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J. Gattacceca, A. Lamali, P. Rochette, M. Boustie, L. Berthe. The effects of explosive-driven shocks on the natural remanent magnetization and the magnetic properties of rocks. *Physics of the Earth and Planetary Interiors*, 2007, 162 (1-2), pp.85-98. <10.1016/j.pepi.2007.03.006>. <hal-00532102>

**HAL Id: hal-00532102**

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Submitted on 4 Nov 2010

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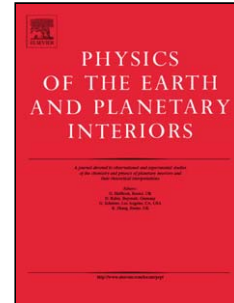


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## Accepted Manuscript

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PII: S0031-9201(07)00058-1  
DOI: doi:10.1016/j.pepi.2007.03.006  
Reference: PEPI 4803

To appear in: *Physics of the Earth and Planetary Interiors*

Received date: 20-10-2006  
Revised date: 16-3-2007  
Accepted date: 21-3-2007

Please cite this article as: Gattacceca, J., Lamali, A., Rochette, P., Boustie, M., Berthe, L., The effects of explosive-driven shocks on the natural remanent magnetization and the magnetic properties of rocks, *Physics of the Earth and Planetary Interiors* (2007), doi:10.1016/j.pepi.2007.03.006

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**The effects of explosive-driven shocks on the natural remanent magnetization and the magnetic properties of rocks**

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## 1 **1. Introduction**

2 The effects of shock waves on the natural remanent magnetization (NRM) and the intrinsic  
3 magnetic properties of geological materials remain poorly known. Still, hypervelocity impact  
4 phenomena are of primary importance in the evolution of many extraterrestrial bodies and of  
5 Earth. Hence the magnetic anomalies associated with impact basins on Mars (Hood et al.,  
6 2003), the Moon (Halekas et al., 2003) or on Earth (Pilkington and Grieve, 1992) cannot be  
7 interpreted with certainty. Also, shock-induced changes of rock magnetic properties should be  
8 taken into account when studying the magnetic signature of terrestrial impact structures.

9 Since the 1970's, efforts have been devoted to the experimental study of remagnetization of  
10 rocks upon impact, whereas little has been done about the shock-induced modifications of  
11 their intrinsic magnetic properties. Different techniques have been used to generate shock  
12 waves in a variety of geological materials: air guns accelerating aluminum or copper  
13 projectiles (Hornemann et al., 1975; Pohl et al., 1975; Martelli and Newton, 1977; Cisowski  
14 and Fuller, 1978; Srnka et al., 1979; Dickinson and Wasilewski, 2000), explosives (Hargraves  
15 and Perkins, 1969; Pesonen et al., 1997), free falling mass (Kletetschka et al., 2004).  
16 Common limitations are the difficulty in calibrating the pressure in the shocked sample and  
17 the recovery of intact samples due to brecciation of the target. On the other hand modern  
18 mechanical impact experiments using gas guns can recover intact samples with good pressure  
19 calibration (e.g. Louzada et al., in press). Recently, the use of laser shocks allowed a good  
20 estimate of the demagnetization of saturation isothermal remanent magnetization up to 5 GPa  
21 (Gattacceca et al., 2006), but the pressure gradient was limited to a few millimeters inside the  
22 sample, making it impossible to study the changes in intrinsic magnetic properties. Despite  
23 these numerous experiments, the effects of shock waves on magnetization and magnetic  
24 properties are not clearly understood in particular because of the numerous parameters that  
25 may have an influence: shock intensity and duration, ambient field, magnetic mineralogy,  
26 rheology, magnetization existing prior to shock, and temperature.

27 In this paper, we present new experiments in which different geological materials were  
28 impacted using a high-order explosive (penthrite). The shock wave was modelled numerically  
29 and we studied the effects of shock on the NRM and on the intrinsic magnetic properties of  
30 the rocks.

## 31 **32 2. Material and methods**

33 Four different lithologies were studied: a Quaternary alkaline basalt (from the Bas-Vivarais  
34 area, France, described in Rochette et al. (1993), a Tertiary quartzitic microdiorite

35 ("esterellite", from the Estérel range near the southeastern coast of France), a Permian  
36 rhyolitic pyroclastic rock from the same area (Vlag et al., 1997), and a Precambrian-  
37 Ordovician pyrrhotite-rich metamorphic schist from Wilson Terrane (Northern Victoria Land,  
38 Antarctica). This choice was guided by the differences in magnetic mineralogy and magnetic  
39 properties among these four lithologies.

40 The shock wave was driven into the samples using a high-order explosive (penthrite, PETN).  
41 The shocked samples were large blocks in the range 10-20 kg. A flat surface was cut using a  
42 circular saw. The sample was then placed in a large (0.3 m<sup>3</sup>) plastic container half-filled with  
43 gravel. A penthrite detonator (0.2 g of penthrite contained in an aluminium cylinder of  
44 external diameter 7.3 mm and thickness 0.5 mm) was fixed in direct contact with the flat,  
45 locally polished surface of the sample. The container was then filled with gravel. The  
46 penthrite was detonated remotely using a non-electric detonating wire. The outdoor  
47 experimental setting contained no metallic parts except the aluminium cylinder containing the  
48 penthrite, and no electromagnetic signal is expected during the detonation. Therefore we did  
49 not use any system of magnetic field control and assume that the magnetic field seen by the  
50 sample is the local geomagnetic field ( $I = 59^\circ$ , 46  $\mu\text{T}$ ). Each sample was oriented so that its  
51 original NRM made an angle of about  $90^\circ$  with the local field. After recovery of the shocked  
52 samples, we drilled an oriented core ( $\varnothing$  2.5 cm) perpendicular to the flat surface and centred  
53 on the impact. This core was in turn cut in 10 oriented  $4 \times 4 \times 3 \text{ mm}^3$  parallelepipeds with a  
54 wire saw (Fig. 1). Note that the upper few millimetres have been lost during the explosion for  
55 the basalt (3 mm) and the rhyolite (8 mm). In the following, the sample numbers correspond  
56 to the depth (in mm) of the centre of the 3 mm-high parallelepipeds from the original surface.  
57 All magnetic analyses were performed at CEREGE (Aix-en-Provence, France). The  
58 magnetization measurements were performed with a 2G Enterprises DC SQUID  
59 magnetometer and data processed with Paleomac software (Cogné, 2003). Magnetic  
60 susceptibility measurements and thermomagnetic analyses were performed with an Agico  
61 KLY2-CS2 apparatus. Hysteresis loops were studied with a Princeton Micromag VSM  
62 apparatus.

63

### 64 **3. Shock modelling**

65 In order to estimate the stress field induced in the samples by the explosions, a numerical  
66 simulation of the experiments was performed. For this purpose, the Radioss code  
67 [www.mecalog-group.fr] has been used in a 2D axisymmetric configuration. The explosive  
68 PETN has been modelled by a Johnson Wilkins Lee law (Wilkins, 1999):

$$69 \quad P = A \left( 1 - \frac{\omega}{R_1 V} \right) \exp(-R_1 V) + B \left( 1 - \frac{\omega}{R_2 V} \right) \exp(-R_2 V) + \frac{\omega E}{V}$$

70 where  $P$  is the pressure,  $A$ ,  $B$ ,  $R_1$  and  $R_2$  are empirical parameters,  $\omega$  is the Grüneisen  
71 coefficient,  $E$  is the internal energy and  $V$  is the relative volume.

72 The modelled target material is a basalt described by an elasto-plastic constitutive law  
73 coupled to the Mie-Grüneisen equation of state with reference to the Hugoniot through a  
74 linear relationship between shock speed  $D$  and particle velocity  $u$ :  $D = C_0 + s \cdot u$ , where  $C_0$  is  
75 the sound speed and  $s$  is an empirical parameter. All the parameters used in the simulation are  
76 listed in table 1. In the model, the diameter of the basalt has been taken large enough (6 cm)  
77 to avoid lateral reverberations on the free surface during the duration of observation of the  
78 phenomena (first pass of the shock).

79 The modelled evolution of stress versus time at various depths is given in Fig. 2. The shock  
80 applied by the explosion has an almost triangular shape, with a peak pressure of 30 GPa and a  
81 duration of about 1  $\mu$ s. The peak pressure decays to 5 GPa at only 20 mm. At depths greater  
82 than 5 mm, following the compression the material is submitted to a tensile stress, due to the  
83 lateral release initiated at the rim of the explosive charge, which generates expansion when  
84 crossing behind the main shock.

85 According to the assumptions made for this simulation (homogeneous, isotropic and non  
86 porous rock), and the variability of the intrinsic petrophysical parameters among the four  
87 different samples (for instance bulk density is 2270 kg.m<sup>-3</sup> for the rhyolite, 2700 kg.m<sup>-3</sup> for  
88 the microdiorite and the schist, and 3000 kg.m<sup>-3</sup> for the basalt), we provide here only an  
89 estimate of the stress field into the samples. Based on this estimate, the depth scale for each  
90 shocked sample can be converted to a peak pressure scale (Fig. 3).

91

#### 92 **4. Thermal modelling**

93 Because of the generation of the shock by a high explosive, an important heat flux is  
94 transferred to the target. The temperature profile applied to the target by this explosion has  
95 been calculated using the thermomechanical code Carte (Auroux and Deleignies, 2003) from  
96 the Commissariat à l'Energie Atomique (CEA, France). A maximum temperature of about  
97 4500°C is applied during about 1  $\mu$ s. A simulation of the dissipation of this heat into the  
98 target has been performed using a heat conduction simulation with FEMLAB 3.1 code (from  
99 COMSOL company). Physical properties introduced in the code are given in Table 1. The  
100 temperature profile with depth is given in figure 4 and shows that the zone affected by a  
101 significant heating (temperature > 50 °C) is restricted to the first upper micrometers. As a

102 consequence, no thermal effect is expected at all on the magnetization and the magnetic  
103 properties of the studied samples.

104 On the other hand, the shock wave itself is accompanied by a transient increase of  
105 temperature. For pressure below the elastic limit of the rock ( $\sim 5$  GPa for a basalt), this  
106 increase is negligible ( $\sim 10^\circ\text{C}$ ). For pressure greater than the elastic limit, the increase in  
107 temperature can be estimated to about  $50^\circ\text{C}$  at 10 GPa and  $100^\circ\text{C}$  at 20 GPa (e.g. Stöffler et  
108 al., 1991 and references therein).

