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# Statistical properties of the final state in one-dimensional ballistic aggregation

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We investigate the long time behavior of the one-dimensional ballistic aggregation model that represents a sticky gas of  $N$  particles with random initial positions and velocities, moving deterministically, and forming aggregates when they collide. We obtain a closed formula for the stationary measure of the system which allows us to analyze some remarkable features of the final “fan” state. In particular, we identify universal properties which are independent of the initial position and velocity distributions of the particles. We study cluster distributions and derive exact results for extreme value statistics (because of correlations these distributions do not belong to the Gumbel-Fréchet-Weibull universality classes). We also derive the energy distribution in the final state. This model generates dynamically many different scales and can be viewed as one of the simplest exactly solvable model of  $N$ -body dissipative dynamics.

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## I. INTRODUCTION

The ballistic aggregation model is one of the simplest interacting particles processes of nonequilibrium statistical mechanics and as such has attracted a lot of attention in the past decades [1]. This model represents a gas of  $N$  unit-mass particles forming clusters through perfectly inelastic adhesive collisions. The motion of a particle between two collisions is deterministic and free (i.e., ballistic) and the total mass and momentum in a collision are conserved whereas the kinetic energy is dissipated. The stochasticity in this model is due only to the initial configuration which consists of single particles randomly located with uncorrelated random velocities drawn from a continuous distribution. This dissipative system, usually referred to as ballistic aggregation or sticky gas, appears as a minimal model of cluster formation and provides a relevant statistical description of the merger of coherent structures such as vortices, thermal plumes, flowing granular media [2], or the accumulation of cosmic dust into planetoids. The ballistic aggregation model also plays a role in the study of the large-scale structure of the universe because the motion of self-gravitating matter in the expanding universe is similar to that of matter moving solely by inertia [3–5]. Another noteworthy feature of this model stems from its connection with the Burgers equation, an important toy model in the study of turbulence: it can be shown that at very high Reynolds number, the solution of the Burgers equation in the long time limit consists of a series of shocks which follow exactly the dynamics of the ballistic aggregation model [6–8].

Ballistic aggregation exhibits at long time  $t$  a self-similar coarsening behavior first studied by Carnevale, Pomeau, and Young [1]. In one dimension, the coarsening regime occurs for intermediate times satisfying  $t \sim N^{3/2}$ ,  $N$  being the total mass or, equivalently, the total number of initial particles in

the system. For longer times, and in the absence of any boundary, the system ultimately reaches a state where no more collisions are possible: the particles are grouped into clusters of different sizes (or masses) with velocities increasing from left to right. This final ordered state will be called the “fan” state. Once this state is reached, there is no further loss of energy and the system becomes stationary.

The aim of this paper is to investigate the properties of this fan state. We find an exact analytic formula for the joint probability distribution of the sizes and velocities of the clusters by mapping the final state of the ballistic aggregation model to the convex minorant of a one-dimensional random walk. We will show that from this invariant measure that fully characterizes the set of all possible fan states, various statistical properties of the model in the long time limit can be derived. For example, by partially tracing over the velocities we shall retrieve the known fact [11–13] that the probability of obtaining a fan state with exactly  $k$  clusters is a purely combinatorial factor (which is universal in the sense that it does not depend upon the initial velocity distribution of the particles). In addition, we also show that the cluster distribution in the fan state is identical to the statistics of cycles in a random permutation of  $N$  objects [14,15], both problems having a common underlying combinatorial structure related to the convex minorant construction [16–18]. This will allow us to characterize the different scales that are dynamically generated in the system: In the large  $N$  limit, the size of a typical cluster is of order  $N/\ln N$ , the largest cluster contains a finite fraction of the total mass  $N$  and hence grows linearly with  $N$ , the smallest cluster scales as  $\ln N$ , and the rightmost cluster, which has the leading velocity, has a size of order  $\sqrt{N}$ . We shall also derive the (nonuniversal) joint distribution of mass and velocity of this “leader” which is very different from the usual extreme-value distributions obtained for uncorrelated random variables. Finally, we shall

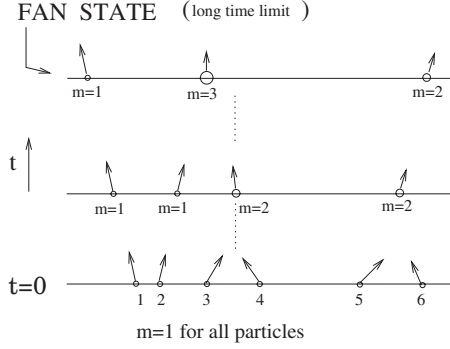


FIG. 1. An example of a configuration of six particles with identical initial mass  $m_i=1$  and varying initial velocities, evolving in the long time limit into the fan state: the velocities are now increasing from left to right and there are no more collisions.

show that the energy remaining in the system after all collisions have ended is of the order  $\ln N$  and shall calculate the distribution of the scaling variable  $E/\ln N$ .

The paper is organized as follows. In Sec. II we define the ballistic aggregation model and focus on its final fan state. In particular, we show how the statistical properties in the fan state are related to those of the convex minorant of an associated one-dimensional random walk problem. This mapping allows us to derive explicitly the joint distribution of the cluster sizes and their velocities in the fan state which is one of the key results of this paper. In Sec. III we show how this joint distribution can be used to derive the cluster size distribution in the fan state which turns out to be universal (with respect to the initial velocity distribution). This universal property is shown to be related to the statistics of cycles lengths in a random permutation problem. In Sec. IV, we further exploit the joint distribution to compute exact asymptotic results for certain extreme variables in the fan state such as the size and the velocity distribution of the right-most (“leader”) cluster. In Sec. V, we derive the exact asymptotic distribution of energy in the fan state. Finally we conclude in Sec. VI with a summary and some open problems.

## II. STATISTICAL DESCRIPTION OF THE STATIONARY STATE

### A. Definition of the model

The ballistic aggregation problem in one dimension is an elementary interactive particles process, the dynamics of which is defined as follows. At the initial time  $t=0$ , we start with  $N$  particles randomly located on an infinite line. We label the particles as  $i=1, 2, \dots, N$  according to the order of their initial positions from left to right (see Fig. 1 where we have taken  $N=6$ ). We denote by  $v_i$  and  $m_i$ , respectively, the initial velocity and the mass of the  $i$ th particle. For  $t>0$ , each particle moves ballistically at constant velocity. When two particles meet they coalesce and form a single new particle whose mass is the sum of the initial masses and whose momentum is the sum of the initial momenta. The only randomness in the model lies in the initial conditions, i.e., the

initial positions and velocities of each particle. We shall assume that the initial velocities  $\{v_i\}$ ’s are uncorrelated random variables, drawn from a common continuous probability density function (PDF)  $\phi(v)$ . For simplicity, we assume that all the initial masses are the same and take them to be unity. Once the initial conditions are given, the evolution of a given realization is deterministic and fully determined by the mass and momentum conservation laws. The shocks being totally inelastic, energy is dissipated at each collision.

At late times  $t \gg O(1)$  (but before the system feels the finiteness of the mass  $N$ ), the ballistic aggregation model exhibits a coarsening regime where typical cluster size grows with time as a power law, as studied first by Carnevale, Pomeau, and Young [1] using scaling arguments and numerical simulations. For example, in this regime, the average size (or mass) of a particle-cluster grows as  $t^{2D/(2+D)}$  and its velocity decreases as  $t^{-D/(2+D)}$ ,  $D$  being the spatial dimension. The global mass of the system being conserved, the total energy of the system decays as  $t^{-2D/(2+D)}$ . These exponents derived in Ref. [1] turn out to be correct in one dimension where the problem is exactly solvable [4,9,10]. This, however, turns out to be a bit fortuitous. In higher dimensions, the exponents predicted in Ref. [1] turn out to be incorrect due to strong correlations between the velocities of the colliding clusters at late times [19]. In one dimension, it is possible to calculate exactly the mass distribution of the clusters [4,9,10] in the scaling limit when  $N \rightarrow \infty$  ( $N$  being the total mass of the system) and  $t \rightarrow \infty$ , keeping the ratio  $t/N^{3/2}$  finite [4,9,10]. However, these scaling results for the coarsening regime are valid only for intermediate times satisfying  $t \sim N^{3/2}$ . When  $t \gg N^{3/2}$ , the system evolves into a stationary state in which no more collisions can occur: the particles are grouped in  $k$  disjoint clusters of different masses, where  $k$  itself is a stochastic variable. Each cluster moves at a constant velocity, and the velocity of a given cluster is larger than that of its left neighbor (if any) and less than that of its right neighbor (if any). In this ultimate state the clusters keep on moving farther apart, i.e., they fan out from each other with increasing time, thus justifying the name “fan” state (see Fig. 1). Once this fan state is reached, there is no further collision and hence no further dissipation of energy.

