilbert space \mathbb{H}^{-1} via the duality given by the \mathbb{H} norm, that is

$$\|f\|_{-1}^2 = \sup_g \{ 2 \ll f, g \gg - \ll g, g \gg_1 \}$$

where the supremum is taken over local bounded functions g .

Then we decompose the solution f_z of the resolvent equation $(z - \mathcal{L})f_z = g$ into $f_z = f_z^- + f_z^+$, $f_z^+ \in \mathbb{H}^a$ and $f_z^- \in \mathbb{H}^s$.

Since \mathcal{A} inverts the parity and \mathcal{S} preserves it, we have, for any $\mu, \nu > 0$,

$$\begin{aligned} \nu \ll f_\mu^+, f_\nu^+ \gg - \ll f_\mu^+, \mathcal{A}f_\nu^- \gg + \gamma \ll f_\mu^+, f_\nu^+ \gg_1 &= \ll f_\mu, g^+ \gg \\ \mu \ll f_\nu^-, f_\mu^- \gg - \ll f_\nu^-, \mathcal{A}f_\mu^+ \gg + \gamma \ll f_\mu^-, f_\nu^- \gg_1 &= \ll f_\nu, g^- \gg. \end{aligned} \quad (2.21)$$

Summing the above equations we have

$$\nu \ll f_\mu^+, f_\nu^+ \gg + \mu \ll f_\nu^-, f_\mu^- \gg + \gamma \ll f_\mu, f_\nu \gg_1 = \ll f_\mu, g^+ \gg + \ll f_\nu, g^- \gg \quad (2.22)$$

Putting $\mu = \nu$ we get

$$\nu \ll f_\nu, f_\nu \gg + \gamma \ll f_\nu, f_\nu \gg_1 \leq \|f_\nu\|_1 \|g\|_{-1}.$$

Hence f_ν is uniformly bounded in \mathbb{H}^1 and by the spectral gap property so is f_ν^- in \mathbb{H} . Moreover, νf_ν converges strongly to 0 in \mathbb{H} . We can then extract weakly convergent subsequences. Taking first the limit, in (2.22), $\nu \rightarrow 0$ and then $\mu \rightarrow 0$ along one such subsequences (converging to f_*) we have

$$\gamma \ll f_*, f_* \gg_1 = \ll f_*, g \gg.$$

Taking again the limit along such subsequence, with $\mu = \nu$, we have then

$$\lim_{\nu \rightarrow 0} \nu \ll f_\nu, f_\nu \gg = 0. \quad (2.23)$$

Next, taking the limit along different weakly convergent subsequences (let f^* be the other limit) we have

$$\gamma \ll f_*, f^* \gg_1 = \ll f_*, g^+ \gg + \ll f^*, g^- \gg$$

and, exchanging the role of the two sequences

$$2\gamma \ll f_*, f^* \gg_1 = \ll f_*, g \gg + \ll f^*, g \gg = \gamma \ll f_*, f_* \gg_1 + \gamma \ll f^*, f^* \gg_1$$

which implies $f_* = f^*$, that is all the subsequences have the same limit. Finally, arguing similarly to the above, for $\nu(f_\nu - f_\mu) + (\nu - \mu)f_\mu - \mathcal{L}(f_\nu - f_\mu) = 0$ we obtain that the convergence takes place in the strong norm. Thus f_ν^- converges strongly to some $f_0^- \in \mathbb{H}^a$. Since $\ll g, f_\nu \gg = \ll g, f_\nu^- \gg$ we conclude that $\ll g, f_\nu \gg$ converges as ν goes to 0.

□

In the harmonic case, $V(r) = r^2/2$, much more is known. Indeed one easily checks that the exact microscopic fluctuation-dissipation equation (2.2) holds with

$$h_x = \frac{1}{2\gamma} r_{x+1} (p_x + p_{x+1}) - \frac{r_{x+1}^2}{4}, \quad \varphi_x = -\frac{1}{2\gamma} (r_x r_{x+1} + p_x^2). \quad (2.24)$$

It follows that the conductivity matrix is given by

$$\kappa(\tau, T) = \begin{pmatrix} \frac{1}{2\gamma} & 0 \\ 0 & \frac{1}{\gamma} \end{pmatrix}.$$

2.1.3 Expansion of the Green-Kubo formula in the weak coupling limit

In the previous subsection we proved the existence of the Green-Kubo formula showing that the transport coefficient is well defined if some noise is added to the deterministic dynamics. We are now interested in the behavior of the Green-Kubo formula as the noise vanishes. We investigate this question in the weak coupling limit, i.e. assuming that the interaction potential is in the form εV where $\varepsilon \ll 1$ is the (small) coupling parameter. For simplicity we fix $\tau = 0$, $d = 1$ and V and its derivatives all smooth and bounded. In order to emphasize the dependence of κ^ε (denoted in the sequel by κ) in the coupling parameter ε and the noise intensity γ , we denote κ by $\kappa(\varepsilon, \gamma)$. Here we propose a formal expansion of the conductivity κ in the form

$$\kappa(\varepsilon, \gamma) = \sum_{n \geq 2} \kappa_n(\gamma) \varepsilon^n. \quad (2.25)$$

Then we study rigorously the first term of this expansion $\kappa_2(\gamma)$.

When the system is uncoupled ($\varepsilon = 0$), the dynamics is given by the generator $\mathcal{L}_0 = \mathcal{A}_0 + \gamma \mathcal{S}$ with \mathcal{S} the flip noise defined by (1.11) and

$$\mathcal{A}_0 = \sum_{x \in \mathbb{Z}} p_x \partial_{q_x} - W'(q_x) \partial_{p_x}.$$

When $\varepsilon > 0$, the generator of the coupled dynamics is denoted by

$$\mathcal{L}_\varepsilon = \mathcal{L}_0 + \varepsilon \mathcal{G} \quad (2.26)$$

where

$$\mathcal{G} = V'(q_x - q_{x-1})(\partial_{p_{x-1}} - \partial_{p_x}).$$

The energy of each cell, which is the sum of the internal energy and of the interaction energy, is defined by

$$\mathcal{E}_x^\varepsilon = \mathcal{E}_x + \frac{\varepsilon}{2} (V(q_{x+1} - q_x) + V(q_x - q_{x-1})), \quad \mathcal{E}_x = \frac{p_x^2}{2} + W(q_x). \quad (2.27)$$

Observe that $\mathcal{E}_x = \mathcal{E}_x^0$ is the energy of the isolated system x . The dynamics generated by \mathcal{L}_0 preserves all the individual energies \mathcal{E}_x . The dynamics generated by \mathcal{L}_ε conserves the total energy. The corresponding energy currents $\varepsilon j_{x,x+1}$, defined by the local conservation law

$$\mathcal{L}_\varepsilon \mathcal{E}_x^\varepsilon = \varepsilon (j_{x-1,x} - j_{x,x+1})$$

are given by

$$\varepsilon j_{x,x+1} = -\frac{\varepsilon}{2} (p_x + p_{x+1}) \cdot \nabla V(q_{x+1} - q_x). \quad (2.28)$$

Let us denote by $\mu_{\beta,\varepsilon} = \langle \cdot \rangle_{\beta,\varepsilon}$ the canonical Gibbs measure at temperature $\beta^{-1} > 0$ (the pressure $\tau = 0$ if the system is unpinned) defined by the Dobrushin-Lanford-Ruelle equations, which of course depends on the interaction εV . We shall assume in all the cases considered that $\mu_{\beta,\varepsilon}$ is analytical in ε for sufficiently small ε (when applied to local functions). In particular we assume that the potentials V and W are such that the Gibbs state is unique and has spatial exponential decay of correlations.

In order to emphasize the dependence in ε we reintroduce some notations. For any given local functions f, g , define the semi-inner product

$$\ll f, g \gg_{\beta,\varepsilon} = \sum_{x \in \mathbb{Z}} [\langle \theta_x f, g \rangle_{\beta,\varepsilon} - \langle f \rangle_{\beta,\varepsilon} \langle g \rangle_{\beta,\varepsilon}]. \quad (2.29)$$

We recall that θ_x is the shift operator by x . The sum is finite in the case $\varepsilon = 0$, and converges for $\varepsilon > 0$ thanks to the exponential decay of the spatial correlations. Denote by $\mathbb{H}_\varepsilon = \mathbb{L}^2(\ll \cdot, \cdot \gg_{\beta,\varepsilon})$ the corresponding closure. We define the subspace of the antisymmetric functions in the velocities

$$\mathbb{H}_\varepsilon^a = \{f \in \mathbb{H}_\varepsilon : f(\mathbf{q}, -\mathbf{p}) = -f(\mathbf{q}, \mathbf{p})\}. \quad (2.30)$$

where $\mathbf{q} = \{\mathbf{q}_x \in \mathbb{R}\}_{x \in \mathbb{Z}}$, $\mathbf{p} = \{\mathbf{p}_x \in \mathbb{R}\}_{x \in \mathbb{Z}}$. Similarly we define the subspace of the symmetric functions in \mathbf{p} as \mathbb{H}_ε^s . On local functions this decomposition is independent of ε . Let us denote by $\mathcal{P}_\varepsilon^a$ and $\mathcal{P}_\varepsilon^s$ the corresponding orthogonal projections, whose definition in fact does not depend on ε . Therefore we sometime omit the index ε in the notation. Finally, for any function $f \in \mathbb{L}^2(\mu_{\beta,\varepsilon})$, define

$$(\Pi_\varepsilon f)(\mathcal{E}) = \mu_{\beta,\varepsilon}(f|\mathcal{E}), \quad Q_\varepsilon = \text{Id} - \Pi_\varepsilon$$

where $\mathcal{E} := \{\mathcal{E}_x ; x \in \mathbb{Z}\}$. According to Theorem 2.1.2 the conductivity is defined by

$$\kappa(\varepsilon, \gamma) = \varepsilon^2 \lim_{\nu \rightarrow 0} \ll j_{0,1}, (\nu - \mathcal{L}_\varepsilon)^{-1} j_{0,1} \gg_{\beta,\varepsilon}. \quad (2.31)$$

It turns out that, for calculating the terms in the expansion (2.25), it is convenient to choose $\nu = \varepsilon^2 \lambda$ in (2.31), for a $\lambda > 0$, and solve the resolvent equation

$$(\lambda \varepsilon^2 - \mathcal{L}_\varepsilon) u_{\lambda,\varepsilon} = \varepsilon j_{0,1} \quad (2.32)$$

for the unknown function $u_{\lambda,\varepsilon}$. The factor ε^2 is the natural scaling in view of the subsequent computations. We assume that a solution of (2.32) is in the form

$$u_{\lambda,\varepsilon} = \sum_{n \geq 0} U_{\lambda,n} \varepsilon^n = \sum_{n \geq 0} (v_{\lambda,n} + w_{\lambda,n}) \varepsilon^n, \quad (2.33)$$

where $\Pi v_{\lambda,n} = Q w_{\lambda,n} = 0$. Here $\Pi = \Pi_0$ and $Q = Q_0$ refer to the uncoupled measure $\mu_{\beta,0}$ ³. Given such an expression we can, in principle, use it in (2.31) to write

$$\begin{aligned} \kappa(\varepsilon, \gamma) &= \lim_{\lambda \rightarrow 0} \sum_{n \geq 0} \varepsilon^{n+1} \ll j_{0,1}, v_{\lambda,n} + w_{\lambda,\varepsilon} \gg_{\beta,\varepsilon} \\ &= \sum_{n \geq 1} \lim_{\lambda \rightarrow 0} \varepsilon^n \ll j_{0,1}, v_{\lambda,n-1} \gg_{\beta,\varepsilon} \end{aligned} \quad (2.34)$$

where we have used the fact that $\ll j_{0,1}, w_{\lambda,\varepsilon} \gg_{\beta,\varepsilon} = 0$ and we have, arbitrarily, exchanged the liminf with the sum. Note that this is not yet of the type (2.25) since the terms in the expansion depend themselves from ε . To identify the coefficients κ_n we would need to expand in ε also the expectations. This is not obvious since the functions $v_{\lambda,n}$ are non local.

Let us consider the operator $\mathfrak{L} = \Pi \mathcal{G} \mathcal{P}^a (-\mathcal{L}_0)^{-1} \mathcal{G} \Pi$. We will show below that the operator \mathfrak{L} is a generator of a Markov process so that $(\lambda - \mathfrak{L})^{-1}$ is well defined for $\lambda > 0$. Plugging (2.33) in (2.32) we obtain the following hierarchy

$$\begin{aligned} w_{\lambda,0} &= (\lambda - \mathfrak{L})^{-1} \Pi \mathcal{G} \mathcal{P}^a (-\mathcal{L}_0)^{-1} j_{0,1} \\ v_{\lambda,1} &= (-\mathcal{L}_0)^{-1} [j_{0,1} + \mathcal{G} w_{\lambda,0}], \\ w_{\lambda,n} &= (\lambda - \mathfrak{L})^{-1} \Pi \mathcal{G} (-\mathcal{L}_0)^{-1} [-\lambda v_{\lambda,n-1} + Q \mathcal{G} v_{\lambda,n}], \quad n \geq 1 \\ v_{\lambda,n+1} &= (-\mathcal{L}_0)^{-1} [-\lambda v_{\lambda,n-1} + \mathcal{G} w_{\lambda,n} + Q \mathcal{G} v_{\lambda,n}], \quad n \geq 1. \end{aligned} \quad (2.35)$$

Observe that in the previous equations the (formal) operator $(-\mathcal{L}_0)^{-1}$ is always applied to functions f such that $\Pi f = 0$ (this is the minimal requirement to have non-silly equations). This is however not sufficient to make sense of the functions $v_{\lambda,n}$ and $w_{\lambda,n}$. Nevertheless, by using Theorem 2.1.2, we have that the local operator \mathcal{T}_0 on \mathbb{H}^a defined by

$$\mathcal{T}_0 f = \lim_{\nu \rightarrow 0} \mathcal{P}^a (\nu - \mathcal{L}_0)^{-1} f, \quad f \in \mathbb{H}_0^a,$$

is well defined. Therefore, it is possible to make sense, as a distribution, of

$$\alpha_{01} = \Pi \mathcal{G} \mathcal{P}^a (-\mathcal{L}_0)^{-1} j_{0,1} := \Pi \mathcal{G} \mathcal{T}_0 j_{0,1}. \quad (2.36)$$

Nevertheless, the function $w_{\lambda,0}$ is still not well defined since we are not sure that $\mathcal{T}_0 j_{0,1}$ is in the domain of \mathcal{G} .

Even if the previous computations are formal a remarkable fact is that the operator \mathfrak{L} , when applied to functions of the internal energies, coincides with the

³The reason to use the orthogonal decomposition of $U_{\lambda,n} = v_{\lambda,n} + w_{\lambda,n}$ is that at some point we will have to consider, for a given function f , the solution h to the Poisson equation $\mathcal{L}_0 h = f$. The minimal requirement for the existence of h is that $\Pi f = 0$.

