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Existence of global strong solution and vanishing capillarity-viscosity limit in one dimension for the Korteweg system

Frédéric Charve*, Boris Haspot †‡

Abstract

In the first part of this paper, we prove the existence of global strong solution for Korteweg system in one dimension. In the second part, motivated by the processes of vanishing capillarity-viscosity limit in order to select the physically relevant solutions for a hyperbolic system, we show that the global strong solution of the Korteweg system converges in the case of a γ law for the pressure ($P(\rho) = a\rho^\gamma$, $\gamma > 1$) to entropic solution of the compressible Euler equations. In particular it justifies that the Korteweg system is suitable for selecting the physical solutions in the case where the Euler system is strictly hyperbolic. The problem remains open for a Van der Waals pressure because in this case the system is not strictly hyperbolic and in particular the classical theory of Lax and Glimm (see [21, 11]) can not be used.

1 Introduction

We are concerned with compressible fluids endowed with internal capillarity. The model we consider originates from the XIXth century work by Van der Waals and Korteweg [38, 22] and was actually derived in its modern form in the 1980s using the second gradient theory, see for instance [9, 20, 37]. The first investigations begin with the Young-Laplace theory which claims that the phases are separated by a hypersurface and that the jump in the pressure across the hypersurface is proportional to the curvature of the hypersurface. The main difficulty consists in describing the location and the movement of the interfaces. Another major problem is to understand whether the interface behaves as a discontinuity in the state space (sharp interface) or whether the phase boundary corresponds to a more regular transition (diffuse interface, DI). The diffuse interface models have the advantage to consider only one set of equations in a single spatial domain (the density takes into account the different phases) which considerably simplifies the mathematical and numerical study (indeed in the case of sharp interfaces, we have to treat a problem with free boundary).

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Let us consider a fluid of density $\rho \geq 0$, velocity field $u \in \mathbb{R}$, we are now interested in the following compressible capillary fluid model, which can be derived from a Cahn-Hilliard like free energy (see the pioneering work by J.- E. Dunn and J. Serrin in [9] and also in [1, 3, 12, 17]). The conservation of mass and of momentum write:

$$\begin{cases} \frac{\partial}{\partial t} \rho^\epsilon + \partial_x(\rho^\epsilon u^\epsilon) = 0, \\ \frac{\partial}{\partial t}(\rho^\epsilon u^\epsilon) + \partial_x(\rho^\epsilon (u^\epsilon)^2) - \epsilon \partial_x(\rho^\epsilon \partial_x u^\epsilon) + \partial_x(a(\rho^\epsilon)^\gamma) = \epsilon^2 \partial_x K, \end{cases} \quad (1.1) \quad \boxed{\text{3systeme}}$$

where the Korteweg tensor reads as following:

$$\operatorname{div} K = \partial_x(\rho^\epsilon \kappa(\rho^\epsilon) \partial_{xx} \rho^\epsilon + \frac{1}{2}(\kappa(\rho^\epsilon) + \rho^\epsilon \kappa'(\rho^\epsilon)) |\partial_x \rho^\epsilon|^2) - \partial_x(\kappa(\rho^\epsilon) (\partial_x \rho^\epsilon)^2). \quad (1.2) \quad \boxed{\operatorname{div} K}$$

κ is the coefficient of capillarity and is a regular function of the form $\kappa(\rho) = \epsilon^2 \rho^\alpha$ with $\alpha \in \mathbb{R}$. In the sequel we shall assume that $\kappa(\rho) = \frac{\epsilon^2}{\rho}$. The term $\partial_x K$ allows to describe the variation of density at the interfaces between two phases, generally a mixture liquid-vapor. $P = a\rho^\gamma$ with $\gamma \geq 1$ is a general γ law pressure term. ϵ corresponds to the controlling parameter on the amplitude of the viscosity and of the capillarity. When we set $v^\epsilon = u^\epsilon + \epsilon \partial_x(\ln \rho^\epsilon)$, we can write (1.1) on the following form (we refer to [13] for the computations):

$$\begin{cases} \frac{\partial}{\partial t} \rho^\epsilon + \partial_x(\rho^\epsilon v^\epsilon) - \epsilon \partial_{xx} \rho^\epsilon = 0, \\ \frac{\partial}{\partial t}(\rho^\epsilon v^\epsilon) + \partial_x(\rho^\epsilon (v^\epsilon)^2) - \epsilon \partial_x(\rho^\epsilon \partial_x v^\epsilon) + \partial_x(a(\rho^\epsilon)^\gamma) = 0, \end{cases} \quad (1.3) \quad \boxed{1.1}$$

We now consider the Cauchy problem of (1.3) when the fluid is away from vacuum. Namely, we shall study (1.3) with the following initial data:

$$\rho^\epsilon(0, x) = \rho_0^\epsilon(x) > 0, \quad u^\epsilon(0, x) = u_0^\epsilon(x), \quad (1.4) \quad \boxed{1.2}$$

such that:

$$\lim_{x \rightarrow +, -\infty} (\rho_0^\epsilon(x), u_0^\epsilon(x)) = (\rho^{+, -}, u^{+, -}), \quad \text{with } \rho^{+, -} > 0.$$

We would like to study in the sequel the limit process of system (1.3) when ϵ goes to 0 and to prove in particular that we obtain entropic solutions of the Euler system:

$$\begin{cases} \frac{\partial}{\partial t} \rho + \partial_x(\rho v) = 0, \\ \frac{\partial}{\partial t}(\rho v) + \partial_x(\rho v^2) + \partial_x(a\rho^\gamma) = 0, \end{cases} \quad (1.5) \quad \boxed{1.3}$$

Let us now explain the interest of the capillary solutions for the hyperbolic systems of conservation laws.

1.1 Viscosity capillarity processes of selection for the Euler system

In addition of modeling a liquid-vapour mixture, the Korteweg also shows purely theoretical interests consisting in the selection of the physically relevant solutions of the

Euler model (in particular when the system is not strictly hyperbolic). The typical case corresponds to a Van der Waals pressure: indeed in this case the system is not strictly hyperbolic in the elliptic region (which corresponds to the region where the phase change occurs).

In the adiabatic pressure framework ($P(\rho) = \rho^\gamma$ with $\gamma > 1$), the system is strictly hyperbolic and the theory is classical. More precisely we are able to solve the Riemann problem when the initial Heaviside data is small in the BV space. Indeed we are in the context of the well known Lax result as the system is also genuinely nonlinear (we refer to [Lax]). It means we have existence of global \mathcal{C}^1 -piecewise solutions which are unique in the class of the entropic solutions.

This result has been extended by Glimm in the context of small initial data in the BV-space by using a numerical scheme and approximating the initial BV data by a \mathcal{C}^1 -piecewise function (which implies to locally solve the Riemann problem via the Lax result). For the uniqueness of the solution we refer to the work of Bianchini and Bressan ([BB1]) who use a viscosity method.

In the setting of the Van der Waals pressure, the existence of global solutions and the nature of physical relevant solutions remain completely open. Indeed the system is not strictly hyperbolic anymore.

If we rewrite the compressible Euler system in Lagrangian coordinates by using the specific volume $\tau = 1/\rho$ in $(\frac{1}{b}, \infty)$ and the velocity u , the system satisfies in $(0, +\infty) \times \mathbb{R}$ the equations:

$$\begin{aligned} \partial_t \tau - \partial_x u &= 0, \\ \partial_t u - \partial_x (\tilde{P}(\tau)) &= 0, \end{aligned} \tag{1.6} \quad \boxed{\text{euler}}$$

with the function $\tilde{P} : (\frac{1}{b}, \infty) \rightarrow (0, \infty)$ given by:

$$\tilde{P}(\tau) = P\left(\frac{1}{\tau}\right), \quad \tau \in \left(\frac{1}{b}, \infty\right).$$

The two eigenvalues of the system are:

$$\lambda_1(\tau, v) = -\sqrt{-\tilde{P}'(\tau)}, \quad \lambda_2(\tau, v) = -\sqrt{-\tilde{P}'(\tau)}. \tag{1.7} \quad \boxed{\text{vp}}$$

The corresponding eigenvectors r_1, r_2 are:

$$w_1(\tau, v) = \begin{pmatrix} 1 \\ \sqrt{-\tilde{P}'(\tau)} \end{pmatrix}, \quad w_2(\tau, v) = \begin{pmatrix} 1 \\ -\sqrt{-\tilde{P}'(\tau)} \end{pmatrix} \tag{1.8}$$

Furthermore by calculus we obtain:

$$\nabla \lambda_1(\tau, v) \cdot w_1(\tau, v) = \frac{\tilde{P}''(\tau)}{2\sqrt{-\tilde{P}'(\tau)}}, \quad \nabla \lambda_2(\tau, v) \cdot w_2(\tau, v) = \frac{-\tilde{P}''(\tau)}{2\sqrt{-\tilde{P}'(\tau)}} \tag{1.9}$$

We now recall the definition of a *standard conservation law* in the sense of Lax (it means entropy solutions):

- The system is **strictly entropic** if the eigenvalues are distinct and real.

