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# The topological gradient method for semi-linear problems and application to edge detection and noise removal. 

Audric Drogoul and Gilles Aubert

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Audric Drogoul<br>Univ. Nice Sophia Antipolis, CNRS, LJAD, UMR 7351, 06100 Nice, France<br>drogoula@unice.fr<br>Gilles Aubert<br>Univ. Nice Sophia Antipolis, CNRS, LJAD, UMR 7351, 06100 Nice, France<br>gaubert@unice.fr


#### Abstract

Résumé The goal of this paper is to apply the topological gradient method for segmenting/restoring images degraded by various noises and blurs. First applied by Hintermuller [Control and Cybernetics, 34 (2005), pp. 305-324] and Belaid etal [C. R. Acad. Sci. Paris, 342 (2006), pp.313-318] to restore images degraded by a Gaussian noise, we propose here to extend the method to blurred images contaminated either by an additive Gaussian noise, or a multiplicative noise of gamma law and to blurred Poissonian images. We compute, both for perforated and cracked domains, the topological gradient for each noise model. Then we present a segmentation/restoration algorithm based on this notion and we apply it to the three degradation models previously described. We compare our method with the Ambrosio-Tortorelli approximation of the Mumford-Shah functional and with those given by a classical TV restoration process. Many experimental results showing the efficiency, the robustness and the rapidity of the approach are presented.


## 1 Introduction

An important problem in image analysis is the reconstruction of an original image $u$ from an observed image f. In general this includes restoration and segmentation processes. The transformation between $f$ and $u$ originates from two phenomena. The first phenomenon is related to the acquisition process (blur created by a wrong lens adjustement or by a movement, Poissonian photons emission rates ...) and the second is due to the signal transmission. A lot of methods to reconstruct such degraded images exist : stochastic methods [22, 13], wavelets decomposition [31, 18], morphological methods [36]. Here we are interested with variational approaches [7]. In this context, the most famous model is the Mumford-Shah functional [33] (1989) but other works based on variational methods do exist ([7]). Among more recent papers, we can cite [8] (2008) for restoration of images contaminated by speckle noise, [16] (2014) for blind restoration of Poissonian images, and [35] (2013) for an overview of image restoration degraded by different type of noise.

In this paper we tackle the segmentation problem by using the topological gradient method. First introduced for cracks detection by Sokolowski and Zochowski [37], Masmoudi [32] and applied in
optimal design and mechanics ([3], [2], [4]), this notion consists in the study of the variations of a cost function $j(\Omega)=J_{\Omega}\left(u_{\Omega}\right)$ with respect to a topological variation, where $J_{\Omega}(u)$ is of the form $J_{\Omega}(u)=\int_{\Omega} F\left(u, \nabla u, \nabla^{2} u, \ldots\right)$ and $u_{\Omega}$ is a solution of a PDE defined on the image domain $\Omega$. In order to compute the topological gradient, we remove to $\Omega$ a small object $\omega_{\varepsilon}$ of size $\varepsilon \rightarrow 0$ centered at $x_{0} \in \Omega$ (generally a ball or a curve) and we set $\Omega_{\varepsilon}=\Omega \backslash \overline{\omega_{\varepsilon}}$. Two typical examples are : for small $\varepsilon>0$ (a) $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon B\right\}}$ and (b) $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon \sigma(n)\right\}}$, where $B=B(O, 1)$ is the unit ball of $\mathbb{R}^{2}$ and $\sigma(n)$ is a straight segment with normal $n$ (a crack). We compute the limit : $\mathscr{I}\left(x_{0}\right)=\lim _{\varepsilon \rightarrow 0} \frac{j\left(\Omega_{\varepsilon}\right)-j(\Omega)}{\varepsilon^{d}}$ where $d$ is the dimension of the ambiant space. $\mathscr{I}\left(x_{0}\right)$ is called the topological gradient at $x_{0}$. It measures the energy contained by a perturbation centered at $x_{0}$ and so the structures that we want to detect correspond to the points $x_{0}$ where $\mathscr{I}\left(x_{0}\right)$ is the largest. The type of structure to be detected depends on the choice of the cost function $J_{\Omega}(u)$. Recently this notion has been used in image processing and to the best of our knowledge the first works in this direction are those by [25], [30], [14]. Then other imaging problems such as inpainting, classification, demosaicing, super resolution, have been adressed using a topological gradient approach ([9, 15], [12, 11], [10], [28], [29]). In [14], only Gaussian additive noise is considered and no blur has been introduced. Blur has been introduced in [29]. In fact, in [29] more general degradations have been taken into account. The authors consider model of the form $f=L u+b$ where $f$ is the observed image, $L$ a linear operator and $b$ a Gaussian additive noise. They compute in the case of a crack the topological gradient and illustrate their approach for various imaging problems (segmentation/restoration, super resolution, demosaicing). Note also that topological gradient methods have been also applied for fine structures detection (e.g. points and filaments) [6], [5], [20]. In this case the cost function is based on second order derivatives. Restoration/segmentation in imaging are in general ill-posed inverse problems and one way to overcome this difficulty is to regularize them. A classical framework to do that is to use a Bayesian formulation which leads to the minimization of an energy consisting in two terms. The first one is a data fidelity term which takes into account both the statistic of the noise and the blur and the second one is an adequate regularizing term. For example if we suppose that the acquisition model is of the form $f=u+b$ where $b$ is Gaussian noise then an anti-log-likelihood estimator amounts to choose as a data fidelity term the $L^{2}$ norm $\|u-f\|_{L^{2}(\Omega)}$. If the noise follows another statistic, of course this term changes. The regularizing term is often based on an $L^{p}$ norm of the gradient. Our main contribution is to generalize the results given in [29] to blurred images contaminated by speckle noise and Poissonian statistic and to give the different expression of the topological gradient associated to the cost function $J_{\Omega}(u)=\int_{\Omega}|\nabla u|^{2}$. More precisely we will consider variational problems of the following form

Speckle and Gaussian noise model :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega)} \frac{\gamma}{2} J_{\Omega}(u)+\int_{\Omega} \psi(x, K u) \tag{1}
\end{equation*}
$$

Poissonian model :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega)} \frac{\gamma}{2} J_{\Omega}(u)+\sum_{j \in I_{\text {ind }}(\Omega)} \psi_{j}\left(\int_{R_{j}} K u\right) \tag{2}
\end{equation*}
$$

where $I_{\text {ind }}(\Omega)$ is the indices set of pixels, $\gamma>0$ is a parameter, $R_{j}$ is a regular domain modeling pixel $j$ such that $\Omega$ is the disjoint union of $\left(R_{i}\right)_{i \in I_{\text {ind }}(\Omega)}, K: L^{2}(\Omega) \longrightarrow L^{2}(\Omega)$ is a convolution operator (generally positive and such that $K \mathbb{1} \neq 0$ ) representing the blur. The functions $\psi(x, u)$ and $\psi_{j}(v)$ will be specified in section 5 and section 6 . Note that problems (1) and (2) are semi-linear and one of our contribution is to show they are well-posed and verify some maximum principles. Speckle noise is a multiplicative noise of gamma law, which is present in SAR images, laser images, microscope images
[27, 24, 38]. A Poisson statistic occurs in confocal microscopy [19], emission tomography [39] and single-photon emission computed tomography [23].

In section 2, we recall the classical rationale justifying the modelization of the data fidelity term in a Bayesian approach. In section 3 we set the variational problem taking into account the blurring. Then in section 4, we give the topological gradient for a blurred and Gaussian noisy image. In section 5 we show that problem (1) is well-posed and compute the associated topological gradient both for perforated and cracked domains (in fact, we study a more general class of problem (1)). In section 6 we treat the Poissonian model (2) whose energy is not standard. We summarize in Table 1 (section 7) all the expressions of the topological gradient according to the type of noise and to the infinitesimal perturbation. In section 8 we show how to apply the notion of topological gradient to restore degraded images. Finally in section 9 , we present, for all the models, a detailed numerical analysis for computing the topological gradient and we display various experimental results illustrating each of them.

We conclude this section by giving some notations and assumptions : we suppose to simplify that $x_{0}=0$ and we denote by $\|u\|_{m, \Omega}$ the $H^{m}(\Omega)$-norm of the Sobolev space $H^{m}(\Omega)=\left\{u, D^{\alpha} u \in L^{2}(\Omega),|\alpha| \leq m\right\}$ and by $\|u\|_{H^{1}(\Omega) / \mathbb{R}}$ the norm on the quotient space $H^{1}(\Omega) / \mathbb{R}$. We set $J_{\Omega}(u)=\int_{\Omega}|\nabla u|^{2}$ and $J_{\Omega_{\varepsilon}}(u)=$ $J_{\mathcal{E}}(u)$.

Only the proof for a perforated domain is performed since for a cracked domain the explicit dependency on the data is killed by the fact that the crack has a null Lebesgue measure. Hence we just give the topological gradient expression for a cracked domain (b) and develop the full proof for a perforated domain (a).

## 2 A Bayesian approach

In this section we recall the classical Bayesian approach allowing to deduce the suitable variational model for restoring noisy images. We denote $N$ the number of pixels in the support of the image $\Omega$. The discrete domain is denoted $\Omega^{N}$. We set $u^{N}$ (respectively $f^{N}$ ) the discrete version of the image $u$ to recover (respectively of the observed image $f$ ). For each pixels $p_{x} \in \Omega^{N}, f^{N}\left(p_{x}\right)$ and $u^{N}\left(p_{x}\right)$ can be viewed as a realization of the random variables $U^{N}\left(p_{x}\right)$ and $F^{N}\left(p_{x}\right)$ where $U^{N}$ and $F^{N}$ stand for the random vector formed by these variables at each pixel. We suppose that they are identically distributed and independent. The reasoning is as follows : we express the a priori density probability $g_{U^{N} \mid F^{N}}$ and then we search for $u^{N}$ as the value maximizing this density probability (a Maximum A Priori estimator). A discrete model associated to a discrete energy is deduced and then passing to the limit when $N \rightarrow \infty$ we get the continuous variational model. Let $g_{U^{N} \mid F^{N}}$ be the a posteriori density probability that we want to maximize with respect to $u^{N}$. Thanks to the Bayes rule, $g_{U^{N} \mid F^{N}}$ expresses as :

$$
g_{U^{N} \mid F^{N}}=\left\{\begin{array}{l}
\frac{g_{\left(F^{N}, U^{N}\right)}}{g_{F^{N}}}=\frac{g_{F^{N} \mid U^{N}} g_{U^{N}}}{g_{F^{N}}}, \text { if } g_{F^{N}}>0 \\
0, \text { otherwise }
\end{array}\right.
$$

$g_{F^{N} \mid U^{N}}$ depends on the noise model and $g_{U^{N}}$ is an a priori density probability. Writing that $u^{N}$ is a minimum of $-\log \left(g_{U^{N} \mid F^{N}}\right)$ we get

$$
u^{N}=\underset{u}{\operatorname{argmin}} E^{N}(u)
$$

where

$$
\begin{equation*}
E^{N}(u)=-\log \left(g_{F^{N} \mid U^{N}}(u, f)\right)-\log \left(g_{U^{N}}(u)\right) \tag{3}
\end{equation*}
$$

The a priori density $g_{U^{N}}$ has to be determined, it will play the role of a regularizing term. In analogy to statistical mechanics, a priori densities are frequently Gibbs functions [22] of the form :

$$
g_{U^{N}}(u)=C \times e^{-\frac{\gamma}{2} J^{N}(u)}, \quad \gamma>0
$$

where $J^{N}(u)$ is a discrete version of a non negative energy functional $J_{\Omega}(u)$ and $C$ is a constant. The choice of the density probability $g_{F^{N} \mid U^{N}}$ depends on the statistic of the model to be considered. Below we review respectively the Gaussian model, the speckle model and finally the Poisson model.

## Gaussian model

A classical modeling of image formations is : $F^{N}=U^{N}+G^{N}$ where $U^{N}$ is the discrete version of the image to recover and $G^{N}$ a Gaussian noise of mean 0 and of standard deviation $\sigma$. The density of the Gaussian noise is $g_{G^{N}}(x)=\frac{1}{\sqrt{2 \pi} \sigma} e^{-\frac{x^{2}}{\sigma^{2}}}$. To simplify we still denote by $F^{N}, S^{N}$ and $U^{N}$ the random variables $F^{N}\left(p_{x}\right), S^{N}\left(p_{x}\right)$ and $U^{N}\left(p_{x}\right)$. Let us express the conditional probability density $g_{F^{N} \mid U^{N}}$. From the definition of the conditional probability we have :

$$
\begin{equation*}
P\left(F^{N} \in \mathscr{A} \mid U^{N}=u\right)=\int_{\mathbb{R}^{\prime}} g_{F^{N} \mid U^{N}}(f \mid u) \mathbb{1}_{f \in \mathscr{A}} d f \tag{4}
\end{equation*}
$$

The conditional probability density $g_{F^{N} \mid U^{N}}(f \mid u)$ is a function of the variable $f$ and depending on a parameter $u$. From the model $F^{N}=U^{N}+G^{N}$, the independency of $U^{N}$ and $G^{N}$ and a change of variable we get

$$
\begin{aligned}
P\left(F^{N} \in \mathscr{A} \mid U^{N}=u\right) & =P\left(U^{N}+G^{N} \in \mathscr{A} \mid U^{N}=u\right) \\
& =P\left(G^{N} \in \mathscr{A}-U^{N} \mid U^{N}=u\right) \\
& =P\left(G^{N} \in \mathscr{A}-u\right) \\
& =\int_{\mathbb{R}} g_{G^{N}}(x) \mathbb{1}_{x \in \mathscr{A}-u} d x \\
& =\int_{\mathbb{R}} g_{G^{N}}(f-u) \mathbb{1}_{f \in \mathscr{A}} d f
\end{aligned}
$$

Hence by identification with (4) we deduce that $g_{F^{N} \mid U^{N}}(f \mid u)=g_{G^{N}}(f-u)$. Thanks to the independency hypothesis, the density of $F^{N} \mid U^{N}$ is the product with respect to each pixel $p_{x}$ of the densities $F^{N}\left(p_{x}\right) \mid U^{N}\left(p_{x}\right)$. So the energy given in (3) rewrites in this case (up to a multiplicative constant ) as

$$
E^{N}(u)=\sum_{p_{x} \in \Omega^{N}} \frac{1}{\sigma^{2}}\left(f^{N}-u\right)^{2}+\gamma J^{N}(u)+C
$$

with $C$ a constant non depending on $u$. The constant $\sigma^{2}$ can be neglected in the model because it can be scaled with the regularization parameter $\gamma$. By passing to the limit when $N \rightarrow+\infty$, we get the following continuous energy

$$
E(u)=\int_{\Omega}(f-u)^{2}+\gamma J_{\Omega}(u)
$$

## Speckle model

For SAR images, the classical modeling is (see [38]) : $F^{N}=S^{N} U^{N}$ where $U^{N}$ is the reflectance of the scene (which is to be recovered) and $S^{N}$ the speckle noise. Let us explicit the law of $S^{N}$. SAR
images are constructed from $L \in \mathbb{N}$ observations $F_{k}^{N}$ for $1 \leq k \leq L$ and for each observations we have $F_{k}^{N}=G_{k}^{N} U^{N}$. Generally $G_{k}^{N}$ is a random variable which follows a negative exponential law with mean 1 and with density $g_{G}(x)=e^{-x} 1_{\{x>0\}}$. Then, the observed image $F^{N}$ is construct from this $L$ observations as : $F^{N}=\frac{1}{L} \sum_{k=1}^{N} F_{k}^{N}=\left(\frac{1}{L} \sum_{k=1}^{N} G_{k}^{N}\right) U^{N}$. We set $S^{N}=\frac{1}{L} \sum_{k=1}^{N} G_{k}^{N} ; S^{N}$ follows a gamma law with density $g_{S^{N}}(x)=\frac{L^{L}}{\Gamma(L)} L^{L-1} e^{-L x} 1_{\{x \geq 0\}}$ with $\Gamma(L)=(L-1)$ ! (the mean of $S^{N}$ is 1 and its variance $\frac{1}{L}$ ). Now we can express the density $g_{F^{N} \mid U^{N}}$. To simplify we still denote by $F^{N}, S^{N}$ and $U^{N}$ the random variables $F^{N}\left(p_{x}\right), S^{N}\left(p_{x}\right)$ and $U^{N}\left(p_{x}\right)$. We start from the definition of the conditional probability :

$$
\begin{equation*}
P\left(F^{N} \in \mathscr{A} \mid U^{N}=u\right)=\int_{\mathbb{R}} g_{F^{N} \mid U^{N}}(f \mid u) \mathbb{1}_{\{f \in \mathscr{A}\}} d f \tag{5}
\end{equation*}
$$

where $g_{F^{N} \mid U^{N}}(f \mid u)$ is a function of the variable $f$ and depending on a parameter $u$. Then, from the model $F^{N}=S^{N} U^{N}$ and the independency of $U^{N}$ and $S^{N}$ we have :

$$
\begin{aligned}
P\left(F^{N} \in \mathscr{A} \mid U^{N}=u\right) & =P\left(S^{N} U^{N} \in \mathscr{A} \mid U^{N}=u\right) \\
& =P\left(\left.S^{N} \in \frac{\mathscr{A}}{U^{N}} \right\rvert\, U^{N}=u\right)=P\left(S^{N} \in \frac{\mathscr{A}}{u}\right)
\end{aligned}
$$

Thanks to the definition of $P$ and by a change of variable we get:

$$
P\left(S^{N} \in \frac{\mathscr{A}}{u}\right)=\int_{\mathbb{R}^{( }} g_{S^{N}}(s) \mathbb{1}_{\left\{s \in \frac{\mathscr{\theta}}{u}\right\}}=\int_{\mathbb{R}} \frac{1}{u} g_{S^{N}}\left(\frac{f}{u}\right) \mathbb{1}_{\{f \in \mathscr{A}\}} d f
$$

Then by identification with (5), we deduce that

$$
\begin{equation*}
g_{F^{N} \mid U^{N}}(f \mid u)=\frac{1}{u} g_{S^{N}}\left(\frac{f}{u}\right) \tag{6}
\end{equation*}
$$

From the independency $F^{N}\left(p_{x}\right)$ and $U^{N}\left(p_{x}\right)$, the density of $F^{N} \mid U^{N}$ is the product with respect to pixels $p_{x}$ of the densities $F^{N}\left(p_{x}\right) \mid u^{N}\left(p_{x}\right)$. By taking the $-\log$ function we deduce that (3) rewrites in this case as

$$
E^{N}(u)=L \sum_{p_{x} \in \Omega^{N}}\left(\frac{f^{N}}{u}+\log (u)\right)+\frac{\gamma}{2} J^{N}(u)+C
$$

for $u \in \mathbb{R}^{N}$ and $u>0$, where C denotes a constant independent of $u$. The factor L can be neglected since it can be scaled with the constant $\gamma$. Passing to the limit as $N \rightarrow \infty$, we deduce the following continuous energy

$$
\begin{equation*}
E(u)=\int_{\Omega}\left(\frac{f}{u}+\log (u)\right) d x+\frac{\gamma}{2} J_{\Omega}(u) \tag{7}
\end{equation*}
$$

## Speckle with Log of the image (Speckle-Log model)

One drawback of (7) is that it is nonconvex. By setting, $v=\log (u)$ and $g=\log (f)$, we deduce from (7) the expression of the data fidelity term in function of $v$ and we introduce the following energy function :

$$
\begin{equation*}
E(v)=\int_{\Omega} v+e^{-(v-g)}+\frac{\gamma}{2} J_{\Omega}(v) \tag{8}
\end{equation*}
$$

which is now a convex function of $v$. The recovered image is then $u=e^{v}$.

