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The high-resolution far-infrared spectrum of methane at the SOLEIL synchrotron[☆]

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Abstract

As a tetrahedral molecule, methane has no permanent dipole moment. Its spectrum, however, displays faint absorption lines in the THz region, due to centrifugal distortion effects. This is important for planetary applications since this region is used to measure methane concentration in some planetary atmospheres, in particular on Titan. Up to now, all measurements relied either on some old low resolution infrared absorption spectra, or on high resolution Stark measurements for low J values only. Even if these results have been reexamined recently [E. H. Wishnow, G. S. Orton, I. Ozier and H. P. Gush, *J. Quant. Spectrosc. Radiat. Transfer* **103**, 102–117 (2007)], it seemed highly desirable to obtain much more precise laboratory data.

The high-intensity synchrotron radiation, combined with a 151.75 ± 0.1 m optical path in a White cell and a Bruker IFS 125 HR FTIR spectrometer at the AILES beamline of SOLEIL, enabled us to record this very weak spectrum at high resolution for the first time. Spectra were obtained in the 50–500 cm^{-1} wavenumber range at 296 K and 9.91, 20, 50 and 100 mbar with a resolution of 0.00074, 0.00134, 0.0034 and 0.0067 cm^{-1} (FWHM of the sinc function),

[☆]This paper is dedicated to Dr. Larry Rothman on the occasion of his 70th birthday.

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respectively. The rotational clusters are fully resolved and the good signal-to-noise ratio has enabled precise measurements of transition intensities (92 cold band lines and 96 Dyad–Dyad hot band lines, with normal abundance intensities in the range 2.10^{-26} to 1.10^{-24} $\text{cm}^{-1}/(\text{molecule}\cdot\text{cm}^{-2})$), yielding an accurate determination of the dipole moment derivatives. Such results should allow a better determination of CH_4 concentration in planetary objects.

Keywords: Methane; Synchrotron radiation; Far infrared; Line intensities; Titan

1. Introduction

Besides being one of the major greenhouse gases on Earth [1, 2], methane (CH_4) is an important molecule for many planetary atmospheres, either in the Solar System (Mars [3], Jupiter [4, 5], Saturn [6] and its main satellite Titan [7, 8, 9, 10, 11], Uranus [12], Neptune [13] and its main satellite Triton [14], Pluto [14, 15]) or also farther away, since it has now been identified in the gas envelop of at least two giant extrasolar planets (exoplanets HD 189733b [16] and HD 209458b [17]). In some cases, the presence of methane raises important questions about its origin (Mars [18, 19], Titan [20], exoplanets, ...) or leads to very interesting chemical processes (Titan [21]). Spectroscopy being the best (if not the sole) mean for studying such distant atmospheres, it is thus essential to have an accurate model of methane's absorption spectrum over a wide spectral range, from microwave to visible regions.

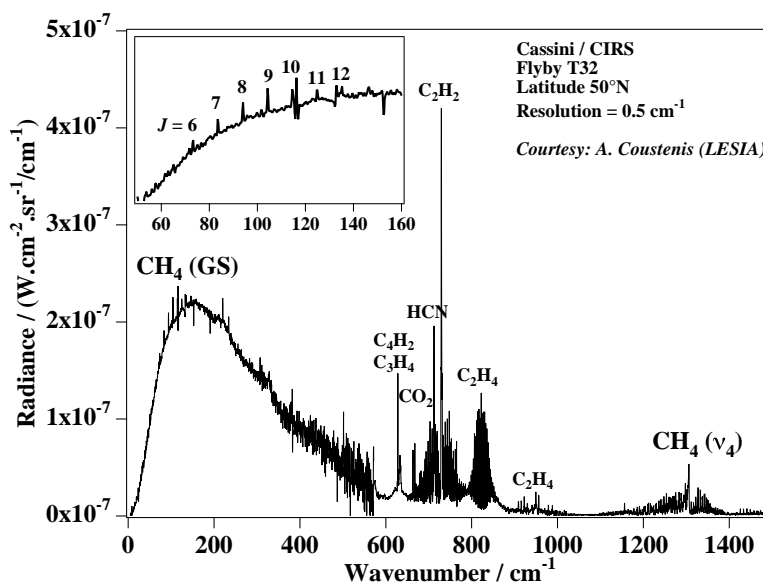


Figure 1: Example of a Cassini CIRS spectrum showing far infrared emission lines of methane.

Far-infrared (THz) emission lines are for instance used to determine molecular abundances in the upper atmosphere (stratosphere) of several objects. This is especially the case for methane lines as observed on Titan (and Saturn) by CIRS (Composite InfraRed Spectrometer) aboard the Cassini spacecraft [22] (see Figure 1). Line positions in this region of the methane spectrum are already well-known, but it is crucial to also have a very good model for line intensities for such abundance measurements.

The THz spectrum of methane (which is in fact the R branch of its pure rotational spectrum) is very weak, since it is induced by centrifugal distortion effects, as this highly-symmetrical molecule does not have a permanent dipole moment. It is thus difficult to study in the laboratory. Because of the importance of this spectrum for planetology, however, several attempts have been made previously to record it and to determine effective dipole moment parameters. A calculation of this induced dipole moment based on contact transformations was performed by Hilico *et al.* [23]. Previously, a low-resolution high-pressure spectrum was recorded at 113.5 K [24]. It has been reinvestigated recently [25], leading to a reevaluation of the intensities and an update included in the 2008 edition of the Hitran database [26]. Up to now, there existed only one partially resolved infrared absorption spectrum, recorded by McKellar in 1988 [27], but these results were never published. Finally, a high resolution and low pressure (100 mTorr) Stark measurement of the induced dipole moment has been reported [28], yielding a quite precise value however relying only on a limited number of low J lines only.

It thus seemed highly desirable to obtain fully resolved high-resolution spectra in order to measure accurate intensities of individual spectral lines, instead of relying on low-resolution profiles of clusters of lines. This necessitates both a long-path spectroscopic cell and a very bright light source in the THz region, in order to maximize the signal-to-noise ratio. Consequently, the synchrotron radiation at the AILES beamline of the French synchrotron SOLEIL combined with a long path White cell is ideally suited for this purpose.

In this article, we firstly describe the experimental setup at SOLEIL and the recorded spectra. We then explain the method used to extract line intensities, the theoretical model used for the analysis and we finally discuss the results.

2. Experimental details

The absorption spectra of methane have been recorded in the 50–500 cm^{-1} spectral range on the far-infrared beamline AILES, located at the Synchrotron facility SOLEIL [29, 30]. The intense far infrared continuum collected by the AILES beamline was focused on the entrance aperture of a high resolution IFS 125 Bruker interferometer, equipped for this experiment with a 6 μm mylar-Silicon composite beamsplitter and a 4.2 K bolometer fitted with a cold 600 cm^{-1} cut-off low pass filter. The scanner optical velocity was set to 2.53 cm/s .

A multipass cell (in a White type arrangement) was connected under vacuum to the interferometer and the optics (2.51 m base length) was adjusted to obtain 151.75(10) m absorption pathlength. In order to limit the intense pure rotation

Table 1: Experimental conditions. The uncertainty on temperature and pressure measurements is given in parentheses, in the unit of the last quoted digit.

Spectrum	Temperature / K	Pressure / mbar	Resolution / cm^{-1}	Number of coadded scans	Integration time / h
1	296(1)	9.91(1)	0.0011	346	31
2	296(1)	20.0(1)	0.002	540	27
3	296(1)	50.0(1)	0.005	356	7.1
4	296(1)	100.0(1)	0.01	360	3.7

absorption lines of water the entire cell was pumped with a secondary pump while heated at about 70 °C for about 48 hours and cooled down for another 48 hours. The cell was isolated from the interferometer by 2 cm diameter, 50 μm thick polypropylene windows. During the acquisition of the spectra, the interferometer and the entire beamline were continuously evacuated by means of three magnetic-bearings turbopumps connected to the interferometer and the beamline, in order to limit the water line absorption outside the cell to negligible levels.

The experimental conditions are detailed in Table 1. We recorded four spectra of pure methane at 23 (1) °C and at 9.91 mbar, 20 mbar, 50 mbar, and 100 mbar (the optical path was identical for the four spectra). The pressures, measured with 10 and 100 mbar full-scale range temperature-controlled capacitance manometers, were constant during the acquisition of the four spectra. The resolutions were set to 0.00074, 0.00134, 0.0034 and 0.0067 cm^{-1} (FWHM of the sinc function), respectively, using boxcar apodization.

Reference spectra were also recorded at different resolutions (0.00074, 0.00134, 0.0034 and 0.0067 cm^{-1}) in order to limit the baseline variations which may cause errors on the measured absolute intensities.

