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Symmetry Breaking of Relativistic Multiconfiguration Methods in the Nonrelativistic Limit

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The *multiconfiguration Dirac-Fock* method allows to calculate the state of relativistic electrons in atoms or molecules. This method has been known for a long time to provide certain wrong predictions in the nonrelativistic limit. We study in full mathematical details the nonlinear model obtained in the nonrelativistic limit for *Be*-like atoms. We show that the method with *sp + pd* configurations in the $J = 1$ sector leads to a symmetry breaking phenomenon in the sense that the ground state is never an eigenvector of L^2 or S^2 . We thereby complement and clarify some previous studies.

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Introduction

Simulations of relativistic systems of Atomic and Molecular Physics and Chemistry are now of widespread use and the need for reliable methods is stronger than ever [24, 14]. The difficulties of doing relativistic simulations are however probably largely underestimated. One of the most common methods is the so-called *multiconfiguration Dirac-Fock* (or *Dirac-Hartree-Fock*) theory. This method has been known for a long time to provide certain wrong predictions in the nonrelativistic limit [31, 23, 13, 15, 11, 14]. The purpose of the present article is to explain in full mathematical details the encountered difficulties.

In atomic relativistic calculations, one usually imposes the total angular momentum $J^2 = (L+S)^2$ whereas in nonrelativistic calculations both L^2 and S^2 are imposed. It has been observed by Kim et al in [15] that a certain multiconfiguration ground state of the symmetry space $J = 1$ for *Be*-like (Beryllium-like) atoms, was converging in the nonrelativistic limit to a state which was not an eigenfunction of S^2 and L^2 . This led to erroneous values of certain transition probabilities like spin-forbidden ones.

In the multiconfiguration methods, the wavefunction is taken to be a linear combination of certain *configurations*. Both the linear coefficients and the orbitals in the configurations are variational parameters, leading to a highly nonlinear problem (with respect to variations of the orbitals). Each configuration is itself a sum of Slater determinants whose coefficients are fixed such that the configuration belongs to a chosen symmetry subspace ($J = 1$ in the case of the example studied in [15]).

It was noticed in [14] that the “error” is slowly disappearing when the number of determinants is increased. This suggests that the phenomenon is purely nonlinear, and that it has nothing to do with the nonrelativistic limit procedure in itself: it is the model obtained in the limit which does not fulfill the usual symmetry properties of nonrelativistic Quantum Chemistry or Physics models. In the limit, one obtains a nonlinear model for which all the configurations have $J = 1$ but they do not necessarily have a fixed orbital angular momentum L^2 or a fixed total spin S^2 . In those cases, the calculated ground state is not an eigenfunction of L^2 or S^2 .

In a linear model, any nondegenerate ground state automatically has the symmetry of the system but in a nonlinear model there could be a *symmetry breaking* phenomenon: although the system has a certain symmetry, the ground state does not necessarily possess this symmetry. One then obtains several minimizers and it is only the set of all ground states which is invariant under the symmetry group.

In this paper we study in detail the nonlinear model obtained in the nonrelativistic limit for *Be*-like atoms in the symmetry $J = 1$, when only *s*, *p* and *d* shells are considered¹, following [15]. Such an atom has four electrons which can only be in the following subshells: $1s_{1/2}$, $2s_{1/2}$, $2p_{1/2}$, $2p_{3/2}$, $3d_{3/2}$ and $3d_{5/2}$. Each subshell has a certain degeneracy but it

¹That is, we only consider the lowest “ungerade” states in the symmetry $J = 1$.

is only described by one radial orbital. The distinction between the $2p_{1/2}$ and the $2p_{3/2}$ subshells is an artefact of the nonrelativistic limit. In nonrelativistic theories, the same radial orbital is used for $2p_{1/2}$ and $2p_{3/2}$. The same holds for $3d_{3/2}$ and $3d_{5/2}$. The fact that the radial orbitals of $2p_{1/2}$ and $2p_{3/2}$ (or $3d_{3/2}$ and $3d_{5/2}$) are allowed to be different is, loosely speaking, similar to enlarging the variational set, which leads to symmetry breaking as will be explained below.

In MCDF theory one considers all the possible configurations of the symmetry $J = 1$ which can be constructed upon these states for four electrons, and one optimizes both the radial orbitals of the subshells and the variational coefficients in front of the configurations. Among all these possible wavefunctions, only few of them are eigenfunctions of both L^2 and S^2 . In the present case, there are three possible symmetries: 1P_1 , 3P_1 and 3D_1 (corresponding to specific eigenvalues of L^2 and S^2 , see Section 2 for a precise definition). In Theorem 2.1 we give necessary and sufficient conditions on the radial orbitals and the coefficients for a wavefunction to be in one of these three symmetries: the radial orbitals of $2p_{1/2}$ and $2p_{3/2}$ (and of $3d_{3/2}$ and $3d_{5/2}$) must be the same and certain relations must hold between the configuration coefficients.

The issue discussed in [15, 11, 14] is whether ‘the’ ground state² of the symmetry $J = 1$ is an eigenfunction of L^2 and S^2 . In other words, does it belong to one of the previous symmetries? To address this question, we look at the ground states obtained by *imposing* each of the symmetries 1P_1 , 3P_1 and 3D_1 , and ask ourselves whether these states can be stationary points and/or local minima of the full model where only $J = 1$ is imposed, or not. Our results (Theorems 2.2, 2.3 and 2.4) are summarized in the following table:

3D_1	is a stationary point, but is not a local minimum.
1P_1	is a stationary point, but is not a local minimum.
3P_1	has lower energy than 3D_1 and 1P_1 , but is not a stationary point.

This shows in particular that enlarging the variational set by only fixing $J = 1$ instead of both L^2 and S^2 leads to symmetry breaking for the wavefunction: the obtained new ground state is never an eigenvector of L^2 or S^2 . It is however essential that all s , p and d orbitals are included in the model (similar results are expected when more shells are added, by the same arguments as the ones presented in the paper). As we will see in Theorem 1.1, there is no symmetry breaking for the ground state if only s and p orbitals are considered.

As noticed first in [15], the question whether the constrained ground states of the symmetries 1P_1 , 3P_1 and 3D_1 are stationary points of the full model or not, is related to certain properties of the occupation numbers of the orbitals. This is explained in detail in Remarks 2.1 and 2.5 below.

The appearance of symmetry breaking is a well-known phenomenon in atomic multiconfiguration methods; it was encountered and rigorously examined in [2]. Similar issues occur in molecular calculations with regard to spacial symmetry, see, e.g., [5, 7, 22, 6]. In this paper we do not propose any practical solution to this phenomenon. Löwdin who emphasized this issue in a famous discussion [21] (after stability results [25, 4, 26, 30, 1] in Hartree-Fock theory) called it a “symmetry dilemma”. Our impression is that it is inherent to the way calculations are currently done. For the model studied in this paper, the only reasonable solution is probably to increase the number of determinants in order to decrease the effect

²We put ‘the’ in quotation marks to emphasize that there is no uniqueness.

of nonlinearities.

We have tried to make our work accessible to both the Mathematics, Quantum Physics and Chemistry communities. In particular, we will state and prove some well-known results (like for instance a special case of the first Hund's rule [16]) for the convenience of the mathematical reader. We hope that our work will help in improving the exchanges between the different communities. On the one hand mathematicians should find the material allowing her/him to understand the models and the typical problems encountered in nonlinear quantum computations. On the other hand we hope physicists and chemists will appreciate our rigorous clarification of the phenomenon they have to deal with.

Notation

We recall that the angular momentum reads $\mathbf{L} = x \times (-i\nabla)$, that $\mathbf{S}_k = \sigma_k/2$ where σ_k are the well known Pauli matrices, and that $\mathbf{J} = \mathbf{L} + \mathbf{S}$. For an N -body system, we still denote by $\mathbf{L} = \sum_{k=1}^N x_k \times (-i\nabla_{x_k})$ the (vector-valued) angular momentum operator of the N particles. A state will be denoted as ${}^{2S+1}L_J$ when it is an eigenfunction of S^2 with eigenvalue $S(S+1)$, of L^2 with eigenvalue $L(L+1)$ (with the identification P, D, F, \dots for $L = 1, 2, 3, \dots$), and of J^2 with eigenvalue $J(J+1)$. We will use the notation ${}^{2S+1}L$ when it is an eigenfunction of S^2 and L^2 (with the same eigenvalues as before), but not necessarily an eigenfunction of J^2 . For more details, we refer to [17, 29, 27].

1. Model with sp configurations only

We consider Be-like atoms, i.e. atoms with 4 electrons, in the symmetry $J = 1$. We start with a simplified multiconfiguration method employing only s and p shells. This means the 4-body wavefunction takes the form [15]

$$\Psi' = a\Phi'(1s_{1/2}^2 2s_{1/2} 2p_{1/2})(R_0, R_1, R_2) + b\Phi'(1s_{1/2}^2 2s_{1/2} 2p_{3/2})(R_0, R_1, R_3)$$

where a and b are configuration-mixing coefficients and R_0, R_1, R_2 and R_3 are the radial functions of, respectively, the shells $1s_{1/2}, 2s_{1/2}, 2p_{1/2}$ and $2p_{3/2}$. The two configurations $\Phi'(1s_{1/2}^2 2s_{1/2} 2p_{1/2})$ and $\Phi'(1s_{1/2}^2 2s_{1/2} 2p_{3/2})$ are some linear combination of Slater determinants made upon the corresponding 4-component shells, with fixed coefficients chosen such that $J = 1$. We will not write the detailed form of the configurations here, but later we will give the precise expression of their nonrelativistic limits (see (1.5) and (1.6)). The Hamiltonian to be used is the Dirac 4-body Coulomb operator which reads

$$H' := \sum_{i=1}^4 \left(D_i^0 - \frac{Z}{|x_i|} \right) + \sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|},$$

where $D^0 = c\boldsymbol{\alpha} \cdot p + mc^2\beta$ is the Dirac operator [28].

Most atomic MCDF packages aim at calculating a certain critical point of the energy $(R_0, R_1, R_2, R_3, a, b) \mapsto \langle \Psi, H\Psi \rangle$, under the following constraints:

$$\int_0^\infty |R_0(r)|^2 r^2 dr = \int_0^\infty |R_1(r)|^2 r^2 dr = \int_0^\infty |R_2(r)|^2 r^2 dr = \int_0^\infty |R_3(r)|^2 r^2 dr = 1, \quad (1.1)$$

$$\int_0^\infty \overline{R_0(r)} R_1(r) r^2 dr = 0, \quad (1.2)$$

$$a^2 + b^2 = 1. \quad (1.3)$$

As the energy is not bounded below due to the negative spectrum of the Dirac operator, it is *a priori* not at all obvious which critical point has to be considered and calculated numerically. Its existence in the infinite dimensional setting is also not clear at all. However, using the methods of [8, 9], one can prove that well-chosen critical points of this energy converge as $c \rightarrow \infty$ to critical points of a certain nonrelativistic model which we will now describe in detail.

