



HAL
open science

Dynamo regimes and transitions in the VKS experiment

Michaël Berhanu, Gautier Verhille, Jean Boisson, Basile Gallet, Christophe Gissinger, Stéphane Fauve, Nicolas Mordant, François Pétrélis, Mickaël Bourgoïn, Philippe Odier, et al.

► **To cite this version:**

Michaël Berhanu, Gautier Verhille, Jean Boisson, Basile Gallet, Christophe Gissinger, et al.. Dynamo regimes and transitions in the VKS experiment. *The European Physical Journal B: Condensed Matter and Complex Systems*, 2010, 77, pp.459. 10.1140/epjb/e2010-00272-5 . hal-00941490

HAL Id: hal-00941490

<https://hal.science/hal-00941490>

Submitted on 13 Apr 2015

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Dynamo regimes and transitions in the VKS experiment

M. Berhanu¹, G. Verhille², J. Boisson⁴, B. Gallet¹, C. Gissinger¹, S. Fauve¹, N. Mordant¹, F. Pétrélis¹, M. Bourgoin³, Ph. Odier², J.-F. Pinton², N. Plihon², S. Aumaitre⁴, A. Chiffaudel⁴, F. Daviaud⁴, B. Dubrulle⁴, and C. Pirat⁴

¹ Laboratoire de Physique Statistique, CNRS & École Normale Supérieure,
24 rue Lhomond, F-75005 Paris, France

² Laboratoire de Physique de l'École Normale Supérieure de Lyon, CNRS & Université de Lyon,
46 allée d'Italie, F-69364 Lyon cedex 07, France

³ Laboratoire des Écoulements Géophysiques et Industriels, CNRS & Université Joseph Fourier,
BP53, F-38041, Grenoble cedex 9, France

⁴ Service de Physique de l'État Condensé, CNRS & CEA Saclay, F-91191 Gif-sur-Yvette, France

April 13, 2015

Abstract. The Von Kármán Sodium experiment yields a variety of dynamo regimes, when asymmetry is imparted to the flow by rotating impellers at different speed F_1 and F_2 . We show that as the intensity of forcing, measured as $F_1 + F_2$, is increased, the transition to a self-sustained magnetic field is always observed via a supercritical bifurcation to a stationary state. For some values of the asymmetry parameter $\theta = (F_1 - F_2)/(F_1 + F_2)$, time dependent dynamo regimes develop. They are observed either when the forcing is increased for a given value of asymmetry, or when the amount of asymmetry is varied at sufficiently high forcing. Two qualitatively different transitions between oscillatory and stationary regimes are reported, involving or not a strong divergence of the period of oscillations. These transitions can be interpreted using a low dimensional model based on the interactions of two dynamo modes.

PACS. 47.35.Tv, 47.65.-d Magnetohydrodynamics in fluids – 05.45.-a Nonlinear dynamical systems – 47.20.Ky Bifurcation, flow instabilities

1 Introduction

While still being quite far from the parameter regime that characterizes natural objects, dynamo experiments using liquid metals have the advantage of having adjustable control parameters. They also display dynamical evolutions that can be recorded over long durations compared to the Joule characteristic time. The Riga [1] and Karlsruhe [2] experiments have established the central role of helicity and shear in the dynamo process, with dynamo characteristics well predicted by laminar models – although the underlying flows are turbulent, with a moderate turbulence rate. In the recent VKS experiment (see [3] and references therein), the situation is different since non axisymmetric velocity components (in the bulk and near the impellers) play a leading role in the magnetic field generation: the observed axisymmetric dynamo mean field cannot be generated by the axisymmetric mean flow alone. Another central observation in the VKS experiment lies in the variety of dynamo regimes observed when the flow driving parameters and magnetic Reynolds number are varied [3]. We report here on the bifurcations observed in the VKS dynamo based on a thorough study of the parameter space. The

von Kármán flow of sodium is generated inside a cylinder by the counter rotation of coaxial impellers at frequencies (F_1, F_2) (see Fig. 1a). When both impellers rotate at the same frequency $F_1 = F_2$, the driving and hence the mean flow structure are symmetric with respect to any rotation \mathcal{R}_π of π around any radial axis in its equatorial plane. When the frequencies F_1 and F_2 are different, this symmetry is broken. One possible variable to quantify this asymmetry is the parameter $\theta = (F_1 - F_2)/(F_1 + F_2)$. In addition, the choice of frequencies (F_1, F_2) imposes a mean shear $F = (F_1 + F_2)/2$. When the parameters (F, θ) are varied, various types of dynamos are observed. We concentrate on the following issues:

- (i) characteristics of the bifurcation to a dynamo regime when F increases, for a fixed value of the asymmetry parameter θ . Our observation is that a (statistically) steady dynamo is always generated first, via a supercritical bifurcation. Time-dependent regimes can develop as a secondary bifurcation when F is increased further for particular intervals of θ .
- (ii) Transition between dynamo regimes (above critical),

in particular changes from stationary to oscillatory dynamics when θ is varied at constant F .

In some cases, we observe a divergence of the period of oscillation during transition or bifurcations. In other cases the period remains finite.

The next section describes the experimental set-up and gives a summary of the dynamo capacity of the various flows configurations studied so far in the VKS experiment. The parameter space and the bifurcations observed when increasing the forcing at a given value of asymmetry are presented in section 3. Section 4 describes how the dynamo undergoes transitions between various regimes as the asymmetry is varied at a given forcing. In section 5, we show that these observations can be understood using the predictions of a low dimensional model, involving the interactions of two dynamo modes. A final discussion is given in section 6.

2 Experimental set-up and configurations

2.1 Set-up

The present set-up is displayed in Fig.1. A von Kármán swirling flow is generated in a cylindrical vessel (radius $R_{\text{vessel}}=289$ mm, length $L=604$ mm) by two counter-rotating impellers 371 mm apart. The flow is surrounded by sodium at rest enclosed between the copper outer vessel and an inner copper cylinder (radius $R_{\text{cyl}} = 206$ mm, length $H = 524$ mm). The impellers are made of soft iron disks (radius $R_{\text{imp}} = 154.5$ mm) fitted with 8 curved blades with height $h = 41.2$ mm. Their rotation rate can be adjusted independently to (F_1, F_2) . The arrows in Fig. 1 define the positive rotation rate $F_1, F_2 > 0$. It corresponds to the case where the curved blades are counter-rotating and “unscoping” the fluid (pushing the fluid with the convex side of the blades). The fluid is liquid sodium (density $\rho = 930$ kg \cdot m $^{-3}$, electrical conductivity $\sigma = 9.6 \cdot 10^6$ $\Omega \cdot$ m $^{-1}$, kinematic viscosity $\nu = 6.7 \cdot 10^{-7}$ m $^2 \cdot$ s $^{-1}$, at 120°C). The driving motor power is 300 kW and cooling by an oil circulation inside the wall of the outer copper vessel allows experimental operation at constant temperature in the range 110 – 140°C. This set-up is a slightly modified version of the one previously described in [3]: the inner copper ring that was located in the mid-plane has been removed. A hydrodynamic study of this configuration has been done in [4] and the influence of the inner ring on the flow has been studied in details in [5].