109

## 110 **5. Pre-shock magnetization and magnetic properties of the studied materials**

111 Thermomagnetic curves (Fig. 5) allow the determination of the main magnetic carrier in  
112 the samples. The basalt has a main Curie temperature ( $T_c$ ) of  $50^\circ\text{C}$  indicating Ti-rich  
113 titanomagnetite of composition  $\text{Fe}_{0.25}\text{Ti}_{0.75}\text{O}_4$  (Dunlop and Özdemir, 1997). A second, less  
114 substituted titanomagnetite population is evidenced by the blocking temperature spectrum of  
115 thermoremanence that extends up to  $350^\circ\text{C}$  interval (Fig. 5). The microdiorite has  $T_c = 575^\circ\text{C}$   
116 indicating almost pure magnetite. This is confirmed by the observation of a Verwey transition  
117 at  $114^\circ\text{K}$ . The rhyolite has  $T_c = 650^\circ\text{C}$  indicating titanohematite. The thermomagnetic curve  
118 for the schist shows the  $\gamma$  transition (antiferromagnetic to ferrimagnetic) of hexagonal  
119 pyrrhotite at  $220^\circ\text{C}$ ,  $T_c = 280^\circ\text{C}$  of hexagonal pyrrhotite and  $T_c = 320^\circ\text{C}$  of monoclinic  
120 pyrrhotite. Hysteresis loops are in agreement with these interpretations (Fig. 6). For the  
121 rhyolite, the wasp-waisted hysteresis loop evidences the existence of a minor amount of spinel  
122 (titanomagnetite or titanomaghemite) in addition to titanohematite.

123 In view of the pressure profile in the samples (§3) and the temperature increase associated  
124 to the shock wave (§4), no thermal effect is expected on the remanent magnetization or the  
125 magnetic mineralogy of the studied rocks. Only for the basalt that has the lowest blocking  
126 temperatures, a partial TRM acquisition (about 5-10% of the total TRM, see Fig. 5) is  
127 possible up to a distance of 10 mm away from the point of explosion, the temperature increase  
128 at this distance being  $\sim 50^\circ\text{C}$ .

129 To study the natural variability of magnetic properties within the studied rocks, we drilled  
130 a  $\varnothing 2.5$  cm core parallel to the direction of the impact and more than 10 cm away from the  
131 impact point. This core was cut in smaller samples whose hysteresis properties ( $B_c$ :  
132 coercivity,  $B_{cr}$ : coercivity of remanence,  $M_s$ : saturation magnetization,  $M_{rs}$ : remanent  
133 magnetization at saturation), magnetic susceptibility, anisotropy of magnetic susceptibility  
134 (AMS) and NRM were measured. The results show no significant variation of the NRM  
135 intensity and magnetic properties with depth or mass of the sample. The only exception is a

136 50% increase of magnetic susceptibility in the upper 5 mm of the rhyolite sample. The basalt,  
137 the microdiorite and the rhyolite are magnetically homogeneous at the sampling scale (~200  
138 mg). The schist is heterogeneous at scale <1g: although the magnetic carriers have identical  
139 hysteresis properties their concentration varies (and hence the magnetic susceptibility and  
140 NRM) due to aggregation of pyrrhotite grains. The pre-shock magnetic properties are given in  
141 Table 2. In summary, we studied four samples with the following dominant magnetic  
142 mineralogy: multidomain magnetite (microdiorite), pseudo-monodomain Ti-rich  
143 titanomagnetite (basalt), pseudomonodomain monoclinic pyrrhotite (schist) and monodomain  
144 titanohematite (rhyolite). The direction of the NRM of each sample was measured on an  
145 oriented fragment before the shock. The direction and intensity of the principal axis of  
146 magnetic susceptibility (noted K1 and K3 for the maximum and minimum axes, respectively)  
147 were evaluated on the same samples.

148

## 149 **6. Magnetic properties of shocked materials**

150 After measurement of the NRM (see next section), hysteresis parameters of shocked  
151 samples were measured (Fig. 6). The plots of the hysteresis properties versus the distance to  
152 the impact point show that for the basalt, the microdiorite and the schist, the coercivity and  
153 the remanent magnetization at saturation increase close to the impact (Fig. 7). The effect,  
154 already observed by Pesonen et al. (1997) is stronger for the microdiorite sample, with a five-  
155 fold increase of both parameters. In all three cases, the intrinsic magnetic properties of all  
156 samples are clearly modified by the shock wave. Various causes could explain the  
157 modifications of magnetic properties: change in magnetic grain size, wall displacement in the  
158 magnetic grains, shock-induced defects. We observed that the shock-induced changes in  
159 hysteresis properties are not removed by stepwise heating up to 580°C. Therefore, these  
160 modifications are not attributable to stress whose effects would be annealed at rather low  
161 temperature (e.g. Van Velzen, 1992). Similarly, application of a strong field (up to 3 T) does  
162 not reset the original magnetic properties. Therefore, wall displacements cannot be the cause  
163 of the hysteresis modification. The only plausible explanation is a permanent modification of  
164 the crystalline structure of the magnetite grains, namely microfractures, lattice defects or  
165 dislocations.

166 The bulk magnetic susceptibility and the AMS (except for the rhyolite and the schist) were  
167 measured after the shock. A slight decrease of magnetic susceptibility is observed for heavily  
168 shocked (titano-)magnetite-bearing samples (Fig. 8), as already evidenced in previous works  
169 (e.g. Hargraves and Perkins, 1969).

170 It is noteworthy that for the basalt (with a pre-shock minimum susceptibility axis K3 that is  
171 about 30° away from the direction of impact), the heavily shocked samples have a higher  
172 AMS degree (Fig. 9a) and their K3 axis is parallel to the direction of impact (Fig. 9b). For the  
173 microdiorite (with a pre-shock K3 axis that is perpendicular to the direction of impact), away  
174 from the impact point the K3 axes rotate from a direction parallel to the impact towards the  
175 pre-impact direction (Fig. 9b). The intensity of the AMS for the microdiorite does not decay  
176 away from the impact point as regularly as for the basalt, probably because we observe the  
177 progressive superimposition of two different fabrics: the original one and the impact-induced  
178 one. These observations show that the impact modified the existing magnetic fabric in the  
179 first cm (peak pressure > 10 GPa) of the shocked basalt and microdiorite sample despite the  
180 absence of visible macroscopic brecciation. The physical phenomena responsible for the  
181 shock-induced magnetic anisotropy may be small-scale fracturing and deformation of  
182 magnetite grains. It is unlikely that domain walls displacement by the shock wave is  
183 responsible for even a small fraction of the shock-induced magnetic anisotropy. Indeed, the  
184 magnetic anisotropy remains unchanged after application of high magnetic fields (3 T) that  
185 would reset any domain wall displacement.

186

## 187 **7. Magnetization of shocked materials**

### 188 7.1. Demagnetization data

189 The NRM of all oriented sub-samples was measured and stepwise demagnetized with  
190 alternating field (AF) up to 140 mT (Fig. 10). We also provide demagnetization data for  
191 unshocked samples.

192 Unshocked basalt samples possess a single component of magnetization that is interpreted  
193 as the original TRM. All shocked basalt samples possess a high coercivity component with a  
194 direction close to the magnetization of the unshocked sample, that represents what is left from  
195 the original TRM. The ~20° discrepancy between the pre- and post-shock high coercivity  
196 directions is easily explained by summing the orientation uncertainties when drilling the pre-  
197 and post-shock cores and when cutting and orienting the small parallelepipeds out of the post-  
198 shock core. All shocked samples, even located at 3 cm from the impact (with an estimated  
199 peak pressure of 2 GPa), have acquired a secondary component that is completely erased at 5  
200 mT. This component of magnetization may be attributed to the shock or to the sawing  
201 process. Basalt samples located close to the impact (< 1 cm, estimated peak pressure  
202 > 10 GPa) possess a more stable secondary component isolated below 10 mT. This  
203 component may be interpreted as a shock remanent magnetization (SRM) acquired in the

204 ambient magnetic field present during the explosion. However a thermoremanent origin  
205 cannot be excluded for this low-coercivity component in view of the low blocking  
206 temperatures of the basalt (see §4). Although poorly defined, the low coercivity directions are  
207 not closely related to the ambient field at the time of impact (Fig. 11). They lie within the  
208 plane defined by the pre-shock NRM and the ambient field.

209 For the microdiorite, the unshocked sample possesses two components of magnetization: a  
210 low coercivity component, isolated below 6 mT, and a high coercivity component that is not  
211 fully demagnetized at 150 mT. This latter component is interpreted as the original TRM.  
212 Despite the low bulk coercivity of the microdiorite ( $B_c = 2$  mT), its natural TRM is carried  
213 by a high-coercivity fraction. The shocked samples still possess most of the original TRM.  
214 The shocked samples located close to the impact ( $< 1.5$  cm, peak pressure  $> 6$  GPa) possess a  
215 secondary component isolated below 30 mT and is interpreted as a SRM. The SRM directions  
216 are scattered.

217 For the rhyolite, the unshocked sample possesses a single component of magnetization  
218 carried by hematite. It is not clear if this original magnetization is a TRM or a chemical  
219 remanent magnetization (see discussion in e.g. Vlag et al., 1997). All shocked samples  
220 possess a high-coercivity component of similar direction than the original magnetization. The  
221 shocked samples closest to the impact point ( $< 1$  cm, peak pressure  $> 10$  GPa) also possess a  
222 low-coercivity component (SRM) isolated below  $\sim 30$  mT. The SRM directions are scattered.

223 The pyrrhotite-bearing schist did not provide stable directions of NRM when demagnetized  
224 by alternating field (Fig. 10). Therefore we will not discuss the possible remagnetization in  
225 the shocked schist.

226

## 227 7.2. Remagnetization and demagnetization by shock

228 In our experiments, since the directions of the original TRM and the SRM are different we  
229 cannot simply compare the intensities of the pre- and post-shock NRM. For the basalt for  
230 instance, the intensity of the total post-shock NRM is much lower than the sum of the  
231 intensities of the two components of magnetization (e.g. sample B5 of Fig. 10). In figure 12,  
232 we plot the scalar sum of the intensities of the stepwise-demagnetized magnetization vectors  
233 for basalt and microdiorite samples. The intensities are normalized to the saturation  
234 magnetization of each sample to take into account the natural variability in ferromagnetic  
235 mineral concentration. After the shock, the sum of the intensities of the different components  
236 of magnetization is weaker than before the shock for the basalt, but higher for the  
237 microdiorite. The basalt is globally demagnetized whereas the microdiorite is globally

238 remagnetized by the shock. To our knowledge, this is the first time that shock experiments  
239 actually increase the original NRM of a rock. The increase of the magnetization of the  
240 microdiorite after the shock does not mean that SRM acquisition is more efficient than TRM  
241 acquisition. This is due to the increased capacity of the microdiorite to acquire a remanent  
242 magnetization after the shock (i.e. increase of  $M_r/M_s$  by a factor up to five) and primarily to  
243 the fact that the original remanence was carried almost entirely by the grains with high  
244 coercivity so that the shock added a secondary low-coercivity SRM to the original high-  
245 coercivity TRM that was not much affected by the shock.