The variational set of 4-body wavefunctions obtained in the nonrelativistic limit contains all the functions having the form

$$\Psi = a\Phi(1s_{1/2}^2 2s_{1/2} 2p_{1/2})(R_0, R_1, R_2) + b\Phi(1s_{1/2}^2 2s_{1/2} 2p_{3/2})(R_0, R_1, R_3) \quad (1.4)$$

but this time the two configurations are 2-component functions, i.e. they only depend on the spin variable. Their relation with the usual nonrelativistic configurations are given by the following formula (see, e.g., [3] page 294)

$$\Phi(1s_{1/2}^2 2s_{1/2} 2p_{1/2})(R_0, R_1, R_2) := -\frac{1}{\sqrt{3}} {}^1P_{sp}(R_0, R_1, R_2) + \frac{\sqrt{2}}{\sqrt{3}} {}^3P_{sp}(R_0, R_1, R_2), \quad (1.5)$$

$$\Phi(1s_{1/2}^2 2s_{1/2} 2p_{3/2})(R_0, R_1, R_3) := \frac{\sqrt{2}}{\sqrt{3}} {}^1P_{sp}(R_0, R_1, R_3) + \frac{1}{\sqrt{3}} {}^3P_{sp}(R_0, R_1, R_3). \quad (1.6)$$

Here ${}^1P_{sp}$ and ${}^3P_{sp}$ are some nonrelativistic configurations chosen such that

$$(S^2 - k(k+1)) {}^{2k+1}P_{sp} = 0, \quad (L^2 - 2) {}^{2k+1}P_{sp} = 0 \quad \text{and} \quad (J^2 - 2) {}^{2k+1}P_{sp} = 0$$

for $k = 0, 1$, and which are made only of s and p orbitals. We should probably rather use the notation ${}^{2k+1}P_{1,sp}$ but we refrain to do so, for the sake of simplicity. The form for the functions ${}^1P_{sp}$ and ${}^3P_{sp}$ (with $J_z = 1$) is provided in Appendix A for the convenience of the reader. For many of our arguments, we will not need this explicit form.

The associated minimization principle reads

$$E_{sp}(J = 1) := \inf_{\substack{R_0, R_1, R_2, R_3, a, b \\ \text{satisfying (1.1), (1.2) and (1.3)}}} \langle \Psi, H\Psi \rangle \quad (1.7)$$

where Ψ takes the form (1.4) and H is the nonrelativistic Hamiltonian

$$H := \sum_{i=1}^4 \left(\frac{-\Delta_i}{2} - \frac{Z}{|x_i|} \right) + \sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|}.$$

Note that the Hamiltonian H is real, hence each eigenfunction can be chosen to be real. For this reason, when passing to the nonlinear case, although in principle H should act on complex functions, we will restrict ourselves to real mixing coefficients a, b and real-valued radial orbitals R_k . The same is done in most Quantum Chemistry or Physics packages. The extension of our arguments to complex functions does not present any difficulty.

The Hamiltonian H commutes with both the total orbital angular momentum \mathbf{L} and the total spin \mathbf{S} . For this reason, in a usual *nonrelativistic* multiconfiguration method, one always restricts the search to ground states of a certain symmetry class. The only configurations made of s and p orbitals satisfying $J = 1$ are the ones corresponding to $L = 1$ and $S = 0$

(singlet) or $S = 1$ (triplet), and which appear in (1.5) and (1.6). Therefore, we will compare (1.7) with the usual nonrelativistic methods described by the following variational problems:

$$E_{sp}({}^1P_1) := \inf_{\substack{R_0, R_1, R_2 \\ \text{satisfying (1.1) and (1.2)}}} \langle {}^1P_{sp}(R_0, R_1, R_2), H {}^1P_{sp}(R_0, R_1, R_2) \rangle, \quad (1.8)$$

$$E_{sp}({}^3P_1) := \inf_{\substack{R_0, R_1, R_2 \\ \text{satisfying (1.1) and (1.2)}}} \langle {}^3P_{sp}(R_0, R_1, R_2), H {}^3P_{sp}(R_0, R_1, R_2) \rangle. \quad (1.9)$$

Note that there is no configuration-mixing coefficient to optimize in the above minimization problems. For this reason, (1.8) and (1.9) should indeed be called *Hartree-Fock* methods.

Our wavefunction (1.4) is always an eigenfunction corresponding to $L = 1$ but it is not necessarily an eigenfunction of S^2 . However, taking $R_2 = R_3$ and choosing $a = b/\sqrt{2}$ (resp. $a = -b\sqrt{2}$), we see that our variational set of functions of the form (1.4) indeed contains all possible functions ${}^3P_{sp}(R_0, R_1, R_2)$ (resp. ${}^1P_{sp}(R_0, R_1, R_2)$). Hence we deduce that

$$\boxed{E_{sp}(J = 1) \leq \min \{E_{sp}({}^3P_1), E_{sp}({}^1P_1)\}}. \quad (1.10)$$

The specific case of (the first) Hund's rule proven below in Theorem 1.2 tells us that indeed $E_{sp}({}^3P_1) < E_{sp}({}^1P_1)$, see Corollary 1.1. In principle, however, the inequality in (1.10) could be strict in which case the minimizer would not be an eigenfunction of S^2 , but instead a linear combination of $S = 0$ and $S = 1$ states.

We will see that this problem indeed does not occur for *sp* mixing, as expressed by the

Theorem 1.1 (The nonrelativistic limit for *sp* is correct). *We have*

$$E_{sp}(J = 1) = E_{sp}({}^3P_1) < E_{sp}({}^1P_1).$$

Additionally, any ground state Ψ for $E_{sp}(J = 1)$ satisfies $a = \varepsilon b\sqrt{2}$, $R_3 = \varepsilon R_2$ for some $\varepsilon = \pm 1$.

The rest of this section is devoted to the (simple) proof of the above theorem. As we will need it in the following, we start by proving a special case of the well-known (first) Hund's rule (for an excellent discussion of Hund's rules, we refer to [16], where a result similar to the following one is proved).

Theorem 1.2 (Hund's rule for singlet/triplet states). *For $K \geq 2$, $M \geq 1$, let $\{g_j\}_{j=1}^{2M} \cup \{f_i\}_{i=1}^K$ be an orthonormal system of $L^2(\mathbb{R}^3, \mathbb{C})$ and³*

$$\Psi_1 = \sum_{j=1}^M c_j \left(\bigwedge_{i=1}^K f_i^\uparrow \wedge f_i^\downarrow \right) \wedge g_{2j-1}^\uparrow \wedge g_{2j}^\downarrow, \quad \Psi_2 = \sum_{j=1}^M c_j \left(\bigwedge_{i=1}^K f_i^\uparrow \wedge f_i^\downarrow \right) \wedge g_{2j-1}^\downarrow \wedge g_{2j}^\uparrow$$

where $f^\tau(x, \sigma) = f(x)\delta_\tau(\sigma)$ and $(c_j) \in \mathbb{C}^M \setminus \{0\}$. Let

$$H := \sum_{i=1}^{2(K+1)} (h \otimes I_2)_i + \sum_{1 \leq i < j \leq 2(K+1)} V(x_i - x_j)$$

be a Hamiltonian where h is a self-adjoint operator on $L^2(\mathbb{R}^3, \mathbb{C})$ and V is a positive real function. Then, if Ψ_1 and Ψ_2 belong to the form domain of H , one has

$$\langle \Psi_1 + \Psi_2, H(\Psi_1 + \Psi_2) \rangle < \langle \Psi_1 - \Psi_2, H(\Psi_1 - \Psi_2) \rangle. \quad (1.11)$$

³Throughout the paper, we use the convention $(f_1 \wedge \cdots \wedge f_N)(x_1, \dots, x_N) = (N!)^{-1/2} \det(f_i(x_j))$.

Remark 1.1. *Let us emphasize that we do not impose any spacial symmetry on the functions f_i and g_j .*

Proof. We have

$$\langle \Psi_1 - \Psi_2, H(\Psi_1 - \Psi_2) \rangle - \langle \Psi_1 + \Psi_2, H(\Psi_1 + \Psi_2) \rangle = -4\Re\langle \Psi_1, H\Psi_2 \rangle = -4\Re\langle \Psi_1, \mathbb{V}\Psi_2 \rangle$$

where \mathbb{V} is the interaction (two-body) potential involving the function V . In the last equality we have used that each Slater determinant appearing in Ψ_1 always contains two functions orthogonal with all the functions in any of the Slater determinants of Ψ_2 , which implies that the one-body term vanishes. Calculating the two-body term one gets

$$\langle \Psi_1, \mathbb{V}\Psi_2 \rangle = - \iint_{\mathbb{R}^3 \times \mathbb{R}^3} V(x-y) \left| \sum_{j=1}^M c_j g_{2j-1}(x) g_{2j}(y) \right|^2 dx dy$$

and the result follows. \square

Corollary 1.1 (Hund's rule for sp mixing). *We have $E_{sp}({}^3P_1) < E_{sp}({}^1P_1)$.*

Proof. Using the methods of proof of [19, 20, 10, 18], one can see that there exists (R_0, R_1, R_2) minimizing $E_{sp}({}^1P_1)$. We now choose and fix these functions. Using formula (A.5) given in Appendix A and Property (A.8), we see that

$$\begin{aligned} & \langle {}^3P_{sp}(R_0, R_1, R_2), H({}^3P_{sp}(R_0, R_1, R_2)) \rangle \\ &= \frac{1}{2} \langle S^-({}^3P_{2,sp}(R_0, R_1, R_2)), HS^-({}^3P_{2,sp}(R_0, R_1, R_2)) \rangle \end{aligned} \quad (1.12)$$

where the state ${}^3P_{2,sp}$ is defined in (A.1). Indeed we have precisely

$$S^-({}^3P_{2,sp}(R_0, R_1, R_2)) = s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_1^\downarrow(R_2) + s^\downarrow(R_1) \wedge p_1^\uparrow(R_2) \right)$$

and, see (A.4),

$${}^1P_{sp}(R_0, R_1, R_2) = \frac{1}{\sqrt{2}} s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_1^\downarrow(R_2) - s^\downarrow(R_1) \wedge p_1^\uparrow(R_2) \right),$$

the notation being that of Appendix A. The result follows from (1.12) and Theorem 1.2. \square

We now give the proof of Theorem 1.1:

Proof. Using formulas (1.5), (1.6), (A.4) and (A.5), we deduce that any trial wavefunction for (1.7) can be written:

$$\Psi = \frac{1}{\sqrt{3}} {}^1P_{sp}(R_0, R_1, -aR_2 + b\sqrt{2}R_3) + \frac{1}{\sqrt{3}} {}^3P_{sp}(R_0, R_1, a\sqrt{2}R_2 + bR_3).$$