The magnetic field is measured with Hall probes inserted inside the fluid, as shown in Fig.1. Unless otherwise stated, the measurements presented in this paper were made using the probe at location 1. Because of the small number of probes that were available for the experimental data studied here, no statement will be made on the spatial distribution of the observed dynamo modes. They will only be distinguished by their amplitude and time dynamics at the point of measurement.

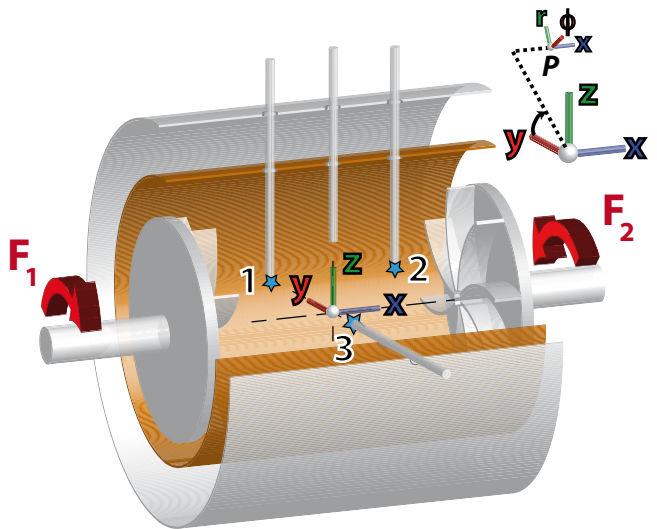


Fig. 1. Experimental setup, showing the location of the Hall probes. x is the axial coordinate directed from impeller 1 to impeller 2, r and ϕ the radial and azimuthal coordinates.

In the following subsection, we give a brief summary of the main results obtained with the different configurations studied so far in VKS: the inner copper wall and ring can be inserted or removed and the impellers material can be varied independently. We define the kinetic Reynolds number of the flow: $Re = 2\pi R_{\text{imp}} F R / \nu$ with $R = R_{\text{cyl}}$ if the inner cylinder is present and $R = R_{\text{vessel}}$ otherwise. The corresponding magnetic Reynolds number is defined in the same way: $R_m = 2\pi R_{\text{imp}} F R \mu_0 \sigma$, where μ_0 is the magnetic permeability of vacuum. This definition is different from the one used in previous publications (it leads to 25% higher values for R_m) but it was chosen here because it contains explicitly the flow scale R and thus allows quantitative comparisons between cases with and without the inner cylinder. Since the asymmetry between the rotation rates of the impellers is a key ingredient in the results presented here, we also define individual magnetic Reynolds numbers based on the velocity of each impeller: $R_{m1,2} = 2\pi R_{\text{imp}} F_{1,2} R \mu_0 \sigma$. Finally, note that the conductivity of sodium is quite sensitive to temperature variations in the vicinity of its melting point ($\pm 2\%$ on R_m for a temperature variation of $\pm 6^\circ\text{C}$ around 125°C). These variations are taken into account in the computation of R_m .

2.2 Configurations and dynamo capacity

The first observation of the dynamo effect in our experiment was made with inner copper wall, inner ring and iron impellers in place. By removing the inner ring, dynamo action is observed for $F_1, F_2 > 0$, but the main change is that no dynamo regime is observed when both impellers still counter-rotate, but in the opposite (“scooping”) direction ($F_1 = F_2 < 0$) up to the maximum operational power. The disappearance of the dynamo regime with “scoop-

impeller 1/impeller 2	inner cylinder+ring ($R_m = 2.42F$ @ 120°C)	inner cylinder alone ($R_m = 2.42F$)	no inner cylinder, no ring ($R_m = 3.40F$)
SS/SS	no dyn. @ $F < 29$ Hz	no dyn. @ $F < 28$ Hz	no dyn. @ $F < 24$ Hz
Iron/SS			A: $F^c = 15$ Hz B: no dyn. @ $F > -22$ Hz C: $F_1^c = 17$ Hz D: no dyn. @ $F_2 < 25$ Hz
Iron/Iron	A: $F^c = 17$ Hz B: $F^c = -18$ Hz C/D: $F_{1,2}^c = 16$ Hz	A: $F^c = 17$ Hz B: no dyn. @ $F > -25Hz$ C/D: $F_{1,2}^c = 16$ Hz	A: $F^c = 12$ Hz B: no data C/D: no data

Table 1. Summary of the dynamo/non dynamo regimes observed in the different VKS configurations. Label “A” corresponds to the “unscooping” exact counter-rotating case ($F_1 = F_2 > 0$), label “B” to the “scooping” exact counter-rotating case ($F_1 = F_2 < 0$), labels “C” and “D” to the cases with a single impeller rotating, respectively $F_1 > 0, F_2 = 0$ and $F_2 > 0, F_1 = 0$, these two cases being different when two impellers of different materials are used (middle line in the table). Light grey cells are cases where no data is available yet. The values indicated are either the observed critical impellers rotation rate when dynamo is present, or the maximal rotation rate achieved, due to power limitations, without observing dynamo. The correspondance between rotation rate and R_m , as defined in this paper, is indicated for each configuration. The dark grey cell corresponds to the configuration presented in this paper.

ing” blades is also observed in kinematic simulations using measured mean velocity fields in an equivalent water experiment [6], although the context is different since the dynamo modes in this case are non-axisymmetric. Note that changing the direction of rotation of the impellers changes the ratio of poloidal over toroidal components of the velocity field because of the curvature of the blades. In the case where the median copper ring is present, dynamo regimes are observed for both directions of rotation of the impellers, although the thresholds differ by about 6%. These observations indicate that some aspects of the flow structure (poloidal/toroidal ratio, position of recirculation loops, stability of the shear layer, level of fluctuations, ...) do play a role in the onset of the instability.

For the configuration with the inner wall removed, the dynamo threshold is reached for a rotation rate 1.4 times lower for exact counter-rotation, compared to the case with inner wall, corresponding to the same critical R_m , as defined in section 2.1.

Other experimental configurations have been studied, in which the iron impellers are replaced by stainless steel (SS) ones (with and without inner walls and ring). In these cases, no dynamo is observed at the highest R_m achievable with our experiment. A hybrid configuration has also been tested in which one of the impellers is made of soft iron and the other of stainless steel (no inner walls or ring). In the exact counter-rotating case, dynamo generation is observed above $F^c = 15$ Hz. When the iron impeller only is rotating and the stainless steel impeller is kept at rest, dynamo is observed above a different critical rate $F^c = 17$ Hz and no dynamical regimes are evidenced in this case. When only the stainless steel impeller is rotating, no dynamo could be observed up to a maximum rotation rate of 25 Hz. In addition, no dynamo is observed in exact corotation.

