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Flow Boiling in Straight Heated Tubes Under Microgravity Conditions

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Keywords: flow boiling, two-phase flow, microgravity, heat transfer, wall friction

Abstract

Forced convective boiling experiments with *HFE-7000* were conducted in earth gravity and under microgravity conditions. The experiment mainly consists in the study of a boiling flow (liquid-vapour flow) through a 6 mm diameter sapphire tube uniformly heated by an *ITO* coating. Measurements of pressure drops (on an adiabatic section), void fraction and wall and liquid temperatures are provided. High-speed movies of the flow are also taken. The data were collected in normal gravity and during a series of parabolic trajectories flown in an airplane. Flow visualizations, temperature, pressure and void fraction measurements are analysed to obtain flow pattern, heat transfer, film thickness and wall friction data.

Introduction

Two-phase thermal systems are broadly used in various industrial applications and engineering fields: flow boiling heat transfer is common in power plants (energy production or conversion...), transport of cryogenic liquids and other chemical or petrochemical processes. Thus, the understanding of boiling mechanisms is of importance for accidentai off-design situations. These systems take advantage of latent heat transportation, which generally enables a good efficiency in heat exchanges.

For that reason, two-phase thermal management systems are considered as extremely beneficiai for space applications. lndeed, in satellites or space-platforms, the major thermal problem is currently to remove the vast heat amount generated by deviees from the inside into the space, in order to ensure suitable environmental and working conditions. Moreover, the growing interest for space applications such as communication satellites and the increasing power requirements of on-board deviees lead to an urgent need of sophisticated management systems capable to deal with larger heat loads. Since the heat transfer capacity associated with phase change is typically large and with a relatively little increase in temperature, this solution could mean decreased size and weight of thermal systems.

But boiling is a complex phenomenon which combines heat and mass transfers, hydrodynamics and interfacial phenomena. Furthermore, gravity consequently affects the fluid dynamics and may lead to unpredictable performances of thermal management systems. It is thus necessary to perform experiments directly in (near) weightless environments. Besides the ISS, microgravity conditions can be simulated by means of a drop-tower, a parabolic flight in aircraft or a sounding-rocket.

Although flow boiling is of great interest for space

applications under microgravity conditions, few experiments have been conducted in low gravity. These experiments provided a partial understanding of the boiling phenomena and have been mostly performed for engineering purposes. Moreover, flow boiling heat transfer experiments in microgravity (referred to as μ -g) are subject to severe restriction in the test apparatus, do not last long and offer few opportunities to repeat measurements for repeatability, which could explain the lack of data and of coherence between existing measurements. Nevertheless, several two-phase flow (gas-liquid flow and boiling flow) experiments have been conducted in the past forty years and enabled to gather data about flow patterns, pressure drops, and heat transfers including critical heat flux and void fraction in thermohydraulic systems. Previous state of the art and data can be found in the papers of Colin et al. (1996), McQuillen et al. (1998), Ohta (2003), and Celata and Zummo (2009).

Several studies have been carried out under microgravity conditions in order to classify adiabatic two-phase flows by various patterns through observation and visualizations of the flow. Various patterns have been identified at different superficial velocities of liquid j_1 and gas j_v , patterns which are also encountered in boiling flows: bubbly flow, slug flow and annular flow. Transitions flows have been studied too: the determination of these transitions is of importance because the wall friction and wall heat transfer are very sensitive to the flow pattern. Colin et al. (1991) and Dukler et al. (1988) drew a map based on void fraction transition criteria to predict patterns in liquid-gas flows. These patterns were also observed in boiling convective for heat transfer smaller than the critical heat flux by Ohta (2003), Lebaigue et al. (1998), Reinarts (1993) and more recently by Celata and Zummo (2009). A general flow pattern map

for bubbly and slug flows based on the value of the Ohnesorge number was proposed by Colin et al. (1996) for air-water flow and also boiling refrigerants. The transition between slug and annular flow has also been investigated by several authors, who proposed criteria based on transition void fraction value (Dukler at al., 1988), critical value of a vapour Weber number (Zhao and Rezkallah, 1993), balance between gas inertia and surface tension (Reinarts, 1993; Zhao and Hu, 2000).

The estimation of void-fraction or averaged gas velocity is a key-point for the calculation of wall and interfacial frictions. It has been shown that the mean gas velocity $U_v = j_v/\alpha$ is well predicted by a drift flux model $U_G = C₀$, *j* for bubbly and slug flow (Colin et al., 1991), j being the mixture velocity and C_0 a coefficient depending on the local void fraction and gas velocity distributions. Few experimental data of film thickness were provided, but only for gas-liquid annular flows (Bousman, 1995; de Jong and Gabriel, $2003...$.