3. Intensity measurements

The individual line intensities were measured using a program developed in Brussels, which adjusts a synthetic spectrum to the observed spectrum of each line using a Levenberg-Marquardt least-squares fitting algorithm. The synthetic spectrum is calculated as the convolution of a monochromatic transmission spectrum with an instrument line shape function, which includes the effects of the finite maximum optical path difference and the finite entrance aperture in the interferometer [31]. The diameter d of the focal point of the synchrotron beam at the entrance of the spectrometer was measured to be equal to about 0.8 mm at 1000 cm^{-1} , decreasing proportionally to the square root of the wavenumber, thus following the relation $d = 0.8\sqrt{1000/\tilde{\nu}}$. The entrance aperture diameter was therefore set accordingly during the measurements. The profile of the lines in the monochromatic spectrum was modeled using a Voigt function [32] and the background was represented by a polynomial expansion, up to the third order. All the measurements performed involved the adjustment of the position, the intensity and the pressure self-broadening parameter of each line together with one to four background parameters. The Gaussian width of the Voigt profile was always held fixed to the value calculated for the Doppler broadening. The initial values of the line parameters required by the least squares fitting procedure were obtained from a peak finding procedure performed using the program “Wspectra” [33]. Altogether, we measured 167, 198, 109, and 133 absolute line intensities in the spectra recorded at sample pressures of 9.91, 20, 50 and 100 mbar, respectively. Fewer measurements could be made in the 50 hPa spectrum because of the somewhat worse signal-to-noise ratio. Intensities measured for the same line were averaged (each data being weighted by the inverse of the square of its estimated measurement uncertainty), rejecting data corresponding

to lines assigned to at least two transitions of similar intensities, having observed intensities differing by more than a factor of 2 from the predicted one, and blended with water vapour lines. This reduced data set amounted to a total of 193 line intensities. They are listed in Table A.1 given in appendix. The uncertainty associated to each of these line intensities was estimated as the square root of the sum of the square of the measurement uncertainty and the square of the deviation of each measured intensity from the corresponding weighted average. The differences, calculated line by line, between the intensities measured at 9.91, 50 and 100 hPa with those obtained at 20 hPa (all being involved in the 193 weighted averages) have mean values and RMS deviations of 0.11/1.6, 0.34/2.9 and 0.61/4.8 10^{-26} $\text{cm}^{-1}/(\text{molecule cm}^{-2})$, respectively. These indicate that the measurements do not exhibit any systematic differences in the individual measurements from the different pressures. Figure 2 shows an example of the measurements.

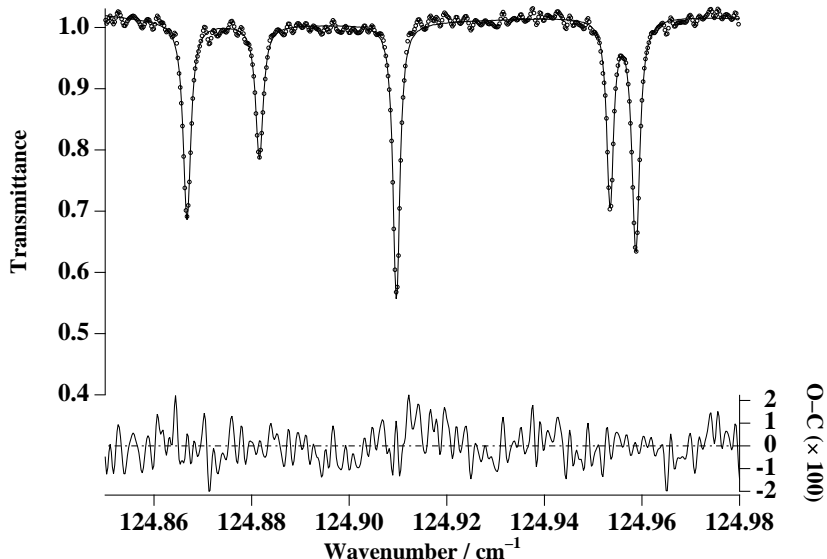


Figure 2: Part of the $R(11)$ region of the far infrared spectrum of $^{12}\text{CH}_4$ observed at 9.91 mbar (circles), and best fit calculated spectrum (solid line) and corresponding residuals (lower panel). Note that a rather wide range is shown here to illustrate the results of the fittings. The actual fits were carried on smaller regions, mainly because of the difficulties to reproduce correctly the residual channeling observed.

As indicated here above, the measurements involved the adjustment of the self-broadening parameter of each line. Indeed, it has never been measured for methane pure rotation lines and self-broadening parameters measured for one band do not necessarily describe accurately the values for another band [34]. So, to avoid biasing the measured line intensities correlated to the line widths, we preferred to adjust the self-broadening parameter of each line instead

of fixing it to an improper value. As the sample pressure was not high enough to allow precise determination of self-broadening parameters, we do not report these measurements. As a rough indication, the values obtained and their J -dependence are close to those observed for the strongest octad bands near 2.27 μm (see Fig. 5 of [34]).

4. Theory

The model we use here is based on the tensorial formalism developed in the Dijon group for methane and other spherical-top molecules [35, 36]. It is based on group theoretical and tensorial operator methods. We do not detail this formalism here, but we outline the elements used in the present study.

Two vibrational polyads of $^{12}\text{CH}_4$ are involved in this work: the Ground State (GS) and the ν_2/ν_4 bending Dyad. The spectra being recorded at room temperature, Dyad energy levels are significantly populated and thus we can observe both cold band lines (GS–GS rotational lines) and hot band lines (Dyad–Dyad rotational lines) in the THz region under investigation.

The effective Hamiltonian for the Ground State and for the Dyad are developed up to order 6 and the parameters are taken from the recent global fit of the first four methane polyads [37]. As we will see below, energy levels in the ground vibrational state of methane are already extremely well known (with a precision of *ca.* 50 kHz, [37]) and thus the line positions in the THz region are perfectly well reproduced by the model. As a consequence, we did not make any attempt to re-fit line positions using the present data, but we rather focused entirely on line intensities.

As a spherical top, CH_4 has no permanent dipole moment. But, when applying to the full dipole moment operator of the molecule the same contact transformation as the one used to isolate the different rovibrational polyads (see [35] for more details), some induced dipole moment terms appear even for pure rotational transitions. We consider here two effective dipole moment operators. The first one is the effective dipole moment for GS–GS transitions which expands as

$$\mu^{\langle\text{GS-GS}\rangle} = \mu_{\{\text{GS-GS}\}}^{\langle\text{GS-GS}\rangle} = \mu_0 R^{2(2,F_2)} + \dots, \quad (1)$$

up to order 1. This term is due to centrifugal distortion effects and depends on the $R^{2(2,F_2)}$ rotational operator (see [36] for definition). μ_0 is a parameter to be determined from the experimental data. It should be noticed that the μ_D parameter used in [25] is related to μ_0 by

$$\mu_D = \sqrt{8}\mu_0. \quad (2)$$

The second effective dipole moment operator we use is the one for Dyad–Dyad transitions,

$$\begin{aligned} \mu^{\langle\text{Dyad-Dyad}\rangle} &= \mu_{\{\text{GS-GS}\}}^{\langle\text{Dyad-Dyad}\rangle} + \mu_{\{\nu_2-\nu_4\}}^{\langle\text{Dyad-Dyad}\rangle} + \mu_{\{\nu_4-\nu_4\}}^{\langle\text{Dyad-Dyad}\rangle} \\ &= \mu_0 R^{2(2,F_2)} + \mu_{2,4} V_{2,4}^{EF_2(F_2)} + \mu_{4,4} V_{4,4}^{F_2F_2(F_2)} + \dots \end{aligned} \quad (3)$$

As explained in [35, 36], the effective dipole moment for hot bands is based on an extrapolation scheme: in the present case, for instance, the Dyad–Dyad effective dipole moment contains the $\mu_{\{\text{GS-GS}\}}^{\langle \text{Dyad-Dyad} \rangle}$ term that is formally identical to the GS–GS effective dipole moment, $\mu_{\{\text{GS-GS}\}}^{\langle \text{GS-GS} \rangle}$. This allows, in principle, to perform a simultaneous analysis of GS–GS and Dyad–Dyad intensities.

The last two terms in (3) are dipole moment derivatives specific to Dyad–Dyad transitions. $V_{2,4}^{EF_2(F_2)}$ and $V_{4,4}^{F_2F_2(F_2)}$ are vibrational operators relative to $\nu_2 - \nu_4$ and $\nu_4 - \nu_4$ transitions, respectively (see [35] for details about the tensorial construction of vibrational operators for tetrahedral molecules), while $\mu_{2,4}$ and $\mu_{4,4}$ are the corresponding parameters to be determined. Figure 3 summarizes schematically the lowest vibrational levels of methane and the effective dipole moment operators implied for rotational transitions. The $\mu_{4,4}$ parameter is related to the $(\mu_z)_{\text{eff}}$ parameter of Mills *et al.* [38] through the relation [39]

$$\mu_{4,4} = -\sqrt{6}(\mu_z)_{\text{eff}}. \quad (4)$$

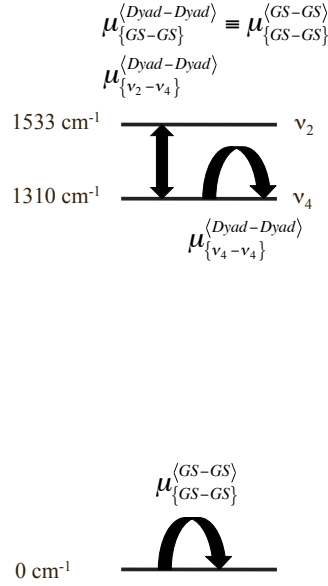


Figure 3: Schema of the lowest vibrational levels of CH₄ with the effective Hamiltonian operators involved in THz rotational transitions.