246 In view of the rather unstable behaviour of the NRM of the microdiorite and the schist  
247 under AF demagnetization and the high resistance of the NRM of the rhyolite to AF  
248 demagnetization (Fig. 10), only the basalt was suitable for a detailed study of the effect of  
249 shock on the remanence. For the basalt that has a simple demagnetization behaviour under  
250 alternating field and a TRM carried by grains spanning the whole coercivity spectrum, we can  
251 compare the magnetization slope ( $dNRM/dAF$ ) for shocked and unshocked samples as a  
252 function of the peak alternating field (Fig. 13). Above 10-15 mT, all shocked samples located  
253 more than 15 mm from the impact point (peak pressure < 6 GPa) closely follow the curve for  
254 the unshocked basalt. Contrarily, heavily shocked samples (B5 to B14) plot below the  
255 unshocked samples between 10 and 40 mT. Above 40 mT all curves are identical. On the  
256 interval 10-40 mT, these heavily shocked samples have a magnetization that has the same  
257 direction as the original TRM. This means that the original TRM of samples B5 to B14 has  
258 been partly demagnetized on the 10-40 mT coercivity window. If a SRM or a TRM has been  
259 acquired on the same coercivity window during the shock, it is small enough not to be  
260 noticeable on the demagnetization plot, i.e. it is at most a few % of the original TRM. We can  
261 quantify the shock demagnetization of the original TRM on the 10-40 mT coercivity interval  
262 by comparing the amount of NRM demagnetized between 10 mT and 40 mT for pre- and  
263 post-shock samples (fig. 14). Up to a distance of 14 mm from the impact, corresponding to a  
264 peak pressure of about 7 GPa (sample B14), the original TRM of shocked samples has been  
265 significantly demagnetized. For sample B5 that has suffered peak pressure around 20 GPa, the  
266 original TRM has been divided by up to a factor 5. It is noteworthy that attempting a  
267 paleointensity measurement on such a shocked sample would be problematic since it is  
268 impossible to notice that part of the TRM has been demagnetized by shock. Although it is not  
269 clear how this demagnetization is dispatched along the blocking temperature spectrum, the  
270 paleointensity experiment would at best provide an underestimated value (up to a factor 5 in  
271 the case of sample B5 for instance).

272 Generally speaking the effect of the explosive-driven shocks on the NRM of the four  
273 studied lithologies appear rather weak compared to a number of previous studies performed  
274 on isothermal remanent magnetization (IRM). Complete demagnetization of the IRM of a  
275 titano-magnetite bearing basalt shocked with a laser pulse was observed at about 2 GPa  
276 (Gattacceca et al. 2006). Shock experiments with a gas gun on pyrrhotite-bearing samples  
277 showed a 80% demagnetization of IRM at only 0.5 GPa (Louzada et al., in press). The large  
278 differences are attributed to the fact that NRM and IRM react differently to shock waves.

279

## 280 **8. Conclusion**

281 Experimental shocks of four rocks with different lithology and magnetic mineralogy show  
282 that, in most cases (only the hematite case is not conclusive), the intrinsic magnetic properties  
283 of the rock are permanently modified by the shock wave. A remarkable effect, not  
284 documented in previous studies, is the capacity of the shock wave to superimpose a new  
285 fabric (with a minimum susceptibility axis parallel to the direction of impact) to the original  
286 magnetic fabric of the rock. Multidomain magnetite-, pseudo-single domain titanomagnetite-  
287 and monoclinic pyrrhotite-bearing rocks show a noticeable increase of their coercivity for  
288 pressure above 10 GPa. These changes are not annealed even at high temperature (580 °C).  
289 They are attributed to fracturing and/or dislocations of the ferromagnetic grains. This  
290 fracturing is also responsible for the appearance of a shock-induced anisotropy of magnetic  
291 susceptibility. These results show that the magnetic properties of meteorites which are  
292 commonly shocked to pressures well above 10 GPa (e.g. Martian meteorites, Nyquist et al.,  
293 2001) may not be representative of the magnetic properties of their parent body.

294 For (titano)magnetite bearing rocks, we observe both a shock demagnetization of the original  
295 TRM for magnetic grains with coercivity up to 40 mT, and a possible shock magnetization for  
296 grains with coercivity up to 10 mT. NRM appears to be much more resistant to shock than  
297 IRM (probably because they have different coercivity spectra), implying that more work is  
298 needed on the effects of shock on natural magnetization. As observed by Cisowski and Fuller  
299 (1978), the demagnetizing effects of the shock wave depend closely on the coercivity  
300 spectrum of the grains carrying the original remanence, and the shock-remagnetizing effect  
301 depend on the presence (or creation by the impact itself) of low-coercivity magnetic grains.  
302 With an impact occurring in an ambient magnetic field of similar intensity to the original  
303 magnetizing field, the post-shock magnetization may be higher or lower than the pre-shock  
304 magnetization depending on these two factors.

305 To summarize, although an impact occurring after dynamo shutdown will indeed demagnetize  
306 the crust to a variable extent, an impact occurring while the dynamo is still active may  
307 demagnetize the crust almost as efficiently or conversely may lead to the situation where the  
308 shocked crust has a stronger magnetization than the surrounding rocks (depending on the  
309 efficiency of SRM acquisition, and on the nature of the original NRM). Therefore, it appears  
310 difficult to draw conclusions about the dynamo history of a planet by studying the magnetic  
311 anomalies above its impact basins, unless the anomalies provide information about the  
312 magnetization of the rocks heated during the impacts (if they are preserved). The decisive clue  
313 to the presence of an active dynamo at the time of impact is the presence or absence of a  
314 thermoremanence carried by the volume of rocks heated above blocking temperatures during  
315 the impact. Similarly, it may be difficult to determine which magnetic phase dominates the  
316 crust based only on the demagnetization pattern around impact basins.

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## Acknowledgements

This work was supported by the Programme National de Planétologie (INSU/CNES, France). We acknowledge very constructive comments by S. T. Stewart-Mukhopadhyay and an anonymous reviewer.

## References

- Auroux, E., and Deleignies, M., 2003. Sensitivity of the EOS of detonation products to the parametrization of the intermolecular potential for pure fluids in thermomechanical calculations. Proceedings of 5<sup>th</sup> international Symposium of high dynamic pressures, CEA eds, p. 281.
- Cogné, J.P., 2003. PaleoMac: a Macintosh™ application for treating paleomagnetic data and making plate reconstructions. *Geochem. Geophys. Geosyst.*, 4, doi:10.1029/2001GC000227.
- Cisowski, S.M., and Fuller, M., 1978. The effect of shock on the magnetism of terrestrial rocks. *J. Geophys. Res.*, 83: 3441-3456.
- Dickinson, T.L., and Wasilewski, P., 2000. Shock magnetism in fine particle iron. *Meteor. Planet. Sci.*, 35: 65-74.
- Dunlop, D.J., and Özdemir, Ö., 1997. *Rock magnetism: fundamentals and frontiers*. Cambridge University Press, Cambridge, 573 pp.
- Gattacceca, J., Boustie, M., Weiss, B.P., Rochette, P., Lima, E., Fong, L.E., and Baudenbacher, F., 2006. Investigating impact demagnetization through laser impacts and SQUID microscopy. *Geology*, 34: 333-336.
- Hargraves, R.B., and Perkins, W.E., 1969. Investigations of the effect of shock on natural remanent magnetism. *J. Geophys. Res.* 74: 2576-2589.
- Halekas, J.S., Lin, R.P., and Mitchell, D.L., 2003. Magnetic fields of lunar multi-ring impact basins. *Meteor. Planet. Sci.*, 38: 565-578.
- Hood, L., Richmond, N.C., Pierazzo, E., and Rochette, P., 2003. Distribution of crustal magnetic fields on Mars: shock effects of basin-forming impacts. *Geophys. Res. Lett.*, 30, doi: 10.1029/2002GL016657.
- Kletetschka, G., Eonnerney, J.E.P., Ness, N.F., and Acuña, M.H., 2004. Pressure effects on martian crustal magnetization near large impact basins. *Meteor. Planet. Sci.*, 39: 1839-1848.

- Louzada, K., Stewart, S.T., and Weiss, B.P., 2005. The effect of shock on the magnetic properties of pyrrhotite, the Martian crust, and meteorites. *Geophys. Res. Lett.*, in press.
- Martelli, G., and Newton, G., 1977. Hypervelocity cratering and impact magnetisation of basalt. *Nature*, 269: 478-480.
- Nyquist, L.E., Bogard, D.D., Shih, C.Y., Greshake, A., Stöffler, D., Eugster, O., 2001. Ages and geologic histories of Martian meteorites. *Space Sci. Reviews*, 96: 105-164.
- Pesonen, L.J., Deutsch, A., Hornemann, U., and Langenhorst, F., 1997. Magnetic properties of diabase samples shocked experimentally in the 4.5 to 35 GPa range. 28<sup>th</sup> Lunar and Planetary Science Meeting: 1087-1088.
- Pilkington, M., and Grieve, R.A.F., 1992. The geophysical signature of terrestrial impact craters. *Rev. Geophysics*, 30: 161-181.
- Pohl, J., Bleil, U., and Hornemann, U., 1975. Shock magnetization and demagnetization of basalt by transient stress up to 10 kbar. *J. Geophysics*, 41: 23-41.
- Rochette, P., Bertrand, H., Braun, C., and Berger, E., 1993. La province volcanique Pléistocène supérieur du Bas-Vivarais (Ardèche, France) : propagation de fentes crustales en échelons ? *C. R. Acad. Sci. Paris*, 316: 913-920.
- Srnka, L.J., Martelli, G., Newton, G., Cisowski, S.M., Fuller, M.D., and Schaal, R.B., 1979. Magnetic field and shock effects and remanent magnetization in a hypervelocity impact experiment. *Earth Planet. Sc. Lett.*, 42: 127-137.
- Stöffler, D., Keil, K., and Scott, E.R.D., 1991. Shock metamorphism of ordinary chondrites. *Geochim. Cosmochim. Acta*, 55: 3845-3867.
- Van Velzen, A.J., and Zijderfeld, J.D.A., 1992. A method to study alterations of magnetic minerals during thermal demagnetization applied to a fine-grained marine marl (Trubi formation, Sicily). *Geophys. J. Int.*, 110: 79-90.
- Vlag, P., Vandamme, D., Rochette, P., and Spinelli, K., 1997. Paleomagnetism in the Esterel rocks: a revisit 22 years after the thesis of H. Zijderfeld. *Geologie en Mijnbouw*, 76: 21-33.
- Wilkins, M.L., 1999. Computer simulation of dynamic phenomena. Springer eds, New York, 246 pp.

