



HAL
open science

Bound and continuum states of molecular anions C₂H and C₃N

Stephen Harrison, Jonathan Tennyson

► **To cite this version:**

Stephen Harrison, Jonathan Tennyson. Bound and continuum states of molecular anions C₂H and C₃N. *Journal of Physics B: Atomic, Molecular and Optical Physics*, 2011, 44 (4), pp.45206. 10.1088/0953-4075/44/4/045206 . hal-00596622

HAL Id: hal-00596622

<https://hal.science/hal-00596622>

Submitted on 28 May 2011

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

Bound and continuum states of molecular anions C_2H^- and C_3N^-

Stephen Harrison and Jonathan Tennyson

Department of Physics and Astronomy, University College London, Gower St.,
London WC1E 6BT, UK

Abstract.

Recently a number of molecular anions, closed-shell linear carbon chains of the form C_nH^- and C_nN^- , have been detected in space. The molecules C_2H^- and C_3N^- are investigated by using the R-Matrix method to consider electron scattering from the corresponding neutral targets. Initial target calculations are conducted and refined in order to produce target state characteristics similar to the experimental data. A number of different scattering models are tested including static exchange and close-coupling models, and the use of Hartree-Fock or natural orbitals in the close-coupling calculations. The calculations concentrate on bound and resonance states for the anions as well as eigenphase sums, elastic cross-sections and electronic excitation cross-sections for electron collisions with the neutral. It is found that electronic resonances are all too high in energy to be important for anion formation in the interstellar medium. However, C_3N^- , unlike C_2H^- , supports a number of very weakly bound excited states, which may well provide the route to electron attachment for this system.

1. Introduction

The first anion detected in space was C_6H^- by McCarthy et al. (2006). Since then C_4H^- (Gupta et al. 2007), C_8H^- (Brunken et al. 2007), C_3N^- (Thaddeus et al. 2008) and C_5N^- (Cernicharo et al. 2008) have also all been detected. CN^- , C_3N^- or C_4H^- , and C_5N^- have also recently been identified in the atmosphere of Titan (Vuitton et al. 2009). Since the present work was completed, CN^- has also been detected in the interstellar medium (ISM) (Agundez et al. 2010). These anions are all linear multiply-bonded carbon chains with closed shell ground state configurations. The prevailing opinion is that these anions are formed by radiative attachment of an electron to the neutral molecule with the same chemical formula (Millar et al. 2007). These neutral precursors are all linear multiply-bonded carbon chains with significant dipole moments which are known to exist in the interstellar medium. The corresponding anions are closed shell molecules whose extra electron is bound by several eV.

Herbst and co-workers (Petrie & Herbst 1997, Terzieva & Herbst 2000, Herbst & Osamura 2008) developed and applied a model for calculating radiative association rates for these C_nH^- and $C_{n+1}N^-$ (n even) species. This model is based on the formation of low-lying resonances in electron collisions and fairly simple phase-space arguments.

Studies suggest that although this model gives anion formation rates of approximately the right magnitude, the rates it provides cannot actually reproduce the observations (Harada & Herbst 2008).

Although there have been a number of electronic structure studies of these anions (see for example Woon (1995), Natterer & Koch (1995), Botschwina & Oswald (2008), Fortenberry et al. (2010)), there appears to have been no previous study of their continuum states. We have therefore performed a series of studies of electron collisions with both C_nH and $C_{n+1}N$ (n even) targets. Initially our studies focussed on electron collisions with C_4H and C_3N as the first member of each series for which the associated anion had been observed in the interstellar medium. However we found it difficult to establish definitive properties for the ground state of C_4H , a problem that is well-documented (Fortenberry et al. 2010). Although we have performed a series of electron collision studies with C_4H , here we report electron collision calculations which focus on C_2H and C_3N as we consider the results of these calculations to be more reliable.

2. Method

Scattering calculations were performed using the R-matrix method (Burke & Berrington 1993), whose use for collision calculations involving molecules has recently been extensively reviewed by one of us (Tennyson 2010). The basic idea of the R-matrix method is the division of space into two regions: a spherical inner region of radius a which is designed to enclose all the charge density of the target, and an outer region where interactions between the electron and the target can be represented by long-range moments. The major advantage of the method is that within the inner region the problem is solved independent of scattering energy. Within this region the wavefunction is written:

$$\psi_k^{N+1} = \mathcal{A} \sum_{ij} a_{ijk} \Phi_i^N(\mathbf{x}_1 \dots \mathbf{x}_N) u_{ij}(\mathbf{x}_{N+1}) + \sum_i b_{ik} \chi_i^{N+1}(\mathbf{x}_1 \dots \mathbf{x}_{N+1}), \quad (1)$$

where the target contains N electrons and functions are labelled as N or $N+1$ according to whether they refer to the target or the compound scattering system respectively. In this work all orbitals underlying the wave functions of eq. (1) were constructed from Gaussian Type Orbitals (GTOs); as discussed below a variety of different orbital sets (Hartree-Fock, CASSCF and Natural orbitals) were tested. In this regard one issue is the requirement to represent all target states included in the calculation with a single orbital set.

In eq. (1), Φ_i^N is the wave function of i^{th} target state. u_{ij} are the extra orbitals introduced to represent the scattering electron (Faure et al. 2002). The operator \mathcal{A} ensures that the product of these two terms obeys the Pauli principle. The second summation in eq. (1) involves L^2 configurations which have no amplitude on the R-matrix boundary and where all electrons are placed in orbitals associated with the target. As we are dealing with electron collision with strongly dipolar systems, it is necessary to allow for the contribution of electron interactions with the long-range target

dipole moment to any cross sections calculated. This was carried out using the program BORNCROS (Baluja et al. 2000), which directly calculates the dipole Born correction to the cross-sections.