5. Analysis and results

We first performed a preliminary simulation using the existing parameters: effective Hamiltonian parameters from [37] and effective dipole moment param-

eters from [25]. As it is shown on Fig. 4, the result is already very convincing. It should be noticed that hot band lines have significant intensities. Water lines are also present but, fortunately, they generally do not overlap too much with methane lines. Figure 5 shows another comparison for different pressures.

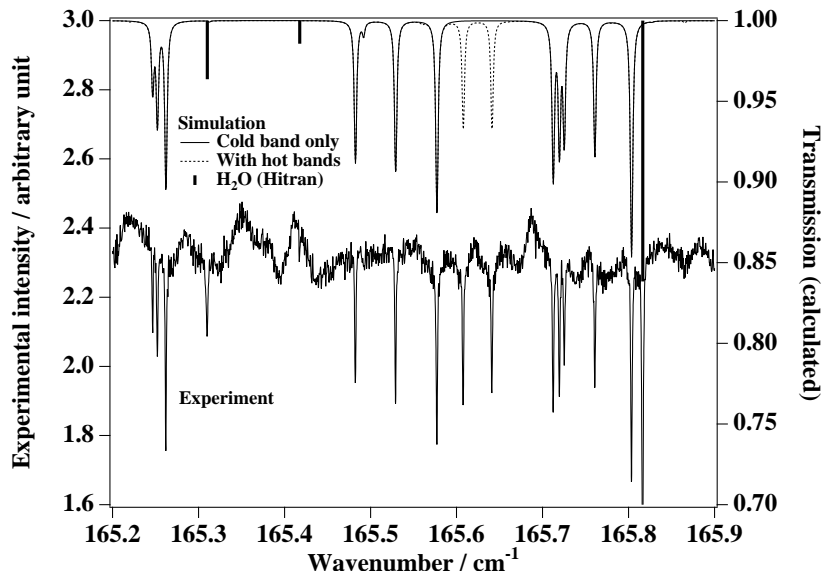


Figure 4: Example of comparison between the experimental spectrum and the simulation: the $R(15)$ region. Two simulations are presented (top): one with cold band lines only (continuous line) and one including hot band lines (dashed line). Three H_2O lines taken from the Hitran database are indicated by vertical sticks.

Line assignments were thus straightforward. They are also given in Table A.1. The lower state rotational quantum number range for assigned lines is $7 \leq J \leq 18$ for the cold band and $2 \leq J \leq 14$ for the hot bands. We could then perform three different fits of the effective dipole moment parameters, using the experimental intensities extracted from the spectra as explained in Section 3. As measurements were performed for natural abundance methane, we multiplied the calculated lines by the the 0.988274 isotopic factor [26]. All lines for which the relative intensity deviation exceeded 40 % were automatically rejected by the fitting program. Examination of the 5 line intensities rejected showed that they all correspond to quite weak (compared to the signal to noise ratio) and/or blended lines. The results are presented in Table 2 and discussed below.

Fit 1 includes only cold band line intensities (*i.e.* GS–GS transitions). Although the root mean square deviation d_{RMS} is somewhat high (*ca.* 9.3 %), we get a μ_0 (or μ_D) value with a good precision, significantly better than in the previous study [25]. The μ_D value appears to be quite close to the calculated value from Hilico *et al.* [23] and thus lower than that of

Table 2: Dipole moment parameters. For fit number 2 (global fit), the two numbers correspond to the cold and hot band lines, respectively. d_{RMS}^J is the root mean square deviation for total intensities of J clusters (see text). The two μ_D values indicated for Wishnow correspond to the so-called low- (0.24 cm^{-1}) and high- (0.06 cm^{-1}) resolution spectra from this reference.

Fit #	Nb. data	σ	$d_{\text{RMS}}/\%$	$d_{\text{RMS}}^J/\%$	$\mu_0/\mu\text{D}$	$\mu_D/\mu\text{D}$	$\mu_{2,4}/\text{mD}$	$\mu_{4,4}/\text{mD}$
1	92	1.468	8.52	3.53	7.976(29)	22.559(82)	–	–
2	92 (GS)	1.468	8.52	3.53	7.976(28)	22.559(79)	–	–
	96 (Dyad)	1.309	8.74	3.98	–	–	7.00(13)	–36.09(16)
3	96	1.306	8.74	3.98	7.9765 [†]	22.559 [†]	7.00(13)	–36.09(15)
Hilico [23]	(calc.)	–	–	–	7.90	22.34	10.8	–34.7
Ozier [28]	(Stark)	–	–	–	–	24.06(45)	–	–
Wishnow [25]	3 [‡]	–	–	–	–	23.82(88)	–	–
						23.9(1.2)		
Ozier [41]	2	–	–	–	–	–	–	–33.8(4.9)
Mourbat [40]	(calc.)	–	–	–	–	–	7.186	–37.47

[†] fixed value, [‡] unresolved clusters

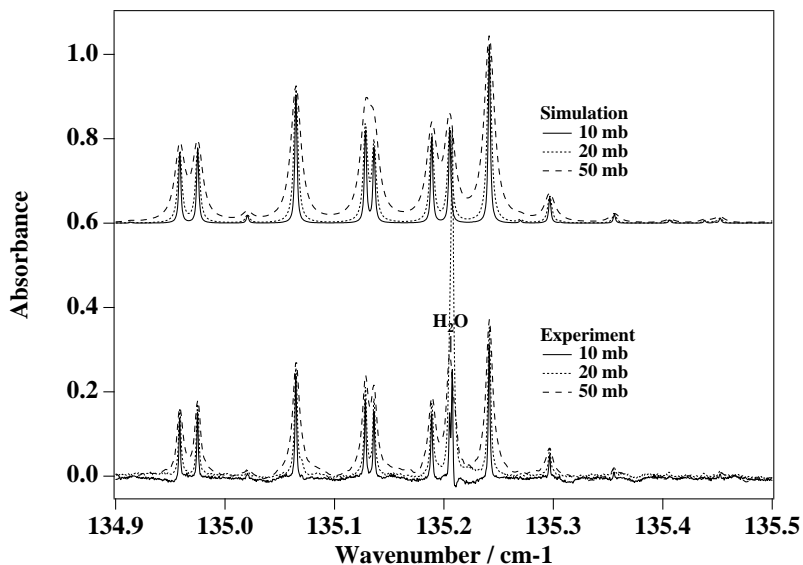


Figure 5: Examples of comparison between experimental spectra at different pressures and the corresponding simulations, in the $R(12)$ region. The most prominent line in the experimental spectra is due to H_2O , as indicated.

Wishnow *et al.* [25], obtained from low-resolution high-pressure spectra. The value from Stark measurements [28], also quite precise, is somewhat higher. This can reflect an underestimated experimental uncertainty on the Stark value. Alternatively, it might arise from the fact that the Stark value characterizes ΔJ matrix elements for $J=2$ only, while the present value characterizes $\Delta J = \pm 1$ matrix elements for J values in the 7 to 18 range.

Fit 2 includes both cold and hot band line intensities (*i.e.* GS–GS and Dyad–Dyad transitions) in order to determine simultaneously μ_0 , $\mu_{2,4}$ and $\mu_{4,4}$. Figure 6 displays the fit residuals in this case. They are also included in Table A.1. This second fit is fully consistent with the previous one, in agreement with the vibrational extrapolation for the effective dipole moment (see previous section). It should be noticed, however, that the GS–GS and Dyad–Dyad intensity datasets are virtually independent; a careful analysis of the fit results shows that μ_0 is determined by GS–GS data only, while the other two parameters are fully determined by Dyad–Dyad data.