Since H commutes with S , the scalar product between the above two eigenfunctions of S^2 (corresponding to different eigenvalues) vanishes, and we get

$$\begin{aligned}
\langle \Psi, H\Psi \rangle &= \frac{1}{3} \left\langle {}^1P_{sp}(R_0, R_1, -aR_2 + b\sqrt{2}R_3), H {}^1P_{sp}(R_0, R_1, -aR_2 + b\sqrt{2}R_3) \right\rangle \\
&\quad + \frac{1}{3} \left\langle {}^3P_{sp}(R_0, R_1, a\sqrt{2}R_2 + bR_3), H {}^3P_{sp}(R_0, R_1, a\sqrt{2}R_2 + bR_3) \right\rangle \\
&\geq \frac{\| -aR_2 + b\sqrt{2}R_3 \|^2}{3} E_{sp}({}^1P_1) + \frac{\| a\sqrt{2}R_2 + bR_3 \|^2}{3} E_{sp}({}^3P_1) \\
&\geq \frac{\| -aR_2 + b\sqrt{2}R_3 \|^2}{3} \{ E_{sp}({}^1P_1) - E_{sp}({}^3P_1) \} + E_{sp}({}^3P_1) \geq E_{sp}({}^3P_1) \quad (1.13)
\end{aligned}$$

where $\| \cdot \|$ denotes the $L^2(r^2 dr)$ norm and we have used that

$$\| -aR_2 + b\sqrt{2}R_3 \|^2 + \| a\sqrt{2}R_2 + bR_3 \|^2 = 3(a^2 + b^2) = 3.$$

By Hund's rule (Corollary 1.1), we have $E_{sp}({}^1P_1) - E_{sp}({}^3P_1) > 0$, hence we get that $E_{sp}(J = 1) \geq E_{sp}({}^3P_1)$. Therefore there must be equality in (1.13) and it holds $aR_2 = b\sqrt{2}R_3$. Taking the square of the previous relation and using that $\int (R_2)^2 = \int (R_3)^2$, we prove the result. \square

2. Model with sp and pd configurations

In the previous section we have seen that the nonrelativistic limit of our model with only sp configurations was "correct". We now study in detail the model which was considered in [15]. The idea is to add configurations by considering d shells. The nonrelativistic wavefunction now takes the form:

$$\begin{aligned}
\Psi &= a\Phi(1s_{1/2}^2 2s_{1/2} 2p_{1/2})(R_0, R_1, R_2) + b\Phi(1s_{1/2}^2 2s_{1/2} 2p_{3/2})(R_0, R_1, R_3) \\
&\quad + c\Phi(1s_{1/2}^2 2p_{1/2} 3d_{3/2})(R_0, R_2, R_4) + d\Phi(1s_{1/2}^2 2p_{3/2} 3d_{3/2})(R_0, R_3, R_4) \\
&\quad + e\Phi(1s_{1/2}^2 2p_{3/2} 3d_{5/2})(R_0, R_3, R_5). \quad (2.1)
\end{aligned}$$

Here functions R_i are accounting for the radial part of each shell orbital, which are normalized like in (1.2). Only R_0 and R_1 have to be orthogonal. As before one can first consider the relativistic model with 4-component wavefunctions and pass to the nonrelativistic limit $c \rightarrow \infty$. One obtains the above form (2.1) of the wavefunction. The first two functions of (2.1) have already been defined in (1.5) and (1.6). The other three functions are given by (see again [3] page 294)

$$\begin{aligned}
\Phi(1s_{1/2}^2 2p_{1/2} 3d_{3/2})(R_0, R_2, R_4) &:= \frac{1}{\sqrt{3}} {}^1P_{pd}(R_0, R_2, R_4) - \frac{1}{\sqrt{6}} {}^3P_{pd}(R_0, R_2, R_4) \\
&\quad + \frac{1}{\sqrt{2}} {}^3D_{pd}(R_0, R_2, R_4), \quad (2.2)
\end{aligned}$$

$$\begin{aligned}
\Phi(1s_{1/2}^2 2p_{3/2} 3d_{3/2})(R_0, R_3, R_4) &:= -\frac{1}{\sqrt{15}} {}^1P_{pd}(R_0, R_3, R_4) + \frac{2\sqrt{2}}{\sqrt{15}} {}^3P_{pd}(R_0, R_3, R_4) \\
&\quad + \frac{\sqrt{2}}{\sqrt{5}} {}^3D_{pd}(R_0, R_3, R_4), \quad (2.3)
\end{aligned}$$

$$\begin{aligned} \Phi(1s_{1/2}^2 2p_{3/2} 3d_{5/2})(R_0, R_3, R_5) := & \frac{\sqrt{3}}{\sqrt{5}} {}^1P_{pd}(R_0, R_3, R_5) + \frac{\sqrt{3}}{\sqrt{10}} {}^3P_{pd}(R_0, R_3, R_5) \\ & - \frac{1}{\sqrt{10}} {}^3D_{pd}(R_0, R_3, R_5). \end{aligned} \quad (2.4)$$

The above nonrelativistic configurations satisfy, for $k = 0, 1$

$$(S^2 - k(k+1)) {}^{2k+1}P_{pd} = (L^2 - 2) {}^{2k+1}P_{pd} = (J^2 - 2) {}^{2k+1}P_{pd} = 0,$$

$$(S^2 - 2) {}^3D_{pd} = (L^2 - 6) {}^3D_{pd} = (J^2 - 2) {}^3D_{pd} = 0.$$

Formulas for these nonrelativistic functions of the pd shells with $J_z = 1$ are given in Appendix A.

2.1. Eigenfunctions of L^2 and S^2

Among functions of the form (2.1), we will be interested in the ones which are eigenfunctions of S^2 and L^2 , i.e. the ones which have the symmetry which is imposed in nonrelativistic calculations. Note that, contrarily to the sp mixing studied in the previous section, our wavefunction is *a priori* not even an eigenfunction of L^2 .

We will write $\Psi \in {}^1P_1$ when Ψ is a linear combination of configurations ${}^1P_{sp}$ and ${}^1P_{pd}$. We use similar notations for $\Psi \in {}^3P_1$ and $\Psi \in {}^3D_1$. The following result will be crucial in our analysis:

Theorem 2.1 (Eigenvectors of L^2 and S^2 of the form (2.1)). *Let Ψ a normalized wavefunction of the form (2.1).*

(1) *We have $\Psi \in {}^3P_1$ if and only if there exists $\varepsilon, \varepsilon' = \pm 1$ such that $R_2 = \varepsilon R_3$, $R_4 = \varepsilon' R_5$, $a = \varepsilon\sqrt{2}b$, $3d = 4\varepsilon'e$ and $4c = -\varepsilon\sqrt{5}d$. In this case*

$$\Psi = \frac{a}{\sqrt{3}} \left(\sqrt{2} + \frac{1}{\sqrt{2}} \right) {}^3P_{sp}(R_0, R_1, R_2) - c\sqrt{6} {}^3P_{pd}(R_0, R_2, R_4), \quad (2.5)$$

(2) *We have $\Psi \in {}^1P_1$ if and only if there exists $\varepsilon, \varepsilon' = \pm 1$ such that $R_2 = \varepsilon R_3$, $R_4 = \varepsilon' R_5$, $a\sqrt{2} = -\varepsilon b$, $3d = -\varepsilon'e$ and $c = -\varepsilon\sqrt{5}d$. In this case*

$$\Psi = -a\sqrt{3} {}^1P_{sp}(R_0, R_1, R_2) + c\sqrt{3} {}^1P_{pd}(R_0, R_2, R_4). \quad (2.6)$$

(3) *We have $\Psi \in {}^3D_1$ if and only if there exist $\varepsilon, \varepsilon', \varepsilon'' = \pm 1$ such that $a = b = 0$, $R_2 = \varepsilon R_3$, $R_4 = \varepsilon' R_5$, $c = \varepsilon''/\sqrt{2}$, $d = \sqrt{2/5}\varepsilon\varepsilon''$, $e = -\sqrt{1/10}\varepsilon\varepsilon''$. In this case*

$$\Psi = \varepsilon'' {}^3D_{pd}(R_0, R_2, R_4). \quad (2.7)$$

Proof. Using formulas (2.1)–(2.4), we get (for the sake of clarity, we omit to mention R_0 which appears in all configurations)

$$\begin{aligned} \Psi = & \frac{1}{\sqrt{3}} {}^3P_{sp}(R_1, a\sqrt{2}R_2 + bR_3) + \frac{1}{\sqrt{30}} \left({}^3P_{pd}(-c\sqrt{5}R_2 + 4dR_3, R_4) + {}^3P_{pd}(3eR_3, R_5) \right) \\ & + \frac{1}{\sqrt{3}} {}^1P_{sp}(R_1, -aR_2 + b\sqrt{2}R_3) + \frac{1}{\sqrt{30}} \left({}^1P_{pd}(c\sqrt{10}R_2 - d\sqrt{2}R_3, R_4) + {}^1P_{pd}(3e\sqrt{2}R_3, R_5) \right) \\ & + \frac{1}{\sqrt{30}} \left({}^3D_{pd}(c\sqrt{15}R_2 + 2\sqrt{3}dR_3, R_4) + {}^3D_{pd}(-e\sqrt{3}R_3, R_5) \right). \end{aligned} \quad (2.8)$$

Hence, using the orthogonality properties of the different configurations, we see that $\Psi \in {}^3P_1$ if and only if

$$\begin{cases} -a R_2 + b\sqrt{2} R_3 = 0, \\ c\sqrt{5}R_2 + (3e - d) R_3 = 3e(1 - (R_4, R_5)) R_3, \\ c\sqrt{5} R_2 + (2d - e) R_3 = -e(1 - (R_4, R_5)) R_3, \\ e(R_5 - (R_4, R_5)R_4) = 0. \end{cases}$$

The last equation tells us that either $e = 0$ or $R_4 = \varepsilon' R_5$ with $\varepsilon' = \pm 1$. If $e = 0$ then the second and third equations imply that $c = d = 0$ and only the first equation remains. If $R_4 = R_5$, then the system reduces to

$$\begin{cases} -a R_2 + b\sqrt{2} R_3 = 0, \\ c\sqrt{5}R_2 + (3e - d) R_3 = 0, \\ c\sqrt{5} R_2 + (2d - e) R_3 = 0, \\ R_4 = R_5. \end{cases}$$

The second and third equations then imply that $3d = 4e$. The rest follows from the normalization of R_2 and R_3 . The proof is similar for $\varepsilon' = -1$, and for 1P_1 and 3D_1 states. \square