The scattering calculations reported here were performed using the UK molecular R-matrix codes (Morgan et al. 1998), mostly starting from the Quantemol-N interface (Tennyson et al. 2007). Target electronic structure calculations were largely performed with MOLPRO (Werner et al. 2008), which was also used to provide many of the sets of target molecular orbitals. For both molecules scattering calculations were performed at two levels: static exchange (SE), which is useful for identifying shape resonances, and close coupling (CC). The presence of very low-lying excited target states meant that the widely used one-state static exchange plus polarisation model was deemed inappropriate for this study, as this method is prone to giving pseudoresonances at energies above the first electronically excited target state.

All calculations were performed at the equilibrium geometry of the neutral target and ignored effects due to nuclear motion. We note that although all molecules considered here are linear, neither the polyatomic R-matrix codes nor MOLPRO use $C_{\infty v}$ symmetry. All calculations were therefore performed using C_{2v} symmetry. In most cases identifying the results in $C_{\infty v}$ is straightforward and these symmetry labels are largely used below.

3. C_2H^-

In choosing appropriate target wavefunctions to include in the scattering calculation a number of criteria are used. Firstly, the target model must be one for which an appropriate, balanced scattering model can be designed (Tennyson 1996). For this a complete active space configuration interaction (CAS-CI) model is often adopted. Secondly the excitation energy of important target electronic states should be approximately correct. Thirdly other target properties, in particular for strongly polar molecules, the permanent dipole moment of the ground state, should be well represented. Satisfying all these criteria for the systems studied here proved to be difficult and led to some compromises being made. Indeed we note that for C_4H^- electronic structure calculations performed without requirement to use the resulting wavefunction in a scattering calculation struggled to satisfy the other requirements (Fortenberry et al. 2010).

3.1. Target calculations

A variety of models were tested when calculating results for the neutral C_2H target, a selection of these are shown in Table 1. Models 1 – 6 were carried out using the R-Matrix software Quantemol-N. These comprised Hartree-Fock (HF) and Complete Active Space Configuration-Interaction (CAS-CI) calculations using each of the 3 basis sets, DZP, cc-pVDZ and cc-pVTZ. The CAS-CI calculations used an orbital space of 7,2,2 orbitals of

Table 1. Selected calculated target properties of C_2H . All results are for the equilibrium geometry of C_2H . The calculated absolute energies of the X $^2\Sigma^+$ ground state is given in Hartree while the vertical excitation energy to the low-lying A $^2\Pi$ is given in eV. μ is the ground state dipole moment, given in Debye. See text for details of models.

Model	Basis	Orbitals	Method	X $^2\Sigma^+$	μ	A $^2\Pi$
This work/1	DZP	SCF	HF	-76.142	0.804	1.186
2	DZP	SCF	CAS-CI	-76.200	0.837	1.204
3	cc-pVDZ	SCF	HF	-76.138	0.789	
4	cc-pVDZ	SCF	CAS-CI	-76.207	0.801	1.304
5	cc-pVTZ	SCF	HF	-76.156	0.797	
6	cc-pVTZ	SCF	CAS-CI	-76.208	0.801	1.248
7	cc-pVTZ	NOs	CAS-CI	-76.296	0.640	1.089
Woon (1995)	cc-pVDZ		RCCSD(T)	-76.401	0.766	0.361
	cc-pVTZ		RCCSD(T)	-76.468	0.774	0.402
	CBS limit		RCCSD(T)	-76.499		0.423
Cui & Morokuma (1998)	cc-pVTZ		CASPT2	-76.458		0.790

symmetry a_1 , b_1 , b_2 respectively which maps to 7 σ and 2 π orbitals. Eight electrons were frozen in the 1 – 4 σ orbitals with five active electrons distributed amongst the 5 – 7 σ and 1 – 2 π . Model 7 used this CAS-CI model but the molecular orbitals used in the calculation were produced by MOLPRO, in this case using a multireference CI (MRCI) calculation, and the R-Matrix calculation was carried out using the R-Matrix codes manually, to allow for use of MOLPRO orbitals. These natural orbitals (NOs) were state averaged from 3 states produced one each from 2A_1 , 2B_1 and 2B_2 , ($^2\Sigma^+$ and $^2\Pi$) MRCI calculations, using the MOLPRO 'MATROP' facility, with a 13:1:1 weighting. This weighting was chosen to create a target model with good characteristics for the scattering calculation i.e. a good reproduction of the excitation energies and dipole moment in comparison with the original MOLPRO calculations. Use of the pure 2A_1 MRCI orbitals was found to make the first vertical excitation energy too high in energy by a factor of 2, and the use of more equally weighted state-averaged orbitals was found to reduce the dipole to a value too small to be acceptable in the scattering calculation by up to a factor of 3. Hence model 7 was deemed acceptable in describing both these characteristics without sacrificing one in favour of the other. Whilst we were able to find some experimental data for the bound state energy of the anion, we could not find any data on the neutral, such as the dipole moment.

There have been a number of previous *ab initio* studies on C_2H ; Table 1 compares our models with the best of these. As we only use a very limited correlation space, the other studies give lower absolute energies even for calculations which use the same target basis. Our dipole moments are in reasonable agreement with those predicted by Woon (1995). However our predictions for the vertical energy of excitation for the very low-lying A $^2\Pi$ state appear to systematically to high.