Table 1 presents a summary of these results, including the dynamo thresholds based on the rotation frequency, obtained in the various configurations which produced a dynamo. The study presented in this paper corresponds to

one specific configuration, where the parameter space has been explored in details. It includes iron impellers and the inner copper cylinder, without the copper ring in the mid-plane. Nevertheless, similar parameter spaces have been observed in all VKS flows driven by 2 iron impellers. For instance, when the inner cylinder is removed, the parameter space still has non-dynamo to dynamo transitions with supercritical bifurcations to statistically stationary magnetic fields and regimes with oscillations and reversal are observed, with the restriction that only one window of time dependent behavior is evidenced.

3 Parameter space and bifurcations

The flow and dynamo configurations spanned in the results reported here are summarized in Fig.2. It shows the (R_{m1}, R_{m2}) parameter space (only in the case $F_1, F_2 > 0$), where each R_m is defined based on the rotation rate of one impeller. Since we observe almost no asymmetry in this parameter space under the exchange $F_1 \leftrightarrow F_2$, the actual experimental points are represented only in the upper left half of the plane (see Fig. 6b for an estimate of the experimental imperfections). Color codes are used, corresponding to the dynamo regime observed at each location, broadly characterized as non dynamo, (statistically) stationary magnetic field and time-dependent regimes. We also indicate the paths that have been followed in the experimental measurements at F or θ constant (note that $\theta=\text{constant}$ also corresponds to $F_1/F_2 = \text{constant}$) along which dynamo transitions are discussed in more details below. In the lower right part of the graph in Fig. 2, a schematic view of the regimes is sketched.

We start by exploring the bifurcations that develop at a constant asymmetry parameter θ as the magnetic Reynolds number is increased. At all investigated values of θ , we observe that increasing R_m , the first instability is a supercritical bifurcation to a stationary dynamo, with time-dependent regimes possibly developing at larger R_m

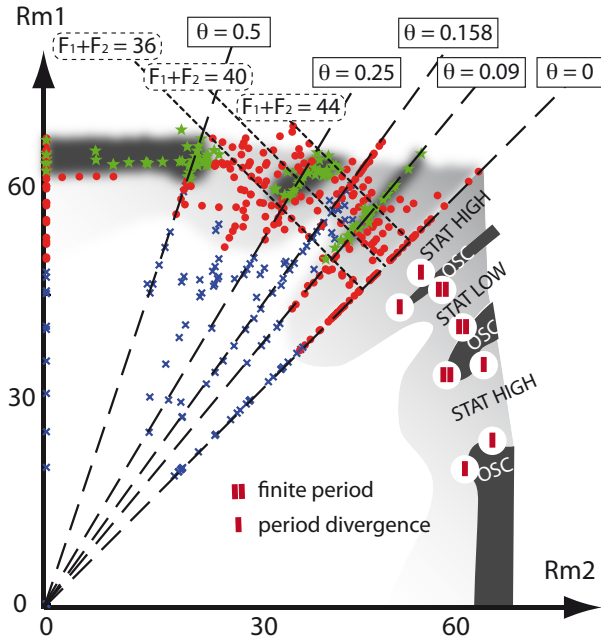


Fig. 2. (R_{m1}, R_{m2}) parameter space explored, varying independently the rotation rate of each impeller (top left, where blue crosses indicate no dynamo regimes, red circles stationary dynamo regimes and green stars time dependent regimes), and schematic of accessible dynamical regimes (bottom right) – boundaries between different regimes are voluntarily smoothed due to lack of experimental resolution. The type of transition is also indicated, as well as some paths along F or θ constant, followed during the experimental investigations.

values. We note that the possibility of a direct bifurcation to time-dependent dynamo cannot be excluded, for instance near $|\theta| \sim 0.12$ where non dynamo and oscillating regimes are observed in Fig. 2 in very close proximity. We first discuss the case where time dependent regimes do not appear, the dynamo field remaining statistically stationary at the highest R_m values achieved in the experiment (due to power/torque limitations). Figure 3 shows examples of this type. The amplitude and standard deviation of the magnetic field are plotted as a function of $R_m = (R_{m1} + R_{m2})/2$. In the case of exact counter-rotation ($\theta = 0$, top figure), a clear transition to dynamo can be observed at $R_m^c \sim 39$ after an initial slowly growing phase ($R_m < 38$), which we interpret as induction from the ambient magnetic field. Above the threshold, the amplitude of the magnetic field is observed to grow linearly with $R_m - R_{m_c}$ up to $R_m \sim 55$, after which the trend is less clear. This is different from the case when the inner ring is present [3], where a best fit of dynamo field growth lead to a power law increase $(R_m - R_{m_c})^\alpha$, with $\alpha \sim 0.77$. Note that, due to the imperfection of the bifurcation curve already discussed in presence of the inner ring, the determination of the exponent depends strongly on the chosen value for the threshold R_m^c , leading to a larger uncertainty. Simple arguments from bifurcation theory [7] would lead one to expect an $\alpha = 1/2$ value. For such a bifurcation

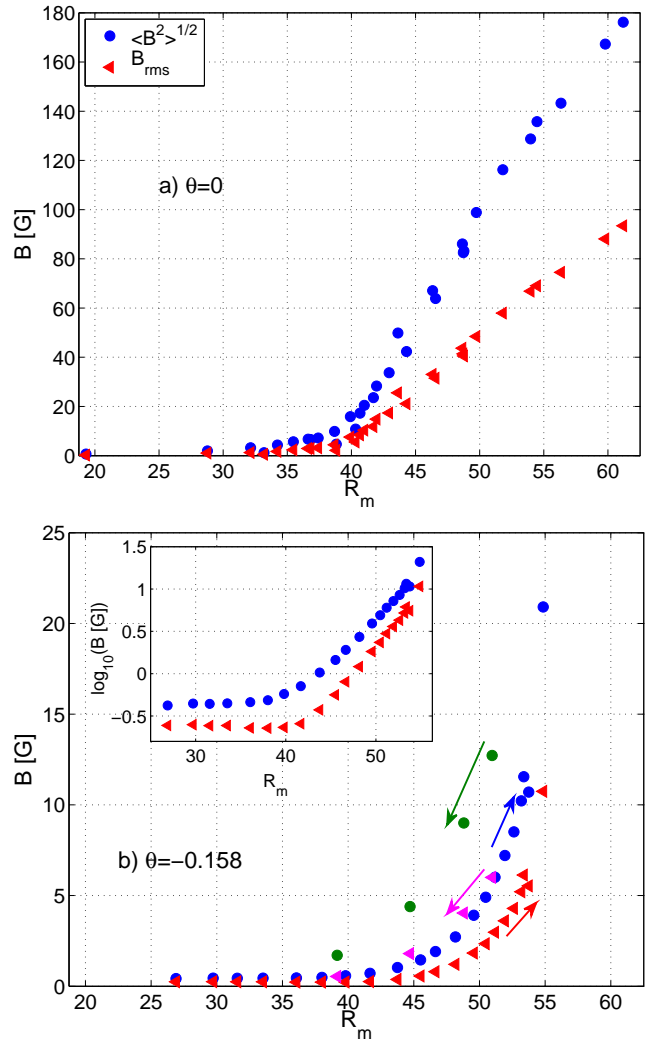


Fig. 3. Bifurcation to stationary dynamos, magnetic field amplitude and standard deviation. a) Exact counter-rotation, $\theta = 0$, STAT HIGH regime; b) $\theta = -0.158$, STAT LOW regime. In this case, as indicated by the arrows, data points measured with increasing R_m are shown in blue (amplitude) and red (standard deviation), while data points measured with decreasing R_m are shown in green (amplitude) and magenta (standard deviation). The inset shows the same plot with vertical log scale, in the case of increasing R_m .

from a turbulent state, the existence and universality of critical exponents is still an open problem.