Markov generator \mathfrak{L}_{GL} of a reversible Ginzburg-Landau dynamics. Let us denote by ρ_β the distribution of the internal energies $\mathcal{E} = \{\mathcal{E}_x; x \in \mathbb{Z}\}$ under the Gibbs measure $\mu_{\beta,0}$. It can be written in the form

$$d\rho_\beta(\mathcal{E}) = \prod_{x \in \mathbb{Z}} Z_\beta^{-1} \exp(-\beta \mathcal{E}_x - U(\mathcal{E}_x)) d\mathcal{E}_x$$

for a suitable function U . We denote the formal sum $\sum_x U(\mathcal{E}_x)$ by $\mathcal{U} := \mathcal{U}(\mathcal{E})$. We denote also, for a given value of the internal energy $\tilde{\mathcal{E}}_x$ in the cell x , by $\nu_{\tilde{\mathcal{E}}_x}^x$ the microcanonical probability measure in the cell x . i.e. the uniform probability measure on the manifold

$$\Sigma_{\tilde{\mathcal{E}}_x} := \{(q_x, p_x) \in \Omega; \mathcal{E}_x(q_x, p_x) = \tilde{\mathcal{E}}_x\}.$$

Then, the generator \mathfrak{L}_{GL} is given by

$$\mathfrak{L}_{\text{GL}} = \sum_x e^{\mathcal{U}} (\partial_{\mathcal{E}_{x+1}} - \partial_{\mathcal{E}_x}) [e^{-\mathcal{U}} \gamma^2(\mathcal{E}_x, \mathcal{E}_{x+1}) (\partial_{\mathcal{E}_{x+1}} - \partial_{\mathcal{E}_x})], \quad (2.37)$$

where

$$\gamma_{0,1}^2 := \gamma^2(\mathcal{E}_0, \mathcal{E}_1) = \int_{\Sigma_{\mathcal{E}_0} \times \Sigma_{\mathcal{E}_1}} (j_{0,1} \mathcal{T}_0 j_{0,1}) d\nu_{\mathcal{E}_0}^0 d\nu_{\mathcal{E}_1}^1. \quad (2.38)$$

The operator \mathfrak{L}_{GL} is well defined only if γ^2 has some regularity properties, that are actually proven in specific examples [41, 23]. We can show that the Dirichlet forms ⁴ associated to \mathfrak{L} and \mathfrak{L}_{GL} coincide. Then in the cases where γ^2 is proven smooth (2.37) is well defined and $\mathfrak{L} = \mathfrak{L}_{\text{GL}}$.

Proposition 2.1.5 ([9]). *For each local smooth functions f, g of the internal energies only we have*

$$\lll g, (-\mathfrak{L})f \ggg_{\beta,0} = \lll g, (-\mathfrak{L}_{\text{GL}})f \ggg_{\beta,0}. \quad (2.39)$$

The operator \mathfrak{L}_{GL} is the generator of a Ginzburg-Landau dynamics which is reversible with respect to ρ_β , for any $\beta > 0$. It is conservative in the energy $\sum_x \mathcal{E}_x$ and the corresponding currents are given by $\theta_x \alpha_{0,1}$ where $\alpha_{0,1}$ has been defined in (2.36). The corresponding finite size dynamics appears in [41, 24] as the weak coupling limit of a finite number N (fixed) of cells weakly coupled by a potential ϵV in the limit $\epsilon \rightarrow 0$ when time t is rescaled as $t\epsilon^{-2}$. Moreover, the hydrodynamic limit of the Ginzburg-Landau dynamics is then given (in the diffusive time scale tN^2 , $N \rightarrow +\infty$), by a heat equation with diffusion coefficient which coincides with κ_2 as given by (2.41) below ([55]). This is summarized in Figure 2.2.

⁴They are well defined even if γ^2 is not regular.

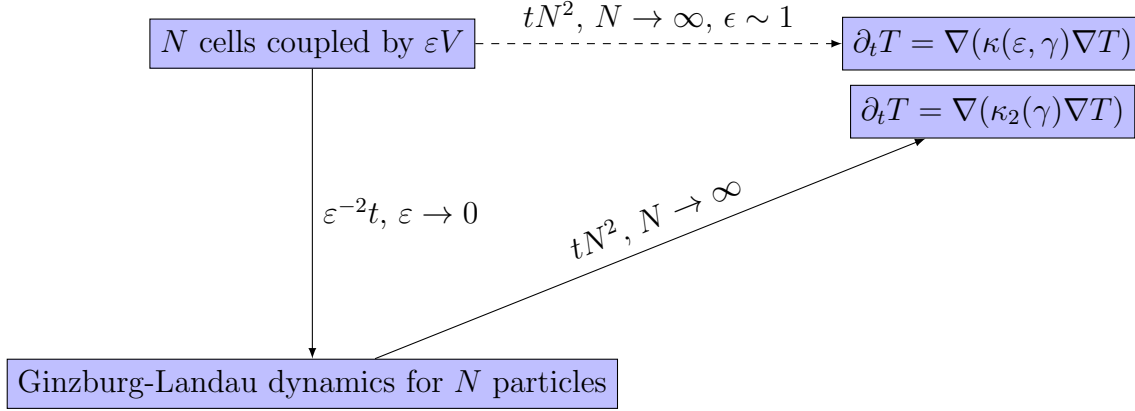


Figure 2.2: The relation between the hydrodynamic limit, the weak coupling limit and the Green-Kubo expansion. The dotted arrow (hydrodynamic limits in the diffusive time scale) has not been proved. The weak coupling limit (vertical arrow) has been proved in [41] (see also [24]) and the diagonal arrow (hydrodynamic limits for a Ginzburg-Landau dynamics) has been obtained in [55] in some cases which however do not cover our cases. In [9] it is argued that $\kappa(\varepsilon, \gamma) \sim \varepsilon^2 \kappa_2(\gamma)$ as $\varepsilon \rightarrow 0$.

According to the previous expansion it makes sense to define $\kappa_2(\gamma)$ by

$$\kappa_2(\gamma) = \lim_{\varepsilon \rightarrow 0} \lim_{\lambda \rightarrow 0} \{ \ll j_{0,1}, \mathcal{T}_0 j_{0,1} \gg_{\beta, \varepsilon} + \ll j_{0,1}, \mathcal{T}_0 \mathcal{G} w_{\lambda,0} \gg_{\beta, \varepsilon} \} \quad (2.40)$$

if the limits exist. In fact, a priori, it is not even clear that the term $\mathcal{T}_0 \mathcal{G} w_{\lambda,0}$ makes sense since $w_{\lambda,0}$ is not well defined. In [9] we argue that

$$\kappa_2(\gamma) = \langle \gamma_{0,1}^2 \rangle_{\beta} - \ll \alpha_{0,1}, (-\mathcal{L}_{GL})^{-1} \alpha_{0,1} \gg_{\beta}. \quad (2.41)$$

Here $\langle \cdot \rangle_{\beta}$ and $\ll \cdot \gg_{\beta}$ refer to the scalar products w.r.t. ρ_{β} . In the special case $W = 0$ ⁵, we prove rigorously in [9] that we can make sense for any λ, ε of the term in the righthandside of (2.40) and that (2.41) is valid, supporting the conjecture that (2.41) is valid in more general situations. Observe that (2.41) is the Green-Kubo formula for the diffusion coefficient of the Ginzburg-Landau dynamics.

In specific examples, it has been possible to study the behavior of $\kappa_2(\gamma)$ defined by (2.41) in the vanishing noise limit $\gamma \rightarrow 0$:

1. Harmonic chain: it is known that the conductivity of the (deterministic) harmonic chain is $\kappa(\varepsilon, 0) = \infty$. If $\gamma > 0$, $\kappa(\varepsilon, \gamma) = c\gamma^{-1}\varepsilon^{-2}$, $c > 0$ a constant, and we get thus that $\lim_{\gamma \rightarrow 0} \kappa_2(\gamma) = \infty$.

⁵If $W = 0$ the variables q_x have to take values in a compact manifold.

2. Disordered pinned harmonic chain: V is quadratic and the one-site potential W is site-dependent given by $W_x(q) = \nu_x q^2$ where $\{\nu_x; x \in \mathbb{Z}\}$ is a sequence of *i.i.d.* positive bounded random variables⁶. It is known ([8]) that $\kappa(\varepsilon, 0) = 0$ so that $\kappa_2(\varepsilon, 0) = 0$. One can prove that $\kappa_2(\gamma)$ vanishes as γ goes to 0.
3. Harmonic chain with quartic pinning potential: V is quadratic and $W(q) = q^4$. Then we show that $\limsup_{\gamma \rightarrow 0} \kappa_2(\gamma) < \infty$. This upper bound does not prevent the possibility that $\lim_{\gamma \rightarrow 0} \kappa_2(\gamma) = 0$.

To prove these results we use the upper bound $\kappa_2(\gamma) \leq \langle \gamma_{0,1}^2 \rangle_\beta$. Recalling (2.38) we see that if we are able to compute $\mathcal{T}_{0,j_{0,1}}$ then we can estimate $\langle \gamma_{0,1}^2 \rangle_\beta$. It is exactly what is done in [9] for the specific cases above.

It would be highly interesting to have a rigorous derivation of the formal expansion above. Bypassing this problem, an other relevant issue is to decide if genuinely $\lim_{\gamma \rightarrow 0} \kappa_2(\gamma)$ is zero or not. Some authors (see [22] and references therein) conjecture that, in some cases, the conductivity of the deterministic chain $\kappa(\varepsilon, 0)$ has a trivial weak coupling expansion ($\kappa(\varepsilon, 0) = \mathcal{O}(\varepsilon^n)$ for any $n \geq 2$). Showing that $\kappa_2(\gamma) \rightarrow 0$ as $\gamma \rightarrow 0$ would support this conjecture.

2.2 Harmonic chain with velocity-flip noise

In this section we assume that $V(r) = r^2/2$.

2.2.1 Hydrodynamic limits

As explained in the beginning of this chapter an interesting problem consists to derive a diffusion equation for a chain of oscillators perturbed by an energy conserving noise. Consider a one dimensional unpinned chain of N harmonic oscillators with periodic boundary conditions perturbed by the velocity flip noise in the diffusive scale. In other words let $\omega(t) = (\mathbf{r}(t), \mathbf{p}(t))_{t \geq 0}$ be the process with generator $N^2 \mathcal{L}_N = N^2 [\mathcal{A}_N + \gamma \mathcal{S}_N]$ where \mathcal{S}_N is given by (1.11), \mathbb{Z}^d being replaced by \mathbb{T}_N , the discrete torus of length N , and \mathcal{A}_N is the Liouville operator of a chain of unpinned harmonic oscillators with periodic boundary conditions. The system conserves two quantities: the total energy $\sum_{x \in \mathbb{T}_N} \mathcal{E}_x$, $\mathcal{E}_x = \frac{p_x^2}{2} + \frac{r_x^2}{2}$, and the total deformation of the lattice $\sum_{x \in \mathbb{T}_N} r_x$. Consequently, the Gibbs equilibrium measures $\nu_{\beta, \tau}$ are indexed

⁶Even if this model does not belong *stricto sensu* to the class of models discussed above it is easy to generalize to this case, at least formally, the previous results.

by two parameters $\beta > 0$, the inverse temperature, and $\tau \in \mathbb{R}$, the pressure. They take the form

$$d\nu_{\beta,\tau}(d\mathbf{r}, d\mathbf{p}) = \prod_{x \in \mathbb{T}_N} \frac{\beta}{2\pi} \exp\{-\beta(\mathcal{E}_x - \tau r_x)\} dr_x dp_x.$$

Observe the following thermodynamic relations

$$\int \mathcal{E}_x d\nu_{\beta,\tau} = \beta^{-1} + \frac{\tau^2}{2}, \quad \int r_x d\nu_{\beta,\tau} = \tau.$$

Definition 2.2.1. Let $\mathbb{T} = [0, 1)$ be the continuous torus. Let $\mathbf{e}_0 : \mathbb{T} \rightarrow \mathbb{R}$ and $\mathbf{r}_0 : \mathbb{T} \rightarrow \mathbb{R}$ be two continuous macroscopic profiles such that $\mathbf{e}_0 > \frac{\mathbf{r}_0^2}{2}$. A sequence of probability measures $(\mu^N)_{N \geq 1}$ on $(\mathbb{R} \times \mathbb{R})^{\mathbb{T}_N}$ is said to be a sequence of Gibbs local equilibrium state associated to the energy profile \mathbf{e}_0 and the deformation profile \mathbf{r}_0 if

$$d\mu^N(d\mathbf{r}, d\mathbf{p}) = \prod_{x \in \mathbb{T}_N} \frac{\beta_0(x/N)}{2\pi} \exp\{-\beta_0(x/N)(\mathcal{E}_x - \tau_0(x/N)r_x)\} dr_x dp_x$$

where the functions β_0 and τ_0 are defined by

$$\tau_0 = \mathbf{r}_0, \quad \beta_0 = \left\{ \mathbf{e}_0 - \frac{\mathbf{r}_0^2}{2} \right\}^{-1}.$$

Once we have the microscopic fluctuation-dissipation equation (see (2.24)) and assuming the propagation of local equilibrium in the diffusive time scale it is easy to guess the hydrodynamic equations followed by the system. In [47] is proved the following theorem.

Theorem 2.2.2 ([47]). Consider the unpinned velocity-flip model with periodic boundary conditions. Let $(\mu^N)_N$ be a sequence of Gibbs local equilibrium state ⁷ associated to a bounded energy profile \mathbf{e}_0 and a deformation profile \mathbf{r}_0 . For every $t \geq 0$, and any test continuous functions $G, H : \mathbb{T} \rightarrow \mathbb{R}$, the random variables

$$\left(\frac{1}{N} \sum_{x \in \mathbb{T}_N} G\left(\frac{x}{N}\right) r_x(tN^2), \frac{1}{N} \sum_{x \in \mathbb{T}_N} H\left(\frac{x}{N}\right) \mathcal{E}_x(tN^2) \right)$$

converge in probability as $N \rightarrow \infty$ to

$$\left(\int_{\mathbb{T}} G(y) \mathbf{r}(t, y) dy, \int_{\mathbb{T}} H(y) \mathbf{e}(t, y) dy \right)$$

where \mathbf{r} and \mathbf{e} are solutions to the hydrodynamical equations

$$\begin{cases} \partial_t \mathbf{r} = \frac{1}{\gamma} \partial_y^2 \mathbf{r}, \\ \partial_t \mathbf{e} = \frac{1}{2\gamma} \partial_y^2 \left[\mathbf{e} + \frac{\mathbf{r}^2}{2} \right] \end{cases}, \quad y \in \mathbb{T} \quad (2.42)$$

with initial conditions $\mathbf{r}(0, y) = \mathbf{r}_0(y)$, $\mathbf{e}(0, y) = \mathbf{e}_0(y)$.

⁷One can consider more general initial states, see [47].

The proof of this theorem is based on Yau's relative entropy method ([56], [44]). The general strategy is simple. Let μ_t^N be the law of the process at time tN^2 starting from μ^N and let $\tilde{\mu}_t^N$ be a sequence of Gibbs local equilibrium state corresponding to the deformation profile $\mathfrak{r}_t(\cdot) := \mathfrak{r}(t, \cdot)$ and energy profile $\mathfrak{e}_t(\cdot) := \mathfrak{e}(t, \cdot)$ solution of (2.42). We expect that since \mathfrak{e} and \mathfrak{r} are the hydrodynamic profiles, the probability measure of the process μ_t^N is close, in some sense, to the local Gibbs state $\tilde{\mu}_t^N$. Yau's relative entropy method consists to show that the entropic distance ⁸

$$H_N(t) := H(\mu_t^N | \tilde{\mu}_t^N) = o(N) \quad (2.43)$$

between the two states is relatively small. Assuming (2.43), in order to prove for example the convergence of the empirical energy, we use the entropy inequality ⁹ which states that for any $\alpha > 0$ and test function ϕ

$$\int \phi d\mu_t^N \leq \frac{H(\mu_t^N | \tilde{\mu}_t^N)}{\alpha} + \frac{1}{\alpha} \log \left(\int e^{\alpha\phi} d\tilde{\mu}_t^N \right). \quad (2.44)$$

We take then $\alpha = \delta N$, $\delta > 0$, and

$$\phi = \left| \frac{1}{N} \sum_{x \in \mathbb{T}_N} H\left(\frac{x}{N}\right) \mathcal{E}_x(tN^2) - \int_{\mathbb{T}} H(y) \mathfrak{e}(t, y) dy \right|.$$

Since $\tilde{\mu}_t^N$ is fully explicit and even product, by using large deviations estimates, it is possible to show that

$$\limsup_{N \rightarrow \infty} \frac{1}{\delta N} \log \left(\int e^{\delta N \phi} d\tilde{\mu}_t^N \right) = I(\delta) \quad (2.45)$$

where $I(\delta) \rightarrow 0$ as $\delta \rightarrow 0$. By using (2.43), we are done. It remains then to prove (2.43) and for this we rely on a Gronwald inequality for the entropy production ($C > 0$ is a constant)

$$\partial_t H_N \leq C H_N(t) + o(N). \quad (2.46)$$

The proof of (2.46) is quite evolved and we refer the interest reader to [47], [10] (see also [36] for some overview on the subject). This is in this step that the macro-ergodicity of the dynamics is used in order to derive the so-called one-block estimate.