In this work, we shall focus on the statistical description of the fan state of the random aggregation process. We shall show that properties of the fan state related to the sizes of the clusters (regardless of their respective ordering and their velocities) are universal as they do not depend on the PDF  $\phi(v)$  from which the initial velocities of the particles are drawn. The cluster statistics can in fact be mapped to the cycle length distribution in random permutations, which are fundamental combinatorial objects. From this observation the distribution of the cluster masses and in particular the typical mass of the largest and the smallest cluster can readily be calculated. When the distribution of the velocities of the clusters is taken into account, universality with respect to  $\phi(v)$  is lost. However, in the large  $N$  limit, we will see that some universality is restored, thanks to the central limit theorem, in the size distribution of the rightmost cluster (the leader) provided the second moment  $\sigma^2 = \int_{-\infty}^{\infty} v^2 \phi(v) dv$  is finite. Finally, we are also able to calculate explicitly the energy distribution in the fan state, but only for the special case when  $\phi(v)$  is Gaussian.

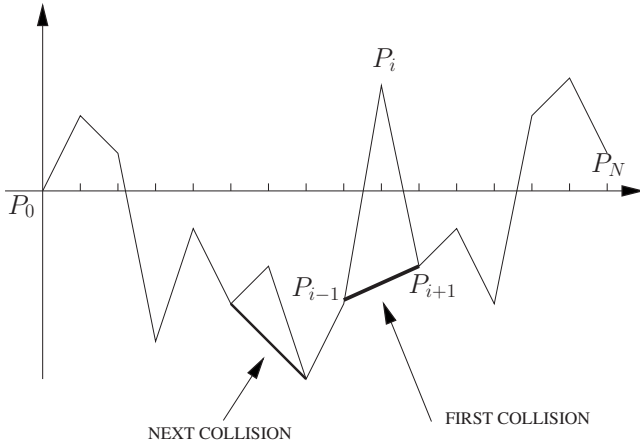


FIG. 2. The initial configuration is represented by a graph (thin line) joining  $P_0$  to  $P_N$  (we have taken  $m_j=1 \forall j$ ). This graph can be interpreted as a random walk in the cumulative momentum and cumulative mass space. After the first collision the particles  $i$  and  $i+1$  aggregate into one cluster: in the associated graph, the angle  $(P_{i-1}P_iP_{i+1})$  is replaced by the segment  $(P_{i-1}P_{i+1})$ .

### B. The fan state as a convex minorant

The fan state of the random aggregation process can be determined geometrically from the initial conditions [12,17,18]. We now explain this graphical interpretation of the fan state as a convex minorant of a random walk. Recall that we label the particles as  $i=1,2,\dots,N$  according to the order of their initial positions from left to right (see Fig. 1) and that we denote by  $v_i$ , respectively, the initial velocity of the  $i$ th particle (all the initial masses being unity). For any realization of the initial state  $\{v_1, v_2, \dots, v_N\}$ , the final state is uniquely determined, i.e., the number of final clusters, their masses and their velocities are unique.

The initial state  $\{v_1, v_2, \dots, v_N\}$  of the system is represented by a broken graph  $(P_0, P_1, \dots, P_N)$  (see Fig. 2) such that (i)  $P_0=(0,0)$  is the origin and (ii) the coordinates of  $P_i$  are given recursively by  $P_i=P_{i-1}+(1, v_i)$  for  $1 \leq i \leq N$ .

In other words, the velocity  $v_i$  of the  $i$ th particle is represented by the slope of the line  $(P_{i-1}P_i)$ . We also emphasize that the horizontal coordinate of  $P_i$  does not correspond to the actual position of the  $i$ th particle but only on its label.

More generally, if we had started with particles having different masses  $\{m_1, m_2, \dots, m_N\}$  the coordinates of  $P_i$  would be  $P_i=P_{i-1}+m_i(1, v_i)$  [12]. In other words, the initial state is drawn as a random walk in the cumulative momentum and cumulative mass space as shown in Figs. 2 and 3.

Suppose that the first collision occurs between particles  $i$  and  $(i+1)$  with velocities  $v_i$  and  $v_{i+1}$ , respectively. These two particles can aggregate if  $v_i > v_{i+1}$ , which means that the slope of the  $(P_{i-1}P_i)$  is larger than the slope of the  $(P_iP_{i+1})$ . Equivalently, this means that  $P_i$  is located above the segment  $(P_{i-1}P_{i+1})$ , i.e., locally, the graph  $(P_{i-1}P_iP_{i+1})$  has a negative curvature. After this collision, a cluster is formed with mass 2 and velocity  $(v_i+v_{i+1})/2$ . This cluster is represented by the vector  $(P_{i-1}P_{i+1})$  (see Fig. 2) with coordinates  $(2, v_i+v_{i+1})$ . We note that the slope of this vector again represents the velocity of the cluster. We now have a system of  $N-1$  “particles,” with  $N-2$  particles of mass 1 and one particle of

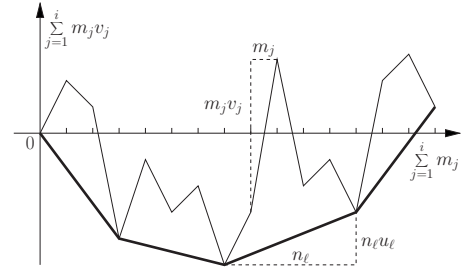


FIG. 3. We draw the convex minorant of the graph representing the same initial configuration as in Fig. 2. This convex minorant represents the final “fan” state in which no more collisions can occur. Each straight segment of the convex minorant corresponds to a cluster of  $n_i$  particles with momentum  $n_i u_i$ .

mass 2. The state of this system is represented by a broken line in which the angle  $(P_{i-1}P_iP_{i+1})$  with downward curvature is replaced by its base (see Fig. 2). Similarly, the next collision is also represented graphically by replacing another angle with downward curvature by its base (see Fig. 2). This process will continue iteratively and the particles will aggregate forming clusters until all angles with downward curvature have been eliminated, i.e., all collisions have occurred. It follows that for any given initial state, the final state is uniquely given by the convex minorant of the corresponding random walk. Each line segment of the convex minorant represents a cluster in the fan state—the horizontal and the vertical components of the segment give, respectively, the mass and the momentum of the cluster, i.e., the slope gives the velocity of the cluster. The geometry of the convex minorant (including the number of segments which represents the number of final clusters), however, differs from one realization of initial conditions to another. Various statistical properties of the ballistic aggregation in the fan state can be found from the statistical properties of the underlying convex minorant.

### C. The invariant measure in the fan state

We now derive the joint probability distribution which describes the fan state. Let  $p_N(k; \{n_i, u_i\})$  denote the joint PDF of having  $k$  clusters in the fan state of masses and velocities  $n_i$  and  $u_i$ , respectively, with  $i=1,2,\dots,k$ . Note that the various segments of the convex minorant (see Fig. 3) are correlated as their slopes must be increasingly ordered from left to right, i.e.,

$$u_1 < u_2 < \dots < u_k. \quad (1)$$

Moreover, the total mass is conserved, i.e.,

$$\sum_{i=1}^k n_i = N. \quad (2)$$

Once these two constraints are specified, the of each final cluster is essentially determined by the momentum conservation law

$$n_l u_l = \sum_{j=N_{l-1}+1}^{N_l} m_j v_j \quad (3)$$

for the  $l$ th cluster with  $l=1, 2, \dots, k$ . We have defined here the cumulated masses  $N_l = n_1 + \dots + n_l$ . The sum in the above equation is thus restricted to the initial velocities of those particles which form the  $l$ th cluster. We also recall that  $m_j = 1 \forall j$ .