Fit 3 is similar to Fit 2, but this time with μ_0 fixed to the value from Fit

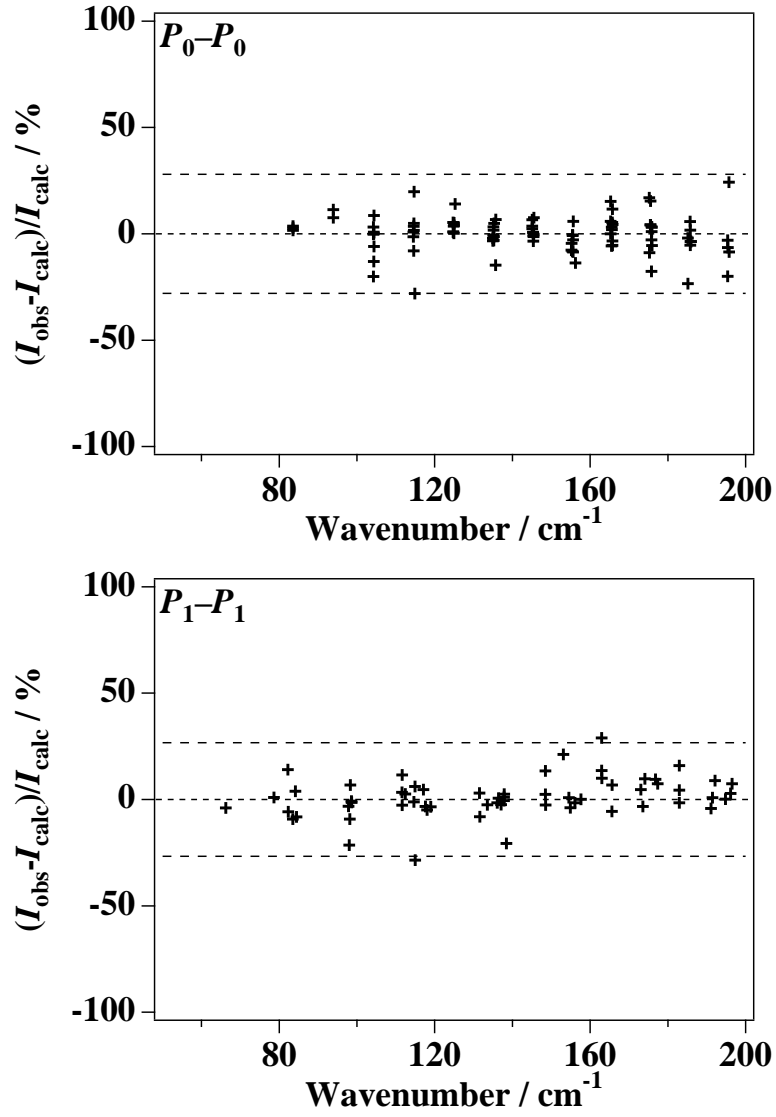


Figure 6: Fit residuals for line intensities for the cold (top) and hot (bottom) band lines in the case of the global fit. The upper and lower dashed lines display the $\pm 3d_{\text{RMS}}$ interval.

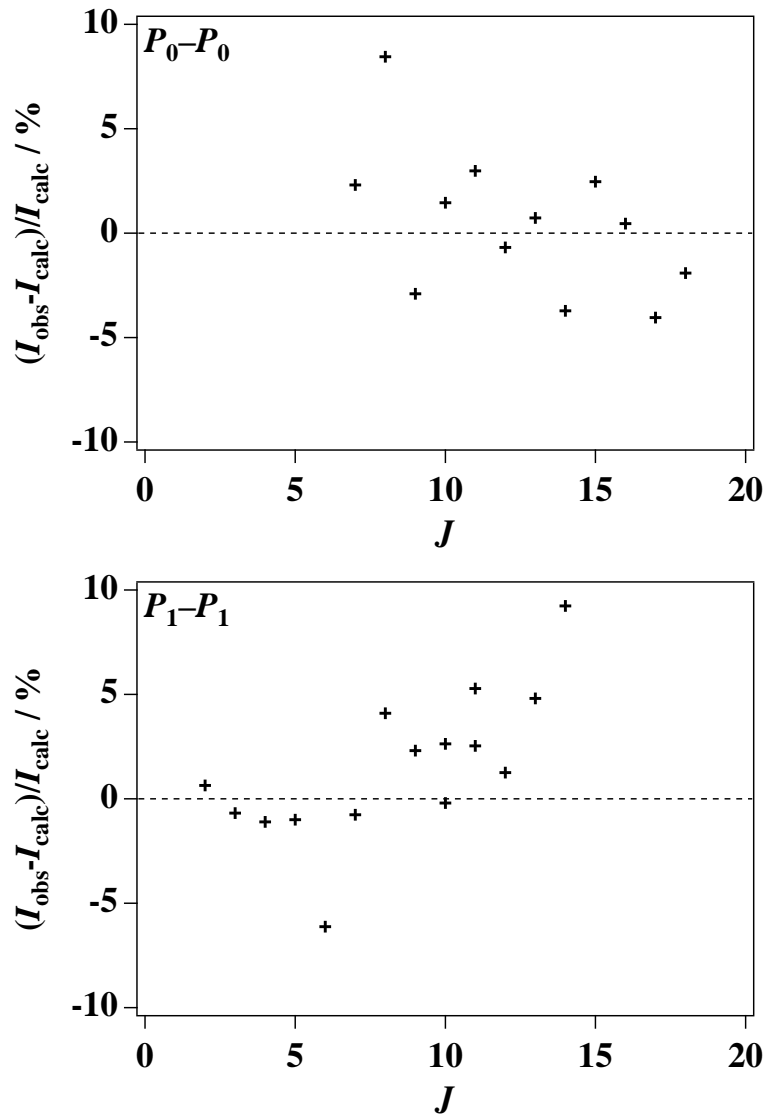


Figure 7: Fit residuals for total intensities of each $R(J)$ cluster in the case of the global fit.

1. This gives $\mu_{2,4}$ and $\mu_{4,4}$ identical to those from Fit 2 (with the same residuals as in Figure 6), confirming the full consistency of the results.

While the root mean square deviation calculated with all individual line intensities is close to 9 %, it appears that the total intensity of each $R(J)$ cluster of lines compares even better to the experiment. Figure 7 shows the fit residuals for intensity sums for each J value. The root mean square deviation calculated for these intensity sums, d_{RMS}^J , is lower than 4 % for both cold and hot band lines (see Table 2). Figure 8 shows the calculated rovibrational energy levels in the Ground State and the Dyad with the levels involved in the observed transitions marked in black. This gives an idea of how the present spectra sample the energy levels.

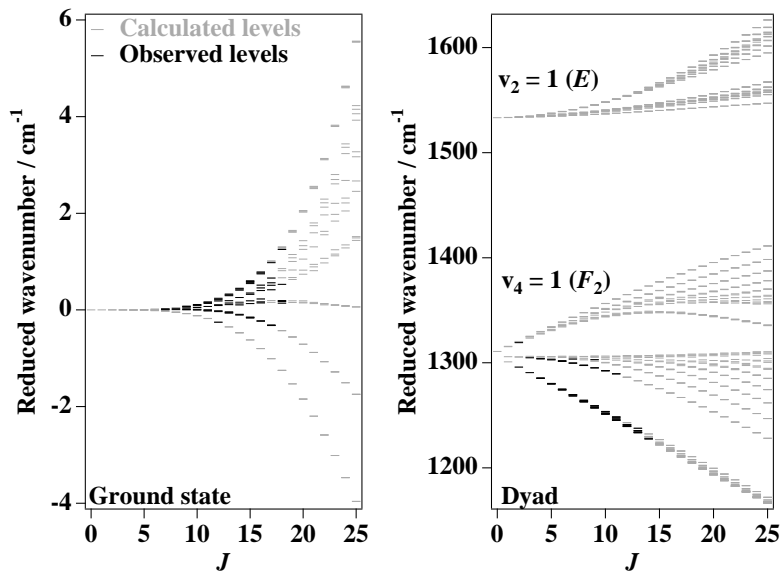


Figure 8: Reduced energy levels showing those reached by observed transitions (in black) corresponding to cold (left) and hot band lines (right).

The main outcome of this work is that we can recommend for μ_D the value

$$\mu_D = 22.559(79) \mu\text{D}, \quad (5)$$

and for the vibrational parameters for the hot bands the values

$$\mu_{2,4} = 7.00(13) \text{ mD}, \quad \mu_{4,4} = -36.09(16) \text{ mD}. \quad (6)$$

The numbers given between parentheses are the standard deviation of the fit, in the units of the last quoted digit. To estimate the effect of the uncertainty on the temperature on the determined value of μ_D , Fit 1 was also done assuming that the line intensities had been measured at 295 K (instead of 296 K). The

obtained value and corresponding uncertainty, $\mu_D = 22.624(82) \mu\text{D}$, show that the variation is within the given standard deviation.

The $\mu_{4,4}$ value we find here is quite consistent with previous theoretical [23, 40] and experimental values [41]. The signs are taken here identical as in the theoretical work of Ref. [23].

It should also be noticed that the hot band lines used in Fit 2 and Fit 3 contain only one $\nu_2 - \nu_4$ line. This means that only one upper level is assigned to a $\nu_2 = 1$ level. But it should be kept in mind that, due to the Coriolis coupling between the $\nu_2 = 1$ and $\nu_4 = 1$ levels, all rovibrational levels are mixed and so-called $\nu_4 = 1$ levels contain a contribution from $\nu_2 = 1$ levels. There is thus an intensity sharing between both hot bands and this allows a correct determination of $\mu_{2,4}$, although this one should probably be considered with some care.