## Poissonian model

This model is classical in astronomical and confocal microscopy images [19]. Poissonian observations originates from the stochastic nature of photons emission. We denote $R_{j}$, for $j \in I_{\text {ind }}(\Omega)$, the domain of $\mathbb{R}^{2}$ modeling pixel $j$ and such that $\Omega$ is the disjoint union of all the $\left(R_{j}\right)_{j \in I_{\text {ind }}(\Omega)}$. We assume that $f$ is a step function constant on each $R_{j}$ and we still denote $f^{N}=f$ the observed image seen as the realization of the random vector $F$.

More precisely for $j \in I_{i n d}(\Omega), f_{j}$ is a realization of a Poisson statistic of mean and variance equal to $\lambda_{j}^{N}=\int_{R_{j}} u^{N}(x) d x$ where $x \mapsto u^{N}(x)$ is a discrete version of $u^{N} \in \mathbb{R}^{N}$ (may be a step function or a bi-linear interpolation). Thanks to the independence of $F_{j}$ and $U_{j}^{N}$, the conditional probability $P\left(F=f \mid U^{N}=u\right)$ is given by :

$$
P\left(F=f \mid U^{N}=u\right)=\prod_{j \in I_{\text {ind }}(\Omega)} \frac{\lambda_{j}^{N f_{j}} e^{-\lambda_{j}^{N}}}{f_{j}!}
$$

and by applying the $-\log$ function, we have :

$$
-\log \left(P\left(F=f \mid U^{N}=u\right)\right)=\sum_{j \in I_{\text {ind }}(\Omega)} \lambda_{j}^{N}-f_{j} \log \left(\lambda_{j}^{N}\right)+C
$$

where $C$ is constant independent of $u$. We deduce that (3) rewrites in this case as

$$
E^{N}(u)=\sum_{j \in I_{\text {ind }}(\Omega)}\left(\lambda_{j}^{N}-f_{j} \log \left(\lambda_{j}^{N}\right)\right)+\frac{\gamma}{2} J^{N}(u)
$$

The dependence of $E^{N}$ with respect to $u$ comes from the definition of $\lambda^{N}$. Passing to the limit we get the continuous energy :

$$
\begin{equation*}
E(u)=\sum_{j \in I_{\text {ind }}(\Omega)}\left(\int_{R_{j}} u(x) d x-f_{j} \log \left(\int_{R_{j}} u(x) d x\right)\right)+\frac{\gamma}{2} J_{\Omega}(u) \tag{9}
\end{equation*}
$$

## 3 Blurring modeling

In most imaging applications the optical material, the motion of the camera or of the target introduce a blur on the observed image (see [34]). Generally spatially invariant blur is modeled as a positive convolution operator $u \mapsto K u$ with $K \mathbb{1} \neq 0$. We denote by $K^{N}$ the $N \times N$ matrix associated to the discrete version of K on $\Omega^{N}$. From section 2 we deduce the following models adapted to each kind of noise and taking into account the blur :

1. Gaussian model : the observed image writes as $F^{N}=K^{N} U^{N}+G^{N}$ and by the same reasoning of section 2 we get the following energy :

$$
\begin{equation*}
E(u)=\int_{\Omega}(f-K u)^{2}+\gamma J_{\Omega}(u) \tag{10}
\end{equation*}
$$

2. Speckle model : the observed image writes as $F^{N}=S^{N} K^{N} U^{N}$ and the energy is

$$
\begin{equation*}
E(u)=\int_{\Omega} \log (K u)+\frac{f}{K u}+\frac{\gamma}{2} J_{\Omega}(u) \tag{11}
\end{equation*}
$$

3. Speckle model with the Log of the image (Speckle-Log model). We recall that the model writes as $G^{N}=V^{N}+T^{N}$ with $V^{N}=\log \left(K^{N} U^{N}\right)$. The deblurring cannot be handled simultaneously with the denoising process. After the denoising step we must solve the problem $V^{N}=\log \left(K^{N} U^{N}\right)$ where the unknown $U^{N}$ can be found by a least square formula :

$$
U^{N}=\left(\left(K^{N}\right)^{T} K^{N}\right)^{-1}\left(K^{N}\right)^{T} e^{V^{N}}
$$

but we know that this problem is ill-posed, particularly when $K$ contains small eigenvalues. For this reason the blurring problem is not handled for speckle noise by our method. In this case, if we only want to correctly restore a blurred and speckled image it is preferable to use (11).
4. Poissonian model : the observed image at pixel $p_{x}$ is a realization of a Poisson statistic of mean $\int_{R_{p_{x}}} K u^{N}(x) d x$, so the energy is

$$
\begin{equation*}
E(u)=\sum_{p_{x} \in I_{\text {ind }}(\Omega)}\left(\int_{R_{p_{x}}} K u(x) d x-f\left(p_{x}\right) \log \left(\int_{R_{p_{x}}} K u(x) d x\right)\right)+\frac{\gamma}{2} J_{\Omega}(u) \tag{12}
\end{equation*}
$$

In the sequel we give the topological gradient for the Gaussian and Poisson models with blur and for the Speckle-Log model without blur.

## 4 Gaussian noise with blurring

We consider problem (1) with the energy (10) :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega)} \gamma J_{\Omega}(u)+\int_{\Omega} \psi(x, K u) \tag{13}
\end{equation*}
$$

with $\psi(x, v)=(f(x)-v)^{2}$. We do not give the computation of the topological gradient here because of the similarity with the case without blurring (see [4], [29]). We just give the topological gradient expression and some experimental results. Note that this expression only needs the resolution of two problems : the direct and the adjoint problems (we will see in the next section why an adjoint problem is necessary). By following the notations used in [4], the direct and the adjoint problems $u_{0}$ and $v_{0}$ are given in this case by :

$$
\left(\mathscr{P}_{0}\right)\left\{\begin{align*}
-\gamma \Delta u_{0}+K^{\star} K u_{0} & =K^{\star} f, \text { in } \Omega  \tag{14}\\
\partial_{n} u_{0} & =0, \text { on } \partial \Omega
\end{align*}\right.
$$

and

$$
\left(\mathscr{Q}_{0}\right)\left\{\begin{align*}
-\gamma \Delta v_{0}+K^{\star} K v_{0} & =K^{\star}\left(2 K u_{0}-f\right), \text { in } \Omega  \tag{15}\\
\partial_{n} v_{0} & =0, \text { on } \partial \Omega
\end{align*}\right.
$$

We can show (see [7] chapter 3) that problems $\left(\mathscr{P}_{0}\right)$ and $\left(\mathscr{Q}_{0}\right)$ are well-posed in $H^{1}(\Omega)$ as soon as $K \mathbb{1} \neq 0$ and $\gamma>0$. The topological gradients at $x_{0} \in \Omega$ for a perforated domain (a) $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon B\right\}}$ and for a cracked domain (b) $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon \sigma\right\}}$, denoted respectively by $I_{\text {Lap }}^{b}\left(x_{0}\right)$ and $I_{\text {Lap }}^{c}\left(x_{0}\right)$ can be easily deduced from the case without blur and are given in the following Theorem.

Theorem 4.1. The topological gradients associated to problems (14) and (15) and to the cost function $J_{\varepsilon}(u)=\int_{\Omega_{\varepsilon}}|\nabla u|^{2}$, for a perforated and a cracked domain are respectively :

$$
\begin{equation*}
I_{\text {Lap }}^{b}\left(x_{0}\right)=-2 \pi \nabla u_{0}\left(x_{0}\right) \cdot \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma}\left(f\left(x_{0}\right)-K u_{0}\left(x_{0}\right)\right)\left(K v_{0}\left(x_{0}\right)-K u_{0}\left(x_{0}\right)\right) \tag{16a}
\end{equation*}
$$

$$
\begin{align*}
& I_{\text {Lap }}^{c}\left(x_{0}\right)=\min _{\|n\|=1} \mathscr{I}\left(x_{0}, \vec{n}\right)  \tag{16b}\\
& \quad \text { with } \mathscr{I}\left(x_{0}, \vec{n}\right)=-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}
\end{align*}
$$

with $u_{0}$ and $v_{0}$ given by (14) and (15) and $\gamma>0$.

## 5 Speckle multiplicative noise

We consider the variational problem (1) with the energy given in (8). More precisely we study the minimization problem :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega)} \frac{\gamma}{2} J_{\Omega}(u)+\int_{\Omega} \psi(x, u) \tag{17}
\end{equation*}
$$

where $\psi(x, u)=u+e^{-(u-g(x))}$ and $g=\log (f)$ is the logarithm of the observed image. We assume that there is no blur i.e. $K$ is the identity operator. To shorten notations we write sometimes $\psi(u)$ instead of $\psi(x, u)$.

Remark 1. In [8] the authors propose a speckle denoising model using the total variation model with the data fidelity term associated to (7).

This section is organized as follows. First, we show that (1) admits a unique solution for a more general class of functions $\psi$ (verifying Hypotheses 1) and we prove that the solution verifies some $\mathrm{min} / \mathrm{max}$ principles. Then we apply the general result to show that problem (17) admits a unique solution with $\psi(x, u)=u+e^{-(u-g(x))}$. Next, we compute the topological gradient for a general function $\psi$ verifying Hypotheses 1 . We still denote $J_{\Omega}(u)=\int_{\Omega}|\nabla u|^{2}$ and $E_{\Omega}(u)=\frac{\gamma}{2} J_{\Omega}(u)+\int_{\Omega} \psi(x, u)$.

### 5.1 Well-posedness of problem (1)

In this subsection we first establish the well-posedness of (1) for a general class of functions $\psi(x, u)$ and then we check that the function $\psi(x, u)$ associated to the speckle-Log model (17) matches these hypotheses. To simplify we suppose that $\gamma=1$ and sometimes we write $\psi(u)$ for $\psi(x, u)$.

Hypotheses 1. Let $\psi: \Omega \times I \longrightarrow \mathbb{R}$ such that
(i) $u \mapsto \psi(x, u) \in C^{3}(I) \forall x \in \bar{\Omega}$
(ii) $x \mapsto D_{u} \psi(x, u) \in C^{0}(\Omega) \forall u \in I$
(iii) $u \mapsto \psi(x, u)$ is stricly convex on I , uniformly with respect to $x \in \Omega$.
(iv) $\psi$ is bounded from below on $\Omega \times I$
(v) $\exists a, b \in I$ such that for all $x, D_{u} \psi(x, a) \leq 0$ and $D_{u} \psi(x, b) \geq 0$ with $[a, b] \subset I$.

Lemma 5.1. Let $\psi(x, u)$ a function verifying Hypotheses 1 , then (1) admits a unique solution $u_{\Omega} \in$ $H^{1}(\Omega)$ such that $a \leq u_{\Omega} \leq b$.

Démonstration. Existence : Let $\left(u_{n}\right)$ a minimizing sequence. There exists a constant $C_{1}$ such that $E_{\Omega}\left(u_{n}\right) \leq C_{1}$. As $\psi(x, u)$ is bounded from below on $\Omega \times I$ there exists a constant $C_{2}$ such that $\int_{\Omega} \psi\left(x, u_{n}\right) \geq$ $C_{2}$. Therefore :

$$
\int_{\Omega}\left|\nabla u_{n}\right|^{2} \leq \max \left(C_{1}, C_{1}-C_{2}\right)
$$

Let $v_{n}=\max \left(u_{n}, a\right)$, and $\Omega_{n}^{-}=\Omega \cap\left\{u_{n} \leq a\right\}$, we have $v_{n} \geq a$ and

$$
E_{\Omega}\left(v_{n}\right)-E_{\Omega}\left(u_{n}\right)=-\int_{\Omega_{n}^{-}}\left|\nabla u_{n}\right|^{2}+\int_{\Omega_{n}^{-}} \psi(a)-\psi\left(u_{n}\right)
$$

By convexity:
$\psi\left(u_{n}\right)-\psi(a) \geq D_{u} \psi(a)\left(u_{n}-a\right)$ and $\int_{\Omega_{n}^{-}}\left(\psi(a)-\psi\left(u_{n}\right)\right) \leq \int_{\Omega_{n}^{-}} D_{u} \psi(a)\left(a-u_{n}\right) \leq 0$.
We easily deduce that $E_{\Omega}\left(v_{n}\right) \leq E_{\Omega}\left(u_{n}\right)$. Thus $v_{n}$ is still a minimizing sequence and $v_{n} \geq a$. Similarly by setting $w_{n}=\min \left(v_{n}, b\right)$, we get $w_{n} \leq b$ and $w_{n}$ is a minimizing sequence. Therefore we can suppose that any minimizing sequence $u_{n}$ verifies $a \leq u_{n} \leq b$. It is easily seen that $u_{n}$ is bounded in $H^{1}(\Omega)$. Thus, up to a subsequence there exists $u \in H^{1}(\Omega)$ such that $u_{n} \xrightarrow{L^{2}(\Omega)} u$ and $u_{n} \xrightarrow{H^{1}(\Omega)} u$ (where ${ }^{H^{1}(\Omega)}$ stands for the weak topology). By using the lower semi-continuity of $J_{\Omega}(u)$ and Fatou's Lemma we get that $u$ is a solution of (1). Moreover we have $a \leq u \leq b$ a.e. on $\Omega$.

Uniqueness: From the existence, we can work on the set $H(\Omega)=\left\{v \in H^{1}(\Omega), a \leq v \leq b\right\}$. Since $\psi(u)$ is strictly convex on $[a, b] \subset I$ and $J_{\Omega}(u)$ is strictly convex on $H^{1}(\Omega)$, we deduce that $E_{\Omega}(u)$ is strictly convex on $H(\Omega)$ which is a convex set and that $E_{\Omega}$ has a unique minimum in $H^{1}(\Omega)$.
We apply below Lemma 5.1 to the speckle-Log model.
Proposition 1. Let $f$ a function such that $\forall x, 0<\alpha \leq f(x) \leq \beta$ where $\alpha$ and $\beta$ are two constants, then problem (17) with $\phi(u)=|\nabla u|^{2}$ and $\psi(x, u)=u+e^{-(u-g(x))}$ where $g=\log (f)$ has a unique solution $u \in H^{1}(\Omega)$ and $\forall x, \log (\alpha) \leq u \leq \log (\beta)$.

Démonstration. A standard computation leads to $D_{u} \psi(u)=1-e^{-(u-g)}$ and $D_{u}^{2} \psi(u)=e^{-(u-g)}>0$. Hence $\psi(u)$ is strictly convex on $]-\infty, \eta] \forall \eta \in \mathbb{R}$. By using that $0<\alpha \leq f \leq \beta$ we get

$$
1-e^{-(u-\log (\beta))} \leq D_{u} \psi(u) \leq 1-e^{-(u-\log (\alpha))}
$$

Let $a=\log (\alpha)$ and $b=\log (\beta)$, the following inequalities hold

$$
D_{u} \psi(b) \geq 0 \text { and } D_{u} \psi(a) \leq 0
$$

From Lemma 5.1, there exists a unique function $u \in H^{1}(\Omega)$ solution of (17). Moreover we have $a \leq u \leq b$.

In the next subsection we detail the computation of the topological gradient for a perforated domain and we just give the result for a cracked domain.

### 5.2 Computation of the topological gradient for a perforated domain

Let $u_{\varepsilon}=u_{\Omega_{\varepsilon}}$ be the solution of problem (1) replacing $\Omega$ by $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon B\right\}}$ and let $J_{\varepsilon}(u)=$ $J_{\Omega_{\varepsilon}}(u)$. In order to establish the topological expansion for a more general class of problems, we assume that $\psi(x, u)$ verifies Hypotheses 1 . By writing that $D E_{\Omega_{\varepsilon}}\left(u_{\varepsilon}\right) . v=0$ for all $v \in H^{1}\left(\Omega_{\varepsilon}\right)$, we obtain the following variational formulation : find $u_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right)$ such that

$$
\begin{equation*}
\int_{\Omega_{\varepsilon}} \gamma \nabla u_{\varepsilon} . \nabla v+D_{u} \psi\left(x, u_{\varepsilon}\right) v=0 \quad \forall v \in H^{1}\left(\Omega_{\varepsilon}\right) \tag{18}
\end{equation*}
$$

Then by an integration by parts, $u_{\varepsilon}$ necessarily verifies the Euler equations :

$$
\left(\mathscr{P}_{\varepsilon}\right)\left\{\begin{align*}
-\gamma \Delta u_{\varepsilon}+D_{u} \psi\left(x, u_{\varepsilon}\right) & =0 \text { on } \Omega_{\varepsilon}  \tag{19}\\
\partial_{n} u_{\varepsilon} & =0 \text { on } \partial \Omega_{\varepsilon}
\end{align*}\right.
$$

If $F_{\varepsilon}$ stands the following functional

$$
\begin{equation*}
F_{\varepsilon}(u, v)=\int_{\Omega_{\varepsilon}} \gamma \nabla u . \nabla v+D_{u} \psi(x, u) v \quad \forall u, v \in H^{1}\left(\Omega_{\varepsilon}\right) \tag{20}
\end{equation*}
$$

Then (18) rewrites as : find $u_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right)$ such that $F_{\varepsilon}\left(u_{\varepsilon}, v\right)=0 \quad \forall v \in H^{1}\left(\Omega_{\varepsilon}\right)$.
We denote by $u_{0}$ the solution of (18) for $\varepsilon=0$ and we state the main result of this section. To simplify notations we perform the proof with $\gamma=1$ (but we state the Theorem with $\gamma>0$ ).

Theorem 5.2. The topological gradient $I_{\text {Lap }}^{b}$ associated to problem (19) with $\psi(x, u)$ verifying Hypotheses 1 and with the cost function $J_{\varepsilon}(u)=\int_{\Omega_{\varepsilon}}|\nabla u|^{2}$ for a perforated domain is

$$
\begin{equation*}
I_{\text {Lap }}^{b}\left(x_{0}\right)=-2 \pi \nabla u_{0}\left(x_{0}\right) \cdot \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma} D_{u} \psi\left(x_{0}, u_{0}\right)\left(u_{0}\left(x_{0}\right)-v_{0}\left(x_{0}\right)\right) \tag{21}
\end{equation*}
$$

with $u_{0}$ and $v_{0}$ given by (19) and (26) for $\varepsilon=0$ and with $\gamma>0$.
Remark 2. The topological sensitivity for some non linear PDEs using as main operator the Laplace operator with Dirichlet boundary conditions (on the boundary of the small perturbation) has been studied by Amstutz [2].

