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The high resolution Fourier-transform chemiluminescence

spectrum of the HS₂ radical

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Abstract

The chemiluminescence spectrum of the HS_2 radical has been recorded with a Fouriertransform spectrometer. The overview spectrum in the region 9000 cm⁻¹ to 4000 cm⁻¹ has been analysed and the vibrational parameters obtained are presented. In addition three of the bands $(0_0^0, 3_0^1 \text{ and } 3_1^0)$ which have been recorded at high resolution have been rotationally analysed and the results of the fit are presented. The results are discussed in the context of previous theoretical and experimental studies.

Keywords: Thiosulfeno radical; rotational analysis; chemiluminescence; Fourier-transform

spectroscopy;

Introduction

The HS₂ radical was first detected in 1950 during early flash-photolysis studies of hydrogen sulphide [1]. The new band systems were red-degraded, transient and diffuse between 315 and 380 nm. Porter's assignment was supported by later work where these new band systems were shown to occur much more strongly in the flash-photolysis of hydrogen disulphide [2]. A detailed theoretical study by Gosavi et al. [3] used a semi-empirical open-shell CNDO calculation to suggest these bands were the lowest-energy ${}^{2}A' \leftarrow {}^{2}A''$ electronic transition of HS₂. This interpretation was questioned by Sannigrahi et al. [4] who pointed out that the analogous transitions of the isovalent HO₂ and HSO radicals had been found at much lower energies experimentally as well as in *ab initio* SCF and CI calculations. These authors carried out detailed calculations of the energies, molecular structure, potential curves, and vibrational constants of HS₂ and DS₂ in their low-lying electronic states and of f-values for various transitions. They assigned the strong UV absorption system to a ${}^{2}A'' \leftarrow \tilde{X} {}^{2}A''$ transition and predicted the lowest energy ${}^{2}A'$ state of HS₂ to lie at 0.85 eV with very similar characteristics as in the isovalent radicals HO₂ and HSO.

Stimulated by these predictions and earlier experimental work on HO₂ [5] and HSO [6], Holstein et al. [7] studied the near-infrared chemiluminescence spectrum of HS₂ excited by energy transfer from metastable singlet oxygen in a fast-flow system. They observed a new band system in the range 1100-1700 nm for both, HS₂ and DS₂, and attributed the bands to the $\tilde{A}^2A' \rightarrow \tilde{X}^2A''$ transitions of the radicals. Initial work was at low resolution but this still enabled the investigators to make several important observations. Not only were the wavenumbers of the v₃ vibrations in both electronic states determined but also the v₂ vibrational wavenumber in the \tilde{X}^2A'' state was observed. In addition band contour analysis allowed estimates of the rotational constants to be made. It was this analysis which first

indicated that a significant degree of unexpected $\Delta K_a = 0$ character was present in the band contour.

The ground state of this radical has been well characterised spectroscopically. First, Yamamoto and co-workers recorded and analysed the spectrum of both HS₂ and DS₂ in the microwave region [8]. Later, HS₂ was discovered by Ashworth et al. during work on the far infra-red laser magnetic resonance (LMR) spectrum of the SH radical [9]. Subsequently further recordings were made and the FIR LMR spectrum analysed [10]. Further work on the ground state has been carried out by Winnewisser and co-workers [11] in the millimetre wave region.

The HS₂ radical continues to be of interest both theoretically and experimentally. On the one hand its analogues HO₂ and HSO are important in a variety of environments, notably combustion and the oxidation of reduced forms of sulfur [12, 13]. On the other it can be regarded as the simplest species with a S–S bond and as such the proton affinity [14] and the matrix isolated photolysis products of H₂S₂ [15] have been investigated. In addition there have been further attempts to predict the equilibrium structure and other properties of the radical *ab initio* [16, 17]. The most recent of these and only the second to calculate properties of the excited state is by Denis [18].

This paper deals with the analysis of both low and high resolution Fourier-transform chemiluminescence spect the HS₂ radical recorded at the University of Wuppertal. The vibrational parameters from the analysis of the low resolution spectra are first presented. The results of the analysis of three of the electronic bands $(0_0^0, 3_0^1 \text{ and } 3_1^0)$ are reported and compared with the earlier low-resolution results.

Experimental

The experiments were carried out in a fast-flow system using a 120 cm long Pyrex tube of 6 cm diameter with quartz windows at both ends as the observation volume. The tube was



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equipped with two glass inlet systems of smaller diameter and a pumping port. In one inlet system, HS₂(DS₂) radicals were generated by reacting H(D) atoms with sulfur vapor (S_x). A mixture of 5-10 Pa H₂(D₂) and 100-150 Pa helium was passed through a 100 W microwave discharge and over molten sulfur contained in a 0.5 L glass vessel which was heated to about 150° C. In the second inlet system, metastable oxygen molecules, O₂(a¹ Δ_g), were generated by passing 100-150 Pa O₂ with a trace of mercury vapor through a second microwave discharge. The film of HgO forming behind the discharge on the wall of the inlet system served to remove most of the oxygen atoms from the gas flow. The O₂(a¹ Δ_g) molecules served as energy carriers and were used to excite the HS₂(DS₂) radicals in the observation tube to their \tilde{A} ²A' state by near-resonant electronic-to-electronic (E-E) energy exchange. The system was pumped with a 500 m³h⁻¹ roots pump in series with a 30 m³h⁻¹ forepump resulting in flow velocities of about 20 ms⁻¹ in the large diameter tube. The total pressure, as measured with a Baratron capacitance manometer, was in the range 200-300 Pa and consisted of about equal partial pressures of O₂ and He and < 20 Pa of sulfur containing compounds.

The near-infrared radiation emitted along the axis of the glass tube was focused onto the entrance iris of a Bruker IFS 120 HR Fourier-transform spectrometer with a quartz lens. A mirror was placed at the far end of the observation tube and aligned so that the signal at the entrance aperture was maximised. Spectra were measured in the range 3 500 – 10 000 cm⁻¹ at nominal resolutions of 2 cm⁻¹ to 0.01 cm⁻¹ using liquid-nitrogen-cooled germanium (Applied Detector Corp. Model 403 S) or InSb (Cincinnati Electronics Corp. Model IDH 100) detectors. The latter was equipped with a cooled 2.8 µm short-pass filter. All gases and chemicals were research grade and were used without further purification. The wavenumber scale of the spectrometer was calibrated by use of reference lines of the O₂(a \rightarrow X) band [19] which showed up in all spectra. The absolute error of the given vacuum wavenumbers is less

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than ± 0.005 cm⁻¹, and the relative precision of wavenumbers of strong lines is on the order of 10^{-4} cm⁻¹.