Figure 9 summarizes the situation for a small portion of the spectrum, showing again that our value is close to the the value of Hilico *et al.* [23] and lower than that of Wishnow *et al.* [25]. The latter reference recommended to multiply the THz methane intensities available in the Hitran 2004 database [42], based on the calculated values from Hilico *et al.* [23], by a factor 1.154. This was done in the Hitran 2008 edition [26]. We recommend here the factor 1.019 obtained both from Fit 1 and Fit 2.

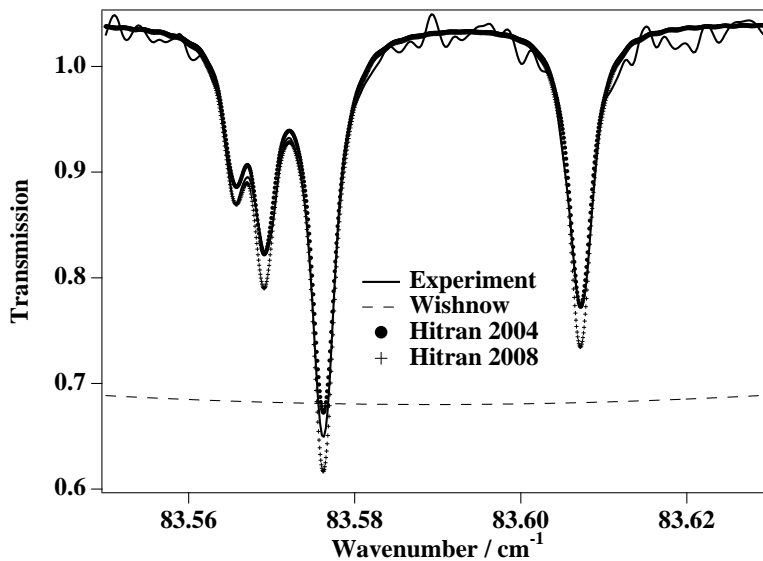


Figure 9: Comparison of the present measurements with the Wishnow *et al.* low resolution spectrum and with simulations using the Hitran 2004 and Hitran 2008 databases.

6. Conclusion

We presented here the first high-resolution spectra of methane ever recorded in the THz region with synchrotron radiation. These high quality spectra allowed to measure line intensities and to obtain precise effective dipole moment parameters for pure rotation transitions in the ground and $v_4 = 1$ states. The intensity information also obtained for the $\nu_2 - \nu_4$ band is probably rather approximate as it is mainly obtained through intensity sharing due to the Coriolis coupling between lines involving the $v_2 = 1$ and $v_4 = 1$ states. The measurements also led to the determination of self-broadening parameters, found to roughly agree with those observed for the strongest bands near $2.27 \mu\text{m}$. They are however not reported because the sample pressure was not high enough to allow their precise determination.

Despite the large improvements in signal-to-noise ratio provided by synchrotron radiation, the root mean square deviation obtained here remains large. Further improvements in the retrieved information could result from higher signal-to-noise ratios recently achieved at the AILES beamline [30] or the use of “multispectrum” line parameters retrieval techniques [43].

We hope that the present results will help improving the interpretation of planetary atmosphere spectra, providing more precise determinations of methane concentrations in such environments. Along this line, a further improvement of the spectroscopic information available for the THz spectrum of methane would involve the study of N_2 collisional broadening effects.

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Appendix A. Line list

Table A.1 : Intensities [in $\text{cm}^{-1}/(\text{molecule cm}^{-2})$ at 296 K] measured for pure rotational lines of $^{12}\text{CH}_4$, in natural abundance. This table is also given in ASCII form as supplementary material to this paper.

Table headings :

#	number of intensities averaged to yield the observed intensity given under heading “Intensity”;
Position	calculated line position, in cm^{-1} ;
Intensity	weighted average of “#” observed line intensities;
Unc.	uncertainty, estimated as the square root of the sum of the square of the measurement uncertainty and the square of the deviation of each measured intensity from “Intensity”;
U/%	$100 * \text{“Unc.”} / \text{“Intensity”}$;
O-C/%	“Intensity” minus the intensity calculated at the end of Fit 2 (see Section 5), expressed as a percentage of the calculated intensity;
E”	calculated lower state energy, in cm^{-1} ;
Bd (Band)	CB = transition within the ground state, H1 = transition within $\nu_4 = 1$, H2 = $\nu_2 - \nu_4$ band;
Br	branch;
J, C, n #vib	total angular momentum quantum number, symmetry (T_d group irreps), multiplicity index and main component for the projection on the initial basis, for the lower (”) and upper states.

Table A.1: Intensities [in $\text{cm}^{-1}/(\text{molecule cm}^{-2})$ at 296 K] measured for pure rotational lines of $^{12}\text{CH}_4$, in natural abundance.

#	Position	Intensity	Unc.	U/%	O-C/%	E"	Bd	Br	J"	C"	n"	#vib"	J	C	n	#vib		
1	59.08187	1.372E-25	9.5E-27	6.9	69.	1327.30839	H1	R	2	F2	1	2	100%	3	F1	3	2	99%
1	66.23860	7.277E-26	4.9E-27	6.7	-4.0	1367.73340	H1	R	3	F2	1	2	99%	4	F1	3	2	100%
1	77.49102	9.177E-26	1.5E-26	17.	65.	1596.72572	H1	R	7	A1	2	2	99%	8	A2	1	2	96%
3	78.65308	1.525E-25	4.4E-26	29.	1.0	1353.52728	H1	R	3	F1	1	2	100%	4	F2	3	2	99%
3	82.21987	2.368E-25	1.8E-26	7.8	14.	1408.48479	H1	R	4	A2	1	2	99%	5	A1	1	2	99%
2	82.31512	8.136E-26	2.1E-26	26.	-5.7	1409.00661	H1	R	4	F2	2	2	99%	5	F1	4	2	100%
1	83.52005	3.329E-26	2.5E-27	7.5	-9.2	1641.04784	H1	Q	8	F1	2	2	99%	8	F2	7	2	99%
2	83.56549	7.909E-26	1.4E-26	18.	2.7	293.16993	CB	R	7	E	1	1	100%	8	E	1	1	100%
2	83.56913	1.366E-25	4.7E-26	34.	3.6	293.16437	CB	R	7	F2	2	1	100%	8	F1	1	1	100%
2	83.57622	2.940E-25	9.8E-27	3.3	1.7	293.15400	CB	R	7	A2	1	1	100%	8	A1	1	1	100%
1	84.14751	5.149E-26	6.2E-27	12.	3.9	1437.32117	H1	R	5	F2	1	2	100%	6	F1	2	2	98%
2	84.47025	1.132E-25	2.0E-26	18.	-8.2	1437.49370	H1	R	5	A2	1	2	100%	6	A1	1	2	98%
3	93.91555	1.851E-25	1.3E-26	6.9	11.	376.80455	CB	R	8	F1	2	1	100%	9	F2	1	1	100%
3	93.93107	2.259E-25	3.6E-26	16.	7.6	376.78565	CB	R	8	F2	1	1	100%	9	F1	1	1	100%
1	93.95225	4.521E-26	5.1E-27	11.	62.	1773.78112	H1	R	9	A2	2	2	98%	10	A1	2	2	94%
4	97.81532	4.147E-25	7.9E-26	19.	-3.2	1389.95202	H1	R	4	A1	1	2	100%	5	A2	2	2	98%
1	98.02992	5.120E-26	5.7E-27	11.	-21.	1494.72267	H1	R	6	F1	1	2	99%	7	F2	2	2	97%
2	98.07789	4.374E-26	2.7E-26	62.	-43.	1494.94792	H1	R	6	F2	2	2	100%	7	F1	3	2	97%
2	98.17557	2.003E-25	3.3E-26	16.	-9.2	1390.13327	H1	R	4	F1	1	2	100%	5	F2	4	2	98%
2	98.30354	1.295E-25	1.2E-26	9.3	6.8	1460.50595	H1	R	5	F1	2	2	99%	6	F2	5	2	100%
1	98.58852	1.256E-25	3.0E-27	2.4	-0.8	1459.91150	H1	R	5	F2	2	2	98%	6	F1	5	2	99%
2	98.62756	1.827E-25	1.6E-26	9.0	-1.4	1390.28526	H1	R	4	F2	1	2	100%	5	F1	3	2	99%
4	104.22470	3.375E-25	4.2E-26	13.	3.2	470.83073	CB	R	9	A1	1	1	100%	10	A2	1	1	100%
2	104.24737	2.187E-25	8.6E-27	3.9	-0.2	470.80505	CB	R	9	F1	2	1	100%	10	F2	1	1	100%
2	104.25229	1.326E-25	4.3E-26	33.	-20.	470.79874	CB	R	9	E	1	1	100%	10	E	1	1	100%
2	104.31507	2.432E-25	1.4E-26	5.8	-13.	470.85477	CB	R	9	F1	3	1	100%	10	F2	2	1	100%
2	104.31924	2.579E-25	2.9E-26	11.	0.1	470.86482	CB	R	9	F2	2	1	100%	10	F1	1	1	100%
4	104.35000	5.639E-25	9.1E-26	16.	0.7	470.87268	CB	R	9	A2	1	1	100%	10	A1	1	1	100%
2	104.36479	5.884E-26	1.4E-26	23.	8.5	470.80505	CB	R	9	F1	2	1	100%	10	F2	2	1	100%
1	104.39473	5.951E-26	6.1E-27	10.	-6.0	470.86482	CB	R	9	F2	2	1	100%	10	F1	2	1	100%
2	111.60251	8.491E-26	1.7E-26	20.	3.4	1562.78797	H1	R	7	F1	2	2	99%	8	F2	3	2	96%
2	111.62766	6.644E-26	1.7E-26	25.	12.	1562.85866	H1	R	7	E	1	2	99%	8	E	2	2	96%
2	111.64262	1.224E-25	2.5E-26	21.	-2.7	1562.57412	H1	R	7	A1	1	2	99%	8	A2	1	2	96%
1	112.46121	4.093E-26	5.5E-27	13.	2.6	1938.19333	H1	Q	11	A1	1	2	99%	11	A2	3	2	99%
3	114.52344	2.267E-25	4.5E-26	20.	0.9	575.18407	CB	R	10	F1	1	1	100%	11	F2	1	1	100%
3	114.53532	2.576E-25	4.5E-26	18.	-1.4	575.16984	CB	R	10	F2	2	1	100%	11	F1	1	1	100%
2	114.61438	1.686E-25	1.2E-26	7.1	-8.0	575.27168	CB	R	10	E	2	1	100%	11	E	1	1	100%
2	114.61714	2.930E-25	1.8E-26	6.1	3.5	575.25955	CB	R	10	F1	2	1	100%	11	F2	2	1	100%
4	114.63941	7.343E-25	2.0E-25	28.	4.9	575.22269	CB	R	10	A1	1	1	100%	11	A2	1	1	100%
2	114.65061	2.374E-25	2.0E-26	8.6	-1.1	1521.96395	H1	R	6	A1	1	2	98%	7	A2	2	2	100%
4	114.67144	3.632E-25	2.9E-26	7.9	1.6	575.28518	CB	R	10	F2	3	1	100%	11	F1	2	1	100%
3	114.69262	6.598E-26	1.6E-26	24.	20.	575.18407	CB	R	10	F1	1	1	100%	11	F2	2	1	100%
1	114.87932	3.423E-26	5.6E-27	16.	-28.	575.16984	CB	R	10	F2	2	1	100%	11	F1	3	1	100%
3	114.92642	1.540E-25	2.1E-26	14.	6.1	1521.46869	H1	R	6	F1	2	2	98%	7	F2	5	2	99%
3	114.97639	7.258E-26	3.7E-26	51.	-29.	1521.28438	H1	R	6	E	2	2	98%	7	E	4	2	99%
4	117.12457	3.327E-25	2.3E-26	7.0	4.7	1436.86503	H1	R	5	F1	1	2	99%	6	F2	4	2	97%
4	117.61930	1.670E-25	2.1E-26	13.	-3.3	1437.20057	H1	R	5	E	1	2	100%	6	E	3	2	98%
4	118.07167	2.186E-25	2.5E-26	12.	-4.9	1437.32117	H1	R	5	F2	1	2	100%	6	F1	4	2	98%
4	118.98777	3.270E-25	3.7E-26	11.	-3.4	1437.49370	H1	R	5	A2	1	2	100%	6	A1	2	2	99%