For non-gradient systems, the previous strategy has to be modified. Indeed, in order to have (2.43) it is necessary to replace the local equilibrium Gibbs state

⁸There is some abuse of language here since the relative entropy is not a distance between probability measures.

⁹It is a trivial consequence of the definition (1.8).

$\tilde{\mu}_t^N$ by a local equilibrium state with a first order correction term in the form

$$\begin{aligned} & d\hat{\mu}_t^N(d\mathbf{r}, d\mathbf{p}) \\ &= Z_{t,N}^{-1} \prod_{x \in \mathbb{T}_N} \exp \left\{ -\beta_t(x/N)(\mathcal{E}_x - \tau_t(x/N)r_x) + \frac{1}{N}F(t, x/N)(\theta_x g)(\mathbf{r}, \mathbf{p}) \right\} dr_x dp_x \end{aligned} \quad (2.47)$$

where $Z_{t,N}$ is a normalization constant,

$$\tau_t = \mathbf{r}_t, \quad \beta_t = \left\{ \mathbf{e}_t - \frac{\mathbf{r}_t^2}{2} \right\}^{-1}$$

and the functions F and g are judiciously chosen. The choice is guided by the fluctuation-dissipation relation (2.24) as explained bellow.

Let $\Omega^N = (\mathbb{R} \times \mathbb{R})^{\mathbb{T}^N}$ be the configurations space and denote

$$\hat{H}_N(t) := H(\mu_t^N | \hat{\mu}_t^N) = \int_{\Omega^N} f_t^N(\omega) \log \frac{f_t^N(\omega)}{\phi_t^N(\omega)} d\nu_*(\omega), \quad (2.48)$$

where f_t^N is the density of μ_t^N with respect to the Gibbs reference measure $\nu_* := \nu_{1,0}$. In the same way, ϕ_t^N is the density of $\hat{\mu}_t^N$ with respect to ν^* (which is fully explicit). The goal is to get (2.46) with H_N replaced by \hat{H}_N .

We begin with the following entropy production bound. Let us denote by $\mathcal{L}_N^* = -\mathcal{A}_N + \gamma \mathcal{S}_N$ the adjoint of \mathcal{L}_N in $\mathbb{L}^2(\nu^*)$.

Lemma 2.2.3.

$$\partial_t \hat{H}_N(t) \leq \int \frac{1}{\phi_t^N} (N^2 \mathcal{L}_N^* \phi_t^N - \partial_t \phi_t^N) f_t^N d\nu_* = \int \left[\frac{1}{\phi_t^N} (N^2 \mathcal{L}_N^* \phi_t^N - \partial_t \phi_t^N) \right] d\mu_t^N.$$

Proof. We have that f_t^N solves the Fokker-Plack equation $\partial_t f_t^N = N^2 \mathcal{L}_N^* f_t^N$. Assuming it is smooth to simplify, we have

$$\begin{aligned} \partial_t \hat{H}_N(t) &= \int \partial_t f_t^N [1 + \log f_t^N] d\nu_* - \int \partial_t f_t^N \log \phi_t^N d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \\ &= N^2 \int \mathcal{L}_N^* f_t^N [\log f_t^N - \log \phi_t^N] d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \\ &= N^2 \int f_t^N \mathcal{L}_N [\log \frac{f_t^N}{\phi_t^N}] d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \\ &= N^2 \int \frac{f_t^N}{\phi_t^N} \mathcal{L}_N [\log \frac{f_t^N}{\phi_t^N}] \phi_t^N d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \\ &\leq N^2 \int \mathcal{L}_N [\frac{f_t^N}{\phi_t^N}] \phi_t^N d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \\ &= N^2 \int \frac{f_t^N}{\phi_t^N} \mathcal{L}_N^* \phi_t^N d\nu_* - \int \partial_t \phi_t^N \frac{f_t^N}{\phi_t^N} d\nu^* \end{aligned}$$

where we used that for any positive function h , $h\mathcal{L}_N(\log h) \leq h$ (this is a consequence of Jensen's inequality). \square

We define $\xi_x := (\mathcal{E}_x, r_x)$ and $\pi(t, q) := (\mathbf{e}(t, q), \mathbf{r}(t, q))$. If f is a vectorial function, we denote its differential by Df .

Proposition 2.2.4 ([47]). *Let (λ, β) be defined by $\beta = (\mathbf{e} - \frac{\mathbf{r}^2}{2})^{-1}$ and $\lambda = -\beta\mathbf{r}$. The term $(\phi_t^N)^{-1} (N^2\mathcal{L}_N^*\phi_t^N - \partial_t\phi_t^N)$ is given by*

$$\begin{aligned} & (\phi_t^N)^{-1} (N^2\mathcal{L}_N^*\phi_t^N - \partial_t\phi_t^N) \\ &= \sum_{k=1}^5 \sum_{x \in \mathbb{T}_N} v_k \left(t, \frac{x}{N} \right) \left[J_x^k - H_k \left(\pi \left(t, \frac{x}{N} \right) \right) - (DH_k) \left(\pi \left(t, \frac{x}{N} \right) \right) \cdot \left(\xi_x - \pi \left(t, \frac{x}{N} \right) \right) \right] \\ & \qquad \qquad \qquad + o(N) \quad (2.49) \end{aligned}$$

where

k	J_x^k	$H_k(\mathbf{e}, \mathbf{r})$	$v_k(t, q)$
1	$p_x^2 + r_x r_{x-1} + 2\gamma p_x r_{x-1}$	$\mathbf{e} + \mathbf{r}^2/2$	$-(2\gamma)^{-1} \partial_q^2 \beta(t, q)$
2	$r_x + \gamma p_x$	\mathbf{r}	$-\gamma^{-1} \partial_q^2 \lambda(t, q)$
3	$p_x^2 (r_x + r_{x-1})^2$	$(2\mathbf{e} - \mathbf{r}^2) (\mathbf{e} + 3\mathbf{r}^2/2)$	$(4\gamma)^{-1} [\partial_q \beta(t, q)]^2$
4	$p_x^2 (r_x + r_{x-1})$	$\mathbf{r} (2\mathbf{e} - \mathbf{r}^2)$	$\gamma^{-1} \partial_q \beta(t, q) \partial_q \lambda(t, q)$
5	p_x^2	$\mathbf{e} - \mathbf{r}^2/2$	$\gamma^{-1} [\partial_q \lambda(t, q)]^2$

Observe that $H_k(e, r)$ is equal to $\int J_x^k d\nu_{\beta, \tau}$ where β, τ are related to e, r by the thermodynamic relations. Thus, the terms appearing in the righthand side of (2.49) can be seen as first order ‘‘Taylor expansion’’. The form of the first order correction in (2.47) plays a crucial role in order to get such expansions. A priori the first term on the right-hand side of (2.49) is of order N , but we want to take advantage of these microscopic Taylor expansions to show it is in fact of order $o(N)$.

First, we need to cut-off large energies in order to work with bounded variables only. To simplify, we assume they are bounded ab initio.

Let ℓ be some integer (dividing N). We introduce some averaging over microscopic blocks of size ℓ and we will let $\ell \rightarrow \infty$ after $N \rightarrow \infty$. We decompose \mathbb{T}_N in a disjoint union of $p = N/\ell$ boxes $\Lambda_\ell(x_j)$ of length ℓ centered in x_j , $j \in \{1, \dots, p\}$. The microscopic averaged profiles in a box of size ℓ around $y \in \mathbb{T}_N$ are defined by

$$\tilde{\xi}_\ell(y) = \frac{1}{\ell} \sum_{x \in \Lambda_\ell(y)} \xi_x.$$

Similarly we define

$$\tilde{J}_\ell^k(y) = \frac{1}{\ell} \sum_{x \in \Lambda_\ell(y)} J_x^k.$$

In (2.49) we rewrite the sum $\sum_{x \in \mathbb{T}_N}$ as $\sum_{j=1}^p \sum_{x \in \Lambda_\ell(x_j)}$ and, by using the smoothness of the function v_k , H_k , it is easy to replace the term

$$\frac{1}{N} \sum_{x \in \mathbb{T}_N} v_k \left(t, \frac{x}{N} \right) \left[J_x^k - H_k \left(\pi \left(t, \frac{x}{N} \right) \right) - (DH_k) \left(\pi \left(t, \frac{x}{N} \right) \right) \cdot \left(\xi_x - \pi \left(t, \frac{x}{N} \right) \right) \right]$$

by

$$\frac{1}{p} \sum_{j=1}^p v_k \left(t, \frac{x_j}{N} \right) \left[\tilde{J}_\ell^k(x_j) - H_k \left(\pi \left(t, \frac{x_j}{N} \right) \right) - (DH_k) \left(\pi \left(t, \frac{x_j}{N} \right) \right) \cdot \left(\tilde{\xi}_\ell(x_j) - \pi \left(t, \frac{x_j}{N} \right) \right) \right]$$

in the limit $N, \ell \rightarrow \infty$ with some error term of order $o(1)$.

Then, the strategy consists to prove the crucial estimate, often called the *one-block estimate*: we replace the empirical averaged current $\tilde{J}_\ell^k(x_j)$ which is averaged over a box centered at x_j by its mean with respect to a Gibbs measure with the parameters corresponding to the microscopic averaged profiles $\tilde{\xi}_\ell(x_j)$, i.e. $H_k(\tilde{\xi}_\ell(x_j))$. This is achieved thanks to some compactness argument and the macro-ergodicity of the dynamics.

Consequently we have to deal with terms in the form

$$\begin{aligned} \frac{1}{p} \sum_{j=1}^p v_k \left(t, \frac{x_j}{N} \right) & \left[H_k \left(\tilde{\xi}_\ell(x_j) \right) - H_k \left(\pi \left(t, \frac{x_j}{N} \right) \right) \right. \\ & \left. - (DH_k) \left(\pi \left(t, \frac{x_j}{N} \right) \right) \cdot \left(\tilde{\xi}_\ell(x_j) - \pi \left(t, \frac{x_j}{N} \right) \right) \right] \end{aligned} \quad (2.50)$$

The final step consists then to apply the entropy inequality (2.44) with respect to $\hat{\mu}_t^N$ with $\phi := \phi_{\ell, N}$ given by (2.50) and $\alpha = \delta N$, $\delta > 0$ fixed but small. This will produce some term of order $\hat{H}_N(t)/N$ plus the term

$$\limsup_{\ell \rightarrow \infty} \limsup_{N \rightarrow \infty} \frac{1}{\delta N} \log \left(\int e^{\delta N \phi} d\hat{\mu}_t^N \right) = I(\delta).$$

By using some large deviations estimates (observe that $\hat{\mu}_t^N$ is explicit and product at first order in N) one can show that $I(\delta)$ is nonpositive for δ sufficiently small. Thus we get the desired Gronwall inequality.

There is some additional difficulty that we hid under the carpet in the sketch of the proof. Since the state space is non compact, a control of high energies is required

for the initial cut-off. This is a highly non trivial problem ¹⁰. In the harmonic case considered here this control is obtained thanks to the following remark: the set of mixtures of Gaussian probability measures ¹¹ is preserved by the (harmonic) velocity-flip model. Since for Gaussian measures all the moments are expressed in terms of the covariance matrix, required bounds can be obtained by a suitable control of the covariance matrices appearing in the mixture.

The extension of this results in the anharmonic case is a challenging open problem (see however [45]).

2.2.2 Fourier's law

Since in the harmonic case an exact fluctuation-dissipation equation is available Fourier's law can be obtained without too much work ¹².

Theorem 2.2.5 ([11, 12]). *Consider the one-dimensional harmonic chain in contact with two heat baths and with forced boundary conditions as in Section 2.1.1. Then Fourier's law holds:*

$$\tilde{J}_s := \lim_{N \rightarrow \infty} N \langle j_{0,1}^e \rangle_{ss} = \frac{1}{2\gamma} \{ (T_\ell - T_r) + (\tau_\ell^2 - \tau_r^2) \} \quad (2.51)$$

and we have

$$\begin{aligned} \hat{J}_\ell &= \lim_{N \rightarrow \infty} N (\langle p_1^2 \rangle_{ss} - T_\ell) = \frac{1}{2\gamma\gamma_\ell} [(T_r - T_\ell) + (\tau_\ell - \tau_r)^2], \\ \hat{J}_r &= \lim_{N \rightarrow \infty} N (T_r - \langle p_N^2 \rangle_{ss}) = \frac{1}{2\gamma\gamma_r} [(T_r - T_\ell) - (\tau_\ell - \tau_r)^2]. \end{aligned} \quad (2.52)$$

It follows that the system can be used as a heater but not as a refrigerator ¹³.

Proof. We divide the proof in two steps:

¹⁰A similar problem appears in [44] where the authors derived Euler equations for a gas perturbed by some ergodic noise. There, to overcome this difficulty, the authors replace ab initio the kinetic energy by the relativistic kinetic energy.

¹¹A Gibbs local equilibrium state is a Gaussian state in the harmonic case.

¹²The a posteriori simple but fundamental remark that an exact fluctuation-dissipation equation exists for the harmonic model is the real contribution of [11].

¹³Assume that $T_r > T_\ell$. The term \hat{J}_ℓ (resp. \hat{J}_r) is the macroscopic heat current from the left reservoir to the system (resp. from the system to the right reservoir). Whatever the values of τ_ℓ, τ_r are, $\hat{J}_\ell > 0$ and we can not realize a refrigerator. But if $(T_r - T_\ell) < (\tau_r - \tau_\ell)^2$ then $\hat{J}_r < 0$ and we realized a heater.

- We first prove that there exists a constant C independent of N such that $|\langle j_{0,1}^e \rangle_{ss}| \leq C/N$. This is obtained by using the fluctuation-dissipation equation and the fact that $\langle j_{x,x+1}^e \rangle_{ss}$ is independent of x :

$$\begin{aligned}
\langle j_{0,1}^e \rangle_{ss} &= \frac{1}{N-3} \sum_{x=2}^{N-2} \langle j_{x,x+1}^e \rangle_{ss} \\
&= -\frac{1}{2\gamma} \frac{1}{N-3} \sum_{x=2}^{N-2} \langle \nabla [p_x^2 + r_x r_{x+1}] \rangle_{ss} \\
&= \frac{1}{2\gamma} \frac{1}{N-3} \{ (\langle p_2^2 \rangle_{ss} + \langle r_2 r_3 \rangle_{ss}) - (\langle p_{N-1}^2 \rangle_{ss} + \langle r_{N-1} r_N \rangle_{ss}) \}.
\end{aligned} \tag{2.53}$$

By using simple computations, one can show that $(\langle p_2^2 \rangle_{ss} + \langle r_2 r_3 \rangle_{ss}) - (\langle p_{N-1}^2 \rangle_{ss} + \langle r_{N-1} r_N \rangle_{ss})$ is uniformly bounded in N by a positive constant.

- Now we have only to evaluate the limit of each term appearing in $(\langle p_2^2 \rangle_{ss} + \langle r_2 r_3 \rangle_{ss}) - (\langle p_{N-1}^2 \rangle_{ss} + \langle r_{N-1} r_N \rangle_{ss})$. Notice that assuming local equilibrium we get easily the result. The first step shows that $\langle j_{0,1}^e \rangle_{ss}$ and $\langle j_{N,N+1}^e \rangle_{ss}$ vanish as $N \rightarrow +\infty$. Since V_s goes to 0 by Lemma 2.1.1, one has that $\langle p_1^2 \rangle_{ss}$ and $\langle p_N^2 \rangle_{ss}$ converge respectively to T_ℓ and T_r . By using some “entropy production bound” one can propagate this local equilibrium information to the particles close to the boundaries and show (2.51).