**Figure captions**

- Figure 1 - Sketch illustrating the different steps of the sampling after explosion. The wire used to saw the parallelepipeds is 220  $\mu\text{m}$  in diameter.
- Figure 2 - Modelled pressure versus time at various depths in the rock straight below the explosion. The modelling was performed with a Johnson Wilkins Lee law for the explosive and the Radioss code for the shock wave propagation in the sample (see text).
- Figure 3 - Peak pressure versus depth in the rock straight below the explosion. The pressure values are deduced from the shock wave modelling presented in Fig. 2. The solid line is computed with a density of 2300  $\text{kg.m}^{-3}$ , the dashed line is computed with a density of 3000  $\text{kg.m}^{-3}$ .
- Figure 4 - Modelled peak temperature versus depth in the sample straight below the explosion, considering the amount of heat transferred to the to by the explosion of the PETN detonator. Dashed line is at 50°C. The modelling was performed using Carte and FEMLAB codes (see text).
- Figure 5 - Thermomagnetic curves for the four studied lithologies (susceptibility vs. temperature). Empty circles indicate the beginning of the heating curves. Heating = thick lines, cooling = thin lines. For the basalt and the microdiorite, progressive TRM acquisition curves are also plotted (dotted line, with standard deviation when available).
- Figure 6 - Representative hysteresis loops for shocked and unshocked samples of the four studied lithologies. The loops have been corrected for the paramagnetic slope.
- Figure 7 - Hysteresis properties vs. depth for the shocked samples ( $M_r$ : remanent magnetization at saturation,  $M_s$ : saturation magnetization,  $B_{cr}$ : coercivity of remanence,  $B_c$ : coercivity). Dashed line is the mean value for unshocked samples (grey band = one standard deviation).
- Figure 8 - Low-field specific magnetic susceptibility of impacted samples. Pre-impact mean are indicated with one standard deviation (grey bands).
- Figure 9 - a-c) Degree of anisotropy of magnetic susceptibility ( $P_{\text{ams}}$  = ratio of maximum to minimum principal susceptibility values) vs. distance to the impact. Pre-impact means are indicated (dashed line) with one standard deviation (grey bands). b-d) Lower-hemisphere equal area stereographic projection of minimum axis of magnetic susceptibility  $K_3$ .
- Figure 10 - Orthogonal demagnetization plots of oriented samples of the shocked rocks. Open and solid symbols represent projections of the magnetization vector on vertical and horizontal planes, respectively. Demagnetization plots for oriented unshocked samples are given for comparison. Demagnetization steps are 5, 10, 15, ..., 50, 60, 70, 80, 100, 120, 140 mT.
- Figure 11 - Equal-area stereographic projection of the directions of magnetization ( $0^\circ$  = North and  $90^\circ$  = East in Fig. 10). Open (solid) dots represent projection onto upper (lower) hemisphere. Open ellipses are the projections of the 95% confidence cones about the directions. Diamond: ambient field during the impact;  $K_1$  (resp.  $K_3$ ) : maximum (resp. minimum) susceptibility principal axis, large labelled circles: low-coercivity components, small circles: high-coercivity components; large labelled star: pre-impact direction of magnetization (determined on a separate core), star: mean direction of the high-coercivity components of shocked samples.
- Figure 12 - Sum of demagnetized vectors (scalar intensity) normalized by saturation magnetization vs. distance from the impact. The dotted line represents an unshocked sample.
- Figure 13 - NRM slope ( $d\text{NRM}/d\text{AF}$ ) as a function of alternating field for shocked basalt samples. The curve for pre-shock NRM (dashed line) is plotted for comparison. Curves are normalized to saturation magnetization ( $M_s$ ) to take into account the possible variability in titanomagnetite content. Data for demagnetizing field below 10 mT are not represented

because the origin of the low coercivity component is not ascertained for the basalt (see text).

Figure 14 - Mean degree of shock demagnetization of the basalt on the 10-40 mT coercivity window as a function of distance to impact. This value is derived from the ratio of the NRM moment demagnetized between 10 mT and 40 mT before and after shock.

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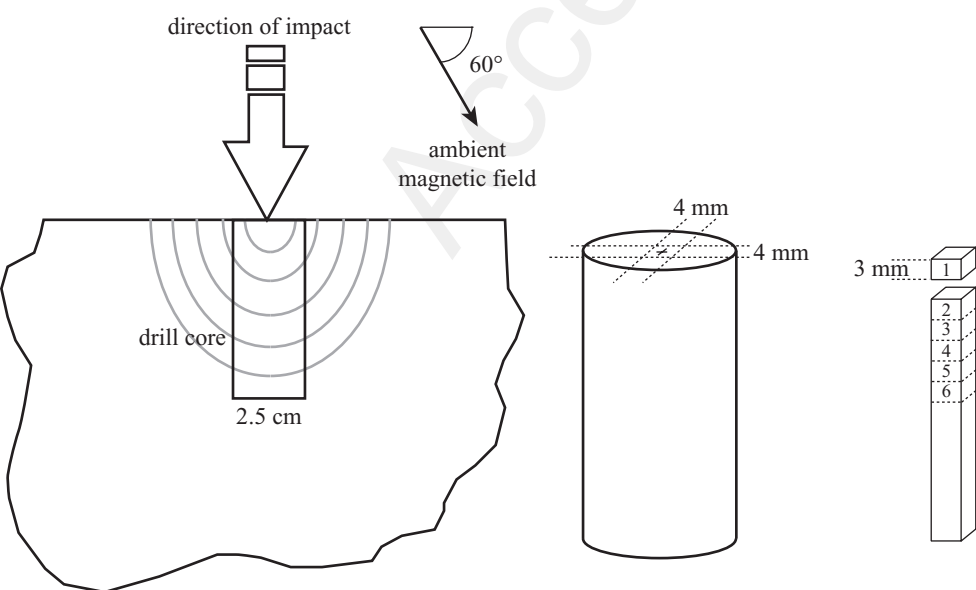


Figure 1

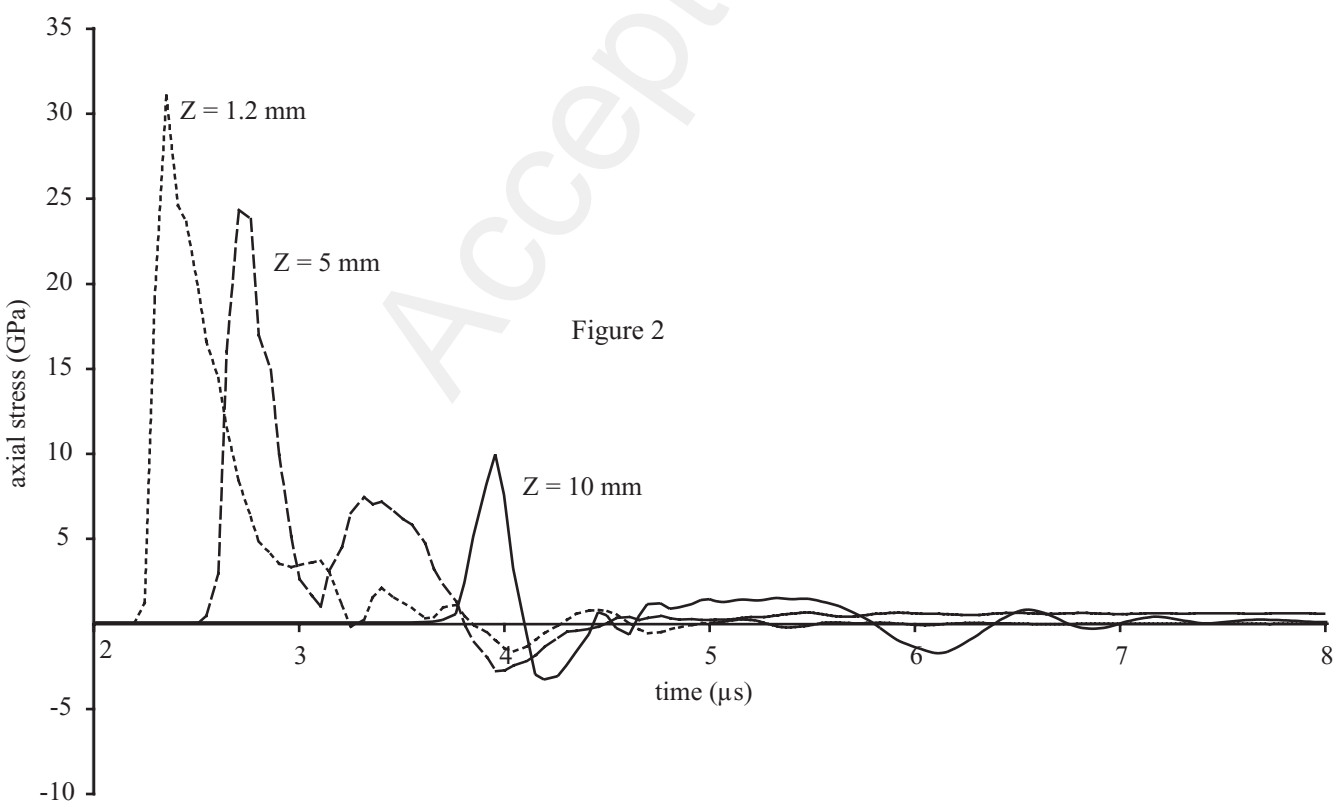
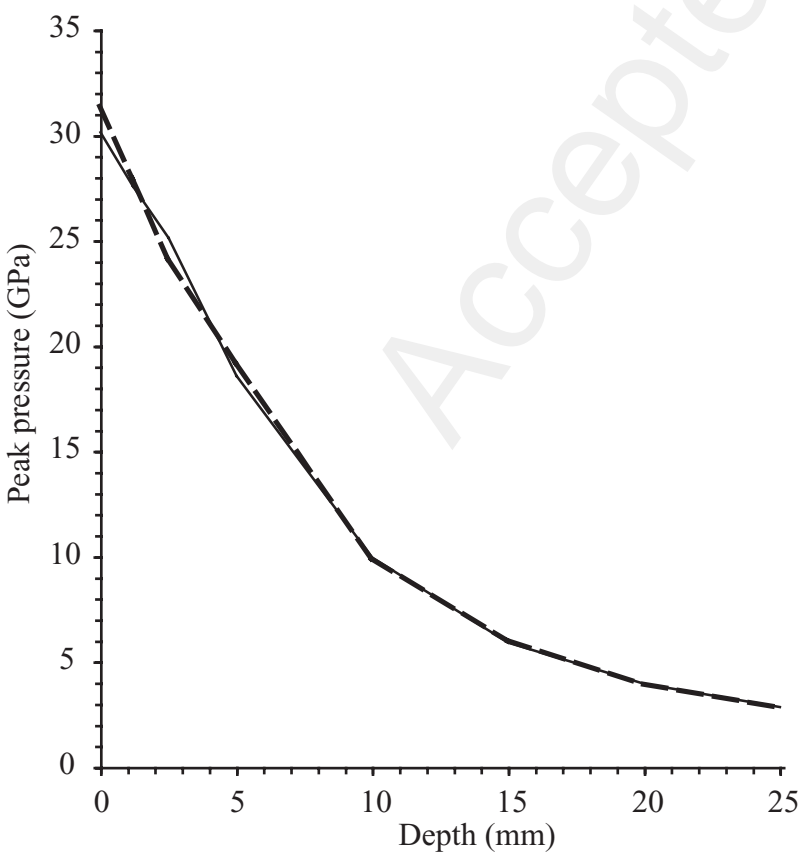