- The characteristics fields are **genuinely nonlinear** if we have for all (τ, v) ,

$$\nabla\lambda_1(\tau, v) \cdot w_1(\tau, v) \neq 0 \text{ and } \nabla\lambda_2(\tau, v) \cdot w_2(\tau, v) \neq 0,$$

for more details we refer to ^{Serre}[33]. The definition of genuine nonlinearity is some kind of extension of the notion of convexity to vector-valued functions (in particular when we consider the specific case of the traveling waves). The previous assumptions aim at ensuring the existence and the uniqueness of the Riemann problem (see ^{Evans}[10] and ^{Serre}[33]).

When P is a Van der Waals pressure, we observe that the first conservation law (^{Serre, Evans}[33, 10]) is far from being a standard hyperbolic system, indeed:

- It is not hyperbolic (but elliptic) in $(\frac{1}{\alpha_1}, \frac{1}{\alpha_2}) \times \mathbb{R}$,
- the characteristic fields are not genuinely nonlinear in the hyperbolic part of the state space.

Here the classical Lax-Glimm theory cannot be applied. In particular there doesn't exist any entropy-flux pair, which suggests that the entropy framework is not adapted for selecting the physically relevant solutions. In order to deal with this problem, Van der Waals and Korteweg began by considering the stationary problem with null velocity, and solving $\nabla P(\rho) = 0$. For more details we refer to ^{Rohde}[31]. It consists in minimizing in the following admissible set

$$A_0 = \{\rho \in L^1(\Omega) / W(\rho) \in L^1(\Omega), \int_{\Omega} \rho(x) dx = m\},$$

the following functional

$$F[\rho] = \int_{\Omega} W(\rho(x)) dx.$$

Unfortunately this minimization problem has an infinity of solutions, and many of them are physically irrelevant. In order to overcome this difficulty, Van der Waals in the XIX-th century was the first to regularize the previous functional by adding a quadratic term in the density gradient. More precisely he considered the following functional:

$$F_{local}^{\epsilon} = \int_{\Omega} (W(\rho^{\epsilon}(x)) + \gamma \frac{\epsilon^2}{2} |\nabla \rho^{\epsilon}|^2) dx,$$

with:

$$A_{local} = H^1(\Omega) \cap A_0.$$

This variational problem has a unique solution and its limit (as ϵ goes to zero) converge to a physical solution of the equilibrium ^{REF}problem for the Euler system with Van der Waals pressure, that was proved by Modica in ^{REF}[28] with the use of gamma-convergence.

By the Euler-Lagrange principle, the minimization of the Van der Waals functional consists in solving the following stationary problem:

$$\nabla P(\rho^{\epsilon}) = \gamma \epsilon^2 \rho^{\epsilon} \nabla \Delta \rho^{\epsilon},$$

where the right-hand side can be expressed as the divergence of the capillarity tensor.

Heuristically, we also hope that the process of vanishing capillarity-viscosity limit selects the physical relevant solutions as it does for the stationary system. This problem actually remains open.

1.2 Existence of global entropic solutions for Euler system

Before presenting the results of this paper let us recall the results on this topic in these last decades. We shall focus on the case of a γ pressure law $P(\rho) = a\rho^\gamma$ with $\gamma > 1$ and a positive. Let us mention that these cases are the only ones well-known (essentially because the system is strictly hyperbolic in this case and that we can exhibit many entropy-flux pairs). Here the Lax-Glimm theory can be applied, however at the end of the 70's, one was interested in relaxing the conditions on the initial data by only assuming ρ_0 and u_0 in L^∞ .

In the beginning of the 80's Di Perna initiated this program, consisting in obtaining global entropic solutions for L^∞ initial data.

Indeed in [7, 8], Di Perna prove the existence of global weak entropy solution of (1.3) for $\gamma = 1 + \frac{2}{2d+1}$ and $\gamma = 2k + \frac{3}{2k} + 1$ (with $k \geq 1$), $d \geq 2$ by using the so-called "compensated compactness" introduced by Tartar in [35]. This result was extended by Chen in [4] in the case $\gamma \in (1, \frac{5}{3}]$ and by Lions et al in [26] in the case $\gamma \in [3, \infty)$. In [25], Lions et al generalize this result to the general case $\gamma \in (1, 3)$, and finally the case $\gamma = 1$ is treated by [18]. We would like to mention that these results are obtained through a vanishing artificial viscosity on both density and velocity.

The problem of vanishing physical viscosity limit of compressible Navier-Stokes equations to compressible Euler equations was until recently an open problem. However Chen and Perepelista in [5] proved that the solutions of the compressible Navier-Stokes system with constant viscosity coefficients converge to a entropic solution of the Euler system with finite energy. This result was extended in [19] to the case of viscosity coefficients depending on the density.

Inspired by [5] and [19], we would like to show that the solution of the Korteweg system (1.3) converges to a entropic solution of the Euler system with finite energy when the pressure is a γ law. To do this, we will prove for the first time up our knowledge the existence of global strong solution for the Korteweg system in one dimension in the case of Saint-Venant viscosity coefficients. By contrast, the problem of global strong solutions for compressible Navier-Stokes equations remains open (indeed one of the main difficulties consists in controlling the vacuum). This result justifies that the Korteweg system allows us to select the relevant physical solutions of the compressible Euler system at least when the pressure is adiabatic ($P(\rho) = a\rho^\gamma$ with $\gamma > 1$). The problem remains open in the case of a Van der Waals pressure.

1.3 Results

Let us now describe our main result. In the first theorem we prove the existence of global strong solution for the Korteweg system (1.3).

Theorem 1.1 *Let $\bar{\rho} > 0$. Assume that the initial data ρ_0 and u_0 satisfy:*

$$0 < m_0 \leq \rho_0 \leq M_0 < +\infty, \rho_0 - \bar{\rho} \in H^1(\mathbb{R}), v_0 \in H^1(\mathbb{R}) \cap L^\infty(\mathbb{R}). \quad (1.10) \quad \boxed{2.5}$$

Then there exists a global strong solution (ρ, v) of (1.3) on $\mathbb{R}^+ \times \mathbb{R}$ such that for every $T > 0$:

$$\begin{aligned} \rho - \bar{\rho} &\in L^\infty(0, T, H^1(\mathbb{R})), \rho \in L^\infty(0, T, L^\infty(\mathbb{R})), \\ v &\in L^\infty(0, T, H^1(\mathbb{R})) \cap L^2(0, T, H^2(\mathbb{R})) \quad \text{and} \quad v \in L^\infty(0, T, L^\infty(\mathbb{R})). \end{aligned}$$

Finally this solution is unique in the class of weak solutions satisfying the usual energy inequality.

theo

Remark 1 We would like to point out that the problem remains open in the case of the Saint-Venant system, which corresponds to system (1.3) without capillarity.

In the following theorem, we are interested in proving the convergence of the global solutions of system (1.3) to entropic solutions of the Euler system (1.5).

Theorem 1.2 Let $\gamma > \frac{5}{3}$ and $(\rho^\epsilon, v^\epsilon)$ with $m_\epsilon = \rho^\epsilon v^\epsilon$ be the global solution of the Cauchy problem (1.3) with initial data $(\rho_0^\epsilon, v_0^\epsilon)$ as in theorem 1.1. Then, when $\epsilon \rightarrow 0$, there exists a subsequence of $(\rho^\epsilon, m^\epsilon)$ that converge almost everywhere to a finite entropy solution $(\rho, \rho v)$ to the Cauchy problem (1.5) with initial data $(\rho_0, \rho_0 v_0)$.

theo1

Remark 1 We would like to point out that Lions et al in [25] had obtained the existence of global entropic solution for $\gamma > 1$ by a viscosity vanishing process, and the considered regularizing system was exactly the Korteweg system modulo the introduction of the effective velocity.