□

The proof of the validity of Fourier’s law for anharmonic chains perturbed by an energy conserving noise is still open.

2.2.3 Macroscopic Fluctuation Theory for the energy conserving harmonic chain

The macroscopic fluctuation theory ([17]) is a general approach developed by Bertini, De Sole, Gabrielli, Jona-Lasinio and Landim to calculate the large deviation functional of the empirical profiles of the conserved quantities of Markov processes in a NESS. Its main interest is that it can be applied to a large class of boundary driven diffusive systems and does not require the explicit form of the NESS but only the knowledge of two thermodynamic macroscopic parameters of the system, the diffusion coefficient $D(\rho)$ and the mobility $\chi(\rho)$. This theory can be seen as an infinite dimensional generalization of the Freidlin-Wentzel theory [26] and is based on the large deviations principle for the hydrodynamics of the system.

In order to explain (roughly) the theory we consider for simplicity a Markovian system $\{\eta(t) := \{\eta_x(t) \in \mathbb{R}; x \in \{1, \dots, N\}\}_{t \geq 0}$ with only one conserved quantity, say the density ρ , in contact with two reservoirs at each extremity. Here N is the size of the system which will be sent to infinity and we denote by μ_{ss}^N the nonequilibrium stationary state of $\{\eta(t)\}_{t \geq 0}$. For any microscopic configuration $\eta := \{\eta_x; x \in \{1, \dots, N\}\}$ let

$$\pi^N(\eta, \cdot) = \sum_{x=1}^{N-1} \eta_x \mathbf{1}_{\left[\frac{x}{N}, \frac{x+1}{N}\right)}(\cdot)$$

be the empirical density profile. In the diffusive time scale, we assume that $\pi^N(\eta(tN^2), \cdot)$ converges as N goes to infinity to $\rho_t(\cdot) := \rho(t, \cdot)$ solution of

$$\begin{cases} \partial_t \rho = \partial_y (D(\rho) \partial_y \rho), & y \in [0, 1], \quad t \geq 0, \\ \rho(t, 0) = \rho_\ell, \quad \rho(t, 1) = \rho_r, & t \geq 0, \\ \rho(0, \cdot) = \rho_0(\cdot) \end{cases}$$

where $\rho_0(\cdot)$ is the initial density profile, $D(\rho)$ is the diffusion coefficient and ρ_ℓ, ρ_r the densities fixed by the reservoirs. As $t \rightarrow \infty$ the solution ρ_t of the hydrodynamic equation converges to a stationary profile $\bar{\rho} : [0, 1] \rightarrow \mathbb{R}$ solution of $D(\bar{\rho}) \partial_y \bar{\rho} = J = \text{const.}$ with $\bar{\rho}(0) = \rho_\ell, \bar{\rho}(1) = \rho_r$. We assume that under μ_{ss}^N , the empirical density profile $\pi^N(\eta, \cdot)$ converges to $\bar{\rho}$. This assumption is nothing but a law of large numbers for the random variables π^N .

We are here interested in the corresponding large deviations principle. Thus, we want to estimate the probability that in the NESS μ_{ss}^N the empirical density profile π^N is close to an atypical macroscopic profile $\rho(\cdot) \neq \bar{\rho}$. This probability will be typically of order $e^{-NV(\rho)}$ where V is the rate function:

$$\mu_{ss}^N(\pi^N(\eta, \cdot) \approx \rho(\cdot)) \approx e^{-NV(\rho)}.$$

The goal of the macroscopic fluctuation theory is to obtain informations about this functional.

The condition filled by the system to apply the theory of Bertini et al. is that it satisfies a *dynamical large deviation principle* with a rate function which takes the following form ¹⁴.

Let us first explain what we mean by dynamical large deviations principle. Imagine we start the system from a Gibbs local equilibrium state corresponding to the macroscopic profile ρ_0 . We want to estimate the probability that the empirical density $\pi^N(\eta(tN^2), \cdot)$ is close during the macroscopic time interval $[0, T]$, T fixed,

¹⁴Such property has been proved to be valid for a large class of stochastic dynamics ([36]).

to a smooth macroscopic profile $\gamma(t, y)$ supposed to satisfy ¹⁵ $\gamma(0, \cdot) = \rho_0$. This probability is exponentially small in N with a rate $I_{[0, T]}(\gamma | \rho_0)$

$$\mathbb{P} \left[\pi^N(\eta(tN^2), y) \approx \gamma(t, y), (t, y) \in [0, T] \times [0, 1] \right] \sim e^{-NI_{[0, T]}(\gamma | \rho_0)}. \quad (2.54)$$

The rate function is given by

$$I_{[0, T]}(\gamma | \rho_0) = \frac{1}{2} \int_0^T dt \int_0^1 dy \chi(\rho(t, y)) [(\partial_y H)(t, y)]^2$$

where $\partial_y H$ is the extra gradient external field needed to produce the fluctuation γ , namely such that

$$\partial_t \gamma = \partial_y [D(\gamma) \partial_y \gamma - \chi(\gamma) \partial_y H]. \quad (2.55)$$

Thus, $I_{[0, T]}(\gamma | \rho_0)$ is the work done by the external field $\partial_y H$ to produce the fluctuation γ in the time interval $[0, T]$. The function χ appearing in (2.55) is the second thermodynamic parameter (with the diffusion coefficient D) mentioned in the beginning of this section.

To show this result the strategy is the following. We perturb the Markov process $\{\eta(t)\}_{t \geq 0}$ thanks to the function $H := H(\gamma)$, which is solution of the Poisson equation (2.55), by adding locally a small space inhomogeneous drift provided by $\partial_y H$. In doing so we obtain a new Markov process $\{\eta^H(t)\}_{t \geq 0}$ such that in the diffusive time scale $\pi^N(\eta^H(tN^2), \cdot)$ converges to $\gamma(\cdot)$. Let \mathbb{P}^H (resp. \mathbb{P}^0) be the probability measure on the empirical density paths space induced by $\{\eta^H(tN^2)\}_{t \in [0, T]}$ (resp. $\{\eta(tN^2)\}_{t \in [0, T]}$). Then, by using hydrodynamic limits techniques similar to the ones explained in Section 2.2.1 we show that in the large N limit, under \mathbb{P}^H , the Radon-Nikodym derivative is well approximated by ¹⁶

$$\frac{d\mathbb{P}^0}{d\mathbb{P}^H}(\pi) \approx \exp \left\{ -NI_{[0, T]}(\pi | \rho_0) \right\}.$$

Here $\pi := \{\pi(t, y); t \in [0, T], y \in [0, 1]\}$ is any space-time density profile. Thus, since

$$\mathbb{P} \left[\pi^N(\eta(tN^2), \cdot) \sim \gamma(t, \cdot), t \in [0, T] \right] = \mathbb{E}^H \left[\frac{d\mathbb{P}^0}{d\mathbb{P}^H}(\pi) \mathbf{1}_{\{\pi(t, \cdot) \sim \gamma(t, \cdot), t \in [0, T]\}} \right]$$

we obtain (2.54).

¹⁵This assumption avoid to take into account the cost to produce the initial profile, cost which is irrelevant for us.

¹⁶We use Girsanov transform to express the Radon-Nikodym derivative. A priori it is note a functional of the empirical density and we need to establish some *replacement lemma* (see [36]).

The macroscopic fluctuation theory claims that the large deviations functional $V(\rho)$ of the empirical density in the NESS coincides with the quasi-potential $W(\rho)$ defined by

$$W(\rho) = \inf_{\substack{\gamma: \gamma(-\infty) = \bar{\rho} \\ \gamma(0) = \rho}} I_{[-\infty, 0]}(\gamma | \bar{\rho}).$$

Here $I_{[-\infty, 0]}$ is obtained from $I_{[0, T]}$ by a shift in time by $-T$, T being sent to $+\infty$ afterwards. In words, the quasi potential determines the cost to produce a fluctuation equal to γ at $t = 0$ when the system is macroscopically in the stationary profile $\bar{\rho}$ at $t = -\infty$.

Thus, the problem is reduced to compute W . It can be shown that W solves (at least formally) the infinite-dimensional Hamilton-Jacobi equation

$$\frac{1}{2} \left\langle \partial_y \left[\frac{\delta W}{\delta \rho} \right], \chi(\rho) \partial_y \left[\frac{\delta W}{\delta \rho} \right] \right\rangle + \left\langle \frac{\delta W}{\delta \rho}, \partial_y [D(\rho) \partial_y \rho] \right\rangle = 0 \quad (2.56)$$

where $\langle \cdot, \cdot \rangle$ denotes the usual scalar product in $\mathbb{L}^2([0, 1])$. Note that there is no uniqueness of solutions ($W = 0$ is a solution) and up to now a general theory of infinite dimensional Hamilton-Jacobi equations is still missing. This implies that we have in fact to solve by hand the variational problem and the solution is only known for few systems. This is an important limitation of the macroscopic fluctuation theory. Even getting interesting qualitative properties on W is difficult.

The rigorous implementation of this long program has only been carried for the boundary driven Symmetric Simple Exclusion Process and extended without less rigor to few other systems (see [15], [18], [25] for rigorous results).

Let us now try to apply this theory for the harmonic chain with velocity-flip noise. Since we have a fully explicit microscopic fluctuation-dissipation equation (also if some harmonic pinning is added) we can easily guess what is the form of the hydrodynamic equations under various boundary conditions by assuming that the propagation of local equilibrium in the diffusive time scale holds. Nevertheless, let us observe that a rigorous derivation is missing, the obstacle being a sufficiently good control of the high energies¹⁷. The boundary conditions we impose to the system are the following. At the left (resp. right) we put the chain in contact with a Langevin bath at temperature T_ℓ (resp. T_r) and consider the system with fixed boundary conditions or with forced boundary conditions with the same force τ at the two boundaries. Then, for the unpinned chain, the equations (2.42) are still valid but they are supplemented by the boundary conditions ([14])

$$\left[\mathbf{e} - \frac{\mathbf{r}^2}{2} \right] (t, 0) = T_\ell, \quad \left[\mathbf{e} - \frac{\mathbf{r}^2}{2} \right] (t, 1) = T_r, \quad (2.57)$$

¹⁷This control is only available in the case of periodic boundary conditions ([47]).

since the Langevin baths fix the temperatures at the boundaries and

$$\partial_y \mathbf{r}(t, 0) = \partial_y \mathbf{r}(t, 1) = 0 \quad (2.58)$$

for fixed boundary conditions (the total length of the chain is constant) and

$$\mathbf{r}(t, 0) = \mathbf{r}(t, 1) = \tau \quad (2.59)$$

for forced boundary conditions.

If the chain is pinned by the harmonic potential $W(q) = \nu q^2/2$ then only the energy is conserved and the macroscopic diffusion equation takes the form

$$\begin{cases} \partial_t \mathbf{e} = \partial_y (\kappa \partial_y \mathbf{e}), \\ \mathbf{e}(0, y) = \mathbf{e}_0(y), \\ \mathbf{e}(t, 0) = T_\ell, \mathbf{e}(t, 1) = T_r, \end{cases} \quad y \in (0, 1) \quad (2.60)$$

where the conductivity κ is equal to ([14])

$$\kappa = \frac{1/\gamma}{2 + \nu^2 + \sqrt{\nu(\nu + 4)}}. \quad (2.61)$$

Assuming a good control of high energies, it is possible to derive the dynamical large deviations function of the empirical conserved quantities. The goal would be to compute the large deviation functional of the NESS which according to the macroscopic fluctuation theory coincides with the quasi potential. We recall that the quasi potential is defined by a variational problem and that it depends only on two thermodynamic quantities, the diffusion coefficient and the mobility (the latter are matrices if several conserved quantities are involved).

Let us first consider the pinned velocity flip model where the energy is the only conserved quantity. It turns out that the mobility is a quadratic function. Consequently, the methods exposed in Theorem 6.5 of [16] apply and the variational formula can be computed. The quasi potential $V(\cdot)$ is given by ([14])

$$V(e) = \int_0^1 dq \left[\frac{e(q)}{F(q)} - 1 - \log \left(\frac{e(q)}{F(q)} \right) - \log \left(\frac{F'(q)}{T_r - T_\ell} \right) \right], \quad (2.62)$$

where F is the unique increasing solution of

$$\begin{cases} \frac{\partial_q^2 F}{(\partial_q F)^2} = \frac{F - e}{F^2}, \\ F(0) = T_\ell, F(1) = T_r. \end{cases} \quad (2.63)$$

Surprisingly, the function V is independent of the pinning value ν and of the intensity of the noise γ . It is thus natural to conjecture that in the NESS of the unpinned velocity flip model the large deviation function of the empirical energy profile coincides with V but we did not succeed to prove it. Observe that at equilibrium ($T_\ell = T_r$), the last term in (2.62) disappears and the quasi potential is local. On the other hand, if $T_\ell \neq T_r$, this is no longer the case and this reflects the presence of long-range correlations in the NESS. In particular, an approximation of the NESS by a Gibbs local equilibrium state in the form (2.8) would not give the correct value of the quasi potential.

For the unpinned chain we have two conserved quantities. Solving the corresponding variational problem is a very difficult problem¹⁸ and is open (see [5] for a partial result for some other stochastic perturbation of the harmonic chain).

¹⁸Here we do not have any exactly solvable model like the Symmetric Simple Exclusion Process which could give us some hints for the form of the quasi-potential.

Chapter 3

Anomalous diffusion

Anomalous large conductivity is observed experimentally in carbon nanotubes and numerically in chains of oscillators without pinning, where numerical evidence shows a conductivity diverging with the size N of the system like N^α , with $\alpha < 1$ in dimension $d = 1$, and like $\log N$ in dimension $d = 2$. If some nonlinearity is present in the interaction, finite conductivity is observed numerically in all pinned case or in dimension $d \geq 3$. Consequently it has been suggested that conservation of momentum is an important ingredient for the anomalous conductivity in low dimensions.

In Chapter 2 we considered chains of oscillators perturbed by a noise conserving only energy and destroying the possible momentum conservation. In the harmonic case we obtained Fourier's law and in the anharmonic case we proved existence of the Green-Kubo formula for the thermal conductivity.

In this chapter the perturbation added conserves both energy and momentum (energy and volume for the Hamiltonian systems considered in Section 1.2). These systems have qualitatively the same behavior as Hamiltonian chains of oscillators (without any noise), i.e. anomalous transport for unpinned chain in dimension $d = 1, 2$ and normal transport otherwise. We could even be more optimistic and hope that they share with the deterministic systems common limits for the energy fluctuation fields, two point correlation functions . . . This is because one expect that the microscopic details of the dynamics are irrelevant. Therefore some *universality* should hold. Recently H. Spohn ([48]), by following ideas of [54], used the nonlinear fluctuating hydrodynamics theory to classify very precisely the different expected universality classes. The nonlinear fluctuating hydrodynamics theory is based on the assumption that the microscopic dynamics evolve in the Euler time scale according to a system of conservation laws. The theory is macroscopic in the sense that all the predictions are done starting from this system of conservation laws without further references to the microscopic dynamics. Since we have seen that the presence of

the energy-momentum conserving noise does not change the form of the hydrodynamic equations, the theory claims in fact that the limit of the fluctuations fields of the conserved quantities for purely deterministic chains of oscillators and for noisy energy-momentum conserving chains are exactly the same.

3.1 Harmonic chains with momentum exchange noise

Getting some informations on the behavior of the energy fluctuation field in the large scale limit remains challenging. So far, satisfactory but not complete results have only been obtained in the harmonic case. The anharmonic case is much more difficult.

In [1], [2] we compute explicitly the time correlation current for a system of harmonic oscillators perturbed by an energy-momentum conserving noise¹ and we find that it behaves, for large time, like $t^{-d/2}$ in the unpinned cases, and like $t^{-d/2-1}$ when on-site harmonic potential is present.