Figure 2



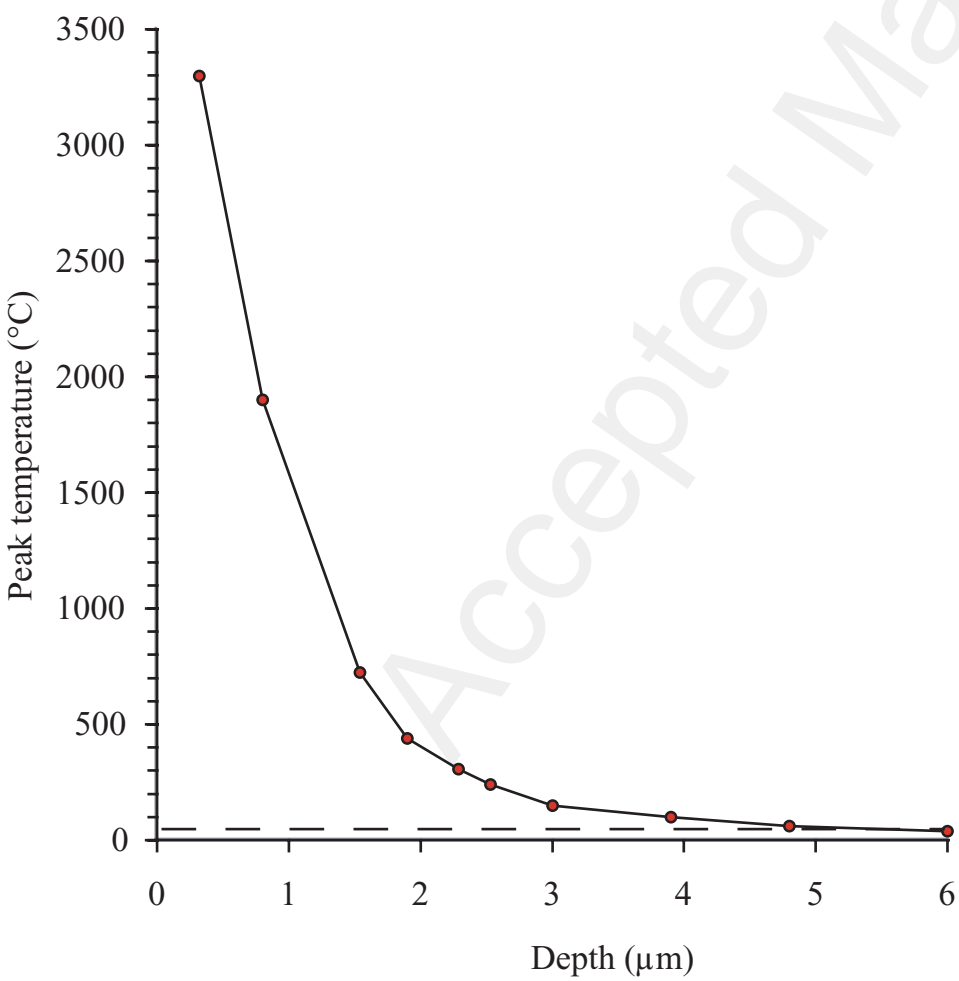
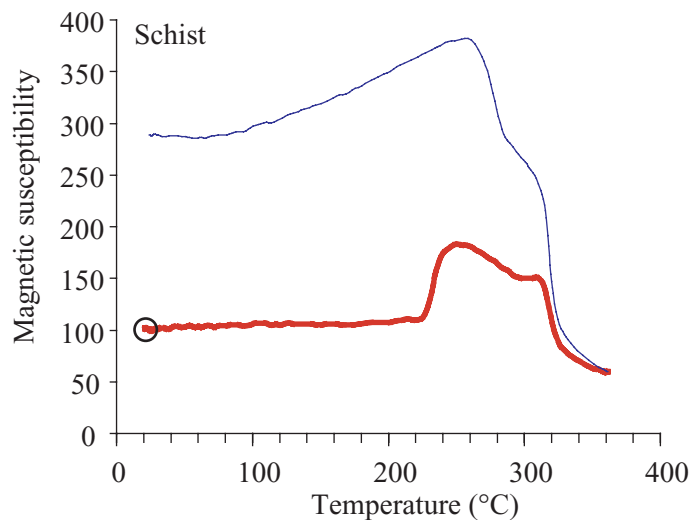
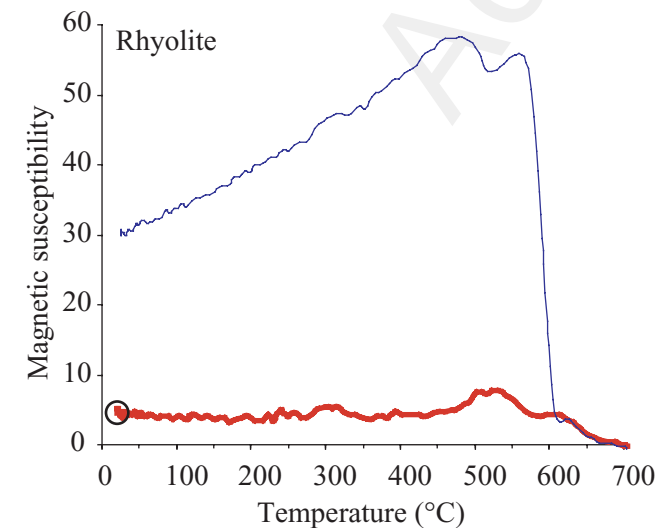
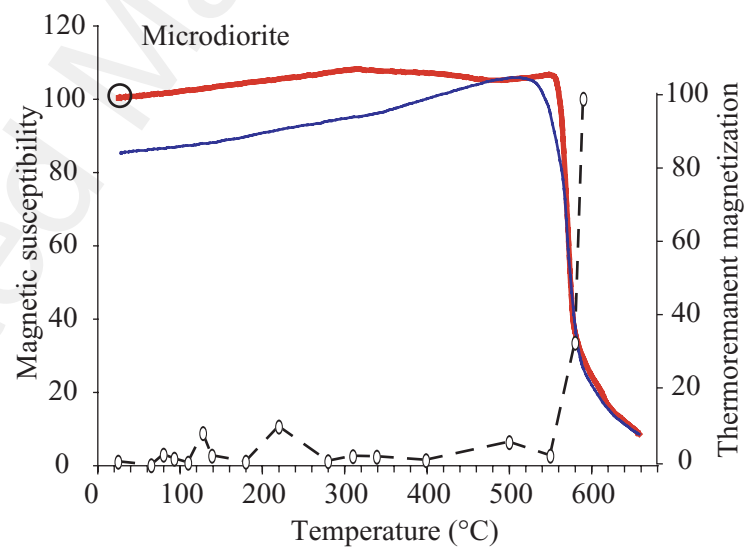
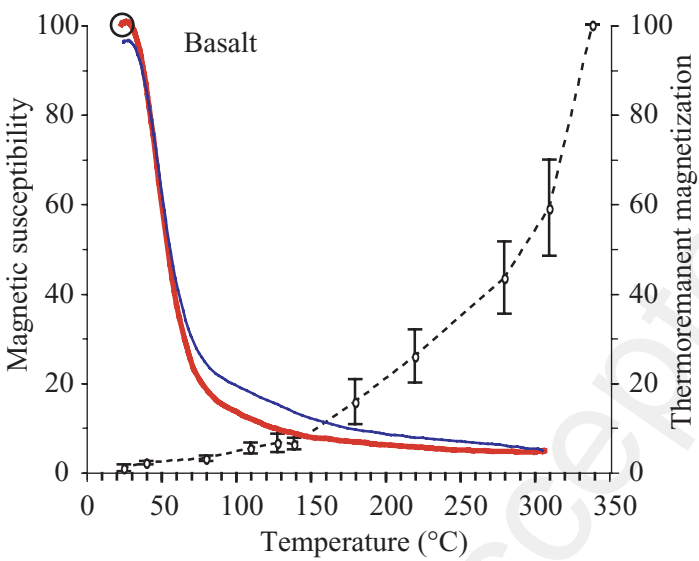
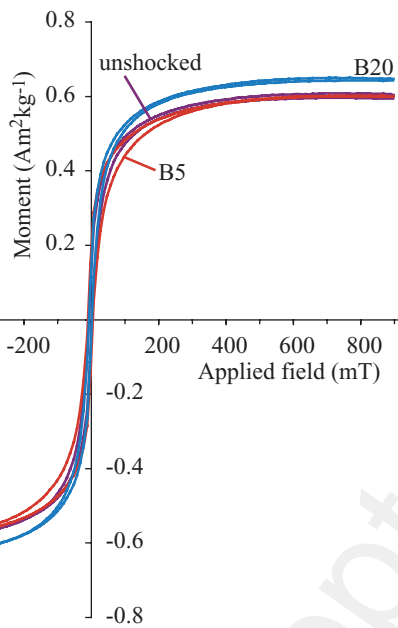


