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Implications for its evolution and recent reactivation**

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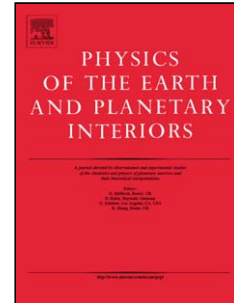
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1 **Shallow structure beneath the Central Volcanic Complex of Tenerife**
2 **from new gravity data: Implications for its evolution and recent**
3 **reactivation**

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17

18

Abstract

19 We present a new local Bouguer anomaly map of the Central Volcanic Complex
20 (CVC) of Tenerife, Spain, constructed from the amalgamation of 323 new high
21 precision gravity measurements with existing gravity data from 361 observations. The
22 new anomaly map images the high-density core of the CVC and the pronounced
23 gravity low centred in the Las Cañadas caldera in greater detail than previously
24 available. Mathematical construction of a subsurface model from the local anomaly
25 data, employing a 3-D inversion based on “growing” the subsurface density
26 distribution via the aggregation of cells, enables mapping of the shallow structure
27 beneath the complex, giving unprecedented insights into the sub-surface architecture.
28 We find the resultant density distribution in agreement with geological and other
29 geophysical data. The modelled subsurface structure supports a vertical collapse
30 origin of the caldera, and maps the headwall of the ca. 180ka Icod landslide, which
31 appears to lie buried beneath the Pico Viejo – Pico Teide stratovolcanic complex. The
32 results allow us to put into context the recorded ground deformation and gravity
33 changes at the CVC during its reactivation in spring 2004 in relation to its dominant
34 structural building blocks. For example, the areas undergoing the most significant
35 changes at depth in recent years are underlain by low-density material and are aligned
36 along long-standing structural entities, which have shaped this volcanic ocean island
37 over the past few million years.

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44 **1. Introduction**

45

46 The Central Volcanic Complex (CVC) of Tenerife (Canary Islands, Spain)), which
47 includes the 16 x 9 km wide Las Cañadas caldera (LCC, Figure 1), reactivated after
48 an almost century-long period of quiescence in spring 2004, with increased seismicity
49 (including felt earthquakes) and subsurface mass addition into its north-western and
50 western portions (García et al., 2006; Tárraga et al., 2006; Gottsmann et al., 2006;
51 Almendros et al., 2007). It has been speculated that earthquake locations and density
52 changes during the reactivation may be controlled by the internal structure of the
53 CVC (Martí et al., 2008a). However, the available geophysical information on its sub-
54 surface architecture is mainly based on rather crude surveys highlighting large
55 wavelength anomalies such as from aero-magnetic data (Araña et al., 2000; García et
56 al., 2007) or gravimetric data with substantial uncertainties (Araña et al., 2000). A
57 detailed assessment of structural controls on the spatio-temporal variations in
58 geophysical parameters during the recent reactivation thus remains ambiguous. At
59 other collapse calderas (e.g. Long Valley, Valles, Toba, Campi Flegrei) geophysical
60 imaging has provided important insights into their internal architecture including the
61 identification of potential magma and hydrothermal reservoirs (e.g. Sanders et al.,
62 1995; Aprea et al., 2002 Guidarelli et al., 2002; Masturyono et al., 2001). Such
63 reservoirs may play an important role during post-collapse processes, including
64 episodes of unrest (see Martí et al., 2008b for a recent review). In light of the recent
65 unrest, knowledge about the shallow subsurface beneath the CVC is of great
66 importance, and provided the motivation for this study. This work also aims to
67 contribute to the discussion as to the origin of the Las Cañadas Caldera (LCC): lateral

68 collapse due to edifice instability (Cantagrel et al., 1999) or vertical collapse due to
69 explosive volcanism (Martí and Gudmundsson, 2000) having been proposed.

70 We present results from a new series of gravity measurements performed at the CVC
71 over the past few years, which, combined with existing gravimetric data, enable us to
72 image its shallow sub-surface density distribution at a higher resolution than
73 previously possible.

74

75 **2. Previous geophysical work**

76

77 While the surface geology of the CVC (Fig. 1) has been a target in the past (e.g.
78 Martí et al., 1994; Ablay et al., 1998; Ablay and Martí, 2000), the complex's internal
79 structure remains rather enigmatic. Previous geophysical studies have focused on
80 obtaining sub-surface images based on large wavelength ($\gg 1$ km) anomalies
81 (MacFarlane and Ridley, 1968; Ablay and Kearey, 2000; Watts et al., 1997; Araña et
82 al., 2000) but have not provided information on the shallow structure beneath the
83 CVC, including the caldera itself. Ablay and Kearey (2000), provided a coherent
84 interpretation of the magmatic and structural evolution of Tenerife from gravity data
85 predominantly collected in the 1980s and 90s by members of the Instituto de
86 Astronomía (CSIC-UCM) and the Instituto Geográfico Nacional (Camacho et al.,
87 1991; Araña et al., 2000). Ablay and Kearey's emphasis lay in the integration of a rich
88 petrological data set with the existing gravimetric data. The original gravity data
89 suffered from significant uncertainties stemming from both terrain correction methods
90 and control of benchmark elevation using topographic maps and barometers, which
91 resulted in uncertainties in the Bouguer anomalies on the order of 5 mGal (1
92 $\text{mGal} = 10^{-5} \text{m/s}^2$) (Ablay and Kearey, 2000). One of the first high-resolution
93 geophysical images to a depth of approximately 1500 m beneath the LCC floor was

94 provided by the audiomagnetotelluric study of Coppo et al. (2008), which enabled the
95 identification of three adjacent bowl-shaped depressions consistent with a multiple-
96 event vertical collapse origin of the LCC.

97

98 **3. Geological background**

99

100

101 The geological and tectonic evolution of the triangular island of Tenerife and the

102 CVC was described in recent papers (Martí et al., 1994; Martí et al., 1998; Araña et

103 al., 2000) and we therefore only provide a short summary. The Canary Islands form a

104 volcanic archipelago with a long-standing history of volcanic activity that began more

105 than 40 million years ago (Araña and Ortíz, 1991). More than a dozen eruptions have

106 occurred on the islands of Tenerife, Lanzarote, and La Palma since the 16th century.

107 Tenerife, the largest of the Canary Islands, has an eruptive history of over 12 million

108 years, including a shield building phase and the construction of a central volcanic

109 structure, the Las Cañadas edifice (LCE) from 3.5 Ma onwards (Martí et al., 1994).

110 Table 1 provides a stratigraphic scheme of the island, and highlights the main

111 constructive and destructive episodes during its formation. The LCE is composed of

112 a mafic post-shield Lower Group (>3.5 Ma – 1.57 Ma), overlain by the Upper Group

113 comprising three felsic formations (1.57 Ma -0.179 Ma).

114 The evolution of the LCE comprises both constructive and destructive phases,

115 including vertical and lateral collapses with volumes on the orders of several to tens

116 of km³ (Martí et al., 1997). Martí et al. (1994) propose at least three vertical collapses

117 during Upper Group times, which resulted in the formation of the LCC (Table 1).

118 High walls nowadays bound the LCC to the southwest, south, east and northeast.

119 Over the past 170-190 ka, the prominent Pico Viejo - Pico Teide (PV—PT) volcanic
120 complex was emplaced inside the caldera depression during predominantly effusive
121 but also explosive activity (Ablay and Martí, 2000; Edgar et al., 2007). The PV-PT
122 complex appears to be fed by both shallow-level (< 5 km) phonolitic magma
123 reservoirs and deeper-seated basaltic magma patches (Ablay and Hürlimann, 2000;
124 Ablay and Martí, 2000). Recent (< 2 ka) volcanic activity was located both on the
125 PV—PT complex (explosive and effusive phonolitic eruptions) and along two
126 extensional structural lineaments referred to as the NW—SE oriented Santiago Rift,
127 and the SW-NE striking Dorsal Ridge (both rifts are dominated by monogenetic mafic
128 eruptions). Historic eruptions at these centres were fed by mafic magmas and
129 occurred in 1704, 1706, 1798, and 1909. To the north of the CVC lie the valleys of
130 Icod and Orotava, which are interpreted to represent scars from large lateral collapses
131 (landslides). The absence of a visible caldera wall to the North has led some workers
132 to infer a lateral collapse origin of the LCC (Cantagrel et al., 1999; Ancochea et al.,
133 1999). Others see clear evidence for vertical collapse origin of the LCC based on the
134 abundance and nature of pyroclastic deposits consistent with explosive caldera
135 formation (Martí et al., 1998; Bryan et al., 1998; Ablay and Hürlimann, 2000; Martí
136 and Gudmundsson, 2000; Brown and Branney, 2004).

137

138 **4. Gravity data**

139 **a) New gravity survey**

140

141 All 323 new gravity measurements were performed by the same operator (JG)
142 at the CVC between May 2004 and November 2006 using LaCoste&Romberg
143 gravimeter G-403 (Fig. 1). The accuracy of individual readings was maintained

144 through regular reoccupation of each station and the reference station. The reference
145 was conveniently located in the caldera depression at 340088.27m Easting and
146 3124255.73m Northing, at an ellipsoidal elevation of 2223.83 m. The reference was
147 occupied between 4 and 8 times per day depending on the design of measurement
148 loops to check for instrument drift and tares. At each benchmark the reading was
149 derived by the averaging of between 5-10 manual readings taken over a few minutes,
150 depending on benchmark stability and ambient noise. The error on the raw readings is
151 less than 0.02 mGal. Location and elevation data were provided by differential GPS
152 measurements (operated by LW), conducted at the same time as the gravity
153 measurements. Leica System 530 receivers and AT502 antennas were used, with data
154 sampling at 1 Hz. A reference station was established (with matching occupation
155 frequency of 1 Hz running for a total of more than 250 hrs), for which absolute
156 WGS84 co-ordinates were derived. A rover GPS receiver, mounted on a 1.9 m-high
157 staff, was placed at the measurement point and recorded for periods of between 60
158 and 180 seconds, depending upon satellite visibility and the distance to the reference.
159 Post-processing of the GPS data was carried out using Leica Geosystems' Ski-Pro
160 software. 2σ errors for most positions were generally under 5 cm in the Z axis and
161 better than 4 cm in the X and Y axes. Fig. 2 shows the location of the new
162 gravity/GPS benchmarks which provide coverage of the LCC, the PV-PT complex
163 and the Santiago Rift. An additional 361 gravity data (from a total set of 975 covering
164 the entire island) obtained earlier are used for investigating the wider area surrounding
165 the CVC (Camacho et al., 1991, Araña et al., 2000). These earlier gravity readings
166 are reported with an uncertainty of 1.2 mGal. The significantly larger error in the
167 earlier gravity data compared to the new data is due to uncertainties in benchmark
168 elevation.

169

170 **b. Terrain density**

171 The terrain density is among the most critical parameters for gravimetric data
172 reduction and choosing the ‘right’, i.e., realistic, terrain density for a whole island
173 (with bimodal magmatism/volcanism spanning several million years and the
174 associated variety in eruptive products) is prone to errors. Terrain effects are
175 significant on Tenerife as the CVC represents the top of a volcanic edifice rising to
176 more than 3.7 km elevation within 12 km from the northern shoreline, and extending
177 more than 3.5 km below sea level. Despite these complexities, it is possible to deduce
178 an optimal value for terrain density from both gravity and elevation data. According
179 to a general assumption (Nettleton, 1939), a suitable density value must produce a
180 minimum correlation between gravity anomaly (Δg) and elevation (h). However, this
181 approach can lead to erroneous density estimates, for example if long wavelength
182 components of both gravity and topography are correlated due to deep structural
183 features. The general terrain on Tenerife follows a convex geometry common to all
184 volcanic islands related to the intrusive and extrusive nature of ocean island building,
185 and thus the gravity anomaly will also follow this pattern. A more realistic value for
186 the terrain density can be obtained by looking for the minimum correlation between
187 the shortest wavelength components of both gravity and topography.

188 For each benchmark P_i , we therefore calculate the short wavelength values $\Delta \tilde{g}_i$ and
189 $\Delta \tilde{h}_i$ for the gravity anomaly and elevation, respectively, by removing regional values
190 determined as mean values between P_i and surrounding benchmarks P_j up to a radius
191 r :

192

$$193 \quad \Delta\tilde{g}_i = \Delta g_i - \frac{1}{n_i} \sum_J \Delta g_j, \quad \Delta\tilde{h}_i = \Delta h_i - \frac{1}{n_i} \sum_J \Delta h_j$$

194

195 where distance $(P_i, P_j) \leq r$.

196 The optimal theoretical terrain density ρ_t corresponds to the minimum correlation
 197 between the resulting $\Delta\tilde{g}_i$ and $\Delta\tilde{h}_i$ values. Fig. 3 gives terrain densities (kg/m^3) as a
 198 function of r . However, we choose a terrain density accounting for both theoretical
 199 constraints and measured densities of exposed lithologies of the CVC. Ablay and
 200 Kearey (2000) report more than 30 measured bulk densities of Tenerife rock types.
 201 Calculating the mean values for their reported rock types (except for Teno and Dorsal
 202 ridge basalts) and weighting the occurrence of effusive and flow deposits against fall
 203 deposits (0.8 vs. 0.2, respectively), to account for exposed CVC lithologies (Table 2),
 204 we derive $2200 \pm 271 \text{ kg/m}^3$. The mean value is consistent with our theoretical
 205 minimum correlation for r of about 3 km, providing a reasonable wavelength for
 206 topographic and gravity anomalies. We thus use a terrain density of 2200 kg/m^3 for
 207 our gravity reduction.

208

209 **c) Gravity data reduction**

210

211 Standard techniques for gravimetric data reduction for the effect of Earth and
 212 ocean tides, latitude and benchmark elevation were employed. Ocean loading effects
 213 were corrected using the Schwiderski global ocean tide model (Schwiderski, 1980).
 214 Continuous gravity observation, performed by an automated Burris gravity meter (B-
 215 28) in 2007 over a period of 8 days in the Las Cañadas caldera, revealed that residual
 216 ocean loading effects amount to less than 0.01 mGal and are thus negligible for our
 217 purpose.