Regarding the measurements of the wall shear stress, most of the studies performed under microgravity conditions concern gas-liquid flow without phase change (Bousman and McQuillen, 1994; Zhao and Rezkallah, 1995; Colin et al., 1996; Choi et al., 2002). Sorne results also exist for liquid-vapour flow (Chen et al., 1991), but only in an adiabatic test section. Recently, Awad and Musychka (2010) and Fang and Xu (2012) proposed modified correlations of Lockhart and Martinelli and found a good agreement with the experimental data. Very few studies reported data on the interfacial shear stress in annular flow (Dukler et al., 1988). This can be explained by the fact that such a measurement is based on pressure drop and liquid film thickness measurements which remain a difficult task.

Few researches on flow boiling heat transfer have been conducted, mainly because of the restrictive experimental conditions. Lui et al. (1994) carried out heat transfer experiments in subcooled flow boiling with *R113* through a tubular tests section (12 mm internai diameter, 914.4 mm length). Heat transfer coefficients were approximately 5 to 20 % higher in microgravity, generally increasing with higher qualities, which was believed to be caused by the greater movement of vapour bubbles on the heater surface. Ohta et al. (1995, 1997, and 2003) studied flow boiling of *FC-72* and *R113* in vertical transparent tubes (4,6 and 8 mm internai diameters), internally coated with a gold film, both on ground and during parabolic flight campaigns, and for a future experiment in the ISS. Authors examined various patterns and the influence of gravity levels on heat transfer coefficients for two-phase forced-convection heat transfer regime. It was noticed that the influence of gravity is not evident for high mass fluxes (G superior to $250 \text{ kg.s}^{-1} \text{.m}^{-2}$). This observation was also made by Baltis, Celata and Zummo (2009) who performed subcooled flow boiling experiments with *FC-72* in Pyrex tubes (2, 4 and 6 mm internai diameters). It was shown that the heat transfer coefficient decreases by up to 30-40% in microgravity in comparison with terrestrial gravity and that an increase of mass or heat flux seems to reduce the influence of gravity. A new technique for the measurement of heat transfer distributions has also been developed: Kim and al. (2012)

used an IR camera to determine the temperature distribution within a multilayer consisting of a silicon substrate coated with a thin insulator. They have not quantified the difference between microgravity and normal gravity yet.

Work has still to be done to confirm and give coherence to the previous results of the literature on flow boiling and to compare the data sets obtained by the different authors.

Objectives

In this work, the authors intend to collect, analyse and compare flow boiling data in normal gravity or under microgravity conditions, thanks to a parabolic flight campaign. The working fluid is the *HFE-7000* which circulates in a heated test section made up of a 6 mm inner diameter sapphire tube with a conductive transparent *ITO* coating. Flow patterns, void fraction, wall friction and heat transfer are studied.

This paper presents the results of the measurement campaigns within three major sections. The first section describes the experimental apparatus and the measurement techniques and accuracy. The data reduction to obtain the mass quality, gas velocity, heat transfer coefficient and wall shear stress is described in a second section. Finally the experimental results obtained in μ -g and 1-g experiments are presented and discussed.

Nomenclature

 τ shear stress (Pa)

Subscripts

coating which enables a uniformly heating by Joule effect and a visual display of the flow.

Figure 2: Schematic of the test section.

Various measurement instruments provide experimental data for the calculation of pressure drops and beat transfer:

adiabatic section: two differential pressure transducers measure the pressure drop along an adiabatic section at the outlet of the test section.

• test section:

- *pressure:* two absolute pressure sensors are used to calculate the saturation temperature at the inlet and outlet of the sapphire tube. No differentiai pressure measurement is performed on this section.

- *temperature:* type K thermocouples are used to measure the flow temperature at the test section inlet and outlet. Two type T thermocouples are also used to measure the temperature difference between a hot junction and a cold junction located at the inlet and outlet of the test section, respectively. This differentiai thermocouple allows a very accurate measurement of the liquid temperature evolution along the test section. Pt100 probes measure the ambient temperature and the temperature of the externat surface of the sapphire tube at four different positions

- *visualization:* a high speed camera provides movies of the flow through the *ITO* coating.

- *void fraction*: specific void fraction probes were designed and built at *IMFT* to provide accurate data of the volume fraction of the vapour phase. These sensors measure the capacitance between two copper electrodes which depends on the fraction of liquid and vapour in the considered volume. One of them is represented in Figure 3. Liquid *HFE-7000* and Teflon roads are used to mimic the annular flow configuration for the calibration.

