



**HAL**  
open science

## The $\nu_3$ band of $^{14}\text{C}^{16}\text{O}_2$ molecule measured by optical-frequency-comb-assisted cavity ring-down spectroscopy

Iacopo Galli, Pablo Cancio Pastor, Gianfranco Di Lonardo, Luciano Fusina, Giovanni Giusfredi, Davide Mazzotti, Filippo Tamassia, Paolo de Natale

► **To cite this version:**

Iacopo Galli, Pablo Cancio Pastor, Gianfranco Di Lonardo, Luciano Fusina, Giovanni Giusfredi, et al.. The  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  molecule measured by optical-frequency-comb-assisted cavity ring-down spectroscopy. *Molecular Physics*, 2011, 109 (17-18), pp.2267-2272. 10.1080/00268976.2011.614284 . hal-00742695

**HAL Id: hal-00742695**

**<https://hal.science/hal-00742695v1>**

Submitted on 17 Oct 2012

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



**The  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  molecule measured by optical-frequency-comb-assisted cavity ring-down spectroscopy**

Journal:	<i>Molecular Physics</i>
Manuscript ID:	TMPH-2011-0214.R1
Manuscript Type:	Special Issue Paper - Dijon HRMS
Date Submitted by the Author:	28-Jul-2011
Complete List of Authors:	Galli, Iacopo; Istituto Nazionale di Ottica-CNR Cancio Pastor, Pablo; Istituto nazionale di Ottica-CNR Di Lonardo, Gianfranco; Università di Bologna, Dipartimento di Chimica Fisica e Inorganica fusina, luciano; Università di Bologna, Dipartimento di Chimica Fisica e Inorganica Giusfredi, Giovanni; Istituto Nazionale di Ottica-CNR Mazzotti, Davide; Istituto Nazionale di Ottica-CNR Tamassia, Filippo; Università di Bologna, Dipartimento di Chimica Fisica e Inorganica De Natale, Paolo; Istituto nazionale di Ottica-CNR
Keywords:	radiocarbon dioxide, infrared spectroscopy, $\nu_3$ fundamental band, optical frequency comb, cavity ring-down
Note: The following files were submitted by the author for peer review, but cannot be converted to PDF. You must view these files (e.g. movies) online.	
14CO2_REW.tex	

SCHOLARONE™  
Manuscripts

For Peer Review Only

1  
2  
3  
4  
5  
6  
7  
8  
9  
10  
11  
12  
13  
14  
15  
16  
17  
18  
19  
20  
21  
22  
23  
24  
25  
26  
27  
28  
29  
30  
31  
32  
33  
34  
35  
36  
37  
38  
39  
40  
41  
42  
43  
44  
45  
46  
47  
48  
49  
50  
51  
52  
53  
54  
55  
56  
57  
58  
59  
60

## RESEARCH ARTICLE

The  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  molecule measured by optical-frequency-comb-assisted cavity ring-down spectroscopy

Iacopo Galli<sup>a,\*</sup>, Pablo Cancio Pastor<sup>a</sup>, Gianfranco Di Lonardo<sup>b</sup>, Luciano Fusina<sup>b</sup>,  
Giovanni Giusfredi<sup>a</sup>, Davide Mazzotti<sup>a</sup>, Filippo Tamassia<sup>b</sup>, Paolo De Natale<sup>a</sup>

<sup>a</sup>*Istituto Nazionale di Ottica - CNR and LENS, Via N. Carrara 1,  
50019 Sesto Fiorentino FI, Italy*

<sup>b</sup>*Dipartimento di Chimica Fisica ed Inorganica, Università di Bologna,  
Viale del Risorgimento 4, 40136 Bologna, Italy*

(July 28, 2011)

The infrared spectrum of the rare isotopologue  $^{14}\text{C}^{16}\text{O}_2$  has been investigated in the range 2190–2250  $\text{cm}^{-1}$  with a frequency comb-referenced cavity ring-down spectrometer. Thirty-three ro-vibrational transitions of the  $\nu_3$  fundamental band have been detected. Their absolute frequency was measured with a relative uncertainty ranging from  $7.1 \times 10^{-8}$  to  $1.1 \times 10^{-8}$ . The experimental frequencies were fitted to the conventional Hamiltonian of a linear molecule and a new set of spectroscopic parameters for the fundamental vibrational state has been improved for this species.