However, most importantly, to form a specific (say  $l$ th) final cluster (i.e., a line segment in the underlying convex minorant) by aggregation of  $n_l$  consecutive initial particles, the portion of the underlying random walk of  $n_l$  steps must be such that it originates from one end of the line segment of convex minorant and finishes at its other end, while staying above the segment in the intermediate steps. Now, consider a random walk of  $n_l$  steps that starts from one end of a line segment and finishes at its other end, but is otherwise allowed to cross the segment in the intermediate steps. For any realization of such a walk, given that there is one unique minimum with respect to the segment, if we consider all the  $n_l$  cyclic permutations of the steps, then out of  $n_l$  different realizations of the walks there is one and only one arrangement where in the intermediate steps the walk stays above the segment. This property is known as Raney's lemma [20,21]. [Here, the uniqueness of the arrangement is guaranteed by the fact that the PDF  $\phi(v)$  is continuous.] Thus, the probability that a random walk of  $n_l$  steps that starts at one end of a line segment and finishes at its other end while staying above the segment in-between equals  $1/n_l$ , i.e.,

$$\left\langle \prod_{n=1}^{n_l-1} \theta\left(\frac{1}{n} \sum_{j=N_{l-1}+1}^{N_l+n} v_j - u_l\right) \right\rangle = \frac{1}{n_l}, \quad (4)$$

where we recall  $N_l = n_1 + \dots + n_l$ . Note that this probability is independent of the slope  $u_l$  of the segment as well as the PDF  $\phi(v)$  of the jump lengths.

Finally, gathering Eqs. (1)–(4) together we write

$$p_N(k; \{n_i, u_i\}) = \delta\left(N - \sum_{i=1}^k n_i\right) \prod_{i=1}^{k-1} \theta(u_{i+1} - u_i) \times \prod_{l=1}^k \frac{1}{n_l} \left\langle \delta\left(u_l - \frac{1}{n_l} \sum_{j=N_{l-1}+1}^{N_l} v_j\right) \right\rangle, \quad (5)$$

where the angle brackets  $\langle \dots \rangle$  denote the averaging over the set of initial velocities  $\{v_j\}$ . This key result provides a full statistical description of the clusters sizes and velocities in the fan state. From this joint probability distribution, all the statistical properties of the final state can be derived.

### III. UNIVERSAL STATISTICAL PROPERTIES OF CLUSTERS IN THE FAN STATE

In this section, we compute the distribution of the number and sizes of the clusters in the fan state: we will show that this distribution, obtained by integrating out the velocities of the clusters from the joint probability distribution given in Eq. (5), is universal with respect to the initial continuous PDF  $\phi(v)$ .

#### A. Statistics of the total number of clusters in the fan state

We first calculate  $p_N(k)$ , the probability that the final state contains  $k$  distinct clusters. The value of  $p_N(k)$  is obtained by summing Eq. (5) over all  $n_l$ 's and integrating it over  $u_l$ 's, keeping  $k$  fixed. Thus,

$$p_N(k) = \sum_{\{n_l \geq 1\}} \left[ \prod_{l=1}^k \frac{1}{n_l} \right] \delta\left(N - \sum_{l=1}^k n_l\right) \int_{-\infty}^{+\infty} \prod_{l=1}^{k-1} \theta(u_{l+1} - u_l) \times \left[ \prod_{l=1}^k du_l \mathcal{P}(n_l, u_l) \right]. \quad (6)$$

In this equation, we have used

$$\begin{aligned} \mathcal{P}(n_l, u_l) &\equiv \left\langle \delta\left(u_l - \frac{1}{n_l} \sum_{j=N_{l-1}+1}^{N_l} v_j\right) \right\rangle \\ &= \int_{-\infty}^{+\infty} \delta\left(u_l - \frac{1}{n_l} \sum_{j=N_{l-1}+1}^{N_l} v_j\right) \prod_{j=N_{l-1}+1}^{N_l} dv_j \phi(v_j), \end{aligned} \quad (7)$$

where  $\phi$  is the distribution of the initial velocities. We note that the function  $\mathcal{P}(n, u)$  is normalized:

$$\int_{-\infty}^{+\infty} \mathcal{P}(n, u) du = 1. \quad (8)$$

To disentangle the constraints in Eq. (6), it is useful to consider the following generating function:

$$\sum_{N=k}^{\infty} p_N(k) z^N = \int_{-\infty}^{+\infty} \left[ \prod_{l=1}^k du_l \rho_z(u_l) \right] \left[ \prod_{l=1}^{k-1} \theta(u_{l+1} - u_l) \right], \quad (9)$$

with

$$\rho_z(u_l) = \sum_{n_l=1}^{\infty} \frac{z^{n_l}}{n_l} \mathcal{P}(n_l, u_l). \quad (10)$$

The integral on the right-hand side (RHS) of Eq. (9) can be evaluated recursively and is found to be

$$\begin{aligned} &\int_{-\infty}^{+\infty} du_k \rho_z(u_k) \int_{-\infty}^{u_k} du_{k-1} \rho_z(u_{k-1}) \cdots \int_{-\infty}^{u_2} du_1 \rho_z(u_1) \\ &= \frac{1}{k!} \left( \int_{-\infty}^{+\infty} \rho_z(u) du \right)^k. \end{aligned} \quad (11)$$

Now using the definition of  $\rho_z(u)$  from Eq. (10) and the normalization property from Eq. (8) we get

$$\int_{-\infty}^{+\infty} \rho_z(u) du = \sum_{n=1}^{\infty} \frac{z^n}{n} = -\ln(1-z). \quad (12)$$

Thus, for any given value of  $k > 0$ , Eq. (9) yields



$$\sum_{N=k}^{\infty} p_N(k) z^N = \frac{1}{k!} [-\ln(1-z)]^k. \quad (13)$$

After extracting the coefficient of  $z^N$  on the RHS of this equation, we conclude that

$$p_N(k) = \frac{1}{k!} \sum_{\{n_l \geq 1\}} \left[ \prod_{l=1}^k \frac{1}{n_l} \right] \delta\left(N - \sum_{l=1}^k n_l\right). \quad (14)$$

This formula can be expressed in terms of some classical combinatorial numbers as follows. It is convenient to introduce a second auxiliary variable  $t$  and to calculate the generating function of Eq. (13) with respect to  $t$ :

$$\sum_{k=0}^{\infty} \sum_{N=k}^{\infty} p_N(k) z^N t^k = (1-z)^{-t}. \quad (15)$$

Considering the series

$$(1-z)^{-t} = 1 + tz + \frac{t(t+1)}{2!} z^2 + \frac{t(t+1)(t+2)}{3!} z^3 + \dots$$

and then identifying the coefficients of  $z^N$  on both sides of Eq. (15) we obtain the identity

$$\sum_{k=1}^N p_N(k) t^k = \frac{t(t+1)(t+2) \cdots (t+N-1)}{N!}. \quad (16)$$

Using the classical formula

$$t(t+1)(t+2) \cdots (t+N-1) = \sum_{k=1}^N S_1(N, k) t^k, \quad (17)$$

that defines the unsigned Stirling numbers of the first kind [22],  $S_1(N, k)$ , we conclude that the probability of finding  $k$  clusters in the fan state for a system starting with  $N$  distinct unit-mass particles is given by

$$p_N(k) = \frac{S_1(N, k)}{N!}, \quad \text{with } k = 1, 2, \dots, N. \quad (18)$$

We thus recover by a different method the result of Ref. [12].