One important basis of our problem for theorem 1.2 is the following compactness theorem established in [5].

Theorem 1.3 (Chen-Perepelitsa [5]) Let $\psi \in C_0^2(\mathbb{R})$, (η^ψ, q^ψ) be a weak entropy pair generated by ψ . Assume that the sequences $(\rho^\epsilon(x, t), v^\epsilon(x, t))$ defined on $\mathbb{R} \times \mathbb{R}_+$ with $m^\epsilon = \rho^\epsilon v^\epsilon$, satisfies the following conditions:

1. For any $-\infty < a < b < +\infty$ and all $t > 0$, it holds that:

$$\int_0^t \int_a^b (\rho^\epsilon)^{\gamma+1} dx d\tau \leq C(t, a, b), \quad (1.11) \quad 1.8$$

where $C(t) > 0$ is independent of ϵ .

2. For any compact set $K \subset \mathbb{R}$, it holds that

$$\int_0^t \int_K ((\rho^\epsilon)^{\gamma+\theta} + \rho^\epsilon |v^\epsilon|^3) dx d\tau \leq C(t, K), \quad (1.12) \quad 1.9$$

where $C(t, K) > 0$ is independent of ϵ .

3. The sequence of entropy dissipation measures

$$\eta^\psi(\rho^\epsilon, m^\epsilon)_t + q^\psi(\rho^\epsilon, m^\epsilon)_x \text{ are compact in } H_{loc}^{-1}(\mathbb{R}_+^2). \quad (1.13) \quad 1.10$$

Then there is a subsequence of $(\rho^\epsilon, m^\epsilon)$ (still denoted $(\rho^\epsilon, m^\epsilon)$) and a pair of measurable functions (ρ, m) such that:

$$(\rho^\epsilon, m^\epsilon) \rightarrow (\rho, m), \text{ a.e as } \epsilon \rightarrow 0. \quad (1.14) \quad 1.11$$

theo2

Remark 2 We would like to recall that the estimate (1.12) was first derived by Lions et al in [26] by relying the moment lemma introduced by Perthame in [30].

The paper is arranged as follows. In section 2 we recall some important results on the notion of entropy-flux pair for Euler system and on the kinetic formulation of Lions et al in [26]. In section 3, we show theorem 1.1 and in the last section 4.1 we prove theorem 1.2.

2 Mathematical tools

Definition 2.1 A pair of functions $(\eta(\rho, v), H(\rho, v))$ or $(\eta(\rho, m), q(\rho, m))$ for $m = \rho v$, is called an entropy-entropy flux pair of system (1.3), if the following holds:

$$[\eta(\rho, v)]_t + [H(\rho, v)]_x = 0,$$

for any smooth solution of (1.3). Furthermore $(\eta(\rho, v))$ is called a weak entropy if:

$$\eta(0, u) = 0, \quad \text{for any fixed } v.$$

Definition 2.2 An entropy $\eta(\rho, m)$ is convex if the Hessian $\nabla^2 \eta(\rho, m)$ is nonnegative definite in the region under consideration.

Such η satisfy the wave equation:

$$\partial_{tt}\eta = \theta^2 \rho^{\gamma-3} \partial_{xx}\eta.$$

From [26], we obtain an explicit representation of any weak entropy (η, q) under the following form:

$$\begin{aligned} \eta^\psi(\rho, m) &= \int_{\mathbb{R}} \chi(\rho, s - v) \psi(s) ds, \\ H^\psi(\rho, m) &= \int_{\mathbb{R}} (\theta s + (1 - \theta)u) \chi(\rho, s - v) \psi(s) ds, \end{aligned} \quad (2.15) \quad \boxed{1.4}$$

where the kernel χ is defined as follows:

$$\chi(\rho, v) = [\rho^{2\theta} - v^2]_+^\lambda, \quad \lambda = \frac{3 - \gamma}{2(\gamma - 1)} > -\frac{1}{2}, \quad \text{and } \theta = \frac{\gamma - 1}{2},$$

and here:

$$\begin{aligned} t_+^\lambda &= t^\lambda \quad \text{for } t > 0, \\ &= 0 \quad \text{for } t \leq 0, \end{aligned}$$

Proposition 2.1 (see [26])

For instance, when $\psi(s) = \frac{1}{2}s^2$, the entropy pair is the mechanical energy and the associated flux:

$$\eta^*(\rho, m) = \frac{m^2}{2\rho} + e(\rho), \quad q^*(\rho, m) = \frac{m^3}{2\rho^2} + e'(\rho), \quad (2.16) \quad \boxed{1.5}$$

where $e(\rho) = \frac{\kappa}{\gamma-1} \rho^\gamma$ represents the gas internal energy in physics.

In the sequel we will work far away of the vacuum that it why we shall introduce equilibrium states such that we avoid the vacuum. Let $(\bar{\rho}(x), \bar{v}(x))$ be a pair of smooth monotone functions satisfying $(\bar{\rho}(x), \bar{v}(x)) = (\rho^{-, +}, v^{-, +})$ when $- + x \geq L_0$ for some large $L_0 > 0$. The total mechanical energy for (1.3) in \mathbb{R} with respect to the pair of reference function $(\bar{\rho}(x), \bar{v}(x))$ is:

$$E[\rho, v](t) = \int_{\mathbb{R}} \left(\frac{1}{2} \rho(t, x) |v(t, x) - \bar{v}(x)|^2 + e^*(\rho(t, x), \bar{\rho}(x)) \right) dx \quad (2.17) \quad \boxed{1.7}$$

where $e^*(\rho, \bar{\rho}) = e(\rho)e(\bar{\rho}) - e'(\bar{\rho}(\rho - \bar{\rho})) \geq 0$. The total mechanical energy for system (1.7) with $\kappa(\rho) = \frac{\kappa}{\rho}$ is:

$$E_1[\rho, u](t) = \int_{\mathbb{R}} \left(\frac{1}{2} \rho(t, x) |u(t, x) - \bar{u}(x)|^2 + e^*(\rho(t, x), \bar{\rho}(x)) + \epsilon^2 (\partial_x \rho^{\frac{1}{2}})^2 \right) dx \quad (2.18) \quad \boxed{1.7}$$

and the total mechanical energy for system (1.3) is:

$$E_2[\rho, v](t) = \int_{\mathbb{R}} \left(\frac{1}{2} \rho(t, x) |v(t, x) - \bar{v}(x)|^2 + e^*(\rho(t, x), \bar{\rho}(x)) \right) dx \quad (2.19) \quad \boxed{1.7}$$

Definition 2.3 Let (ρ_0, v_0) be given initial data with finite-energy with respect to the end states: (ρ^\pm, v^\pm) at infinity, and $E[\rho_0, v_0] \leq E_0 < +\infty$. A pair of measurable functions $(\rho, u) : \mathbb{R}_+^2 \rightarrow \mathbb{R}_+^2$ is called a finite-energy entropy solution of the Cauchy problem (1.3) if the following properties hold:

1. The total energy is bounded in time such that there exists a bounded function $C(E, t)$, defined on $\mathbb{R}^+ \times \mathbb{R}^+$ and continuous in t for each $E \in \mathbb{R}^+$ with for a.e $t > 0$:

$$E[\rho, v](t) \leq C(E_0, t).$$

2. The entropy inequality:

$$\eta^\psi(\rho, v)_t + q^\psi(\rho, v)_x \leq 0,$$

is satisfied in the sense of distributions for all test functions $\psi(s) \in \{\pm 1, \pm s, s^2\}$.

3. The initial data (ρ_0, v_0) are obtained in the sense of distributions.

We now give our main conditions on the initial data (1.4), which is inspired from [10].