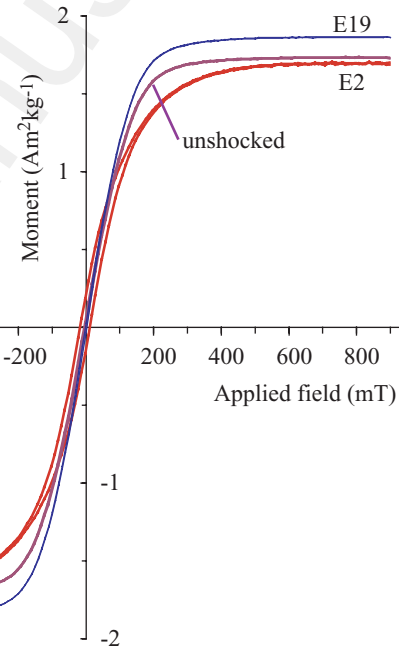
Figure 4



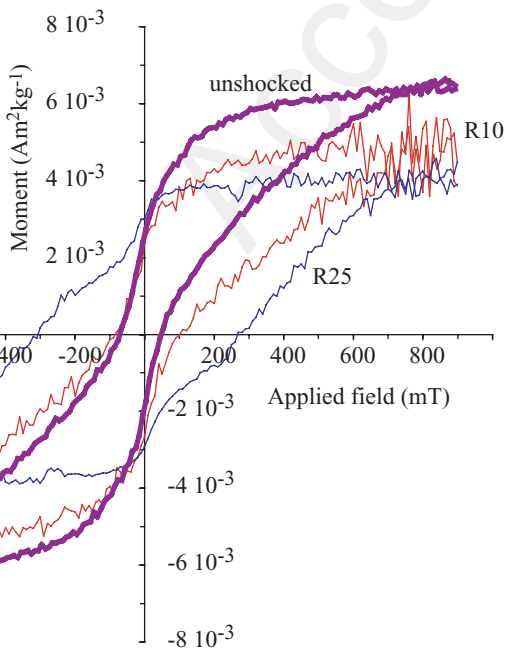
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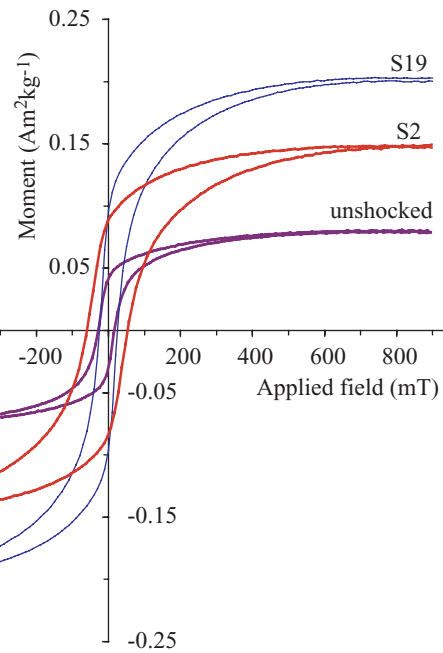
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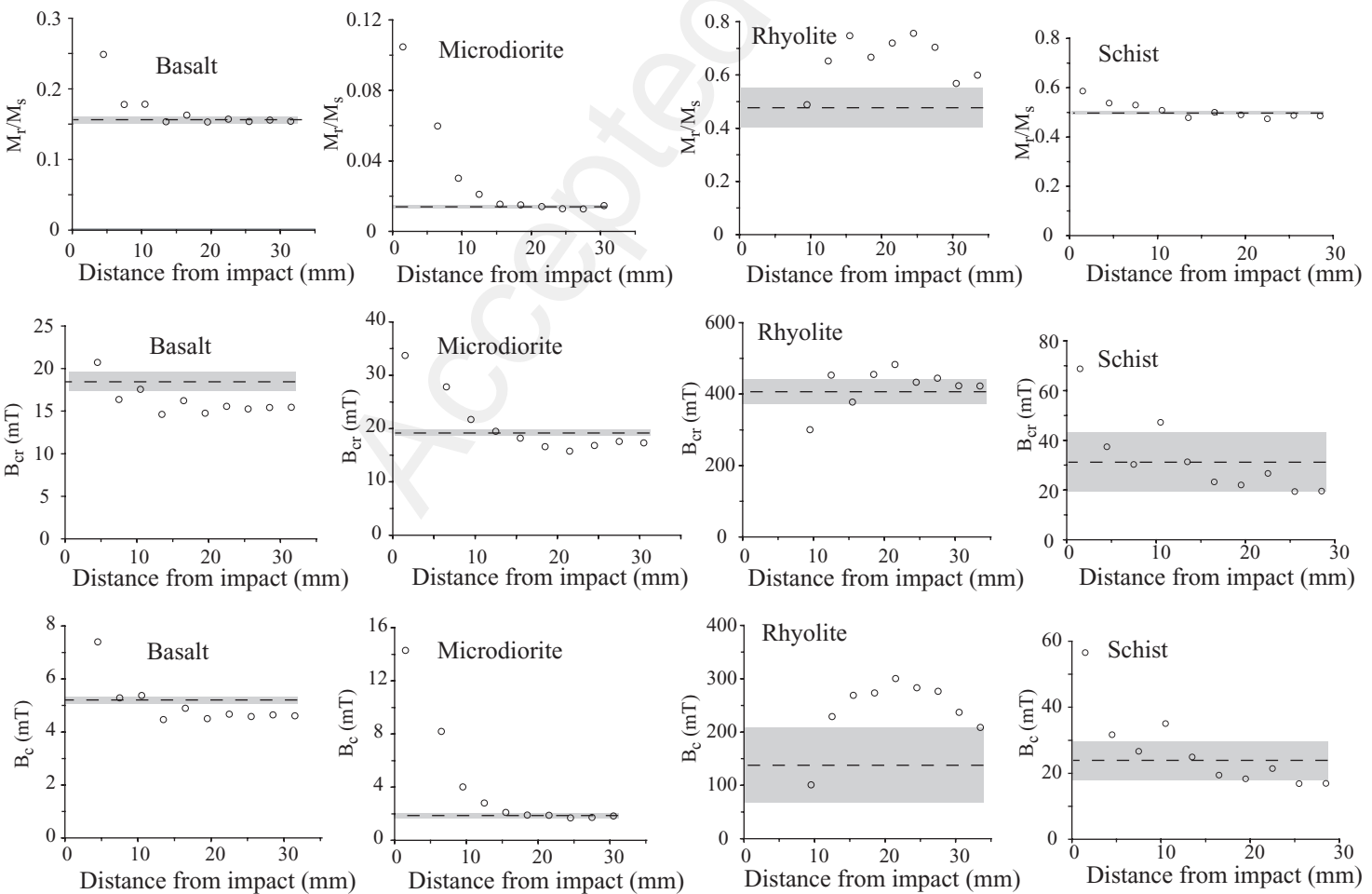


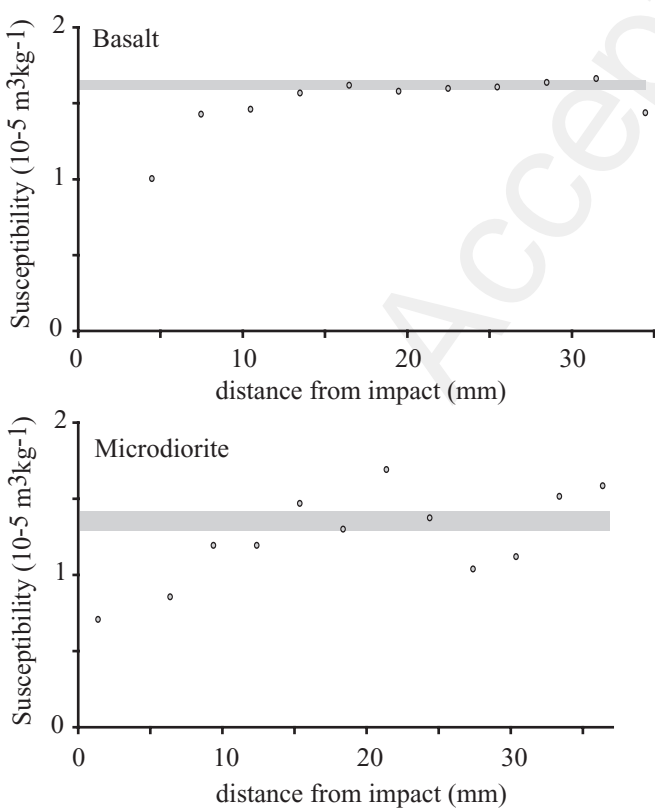
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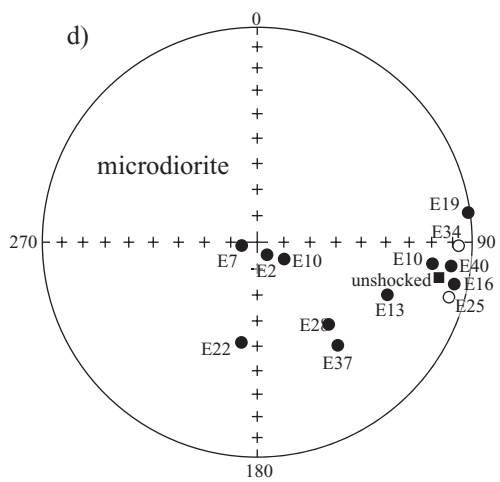
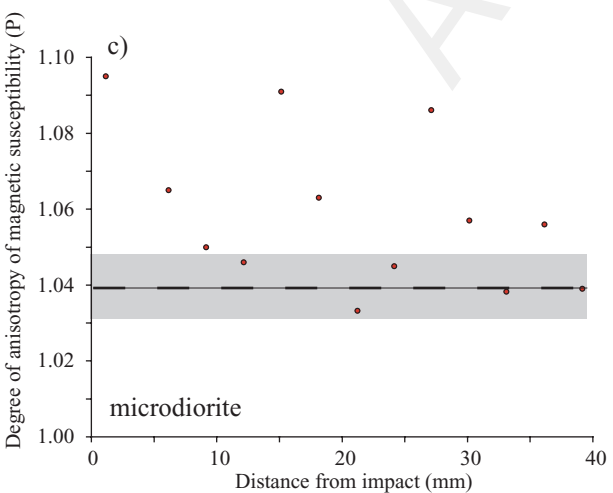
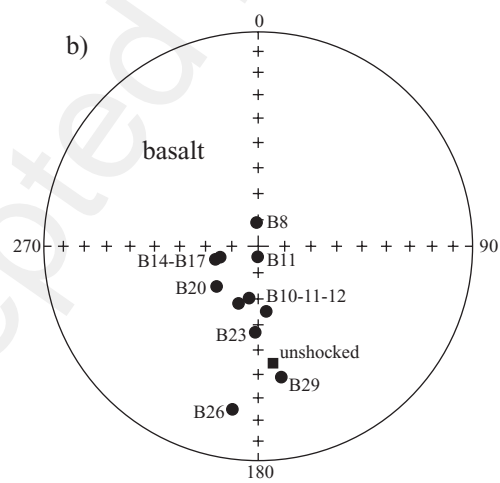
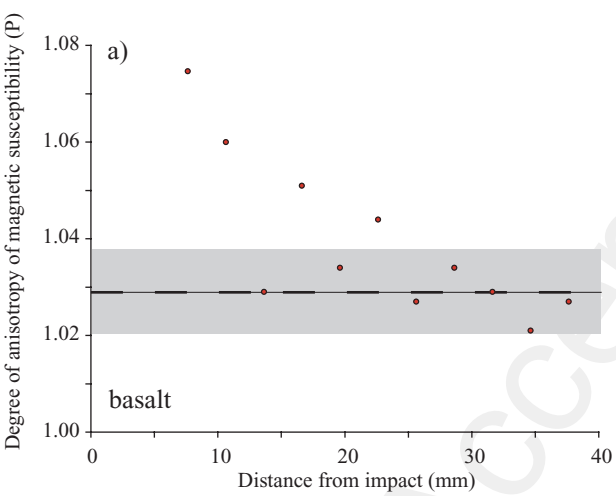