**Keywords:**  $^{14}\text{C}^{16}\text{O}_2$ ; infrared spectroscopy;  $\nu_3$  fundamental band; optical frequency comb; cavity ring-down

## 1. Introduction

Infrared spectroscopy of radiocarbon dioxide ( $^{14}\text{C}^{16}\text{O}_2$ ) has been the subject of very few papers [1–6]. The most recent systematic works on the  $\nu_3$  fundamental band were performed by Sams and De Voe in 1988 [5], and by Dobos *et al.* [6] in 1989. In Ref. [5], they measured twelve P and R-branch transitions by diode laser spectroscopy. The lines were assigned an uncertainty ranging between 0.001 and 0.005  $\text{cm}^{-1}$ . In Ref. [6] a medium resolution (0.18  $\text{cm}^{-1}$ ) FTIR spectrum and a high-resolution diode laser spectrum were recorded. In the latter case, the precision estimated by the authors was better than 0.001  $\text{cm}^{-1}$ . Some of the papers published on this subject dealt with  $\text{CO}_2$  lasers [3, 4], because the use of  $^{14}\text{C}^{16}\text{O}_2$  can broaden the working range of spectrometers using  $\text{CO}_2$  lasers as source of radiation. However, this species is of particular interest for several other reasons. A spectroscopic study can provide physical information on a still poorly investigated isotopologue of  $\text{CO}_2$ . This molecule is produced in nuclear power reactors and it is detected in order to monitor its environmental concentration. This is usually done through accelerated mass spectrometry [7]. Labrie and Reid [8] showed that radiocarbon dating can be performed by infrared spectroscopy. Also,  $^{14}\text{C}$  measurements have been carried out on planetary atmospheres [9].

\*Corresponding author. Email: iacopo.galli@ino.it

In this paper we report on the high-resolution investigation of the fundamental  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  by comb-referenced cavity ring-down spectroscopy. The main advantages of this technique are: high sensitivity ( $\text{CO}_2$  absorptions with line-strength as low as  $5 \times 10^{-32}$  cm can be detected), high resolution (1 kHz) and absolute frequency measurements, which in turn implies that no calibration of the spectra is needed.

## 2. Experimental details

Cavity ring-down (CRD) spectroscopy at  $4.5 \mu\text{m}$  was realized to detect thirty-three ro-vibrational transitions belonging to the  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  molecule. The experimental set-up, described elsewhere [10–12], uses a cw IR coherent radiation at  $4.5 \mu\text{m}$  generated by non-linear frequency difference (DFG) on a periodically-poled  $\text{MgO}:\text{LiNbO}_3$  crystal, placed inside a cw ring Ti:sapphire laser cavity [13]. Such a laser, single-mode operated by optical injection of an external cavity diode laser (ECDL), is the DFG “pump” source, whereas the DFG “signal” light is provided by an external Nd:YAG laser at 1064 nm. A continuous IR tuning between  $3850 \div 4540$  nm ( $2597 \div 2203$   $\text{cm}^{-1}$ ) is allowed thanks to the tuning range of the ECDL-Ti:sapphire laser between  $834 \div 862$  nm, with a maximum IR power of 25 mW. The IR source is frequency stabilized against an octave-spanning ( $500 \div 1100$  nm) fs Ti:sapphire Optical Frequency Comb Synthesizer (OFCS), as “pump” and “signal” frequencies fall within the OFCS spectral operation range [14]. As a consequence, the absolute frequency  $\nu_i$  of the generated “idler” radiation is given by the following equation:

$$\nu_i = \nu_p - \nu_s = \left( \frac{N_p}{N_s} - 1 \right) \nu_s + \nu_\mu \quad (1)$$

where  $N_p$  and  $N_s$  are the integer orders of the closest comb modes to the “pump” and “signal” absolute frequencies,  $\nu_p$  and  $\nu_s$  respectively. As the Nd:YAG laser frequency drift is controlled against the nearest OFCS mode, the absolute frequency of the IR radiation is stable to about  $6 \times 10^{-13}$  at 1 s and accurate to  $10^{-12}$ , both limited by the performance of the Rb/GPS-disciplined 10-MHz quartz reference oscillator of our OFCS. In addition, synthesized IR frequency scans with a Cs-frequency standard traceability can be performed by changing the microwave synthesizer frequency  $\nu_\mu$  in Eq. 1. With this procedure a maximum scan of 450 MHz (a bit less than half the repetition rate of the OFCS) can be performed, wide enough to detect the whole absorption line profiles of the observed transitions for the molecular gas conditions of the present measurements. The generated “idler” IR radiation, sent out of the Ti:sapphire cavity thanks to dispersion at the  $\lambda_p$ -Brewster-cut output facet of the non-linear crystal, is efficiently coupled into a 1-m-long high-finesse ( $F \approx 11600$ ) Fabry-Perot cavity (FPC) for CRD spectroscopy. When the FPC is loaded up to a threshold level, the IR light is quickly switched off-resonance by using an AOM placed on the Nd:YAG laser, and transmitted photons leaking out of the cavity are detected by a InSb photodiode. The CRD decay signal is acquired by a 18-bit digitizing oscilloscope with a sampling rate of 10 MS/s. The measured decay time,  $\tau_c = 1/\gamma_c$  for empty-cavity conditions is about  $12 \mu\text{s}$  ( $\gamma_c \approx 86$   $\text{ms}^{-1}$ ).