218 The earlier gravity data were referenced to the absolute gravity reference
219 IGSN71 (Morelli et al., 1974). In order to merge the data sets, we determined two sets
220 of Bouguer anomalies, one using solely the new data and the second using only the
221 earlier data. We then deduced an offset value, which produced a maximum
222 autocorrelation between the two sets of anomaly data. This gravity offset was then
223 used to combine the two data sets to form a single coherent set of gravity data. This
224 technique involving global overlapping of individual data is superior to techniques
225 based on matching few common benchmarks to calculate an offset value, as local
226 inaccuracies could be propagated, resulting in an erroneous final data set.

227 Free-air and Bouguer corrections, respectively, were performed with respect to
228 theoretical predictions based on GRS80 [Geodetic Reference System 1980] using a
229 standard free-air gradient of -0.308 mGal/m and the above derived terrain density of
230 2200 kg/m³ for the Bouguer correction

231

232 d) Terrain correction

233

234 To correct for the topographic effects (on-shore Tenerife, neighboring islands
235 and off-shore areas) on gravity measurements, we employed an automated algorithm
236 based on dense circular zones (34 zones from 1 m to 100 km radius, 894
237 compartments in total), similar to Hammer zones, around the individual benchmarks
238 (Hammer, 1939). A 10 m digital elevation model (DEM), constructed from
239 topographic maps provided by the Instituto Geografico Nacional (IGN), approximated
240 the immediate on-shore topography up to 895 m distance from each gravity point. For
241 on-shore distances greater than 895 m, we used a 100 m DEM. For the surrounding
242 marine areas, satellite altimetry data (http://topex.ucsd.edu/cgi-bin/get_data.cgi and

243 local charts were used to provide information about the local geoid and bathymetry
244 (Smith and Sandwell, 1997). Similar to the Hammer reduction procedure, average
245 elevations were calculated for each compartment and the terrain effect in each
246 compartment was propagated to obtain the final terrain correction value for each
247 benchmark.

248 The so derived terrain corrections for the gravity benchmarks vary between 8.8
249 mGal and 97.7 mGal, with an average of 19.3 mGal. The maximum correction applies
250 to a benchmark at the base of Teide's summit cone. Possible errors in the choice of
251 the terrain density ρ_t are accounted for in the modelling process. For this, an
252 additional unknown $\delta\rho_t$, which represents a correction to the initially assumed terrain
253 density, is included in the inversion routine as described below.

254

255 **e) Regional gravity trend**

256 A very long-wavelength component of gravity variations can be produced by
257 crustal structures defining a regional trend in the resultant gravity data. This effect
258 needs to be accounted for in creating a local Bouguer anomaly map of the CVC. We
259 used Geosat and ERS 1 satellite altimetry and gravity data ([http://topex.ucsd.edu/cgi-](http://topex.ucsd.edu/cgi-bin/get_data.cgi)
260 [bin/get_data.cgi](http://topex.ucsd.edu/cgi-bin/get_data.cgi); (Sandwell and Smith, 1997)) to quantify the regional gravity trend
261 considering an area up to a distance of 100 km around Tenerife and 3 km radial
262 spacing between individual nodes.

263 Using the free-air satellite anomaly data combined with bathymetry and geoid
264 heights, we determined the regional Bouguer anomaly taking 2900 kg/m^3 as the mean
265 background density for oceanic crust (Carlson and Raskin, 1984). We determined a
266 smoothed linear trend of $0.27 \pm 0.03 \text{ mGal/km}$ with azimuth $N113^\circ E$, which represents
267 the very long wavelength component superimposed on the local anomaly. It is in

268 reasonable agreement with a general average value of 0.2 mGal/km obtained by
269 Bosshard and MacFarlane (1970) extending from west of La Palma and El Hierro to
270 Gran Canaria. This trend is believed to be due to crustal thickening towards the
271 African continent.

272 Subtracting the regional trend from the data, we obtain the local Bouguer
273 anomaly (Fig. 4), which highlights local sub-surface structures. These Bouguer
274 anomaly data are then employed for the mathematical inversion.

275

276

277 **5. Gravity inversion**

278 **a. Theory and sensitivity**

279 The inverse gravimetric problem, namely the determination of a sub-surface
280 density distribution consistent with an observed gravity anomaly, has an intrinsically
281 non-unique solution (e.g. Al-Chalabi, 1971). Moreover, the available data are always
282 insufficient and inaccurate to resolve ambiguities. One can, however, obtain realistic
283 solutions by including additional constraints on model parameters (subsurface
284 structure) and data parameters (statistical properties of inaccurate data, e.g. Gaussian
285 distribution). The non-linear inversion methods that delineate the geometrical
286 properties of anomalous bodies with prescribed density contrasts (e.g. Pedersen, 1979;
287 Barbosa et al., 1997) are of course limited by the underlying hypothesis on source
288 density. However, these techniques provide results worth exploring. For full non-
289 linear treatments, methods based on the exploration of model possibilities often give
290 the best results (Tarantola, 1988). This exploration process can for example be
291 conducted randomly (Silva and Hohmann, 1983).

292 In this study we use the inversion routine presented in Camacho et al. (2000 and
 293 2002), which has also been applied in other gravimetric studies (Camacho et al., 2001,
 294 2007). The inversion process constructs a subsurface model defined by a 3-D
 295 aggregation of M parallelepiped cells, which are filled, in a “growth” process, by means
 296 of prescribed positive and/or negative density contrasts. The design equation to relate
 297 observables, i.e., the gravity anomaly Δg_i at N benchmarks (x_i, y_i, z_i) , with modelling
 298 parameters, and residuals v_i is:

$$299 \quad \Delta g_i = \sum_{j \in J^+} A_{ij} \Delta \rho_j^+ + \sum_{j \in J^-} A_{ij} \Delta \rho_j^- + \delta g_{reg} + \delta g_{top} + v_i \quad , \quad i = 1, \dots, N \quad ,$$

300 (1)

301 where A_{ij} is the vertical attraction for unit density for the j -th parallelepiped cell upon
 302 the i -th observation point, $\Delta \rho_j^-$, $\Delta \rho_j^+$ are prescribed density contrasts (negative and
 303 positive) for the j -th cell, J^+ , J^- are sets of indexes corresponding to the cells filled
 304 with positive or negative density values, and δg_{reg} , δg_{top} are optional terms for
 305 regional trend and additional topographical correction given by

$$306 \quad \delta g_{reg} = p_0 + p_x(x_i - x_M) + p_y(y_i - y_M) \quad , \quad i = 1, \dots, N \quad , \quad (2)$$

$$307 \quad \delta g_{top} = \delta \rho_T C_i \quad . \quad (3)$$

308 Here p_0 , p_x , p_y are parameters for the linear trend, x_M, y_M are averaged coordinates, C_i
 309 are coefficients for the terrain correction and $\delta \rho_T$ is an optional additional value with
 310 respect to the initially adopted terrain density. In Eqns. 1-3, J^+ , J^- (cells filled with
 311 positive and negative density contrast), p_0 , p_x , p_y , and $\delta \rho_T$ (regional parameters) are
 312 the main unknowns to be determined by the inversion.

313 The problem of non-uniqueness is usually approached by numerical
 314 techniques in the form of a generalised matrix inversion including some form of

315 “perturbation” scheme such as the Levenberg-Marquardt method (e.g. Enmark, 1981;
316 García-Abdeslem, 2000; Marinara and Hall, 2001). In short, the “perturbation” idea of
317 the Levenberg-Marquardt method (and other similar methods) is a numerical
318 approach that involves modification of the usual least square solution for the case of
319 singularity or non-uniqueness, by an “ad-hoc” term or perturbation ($\lambda \mathbf{I}$; where \mathbf{I} is an
320 identity matrix), which may transform a singularity into non-singularity. Such
321 solutions aim at minimising an objective function that is a combination of the l_2 -norm
322 of the residuals (“least squares fit”) and the l_2 -norm of the model parameters, using a
323 positive value of λ for balance (the so-called damping factor or Lagrange multiplier).
324 The methodology followed in our study can be considered as a generalisation of this
325 numerical approach as outlined by Tarantola (1988). This author gives a general
326 treatment of least-squares inversion methods by including them into a general theory,
327 which represents the inversion problem as a problem of combination of inaccurate
328 information. For instance, the least-squares inversion approach is presented as the
329 combination of observable data with Gaussian uncertainties (given by a covariance
330 matrix \mathbf{Q}_D) and information on the model, given by a previous model \mathbf{m}_{prev} which is
331 also subject to Gaussian uncertainties (given by a covariance matrix \mathbf{Q}_M). For the
332 case where the previous model is exactly zero ($\mathbf{m}_{prev}=\mathbf{0}$, no particular previous
333 information on the model is available) the formulae of the general methodology of
334 Tarantola are close to the numerical methods pointed out before. For the particular
335 case of gravity inversions aiming at deriving a unique model of density distribution at
336 depth, the general condition of l_2 -minimisation of the model parameters becomes a
337 minimisation condition for the total anomalous mass; i.e. a model involving
338 anomalous bodies with simple and smoothed geometries *and* a minimum total mass is
339 uniquely determined. Accordingly, in order to get unique results from our data

340 inversion, we adopt a mixed minimisation condition, based on model “fitness” (least
 341 square fitness) and “smoothness” (total anomalous mass) (see also Camacho et al.,
 342 2002):

$$343 \quad \mathbf{v}^T \mathbf{Q}_D^{-1} \mathbf{v} + \lambda \mathbf{m}^T \mathbf{Q}_M^{-1} \mathbf{m} = \min, \quad (4)$$

344 where $\mathbf{m} = (\Delta\rho_1, \dots, \Delta\rho_M)^T$ (superscript T denotes transpose of a matrix) are density
 345 contrast values for the M cells of the model, $\mathbf{v} = (v_1, \dots, v_N)^T$ are residual values for the
 346 N data points, \mathbf{Q}_D is an apriori covariance matrix for uncertainties of the gravity data,
 347 \mathbf{Q}_M is an apriori covariance matrix for uncertainties of the model parameters, and λ is
 348 a factor for selected balance fitness/smoothness of the model. For practical
 349 applications λ is selected to produce uncorrelated inversion residuals. In traditional
 350 numerical inversion methods based on the Levenberg-Marquardt smoothing approach,
 351 an identity matrix \mathbf{I} is used in place of \mathbf{Q}_M . However, taking into account that the
 352 cells of the model have different sensitivities (lower for deep or peripheral cells; see
 353 below), using a covariance matrix to quantify cell sensitivity based on its geometry
 354 and location, will produce more balanced and coherent minimization conditions for
 355 the model parameters, and will thus produce more consistent solutions.

356 In interpreting the inversion results a few important points need to be
 357 considered. Although the inversion routine counteracts the problem of non-uniqueness
 358 of results via the mixed minimum condition as explained above, there is still some
 359 degree of ambiguity in the results. Despite the complexity of the inversion, the
 360 resultant models should not be regarded as an exact replication of the sub-surface
 361 architecture, as both data collection and post-processing suffer from inaccuracies and
 362 assumptions. The modelled density contrasts should be seen as mean values for a
 363 particular spatial distribution of cells in the form of a smoothed model. The sensitivity

364 to the data is dependent on the position of each cell within each the model. The
365 sensitivity of very shallow, very deep and very peripheral cells is lower than those
366 located beneath the survey area at depths of about 2 km below the surface. During
367 inversion the size of the cells is variable to account for differences in sensitivity: very
368 deep cells or very peripheral cells are larger then those cells located at shallow depth.
369 Figure 5 shows results from a sensitivity test of the inversion for the particular case of
370 the Tenerife study. We simulate an S-shaped anomalous body with a prescribed
371 density contrast of 800 kg/m^3 located beneath the CVC. Using the same gravity
372 benchmarks employed for the inversion as described in the next section, we calculate
373 the resultant gravity value at each benchmark and then use these data for inversion.
374 The inversion model gives an adequate representation of the simulated body in terms
375 of its geometry. Distortions in the shape of the modelled body with respect to the
376 simulated body are apparent for greater depths as well as in areas of insufficient
377 gravity data. The density contrast is found to be ca. 400 kg/m^3 (i.e., 50% of the
378 prescribed contrast), which represents the value consistent with the minimum total
379 anomalous mass necessary to match the observed Bouguer anomaly (see previous
380 section on inversion theory). In general, while the shape and distribution of the
381 modelled anomalous bodies are indicative of their real features as shown by the above
382 sensitivity analysis and also by the examples given by Camacho et al. (2002), due to
383 the tendency of the methodology to produce models involving a minimum total
384 anomalous mass, the density contrasts are likely to be stronger in reality. It is
385 therefore impossible to attribute a finite “real” value to any density contrast.

386

387 **b. Inversion of data**

388

389 The operational method for solving the design system is a controlled 3-D
 390 growth of anomalous bodies by means of an exploratory approach, subject to the fit
 391 and smoothness conditions explained in the previous section.

392 From the full data set of 361 existing and 323 new gravity measurements we
 393 selected 392 data (Figs. 1 and 2) matching the following conditions: (i) a minimum
 394 distance between neighbouring benchmarks of equal or larger than 200 m for the new
 395 data (core of the data set) and (ii) equal or larger than 1000 m for the earlier data to
 396 account for gravimetric data obtained at the periphery of and outside the CVC.
 397 Outliers in the older gravity data were removed. The Bouguer anomaly value Δg_i at
 398 each benchmark ($i= 1, 2, 3, \dots, 392$) and its coordinates were fed into the automated
 399 inversion. In a first step, the sub-surface volume was partitioned into 16850
 400 parallelepiped cells. The cells had sides ranging from 330 m in the shallowest zone (3
 401 km a.s.l.) to 2023 m at a depth of 24 km b.s.l., with an average side of 700 m. Matrix
 402 \mathbf{Q}_D for the *a priori* covariance of gravity data was considered to be a diagonal matrix
 403 corresponding to an assumed standard deviation of ± 0.4 mGal for the new data and
 404 ± 1.2 mGal for the earlier gravity data. Matrix \mathbf{Q}_M for *a priori* covariance of the cell
 405 distribution geometry was taken as a diagonal normalizing matrix of non-zero elements
 406 that matched the diagonal elements of $\mathbf{A}^T \mathbf{Q}_D^{-1} \mathbf{A}$, where \mathbf{A} is the design matrix with
 407 elements A_{ij} . In a second step, we fitted a rough model based on prescribed (positive and
 408 negative) density values in the range between -250 kg/m^3 and 230 kg/m^3 . Including the
 409 smoothing effect, the resultant anomalous density values for the model are -209 kg/m^3
 410 and 184 kg/m^3 . The value λ is automatically determined by assuming the condition of
 411 zero autocovariance for the final residuals.