Table A.1: Continued

#	Position	Intensity	Unc.	U/%	O-C/%	E"	Bd	Br	J"	C"	n"	#vib"	J	C	n	#vib		
3	124.76275	1.632E-25	1.3E-26	7.8	5.3	689.88606	CB	R	11	E	1	1	100%	12	E	1	1	100%
3	124.77117	2.414E-25	1.8E-26	7.6	0.8	689.87668	CB	R	11	F2	2	1	100%	12	F1	1	1	100%
4	124.78389	4.350E-25	1.6E-26	3.6	0.9	689.86209	CB	R	11	A2	1	1	100%	12	A1	1	1	100%
4	124.86687	2.760E-25	2.4E-26	8.7	3.6	690.01741	CB	R	11	F2	3	1	100%	12	F1	2	1	100%
4	124.90983	3.801E-25	2.1E-26	5.4	0.2	689.95663	CB	R	11	F1	2	1	100%	12	F2	1	1	100%
2	124.95360	2.366E-25	1.2E-26	5.2	3.9	690.03951	CB	R	11	E	2	1	100%	12	E	2	1	100%
2	124.95887	3.365E-25	9.9E-27	2.9	5.2	690.04917	CB	R	11	F1	3	1	100%	12	F2	2	1	100%
4	125.28147	1.230E-25	1.9E-26	16.	14.	689.86209	CB	R	11	A2	1	1	100%	12	A1	2	1	100%
4	131.54208	1.558E-25	1.2E-26	7.5	3.1	1593.02581	H1	R	7	F1	3	2	97%	8	F2	7	2	99%
4	131.62911	1.449E-25	1.0E-26	7.0	-8.1	1592.75259	H1	R	7	F2	2	2	97%	8	F1	6	2	99%
3	133.57281	8.351E-26	7.1E-27	8.5	-2.5	1640.20832	H1	R	8	A1	1	2	99%	9	A2	2	2	98%
3	134.95865	2.026E-25	1.0E-26	5.1	-3.3	814.88428	CB	R	12	F1	2	1	100%	13	F2	1	1	100%
3	134.97498	2.127E-25	3.6E-27	1.7	-2.1	814.86645	CB	R	12	F2	1	1	100%	13	F1	1	1	100%
2	135.06476	3.871E-25	8.0E-27	2.1	-0.5	815.08895	CB	R	12	A2	1	1	100%	13	A1	1	1	100%
3	135.12831	2.870E-25	5.2E-27	1.8	3.2	815.00803	CB	R	12	F2	2	1	100%	13	F1	2	1	100%
3	135.13616	2.268E-25	1.0E-26	4.5	1.7	814.99312	CB	R	12	E	2	1	100%	13	E	1	1	100%
3	135.18891	2.471E-25	1.7E-26	6.8	-3.1	815.11586	CB	R	12	F2	3	1	100%	13	F1	3	1	100%
3	135.24148	5.208E-25	2.4E-26	4.6	-1.3	815.14356	CB	R	12	A1	2	1	100%	13	A2	1	1	100%
2	135.29673	8.246E-26	1.0E-26	12.	4.9	815.00803	CB	R	12	F2	2	1	100%	13	F1	3	1	100%
1	135.65576	3.389E-26	1.7E-27	4.9	-15.	814.86645	CB	R	12	F2	1	1	100%	13	F1	4	1	100%
1	135.73905	3.190E-26	7.6E-28	2.4	6.7	814.64599	CB	R	12	A1	1	1	100%	13	A2	1	1	100%
4	136.07656	3.494E-25	1.7E-26	5.0	-1.5	1494.09862	H1	R	6	F2	1	2	99%	7	F1	5	2	97%
4	136.41416	2.250E-25	2.8E-26	13.	0.6	1494.19605	H1	R	6	E	1	2	99%	7	E	3	2	97%
3	137.08228	2.527E-25	7.4E-27	2.9	-2.4	1494.72267	H1	R	6	F1	1	2	99%	7	F2	4	2	97%
3	137.84405	1.364E-25	5.0E-27	3.6	-0.5	1729.88932	H1	R	9	A2	1	2	99%	10	A1	2	2	94%
3	137.85432	7.932E-26	2.1E-26	26.	1.0	1729.80843	H1	R	9	F2	2	2	99%	10	F1	3	2	94%
3	137.86736	5.267E-26	3.9E-27	7.5	2.4	1729.76129	H1	R	9	E	2	2	99%	10	E	2	2	94%
1	138.43967	1.766E-25	8.6E-27	4.9	-21.	1494.94792	H1	R	6	F2	2	2	100%	7	F1	6	2	98%
2	145.09922	2.920E-25	3.2E-27	1.1	3.6	950.15371	CB	R	13	A1	1	1	100%	14	A2	1	1	100%
3	145.11543	1.761E-25	1.6E-26	9.1	2.4	950.13634	CB	R	13	F1	2	1	100%	14	F2	1	1	100%
3	145.12193	1.244E-25	2.0E-26	16.	6.6	950.12928	CB	R	13	E	1	1	100%	14	E	1	1	100%
4	145.29512	2.128E-25	1.1E-26	5.0	0.7	950.33681	CB	R	13	F2	2	1	100%	14	F1	1	1	100%
4	145.31437	2.656E-25	1.8E-26	6.7	0.7	950.30477	CB	R	13	F1	3	1	100%	14	F2	2	1	100%
3	145.38186	1.930E-25	8.6E-27	4.4	-0.8	950.48723	CB	R	13	F2	3	1	100%	14	F1	2	1	100%
3	145.38969	1.388E-25	1.5E-26	11.	-3.4	950.50491	CB	R	13	E	2	1	100%	14	E	2	1	100%
4	145.44356	5.557E-25	1.5E-26	2.8	-1.3	950.38504	CB	R	13	A2	1	1	100%	14	A1	1	1	100%
4	145.45911	2.502E-25	2.4E-26	9.6	-0.2	950.52221	CB	R	13	F1	4	1	100%	14	F2	3	1	100%
4	145.53228	7.812E-26	8.9E-27	11.	7.6	950.33681	CB	R	13	F2	2	1	100%	14	F1	2	1	100%
4	148.46977	9.857E-26	2.7E-26	27.	2.4	1674.48632	H1	R	8	E	2	2	96%	9	E	5	2	99%
4	148.50492	1.645E-25	3.0E-26	18.	13.	1674.39048	H1	R	8	F2	3	2	96%	9	F1	7	2	99%
4	148.54381	2.412E-25	2.2E-26	9.2	-2.6	1674.21674	H1	R	8	A2	1	2	96%	9	A1	2	2	99%
1	153.10148	7.379E-26	2.7E-27	3.6	21.	1773.78112	H1	R	9	A2	2	2	98%	10	A1	3	2	97%
3	154.52536	6.270E-25	3.9E-26	6.2	0.8	1561.56706	H1	R	7	A2	1	2	99%	8	A1	2	2	95%
4	154.91492	3.362E-25	3.2E-26	9.4	-3.9	1561.71914	H1	R	7	F2	1	2	99%	8	F1	5	2	96%
3	155.20508	1.174E-25	1.5E-26	13.	-7.7	1095.63193	CB	R	14	F1	1	1	100%	15	F2	1	1	100%
3	155.21702	1.238E-25	4.5E-27	3.7	-4.5	1095.61913	CB	R	14	F2	2	1	100%	15	F1	1	1	100%
4	155.43277	1.694E-25	2.3E-26	14.	-8.2	1095.86909	CB	R	14	F1	2	1	100%	15	F2	2	1	100%
4	155.46268	3.485E-25	1.6E-26	4.5	-0.5	1095.82859	CB	R	14	A1	1	1	100%	15	A2	1	1	100%
4	155.51376	1.512E-25	1.1E-26	6.9	-0.8	1096.13287	CB	R	14	F1	3	1	100%	15	F2	3	1	100%