Such a set-up is suitable for linear-absorption CRD spectroscopy [15] and saturated-absorption cavity ring-down (SCAR) spectroscopy [10], with sensitivity figures, expressed in terms of the minimum detectable absorption coefficient  $\alpha_{\min}$ , of  $10^{-9}$  and  $10^{-10}$   $\text{cm}^{-1}\text{Hz}^{-1/2}$ , respectively. Although line-strengths of the ro-vibrational transitions of  $^{14}\text{C}^{16}\text{O}_2$   $\nu_3$  band are expected to be of the same order

of magnitude as other  $\text{CO}_2$  isotopologues (about  $10^{-18}$  cm for the most intense), the very low natural abundance of such species ( $\approx 10^{-12}$ ) needs high-sensitivity spectroscopy techniques, as CRD, to be detected. For the measurements presented here a pure  $^{14}\text{C}^{16}\text{O}_2$  sample would be preferred, to get the highest possible signal without absorption interferences from other  $\text{CO}_2$  lines. Unfortunately, it is very difficult to handle such a sample, taking into account also radioactivity safety reasons. Only spectroscopy of a few orders of magnitude  $^{14}\text{C}$ -enriched samples (with respect to natural abundance) is feasible, which has partly prevented the spectroscopic measurements of such molecule. Even in these enriched samples, ultra-sensitive techniques such as CRD or SCAR are required. As a price to pay, other  $\text{CO}_2$  absorptions lying in the same spectral region are also detected, which are added as an interference background to be accurately evaluated (see fitting procedure in the “results and discussion” section for other details). For the measurements presented here, we used a 2200 times  $^{14}\text{C}$ -enriched  $\text{CO}_2$  gas. Fig. 1 shows the IR spectra of most part of the ro-vibrational transitions between the P(40) and R(30) lines of  $^{14}\text{C}^{16}\text{O}_2$   $\nu_3$  fundamental band.

Lines from the P(30) to R(30) were recorded by linear-absorption CRD spectroscopy with a total  $\text{CO}_2$  pressure of 12 mbar. The FPC was dry-ice cooled in this case to get a gas temperature of 195 K, thus minimizing the interference effects from hot-bands of other  $\text{CO}_2$  isotopologues. Nevertheless, strong interferences are detected, as it can be seen in the R-side of Fig. 1.

Lines from the P(40) to P(32) were recorded by SCAR spectroscopy with a total  $\text{CO}_2$  pressure of 18 mbar<sup>1</sup>. For these transitions, room temperature (296 K) was used to increase the ground state population, and hence the absorption signal. At the same time, the interfering background is increased, but it is still not larger than the  $^{14}\text{C}^{16}\text{O}_2$  transition to be measured. In addition, the SCAR spectroscopic technique independently measures, in the same decay event, both  $\gamma_g$  and  $\gamma_c$ . The former decay rate  $\gamma_g = c\alpha_g$  contains only the saturable part of the gas absorption. The latter includes all other contributions: cavity mirror losses, technical fringes, absorption from non-saturable transitions interfering with the target line.

We plot in Fig. 1  $\alpha_g$  values (left vertical scale) for these SCAR lines, all having a negligible background except P(40), at which frequency the achievable saturation parameter was too low. On the contrary, for the CRD spectra,  $\alpha = (\gamma_g + \gamma_c)/c$  was measured (right vertical scale in Fig. 1, where a constant  $\alpha_c = 2.87 \times 10^{-6}$  cm<sup>-1</sup> due to the empty FPC losses was removed), and residual interfering background is present. In fact, the P(10), R(24) and R(28) transitions were not recorded because of the absorbing background.