Table 2. C_2H^- bound states with energies given in eV. All calculations are for the equilibrium geometry of C_2H

Model	Basis	Orbitals	Method	$^1\Sigma^+$ Binding energy (eV)
This work/1	DZP	SCF	SE	1.186
2	DZP	SCF	CC	2.713
3	cc-pVDZ	SCF	SE	1.114
4	cc-pVDZ	SCF	CC	2.637
5	cc-pVTZ	SCF	SE	1.156
6	cc-pVTZ	SCF	CC	2.617
7	cc-pVTZ	NOs	CC	2.282
Baker et al. (1986)	6-311 ⁺⁺ G	UHF	UHF	1.56
	6-311 ⁺⁺ G	UHF	UCISD	2.62
Janousek et al. (1979)			experiment	2.94±0.10
Ervin & Lineberger (1991)			experiment	2.969±0.006
Taylor & Xu (1998)			experiment	2.956±0.020

3.2. Scattering calculations and bound anion states

Our scattering calculations concentrated on finding bound states and resonances for C_2H^- but also provide low-energy electron – C_2H collision cross sections. Two classes of models were tested: SE based on a single target state and CC models where 48 state calculations (doublet and quartet states, 6 states per symmetry) were performed in the inner region. As discussed below, different numbers of these states were retained in the outer region of the calculation.

Bound states were calculated using the outer region program BOUND (Sarpal et al. 1991). For these calculations only the low-lying X $^2\Sigma^+$ and A $^2\Pi$ were retained in the outer region of the CC calculations, since experience has shown that strongly closed states can cause such outer region calculations to be numerically unstable. These were propagated to a distance of 30.1 a_0 . Table 2 summarizes the results of these studies. In contrast to C_3N^- discussed below, all models found only a single bound state. The SE calculations give a vertical binding energy of about 1.2 eV whereas for the CC calculations this binding energy is increased to about 2.6 eV. These figures are in broad agreement with the electronic structure calculations of Baker et al. (1986). The CC binding energies are some 0.6 eV less than the measured ones; however the measured binding energies are adiabatic and will therefore be higher than our calculated vertical ones.

3.3. Resonances

Outer region calculations were performed retaining those target states which lie below 10 eV for the CC calculations, the Quantemol-N default setting. This means 25 states for

Table 3. C_2H^- low lying resonance positions (and widths) in eV

Model	1	2	3	4	5	6	7
$^1\Sigma^+$							2.957 (0.122)
$^1\Sigma^-$		3.876 (1.115)		4.231 (1.056)		3.908 (1.080)	5.130 (1.596)
$^3\Sigma^+$		2.589 (0.430)		2.768 (0.377)		2.679 (0.444)	2.923 (0.127)
$^3\Pi$	3.522 (1.949)	3.49 a	3.634 (2.038)	3.380 (1.360)	3.562 (2.021)	3.53 a	3.942 (0.125)
$^3\Sigma^-$		3.374 (0.867)		3.625 (0.814)		3.416 (0.854)	4.495 (1.298)

^a Resonance present in eigenphases but not fitted by RESON.

models 2, 4 and 6 but only 16 for model 7. The results reported below are insensitive to increasing these numbers. Resonances were characterised using a Breit-Wigner fit to the eigenphases, which are shown in Figure 1, as implemented in the automated detection and fitting program RESON (Tennyson & Noble 1984). As noted below, resonances which overlap excitation thresholds were not fitted by this procedure. Only low-lying resonances were considered since higher resonances are unlikely to be important in any radiative association process.

Table 3 summarizes the resonances detected for C_2H^- . The SE calculations all detect a single shape resonance of $^3\Pi$ symmetry at about 3.5 eV. At the SE level this is a very broad resonance with a width of about 2 eV. In the CC calculations this resonance position is lowered to about 3.4 eV and becomes considerably narrower with SCF orbitals and broader with MRCI NO's. The CC calculations based on the use of SCF orbitals introduce Feshbach resonances of $^1\Sigma^-$, $^3\Sigma^+$ and $^3\Sigma^-$ symmetry. Use of NOs introduce a further narrow $^3\Sigma^+$ resonance. We note that none of these resonances lie below the A $^2\Pi$ first excited state of C_2H .

3.4. Cross-sections

Figure 2 presents the total elastic cross-section of C_2H for electron scattering energies up to 10 eV, for all 3 models, SCF-SE, SCF-CC and MRCI-CC. For the final MRCI-CC model, a Born correction was applied due to the dipolar nature of the molecule, which acts to increase the cross-section at all energies.

Electron impact electronic excitation cross sections are given by the many-state CC calculations. Figure 3 shows the cross-sections representing the excitation from the $^2\Sigma^+$ ground state of the neutral to the first three excited states, A $^2\Pi$, a $^4\Sigma^+$ and b $^4\Delta$. Here the Born correction is only applied to the $^2\Pi$ excitation, as this is the only dipole-allowed transition. All three excitation cross sections show resonance features. These are particularly pronounced in the CC-MRCI calculations; we note that cross sections approximately average the SCF-CC ones which show fewer resonances. Although it is likely that these cross sections display resonance features, they can only be properly characterised by much larger calculations than the one presented here.

4. C_3N^-

Calculations on C_3N and C_3N^- follow closely those on C_2H and C_2H^- reported above.