Démonstration. The proof is very technical and can be skipped by readers not interested by details. The topological gradient is given by the leading term in the difference $J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right)$. Let us introduce the functional $\widetilde{J}_{\varepsilon}(u)=-\int_{\Omega_{\varepsilon}} D_{u} \psi(x, u) u$; by using (18) with $v=u_{\varepsilon}$, it is straightforward that :

$$
\begin{align*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right) & =\widetilde{J}_{\varepsilon}\left(u_{\varepsilon}\right)-\widetilde{J}_{0}\left(u_{0}\right) \\
& =-\int_{\Omega_{\varepsilon}}\left(D_{u} \psi\left(u_{\varepsilon}\right) u_{\varepsilon}-D_{u} \psi\left(u_{0}\right) u_{0}\right)+\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) u_{0} \\
& =-\int_{\Omega_{\varepsilon}} D_{u} \psi\left(u_{0}\right)\left(u_{\varepsilon}-u_{0}\right)+\left(D_{u} \psi\left(u_{\varepsilon}\right)-D_{u} \psi\left(u_{0}\right)\right) u_{0}  \tag{22}\\
& +\int_{\Omega_{\varepsilon}}\left(D_{u} \psi\left(u_{\varepsilon}\right)-D_{u} \psi\left(u_{0}\right)\right)\left(u_{\varepsilon}-u_{0}\right) \\
& =L_{\varepsilon}\left(u_{\varepsilon}-u_{0}\right)+\mathscr{I}_{\varepsilon}+\mathscr{E}_{1}
\end{align*}
$$

where

$$
\begin{align*}
L_{\varepsilon}(u) & =-\int_{\Omega_{\varepsilon}}\left(D_{u} \psi\left(u_{0}\right) u+D_{u}^{2} \psi\left(u_{0}\right) u_{0} u\right) \\
\mathscr{I}_{\varepsilon} & =\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) u_{0}  \tag{23}\\
\mathscr{E}_{1} & =-\int_{\Omega_{\varepsilon}}\left(\frac{1}{2} D_{u}^{3} \psi\left(u_{\eta_{\varepsilon}}\right) u_{0}+D_{u}^{2} \psi\left(u_{\delta_{\varepsilon}}\right)\right)\left(u_{\varepsilon}-u_{0}\right)^{2}
\end{align*}
$$

and where $u_{\delta_{\varepsilon}}=\theta_{1} u_{0}+\left(1-\theta_{1}\right) u_{\varepsilon}, u_{\eta_{\varepsilon}}=\theta_{2} u_{0}+\left(1-\theta_{2}\right) u_{\varepsilon}$ with $\theta_{i}: \Omega \mapsto \mathbb{R}, 0 \leq \theta_{i} \leq 1$ for $1 \leq i \leq 2$. To compute the term $L_{\varepsilon}\left(u_{\varepsilon}-u_{0}\right)$ in (22) it is classical [3,37] to introduce an adjoint problem. Due to the non linearity of the direct problem, we first make a second order Taylor expansion of $F_{\varepsilon}(u, v)$ with respect to $u$ at point $u_{0}$ :

$$
\begin{align*}
F_{\varepsilon}\left(u_{\varepsilon}, v\right) & =F_{\varepsilon}\left(u_{0}, v\right)+D_{u} F_{\varepsilon}\left(u_{0}, v\right) \cdot\left(u_{\varepsilon}-u_{0}, v\right)+\frac{1}{2} D_{u}^{2} F_{\varepsilon}\left(u_{\delta_{\varepsilon}}, v\right) \cdot\left(u_{\varepsilon}-u_{0}, u_{\varepsilon}-u_{0}\right)  \tag{24}\\
& =F_{\varepsilon}\left(u_{0}, v\right)+b_{\varepsilon}\left(u_{\varepsilon}-u_{0}, v\right)+c_{\varepsilon}\left(u_{\varepsilon}-u_{0}, u_{\varepsilon}-u_{0}, v\right)
\end{align*}
$$

with

$$
\begin{aligned}
b_{\varepsilon}(u, v) & =D F_{\varepsilon}\left(u_{0}, v\right) \cdot u=\int_{\Omega_{\varepsilon}} \nabla u \cdot \nabla v+\int_{\Omega_{\varepsilon}} D_{u}^{2} \psi\left(u_{0}\right) u v \\
c_{\varepsilon}(u, t, v) & =\frac{1}{2} D_{u}^{2} F_{\varepsilon}\left(u_{\delta_{\varepsilon}}, v\right) \cdot(u, t)=\int_{\Omega_{\varepsilon}} \frac{1}{2} D_{u}^{3} \psi\left(u_{\eta_{\varepsilon}}\right) u t v
\end{aligned}
$$

Then the adjoint solution $v_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right)$ is defined as :

$$
\begin{equation*}
b_{\varepsilon}\left(u, v_{\varepsilon}\right)=-L_{\varepsilon}(u) \quad \forall u \in H^{1}\left(\Omega_{\varepsilon}\right) \tag{25}
\end{equation*}
$$

The Euler equations associated with (25) are :

$$
\left(\mathscr{Q}_{\varepsilon}\right)\left\{\begin{align*}
-\gamma \Delta v_{\varepsilon}+D_{u}^{2} \psi\left(u_{0}\right) v_{\varepsilon} & =D_{u} \psi\left(u_{0}\right)+D_{u}^{2} \psi\left(u_{0}\right) u_{0} \text { on } \Omega_{\varepsilon}  \tag{26}\\
\partial_{n} v_{\varepsilon} & =0 \text { on } \partial \Omega_{\varepsilon}
\end{align*}\right.
$$

Remark 3. (i) In the proof, we take $\gamma=1$.
(ii) The adjoint problem $\left(\mathscr{Q}_{\varepsilon}\right)$ is linear and we can notice that the strict convexity of $u \mapsto \psi(x, u)$ is necessary to $\left(\mathscr{Q}_{\varepsilon}\right)$ be coercive. Since $u \mapsto \psi(x, u)$ is $C^{2}(I)$ and thanks to Lemma 5.1 there exist two constants $A, B \in \mathbb{R}$ such that

$$
A<D_{u} \psi\left(u_{0}\right)+D_{u}^{2} \psi\left(u_{0}\right) u_{0}<B
$$

Hence (26) is well-posed and thanks to Lemma 5.1, we have the following inegality

$$
\frac{A}{\sup _{\Omega} D_{u}^{2} \psi\left(u_{0}\right)} \leq v_{\varepsilon} \leq \frac{B}{\inf _{\Omega} D_{u}^{2} \psi\left(u_{0}\right)}
$$

We deduce from (22) and (24) that :

$$
\begin{align*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right) & =-b_{\varepsilon}\left(u_{\varepsilon}-u_{0}, v_{\varepsilon}\right)+\mathscr{I}_{\varepsilon}+\mathscr{E}_{1} \\
& =F_{\varepsilon}\left(u_{0}, v\right)-F_{\varepsilon}\left(u_{\varepsilon}, v_{\varepsilon}\right)+c_{\varepsilon}\left(u_{\varepsilon}-u_{0}, u_{\varepsilon}-u_{0}, v_{\varepsilon}\right)+\mathscr{I}_{\varepsilon}+\mathscr{E}_{1}  \tag{27}\\
& =F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)+\mathscr{E}_{2}+\mathscr{I}_{\varepsilon}+\mathscr{E}_{1}
\end{align*}
$$

with

$$
\begin{equation*}
\mathscr{E}_{2}=c_{\varepsilon}\left(u_{\varepsilon}-u_{0}, u_{\varepsilon}-u_{0}, v_{\varepsilon}\right) \tag{28}
\end{equation*}
$$

By using an integration by parts, the term $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ expresses as :

$$
\begin{aligned}
F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right) & =\int_{\Omega_{\varepsilon}} \nabla u_{0} . \nabla v_{\varepsilon}+D_{u} \psi\left(u_{0}\right) v_{\varepsilon} \\
& =-\int_{\partial B_{\varepsilon}} \partial_{n} u_{0} v_{\varepsilon}+\int_{\Omega_{\varepsilon}}\left(-\Delta u_{0}+D_{u} \psi\left(u_{0}\right)\right) v_{\varepsilon} \\
& =-\int_{\partial B_{\varepsilon}} \partial_{n} u_{0} w_{\varepsilon}-\int_{\partial B_{\varepsilon}} \partial_{n} u_{0} v_{0}
\end{aligned}
$$

with $w_{\varepsilon}=v_{\varepsilon}-v_{0}$. Now for $\varphi \in H^{1 / 2}\left(\partial B_{\varepsilon}\right)$ we introduce the following extension on $B_{\varepsilon}$ :

$$
\left\{\begin{align*}
\Delta l_{\varepsilon}^{\varphi} & =0, \text { on } B_{\varepsilon}  \tag{29}\\
l_{\varepsilon}^{\varphi} & =\varphi, \operatorname{sur} \partial B_{\varepsilon}
\end{align*}\right.
$$

For $v \in H^{1}\left(\Omega_{\varepsilon}\right)$, we denote by $l_{\varepsilon}^{v}$ the harmonic function defined on $B_{\varepsilon}$ such that $l_{\varepsilon}^{v}=v$ on $\partial B_{\varepsilon}$. By integration by parts $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ writes as

$$
\begin{align*}
F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right) & =-\int_{B_{\varepsilon}}\left(\nabla u_{0} . \nabla v_{0}+\Delta u_{0} v_{0}\right)-\int_{B_{\varepsilon}}\left(\nabla u_{0} \cdot \nabla l_{\varepsilon}^{w_{\varepsilon}}+\Delta u_{0} \cdot l_{\varepsilon}^{w_{\varepsilon}}\right) \\
& =-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}} \partial_{n} v_{0}+\int_{B_{\varepsilon}} \widetilde{u_{0}} \Delta v_{0}-\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) v_{0}-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}} \partial_{n} l_{\varepsilon}^{w_{\varepsilon}}-\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) l_{\varepsilon}^{w_{\varepsilon}}  \tag{30}\\
& =-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}}\left(\partial_{n} v_{0}+\partial_{n} l_{\varepsilon}^{w_{\varepsilon}}\right)-\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) v_{0}+\mathscr{E}_{3}+\mathscr{E}_{4} \\
& =\mathscr{J}_{\varepsilon}+\mathscr{K}_{\varepsilon}+\mathscr{E}_{3}+\mathscr{E}_{4}
\end{align*}
$$

with $\widetilde{u_{0}}=u_{0}-u_{0}(0)$ and

$$
\begin{align*}
& \mathscr{J}_{\varepsilon}=-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}}\left(\partial_{n} v_{0}+\partial_{n} l_{\varepsilon}^{w_{\varepsilon}}\right) \quad \mathscr{K}_{\varepsilon}=-\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) v_{0} \\
& \mathscr{E}_{3}=\int_{B_{\varepsilon}} \widetilde{u_{0}} \Delta v_{0} \quad \mathscr{E}_{4}=-\int_{B_{\varepsilon}} D_{u} \psi\left(u_{0}\right) l_{\varepsilon}^{w_{\varepsilon}} \tag{31}
\end{align*}
$$

The computation of $\mathscr{J}_{\varepsilon}, \mathscr{E}_{1}, \mathscr{E}_{2}$ and $\mathscr{E}_{4}$ needs to approximate $w_{\varepsilon}$ and $X_{\varepsilon}=u_{\varepsilon}-u_{0}$ in the $H^{1}\left(\Omega_{\varepsilon}\right)$ sense. From Lemma 10.3 (see Appendix A), we can show that $\left\|X_{\mathcal{\varepsilon}}\right\|_{0, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$ and that $w_{\varepsilon}=\varepsilon Q\left(\frac{x}{\varepsilon}\right)+r_{\varepsilon}$ where Q is given by (57) with $g=-\nabla v_{0}(0) \cdot \vec{n}$ and $\left\|r_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$.

Proposition 2. Let $\mathscr{I}_{\varepsilon}, \mathscr{J}_{\varepsilon}, \mathscr{K}_{\varepsilon}, \mathscr{E}_{1}, \mathscr{E}_{2}, \mathscr{E}_{3}$ and $\mathscr{E}_{4}$ given by (23), (28) and (31), we have the following estimations :

$$
\begin{align*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right) & =\mathscr{I}_{\varepsilon}+\mathscr{J}_{\varepsilon}+\mathscr{K}_{\varepsilon}+\sum_{i=1}^{4} \mathscr{E}_{i} \\
\mathscr{I}_{\varepsilon} & =\pi \varepsilon^{2} D_{u} \psi\left(0, u_{0}\right) u_{0}(0)  \tag{32}\\
\mathscr{J}_{\varepsilon} & =-2 \pi \varepsilon^{2} \nabla u_{0}(0) . \nabla v_{0}(0) \\
\mathscr{K}_{\varepsilon} & =-\pi \varepsilon^{2} D_{u} \psi\left(0, u_{0}\right) v_{0}(0)
\end{align*}
$$

and $\mathscr{E}_{i} \sim o\left(\varepsilon^{2}\right)$ for $i \in[\lfloor 1 . .4]$
Démonstration. The first equality is straightforward. Lemma 10.3, the regularity of $\psi(x, u)$ and Proposition 1 allow to say that $\mathscr{E}_{1}$ and $\mathscr{E}_{2}$ are $O\left(\left\|u_{\varepsilon}-u_{0}\right\|_{0, \Omega_{\varepsilon}}^{2}\right)$ and then $\sim O\left(-\varepsilon^{4} \log (\varepsilon)\right)$. We recall the notation $\psi(u)$ means $\psi(x, u(x))$. A change of variable (CV) and the continuity of $x \mapsto D_{u} \psi\left(x, u_{0}(x)\right)$ lead to

$$
\begin{aligned}
& \mathscr{I}_{\varepsilon}=\varepsilon^{2} \int_{B} D_{u} \psi\left(\varepsilon X, u_{0}(\varepsilon X)\right) u_{0}(\varepsilon X)=\varepsilon^{2} \pi D_{u} \psi\left(0, u_{0}\right) u_{0}(0)+o\left(\varepsilon^{2}\right) \\
& \mathscr{K}_{\varepsilon}=\varepsilon^{2} \int_{B} D_{u} \psi\left(\varepsilon X, u_{0}(\varepsilon X)\right) v_{0}(\varepsilon X) d X=\varepsilon^{2} \pi D_{u} \psi\left(0, u_{0}\right) v_{0}(0)+o\left(\varepsilon^{2}\right)
\end{aligned}
$$

Again with a CV, the equality $\Delta v_{0}=D_{u}^{2} \psi\left(u_{0}\right)\left(v_{0}-u_{0}\right)-D_{u} \psi\left(u_{0}\right)$ and the regularity of $\psi(u)$, we get $\mathscr{E}_{3} \sim O\left(\varepsilon^{3}\right)$. By using Lemma 10.4 and Lemma 10.1 (see Appendix A) we have :

$$
\begin{aligned}
\mathscr{J}_{\varepsilon} & =\int_{\partial B}\left(u_{0}-u_{0}(0)\right)\left(\partial_{n} l_{\varepsilon}^{p_{\varepsilon}}+\partial_{n} v_{0}\right)+\int_{\partial B_{\varepsilon}}\left(u_{0}-u_{0}(0)\right) \partial_{n} l_{\varepsilon}^{e_{\varepsilon}} \\
& =\varepsilon^{2} \nabla u_{0}(0) \cdot \int_{\partial B} \lambda(x) x+\mathscr{F}_{1}+\mathscr{F}_{2}
\end{aligned}
$$

with

$$
\begin{aligned}
\lambda(x) & =-2 \nabla v_{0}(0) \cdot \vec{n} \\
\mathscr{F}_{1} & =\int_{\partial B_{\varepsilon}}\left(u_{0}-u_{0}(0)\right)\left(\partial_{n} v_{0}-\nabla v_{0}(0) \cdot \vec{n}\right) \\
\mathscr{F}_{2} & =\int_{\partial B_{\varepsilon}}\left(u_{0}-u_{0}(0)\right) \partial_{n} l_{\varepsilon}^{e_{\varepsilon}}
\end{aligned}
$$

A CV and a Taylor expansion of $u_{0}$ and $v_{0}$ at 0 lead to $\mathscr{F}_{1}=O\left(\varepsilon^{3}\right)$ and $\mathscr{F}_{2} \leq C \varepsilon^{2}\left\|\partial_{n} l_{\varepsilon}^{e_{\varepsilon}}(\varepsilon X)\right\|_{-1 / 2, \partial B}$. For $\mathscr{F}_{2}$ it suffices to make a CV and use the trace Theorem on $B_{2} \backslash \bar{B}$ :

$$
\begin{aligned}
\left\|\partial_{n} l_{\varepsilon}^{e_{\varepsilon}}(\varepsilon X)\right\|_{-1 / 2, \partial B} & \leq \frac{1}{\varepsilon}\left|l_{\varepsilon}^{e_{\varepsilon}}(\varepsilon X)\right|_{1, B}=\frac{1}{\varepsilon}\left|l^{e_{\varepsilon}(\varepsilon X)}(X)\right|_{1, B} \\
& \leq \frac{C}{\varepsilon}\left\|e_{\varepsilon}(\varepsilon X)\right\|_{H^{1 / 2}(\partial B) / \mathbb{R}} \leq C\left\|e_{\varepsilon}(\varepsilon X)\right\|_{H^{1}\left(B_{2} \backslash \bar{B}\right) / \mathbb{R}}
\end{aligned}
$$

Now from the equivalency of the $H^{1}\left(B_{2} \backslash \bar{B}\right) / \mathbb{R}$-norm with the semi norm and a CV we get $\left\|\partial_{n} l_{\varepsilon}^{e_{\varepsilon}}(\varepsilon X)\right\|_{-1 / 2, \partial B} \leq$ $C\left|e_{\varepsilon}\right|_{1, \Omega_{\varepsilon}}$. By using Lemma 10.4, we obtain $\mathscr{F}_{2}=O\left(\varepsilon^{3} \sqrt{-\log (\varepsilon)}\right)$. Finally by using a CV, the continuity of $\varphi \mapsto l^{\varphi}$ from $H^{1 / 2}(\partial B)$ to $H^{1}(B)$, the trace Theorem on $B_{2} \backslash \bar{B}$, again a CV and Lemma 10.4 we have :

$$
\begin{aligned}
\left|\mathscr{E}_{4}\right| & \leq C \varepsilon^{2}\left\|l_{\varepsilon}^{w_{\varepsilon}}(\varepsilon X)\right\|_{0, B}=C \varepsilon^{2}\left\|l^{w_{\varepsilon}(\varepsilon X)}(X)\right\|_{0, B} \\
& \leq C \varepsilon^{2}\left\|w_{\varepsilon}(\varepsilon X)\right\|_{1 / 2, \partial B} \leq C\left\|w_{\varepsilon}(\varepsilon X)\right\|_{1, B_{2} \backslash \bar{B}} \\
& \leq C \varepsilon^{2}\left(\frac{1}{\varepsilon}\left\|w_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}+\left|w_{\varepsilon}\right|_{1, \Omega_{\varepsilon}}\right) \leq C \varepsilon^{3} \sqrt{-\log (\varepsilon)}
\end{aligned}
$$

The topological expression is easily deduced from Proposition 2 and Theorem 5.2 is proven.