In view of the above result, we now introduce the nonrelativistic ground state energies with $sp + pd$ mixing

$$E_{sp+pd}({}^3P_1) := \inf_{\Psi \text{ of the form (2.5)}} \langle \Psi, H\Psi \rangle, \quad (2.9)$$

$$E_{sp+pd}({}^1P_1) := \inf_{\Psi \text{ of the form (2.6)}} \langle \Psi, H\Psi \rangle, \quad (2.10)$$

$$E_{pd}({}^3D_1) := \inf_{\Psi \text{ of the form (2.7)}} \langle \Psi, H\Psi \rangle. \quad (2.11)$$

Our goal is to compare these nonrelativistic energies with the one obtained in the nonrelativistic limit:

$$E_{sp+pd}(J = 1) := \inf_{\Psi \text{ of the form (2.1)}} \langle \Psi, H\Psi \rangle. \quad (2.12)$$

By definition, we of course have

$$\boxed{E_{sp+pd}(J = 1) \leq \min \left\{ E_{sp+pd}({}^3P_1), E_{sp+pd}({}^1P_1), E_{sp+pd}({}^3D_1) \right\}.} \quad (2.13)$$

The phenomenon which was observed by Kim *et al.* in [15] was precisely that ‘the’ ground state for $E_{sp+pd}(J = 1)$ (i.e. the nonrelativistic limit of ‘the’ MCDF ground state) was *not* an eigenfunction of L and S , hence it was not a solution of any of the problems $E_{sp+pd}({}^3P_1)$, $E_{sp+pd}({}^1P_1)$ or $E_{sp+pd}({}^3D_1)$. This means that there must be a strict inequality $<$ in (2.13). This relaxation phenomenon is itself the reason for the deficiency of the nonrelativistic limit of MCDF theory. It is a typical nonlinear phenomenon.

We will now study with more details the three minimizers for (2.10), (2.11) and (2.9). We will in particular give some conditions under which (2.13) is a strict inequality, and we will check these conditions numerically in Section 3.

2.2. Study of the 1P_1 state

In this section, we prove that ‘the’ 1P_1 state minimizing $E_{sp+pd}({}^1P_1)$ is *never* a ground state for $J = 1$, although it is always a stationary point of the associated energy functional.

Theorem 2.2 (The 1P_1 state). *Let $\Psi \in {}^1P_1$ be a function of the form (2.6) minimizing $E_{sp+pd}({}^1P_1)$ defined in (2.10).*

(1) *The associated mixing coefficients and orbitals $(a, \dots, e, R_0, \dots, R_5)$ satisfying the relations of Theorem 2.1 (Assertion 2) provide a stationary point of the total energy functional*

$$(a, \dots, e, R_0, \dots, R_5) \mapsto \langle \Psi, H\Psi \rangle$$

where Ψ takes the form (2.1).

(2) *However, Ψ is never a local minimum of this functional.*

Proof. We first show that Ψ is a critical point of the total energy functional. For simplicity we assume that $c \neq 0$, the proof being the same otherwise. We also assume for simplicity that $\varepsilon = \varepsilon' = 1$, hence $R_2 = R_3$ and $R_4 = R_5$. We have to consider both variations with respect to mixing coefficients, and to orbitals. The vanishing of the variation with respect to the mixing coefficients (a, \dots, e) is a simple consequence of the fact that there is no overlap between states belonging to different symmetry spaces,

$$\langle H\Psi, {}^3P_{sp}(\dots) \rangle = \langle H\Psi, {}^3P_{pd}(\dots) \rangle = \langle H\Psi, {}^3D_{pd}(\dots) \rangle = 0,$$

for any radial functions in the corresponding states.

We now turn to the variation with respect to orbitals. By extremality of the singlet function Ψ among 1P_1 states and due to the constraints on R_0, \dots, R_5 , we have⁴

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_i} \in \text{span}(R_0, R_1) \quad \text{for } i = 0, 1, \quad (2.14)$$

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_2} + \frac{\partial \mathcal{E}(\Psi)}{\partial R_3} \Big|_{R_3=R_2} \in \text{span}(R_2) \quad (2.15)$$

and

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_4} + \frac{\partial \mathcal{E}(\Psi)}{\partial R_5} \Big|_{R_5=R_4} \in \text{span}(R_4). \quad (2.16)$$

Using $\varepsilon = 1$, $a\sqrt{2} = -b$, $3d = -e$ and $c = -\sqrt{5}d$ as given by Theorem 2.1, we see that at Ψ we have

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_2} = \frac{1}{2} \frac{\partial \mathcal{E}(\Psi)}{\partial R_3} \Big|_{R_3=R_2}, \quad (2.17)$$

and

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_4} = \frac{2}{3} \frac{\partial \mathcal{E}(\Psi)}{\partial R_5} \Big|_{R_5=R_4}, \quad (2.18)$$

⁴The derivatives appearing below are the coordinates of the gradient of the energy with respect to the scalar product of $L^2([0, \infty), r^2 dr)$.

From (2.15)-(2.17) and (2.16)-(2.18), we find

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_i} \in \text{span}(R_i) \quad \text{for } i = 2, \dots, 5 \quad (2.19)$$

which ends the proof of the criticality of Ψ .

Remark 2.1. *The exceptional relation (2.17) holds true because, as was noticed first in [15], for the singlet state the ratio between the occupation numbers of the $p_{1/2}$ function R_2 and the occupation number of the $p_{3/2}$ function R_3 in the ${}^1P_{sp}$ state, is the same as the corresponding ratio for the ${}^1P_{pd}$ state. Let us explain this with more details. Considering a variation $\delta R \in (R_2 = R_3)^\perp$, we find that the variations of the total energy functional are, using (2.8),*

$$\begin{aligned} \frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) &= 2 \left\langle H\Psi, \left(\frac{-a}{\sqrt{3}} {}^1P_{sp}(R_1, \delta R) + \frac{c\sqrt{10}}{\sqrt{30}} {}^1P_{pd}(\delta R, R_4) \right) \right\rangle, \\ \frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) &= 2 \left\langle H\Psi, \left(\frac{b\sqrt{2}}{\sqrt{3}} {}^1P_{sp}(R_1, \delta R) + \frac{-d\sqrt{2} + 3e\sqrt{2}}{\sqrt{30}} {}^1P_{pd}(\delta R, R_4) \right) \right\rangle. \end{aligned}$$

When the matrix

$$\begin{pmatrix} \frac{-a}{\sqrt{3}} & \frac{c\sqrt{10}}{\sqrt{30}} \\ \frac{b\sqrt{2}}{\sqrt{3}} & \frac{-d\sqrt{2} + 3e\sqrt{2}}{\sqrt{30}} \end{pmatrix}$$

is not invertible, its columns are colinear and we have $\frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) = k \frac{\partial \mathcal{E}(\Psi)}{\partial R_3}(\delta R)$ for some k ($k = 1/2$ in our case). Hence

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) + \frac{\partial \mathcal{E}(\Psi)}{\partial R_3}(\delta R) = 0 \iff \begin{cases} \frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) = 0, \\ \frac{\partial \mathcal{E}(\Psi)}{\partial R_3}(\delta R) = 0, \end{cases}$$

as we want. This holds true when

$$\frac{\frac{-a}{\sqrt{3}}}{\frac{b\sqrt{2}}{\sqrt{3}}} = \frac{\frac{c\sqrt{10}}{\sqrt{30}}}{\frac{-d\sqrt{2} + 3e\sqrt{2}}{\sqrt{30}}}.$$

Using the relations between a, \dots, e provided by Theorem 2.1, we see that the above equality reduces to (when $a, c \neq 0$)

$$\frac{1}{3} = \frac{10}{30} \\ \frac{2}{3} = \frac{20}{30}$$

which is precisely the ratio between the occupation numbers as mentioned before. The argument is the same for $R_4 = R_5$.

We now turn to the proof that Ψ is never a local minimum, which will simply follow from Hund's rule. Note that we can write the relations between a, b, \dots, e as

$$\begin{pmatrix} a \\ b \end{pmatrix} = -a\sqrt{3}U_{sp}^* \begin{pmatrix} 0 \\ 1 \end{pmatrix}, \quad \begin{pmatrix} c \\ d \\ e \end{pmatrix} = c\sqrt{3}U_{pd}^* \begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix}$$

where U_{sp} and U_{pd} are the Condon-Shortley unitary matrices [3]

$$U_{sp} = \frac{1}{\sqrt{3}} \begin{pmatrix} \sqrt{2} & 1 \\ -1 & \sqrt{2} \end{pmatrix}, \quad U_{pd} = \frac{1}{\sqrt{30}} \begin{pmatrix} -\sqrt{3} & 3 & 3\sqrt{2} \\ 2\sqrt{3} & 4 & -\sqrt{2} \\ \sqrt{15} & -\sqrt{5} & \sqrt{10} \end{pmatrix}.$$

Now we define the following new mixing coefficients

$$\begin{pmatrix} a' \\ b' \end{pmatrix} = -a\sqrt{3}U_{sp}^* \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \begin{pmatrix} c' \\ d' \\ e' \end{pmatrix} = c\sqrt{3}U_{pd}^* \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix}$$

and note that by construction (a', \dots, e') is orthogonal to (a, \dots, e) for the scalar product of \mathbb{R}^5 . Also we have

$$\begin{aligned} \Psi' &:= \Psi(a', b', c', d', e', R_0, R_1, R_2 = R_3, R_4 = R_5) \\ &= -a\sqrt{3}{}^3P_{sp}(R_0, R_1, R_2) + c\sqrt{3}{}^3P_{pd}(R_0, R_2, R_4), \end{aligned}$$

i.e. it takes exactly the same form as Ψ but with triplet states instead of singlet states. Now we vary the mixing coefficients as follows $\sqrt{1-t^2}(a, \dots, e) + t(a', \dots, e')$, which results into a variation for the wavefunction of the form $\sqrt{1-t^2}\Psi + t\Psi'$. Calculating the energy of this new wavefunction we find

$$\mathcal{E}(\sqrt{1-t^2}\Psi + t\Psi') = \mathcal{E}(\Psi) + t^2(\mathcal{E}(\Psi') - \mathcal{E}(\Psi)).$$

Note that there is no first order term since Ψ is a stationary state as shown before (or simply because Ψ and Ψ' belong to different symmetry spaces, hence $\langle H\Psi, \Psi' \rangle = 0$).