Figure 3: Schematic of a void fraction probes.

Experiments were conducted both on ground and under microgravity conditions. A weightless situation is simulated during a parabolic flight campaign which consists of three flights with around 31 parabolas per flight. Each parabola

Experimental test setup

The experimental set-up mainly consists of a hydraulic loop which is represented in Figure 1. In this pressurized circuit, the working fluid is the refrigerant 1-methoxyheptafluoropropane (C₃F₇OCH₃), which will be referred as *HFE-7000*. It is first pumped at liquid state by a gear pump while the liquid flow rate is measured by a Coriolis flowmeter. The fluid is heated to its boiling point and partially vaporized in two seriai heaters. Then it enters a stainless steel tube just upstream the test section. In the test section, the *HFE-7000* is further vaporized in a 6 mm diameter sapphire tube heated through an outside *ITO* coating (upward flow). The fluid is then condensed and cooled down 10°C below its saturation temperature into four cold plates including Peltier modules and fans before it enters the pump again. The pressure is adjusted in the circuit via a volume compensator, whose bellow can be pressurized by air.

Figure **1:** Schematic of the experimental apparatus.

The *HFE-7000* has been chosen as working fluid for safety reasons and because of its low saturation temperature at atmospheric pressure (34°C at 1 bar) and its low latent beat of vaporization. In the circuit, the *HFE-7000* may be in a liquid or a liquid-vapour state depending on the portion of the hydraulic circuit, but it is never in a pure vapour state. The loop pressure is set from 1 to 2 bars and the fluid circulates with mass fluxes *G* between 50 and 1200 kg.s^{-1} .m⁻². A wide range of flow boiling regimes is studied, from subcooled flow boiling to saturated flow boiling, by adjusting the power input of the heaters (vapour mass qualities up to 0.9) and the power of the *ITO* coating (wall heat flux up to 45 000 W.m^{-2}).

The test section is represented in Figure 2. It mainly consists of a 20 cm long sapphire tube with a 6 mm inner diameter and a 1 mm thickness. The outer surface is almost totally coated with *ITO,* an electrical conductive and transparent provides up to 22 seconds of microgravity with a gravity level smaller than 3.10^{-2} g.

During the on-ground measurement campaign, relevant parabolas were reproduced in order to compare data obtained in normal gravity and under microgravity conditions. A series of parametric runs bas also been conducted to complete the database.

Data reduction

Hereafter, the calculations of beat transfer coefficients, vapour quality and wall friction are presented. These values are deduced from the measurements of wall and liquid temperatures, beat flux, pressure drop and void fraction by using mass, momentum and enthalpy balance equations.

> Calculation of the beat transfer coefficient

A cross-section of the sapphire tube is represented in Figure 4. *Tiw* and *Tow* are the inner and outer temperatures of the sapphire tube wall. respectively. $T_{\rm esc}$ is the external temperature far from the tube ("at infinity") and $T_{i\infty}$ is the fluid bulk temperature in **Figure 4:** Notations. the tube far from the wall.

 T_{in} and T_{out} are the liquid bulk temperatures at the inlet and outlet of the sapphire tube. The inner and outer radii of the sapphire tube are denoted by R_i and R_o , respectively. The sapphire thermal conductivity is denoted by *k.* The *ITO* coating on the external surface of the test section provides a beat flux *qow·* The beat flux *qiw* delivered to the fluid is considered as equal to q_{ow} corrected by the radii ratio R_o/R_i . The beat transfer between the internai flow and the internai wall of the sapphire tube is characterized by the beat transfer coefficient *hi.* The beat transfer between the environment and the external wall of the sapphire tube (thermal losses) is characterized by the beat transfer coefficient *he.*

Following hypotheses are made: 1) Temperature profiles are axisymmetric. 2) The axial conduction can be neglected. 3) The beat transfer by radiation and the beat transfer with the environment can be neglected.

An energy balance in steady state between the inner and outer walls of the tube leads to expressions (l) and (2).

$$
T_{\text{ow}} - T_{\text{iw}} = [q_{\text{w}} - h_{\text{e}} \cdot (T_{\text{ow}} - T_{\text{e}\infty})] \cdot \frac{\ln\left(\frac{R_0}{R_1}\right) \cdot R_0}{k} \tag{1}
$$

$$
h_{\rm i}.(T_{\rm iw}-T_{\rm i\infty})=\frac{R_{\rm o}}{R_{\rm i}}.\left[q_{\rm ow}-h_{\rm e}.(T_{\rm ow}-T_{\rm e\infty})\right]
$$
 (2)

The temperature evolution between T_{in} and T_{out} is considered as linear or parabolic, which enables to calculate T_{iso} . If thermal losses are neglected (on-ground experiments confirm this hypothesis), the heat transfer coefficient *h*_i can be expressed in function of q_{iw} , T_{ow} and T_{iw} .