### 5.3 Expression of the topological gradient for a cracked domain

For the cracked domain $\Omega_{\varepsilon}=\Omega \backslash \overline{x_{0}+\varepsilon \sigma(\vec{n})}$, computations are similar. The term $\mathscr{I}_{\varepsilon}$ of (27) is zero and the term $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ expresses as $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)=-\int_{B_{\varepsilon}} \partial_{n} u_{0}\left[w_{\varepsilon}\right]$. The asymptotic expansion of $u_{\varepsilon}$ and $v_{\varepsilon}$ are similar and then the computation of the topological gradient is the same as in the linear case (see [4, 21] for more details).

Theorem 5.3. The topological gradient $I_{\text {Lap }}^{c}$ associated to problem (19) and to the cost function $J_{\mathcal{E}}(u)=\int_{\Omega_{\varepsilon}}|\nabla u|^{2}$ for a cracked domain is

$$
\begin{align*}
& I_{\text {Lap }}^{c}\left(x_{0}\right)=\min _{\|\vec{n}\|=1} \mathscr{I}\left(x_{0}, \vec{n}\right)  \tag{33}\\
& \quad \text { with } \mathscr{I}\left(x_{0}, \vec{n}\right)=-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}
\end{align*}
$$

with $u_{0}$ and $v_{0}$ given by (19) and (26) for $\varepsilon=0$ and with $\gamma>0$ in front of the Laplacian.

## 6 Poissonian model with blurring

We recall that we model the observed image $f$ as a step function defined by $f_{j} \in \mathbb{R}$ on $R_{j}$ for $j \in I_{\text {ind }}(\Omega)$, where $R_{j}$ is a regular domain of $\mathbb{R}^{2}$ modeling pixel $j$ and we denote $N=\left|I_{\text {ind }}(\Omega)\right|$. To simplify, we suppose that $\left|R_{j}\right|=1$ and that $\Omega$ is the disjoint union of the $\left(R_{i}\right)_{i \in I_{i n d}(\Omega)}$. We assume that
$\min _{j \in I_{\text {ind }}(\Omega)} f_{j}>0$. We recall the general minimization problem associated to the Poisson model given in (12) :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega)} \frac{\gamma}{2} J_{\Omega}(u)+\sum_{j \in I_{\text {ind }}(\Omega)} \psi_{j}\left(\int_{R_{j}} K u\right) \tag{34}
\end{equation*}
$$

with $\psi_{j}(x)=x-f_{j} \log (x)$. We still denote in this section $J_{\Omega}(u)=\int_{\Omega}|\nabla u|^{2}$ and $E_{\Omega}(u)=\frac{\gamma}{2} J_{\Omega}(u)+$ $\sum_{j \in I_{i n d}(\Omega)} \psi_{j}\left(\int_{R_{j}} K u\right)$.

First we show that problem (34) is well-posed, then we compute the topological gradient for a perforated domain (a) : $\Omega_{\varepsilon}=\Omega \backslash \overline{x_{0}+\varepsilon B}$, and we give the expression for a cracked domain (b) : $\Omega_{\varepsilon}=\Omega \backslash \overline{x_{0}+\varepsilon \sigma}$ without proof.

### 6.1 Well-posedness of problem (34)

Proposition 3. Let $f$ a step function such that $\min _{i} f_{i}>0$ and $\max _{i} f_{i}<+\infty$, then problem (34) with $\psi_{j}(x)=x-f_{j} \log (x)$ for $j \in I_{\text {ind }}(\Omega)$ admits a unique solution $u \in H^{1}(\Omega)$.
Moreover this solution verifies $\alpha \leq \int_{R_{i}} u \leq \beta, \forall i \in I_{\text {ind }}(\Omega)$ with $\alpha=\frac{\min _{i} f_{i}}{N}$ and $\beta=\int_{\Omega} f=\sum_{i \in I_{\text {ind }}(\Omega)} f_{i}$.

Démonstration. Existence : To simplify the proof we suppose that $K$ is the identity operator and $\gamma=1$. The proof for the general case is quite similar. For more details see chapter 3 of [7].
We must add a priori to (34) the condition $\int_{R_{i}} u>0, \forall i \in I_{\text {ind }}(\Omega)$.
We set $H=\left\{u \in H^{1}(\Omega), \int_{R_{i}} u>0 \forall i \in I_{\text {ind }}(\Omega)\right\}$. Then (34) rewrites as :

$$
\min _{u \in H(\Omega)} \int_{\Omega}|\nabla u|^{2}+\sum_{j} \psi_{j}\left(\int_{R_{j}} u\right)
$$

Let $\left(u_{n}\right)_{n}$ a minimizing sequence of $E_{\Omega}(u)$ in $H(\Omega)$. There exists a constant $D>0$ such that $J_{\Omega}\left(u_{n}\right) \leq$ $D$.
If $C=\sum_{j} \min _{x \in] 0,+\infty[ } \psi_{j}(x)=\sum_{j} f_{j}-f_{j} \log \left(f_{j}\right)>-\infty$, then :

$$
0 \leq \int_{\Omega}\left|\nabla u_{n}\right|^{2} \leq \max (D, D-C)
$$

By using the positiveness of $\int_{\Omega}\left|\nabla u_{n}\right|^{2}$, we deduce that $\sum_{j} \psi_{j}\left(\int_{R_{j}} u_{n}\right) \leq D$. Setting $K_{i}=\sum_{j \neq i} \min _{x} \psi_{j}$, it is straightforward that $\psi_{i}\left(\int_{R_{i}} u_{n}\right) \leq D-K_{i}$ and then

$$
\begin{equation*}
0<\underline{E_{i}} \leq \int_{R_{i}} u_{n} \leq \overline{E_{i}} \tag{35}
\end{equation*}
$$

with $\overline{E_{i}}=\max \left\{\psi_{i}^{-1}\left(D-K_{i}\right)\right\}$ and $\underline{E_{i}}=\min \left\{\psi_{i}^{-1}\left(D-K_{i}\right)\right\}\left(\right.$ we recall the notation $\left.\psi_{i}^{-1}(b)=\{x \in] 0, \infty\left[\psi_{i}(x)=b\right\}\right)$ . Hence the constraint $\int_{R_{i}} u_{n}>0$ is fulfilled. We deduce that $\sum_{i} \underline{E_{i}} \leq \int_{\Omega} u_{n}=\sum_{i} \int_{R_{i}} u_{n} \leq \sum_{i} \overline{E_{i}}$ and thanks to Poincaré-Wirtinger Lemma we get that $u_{n}$ is bounded in $L^{2}(\Omega)$. So, there exist a subsequence $u_{n_{k}}$ (still denoted $u_{n}$ ) and $u \in H^{1}(\Omega)$ such that $u_{n} \xrightarrow{L^{2}(\Omega)} u$ and $u_{n} \xrightarrow{H^{1}(\Omega)^{\star}} u$. We deduce that $J_{\Omega}(u) \leq \liminf J_{\Omega}\left(u_{n}\right)$ and thanks to (35) and Bolzano-Weierstrass Lemma, we can extract a subsequence $u_{n}$ such that $\int_{R_{i}} u_{n} \rightarrow l_{i} \in \mathbb{R}$. By continuity, we have $\psi_{i}\left(\int_{R_{i}} u_{n}\right) \rightarrow \psi_{i}\left(\mathfrak{1}_{i}\right) \forall i \in I_{\text {ind }}(\Omega)$ and

$$
E_{\Omega}(u) \leq \liminf E_{\Omega}\left(u_{n}\right)
$$

which proves that $u$ is a minimizer of $E_{\Omega}(u)$ in $H^{1}(\Omega)$.
Bounds : If $u \in H^{1}(\Omega)$ is the solution of (34) then $D E_{\Omega}(u) . v=0 \forall v \in H^{1}(\Omega)$ i.e.

$$
\begin{equation*}
\int_{\Omega} \nabla u \cdot \nabla v+\sum_{j \in I_{\text {ind }}(\Omega)} D \psi_{j}\left(\int_{R_{j}} u\right) \int_{R_{j}} v=0, \quad \forall v \in H^{1}(\Omega) \tag{36}
\end{equation*}
$$

with $D \psi_{j}\left(\int_{R_{j}} u\right)=1-\frac{f_{j}}{\int_{R_{j}} u}$.
(i) By taking $v=1$, we get the egality $N=\sum_{j} \frac{f_{j}}{\int_{R_{j}} u}$. As $\frac{f_{j}}{\int_{R_{j}} u} \geq 0, \forall j \in I_{\text {ind }}(\Omega)$ and if $i_{0}=\underset{i}{\operatorname{argmin}} \int_{R_{i}} u$, we have $N \geq \frac{f_{i_{0}}}{\int_{R_{i_{0}}} u}$ which leads to $\int_{R_{i_{0}}} u \geq \frac{f_{i_{0}}}{N} \geq \frac{\min _{i} f_{i}}{N}$.
(ii) By taking $v=u$, we get the inequality $\sum_{i} \int_{R_{i}} u-f_{i} \leq 0$ which leads to $\max _{i} \int_{R_{i}} u \leq \sum_{i} \int_{R_{i}} u \leq \sum_{i} f_{i}$.

Uniqueness : From the two previous points we can consider the minimization space $H(\Omega)=\left\{u \in H^{1}(\Omega), \alpha \leq \int_{R_{j}} u \leq \beta\right\}$. Since $\psi_{j}(X)$ is strictly convex on $[\alpha, \beta]$ for all $j \in I_{\text {ind }}(\Omega)$, we get from the linearity of the integral that $u \mapsto \psi_{j}\left(\int_{R_{j}} u\right)$ is strictly convex on $H(\Omega)$. As $J_{\Omega}$ is strictly convex, we deduce that $E_{\Omega}(u)$ is strictly convex on the convex set $H(\Omega)$ and so that $E_{\Omega}$ has a unique minimum in $H^{1}(\Omega)$.

Remark 4. (i) Under the same hypotheses on f , we get the existence and uniqueness of a solution $u_{\varepsilon}$ in $H^{1}\left(\Omega_{\varepsilon}\right)$ for (19). For $\varepsilon$ small enough, we still have $\min _{i} \int_{R_{i}^{\varepsilon}} u_{\varepsilon} \geq \frac{\min _{i} f_{i}}{N}$ and $\max _{i} \int_{R_{i}^{\varepsilon}} u_{\varepsilon} \leq$ $\int_{\Omega} f=\sum_{i=1}^{N} f_{i}, \forall i \in I_{\text {ind }}(\Omega)$.
(ii) When $K \neq I$ is such that $K \mathbb{1} \neq 0$, we can show that problem (34) is well-posed in $H^{1}(\Omega)$. For more details we refer the reader to [7] chapter 3.
(iii) We can show that Proposition 3 holds as soon as $\psi_{j}$ are bounded from below for $j \in I_{\text {ind }}(\Omega)$ and strictly convex on $I \subset \mathbb{R}$. In the general case $\alpha$ and $\beta$ are implicitly defined in function of $\psi_{j}$.

### 6.2 Computation of the topological gradient for a perforated domain

In this section we compute the topological gradient for a perforated domain $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon B\right\}}$. Let $j_{0} \in I_{\text {ind }}(\Omega)$ be such that $R_{j_{0}} \supset B_{\varepsilon}\left(x_{0}\right)$ where $B_{\varepsilon}\left(x_{0}\right)$ is the ball centered at $x_{0}$ and of radius $\varepsilon$. For $j \in I_{\text {ind }}(\Omega)$, let $R_{j}^{\varepsilon}$ be the domain equal to $R_{j_{0}} \backslash \overline{B_{\varepsilon}\left(x_{0}\right)}$ if $j=j_{0}$ and $R_{j}$ otherwise. Now let us denote $I_{j}^{\varepsilon}(u)=\int_{R_{j}^{\varepsilon}} u, I_{j}(u)=\int_{R_{j}} u, J_{\varepsilon}(u)=J_{\Omega_{\varepsilon}}(u)$ and $u_{\varepsilon}$ the solution of (34) replacing $\Omega$ by $\Omega_{\varepsilon}$.

By writing $D E_{\Omega_{\varepsilon}} . v=0 \forall v \in H^{1}\left(\Omega_{\varepsilon}\right)$, we deduce the following variational formulation of (34) :

$$
\text { find } u_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right) \text { such that } F_{\varepsilon}\left(u_{\varepsilon}, v\right)=0 \quad \forall v \in H^{1}\left(\Omega_{\varepsilon}\right)
$$

where $F_{\varepsilon}(u, v)$ is the following functional on $H^{1}\left(\Omega_{\varepsilon}\right) \times H^{1}\left(\Omega_{\varepsilon}\right)$ :

$$
\begin{equation*}
F_{\varepsilon}(u, v)=\int_{\Omega_{\varepsilon}} \gamma \nabla u \cdot \nabla v+\sum_{j \in I_{\text {ind }}(\Omega)} D \psi_{j}\left(I_{j}^{\varepsilon}(K u)\right) \int_{R_{j}^{\varepsilon}} K v \tag{37}
\end{equation*}
$$

By taking $v \in \mathscr{D}\left(R_{j}^{\varepsilon}\right)$, the space of $C^{\infty}\left(R_{j}^{\varepsilon}\right)$ functions with compact support in $R_{j}^{\varepsilon}$, we obtain $-\gamma \Delta u_{\varepsilon}+$ $D \psi_{j}\left(I_{j}^{\varepsilon}\left(K u_{\varepsilon}\right)\right) K^{\star} \mathbb{1}=0$ on $R_{j}^{\varepsilon}, \forall j \in I_{\text {ind }}(\Omega)$. Then if $v$ is any test function $v \in H^{1}(\Omega)$, we get $\partial_{n} u_{\varepsilon}=0$ on $\partial \Omega_{\varepsilon}$. Therefore :

$$
\left(\mathscr{P}_{\varepsilon}\right)\left\{\begin{array}{l}
-\gamma \Delta u_{\varepsilon}+D \psi_{j}\left(I_{j}^{\varepsilon}\left(K u_{\varepsilon}\right)\right) K^{\star} \mathbb{1}=0, \text { on } R_{j}^{\varepsilon}  \tag{38}\\
\partial_{n} u_{\varepsilon}=0 \text { on } \partial \Omega_{\varepsilon} \quad \text { and } \quad\left[u_{\varepsilon}\right]_{\partial R_{j}}=0
\end{array}\right.
$$

where $\left[u_{\varepsilon}\right]_{\partial R_{j}}$ denotes the jump of $u_{\varepsilon}$ through $\partial R_{j}$.
We now give the main result of this section. The proof is performed in the case $\gamma=1$ and $K=I$ but the proof is the same in the general case.

Theorem 6.1. The topological gradient $I_{\text {Lap }}^{b}$ associated to problem (38) and to the cost function $J_{\varepsilon}(u)=\int_{\Omega_{\varepsilon}}|\nabla u|^{2}$ for a perforated domain is

$$
\begin{align*}
I_{\text {Lap }}^{b}\left(x_{0}\right) & =-2 \pi \nabla u_{0}\left(x_{0}\right) . \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma} D \psi_{j_{0}}\left(I_{j_{0}}\left(K u_{0}\right)\right)\left(K u_{0}\left(x_{0}\right)-K v_{0}\left(x_{0}\right)\right) \\
& +\frac{\pi}{\gamma} D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(K u_{0}\right)\right) K u_{0}\left(x_{0}\right)\left(I_{j_{0}}\left(K u_{0}\right)-I_{j_{0}}\left(K v_{0}\right)\right) \tag{39}
\end{align*}
$$

with $u_{0}$ and $v_{0}$ given by (38) and (44) for $\varepsilon=0$.
Démonstration. We set

$$
\widetilde{J}_{\mathcal{E}}(u)=-\sum_{j \in I_{\text {ind }}(\Omega)} \int_{R_{j}^{\varepsilon}} D \psi_{j}\left(I_{j}^{\varepsilon}(u)\right) u=-\sum_{j \in I_{\text {ind }}(\Omega)} D \psi_{j}\left(I_{j}^{\varepsilon}(u)\right) I_{j}^{\varepsilon}(u)
$$

The difference $J_{\mathcal{E}}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right)$ is :

$$
\begin{equation*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right)=\widetilde{J}_{\varepsilon}\left(u_{\varepsilon}\right)-\widetilde{J}_{\varepsilon}\left(u_{0}\right)=L_{\varepsilon}\left(u_{\varepsilon}-u_{0}\right)+\mathscr{I}_{\varepsilon} \tag{40}
\end{equation*}
$$

with

$$
\begin{align*}
L_{\varepsilon}(u) & =-\sum_{j \in I_{\text {ind }}(\Omega)} \int_{R_{j}^{\varepsilon}}\left(D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right)+D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) I_{j_{0}}\left(u_{0}\right)\right) u \\
& =-\sum_{j \in I_{\text {ind }}(\Omega)} \int_{R_{j}^{\varepsilon}} u  \tag{41}\\
\mathscr{I}_{\varepsilon} & =\int_{B_{\varepsilon}}\left(D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right)+D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) I_{j_{0}}\left(u_{0}\right)\right) u_{0} \\
& =\int_{B_{\varepsilon}} u_{0}
\end{align*}
$$

Then in order to introduce the adjoint problem, we make a second order Taylor expansion with respect to $u$ for $F_{\mathcal{E}}(u, v)$ :

$$
\begin{equation*}
F_{\varepsilon}\left(u_{\varepsilon}, v\right)=G_{\varepsilon}\left(u_{0}, v\right)+b_{\varepsilon}\left(u_{\varepsilon}-u_{0}, v\right)+c_{\varepsilon}(v)+d_{\varepsilon}(v) \tag{42}
\end{equation*}
$$

with

$$
\begin{aligned}
G_{\varepsilon}(u, v) & =\int_{\Omega_{\varepsilon}} \nabla u \cdot \nabla v+\sum_{j \in P(\Omega)} D \psi_{j}\left(I_{j}(u)\right) \int_{R_{j}^{\varepsilon}} v \\
b_{\varepsilon}(u, v) & =\int_{\Omega_{\varepsilon}} \nabla u \cdot \nabla v+\sum_{j \in I_{\text {ind }}(\Omega)} D^{2} \psi_{j}\left(I_{j}\left(u_{0}\right)\right) \int_{R_{j}^{\varepsilon}} u \int_{R_{j}^{\varepsilon}} v \\
c_{\varepsilon}(v) & =\frac{1}{2} \sum_{j \in I_{\text {ind }}(\Omega)} \int_{R_{j}^{\varepsilon}} D^{3} \psi_{j}\left(\xi_{\varepsilon}\right)\left(I_{j}^{\varepsilon}\left(u_{\varepsilon}\right)-I_{j}^{0}\left(u_{0}\right)\right)^{2} \int_{R_{j}^{\varepsilon}} v \\
d_{\varepsilon}(v) & =-D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} u_{0} \int_{R_{j_{0}}^{\varepsilon}} v
\end{aligned}
$$

where $\xi_{\varepsilon}=\theta \int_{R_{j}^{\varepsilon}} u_{\varepsilon}+(1-\theta) \int_{R_{j}} u_{0} \in[\alpha, \beta]$ with $0<\theta<1$. Next, we introduce the adjoint problem : find $v_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right)$ solution of

$$
\begin{equation*}
b_{\varepsilon}\left(u, v_{\varepsilon}\right)=-L_{\varepsilon}(u) \quad \forall u \in H^{1}\left(\Omega_{\varepsilon}\right) \tag{43}
\end{equation*}
$$