Definition 2.4 Let $(\bar{\rho}(x), \bar{v}(x))$ be some pair of smooth monotone functions satisfying $(\bar{\rho}(x), \bar{v}(x)) = (\rho^{-, +}, v^{-, +})$ when $-x \geq L_0$ for some large $L_0 > 0$. For positive constant C_0, C_1 and C_2 independent of ϵ , we say that the initial data $(\rho_0^\epsilon, v_0^\epsilon)$ satisfy the condition \mathcal{H} if they verify the following properties:

- $\rho_0^\epsilon > 0$, $\int_{\mathbb{R}} \rho_0^\epsilon(x) |u_0^\epsilon(x) - \bar{u}(x)| \leq C_0 < +\infty$,
- The energy is finite:

$$\int_{\mathbb{R}} \left(\frac{1}{2} \rho_0^\epsilon(x) |v_0^\epsilon(x) - \bar{v}(x)|^2 + e^*(\rho_0^\epsilon(x), \bar{\rho}(x)) \right) dx \leq C_1 < +\infty,$$

•

$$\epsilon^2 \int_{\mathbb{R}} \frac{|\partial_x \rho_0^\epsilon(x)|^2}{\rho_0^\epsilon(x)^{3-2\alpha}} dx \leq C_2 < +\infty.$$

In this section, we would like to recall some properties on the pair of entropy for the system (1.3). Smooth solutions of (1.3) satisfy the conservation laws:

$$\partial_t \eta(\rho, u) + \partial_x H(\rho, u) = 0,$$

if and only if:

$$\eta_{\rho\rho} = \frac{P'(\rho)}{\rho^2} \eta_{uu}. \quad (2.20) \quad \boxed{\text{ondes}}$$

We supplement the equation $\overline{2.20}$ by giving initial conditions:

$$\eta(0, u) = 0, \eta_\rho(0, u) = \psi(u). \quad (2.21) \quad \boxed{\text{initial}}$$

We are now going to give a sequel of proposition on the properties of η , we refer to $\overline{[26]}$ for more details.

$\boxed{\text{proutile}}$

Proposition 2.2 For $\rho \geq 0, u, \omega \in \mathbb{R}$,

- The fundamental solution of $\overline{(2.20)}$ - $\overline{(2.21)}$ is the solution corresponding to $\eta_\rho(0, u) = \delta(u)$ is given by:

$$\chi(\rho, \omega) = (\rho^{\gamma-1} - \omega^2)_+^\lambda \quad \text{with } \lambda = \frac{3-\gamma}{2(\gamma-1)}. \quad (2.22)$$

- The solution of $\overline{(2.20)}$ - $\overline{(2.21)}$ is given by:

$$\eta(\rho, u) = \int_{\mathbb{R}} \psi(\xi) \chi(\rho, \xi - u) d\xi, \quad (2.23) \quad \boxed{\text{entropie}}$$

- η is convex in $(\rho, \rho u)$ for all ρ, u if and only if g is convex.
- The entropy flux H associated with η is given by:

$$H(\rho, u) = \int_{\mathbb{R}} \psi(\xi) [\theta \xi + (1-\theta)\xi] \chi(\rho, \xi - u) d\xi \quad \text{where } \theta = \frac{\gamma-1}{2}. \quad (2.24) \quad \boxed{\text{flux}}$$

We now give a important result on the entropy pair (see $\overline{[26]}$, lemma 4) .

$\boxed{\text{pair35}}$

Proposition 2.3 Taking $\psi(s) = \frac{1}{2}s|s|$, then there exists a positive constant $C > 0$, depending only on $\gamma > 1$, such that the entropy pair (η^ψ, H^ψ) satisfies:

$$\begin{aligned} |\eta^\psi(\rho, u)| &\leq (\rho|u|^2 + \rho^\gamma), \\ H^\psi(\rho, u) &\geq C^{-1}(\rho|u|^3 + \rho^{\gamma+\theta}), \quad \text{for all } \rho \geq 0 \text{ and } u \in \mathbb{R}, \\ |\eta_m^\psi(\rho, u)| &\leq (\rho|u| + \rho^\theta), \\ |\eta_{mm}^\psi(\rho, u)| &\leq C\rho^{-1}. \end{aligned} \quad (2.25) \quad \boxed{2.37}$$

We are now going to give recent results on the entropy pair (η^ψ, q^ψ) generated by $\psi \in C_0^2(\mathbb{R})$ (we refer to $\overline{[5]}$ for more details).

$\boxed{\text{propChen}}$

Proposition 2.4 For a C^2 function $\psi : \mathbb{R} \rightarrow \mathbb{R}$, compactly supported on the interval $[a, b]$, we have:

$$\text{supp}(\eta^\psi), \text{supp}(q^\psi) \subset \{(\rho, m) = (\rho, \rho u) : u + \rho^\theta \geq a, u - \rho^\theta \leq b\} : \quad (2.26) \quad \boxed{3.2}$$

Furthermore, there exists a constant C_ψ such that, for any $\rho \geq 0$ and $u \in \mathbb{R}$, we have:

- For $\gamma \in (1, 3]$,

$$|\eta^\psi(\rho, m)| + |q^\psi(\rho, m)| \leq C_\psi \rho. \quad (2.27) \quad \boxed{3.3}$$

- For $\gamma \in (3, +\infty)$,

$$|\eta^\psi(\rho, m)| \leq C_\psi \rho, \quad |q^\psi(\rho, m)| \leq C_\psi(\rho + \rho^{\theta+1}). \quad (2.28) \quad \boxed{3.4}$$

- If η^ψ is considered as a function of (ρ, m) , $m = \rho u$ then

$$|\eta_m^\psi(\rho, m)| + |\rho \eta_{mm}^\psi(\rho, m)| \leq C_\psi, \quad (2.29) \quad \boxed{3.5}$$

and, if η_m^ψ is considered as a function of (ρ, u) , then

$$|\eta_m^\psi(m, u)| + |\rho^{1-\theta} \eta_{m\rho}^\psi(\rho, \rho u)| \leq C_\psi. \quad (2.30) \quad \boxed{3.6}$$

We now would like to express the kinetic formulation of [\(1.5\)](#) introduced in [\[26\]](#).

Theorem 2.4 *Let $(\rho, \rho v) \in L^\infty(\mathbb{R}^+, L^1(\mathbb{R}))$ have finite energy and $\rho \geq 0$, then it is an entropy solution of [\(1.5\)](#) if and only if there exists a non-positive bounded measure m on $\mathbb{R}^+ \times \mathbb{R}^2$ such that the function $\chi(\rho, \xi - u)$ satisfies:*

$$\partial_t \chi + \partial_x [(\theta \xi + (1 - \theta)u)\chi] = \partial_{\xi\xi} m(t, x, \xi). \quad (2.31)$$

cinetique

3 Proof of theorem [1.1](#)^{theo}

section3

We would like to start with recalling an important result due to Solonnikov (see [\[34\]](#)). Let ρ_0 the initial density such that:

$$0 < m_0 \leq \rho_0 \leq M_0 < +\infty. \quad (3.32) \quad \boxed{\text{initiald}}$$

When the viscosity coefficient $\mu(\rho)$ satisfies:

$$\mu(\rho) \geq c > 0 \text{ for all } \rho \geq 0, \quad (3.33) \quad \boxed{\text{visco}}$$

we have the existence of strong solution for small time. More exactly, we have:

Proposition 3.5 *Let (ρ_0, v_0) satisfy [\(3.32\)](#)^{initiald} and assume that μ satisfies [\(3.33\)](#)^{visco}, then there exists $T_0 > 0$ depending on $m_0, M_0, \|\rho_0 - \bar{\rho}\|_{H^1}$ and $\|v_0\|_{H^1}$ such that [\(1.3\)](#)^{1.1} has a unique solution (ρ, v) on $(0, T_0)$ satisfying:*

$$\begin{aligned} \rho - \bar{\rho} &\in L^\infty(H^1(\mathbb{R})), \quad \partial_t \rho \in L^2((0, T_1) \times \mathbb{R}), \\ v &\in L^2(0, T_1, H^2(\mathbb{R})), \quad \partial_t v \in L^2((0, T_1) \times \mathbb{R}) \end{aligned}$$

for all $T_1 < T_0$.