Table A.1: Continued

#	Position	Intensity	Unc.	U/%	O-C/%	E"	Bd	Br	J"	C"	n"	#vib"	J	C	n	#vib		
4	155.60915	2.216E-25	4.3E-26	19.	-8.5	1095.98132	CB	R	14	F2	3	1	100%	15	F1	2	1	100%
4	155.62286	1.143E-25	2.5E-26	22.	-2.9	1096.15684	CB	R	14	E	3	1	100%	15	E	2	1	100%
4	155.63624	1.896E-25	2.2E-26	11.	5.8	1096.17086	CB	R	14	F2	4	1	100%	15	F1	3	1	100%
4	156.07978	4.081E-25	5.2E-26	13.	-1.6	1562.57412	H1	R	7	A1	1	2	99%	8	A2	2	2	97%
1	156.23286	5.215E-26	3.2E-27	6.2	-14.	1095.82859	CB	R	14	A1	1	1	100%	15	A2	2	1	100%
4	157.64383	2.107E-25	1.9E-26	8.9	0.1	1562.78797	H1	R	7	F1	2	2	99%	8	F2	6	2	98%
2	162.96860	3.754E-26	9.1E-27	24.	14.	1938.25388	H1	R	11	E	2	2	99%	12	E	3	2	91%
2	162.97687	6.350E-26	1.0E-26	16.	29.	1938.23462	H1	R	11	F1	3	2	99%	12	F2	4	2	91%
3	162.99640	8.925E-26	2.5E-26	28.	10.0	1938.19333	H1	R	11	A1	1	2	99%	12	A2	1	2	91%
2	165.24669	6.758E-26	7.2E-27	11.	15.	1251.30765	CB	R	15	E	1	1	100%	16	E	1	1	100%
2	165.25216	9.349E-26	8.7E-27	9.3	5.9	1251.30186	CB	R	15	F2	2	1	100%	16	F1	1	1	100%
2	165.26211	1.488E-25	9.1E-27	6.1	-0.0	1251.29127	CB	R	15	A2	1	1	100%	16	A1	1	1	100%
3	165.48246	1.189E-25	1.9E-26	16.	-5.7	1251.64663	CB	R	15	F2	3	1	100%	16	F1	2	1	100%
3	165.52922	1.403E-25	2.3E-26	16.	3.1	1251.59047	CB	R	15	F1	2	1	100%	16	F2	1	1	100%
3	165.57721	1.793E-25	1.6E-26	9.1	1.9	1252.00188	CB	R	15	A1	1	1	100%	16	A2	1	1	100%
3	165.60785	1.169E-25	1.8E-26	16.	-5.7	1766.04623	H1	R	9	F1	3	2	95%	10	F2	8	2	99%
3	165.64127	1.331E-25	2.6E-26	20.	6.8	1765.92945	H1	R	9	F2	3	2	95%	10	F1	8	2	99%
2	165.71246	1.466E-25	2.8E-26	19.	5.4	1251.80710	CB	R	15	F1	3	1	100%	16	F2	2	1	100%
2	165.71962	1.055E-25	1.0E-26	9.8	-5.5	1251.77970	CB	R	15	E	2	1	100%	16	E	2	1	100%
2	165.72524	1.011E-25	1.4E-26	14.	-3.4	1252.02761	CB	R	15	F1	4	1	100%	16	F2	3	1	100%
3	165.76086	1.348E-25	1.2E-26	8.8	12.	1252.04625	CB	R	15	F2	4	1	100%	16	F1	3	1	100%
3	165.80353	2.266E-25	2.4E-26	10.	4.4	1252.06145	CB	R	15	A2	2	1	100%	16	A1	2	1	100%
4	173.04869	3.489E-25	3.6E-26	10.	4.6	1639.50861	H1	R	8	F2	1	2	99%	9	F1	5	2	94%
3	173.54862	1.986E-25	1.5E-26	7.5	-3.3	1639.74547	H1	R	8	E	1	2	99%	9	E	4	2	95%
4	174.06745	3.047E-25	3.7E-26	12.	9.7	1639.90401	H1	R	8	F1	1	2	99%	9	F2	6	2	95%
2	175.23026	6.572E-26	8.4E-27	13.	17.	1417.12909	CB	R	16	F1	2	1	100%	17	F2	1	1	100%
2	175.23920	5.148E-26	4.8E-27	9.2	-8.9	1417.11969	CB	R	16	F2	1	1	100%	17	F1	1	1	100%
3	175.47690	1.414E-25	1.6E-26	11.	4.2	1417.57909	CB	R	16	A2	1	1	100%	17	A1	1	1	100%
3	175.52873	8.728E-26	2.8E-26	32.	4.2	1417.51956	CB	R	16	F2	2	1	100%	17	F1	2	1	100%
2	175.54531	6.748E-26	2.8E-26	42.	15.	1417.49932	CB	R	16	E	2	1	100%	17	E	1	1	100%
2	175.75462	8.168E-26	1.7E-26	20.	-2.9	1417.80711	CB	R	16	F1	3	1	100%	17	F2	2	1	100%
1	175.77014	1.052E-25	3.9E-27	3.7	3.3	1417.75285	CB	R	16	F2	3	1	100%	17	F1	3	1	100%
2	175.77944	5.211E-26	7.4E-27	14.	-18.	1418.09887	CB	R	16	F1	4	1	100%	17	F2	3	1	100%
2	175.80961	4.971E-26	1.0E-26	21.	1.1	1418.11850	CB	R	16	E	3	1	100%	17	E	2	1	100%
2	175.88877	8.307E-26	8.5E-27	10.	3.1	1418.13715	CB	R	16	F2	4	1	100%	17	F1	4	1	100%
2	175.91944	1.914E-25	1.4E-26	7.6	-5.5	1417.86498	CB	R	16	A1	2	1	100%	17	A2	1	1	100%
4	176.79514	1.983E-25	2.5E-26	13.	9.4	1641.04784	H1	R	8	F1	2	2	99%	9	F2	7	2	97%
4	177.35729	1.931E-25	9.0E-27	4.7	7.4	1641.16444	H1	R	8	F2	2	2	99%	9	F1	6	2	97%
3	182.92118	1.595E-25	1.3E-26	8.0	-1.5	1867.73337	H1	R	10	A1	2	2	94%	11	A2	3	2	99%
2	182.93861	1.017E-25	7.3E-27	7.2	4.4	1867.66276	H1	R	10	F1	3	2	94%	11	F2	8	2	99%
2	182.94507	7.552E-26	4.6E-27	6.0	16.	1867.62865	H1	R	10	E	2	2	94%	11	E	6	2	99%
1	185.14892	5.447E-26	1.2E-27	2.2	-2.0	1593.05599	CB	R	17	A1	1	1	100%	18	A2	1	1	100%
1	185.15627	2.560E-26	1.2E-27	4.7	-23.	1593.04830	CB	R	17	F1	2	1	100%	18	F2	1	1	100%
1	185.15978	3.178E-26	1.9E-27	6.0	42.	1593.04462	CB	R	17	E	1	1	100%	18	E	1	1	100%
1	185.72296	3.828E-26	1.1E-27	3.0	5.7	1593.92810	CB	R	17	E	2	1	100%	18	E	2	1	100%
1	185.75272	5.439E-26	2.0E-27	3.8	-5.4	1593.87831	CB	R	17	F2	3	1	100%	18	F1	2	1	100%
1	185.77760	4.182E-26	2.3E-27	5.6	1.8	1594.34467	CB	R	17	F2	4	1	100%	18	F1	3	1	100%
3	185.81261	1.109E-25	6.3E-27	5.7	-3.7	1593.78442	CB	R	17	A2	1	1	100%	18	A1	1	1	100%
2	185.95843	6.835E-26	1.1E-26	16.	-3.6	1594.02591	CB	R	17	F1	4	1	100%	18	F2	3	1	100%