By taking $u \in \mathscr{D}\left(R_{j}^{\varepsilon}\right)$ and by integrating by parts, we deduce the following Euler equations associated with (43)

$$
\left\{\begin{array}{c}
-\Delta v_{\varepsilon}+D^{2} \psi_{j}\left(I_{j}\left(u_{0}\right)\right) \int_{R_{j}^{\varepsilon}} v_{\varepsilon}=\mathbb{1} \text { on } R_{j}^{\varepsilon} \\
\partial_{n} v_{\varepsilon}=0 \text { on } \partial \Omega_{\varepsilon} \quad \text { and } \quad\left[v_{\varepsilon}\right]_{\partial R_{j}}=0
\end{array}\right.
$$

Remark 5. For $\gamma \neq 1$ and $K \neq I$, the adjoint problem is defined by

$$
\left\{\begin{array}{l}
-\gamma \Delta v_{\varepsilon}+D^{2} \psi_{j}\left(I_{j}\left(K u_{0}\right)\right) \int_{R_{j}^{\varepsilon}} K v_{\varepsilon}=K^{\star} \mathbb{1} \text { on } R_{j}^{\varepsilon}  \tag{44}\\
\partial_{n} v_{\varepsilon}=0 \text { on } \partial \Omega_{\varepsilon} \quad \text { and } \quad\left[v_{\varepsilon}\right]_{\partial R_{j}}=0
\end{array}\right.
$$

From (40) and (42), we obtain

$$
\begin{align*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right) & =-L_{\varepsilon}\left(u_{\varepsilon}-u_{0}\right)+\mathscr{I}_{\varepsilon}=-b_{\varepsilon}\left(u_{\varepsilon}-u_{0}, v_{\varepsilon}\right)+\mathscr{I}_{\varepsilon} \\
& =G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)+c_{\varepsilon}\left(v_{\varepsilon}\right)+d_{\varepsilon}\left(v_{\varepsilon}\right)+\mathscr{I}_{\varepsilon}  \tag{45}\\
& =G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)+\mathscr{I}_{\varepsilon}+\mathscr{J}_{\varepsilon}+\mathscr{E}_{1}
\end{align*}
$$

with

$$
\begin{equation*}
\mathscr{E}_{1}=c_{\varepsilon}\left(v_{\varepsilon}\right) \quad \text { and } \quad \mathscr{J}_{\varepsilon}=d_{\varepsilon}\left(v_{\varepsilon}\right) \tag{46}
\end{equation*}
$$

By using an integration by parts, $G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ expresses as

$$
\begin{aligned}
G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right) & =-\int_{\partial B_{\varepsilon}} \partial_{n} u_{0} v_{\varepsilon}-\int_{\Omega_{\varepsilon}} \Delta u_{0} v_{\varepsilon}+\sum_{j \in I_{\text {ind }}(\Omega)} D \psi_{j}\left(I_{j}\left(u_{0}\right)\right) \int_{R_{j}^{\varepsilon}} v_{\varepsilon} \\
& =-\int_{\partial B_{\varepsilon}} \partial_{n} u_{0} v_{\varepsilon}
\end{aligned}
$$

With a similar computation to the one made in (30) for the speckle model, $G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ rewrites as :

$$
\begin{aligned}
G_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right) & =-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}}\left(\partial_{n} v_{0}+\partial_{n} l_{\varepsilon}^{w_{\varepsilon}}\right)-D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} v_{0}+\int_{B_{\varepsilon}} \widetilde{u_{0}} \Delta v_{0}-D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} l_{\varepsilon}^{w_{\varepsilon}} \\
& =\mathscr{K}_{\varepsilon}+\mathscr{L}_{\varepsilon}+\mathscr{E}_{2}+\mathscr{E}_{3}
\end{aligned}
$$

with $\tilde{u_{0}}=u_{0}-u_{0}(0), w_{\varepsilon}=v_{\varepsilon}-v_{0}$ and

$$
\begin{align*}
\mathscr{K}_{\varepsilon} & =-\int_{\partial B_{\varepsilon}} \widetilde{u_{0}}\left(\partial_{n} v_{0}+\partial_{n} l_{\varepsilon}^{w_{\varepsilon}}\right) \\
\mathscr{L}_{\varepsilon} & =-D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} v_{0}  \tag{47}\\
\mathscr{E}_{2} & =\int_{B_{\varepsilon}} \widetilde{u_{0}} \Delta v_{0} \quad, \quad \mathscr{E}_{3}=-D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} l_{\varepsilon}^{w_{\varepsilon}}
\end{align*}
$$

In the following proposition we give the asymptotic expansion of the previous quantities.

Proposition 4. Let $\mathscr{I}_{\varepsilon}, \mathscr{J}_{\varepsilon}, \mathscr{K}_{\varepsilon}, \mathscr{L}_{\varepsilon}, \mathscr{E}_{1}, \mathscr{E}_{2}$ and $\mathscr{E}_{3}$ given by (41), (46) and (47), then we have the following estimations :

$$
\begin{align*}
J_{\varepsilon}\left(u_{\varepsilon}\right)-J_{0}\left(u_{0}\right) & =\mathscr{I}_{\varepsilon}+\mathscr{J}_{\varepsilon}+\mathscr{K}_{\varepsilon}+\mathscr{L}_{\varepsilon}+\sum_{i=1}^{3} \mathscr{E}_{i} \\
\mathscr{I}_{\varepsilon} & =\pi \varepsilon^{2}\left(D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right)+D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) I_{j_{0}}\left(u_{0}\right)\right) u_{0}(0)+o\left(\varepsilon^{2}\right) \\
\mathscr{J}_{\varepsilon} & =-\pi \varepsilon^{2} D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) u_{0}(0) I_{j_{0}}\left(v_{0}\right)+o\left(\varepsilon^{2}\right)  \tag{48}\\
\mathscr{K}_{\varepsilon} & =-2 \pi \varepsilon^{2} \nabla u_{0}(0) . \nabla v_{0}(0)+o\left(\varepsilon^{2}\right) \\
\mathscr{L}_{\varepsilon} & =-\pi \varepsilon^{2} D \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) v_{0}(0)+o\left(\varepsilon^{2}\right)
\end{align*}
$$

and $\mathscr{E}_{i} \sim o\left(\varepsilon^{2}\right)$ for $\left.i \in \llbracket[1 . .3]\right]$.
Démonstration. The first equality is straightforward. A Taylor expansion of $u_{0}$ at 0 gives the first estimation. Again a Taylor expansion of $u_{0}$ at 0 , Lemma 10.6 (see Appendix B) and the fact that $R_{j_{0}}^{\varepsilon} \xrightarrow{\varepsilon \rightarrow 0} R_{j_{0}}$ give the second estimation. For $\mathscr{K}_{\varepsilon}$ we refer the reader to the proof of Proposition 2 . For $\mathscr{E}_{1}$, we use Lemma 10.5 , the regularity of $\psi(x)$ and that $\xi_{\varepsilon} \in[\alpha, \beta]$ :

$$
\begin{aligned}
\left|\mathscr{E}_{1}\right| & \leq C \sum_{j \in v}\left(\int_{R_{j}^{\varepsilon}} u_{\varepsilon}-u_{0}\right)^{2}+C\left(\int_{B_{\varepsilon}} u_{0}\right)^{2} \\
& \leq C\left\|u_{\varepsilon}-u_{0}\right\|_{0, \Omega_{\varepsilon}}^{2}+C \varepsilon^{4}=O\left(\varepsilon^{4} \log (\varepsilon)\right)
\end{aligned}
$$

By using that $\Delta v_{0}=D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) I_{j_{0}}\left(v_{0}\right)$ and a Taylor expansion of $u_{0}$ at 0 we get $\mathscr{E}_{2}=O\left(\varepsilon^{3}\right)$. For $\mathscr{E}_{3}$, from a change of variable and Lemma 10.6, we get

$$
\left|\mathscr{E}_{3}\right| \leq C \varepsilon^{2}\left\|l_{\varepsilon}^{w_{\varepsilon}}(\varepsilon X)\right\|_{0, B} \leq C \varepsilon^{3} \sqrt{-\log (\varepsilon)}
$$

(see Proposition 2).

Denoting by $j_{0}$ the integer such that $R_{j_{0}} \ni x_{0}$, we deduce the expression given in Theorem 6.1.

### 6.3 Expression of the topological gradient for a cracked domain

For a cracked domain $\Omega_{\varepsilon}=\Omega \backslash \overline{\left\{x_{0}+\varepsilon \sigma(\vec{n})\right\}}$, the computations are similar. The term $\mathscr{I}_{\varepsilon}$ of (45) is zero and the term $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)$ expresses as $F_{\varepsilon}\left(u_{0}, v_{\varepsilon}\right)=-\int_{\sigma_{\varepsilon}} \partial_{n} u_{0}\left[w_{\varepsilon}\right]$. The topological expansion of $u_{\varepsilon}$ and $v_{\varepsilon}$ are similar with the perforated domain and the computation of the topological gradient is the same as the linear case (see [4, 21] for more details).
Theorem 6.2. The topological gradient $I_{\text {Lap }}^{c}$ associated to problem (38) and to the cost function $J_{\varepsilon}(u)=\int_{\Omega_{\varepsilon}}|\nabla u|^{2}$ for a cracked domain is

$$
\begin{align*}
& I_{\text {Lap }}^{c}\left(x_{0}\right)=\min _{|n|=1} \mathscr{I}\left(x_{0}, \vec{n}\right)  \tag{49}\\
& \quad \text { with } \mathscr{I}\left(x_{0}, \vec{n}\right)=-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}
\end{align*}
$$

with $u_{0}$ and $v_{0}$ given by (38) and (44) for $\varepsilon=0$.
Remark 6. The topological gradient is the same in the general case of functions $\psi_{j} \in C^{3}(I)$, strictly convex on I and bounded from below on I. Just in the right hand-side of (44), $K^{\star} \mathbb{1}$ must be replaced by $\left(D^{2} \psi_{j}\left(I_{j}\left(K u_{0}\right)\right) I_{j}\left(K u_{0}\right)+D \psi_{j}\left(I_{j}\left(K u_{0}\right)\right)\right) K^{\star} \mathbb{1}$.

## 7 Summary table of the topological gradient expressions

We summarize in Table 1, all the expressions of the topological gradient according to the type of noise and to the infinitesimal perturbation.

|  | Ball | Crack |
| :---: | :---: | :---: |
| Gauss <br> $(K \neq I)$ | $-2 \pi \nabla u_{0}\left(x_{0}\right) \cdot \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma}\left(f\left(x_{0}\right)-K u_{0}\left(x_{0}\right)\right)\left(K v_{0}\left(x_{0}\right)-K u_{0}\left(x_{0}\right)\right)$ | $-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}$ |
| Speckle-Log <br> $(K=I)$ | $-2 \pi \nabla u_{0}\left(x_{0}\right) . \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma} D_{u} \psi\left(x_{0}, u_{0}\right)\left(u_{0}\left(x_{0}\right)-v_{0}\left(x_{0}\right)\right)$ <br> with $\psi(x, u)=u-\log (f(x))+f(x) e^{-u}$ | $-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}$ |
| Poisson <br> $(K \neq I)$ | $-2 \pi \nabla u_{0}\left(x_{0}\right) . \nabla v_{0}\left(x_{0}\right)+\frac{\pi}{\gamma} D \psi_{j_{0}}\left(I_{j_{0}}\left(K u_{0}\right)\right)\left(K u_{0}\left(x_{0}\right)-K v_{0}\left(x_{0}\right)\right)$ <br> $+\frac{\pi}{\gamma} D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(K u_{0}\right)\right) K u_{0}\left(x_{0}\right)\left(I_{j_{0}}\left(K u_{0}\right)-I_{j_{0}}\left(K v_{0}\right)\right)$ <br> with $\psi_{j_{0}}(x)=x-f_{j_{0}} \log (x)$ and $I_{j_{0}}(v)=\int_{R_{j_{0}}} v$, where $R_{j_{0}} \ni x_{0}$ | $-\pi \nabla u_{0}\left(x_{0}\right) \cdot \vec{n} \nabla v_{0}\left(x_{0}\right) \cdot \vec{n}$ |

TABLE 1 - Summary of the topological gradient expressions

## 8 Restoration using the topological gradient for a cracked domain

As a by product the computation of the topological gradient $I_{\text {Lap }}^{c}$ for a cracked domain allows to restore images degraded by blur or/and various noise statistics. Once $I_{\text {Lap }}^{c}$ is computed, we define for a fixed threshold $\delta>0$, the set $E_{\delta}=\left\{x \in \Omega ;\left|I_{\text {Lap }}^{c}(x)\right| \geq \delta\right\}$ and the approximated characteristic function

$$
\chi_{\eta}(x)=\left\{\begin{array}{l}
\eta \text { if } x \in E_{\eta} \\
1 \text { otherwise }
\end{array}\right.
$$

where $\eta$ is a small positive parameter. From the computation of $I_{\text {Lap }}^{c}$ we also get the normalized direction $\vec{\tau}=\vec{n}^{\perp}$ of the edge. If $n=(\cos (\varphi), \sin (\varphi))$ is the normal to the crack given by $I_{\text {Lap }}^{c}$, we have $\vec{\tau}=(\sin (\varphi),-\cos (\varphi))$. Then, if $f$ is the degraded observed image, we want to find a restored version $u$ of $f$ as the solution of the following anisotropic PDE :

$$
\left\{\begin{align*}
-\operatorname{div}\left(\gamma P_{\eta}^{\varphi}(x) \nabla u\right)+K^{\star} D \psi(K u) & =0 \text { on } \Omega  \tag{50}\\
\partial_{n} u & =0 \text { on } \Gamma
\end{align*}\right.
$$

with

$$
\psi(x, u)=\left\{\begin{array}{l}
\frac{1}{2}(f-u)^{2} \quad(\text { Gaussian model })  \tag{51}\\
\sum_{j \in I_{\text {Ind }}}\left(\int_{R_{j}} u-f_{j} \log \left(\int_{R_{j}} u\right)\right) \mathbb{1}_{R_{j}}(x) \quad \text { (Poisson model) } \\
\log (u)+\frac{f}{u} \quad \text { (Speckle model) }
\end{array}\right.
$$

and where $P_{\eta}^{\varphi}(x)$ is a tensor constructed from $\varphi(x)$ and $\chi_{\eta}(x)$ and $\gamma$ is a parameter to tune. More precisely, we choose $P_{\eta}^{\varphi}(x) \nabla u(x)=(\nabla u . \vec{\tau}) \vec{\tau}+\chi_{\eta}(x)(\nabla u \cdot \vec{n}) \vec{n}$. A simple identification shows that $P_{\eta}^{\varphi}(x)$ is the matrix

$$
P_{\eta}^{\varphi}(x)=\left(\begin{array}{ll}
n_{2}^{2}+\chi_{\eta}(x) n_{1}^{2} & n_{1} n_{2}\left(\chi_{\eta}(x)-1\right)  \tag{52}\\
n_{1} n_{2}\left(\chi_{\eta}(x)-1\right) & n_{1}^{2}+\chi_{\eta}(x) n_{2}^{2}
\end{array}\right)
$$

where $n_{1}=\cos (\varphi(x))$ and $n_{2}=\sin (\varphi(x))$. The interpretation of this matrix $P_{\eta}^{\varphi}(x)$ is as follows :
(i) if x belongs to the background, thanks to the definition of $\chi_{\eta}(x), P_{\eta}^{\varphi}(x)$ writes as $P_{\eta}^{\varphi}(x)=I$, so $\operatorname{div}\left(P_{\eta}^{\varphi}(x) \nabla u\right)=\Delta u$ and the smoothing is isotropic.
(ii) if x belongs to an edge (i.e. $\left.x \in E_{\delta}\right)$, then $\chi_{\eta}(x)$ is close to zero and $P_{\eta}^{\varphi}(x) \nabla u(x) \approx(\nabla u . \vec{\tau}) \vec{\tau}$ and the diffusion is in the direction of the edge. As we will see in section 9 on numerical examples, the restoration results obtained when applying equation (50) are very good.

## 9 Numerical application to 2D imaging

In this section we illustrate the theory of the topological gradient by giving various experimental results for models (13), (17) and (34).

The topological gradient expressions for the three models are stated in section 4,5 and 6 and are summarized in section 7.

For each model, to compute the topological gradient (TG) we use Algorithm 1. The computation of the direct and adjoint solutions is specific to each model. The topological gradients are summarized in Table 1.

```
Algorithm 1 Computation of the topological gradient
    Computation of \(u_{0}\) and \(v_{0}\) by using either Algorithm 2 or 3 according to the model.
    Computation of the derivatives of \(u_{0}\) by convolution with derivative filters.
    Computation of the TG relatively to the model by using Table 1 and/or Theorems of section 4, 5
    and 6.
```

Remark 7. For a cracked domain, indicators $I_{\text {Lap }}^{c}$ (16b), (33) and (49) are given by the minimal eigenvalue of a $2 \times 2$ symmetric matrix :

$$
I_{\text {Lap }}^{c}=\lambda_{\min }\left(M_{0}\right)
$$

with

$$
M_{0}=-\pi \nabla u_{0} \nabla v_{0}^{T}
$$

First, since equations (14) and (15) are linear, we develop the numerical analysis in a specific subsection. Then we perform the discretization of problems (17) and (34) and finally we give the experimental results. As the adjoint problems (26) and (44) are linear with non constant coefficients we discretize them by a finite difference scheme and we compute the discrete solution by using a sparse solver.

### 9.1 Numerical analysis for Gaussian model with blurring

To discretize (14) and (15) we use a DCT1 (discrete cosine transform of type 1) thanks to the symmetry properties and the fact that the DCT1 of a convolution product of two vectors is the product of the DCT1 of each vector. We choose this discretization because of the symmetry properties guaranteed by the algorithm. A DCT1 of $N$ points is equivalent to a DFT (discrete Fourier transform) of $2 \mathrm{~N}-2$ points. For example in 1D for $N=4$ a DCT1 of $\left[x_{0}, x_{1}, x_{2}, x_{3}\right]$ is equivalent to a DFT of $\left[x_{0}, x_{1}, x_{2}, x_{3}, x_{2}, x_{1}\right]$. We use the FFT (fast Fourier transform) to perform the DFT. The computation time of a FFT is a $\mathscr{O}(N \log (N))$. A numerical study shows that for $N \leq 10^{10}$, FFT is faster than a finite difference scheme. Algorithm 2 gives the different steps to compute the discrete solutions (14) and (15).