$>$ Calculation of vapour quality

The vapour quality can be calculated by using the total enthalpy conservation equation in steady state. q_{iw} is the wall heat flux delivered to the fluid by Joule effect through the *ITO* coating, and D_i is the inner diameter of the sapphire tube.

• for saturated boiling regimes: if T_1 is the liquid temperature, $T_1 = T_{\text{sat}}$, which leads to the expression (3). The evolution of the mass quality can be directly calculated with the wall beat flux and the vapour quality at the inlet of the test section is considered as equal to the vapour quality at the outlet of the heaters.

$$
\frac{4. q_{\text{iw}}}{D_{\text{i}}} = G. \Delta h_{\text{ly}} \cdot \frac{\mathrm{d}x}{\mathrm{d}z} \tag{3}
$$

Considering that the wall beat flux is constant, a linear evolution of vapour quality is observed. The vapour quality at the inlet of the test section is deduced from an energy balance in the preheaters. In this case, the vapour quality calculated with the total enthalpy conservation equation is equal to the classical thermodynamic vapour quality.

• for subcooled boiling regimes, we have $T_1 < T_{\text{sat}}$ and the vapour temperature is assumed to be equal to the saturation temperature. The enthalpy balance equation for the mixture can be written:

can be written:
\n
$$
\frac{4. q_{iw}}{D_i} = \frac{d}{dz} (\rho_v. \alpha. U_v. h_{v, sat} + \rho_l. (1 - \alpha). U_L. h_L)
$$
\n
$$
= \frac{d}{dz} \left(\frac{G. [x. h_{v, sat} + O_l. (T_1 - T_{sat})]}{(1 - x). (h_{l, sat} + C p_l. (T_1 - T_{sat}))} \right)
$$
\n(4)

where h_1 is the liquid enthalpy and $h_{v, \text{sat}}$ and $h_{\text{1,sat}}$ are the vapour and liquid enthalpies at saturation temperature, respectively.

The wall beat flux leads to an increase of the total enthalpy of the mixture, both by phase change and by increasing the liquid temperature:

$$
\frac{4. q_{\text{iw}}}{D_{\text{i}}} = G. \left(\Delta h_{\text{1,v}} + C p_{\text{l}}. (T_{\text{sat}} - T_{\text{l}}) \right) \cdot \frac{\partial x}{\partial z} + G. C p_{\text{l}}. (1 - x) \cdot \frac{\partial T_{\text{l}}}{\partial z}
$$
(5)

For the experiments in subcooled boiling, a single-phase flow is observed at the inlet of the test section. Indeed, a 22 cm long stainless steel tube enables to condensate potential bubbles coming from the two seriai heaters. Then the inlet quality is 0 and equation (5) leads to the expression of *x* through a global balance:

$$
x_{\text{out}} = \frac{4. L_{\text{heated}} \cdot q_{\text{iw}} - G. D_{\text{i}}. C p_{\text{l}}. (T_{\text{out}} - T_{\text{in}})}{G. D_{\text{i}}. [\Delta h'|_{\text{iv}} - C p_{\text{l}} (T_{\text{out}} - T_{\text{in}})]}
$$
(6)

where $\Delta h'_{1,v} = \Delta h_{1,v} + C p_1 \cdot (T_{sat} - T_1)$ (7) For both cases, the fluid temperature is measured at the inlet and outlet of the test section and the temperature evolution

between these two points is considered as linear. The calculation of the vapour quality in subcooled boiling is tricky because of the order of magnitude of *x* and of

measurement uncertainties. We can define measurement errors Δq_w on the measured wall heat flux and ΔT_1 on the measured liquid temperature. Measurement errors on mass flux G, and geometrical and physical properties are neglected. By considering Equation (5), the error Δx on a low vapour quality can be expressed:

$$
\Delta x \approx \frac{4. \Delta q_{\rm w}}{D_{\rm i}} \cdot \frac{L_{\rm heated}}{G. \Delta h_{\rm l,v}} + \frac{C p_{\rm l}}{\Delta h_{\rm l,v}} \cdot \Delta T_{\rm l}
$$
(8)

$$
\Delta x = 4. \frac{10^3}{6. 10^{-3}} \cdot \frac{16. 10^{-2}}{8. 10^2. 14. 10^4} + \frac{13. 10^2}{14. 10^4} \cdot 2. 10^{-1}
$$

$$
\Delta x \propto 2. 10^{-3}
$$

The error on the vapour quality is 2.10^{-3} , which is the order of magnitude of x itself. This error was confirmed by an analysis of flow videos by comparing the mean hubble velocity measured from the images and those calculated using the void fraction and the quality values.