Results and Analysis

Figures 1a and 1b show overview spectra of the $\tilde{A}^2 A' \rightarrow \tilde{X}^2 A''$ transitions of HS₂ and DS₂ recorded at low spectral resolution (2 cm^{-1}) in the sensitivity range of the Ge detector. The dominant features in both spectra are some short sequences of $00v'_3 - 00v''_3$ bands and the 0 -0 band of the $b^1\Sigma_g^+ \rightarrow X^3\Sigma_g^-$ transition of S₂ near 8 000 cm⁻¹. As in the low-resolution work of Holstein *et al.* [7] some weaker features which show up in the spectrum of HS₂ but not in that of DS₂ are tentatively assigned to bands involving transitions in the bending mode. For HS₂, the spectrum has also been measured in the range 4 000 – 9 000 cm⁻¹ with the less sensitive InSb detector. As is seen from Figure 2, two more sequences of the $\Delta v_3 = \text{const.}$ series are observed. A weak band at 4 560 cm⁻¹ which does not belong to this series is tentatively assigned to the $\tilde{A}^2A'(000) \rightarrow \tilde{X}^2A''(100)$ transition.

As is seen from the figures, the $00v'_3 - 00v''_3$ bands show characteristic sharp peaks at their short wavelength sides. By comparison with high-resolution spectra which have been rotationally analysed (see below) it is seen that the peaks consist mostly of the R lines and low-*N* Q and P lines of the $\Delta K_a = +1$ sub-bands and possibly of weak but very dense R lines of $\Delta K_a = 0$ sub-bands. Such $\Delta K_a = 0$ bands, which are "forbidden" in perpendicular transitions, have been identified in the analogous bands of HO₂ and were found to be mainly due to magnetic dipole transitions [20, 21]. In band contour analyses of the low-resolution bands, $\Delta K_a = 0$ sub-bands were also needed for HS₂ to fit the calculated and experimental band contours [7]. For the 001-000, 000-000, and 000-001 bands, which have been measured



Figure 1 about here

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 $K_a = 0 - 1$ subbands respectively.

at high resolution and rotationally analysed (see below), the wavenumbers of the peaks are respectively 8.2, 8.5, and 9.0 cm⁻¹ larger than those of the corresponding band origins. Wavenumbers of the sharp peaks have been measured for 18 bands of HS₂ and 7 bands of DS_2 , and were fitted to Equation (1) [22]____

$$v(v'_{3} - v''_{3}) = v_{0} + \omega_{3}^{\prime 0}v'_{3} + x_{33}^{\prime 0}v'_{3}^{\prime 2} - \omega_{3}^{\prime 0}v''_{3} - x_{33}^{\prime 0}v''_{3}^{\prime 2}$$

where ν_0 is the energy separation of the $\tilde{A}^2A'(000)$ and	$\tilde{X}^{2}A''(000)$ vibration	nal ground
states and ω_3^0 and x_{33}^0 are parameters the sum of which yields	elds the vibrational qu	tantum v_3 .
The wavenumbers of the bands and the observed-calculated of	lifferences are given i	n Table 1.
The results of the fit and other data deduced from the low-re	solution spectra are c	ollected in
Table 2.	Table 1 about here	Table 2 about here

The spectrum of HS₂ in the range 6400 - 8100 cm⁻¹ was measured at a nominal resolution of 0.01 cm⁻¹ and showed apodized line half-widths of 0.016 cm⁻¹. An overview spectrum of the 000-000 and 001-001 bands is shown in Figure 3. Neither individual rotational lines nor coarse K_a structure are obvious in this spectrum. An expanded portion of the 001-000 band is shown in Figure 4. The most obvious features in each band are two series of strong doublets, where the splitting of each pair of lines slowly decreases as one moves to lower wavenumbers. The spacing between doublets is, however, steadily increasing. At the start of the analysis these two features were assumed to be the Q-branch lines of the $K_a = 1 - 0$ and

> Figure 3 about here Figure 4 about here

The ground state of the HS_2 radical has already been well characterised [8, 10, 11] and the energy levels for the ground state can be calculated with confidence. The data in this work have been analysed using a program based on the asymmetric rotor program originally developed by Brown and Sears [23] and released into the public domain as ASYTOP [24]

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Having calculated the position of the levels with $K_a = 1$ in the lower state, common differences were calculated and each Q–branch line tested for the presence of associated P–and R–branch transitions.

The same procedure could not be used for the $K_a = 1-0$ transitions. The lack of *K*-doubling in the $K_a = 0$ level of the lower electronic state meant that the transitions observed to levels with $K_a = 1$ in the upper state would show residual *K*-doubling and the magnitude of this *K*doubling in the upper electronic state was unknown. Fortunately, the parameters determined from the $K_a = 0-1$ band were sufficient to allow $K_a = 1-0$ band to be assigned quite straightforwardly.

At this level of resolution second order effects tend make certain misassignments obvious. For example the off-diagonal element of the spin-tensor, $\frac{1}{2}|\varepsilon_{ab} + \varepsilon_{ba}|$, describes the interaction of levels differing in *N* and *K* by one unit. Thus when one spin component, N = n, in the K = k stack has approximately the same energy as the spin component with N = n-1, K = k+1 the two levels interact. The ensuing interaction is described by the off-diagonal element in the spin-tensor. This matrix element has only weak, if any, *K* dependence and thus tends to pin-point the crossing of the two *K* stacks. At the present resolution this is a very sensitive way to determine both the *A* rotational constant and the *N* numbering of the transitions. This particular perturbation can be seen in the residuals for one of the spin components. At J = 17 in the $v'_3 = 1 \rightarrow v''_3 = 0$ band the perturbation is a maximum with a positive residual. The residuals get progressively larger from lower *J* and suddenly become negative. This pattern was reproduced with residuals of the same magnitude in the corresponding P- and R- branches.

The fit was extended in an iterative procedure by predicting additional transitions and searching for appropriate matches. To a great extent the fits were limited to Q-branch

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transitions. This was because, within any one *K*-sub-band, the Q – branch transitions were significantly more intense than the P – and R – branch transitions. Transitions which lie at the high frequency end of the spectrum have also been, in some cases, difficult to assign. This proved particularly so in the $K_a = 3-2$ sub-band of the $1_0^0 2_0^0 3_0^1$ vibrational band. The most likely reason is that the energy levels are perturbed in some manner. This is not surprising as the analogous band system in the HO₂ radical has been found [20] to be highly perturbed. Given that the HS₂ radical has a higher density of states than HO₂ perturbations are more likely. The relatively low signal to noise ratio and high line density of the present spectra means that perturbations are neither easy to pick out nor to confirm by comparing other transitions with common energy levels.

The $1_0^0 2_0^0 3_0^1$ vibrational band was analysed first, simply because it was the strongest band in which the transitions terminated in the vibrationless levels of all three vibrational modes. A separate analysis was performed for the $1_0^0 2_0^0 3_0^0$ vibrational band. In both cases the microwave and millimetre wave data from references 8 and 11 were included with weights appropriate to the uncertainties of the measurements. The ground state parameters were allowed to float in the analysis of the $1_0^0 2_0^0 3_0^0$ vibrational band but not in that of the $1_0^0 2_0^0 3_0^1$ vibrational band. This allowed the ground state parameters to reflect the much larger values of *N* which were assigned in the present work. A selection of the strongest Q–branch lines from the $\Delta K_a = +1$ and $\Delta K_a = -1$ sub-bands in each of the three bands analysed here are given in Table 3, along

with the residuals from the fit of each of the bands.