At $\theta = -0.158$ (bottom figure), we also observe a supercritical bifurcation, but the amplitude of the dynamo field is much lower, by almost an order of magnitude. We thus call this regime “STAT LOW” as opposed to the “STAT HIGH” regime observed in exact counter-rotation. The change in behavior from an induction regime to a low field dynamo regime is best seen in the lin/log plot in the inset, showing a clear break in the slopes at about $R_m \sim 41 - 42$. It is a low amplitude dynamo, and the effect of the slight magnetization of the iron disks cannot be

ignored; hence the magnetic field is different if the measurement is made after a dynamo run (green and magenta symbols in 3b – measured when decreasing R_m) or following a non dynamo configuration (blue and red symbols – measured when increasing R_m). Note that induction effects from the ambient field are also present, which can add up to the starting dynamo. Because in this case the dynamo is a low amplitude one, both effects can be of the same order of magnitude, so that the behavior near threshold is non trivial. The standard deviations follow the same evolution, with an amplitude about half of that of the mean fields.

Fig. 4a,d,g shows bifurcation curves for other values of θ , for which we observe the development of time dependent regimes as R_m is increased. A stationary dynamo is first generated at $R_{mc1} \sim 25 - 30$, then oscillations develop as $R_m > R_{mc2} \sim 30 - 40$ (examples of the time signals shown in Fig. 4c,f,i). For the larger values of θ , the dynamical change is also associated with a discontinuity in the amplitude of the magnetic field. It indicates that this second bifurcation is not supercritical, although the observation of the discontinuity may depend on the choice of the order parameters. When the regime becomes oscillatory, the mean value vanishes for each component, as can be seen by the equality between the field amplitude and its standard deviation. Another interesting observation, evidenced in Fig. 4b,e,h, is that at low ($\theta = -0.09$) and high ($\theta = -0.5$) asymmetry, the period diverges at the secondary bifurcation threshold. At the intermediate value $\theta = -0.25$ the oscillations develop with a finite period.

4 Transitions between regimes

The observation of a variety of dynamo regimes when the asymmetry is varied can also be explored in the parameter space along lines of constant F , at varying θ values. This corresponds to a fairly constant global magnetic Reynolds number R_m , save for small changes due to variations in the temperature of the sodium. We first consider the behavior for $F_1 + F_2 = 36$ Hz ($R_m = 34.8 \pm 0.6$) as $|\theta|$ increases – shown in Fig.5a. Close to exact counter-rotation ($\theta \approx 0$), the dynamo is a STAT HIGH type. Then, the region $\theta \in [-0.25, -0.1]$ corresponds to the STAT LOW regime (identified as a STAT LOW dynamo when the field amplitude is larger than 3 G), before reaching again a STAT HIGH regime, as θ increases. At higher values ($\theta < -0.45$) the stationary dynamo loses its stability for a regime with random reversals before turning into an oscillating one. As shown in the inset, as the oscillating/reversing regime approaches the stationary regime (decreasing $|\theta|$), the period at the transition diverges. As R_m is increased, the main change is the occurrence of reversal and oscillating regimes on each side of the STAT LOW region. This is confirmed in Fig. 5b and c, where the transition plots are shown respectively in the case $F_1 + F_2 = 40$ Hz ($R_m = 38.6 \pm 0.5$) and $F_1 + F_2 = 44$ Hz ($R_m = 41.8 \pm 0.7$). Note that the oscillating region at high θ observed in Fig. 5a could also be present at higher forcing, but it was not possible to

reach regimes at $\theta < -0.4$ for $F \geq 40$ Hz, due to torque limitations on the motors driving the impellers. It is interesting to observe that these transition plots display a similar shape. Actually, a qualitative collapse of the 3 plots shown in Fig. 5 can be obtained (see Fig. 6a) if the variable $\Delta R_m = R_{m2} - R_{m1}$ is used for the abscissa and the amplitude of the magnetic field in the ordinate is normalized by $(1/\sigma R_{cyl})\sqrt{\rho/\mu_0}(R_m - R_m^c)$ (with $R_m^c = 39$), as suggested from the linear growth observed in Fig. 3a. Similarly to the θ variable, ΔR_m is also representative of the amount of asymmetry in the flow and seems to provide a better collapse of these data.

As was observed earlier, windows of oscillating dynamo can be present between two stationary modes. Fig. 6b displays the periods measured in all the time dependent cases of the parameter space, as a function of $|\Delta R_m|$. Using this variable, the evolution of the oscillation periods collapses on a single plot. Note that the collapse is slightly better when only the points related to a given sign of ΔR_m are plotted. This traces back to experimental imperfections that slightly break the symmetry of the parameter space when F_1 and F_2 are exchanged. In the cases of transition from time dependent to STAT HIGH regions, the oscillation period diverges, displaying a $(|\Delta R_m| - |\Delta R_m^c|)^{-1/2}$ behavior, where ΔR_m^c is defined as the value of ΔR_m at the onset of each oscillating regime. On the other hand, when oscillatory regimes approach the STAT LOW region, the transition develops with a finite period.

5 Low dimensional dynamics

5.1 Observations

In order to understand better the observed transitions, we can display a cut in the phase space for the magnetic field recorded in one location, by representing a component of the field as it evolves in time, versus the same component delayed by a time τ . Fig. 7a shows examples of such trajectories in the phase space for different regimes: two fixed points are observed, corresponding to the STAT HIGH regime (green curve) and to the STAT LOW regime (blue, where a transient regime can also be seen). Then the two limit cycles (magenta and red) correspond respectively to a periodically oscillating regime and a randomly reversing one. Examples of the 4 types of dynamo regimes are shown in Fig. 7b to e. The transition between the STAT LOW and the STAT HIGH regimes, via the two time dependent regimes, corresponds to the rising branch in Fig. 5c, in the region of $0.2 < \theta < 0.27$.