Remark 3 *The main point in this theorem is that the time of existence T_0 depends only of the norms of ρ_0 which gives us a low bounds on T_0 of the system [\(1.3\)](#)^{1.1}.*

In view of this proposition, we see that if we introduce a truncated viscosity coefficient $\mu_n(\rho)$:

$$\mu_n(\rho) = \max(\rho, \frac{1}{n}),$$

then there exists approximated solutions (ρ_n, v_n) defined for small time $(0, T_0)$ of the system (1.3). In order to prove theorem 1.1, we only have to show that (ρ_n, v_n) satisfies the following bounds uniformly with respect to n and T large:

$$\begin{aligned} 0 < m_0 \leq \rho_n \leq M_0 < +\infty, \quad \forall t \in [0, T], \\ \rho_n - \bar{\rho} &\in L_T^\infty(H^1(\mathbb{R})), \\ v_n &\in L_T^\infty(H^1(\mathbb{R})). \end{aligned} \tag{3.34}$$

We are going to follow the method of Lions et al in [25], indeed the main point is to prove that we can extend the notion of Riemann invariant or more precisely the kinetic formulation of proposition 2.4 to the system (1.3). We recall that system (1.3) has the following form:

$$\begin{cases} \frac{\partial}{\partial t} \rho_n + \partial_x(\rho_n v_n) - \epsilon \partial_{xx} \rho_n = 0, \\ \frac{\partial}{\partial t}(\rho_n v_n) + \partial_x(\rho_n v_n v_n) - \epsilon \partial_x(\partial_x \rho_n v_n) - \epsilon \partial_x(\rho_n \partial_x v_n) + \partial_x(a(\rho_n)^\gamma) = 0, \end{cases} \tag{3.35} \quad \boxed{1.1a}$$

and we have finally:

$$\begin{cases} \frac{\partial}{\partial t} \rho_n + \partial_x(\rho_n v_n) - \epsilon \partial_{xx} \rho_n = 0, \\ \frac{\partial}{\partial t}(\rho_n v_n) + \partial_x(\rho_n v_n v_n) - \epsilon \partial_x \partial_x(\rho_n v_n) + \partial_x(a(\rho_n)^\gamma) = 0, \end{cases} \tag{3.36} \quad \boxed{1.1b}$$

Following [25] and setting $m_n = \rho_n v_n$ we have for any pair of entropy flux $(\eta(\rho, u), H(\rho, u))$ defined by (2.23) and (2.24) where η is a convex function of (ρ_n, m_n) . We write $\eta = \bar{\eta}(\rho_n, m_n)$:

$$\begin{aligned} \partial_t \eta + \partial_x H &= \epsilon \bar{\eta}_\rho \partial_{xx} \rho_n + \epsilon \bar{\eta}_m \partial_{xx} m_n, \\ &= \epsilon \partial_{xx} \eta - \epsilon (\bar{\eta}_{\rho\rho} (\partial_x \rho_n)^2 + 2 \bar{\eta}_{\rho m} (\partial_x \rho_n) (\partial_x m_n) + \bar{\eta}_{mm} (\partial_x m_n)^2). \end{aligned}$$

Here we define μ_n such that:

$$\mu_n = \bar{\eta}_{\rho\rho} (\partial_x \rho_n)^2 + 2 \bar{\eta}_{\rho m} (\partial_x \rho_n) (\partial_x m_n) + \bar{\eta}_{mm} (\partial_x m_n)^2$$

By proposition 2.2, we can check that $\mu_n \geq 0$. We obtain then that:

$$\partial_t \eta(\rho_n, v_n) + \partial_x H((\rho_n, v_n)) - \epsilon \bar{\eta}_\rho \partial_{xx} \rho_n \leq 0 \quad \text{in } \mathbb{R} \times (0, +\infty).$$

By applying the same method than for proving the theorem 2.4, we obtain the following kinetic formulation:

$$\partial_t \chi + \partial_x([\theta \xi + (1 - \theta)v_n] \chi) - \partial_{xx} \chi = \partial_{\xi\xi} \bar{m}_n \quad \text{on } \mathbb{R}^2 \times (0, +\infty), \tag{3.37} \quad \boxed{\text{riemann}}$$

where \bar{m}_n is a nonpositive bounded measure on $\mathbb{R}^2 \times (0, +\infty)$. Finally we recover the classical maximum principle by multiplying (3.37) by the convex functions $g(\xi) = (\xi - \xi_0)_+$ and $g(\xi) = (\xi - \xi_0)_-$ and integrating over $\mathbb{R}^2 \times (0, +\infty)$. Indeed as we have that:

$$-C \leq \min_x (v_0 - \rho_0^\theta) \leq \max_x (v_0 + \rho_0^\theta) \leq C,$$

and that:

$$\text{supp}\xi = [v - \rho^\theta, v + \rho[\theta].$$

For ξ_0 large enough, we can show that:

$$\text{supp}\xi_0 \cap \text{supp}\chi = \emptyset.$$

We have obtain then that:

$$-C \leq \min_x (v_0 - \rho_0^\theta) \leq v_n - \rho_n^\theta \leq v_n + \rho_n^\theta \leq \max_x (v_0 + \rho_0^\theta) \leq C.$$

In particular we obtained that ρ_n and v_n are uniformly bounded in $L^\infty(0, T_n, L^\infty(\mathbb{R}))$ or:

$$\sup_{x \in \mathbb{R}, t \in (0, T_n)} (|\rho_n(t, x)| + |v_n(t, x)|) \leq C_0, \quad (3.38) \quad \boxed{\text{imp2}}$$

4 Proof of theorem ^{theo1}1.2

4.1 Uniform estimates for the solutions of ^{1.1}(1.3)

section4

First we assume that $(\rho^\epsilon, v^\epsilon)$ is the global solutions of Korteweg's equations ^(1.1)(1.3) constructed in theorem ^{theo}1.1 and satisfying:

$$\rho^\epsilon(t, x) \geq c^\epsilon(t), \text{ for some } c^\epsilon(t) > 0, \quad (4.39) \quad \boxed{2.1}$$

and

$$\lim_{x \rightarrow \pm\infty} (\rho^\epsilon, v^\epsilon)(x, t) = (\rho^\pm, u^\pm). \quad (4.40) \quad \boxed{2.2}$$

Here we are working around a non constant state $(\bar{\rho}, \bar{v})$ with:

$$\lim_{x \rightarrow \pm\infty} (\bar{\rho}, \bar{v})(x, t) = (\rho^\pm, u^\pm).$$

It is a simple extension of theorem ^{theo}1.1. Our goal is now to check the properties ^(1.8)(1.11), ^(1.9)(1.12) and ^(1.10)(1.13) in order to use the theorem ^{theo2}1.3 of Chen and Perepelista (see ^(1.10)[5]) in order to prove the theorem ^{theo1}1.2.

For simplicity, throughout this section, we denote $(\rho, v) = (\rho^\epsilon, v^\epsilon)$ and $C > 0$ denote the constant independent of ϵ .

We start with recalling the inequality energy for system ^(1.1)(1.3), indeed by the introduction of the effective velocity we obtain new entropies (see ^{Prepa}[13]).

Lemma 1 *Suppose that $E_1[\rho_0, u_0] \leq E_0 < +\infty$ for some $E_0 > 0$ independent of ϵ . It holds that:*

$$\sup_{0 \leq \tau \leq t} E_1[\rho, u](\tau) + \epsilon \int_0^t \int_{\mathbb{R}} \rho u_x^2 dx d\tau \leq C(t), \quad (4.41) \quad \boxed{2.3a}$$

and:

$$\sup_{0 \leq \tau \leq t} E_2[\rho, v](\tau) + \epsilon \int_0^t \int_{\mathbb{R}} \rho v_x^2 dx d\tau + \epsilon \int_0^t \int_{\mathbb{R}} \rho^{\gamma-2} \rho_x^2 dx d\tau \leq C(t), \quad (4.42) \quad \boxed{2.3}$$

lemma1

where $C(t)$ depends on E_0 , t , $\bar{\rho}$, and \bar{u} but not on ϵ .

Proof: It suffices to write the energy inequalities for system (I.3) and from system (I.5) . More exactly we have:

$$\frac{d}{dt} \int_{\mathbb{R}} (\eta(\rho, m) - \eta(\bar{\rho}, \bar{\rho}\bar{u})) dx + \epsilon \int_{\mathbb{R}} \rho u_x^2 dx = q(\rho^-, m^-) - q(\rho^+, m^+),$$

with the entropy pair:

$$\eta(\rho, m) = \frac{m^2}{2\rho} + e(\rho), \quad q(\rho, m) = \frac{m^3}{2\rho^2} + m e'(\rho),$$

with $e(\rho) = \frac{a}{\gamma-1} \rho^\gamma$. Since we have:

$$\epsilon(\rho, \bar{\rho}) \geq \rho(\rho^\theta - \bar{\rho}^\theta)^2, \quad \theta = \frac{\gamma-1}{2},$$

we can classically bootstrap on the left hand-side the term $q(\rho^-, m^-) - q(\rho^+, m^+)$.

Remark 4 Since vacuum could occur in our solution, the inequality

$$\int_0^t \int_{\mathbb{R}} \rho u_x^2 dx d\tau \leq C(t),$$

in (I.42) is much weaker than the corresponding one in (I.5) . That is why lemma I.2 will be more tricky to obtain.