Now we claim that $\mathcal{E}(\Psi') < \mathcal{E}(\Psi)$, which will clearly imply that Ψ cannot be a local minimum. By (A.8) and (A.15) in Appendix A, we know that $\mathcal{E}(\Psi') = \mathcal{E}(\Psi'')$ where

$$\Psi'' = -\frac{a\sqrt{3}}{\sqrt{2}}S^-{}^3P_{2,sp}(R_0, R_1, R_2) + \frac{c\sqrt{3}}{\sqrt{2}}S^-{}^3P_{2,pd}(R_0, R_2, R_4)$$

which is the simple triplet state taking the same form as Ψ but with the adequate signs reversed. The last step is to apply Theorem 1.2, with the following functions: $f_1 = s(R_0)$, $g_1 = p_1(R_2)$, $g_2 = (c^2/20 + a^2)^{-1/2}(c/(2\sqrt{5})d_0(R_4) + as(R_1))$, $g_3 = p_{-1}(R_2)$, $g_4 = d_2(R_4)$, $g_5 = p_0(R_2)$ and $g_6 = d_1(R_4)$. This ends the proof of Theorem 2.2. \square

Remark 2.2. When $R_2 = \pm R_3$ and $R_4 = R_5$, the problem consisting of varying only the mixing coefficient essentially reduces to that of finding the eigenvalues of the nonrelativistic Hamiltonian matrix, i.e. the matrix of H in the space spanned by the 5 configurations built upon the orbitals:

$${}^1P_{sp}(R_0, R_1, R_2), \quad {}^1P_{pd}(R_0, R_2, R_4), \quad {}^3P_{sp}(R_0, R_1, R_2), \quad {}^3P_{pd}(R_0, R_2, R_4), \quad {}^3D_{pd}(R_0, R_2, R_4).$$

This 5×5 matrix is block diagonal and its eigenvalues are:

$$\lambda_1({}^kP_1) = \inf_{\alpha^2 + \beta^2 = 1} \mathcal{E}(\alpha {}^kP_{sp}(R_0, R_1, R_2) + \beta {}^kP_{pd}(R_0, R_2, R_4))$$

$$\lambda_2({}^kP_1) = \sup_{\alpha^2 + \beta^2 = 1} \mathcal{E}(\alpha {}^kP_{sp}(R_0, R_1, R_2) + \beta {}^kP_{pd}(R_0, R_2, R_4))$$

$$\lambda({}^3D_1) = \langle H {}^3D_{pd}(R_0, R_2, R_4), {}^3D_{pd}(R_0, R_2, R_4) \rangle.$$

What we have used in this second part is that $\lambda_1({}^3P_1) < \lambda_1({}^1P_1)$, by Hund's rule (indeed, we even have that the 2×2 Hamiltonian matrix of 3P_1 states is smaller than the one of 1P_1 states, in the sense of quadratic forms). However although it is expected that in many cases $\lambda_1({}^1P_1) < \lambda({}^3D_1)$, there is no general rule: this may depend on the orbitals R_0, \dots, R_5 .

2.3. Study of the 3D_1 state

In this section, we prove that 'the' 3D_1 state minimizing $E_{pd}({}^3D_1)$ is also always a stationary point of the associated energy functional and we give a condition implying that it is not a local minimum.

Theorem 2.3 (The 3D_1 state). *Let $\Psi \in {}^3D_1$ be a function of the form (2.7) minimizing $E_{pd}({}^3D_1)$ defined in (2.11).*

(1) *The associated mixing coefficients and orbitals $(a, \dots, e, R_0, \dots, R_5)$ given by Theorem 2.1 provide a stationary point of the total energy functional*

$$(a, \dots, e, R_0, \dots, R_5) \mapsto \langle \Psi, H\Psi \rangle \quad (2.20)$$

where Ψ takes the form (2.1).

(2) *If moreover*

$$\inf_{\substack{\alpha^2 + \beta^2 = 1 \\ R_1 \in \{R_0\}^\perp, \|R_1\|_{L^2(r^2 dr)} = 1}} \mathcal{E}(\alpha {}^3P_{sp}(R_0, R_1, R_2) + \beta {}^3P_{pd}(R_0, R_2, R_4)) < E_{pd}({}^3D_1), \quad (2.21)$$

where R_0, R_2, R_4 are the radial functions of Ψ , then the stationary state Ψ is not a local minimum of the functional (2.20).

Remark 2.3. *Our assumption (2.21) exactly means that $\lambda({}^3D_1) = E_{pd}({}^3D_1)$ is not the lowest eigenvalue of the Hamiltonian matrix as explained above in Remark 2.2. It is not clear how to prove (2.21) rigorously. In Section 3 we will verify it numerically on the approximated solutions provided by the ATSP HF and MCHF program by Froese-Fischer [12].*

Proof. We do not give all the details of the proof which is very similar to that of the 1P_1 state. The fact that Ψ is a stationary point is seen exactly as in Theorem 2.2.

The proof that Ψ is not a local minimum is also very similar to that of the 1P_1 state, with the difference that we do not have a general Hund's rule for 3D_1 states, hence we need to require condition (2.21), which is expected to be true in many cases. This time we have (assuming again $\varepsilon = \varepsilon' = \varepsilon'' = 1$ for simplicity)

$$\begin{pmatrix} a \\ b \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}, \quad \begin{pmatrix} c \\ d \\ e \end{pmatrix} = U_{pd}^* \begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}.$$

The result is then obtained by arguing as before with, this time,

$$\begin{pmatrix} a' \\ b' \end{pmatrix} = \alpha U_{sp}^* \begin{pmatrix} 0 \\ 1 \end{pmatrix}, \quad \begin{pmatrix} c' \\ d' \\ e' \end{pmatrix} = \beta U_{pd}^* \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix},$$

where α and β are chosen to minimize the left side of (2.21). □

2.4. Study of the 3P_1 state

In this section, we give a simple condition implying that ‘the’ 3P_1 state minimizing $E_{sp+pd}({}^3P_1)$ is *not* a stationary point of the total energy functional.

Theorem 2.4 (The 3P_1 state). *Let $\Psi \in {}^3P_1$ be a function of the form (2.5) minimizing $E_{sp+pd}({}^3P_1)$ defined in (2.9), and denote by $(a, \dots, e, R_0, \dots, R_5)$ the associated mixing coefficients and orbitals satisfying the relations of Theorem 2.1.*

If there exists $\delta R \in (R_3)^\perp$ such that

$$c\langle H\Psi, {}^3P_{pd}(R_0, \delta R, R_4) \rangle \neq 0 \quad \text{or} \quad a\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle \neq 0, \quad (2.22)$$

then Ψ does not provide a stationary point of the total energy functional

$$(a, \dots, e, R_0, \dots, R_5) \mapsto \langle \Psi, H\Psi \rangle.$$

Remark 2.4. *Condition (2.22) is very intuitive. It indeed implies that*

$$E_{sp+p'd}({}^3P_1) < E_{sp+pd}({}^3P_1), \quad (2.23)$$

where

$$E_{sp+p'd}({}^3P_1) = \inf_{\substack{\alpha^2 + \beta^2 = 1, \\ (R_0, \dots, R_4) \\ \text{satisfying (1.1) and (1.2)}}} \mathcal{E}(\alpha {}^3P_{sp}(R_0, R_1, R_2) + \beta {}^3P_{pd}(R_0, R_3, R_4)).$$

The relation between (2.22) and (2.23) was already noticed in [11].

Note that when $a, c \neq 0$, one can prove that $\langle H\Psi, {}^3P_{pd}(R_0, \delta R, R_4) \rangle \neq 0$ or that $\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle \neq 0$ for some δR and these two conditions are indeed equivalent. In Section 3, the condition $\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle \neq 0$ is verified numerically.

We now give the

Proof. As before, it can easily be seen that our state Ψ is indeed a stationary state with respect to variations of the mixing coefficients only. The non-criticality will come from the variations of the orbitals, as suggested by (2.22). As before we assume for simplicity that $\varepsilon = \varepsilon' = 1$, hence $R_2 = R_3$ and $R_4 = R_5$ for the ground state Ψ .

By extremality of the triplet function Ψ among 3P_1 states, we have similarly as before

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_i} \in \text{span}(R_0, R_1) \quad \text{for } i = 0, 1, \quad (2.24)$$

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_2} + \frac{\partial \mathcal{E}(\Psi)}{\partial R_3} \Big|_{R_3=R_2} \in \text{span}(R_2) \quad (2.25)$$

and

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_4} + \frac{\partial \mathcal{E}(\Psi)}{\partial R_5} \Big|_{R_5=R_4} \in \text{span}(R_4). \quad (2.26)$$

This time we find using the relations of Theorem 2.1 that

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_4} = \frac{\partial \mathcal{E}(\Psi)}{\partial R_5} \Big|_{R_5=R_4},$$

hence only variations with respect to R_2 and R_3 remain to be considered. The main point is that there is *a priori* no relation between $\frac{\partial \mathcal{E}(\Psi)}{\partial R_2}$ and $\frac{\partial \mathcal{E}(\Psi)}{\partial R_3}$. More precisely, let us consider a variation $\delta R \in (R_2 = R_3)^\perp$. We have

$$\frac{\partial \mathcal{E}(\Psi)}{\partial R_2}(\delta R) = 2 \left\langle H\Psi, \left(\frac{a\sqrt{2}}{\sqrt{3}} {}^3P_{sp}(R_1, \delta R) + \frac{-c\sqrt{5}}{\sqrt{30}} {}^3P_{pd}(\delta R, R_4) \right) \right\rangle$$

and

$$\begin{aligned} \frac{\partial \mathcal{E}(\Psi)}{\partial R_3}(\delta R) &= 2 \left\langle H\Psi, \left(\frac{b}{\sqrt{3}} {}^3P_{sp}(R_1, \delta R) + \frac{4d+3e}{\sqrt{30}} {}^3P_{pd}(\delta R, R_4) \right) \right\rangle \\ &= 2 \left\langle H\Psi, \left(\frac{a}{\sqrt{6}} {}^3P_{sp}(R_1, \delta R) + \frac{-5c\sqrt{5}}{\sqrt{30}} {}^3P_{pd}(\delta R, R_4) \right) \right\rangle, \end{aligned}$$

where we have used the relations of Theorem 2.1 in the last line. The main difference with the 1P_1 state is now that the matrix

$$\begin{pmatrix} \frac{\sqrt{2}}{\sqrt{3}} & -\frac{\sqrt{5}}{\sqrt{30}} \\ \frac{1}{\sqrt{6}} & -\frac{5\sqrt{5}}{\sqrt{30}} \end{pmatrix} \quad (2.27)$$

is invertible. Hence we get that Ψ is stationary with respect to variations of R_2 and R_3 independently if and only if

$$\forall \delta R \in (R_2 = R_3)^\perp, \quad a \langle H\Psi, {}^3P_{sp}(R_1, \delta R) \rangle = c \langle H\Psi, {}^3P_{pd}(\delta R, R_4) \rangle = 0.$$

This clearly leads to a contradiction when (2.22) holds true. \square

Remark 2.5. *The fact that the matrix (2.27) is invertible can be interpreted in saying that the ratio between the occupation number of the $p_{1/2}$ function R_2 and the one of the $p_{3/2}$ function R_3 in the *sp* configuration is not the same as the one of the *pd* configuration [15].*

2.5. Conclusion

In the previous sections we have studied the states of the nonrelativistic symmetries 3P_1 , 1P_1 and 3D_1 . As a consequence of our results we obtain the

Corollary 2.1 (Occurrence of symmetry breaking for $sp + pd$). *Assume that (2.21) holds for a 3D_1 ground state and that condition (2.22) holds for a 3P_1 ground state. Then we have*

$$E_{sp+pd}(J=1) < \min \left\{ E_{sp+pd}({}^3P_1), E_{sp+pd}({}^1P_1), E_{sp+pd}({}^3D_1) \right\}. \quad (2.28)$$

Additionally a ground state Ψ for $E_{sp+pd}(J=1)$ is never an eigenfunction of L^2 and neither of S^2 .