These results are given in the Green-Kubo formalism. Their counterpart in the NESS formalism have been considered in [40] but a rigorous proof is still missing. Several variations of the Green-Kubo formula can be found in the literature: one can start with the infinite system in the canonical ensemble, as we did in Subsection 2.1.2, or with a finite system, in the canonical or micro-canonical ensemble, sending the size of the system to infinity. It is widely believed that all these definitions coincide (also in the case of infinite conductivity). As shown in [2], this is essentially true for the energy-momentum conserving harmonic chain. Here we consider the simplest possible definition avoiding to discuss the rigorous definition of the canonical ensemble in infinite volume and the equivalence of ensembles problem.

The set-up is the following. We consider a chain perturbed by the energy-momentum conserving noise (see (1.10)) with periodic boundary conditions. Its Hamiltonian is given by $\mathcal{H}_N = \sum_{x \in \mathbb{T}_N^d} \mathcal{E}_x$ where the energy \mathcal{E}_x of atom x is

$$\mathcal{E}_x = \frac{|p_x|^2}{2} + W(q_x) + \frac{1}{2} \sum_{|y-x|=1} V(q_x - q_y). \quad (3.1)$$

The system is considered at equilibrium under the Gibbs grand-canonical mea-

¹It is straightforward to adapt the proofs given in [1], [2] to the case of the momenta exchange noise.

sure

$$d\mu_{N,T} = \frac{e^{-\mathcal{H}_N/T}}{Z_{N,T}} d\mathbf{q}d\mathbf{p}$$

where $Z_{N,T}$ is the renormalization constant.

The Green-Kubo formula for the thermal conductivity in the direction e_k , $1 \leq k \leq d$, is ² the limiting variance of the energy current $J_{x,x+e_k}^{e,\gamma}([0,t])$ up to time t in the direction e_k in a space-time box of size $N \times t$:

$$\kappa(T) = \frac{1}{2T^2} \lim_{t \rightarrow +\infty} \lim_{N \rightarrow \infty} \mathbb{E}_{\mu_{N,T}} \left[\left(\frac{1}{\sqrt{N^d t}} \sum_{x \in \mathbb{T}_N^d} J_{x,x+e_1}^{e,\gamma}([0,t]) \right)^2 \right]. \quad (3.2)$$

The energy currents $\{J_{x,x+e_k}^{e,\gamma}([0,t]); k = 1, \dots, d\}$ are defined by the energy conservation law

$$\mathcal{E}_x(t) - \mathcal{E}_x(0) = \sum_{k=1}^d (J_{x-e_k,x}^{e,\gamma}([0,t]) - J_{x,x+e_k}^{e,\gamma}([0,t])).$$

The energy current up to time t can be written as

$$J_{x,x+e_k}^{e,\gamma}([0,t]) = \int_0^t j_{x,x+e_k}^{e,\gamma}(s) ds + M_{x,x+e_k}(t) \quad (3.3)$$

where $M_{x,x+e_k}(t)$ is a martingale and $j_{x,x+e_k}^{e,\gamma}$ is the instantaneous current which has the form

$$j_{x,x+e_k}^{e,\gamma} = \tilde{j}_{x,x+e_k}^e + \frac{\gamma}{2} [p_{x+e_k}^2 - p_x^2], \quad \tilde{j}_{x,x+e_k}^e = -\frac{1}{2} V'(q_{x+e_k} - q_x)(p_{x+e_k} + p_x). \quad (3.4)$$

The term $\tilde{j}_{x,x+e_k}^e$ is the Hamiltonian contribution while the gradient term is due to the noise.

We expand now the square in (3.2). Notice first that since we have periodic boundary conditions the gradient term appearing in (3.4) does not contribute. By a time reversal argument one can show that the cross term between the martingale and the time integral of the instantaneous current vanishes. Moreover a simple computation shows that the square of the martingale term gives a contribution equal to γ (see [2] for details). Thus we obtain

$$\begin{aligned} \kappa(T) &= T^{-2} \lim_{t \rightarrow +\infty} \lim_{N \rightarrow \infty} \frac{1}{2N^d t} \mathbb{E}_{\mu_{N,T}} \left[\left(\sum_{x \in \mathbb{T}_N^d} \int_0^t \tilde{j}_{x,x+e_k}^e(s) ds \right)^2 \right] + \gamma \\ &= T^{-2} \lim_{t \rightarrow +\infty} \lim_{N \rightarrow \infty} \sum_{x \in \mathbb{T}_N^d} \int_0^{+\infty} ds \left(1 - \frac{s}{t}\right)^+ \mathbb{E}_{\mu_{N,T}} [\tilde{j}_{0,e_k}^e(0) \tilde{j}_{x,x+e_k}^e(s)] ds + \gamma \end{aligned} \quad (3.5)$$

²By symmetry arguments this is independent of k .

where the last line is obtained by time and space stationarity of the Gibbs measure and u^+ denotes $\max(u, 0)$ ³. It is then clear that the divergence of the Green-Kubo formula, i.e. anomalous transport, is due to a slow decay of the time correlation function $C(t)$ defined by

$$C(t) = \lim_{N \rightarrow \infty} \sum_{x \in \mathbb{T}_N^d} \mathbb{E}_{\mu_{N,T}} [\tilde{j}_{0,e_k}^e(0) \tilde{j}_{x,x+e_k}^e(t)]. \quad (3.6)$$

Theorem 3.1.1 ([2]). *Consider the harmonic case: $V(r) = \alpha r^2$, $W(q) = \nu q^2$ where $\alpha > 0$ and $\nu \geq 0$.*

Then the limit defining $C(t)$ in (3.6) exists and can be computed explicitly. In particular, we have that $C(t) \sim t^{-d/2}$ if $\nu = 0$ and $C(t) \sim t^{-d/2-1}$ if $\nu > 0$.

Consequently, the limit (3.5) exists in $(0, +\infty]$ and is finite if and only if $d \geq 3$ or $\nu > 0$. When finite, $\kappa(T)$ is independent of T and can be computed explicitly.

Proof. We compute the Laplace transform $L_N(z) = \int_0^{+\infty} e^{-zt} C_N(t) dt$, $z > 0$, of $C_N(t) = \sum_{x \in \mathbb{T}_N^d} \mathbb{E}_{\mu_{N,T}} [\tilde{j}_{0,e_k}^e(0) \tilde{j}_{x,x+e_k}^e(t)]$. Since we have

$$L_N(z) = N^{-1} \mu_{N,T} \left[\left(\sum_{x \in \mathbb{T}_N^d} \tilde{j}_{x,x+e_k}^e \right) (z - \mathcal{L}_N)^{-1} \left(\sum_{x \in \mathbb{T}_N^d} \tilde{j}_{x,x+e_k}^e \right) \right]$$

it is equivalent to solve the resolvent equation $(z - \mathcal{L}_N)h_N = \sum_{x \in \mathbb{T}_N^d} \tilde{j}_{x,x+e_k}^e$. Notice that \mathcal{L}_N maps polynomial functions of degree 2 into polynomial functions of degree 2 and that $\sum_x \tilde{j}_{x,x+e_k}^e$ is a polynomial function of degree 2. Thus, the function h_N is a polynomial function of degree 2. Moreover it has to be space translation invariant since $\sum_x \tilde{j}_{x,x+e_k}^e$ is. Therefore we can look for a function h_N in the form

$$h_N = \sum_{x,y} a(y-x) p_x p_y + \sum_{x,y} b(y-x) p_x q_y + \sum_{x,y} c(y-x) q_x q_y$$

where a, b and c are functions from \mathbb{T}_N^d into \mathbb{R} . We compute explicitly a, b and c and we get $a = c = 0$ and b is the solution to

$$(z + 2\nu - \gamma\Delta)b = -\alpha(\delta_{e_k} - \delta_{-e_k})$$

where Δ is the discrete Laplacian. Then we deduce $L_N(z)$, then $C_N(t)$, and the limit $C(t) = \lim_{N \rightarrow +\infty} C_N(t)$. \square

³Observe that replacing $(1 - \frac{s}{t})^+$ by $e^{-s/t}$ and $\lim_{N \rightarrow \infty} \sum_{x \in \mathbb{T}_N^d}$ by $\sum_{x \in \mathbb{Z}^d}$ we get formally an expression similar to the Green-Kubo formula of Theorem 2.1.2.

Consequently in the unpinned harmonic cases in dimension $d = 1$ and 2 , the conductivity of our model diverges. Otherwise it converges. In the anharmonic case we obtained some upper bounds showing that the divergence cannot be worse than in the harmonic case. These upper bounds also show that the conductivity cannot be infinite if $d \geq 3$ (see [2] for details and precise statements).

3.2 A class of perturbed Hamiltonian systems

In [13] is proposed a class of models for which anomalous diffusion is observed. These models have been introduced in Section 1.2 of Chapter 1. The goal of [13] was to show that these systems have a behavior very similar to that of the standard one-dimensional chains of oscillators conserving momentum ⁴.

3.2.1 Definition of thermodynamic variables

Let us fix a potential V and consider the stochastic energy-volume conserving model defined by the generator $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$, $\gamma \geq 0$, where \mathcal{A} and \mathcal{S} are given by (1.20). Recall that the Gibbs grand-canonical probability measures $\mu_{\beta,\lambda}$, $\beta > 0$, $\lambda \in \mathbb{R}$, defined on Ω by

$$d\mu_{\beta,\lambda}(\eta) = \prod_{x \in \mathbb{Z}} Z(\beta, \lambda)^{-1} \exp\{-\beta V(\eta_x) - \lambda \eta_x\} d\eta_x$$

form a family of invariant probability measures for the infinite dynamics. We assume that the partition function Z is well defined on $(0, +\infty) \times \mathbb{R}$. The following thermodynamic relations relate the chemical potentials β, λ to the mean volume v and the mean energy e under $\mu_{\beta,\lambda}$:

$$\begin{aligned} v(\beta, \lambda) &= \mu_{\beta,\lambda}(\eta_x) = -\partial_\lambda \left(\log Z(\beta, \lambda) \right), \\ e(\beta, \lambda) &= \mu_{\beta,\lambda}(V(\eta_x)) = -\partial_\beta \left(\log Z(\beta, \lambda) \right). \end{aligned} \tag{3.7}$$

These relations can be inverted by a Legendre transform to express β and λ as a function of e and v . Define the thermodynamic entropy $S : (0, +\infty) \times \mathbb{R} \rightarrow [-\infty, +\infty)$ as

$$S(e, v) = \inf_{\lambda \in \mathbb{R}, \beta > 0} \left\{ \beta e + \lambda v + \log Z(\beta, \lambda) \right\}.$$

Let \mathcal{U} be the convex domain of $(0, +\infty) \times \mathbb{R}$ where $S(e, v) > -\infty$ and $\mathring{\mathcal{U}}$ its interior. Then, for any $(e, v) := (e(\beta, \lambda), v(\beta, \lambda)) \in \mathring{\mathcal{U}}$, the parameters β, λ can be obtained

⁴They could be defined in any dimension.

as

$$\beta = (\partial_e S)(e, v), \quad \lambda = (\partial_v S)(e, v). \quad (3.8)$$

We also introduce the tension $\tau(\beta, \lambda) = \mu_{\beta, \lambda}(V'(\eta_0)) = -\lambda/\beta$. The microscopic energy current $j_{x, x+1}^{e, \gamma}$ and volume current $j_{x, x+1}^{v, \gamma}$ are given by

$$\begin{aligned} j_{x, x+1}^{e, \gamma} &= -V'(\eta_x)V'(\eta_{x+1}) - \gamma \nabla[V(\eta_x)], \\ j_{x, x+1}^{v, \gamma} &= -[V'(\eta_x) + V'(\eta_{x+1})] - \gamma \nabla[\eta_x]. \end{aligned} \quad (3.9)$$

With these notations we have

$$\mu_{\beta, \lambda}(j_{x, x+1}^{e, \gamma}) = -\tau^2, \quad \mu_{\beta, \lambda}(j_{x, x+1}^{v, \gamma}) = -2\tau. \quad (3.10)$$

In the sequel, with a slight abuse of notation, we also write τ for $\tau(\beta(e, v), \lambda(e, v))$ where $\beta(e, v)$ and $\lambda(e, v)$ are defined by relations (3.8).

3.2.2 Hydrodynamic limits

Consider the finite *closed* stochastic energy-volume dynamics with periodic boundary conditions, that is the dynamics generated by $\mathcal{L}_{N, \text{per}} = \mathcal{A}_{N, \text{per}} + \gamma \mathcal{S}_{N, \text{per}}$ where

$$\left(\mathcal{A}_{N, \text{per}} f \right) (\eta) = \sum_{x \in \mathbb{T}_N} [V'(\eta_{x+1}) - V'(\eta_{x-1})] \partial_{\eta_x} f(\eta), \quad (3.11)$$

and

$$\left(\mathcal{S}_{N, \text{per}} f \right) (\eta) = \sum_{x \in \mathbb{T}_N} [f(\eta^{x, x+1}) - f(\eta)].$$

We choose to consider the dynamics on \mathbb{T}_N rather than on \mathbb{Z} to avoid (nontrivial) technicalities. We are interested in the macroscopic behavior of the two conserved quantities on a macroscopic time-scale Nt as $N \rightarrow \infty$.

We assume that the system is initially distributed according to a local Gibbs equilibrium state corresponding to a given energy-volume profile $X_0 : \mathbb{T} \rightarrow \mathcal{U}$:

$$X_0 = \begin{pmatrix} \mathbf{e}_0 \\ \mathbf{v}_0 \end{pmatrix},$$

in the sense that, for a given system size N , the initial state of the system is described by the following product probability measure:

$$d\mu_{\mathbf{e}_0, \mathbf{v}_0}^N(\eta) = \prod_{x \in \mathbb{T}_N} \frac{\exp \{ -\beta_0(x/N)V(\eta_x) - \lambda_0(x/N)\eta_x \}}{Z(\beta_0(x/N), \lambda_0(x/N))} d\eta_x, \quad (3.12)$$

where $(\beta_0(x/N), \lambda_0(x/N))$ is actually a function of $(\mathbf{e}_0(x/N), \mathbf{v}_0(x/N))$ through relations (3.8).

Starting from such a state, we expect the state of the system at time Nt to be close, in a suitable sense, to a local Gibbs equilibrium measure corresponding to an energy-volume profile

$$X(t, \cdot) = \begin{pmatrix} \mathbf{e}(t, \cdot) \\ \mathbf{v}(t, \cdot) \end{pmatrix},$$

satisfying a suitable partial differential equation with initial condition X_0 at time $t = 0$. In view of (3.10), and assuming propagation of local equilibrium, it is not difficult to show that the expected partial differential equation is the following system of two conservation laws:

$$\begin{cases} \partial_t \mathbf{e} - \partial_q \tau^2 = 0, \\ \partial_t \mathbf{v} - 2\partial_q \tau = 0, \end{cases} \quad (3.13)$$

with initial conditions $\mathbf{e}(0, \cdot) = \mathbf{e}_0(\cdot)$, $\mathbf{v}(0, \cdot) = \mathbf{v}_0(\cdot)$. We write (3.13) more compactly as

$$\partial_t X + \partial_q \mathfrak{J}(X) = 0, \quad X(0, \cdot) = X_0(\cdot),$$

with

$$\mathfrak{J}(X) = \begin{pmatrix} -\tau^2(\mathbf{e}, \mathbf{v}) \\ -2\tau(\mathbf{e}, \mathbf{v}) \end{pmatrix}. \quad (3.14)$$

The system of conservation laws (3.13) has other nontrivial conservation laws. In particular, the thermodynamic entropy S is conserved along a smooth solution of (3.13):

$$\partial_t S(\mathbf{e}, \mathbf{v}) = 0. \quad (3.15)$$

Since the thermodynamic entropy is a strictly concave function on $\mathring{\mathcal{U}}$, the system (3.13) is strictly hyperbolic on $\mathring{\mathcal{U}}$ (see [46]). The two real eigenvalues of $(D\mathfrak{J})(\bar{\xi})$ are 0 and $-\partial_e(\tau^2) + 2\partial_v(\tau)$, corresponding respectively to the two eigenvectors

$$\begin{pmatrix} -\partial_v \tau \\ \partial_e \tau \end{pmatrix}, \quad \begin{pmatrix} \tau \\ 1 \end{pmatrix}. \quad (3.16)$$

It is well known that classical solutions to systems of $n \geq 1$ conservation laws develop shocks, even when starting from smooth initial conditions. If we consider weak solutions rather than classical solutions, then a criterion is needed to select a unique, relevant solution among the weak ones. For scalar conservation laws ($n = 1$), this criterion is furnished by the so-called entropy inequality and existence and uniqueness of solutions is fully understood. If $n \geq 2$, only partial results exist (see [46]). This motivates the fact that we restrict our analysis to smooth solutions before the appearance of shocks.