Once reliable parameters for the $1_0^0 2_0^0 3_0^1$ vibrational band had been determined an analysis of the $1_0^0 2_0^0 3_1^0$ band was undertaken. This proved to be relatively straightforward b molecular parameters determined in the $v_3 = 1$ vibrational level are far less precise than those in the $v_3 = 0$ level because the uncertainties of the Fourier transform data have higher Comment [SHA16]: Numbering changed

Comment [SHA17]: "shown" replaced by "given in Table 3"

Table 3 about here

uncertainty than the microwave and millimetre wave data used to determine the parameters in

the $v_3 = 0$ level.

The *K*-sub-band structure in all the bands which have been observed is predominantly perpendicular in nature ($\Delta K_a = \pm 1$). There is, however, evidence at low resolution [7] that shows that significant parallel transition ($\Delta K_a = 0$) intensity must also be invoked in order to model the band contour correctly. Parallel transitions have been sought in the high resolution spectrum but no consistent results have been obtained. The line density is sufficiently high that there are many coincidences between calculated and observed transitions but no branches with $\Delta K_a = 0$ have been identified that can be traced with any certainty. Given that all the necessary parameters for both the ground and excited states have been determined, the parallel transitions should be straightforward to predict. It seems, therefore, that at the present signal to noise ratio the parallel transitions are too weak to be picked out reliably, even though sufficient intensity is present that, when integrated in a low resolution scan or when a high resolution scan is convoluted, the overall band contour shows distinct parallel character.

The usual expressions for the energy levels in asymmetric top molecules refer to five centrifugal distortion parameters. This formulation has been shown to be over-determined in the case of a planar molecule [25, 26] in which case only four are required. In the course of the analysis the program was adapted to allow the fifth constant to be evaluated from the other four. Although this did not allow us to determine more parameters than were varied, it gave an important cross-check which refer useful when the higher-order terms were not well determined.

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Table 4 about here

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Discussion

The data in Table 2 show that there is a very good agreement between the present low resolution data and the results of the high resolution analysis. The implication is that the maximum of the band at low resolution is in a consistent position with respect to the band origin. These are both consistent with earlier low resolution data [7]. It is obvious from Figure 1a that the bands involving excitation of v_2 are much weaker and less well defined than the bands involving v_3 and the uncertainty is correspondingly large.

One of the bands, shown in Figure 2 is assigned as having one quantum of vibration of v_1 excited in the ground state. This is consistent with the calculated wavenumber of 2 607 cm⁻¹ from reference [18]. On the other hand an experimental wavenumber of 2 463 cm⁻¹ has been measured in a matrix [15] and the Franck-Condon factors reported in reference [4] show that the transitions involving v_3 would be about one hundred times more intense than those involving v_1 . This implies the variansitions exciting v_1 would not be observable in the current experiment. Although it could be argued that matrix effects have perturbed the frequency of the vibration there is no obvious reason for the far greater than expected intensity. It is, however, hard to envisage another assignment for the band assigned l_1^0 here.

 HS_2 has two heavy sulphur atoms which because of their mass lie very nearly on a principal inertial axis. This has the effect that although HS_2 is an asymmetric top it is very nearly at the limit of a prolate symmetric top. This can be shown by considering a measure such as Ray's asymmetry parameter κ , which is given by,

$$\kappa = \frac{\left(2B - A - C\right)}{\left(A - C\right)}.$$

 κ runs from -1 when the top is near prolate through zero in the very asymmetric case to +1 at the oblate limit. In this case for the lower state $\kappa = -0.99848$ and for the upper state $\kappa = -0.99868$ which are both almost indistinguishable from -1. Hence the deviation from **Comment [SHA19]:** Numbering changed.

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prolate symmetric top energy levels would be expected to be very small indeed. On the other hand the high density of states means that many unpredictable perturbations are possible. Rotational constants have been determined for both $v_3 = 0$ and $v_3 = 1$ in both upper and lower states. This has allowed the determination of the vibrational dependence of the rotational constants on excitation of v_3 . The derived constants are given in Table 5. The derived parameters are denoted B_{00e} to indicate that the vibrational dependence of these parameters with respect to the v_1 and v_2 vibrations has not been taken into account. Table 5 about here There have been many attempts to determine the structure of the HS_2 radical. Most of these have been through ab initio calculation [3, 16-18]. Saito et al. [8] have used the microwave spectra of HS₂ and DS₂ to try to determine both r_{e_1} and r_z structures. Of all previous structural calculations the rotational constants calculated in the most recent publication [18] are the closest to those determined here. In addition the structure of the radical in the \tilde{A}^2A' electronic state was calculated. The results for the excited state are qualitatively similar results to those calculated earlier by Sannigrahi et al. [4] but yield rotational constants which are closer to those observed. Obviously these parameters are not directly comparable as the ab initio calculation generates an equilibrium structure and in the present work only the vibrational dependence of the v_3 vibration has been taken into account.

The same study [18] also calculated the vibrational quanta in both $\tilde{A}^{2}A'$ and $\tilde{X}^{2}A''$ states. These are given in Table 5 in comparison with the vibrational quanta determined here. It is interesting to note that the v_3 vibrational wavenumber decreases by nearly 20% whereas the rotational constants B_{00e} and C_{00e} decrease by only 10% and A_{00e} decreases by less than 3% on electronic excitation. This is consistent with the prediction [18] that the electronic excitation results primarily in the contraction of the bond angle and the lengthening of the SS Comment [SHA23]: Numbering

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bond, the SH distance being essentially unchanged. It also accounts neatly for the fact that

the spectrum is primarily a progression in the v_3 vibration.

Acknowledgements

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Figure Captions

- <text> Fig. 1. Overview spectra of the $\tilde{A}^2 A' \rightarrow \tilde{X}^2 A''$ transitions of HS₂ (a) and DS₂ (b) in the sensitivity range of the Ge detector. The resolution is 2 cm^{-1} .
- Fig. 2. Section of the overview spectrum of HS_2 measured with the InSb detector at a resolution of 2 cm⁻¹.
- The asterisks mark strong atomic lines.
- Fig. 3. Spectrum of the 000-000 and 001-001 bands of the $\tilde{A}^2A' \rightarrow \tilde{X}^2A''$ system of HS₂ at a nominal
- resolution of 0.01 cm⁻¹. The apodized line width is 0.016 cm⁻¹.
- Fig. 4. Section of the high-resolution spectrum of the 001-000 band with rotational assignment of some Qlines of the $K_a = 1 - 0$ sub-band.