From Eq. (17), we observe that  $\sum_{k=1}^N S_1(N, k) = N!$ . Thus,  $p_N(k)$  and, consequently, the joint PDF given by Eq. (5) are both clearly normalized as it must. Furthermore, taking the derivative with respect to  $t$  in Eq. (16) and putting  $t=1$ , one finds the mean number of clusters is given, exactly for all  $N$ , by

$$\langle k \rangle = 1 + \frac{1}{2} + \frac{1}{3} + \dots + \frac{1}{N}. \quad (19)$$

Clearly, for large  $N$ ,  $\langle k \rangle \approx \ln(N)$ . Using Eq. (16) one can also compute the average size of a cluster  $\langle M \rangle = \langle \frac{N}{k} \rangle$ . Indeed, dividing both sides of Eq. (16) by  $t$  and then integrating over  $t$  in  $[0, 1]$  one gets

$$\langle M \rangle = \left\langle \frac{N}{k} \right\rangle = \frac{N}{N!} \int_0^1 (t+1)(t+2) \cdots (t+N-1) dt. \quad (20)$$

The integral on the RHS can be easily estimated for large  $N$  by noting that the dominant contribution for large  $N$  arises from the vicinity of  $t=1$ . Setting  $t=1-\epsilon$ , one finds that for small  $\epsilon$ , the integrand can be approximated as  $\approx N! \exp[-\epsilon(1/2 + 1/3 + \dots + 1/N)]$ . Subsequently performing the integral, one gets the average cluster size for large  $N$

$$\langle M \rangle \approx \frac{N}{\ln(N)}. \quad (21)$$

In fact, it is well known [22] that the unsigned Stirling numbers  $S_1(N, k)$  enumerate the number of permutations of  $N$  elements with  $k$  disjoint cycles exactly: thus,  $p_N(k)$  is identical to the probability distribution of number of cycles in random permutations with uniform measure. Many results have been derived concerning the statistics of random permutations. For example, it has been shown in the context of permutation cycles [15] that for large  $N$ ,  $p_N(k)$  has a Gaussian form around its mean  $\langle k \rangle \sim \ln N$ , with a variance which is also of order  $\ln N$ . It follows that since  $k$  is peaked around its mean  $\ln(N)$  for large  $N$ , the average cluster size  $\langle M \rangle = \langle \frac{N}{k} \rangle \approx \frac{N}{\langle k \rangle} \approx N/\ln(N)$  which is perfectly consistent with the direct calculation of  $\langle M \rangle$  in Eq. (21). Using this together with the  $t^{2/3}$  asymptotic growth of cluster sizes before the fan state is reached, the characteristic time to reach this state is estimated to be  $t_c \sim (N/\ln N)^{3/2}$ .

## B. Distribution of clusters sizes including the biggest and the smallest

We now consider a more refined observable. It may be tempting to calculate the cluster size distribution, i.e., to integrate out the velocities from Eq. (5) and determine the marginal PDF  $p_N(k; \{n_i\})$ . Unfortunately, this marginal distribution is not universal: it depends on the choice of the initial PDF  $\phi(v)$  (this fact can be verified by working out explicit examples). However, if one considers the statistics of the clusters according to their sizes regardless of their spatial ordering then the result is again  $\phi(v)$  independent.

More precisely, let  $c_j$  for  $j=1, 2, \dots, N$  denote the number of clusters of size  $j$  in the fan state. By mass conservation, we have  $\sum_j j c_j = N$ . Then, it can be shown that the joint probability  $P_N(c_1, c_2, \dots, c_N) = \text{Prob}\{c_1, c_2, \dots, c_N\}$  is given by

$$P_N(c_1, c_2, \dots, c_N) = \delta\left(N - \sum_{j=1}^N j c_j\right) \left[ \prod_{j=1}^N \frac{1}{j^{c_j} c_j!} \right]. \quad (22)$$

The derivation of this formula from the invariant measure in the fan state, Eq. (5), is given in Appendix A. This expression is identical to the distribution of cycles in random permutations where  $c_j$ 's are to be identified with the number of cycles of length  $j$  [22].

The generating function associated with this probability distribution is defined as follows:

$$G_N(t_1, \dots, t_N) = \sum_{c_1, \dots, c_N} P_N(c_1, c_2, \dots, c_N) t_1^{c_1}, \dots, t_N^{c_N}. \quad (23)$$

This function  $G_N$  can itself be embedded in a “grand-canonical” generating function given by

$$\tilde{G}(z, \{t_k\}) = \sum_N G_N(t_1, \dots, t_N) z^N = \exp\left(\sum_{k=1}^{\infty} \frac{z^k t_k}{k}\right), \quad (24)$$

where the last equality is obtained by using Eq. (22). This expression embodies all the necessary information and will be useful for further calculations.

For example, we can study the extreme cluster sizes. Let  $Q_N(L)$  be the probability that the largest cluster size (mass)  $n_{\max} \leq L$  and  $R_N(S)$  the probability that the smallest cluster size  $n_{\min} \geq S$ . It follows from these definitions that the probability that the largest cluster is of size  $L$  is given by

$$\text{Prob}(n_{\max} = L) = Q_N(L) - Q_N(L-1) \quad (25)$$

and, similarly,

$$\text{Prob}(n_{\min} = S) = R_N(S) - Q_N(S+1). \quad (26)$$

Using Eq. (22) we find

$$\begin{aligned} Q_N(L) &= \sum_{c_1, \dots, c_L} \text{Prob}\{c_1, c_2, \dots, c_L, 0, \dots, 0\} \\ &= G_N(t_1 = 1, \dots, t_L = 1, 0, \dots, 0) \end{aligned} \quad (27)$$

and, similarly,

$$\begin{aligned} R_N(S) &= \sum_{c_S, \dots, c_N} \text{Prob}\{0, \dots, 0, c_S, \dots, c_N\} \\ &= G_N(0, \dots, 0, t_S = 1, \dots, t_N = 1). \end{aligned} \quad (28)$$

Using the grand-canonical generating function defined in Eq. (24), we readily obtain

$$\tilde{Q}_L(z) = \sum_N Q_N(L) z^N = \exp\left(\sum_{n=1}^L \frac{z^n}{n}\right)$$

and

$$\tilde{R}_L(z) = \sum_N R_N(S) z^N = \exp\left(\sum_{n=S}^{\infty} \frac{z^n}{n}\right). \quad (29)$$

The mean sizes of the largest and the smallest clusters in the fan state containing  $N$  particles are given by

$$\langle n_{\max} \rangle_N = \sum_L L [Q_N(L) - Q_N(L-1)] = \sum_L [1 - Q_N(L)] \quad (30)$$

and

$$\langle n_{\min} \rangle_N = \sum_S S [R_N(S) - R_N(S+1)] = \sum_S R_N(S). \quad (31)$$

For large  $N$ , the asymptotic behavior of  $\langle n_{\max} \rangle_N$  and  $\langle n_{\min} \rangle_N$  can be obtained by analyzing the associated generating functions in the  $z \rightarrow 1$  limit. For example, using Eqs. (29) we obtain

$$\begin{aligned} \sum_N \langle n_{\max} \rangle_N z^N &= \sum_L \left\{ \frac{1}{1-z} - \exp\left(\sum_{n=1}^L \frac{z^n}{n}\right) \right\} \\ &= \frac{1}{1-z} \sum_L \left\{ 1 - \exp\left(-\sum_{n=L+1}^{\infty} \frac{z^n}{n}\right) \right\}. \end{aligned} \quad (32)$$

When  $z \rightarrow 1$ , the sums on the right-hand sides can be transformed into integrals and the desired asymptotics are obtained by extracting the leading singularity at  $z=1$ . Thus, for large  $N$ , we obtain

$$\begin{aligned} \langle n_{\max} \rangle &\sim \lambda N \text{ with } \lambda = \int_0^{\infty} \exp\left(-x - \int_x^{\infty} \frac{e^{-y}}{y} dy\right) dx \\ &= 0.6243299885 \dots \end{aligned} \quad (33)$$

This constant is known as the Golomb-Dickman constant [23]. Similarly, we find,  $\langle n_{\min} \rangle \sim e^{-C} \ln N$ , where  $C = 0.57721566 \dots$  is Euler's constant, and we again note that,  $\langle n_{\min} \rangle$  and  $\langle n_{\max} \rangle$  are identical to the mean lengths of the shortest and the longest cycles, respectively, of random permutations and were already calculated in this context [15,24]. Recently, the same constants also appeared in the context of the statistics of records of independent and identically distributed random variables [25,26]. In fact, the statistics of the lengths of cycles in random permutation is known to be related to the statistics of intervals of time between successive records [27,28]. Thus, the longest (shortest) cycle of random permutations has the same statistics as the longest (shortest) interval between two successive records.