The following higher order integrability estimate is crucial in compactness argument.

lemma2

Lemma 2 If the conditions of lemma I.1 hold, then for any $-\infty < a < b < +\infty$ and all $t > 0$, it holds that:

$$\int_0^t \int_a^b \rho^{\gamma+1} dx d\tau \leq C(t, a, b), \quad (4.43) \quad \boxed{2.21}$$

where $C(t) > 0$ depends on E_0 , a , b , γ , t , $\bar{\rho}$, \bar{u} but not on ϵ .

Remark 5 The proof follows the same ideas than in the case of compressible Navier-Stokes equations when we wish to obtain a gain of integrability on the density. We refer to (I.2) for more details. The proof is also inspired from Huang et al in (I.9) .

Proof. Choose $\omega \in C_0^\infty(\mathbb{R})$ such that:

$$0 \leq \omega(x) \leq 1, \quad \omega(x) = 1 \quad \text{for } x \in [a, b], \quad \text{and } \text{supp} \omega = (a-1, b+1).$$

By the momentum equation of (I.3) and by localizing, we have

$$(P(\rho)\omega)_x = -(\rho uv\omega)_x + (P(\rho) + \rho uv)\omega_x - (\rho v)_t \omega + \epsilon(\rho v_x \omega)_x - \epsilon \rho v_x \omega_x. \quad (4.44) \quad \boxed{2.22}$$

Integrating $(\frac{2.22}{4.44})$ with respect to spatial variable over $(-\infty, x)$, we obtain:

$$P(\rho)\omega = -\rho uv\omega + \epsilon(\rho v_x\omega)_x - \left(\int_{-\infty}^x \rho v \omega dy\right)_t + \int_{-\infty}^x [(\rho uv + P(\rho))\omega_x - \epsilon\rho v_x\omega_x] dx. \quad (4.45) \quad \boxed{2.23}$$

Multiplying $(\frac{2.23}{4.45})$ by $\rho\omega$, we have

$$\begin{aligned} \rho P(\rho)\omega^2 &= -\rho^2 uv\omega^2 + \epsilon\rho^2 v_x\omega^2 - (\rho\omega \int_{-\infty}^x \rho v \omega dy)_t \\ &\quad - (\rho u)_x\omega \left(\int_{-\infty}^x \rho u \omega dy\right) + \rho\omega \int_{-\infty}^x [(\rho uv + P(\rho))\omega_x - \epsilon\rho v_x\omega_x] dx, \\ &= \epsilon\rho^2 v_x\omega^2 - (\rho\omega \int_{-\infty}^x \rho v \omega dy)_t - (\rho u\omega \int_{-\infty}^x \rho v \omega dy)_x \\ &\quad + \rho u\omega_x \int_{-\infty}^x \rho v \omega dy + \rho\omega \int_{-\infty}^x [(\rho uv + P(\rho))\omega_x - \epsilon\rho v_x\omega_x] dx, \end{aligned} \quad (4.46) \quad \boxed{2.24}$$

We now integrate $(\frac{2.24}{4.46})$ over $(0, t) \times \mathbb{R}$ and we get:

$$\begin{aligned} \int_0^t \int_{\mathbb{R}} a\rho^{\gamma+1}\omega^2 dx d\tau &= \epsilon \int_0^t \int_{\mathbb{R}} \rho^2 v_x\omega^2 - \int_{\mathbb{R}} (\rho\omega \int_{-\infty}^x \rho v \omega dy) dx \\ &+ \int_{\mathbb{R}} (\rho_0\omega \int_{-\infty}^x \rho_0 v_0 \omega dy) dx + \int_0^t \int_{\mathbb{R}} (\rho u\omega_x \int_{-\infty}^x \rho v \omega dy) dx d\tau \\ &\quad + \int_0^t \int_{\mathbb{R}} (\rho\omega \int_{-\infty}^x [(\rho uv + P(\rho))\omega_x - \epsilon\rho v_x\omega_x] dx) dx d\tau. \end{aligned} \quad (4.47) \quad \boxed{2.25}$$

Let

$$A = \{x : \rho(t, x) \geq \underline{\rho}\}, \quad \text{where } \underline{\rho} = 2 \max(\rho_+, \rho_-), \quad (4.48) \quad \boxed{2.26}$$

then we have the following estimates by $(\frac{2.3}{4.42})$:

$$|A| \leq \frac{C(t)}{e^*(2\underline{\rho}, \underline{\rho})} = d(t). \quad (4.49) \quad \boxed{2.27}$$

By $(\frac{2.26}{4.48})$, for any (t, x) there exists a point $x_0 = x_0(t, x)$ such that $|x - x_0| \leq d(t)$ and $\rho(t, x_0) = \underline{\rho}$. Here we choose $\beta = \frac{\gamma+1}{2} > 0$,

$$\begin{aligned} \text{supp}_{x \in \text{supp}(\omega)} \epsilon\rho^\beta(t, x) &\leq \epsilon\underline{\rho}^\beta + \text{supp}_{x \in \text{supp}(\omega) \cap A} \epsilon\rho^\beta(t, x), \\ &\leq 2\epsilon\underline{\rho}^\beta + \text{supp}_{x \in \text{supp}(\omega) \cap A} |\epsilon\rho^\beta(t, x) - \epsilon\rho^\beta(t, x_0)|, \\ &\leq 2\epsilon\underline{\rho}^\beta + \text{supp}_{x \in \text{supp}(\omega) \cap A} \int_{x_0-d(t)}^{x_0+d(t)} |\beta| |\epsilon\rho^{\beta-1}(t, x) \rho_x| dx, \\ &\leq 2\epsilon\underline{\rho}^\beta + \int_{a-1-2d(t)}^{b+1+2d(t)} |\beta| \rho^{2\beta-1} dx + \int_{\mathbb{R}} \epsilon^2 \rho^{-1} \rho_x^2 dx, \\ &\leq C(t) + \int_{a-1-2d(t)}^{b+1+2d(t)} \rho^\gamma dx, \\ &\leq C(t). \end{aligned} \quad (4.50) \quad \boxed{2.28}$$

Using (2.28) , (4.50) , Young inequalities and Hölder's inequalities, the first term on the right hand side of (4.47) is treated as follows:

$$\begin{aligned}
& \int_0^t \int_{\mathbb{R}} \rho^2 v_x \omega^2 dx d\tau \\
& \leq \frac{1}{2} \epsilon \int_0^t \int_{\mathbb{R}} \rho^3 \omega^4 dx d\tau + \frac{1}{2} \epsilon \int_0^t \int_{\mathbb{R}} \rho v_x^2 dx d\tau, \\
& \leq C(t) + \epsilon \int_0^t \int_{\mathbb{R}} \rho^3 \omega^2 dx d\tau, \\
& \leq C(t) + C(t) \int_0^t \int_{\mathbb{R}} \rho^{4-\beta} \omega^2 dx d\tau, \\
& \leq C(t) + \delta \int_0^t \int_{\mathbb{R}} \rho^{\gamma+1} \omega^2 dx d\tau,
\end{aligned} \tag{4.51} \quad \boxed{2.29}$$

Here we have used the fact that $\gamma > \frac{5}{3}$. By lemma $\boxed{1}$ and the Hölder inequality, we obtain

$$\begin{aligned}
\left| \int_{-\infty}^x \rho v \omega dy \right| & \leq \int_{\text{supp}(\omega)} |\rho v| dy, \\
& \leq \left(\int_{\text{supp}(\omega)} \rho dy \right)^{\frac{1}{2}} \left(\int_{\text{supp}(\omega)} \rho v^2 dy \right)^{\frac{1}{2}} \leq C(t).
\end{aligned} \tag{4.52} \quad \boxed{2.30}$$

Then:

$$\begin{aligned}
& \left| \int_{\mathbb{R}} (\rho \omega \int_{-\infty}^x \rho v \omega dy) dx \right| + \left| \int_{\mathbb{R}} (\rho_0 \omega \int_{-\infty}^x \rho_0 v_0 \omega dy) dx \right| \\
& \quad + \left| \int_0^t \int_{\mathbb{R}} (\rho \omega_x \int_{-\infty}^x \rho v \omega dy) dx d\tau \right| \leq C(t).
\end{aligned} \tag{4.53} \quad \boxed{2.32}$$

Similarly, we have:

$$\left| \int_0^t \int_{\mathbb{R}} (\rho \omega \int_{-\infty}^x (\rho v v + P(\rho)) \omega_x dy) dx d\tau \right| \leq C(t), \tag{4.54} \quad \boxed{2.33}$$

and

$$\begin{aligned}
& \epsilon \left| \int_0^t \int_{\mathbb{R}} (\rho \omega \int_{-\infty}^x \rho v_x \omega_x dy) dx d\tau \right| \\
& \leq \epsilon \left| \int_0^t \int_{\mathbb{R}} (\rho \omega \int_{\mathbb{R}} \rho |v_x| |\omega_x| dy) dx d\tau \right|, \\
& \leq \epsilon \left| \int_0^t \left(\int_{\mathbb{R}} \rho \omega dx \right) \left(\int_{\mathbb{R}} \rho v_x^2 dy + \int_{\mathbb{R}} \rho \omega_x^2 dy \right) d\tau \right|, \\
& \leq C(t).
\end{aligned} \tag{4.55} \quad \boxed{2.34}$$

Substituting (2.29) , (4.51) , (2.32) , (4.53) - (2.34) , (4.55) into (4.47) and noticing the smallness of δ , we proved lemma $\boxed{2}$.