412 The former model is obtained without any particular assumption towards the
 413 physical properties of the surrounding non-anomalous medium. The anomalous

414 densities need to be added to some non-anomalous medium density to give the final
415 finite subsurface density distribution. Assuming a homogeneous non-anomalous
416 medium, the final model will be similar to the adjusted model. However, the bimodal
417 evolution of Tenerife, with a basic oceanic shield-building phase succeeded by an
418 evolved stratovolcano building phase warrants the assumption of a significant density
419 contrast between shallower and deeper portions of the island and the crust. As a
420 consequence, large positive density contrasts ($>250 \text{ kg/m}^3$) at greater depth, may not
421 be anomalous but rather reflect vertical and horizontal discontinuities with strong
422 density contrasts with respect to shallower depths. A way to accommodate density
423 stratification with increasing depth is to assume a stratified background medium
424 during inversion. In doing so, we followed the hypothesis of a general background
425 stratification, defined by a continuous exponential law for density increase with
426 respect to depth z : $\Delta\rho = \rho_0 + a \exp(-b z)$, where ρ_0 is 2200 kg/m^3 , a is 580 kg/m^3 and
427 b is $54 \times 10^{-6} \text{ m}^{-1}$. Values for a and b were calculated such as to avoid background
428 density inversion; i.e., high density material overlying low density material in the
429 model. We chose an exponential density increase rather than a linear increase to
430 account for the fact that most of the interior of Tenerife is composed of mafic
431 volcanics, while evolved (phonolitic) material was only erupted with the construction
432 of both the Las Cañadas edifice from 3.5 Ma onwards and the PV-PT complex (from
433 179 ka onwards). We estimate a minimum volume of Tenerife of 15000 km^3 , with
434 most of this volume associated with the mafic shield-building phase (ca. 90%). The
435 remaining 10 % of volume includes the Las Cañadas edifice and the PV-PT complex.
436 The maximum estimate for the volume of phonolitic material recognisable at the
437 surface is $250\text{-}300 \text{ km}^3$, which mostly corresponds to the Las Cañadas phonolites. We

438 thus assume a maximum proportion of 2% of evolved (phonolitic) as opposed to 98%
439 of mafic material.

440 To produce quasi-homogeneous anomalous structures, we applied the same
441 inversion process as explained above, but account for an anomalous positive density
442 contrast decrease as a function of z : $\Delta\rho = \rho_0 - a \exp(-b z)$. The adjusted inversion
443 model has a “bell” shape with anomalous positive density decreasing, while the
444 background medium shows an exponentially increasing density with depth, and the
445 addition of both results in quasi-homogeneous structures embedded in a stratified
446 medium.

447 While we discuss results from both homogenous and stratified background
448 models, we choose to restrict our interpretations of the inversion results to those
449 obtained from models based on a density-stratification of the background medium.
450 We feel that, despite our simplified stratification, these models provide a more
451 realistic account of the “real” density distribution beneath the CVC at greater depth
452 compared to a homogeneous medium. The modelled shallow structures to depths of
453 about 1000 m bsl differ insignificantly in either model, providing a sound basis for the
454 interpretation of the immediate sub-surface, while accounting for stratification may
455 more realistically image density variations at greater depth. A general value of 41
456 kg/m^3 can be assumed as the mean sensitivity of the cells in either model.

457

458 **6. Results**

459

460 **a. Bouguer and local anomaly**

461

462 Fig. 2 shows the Bouguer gravity anomaly map. The anomaly is centred
463 around an average value of 256 mGal with a standard deviation of 25 mGal. The
464 central volcanic complex is dominated by higher than average values, which produce
465 a local positive anomaly (Fig. 4) of more than 45 mGal centred at the western edge of
466 the Las Cañadas caldera, an area known as Boca Tauce (see Fig. 1C and Table 1). A 5
467 km-wide curvilinear gravity high follows the margin of the LCC wall and connects
468 Boca Tauce with La Fortaleza at the eastern boundary of the LCC. The Boca Tauce
469 gravity high correlates well with a pronounced magnetic high in the same area (et al.,
470 2007). In general, the gravity highs of the CVC correlate well with observed magnetic
471 highs extending from Boca Tauce along the LCC towards La Fortaleza (García et al.,
472 2007), as well as with a high-velocity zone (5-6 km/s) located south of the PV-PT
473 complex (Araña et al., 2000; Watts et al., 1997). The relative gravity low over PT
474 (and extending northwards) correlates with a low velocity zone (3-5 km/s) and a
475 magnetic low (García et al., 2007). Somewhat disconnected from the central high, the
476 Teno massif shows an isolated gravity high with values similar to those observed at
477 the intersection between the Dorsal ridge and the LCC. The overall impression is that
478 the CVC, and Tenerife in general, are formed by a high density core with a
479 pronounced gravity low centred in the LCC beneath Pico Teide. This area is
480 dominated by the strongest horizontal gravity gradient of about 8 mGal/km and a
481 N45°E azimuth. The northern boundary of the Boca Tauce high coincides with a
482 break in topographic slope between the Ucanca plain and the start of the Icod valley.
483 If one were to project the north-westernmost part of the LCC wall towards the PV-PT
484 complex, the imaginary line would follow the Boca Tauce gravity high's
485 northernmost border (50 mGal isogal). Gravity values around 30 mGal dominate the

486 northern part of the PV-PT complex, an almost 50% decrease compared to the gravity
487 maximum.

488 b. Gravity inversion

489 i. Homogenous background

490
491 Inversion results shown in Figures 6 and 7 are obtained using the new gravity
492 data only. The root mean square error in the inversion is 0.42 mGal (Fig. 8).
493 Horizontal sections through the 3-D model between 2000 m elevation and 2000 m
494 below sea level (bsl) show a central low density body located beneath the PV-PT
495 complex (Fig. 6). The low density feature with a decrease in density of almost 300
496 kg/m^3 from background values, appears to be irregularly shaped and inclined from
497 Pico Viejo towards the NE, following the strike of the Dorsal Ridge. The body
498 reaches its maximum width of ca. 6 km at around sea level and also extends beneath
499 Montaña Blanca and the entire northern slope of Pico Teide (Fig. 6). Short-
500 wavelength positive densities down to ca. 1500 m elevation appear to follow the LCC
501 wall from its northwesternmost part anticlockwise towards La Fortaleza. This high-
502 density structure becomes more or less coherent at 1000 to 500 m elevation, from
503 where the structure extends to at least 1000 m below sea level. From there onwards
504 the high-density body seems to be restricted beneath and beyond (westwards of) the
505 Ucanca sector of the LCC. Minimal density contrasts dominate the centre of the LCC
506 and the adjacent Santiago rift as well as along a zone opening towards the SW from
507 the southwestern part of the LCC (particularly prominent between +500 m and -1000
508 m elevation). This lower density zone in the Boca Tauce high-density body (Figs. 6,
509 9) may have important bearings on gravity and GPS time series collected over the past
510 years (see section 7e). Low density bodies invoked at the western and northern
511 periphery of the surveyed area suffer from inaccuracies in the computational domain

512 (wall effects) and are thus ignored for the moment. We shall see that employing the
513 complete data set leads to a better resolution of these anomalous bodies Northwest
514 and North of the PV-PT complex.

515 Inversion of the complete data (Figs. 9 and 10) confirms the results for the
516 LCC and the PV-PT complex obtained from inversion of the new data only. The main
517 change compared with the previous inversion relates to a deterioration in inversion
518 residuals to 0.67 mGal (an almost 80% increase; Fig. 11). Again the horseshoe shaped
519 high-density body, which opens towards the NW, dominates the CVC. This body
520 follows the topographic expression of the LCC wall to a depth of ca. 1000 m bsl
521 surrounding a low-density body located beneath the PV-PT complex. Below 1000 m
522 bsl neutral densities dominate the centre of the CVC, before the high-density body
523 forms a distorted heart-shape in plan view at a depth of ca 6000 m bsl, underlying the
524 major part of the CVC. A set of high density bodies are inferred at the northern slopes
525 of the PV-PT complex, which align to form a second horseshoe shape extending from
526 1000 m elevation to at least sea level. These bodies form a more or less coherent
527 western, southern and eastern boundary of neutral density material extending
528 northward to the shoreline. The root of this body appears to extend to at least 3000 m
529 below sea level N of PT and NW of La Fortaleza forming the northern boundary of
530 the prominent central low-density body located to the south (Fig. 10). This northern
531 horseshoe-shaped body is separated (at least above sea level) from the Boca Tauce
532 body by a neutral density feature aligned along the Santiago rift. Three other
533 individual shallow low-density features with densities similar to those of the central
534 CVC are worth highlighting. One is located in the Santiago rift close to the location of
535 the most recent eruption on Tenerife (Chinyero in 1909) and directly adjacent to
536 another similar-sized low-density body to the north, whose location coincides roughly

537 with the inferred western edge of the Icod valley. The second is located beneath the
538 western wall of the La Orotava valley and the third south-east of Guajara peak, along
539 the strike of the Rocas de García spur of the LCC (see Fig. 1C). A particular common
540 feature, despite their very low densities, is that these zones are either only shallow
541 rooted or they narrow at depth to below the spatial resolution of the inversion. Below
542 8000 m bsl the up to 35 km wide high-density core dominates central Tenerife
543 mimicking today's shape of the island at depth, while the Teno gravity high is at its
544 largest between depths of 2000 and 6000 m bsl. Both high-density bodies stand in
545 stark contrast to the surrounding medium with neutral densities at depths in excess of
546 6000 m bsl. However, this contrast is highly biased and exaggerated by assuming a
547 background of homogeneous density.

548

549 **ii. Density-stratified background**

550 Effects of a density stratified background on inversion results become
551 significant below sea level (Figs. 12 and 13). We obtain density contrasts in excess of
552 600 kg/m³ between low and high densities, being significantly skewed towards a
553 higher density contrast for the central core of the island. As expected, at greater depth
554 the density contrast between the background medium and the dense core decreases
555 significantly compared to the homogeneous medium inversion results (50 vs. 220
556 kg/m³, respectively) and is assumed to be closer to reality. Particular anomalous
557 features, highlighted in the previous section up to about sea level, are also found in
558 the new inversion and are hence not repeated here. The most obvious difference
559 relates to the low density “depression” modelled beneath the LCC and PV-PT
560 complex below sea level. While the homogeneous inversion gives a contrast of more
561 than 200 kg/m³, creating the impression of a bowl-shaped density decrease within the

562 high-density core of the island up to 5000 m bsl. (profiles f-i in Fig. 10), the stratified
563 inversion significantly decreases the maximum depth of this depression. Still
564 maintaining both a bowl-shape and a similar density contrast against the high-density
565 core, the structure now only extends to about 2000 m bsl (profiles f-i in Fig. 13).

566

567 **7. Gravity model interpretation and discussion**

568

569 **a. General considerations**

570

571 We limit our interpretation of the data to the immediate CVC and from depths
572 of 4000 m bsl upwards. The model results do not significantly alter inferences on the
573 deep structures of the CVC (> 4 km bsl), the Teno massif or the Dorsal Ridge as put
574 forth by previous studies (Ablay and Kearey, 2000; Camacho et al., 1991), since all of
575 these were based on the same gravity data set collected outside the CVC. A
576 fundamental controversy on the deep structure relates to whether the core of Tenerife
577 is formed by an uplifted basement block as suggested by Araña et al. (2000) or as a
578 result of intrusion of mafic magma forming dense gabbroic plutons (Ablay and
579 Kearey, 2000). The different theories brought forward are potentially heavily biased
580 towards the employed data modelling (full inversion in the former case as opposed to
581 forward modelling and parameter estimation in the latter), as both studies rely on the
582 same onshore gravimetric data set. It would appear necessary to include marine
583 gravity data in modelling the deep structure beneath Tenerife to obtain more
584 conclusive results.

585 The new gravity data in combination with a cleaned set of the existing data,
586 however, enables the assessment of the shallow and intermediate sub-surface

587 architecture of the CVC and its evolution in unprecedented detail and this shall be the
588 focus for the remaining part of the paper.

589 Fig. 14 identifies particular geometric features in the sub-surface density
590 distribution beneath the CVC.

591 First and foremost, the complex appears to be built on two intersecting
592 structural features:

- 593 i) a NE-SW striking high density ridge forming the southern part of the
594 Las Cañadas edifice with a direct link to the Dorsal Ridge and
- 595 ii) a NW-SE striking structure linking the Teno massif to the CVC via
596 the Santiago rift.

597 At their intersection at the SW part of the Ucanca caldera, a graben-like structure is
598 modelled by the inversion (Figs. 9, 14). This structure may have played an important
599 role during the reactivation of the complex in 2004, as explained in section 7e. This
600 graben is following the NE-SW trend described above and may be linked to a graben
601 structure exposed in deposits of the Ucanca formation in the caldera wall close to the
602 intersection of the Roques de García and the caldera wall (Galindo et al., 2005). It is
603 obvious that the intersection of these two lineaments controlled the evolution of the
604 CVC. It is interesting to note that the Dorsal Ridge appears to be linked directly with
605 the older Las Cañadas edifice and it is therefore not extending beneath the PV-PT
606 complex to intersect the Santiago rift beneath Pico Viejo as proposed earlier (Ablay et
607 al., 1998; Carracedo, 1994). The intersection is likely to be situated further south
608 within the Ucanca sector of the LCC.