> Calculation of the wall friction

The two differentiai pressure transducers measure the pressure drop ΔP along an adiabatic section at the outlet of the test section, which enables to write the momentum balance equation in steady state without the acceleration term depending of the vapour quality:

$$
\frac{\mathrm{d}P}{\mathrm{d}z} = \frac{4}{D} \cdot \tau_{\mathrm{w}} - g \cdot [\rho_{\mathrm{V}}.\alpha + \rho_{\mathrm{L}}.(1-\alpha)] \tag{9}
$$

By integrating Equation (9) between the inlet and the outlet of the test section, the wall friction can be expressed according to the pressure difference along the adiabatic section and the void fraction:

$$
\Delta P = -\frac{4}{D} \int_{\text{inlet}}^{\text{outer}} \tau_{\text{w}} \, \text{d}z + g \int_{\text{inlet}}^{\text{outlet}} (\rho_{\text{V}} \, \alpha(z) + \rho_{\text{L}} \cdot (1 - \alpha(z))). \, \text{d}z \tag{10}
$$

 ΔP is measured by the differential transducers (there is no vapour in transmissions lines) and α is measured at the inlet and outlet of the test section by two void fraction probes.

In μ -g, the last term of equation (10) can be neglected. On ground, in vertical flow, the last term is dominant, thus the accuracy on the wall shear stress measurement is directly linked to the accuracy on the void fraction measurement itself.

Measurements for single-phase flows in normal gravity have enabled to validate the measurement technics and calculation protocols by confronting the data to standard correlations.

> Pressure drops validation

For the wall friction coefficient f_w in single-phase flow in normal gravity, the Blasius correlation based on the global Reynolds number *Re* is used:

 $f_{\mathbf{w}} = 0.0791$. Re $^{-0}$ 25 (11) Figure *5* shows the measurements obtained with the two differentiai transducers in 1-g for single-phase flows with various mass fluxes and the Blasius' correlation with error bars at $+10$ % and -10 %.

Figure 5: Wall friction coefficient in single-phase flow

> Heat transfer coefficient validation

For a fully developed turbulent single-phase flow, the Nüsselt number Nu_{∞} can be calculated with the Dittus and Boelter's correlation (Table 1) and the Gnielinski's correlation (Table 1), based on the Reynolds number *Re* and the Prandtl number *Pr*. The wall friction coefficient f_w is calculated with the Blasius' correlation. *z* is the distance between the probe and the inlet of the heated section and D_h is the hydraulic diameter which is equal to the inner diameter D_i .

Invest.	Correlations and application range
	Dittus and $Nu_{\infty} = 0.024$. $Re^{0.8}$. $Pr^{0.4}$
Boelter	Fully developed turbulent flow, $L/d > 60$ $0.7 \le Pr \le 120$ and $2500 \le Re \le 1.24.10^5$
Gnielinski	$Nu_{\infty} = \frac{(f_{\text{w}}/2) . (Re - 1000). Pr}{1 + 12.7 . (f_{\text{w}}/2)^{1/2} . (Pr^{2/3} - 1)}$
	Fully developed turbulent flow
	$0.5 \le Pr \le 2000$ and $2300 \le Re \le 5.10^{\circ}$
Al-Arabi	$\frac{\overline{Nu_{\rm m}}}{\overline{Nu_{\infty}}} = 1 + \frac{C}{\overline{z}/D_{\rm h}}$
	$C = \frac{\left(\frac{Z}{D_h}\right)^{0.1}}{P_{H} \cdot 1/6} \left(0.68 + \frac{3000}{P_{H} \cdot 0.81}\right)$
	Thermally developing flow Pr > 0.2

Table 1: Nüsselt correlations for fully developed and thermally developing flows in smooth and circular ducts.

Nüsselt numbers corresponding to the four Pt100 probes located on the outer surface of the sapphire tube for various mass fluxes in single-phase flows are calculated and compare to the Gnielinski's correlation. The deviation from the correlation is in inverse proportion to the distance between the temperature sensors (1, 2, 3 or 4) and the inlet of the heated section; hence the hypothesis of a thermally developed flow in the sapphire tube is not satisfied.