#### IV. LEADER STATISTICS

In this section we compute the statistics of the size and the velocity of the “right-most cluster” in the fan state. This right-most cluster will be referred to as the “leader” since it moves with the highest velocity. Let  $L_N(n, v)$  denote the joint PDF of the leader's size  $n$  and velocity  $v$  in the fan state. To obtain this PDF, we start from the basic microscopic joint PDF of all cluster sizes and their velocities in Eq. (5) and integrate out the sizes and velocities of all clusters except the leader. As usual, it is convenient to consider the generating function which then reads

$$\begin{aligned} \sum_N L_N(n, v) z^N &= \frac{z^n}{n} \mathcal{P}(n, v) \sum_{k=1}^{\infty} \int \left[ \prod_{i=1}^{k-1} du_i \rho_z(u_i) \right] \theta(v - u_{k-1}) \\ &\quad \times \left[ \prod_{i=1}^{k-2} \theta(u_{i+1} - u_i) \right], \end{aligned} \quad (34)$$

where  $\rho_z(u)$  and  $\mathcal{P}(n, v)$  are defined, respectively, in Eqs. (10) and (7). The  $(k-1)$ -fold integral over the  $u_i$ 's can be recursively evaluated as in Eq. (11) and we get

$$\int \left[ \prod_{i=1}^{k-1} du_i \rho_z(u_i) \right] \theta(v - u_{k-1}) \left[ \prod_{i=1}^{k-2} \theta(u_{i+1} - u_i) \right] \\ = \frac{1}{(k-1)!} \left( \int_{-\infty}^v \rho_z(u) du \right)^{k-1}. \quad (35)$$

Substituting this in Eq. (34) and summing over  $k$  we get a somewhat compact expression

$$\sum_N L_N(n, v) z^N = \frac{z^n}{n} \mathcal{P}(n, v) \exp \left[ \int_{-\infty}^v \rho_z(u) du \right]. \quad (36)$$

For the subsequent asymptotic analysis, it turns out to be convenient to rewrite the expression in Eq. (36) in a slightly different form. By putting  $k=1$  in Eq. (11) one gets  $\int_{-\infty}^v \rho_z(u) du = -\ln(1-z)$  implying

$$\int_{-\infty}^v \rho_z(u) du = -\ln(1-z) - \int_v^{\infty} \rho_z(u) du. \quad (37)$$

Using this in Eq. (36) gives

$$\sum_N L_N(n, v) z^N = \frac{z^n}{n(1-z)} \mathcal{P}(n, v) \exp \left[ - \int_v^{\infty} \rho_z(u) du \right]. \quad (38)$$

From the joint PDF  $L_N(n, v)$  of the leader's size and velocity in Eq. (38) one can then obtain the marginals  $C_N(n) = \int_{-\infty}^{\infty} L_N(n, v) dv$  for the size PDF and  $D_N(v) = \sum_{n=1}^{\infty} L_N(n, v)$  for the velocity PDF of the leader. It is evident from Eq. (38) that  $L_N(n, v)$ , for any finite  $N$ , depends explicitly on the initial velocity distribution  $\phi(v)$  since both  $\mathcal{P}(n, v)$  as well as  $\rho_z(u)$  depends on  $\phi(v)$ . This is unlike the distribution of the cluster sizes as derived in Sec. III which is completely universal, i.e., independent of  $\phi(v)$  for any finite  $N$ . So, a natural question is "to what extent are the leader size and velocity distributions universal, i.e., independent of the details of  $\phi(v)$ ?" We will see below that the universal property holds for the leader's size distribution, but only in the large  $N$  limit and for continuous and symmetric  $\phi(v)$  with a finite variance  $\sigma^2 = \int_{-\infty}^{\infty} v^2 \phi(v) dv$ . In that case, for large  $N$ , one can use the central limit theorem and the limiting marginal size distribution  $C_N(n)$  essentially becomes universal, though the marginal velocity distribution  $D_N(v)$  still remains nonuniversal even in the large  $N$  limit. A natural question is how these results get modified when  $\phi(v)$  is such that its variance  $\sigma^2$  is infinite. A particular example of this class is the Cauchy distribution for  $\phi(v)$ . We will show that the Cauchy case is exactly solvable and one can derive the size and the velocity distribution of the leader explicitly. In Secs. IV A and IV B below, we consider, for finite  $\sigma^2$ , the limiting size and the velocity distribution of the leader respectively. In Sec. IV C, we derive the explicit distributions for the Cauchy case.

#### A. The limiting leader size distribution

In this subsection we compute the limiting size distribution  $C_N(n)$  of the leader for large  $N$ . Integrating Eq. (38) over  $v$  and writing  $z=e^{-s}$  we get

$$\sum_N C_N(n) e^{-sN} = \frac{e^{-sn}}{n(1-e^{-s})} \int_{-\infty}^{\infty} dv \mathcal{P}(n, v) \exp[-Y(s, v)], \quad (39)$$

where

$$Y(s, v) = \sum_{k=1}^{\infty} \frac{e^{-sk}}{k} \int_v^{\infty} du \mathcal{P}(k, u) \quad (40)$$

with  $\mathcal{P}(k, u)$  defined in Eq. (7). Now, for large  $N$ , we need to investigate the behavior of the RHS of Eq. (39) in the limit  $s \rightarrow 0$ . It is evident that to obtain a sensible limit of the RHS in Eq. (39) one needs to consider the scaling limit  $s \rightarrow 0$  but  $n \rightarrow \infty$  keeping the product  $sn$  finite. In this scaling limit, it is not difficult to see that the dominant contribution to the sum  $Y(s, v)$  in Eq. (40) comes from those terms where the product  $sk$  is finite in the limit  $s \rightarrow 0$ , i.e., terms where  $k \sim 1/s$  is large.

Now, for large  $k$ , it follows from the definition in Eq. (7) that for a symmetric  $\phi(v)$  with zero mean and a finite variance  $\sigma^2 = \int_{-\infty}^{\infty} v^2 \phi(v) dv$  is finite, one can invoke the central limit theorem to assert that  $\mathcal{P}(k, u)$  has a Gaussian distribution, i.e.,

$$\mathcal{P}(k, u) \approx \sqrt{\frac{k}{2\pi\sigma^2}} \exp \left[ -\frac{ku^2}{2\sigma^2} \right]. \quad (41)$$

Substituting this result in Eq. (40) and replacing the resulting sum by an integral in the small  $s$  limit, we find that an appropriate scaling limit of  $Y(s, v)$  exists when  $s \rightarrow 0$ ,  $v \rightarrow 0$  keeping the ratio  $v/\sqrt{s}$  fixed. In this scaling limit, we get

$$Y(s, v) \approx \frac{1}{2} \int_s^{\infty} \frac{dy}{y} e^{-y} \operatorname{erfc} \left( \frac{v\sqrt{y}}{\sigma\sqrt{2s}} \right). \quad (42)$$

The integral on the RHS of Eq. (42) can be done explicitly and one gets to leading order in small  $s$

$$Y(s, v) \approx -\frac{1}{2} [\mathbf{C} + \ln(s)] - \ln[w + \sqrt{w^2 + 1}], \quad (43)$$

where  $\mathbf{C}=0.57721566\cdots$  is Euler's constant and  $w = \frac{v}{\sigma\sqrt{2s}}$  is the scaling variable. Finally we substitute this expression for  $Y(s, v)$  and the limiting Gaussian form of  $\mathcal{P}(n, v)$  from Eq. (41) onto the RHS of Eq. (39), perform the subsequent integration, and finally arrive at the following expression in the scaling limit  $s \rightarrow 0$ ,  $n \rightarrow \infty$  but keeping the product  $ns$  fixed:

$$\sum_N C_N(n) e^{-sN} \approx \frac{1}{\sqrt{n}} G(ns), \quad (44)$$

where the scaling function  $G(x)$  is given by

$$G(x) = \frac{2b}{\sqrt{\pi}} \frac{e^{-x}}{\sqrt{x}} \int_0^{\infty} dy e^{-y^2} \sqrt{1 + \frac{y^2}{x}} \quad (45)$$

and

$$b = e^{C/2} = 1.33456\cdots. \quad (46)$$

We notice one remarkable fact: even though we needed to have  $\sigma^2$  finite in arriving at the result in Eq. (45),  $\sigma^2$  has



dropped out of the final expression in Eq. (45). Thus for any finite  $\sigma^2$ , the scaling function  $G(x)$  is universal and is actually independent of the value of  $\sigma^2$ .