We assume that the potential V satisfies the following

Assumption 3.2.1. *The potential V is a smooth, non-negative function such that the partition function $Z(\beta, \lambda) = \int_{-\infty}^{\infty} \exp(-\beta V(r) - \lambda r) dr$ is well defined for $\beta > 0$ and $\lambda \in \mathbb{R}$ and there exists a positive constant C such that*

$$0 < V''(r) \leq C, \quad (3.17)$$

and

$$\limsup_{|r| \rightarrow +\infty} \frac{rV'(r)}{V(r)} \in (0, +\infty), \quad (3.18)$$

$$\limsup_{|r| \rightarrow +\infty} \frac{[V'(r)]^2}{V(r)} < +\infty. \quad (3.19)$$

Provided we can prove that the infinite volume dynamics is macro-ergodic, then we can rigorously prove (even if $\gamma = 0$), using the relative entropy method of Yau ([56]), that (3.13) is indeed the hydrodynamic limit in the smooth regime, *i.e.* for times t up to the appearance of the first shock (see for example [36, 52]). Observe that the expected hydrodynamic limits do not depend on γ . We need to assume $\gamma > 0$ to ensure the macro-ergodicity of the dynamics.

Remark 3.2.2. *As argued in [52], it turns out that when there is more than one conservation law, the conservation of thermodynamic entropy (3.15) must hold for the hydrodynamic limit to be well defined. This relation is indeed fundamental for Yau's method where, in the expansion of the time derivative of relative entropy, the cancelation of the linear terms is a consequence of the preservation of the thermodynamic entropy.*

Averages with respect to the empirical energy-volume measure are defined, for continuous functions $G, H : \mathbb{T} \rightarrow \mathbb{R}$, as

$$\begin{pmatrix} \mathcal{E}_N(t, G) \\ \mathcal{V}_N(t, H) \end{pmatrix} = \begin{pmatrix} \frac{1}{N} \sum_{x \in \mathbb{T}_N} G\left(\frac{x}{N}\right) V(\eta_x(t)) \\ \frac{1}{N} \sum_{x \in \mathbb{T}_N} H\left(\frac{x}{N}\right) \eta_x(t) \end{pmatrix}.$$

We can then state the following result.

Theorem 3.2.3 ([13]). *Fix some $\gamma > 0$ and consider the dynamics on the torus \mathbb{T}_N generated by $\mathcal{L}_{N, \text{per}}$ where the potential V satisfies Assumption 3.2.1. Assume that the system is initially distributed according to a local Gibbs state (3.12) with smooth energy profile \mathbf{e}_0 and volume profile \mathbf{v}_0 . Consider a positive time t such that the solution (\mathbf{e}, \mathbf{v}) to (3.13) belongs to \mathcal{U} and is smooth on the time interval $[0, t]$.*

Then, for any continuous test functions $G, H : \mathbb{T} \rightarrow \mathbb{R}$, the following convergence in probability holds as $N \rightarrow +\infty$:

$$\left(\mathcal{E}_N(tN, G), \mathcal{V}_N(tN, H) \right) \longrightarrow \left(\int_{\mathbb{T}} G(q) \mathbf{e}(t, q) dq, \int_{\mathbb{T}} H(q) \mathbf{v}(t, q) dq \right).$$

The derivation of the hydrodynamic limits beyond the shocks for systems of $n \geq 2$ conservation laws is very difficult and is one of the most challenging problems in the field of hydrodynamic limits. The first difficulty is of course our poor understanding of the solutions to such systems. Recently, J. Fritz proposed in [27] to derive hydrodynamic limits for hyperbolic systems (in the case $n = 2$) by some extension of the compensated-compactness approach [50] to stochastic microscopic models. This program has been achieved in [30] (see also the recent paper [28]), where the authors derive the classical $n = 2$ Leroux system of conservation laws. In fact, to be exact, only the convergence to the set of entropy solutions is proved, the question of uniqueness being left open. It remains nonetheless the best result available at this time. The proof is based on a strict control of entropy pairs at the microscopic level by the use of logarithmic Sobolev inequality estimates. It would be very interesting to extend these methods to systems such as the ones considered here.

3.2.3 Anomalous diffusion

We investigate now the problem of anomalous diffusion of energy for these models.

If $V(r) = r^2$ then Theorem 3.1.1 is *mutatis mutandis* valid and we get the same conclusions: the time-space correlations for the current behave for large time t like $t^{-1/2}$. Thus the system is super-diffusive (see [13] for the details).

For generic anharmonic potentials, we can only provide numerical evidence of the super-diffusivity. However, it is difficult to estimate numerically the time autocorrelation functions of the currents because of their expected long-time tails, and because statistical errors are very large (in relative value) when t is large. Also, for finite systems (the only ones we can simulate on a computer), the autocorrelation is generically exponentially decreasing for anharmonic potentials, and, to obtain meaningful results, the thermodynamic limit should be taken before the long-time limit.

A more tenable approach consists in studying a nonequilibrium system in its steady-state. We consider a finite system of length $2N + 1$ in contact with two thermostats which fix the value of the energy at the boundaries. The generator of the dynamics is given by

$$\mathcal{L}_N = \mathcal{A}_N + \gamma \mathcal{S}_N + \lambda_\ell \mathcal{B}_{-N, T_\ell} + \lambda_r \mathcal{B}_{N, T_r}, \quad (3.20)$$

where \mathcal{A}_N and \mathcal{S}_N are defined by

$$\begin{aligned} (\mathcal{A}_N f)(\eta) &= \sum_{x=-(N-1)}^{N-1} \left(V'(\eta_{x+1}) - V'(\eta_{x-1}) \right) (\partial_{\eta_x} f)(\eta) \\ &\quad - V'(\eta_{N-1}) (\partial_{\eta_N} f)(\eta) + V'(\eta_{-N+1}) (\partial_{\eta_{-N}} f)(\eta), \\ (\mathcal{S}_N f)(\eta) &= \sum_{x=-N}^{N-1} [f(\eta^{x,x+1}) - f(\eta)], \end{aligned}$$

and $\mathcal{B}_{x,T} = T\partial_{\eta_x}^2 - V'(\eta_x)\partial_{\eta_x}$. The positive parameters λ_ℓ and λ_r are the intensities of the thermostats and T_ℓ, T_r the ‘‘temperatures’’ of the thermostats.

The generator $\mathcal{B}_{x,T}$ is a thermostating mechanism. In order to fix the energy at site $-N$ (resp. N) to the value e_ℓ (resp. e_r), we have to choose $\beta_\ell = T_\ell^{-1}$ (resp. $\beta_r = T_r^{-1}$) such that $e(\beta_\ell, 0) = e_\ell$ (resp. $e(\beta_r, 0) = e_r$). We denote by $\langle \cdot \rangle_{\text{ss}}$ the unique stationary state for the dynamics generated by \mathcal{L}_N .

The energy currents $j_{x,x+1}^{e,\gamma}$, which are such that $\mathcal{L}_{N,\text{open}}(V(\eta_x)) = -\nabla j_{x-1,x}^{e,\gamma}$ (for $x = -N, \dots, N+1$), are given by the first line of (3.9) for $x = -N+1, \dots, N-1$ while

$$\begin{aligned} j_{-N-1,-N}^{e,\gamma} &= \lambda_\ell [T_\ell V''(\eta_{-N}) - (V'(\eta_{-N}))^2], \\ j_{N,N+1}^{e,\gamma} &= -\lambda_r [T_r V''(\eta_N) - (V'(\eta_N))^2]. \end{aligned}$$

Since $\langle \mathcal{L}_{N,\text{open}}(V(\eta_x)) \rangle_{\text{ss}} = 0$, it follows that, for any $x = -N, \dots, N+1$, $\langle j_{x,x+1}^{e,\gamma} \rangle_{\text{ss}}$ is equal to a constant $J_N^\gamma(T_\ell, T_r)$ independent of x . In fact,

$$J_N^\gamma(T_\ell, T_r) = \langle \mathcal{J}_N^\gamma \rangle_{\text{ss}}, \quad \mathcal{J}_N^\gamma = \frac{1}{2N} \sum_{x=-N}^{N-1} j_{x,x+1}^{e,\gamma}. \quad (3.21)$$

The latter equation is interesting from a numerical viewpoint since it allows to perform some spatial averaging, hence reducing the statistical error of the results. We estimate the exponent $\delta \geq 0$ such that

$$\kappa(N) := NJ_N^\gamma \sim N^\delta \quad (3.22)$$

using numerical simulations. If $\delta = 0$, the system is a normal conductor of energy. If on the other hand $\delta > 0$, it is a superconductor.

The numerical simulations giving the value of δ are summarized in table 4.1. They have been performed for the harmonic chain $V(r) = r^2/2$, the quartic potential $V(r) = r^2/2 + r^4/4$ and the exponential potential $V(r) = e^{-r} + r - 1$. In Section 3.4 we will explain our interest in the exponential potential.

γ	$V(r) = r^2/2$	$V(r) = r^2/2 + r^4/4$	$V(r) = e^{-r} + r - 1$
0	1	0.13	1
0.01	–	0.14	0.12
0.1	0.50	0.27	0.25
1	0.50	0.43	0.33

Table 3.1: Conductivity exponents

Exponents in the harmonic case agree with their expected values. For nonlinear potentials, except for the singular value $\delta = 1$ when $\gamma = 0$ and $V(r) = e^{-r} + r - 1$, the exponents seem to be monotonically increasing with γ . A similar behavior of the exponents is observed for Toda chains [33] with a momentum conserving noise. As explained below this strange behavior is probably due to a numerical simulations defect.

Note also that the value found for $\gamma = 0$ with the anharmonic FPU potential $V(r) = r^2/2 + r^4/4$ is smaller than the corresponding value for standard oscillators chains, which is around 0.33 (see [42]). We performed also numerical simulations for a “rotor” model, $V(r) = 1 - \cos(r)$, and we found $\delta \approx 0.02$, i.e. a normal conductivity. A similar picture is observed for the usual rotor model which is composed of a chain of unpinned oscillators with interaction potential $V(r) = 1 - \cos(r)$. The normal behavior is conjectured to be due to the compactness of the state space ⁵.

3.3 Fractional superdiffusion for a harmonic chain with bulk noise

In this section we consider the energy-volume conserving model with quadratic potential. Fix $\lambda \in \mathbb{R}$ and $\beta > 0$, and consider the process $\{\eta(t); t \geq 0\}$ generated by (1.20) with $V(\eta) = \eta^2/2$ and with initial distribution $\mu_{\beta, \lambda}$. Notice that $\{\eta(t) + \rho; t \geq 0\}$ has the same distribution of the process with initial measure $\mu_{\beta, \rho + \lambda}$. Therefore, we can assume, without loss of generality, that $\lambda = 0$. We will write $\mu_\beta = \mu_{\beta, 0}$ to simplify notation. We will denote by \mathbb{P} the law of $\{\eta(t); t \geq 0\}$ and by \mathbb{E} the expectation with respect to \mathbb{P} . The *energy correlation function* $\{S_t(x); x \in \mathbb{Z}, t \geq 0\}$ is defined as

$$S_t(x) = \frac{\beta^2}{2} \mathbb{E} \left[\left(\eta_0(0)^2 - \frac{1}{\beta} \right) \left(\eta_x(t)^2 - \frac{1}{\beta} \right) \right] \quad (3.23)$$

⁵The variable r has to be interpreted as an angle and belongs to the torus $2\pi\mathbb{T}$.

for any $x \in \mathbb{Z}$ and any $t \geq 0$. The constant $\frac{\beta^2}{2}$ is just the inverse of the variance of $\eta_x^2 - \frac{1}{\beta}$ under μ_β . By translation invariance of the dynamics and the initial distribution μ_β , we see that

$$\frac{\beta^2}{2} \mathbb{E}\left[\left(\eta_x(0)^2 - \frac{1}{\beta}\right)\left(\eta_y(t)^2 - \frac{1}{\beta}\right)\right] = S_t(y - x) \quad (3.24)$$

for any $x, y \in \mathbb{Z}$.

Theorem 3.3.1 ([7]). *Let $f, g : \mathbb{R} \rightarrow \mathbb{R}$ be smooth functions of compact support. Then,*

$$\lim_{n \rightarrow \infty} \frac{1}{n} \sum_{x, y \in \mathbb{Z}} f\left(\frac{x}{n}\right) g\left(\frac{y}{n}\right) S_{tn^{3/2}}(x - y) = \iint f(x) g(y) P_t(x - y) dx dy, \quad (3.25)$$

where $\{P_t(x); x \in \mathbb{R}, t \geq 0\}$ is the fundamental solution of the fractional heat equation

$$\partial_t u = -\frac{1}{\sqrt{2}} \{(-\Delta)^{3/4} - \nabla(-\Delta)^{1/4}\} u. \quad (3.26)$$

A fundamental step in the proof of this theorem will be the analysis of the correlation function $\{S_t(x, y); x \neq y \in \mathbb{Z}, t \geq 0\}$ given by

$$S_t(x, y) = \frac{\beta^2}{2} \mathbb{E}\left[\left(\eta_x(0)^2 - \frac{1}{\beta}\right)\eta_x(t)\eta_y(t)\right] \quad (3.27)$$

for any $t \geq 0$ and any $x \neq y \in \mathbb{Z}$. Notice that this definition makes perfect sense for $x = y$ and, in fact, we have $S_t(x, x) = S_t(x)$. For notational convenience we define $S_t(x, x)$ as equal to $S_t(x)$. However, these quantities are of different nature, since $S_t(x)$ is related to *energy fluctuations* and $S_t(x, y)$ is related to *volume fluctuations* (for $x \neq y$).

Remark 3.3.2. *It is not difficult to see that with a bit of technical work our techniques actually show that the distributions valued process $\{\mathcal{E}_t^n(\cdot); t \geq 0\}$ defined for any test function f by*

$$\mathcal{E}_t^n(f) = \frac{1}{\sqrt{n}} \sum_{x \in \mathbb{Z}} f\left(\frac{x}{n}\right) \left\{ \eta_x(tn^{3/2})^2 - \frac{1}{\beta} \right\}$$

converges as n goes to infinity to an infinite dimensional 3/4-fractional Ornstein-Uhlenbeck process, i.e. the centered Gaussian process with covariance prescribed by the right hand side of (3.25).