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Tables

Table 1 Wavenumbers of band maxima of the $\tilde{A}^2 A' \rightarrow \tilde{X}^2 A''$ transitions of HS₂ and DS₂.

Band	HS_2	DS_2
$00v'_3 - 00v''_3$	v / cm^{-1}	v / cm^{-1}
0 - 0	7262.5(-1) ^a	7270.0(3)
0 - 1	6666.8(-1)	6674.9(-4)
0 - 2	6076.5(-1)	6087.6(-0)
0 – 3	5491.8(2)	
0 - 4	4912.0(-0)	
1 - 0	7766.9(2)	7772.2(-2)
1-1	7171.2(2)	7178.4(4)
1 - 2	6580.7(0)	
1 – 3	5995.5(-2)	
1 – 4	5416.3(2)	
1 – 5	4842.0(0)	
2 - 0	8266.2(-0)	8270.6(-2)
2 - 2	7080.0(-2)	
2-5	5341.3(-1)	
2 - 6	4772.8(2)	
3 - 0	8761.0(-0)	8765.2(1)
3 – 7	4703.9(-1)	
4 - 1	8655.5(1)	

^a Numbers in parentheses are the observed-calculated differences in units of the last digit.

		HS		DS ₂		
		Experiment	Theory [7] ^a	Experiment	Theory [7]	
$\tilde{A}^{2}A' \rightarrow \tilde{X}^{2}A''$	000 - 000	7254(1) ^b 7254 035(5) ^c	6936	7262(2) 7264(15) ^d	6936	_
	000 - 001	7255(7) ^d 6658(1)		6667(2)		
	001 - 000	6657.755(5) ^c 7758(1)		7764(2)		
	001 - 010	7758.693(5) ^c 6858(10)				
	000 - 010	6352(10)				
	010 - 020 000 - 100	6268(10) 4566(10)				
$\tilde{X}^{2}A''$	ω_3^0	598.42(6)		597.7(9)		
	x_{33}^{0}	- 2.692(9)		- 3.33(48)		
	v ₃	595.73(7)	559	594.4(10)	555	
		$596.27996(44)^{\circ}_{-}$		591(10) ^a		Comment [SHA28]: Correction m to value
	v_2	892(10)	953	696(20) ^d	674	
		904(8) ^d				
	ν_1	903° 2688(10) 2463°	2748		1966	
ã 2 A I	0	506 46(9)		504.8(6)		
A A	ω_3	2 222(25)		2 10(21)		
	л ₃₃ V2	-2.555(25) 504.13(10)	514	-2.10(21) 502.7(8)	511	
	•3	504.533914(44) ^c _ 504(4) ^d		502(15) ^d		Comment [SHA29]: Correction m to value
	v_2	808(10)	838		593	
	ν_1		2775		1986	
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More theoretical da Numbers in parentl Band origins of the	ata are found in Ref heses are estimated low–resolution ban	error limits or standard deviat ds have been estimated by sul	tions of the paramete otracting 8.5 cm ⁻¹ fro	om the wavenumbers of t	he	Comment [SHA31]: Numbering
More theoretical da Numbers in parent Band origins of the characteristic peaks	ata are found in Ref heses are estimated low–resolution ban	error limits or standard deviat ds have been estimated by sul	tions of the paramete ptracting 8.5 cm ⁻¹ fro	om the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parentl Band origins of the characteristic peaks From high–resoluti Results of the low-	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [error limits or standard deviat ds have been estimated by sub	tions of the parameter btracting 8.5 cm^{-1} from the second secon	om the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parentl Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	error limits or standard deviat ds have been estimated by sub 7]. 7 in Ar matrices [15].	tions of the paramete otracting 8.5 cm ⁻¹ fro	om the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	 retror limits or standard deviat ds have been estimated by suf 7]. 7 in Ar matrices [15]. 	tions of the paramete btracting 8.5 cm ⁻¹ from	m the wavenumbers of t		Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	 error limits or standard deviat ds have been estimated by sul 7]. r in Ar matrices [15]. 	tions of the paramete otracting 8.5 cm ⁻¹ fro	m the wavenumbers of t		Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban on analyses. -resolution studies [red absorption study	 rin Ar matrices [15]. 	tions of the paramete otracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban on analyses. -resolution studies [red absorption study	 rerror limits or standard deviat ds have been estimated by suf 7]. 7 in Ar matrices [15]. 	tions of the paramete otracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban resolution studies [red absorption study	 error limits or standard deviat ds have been estimated by suf 7]. 7 in Ar matrices [15]. 	ions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high-resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	7]. in Ar matrices [15].	ions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban on analyses. -resolution studies [red absorption study	7]. in Ar matrices [15].	tions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban resolution studies [red absorption study	7]. 7]. 7]. 7].	ions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high-resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	7]. in Ar matrices [15].	tions of the paramete btracting 8.5 cm ⁻¹ from the parameters of	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban .resolution studies [red absorption study	7]. in Ar matrices [15].	ions of the paramete otracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	ine characteristic contracteristic	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban resolution studies [red absorption study	7]. in Ar matrices [15].	ions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban resolution studies [red absorption study	7]. 7]. 7]. 7]. 7]. 7]. 7]. 7].	ions of the paramete btracting 8.5 cm ⁻¹ fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.
More theoretical da Numbers in parent Band origins of the characteristic peaks From high–resoluti Results of the low- Results of an infrar	ata are found in Ref heses are estimated low–resolution ban ion analyses. -resolution studies [red absorption study	7]. 7]. 7]. 7]. 7]. 7]. 7]. 7].	ions of the paramete otracting 8.5 cm ^{−1} fro	m the wavenumbers of t	he	Comment [SHA31]: Numbering changed. Comment [SHA32]: Numbering changed.