Whenever the flow is thermally developing, a local heat transfer coefficient corresponding to a mean Nüsselt number $Nu_{\rm m}$ is measured. This number must be corrected according to the distance between the measurement point and the inlet of the heated section in order to calculate the fully developed flow Nüsselt number Nu_{∞} and to compare it with correlations. Figure 6 shows the measurements corrected with Al-Arabi's correlation (Table 1) and compared to the Gnielinski's formula.

The experimental data corrected according to the sensor position meet the Gnielinski's correlation with a maximal error of ±17%. The error depends of the mass flux (that varies the thermal entrance length) and of the probe position. The precision between measurements and correlations is satisfying for the whole set of experiments in single-phase flow. It also confirms the weak impact of externat thermal heat losses on the measurements.

Figure 6: Measured Nüsselt number with correction.

Results and Discussion

Experimental results about flow patterns, wall friction and heat transfer coefficients are presented. Preliminary data concerning the evolution of the liquid film thickness in annular flow are also plotted.

Flow patterns

The high speed camera has enabled to visualize flow patterns for various mass fluxes G, vapour qualities *x* at the inlet of the test section and heat fluxes q_{ow} through the *ITO* coating. Three main flow patterns have been observed in both normal gravity and microgravity conditions: annular flow, slug flow and bubby flow were identified on the videos. Another flow pattern is also referred in the literature as slug/annular transition.

Figure 7 shows annular flow and slug flow with nucleated boiling in 1-g and μ -g, and Figure 8 shows a comparison between bubbly flows in 1-g and μ -g for the same parameters $(G, x \text{ and } q_{ow})$. Bubbly flows correspond to subcooled regimes ($T_1 < T_{sat}$). The impact of gravity level on the bubbles size and shape can be seen in the videos for mass fluxes lower than $400 \text{ kg.s}^{-1} \text{.m}^{-2}$: under microgravity conditions, bubbles are larger than in normal gravity and are not deformed by the gravity field. The larger hubble size in microgravity can be explained by both the larger hubble diameter at detachment and the higher rate of coalescence. A vapour quality increase leads through a coalescence phenomenon to slug flow which alternates between long Taylor bubbles and liquid plugs that can be aerated by small bubbles. Coalescence of long Taylor bubbles indicates the transition to annular flows which almost always correspond to saturated regimes $(T_1 = T_{sat})$. The precision of the flow parameters setting and the camera spatial resolution do not enable to see clearly differences between $1-g$ and $\mu-g$ in the videos for annular flows.

Figure 8: Flow visualizations bubbly flows, $\Delta T_{\text{sub}} = 12 \text{ °C}, q = 2 \text{ W.cm}^{-2}$ $G = 540 \text{ kg.s}^{-1} \text{ m}^{-2}$ (a) in 1-g, (b) in μ -g $G = 220 \text{ kg.s}^{-1} \text{ m}^{-2}$ (c) in 1-g, (d) in μ -g

Figures 9 and 10 show flow patterns maps for the experiments performed in normal gravity and under microgravity conditions, respectively. Regimes are indicated according to the superficial vapour and liquid velocities, and isocurves for G and *x* are added.

Figure 9: Flow patterns map in 1-g

Figure 10: Flow patterns map in μ -g

Because of the uncertainty about the calculation of the vapour quality, experimental points for *x* inferior to 0.01 are not shown. Figure 10 presents relevant parabolas performed during the first and second flight campaigns for mass fluxes up to $400 \text{ kg.s}^{-1} \text{.m}^{-2}$. Indeed, a preliminary data reduction just after the first measurement campaign has underlined the fact that gravity effect was not clearly quantifiable for mass fluxes G superior to $400 \text{ kg.s}^{-1} \text{ m}^{-2}$. Therefore, no experiments and parametric runs were performed in 1-g and μ -g for G > 400 kg.s⁻¹.m⁻² afterwards.

In normal gravity, the transition between bubbly and annular flow occurs very fast for low mass fluxes, typically at x around 0.1. For higher mass flux, this transition seems to last within a larger x range and some subcooled annular flows can be observed, but these regimes were not further investigated. Bubbly and slug flows evolve toward annular flow earlier in microgravity (transition around $x = 0.08$ -0.09 for Iower mass fluxes) than in normal gravity (transition around $x = 0.1$) because of the more important coalescence phenomenon.