For all the lines shown in Fig. 1, the acquisition procedure was the same independently of the CRD or SCAR technique used. For each IR absolute frequency, 128 consecutive decay events are averaged, and fitted to the expected CRD or SCAR decay function to obtain as a result  $\alpha$  or  $\alpha_g$ , respectively. Ten consecutive  $\alpha$  (or  $\alpha_g$ ) values are measured before a  $\pm 10$ -MHz IR frequency step is applied and the FPC length adjusted to retrieve the resonance condition. Each full 400 MHz scan is performed by stepping the frequency once forward (+ sign) and once backward (- sign). After each scan, the  $\alpha$  (or  $\alpha_g$ ) plotted value at each scan frequency is the weighted mean of the 20 measured values (ten + and ten -). This procedure takes about 5 min for each complete scan. In the present gas conditions, linewidths are limited by Gaussian thermal broadening (FWHM 98 MHz and 120 MHz for 195 K and 296 K respectively) and Lorentzian pressure broadening [16]. Hence, a total

<sup>1</sup>With respect to Ref. [10], valid in the inhomogeneous broadening regime, the theoretical model to fit the SCAR signals was adjusted to work at these pressure conditions [12].

1 400 MHz scan is adequate to record a whole line profile. We note finally that in  
2 Fig. 1 all spectra were arranged together, one close to the other in a relative fre-  
3 quency scale, for the sake of clarity. For the analyzed spectra, the frequency scale  
4 is absolute and reproducible against the OFCS, as described above.  
5  
6  
7

### 8 3. Results and Discussion

9  
10 Only thirty-three transitions, plotted in Fig. 1, were detected in our experi-  
11 ments. The main limitations to the detection of transitions with higher values of  
12  $J$  arose from the operating frequency range of the spectrometer, the weakness of  
13 the high- $J$  transitions, even at room temperature, as well as the overlap with  
14 lines of more abundant isotopologues. The central frequency of each transition was  
15 measured by fitting the line profile to the expected Voigt profile, plus absorbing  
16 background. The fit residuals, both for CRD and SCAR measurements, are flat  
17 within noise, with no evidence of collisional narrowing [17]. Initial parameters for  
18 frequencies and intensities of the interfering transitions were obtained from the  
19 HITRAN database [16]. Unfortunately, the uncertainty of these data, if compared  
20 with the frequency precision and sensitivity of our measurements, and the presence  
21 of CO<sub>2</sub> transitions not listed by the database made difficult, and sometimes im-  
22 possible, to evaluate the absorbing background. This effect is more important for  
23 CRD spectra, and the P(22), P(16), R(0) and R(10) transitions are an example of  
24 that; in addition, background due to technical fringes was also considered. For the  
25 Voigt profile of the <sup>14</sup>C<sup>16</sup>O<sub>2</sub> transitions, Gaussian linewidth was always fixed to  
26 the required value by each temperature condition. The other line parameters, area,  
27 Lorentzian linewidth, and center frequency were left free. In Fig. 1, the agreement  
28 between experimental data and fit function (solid line trace) is evident in both  
29 CRD and SCAR experimental methodologies.  
30  
31

32 For each transition, several fit attempts were done with different absorbing back-  
33 ground parameters, either fixing or not the Lorentzian linewidth, in order to get the  
34 influence of such change in the final line center uncertainty. We have also checked  
35 for possible systematic errors due to use of two different recording methods (CRD  
36 and SCAR spectroscopy) by using the P(26) transition as a trial test. The good  
37 agreement between different recordings is evident in Fig. 1, where this line is plot-  
38 ted in the two cases. In particular, the P(26) measured by CRD is affected only  
39 by technical fringes and background offset (no interfering lines), while the SCAR  
40 measurement exhibits no background at all. It is worth noting that this is one of  
41 the main advantages intrinsic to the SCAR technique. The different spectral ar-  
42 eas are simply due to different temperature conditions. The results, including final  
43 standard errors are listed in Table 1. The line center uncertainty from the fit is  
44 several times lower (one order of magnitude for some transitions) than the final  
45 error reported in Table 1. However, we believe that the background parameters'  
46 uncertainties derived from the line profile fits were underestimated. We thought  
47 that a more realistic approach was to correct the fits uncertainties by individual  
48 factors, based on the goodness of the profile and the influence of contributions such  
49 as overlaps with other transitions and interference fringes.  
50  
51