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2								1	Table 3					
3		W	avenu	mbers c	of select	ed line	s of Q–	branche	s observed in the	K = 1 - 0	and $K = 0-1$ ba	nds of th	e	
4	che	emilum	inesce	nce spe	ctrum o	of HS ₂ .	The de	viations	from the calcula	ted valu	es are included a	is an indi	cation as	
5							1	the good	iness of the fit.			0		0
6									$v'_3 = 0 \rightarrow v''_3$	= 1	$v_3' = 0 \rightarrow v_3''$	=0	$v'_3 = 1 \rightarrow v''_3 =$	= 0
7	N'	K'_a	K_c'	J'	N''	K_a''	K_c''	J''	Obs ^a	o-c ^b	Obs	o–c	Obs	о-с
8	1	0	1	1.5	1	1	1	1.5	6648.69514*		7244.93959*		7749.45549*	
9	1	0	1	0.5	1	1	1	0.5	6647.28568 *		7243.58261 *		7748.09816*	
10	2	0	2	2.5	2	1	2	2.5	6648.48291 *		7244.72342 *		7749.23356*	
11	2	0	2	1.5	2	1	2	1.5	6647.36484 *		7243.65354 *		7748.16310*	
12	3	0	3	3.5	3	1	3	3.5	6648.26163	2.14	7244.49236	-0.51	7748.99432*	
13	3	0	3	2.5	3	1	3	2.5	6647.35163	1.46	7243.62648 *		7748.12714*	
1/	4	0	4	4.5	4	1	4	4.5	6648.01306	-1.15	7244.23485	-2.03	7748.72672*	
14	4	0	4	3.5	4	1	4	3.5	6647.25791	1.75	7243.51674 *		7748.00556*	
10	5	0	5	5.5	5	1	5	5.5	6647.74053	0.89	7243.94795 *		7748.42325*	
10	5	0	5	4.5	5	1	5	4.5	6647.09379*		7243.33559*		7747.80966*	
17	6	0	6	6.5	6	1	6	6.5	6647.43222	1.33	7243.62122 *		7748.07903*	
18	6	0	6	5.5	6	1	6	5.5	6646.87180	1.37	7243.09054 *		7747.54691*	
19	7	0	7	7.5	7	1	7	7.5	6647.08646	1.73	7243.25350 *		7747.69091*	
20	7	0	7	6.5	7	1	7	6.5	6646.59187	0.92	7242.78637 *		7747.22213*	
21	8	0	8	8.5	8	1	8	8.5	6646.69867	-0.34	7242.84271	0.00	7747.25720	0.45
22	8	0	8	7.5	8	1	8	7.5	6646.25847 *		7242.42621 *		7746.83841 *	
23	9	0	9	9.5	9	1	9	9.5	6646.27224 *		7242.38656	-0.82	7746,77798	2.83
24	9	0	9	8.5	9	1	9	8.5	6645 87319	-1.90	7242 01408	1 99	7746 39780 *	2.05
25	10	0	10	10.5	10	1	10	10.5	6645 80201	-1 30	7241 88786	1.30	7746 24668	1 50
26	10	0	10	9.5	10	1	10	9.5	6645 44209	-0.15	7241 54596	0.58	7745 90166 *	1.57
27	11	0	11	9.5 11 5	10	1	11	9.5	6645 20073	-0.15	7241.34390	0.38	7745.50100	1 16
28	11	0	11	10.5	11	1	11	10.5	6644 06054	-0.99	7241.33994	0.45	7745.00470	-1.10
29	11	0	11	10.5	11	1	12	10.5	6644.90034	-0.34	7241.02702 *		7745.33093	1.65
30	12	0	12	12.5	12	1	12	12.5	6644.73000	0.42	7240.74303	0.22	7743.03630	1.05
31	12	0	12	11.5	12	1	12	11.5	0044.43213	0.43	7240.43790	0.22	7744.74028**	1.40
22	13	0	13	13.5	13	1	13	13.5	0044.13/18	-0.61	7240.10308	0.44	7744.33931	1.49
3Z 22	13	0	13	12.5	13	1	13	12.5	6643.85463	-0.57	7239.83782*	0.67	7742.08818*	0.00
33 04	14	0	14	14.5	14	1	14	14.5	6643.49414	-0.63	7239.41546	-0.67	7743.62822	-0.09
34	14	0	14	13.5	14	1	14	13.5	6643.23199	0.26	7239.16827	0.48	//43.3/696*	0.20
35	15	0	15	15.5	15	1	15	15.5	6642.80772	0.36	7238.67989*		7742.84839	0.30
36	15	0	15	14.5	15	1	15	14.5	6642.56182	0.26	7238.44809	0.24	7742.61287*	
37	16	0	16	16.5	16	I	16	16.5	6642.07562	0.30	7237.89549	-0.21	7742.01727	0.28
38	16	0	16	15.5	16	1	16	15.5	6641.84381	-1.08	7237.67889	0.72	7/41./9610*	
39	17	0	17	17.5	17	1	17	17.5	6641.29978	1.29	7237.06214	-1.25	7741.13404	-0.76
40	17	0	17	16.5	17	1	17	16.5	6641.08144	-0.41	7236.85859	-0.29	7740.92677 *	
41	18	0	18	18.5	18	1	18	18.5	6640.47753	0.82	7236.18263	-0.18	7740.20333	1.94
42	18	0	18	17.5	18	1	18	17.5	6640.27072	-1.81	7235.98878	-1.30	7740.00497 *	
43	19	0	19	19.5	19	1	19	19.5	6639.60913	-0.71	7235.25101	-2.83	7739.21756	0.95
44	19	0	19	18.5	19	1	19	18.5	6639.41669	-0.34	7235.07157	-0.26	7739.03078*	
45	20	0	20	20.5	20	1	20	20.5	6638.69736	-0.40	7234.27612	-0.22	7738.18049	0.12
46	20	0	20	19.5	20	1	20	19.5	6638.51537 *		7234.10342	-0.76	7738.00424 *	
40 //7	21	0	21	21.5	21	1	21	21.5	6637.74098	0.63	7233.25114	0.92	7737.09271	0.17
47 70	21	0	21	20.5	21	1	21	20.5	6637.56759*		7233.08715 *		7736.92536*	
40	22	0	22	22.5	22	1	22	22.5	6636.73739	-0.15	7232.17547	0.09	7735.95485	1.83
49 50	22	0	22	21.5	22	1	22	21.5	6636.57544	1.73	7232.02363	2.87	7735.79526	1.10
50	23	0	23	23.5	23	1	23	23.5	6635.68965	0.45	7231.05127	-0.46	7734.76156	-0.17
51														
52														