Lemma 3 Suppose that $(\rho_0(x), v_0(x))$ satisfy the conditions in the lemmas $\boxed{1}$. Furthermore there exists $M_0 > 0$ independent of ϵ , such that

$$\int_{\mathbb{R}} \rho_0(x) |v_0(x) - \bar{v}(x)| dx \leq M_0 < +\infty, \tag{4.56} \quad \boxed{2.35}$$

then for any compact set $K \subset \mathbb{R}$, it holds that:

$$\int_0^t \int_K (\rho^{\gamma+\theta} + \rho|v|^3) dx d\tau \leq C(t, K), \quad (4.57) \quad \boxed{2.36}$$

where $C(t, K)$ is independent of ϵ .

Remark 6 In order to prove the inequality $\frac{2.36}{(4.57)}$, we will use the same ingredients than in $\frac{35}{[26]}$ where this inequality was obtained for the first time.

Proof. We are now working with the function ψ of proposition $\frac{\text{pair35}}{2.3}$. If we consider η_m^ψ as a function depending of (ρ, v) , we have for all $\rho \geq 0$ and $v \in \mathbb{R}$:

$$\begin{cases} |\eta_{mv}^\psi(\rho, v)| \leq C, \\ |\eta_{m\rho}^\psi(\rho, v)| \leq C\rho^{\theta-1}. \end{cases} \quad (4.58) \quad \boxed{2.38}$$

For this weak entropy pair (η^ψ, H^ψ) , we observe that:

$$\eta^\psi(\rho, 0) = \eta_\rho^\psi(\rho, 0) = 0, \quad H^\psi(\rho, 0) = \frac{\theta}{2}\rho^{3\theta+1} \int_{\mathbb{R}} |s|^3 [1 - s^2]_+^\lambda ds,$$

and:

$$\eta_m^\psi(\rho, 0) = \beta\rho^\theta \quad \text{with } \beta = \int_{\mathbb{R}} |s|[1 - s^2]_+^\lambda ds.$$

By Taylor formula, we have:

$$\eta^*(\rho, m) = \beta\rho^\theta m + r(\rho, m), \quad (4.59) \quad \boxed{2.39}$$

with:

$$r(\rho, m) \leq C\rho v^2, \quad (4.60) \quad \boxed{2.40}$$

for some constant $C > 0$. Now we introduce a new entropy pair $(\hat{\eta}, \hat{H})$ such that,

$$\hat{\eta}(\rho, m) = \eta^\psi(\rho, m - \rho v^-), \quad \hat{H}(\rho, m) = H^\psi(\rho, m - \rho v^-) + v^- \eta^\psi(\rho, m - \rho v^-),$$

with $m = \rho v$ which satisfies:

$$\begin{cases} \hat{\eta}(\rho, m) = \beta\rho^{\theta+1}(v - v^-) + r(\rho, \rho(v - v^-)), \\ r(\rho, \rho(v - v^-)) \leq C\rho(v - v^-)^2. \end{cases} \quad (4.61) \quad \boxed{2.41}$$

Integrating $\frac{1.1}{(1.3)}_1 \times \hat{\eta}_\rho + \frac{1.1}{(1.3)}_2 \times \hat{\eta}_m$ over $(0, t) \times (-\infty, x)$, we have:

$$\begin{aligned} & \int_{-\infty}^x (\hat{\eta}(\rho, m) - \hat{\eta}(\rho_0, m_0)) dy + \int_0^t q^*(\rho, \rho(v - v^-)) + v^- \eta^*(\rho, \rho(v - v^-)) d\tau \\ & = tq^*(\rho^-, 0) + \epsilon \int_0^t \hat{\eta}_m \rho v_x d\tau - \epsilon \int_0^t \int_{-\infty}^x (\hat{\eta}_{mu} \rho v_x^2 + \hat{\eta}_{m\rho} \rho \rho_x v_x) dy d\tau. \end{aligned} \quad (4.62) \quad \boxed{2.42}$$

By using $\frac{2.38}{(4.58)}$, we obtain:

$$|\epsilon \int_0^t \int_{-\infty}^x \hat{\eta}_{mu} \rho v_x^2 dy d\tau| \leq C\epsilon \int_0^t \int_{\mathbb{R}} \rho v_x^2 dy d\tau \leq C(t), \quad (4.63) \quad \boxed{2.43}$$

$$\begin{aligned}
|\epsilon \int_0^t \int_{-\infty}^x \widehat{\eta}_{m\rho} \rho \rho_x v_x dy d\tau| &\leq C\epsilon \int_0^t \int_{\mathbb{R}} \rho^{\theta-1} \rho |\rho_x v_x| dy d\tau \leq C(t), \\
&\leq C\epsilon \int_0^t \int_{\mathbb{R}} \rho v_x^2 dy d\tau + C\epsilon \int_0^t \int_{\mathbb{R}} \rho^{\gamma-2} \rho_x^2 dy d\tau \leq C(t).
\end{aligned} \tag{4.64} \quad \boxed{2.44}$$

Substituting $\frac{2.43}{(4.63)}$ and $\frac{2.44}{(4.64)}$ into $\frac{2.42}{(4.62)}$, then integrating over K and using $\frac{2.37}{(2.25)}$, we obtain:

$$\begin{aligned}
&\int_0^t \int_K \rho^{\theta+\gamma} + \rho |v - v^-|^3 dx d\tau \\
&\leq C(t) + C \int_0^t \int_K |\eta^*(\rho, \rho(v - v^-))| dx d\tau + C\epsilon \int_0^t \int_K \rho |v| |v_x| dx d\tau \\
&+ C\epsilon \int_0^t \int_K \rho^{1+\theta} |v_x| dx d\tau + 2 \sup_{\tau \in [0, t]} \left| \int_K \left(\int_{-\infty}^x \widehat{e\tau a}(\rho(y, \tau), (\rho v)(y, \tau)) dy \right) dx \right|.
\end{aligned} \tag{4.65} \quad \boxed{2.45}$$

Applying lemma $\frac{\text{lemma 1}}{\text{I}}$, we have:

$$\int_0^t \int_K |\eta^*(\rho, \rho(v - v^-))| dx d\tau \leq C(t). \tag{4.66} \quad \boxed{2.46}$$

By Hölder's inequality and $\frac{2.28}{(4.50)}$, we get:

$$\begin{aligned}
\epsilon \int_0^t \int_K \rho^{1+\theta} |v_x| dx d\tau &\leq C\epsilon \int_0^t \int_K \rho v_x^2 dx d\tau + C\epsilon \int_0^t \int_K \rho^{1+2\theta} dx d\tau, \\
&\leq C(t) + C(t) \int_0^t \int_K \rho^\theta dx d\tau, \\
&\leq C(t).
\end{aligned} \tag{4.67} \quad \boxed{2.47}$$

We have now:

$$\begin{aligned}
\epsilon \int_0^t \int_K \rho |v| |v_x| dx d\tau &\leq \frac{1}{2}\epsilon \int_0^t \int_K \rho v_x^2 dx d\tau + \frac{1}{2}\epsilon \int_0^t \int_K \rho v^2 dx d\tau, \\
&\leq C(t).
\end{aligned} \tag{4.68} \quad \boxed{2.48}$$

Now we are going to deal with the last term on the right hand side of $\frac{2.45}{(4.65)}$. $\frac{1.1}{(1.3)}$ implies that:

$$(\rho v - \rho v^-)_t + (\rho v^2 + P(\rho) - \rho u u^-)_x = \epsilon (\rho v_x)_x. \tag{4.69} \quad \boxed{2.49}$$