5.2 Comparison with a model

Several studies have shown how features of the Earth paleomagnetic records [8,9] or sunspots activity [10,11], can be described as the dynamics of a low dimensional system. In the VKS experiment, the low dimensional nature of the dynamics of the magnetic field has been emphasized in [12]. A model based on the interactions of two

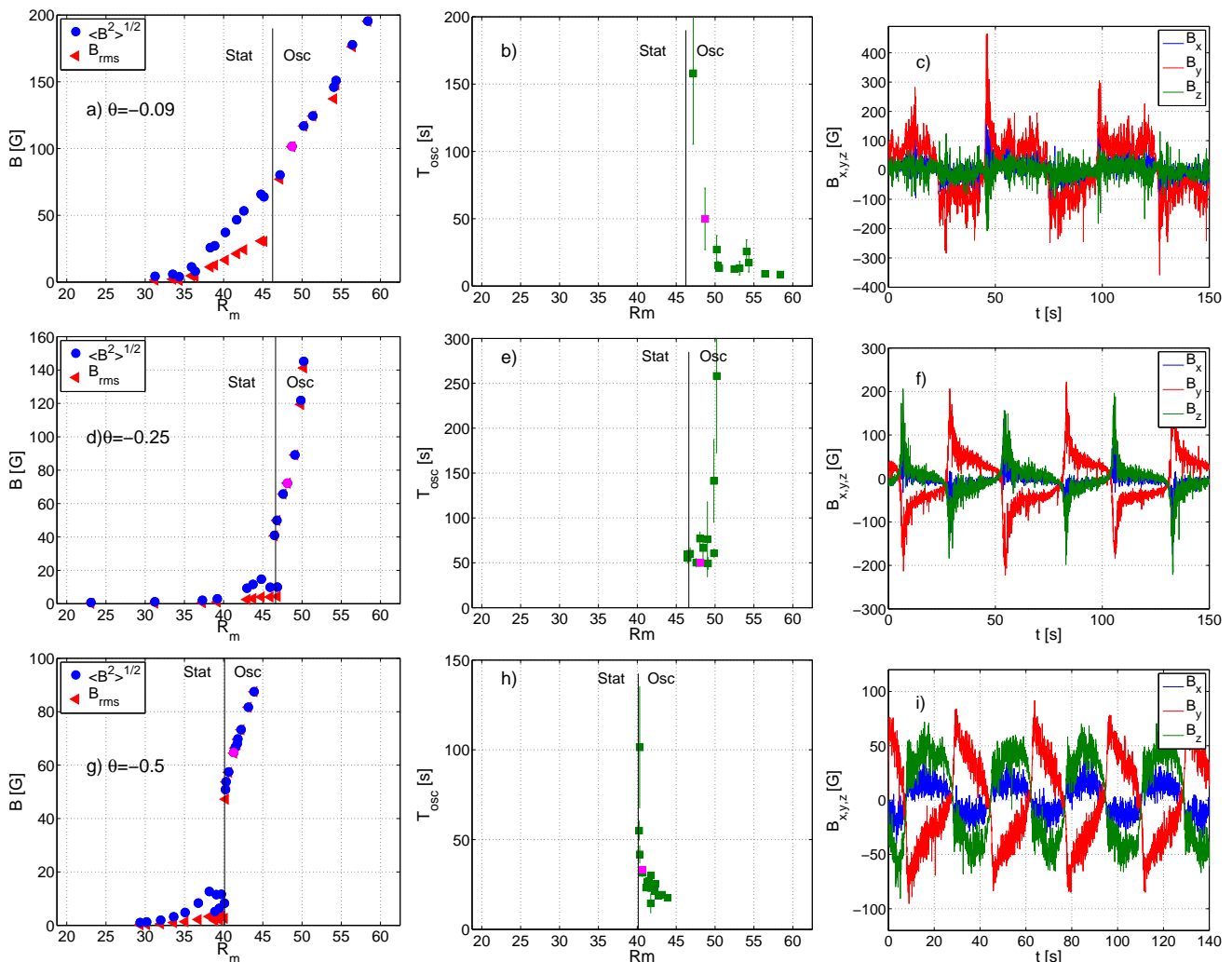


Fig. 4. Examples of dynamo bifurcations with a secondary bifurcation to a time-dependent regime (left) and corresponding periods of oscillation (middle). The error bars represent the standard deviation for the series of measurements. In the case where only one (or one half) period is observed, the error has been arbitrarily set to one third of the period. The separation between stationary and oscillatory regimes is indicated by a vertical line. For each case, one point (magenta symbol in each of the left and middle plots) has been chosen in the oscillating region, for which the corresponding time signal is shown in the right plots.

magnetic modes having the symmetry of a dipole and of a quadrupole has been introduced in [13]. In this framework two types of bifurcations from stationary to oscillatory dynamo regimes have been introduced. When the stationary state is far enough from the dynamo threshold, the transition to an oscillatory regime occurs via a saddle-node bifurcation [13]. When both instability modes are nearly marginal, *i.e.* in the vicinity of a codimension-two bifurcation, bistability of stationary and oscillatory states occurs and the dynamics is more complex [14]. We expect that the STAT HIGH and STAT LOW stationary dynamos bifurcate to time dependent regimes in related ways.

The STAT HIGH regime is at finite distance from the dynamo threshold when it undergoes a time dependent instability, and we ascribe this transition to a saddle-node bifurcation. The fixed point related to STAT HIGH collides with an unstable fixed point and a limit cycle is gen-

erated (see figure 7a), which connects the system to the opposite polarity. As expected in the vicinity of a saddle-node bifurcation, a divergence of the period of oscillation is observed with a $(\Delta R_m - \Delta R_m^c)^{-1/2}$ behavior (see figure 6b). Random reversals (figure 7d) appear at the border between the STAT HIGH stationary state (figure 7e) and the periodic regime (figure 7c). Indeed, slightly before the saddle-node bifurcation, small fluctuations are enough to push the system beyond the unstable fixed point and thus generate a field reversal [15]. Although fluctuations are necessary to escape from the metastable fixed point STAT HIGH, most of the trajectory that connects this point to its opposite in phase space is driven by the deterministic low dimensional dynamics. This explains why trajectories related to different reversals are robust and can be superimposed [3]. But the time between two reversals is random because turbulent fluctuations acting as