Film thickness

The data set used to study the film thickness contains runs whose points are in both annular and transitional flows (near the slug-to-annular region). The calculations are made by assuming that ali the liquid is included in the liquid film surrounding the vapour core (the droplets entrainment phenomenon is neglected). With this hypothesis, the film thickness is directly obtained with the void fraction data, as a time averaged value.

Results are shown in Figure 11 for both microgravity and normal gravity conditions. The liquid film thickness δ is plotted as a function of the vapour quality for two different liquid mass fluxes $(G = 400 \text{ kg.s}^{-1} \text{ m}^{-2} \text{ and } 200 \text{ kg.s}^{-1} \text{ m}^{-2})$.

Figure 11: Liquid film thickness in 1-g and μ -g

As can be seen from Figure 11, changing the vapour quality has a significant influence on the liquid film thickness which decreases as *x* increases through a non-linear process, both in μ -g and 1-g (an asymptote can be observed as x increases to high qualities). Changing the liquid mass flux has less effect on ô, even if an increase of the mass flux G leads to an increase of the film thickness at a constant vapour quality. This is consistent with the fact that annular flow is dominated by the increasing effect of liquid and gas inertial forces compared to the gravity forces.

Moreover, it can be seen that for given vapour quality and mass flux, a change in the gravity level leads to important changes in the film thickness. Even if the difference between 1-g and μ -g depends on the mass flux, there is a clear tendency towards a thinner liquid film in microgravity. On ground, the gravity force tends to drive the liquid film downwards, leading to a larger film thickness.

Wall friction

In microgravity, the gravity term of Equation (9) can be neglected. Therefore, the wall friction can be directly deduced from the pressure drop measurement.

During the first campaign, the Validyne transducers were located on the heated test section. Preliminary reduction using Martinelli's relationship for the determination of α has shown that the inertia term represents up to 20% of the pressure drop and cannot be neglected in the global balance. By taking this consideration into account, the differentiai pressure measurement was moved to an adiabatic, which enables to neglect the inertia term.

Figure 12 and Figure 13 represent the experimental two-phase frictional multiplier Φ_{LO} in normal gravity and microgravity respectively, compared to the one predicted by the Lockhart and Martinelli's correlation (1949). The experimental Φ_{LO} is calculated from the wall shear stress measurements, and plotted versus the Martinelli's parameter X given in Equation 11.

$$
X = \frac{j_1}{j_v} \cdot \sqrt{\frac{\rho_1}{\rho_v} \cdot \frac{f_{\text{wl}}}{f_{\text{wv}}}}
$$
(11)

$$
\Phi_{LO}^2 = \left(1 + \frac{C}{X} + \frac{1}{X^2}\right) \tag{12}
$$

Figure 12: Pressure drops in normal gravity compared to Lockhart and Martinelli's correlation

Figure 12: Pressure drops in microgravity compared to Lockhart and Martinelli's correlation

The significant error on the superficial vapour velocity for low vapour qualities (subcooled data) leads to an underestimation of the associated two-phase multiplier for the bubbly flows. On the other hand, the saturated data (corresponding to annular flows) are in very good agreement with the Lockhart and Martinelli's correlation in 1-g. The same trend can be observed with the μ -g data that fit the correlation with a 5% mean deviation for the annular regime data and 27% mean deviation for the bubbly regime data.

Heat transfer coefficients

Heat transfer coefficients in 1-g and in μ -g for saturated boiling regimes and subcooled boiling regimes are presented in Figures 13, 14 and 15.

The precision of the calculation of h_i (also denoted h) is directly linked to the measurement accuracy of q_{row} and of temperatures (wall and liquid bulk temperatures) and to the chosen evolution of the temperature in the sapphire tube between T_{in} and T_{out} . By using a linear evolution of the temperature in the heated section and by considering uncertainties $\Delta q_{\rm w}$ and ΔT on the heat flux and temperatures, respectively, heat transfer coefficients can be calculated with an error bar at ±14%.

Results about heat transfers are difficult to summarized because the coefficient h_i and the different correlations depends on various parameters such as the vapour quality, mass flux, heat flux or gravity level.

Figure 13 shows heat transfer coefficients according to the vapour quality for a wall heat flux $q_i = 2$ W.cm⁻² and mass fluxes between $G = 100 \text{ kg.s}^{-1} \text{ m}^{-2}$ and $G = 400 \text{ kg.s}^{-1} \text{ m}^{-2}$, for saturated boiling regimes.