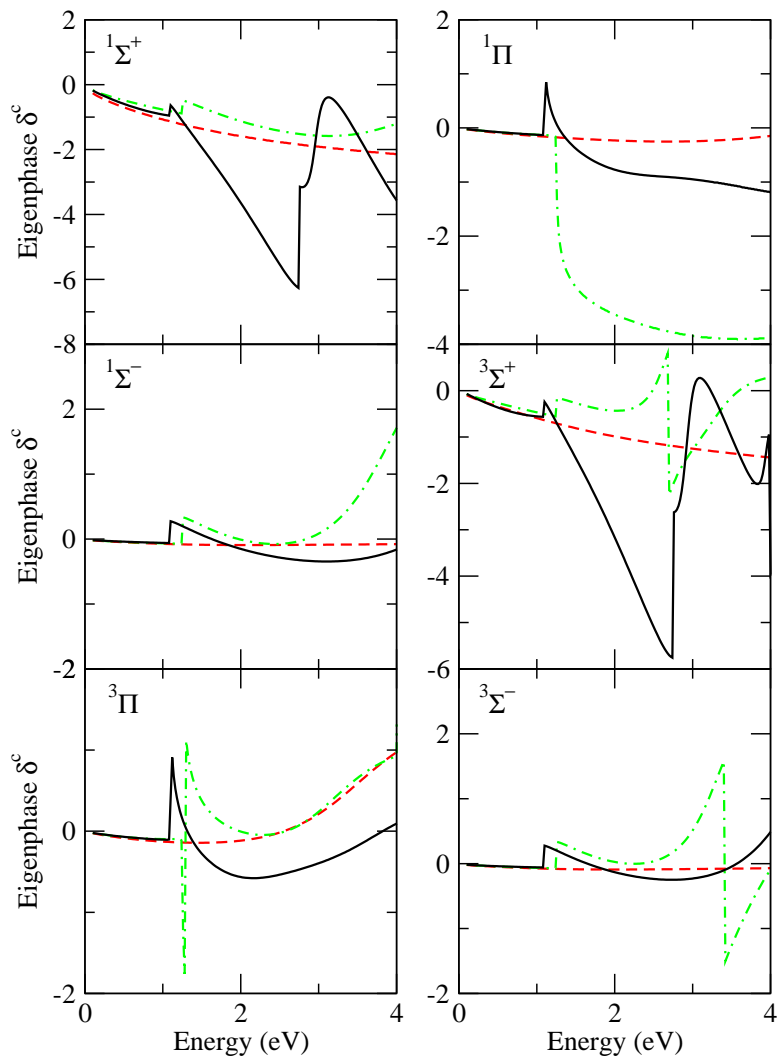


Figure 1. Electron – C_2H eigenphase sums for different symmetries for the SCF-SE(dash), SCF-CC(dash-dot) and MRCI-CC(solid) models.

4.1. Target calculations

Quantemol-N was used to carry out HF and CAS-CI calculations using DZP and cc-pVDZ basis sets. Tests were performed with several other basis sets, including the ANO basis recommended by Widmark et al. (1990) for the C_3N but for which we were unable to get satisfactory results.

Models 2 and 4, used a 10σ , 3π orbital space, with 8 electrons frozen in the $1 - 4\sigma$ orbitals and the remaining 17 electron distributed among the $5 - 10\sigma$ and all the π orbitals. This CAS-CI model was retained for use in model 5, which made use of MOLPRO MRCI orbitals; these NOs were obtained by the state-averaging the lowest 2A_1 , 2B_1 , 2B_2 , and 2A_2 MRCI states in a 5:2:2:1 weighting, manually chosen in order to produce good target characteristics for the scattering run.

Table 4 compares our target models with the results of previous electronic structure calculations. C_3N is a more complicated system than C_2H and there is less agreement

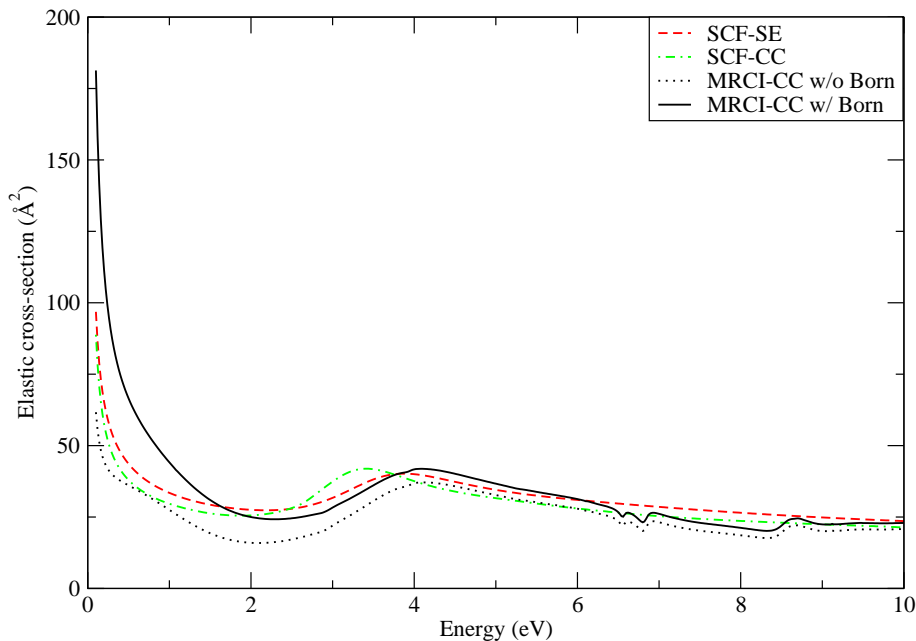


Figure 2. Elastic Cross-Section for electron C_2H collisions

Table 4. Selected calculated target properties of C_3N . All results are for the equilibrium geometry of C_3N . The calculated absolute energies of the $X \ ^2\Sigma^+$ ground state is given in Hartree while the vertical excitation energy to the low-lying $A \ ^2\Pi$ is given in eV. μ is the ground state dipole moment, given in Debye. See text for details of models.