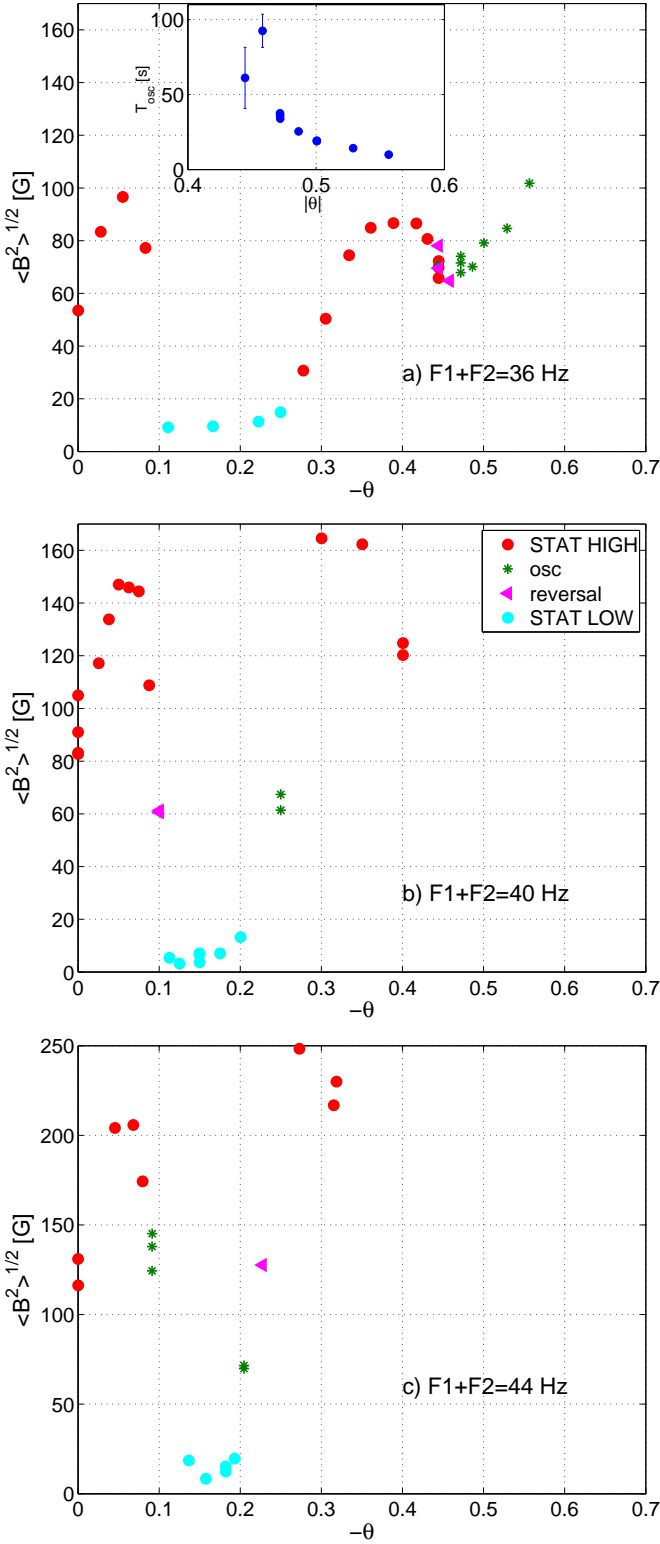


Fig. 5. (a) Transition plots as asymmetry (θ) is varied at fixed F . a) $F_1 + F_2 = 36$ Hz. Color code as in (b). The inset shows the corresponding evolution of the period of oscillation in the oscillating/reversing regimes. (b) $F_1 + F_2 = 40$ Hz. (c) $F_1 + F_2 = 44$ Hz. Color code as in (b).

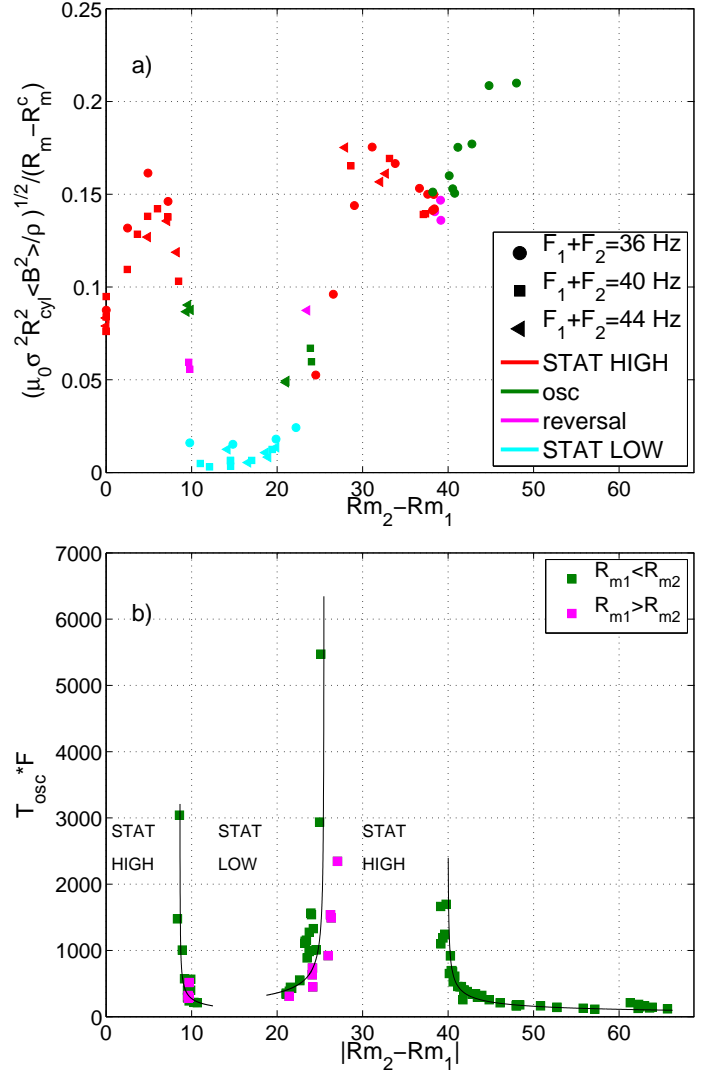


Fig. 6. a) Rescaling of the 3 transition plots of figure 5, with $\Delta R_m = R_{m_2} - R_{m_1}$ for the abscissa and $|B| \sigma R_{cyl} \sqrt{\mu_0 / \rho} / (R_m - R_m^c)$ for the ordinate ($R_m^c = 39$). Symbols correspond to the different values of R_m , colors show the different dynamo regimes. b) Evolution of the period of oscillations on either side of several windows of stationary dynamos. In this plot, since the oscillations period are not expected to depend on the sign of ΔR_m all data points are shown, using different colors for points corresponding to ΔR_m positive and negative. The solid lines are $(|\Delta R_m| - |\Delta R_m^c|)^{-1/2}$ fits, where ΔR_m^c is the value of ΔR_m at the onset of each of the 3 transitions that display a period divergence.

noise trigger the escape from STAT HIGH. This waiting time is governed by the distance to the saddle-node bifurcation and by the intensity of the fluctuations. It can be very long compared to the duration of a reversal when the distance to the saddle node bifurcation increases because the mean exit time depends exponentially on the system parameters (see equation (6) in [15]).