From the expression in Eq. (45) it is easy to compute the moments of the leader size distribution for large  $N$ . For example, to compute the  $k$ th moment  $\langle n^k \rangle_N$  for large  $N$ , we multiply both sides of Eq. (45) by  $n^k$  and sum over all  $n$ . This gives

$$\sum_N \langle n^k \rangle_N e^{-sN} \approx \frac{2b}{\sqrt{\pi s}} \sum_n e^{-sn} n^{k-1} \int_0^\infty dy e^{-y^2} \sqrt{1+y^2/ns}. \quad (47)$$

We next make a change of variable  $y = \sqrt{sn}u$  in the integral and carry out the sum over  $n$ . In the scaling limit, this sum can be replaced by an integral which can be explicitly performed and we get for small  $s$

$$\begin{aligned} \sum_N \langle n^k \rangle_N e^{-sN} &\approx \frac{2b\Gamma(k+1/2)}{\sqrt{\pi s^{k+1/2}}} \int_0^\infty \frac{du}{(1+u^2)^k} \\ &= b \frac{\Gamma(k-1/2)\Gamma(k+1/2)}{\Gamma(k)} s^{-(k+1/2)}. \end{aligned} \quad (48)$$

Note that this expression is strictly valid for  $k \geq 1$ . For  $k=0$ , one can show independently that  $\sum_N C_N(n) e^{-sN} \rightarrow 1/s$  as  $s \rightarrow 0$  as expected since the PDF  $C_N(n)$  is normalized to unity. Inverting the Laplace transform in Eq. (48), it then follows that for large  $N$ , the  $k$ th moment of the leader size behaves for  $k \geq 1$ ,

$$\langle n^k \rangle_N \approx A_k N^{k-1/2}, \text{ where } A_k = \frac{b\Gamma(k-1/2)}{\Gamma(k)}. \quad (49)$$

In particular the average leader size ( $k=1$ ) behaves for large  $N$  as

$$\langle n \rangle_N \approx b\sqrt{\pi} N^{1/2} = (2.63533 \dots) N^{1/2}. \quad (50)$$

From the expression for the moments in Eq. (48), it follows that for large  $N$ , the leader size distribution  $C_N(n)$  has the following scaling form:

$$C_N(n) \approx \frac{1}{N^{3/2}} W\left(\frac{n}{N}\right), \quad (51)$$

where the scaling function  $W(x)$ , for  $0 \leq x \leq 1$  is such that, for all  $k \geq 1$ ,

$$\int_0^1 W(x) x^k dx = b \frac{\Gamma(k-1/2)}{\Gamma(k)}. \quad (52)$$

Finally, one can analytically continue the Eq. (52) for non-integer  $k$  and then invert the equation to obtain the following explicit expression of  $W(x)$  valid for  $0 \leq x \leq 1$ :

$$W(x) = \frac{b}{\sqrt{\pi}} x^{-3/2} (1-x)^{-1/2}. \quad (53)$$

It is easy to verify that for all  $k \geq 1$ , the above expression for  $W(x)$  does indeed satisfy Eq. (52). Note that the scaling in Eq. (51) is valid only for  $1 \leq n \sim N$ . In particular it does not hold when  $n \sim O(1)$ . This is also evident from the explicit

form of the scaling function  $W(x)$  in Eq. (53) which diverges as  $x^{-3/2}$  as  $x \rightarrow 0$ , making it non-normalizable. Indeed, for small  $n$ , one has to keep the full expression of the generating function to recover the correct normalization. Thus the main result of this subsection is the universal scaling form of the leader size distribution in Eqs. (51)–(53) valid for any continuous and symmetric  $\phi(v)$  with a finite  $\sigma^2$  and remarkably it is independent of the actual value of the  $\sigma^2$ .

## B. The limiting leader velocity distribution

The marginal velocity PDF of the leader  $D_N(v)$  can be obtained by summing Eq. (38) over  $n$ . This gives

$$\sum_{N=1}^\infty D_N(v) z^N = \frac{1}{1-z} \rho_z(v) \exp\left[-\int_v^\infty \rho_z(u) du\right]. \quad (54)$$

A somewhat more convenient quantity is the cumulative velocity distribution of the leader  $F_N(v) = \int_{-\infty}^v D_N(v') dv'$ . One can easily derive its generating function by integrating Eq. (54)

$$\sum_{N=1}^\infty F_N(v) z^N = \frac{\exp[-\int_v^\infty \rho_z(u) du] - (1-z)}{1-z}. \quad (55)$$

Note that as  $v \rightarrow \infty$ , the RHS of Eq. (55) becomes  $z/(1-z)$ . This is consistent with the LHS of Eq. (55) since  $F_N(v) \rightarrow 1$  as  $v \rightarrow \infty$  and thus the generating function in that limit is precisely  $z/(1-z)$ .

The result for the cumulative velocity distribution of the leader in Eq. (55) is exact for all  $N$ . We next address the question of the limiting leader velocity distribution as  $N \rightarrow \infty$ . For large  $N$ , one needs to investigate the RHS of Eq. (55) in the limit  $z \rightarrow 1$ . In that limit, the RHS scales as  $\exp[-\int_v^\infty \rho_1(u) du]/(1-z)$ , provided the numerator is finite. In such cases, it follows from Eq. (55) that  $F_N(v)$  tends to a limiting  $N$ -independent distribution  $F_\infty(v)$  given by

$$F_\infty(v) = \exp\left[-\int_v^\infty \rho_1(u) du\right] = \exp\left[-\sum_{n=1}^\infty \frac{1}{n} \int_v^\infty \mathcal{P}(n, u) du\right]. \quad (56)$$

Note that for some  $\phi(v)$  this limiting distribution may not exist. In fact, we will see later that when  $\phi(v)$  has the Cauchy distribution, which is symmetric and continuous, there is no  $N$ -independent limiting distribution. It is evident from Eq. (56) that, unlike the limiting leader size distribution, the limiting leader velocity distribution, whenever it exists, is highly nonuniversal and depends explicitly on the initial velocity distribution  $\phi(v)$ .

To see how this limiting distribution in Eq. (56) looks, we first consider the Gaussian distribution  $\phi(v) = \exp(-v^2/2\sigma^2)/\sqrt{2\pi\sigma^2}$ , which is physically most relevant since it represents the thermal Boltzmann distribution. In this case,  $\mathcal{P}(n, u) = \sqrt{\frac{n}{2\pi\sigma^2}} \exp[-\frac{nu^2}{2\sigma^2}]$  exactly for all  $u$ . Substituting on the RHS of Eq. (56) and performing the integral one can express the limiting distribution  $F_\infty(v)$  as a function of the dimensionless variable  $v/\sigma$ ,

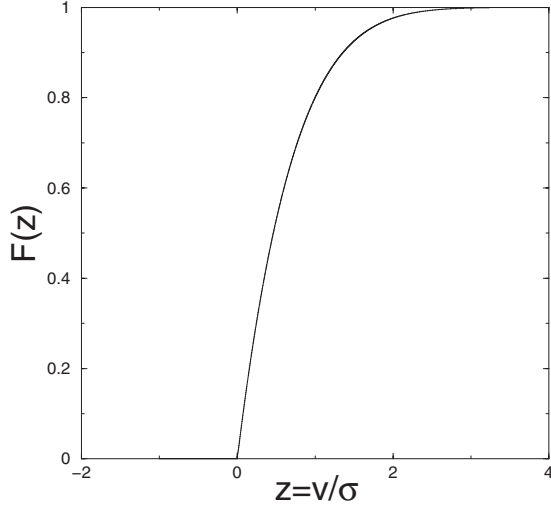


FIG. 4. The limiting cumulative velocity distribution  $F_\infty(v) = F(z=v/\sigma)$  plotted against  $z$ , for the Gaussian distribution  $\phi(v) = \exp(-v^2/2\sigma^2)/\sqrt{2\pi\sigma^2}$ .

$$F_\infty(v) = F(v/\sigma), \text{ where } F(z) = \exp\left[-\frac{1}{2}\sum_{n=1}^{\infty}\frac{1}{n}\operatorname{erfc}\left(\sqrt{\frac{n}{2}}z\right)\right]. \quad (57)$$

While we were unable to perform the sum exactly in Eq. (57), the function  $F(z)$  can be easily plotted using MATHEMATICA and is shown in Fig. 4. It is easy to show from Eq. (57) that the function  $F(z)$  has finite nonzero support only for positive  $z$ . It is identically zero for all  $z \leq 0$ . Furthermore, the asymptotics of  $F(z)$  as  $z \rightarrow 0$  and  $z \rightarrow \infty$  can also be derived. We find

$$F(z) \approx \sqrt{2}z \text{ as } z \rightarrow 0 \quad (58)$$

$$\approx 1 - \frac{1}{\sqrt{2\pi}z} e^{-z^2/2} \text{ as } z \rightarrow \infty. \quad (59)$$

The asymptotic behavior as  $z \rightarrow \infty$  is easy to derive, as in this limit the dominant contribution comes from the  $n=1$  term of the sum in Eq. (57). In contrast, the other limit  $z \rightarrow 0$  in Eq. (58) is more tricky to derive. We provide a derivation in Appendix B. Let us make a couple of interesting observations on the general formula for the limiting velocity distribution in Eq. (56) when it exists.