609

610 **b. The core of the CVC**

611

612 The general picture emerging from the inversion results is that a high density
613 body extending from depth (in excess of 4 km bsl) to shallow levels forms the core of
614 CVC. This core appears to be partly associated with a fundamental mafic shield
615 building phase of Tenerife similar to other volcanic ocean islands. Remnants of the
616 shield are exposed at the Roques del Conde sequence, and the Teno and Anaga
617 massifs up to elevations of nearly 1000 m. Shallowing of the structure to near surface
618 levels to the west of the LCC may be attributed to the formation of the more or less
619 conical Boca Tauce volcano (Fig 12, for $z = 1000\text{m}$). Eruptive products of the Boca
620 Tauce volcano include lavas and pyroclastic rocks rich in mafic minerals and
621 plagioclase cumulates. Rock densities for this sequence can exceed 2900 kg/m^3 and
622 represent the densest rocks exposed nowadays at the CVC. It is thus likely that the
623 modelled density contrasts, in excess of 370 kg/m^3 , are indicative of these dense rocks
624 and even denser deeper-seated plutonic rocks including cumulates (see Ablay and
625 Kearey, 2000, for estimates on cumulate densities). The long-wavelength positive
626 magnetic anomaly detected in this area is consistent with the interpretation of an old
627 mafic volcanic edifice (Araña et al., 2000; García et al., 2007). The Boca Tauce
628 lithologies with Lower Group ages of up to 3.5 Ma are exposed in the South western
629 part of the LCC's Ucanca sector and adjacent valleys. These rocks occur also as
630 accidental fragments at the PV summit, erupted during a phreatic phase in the central
631 crater as part of the 1792 Narices del Teide fissure eruption. This highlights that a pre-
632 Las Cañadas edifice also extended beneath what is now the CVC early in the
633 evolution of Tenerife and contradicts inferences of three isolated proto-Tenerife
634 islands (Teno, Anaga and Roques del Conde; see Ancochea et al., 1990; Walter,
635 2003). As a consequence, the high-density body, modelled to extend to shallow depth
636 beneath PV, may indeed be part of the Boca Tauce edifice (Fig. 13, profile f).

637 Including both Lower and Upper Group lithologies, and thus also the Boca Tauce as
638 its westernmost entity, the Las Cañadas edifice was constructed up to 179ka. We
639 interpret the horseshoe-shaped gravity ridge, with positive density contrasts of up to
640 310 kg/m^3 following the LCC wall, as remnants of this dominant edifice. A rather
641 dramatic petrological and volcanological change occurred with the formation of
642 Upper Group lithologies which resulted from an increased in the eruption of evolved
643 phonolitic magmas and thus increasingly highly explosive volcanism. The slight
644 reduction in modelled densities along the ridge may reflect an overall reduction in the
645 bulk density of Upper Group lithologies due to the occurrence of lower density
646 magmas and lavas and a higher proportion of pyroclastic deposits during the
647 construction of the LCC.

648

649 c. The shallow structure of the CVC

650

651 One of the most important inferences from the models is that the shallow
652 subsurface beneath the CVC is not entirely composed of high-density material.

653 Four zones of low density appear to characterise the shallow structure beneath
654 the CVC up to surface levels:

- 655 i) the bowl-shaped density low beneath the PV-PT complex (discussed
656 in detail in section 7d),
- 657 ii) a NNE-SSW trending low density region made up by two gravity
658 lows located in the Icod valley and crossing the Santiago rift at an
659 angle of ca. 80° (Fig. 12 for $z = 1000$ to -1000m , and Fig. 14 for $z =$
660 500m),

- 661 iii) a linear alignment of low density bodies connecting the Icod low
662 density body with a low beneath the Tigaiga massif at the western
663 border of the Orotava Valley (Figs. 12, 14), and
664 iv) two density lows southeast of Guajara Peak, at the southern slopes
665 of the Las Cañadas edifice.

666 All four zones show similar negative density contrasts, mainly due to the general
667 assumptions behind the inversion approach. Except for the Icod anomaly, all low
668 density bodies appear to be shallow rooted, with maximum extension to around sea
669 level (Figs. 12 and 14). The Icod low density body appears to be deep-rooted,
670 extending from the ground surface to around 2000m below sea level. The area
671 underlain by this body is dominated by mafic dykes and normal faults exposed in road
672 cuts, which document the extensional stress field mirrored also by the abundance of
673 young mafic cones along the Santiago rift, including the 1909 Chinyero cone. It is
674 therefore likely that this body has a deep structural control such as a fault or damage
675 zone potentially associated with dykes. Low densities may result from hydrothermal
676 fluid migration creating secondary void space and alteration, corrosion and/or active
677 fracturing and faulting along these structural elements. It also appears that this zone
678 played an dominant role during the recent reactivation of the CVC, as discussed in
679 more detail in section 7e.

680 The North-South profile A in Fig. 13, reveals the pronounced asymmetry in
681 the distribution of high-density material between surface levels and 3000 bsl. Whereas
682 the high density body reaches the near-surface beneath the southern portion of the
683 LCC, the area below and north to the PV-PT complex is dominated by lower density
684 material except for the shallow rooted horseshoe alignment of high-density material
685 forming the northern border of the low-density interior of the PV-PT complex (Fig.

686 13). The strong N-S asymmetry has also been highlighted in seismic velocity (Watts
687 et al., 1997) and aeromagnetic data (Araña et al., 2000; García et al. 2007). It has been
688 suggested by both Araña et al. (2000) and Ablay and Kearey (2000), that these
689 anomalies are due to low density, partly chaotic and unconsolidated volcanic rocks
690 (mainly mafic and phonolitic lavas), erupted from the PV-PT complex over its
691 lifetime from 179 ka onwards. These deposits represent the major infill of the Icod
692 valley. Our model clearly supports these earlier inferences, yet we can put further
693 constraints on the destruction of the Las Cañadas edifice and the formation of the Las
694 Cañadas caldera.

695 The absence of a surface expression of a northern caldera wall has led various
696 authors to suggest that the LCC may indeed have formed as a result of a large lateral
697 mass wasting (Cantagrel et al., 1999; Ancochea et al., 1999; Watts and Masson,
698 2001). In their conceptual model, the LCC wall is the headwall of the Icod landslide
699 due to its “striking resemblance to the head scars of large landslides” (Watts and
700 Masson, 2001). While landslides undoubtedly played an important role in the
701 geological evolution of Tenerife (Ablay and Hürlimann, 2000), our investigation does
702 not find any evidence for a lateral collapse origin of the LCC. Quite the opposite, we
703 find clear evidence supporting its formation by vertical collapse.

704 First, neutral density material is not only infilling the Las Cañadas depression but also
705 extending to considerable (>2000 m bsl) depth, significantly displacing the underlying
706 high density body. This displacement requires a deep structural perturbation unlikely
707 to be induced by a surficial landslide with a maximum penetration depth of 1000 m as
708 inferred for the Icod case (Hürlimann et al., 2000). The perturbation is consistent with
709 a down faulting of a sizeable part of the high density interior of the Las Cañadas

710 edifice and possible alteration in the form of lateral landslides directed to the north
711 during Lower Group times.

712 Second, short wavelength high-density bodies in the Diego Hernández sector, shown
713 in Fig. 6 (for $z=1500\text{m}$), delineate an elliptic pattern surrounding a central neutral
714 density body and may be interpreted as shallow intrusions. The inferred semi-major
715 axis of this feature strikes NNW-SEE and matches electromagnetic observations in
716 this area (Coppo et al. 2008), which identified a NNW-SSE elongated funnel-shaped
717 conductive layer at depth. This area is also characterised by a well defined magnetic
718 low as revealed by recent aeromagnetic data (García et al., 2007). The semi-minor
719 axis of the elliptic alignment of the high-density bodies matches the semi-minor axis
720 of the electromagnetic resistivity print as well as the magnetic low and we propose
721 that these features are related, i.e., have a common source. The distribution of the
722 high-density bodies resembles intrusions along a ring fault, which is consistent with
723 the interpretation of the conductive layer representing the remnant of an older edifice
724 affected by vertical collapse. The resistivity prints and the morphology of the
725 conductive layer led Coppo et al. (2008) to propose the initiation of the collapse in the
726 SE part of the Diego Hernández sector. The abundance of high-density bodies in that
727 part of the caldera depression is in perfect agreement with ring fault formation and
728 initiation of the vertical collapse there. The combination of results from these
729 potential field investigations (gravity, magnetotellury and magnetics) provides
730 irrefutable evidence for a vertical collapse origin of the LCC, which is also supported
731 by the abundance of voluminous products of phonolitic explosive (caldera-forming)
732 volcanism on the island (Martí et al., 1994; Edgar et al., 2007; Pittari et al, 2008)

733 It has been suggested by various authors that the head of the Icod valley is
734 buried beneath eruptive products of the PV-PT complex (Martí et al., 1997; Martí et

735 al., 1998; Ablay and Hürlimann, 2000). The horseshoe-shaped alignment of shallow
736 rooted high density bodies to the north of the complex gives the impression of a
737 structural barrier between the complex and the topographic expression of the Icod
738 valley, coinciding with a significant break in slope (Fig. 12, for $z = 1000\text{m}$; Fig. 13,
739 profile A). One could suspect that parts of this structure represent the northern
740 (buried) caldera wall. It is furthermore interesting to note that the maximum opening
741 of the structure matches the maximum width of the Icod valley and we thus interpret
742 that structure to be associated with the formation of the landslide valley. Shallow
743 AMT resistivity prints at the northern slope of the PV-PT complex identify a dramatic
744 shallowing of a conductive body below a resistive layer in the areas highlighted by
745 these high density bodies (Coppo, 2007). This author interprets the conductive layer
746 as the décollement of the Icod landslide, most probably coinciding with the El
747 Mortalon layer, which is thought to be the gliding plane of the slide (Bravo, 1962;
748 Cantagrel et al., 1999). We propose that some of the high density bodies represent
749 intrusives along the (sub)vertical lateral failure planes of the slide, marking its side
750 walls. Similar features are observed elsewhere in the form of basic magmatic
751 intrusions along structures resulting from lateral mass wasting on ocean islands, such
752 as Hawaii or Reunion (Oehler et al., 2004). The headwall may also be either marked
753 by an intrusion or coincide with the remnant of the northern caldera wall. We cannot
754 provide a conclusive answer as to the fate of the northern caldera wall and its role
755 during the formation of the Icod valley, but all evidence points towards a remnant
756 structure below the PV-PT complex whether it be the Icod head scar or the northern
757 caldera wall or (probably most likely) a combination of both. The key finding here is
758 that the headwall of the Icod valley is located well outside (to the N of) the current

759 caldera depression and cannot thus have had any role in the formation of the LCC
760 wall.

761

762 d. The interior of the PV-PT complex

763

764 The interior of the PV-PT complex appears to be made up of low density
765 material extending from close to the surface to about sea level. The horizontal extent
766 of the density low is from Pico Viejo in the west to Montaña Blanca in the east of the
767 complex. This marked density low has an overall irregular shape with a finger-like
768 protrusion below PV. There are several possibilities to explain the low densities
769 below the complex.

770 First, the complex is dominated by phonolitic eruptive products; mostly lava
771 flows and pumice fall deposits with the lowest bulk densities (as low as about 1100
772 m³/kg as measured in the field; see Table 2 and Ablay and Kearey (2000), as well as
773 more primitive highly-vesiculated lavas, also with bulk densities lower than the
774 background density.

775 Second, the complex may, to some extent, be hydrothermally altered and
776 internally eroded causing secondary void space leading to lower bulk densities.
777 Surface expressions of hydrothermal alteration zones are restricted to the area of the
778 Roques de García, the lower part of the caldera wall and within the youngest Pico
779 Teide summit crater. Remnants of two older Teide craters, exposed at ca. 3500 m
780 close to the upper cable car stop, exhibit fumarolic activity at the time of this writing.
781 The bulk density of the highly altered summit crater deposits is well below the
782 background density. Assuming that at least parts of the interior of the complex have
783 undergone a similar degree of alteration, one would expect lower densities for the
784 interior of the complex, as perhaps best exemplified by the conduit-shaped low

785 density feature beneath PV extending upwards and to the west from the centre of the
786 density low (Figs. 6, 7 and 13). In fact, recent electromagnetic data show a highly
787 conductive body beneath Montaña Blanca with electrical resistivities of less than 20
788 ohm m (Coppo et al., 2008), a clear indication for either the presence of shallow
789 hydrothermal alteration or a geothermal system. With the water table being located
790 well below 400 m below the caldera floor (Gottsmann et al., 2006 and references
791 therein), i.e. < 1800 m asl, it is conceivable that large parts of the immediate
792 subsurface beneath Montaña Blanca (summit at 2800 m asl) is hydrothermally altered
793 rather than hosting saline ground water. The total extent of hydrothermal alteration of
794 the complex is not well defined but both our gravimetric as well as the AMT data
795 indicate that a rather substantial volume of the total volume of the complex of ca. 9
796 km³ may be substantially altered and structurally weakened. Of course these findings
797 have important bearings on the stability of the complex, with northern slopes showing
798 angles of inclination of 38° or more. With a history of several landslides, the northern
799 flank of the complex is a likely locus for future mass wasting facilitated by internal
800 corrosion of the complex.

801 Petrological evidence points towards the existence of shallow magma
802 chambers (at depths of around sea level) feeding past phonolitic eruptions at the
803 complex (Ablay et al., 1998; Triebold et al., 2006; Andujar, 2007). It is possible that
804 the rather clearly defined lower boundary of the low-density body, resolved by the
805 stratified background model (Figs. 12 and 13), marks the transition from
806 hydrothermally altered material to remnant magma reservoirs. At a depth of about
807 1000 m bsl a neutral density contrast is modelled which increases with depth to a
808 contrast of ca 100 kg/m³, forming a sizable (several km³ in volume) bowl-shaped
809 body beneath the complex, clearly separated from the high density material forming

810 the core of the Las Cañadas edifice. One would expect that partially crystallised
811 phonolitic magma would have a density similar to background densities and one could
812 propose this region to at least host remnants of highly evolved magmatic bodies.
813 Whether or not these bodies contain eruptable material, if defined by a critical
814 abundance of melt, may not be resolved by the available gravimetric data.

815
816

817 e. Relation to recent unrest

818

819 The sub-surface density distribution allows us to put these new findings in
820 relation to geodetic and seismic observations performed during the recent reactivation
821 of the CVC in spring 2004 and onwards. Details on investigation strategies and results
822 can be found in the recent papers by Gottsmann et al. (2006), Almendros et al.
823 (2007), Fernández et al. (2007), and Martí et al. (2008a). Two features are worth
824 highlighting.