It consists in the following steps :

- Symmetric extension of the initial $N_{y} \times N_{x}$ image in an $2\left(N_{y}-1\right) \times 2\left(N_{x}-1\right)$ image and extension of the $2 n_{y}+1 \times 2 n_{x}+1$ kernel in a $2\left(N_{y}-1\right) \times 2\left(N_{x}-1\right)$ kernel. To fix ideas in 1D and for $n_{x}=2$, the extension of the discrete kernel $\left[x_{-2} x_{-1} x_{0} x_{1} x_{2}\right]$ is $\left[x_{0} x_{1} x_{2} 0 \ldots 0 x_{-2} x_{-1}\right]$.
- Computation of the DFT of the image and of the kernels.
- Algebraic inversion in the Fourier domain.
— Computation of the solution by inverse FFT.
An important point is the choice of the frequency domain. Indeed the natural definition of the frequency domain would be $\left\{0, \frac{2 \pi}{N_{x}}, \ldots, \frac{2 \pi\left(N_{x}-1\right)}{N_{x}}\right\} \times\left\{0, \frac{2 \pi}{N_{y}}, \ldots, \frac{2 \pi\left(N_{y}-1\right)}{N_{y}}\right\}$, but it is not a good choice. This fact is explained in [26]. Let us give the reasoning in 1D for a 1-periodic function. The trigonometric function associated with the vector of DFT coefficients is

$$
u^{N}(x)=\sum_{k=0}^{N-1} \widehat{y_{k}^{N}} e^{\frac{2 \pi i}{N}\left(k+m_{k} N\right)}
$$

where $\widehat{y^{N}} \in \mathbb{R}^{N}$ is the vector such that $\widehat{y^{N}}=\operatorname{DFT}\left(y^{N}\right), y^{N}=\left(u^{N}(l / N)\right)_{0 \leq l \leq N-1}$ and $m_{k} \in \mathbb{Z}$ are coefficients which do not change function $u$ at points $x_{l}=\frac{l}{N}$, but they greatly modify $u$ between these points (aliasing phenomenon appears). If we compute the $L^{2}(0,1)$-norm of the first derivative, we get :

$$
\begin{equation*}
\left\|u^{N}\right\|_{L^{2}(0,1)}^{2}=\left.(2 \pi)^{2} \sum_{k=0}^{N-1} \widehat{y_{k}^{N}}\right|^{2}\left(k+m_{k} N\right)^{2} \tag{53}
\end{equation*}
$$

From (53), we see that the $m_{k}$ coefficients change considerably the $L^{2}(0,1)$-norm of the first derivative and the good choice for $m_{k}$ is the value minimizing $\left(k+m_{k} N\right)^{2}$. If $0 \leq k<N / 2$ then $\left(k+m_{k} N\right)^{2}$ is minimized for $m_{k}=0$ and if $N / 2 \leq k<N$ then $\left(k+m_{k} N\right)^{2}$ is minimized for $m_{k}=-1$. A special consideration is made for $k=N / 2$ when $N$ is even because of the two possible choices $\left(m_{k}=-1\right.$ or $m_{k}=0$, see [26] for more details). By following these considerations we define the frequency domain $E=\left\{\left(\frac{\pi k_{x}}{\left(N_{x}-1\right)}, \frac{\pi k_{y}}{\left(N_{y}-1\right)}\right),\left(k_{x}, k_{y}\right) \in E_{x} \times E_{y}\right\}$, with $E_{x}=\left\{0, \ldots, N_{x}-1,-\left(N_{x}-2\right), \ldots,-1\right\}$ and
$E_{y}=\left\{0, \ldots, N_{y}-1,-\left(N_{y}-2\right), \ldots,-1\right\}$. We denote by $\Lambda=\left(\Lambda_{x}, \Lambda_{y}\right)$ the $2\left(N_{y}-1\right) \times 2\left(N_{x}-1\right)$ mesh grid associated to this discrete space. The vector of Fourier coefficients associated to a discrete signal $x \in \mathbb{R}^{N}$ is denoted by X .

## Algorithm 2 Computation of the direct and adjoint solutions

1. Given an image $f_{i j}$ defined for $(i, j) \in\left[\left[0, N_{y}-1\right)\right] \times\left[\left[0, N_{x}-1\right]\right.$, extend it to a periodic and symmetric image defined on $\left[\left[0,2\left(N_{y}-1\right)-1\right] \times\left[\left[0,2\left(N_{x}-1\right)-1\right]\right.\right.$.
2. Given a blurring convolution kernel $k_{i j}^{1}$ defined for $(i, j) \in\left[\left[0,2 n_{y}\right]\right] \times\left[\left[0,2 n_{x}\right]\right]$, use the procedure described in section 9.1 to calculate its symmetric extension $k_{m n}$ for $0 \leq m<2(N y-1)$ and $0 \leq n<2\left(N_{x}-1\right)$.
3. Use an FFT to compute $F_{k l}$ and $K_{k l}$ for $(k, l) \in\left[\left[0,2\left(N_{y}-1\right)-1\right]\right] \times\left[\left[0,2\left(N_{x}-1\right)-1\right]\right.$.
4. Given $\Lambda=\left(\Lambda_{x}, \Lambda_{y}\right)$, the meshgrid associated to the frequencies domain described in section 9.1, compute
$U_{k l}=\frac{\overline{K_{k l}} F_{k l}}{\left|K_{k l}\right|^{2}+\gamma \Lambda_{k l}^{2}}$ and $V_{k l}=\frac{\overline{K_{k l}}\left(2 K_{k l} U_{k l}-F_{k l}\right)}{\left|K_{k l}\right|^{2}+\gamma \Lambda_{k l}^{2}}$.
5. Use an inverse FFT to compute $u_{i j}$ and $v_{i j}$ for $(i, j) \in\left[\left[0,2\left(N_{y}-1\right)-1\right] \times\left[\left[0,2\left(N_{x}-1\right)-1\right]\right.\right.$.

### 9.2 Numerical analysis for Poisson and Speckle models

If $\operatorname{Inff}>0$, by Proposition 1 and Proposition 3, problems (17) and (34) are well-posed and can be discretized as :

$$
\begin{align*}
\text { (Speckle-Log model) } & \min _{x \geq \log \left(\alpha_{s}\right)} J_{s}(x), \quad \alpha_{s}>0  \tag{54a}\\
\text { (Poisson model) } & \min _{x \geq \alpha_{p}} J_{p}(x), \quad \alpha_{p}>0 \tag{54b}
\end{align*}
$$

where $\alpha_{s}=\min \left(f^{N}\right)>0$ and $\alpha_{p}=\frac{\min \left(f^{N}\right)}{N} . f^{N}$ is a discretization of $f ; J_{p}(x)$ and $J_{s}(x)$ are respectively the discrete versions of the energy functions (12) and (8). We choose a simple discretization : $u^{N}(x)$ is the step function equals to $u(j)$ on pixel $j$, and we represent $u^{N}$ by a vector of $\mathbb{R}^{N}$.

During the construction of the sequence $x^{(k)}$, the condition $x^{(k)} \geq \alpha_{p}$ for the Poisson model (respectively $x^{(k)} \geq \log \left(\alpha_{s}\right)$ for the speckle model) must be fulfilled at each step. Hence a projection ensures this condition. To solve these problems we use an iterative algorithm based on the descent method called the SGP algorithm [17] (scaled gradient projection). Let us write the discrete cost functions :

$$
\begin{aligned}
J_{p}(x) & =-\frac{\gamma}{2} x^{T} A x+\sum_{i=1}^{N}(K x)_{i}-f_{i} \log \left((K x)_{i}\right) \\
J_{s}(x) & =-\frac{\gamma}{2} x^{T} A x+\sum_{i=1}^{N}\left(u_{i}-g_{i}+e^{-\left(u_{i}-g_{i}\right)}\right)
\end{aligned}
$$

where $A$ is the Neumann Laplacian matrix, K is a discretization of the blurring operator (circulant block matrix if we assume that the image is periodic) and we recall that $g_{i}=\log \left(f_{i}\right)$. Let us give the main ideas of the SGP algorithm. The discrete energies $J_{p}$ and $J_{s}$ are denoted by $J$ as soon as we
do not use their expression and $\delta$ wiil be the number equal to $\alpha$ for the Poisson model and equal to $\log (\alpha)$ for the Speckle-Log model. We set $\Lambda=\left\{x \in \mathbb{R}^{N}, x \geq \delta\right\}$. We want to find $x^{\star} \in \Lambda$ such that $\nabla J\left(x^{\star}\right)=0$. At step k , a first order Taylor expansion at point $x=x^{(k)}$ leads to the following equation

$$
\nabla J\left(x^{(k)}\right)+\nabla^{2} J\left(x^{(k)}\right)\left(x-x^{(k)}\right)=0
$$

If $\operatorname{det}\left(\nabla^{2} J\left(x^{(k)}\right)\right) \neq 0$, we get $x=x^{(k)}-\nabla^{2} J\left(x^{(k)}\right)^{-1} \nabla J\left(x^{(k)}\right)$. We deduce by this reasoning that the direction of the descent algorithm can be given by $\nabla^{2} J\left(x^{(k)}\right)^{-1} \nabla J\left(x^{(k)}\right)$, but we see that the computation of this direction is very costly. We denote by $\mathscr{D}_{L}$ the compact set of the symmetric positive definite $N \times N$ matrices such that $\|D\| \leq L$ and $\left\|D^{-1}\right\| \leq \frac{1}{L}$. The main idea of the SGP algorithm is to construct two sequences $\alpha_{k}$ and $D_{k} \in \mathscr{D}_{L}$ such that $\alpha_{k} D_{k}$ approximates $\nabla^{2} J\left(x^{(k)}\right)$ and to project each iterate on $\Lambda$ with respect to the norm $\|x\|_{D}=\sqrt{x^{T} D x}$. We set $P_{\Lambda, D^{-1}}$ for $D \in \mathscr{D}_{L}$ the projector on $\Lambda$ with respect to the norm $\|\cdot\|_{D}$.

We recall the SGP algorithm in Algorithm 3 ( see [17]).

```
Algorithm 3 SGP algorithm
    Set \(\left.x^{(0)} \geq \alpha, \beta, \theta \in\right] 0,1\left[, 0<\alpha_{\min }<\alpha_{\max }, L>0\right.\), and fix a positive integer M .
    for \(k=0: k_{\text {max }}\) do
        Choose the parameter \(\alpha_{k} \in\left[\alpha_{\min }, \alpha_{\max }\right]\) and the scaling matrix \(D_{k} \in \mathscr{D}_{L}\)
        Projection : \(y^{(k)} \leftarrow P_{\Lambda, D_{k}^{-1}}\left(x^{(k)}-\alpha_{k} D_{k} \nabla f\left(x^{(k)}\right)\right.\)
        if \(y^{(k)}=x^{(k)}\) then
            Stop, \(x^{(k)}\) is a stationary point.
        end if
        Descent direction \(d^{(k)}=y^{(k)}-x^{(k)}\);
        \(\lambda_{k} \leftarrow 1\) and \(J_{\max } \leftarrow \max _{0 \leq j \leq \min (k, M-1)} J\left(x^{(k-j)}\right)\)
        \(\lambda_{k}\) fixed by backtracking :
        while \(f\left(x^{(k)}+\lambda_{k} d^{(k)}\right) \leq J_{\max }+\beta \lambda_{k} \nabla J\left(x^{(k)}\right)^{T} d^{(k)}\) do
            \(\lambda_{k} \leftarrow \theta \lambda_{k}\)
        end while
        \(x^{(k+1)} \leftarrow x^{(k)}+\lambda_{k} d^{(k)}\)
    end for
```

The construction of the sequences $D_{k}$ and $\alpha_{k}$ needs some explanations. We choose $D_{k}=\operatorname{diag}\left(d_{i}^{k}\right)$ with $d_{i}^{k}=\min \left(L, \max \left(\frac{1}{L}, \frac{\partial^{2} J}{\partial x_{i}^{2}}\left(x^{(k)}\right)^{-1}\right)\right)$. The approximation of the Hessian matrix $\nabla^{2} J\left(x^{(k)}\right)$ is $B\left(\alpha_{k}\right)=$ $\alpha_{k} D_{k}$. By using a first order Taylor expansion of $\nabla J(x)$ at point $x^{(k-1)}$ we get that

$$
\nabla J\left(x^{(k)}\right)-\nabla J\left(x^{(k-1)}\right)=\nabla^{2} J\left(x^{(k)}\right) \cdot\left(x^{(k)}-x^{(k-1)}\right)+o\left(\left(x^{(k)}-x^{(k-1)}\right)^{2}\right)
$$

Hence two possible choices of $\alpha_{k}$ can be made :

$$
\begin{aligned}
& \alpha_{k}^{1}=\arg \min _{\alpha}\left\|B\left(\alpha_{k}\right) s^{(k-1)}-z^{(k-1)}\right\|_{D_{k}}=\frac{s^{(k-1)^{T}} D_{k}^{-1} D_{k}^{-1} s^{(k-1)}}{s^{(k-1)^{T}} D_{k}^{-1} z^{(k-1)}} \\
& \alpha_{k}^{2}=\arg \min _{\alpha}\left\|s^{(k-1)}-B\left(\alpha_{k}\right)^{-1} z^{(k-1)}\right\|_{D_{k}}=\frac{s^{(k-1)^{T}} D_{k} z^{(k-1)}}{z^{(k-1)^{T}} D_{k} D_{k} z^{(k-1)}}
\end{aligned}
$$

where $s^{(k-1)}=x^{(k)}-x^{(k-1)}$ and $z^{(k-1)}=\nabla J\left(x^{(k)}\right)-\nabla J\left(x^{(k-1)}\right)$. In [17] the choice of $\alpha_{k}$ is the output of an algorithm called SGP-SS Algorithm (SGP step length selection) which uses two thresholds $0<\alpha_{\text {min }}<\alpha_{\text {max }}$. Let us give the derivative of the discrete cost functions $J_{p}$ and $J_{s}$ :

$$
\begin{aligned}
\nabla J_{p} & =-\gamma A x-K^{T} \frac{f}{K x}+K^{T} \mathbb{1} \\
\nabla^{2} J_{p} & =-\gamma A+K^{T} \operatorname{diag}\left(\frac{f}{(K x)^{2}}\right) K \\
\nabla J_{s} & =-\gamma A x+1-e^{-(x-f)} \\
\nabla^{2} J_{s} & =-\gamma A+\operatorname{diag}\left(e^{-(x-f)}\right)
\end{aligned}
$$

where $\mathbb{1} \in \mathbb{R}^{N}$ denotes the vector with each coefficient equal to $1, \operatorname{diag}(x)$ for $x \in \mathbb{R}^{N}$ is the diagonal matrix with diagonal entries equal to $x$. For $x \in \mathbb{R}^{N}$ and $\varphi: \mathbb{R} \longrightarrow \mathbb{R}$ a function, $\varphi(x)$ stands for the vector $\left(\varphi\left(x_{i}\right)\right)_{1 \leq i \leq N}$. The choice of the parameters in Algorithm 3 is the following : $\beta=10^{-4}$, $\theta=0.4, k_{\max }=600, M=1$ and for the Poisson model (34) we set $\alpha_{\min }=10^{-10}, \alpha_{\max }=10^{5}$ while for the Speckle-Log model (17) we set $\alpha_{\min }=10^{-5}, \alpha_{\max }=10^{15}$. The initial value of $x^{(0)}$ is either the observed image for the Poisson model or its logarithm for the Speckle-Log model. Let us note that in the case of the speckle model, (50) is performed with $\psi(u)=\log (u)+\frac{f}{u}$.

### 9.3 Comparison of our method with some classical models

As said in the introduction other variational methods exist for segmenting/restorating images.
We will compare the topological gradient segmentation process with the one performed by the Mumford-Shah model. We will also compare the restoration proposed in (50) with the ones given by the Mumford-Shah restoration and by the TV restoration.

## Mumford-Shah model of segmentation/restoration and its approximation

Let $u$ the image of support $\Omega$, the functional introduced by Mumford and Shah in 1989 (see [33]) is :

$$
F(u, \gamma)=\int_{\Omega}\left|u-u_{0}\right|^{2}+\lambda \int_{\Omega \backslash \bar{\gamma}}|\nabla u|^{2}+\alpha \mathscr{H}^{1}(\gamma)
$$

where $f$ is the observed image, $u$ is a function defined on $\Omega \backslash \bar{\gamma}$ (the restored version of $f$ ) and $\gamma \subset \Omega$ is the set of discontinuity of $u . \mathscr{H}^{1}$ is the Hausdorff measure of $\gamma, \lambda$ and $\alpha$ are positive parameters. The difficulty was that the unknown are not of same nature : $u$ is a function and $\gamma$ is a set. Ambrosio and Tortorelli [1] proposed an approximation of this functional as follows :

$$
\mathscr{F}_{\varepsilon}(u, b)=\int_{\Omega}\left[|u-f|^{2}+\lambda b^{2}|\nabla u|^{2}+\alpha\left(\varepsilon|\nabla b|^{2}+\frac{(b-1)^{2}}{4 \varepsilon}\right)\right]
$$

We will change the data fidelity term $|u-f|^{2}$ according to the a priori model (Gaussian, Poissonian and speckle model) i.e. the model that we will compare with (50) is (see [35]) :

$$
\begin{equation*}
\min _{u \in H^{1}(\Omega), b \in H^{1}(\Omega)} \int_{\Omega}\left[\psi(x, K u)+\lambda b^{2}|\nabla u|^{2}+\alpha\left(\varepsilon|\nabla b|^{2}+\frac{(b-1)^{2}}{4 \varepsilon}\right)\right] \tag{55}
\end{equation*}
$$

where $K$ is the blur operator, $u(x)$ is the restored image, $1-b(x) \approx 0$ is the characteristic function of the edges and $\psi(x, u)$ is given in (51). We will call this model the Mumford-Shah model (MS).

## TV model of restoration

The TV model is well-known : we search for a restored version $u$ minimizing an energy functional which is the sum of the total variation $\|D u\|$ and a data fidelity term which depends on the a priori model (see [8] for the speckle model and [16] for the Poisson one). Thus, we will compare our model (50) to :

$$
\begin{equation*}
\min _{u \in \mathscr{B} \mathscr{V}(\Omega)} \int_{\Omega}|D u|+\lambda \psi(x, K u) \tag{56}
\end{equation*}
$$

where $\lambda$ is a parameter, $K$ the blur operator, and $\psi(u)$ is given in (51). In the sequel we call this model the TV model.