(i) For any continuous and symmetric initial velocity distribution  $\phi(v)$  for which the limiting velocity distribution  $F_\infty(v)$  exists,  $F_\infty(v)=0$  for any  $v \leq 0$ . This follows from the fact for  $v=0$ , by symmetry,  $\int_0^\infty \mathcal{P}(n,u)du = 1/2$ . Thus, the sum in Eq. (56) diverges for  $v=0$  and  $F_\infty(0)=0$ . Since  $F_\infty(v)$  is a cumulative distribution, it is a nondecreasing function of  $v$ . Hence it follows that if  $F_\infty(v=0)=0$  then  $F_\infty(v)=0$  for all  $v \leq 0$ . The implication of this result is rather interesting. It implies that even though initial velocity distribution may be symmetric in  $v$  and there may be a lot of particles initially with a negative velocity, the eventual velocity of the leader, i.e., the right-most cluster is always positive.

(ii) Note that the velocity of the leader is also the maximum of the final velocities. Now if the collisions were elastic, the particles would have merely interchanged the velocities in each collision (this is called the Jepsen gas), and in the final state the velocity of the leader would have been the maximum of all the initial velocities [30]. Therefore, in this case the distribution of the leader velocity would have been given by the usual extreme value statistics of uncorrelated random variables [31]. The velocity of the final leader of elastic Jepsen gas scales as  $u \approx a_N w + b_N$  for large  $N$ , and the limiting distribution of the scaled velocity  $w$  has only three possible forms depending on the tail of  $\phi(v)$  [30]. However, the scaling parameters  $a_N$  and  $b_N$  depends on  $N$  as well as on the complete form of  $\phi(v)$ . For example, when  $\phi(v)$  is Gaussian,  $a_N = (2 \ln N)^{-1/2}$  and  $b_N = (2 \ln N)^{1/2} - (2 \ln N)^{-1/2}(\ln \ln N + \ln 4\pi)/2$ . The limiting velocity distribution is Gumbel in this case. Thus, unlike the elastic gas where the leader velocity increases with  $N$ , it follows from Eq. (56) (when the limiting distribution exists) that in the sticky gas, the leader velocity remains of  $O(1)$  in the large  $N$  limit. This is due to the strong correlations generated between the velocities of different clusters in the fan state. Thus Eq. (56) provides an exactly solvable case of extreme value statistics of correlated random variables.

### C. Exact leader statistics for the Cauchy distribution

In this subsection, we consider the special case of the symmetric Cauchy distribution for the initial velocity

$$\phi(v) = \frac{1}{\pi} \frac{a}{v^2 + a^2}. \quad (60)$$

For this distribution, the mean is zero, but the second moment  $\sigma^2$  diverges. So, the results of the previous two subsections, where it was assumed  $\sigma^2$  is finite, do not hold. However, the principal result in Eq. (38) is still valid and we show here that the Cauchy distribution preents an exactly solvable case in the sense that the RHS of Eq. (38) can be exactly evaluated in closed form.

The first simplification occurs when one calculates  $\mathcal{P}(n,u)$  defined in Eq. (7). Since the Cauchy distribution is stable, it turns out (as can be easily proved) that

$$\mathcal{P}(n,u) = \phi(u) = \frac{1}{\pi} \frac{a}{u^2 + a^2}. \quad (61)$$

Thus  $\mathcal{P}(n,u)$  is completely independent of  $n$ . This particular feature of the Cauchy distribution was also used recently to obtain exact results for the statistics of records in a sequence of random walk in presence of a drift [29]. Using this independence on  $n$ , it then follows from the definition in Eq. (10) that

$$\rho_z(u) = -\frac{1}{\pi} \frac{a}{u^2 + a^2} \ln(1-z). \quad (62)$$

One can then trivially do the integral

$$\int_v^\infty \rho_z(u) du = - \left[ \frac{1}{2} - \frac{1}{\pi} \arctan(v/a) \right] \ln(1-z). \quad (63)$$

Substituting these results in Eq. (38) gives the following exact expression:

$$\sum_{N=1}^\infty L_N(n,v) z^N = \frac{1}{\pi} \frac{a}{v^2 + a^2} \frac{z^n}{n} (1-z)^{-\mu(v)}, \quad \text{where } \mu(v) = \frac{1}{2} + \frac{1}{\pi} \arctan(v/a). \quad (64)$$

Comparing coefficient of  $z^N$  gives us the exact joint distribution of size and the velocity of the leader

$$L_N(n,v) = \frac{1}{\pi} \frac{a}{v^2 + a^2} \frac{\Gamma(\mu(v) + N - n)}{n \Gamma(N - n + 1) \Gamma[\mu(v)]}. \quad (65)$$

*Marginal size distribution.* To compute the marginal size distribution  $C_N(n) = \int_{-\infty}^\infty L_N(n,v) dv$  of the leader, it is convenient to do the integration over  $v$  in Eq. (64). This gives the exact generating function

$$\sum_{N=1}^\infty C_N(n) z^N = - \frac{z^{n+1}}{n(1-z) \ln(1-z)}. \quad (66)$$

From this generating function, it is easy to calculate the moments  $\langle n^k \rangle_N$  of the leader size for large  $N$  by investing the singularity of the RHS in Eq. (66) near  $z=1$ . Skipping the details, we find that for large  $N$  and for  $k \geq 1$

$$\langle n^k \rangle_N \approx \frac{1}{k} \frac{N^k}{\ln(N)}. \quad (67)$$

This result should be compared to the case in Eq. (49) where  $\sigma^2$  is finite. In particular, the average leader size grows as  $\langle n \rangle_N \sim N / \ln N$ , much faster than the  $N^{1/2}$  growth for the case where  $\sigma^2$  is finite.

*Marginal velocity distribution.* Similarly, the marginal velocity PDF of the leader  $D_N(v) = \sum_n L_N(n,v)$  can be computed conveniently directly from Eq. (64). We get

$$\sum_{N=1}^\infty D_N(v) z^N = \frac{1}{\pi} \frac{a}{v^2 + a^2} [-\ln(1-z)] (1-z)^{-\mu(v)}, \quad (68)$$

from which one can compute the generating function for the cumulative velocity distribution  $F_N(v) = \int_{-\infty}^v D_N(v') dv'$ . We get a nice compact expression

$$\sum_{N=1}^\infty F_N(v) z^N = (1-z)^{-\mu(v)} - 1. \quad (69)$$

Comparing powers of  $z^N$  gives a very simple but nontrivial distribution valid for all  $N \geq 1$

$$F_N(v) = \frac{\Gamma(\mu(v) + N)}{\Gamma(N+1)}, \quad \text{where } \mu(v) = \frac{1}{2} + \frac{1}{\pi} \arctan(v/a). \quad (70)$$

Note that when  $v \rightarrow \infty$ ,  $\mu(v) \rightarrow 1$  and when  $v \rightarrow -\infty$ ,  $\mu(v) \rightarrow 0$ . Thus, in the limit  $v \rightarrow \infty$ ,  $F_N(v) \rightarrow 1$  and when  $v \rightarrow -\infty$ ,

$F_N(v) \rightarrow 0$  as expected. Also, for  $N=1$ , we get  $F_1(v) = \mu(v)$  which is just the cumulative Cauchy distribution. This is expected, because the velocity of a single particle remains unchanged as there is no collision.