Proof. The strict inequality (2.28) is an obvious consequence of the previous results. It implies that a ground state for the minimization problem $E_{sp+pd}(J=1)$ cannot be a common eigenfunction of L^2 and S^2 . What remains to be proven is that it cannot even be an eigenfunction of L^2 or of S^2 separately. This means that any ground state must have a nonvanishing projection in each of the symmetries 1P_1 , 3P_1 and 3D_1 .

Assume for instance that Ψ is an eigenfunction of L^2 . It cannot be a 3D_1 state by (2.28), hence one must have $\Psi \in {}^1P_1 + {}^3P_1$. Using (2.8), one obtains the conditions

$$\begin{cases} c\sqrt{15}R_2 + (2d\sqrt{3} - e\sqrt{3}(R_4, R_5))R_3 = 0, \\ e(R_5 - (R_4, R_5)R_4) = 0. \end{cases}$$

If $e = 0$ we get $R_2 = \varepsilon R_3$ and $c\sqrt{5} = -2\varepsilon d$ for some $\varepsilon \in \{\pm 1\}$. Hence

$$\begin{aligned} \Psi = \frac{a\sqrt{2} + \varepsilon b}{\sqrt{3}} {}^3P_{sp}(R_1, R_2) + \frac{-c\sqrt{2}}{\sqrt{3}} {}^3P_{pd}(R_2, R_4) \\ + \frac{-a + b\varepsilon\sqrt{2}}{\sqrt{3}} {}^1P_{sp}(R_1, R_2) + \frac{c\sqrt{3}}{2} {}^1P_{pd}(R_2, R_4). \end{aligned}$$

By Hund's rule we get

$$\begin{aligned} \mathcal{E}(\Psi) \geq & \left\| \frac{a\sqrt{2} + \varepsilon b}{\sqrt{3}} {}^3P_{sp}(R_1, R_2) + \frac{-c\sqrt{2}}{\sqrt{3}} {}^3P_{pd}(R_2, R_4) \right\|^2 E_{sp+pd}({}^3P_1) \\ & + \left\| \frac{-a + b\varepsilon\sqrt{2}}{\sqrt{3}} {}^1P_{sp}(R_1, R_2) + \frac{c\sqrt{3}}{2} {}^1P_{pd}(R_2, R_4) \right\|^2 E_{sp+pd}({}^1P_1) \geq E_{sp+pd}({}^3P_1) \end{aligned}$$

which contradicts (2.28). The argument is the same if $R_4 = \pm R_5$. Hence we have shown that Ψ cannot be an eigenfunction of L^2 . The proof that Ψ cannot be an eigenfunction of S^2 is very similar. \square

3. Numerical verification of (2.21) and (2.22)

In this section we verify numerically the two assumptions made in the previous section. We use the package ATSP of Froese-Fischer [12] to get approximations of the nonrelativistic states in the different symmetry spaces. The program uses 220 discretization points on a logarithmic grid and a finite difference method. We have then verified conditions (2.21) and (2.22) with the help of *Mathematica*.

3.1. Verification of (2.22)

First we start with the MCHF calculation for the configuration ${}^3P_{sp+pd}$ for the Beryllium atom. This yields constants a, c and functions R_0, R_1, R_2, R_4 , which are numerical approximations of the real ones. We can also run the program for the simpler Hartree-Fock cases ${}^3P_{sp}$, ${}^3P_{pd}$ and we get the following total energies (in Hartree).

$$\begin{aligned} E_{sp}({}^3P_1) &\simeq -14.5115, \\ E_{sp+pd}({}^3P_1) &\simeq -14.5166. \end{aligned}$$

Also we obtain for the ${}^3P_{sp+pd}$ configuration the following numerical values for a and c :

$$a \simeq 0.9951963, \quad c \simeq -0.0978997.$$

The plots of the radial parts of the orbitals $R_0, R_1, R_2 = R_3$ and $R_4 = R_5$ (corresponding respectively to the shells $1s, 2s, 2p$ and $3d$) are displayed in Figure 1 below.

Now, we are going to verify (2.22) by showing the existence of a function δR , orthogonal to R_2 , such that

$$\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle \neq 0. \quad (3.1)$$

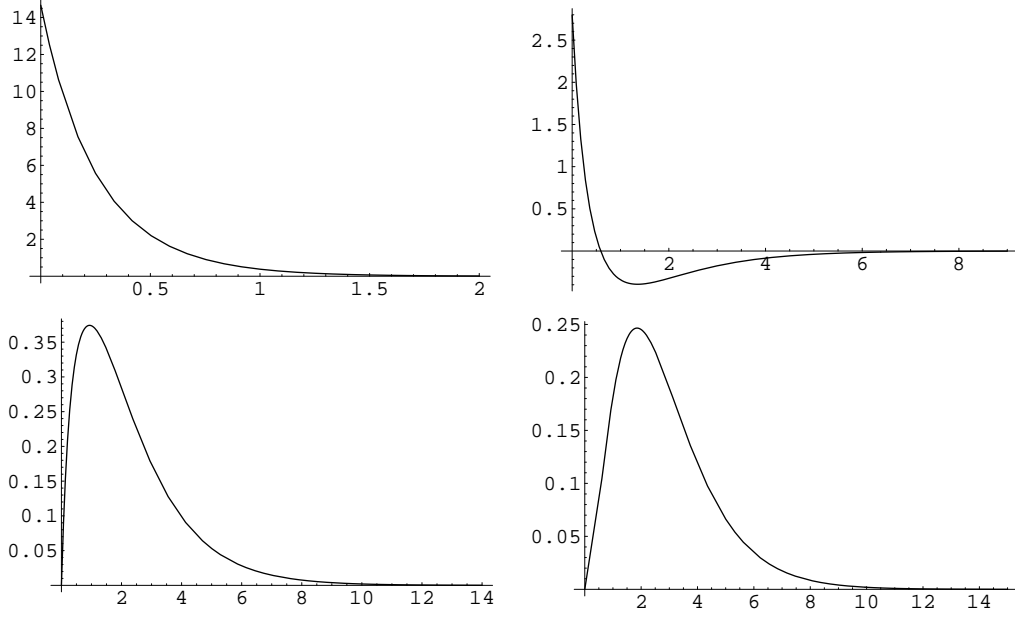


Fig. 1. Plots of R_0 (top left), R_1 (top right), R_2 (bottom left) and R_4 (bottom right) for the optimal ${}^3P_1(sp+pd)$ configuration of the Beryllium atom, obtained with the package ATSP.

As explained in detail in Appendix B, we can calculate the exact expression of (3.1) in terms of the radial functions only. We obtain

$$\begin{aligned}
\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle &= a \int_0^\infty \left(\frac{s^4}{2} \left(\frac{R_2(s)}{s} \right)' \left(\frac{\delta R(s)}{s} \right)' - 4s R_2(s) \delta R(s) \right) ds \\
&\quad + a \int_0^\infty \int_0^\infty (2|R_0(s)|^2 + |R_1(s)|^2) R_2(t) \delta R(t) \frac{s^2 t^2}{\max\{s, t\}} ds dt \\
&\quad - a \int_0^\infty \int_0^\infty R_0(s) \delta R(s) R_2(t) R_0(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt \\
&\quad - a \int_0^\infty \int_0^\infty R_1(s) \delta R(s) R_2(t) R_1(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt \\
&\quad - c \sqrt{2} \int_0^\infty \int_0^\infty R_1(s) R_2(s) \delta R(t) R_4(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt.
\end{aligned}$$

Finding a function δR , orthogonal to R_2 , for which the above formula is away from 0 is less convincing than arguing as follows. The above formula is linear with respect to δR hence it can be written

$$\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle = \int_0^\infty F(r) \delta R(r) dr$$

with

$$\begin{aligned}
F(r) := & a \left(-\frac{r^2}{2} R_2''(r) - r R_2'(r) + R_2(r) - 4r R_2(r) \right) + a \int_0^\infty (2|R_0(s)|^2 + |R_1(s)|^2) \frac{R_2(r) s^2 r^2}{\max\{s, r\}} ds \\
& - a \int_0^\infty R_0(r) R_2(s) R_0(s) \frac{r^2 s^2 \min\{r, s\}}{3 \max^2\{r, s\}} ds - a \int_0^\infty R_1(r) R_2(s) R_1(s) \frac{r^2 s^2 \min\{r, s\}}{3 \max^2\{r, s\}} ds \\
& - c \sqrt{2} \int_0^\infty R_1(s) R_2(s) R_4(r) \frac{s^2 r^2 \min\{s, r\}}{3 \max^2\{r, t\}} ds.
\end{aligned}$$

Then we note that the vanishing of $\int_0^\infty F(r) \delta R(r) dr$ for all δR orthonormal to R_2 , is indeed equivalent to the existence of a Lagrange multiplier λ such that $F = \lambda R_2$. Hence it suffices to show that $F(r)/R_2(r)$ is not constant in order to verify (2.22).

We compute numerically F/R_2 and get the following graph:

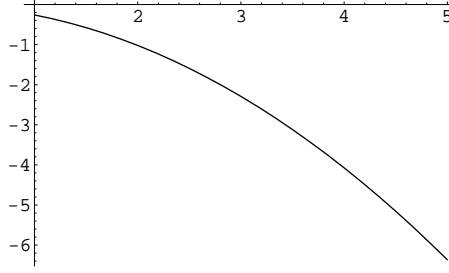


Fig. 2. Plot of $F(r)/R_2(r)$ for r between 1 and 5.

This clearly shows that the function F/R_2 is not constant. Hence (2.22) holds true.