1														
2	23	0	23	22.5	23	1	23	22.5	6635.53504	1.31	7230.90622	1.22	7734.61164	1.00
3	24	0	24	24.5	24	1	24	24.5	6634.59529	0.04	7229.87917 *		7733.51970	1.13
4	24	0	24	23.5	24	1	24	23.5	6634 44925	1.62	7229.74104	1.18	7733.37569	0.90
5	25	Ő	25	25.5	25	1	25	25.5	6633 45620	0.60	7228.65792	0.29	7732.22299	-0.48
6	25	0	25	24.5	25	1	25	24.5	6633 31596	0.55	7228 52658	1.25	7732.08397	-2.64
7	26	0	26	26.5	26	1	26	26.5	6632 26992	-0.26	7227 38700 *	1.20	7730 87768	1 34
8	26	0	26	25.5	26	1	26	25.5	6632 13703 *	0.20	7227.36106		7730 74683	0.79
9	20	0	20	27.5	20	1	20	27.5	6631 03875	-0.13	7226.06648	-0 74	7729 47776	0.77
10	27	0	27	26.5	27	1	27	26.5	6630 91368	1 22	7225 94870	0.74	7729 35315	0.09
11	28	0	28	28.5	28	1	28	28.5	6629 76165 *	1.22	7224 69883	0.75	7728 02705	1 39
12	20	0	28	20.5	20	1	28	20.5	6629 64260	0.94	7224.09003	0.04	7727 90924	1.59
12	20	0	20	20.5	20	1	20	20.5	6628 43682	-1.56	7224.38362	-0.35	7726 52301	1.00
14	29	0	29	29.5	29	1	29	29.5	6628 32452	-0.07	7223.27930	0.55	7726.41219	2 44
14	30	0	30	20.5	29 30	1	30	20.5	6627.06894	-0.07	7223.17343	0.69	7724 96970	2.44
10	30	0	30	20.5	30	1	30	20.5	6626.06057	-0.05	7221.81202	0.52	7724.90970	1.80
10	31	0	31	29.5	30	1	31	29.5	6625 65181	1.60	7221.71030		7723 35850	1.00
17	31	0	31	30.5	31	1	31	30.5	6625 55137	0.10	7220.29487		7723 25800	2.54
18	22	0	22	22.5	22		22	22.5	6624 10165	-0.10	7220.19879	0.01	7721 60687	0.25
19	32	0	32 22	52.5 21.5	32		22	32.3 21.5	6624.00585	0.10	7210.72710	-0.91	7721.09087	0.55
20	32 22	0	32 22	22.5	32	1	22	22.5	6622 68407	1.64	7218.03738	0.46	7710.08440	1 20
21	22	0	22	22.5	22	1	22	22.5	6622 50227	0.72	7217.11110	-0.40	7710 20242	0.02
22	24	0	24	52.5 24.5	33 34	1	24	52.5 24.5	6621 12026	0.75	7217.02021	1.07	7718 21010	-0.02
23	54 24	0	24	54.5 22.5	34 24	1	24	22.5	6621.04507	0.59	7215.44415	-1.27	7718 12480	2.23
24	54 25	0	25	25.5 25.5	25	1	54 25	25.5	6610 52711	2.55	7213.30321	0.72	7716.13460	0.11
25	25	0	25	24.5	25	1	25	24.5	6610 44752	-0.50	7213.72009	-0.72	7716 22102	-0.11
26	26	0	26	26.5	35	1	26	26.5	6617 87034	-0.11	7213.03432	0.20	7714 52015	2.26
27	26	0	26	25.5	30	1	26	25.5	6617 80527	-0.52	7211.90200	-0.64	7714.52915	2.70
28	27	0	27	33.3 27.5	27	1	27	33.3 27 5	0017.80327	0.14	7211.69400	0.02	7712.60259	5.08
29	27	0	27	265	27	1	27	37.5	6616 11615	0.25	7210.14739	1.40	7712.00238	0.07
30	20	0	20	20.5	20	1	20	20.5	6614 44105	1.01	7210.06595	1.40	7710 62574	2.17
31	20 20	0	20	20.5 27.5	20 20	1	20	20.5 27.5	6614 28118	-1.91	7208.22275	1.20	7710.02374	1.22
32	20	0	20	20.5	30 20	1	20	20.5	6612 65628	1.54	7206.26139	0.19	7710.30110	1.90
33	39 20	0	39 20	39.3 20 5	39	1	20	39.3 20 5	6612 50726	0.08	7200.30323	-0.18	7708.39283	-1.20
34	39 40	0	39 40	30.3 40.5	39 40	1	39 40	30.3 40.5	6610 82056	0.56	7200.51162	1.49	7706 50605	1.71
35	40	0	40	40.5	40	1	40	40.5 20.5	6610 76710	-0.01	7204.39901	2 16	7706 45 42 1	-5.75
36	40	0	40	39.3 41.5	40	1	40	39.3 41.5	6608 02770	0.24	7204.55201	5.10	7704.27558	-1.05
37	41	0	41	41.5	41	1	41	41.5	6608 88008	-0.71	7202.38223	0.27	7704.37338	1.45
38	41	0	41	40.5	41	1	41	40.5	6607.00010*	-0.89	7202.33073	-0.27	7702 18400	0.22
30	42	0	42	42.5	42	1	42	42.5	6606.06550	0.20	7200.31304 *		7702.18409	-0.55
40	42	0	42	41.5	42	1	42	41.5	6605 02258 *	-0.28	7200.27470	0.46	7702.13793	-1.05
40 //1	43	0	45	43.5	43	1	43	43.5	6604 00455 *		7198.19770	0.40	7099.94284	1.40
41	45	0	45	42.5	45	1	45	42.5	6602 00050	0.76	7196.10056	-1.00	7099.89085	-4.14
42	44	0	44	44.5	44	1	44	44.5	6602.07621	0.70	7190.02993	1.02	7697.04108	-5.46
43	44	1	44	45.5	44	0	44	45.5	6668 51602 *	0.57	7195.99800 *		7097.00003	-3.05
44	1	1	1	1.5	1	0	1	1.5	6666 08221 *		7204.79237*		7766 92166 *	
45	1	1	1	0.5	1	0	1	0.5	0000.08331*		7262.33901 *		7700.83100*	
46	2	1	2	2.3	2	0	2	2.3	6665 64700 *		7261.01604 *		7766 20040 *	
47	2	1	2	1.5	2	0	2	1.5	6666 60208 *		7261.91094 *		7767.22220 *	
48	3 2	1	3 2	3.3 2.5	3 2	0	3 2	3.3 2.5	6667 20422 *		1202.00300*		7768 00205 *	
49	э 1	1	3 1	2.3 1 5	Э 4	0	3 1	2.3 1 5	6666 44106 *		1203.4041/*	044	7767 15146	0.20
50	4 1	1	4 1	4.5	4 1	0	4 1	4.5	6666 02404 *		7262 12072 *	0.00	7767 70797 *	0.50
51	4	1	4	<i>J.J</i>	4	U	4	<i>s.s</i>	0000.93404*		1203.10012*		1101.10201*	
52														