Model	Basis	Orbitals	Method	$X \ ^2\Sigma^+$	μ	$A \ ^2\Pi$
This work/1	DZP	SCF	HF	-167.870	2.332	
2	DZP	SCF	CAS-CI	-167.933	1.288	1.038
3	cc-pVDZ	SCF	HF	-167.871	2.749	
4	cc-pVDZ	SCF	CAS-CI	-167.937	2.270	1.235
5	cc-pVTZ	NOs	CAS-CI	-168.003	3.655	0.406
Feher (1992)	6-31G*		UMP2	-168.345	3.270	
Sadlej & Roos (1991)	ANO	CAS-SCF	CAS-SCF		2.920	0.750
		CAS-SCF	MRCI		3.040	0.400
McCarthy et al. (1995)	cc-pVTZ		RCCSD(T)	-167.440		0.266
			RHF	-168.071	3.254	
			RCCSD	-167.439	2.887	
			RCCSD(T)	-167.400	2.852	0.270

between these calculations on important properties such as the lowest vertical excitation energy and the permanent target dipole moment. Despite a thorough search no experimental dipole moments could be found for the neutral molecule

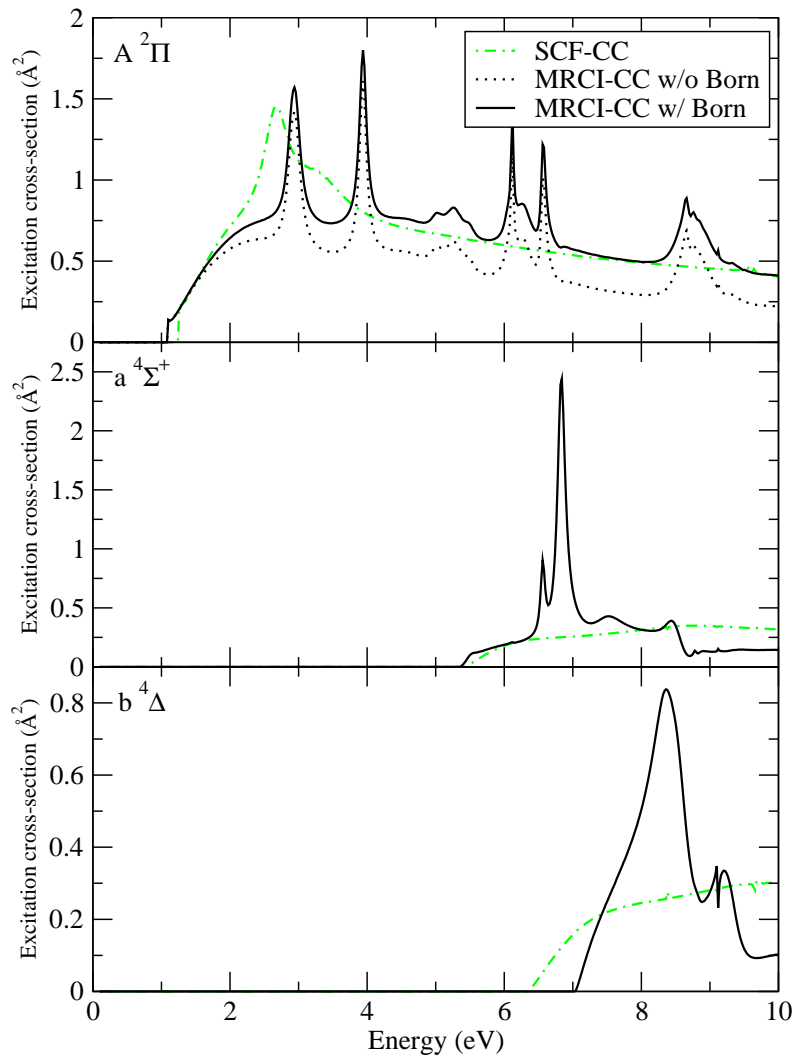


Figure 3. Electron impact electronic excitation cross sections for C_2H final state: $\text{A } ^2\Pi$, $\text{a } ^4\Sigma^+$ and $\text{b } ^4\Delta$.

4.2. Scattering calculations and bound anion states

Again a 48-state calculation was carried out (doublet and quartet states with 6 states per symmetry being included). The 40 target states below 10eV retained in the outer region. 46 in Quantemol-N jobs. As for C_2H^- the bound state calculations retained only the $\text{X } ^2\Sigma^+$ and $\text{A } ^2\Pi$ target states, and were propagated to $30.1 a_0$.

Table 5 presents the calculated electronic bound states for all models, as well as previous results from literature. All models find the deeply bound $\text{X } ^1\Sigma^+$ state, with CC calculations finding the state between 1 and 1.5 eV deeper, furthermore use of the MRCI-CC calculation bring the bound state even deeper to about 4 eV, closer in line with data from previous results in the literature. This model also finds numerous other weakly bound states of C_3N^- , presented in Table 6. It should be noted that model 3 (SCF-SE calculation using the cc-pVDZ basis set) also finds the $^1,^3\Pi$ bound state at 0.0002 eV, although this was not found in any other models that use SCF orbitals.

Table 5. C_3N^- bound states with energies given in eV. All results are for the equilibrium geometry of C_3N

Model	Basis	Orbitals	Method	X $^1\Sigma^+$
This work/1	DZP	SCF	SE	2.793
2	DZP	SCF	CC	3.896
3	cc-pVDZ	SCF	SE	2.248
4	cc-pVDZ	SCF	CC	3.788
5	cc-pVTZ	MRCI	CC	4.009
Botschwina & Oswald (2008)	VQZ+	RHF	RHF	3.651
	VQZ+	RHF	RCCSD	4.343
	VQZ+	RHF	RCCSD(T)	4.395
Kolos et al. (2008)	aug-cc-pCV5Z	RHF	RHF	3.650
	aug-cc-pCV5Z	RHF	RCCSD	4.363
	aug-cc-pCV5Z	RHF	RCCSD(T)	4.417
Zhan & Iwata (1996)	6-31G+(d)	HF	HF	2.519

Table 6. Binding energies, in eV, of the states of C_3N^- calculated for the MRCI-CC scheme (model 5).