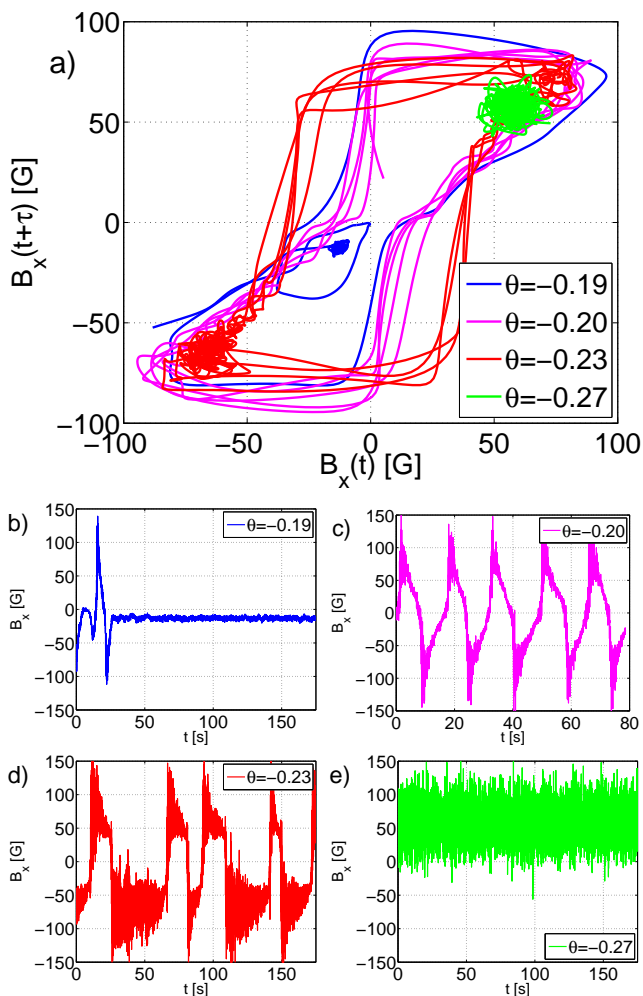


Fig. 7. a) Phase space for the axial component recorded at the location #2, for several kinds of dynamos obtained for different values of θ . The time signal were low-pass filtered at 2 Hz and $\tau = 1$ s. Examples of the corresponding time signals for: b) the low stationary dynamo with a part of transient regime (blue in the phase space plot) c) a periodic reversing dynamo (magenta), d) a random reversing dynamo (red) and e) a high stationary dynamo (green). For these cases R_m is roughly constant around 41.2 ($F_1 + F_2 = 44$ Hz).

We propose that the STAT LOW regime bifurcates to a time dependent regime in a different way, similar to the transitions observed near the regime where only one impeller is rotating, a situation described in details in [14]. The transition from STAT LOW occurs in the vicinity of non-dynamo regimes, as shown in figures 1 and 4. Ascribing the bifurcation to the features of a codimension 2 point, one has the existence of two stationary states with opposite polarities which are encircled in phase space by two limit cycles, an unstable one and a stable one. The unstable orbit separates the stationary states and the stable oscillatory one. When R_m is increased from the stationary state in the bistable regime, the unstable limit cycle bifurcates via a double saddle homoclinic connection with

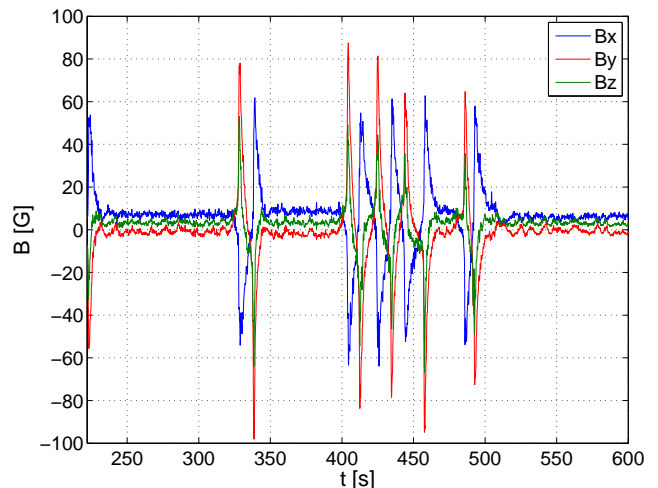


Fig. 8. Example of time signal at $F_1 = 16.75$ Hz and $F_2 = 25.25$ Hz ($\theta = 0.2$).

zero, generating two unstable limit cycles located around each stable fixed point. Then, these limit cycles shrink on the fixed points which become unstable. The system jumps to the stable limit cycle with finite oscillation period. When the forcing is decreased from the oscillatory state in the bistable regime, the system stays on the stable limit cycle until it collides with the unstable one and disappears. The magnetic field then jumps to one of the stationary states. This scenario is in agreement with our observations. It also predicts a region of bistability, clearly evidenced for the one-impeller flow (at $\theta=1$) [14] and also observed in a narrow range of parameters around $\theta = 0.2$, as can be seen in Fig. 8, showing a time signal where the dynamo spontaneously jumps from a STAT LOW regime to an oscillatory regime [16].

6 Discussion

We first consider several aspects related to dynamo generation and then discuss the modeling of the dynamical regimes. Our results indicate that the iron impellers play a crucial role. As shown in Table 1, no dynamo is generated when driving the flow by impellers made of other materials. In addition, when the flow is driven by an iron impeller and a stainless steel one, only the rotation of the iron impeller gives rise to a dynamo and no further time dependent regime is observed. The exact influence of the iron impellers is still an open question, although it has been emphasized in several studies: modeling [17], experimental [18] and numerical [19,20]. New experimental runs, using impellers with disks and blades made of different materials are underway in the VKS set-up, and will hopefully contribute to the understanding of this issue.

We have also observed that the dynamo threshold can depend on the flow characteristics. For instance, without the inner ring, the threshold has not been reached when the impellers are rotated in the scooping direction. This observation, together with others mentioned in section 2.2,

implies that certain aspects of the flow characteristics play a role in the onset of the instability.

Once the threshold is reached for which a dynamo is generated, we have shown that owing to the asymmetry that can be introduced by rotating both impellers at different frequencies, a rich variety of dynamical regimes arises. Experiments in an equivalent water device have been performed in order to check whether the dynamics of the magnetic field is related to instabilities of the flow in the non dynamo regime. They showed that a hysteretic bifurcation is observed at $|\theta| = 0.09$, between a flow with two recirculation cells and a flow with one recirculation cell [5]. In the same set-up, adding the inner ring in the median plane, this hydrodynamic bifurcation is moved to a larger asymmetry ($|\theta| = 0.16$), showing that the presence or absence of the inner ring can strongly change the flow characteristics. As to the magnetic bifurcations observed in the VKS experiment, we have seen that at $|\theta| = 0.09$, a transition between high amplitude and low amplitude steady dynamo occurs, sometimes via a time-dependent one. But this feature is robust whether the inner ring is present or not (see figure 18 in [3]). Other magnetic bifurcations are present around $|\theta| = 0.16$ with the inner ring and, without inner ring, at higher $|\theta|$ values ($0.2 < |\theta| < 0.3$, dependent on R_m) (see Fig. 5).