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2	5	1	5	5.5	5	0	5	5.5	6666.21484*		7262.44726	2.05	7766.89370	-4.02
3	5	1	5	4.5	5	0	5	4.5	5 6666.63004 *		7262.86027 *		7767.36153	-2.23
4	6	1	6	6.5	6	0	6	6.5	5 6665.92857 *		7262.13868	-0.54	7766.57684	-0.51
5	6	1	6	5.5	6	0	6	5.5	5 6666.28582*		7262.49734	1.03	7766.98008	1.06
6	7	1	7	7.5	7	0	7	7.5	5 6665.58479*		7261.77244 *		7766.19018	-2.58
7	7	1	7	6.5	7	0	7	6.5	6665.89749*		7262.08732	2.35	7766.54462*	
8	8	1	8	8.5	8	0	8	8.5	5 6665.18724	2.10	7261.34651 *		7765.74056	-5.13
9	8	1	8	7.5	8	Ő	8	7.5	5 6665 46282 U	0.14	7261.62319	-0.64	7766.05480	-3.23
10	9	1	9	9.5	9	Ő	9	9.5	5 6664.73217	1.45	7260 86251 *	0101	7765.23493	-2.41
11	9	1	9	8.5	9	Ő	9	8.5	5 6664 97811 -	1.73	7261.11135	-0.06	7765.51816	0.54
12	10	1	10	10.5	10	0	10	10.5	5 6664.22216 -	0.11	7260.32120*	0.00	7764.66534	-3.16
13	10	1	10	9.5	10	Ő	10	9.5	5 6664 44873 (0.72	7260.54818	1.48	7764.92226	-0.07
14	11	1	11	11.5	11	0	11	11.5	5 6663 66023 -	0.09	7259 72312 *	1110	7764 03780	-1.98
15	11	1	11	10.5	11	Ő	11	10.5	5 6663 86769	1 15	7259 92905 *		7764 26576	-5.65
16	12	1	12	12.5	12	Ő	12	12.5	5 6663 04399 -	1 30	7259.06907	0.42	7763 35070	-0.89
10	12	1	12	11.5	12	0	12	11.5	5 6663 23674	1.78	7259.00007	-0.10	7763 56426	-0.13
17	12	1	13	13.5	12	0	13	13.5	5 6662 37691 -	0.55	7258 35862	0.10	7762 60471	0.13
18	13	1	13	12.5	13	0	13	12.5	5 6662.55565	7.68	7258 53403	0.51	7762.00471	3 3 2
19	14	1	14	14.5	14	0	14	14.5	5 6661 65738 I	2.00 1.20	7257 59163	-0.10	7761 79623	-1.87
20	14	1	14	13.5	14	0	14	13.5	5 6661 82134 (0.29	7257 75466 *	-0.10	7761 98132	-1.07
21	15	1	15	15.5	15	0	15	15.5	6660 88303	1 33	7256 76962	0.11	7760.03105	1.31
22	15	1	15	14.5	15	0	15	14.5	6661 03847	1.33	7256 02358	-0.11	7761 10205	-1.51
23	15	1	15	14.5	15	0	15	16.5	5 6660 05900	1.40 1.45	7255 80224 *	1.00	7760.00050	-0.08
24	16	1	16	15.5	16	0	16	15.5	6660 20508).4J	7255.03224		7760 17062	1.05
25	17	1	17	17.5	10	0	17	17.5	6659 18220	0.30	7254 05880	0.54	7750.03128	3.01
26	17	1	17	16.5	17	0	17	16.5	6650 31002	0.30 7 3 7	7255 00751	3 30	7759.03128	1.66
27	19	1	19	18.5	17	0	19	18.5	6658 25222	2.32	7253.09731	5.59	7759.18009	0.64
28	10	1	10	10.5	10	0	10	10.5	5 6658 28150 *	0.50	7253.97353		7758 12017 *	-0.04
29	10	1	10	10.5	10	0	10	10.5	6657 27215	D 10	7252 02780	0.52	7756 88760	2 78
30	19	1	19	19.5	19	0	19	19.5	5 6657 20442 *	5.19	7252.92780	-0.55	7750.88709	-2.78
31	20	1	20	20.5	20	0	20	20.5	5 6656 24052	00	7251 83036	0.08	7755 73426	0.30
32	20	1	20	10.5	20	0	20	10.5	5 6656 35640 *	0.09	7251.03050	0.08	7755 86222 *	-0.50
33	20	1	20	21.5	20	0	20	21.5	5 6655 15703	0.35	7250 67570	1 56	7754 51803	2.76
34	21	1	21	21.5	21	0	21	21.5	6655 26512	0.33 7 3 7	7250 78532	-1.50	7754.51805	-2.70
35	21	1	21	20.5	21	0	21	20.5	5 6654.02070 -1	0.57	7249 46907	-0.58	7753 24567	-3 58
36	22	1	22	21.5	22	0	22	21.5	5 6654.12590 -	1.68	7249.57344	-0.50	7753 36161	-4.12
37	22	1	22	21.5	22	0	22	21.5	5 6652 83458 I	1.00 1.1.1	7248 20694	0.35	7751 01813	1.02
38	23	1	23	23.5	23	0	23	23.5	5 6652.03600 (0.11	7248.20094	-0.55	7752 02834	-1.92
39	23	1	23	24.5	23	0	23	22.5	5 6651 59592 -(0.76	7246.30070	0.93	7750 53197	-1.31
40	24	1	24	24.5	24 24	0	24	24.5	5 6651.69446 -1	0.40	7246.09127	-1.93	7750 63611	-4.84
41	25	1	25	25.5	24	0	25	25.5	5 665030631 - 1	0.00	7245 51858	-0.36	7749.08560	-3.44
4 2	25	1	25	23.5	25	0	25	23.5	5 6650.40254 -(0.00 D.40	7245.61308	-0.50	7749 18846	-4 59
12	25	1	26	26.5	25	0	26	24.5	5 6648 96785	1 1 1	7243.01366	0.48	7747 58427	-3.16
43	20	1	26	20.5	26	0	26	20.5	5 6649.05912 -	1.11	7244.07304	-1.41	7747 68639	-1.85
44 15	20	1	20	25.5	20	0	20	25.5	5 6647.57490 (0.75	7244.10445	-1.41	7746 02706	0.58
40 46	27	1	27	21.5	27	0	27	27.5	5 6647.66530	1.05	7242.01373	0.00	7746 12371	-0.56
40 47	28	1	21	20.5 28 5	27	0	21 28	20.5	5 6646 13303 I	1.05	7241 07883 *		7744 41356	-2.00
4/ 10	28	1	20 28	20.5 27.5	20	0	20 28	20.5	5 6646 22357	1 27	7241 16721	-0.14	7744 50703	_0.10
40 40	20	1	20 20	20.5	20 20	0	20 20	20.5	5 664 <u>4</u> 63060	1.27 134	7239 40065	0.14	7742 74087	1.52
49 50	29 29	1	29 29	29.5 28.5	29 29	0	29 29	29.5 28.5	5 6644 72768 *	J.J+	7239 57690	-0.54	7742 83297 *	1.52
50	29 30	1	29 30	20.5 30 5	29 30	0	29 30	20.5	5 6643 00533 1	104	7237 8/601	-0.54	7741 00080	0.56
51	50	1	50	50.5	50	0	30	50.5	· · · · · · · · · · · · · · · · · · ·	0.74	1401104071	-1.23	111.00700	0.50