Remark 3.3.3. *It is interesting to notice that P_t is the maximally asymmetric 3/2-Levy distribution. It has power law as $|x|^{-5/2}$ towards the diffusive peak and stretched exponential as $\exp[-|x|^3]$ towards the exterior of the sound cone ([53] Chapter 4). As noticed to us by H. Spohn, this reflects the expected physical property that no propagation beyond the sound cone occurs.*

Remark 3.3.4. *With a bit of technical work the proof of this theorem can be adapted to obtain a similar statement for a harmonic chain perturbed by the momentum exchanging noise (see [34] where such statement is proved for the Wigner function). In this case the skew fractional heat equation is replaced by the (symmetric) fractional heat equation.*

Proof. Denote by $\mathcal{C}_c^\infty(\mathbb{R}^d)$ the space of infinitely differentiable functions $f : \mathbb{R}^d \rightarrow \mathbb{R}$ of compact support. Then, $\|f\|_{2,n}$ denotes the weighted $\ell^2(\mathbb{Z}^d)$ -norm

$$\|f\|_{2,n} = \sqrt{\frac{1}{n^d} \sum_{x \in \mathbb{Z}^d} f\left(\frac{x}{n}\right)^2}. \quad (3.28)$$

Let $g \in \mathcal{C}_c^\infty(\mathbb{R})$ be a fixed function. For each $n \in \mathbb{N}$ we define the field $\{\mathcal{S}_t^n; t \geq 0\}$ as

$$\mathcal{S}_t^n(f) = \frac{1}{n} \sum_{x,y \in \mathbb{Z}} g\left(\frac{x}{n}\right) f\left(\frac{y}{n}\right) S_{tn^{3/2}}(y-x) \quad (3.29)$$

for any $t \geq 0$ and any $f \in \mathcal{C}_c^\infty(\mathbb{R})$. By the Cauchy-Schwarz inequality we have the *a priori* bound

$$|\mathcal{S}_t^n(f)| \leq \|g\|_{2,n} \|f\|_{2,n} \quad (3.30)$$

for any $t \geq 0$, any $n \in \mathbb{N}$ and any $f, g \in \mathcal{C}_c^\infty(\mathbb{R})$. For a function $h \in \mathcal{C}_c^\infty(\mathbb{R}^2)$ we define $\{Q_t^n(h); t \geq 0\}$ as

$$Q_t^n(h) = \frac{1}{n^{3/2}} \sum_{x,y,z \in \mathbb{Z}} g\left(\frac{x}{n}\right) h\left(\frac{y}{n}, \frac{z}{n}\right) S_{tn^{3/2}}(y-x, z-x). \quad (3.31)$$

Notice as well that $Q_t^n(h)$ depends only on the symmetric part of the function h . Therefore, we will always assume, without loss of generality, that $h(x, y) = h(y, x)$ for any $x, y \in \mathbb{Z}$. We point out that $Q_t^n(h)$ does not depend on the values of h at the diagonal $\{x = y\}$. We have the *a priori* bound

$$|Q_t^n(h)| \leq 2\|g\|_{2,n} \|\tilde{h}\|_{2,n}, \quad (3.32)$$

where \tilde{h} is defined by $\tilde{h}\left(\frac{x}{n}, \frac{y}{n}\right) = h\left(\frac{x}{n}, \frac{y}{n}\right) \mathbf{1}_{x \neq y}$.

For a function $f \in \mathcal{C}_c^\infty(\mathbb{R})$, we define $\Delta_n f : \mathbb{R} \rightarrow \mathbb{R}$ as

$$\Delta_n f\left(\frac{x}{n}\right) = n^2 \left(f\left(\frac{x+1}{n}\right) + f\left(\frac{x-1}{n}\right) - 2f\left(\frac{x}{n}\right) \right). \quad (3.33)$$

In other words, $\Delta_n f$ is a discrete approximation of the second derivative of f . We also define $\nabla_n f \otimes \delta : \frac{1}{n}\mathbb{Z}^2 \rightarrow \mathbb{R}$ as

$$(\nabla_n f \otimes \delta)\left(\frac{x}{n}, \frac{y}{n}\right) = \begin{cases} \frac{n^2}{2} \left(f\left(\frac{x+1}{n}\right) - f\left(\frac{x}{n}\right) \right); & y = x+1 \\ \frac{n^2}{2} \left(f\left(\frac{x}{n}\right) - f\left(\frac{x-1}{n}\right) \right); & y = x-1 \\ 0; & \text{otherwise.} \end{cases} \quad (3.34)$$

Less evident than the interpretation of $\Delta_n f$, $\nabla_n f \otimes \delta$ turns out to be a discrete approximation of the distribution $f'(x) \otimes \delta(x = y)$, where $\delta(x = y)$ is the δ of Dirac at the line $x = y$. We have that

$$\frac{d}{dt} \mathcal{S}_t^n(f) = -2Q_t^n(\nabla_n f \otimes \delta) + \mathcal{S}_t^n\left(\frac{1}{\sqrt{n}} \Delta_n f\right). \quad (3.35)$$

In this equation we interpret the term $Q_t^n(\nabla_n f \otimes \delta)$ in the obvious way. By the *a priori* bound (3.30), the term $\mathcal{S}_t^n\left(\frac{1}{\sqrt{n}} \Delta_n f\right)$ is negligible, as $n \rightarrow \infty$. We describe now the equation satisfied by $Q_t^n(h)$. For this we need some extra definitions. For $h \in \mathcal{C}_c^\infty(\mathbb{R}^2)$ we define $\Delta_n h : \mathbb{R}^2 \rightarrow \mathbb{R}$ as

$$\Delta_n h\left(\frac{x}{n}, \frac{y}{n}\right) = n^2 \left(h\left(\frac{x+1}{n}, \frac{y}{n}\right) + h\left(\frac{x-1}{n}, \frac{y}{n}\right) + h\left(\frac{x}{n}, \frac{y+1}{n}\right) + h\left(\frac{x}{n}, \frac{y-1}{n}\right) - 4h\left(\frac{x}{n}, \frac{y}{n}\right) \right). \quad (3.36)$$

In other words, $\Delta_n h$ is a discrete approximation of the Laplacian of h . We also define $\mathcal{A}_n h : \mathbb{R} \rightarrow \mathbb{R}$ as

$$\mathcal{A}_n h\left(\frac{x}{n}, \frac{y}{n}\right) = n \left(h\left(\frac{x}{n}, \frac{y-1}{n}\right) + h\left(\frac{x-1}{n}, \frac{y}{n}\right) - h\left(\frac{x}{n}, \frac{y+1}{n}\right) - h\left(\frac{x+1}{n}, \frac{y}{n}\right) \right). \quad (3.37)$$

The function $\mathcal{A}_n h$ is a discrete approximation of the directional derivative $(-2, -2) \cdot \nabla h$. Let us define $\mathcal{D}_n h : \frac{1}{n}\mathbb{Z} \rightarrow \mathbb{R}$ as

$$\mathcal{D}_n h\left(\frac{x}{n}\right) = n \left(h\left(\frac{x}{n}, \frac{x+1}{n}\right) - h\left(\frac{x-1}{n}, \frac{x}{n}\right) \right) \quad (3.38)$$

and $\tilde{\mathcal{D}}_n h : \frac{1}{n}\mathbb{Z}^2 \rightarrow \mathbb{R}$ as

$$\tilde{\mathcal{D}}_n h\left(\frac{x}{n}, \frac{y}{n}\right) = \begin{cases} n^2 \left(h\left(\frac{x}{n}, \frac{x+1}{n}\right) - h\left(\frac{x}{n}, \frac{x}{n}\right) \right); & y = x + 1 \\ n^2 \left(h\left(\frac{x-1}{n}, \frac{x}{n}\right) - h\left(\frac{x-1}{n}, \frac{x-1}{n}\right) \right); & y = x - 1 \\ 0; & \text{otherwise.} \end{cases} \quad (3.39)$$

The function $\mathcal{D}_n h$ is a discrete approximation of the directional derivative of h along the diagonal $x = y$, while $\tilde{\mathcal{D}}_n h$ is a discrete approximation of the distribution $\partial_y h(x, x) \otimes \delta(x = y)$. Finally we can write down the equation satisfied by the field $Q_t^n(h)$:

$$\frac{d}{dt} Q_t^n(h) = Q_t^n(n^{-1/2} \Delta_n h + n^{1/2} \mathcal{A}_n h) - 2\mathcal{S}_t^n(\mathcal{D}_n h) + 2Q_t^n(n^{-1/2} \tilde{\mathcal{D}}_n h). \quad (3.40)$$

Given $f \in \mathcal{C}_c^\infty(\mathbb{R})$, if we choose $h := h_n(f)$ such that

$$n^{-1/2} \Delta_n h + n^{1/2} \mathcal{A}_n h = 2\nabla_n f \otimes \delta$$

then summing (3.35) and (3.40) we get

$$\frac{d}{dt} \mathcal{S}_t^n(f) = -\frac{d}{dt} Q_t^n(h) + \mathcal{S}_t^n\left(\frac{1}{\sqrt{n}} \Delta_n f\right) - 2\mathcal{S}_t^n(\mathcal{D}_n h) + 2Q_t^n(n^{-1/2} \tilde{\mathcal{D}}_n h).$$

We integrate in time the previous expression. By the a priori bounds, the term $\mathcal{S}_t^n(\frac{1}{\sqrt{n}}\Delta_n f)$ is small as well as $Q_t^n(h) - Q_0^n(h)$. The term $Q_t^n(n^{-1/2}\tilde{\mathcal{D}}_n h)$ is quite singular since it involves an approximation of a distribution but it turns out to be negligible but this not follow directly from the a priori bounds (see [7]). By using Fourier transform one can see that $-2\mathcal{D}_n h$ converges to $-\frac{1}{\sqrt{2}}\{(-\Delta)^{3/4} - \nabla(-\Delta)^{1/4}\}f$ and we are done. \square

3.4 Anomalous diffusion for a perturbed Hamiltonian system with exponential interactions

We investigate here in more details the exponential case $V_{exp}(r) = e^{-r} - 1 + r$. The deterministic system with generator (3.11) and with the exponential potential above is well known in the integrable systems literature ⁶. It has been introduced in [35] by Kac and van Moerbeke and was shown to be completely integrable. Consequently, using Mazur's inequality, it is easy to show that the energy transport is ballistic ([13]).

As we will see this is different when the noise is added: the energy transport is no more ballistic but superdiffusive. Thus the situation is similar to the harmonic case. Nevertheless we expect that the time autocorrelation of the current decay like $t^{-2/3}$. We are not able to show this but we proved in [6] lower bounds sufficient to imply superdiffusivity.

The results are stated in infinite volume: we consider the stochastic energy-volume conserving dynamics $\{\eta(t)\}_{t \geq 0}$ with potential $V := V_{exp}$. Its generator is given by $\mathcal{L} = \mathcal{A} + \gamma\mathcal{S}$ where \mathcal{A} and \mathcal{S} are defined by (1.20). Since the exponential potential grows very fast as $r \rightarrow -\infty$, some care has to be taken to show that the infinite dynamics is well defined (see [6]). We recall that grand canonical Gibbs measures are denoted by $\mu_{\beta,\lambda}$ and take the form

$$d\mu_{\beta,\lambda}(\eta) = \prod_{x \in \mathbb{Z}} \frac{e^{-\beta V(\eta_x) - \lambda \eta_x}}{Z(\beta, \lambda)} d\eta_x, \quad \beta > 0, \lambda + \beta < 0.$$

In this section, β and λ are fixed and we denote by e (resp. v) the average energy (resp. volume) w.r.t. $\mu_{\beta,\lambda}$ (see (3.7)).

⁶It seems that despite the Hamiltonian structure of the Kac-van-Moerbeke system was known, the interpretation of the latter as a chain of oscillators with exponential kinetic energy and exponential interaction was not observed.

The microscopic energy current $j_{x,x+1}^{e,\gamma}$ and volume current $j_{x,x+1}^{v,\gamma}$ are given by

$$j_{x,x+1}^{e,\gamma}(\eta) = -e^{-(\eta_x + \eta_{x+1})} + (e^{-\eta_x} + e^{-\eta_{x+1}}) - \gamma \nabla(V(\eta_x))$$

and

$$j_{x,x+1}^{v,\gamma}(\eta) = e^{-\eta_x} + e^{-\eta_{x+1}} - \gamma \nabla \eta_x.$$

We will use the compact notations

$$\omega_x = \begin{pmatrix} V(\eta_x) \\ \eta_x \end{pmatrix}, \quad J_{x,x+1} = \begin{pmatrix} j_{x,x+1}^{e,\gamma} \\ j_{x,x+1}^{v,\gamma} \end{pmatrix}.$$

In the hyperbolic scaling, the hydrodynamical equations for the energy profile \mathbf{e} and the volume profile \mathbf{v} take the form

$$\begin{cases} \partial_t \mathbf{e} - \partial_q((\mathbf{e} - \mathbf{v})^2) = 0 \\ \partial_t \mathbf{v} + 2 \partial_q(\mathbf{e} - \mathbf{v}) = 0. \end{cases} \quad (3.41)$$

They can be written in the compact form $\partial_t X + \partial_q \mathfrak{J}(X) = 0$ with

$$X = \begin{pmatrix} \mathbf{e} \\ \mathbf{v} \end{pmatrix}, \quad \mathfrak{J}(X) = \begin{pmatrix} -(\mathbf{e} - \mathbf{v})^2 \\ 2(\mathbf{e} - \mathbf{v}) \end{pmatrix}. \quad (3.42)$$

The differential matrix of \mathfrak{J} is given by

$$(\nabla \mathfrak{J})(X) = 2 \begin{pmatrix} -(\mathbf{e} - \mathbf{v}) & \mathbf{e} - \mathbf{v} \\ 1 & -1 \end{pmatrix}.$$

For given (e, v) we denote by $(T_t^+)_{t \geq 0}$ (resp. $(T_t^-)_{t \geq 0}$) the semigroup on $S(\mathbb{R}) \times S(\mathbb{R})$ generated by the linearized system

$$\partial_t \varepsilon + M^T \partial_q \varepsilon = 0, \quad (\text{resp. } \partial_t \varepsilon - M^T \partial_q \varepsilon = 0), \quad (3.43)$$

where

$$M := M(e, v) = [\nabla \mathfrak{J}](\omega), \quad \omega = \begin{pmatrix} e \\ v \end{pmatrix}.$$

We omit the dependence of these semigroups on (e, v) for lightness of the notations. Above $S(\mathbb{R})$ denotes the Schwartz space of smooth rapidly decreasing functions.

The first result of [6] gives a lower bound on the time-scale for which a non-trivial evolution of the energy-volume fluctuation field can be observed.

We take the infinite system at equilibrium under the Gibbs measure $\mu_{\bar{e}, \bar{v}}$ corresponding to a mean energy \bar{e} and a mean volume \bar{v} . Our goal is to study the energy-volume fluctuation field in the time-scale $tn^{1+\alpha}$, $\alpha \geq 0$:

$$\mathcal{Y}_t^{n,\alpha}(\mathbf{G}) = \frac{1}{\sqrt{n}} \sum_{x \in \mathbb{Z}} \mathbf{G}(x/n) \cdot (\omega_x(tn^{1+\alpha}) - \omega), \quad (3.44)$$

where for $q \in \mathbb{R}$, $x \in \mathbb{Z}$,

$$\mathbf{G}(q) = \begin{pmatrix} G_1(q) \\ G_2(q) \end{pmatrix}, \quad \omega_x = \begin{pmatrix} V(\eta_x) \\ \eta_x \end{pmatrix}$$

and G_1, G_2 are test functions belonging to $S(\mathbb{R})$.

We need to introduce several notations. For each integer $z \geq 0$, let $H_z(x) = (-1)^z e^{x^2} \frac{d^z}{dx^z} e^{-x^2}$ be the Hermite polynomial and $h_z(x) = (z! \sqrt{2\pi})^{-1} H_z(x) e^{-x^2}$ the Hermite function. The set $\{h_z, z \geq 0\}$ is an orthonormal basis of $\mathbb{L}^2(\mathbb{R})$. Consider in $\mathbb{L}^2(\mathbb{R})$ the operator $K_0 = x^2 - \Delta$, Δ being the Laplacian on \mathbb{R} . For an integer $k \geq 0$, denote by \mathbb{H}_k the Hilbert space induced by $S(\mathbb{R})$ and the scalar product $\langle \cdot, \cdot \rangle_k$ defined by $\langle f, g \rangle_k = \langle f, K_0^k g \rangle_0$, where $\langle \cdot, \cdot \rangle_0$ denotes the inner product of $\mathbb{L}^2(\mathbb{R})$ and denote by \mathbb{H}_{-k} the dual of \mathbb{H}_k , relatively to this inner product. Let $\langle \cdot \rangle$ represent the average with respect to the Lebesgue measure.