825 First, the low-density body identified inside and crossing the Icod valley
826 southwestwards (Fig. 14) underlies the zone in which the most drastic residual gravity
827 changes were documented in the first few months of the reactivation. It is fair to
828 assume a relationship between this deeply rooted negative density anomaly and the
829 documented sub-surface mass changes. This area has also been recognised to undergo
830 periodic ground deformation, as shown by recent DGPS and InSAR observations (
831 Fernández et al., 2004). The spatio-temporal gravity changes as well as the seismic
832 data have been interpreted to reflect aqueous fluid migration at depth during the
833 period of reactivation. We therefore propose that the perturbation of the CVC, as
834 documented by the various geophysical techniques, may be rooted in the low-density
835 body. Upward movement of pressurised aqueous liquids may have been controlled by

836 pre-existing structural weaknesses. Mass addition was also documented in the Ucanca
837 sector of the LCC along the Santiago Rift (Gottsmann et al., 2006). Accounting for
838 the dominant NW-SE structural control of the CVC, linking the LCC to the Teno
839 massif, it appears reasonable that this structure controlled most of the changes at
840 depth. In fact, results from the inversion of integrated stacked DInSAR and GPS data
841 collected between 2003 and 2006 are in very good agreement with this hypothesis
842 (Fernández et al., 2008).

843 Second, a joint GPS and gravity benchmark located just outside the
844 topographical expression of the Ucanca wall underwent significant ground
845 deformation and gravity changes during the period of reactivation, which were
846 difficult to interpret (Gottsmann et al., 2006). The new gravity inversion results put
847 this benchmark inside an inferred graben structure at the intersection of the two
848 dominant structural lineaments as documented above (Figs. 6 and 9). Indeed from a
849 topographical perspective a sizeable part (> 200m length) of the Ucanca caldera wall
850 is missing in this particular area . Its dissected nature as well as the abundance of
851 slickensides on (sub)vertical surface outcrops provides geomorphological and
852 geological evidence for a tectonic control on the evolution of this part of the LCC.
853 These features are indicative for relative extension, possible graben formation and
854 structural weakening. The ground deformation and density changes observed between
855 2004 and 2005 in this area are consistent with localised
856 pressurisation/depressurisation cycles and associated mass and density changes over
857 periods of a few months or less. Fernández et al. (2008) found deformation sources
858 whose position, depth, radius, pressure and mass evolve over time. Displacements
859 occur particularly along low-density zones (including the southwestern part of the
860 Ucanca depression) and areas surrounding high-density bodies (see Fig. 14). It

861 appears that again shallow and localised sub-surface dynamics were controlled by the
862 dominant structural entities of the CVC at the time of its reactivation.

863

864

865 **8. Summary and Conclusion**

866

867 We present a new set of gravimetric data from the central volcanic complex on
868 Tenerife, which, coupled with a full 3-D data inversion, enabled unprecedented
869 insights into the shallow sub-surface structure beneath the complex.

870 1. The centre of the island is dominated by a high-density core
871 expressed by a local gravity high, which is formed by the
872 intersection of two principal structural lineaments oriented at about
873 right angles: the Teno-Santiago rift and the Dorsal ridge. These
874 structures have controlled the morphological and volcanic evolution
875 of both the island and the central complex over the past few million
876 years and appears to continue controlling sub-surface dynamics as
877 shown during the recent reactivation of the complex in 2004.

878 2. Based on a combination of gravimetric and geological data we
879 propose the existence of a central volcanic complex early in the
880 evolution of Tenerife. The phonolitic complex, beneath the centre of
881 the island, most likely formed within an early shield structure.
882 Construction of isolated volcanic edifices forming the three corners
883 of the island is not supported by our data. The asymmetric density
884 distribution below sea level is consistent with the removal of

885 significant proportions of the early central edifice in the form of
886 lateral landslides directed to the north during Lower Group times.

887 3. We find strong evidence for a vertical collapse origin of the Las
888 Cañadas caldera both from gravimetric data alone as well as from its
889 excellent correlation with results from recent electromagnetic and
890 aeromagnetic investigations. The distribution of small and shallow
891 high-density bodies in the Diego Hernández sector is consistent with
892 intrusions along a ring fault marking the periphery of a funnel-
893 shaped conductive layer interpreted to represent the base of the
894 Diego Hernández caldera.

895 4. A horseshoe-shaped alignment of high-density bodies at shallow
896 depth to the north of the Pico Viejo - Pico Teide complex is
897 interpreted to represent shallow intrusions into a scar left behind by
898 the Icod landslide, marking the horizontal extent of the slide as well
899 as its head wall. The head wall is nowadays buried beneath eruptive
900 products of the youngest volcanic complex on the island, and thus
901 does not relate in any way to the currently exposed LCC wall. The
902 Icod valley head wall may represent parts of the northern caldera
903 wall, formed by the superposition of three vertical collapses over the
904 past 1 Ma.

905 5. The interior of the PV-PT complex is characterised by a cylindrically
906 shaped body extending from depths of ca. 3000 m bsl to about sea
907 level. This body of neutral density is interpreted to represent the
908 current plumbing system of the complex. Above sea level, low
909 density material dominates the shallow interior of the complex and is

910 thought to result from a combination of evolved eruption products
911 (dominantly phonolitic air-fall and intermediate to evolved
912 vesiculated lavas) with a significant degree of hydrothermal
913 alteration leading to internal corrosion.

914

915 Future geophysical investigations at the CVC aiming at resolving the temporal
916 evolution of the sub-surface in light of the recent reactivation should take into account
917 the results presented here. Two points are worth mentioning. First, the degree of
918 internal alteration of the PV-PT complex needs to be better assessed and implications
919 for slope stability need to be considered. Edifice collapse may have severe impacts in
920 its own right but may also impact on magmatic systems at depth (Reid, 2004). Such
921 scenarios need to be accounted for in quantitative risk assessment and monitoring
922 programs need to provide critical baseline data.

923 Second, the control of the two dominant structural lineaments on the evolution
924 of the central volcanic complex is evident throughout its history. These structures
925 appear to define preferential pathways for both aqueous fluid and magma migration.
926 Monitoring networks should hence include coverage of these structures in addition to
927 the immediate PV-PT complex.

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Acknowledgments

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Table and figure captions

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1156 Table 1: Simplified stratigraphic scheme for the island of Tenerife, showing the major

1157 constructive and destructive episodes. * indicates major vertical collapse, †

1158 indicates major lateral collapse (landslide). Modified and simplified after

1159 Edgar (2003).

1160

1161 Table 2: Measured bulk densities of lithologies exposed at the Central Volcanic

1162 Complex. Values taken from Ablay and Kearey (2000).

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1166 Figure 1. A) Location of Canary Islands off African West coast. B) Topographic

1167 models of Tenerife (28.32°N, 16.57° W) showing main structural units (B and C) and

1168 locations of new gravity benchmarks (squares in C). Coordinates in B) and C) are

1169 given in UTM (m). Key to (B): Dorsal R. = Dorsal Ridge, Icod V. = Icod Valley,

1170 LCC = Las Cañadas caldera, OV = La Orotava Valley, Santiago R. = Santiago Rift, R.

1171 del Conde = Roque del Conde, PV-PT = Pico Viejo – Pico Teide complex, Tg =

1172 Tigaiga massif.

1173 Key to (C): BT = Boca Tauce, GP = Guajara Peak, LF = La Fortaleza, MB = Montaña

1174 Blanca, PT = Pico Teide, PV = Pico Viejo, RdG = Roques de García, SR = Santiago,

1175 Rift, sectors of the Las Cañadas caldera (with decreasing age): UC = Ucanca, GJ =

1176 Guajara, DH = Diego Hernández.

1177

1178

1179 Figure 2. Bouguer anomaly (mGal) map of Central Tenerife. Points indicate location

1180 of benchmarks selected for inversion of gravity data. Coordinates given in UTM (m).

1181

1182

1183 Figure 3. The optimal theoretical terrain density corresponds to the density giving the
1184 minimum correlation between short-wave components of topography and gravity for
1185 distances up to a radius r around any given benchmark. The graph shows the
1186 relationship between terrain density and average radius r (bold line). A terrain density
1187 value of $2200 \text{ (kg/m}^3\text{)}$, indicated by the thick line, is finally chosen, accounting for
1188 both the low correlation between short wave components and measured rock densities
1189 reported in Table 2.

1190

1191 Figure 4. Local gravity anomaly for the central volcanic complex, obtained by
1192 subtracting a general linear SE-NW trend of 0.27 mGal/km with azimuth $\text{N}113^\circ\text{E}$
1193 from the local Bouguer anomaly. This local anomaly data is employed for data
1194 inversion.

1195

1196 Figure 5. Sensitivity test of the inversion methodology based on a simulation of an
1197 anomalous density body located beneath the CVC. a) shows Bouguer gravity anomaly
1198 caused by the anomalous body at the benchmarks shown by black and grey circles. b)
1199 Plan view and vertical cross section of the simulated body with vertical S-shape
1200 geometry. c) Resulting 3D model from the inversion. Tick mark separation in a) –c)
1201 on all axes is 2000 m . Distortions of the inverted structure compared to the simulated
1202 body arise from data distribution and from the tendency of the inversion method to
1203 produce rounded, smoothed structures, particularly for deep ($> 6 \text{ km}$) and peripheral
1204 zones, for which cell sensitivity is lower.

1205

1206

1207 Figure 6. Horizontal sections at selected depths through the 3-D model of density
1208 contrasts (kg/m^3) beneath the CVC employing local gravity anomaly data from 323
1209 new gravity measurements. A homogeneous background density is assumed for data
1210 inversion. Circle marks graben-like structure within the Boca Tauce high density body
1211 (see section 7 for discussion).

1212

1213 Figure 7: Vertical W-E profiles (a to g) and one S-N profile (A) through the 3D model
1214 of density contrasts (kg/m^3) beneath the CVC from inversion of 323 new gravity
1215 measurements. Horizontal section indicates trail lines of profiles. Model is based on
1216 assuming a homogeneous density background.

1217

1218 Figure 8: Inversion statistics for inversion of data from 323 new gravity
1219 measurements presented in Figs. 6 and 7. Root mean square inversion residual is 420
1220 $\mu\text{Gal} = 0.42 \text{ mGal}$.

1221

1222 Figure 9: Horizontal sections at selected depths through the 3-D model of density
1223 contrasts (kg/m^3) beneath the CVC. A homogeneous background density is assumed
1224 for data inversion. Circle marks graben-like structure within the Boca Tauce high
1225 density body (see section 7 for discussion).

1226

1227 Figure 10. Vertical W-E profiles (a to k) and one S-N profile (A) through the 3D
1228 model of density contrasts (kg/m^3) beneath the CVC. Horizontal section indicates trail
1229 lines of profiles.

1230

1231 Figure 11. Inversion statistics for inversion results presented in Fig. 9. Root mean
1232 square inversion residual is 0.67 mGal.

1233

1234
1235 Figure 12. Horizontal sections at selected depths through the 3-D model of density
1236 contrasts (kg/m^3) derived by assuming a density-stratified background beneath the
1237 CVC.

1238

1239
1240 Figure 13. Vertical W-E profiles (a to k) and one S-N profile (A) through the 3D
1241 model of density contrasts (kg/m^3) beneath the CVC. Horizontal section indicates trail
1242 lines of profiles. Model is based on assuming a density-stratified background.

1243

1244
1245
1246
1247 Figure 14: Dominant density anomalies characterising the shallow parts beneath the
1248 CVC include i) a NE-SW striking high density ridge forming the southern part of the
1249 Las Cañadas edifice with a direct link to the Dorsal Ridge (DR) and a NW-SE striking
1250 structure linking the Teno massif to the CVC via the Santiago rift (both marked by
1251 lines); ii) two low density bodies (encircled) including (i) the bowl-shaped density
1252 low beneath the PV-PT complex, (ii) a NNE-SSW trending low density region located
1253 in the Icod valley and crossing the Santiago rift, (iii) a line density lows connecting
1254 the Icod low to the Tigaiga massif (Tg), and (iv) two density lows to the south of the
1255 caldera wall. At depths > 3000 m bsl the high-density core of the CVC resembles the
1256 outline of the current island.

1257 Squares in image $Z=-1000$ m indicate surface projections of positions of ground
1258 deformation sources between 2003 and 2006 (Fernández, et al., 2008). It is interesting

- 1259 to note that most source locations are found in areas defined by shallow subsurface
1260 negative density anomalies and/or along major structural building blocks of the CVC.
1261 See Fig. 12 for scale of density contrasts.