3.2. Verification of (2.21)

Let us now pass to the numerical verification of (2.21). In order to do so, we again run the Froese-Fischer ATSP program [12] for the HF calculation of the ${}^3D_{pd}$ configuration for the Beryllium atom. This provides us with radial functions R_0 , R_2 and R_4 for the orbitals $1s$, $2p$, $3d$. For the $2s$ orbital, we construct a function R_1 by taking the one obtained in the calculation for the ${}^3P_{sp+pd}$ configuration and projecting it on the orthogonal to the space generated by R_0 . We get the following numbers for the total energies (in Hartree):

$$\begin{aligned}
E_{pd}({}^3D_1) & \simeq -14.1889, \\
\mathcal{E}({}^3P_{sp}(R_0, R_1, R_2)) & \simeq -14.4998,
\end{aligned}$$

thus yielding a numerical verification of condition (2.21).

Appendix A. Nonrelativistic configurations in the sector $J = 1$ and $J_z = 1$

For the convenience of the reader, we quickly explain how to construct the nonrelativistic configurations used in the text. In the whole appendix we use the notation

$$s^\tau(R)(x, \sigma) = R(|x|) \delta_\tau(\sigma), \quad p_m^\tau(R)(x, \sigma) = R(|x|) Y_1^m \left(\frac{x}{|x|} \right) \delta_\tau(\sigma),$$

$S_z \backslash L_z$	-1	0	1
-1	${}^3P_2 = 4 s^\downarrow \wedge p_{-1}^\downarrow$	$2\sqrt{2} s^\downarrow \wedge p_0^\downarrow$	$2 s^\downarrow \wedge p_1^\downarrow$
0	${}^1P_1 = 2 \left(s^\uparrow \wedge p_{-1}^\downarrow - s^\downarrow \wedge p_{-1}^\uparrow \right)$ $2 \left(s^\uparrow \wedge p_{-1}^\downarrow + s^\downarrow \wedge p_{-1}^\uparrow \right)$	${}^1P_1 = \sqrt{2} \left(s^\uparrow \wedge p_0^\downarrow - s^\downarrow \wedge p_0^\uparrow \right)$ $\sqrt{2} \left(s^\uparrow \wedge p_0^\downarrow + s^\downarrow \wedge p_0^\uparrow \right)$	${}^1P_1 = s^\uparrow \wedge p_1^\downarrow - s^\downarrow \wedge p_1^\uparrow$ $s^\uparrow \wedge p_1^\downarrow + s^\downarrow \wedge p_1^\uparrow$
1	$2 s^\uparrow \wedge p_{-1}^\uparrow$	$\sqrt{2} s^\uparrow \wedge p_0^\uparrow$	${}^3P_2 = s^\uparrow \wedge p_1^\uparrow$

$J_z = 1$
 $J_z = 2$

Table 1. Derivation of all the common eigenvectors of L^2 , S^2 , L_z and S_z , starting from the ${}^3P_{2,sp}$ state having all the highest possible quantum numbers, at the bottom right of the table. For the sake of clarity, we have not indicated the closed shell $s_0^\uparrow(R_0) \wedge s_0^\downarrow(R_0) \wedge \dots$ which appears in front of all the configurations. We have also refrained from mentioning the radial functions R_1 and R_2 .

$$d_m^\tau(R)(x, \sigma) = R(|x|) Y_2^m \left(\frac{x}{|x|} \right) \delta_\tau(\sigma)$$

for $\tau \in \{\uparrow, \downarrow\}$ and $m = -\ell, \dots, \ell$. Here Y_ℓ^m are the usual eigenfunctions of L^2 and L_z (orbital angular momentum), normalized in the Hilbert space $L^2(S^2)$, such that

$$L^2 Y_\ell^m = \ell(\ell+1) Y_\ell^m, \quad L_z Y_\ell^m = m Y_\ell^m,$$

and which can be expressed in terms of Legendre polynomials [3].

sp configurations

We start with *sp* configurations. Table 1 summarizes the different possible states which are eigenfunctions of both L^2 and S^2 , classified with respect to their L_z and S_z . The usual method is to start from the state having the highest possible quantum numbers ($S = 1$, $S_z = 1$, $L = 1$, $L_z = 1$, $J = 2$ and $J_z = 2$) which is located at the lower right corner:

$${}^3P_{2,sp}(R_0, R_1, R_2) := s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_1^\uparrow(R_2) \right). \quad (\text{A.1})$$

Next we apply the lowering operators $L^- = L_x - iL_y$ and $S^- = S_x - iS_y$ to get two states having the same $S = 1$ and $L = 1$ but without a precise J :

$$S^-({}^3P_{2,sp}(R_0, R_1, R_2)) = s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_1^\downarrow(R_2) + s^\downarrow(R_1) \wedge p_1^\uparrow(R_2) \right), \quad (\text{A.2})$$

$$L^-({}^3P_{2,sp}(R_0, R_1, R_2)) := \sqrt{2} s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_0^\uparrow(R_2) \right). \quad (\text{A.3})$$

A usual antisymmetrization of (A.2) gives our (normalized) singlet state having $S = 0$, hence $J = L = 1$:

$${}^1P_{sp}(R_0, R_1, R_2) = \frac{1}{\sqrt{2}} s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(s^\uparrow(R_1) \wedge p_1^\downarrow(R_2) - s^\downarrow(R_1) \wedge p_1^\uparrow(R_2) \right). \quad (\text{A.4})$$

In the subspace $J_z = 1$, we can construct a 3P_2 state by applying the lowering operator $J^- = L^- + S^-$:

$$J^- ({}^3P_{2,sp}(R_0, R_1, R_2)) = L^- ({}^3P_{2,sp}(R_0, R_1, R_2)) + S^- ({}^3P_{2,sp}(R_0, R_1, R_2)).$$

We know that the subspace $J_z = 1$ is of dimension 3. Hence we deduce by orthogonality that our (normalized) 3P_1 state with $J = 1$ and $J_z = 1$ is

$${}^3P_{sp}(R_0, R_1, R_2) = \frac{(L^- - S^-)}{2} {}^3P_{2,sp}(R_0, R_1, R_2). \quad (\text{A.5})$$

Note that

$$L^+ L^- ({}^3P_{2,sp}(R_0, R_1, R_2)) = 2 ({}^3P_{2,sp}(R_0, R_1, R_2)), \quad (\text{A.6})$$

$$S^+ S^- ({}^3P_{2,sp}(R_0, R_1, R_2)) = 2 ({}^3P_{2,sp}(R_0, R_1, R_2)), \quad (\text{A.7})$$

which explains why $1/2$ is the right normalization in (A.5). From (A.6) and (A.7) we see that for any observable A commuting with L and S (for instance $A = H$, our nonrelativistic Hamiltonian), then we have for all a, b and all R_0, R_1, R_2 and R'_0, R'_1, R'_2

$$\begin{aligned} & \left\langle (aL^- + bS^-) ({}^3P_{2,sp}(R_0, R_1, R_2)), A (aL^- + bS^-) ({}^3P_{2,sp}(R'_0, R'_1, R'_2)) \right\rangle_{L^2(\mathbb{R}^3)^4} \\ &= (a^2 + b^2) \left\langle L^- ({}^3P_{2,sp}(R_0, R_1, R_2)), A L^- ({}^3P_{2,sp}(R'_0, R'_1, R'_2)) \right\rangle_{L^2(\mathbb{R}^3)^4} \\ &= (a^2 + b^2) \left\langle S^- ({}^3P_{2,sp}(R_0, R_1, R_2)), A S^- ({}^3P_{2,sp}(R'_0, R'_1, R'_2)) \right\rangle_{L^2(\mathbb{R}^3)^4} \\ &= 2(a^2 + b^2) \left\langle {}^3P_{2,sp}(R_0, R_1, R_2), A {}^3P_{2,sp}(R'_0, R'_1, R'_2) \right\rangle_{L^2(\mathbb{R}^3)^4}. \quad (\text{A.8}) \end{aligned}$$

pd configurations

We now switch to the calculation of the *pd* configurations.

Table 2 shows the dimension of the different common eigenspaces of L_z and S_z for *pd* configurations. As before we may start from the 3F_4 state having the highest quantum numbers and derive our states in the $J = 1, J_z = 1$ sector by applying successively the lowering operators L^-, S^- and J^- .

The final result is as follows. The 1P_1 state in the $J_z = 1$ sector reads

$$\begin{aligned} {}^1P_{pd}(R_0, R_1, R_2) &= \frac{1}{2\sqrt{5}} s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(\sqrt{6} p_{-1}^\uparrow(R_1) \wedge d_2^\downarrow(R_2) - \sqrt{3} p_0^\uparrow(R_1) \wedge d_1^\downarrow(R_2) \right. \\ &\quad \left. + p_1^\uparrow(R_1) \wedge d_0^\downarrow(R_2) - \sqrt{6} p_{-1}^\downarrow(R_1) \wedge d_2^\uparrow(R_2) + \sqrt{3} p_0^\downarrow(R_1) \wedge d_1^\uparrow(R_2) - p_1^\downarrow(R_1) \wedge d_0^\uparrow(R_2) \right). \quad (\text{A.9}) \end{aligned}$$

$S_z \backslash L_z$	-3	-2	-1	0	1	2	3
-1	1	2	3	3	3	2	1
0	2	4	6	6	6	4	2
1	1	2	3	3	3	2	1

\$J_z = 1\$ \$J_z = 2\$ \$J_z = 3\$ \$J_z = 4\$

Table 2. Dimensions of the different common eigenspaces of L_z and S_z for pd configurations. In each cell, we have indicated the type of state spanning the corresponding space.