52 Systematic corrections and uncertainties were considered in the transition fre-  
53 quencies reported in Table 1. Concerning the pressure shift, there are no literature  
54 data for transitions of this molecule, and we considered the contribution to be equal  
55 to that of the transitions of the  $\nu_3$  band of <sup>13</sup>C<sup>16</sup>O<sub>2</sub> [16]. In addition, we did not  
56 consider the  $J$  dependence of such shift, but we preferred to apply a more conserva-  
57 tive procedure. A mean pressure shift correction value was applied for all measured  
58 transitions, as a result of the average of the tabulated values for the correspond-  
59  
60

ing transitions of  $^{13}\text{C}^{16}\text{O}_2$ . The uncertainties of these corrections are given by the standard deviation of such averages and it was added to the estimated line center fit error. As a result, the CRD measurements were blue-shifted by 0.9 MHz and their uncertainty increased by 0.5 MHz, while the SCAR measurements were blue-shifted by 1.3 MHz and their uncertainty increased by 0.7 MHz. Finally, frequency calibration was not required, thanks to the primary Cs-standard traceability of our measurements. Also, the uncertainty due to the OFCS accuracy was not considered because, even in the worst case (67 Hz for R(30) transition), it is negligible with respect to other contributions.

This  $\nu_3$  fundamental infrared band of  $\text{CO}_2$  is a  $\Sigma - \Sigma$  parallel band with the usual P- and R-branch structure and lacking a Q-branch. The data were analysed using the basic Hamiltonian of a linear molecule, with centrifugal distortion corrections up to the quartic term for the ground state and to the octic term for the upper vibrational state. The ro-vibrational energy term values for non-degenerate vibrational modes of a linear molecule are given by

$$T(v, J) = G_v + B_v J(J+1) - D_v [J(J+1)]^2 + H_v [J(J+1)]^3 + L_v [J(J+1)]^4 \quad (2)$$

The band centre is defined as

$$\nu_0 = G_{v'} - G_{v''} \quad (3)$$

where  $G_{v'}$  and  $G_{v''}$  represent the purely vibrational energy of the upper and lower state, respectively. The parameters for the upper state  $v_3 = 1$ ,  $B_1$ ,  $D_1$ ,  $H_1$  and  $L_1$ , were held fixed at the very accurate values of Ref. [4] while the ground state constants  $B_0$  and  $D_0$  were fitted, together with the band centre  $\nu_0$ . The results of the fit are shown in Table 2. In the fitting procedure the lines were given a weight according to the uncertainties reported in Table 1. The tolerance for rejection was chosen line by line and fixed at twice the standard error associated to the lines. Only four lines were rejected from the fit.

The parameters obtained from the fit are well determined and are considerably improved in terms of both accuracy and precision with respect to the best constants available in literature [6]. The great precision obtained in this work for the band centre, together with the very precise constants of Ref. [4] allowed the determination of the vibrational energy of the levels of the Fermi dyad involved in the  $\text{CO}_2$  laser excitation mechanism. In Ref. [4], they reported  $\nu(001 - I) = 2.5965917618(13) \times 10^7$  MHz and  $\nu(001 - II) = 2.94600023911(49) \times 10^7$  MHz. We eventually derived  $\nu(I) = 4.076192962(51) \times 10^7$  MHz and  $\nu(II) = 3.726784485(51) \times 10^7$  MHz.

#### 4. Conclusions

The detection of the  $\nu_3$  fundamental band of  $^{14}\text{C}^{16}\text{O}_2$  has been performed with a OFCS-referenced cavity ring-down spectrometer. Several ro-vibrational transitions ranging from P(40) to R(30) have been detected and analyzed. The band origin  $\nu_0$  and the ground state parameters  $B_0$  and  $D_0$  have been determined with accuracy improved by about one order of magnitude. This experiment shows the potential of this technique in terms of accuracy and sensitivity and provides important information on a rare isotopologue of  $\text{CO}_2$ .



## Acknowledgements

This work was partially supported by Ente Cassa di Risparmio di Firenze and by Regione Toscana through the projects CTOTUS and SIMPAS, in the framework of POR–CReO FESR 2007–2013.