30	1	30	29.5	30	0	30	29.5	6643.18287	0.18	7237.93264	-1.22	7741.10172	0.51
31	1	31	31.5	31	0	31	31.5	6641.50202	0.11	7236.15145	-0.49	7739.22228 *	
31	1	31	30.5	31	0	31	30.5	6641.58772	0.32	7236.23572	-0.97	7739.30934	-3.57
32	1	32	32.5	32	0	32	32.5	6639.85673	-0.32	7234.40171	-0.24	7737.37673	-1.83
32	1	32	31.5	32	0	32	31.5	6639.94105	-0.86	7234.48544	-0.60	7737.46910	0.92
33	1	33	33.5	33	0	33	33.5	6638.16240	0.60	7232.59979	1.53	7735.47722	-0.98
33	1	33	32.5	33	0	33	32.5	6638.24693	0.63	7232.68186	-0.12	7735.56578	-1.35
34	1	34	34.5	34	0	34	34.5	6636.41598	-0.29	7230.74102 *		7733.51970	-1.61
34	1	34	33.5	34	0	34	33.5	6636.50162	0.93	7230.82394	-0.68	7733.60953	-0.30
35	1	35	35.5	35	0	35	35.5	6634.62056*		7228.83034	0.04	7731.50760	-0.40
35	1	35	34.5	35	0	35	34.5	6634.70615	0.99	7228.91344	-0.62	7731.59499	-1.43
36	1	36	36.5	36	0	36	36.5	6632.77480*		7226.86622 *		7729.43710	-1.31
36	1	36	35.5	36	0	36	35.5	6632.86020	0.37	7226.95128	0.89	7729.52665	-0.35
37	1	37	37.5	37	0	37	37.5	6630.87939	0.32	7224.84889 *		7727.31265 *	
37	1	37	36.5	37	0	37	36.5	6630.96457	-0.23	7224.93372 *		7727.40075	-0.94
38	1	38	38.5	38	0	38	38.5	6628.93204	-1.46	7222.77842 *		7725.13085 *	
38	1	38	37.5	38	0	38	37.5	6629.01881	-1.36	7222.86504	0.88	7725.21946	-1.14
39	1	39	39.5	39	0	39	39.5	6626.93821 *		7220.65491	-0.01	7722.89421	1.08
39	1	39	38.5	39	0	39	38.5	6627.02605 *		7220.74182 *		7722.98208	-1.80
40	1	40	40.5	40	0	40	40.5	6624.89331*		7218.47766	-0.84	7720.60012	0.48
40	1	40	39.5	40	0	40	39.5	6624.98330	0.74	7218.56679 *		7720.68955	-2.08
41	1	41	41.5	41	0	41	41.5	6622.80073	1.83	7216.24867	-0.61	7718.25141	0.91
41	1	41	40.5	41	0	41	40.5	6622.88979*		7216.33888	-0.32	7718.34513	1.13
42	1	42	42.5	42	0	42	42.5	6620.65503	-0.09	7213.96713	-0.23	7715.84651	0.65
42	1	42	41.5	42	0	42	41.5	6620.74788*		7214.05914 *		7715.94110*	
43	1	43	43.5	43	0	43	43.5	6618.46138	-0.69	7211.63088	-1.99	7713.38559	-0.27
43	1	43	42.5	43	0	43	42.5	6618.55694 *		7211.72726	0.52	7713.48611	3.01
44	1	44	44.5	44	0	44	44.5	6616.21987 *		7209.24592 *		7710.87063 *	
44	1	44	43.5	44	0	44	43.5	6616.31577	-1.29	7209.34209 *		7710.97012*	

a The observed wavenumber of the lines in $\mbox{ cm}^{-1}$

b The value of the (obs – calc) $\times 10^3$ in units of cm⁻¹. An asterisk indicates the line was given zero weight in the fit due to

overlap or other perturbation and the quoted value is the calculated wavenumber.

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Demonster	Ã	² A′	$ ilde{\mathbf{X}}^2 \mathbf{A}$	$\tilde{\mathbf{X}}$ ² A "				
Parameter	000	001	000	001				
$v_0 / { m cm}^{-1}$	7254.035347(19) ^a	7758.569314(26)	0	6657.75539(44)				
A B C	289615.1(74) 7198.30(30) 7011.40(27)	289500.8(37) 7152.4363(56) 6968.080(35)	296978.9619(74) 7996.36601(78) 7776.73845(80)	296726.8(21) 7946.856(82) 7727.770(71)				
$\Delta_N \times 10^3$	6.182(82)	5.909(17)	5.88717(17)	5.875(22)				
$\Delta_{\scriptscriptstyle NK}$	0.2244(97)		0.2338296(85)	0.2272(26)				
Δ_{κ}	17.90(27)	18.36(29)	24.3299(11)	23.861(60)				
$\delta_N \times 10^3$	0.20(10)	0.48(18)	0.155935(51)	0.172(27)				
$\delta_{\scriptscriptstyle K}$			0.13091(37)					
$\Phi_{_{NK}} \times 10^6$			0.2479(74)					
$\Phi_{KN} \times 10^3$			0.03595(28)					
$\Phi_{K} \times 10^{3}$			8.32 ^b					
\mathcal{E}_{aa}	62440(72)	65120(429)	-45926.778(13)	-47556(18)				
\mathcal{E}_{bb}	61.7(67)	67.7(12)	-424.287(13)	-420.2(18)				
\mathcal{E}_{cc}	-271.5(61)	-265.41(92)	10.096(13)	8.8(14)				
$\frac{1}{2}\left \mathcal{E}_{ab}+\mathcal{E}_{ba}\right $			235.416(74)	235.416 ^c				
Δ^S_K	-450(31)	-452(29)	6.7205(22)	7.12(71)				
$\Delta_{\scriptscriptstyle NK}^{\scriptscriptstyle S}+\Delta_{\scriptscriptstyle KN}^{\scriptscriptstyle S}$		-60(3)	-0.202128	0.15(13)				
$\Delta_{NK}^{S} \times 10^{3}$			-9.528(66)					
$\Delta_N^S \times 10^3$			0.1787(33)					
$\delta^{\scriptscriptstyle S}_{\scriptscriptstyle K}$			0.0976(67)					
$\delta_{\scriptscriptstyle N}^{\scriptscriptstyle S} imes 10^3$			6.18(23)					
Φ^s_{κ}			-1.94(20)					

^a Numbers in parentheses are the standard deviations of the parameters in units of the last quoted decimal place.

b Constrained at the value from [8].

c Constrained at the value from the ground state.

Table 4

Molecular constants for the 000 and 001 levels of the \tilde{A}^2A' and \tilde{X}^2A'' states of HS₂ (in MHz) except where indicated.

Table 5	
Vibrational spacings and vibrational dependence of the rotational constants.	All parameters in MHz except
where indicated.	

	$\tilde{A}^{2}A'$	$\tilde{\mathbf{X}}^{2}\mathbf{A''}$
$\Delta G(v_3 = 1 - 0) / \text{cm}^{-1}$	504.533914(44) ^a	596.27996(44)
$\Delta G(v_3 = 1 - 0) / \text{cm}^{-1}[18]$	501	592
$lpha_{_A}$	368.8(50)	252.1(21)
$\alpha_{\scriptscriptstyle B}$	45.85(60)	49.505(83)
α_c	43.98(61)	48.969(72)
A_{00e}	289672.(32)	297105.0(11)
B_{00e}	7221.23(60)	8021.112(42)
C _{00e}	7033.39(58)	7801.223(37)

a Uncertainty given in units of last quoted decimal place.

t oncertainty given in units of last quoted decimal place









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