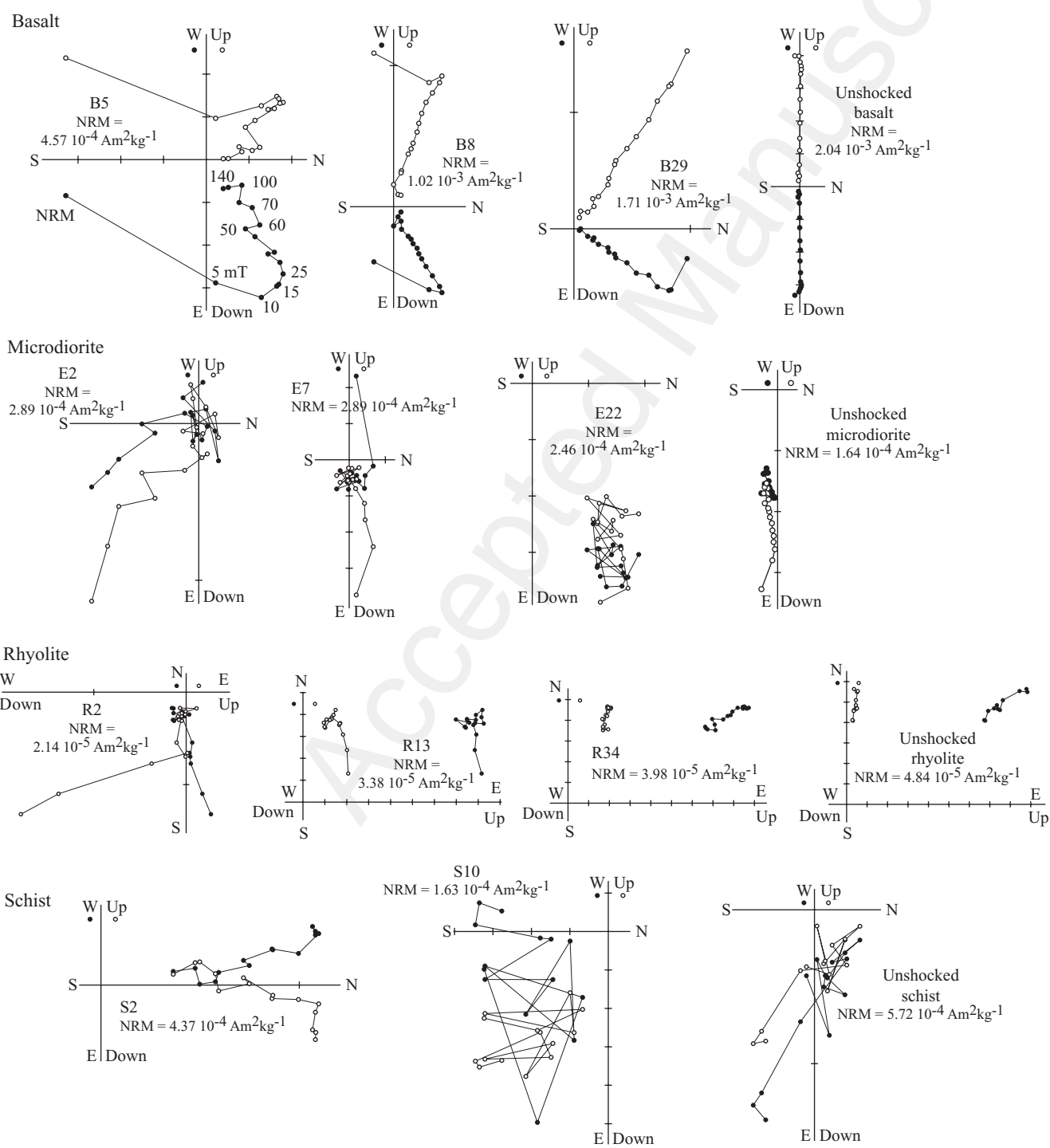
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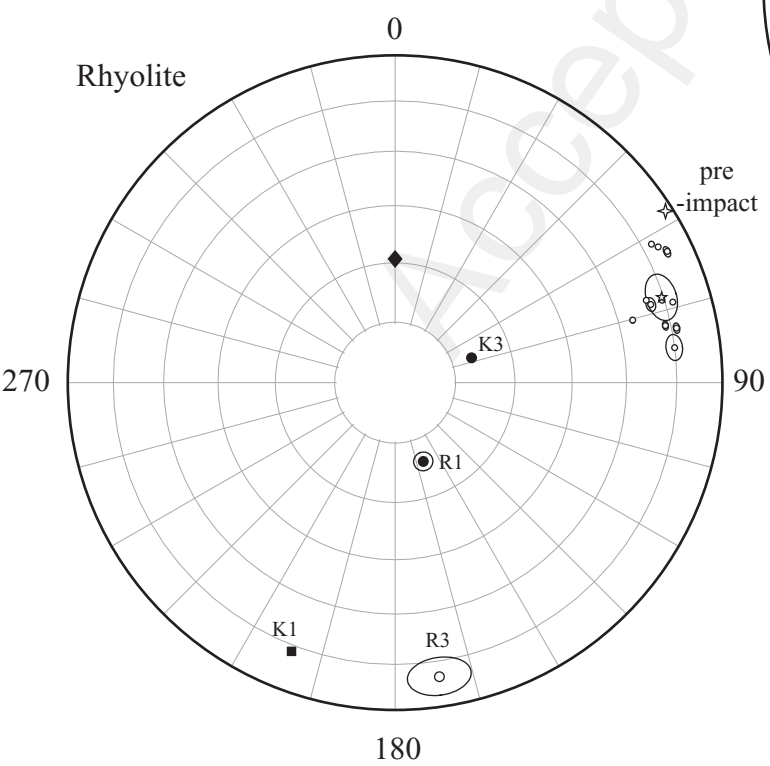
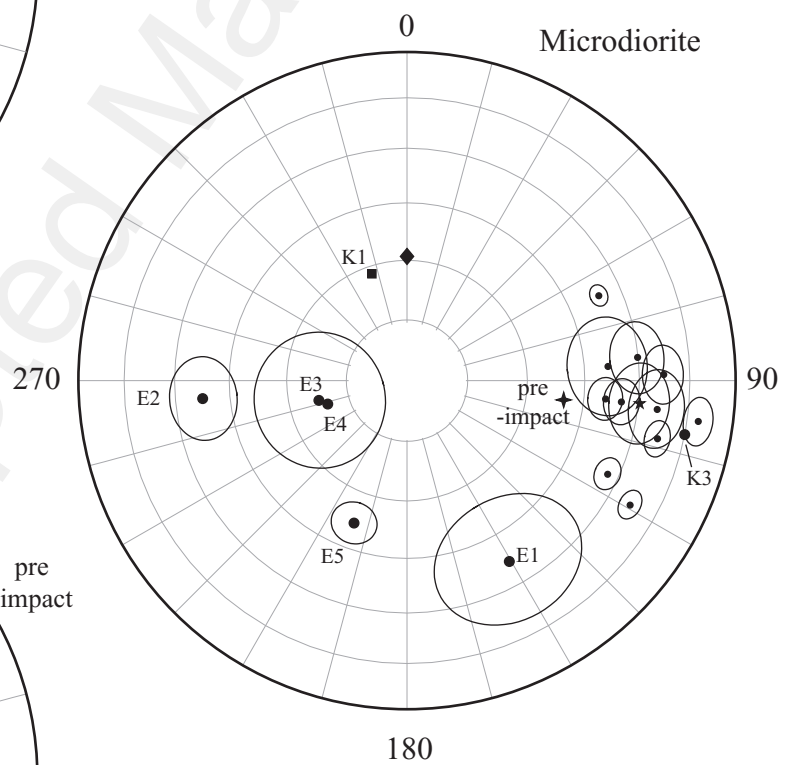
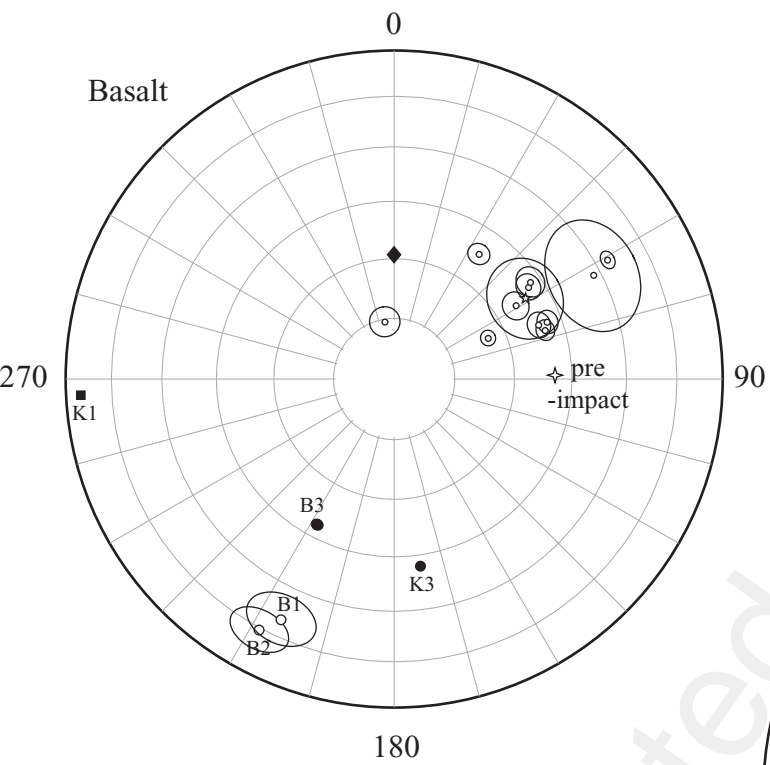