## References

- [1] M. Wahlen, R. S. Eng and K. W. Nill, *Appl. Opt.* **16**, 2350 (1977).
- [2] R. S. Eng, K. W. Nill and M. wahlen, *Appl. Opt.* **16**, 3072 (1977).
- [3] C. Freed, L. C. Bradley and R. G. O'Donnell, *IEEE J. Quantum Electron.* **16**, 1195 (1980).
- [4] L. C. Bradley, K. L. Soohoo and C. Freed, *IEEE J. Quantum Electron.* **22**, 234 (1986).
- [5] R. L. Sams and J. R. DeVoe, *J. Mol. Spectrosc.* **128**, 296 (1988).
- [6] S. Dobos, G. Winnewisser, F. Kling and J. Mink *Z. Naturforsch. A* **44**, 633 (1989).
- [7] C. Melazo Dias, K. Stenström, I. L. Bacelar Lêao, R. Ventura Santos, I. Gomes Nícoli, G. Skog, P. Ekström, R. da Silveira Corrêa and J. Environ. Radioact. **100**, 574 (2009).
- [8] D. Labrie and J. Reid, *Appl. Phys. A* **24**, 381 (1981).
- [9] B. M. Jakosky, R. C. Reedy and J. Masarik, *J. Geophys. Research* **101**, 2247 (1996).
- [10] G. Giusfredi, S. Bartalini, S. Borri, P. Cancio, I. Galli, D. Mazzotti and P. De Natale, *Phys. Rev. Lett.* **104**, 110801 (2010).
- [11] P. Cancio, S. Bartalini, S. Borri, I. Galli, G. Gagliardi, G. Giusfredi, P. Maddaloni, P. Malara, D. Mazzotti and P. De Natale, *Appl. Phys. B* **102**, 255 (2011)
- [12] I. Galli, S. Bartalini, S. Borri, P. Cancio, D. Mazzotti, P. De Natale and G. Giusfredi, *Molecular gas sensing down to parts per quadrillion: Radiocarbon-dioxide optical detection* (2011), submitted to *Phys. Rev. Lett.*
- [13] I. Galli, S. Bartalini, S. Borri, P. Cancio, G. Giusfredi, D. Mazzotti and P. De Natale, *Opt. Lett.* **35**, 3616 (2010).
- [14] I. Galli, S. Bartalini, P. Cancio, G. Giusfredi, D. Mazzotti and P. De Natale, *Opt. Express* **17**, 9582 (2009).
- [15] D. Mazzotti, P. Cancio, A. Castrillo, I. Galli, G. Giusfredi and P. De Natale, *J. Opt. A* **8**, S490 (2006).
- [16] L. S. Rothman, I. E. Gordon, A. Barbe, D. C. Benner, P. F. Bernath, M. Birk, V. Boudon, L. R. Brown, A. Campargue, J.-P. Champion, K. Chance, L. H. Coudert, V. Dana, V. M. Devi, S. Fally, J.-M. Flaud, R. R. Gamache, A. Goldman, D. Jacquemart, I. Kleiner, N. Lacome, W. J. Lafferty, J.-Y. Mandin, S. T. Massie, S. N. Mikhailenko, C. E. Miller, N. Moazzen-Ahmadi, O. V. Naumenko, A. V. Nikitin, J. Orphal, V. I. Perevalov, A. Perrin, A. Predoi-Cross, C. P. Rinsland, M. Rotger, M. Simeckova, M. A. H. Smith, K. Sung, S. A. Tashkun, J. Tennyson, R. A. Toth, A. C. Vandaele and J. Vander Auwera, *J. Quant. Spectrosc. Radiat. Transfer* **110**, 533 (2009).
- [17] P. L. Varghese and R. K. Hanson, *Appl. Opt.* **23**, 2376 (1984).

The  $\nu_3$  band of  $^{14}\text{C}^{16}\text{O}_2$  molecule

7

Table 1. Transitions frequencies, uncertainties and residuals from the fit of the lines measured for the  $\nu_3$  fundamental band of  $^{14}\text{C}^{16}\text{O}_2$ . All numbers are in MHz.

Transition	Obs. Freq.	Uncertainty <sup>a</sup>	Obs.-Calc. <sup>b</sup>
R(30)	67367614.181	4.795	1.759
R(26)	67294430.381	0.825	2.317*
R(22)	67218457.563	0.875	0.047
R(20)	67179428.939	0.743	-0.260
R(18)	67139706.163	0.731	-0.314
R(16)	67099289.625	1.017	-0.450
R(14)	67058180.523	0.877	-0.197
R(12)	67016379.675	0.828	0.529
R(8)	66930706.117	1.178	3.834*
R(6)	66886832.920	1.111	4.442*
R(4)	66842264.775	1.112	-0.645
R(2)	66797014.526	2.003	0.666
P(2)	66680873.788	2.765	-3.952
P(4)	66633216.481	3.176	-5.929
P(6)	66584878.904	1.132	-3.113*
P(8)	66535859.047	1.443	1.710
P(12)	66435758.242	0.996	-0.005
P(14)	66384685.331	2.458	-0.078
P(18)	66280498.427	0.801	1.310
P(20)	66227383.132	0.865	-0.128
P(24)	66119121.905	1.715	1.758
P(26)	66063972.706	0.755	0.194
P(28)	66008145.242	1.749	-3.339
P(32)	65894474.448	0.747	-0.680
P(34)	65836627.418	0.742	0.147
P(36)	65778106.638	0.729	0.186
P(38)	65718914.039	0.752	0.510
P(40)	65659047.695	1.795	-1.677

<sup>a</sup> See text for a discussion about the uncertainty of the lines.