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TABLE 1

Historic eruptions (e.g. Chinyero 1909)				
Post-shield Volcanism	Flank Eruptions		Central Volcanic Complex	
	Dorsal Series (1.8 Ma-)	Santiago Series (0.3 Ma-)	PV-PT Formation	Pico Viejo series 2
				Pico Teide series 2
	Shield construction	Las Cañadas Edifice	Upper Group	Pico Viejo series 1
				Pico Teide series 1
				Early episode
				* Climatic caldera collapse (0.18 Ma) † Icod landslide (0.2-0.13 Ma)
			Lower Group	Diego Hernández Formation (0.57 - 0.18 Ma)
				* Climatic caldera collapse (0.57 Ma) † La Orotava landslide (0.57 -0.54 Ma)
				Guajara Formation (0.85 - 0.57 Ma) † Güimar landslide (0.83 -0.57 Ma)
* Climatic caldera collapse (1.07 Ma?)				
Lower Group sequence (3.8 - 1.57 Ma)			Ucanca Formation (1.57 - 1.07 Ma)	
			Roques de Gracia	
			Montón de Trigo	
			Las Pilas	
			El Cedro	
			El Cabezón	
			Boca Tauce volcano (3.8 - 3.5 Ma)	
Old Basaltic Series				
Roque Del Conde Massif (11.6 - 3.5 Ma)		Teno Massif (6.7 - 5 Ma)	Anaga massif (6.5 - 3.3 Ma)	
		† Teno landslide (5.6 Ma?)	† Anaga landslide (~ 6 Ma)	

TABLE 2

Rock type	Bulk density (m ³ /kg)
phonolitic pumice fall	1080
phonolitic pumice	1390
phonolitic ignimbrite	2540
phonolitic lava	2350
phonolitic obsidian	2200
intermediate lava	2230
historic basalt	2890
mean	2201
stdev	271

FIGURES

Figure 1a

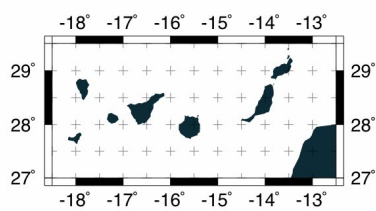


Figure 1b

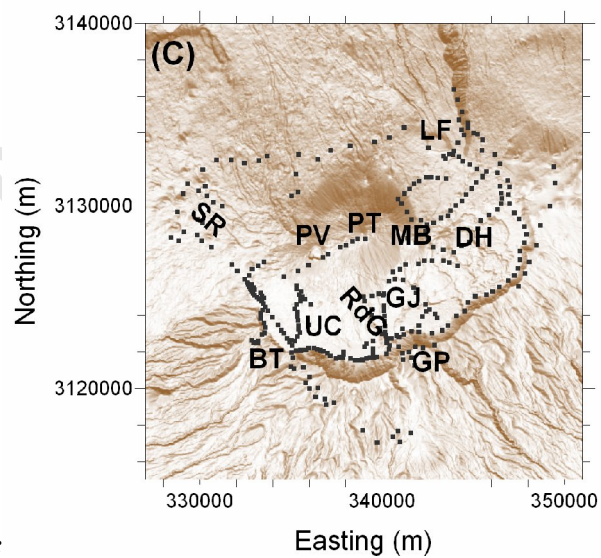
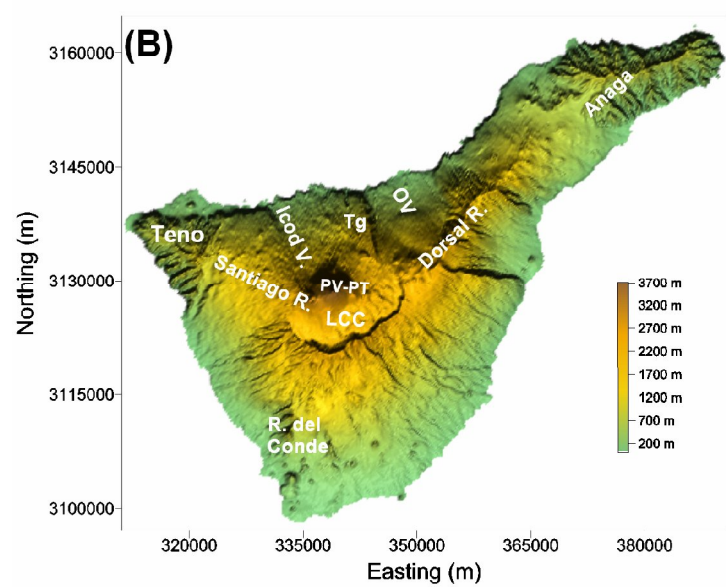


Figure 1c

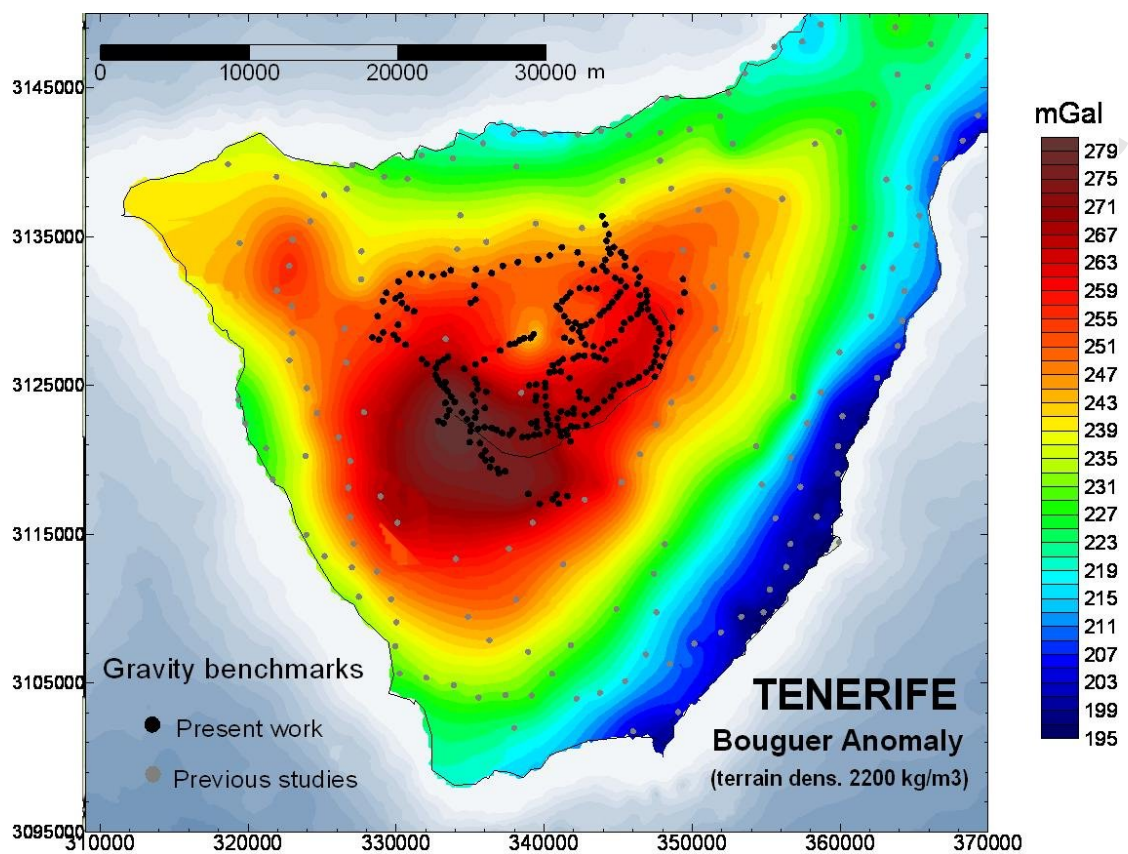


Figure 2

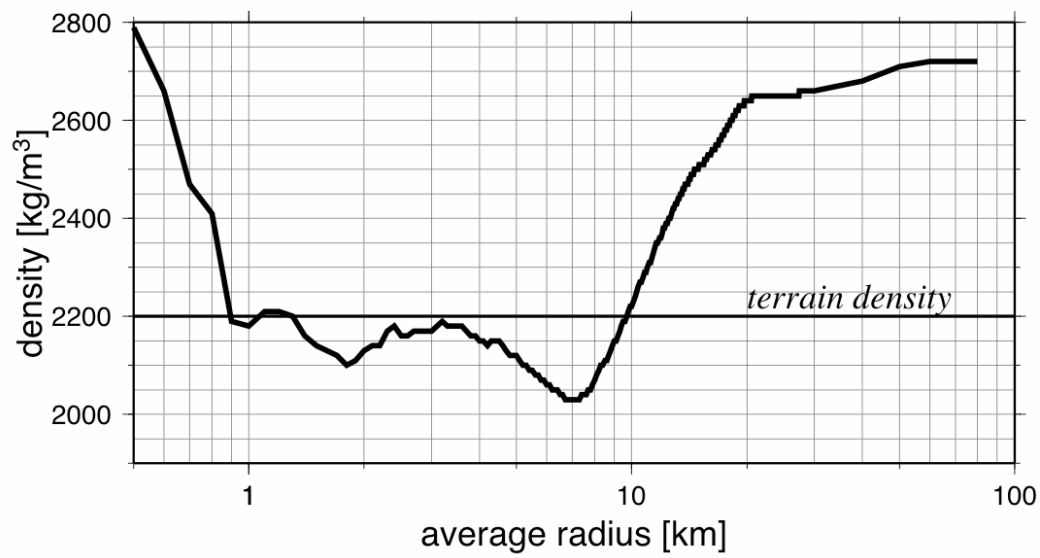


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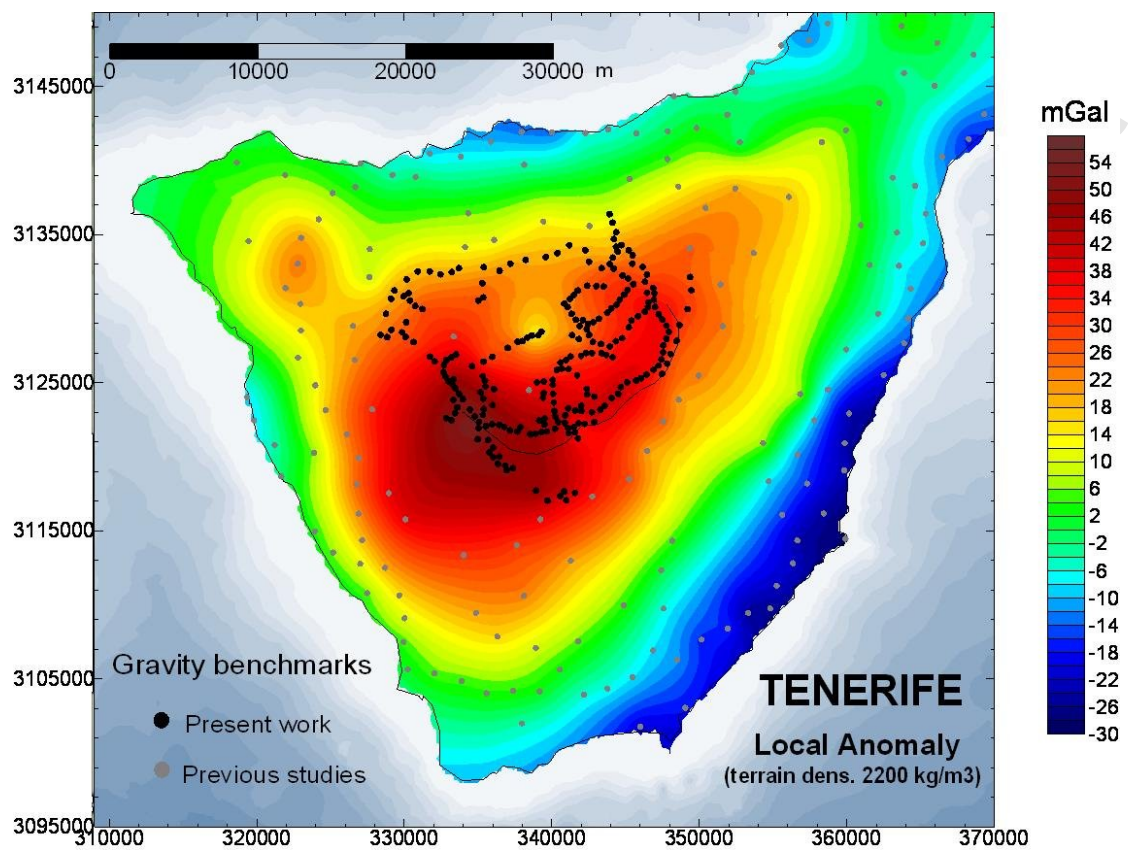
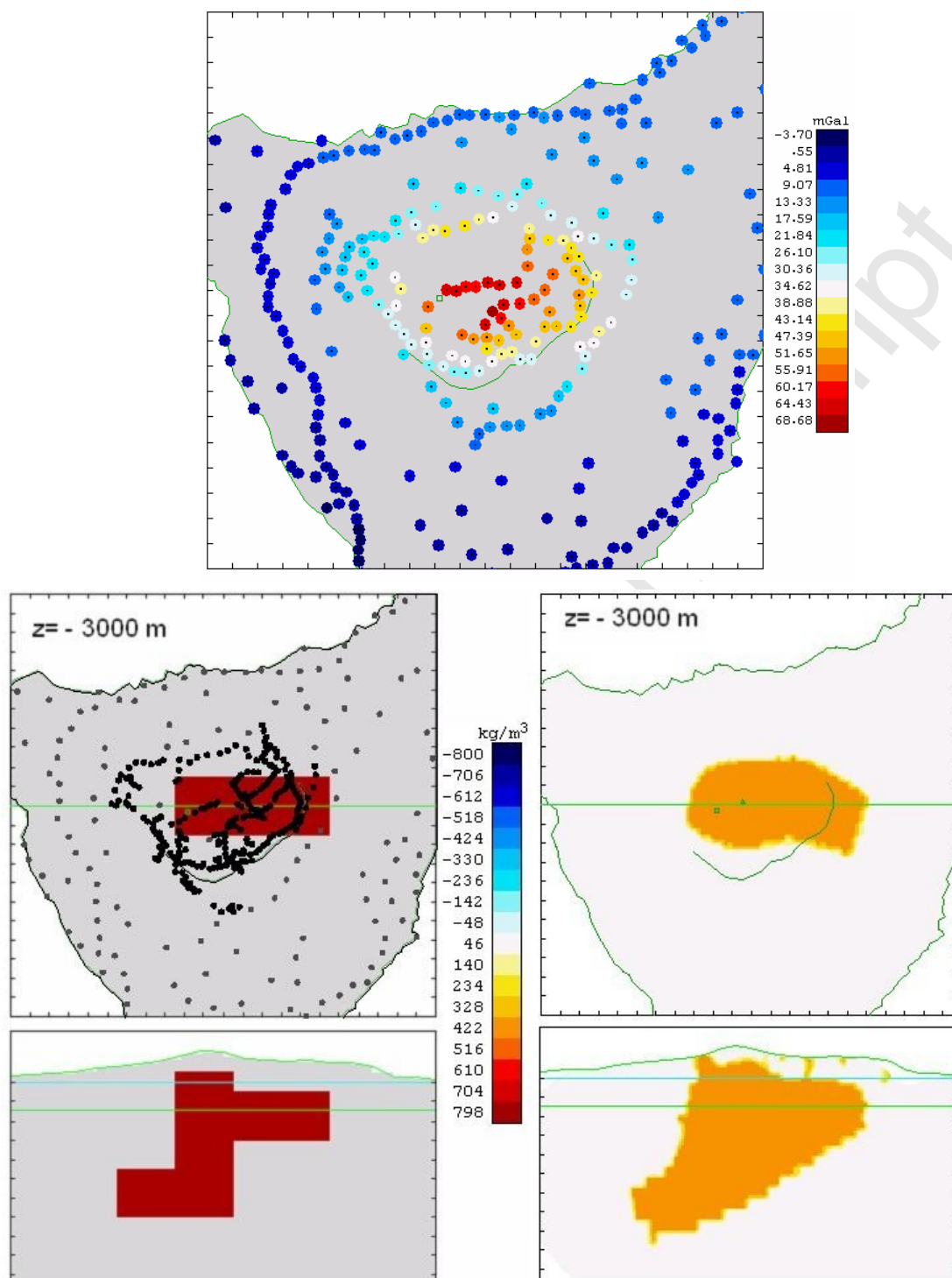