	Singlet	Triplet
Σ^+		0.0012
	0.5490	0.5490
		0.6241
		0.6220
	4.0094	4.0094
Π	0.0003	0.0003

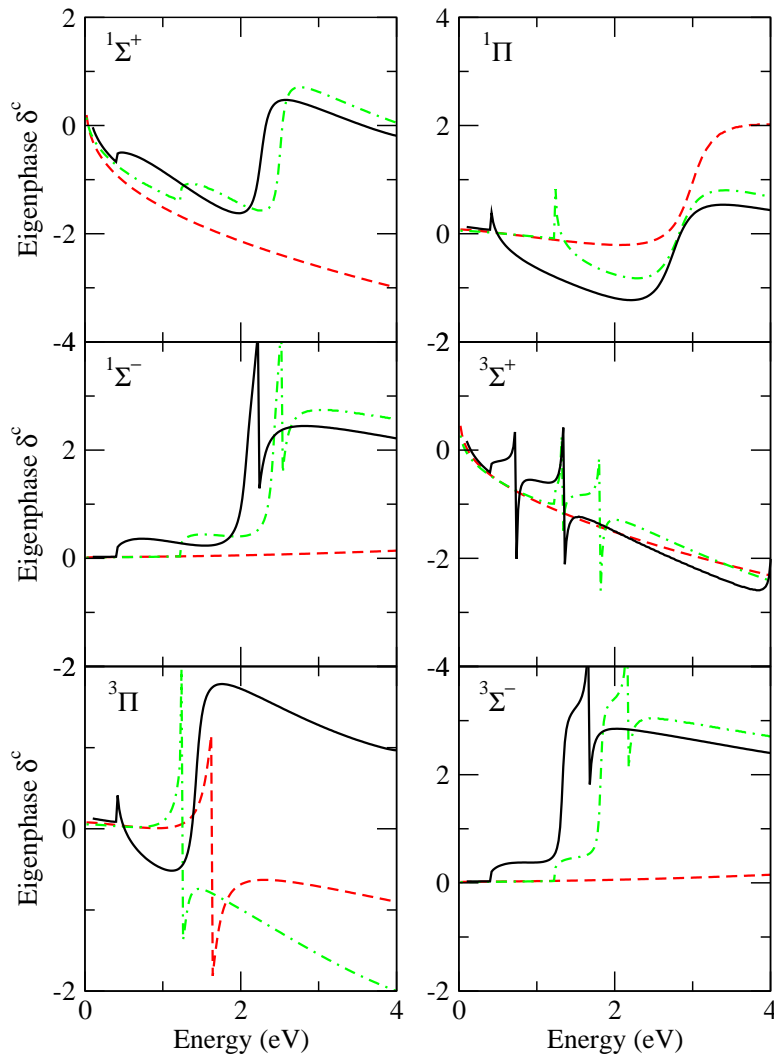
There appears to be no prediction of bound states other than the X $^1\Sigma^+$ states in the literature.

4.3. Resonances

Figure 4 presents the eigenphase sums for C_3N^- calculations. Table 7 summarizes the resonance data for C_3N^- calculations, with the same method being employed as previously described for the C_2H^- resonances. The SE models all find shape resonances of $^1\Pi$ and $^3\Pi$ symmetries. These appear in all CC calculations at about 0.1 eV lower in electron scattering energy, with the width remaining quite stable for the $^1\Pi$ resonance, and at about 0.2 – 0.4 eV lower scattering energy with a narrower width for the $^3\Pi$ resonance. In all CC calculations, Feshbach resonances appear in $^1\Sigma^-$, $^1\Delta$, $^3\Sigma^+$, $^3\Sigma^-$ and $^3\Delta$ symmetries. These resonances appear systematically at higher energies for the larger cc-pVTZ basis compared to cc-pVDZ when using SCF orbitals. However the use

Table 7. C_3N^- low-lying resonance positions (and widths) in eV

Model	1	2	3	4	5
$^1\Pi$	2.905 (0.471)	2.716 (0.452)	2.989 (0.444)	2.822 (0.475)	2.749 (0.437)
$^1\Sigma^-$		2.29 a		2.191 (0.158)	2.094 (0.153)
$^1\Delta$		2.307 (0.161)		2.248 (0.176)	2.167 (0.184)
$^3\Sigma^+$		0.987 (0.016)		1.327 (0.023)	0.741 (0.034)
$^3\Pi$	1.342 (0.204)	1.167 (0.121)	1.632 (0.220)	1.25 a	1.426 (0.141)
$^3\Sigma^-$		1.705 (0.052)		2.162 (0.062)	1.666 (0.075)
$^3\Delta$					0)

^a Resonan**Figure 4.** Electron – C_3N eigenphase sums: for SCF-SE(dash), SCF-CC(dash-dot) and MRCI-CC(solid) models.

of MRCI NO's leads to lower resonance positions than either SCF-CC models for all symmetries. The Feshbach resonance widths remain stable across all CC models.