If E is a Polish space then $D(\mathbb{R}^+, E)$ (resp. $C(\mathbb{R}^+, E)$) denotes the space of E -valued functions, right continuous with left limits (resp. continuous), endowed with the Skorohod (resp. uniform) topology. Let $Q^{n,\alpha}$ be the probability measure on $D(\mathbb{R}^+, \mathbb{H}_{-k} \times \mathbb{H}_{-k})$ induced by the fluctuation field $\mathcal{Y}_t^{n,\alpha}$ and $\mu_{\beta,\lambda}$. Let $\mathbb{P}_{\mu_{\beta,\lambda}}$ denote the probability measure on $D(\mathbb{R}^+, \mathbb{R}^{\mathbb{Z}})$ induced by $(\eta(t))_{t \geq 0}$ and $\mu_{\beta,\lambda}$. Let $\mathbb{E}_{\mu_{\beta,\lambda}}$ denote the expectation with respect to $\mathbb{P}_{\mu_{\beta,\lambda}}$.

Theorem 3.4.1 ([6]). *Fix an integer $k > 2$. Denote by Q the probability measure on $C(\mathbb{R}^+, \mathbb{H}_{-k} \times \mathbb{H}_{-k})$ corresponding to a stationary Gaussian process with mean 0 and covariance given by*

$$\mathbb{E}_Q [\mathcal{Y}_t(\mathbf{H}) \mathcal{Y}_s(\mathbf{G})] = \langle T_t^- \mathbf{H} \cdot \chi T_s^- \mathbf{G} \rangle$$

for every $0 \leq s \leq t$ and \mathbf{H}, \mathbf{G} in $\mathbb{H}_k \times \mathbb{H}_k$. Here $\chi := \chi(\beta, \lambda)$ is the equilibrium covariance matrix of ω_0 . Then, the sequence $(Q^{n,0})_{n \geq 1}$ converges weakly to the probability measure Q .

The theorem above means that in the hyperbolic scaling the fluctuations are trivial: the initial fluctuations are transported by the linearized system of (3.41). To see a nontrivial behavior we have to study, in the transport frame, the fluctuations at a longer time scale $tn^{1+\alpha}$, with $\alpha > 0$. Thus, we consider the fluctuation field $\widehat{\mathcal{Y}}^{n,\alpha}$, $\alpha > 0$, defined, for any $\mathbf{G} \in S(\mathbb{R}) \times S(\mathbb{R})$, by

$$\widehat{\mathcal{Y}}_t^{n,\alpha}(\mathbf{G}) = \mathcal{Y}_t^{n,\alpha}(T_{tn^\alpha}^+ \mathbf{G}). \quad (3.45)$$

Our second main theorem shows that the correct scaling exponent α is greater than $1/3$:

Theorem 3.4.2 ([6]). *Fix an integer $k > 1$ and $\alpha < 1/3$. Denote by Q the probability measure on $C(\mathbb{R}^+, \mathbb{H}_{-k} \times \mathbb{H}_{-k})$ corresponding to a stationary Gaussian process with mean 0 and covariance given by*

$$\mathbb{E}_Q [\mathcal{Y}_t(\mathbf{H}) \mathcal{Y}_s(\mathbf{G})] = \langle \mathbf{H} \cdot \chi \mathbf{G} \rangle$$

for every $0 \leq s \leq t$ and \mathbf{H}, \mathbf{G} in $\mathbb{H}_k \times \mathbb{H}_k$. Then, the sequence $(Q^{n,\alpha})_{n \geq 1}$ converges weakly to the probability measure Q .

The proofs of these theorems can be reduced to the proof of a so-called *equilibrium Boltzmann-Gibbs principle*. Let us explain what it means. Observables can be divided into two classes: non-hydrodynamical and hydrodynamical. The first ones are the non conserved quantities and they fluctuate on a much faster scale than the conserved ones. Hence, they should average out and only their projection on the hydrodynamical variables should persist in the scaling limit. For any local function $g := g(\eta)$, the projection $\mathcal{P}_{e,v} g$ of g on the fields of the conserved quantities is defined by

$$(\mathcal{P}_{e,v} g)(\eta) = g(\eta) - \tilde{g}(e, v) - (\nabla \tilde{g})(e, v) \cdot (\omega_0 - \omega)$$

where $\tilde{g}(e, v) = \langle g \rangle_{\mu_{\beta,\lambda}}$ and $\nabla \tilde{g}$ is the gradient of the function \tilde{g} . As explained above we expect that in the Euler time scale, for any test function $\mathbf{H} \in S(\mathbb{R}) \times S(\mathbb{R})$, the space-time variance

$$\lim_{n \rightarrow \infty} \mathbb{E}_{\mu_{\beta,\lambda}} \left[\left(\int_0^t \frac{1}{\sqrt{n}} \sum_{x \in \mathbb{Z}} \mathbf{H}(x/n) \cdot [\theta_x \mathcal{P}_{e,v} g(\eta(sn))] ds \right)^2 \right] = 0 \quad (3.46)$$

vanishes as n goes to infinity. In fact we have only to show (3.46) for the function $g = J_{0,1}$ ⁷. Thus let us first define the *normalized* currents by

$$\hat{J}_{x,x+1} = [\theta_x \mathcal{P}_{\bar{e},\bar{v}} J_{0,1}] = \begin{pmatrix} j_{x,x+1}^{e,\gamma}(\eta) \\ j_{x,x+1}^{v,\gamma}(\eta) \end{pmatrix} - \mathfrak{J}(\omega) - (\nabla \mathfrak{J})(\omega) \begin{pmatrix} V_b(\eta_x) - e \\ \eta_x - v \end{pmatrix}. \quad (3.47)$$

To estimate the space-time variance involved we use the following inequality (see [37]):

$$\mathbb{E}_{\mu_{\beta,\lambda}} \left[\left(\int_0^t f(\eta(sn^{1+\alpha})) \right)^2 \right] \leq \frac{Ct}{n^{1+\alpha}} \left\langle f, \left(\frac{1}{tn^{1+\alpha}} - \gamma \mathcal{S} \right)^{-1} f \right\rangle_{\mu_{\beta,\lambda}} \quad (3.48)$$

⁷For Theorem 3.4.2, the Boltzmann-Gibbs principle has to be proved in the longer time scale $tn^{1+\alpha}$ and in the transport frame.

where $f = \frac{1}{\sqrt{n}} \sum_{x \in \mathbb{Z}} \mathbf{H}(x/n) \cdot \hat{J}_{x,x+1}$. Due to the very simple form of the operator \mathcal{S} the RHS of (3.48) can be estimated and shown to vanish as n goes to infinity. Nevertheless it has to be done with some care since \mathcal{S} is very degenerate so that without the term $\frac{1}{tn^{1+\alpha}}$ the RHS of (3.48) blows up.

Theorem 3.4.2 does not exclude the possibility of normal fluctuations, i.e. the convergence in law of the fluctuation field of the two conserved quantities to an infinite dimensional Ornstein-Uhlenbeck process in the diffusive time scale ($\alpha = 1$). To see that it is not the case we will show that the diffusion coefficient $\mathcal{D} := \mathcal{D}(e, v)$ appearing in this hypothetical limiting process would be infinite, excluding thus this possibility. Up to a constant matrix coming from a martingale term (due to the noise) and thus irrelevant for us (see [2], [13]), the matrix coefficient \mathcal{D} is defined by the Green-Kubo formula

$$\mathcal{D} = \int_0^\infty \mathbb{E}_{\mu_{\beta,\lambda}} \left[\sum_{x \in \mathbb{Z}} \hat{J}_{x,x+1}(t) \left[\hat{J}_{0,1}(0) \right]^T \right] dt. \quad (3.49)$$

The signature of the superdiffusive behavior of the system is seen in the divergence of \mathcal{D} , i.e. in a slow decay of the current-current correlation function. To study the latter we introduce its Laplace transform

$$\mathcal{F}(\gamma, z) = \int_0^\infty e^{-zt} \mathbb{E}_{\mu_{\beta,\lambda}} \left[\sum_{x \in \mathbb{Z}} \hat{J}_{x,x+1}(t) \left[\hat{J}_{0,1}(0) \right]^T \right] dt$$

which is well defined for any $z > 0$. This can be rewritten as

$$\mathcal{F}(\gamma, z) = \ll \hat{J}_{0,1}, (z - \mathcal{L})^{-1} \hat{J}_{0,1} \gg_{\beta,\lambda}$$

where $\ll \cdot, \cdot \gg_{\beta,\lambda}$ is the scalar product defined with respect to $\mu_{\beta,\lambda}$ in the same way as in (2.13).

Our third theorem is the following lower bound on $\mathcal{F}(\gamma, z)$. Observe that $\mathcal{F}(\gamma, z)$ is a square matrix of size 2 whose entry (i, j) is denoted by $\mathcal{F}_{i,j}$.

Theorem 3.4.3 ([6]). *Fix $\gamma > 0$. There exists a positive constant $c := c(\gamma) > 0$ such that*

$$\mathcal{F}_{1,1}(\gamma, z) \geq cz^{-1/4}$$

and

$$\mathcal{F}_{i,j}(\gamma, z) = 0, \quad (i, j) \neq (1, 1).$$

Moreover, there exists a positive constant $C := C(\gamma)$ such that for any $z > 0$,

$$C^{-1} \mathcal{F}_{1,1}(1, z/\gamma) \leq \mathcal{F}_{1,1}(\gamma, z) \leq C \mathcal{F}_{1,1}(1, z/\gamma). \quad (3.50)$$

The last part of the theorem follows easily by a scaling argument and is in fact also valid for general potentials V and for generic “standard” anharmonic chains of oscillators. In [3, 33, 13], numerical simulations indicate a strange dependence w.r.t. the noise intensity $\gamma > 0$ of the exponent δ in the energy transport coefficient $\kappa(N) \sim N^\delta$ (N is the system size, see (3.22) for the definition of $\kappa(N)$): $\delta := \delta(\gamma) > 0$ is increasing with the noise intensity γ . This is very surprising since the more stochasticity in the model is introduced, the less the system is diffusive! The inequality (3.50) shows that the time decay of the current autocorrelation function is independent of γ (up to possible slowly varying functions corrections, i.e. in a Tauberian sense). Therefore the numerical simulations do not reflect the correct behavior of the system ⁸

The proof of the first part of Theorem 3.4.3 is based on the three following arguments.

- The first idea consists to perform the microscopic change of variables $\xi_x = e^{-\eta x}$, $x \in \mathbb{Z}$, that defines a new Markov process $\{\xi(t)\}_{t \geq 0} = \{\xi_x(t); x \in \mathbb{Z}\}_{t \geq 0}$ with state space $(0, +\infty)^{\mathbb{Z}}$ and conserving $\sum_x \xi_x$ and $\sum_x \log \xi_x$. Its generator is given by $\tilde{\mathcal{L}} = \tilde{\mathcal{A}} + \gamma \tilde{\mathcal{S}}$ where for any local differentiable function f ,

$$\begin{aligned} (\tilde{\mathcal{A}}f)(\xi) &= \sum_{x \in \mathbb{Z}} \xi_x (\xi_{x+1} - \xi_{x-1}) (\partial_{\xi_x} f)(\xi), \\ (\tilde{\mathcal{S}}f)(\xi) &= \sum_{x \in \mathbb{Z}} [f(\xi^{x, x+1}) - f(\xi)]. \end{aligned}$$

The invariant measures for $(\xi(t))_{t \geq 0}$ are obtained from the Gibbs measures $\mu_{\beta, \lambda}$ by the change of variables above. They form a family $\{\nu_{\rho, \theta}\}_{\rho, \theta}$ of translation invariant product measures indexed by two parameters ρ and θ which satisfy

$$\rho = \nu_{\rho, \theta}(\xi_x), \quad \theta = \nu_{\rho, \theta}(\log \xi_x).$$

In fact the marginal of $\nu_{\rho, \theta}$ is a Gamma distribution. The parameters (ρ, θ) are in a one to one explicit correspondence with the parameters (e, v) .

Rewriting $\hat{J}_{x, x+1}$ with these new variables we see that it is sufficient to prove a similar statement for the process $(\xi(t))_{t \geq 0}$ under the equilibrium probability measure $\nu_{\rho, \theta}$. Introducing the inner product $\ll \cdot, \cdot \gg$ defined, for any local functions f, g on $(0, +\infty)^{\mathbb{Z}}$ by

$$\ll f, g \gg = \sum_{x \in \mathbb{Z}} \{\nu_{\rho, \theta}(f \theta_x g) - \nu_{\rho, \theta}(f) \nu_{\rho, \theta}(g)\}$$

⁸It would be very interesting to understand why the numerical simulations are so sensitive to the noise. Of course we could imagine that the exponent δ appearing in the transport coefficient defined by the non-equilibrium stationary state $\kappa(N)$ depends on the intensity of the noise and that it is not the case for the rate decay of the autocorrelation function but this is very unlikely.

we can show that the proof of the first claim of Theorem 3.4.3 is reduced to show there exists a positive constant c such that for any $z > 0$,

$$\ll W_{0,1}, (z - \tilde{\mathcal{L}})^{-1} W_{0,1} \gg \geq cz^{1/4} \quad (3.51)$$

where $W_{0,1}(\xi) = (\xi_0 - \rho)(\xi_1 - \rho)$.

- The second step consists in using a variational formula to express the LHS of (3.51). Indeed we have

$$\begin{aligned} \ll W_{0,1}, (z - \tilde{\mathcal{L}})^{-1} W_{0,1} \gg &= \sup_g \left\{ 2 \ll W_{0,1}, g \gg - \ll g, (z - \gamma \tilde{\mathcal{S}})g \gg \right. \\ &\quad \left. - \ll \tilde{\mathcal{A}}g, (z - \gamma \tilde{\mathcal{S}})^{-1} \tilde{\mathcal{A}}g \gg \right\} \end{aligned}$$

where the supremum is taken over local compactly supported smooth functions g . To get a lower bound it is sufficient to find a function g for which one can show that

$$2 \ll W_{0,1}, g \gg - \ll g, (z - \gamma \tilde{\mathcal{S}})g \gg - \ll \tilde{\mathcal{A}}g, (z - \gamma \tilde{\mathcal{S}})^{-1} \tilde{\mathcal{A}}g \gg \geq cz^{-1/4}.$$

- Let \mathbb{H} be the Hilbert space obtained by completion of the set of local functions w.r.t. the inner product $\ll \cdot, \cdot \gg$. Since $\nu_{\rho, \theta}$ is a product of Gamma distributions, the set of multivariate Laguerre polynomials form an orthogonal basis of \mathbb{H} . It is then possible to decompose \mathbb{H} as an orthogonal sum $\bigoplus_{n \in \mathbb{N}} \mathbb{H}_n$ of subspaces \mathbb{H}_n such that

$$\tilde{\mathcal{S}} : \mathbb{H}_n \rightarrow \mathbb{H}_n, \quad \tilde{\mathcal{A}} : \mathbb{H}_n \rightarrow \mathbb{H}_{n-1} \oplus \mathbb{H}_n \oplus \mathbb{H}_{n+1}.$$

The function $W_{0,1}$ belongs to \mathbb{H}_2 . Then we restrict the variational formula to functions $g \in \mathbb{H}_2$ and we estimate the corresponding new variational problem which is still infinite dimensional but involves only functions belonging to $\mathbb{H}_1 \oplus \mathbb{H}_2 \oplus \mathbb{H}_3$. To solve this variational problem we adapt ideas developed first in the context of Asymmetric Simple Exclusion Process ([4], [38]) and exploited later for other models. One of the difficulties comes again from the fact that the noise is degenerate.

The extension of Theorem 3.4.3 to others interacting potentials is a challenging problem. The general strategy presented here could be carried but the orthogonal basis (formed by Laguerre polynomials in the exponential case) is no longer explicit and only defined by some recurrence relations.

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