Figure 4

Figure 5



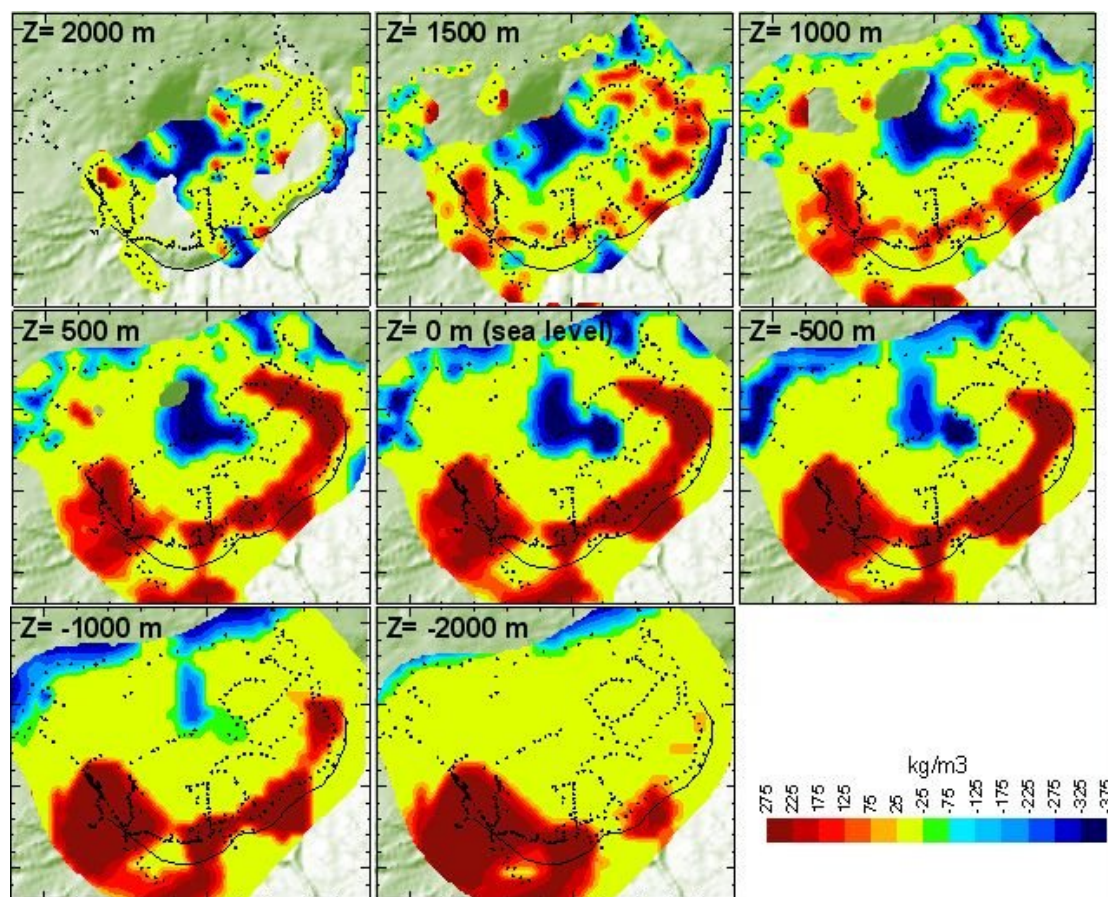


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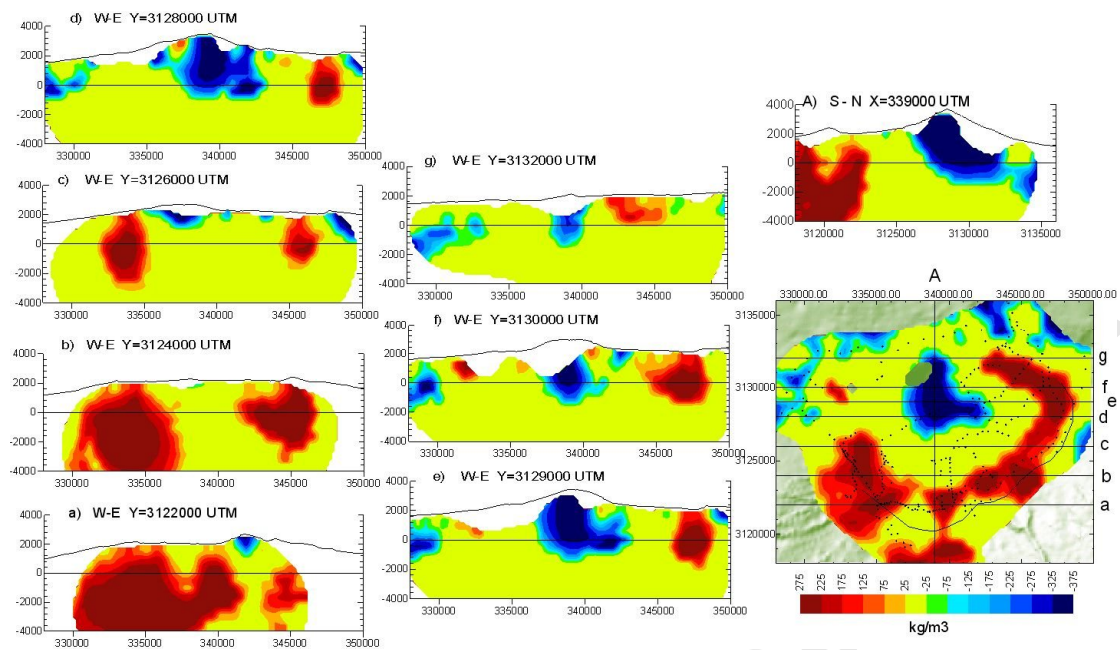