Measurements in normal gravity are compared with two boiling heat transfer correlations: Chen's correlation (1966) and Kandlikar's correlation (1989) which brings into play the boiling number and a coefficient linked to the Martinelli's parameter. At a given mass flux, the evolution of h_i on ground is the expected one: the heat transfer coefficient increases with the vapour quality at fixed heat flux, and increases with heat flux at fixed vapour quality. The heat transfer coefficients also increase with *G* on the chosen mass fluxes range. Measurements in normal gravity fit the empirical correlations (calculated for given *q* and G) with a mean deviation from Kandlikar's correlation that is around 16%.

At $G = 200 \text{ kg.s}^{-1} \text{ m}^{-2}$, the effect of gravity level on boiling heat transfer is more quantifiable than for higher mass rates (results from the frrst parabolic flight campaign have shown that at $G = 500 \text{ kg.s}^1 \text{ m}^2$, the mean difference on h_i between 1-g and μ -g is $\pm 3.5\%$). For that reason, Figure 14 shows experimental heat transfer coefficients in microgravity for mass fluxes G ranging from 100 kg.s⁻¹.m⁻² to 300 kg.s⁻¹.m⁻². Heat transfer data are given for a heat flux $q_i = 2$ W.cm⁻² and compared with Chen and Kandlikar's correlations.

As can be seen from Figure 14, it seems that the beat transfer coefficients are lower in microgravity than in normal gravity for low vapour qualities $(x < 0.2$ at $G = 200$ $kg.s^{-1}.m^{-2}$). For these qualities, Kandlikar's correlation overestimates the μ -g heat transfer coefficient. On the other hand, the heat transfer coefficients are higher in μ -g than in 1-g for high vapour qualities ($x > 0.2$ at $G = 200$ kg.s⁻¹.m⁻²) and Kandlikar's correlation underestimates *h* in this case. The experiments performed in subcooled boiling flows tend to confirm this trend, as can be seen from Figure 15.

Figure 15: Heat transfer coefficients in 1-g and μ -g for subcooled boiling flows (bubbly flows) for various beat fluxes and mass fluxes between 100 and 400 kg.s⁻¹.m⁻²

The subcooled data points (bubbly flows data in Figure 15) correspond to vapour qualities inferior to 20% ($x < 0.2$). For ali these points (whatever the beat flux, mass flux or vapour quality is), the heat transfer coefficient is lower in microgravity than in normal gravity (by 20% on average). These results are consistent with the one in saturated flow regimes. This trend bas to be confirmed by additional experiments. In bubbly flow, the hubble slip velocity is close to zero in microgravity, whereas it is significant in vertical flow on ground. This relative velocity induces a distorsion of the velocity profile near the wall and an increase of the velocity gradient and turbulence level in the liquid flow (Colin et al., 2012). This effect is also responsible for an increase of the wall friction in bubbly flow and may explain the larger beat transfer coefficient in normal gravity.

Conclusions

This paper presents the results of flow boiling experiments performed under microgravity conditions during two parabolic flight campaigns and compared to parametric runs conducted on ground. The objective was to collect beat transfer, void fraction and wall friction data along a heated test section consisting of a 6 mm inner diameter heated sapphire tube, using *HFE-7000* as working fluid.

Special attention was paid to the calculation of the vapour quality in arder to characterize properly the subcooled boiling regimes, but it remains difficult to investigate flows with very low superficial vapour velocities .

Annular flow, slug flow and bubbly flow have been observed in videos according to the vapour quality and the mass fluxes. The results show that the gravity level bas little impact on the flow for mass fluxes superior to 400 kg.s⁻¹.m⁻² whatever the flow pattern is. That is the reason why lower mass fluxes were investigated in this article.

The transition between slug and annular flow seems to occur at lower qualities in microgravity.

Experimental pressure drops data fit the Lockhart and Martinelli's correlation with a good agreement both in normal gravity and microgravity.

On-ground beat transfer values are in good agreement with classical correlations of the literature. However, significant discrepancies with these correlations are found in microgravity. Flow boiling beat transfer at low mass fluxes under microgravity conditions can be either increased for high vapour qualities or reduced by up to 30% for vapour qualities inferior to 20%, which is consistent with the film thickness data that show that the liquid film is thinner under microgravity conditions.

Additional experiments at lower mass fluxes should be conducted in order to validate this trend. Another parabolic flight campaign will be the opportunity to perform new experiments in microgravity. Future test matrix plans to conduct runs at lower mass fluxes by adapting the hydraulic loop and to improve the accuracy on the temperature measurement for the calculation of the vapour quality. Further data processing may enable to access the interfacial friction and new heat transfer model.

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