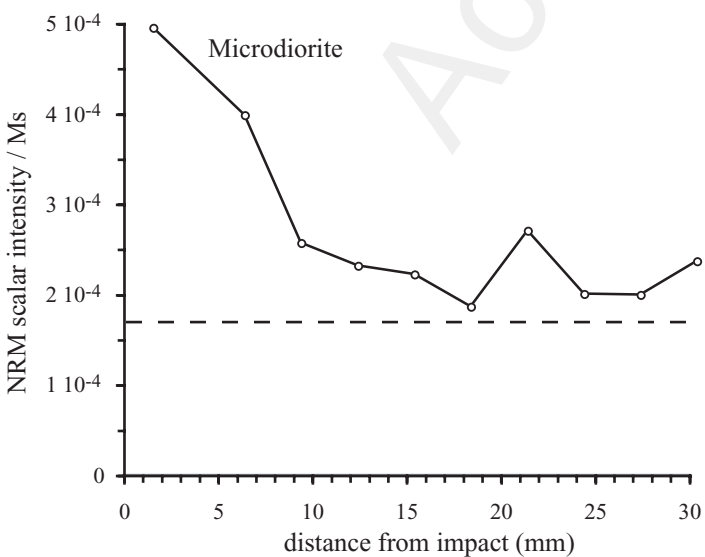
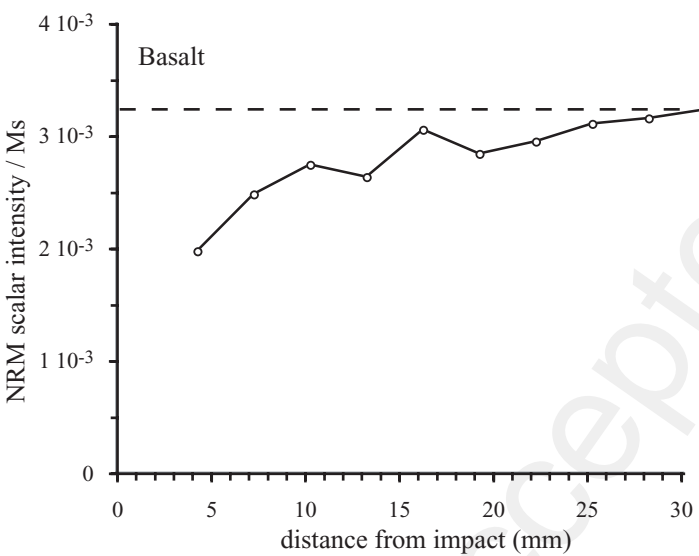


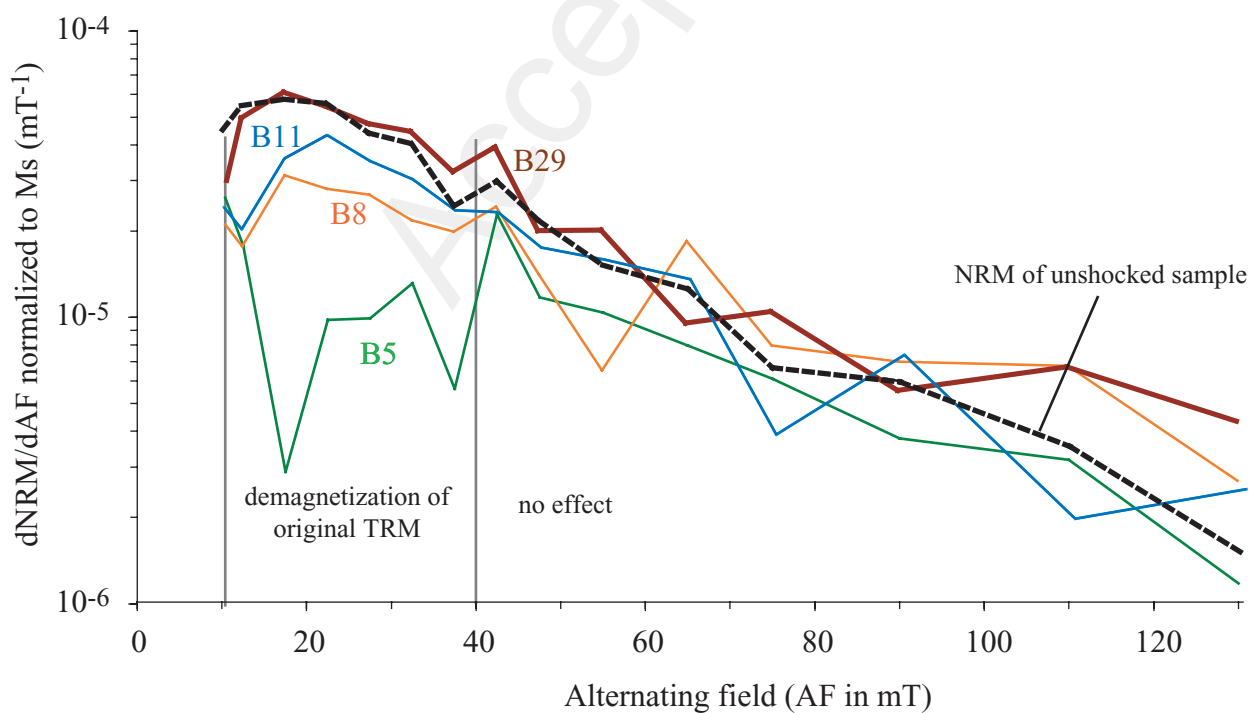


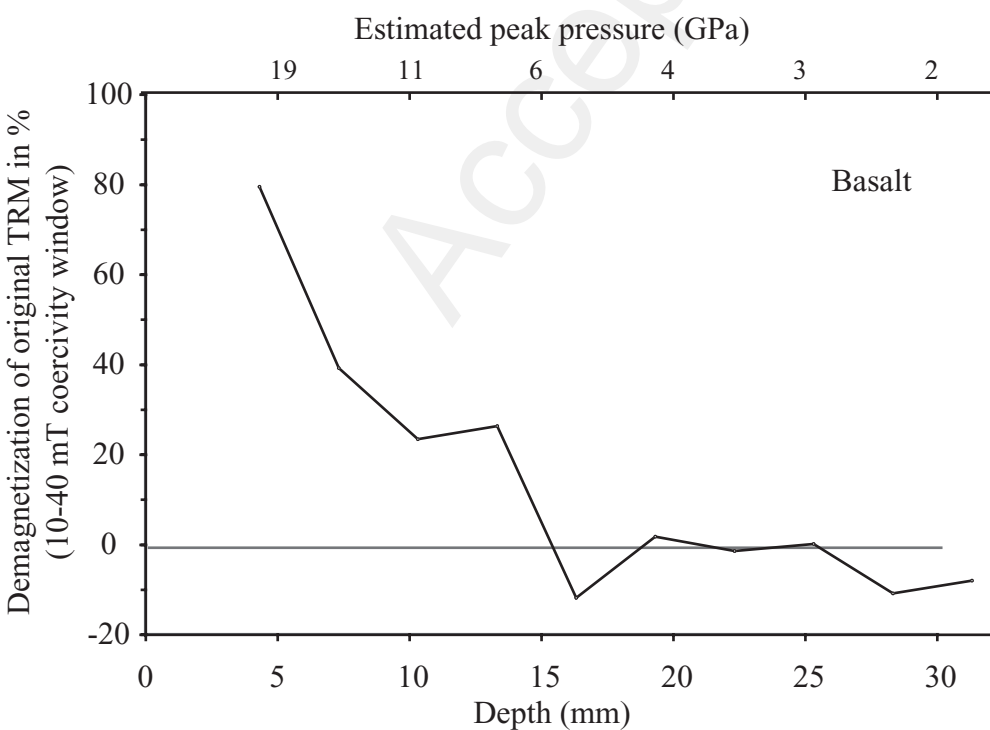












**Tables**

Table 1 - Physical parameters used in the shock and thermal simulations.

Explosive (PETN)	$\rho_0$ (g.cm <sup>-3</sup> )	D (m.s <sup>-1</sup> )	P <sub>CJ</sub> (GPa)	A (GPa)	B (GPa)	R <sub>1</sub>	R <sub>2</sub>	$\omega$
	1.77	8300	33.35	617	16.93	4.4	1.2	0.25
Basalt	$\rho_0$ (g.cm <sup>-3</sup> )	C <sub>0</sub> (m.s <sup>-1</sup> )	S	$\Gamma$	Y <sub>0</sub> (GPa)	k (W.m <sup>-1</sup> .K <sup>-1</sup> )	C (J.kg.K <sup>-1</sup> )	
	2.88	4800	1.34	2	0.3	1.99	1000	

$\rho_0$ : density; D: shock speed; P<sub>CJ</sub>: Chapman-Jouguet pressure; A, B, R<sub>1</sub>, R<sub>2</sub>: JWL parameters associated to the JWL equation mentioned in the text (Wilkins, 1999);  $\omega$ : Grüneisen coefficient; C<sub>0</sub>: sound speed; S: empirical parameter;  $\Gamma$ : Grüneisen coefficient; Y<sub>0</sub>: yield stress; k: thermal conductivity; C: heat capacity.

Table 2 - Magnetic properties of materials before shock

	Basalt	Microdiorite	Rhyolite	Schist
<b>Hysteresis</b>	n=10	n=12	n=10	n=10
Mr (Am <sup>2</sup> kg <sup>-1</sup> )	9.78 ± 0.67 10 <sup>-2</sup>	2.38 ± 0.20 10 <sup>-2</sup>	2.31 ± 0.22 10 <sup>-3</sup>	4.48 ± 2.90 10 <sup>-2</sup>
Ms (Am <sup>2</sup> kg <sup>-1</sup> )	6.22 ± 0.46 10 <sup>-1</sup>	1.28 ± 0.14 10 <sup>-1</sup>	4.91 ± 0.72 10 <sup>-3</sup>	9.14 ± 6.02 10 <sup>-2</sup>
Mr/Ms	1.57 ± 0.03 10 <sup>-1</sup>	1.39 ± 0.07 10 <sup>-2</sup>	4.77 ± 0.73 10 <sup>-1</sup>	4.99 ± 0.28 10 <sup>-1</sup>
Bc (mT)	5.19 ± 0.16	1.86 ± 0.08	137 ± 69	23.8 ± 5.9
Bcr (mT)	18.4 ± 1.1	19.1 ± 0.7	406 ± 35	31.3 ± 11.7
Bcr/Bc	3.54 ± 0.14	10.3 ± 0.4	3.56 ± 1.47	1.29 ± 0.12
<b>Susceptibility</b>	n=22	n=28	n=9	n=7*
Susceptibility (m <sup>3</sup> kg <sup>-1</sup> )	1.61 ± 0.03 10 <sup>-5</sup>	1.35 ± 0.07 10 <sup>-5</sup>	3.31 ± 0.58 10 <sup>-8</sup>	3.19 ± 0.20 10 <sup>-7</sup>
<b>AMS</b>	n=8	n=8		
Anisotropy degree	1.029 ± 0.009	1.040 ± 0.008		
<b>Magnetization</b>	n=14	n=13	n=19	n=7*
NRM (Am <sup>2</sup> kg <sup>-1</sup> )	1.70 ± 0.22 10 <sup>-3</sup>	1.71 ± 0.12 10 <sup>-4</sup>	4.19 ± 0.44 10 <sup>-5</sup>	6.60 ± 2.84 10 <sup>-5</sup>

For each magnetic property, the number of measured samples (n) is indicated.

\* For susceptibility and magnetization of the schist, only samples with mass > 1g have been taken into account due to small-scale heterogeneities in pyrrhotite concentration.