Figure 7

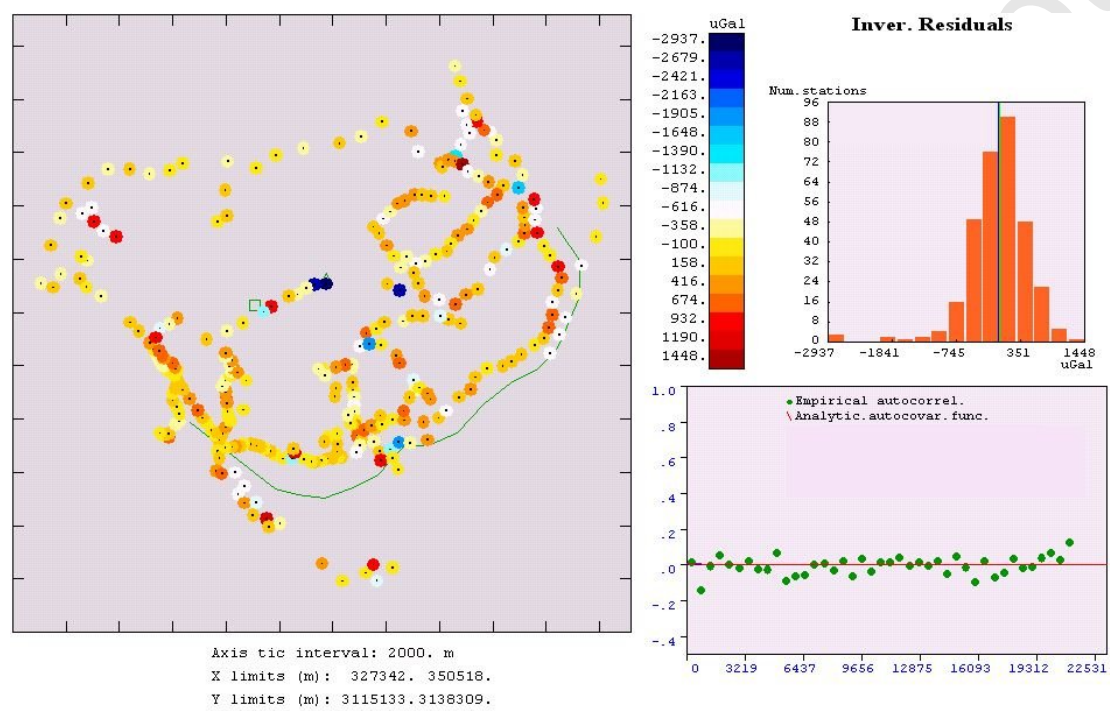


Figure 8

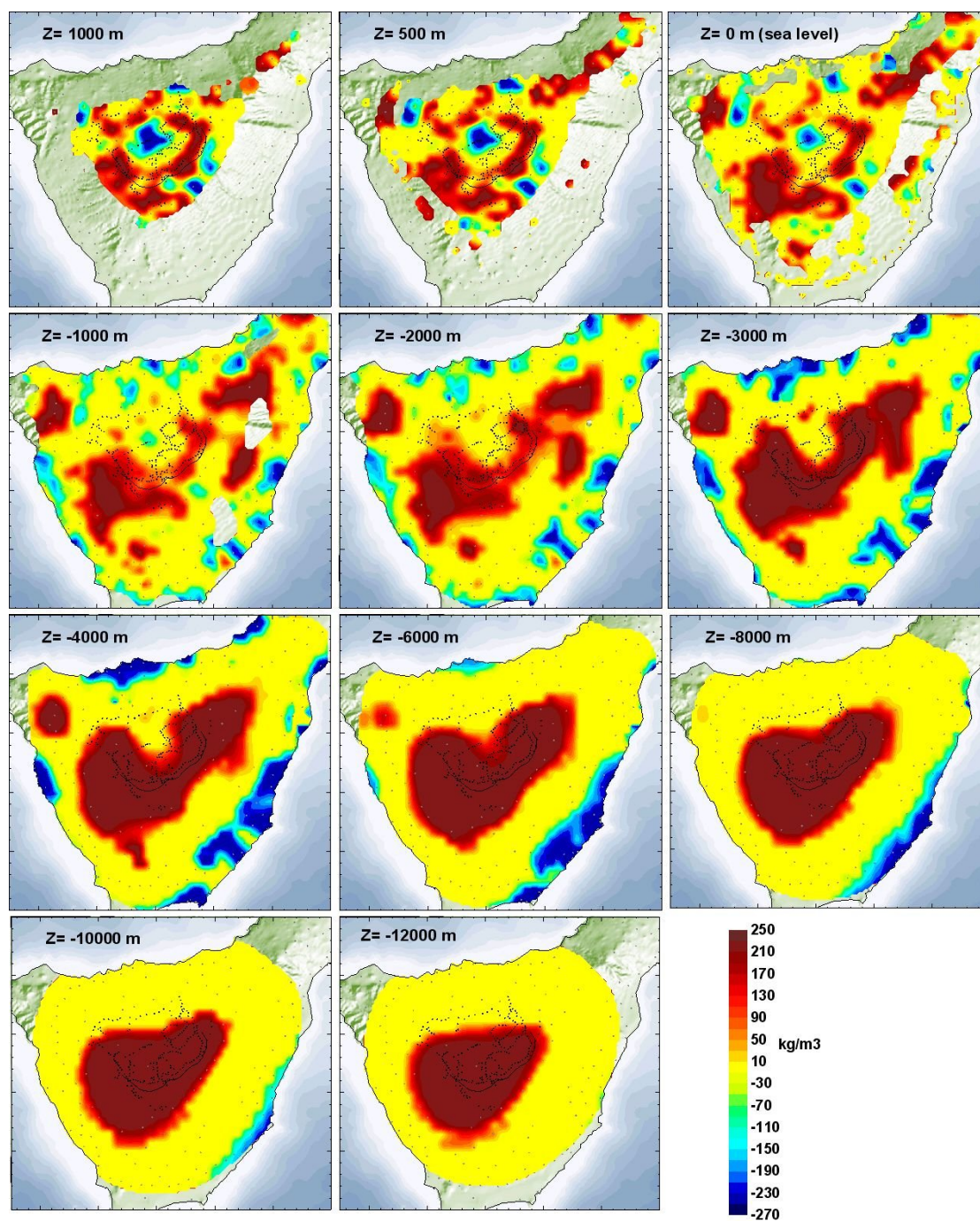


Figure 9

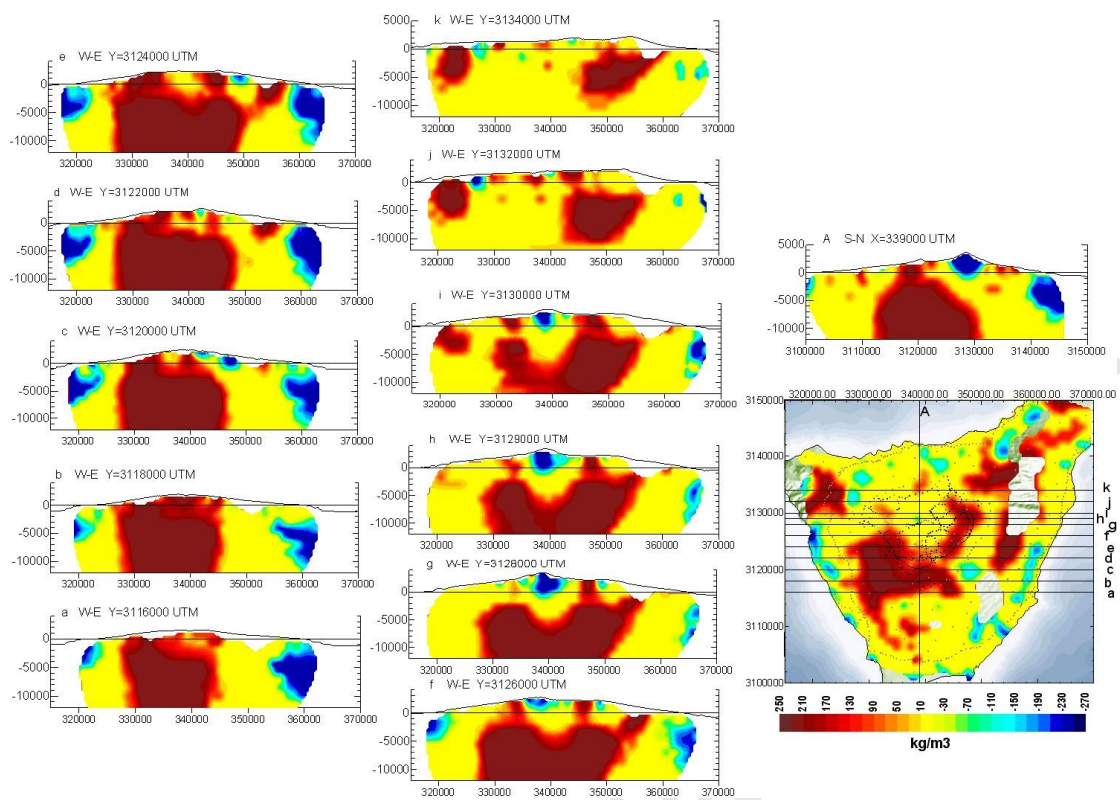


Figure 10

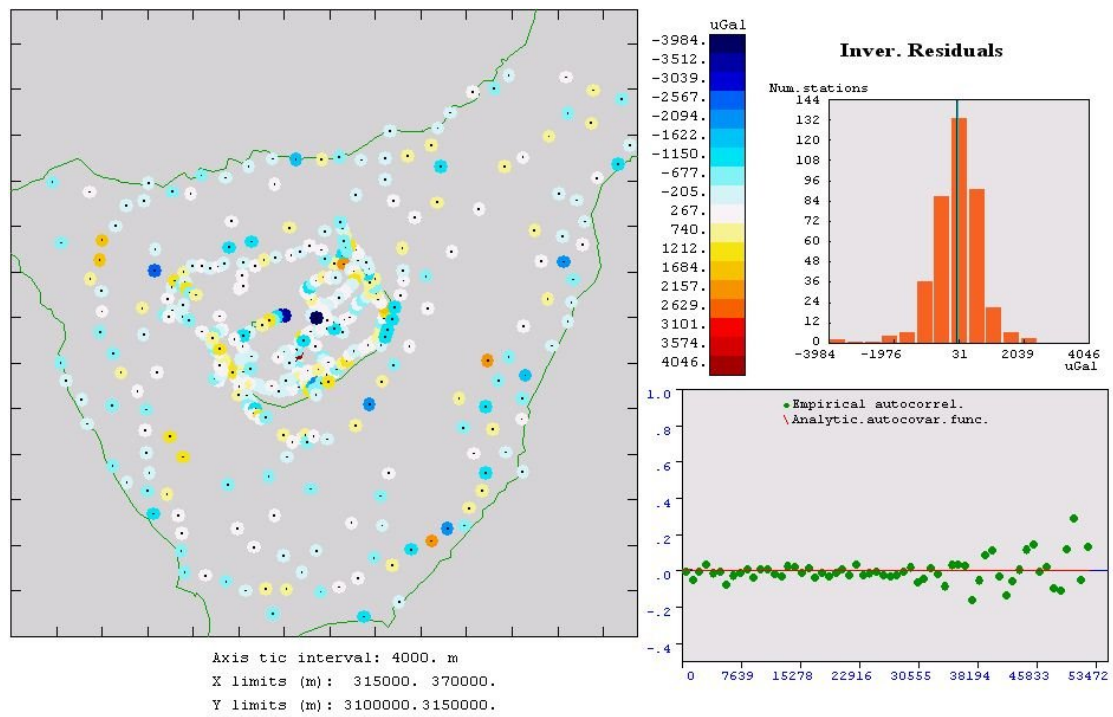


Figure 11

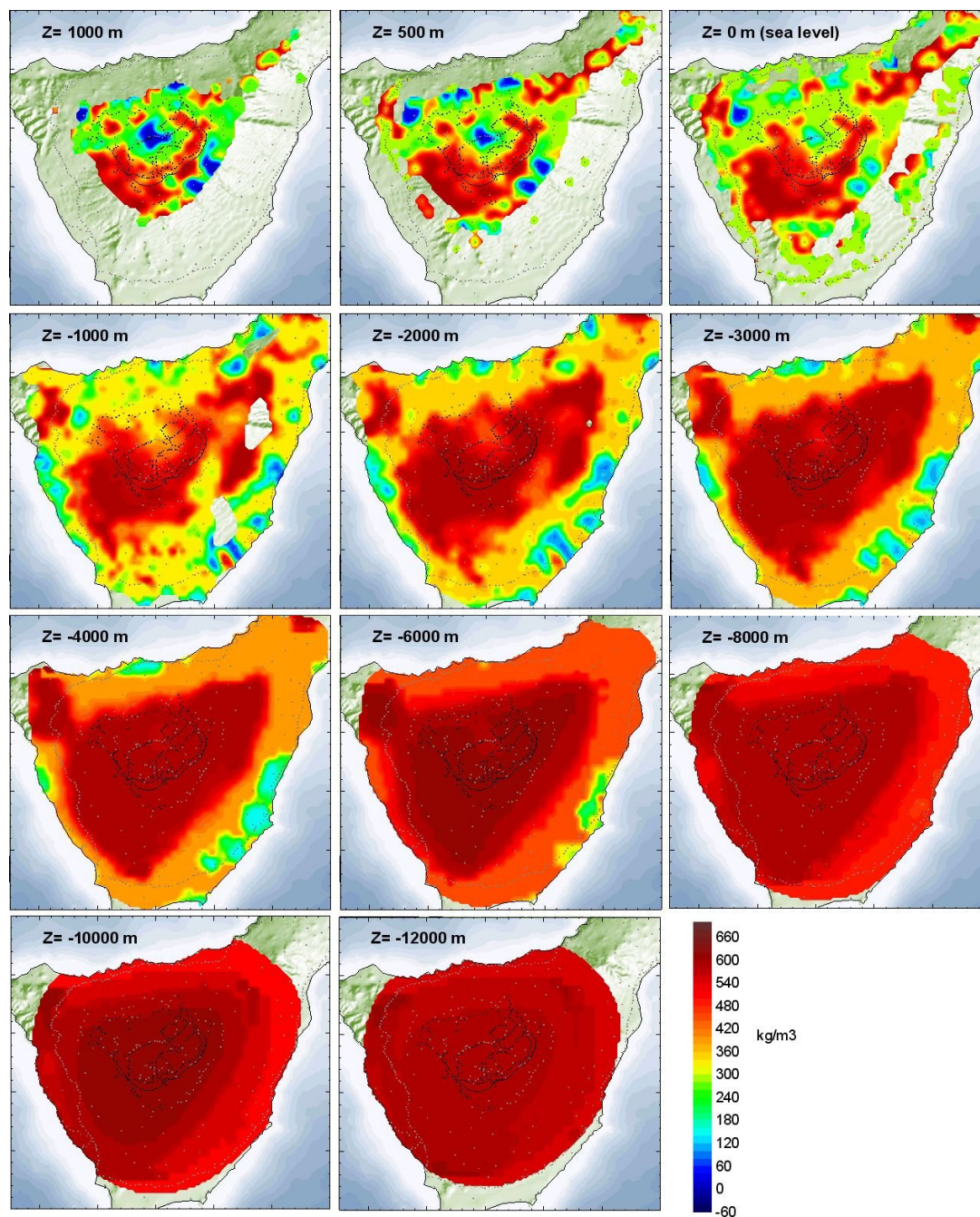


Figure 12

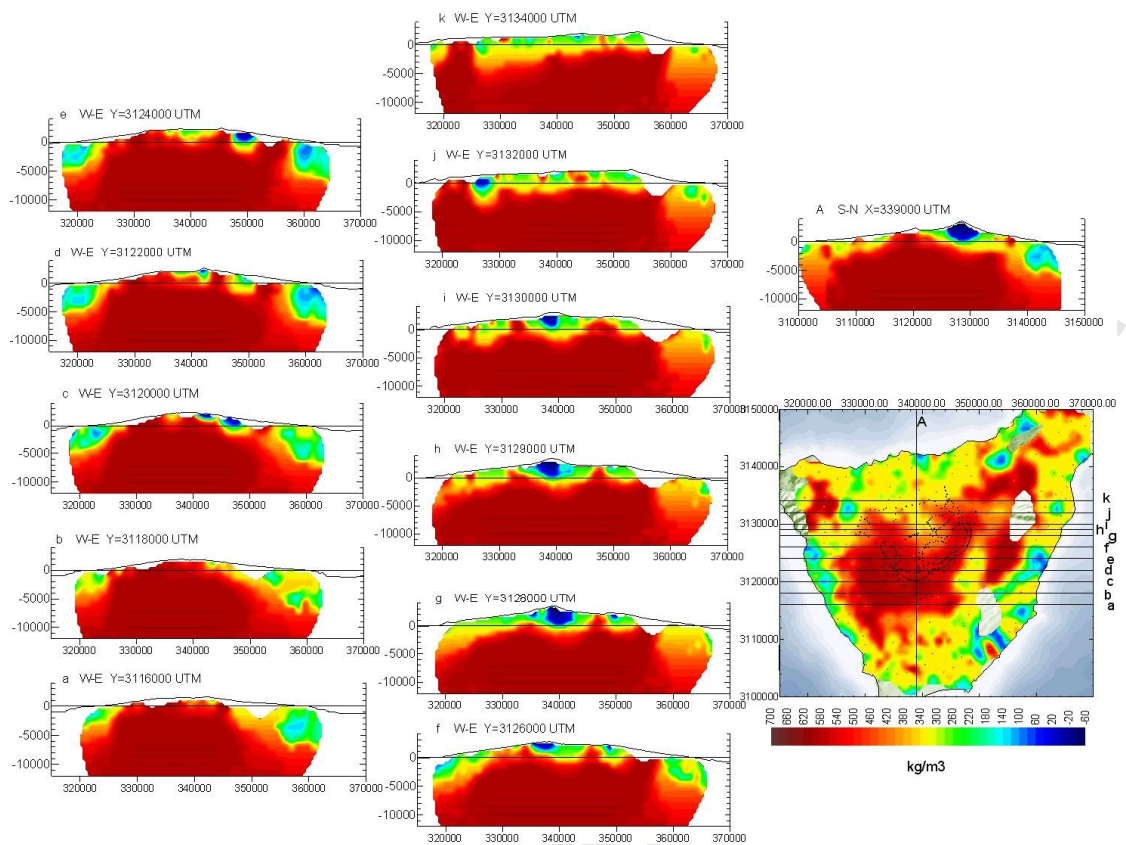


Figure 13

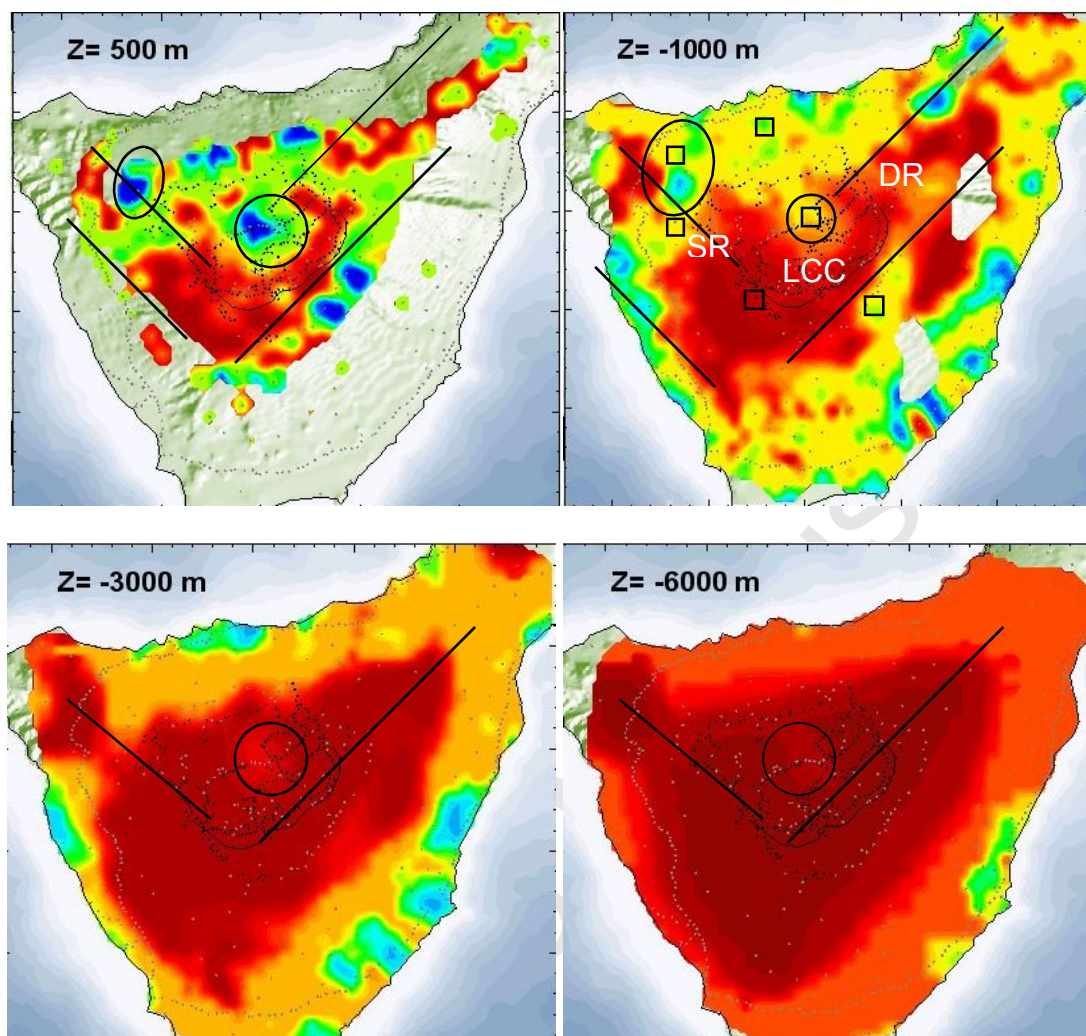


Figure 14