<sup>b</sup> Lines marked with an asterisk exceeded the tolerance for rejection and were excluded from the fit.

Table 2. Spectroscopic parameters (in  $\text{cm}^{-1}$ ) obtained for the  $\nu_3$  fundamental band of  $^{14}\text{C}^{16}\text{O}_2$ .

Parameter	This work <sup>a</sup>	Ref. [6]
$\nu_0$	2225.801399(17)	2225.80239(16)
$B_0$	0.390253082(58)	0.39025488(18)
$D_0 \times 10^{-7}$	1.33144(37)	1.3372(20)
$B_1$	0.38739025628 <sup>b</sup>	
$D_1 \times 10^{-7}$	1.327523 <sup>b</sup>	
$H_1 \times 10^{-14}$	-3.08 <sup>b</sup>	
$L_1 \times 10^{-17}$	1.244 <sup>b</sup>	
$\sigma_{FIT} \times 10^{-5}$	3.1	

<sup>a</sup> The quantity in parentheses is one standard deviation in units of the last quoted digit.

<sup>b</sup> Fixed at the values of Ref. [4]

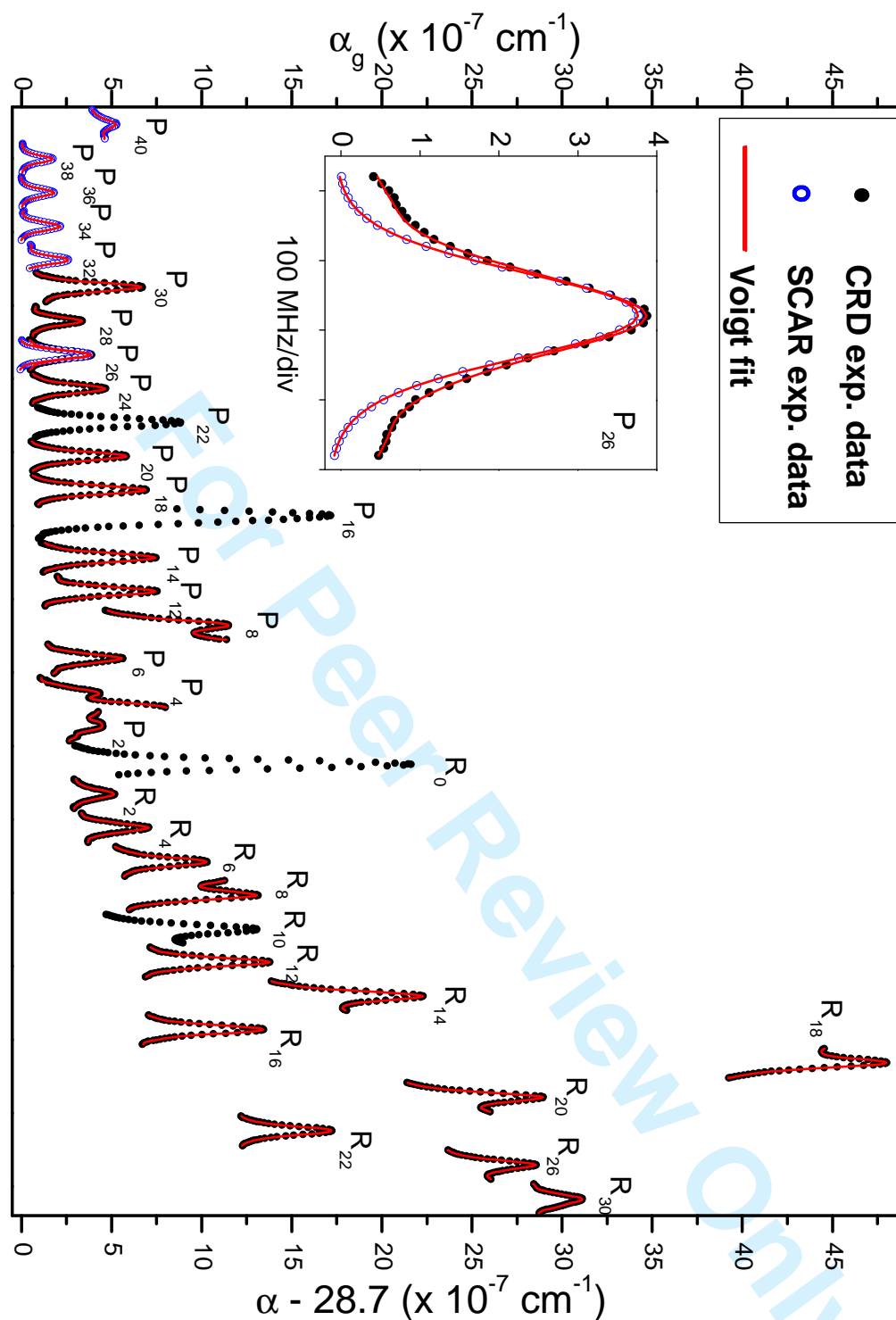
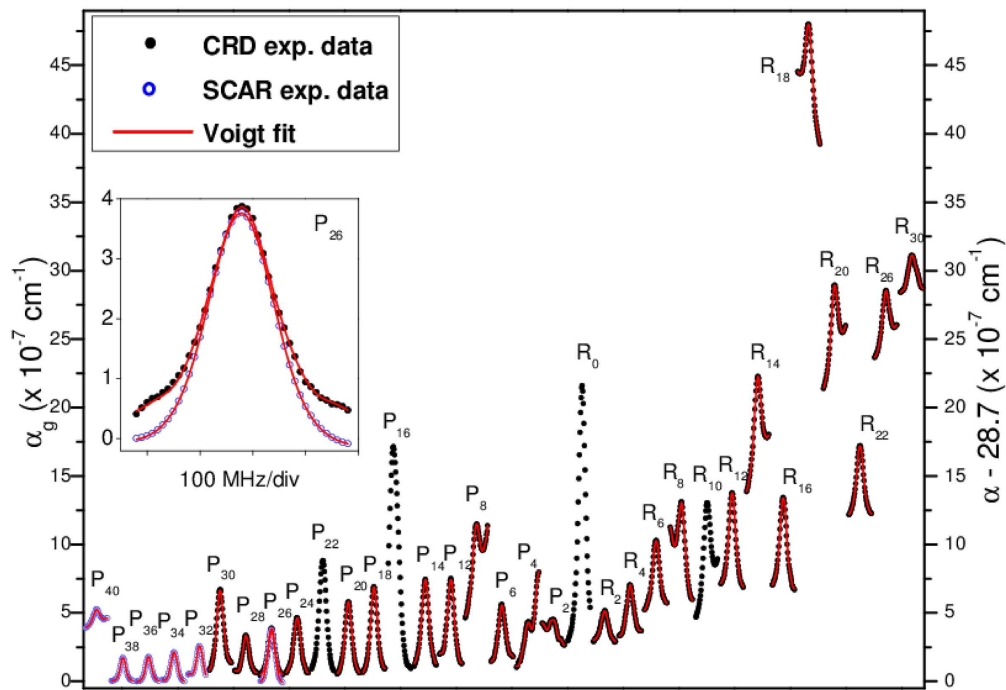


Figure 1. Overview of the infrared spectrum of  $^{14}\text{C}^{16}\text{O}_2$  molecule in the  $\nu_3$  band region. P(40)-P(32) lines were recorded by SCAR spectroscopy, while P(30)-R(30) lines were recorded by CRD spectroscopy. The fit of all measured transitions to a Voigt profile, that takes into account also the background from other  $\text{CO}_2$  lines, is also shown. The spectra are plotted in a relative horizontal frequency scale, and frequency gaps between transitions are not shown for the sake of clarity. The right side vertical scale applies to CRD spectra and the left side vertical scale to SCAR spectra. An empty cavity loss rate of  $\alpha_c = 2.87 \times 10^{-6} \text{ cm}^{-1}$  was removed from the measured rates in the case of CRD spectra to get a comparable scale with the SCAR ones. The  $\text{CO}_2$  pressure and temperature are 18 mbar and 296 K for the SCAR spectra, 12 mbar and 195 K for the CRD ones. The acquisition conditions and fit procedure are described in the text.



201x141mm (300 x 300 DPI)

view Only