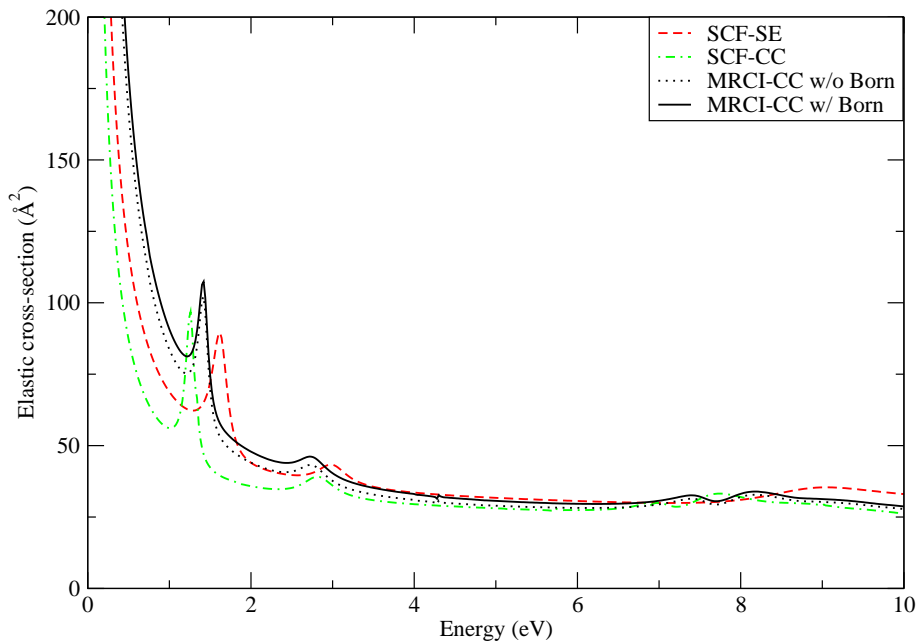


Figure 5. C_3N Elastic Cross-Section

4.4. Cross-sections

Figure 5 presents the total elastic cross-section of C_3N for electron scattering energies up to 10 eV, for three models, SCF-SE, SCF-CC and MRCI-CC. Born corrected results for the MRCI-CC model are also shown.

Figure 6 shows inelastic cross-sections representing the excitation from the $^2\Sigma^+$ ground state of the neutral to the first two excited states, both of which are $^2\Pi$; the Born correction is applied to both excitations. The comparison between the SCF-CC and the MRCI-CC models shows us that the use of MRCI orbitals increases the cross-sections at all energies, although both models share similar features (such as the threshold spike in B $^2\Pi$ excitation). Again the extra resonances shown by the MRCI-CC model should be treated with caution.

5. Conclusions

The presence of electronic resonance features above 1 eV for both molecules suggests that formation of temporary negative ions by electrons trapped in these resonance states is an unlikely mechanism for radiative electron attachment in the ISM. This is because 1 eV converts to a much higher temperature than electrons are found at in the ISM, making it highly unlikely electrons will reach the energies required to become temporarily trapped in the resonance states.

The more likely possibility is that the molecular anions observed may be arise from the very weakly bound anionic states, as found for C_3N^- . These states will support nuclear excited states which lie in the continuum and therefore are resonances. We

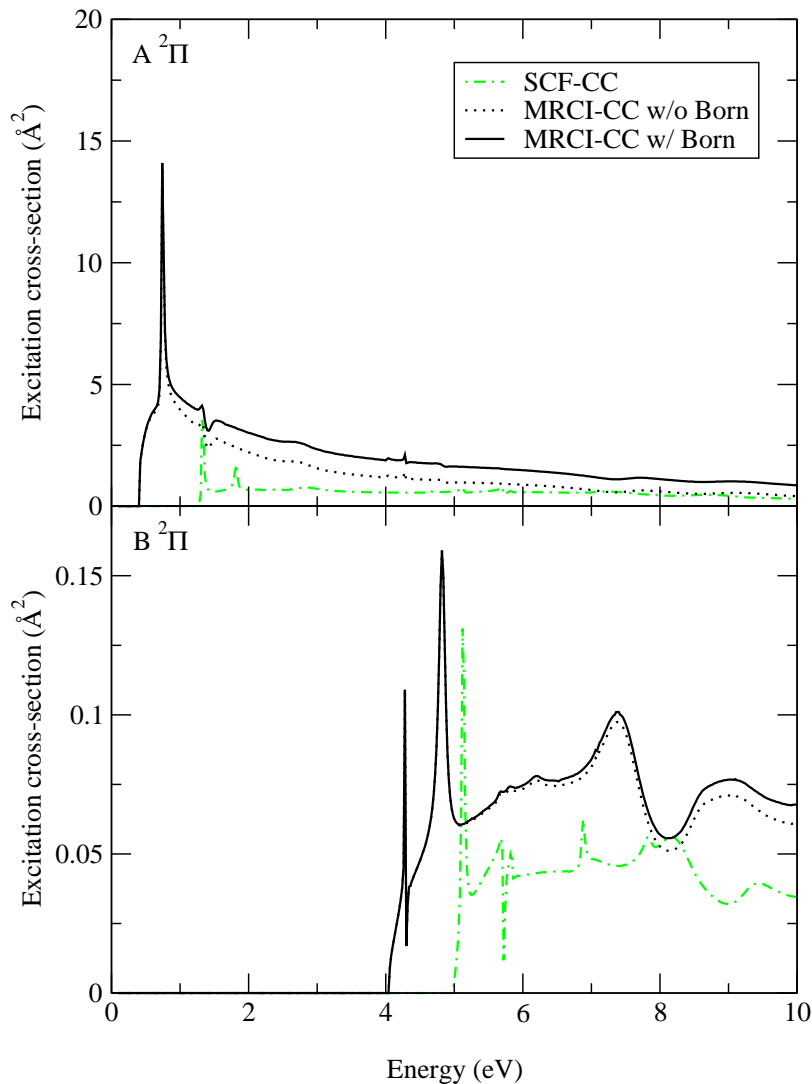


Figure 6. C_3N electron impact electronic excitation cross sections to $A^2\Pi$ and $B^2\Pi$.

note that C_2H^- appears not to support such states. If this mechanism does explain the formation of such anions, C_3N^- would be observed in large quantities but C_2H^- would not. This prediction appears to match current observations. We will investigate this proposal further in future work, both for the molecules in this paper, and by investigating other molecules in the C_nN^- and C_nH^- series.