We also evidenced other situations where some transitions take place that are not driven by hydrodynamics instabilities: for example, we have observed that a periodically oscillating regime observed at $T = 145^\circ\text{C}$ bifurcates to a chaotically reversing one when the temperature drops to $T = 120^\circ\text{C}$ (see figure 22 in [3]). In this case the flow presumably remains unchanged ($|\theta| = 0.16$), while R_m is changed from 42.3 to 44.4 via variations of the sodium electrical conductivity. More generally, one may argue that flows driven at a constant $|\theta|$, for which bifurcations of magnetic behaviors are observed, as discussed in section 3, evolve with the same geometrical characteristics. In VK flows, erratic changes in the structure of the flow have however been reported even when the driving (and everything else such as temperature) had been kept to steady values [5,21]. Thus, one can infer that hydrodynamic bifurcations are not necessary for a magnetic bifurcation to occur, but in some cases, a correlation can exist.

As explained above, the dynamics of the magnetic field reported in this study can be captured by a minimal model which involves the interactions of two magnetic modes. These two modes have the symmetries of an axial dipole and of a quadrupole. The dynamics resulting from the non-linear interactions of two competing modes [13] describes the transitions between neighboring regimes, if one assumes that the two modes are simultaneously marginally stable. Dipole and quadrupole modes have been observed to have nearly the same threshold, in analytical studies of earthlike systems [22] as well as in numerical simulations of the Earth dynamo [23]. In addition, numerical studies aimed at modelling the VKS experiment show that axial dipolar and quadrupolar modes have nearly the same threshold and that their interaction leads to a time

dependent regime when the impellers rotate at different rates [20].

In this description using amplitude equations for the dipolar and quadrupolar components of the magnetic field, the fluid parameters, the flow characteristics and the boundary conditions determine the values of the coefficients of the equations. The level of turbulent fluctuations, which changes with θ [5], is taken into account through multiplicative noise in the model. However, the deterministic part of the dynamics does not explicitly involve velocity modes. Besides describing the bifurcations reported in this paper, this two-dimensional phase space of the deterministic dynamics is crucial to explain why the mean value of the magnetic field should vanish in the time periodic regime, and to correctly predict the shape of the reversals [13]. Including velocity modes will modify the geometry of the phase space and this will be likely to generate behaviors in disagreement with the experimental observations. More generally, no experimental evidence about the dynamics of the magnetic field requires the inclusion of any additional velocity mode in the framework of this model.

There are of course other issues in the VKS experiment besides the dynamical regimes and transitions described here. For instance, the detailed mechanisms of magnetic field generation and saturation still need to be clarified. A future study combining informations from torque measurements and local velocity measurements (using either potential measurements [24] or Doppler velocimetry [25, 26]) could contribute to a better understanding of magnetohydrodynamics features in the VKS dynamo.

Acknowledgements: We thank M. Moulin, C. Gasquet, J.-B Luciani, A. Skiara, N. Bonnefoy, D. Courtiade, J.-F. Point, P. Metz and V. Padilla for their technical assistance. This work is supported by ANR-08-BLAN-0039-02, Direction des Sciences de la Matière et Direction de l’Energie Nucléaire of CEA, Ministère de la Recherche and CNRS. The experiment is operated at CEA/Cadarache DEN/DTN.

References

1. A. Gailitis, O. Lielausis, E. Platacis, S. Dement’ev, A. Cifersons, G. Gerbeth, T. Gundrum, F. Stefani, M. Christen, G. Will, *Physical Review Letters* **86**, 3024 (2001)
2. R. Stieglitz, U. Muller, *Phys. Fluids* **13**, 561 (2001)
3. R. Monchaux, M. Berhanu, S. Aumaitre, A. Chiffaudel, F. Daviaud, B. Dubrulle, F. Ravelet, S. Fauve, N. Mordant, F. Pétrélis et al., *Phys. Fluids* **21**, 035108 (2009)
4. F. Ravelet, A. Chiffaudel, F. Daviaud, J. Léorat, *Phys. Fluids* **17**, 117104 (2005)
5. P. Cortet, P. Diribarne, R. Monchaux, A. Chiffaudel, F. Daviaud, B. Dubrulle, *Phys. Fluids* **21**, 025104 (2009)
6. A. Pinter, B. Dubrulle, F. Daviaud, *Euro. Phys. J. B* **74**, 165 (2010)
7. S.H. Strogatz, *Nonlinear Dynamics and Chaos* (Perseus Books Publishing, Cambridge, MA, 1994)

8. P.M. Melbourne I., R. AM., A heteroclinic model of geodynamo reversals and excursions, in Dynamo and Dynamics, a mathematical Challenge, edited by D. Armbruster, P. Chossat, I. Oprea (NATO Sciences Series - Kluwer Academic Publishers, 2001), Vol. 26, pp. 363–370
9. D.A. Ryan, G.R. Sarson, *Europhys. Lett.* **83**, 49001 (2008)
10. E. Knobloch, A. Landsberg, *Monthly Notices of the Royal Astronomical society* **278**, 294 (1996)
11. P. Mininni, D. Gomez, G. Mindlin, *Solar Physics* **201**, 203 (2001)
12. F. Ravelet, M. Berhanu, R. Monchaux, S. Aumaitre, A. Chiffaudel, F. Daviaud, B. Dubrulle, M. Bourgoin, P. Odier, N. Plihon et al., *Phys. Rev. Lett.* **101**, 074502 (2008)
13. F. Pétrélis, S. Fauve, *Journal of Physics - Condensed Matter* **20**, 494203 (2008)
14. M. Berhanu, B. Gallet, R. Monchaux, M. Bourgoin, P. Odier, J.F. Pinton, N. Plihon, R. Volk, S. Fauve, N. Mordant et al., *J. Fluid Mech.* **641**, 217 (2009)
15. F. Pétrélis, S. Fauve, E. Dormy, J. Valet, *Phys. Rev. Lett.* **102**, 144503 (2009)
16. M. Berhanu, Ph.D. thesis, Université Pierre et Marie Curie (2008)
17. F. Pétrélis, N. Mordant, S. Fauve, *Geophysical and Astrophysical Fluid Dynamics* **101**, 289 (2007)
18. G. Verhille, N. Plihon, M. Bourgoin, P. Odier, J.F. Pinton, *New Journal of Physics* **12**, 033006 (2010)
19. A. Giesecke, F. Stefani, G. Gerbeth, *Phys. Rev. Lett.* **104**, 044503 (2010)
20. C.J.P. Gissinger, *Europhys. Lett.* **87**, 39002 (2009)
21. A. de la Torre, J. Burguete, *Phys. Rev. Lett.* **99**, 054101 (2007)
22. M. Proctor, *Geophys. Astrophys. Fluid Dyn.* **8**, 311 (1977)
23. U. Christensen, J. Aubert, P. Cardin, E. Dormy, S. Gibbons, G. Glatzmaier, E. Grote, Y. Honkura, C. Jones, M. Kono et al., *Phys. Earth Planet Inter.* **128**, 25 (2001)
24. R. Ricou, C. Vives, *International Journal Of Heat And Mass Transfer* **25**, 1579 (1982)
25. Y. Takeda, *Nucl. Technol.* **79**, 120 (1987)
26. H.C. Nataf, T. Alboussiere, D. Brito, P. Cardin, N. Gagniere, D. Jault, D. Schmitt, *Phys. Earth Planet Inter.* **170**, 60 (2008)