For more details on these models we refer the reader to [16], [8], [35] and [33].
For restoration comparisons on synthetic images we use two indicators :

- the PSNR which is defined for a noisy observation $I$ of an image $I_{0}$ by

$$
\operatorname{PSNR}(I)=10 \log \left(\frac{255^{2}}{\left\|I-I_{0}\right\|_{2}^{2} / N}\right)
$$

where $N$ is the number of pixels in the image.
— the SSIM defined as

$$
\operatorname{SSIM}(I)=\frac{\left(2 \mu_{I} \mu_{I_{0}}+c_{1}\right)\left(2 \operatorname{cov}\left(I, I_{0}\right)+c_{2}\right)}{\left(\mu_{I}^{2}+\mu_{I_{0}}^{2}+c_{2}\right)}
$$

where $\mu_{x}$ stands for the mean of $x, \operatorname{cov}$ is the covariance operator and $c_{1}$ and $c_{2}$ are constants given for RGB images by $c_{1}=\left(255 k_{1}\right)^{2}$ and $c_{2}=\left(255 k_{2}\right)^{2}$ with $k_{1}=0.01$ and $k_{2}=0.03$.

### 9.4 Numerical results for the Gaussian model

An interesting study is the comparison of formula (16a) and (16b) giving the topological gradient for respectively a perforated and a cracked domain. A priori $I_{L a p}^{b}$ (16a), associated to a perforated domain, would be adapted for the detection of isotropic structures while $I_{L a p}^{c}(16 \mathrm{~b})$ would seem more adapted to detect edges straight. Fig. 1 compares, for different values of $\gamma, I_{\text {Lap }}^{c}$ and $I_{\text {Lap }}^{b}$ to $b(x)$ obtained by minimizing MS (55). For small $\gamma, I_{\text {Lap }}^{c}$ seems the most robust indicator to detect the black spots contours of the cheetah. We remark by increasing $\gamma$ that $I_{\text {Lap }}^{b}$ becomes smooth and singular on the entire black spots and not only on its contours, while $I_{\text {Lap }}^{c}$ better detects the border of the cheetah.

We deduce that $\gamma$ must be tuned with respect to the noise but also by taking into account the size of structures to detect.


FIGURE 1 - Comparison of the two formula given in (16b) and (16b) for different value of $\gamma$ with the MS model $(55)(b \approx 0)$ for a Gaussian noisy image ( $\mathrm{PSNR}=16 \mathrm{~dB}$ ) containing mainly isotropic small structures.

Fig. 2 shows $I_{\text {Lap }}^{c}$ and the MS result for a Gaussian noisy and Gaussian blurred image. Results are similar but $I_{\text {Lap }}^{c}$ has the advantage of being very fast (for this image the computation time is about one second on a computer equipped with a processor Intel Core 1.9 GHz , see section 9.7 for more details).

(a) Initial image


Figure 2 - Comparison of the topological gradient (16b) with the MS segmentation (55) on a Gaussian noisy and blurred image ( $\mathrm{PSNR}=16 \mathrm{~dB}, \sigma=3$ ).

Fig. 3 shows the restored version performed by (50), (55) and (56) on a Gaussian noisy and

Gaussian blurred image. We see that the restoration given by MS (55) degrades contours and does not remove completely the blur. Restorations computed by (50) and (56) are quite similar but the computation time is shorter for (50).


Figure 3 - Comparison of restored versions (50), (55) and (56) for a synthetic Gaussian noisy and Gaussian blurred ( $\sigma=3$ ) image.

Fig. 4 displays the 1 D profiles of the image to recover, its degraded versions (blurred, blurred+noisy), the restored version (50) and $I_{\text {Lap }}^{c}$ (16b) across an edge. We see that the restored version matches very well the image to recover and $I_{L a p}^{c}$ detects quite well the edge.


Figure 4 - (a) A transverse cut displaying the image to recover, the Gaussian blurred version ( $\sigma=$ 3 ), the blurred and noisy version (PSNR=16dB), (b) the restored version (50) $(\gamma=5)$ and (c) the Topological gradient (16b) $(\gamma=1)$.

### 9.5 Numerical results for the speckle-log model. Comparisons

In this section we still illustrate the segmentation given by $I_{\text {Lap }}^{c}$ (33), $I_{\text {Lap }}^{b}(21)$ and $b(x)$ computed with the MS model (55). We also display the restoration performed by (50), the MS model (55) and the TV model (56).

On Fig. 5, we compare $I_{\text {Lap }}^{c}$ (33) and $I_{\text {Lap }}^{b}$ (21) for different values of $\gamma$ with the function $b(x)$ given by the MS model for a synthetic speckled image. Similarly to the Gaussian case, $I_{\text {Lap }}^{b}$ seems more adapted to detect isotropic structures and we still deduce that $\gamma$ must be tuned with respect to
the noise and to the size of structures to detect. Comparing $I_{\text {Lap }}^{c}$ and MS, similar results are obtained for the cheetah.


Figure 5 - Comparison of the two formula (21) and (33) for different values of $\gamma$ with the MS segmentation (55) for an initial speckled image $(L=6)$ containing mainly isotropic small structures.

The result given in Fig. 6 for a real SAR image is similar to the one of Fig. 5. Here we see that $I_{\text {Lap }}^{b}$ and $I_{\text {Lap }}^{c}$ can be used for different objectives : particularly on small isotropic structures we see that $I_{\text {Lap }}^{b}$ detects the entire object while $I_{\text {Lap }}^{c}$ detects its edges.


Figure 6 - Comparison of the two formula (21) and (33) with the MS segmentation (55) for a real SAR image.

Fig. 7 and Fig. 9 compare the restoration performed by (50), (55) and (56) respectively on a
real SAR image and on a synthetic speckled image. Restoration given by the MS degrades contours while (50) and the TV model (56) are nearly equivalent. However, on Fig. 9, we can notice that the restoration performed by (50) is better than the TV one and computation times are equivalent (see section 9.7).


Figure 7 - Comparison of restored versions (50), (55) and (56) for a real SAR image (Zoom on Fig. 6-(a)).

On Fig. 8 we compare $I_{\text {Lap }}^{c}$ and $b(x)$ computed with MS (55) for a very noisy synthetic image ( $L=1$ i.e. the worst case for this model). Here $I_{\text {Lap }}^{c}$ gives a quite good result with respect to MS where edges are spread out.


Figure 8 - Comparison of the topological gradient (33) with the MS segmentation (55) for a synthetic speckled image ( $L=1$ ).


Figure 9 - Comparison of restored versions (50), (55) and (56) for a speckled synthetic image ( $L=$ 1).

Finally Fig. 10 shows the 1D profiles of the image to recover, its noisy version, the restored version (50) and $I_{\text {Lap }}^{c}$ (33) across an edge. The restored version matches very well the image to recover and edges are not degraded. This shows that (50) is a good restoration process.


Figure 10 - (a) A transverse cut displaying the image to recover and the speckled image, (b) the restored version (50) $(\gamma=0.002)$ and the image to recover, (c) $I_{\text {Lap }}^{c}(33)(\gamma=1.8)$.

### 9.6 Numerical results for the Poisson model. Comparisons

In this section we compare the segmentation performed by $I_{\text {Lap }}^{c}(49), I_{\text {Lap }}^{b}$ (39) and the MS model (55). We also compare the restoration computed with (50), the MS and the TV model.

Fig. 11 and Fig. 12 show respectively the segmentation results in the case of a synthetic Poissonian image and of a real confocal image of a rat's neuron. $I_{\text {Lap }}^{c}$ (49) detects edges quite well compared to the MS model. We see that $I_{\text {Lap }}^{b}$ fills small structures (the size of these structures is related to $\gamma$ as for the Gaussian and Poissonian cases).


Figure 11 - Comparison of the two formula (39) and (49) for different values of $\gamma$ with the MS segmentation (55) for a Poissonian image containing mainly isotropic small structures.
(b) $I_{\text {Lap }}^{b}(39)\left(\gamma=10^{-1}\right)$

(e) $I_{L a p}^{c}(49)(\gamma=3)$

(a) Initial image

(c) $I_{L a p}^{c}(49)\left(\gamma=10^{-1}\right)$
(d) $I_{\text {Lap }}^{b}(39)(\gamma=3)$

$$
\left.\varepsilon=10^{-6}, \alpha=1\right)
$$

Figure 12 - Comparison of the two formula (39) and (49) for different values of $\gamma$ with the MS segmentation (55) for a Poissonian real image containing rat's neurons.

Fig. 13 and Fig. 15 display the restoration computed by (50), (55) and (56) on respectively a real confocal image and a synthetic Poissonian image blurred by Gaussian convolution. We notice that (50) and (56) restore very well the image preserving edges unlike to the MS model (55) which degrades contours and which does not annihilate the blur effect.


Figure 13 - Comparison of the restored versions (50), (55) and (56) for a real Poissonian image of rat's neurons.

Fig. 14 compares $I_{\text {Lap }}^{c}$ (49) with $b(x)$ performed by the MS model (55) for a Poissonian image blurred by a Gaussian convolution.

(a) Initial image

(b) $I_{\text {Lap }}^{c}(49)(\gamma=0.001)$

> (c) MS $(55)(\lambda=0.07$ $\left.\varepsilon=10^{-6}, \alpha=1\right)$
(d) Zoom on $I_{\text {Lap }}^{c}$ (49)
(e) MS (55) : Zoom

Figure 14 - Comparison of $I_{\text {Lap }}^{c}$ (49) with the MS segmentation (55) on a synthetic Poissonian image.


Figure 15 - Comparison of the restored versions for a Poissonian image blurred by Gaussian convolution ( $\sigma=3$ ).

Finally, Fig. 16 shows the 1D profiles of the image to recover, its degraded versions (blurred, blurred + Poissonian process), the restored version (50) and $I_{\text {Lap }}^{c}(49)$ across an edge. We see that (50) allows to recover the initial image and that $I_{\text {Lap }}^{c}$ detects very well the edge.


Figure 16 - (a) A transverse cut displaying the image to recover, the Gaussian blurred version ( $\sigma=$ 3), the blurred and noisy version (PSNR=16dB), (b) the restored version (50) ( $\gamma=0.005$ ) and (c) $I_{\text {Lap }}^{c}$ (16b) $(\gamma=0.001)$.

### 9.7 Computational time comparisons for the three methods

On Fig. 17 and Fig. 18 are compared the three methods for $K=I$ (no blur) : Mumford-Shah (MS), TV and the topological gradient method (TG). The experiments are performed on a computer equipped with a processor Intel Core 1.9 GHz and all algorithms are implemented in Matlab. In the Gaussian case, since all the equations are linear, the topological gradient method is the fastest method both for segmentation and restoration. To fix ideas, for an image containing $6.4 \times 10^{5}$ pixels, the computational time for the segmentation performed by the topological gradient is less than one second and the restoration is about four seconds.


Figure 17 - Comparisons of the computational time of the three methods in function of the number of pixels in the Gaussian case

For the nonlinear cases (Poissonian and speckle models), the computation of the topological gradient still remains the fastest but the restoration step performed by (50) takes approximately the same time as for a restoration given by a TV model. Let us notice that the TV model is implemented by using an iterative algorithm (explicit schema with fixed step length with at maximum 1000 iterations) and that the Mumford-Shah model solution is computed by minimizing the approximate functional alternatively with respect to $u$ and $b$. When this latter model is associated to non quadratic data fidelity terms (Poisson and speckle), the minimization with respect to $u$ (at $b$ fixed) is made by using Algorithm 3 and we perform 15 iterations. Finally let us precise that we present on Fig. 18 the computation time only for the Poissonian model because similar results are obtained from the speckle one.


Figure 18 - Comparisons of the computational time of the three methods of restoration in function of the number of pixels in the Poissonian case

## 10 Appendices

In these appendices we give the asymptotic expansion of the differences $u_{\varepsilon}-u_{0}$ and $v_{\varepsilon}-v_{0}$ for the non linear problems (Poisson and Speckle-log models). Some proofs are similar to the linear case and so we will refer the reader to [4]. To establish these asymptotic expansions we need the following exterior problem

$$
\left(\mathscr{P}_{\text {ext }}\right)\left\{\begin{align*}
\Delta H & =0, \text { on } \mathbb{R}^{2} \backslash \bar{B}  \tag{57}\\
\partial_{n} H & =g, \text { on } \partial B \\
H & \longrightarrow 0, \text { at } \infty
\end{align*}\right.
$$

where $g \in H^{-1 / 2}(\partial B)$ and $\int_{\partial B} g d \sigma=0$. For the computation of the topological gradient we will use the two following lemma. We omit the proofs and we refer the reader to [21] for more details

Lemma 10.1. The solution of (57) expresses as a simple layer potential :

$$
H(x)=\int_{\partial B} \lambda(y) E(x-y) d \sigma
$$

with $E(x)=-\frac{1}{2 \pi} \log (|x|)$ is the fundamental solution of the Laplace operator and $\lambda(y)=-2 g(y)$. Denoting by $l^{H}$ the solution of (29), we have the jump relations through $\partial B$

$$
\begin{aligned}
H-l^{H} & =0 \\
\partial_{n} H-\partial_{n} l^{H} & =-\lambda
\end{aligned}
$$

and $l^{u}$ expresses also as $l^{H}(x)=\int_{\partial B} \lambda(y) E(x-y) d \sigma$.

The following asymptotic estimations holds.
Lemma 10.2. Let $H$ the solution of (57), then :

$$
\begin{array}{rlrl}
|H(x)| & \leq \frac{C}{|x|}, & & |\nabla H(x)| \leq \frac{C}{|x|^{2}} \\
\left\|H\left(\frac{x}{\varepsilon}\right)\right\|_{0, \Omega_{\varepsilon}} & =O(\sqrt{-\log (\varepsilon)}), & \left\|\nabla H\left(\frac{x}{\varepsilon}\right)\right\|_{0, \Omega_{\varepsilon}}=O(\varepsilon)
\end{array}
$$

### 10.1 Appendix A

In this appendix all the proofs are performed assuming $\gamma=1$, and when (19) is referenced we suppose that $\gamma=1$. Moreover we suppose that $\psi$ fulfills Hypotheses 1 .
Lemma 10.3. Let $X_{\varepsilon}=u_{\varepsilon}-u_{0}$ where $u_{\varepsilon}$ and $u_{0}$ are respectively given by (19) for $\varepsilon>0$ and $\varepsilon=0$, and let $P$ be the solution of (57) with $g=-\nabla u_{0}(0) . n$, then we have the following asymptotic expansion :

$$
X_{\varepsilon}=\varepsilon P\left(\frac{x}{\varepsilon}\right)+e_{\varepsilon}
$$

with $\left\|e_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$. Moreover we have the following estimation

$$
\left\|X_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)
$$

Démonstration. First by substracting equations (19) for $\varepsilon>0$ and $\varepsilon=0$ :

$$
\left\{\begin{align*}
-\Delta X_{\varepsilon}+D_{u} \psi\left(x, u_{\varepsilon}\right)-D_{u} \psi\left(x, u_{0}\right) & =0, \text { on } \Omega_{\varepsilon}  \tag{58}\\
\partial_{n} X_{\varepsilon} & =-\partial_{n} u_{0}, \text { on } \partial B_{\varepsilon} \\
\partial_{n} X_{\varepsilon} & =0, \text { on } \Gamma
\end{align*}\right.
$$

Then an integration by parts gives :

$$
\begin{equation*}
\int_{\Omega_{\varepsilon}} \nabla X_{\varepsilon} \cdot \nabla v+\int_{\Omega_{\varepsilon}}\left(D_{u} \psi\left(x, u_{\varepsilon}\right)-D_{u} \psi\left(x, u_{0}\right)\right) v=-\int_{\partial B_{\varepsilon}} \nabla u_{0}(0) . \vec{n} v \quad \forall v \in H^{1}\left(\Omega_{\varepsilon}\right) \tag{59}
\end{equation*}
$$

With a similar manner we integrate by parts the Euler equation checked by $\varepsilon P\left(\frac{x}{\varepsilon}\right)$ on $\Omega_{\varepsilon}$ :

$$
\begin{equation*}
\int_{\Omega_{\varepsilon}} \nabla P\left(\frac{x}{\varepsilon}\right) \nabla v=-\int_{\partial B_{\varepsilon}} \nabla u_{0}(0) \cdot \vec{n}-\int_{\Gamma} \partial_{n} P\left(\frac{x}{\varepsilon}\right) v \tag{60}
\end{equation*}
$$

By setting $e_{\varepsilon}=X_{\varepsilon}-\varepsilon P\left(\frac{x}{\varepsilon}\right)$ and substracting (59) to (60), we get

$$
\int_{\Omega_{\varepsilon}} \nabla e_{\varepsilon} \cdot \nabla v+\int_{\Omega_{\varepsilon}}\left(D_{u} \psi\left(x, u_{\varepsilon}\right)-D_{u} \psi\left(x, u_{0}\right)\right) v=-\int_{\Gamma} \partial_{n} P\left(\frac{x}{\varepsilon}\right) v-\int_{\partial B_{\varepsilon}}\left(\partial_{n} u_{0}-\nabla u_{0}(0) . \vec{n}\right) v
$$

Then, thanks to a Taylor expansion, we rewrite the second term on the right hand-side of the above equality : $D_{u} \psi\left(x, u_{\varepsilon}\right)-D_{u} \psi\left(x, u_{0}\right)=D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right)\left(u_{\varepsilon}-u_{0}\right)$ with $u_{\delta_{\varepsilon}}=\theta u_{0}+(1-\theta) u_{\varepsilon}, \theta: \Omega \longrightarrow \mathbb{R}$, $0 \leq \theta \leq 1$. We can bound from below this term by using Lemma 5.1 and Hypotheses 1. Indeed, since $a \leq u_{\delta_{\varepsilon}} \leq b$ and $\psi(u)$ is strictly convex on $[a, b] \subset I$, we get that there exists $\delta>0$ not depending on $\varepsilon$ such that $D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right) \geq \delta>0$. Thus $e_{\varepsilon}$ is solution of the following well-posed variational problem : find $e_{\varepsilon} \in H^{1}\left(\Omega_{\varepsilon}\right)$ such that

$$
\begin{aligned}
\int_{\Omega_{\varepsilon}} \nabla e_{\varepsilon} . \nabla v+\int_{\Omega_{\varepsilon}} D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right) e_{\varepsilon} v= & -\int_{\Gamma} \partial_{n} P\left(\frac{x}{\varepsilon}\right) v-\int_{\partial B_{\varepsilon}}\left(\partial_{n} u_{0}-\nabla u_{0}(0) . \vec{n}\right) v \\
& -\int_{\Omega_{\varepsilon}} D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right) \varepsilon P\left(\frac{x}{\varepsilon}\right) v
\end{aligned}
$$

for all $v \in H^{1}\left(\Omega_{\varepsilon}\right)$. Now we split $e_{\varepsilon}$ in $e_{\varepsilon}=e_{\varepsilon}^{1}+e_{\varepsilon}^{2}$ with
(i) $e_{\varepsilon}^{1} \in H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}$ solution of

$$
\int_{\Omega_{\varepsilon}} \nabla e_{\varepsilon}^{1} \cdot \nabla v=-\int_{\partial B_{\varepsilon}}\left(\partial_{n} u_{0}-\nabla u_{0}(0) \cdot \vec{n}\right) v \quad \forall v \in H^{1}\left(\Omega_{\varepsilon}\right)
$$

(ii) $e_{\varepsilon}^{2} \in H^{1}\left(\Omega_{\varepsilon}\right)$ solution of

$$
\int_{\Omega_{\varepsilon}} \nabla e_{\varepsilon}^{2} \cdot \nabla v+\int_{\Omega_{\varepsilon}} D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right) e_{\varepsilon}^{2} v=-\int_{\Gamma} \partial_{n} P\left(\frac{x}{\varepsilon}\right) v-\int_{\Omega_{\varepsilon}} D_{u}^{2} \psi\left(x, u_{\delta_{\varepsilon}}\right)\left(\varepsilon P\left(\frac{x}{\varepsilon}\right)+e_{\varepsilon}^{1}\right)
$$

Then, by using a change of variable (CV), a trace theorem on $B_{2} \backslash \bar{B}$ (where $B_{2}$ is the ball of radius 2, centered at 0 ), and the equivalency of the $H^{1}\left(B_{2} \backslash \bar{B}\right)$-norm with the semi-norm and a CV again, we get that $\left\|e_{\varepsilon}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}=O\left(\varepsilon^{2}\right)$. Next, by using Lemma 10.2, a trace Theorem on $\Omega \backslash \bar{B}$ and the fact that $e_{\varepsilon}^{1} \in H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}$, we get that $\left\|e_{\varepsilon}^{2}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$. The $H^{1}\left(\Omega_{\varepsilon}\right)$-norm estimation of $e_{\varepsilon}$ is then straightforward by using these two estimations and the following inequality $\left\|e_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}} \leq$ $\left\|e_{\varepsilon}^{1}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}+\left\|e_{\varepsilon}^{2}\right\|_{1, \Omega_{\varepsilon}} ;$ the estimation of $\left\|X_{\mathcal{E}}\right\|_{0, \Omega_{\varepsilon}}$ comes from $\left\|X_{\mathcal{\varepsilon}}\right\|_{0, \Omega_{\varepsilon}} \leq\left\|\varepsilon P\left(\frac{x}{\varepsilon}\right)\right\|_{0, \Omega_{\varepsilon}}+\left\|e_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}$ and the previous inequalities. This ends the proof. For more details we refer the reader to [21].