The 3P_1 state reads

$${}^3P_{pd}(R_0, R_1, R_2) = \frac{L^- - S^-}{2} {}^3P_{2,pd}(R_0, R_1, R_2) \quad (\text{A.10})$$

where

$${}^3P_{2,pd}(R_0, R_1, R_2) = s^\dagger(R_0) \wedge s^\dagger(R_0) \wedge \left(\frac{\sqrt{6}}{\sqrt{10}} p_{-1}^\dagger(R_1) \wedge d_2^\dagger(R_2) - \frac{\sqrt{3}}{\sqrt{10}} p_0^\dagger(R_1) \wedge d_1^\dagger(R_2) + \frac{1}{\sqrt{10}} p_1^\dagger(R_1) \wedge d_0^\dagger(R_2) \right) \quad (\text{A.11})$$

is the unique 3P_2 configuration having $J_z = 2$. Note that, like in the sp case, one has

$$L^+ L^- ({}^3P_{2,pd}(R_0, R_1, R_2)) = 2 ({}^3P_{2,pd}(R_0, R_1, R_2)), \quad (\text{A.12})$$

$$S^+ S^- ({}^3P_{2,pd}(R_0, R_1, R_2)) = 2 ({}^3P_{2,pd}(R_0, R_1, R_2)). \quad (\text{A.13})$$

Therefore we deduce that (A.8) still holds for linear combinations of the triplets sp and pd

and that for any observable A commuting with L and S ,

$$\begin{aligned} & \left\langle a {}^3P_{sp}(R_0, R_1, R_2) + b {}^3P_{pd}(R_0, R_3, R_4), A \left(a' {}^3P_{sp}(R'_0, R'_1, R'_2) + b' {}^3P_{pd}(R'_0, R'_3, R'_4) \right) \right\rangle_{L^2(\mathbb{R}^3)^4} \\ &= \frac{1}{2} \left\langle a S^- ({}^3P_{2,sp}(R_0, R_1, R_2)) + b S^- ({}^3P_{2,pd}(R_0, R_3, R_4)), \right. \\ & \quad \left. A \left(a' S^- ({}^3P_{2,sp}(R'_0, R'_1, R'_2)) + b' S^- ({}^3P_{2,pd}(R'_0, R'_3, R'_4)) \right) \right\rangle_{L^2(\mathbb{R}^3)^4} \end{aligned} \quad (\text{A.14})$$

$$\begin{aligned} &= \left\langle a {}^3P_{2,sp}(R_0, R_1, R_2) + b {}^3P_{2,pd}(R_0, R_3, R_4), \right. \\ & \quad \left. A \left(a' {}^3P_{2,sp}(R'_0, R'_1, R'_2) + b' {}^3P_{2,pd}(R'_0, R'_3, R'_4) \right) \right\rangle_{L^2(\mathbb{R}^3)^4}. \end{aligned} \quad (\text{A.15})$$

Lastly the 3D_1 state is obtained through the more complicated formula

$${}^3D_{pd}(R_0, R_1, R_2) = \frac{1}{6\sqrt{2}} (L^- L^- - 3L^- S^- + 6S^- S^-) {}^3D_{3,pd}(R_0, R_1, R_2) \quad (\text{A.16})$$

where ${}^3D_{3,pd}(R_0, R_1, R_2)$ is the unique 3D_3 state having $J_z = 3$, given by

$${}^3D_{3,pd}(R_0, R_1, R_2) = s^\uparrow(R_0) \wedge s^\downarrow(R_0) \wedge \left(\frac{\sqrt{2}}{\sqrt{3}} p_0^\uparrow(R_1) \wedge d_2^\uparrow(R_2) - \frac{1}{\sqrt{3}} p_1^\uparrow(R_1) \wedge d_1^\uparrow(R_2) \right). \quad (\text{A.17})$$

Appendix B. Energy expressions as functions of the radial components

In this appendix, we provide the formulas of the energy written in terms of the radial components of the orbitals. We will first need the

Lemma B.1. *Assume that we have 6 mutually orthogonal functions, f_1, \dots, f_6 . Then,*

$$\left\langle f_1 \wedge f_2 \wedge f_3 \wedge f_4, \sum_{1 \leq i < j \leq 4} \left(\frac{1}{|x_i - x_j|} \right) f_1 \wedge f_2 \wedge f_5 \wedge f_6 \right\rangle = \left\langle f_3 \wedge f_4, \frac{1}{|x - y|} f_5 \wedge f_6 \right\rangle, \quad (\text{B.18})$$

$$\begin{aligned} & \left\langle f_1 \wedge f_2 \wedge f_3 \wedge f_4, \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) f_1 \wedge f_2 \wedge f_3 \wedge f_6 \right\rangle \\ &= \left\langle f_1 \wedge f_4, \frac{1}{|x - y|} f_1 \wedge f_6 \right\rangle + \left\langle f_2 \wedge f_4, \frac{1}{|x - y|} f_2 \wedge f_6 \right\rangle + \left\langle f_3 \wedge f_4, \frac{1}{|x - y|} f_3 \wedge f_6 \right\rangle, \end{aligned} \quad (\text{B.19})$$

and

$$\begin{aligned} & \left\langle f_1 \wedge f_2 \wedge f_3 \wedge f_4, \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) f_1 \wedge f_2 \wedge f_3 \wedge f_4 \right\rangle \\ &= \sum_{1 \leq i < j \leq 4} \left\langle f_i \wedge f_j, \frac{1}{|x - y|} f_i \wedge f_j \right\rangle. \end{aligned} \quad (\text{B.20})$$

The proof of the above lemma is based on long but straightforward computations. We use the above result to compute some quantities that are needed in order to perform the numerical computations of Section 3.

We start with the computation of $\langle {}^3P_{sp}(R_0, R_1, R_2), H {}^3P_{sp}(R_0, R_1, R_2) \rangle$ which is needed to verify (2.21). By (A.8) we have

$$\langle {}^3P_{sp}(R_0, R_1, R_2), H {}^3P_{sp}(R_0, R_1, R_2) \rangle = \langle {}^3P_{2,sp}(R_0, R_1, R_2), H {}^3P_{2,sp}(R_0, R_1, R_2) \rangle.$$

A simple calculation shows that

$$\begin{aligned} & \langle {}^3P_{2,sp}(R_0, R_1, R_2), H {}^3P_{2,sp}(R_0, R_1, R_2) \rangle \\ &= \int_{\mathbb{R}^3} \left(|\nabla s_0(R_0)(x)|^2 + \frac{|\nabla s_0(R_1)(x)|^2 + |\nabla p_1(R_2)(x)|^2}{2} \right) dx \end{aligned} \quad (\text{B.21})$$

$$- \int_{\mathbb{R}^3} \frac{4}{|x|} (2|s_0(R_0)(x)|^2 + |s_0(R_1)(x)|^2 + |p_1(R_2)(x)|^2) dx \quad (\text{B.22})$$

$$+ \left\langle {}^3P_{2,sp}(R_0, R_1, R_2), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{2,sp}(R_0, R_1, R_2) \right\rangle. \quad (\text{B.23})$$

The well-known properties of the harmonic spherical functions show that

$$\begin{aligned} (\text{B.21}) + (\text{B.22}) &= \int_0^\infty \left(r^2 R_0'(r)^2 + \frac{r^2 R_1'(r)^2}{2} + \frac{r^4}{2} \left| \left(\frac{R_2(r)}{r} \right)' \right|^2 \right) dr \\ &\quad - \int_0^\infty 4r (2R_0(r)^2 + R_1(r)^2 + R_2(r)^2) dr. \end{aligned}$$

On the other hand, using Lemma B.1, the last integral (B.23) is equal to

$$\begin{aligned} (\text{B.23}) &= \iint_{\mathbb{R}^3 \times \mathbb{R}^3} |s_0(R_0)(x)|^2 |s_0(R_0)(y)|^2 + (2|s_0(R_0)(x)|^2 + |p_1(R_2)(x)|^2) |s_0(R_1)(y)|^2 dx dy \\ &\quad + \iint_{\mathbb{R}^3 \times \mathbb{R}^3} 2|s_0(R_0)(x)|^2 |p_1(R_2)(y)|^2 dx dy \\ &\quad - \iint_{\mathbb{R}^3 \times \mathbb{R}^3} (s_0(R_0)(x)s_0(R_0)(y) + \overline{p_1(R_2)(x)} p_1(R_2)(y)) s_0(R_1)(x)s_0(R_1)(y) dx dy \\ &\quad - \iint_{\mathbb{R}^3 \times \mathbb{R}^3} s_0(R_0)(x)s_0(R_0)(y) \overline{p_1(R_2)(x)} p_1(R_2)(y) dx dy. \end{aligned}$$

Using the well known formulae that can be found for instance in Slater's book [27] (Section 13-3 and Appendix 20a), this can be rewritten as

$$\begin{aligned} (\text{B.23}) &= \int_0^\infty \int_0^\infty \left(R_0(s)^2 R_0(t)^2 + (2R_0(s)^2 + R_2(s)^2) R_1(t)^2 + 2R_0(s)^2 R_2(t)^2 \right) \frac{s^2 t^2 ds dt}{\max\{s, t\}} \\ &\quad - \int_0^\infty \int_0^\infty R_0(s) R_0(t) R_1(s) R_1(t) \frac{s^2 t^2 ds dt}{\max\{s, t\}} \\ &\quad - \int_0^\infty \int_0^\infty (R_2(s)(R_2(t) (R_1(s)R_1(t) + R_0(s)R_0(t))) \frac{s^2 t^2 \min\{s, t\} ds dt}{3 \max^2\{s, t\}}. \end{aligned}$$

We then go on to calculate the expression of $\langle H\Psi, {}^3P_{sp}(R_0, R_1, \delta R) \rangle$, appearing in (2.22). As a corollary of Lemma B.1, of (A.5), and of (A.8), we obtain

$$\begin{aligned} & \left\langle {}^3P_{sp}(R_0, R_1, R_2), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{sp}(R_0, R_1, \delta R) \right\rangle \\ &= \left\langle {}^3P_{2,sp}(R_0, R_1, R_2), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{2,sp}(R_0, R_1, \delta R) \right\rangle \\ &= \iint_{\mathbb{R}^3 \times \mathbb{R}^3} (2|s_0(R_0)(x)|^2 + |s_0(R_1)(x)|^2) p_1(R_2)(y) p_1(\delta R)(y) dx dy \\ &\quad - \iint_{\mathbb{R}^3 \times \mathbb{R}^3} s_0(R_0)(x) p_1(\delta R)(x) p_1(R_2)(y) s_0(R_0)(y) dx dy \\ &\quad - \iint_{\mathbb{R}^3 \times \mathbb{R}^3} s_0(R_1)(x) p_1(\delta R)(x) p_1(R_2)(y) s_0(R_1)(y) dx dy. \end{aligned}$$

Using again the formulae in Slater's book [27] (Section 13-3 and Appendix 20a), we finally deduce that

$$\begin{aligned} & \left\langle {}^3P_{sp}(R_0, R_1, R_2), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{sp}(R_0, R_1, \delta R) \right\rangle \\ &= \int_0^\infty \int_0^\infty (2|R_0(s)|^2 + |R_1(s)|^2) R_2(t) \delta R(t) \frac{s^2 t^2}{\max\{s, t\}} ds dt \\ &\quad - \int_0^\infty \int_0^\infty R_0(s) \delta R(s) R_2(t) R_0(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt \\ &\quad - \int_0^\infty \int_0^\infty R_1(s) \delta R(s) R_2(t) R_1(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt. \end{aligned}$$

On the other hand, by (A.15), and using the same methods as above, we find

$$\begin{aligned} & \left\langle {}^3P_{pd}(R_0, R_2, R_4), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{sp}(R_0, R_1, \delta R) \right\rangle \\ &= \left\langle {}^3P_{2,pd}(R_0, R_2, R_4), \left(\sum_{1 \leq i < j \leq 4} \frac{1}{|x_i - x_j|} \right) {}^3P_{2,sp}(R_0, R_1, \delta R) \right\rangle \\ &= -\sqrt{2} \int_0^\infty \int_0^\infty R_1(s) R_2(s) \delta R(t) R_4(t) \frac{s^2 t^2 \min\{s, t\}}{3 \max^2\{s, t\}} ds dt. \end{aligned}$$

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