Acknowledgements

The first author would like to thank R. Harrison and E. Brunetti for their support and STFC for a studentship.

References

- Agundez M, Cernicharo J, Guelin M, Kahane C, Roueff E, Klos J, Aoiz F J, Lique F, Marcelino N, Goicoechea J R, Garcia M G, Gottlieb C A, McCarthy M C & Thaddeus P 2010 *Astron. Astrophys. Lett.* **517**, L2.
- Baker J, Nobes R H & Radom L 1986 *J. Comp. Chem.* **7**, 349–358.
- Baluja K L, Mason N J, Morgan L A & Tennyson J 2000 *J. Phys. B: At. Mol. Opt. Phys.* **33**, L677–L684.
- Botschwina P & Oswald R 2008 *J. Chem. Phys.* **129**, 044305.
- Brunken S, Gupta H, Gottlieb C A, McCarthy M C & Thaddeus P 2007 *Astrophys. J.* **664**, L43–L46.
- Burke P G & Berrington K A, eds 1993 *Atomic and Molecular Processes, an R-matrix Approach* Institute of Physics Publishing Bristol.
- Cernicharo J, Guelin M, Agundez M, McCarthy M C & Thaddeus P 2008 *Astrophys. J.* **688**, L83–L86.
- Cui Q & Morokuma K 1998 *J. Chem. Phys.* **108**, 626–636.
- Ervin K M & Lineberger W C 1991 *J. Chem. Phys.* **95**, 1167.
- Faure A, Gorfinkiel J D, Morgan L A & Tennyson J 2002 *Computer Phys. Comms.* **144**, 224–241.
- Feher M 1992 *Chem. Phys. Lett.* **188**, 609–612.
- Fortenberry R C, King R A, Stanton J F & Crawford T D 2010 *J. Chem. Phys.* **132**.
- Gupta H, Brunken S, Tamassia F, Gottlieb C A, McCarthy M C & Thaddeus P 2007 *Astrophys. J.* **655**, L58–L60.
- Harada N & Herbst E 2008 *Astrophys. J.* **685**, 272–280.
- Herbst E & Osamura Y 2008 *Astrophys. J.* **679**, 1670–1679.
- Janousek B K, Brauman J I & Simons J 1979 *J. Chem. Phys.* **71**, 2057.
- Kolos R, Gronowski M & Botschwina P 2008 *J. Chem. Phys.* **128**, 154305.
- McCarthy M C, Gottlieb C A, Gupta H & Thaddeus P 2006 *Astrophys. J.* **652**, L141–L144.
- McCarthy M C, Gottlieb C A, Thaddeus P, Horn M & Botschwina P 1995 *J. Chem. Phys.* **103**, 7821–7827.
- Millar T J, Walsh C, Cordiner M A, Chumin R N & Herbst E 2007 *Astrophys. J.* **662**, L87–L90.
- Morgan L A, Tennyson J & Gillan C J 1998 *Computer Phys. Comms.* **114**, 120–128.
- Natterer J & Koch W 1995 *Molecular Physics* **84**, 691–706.
- Petrie S & Herbst E 1997 *Astrophys. J.* **491**, 210–215.
- Sadlej J & Roos B O 1991 *Chem. Phys. Lett.* **180**, 81–87.
- Sarpal B K, Branchett S E, Tennyson J & Morgan L A 1991 *J. Phys. B: At. Mol. Opt. Phys.* **24**, 3685–3699.
- Taylor T R & Xu C S 1998 *J. Chem. Phys.* **108**, 10018–10026.
- Tennyson J 1996 *J. Phys. B: At. Mol. Opt. Phys.* **29**, 6185–6201.
- Tennyson J 2010 *Phys. Rep.* **491**, 29–76.
- Tennyson J, Brown D B, Munro J J, Rozum I, Varambhia H N & Vinci N 2007 *J. Phys. Conf. Series* **86**, 012001.
- Tennyson J & Noble C J 1984 *Computer Phys. Comms.* **33**, 421–424.
- Terzieva R & Herbst E 2000 *International Journal of Mass Spectroscopy* **201**, 135–142.
- Thaddeus P, Gottlieb C A, Gupta H, Brunken S, McCarthy M C, Agundez M, Guelin M & Cernicharo J 2008 *Astrophys. J.* **677**, 1132–1139.
- Vuitton V, Lavvas P, Yelle R V, Galand M, Wellbrock A, Lewis G R, Coates A J & Wahlund J E 2009 *Planet. Space Sci.* **57**.
- Werner H J, Knowles P J, Lindh R, Manby F R, Schütz M, Celani P, Korona T, Mitrushenkov A, Rauhut G, Adler T B, Amos R D, Bernhardsson A, Berning A, Cooper D L, Deegan M J O, Dobbyn A J, Eckert F, Goll E, Hampel C, Hetzer G, Hrenar T, Knizia G, Köppl C, Liu Y, Lloyd A W, Mata R A, May A J, McNicholas S J, Meyer W, Mura M E, Nicklass A, Palmieri P, Pflüger K, Pitzer R, Reiher M, Schumann U, Stoll H, Stone A J, Tarroni R, Thorsteinsson T, Wang M & Wolf A 2008 ‘Molpro, version 2008.3, a package of ab initio programs’. see <http://www.molpro.net>.

Widmark P O, Malmqvist P A & Roos B O 1990 *Theor Chim Acta* **77**, 291–306.

Woon D E 1995 *Chem. Phys. Let.* **244**, 45 – 52.

Zhan C G & Iwata S 1996 *J. Chem. Phys.* **104**, 9058–9064.