Lemma 10.4. Let $w_{\varepsilon}=v_{\varepsilon}-v_{0}$ where $v_{\varepsilon}$ and $v_{0}$ are respectively given by (25) for $\varepsilon>0$ and $\varepsilon=0$, and let $Q$ be the the solution of (57) with $g=-\nabla v_{0}(0) . n$, then we have the following asymptotic expansion :

$$
w_{\varepsilon}=\varepsilon Q\left(\frac{x}{\varepsilon}\right)+r_{\varepsilon}
$$

with $\left\|r_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$. Moreover we have

$$
\left\|w_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right) \quad, \quad\left|w_{\varepsilon}\right|_{1, \Omega_{\varepsilon}}=O(\varepsilon)
$$

Démonstration. The problem is linear. From Lemma 5.1 and Hypotheses 1, we get that there exist $\delta_{1}>0$ and $\delta_{2}>0$ such that $\delta_{2} \geq D_{u}^{2} \psi\left(u_{0}\right) \geq \delta_{1}$. The well-posedness of the problem is then straightforward. Then the proof can be easily deduced from the proof of Theorem 10.3 or from the linear case [4].

### 10.2 Appendix B

In this appendix we consider problem (38) with $K=I$ and $\gamma=1$. The general case (when $K$ is a convolution operator such that $K \mathbb{1} \neq 0$ ) can be easily deduced (see [7] chapter 3 for the wellposedness).

Lemma 10.5. Let $X_{\varepsilon}=u_{\varepsilon}-u_{0}$ where $u_{\varepsilon}$ and $u_{0}$ are respectively given by (38) for $\varepsilon>0$ and $\varepsilon=0$, then we have :

$$
X_{\varepsilon}=\varepsilon P\left(\frac{x}{\varepsilon}\right)+e_{\varepsilon}
$$

where $P$ is defined by (57) with $g=-\nabla u_{0}(0) . \vec{n}$ and where $\left\|e_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2}\right)$. Moreover we have the estimation :

$$
\left\|X_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)
$$

Démonstration. First, let us write the Euler equations checked by $X_{\mathcal{\varepsilon}}$. By substracting equations (38) for $\varepsilon>0$ and for $\varepsilon=0$, we get for $j \in I_{\text {ind }}(\Omega)$ :

$$
\left(\mathscr{X}_{\varepsilon}\right)\left\{\begin{align*}
&-\Delta X_{\varepsilon}+D \psi_{j}\left(\int_{R_{j}^{\varepsilon}} u_{\varepsilon}\right)-D \psi_{j}\left(\int_{R_{j}} u_{0}\right)=0  \tag{61}\\
& \text { on } R_{j}^{\varepsilon}, j \in I_{i n d}(\Omega) \\
& \partial_{n} X_{\varepsilon}=-\partial_{n} u_{0}, \text { on } \partial B_{\varepsilon} \\
& \partial_{n} X_{\varepsilon}=0, \text { on } \Gamma
\end{align*}\right.
$$

Then, by a Taylor expansion there exists $\xi_{\varepsilon}=\theta \int_{R_{j}^{\varepsilon}} u_{\varepsilon}+(1-\theta) \int_{R_{j}} u_{0}$ with $0<\theta<1$ such that

$$
D \psi_{j}\left(\int_{R_{j}^{\varepsilon}} u_{\varepsilon}\right)-D \psi_{j}\left(\int_{R_{j}} u_{0}\right)=D^{2} \psi_{j}\left(\xi_{\varepsilon}\right)\left(\int_{R_{j}^{\varepsilon}} u_{\varepsilon}-\int_{R_{j}} u_{0}\right)
$$

From Proposition 3, it is straightforward that $0<\alpha \leq \xi_{\varepsilon} \leq \beta$ where $\alpha=\frac{\min _{i} f_{i}}{N}$ and $\beta=\sum_{i} f_{i} .\left(\mathscr{X}_{\varepsilon}\right)$ rewrites for $j \in I_{i n d}(\Omega)$ as

$$
\left(\mathscr{X}_{\varepsilon}\right)\left\{\begin{aligned}
-\Delta X_{\varepsilon}+D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}} X_{\varepsilon} & =\int_{R_{j} \backslash \overline{R_{j}^{\varepsilon}}} u_{0} \text { on } R_{j}^{\varepsilon} \\
\partial_{n} X_{\varepsilon} & =-\partial_{n} u_{0}, \text { on } \partial B_{\varepsilon} \\
\partial_{n} X_{\varepsilon} & =0, \text { on } \Gamma
\end{aligned}\right.
$$

with $\int_{R_{j} \backslash \overline{R_{j}^{\varepsilon}}} u_{0}=\delta_{j_{0}}(j) \int_{B_{\varepsilon}} u_{0}$, where $\delta$ is the Dirac function. Let $e_{\varepsilon}=X_{\varepsilon}-\varepsilon P\left(\frac{x}{\varepsilon}\right)$ where P is defined by (57) with $g=-\nabla u_{0}(0) . \vec{n} . e_{\varepsilon}$, then $e_{\varepsilon}$ verifies the following equation

$$
\left(\mathscr{E}_{\varepsilon}\right)\left\{\begin{array}{c}
-\Delta e_{\varepsilon}+D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}} e_{\varepsilon}=-\varepsilon D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}} P\left(\frac{x}{\varepsilon}\right)  \tag{62}\\
\quad \text { on } R_{j}^{\varepsilon} j \in I_{\text {ind }}(\Omega) \\
\partial_{n} e_{\varepsilon}=-\left(\partial_{n} u_{0}-\nabla u_{0}(0) . \vec{n}\right)=\varphi_{\varepsilon}(x)=O(|x|) \text { on } \partial B_{\varepsilon} \\
\partial_{n} e_{\varepsilon}=-\partial_{n} P\left(\frac{x}{\varepsilon}\right)=\phi_{\varepsilon}(x)=O\left(\frac{\varepsilon^{2}}{|x|^{2}}\right) \text { on } \Gamma
\end{array}\right.
$$

We set :
$K_{\varepsilon}^{j}=\left\{\begin{aligned}-\varepsilon D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}} P\left(\frac{x}{\varepsilon}\right) & =O\left(\varepsilon^{3}\right), \text { for } j \neq j_{0} \\ \int_{B_{\varepsilon}} u_{0}-\varepsilon D^{2} \psi_{j_{0}}\left(\xi_{\varepsilon}\right) \int_{R_{j_{0}}^{\varepsilon}} P\left(\frac{x}{\varepsilon}\right) & =O\left(\varepsilon^{2}\right), \text { for } j=j_{0}\end{aligned}\right.$
Now we split $e_{\varepsilon}$ in the sum $e_{\varepsilon}=e_{\varepsilon}^{1}+e_{\varepsilon}^{2}+e_{\varepsilon}^{3}$ with
(i) $e_{\varepsilon}^{1} \in H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}$ solution of

$$
\left\{\begin{aligned}
-\Delta e_{\varepsilon}^{1} & =0, \text { on } \Omega_{\varepsilon} \\
\partial_{n} e_{\varepsilon}^{1} & =\varphi_{\varepsilon}(x), \text { on } \partial B_{\varepsilon} \\
\partial_{n} e_{\varepsilon}^{1} & =0, \text { on } \Gamma
\end{aligned}\right.
$$

(ii) $e_{\varepsilon}^{2} \in H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}$ solution of

$$
\left\{\begin{aligned}
\Delta e_{\varepsilon}^{2} & =0, \text { on } \Omega_{\varepsilon} \\
\partial_{n} e_{\varepsilon}^{2} & =0, \text { on } \partial B_{\varepsilon} \\
\partial_{n} e_{\varepsilon}^{2} & =\phi_{\varepsilon}(x), \text { on } \Gamma
\end{aligned}\right.
$$

(iii) $e_{\varepsilon}^{3} \in H^{1}\left(\Omega_{\varepsilon}\right)$ solution of

$$
\left(\mathscr{E}_{\varepsilon}^{3}\right)\left\{\begin{array}{c}
-\Delta e_{\varepsilon}^{3}+D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}} e_{\varepsilon}^{3}=K_{\varepsilon}^{j}-D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}}\left(e_{\varepsilon}^{2}+e_{\varepsilon}^{2}\right) \\
\text { on } R_{j}^{\varepsilon}, j \in I_{\text {ind }}(\Omega) \\
\partial_{n} e_{\varepsilon}^{3}=0, \text { on } \partial B_{\varepsilon} \\
\partial_{n} e_{\varepsilon}^{3}=0, \text { on } \Gamma
\end{array}\right.
$$

Standard computations (see [4, 21] for more details) lead to the following estimations :

$$
\left\|e_{\varepsilon}^{1}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}} \leq C \varepsilon^{2} \quad\left\|e_{\varepsilon}^{2}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}} \leq C \varepsilon^{2}
$$

To estimate $e_{\varepsilon}^{3}$, we take the variational formulation of $\left(\mathscr{E}_{\varepsilon}^{3}\right)$ :

$$
\begin{equation*}
\int_{\Omega_{\varepsilon}} \nabla e_{\varepsilon}^{3} \cdot \nabla v+\sum_{j \in I_{\text {ind }}(\Omega)} D^{2} \psi_{j}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}} e_{\varepsilon}^{3} \int_{R_{j}^{\varepsilon}} v=\sum_{j \in I_{\text {ind }}(\Omega)} K_{\varepsilon}^{j} \int_{R_{j}^{\varepsilon}} v-D^{2} \psi_{j}^{\varepsilon}\left(\xi_{\varepsilon}\right) \int_{R_{j}^{\varepsilon}}\left(e_{\varepsilon}^{1}+e_{\varepsilon}^{2}\right) \int_{R_{j}^{\varepsilon}} v \tag{63}
\end{equation*}
$$

An easy computation of $D^{2} \psi_{j}$ and Proposition 3 give for $\varepsilon \leq \varepsilon_{0}$

$$
\frac{\max _{j \in I_{\text {ind }}(\Omega)} f_{j}}{\alpha^{2}} \geq D^{2} \psi_{j}\left(\xi_{\varepsilon}\right)=\frac{f_{j}}{\xi_{\varepsilon}^{2}} \geq \frac{\min _{j \in I_{\text {ind }}(\Omega)} f_{j}}{\beta^{2}}>0
$$

where $N=\left|I_{\text {ind }}(\Omega)\right|$. By taking as test function $v=e_{\varepsilon}^{3}$ in (63), we deduce the following estimations :

$$
\begin{aligned}
& \int_{\Omega_{\varepsilon}}\left|\nabla e_{\varepsilon}^{3}\right|^{2}+C \sum_{j \in I_{\text {ind }}(\Omega)}\left(\int_{R_{j}^{\varepsilon}} e_{\varepsilon}^{3}\right)^{2} \\
& \leq \sum_{j \in I_{\text {ind }}(\Omega)}\left|K_{\varepsilon}^{j}\right|\left|\int_{R_{\varepsilon}^{j}} e_{\varepsilon}^{3}\right|+\left|D^{2} \psi_{j}^{\varepsilon}\left(\xi_{\varepsilon}\right)\right| \int_{R_{j}^{\varepsilon}}\left(\left|e_{\varepsilon}^{1}\right|+\left|e_{\varepsilon}^{2}\right|\right)\left|\int_{R_{j}^{\varepsilon}} e_{\varepsilon}^{3}\right| \\
& \leq C\left(\varepsilon^{2}+\left\|e_{\varepsilon}^{1}\right\|_{L^{2}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}+\left\|e_{\varepsilon}^{2}\right\|_{L^{2}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}\right) \sum_{j \in I_{\text {ind }}(\Omega)}| | \int_{R_{\varepsilon}^{j}} e_{\varepsilon}^{3} \mid \\
& \leq C \varepsilon^{2} \sum_{j \in I_{\text {ind }}(\Omega)}\left|\int_{R_{\varepsilon}^{j}} e_{\varepsilon}^{3}\right|
\end{aligned}
$$

Then, thanks to the following inequality which stands for any sequence of real numbers $\left(a_{i}\right)_{i}$

$$
\left(\sum_{i \in I_{\text {ind }}(\Omega)}\left|a_{j}\right|\right)^{2} \leq\left|I_{\text {ind }}(\Omega)\right| \sum_{i \in I_{\text {ind }}(\Omega)}\left|a_{i}\right|^{2}
$$

and the positiveness of $\int_{\Omega_{\varepsilon}}\left|\nabla e_{\varepsilon}^{3}\right|^{2}$, we obtain

$$
\sum_{j \in I_{\text {ind }}(\Omega)}\left|\int_{R_{\varepsilon}^{j}} e_{\varepsilon}^{3}\right| \leq C \varepsilon^{2}
$$

and then $\left|e_{\varepsilon}^{3}\right|_{1, \Omega_{\varepsilon}} \leq C \varepsilon^{2}$. By using the Poincaré-Wirtinger inequality we get :

$$
\begin{aligned}
& \left\|e_{\varepsilon}^{3}\right\|_{1, \Omega_{\varepsilon}} \leq\left\|e_{\varepsilon}^{3}-\frac{1}{|\Omega|} \int_{\Omega_{\varepsilon}} e_{\varepsilon}^{3}\right\|_{1, \Omega_{\varepsilon}}+\left|\int_{\Omega_{\varepsilon}} e_{\varepsilon}^{3}\right| \\
& \leq C\left|e_{\varepsilon}^{3}\right|_{1, \Omega_{\varepsilon}}+C \sum_{j \in I_{\text {ind }}(\Omega)}\left|\int_{R_{\varepsilon}^{j}} e_{\varepsilon}^{3}\right| \leq C \varepsilon^{2}
\end{aligned}
$$

From the inequality $\left\|e_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}} \leq\left\|e_{\varepsilon}^{1}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}+\left\|e_{\varepsilon}^{2}\right\|_{H^{1}\left(\Omega_{\varepsilon}\right) / \mathbb{R}}+\left\|e_{\varepsilon}^{3}\right\|_{1, \Omega_{\varepsilon}}$, we get the result. For the $L^{2}\left(\Omega_{\varepsilon}\right)$-norm estimation of $X_{\varepsilon}$, it suffices to take the $L^{2}\left(\Omega_{\varepsilon}\right)$-norm of its asymptotic expansion and to use the first point of Lemma 10.5 and Lemma 10.2.

Lemma 10.6. Let $w_{\varepsilon}=v_{\varepsilon}-v_{0}$ where $v_{\varepsilon}$ and $v_{0}$ are given by (43) for $\varepsilon>0$ and $\varepsilon=0$, then we have :

$$
w_{\varepsilon}=\varepsilon Q\left(\frac{x}{\varepsilon}\right)+r_{\varepsilon}
$$

where $Q$ is defined by (57) with $g=-\nabla v_{0}(0)$.n, and where $\left\|r_{\varepsilon}\right\|_{1, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}\right)$. Moreover we have :

$$
\left\|w_{\varepsilon}\right\|_{0, \Omega_{\varepsilon}}=O\left(\varepsilon^{2} \sqrt{-\log (\varepsilon)}, \quad\left|w_{\mathcal{\varepsilon}}\right|_{1, \Omega_{\varepsilon}}=O(\varepsilon)\right.
$$

Démonstration. By substracting equations (44) for $\varepsilon>0$ and for $\varepsilon=0$, the Euler equations associated to $w_{\varepsilon}$ are :

$$
\left(\mathscr{W}_{\varepsilon}\right)\left\{\begin{align*}
-\Delta w_{\varepsilon}+D^{2} \psi_{j}\left(I_{j}\left(u_{0}\right)\right) \int_{R_{j}^{\varepsilon}} w_{\varepsilon} & =0 \text { on } R_{j} \text { with } j \neq j_{0}  \tag{64}\\
-\Delta w_{\varepsilon}+D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{R_{j_{0}}^{\varepsilon}} w_{\varepsilon} & \\
=D^{2} \psi_{j_{0}}\left(I_{j_{0}}\left(u_{0}\right)\right) \int_{B_{\varepsilon}} v_{0} & =O\left(\varepsilon^{2}\right) \text { on } R_{j_{0}}^{\varepsilon} \\
\partial_{n} w_{\varepsilon} & =-\partial_{n} v_{0} \text { on } \partial B_{\varepsilon} \\
\partial_{n} w_{\varepsilon} & =0 \text { on } \Gamma
\end{align*}\right.
$$

This problem is linear and from (3) we have:

$$
\frac{\max _{j \in I_{\text {ind }}(\Omega)} f_{j}}{\alpha^{2}} \geq D^{2} \psi_{j}\left(I_{j}\left(u_{0}\right)\right)=\frac{f_{j}}{I_{j}\left(u_{0}\right)^{2}} \geq \frac{\min _{j \in I_{\text {ind }}(\Omega)} f_{j}}{\beta^{2}}
$$

Then the topological expansion of $w_{\varepsilon}$ can be deduced from the proof of Lemma 10.5 or from the linear case [4].
The two last estimations are straightforward by using the topological expansion of $w_{\varepsilon}$ and Lemma 10.2.

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