Table A.1: Continued

#	Position	Intensity	Unc.	U/%	O-C/%	E"	Bd	Br	J"	C"	n"	#vib"	J	C	n	#vib		
1	191.09868	2.669E-25	3.8E-27	1.4	-4.2	1727.74303	H1	R	9	F1	1	2	98%	10	F2	6	2	93%
4	191.49164	1.789E-25	1.3E-26	7.0	0.8	1727.84465	H1	R	9	E	1	2	98%	10	E	4	2	93%
4	192.11636	2.631E-25	1.8E-26	7.0	8.9	1728.19007	H1	R	9	F2	1	2	99%	10	F1	6	2	94%
4	194.81652	2.095E-25	1.1E-26	5.3	0.2	1728.59555	H1	R	9	F1	2	2	99%	10	F2	7	2	96%
1	195.36349	1.843E-26	9.4E-28	5.1	-3.0	1779.65106	CB	R	18	E	2	1	100%	19	E	1	1	100%
1	195.38122	2.301E-26	9.9E-28	4.3	-20.	1779.63103	CB	R	18	F1	2	1	100%	19	F2	2	1	100%
2	195.41070	4.622E-26	4.7E-27	10.	-6.5	1779.59703	CB	R	18	A1	1	1	100%	19	A2	1	1	100%
1	195.69065	4.795E-26	1.6E-27	3.3	24.	1780.70698	CB	R	18	A2	2	1	100%	19	A1	1	1	100%
2	195.74041	3.295E-26	0.0E+00	0.0	-8.5	1779.98434	CB	R	18	F2	3	1	100%	19	F1	2	1	100%
4	196.15405	9.896E-26	1.2E-26	12.	2.9	1729.76129	H1	R	9	E	2	2	99%	10	E	5	2	97%
4	196.49445	1.521E-25	1.5E-26	9.8	7.4	1729.80843	H1	R	9	F2	2	2	99%	10	F1	7	2	97%
2	200.41326	7.569E-26	6.6E-27	8.8	7.2	1979.42855	H1	R	11	F1	4	2	93%	12	F2	0	2	99%
2	200.41852	7.373E-26	5.1E-27	6.8	4.2	1979.38906	H1	R	11	F2	3	2	93%	12	F1	9	2	99%
4	208.57552	3.623E-25	9.2E-27	2.5	-1.5	1826.21678	H1	R	10	A1	1	2	98%	11	A2	2	2	92%
4	208.98398	2.236E-25	1.6E-26	7.2	5.6	1826.33783	H1	R	10	F1	1	2	98%	11	F2	6	2	92%
4	209.48148	2.040E-25	1.7E-26	8.4	3.7	1826.49407	H1	R	10	F2	1	2	98%	11	F1	7	2	92%
4	210.08444	3.147E-25	8.1E-27	2.6	2.0	1826.88440	H1	R	10	A2	1	2	98%	11	A1	3	2	93%
4	213.24790	1.529E-25	7.4E-27	4.8	-1.0	1827.29191	H1	R	10	F2	2	2	98%	11	F1	8	2	95%
3	213.67868	1.068E-25	1.1E-26	10.	3.1	1827.39616	H1	R	10	E	1	2	98%	11	E	5	2	95%
1	214.27081	5.640E-26	1.5E-27	2.7	0.7	1350.69110	H2	Q	2	A1	1	2	99%	2	A2	1	1	100%
4	215.51504	9.341E-26	6.5E-27	6.9	-10.	1828.82669	H1	R	10	F1	2	2	99%	11	F2	7	2	96%
4	216.06227	1.046E-25	1.1E-26	10.	3.0	1828.88901	H1	R	10	F2	3	2	99%	11	F1	9	2	97%
4	226.08668	1.638E-25	2.2E-26	13.	5.4	1935.10815	H1	R	11	F1	1	2	98%	12	F2	7	2	90%
4	226.56253	9.408E-26	1.7E-26	19.	-5.3	1935.27179	H1	R	11	E	1	2	98%	12	E	5	2	91%
3	227.06614	1.441E-25	6.3E-27	4.4	4.6	1935.41678	H1	R	11	F2	1	2	98%	12	F1	7	2	91%
3	231.61144	1.136E-25	1.4E-26	13.	6.8	1936.35196	H1	R	11	F2	2	2	98%	12	F1	8	2	94%
4	232.40061	1.176E-25	1.6E-26	13.	11.	1936.53105	H1	R	11	F1	2	2	98%	12	F2	8	2	94%
4	234.60288	1.240E-25	1.8E-26	15.	5.5	1938.19333	H1	R	11	A1	1	2	99%	12	A2	3	2	96%
2	235.05757	6.229E-26	1.4E-26	23.	-9.6	1938.23462	H1	R	11	F1	3	2	99%	12	F2	9	2	96%
2	235.23171	4.591E-26	1.8E-26	40.	0.1	1938.25388	H1	R	11	E	2	2	99%	12	E	6	2	96%
4	243.05250	1.100E-25	9.2E-27	8.4	5.9	2054.25035	H1	R	12	F2	1	2	97%	13	F1	7	2	89%
4	243.45937	7.103E-26	1.2E-26	17.	6.9	2054.33826	H1	R	12	E	1	2	97%	13	E	5	2	89%
4	243.99743	1.081E-25	1.1E-26	10.	13.	2054.55643	H1	R	12	F1	1	2	98%	13	F2	8	2	89%
3	248.73132	7.714E-26	1.1E-26	14.	-5.4	2055.08303	H1	R	12	F2	2	2	98%	13	F1	8	2	92%
1	250.14194	4.012E-26	1.9E-27	4.6	-13.	2055.82977	H1	R	12	E	2	2	98%	13	E	6	2	93%
2	250.73599	6.427E-26	1.4E-26	21.	-4.9	2055.91682	H1	R	12	F1	2	2	98%	13	F2	9	2	94%
4	251.65046	1.089E-25	2.2E-26	20.	-2.2	2056.05538	H1	R	12	A1	1	2	98%	13	A2	4	2	94%
4	259.42717	1.137E-25	9.7E-27	8.5	3.6	2183.60734	H1	R	13	A2	1	2	97%	14	A1	3	2	87%
2	259.82624	6.418E-26	8.8E-27	14.	0.5	2183.69686	H1	R	13	F2	1	2	97%	14	F1	8	2	87%
4	260.79244	1.053E-25	8.1E-27	7.7	5.6	2184.03721	H1	R	13	A1	1	2	97%	14	A2	3	2	88%
1	266.24801	5.998E-26	1.7E-27	2.8	22.	2184.55750	H1	R	13	F1	2	2	97%	14	F2	9	2	91%
2	268.76924	3.890E-26	1.6E-26	41.	-4.9	2185.60083	H1	R	13	F2	2	2	98%	14	F1	9	2	93%
2	269.90091	4.170E-26	9.7E-27	23.	4.7	2185.72233	H1	R	13	F1	3	2	98%	14	F2	0	2	93%
3	273.44688	4.447E-26	5.1E-27	12.	4.2	2187.85038	H1	R	13	A2	2	2	98%	14	A1	4	2	95%
1	275.81160	3.924E-26	1.2E-27	3.2	3.5	2323.29696	H1	R	14	F2	1	2	97%	15	F1	9	2	85%
1	276.24110	2.968E-26	1.9E-27	6.2	20.	2323.40587	H1	R	14	E	1	2	97%	15	E	6	2	86%

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