Integrating $\frac{2.49}{(4.69)}$ over $[0, t] \times (-\infty, x)$ for $x \in K$, we get:

$$\begin{aligned}
\int_{-\infty}^x \rho(v - \rho v^-) dy &= \int_{-\infty}^x \rho_0(v_0 - \rho v^-) dy - \int_0^t (\rho v^2 + P(\rho) - \rho u u^- - P(\rho^-)) \\
&+ \epsilon \int_0^t \rho v_x d\tau.
\end{aligned} \tag{4.70} \quad \boxed{2.50}$$

Furthermore:

$$\begin{aligned}
& \left| \int_{-\infty}^x \widehat{\eta}((\rho(y, \tau), (\rho v)(y, \tau))) dy \right| \\
& \leq \left| \int_{-\infty}^x (\widehat{\eta}(\rho \rho v) - \beta \rho^{\theta+1}(v - \bar{v})) dy \right| + \left| \int_{-\infty}^x \beta \rho^{\theta+1}(v - \bar{v}) dy \right| \\
& \leq \left| \int_{-\infty}^x (r(\rho \rho(v - \bar{v}))) dy \right| + \left| \int_{-\infty}^x \beta(\rho^\theta - (\rho^-)^\theta) \rho(v - \bar{v}) dy \right| \\
& \qquad \qquad \qquad + \beta(\rho^-)^\theta \left| \int_{-\infty}^x \rho(v - \bar{v}) dy \right|, \\
& \leq C(t) + \beta(\rho^-)^\theta \left| \int_{-\infty}^x \rho(v - \bar{v}) dy \right|.
\end{aligned} \tag{4.71} \quad \boxed{2.51}$$

By using [\(2.35\)](#), lemma [1](#) and [2](#), [\(4.70\)](#) and [\(4.71\)](#) we conclude the proof of the lemma.

4.2 $H_{loc}^{-1}(\mathbb{R}_+^2)$ Compactness

In this section we are going to take profit of the uniform estimates obtained in the previous section in order to prove the following lemma, which gives the $H_{loc}^{-1}(\mathbb{R}_+^2)$ -compactness of the Korteweg solution sequence $(\rho^\epsilon, v^\epsilon)$ on a entropy- entropy flux pair.

Lemma 4 *Let $\psi \in C_0^2(\mathbb{R})$, η^ψ, H^ψ be a weak entropy pair generated by ψ . Then for the solutions $(\rho^\epsilon, v^\epsilon)$ with $m^\epsilon = \rho^\epsilon v^\epsilon$ of Korteweg system [\(1.3\)](#), the following sequence:*

$$\eta^\psi(\rho^\epsilon, m^\epsilon)_t + q^\psi(\rho^\epsilon, m^\epsilon)_x \text{ are compact in } H_{loc}^{-1}(\mathbb{R}_+^2) \tag{4.72} \quad \boxed{3.1}$$

lemme4

Proof: Now we are going to prove the lemma. A direct computation on [\(1.3\)](#)₁ \times [\(1.3\)](#)₂ \times $\eta_m^\psi(\rho^\epsilon, m^\epsilon)$ gives:

$$\begin{aligned}
\eta_\rho^\psi(\rho^\epsilon, m^\epsilon)_t + H_\rho^\psi(\rho^\epsilon, m^\epsilon)_x &= \epsilon(\eta_\rho^\psi(\rho^\epsilon, m^\epsilon)(\rho^\epsilon)v_x^\epsilon) - \epsilon\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)(\rho^\epsilon)(v_x^\epsilon)^2 \\
&\quad - \epsilon\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)(\rho^\epsilon)v_x^\epsilon\rho_x^\epsilon.
\end{aligned} \tag{4.73} \quad \boxed{3.7}$$

Let $K \subset \mathbb{R}$ be compact, using proposition [2.4](#) [\(2.30\)](#) and Hölder inequality, we get:

$$\begin{aligned}
& \epsilon \int_0^t \int_K |\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)(\rho^\epsilon)|(v_x^\epsilon)^2 + |\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)(\rho^\epsilon)v_x^\epsilon\rho_x^\epsilon| dx d\tau \\
& \leq C\epsilon \int_0^t \int_K (\rho^\epsilon)|(v_x^\epsilon)^2 dx d\tau + C\epsilon \int_0^t \int_K (\rho^\epsilon)^{\gamma-2}(\rho_x^\epsilon)^2 dx d\tau \\
& \leq C(t).
\end{aligned} \tag{4.74} \quad \boxed{3.8}$$

This shows that:

$$-\epsilon\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)\rho^\epsilon(v_x^\epsilon)^2 - \epsilon\eta_{mu}^\psi(\rho^\epsilon, m^\epsilon)\rho^\epsilon v_x^\epsilon\rho_x^\epsilon \text{ are bounded in } L^1([0, T] \times K), \tag{4.75} \quad \boxed{3.9}$$

and thus it is compact in $W_{loc}^{-1,p_1}(\mathbb{R}_+^2)$, for $1 < p_1 < 2$. Moreover we observe that

$$|\eta_{mu}^\psi(\rho^\epsilon, \rho^\epsilon v^\epsilon)| \leq C_\psi,$$

, then we obtain:

$$\begin{aligned}
& \int_0^t \int_K (\epsilon \eta_m^\psi(\rho^\epsilon, m^\epsilon) \rho^\epsilon v_x^\epsilon)^{\frac{4}{3}} dx dt \\
& \leq \int_0^t \int_K \epsilon^{\frac{4}{3}} |\rho^\epsilon|^{\frac{4}{3}} |v_x^\epsilon|^{\frac{4}{3}} dx dt \\
& \leq C \epsilon^{\frac{4}{3}} \int_0^t \int_K \rho^\epsilon |v_x^\epsilon|^2 dx dt + C \epsilon^{\frac{4}{3}} \int_0^t \int_K (\rho^\epsilon)^2 dx dt \\
& \leq C(t, K) \epsilon^{\frac{1}{3}} + C \epsilon^{\frac{4}{3}} \int_0^t \int_K (\rho^\epsilon)^{\gamma+1} dx dt \xrightarrow{\epsilon \rightarrow 0} 0.
\end{aligned} \tag{4.76} \quad \boxed{3.10}$$

Using $\boxed{3.10}$ and $\boxed{3.9}$, we obtain

$$\eta_\rho^\psi(\rho^\epsilon, m^\epsilon)_t + H_\rho^\psi(\rho^\epsilon, m^\epsilon)_x \text{ are compact in } W_{loc}^{-1, p_1}(\mathbb{R}_+^2) \text{ for some } 1 < p_1 < 2. \tag{4.77} \quad \boxed{3.11}$$

Furthermore by $\boxed{2.27}$ - $\boxed{2.28}$, lemma $\boxed{1.2}$ and $\boxed{4.57}$, we have:

$$\eta_\rho^\psi(\rho^\epsilon, m^\epsilon)_t + H_\rho^\psi(\rho^\epsilon, m^\epsilon)_x \text{ are uniformly bounded in } L_{loc}^{p_3}(\mathbb{R}_+^2) \text{ for } p_3 > 2. \tag{4.78} \quad \boxed{3.12}$$

where $p_3 = \gamma + 1 > 2$ when $\gamma \in (1, 3]$, and $p_3 = \frac{\gamma+\theta}{1+\theta} > 2$ when $\gamma > 3$. By interpolation we conclude the proof of the lemma $\boxed{4}$.

5 Proof of theorem $\boxed{1.2}$

From lemmas $\boxed{1}$, we have verified the conditions (i)-(iii) of theorem $\boxed{1.3}$ for the sequence of solutions $(\rho^\epsilon, m^\epsilon)$. Using theorem $\boxed{1.3}$, there exists a subsequence $(\rho^\epsilon, m^\epsilon)$ and a pair of measurable functions (ρ, m) such that

$$(\rho^\epsilon, m^\epsilon) \rightarrow (\rho, m), \quad \text{a.e } \epsilon \rightarrow 0. \tag{5.79} \quad \boxed{4.1}$$

It is easy to check that (ρ, m) is a finite-energy entropy solution (ρ, m) to the Cauchy problem $\boxed{1.5}$ with initial data $(\rho_0, \rho_0 u_0)$ for the isentropic Euler equations with $\gamma > \frac{5}{3}$. It achieves the proof of theorem $\boxed{1.2}$.

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