We note from this exact velocity distribution in Eq. (70) that the distribution is  $N$  dependent explicitly and does not have any nontrivial  $N$ -independent limiting form as  $N \rightarrow \infty$ , unlike in the previous subsection for the Gaussian case. In fact, by analyzing the singularity near  $z=1$  in Eq. (69), we see that for any fixed  $v$ , as  $N \rightarrow \infty$ ,  $F_N(v)$  decays with  $N$  as a power law with an exponent that depends continuously on  $v$

$$F_N(v) \approx \frac{1}{\Gamma[\mu(v)]} \frac{1}{N^{1-\mu(v)}}. \quad (71)$$

Since the exponent  $1-\mu(v) = 1/2 - \arctan(v/a)/\pi$  decreases monotonically from 1 (as  $v \rightarrow -\infty$ ) to zero (as  $v \rightarrow \infty$ ), it follows that the distribution for smaller values of  $v$  tends to zero faster than the distribution for higher  $v$ . Thus the marginal leader velocity distribution in the Cauchy case is very different from that of its Gaussian counterpart.

## V. DISTRIBUTION OF THE TOTAL ENERGY IN THE FAN STATE

In this section we compute the distribution of the total energy of the clusters in the fan state. For simplicity, we restrict ourselves to the initial Gaussian distribution of the velocities  $\phi(v) = \exp(-v^2/2)/\sqrt{2\pi}$ , where we have also set  $\sigma^2=1$  without any loss of generality. Thus, initially, the total (kinetic) energy of the system,  $E_0 = \sum_{i=1}^N v_i^2/2$ , is distributed according to the PDF

$$P_N(E_0) = \frac{E_0^{N/2-1} e^{-E_0}}{\Gamma(N/2)}. \quad (72)$$

Now, as the system evolves, it dissipates energy through collisions, and once the fan state is reached there is no more energy loss. It is natural to ask how the PDF of the energy evolves with time from its initial form in Eq. (72), and in particular how does this energy PDF look in the final fan state beyond which there is no more dissipation.

Let  $Q_N(E)$  denote the PDF of the total energy in the fan state, i.e.,

$$Q_N(E) = \left[ \delta \left( E - \frac{1}{2} \sum_{i=1}^k n_i u_i^2 \right) \right], \quad (73)$$

where the angle brackets  $\langle \cdots \rangle$  denote the average over the random variables  $k, \{n_i\}$ , and  $\{u_i\}$ , whose joint PDF is given by Eq. (5). It is convenient to consider the Laplace transform of the energy in Eq. (73)

$$\int_0^\infty Q_N(E) e^{-\alpha E} dE = \left[ \exp \left( -\frac{\alpha}{2} \sum_{i=1}^k n_i u_i^2 \right) \right]. \quad (74)$$

To evaluate this average using the measure in Eq. (5), we need to first evaluate  $\mathcal{P}(n,u)$  in Eq. (7). For Gaussian  $\phi(v)$ , this is simple:  $\mathcal{P}(n,u) = \sqrt{\frac{n}{2\pi}} \exp[-\frac{nu^2}{2}]$  exactly for all  $n$  and  $u$ . We substitute this in the measure in Eq. (5) and note that

calculating the average in Eq. (74) just amounts to renormalizing the effective  $\mathcal{P}(n, u) \rightarrow \sqrt{\frac{n}{2\pi}} \exp[-(1+\alpha)\frac{nu^2}{2}]$ . With this input, one can then carry out exactly the same procedure as in Sec III A and one arrives at the final result

$$\int_0^\infty Q_N(E) e^{-\alpha E} dE = \sum_{k=1}^N (1+\alpha)^{-k/2} \left[ \frac{S_1(N, k)}{N!} \right], \quad (75)$$

where we recall  $S_1(N, k)/N!$  from Eq. (18). The inverse Laplace transform yields

$$Q_N(E) = \sum_{k=1}^N \left[ \frac{E^{k/2-1} e^{-E}}{\Gamma(k/2)} \right] \left[ \frac{S_1(N, k)}{N!} \right]. \quad (76)$$

Since in the final state the number of clusters  $k$  is distributed according to  $p_N(k) = S_1(N, k)/N!$ , it is interesting to note that, for a given number  $k$  of final clusters, the distribution of the total energy is identical to that of  $k$  initial particles.

Equation (76) provides an exact expression for  $Q_N(E)$ , for all values of  $E$  and  $N$ . However, it is always desirable to have an closed-form expression, even though its validity would require taking the limit of large  $N$ . To achieve this goal, we first express Eq. (75) in the following form by using the generating function given in Eq. (17):

$$\int_0^\infty Q_N(E) e^{-\alpha E} dE = \frac{1}{N!} \frac{\Gamma(N+1/\sqrt{1+\alpha})}{\Gamma(1/\sqrt{1+\alpha})}. \quad (77)$$

At this point, we make a short detour to compute the mean and variance of the final total energy, as they are readily available from the first and second derivatives of Eq. (77) with respect to  $\alpha$  at  $\alpha=0$ . We find the mean as

$$\langle E \rangle = \frac{1}{2} H_N, \quad (78)$$

where  $H_N = \sum_{k=1}^N k^{-1}$  is the harmonic number and, asymptotically,  $H_N \sim \ln N + C$ , where  $C$  is the Euler's constant. The variance is

$$\langle E^2 \rangle - \langle E \rangle^2 = \frac{3}{4} H_N - \frac{1}{4} H_{N,2}, \quad (79)$$

where  $H_{N,r} = \sum_{k=1}^N k^{-r}$  is generalized harmonic number. Note that  $H_N = H_{N,1}$ . The limiting value of  $H_{N,r}$  as  $N \rightarrow \infty$  is the Riemann zeta function  $H_{\infty,r} = \zeta(r)$ , and  $\zeta(2) = \pi^2/6$ .

We now return to Eq. (77), and proceed along the main course to obtain a closed-form expression of  $Q_N(E)$  for large  $N$ . The inverse Laplace transform of Eq. (77) is formally given by the Bromwich integral

$$Q_N(E) = \frac{1}{2\pi i} \int_{c-i\infty}^{c+i\infty} e^{S(\alpha)} d\alpha, \quad (80)$$

where

$$S(\alpha) = \ln \Gamma(N+1/\sqrt{1+\alpha}) - \ln(N!) - \ln \Gamma(1/\sqrt{1+\alpha}) + \alpha E. \quad (81)$$

Now a saddle point approximation of the integral in Eq. (80) yields

$$Q_N(E) \approx e^{S(\alpha^*)/\sqrt{2\pi|S''(\alpha^*)|}}, \quad (82)$$

in which  $\alpha^*$  is determined implicitly in terms of  $E$  and  $N$  by the saddle point condition  $S'(\alpha^*)=0$ . Enforcing this condition to Eq. (81) gives the saddle point equation for  $\alpha^*$  as

$$E = \frac{\beta^3}{2} [\psi(N+\beta) - \psi(\beta)], \quad (83)$$

where  $\beta = 1/\sqrt{1+\alpha^*}$  and  $\psi(x) = (d/dx) \ln \Gamma(x) = \Gamma'(x)/\Gamma(x)$  is the digamma function. As  $\psi(N+\beta) \sim \ln N$ , Eq. (83) suggests that when both  $E$  and  $N$  are large, the natural scaling variable is  $z = 2E/\ln N$ . In terms of this scaling variable, for large  $N$ , from Eq. (83) we find

$$\beta = z^{1/3} + \frac{z^{1/3}}{3 \ln N} \psi(z^{1/3}) + O\left(\frac{1}{[\ln N]^2}\right). \quad (84)$$

Now, employing Stirling's approximation (for large  $N$ ) in Eq. (81) and using Eq. (84), in terms of  $z$  we have

$$S(\alpha^*) = \left[ \frac{3}{2} z^{1/3} - \frac{z}{2} - 1 \right] \ln N - \ln \Gamma(z^{1/3}) + O\left(\frac{1}{\ln N}\right). \quad (85)$$

Similarly, we find

$$S''(\alpha^*) \sim \frac{3}{4} z^{5/3} \ln N. \quad (86)$$

Therefore, substituting  $S$  and  $S''$  from Eqs. (85) and (86), respectively, in Eq. (82), finally we obtain

$$Q_N\left(\frac{z}{2} \ln N\right) \approx \left(\frac{3\pi}{2} \ln N\right)^{-1/2} \frac{e^{-h(z) \ln N}}{z^{5/6} \Gamma(z^{1/3})}, \quad (87)$$

where the large deviation function  $h(z)$  is given by

$$h(z) = 1 + \frac{z}{2} - \frac{3}{2} z^{1/3}. \quad (88)$$

The saddle-point approximation (87) is indeed a very good for large  $N$ , as we demonstrate in Fig. 5 by comparing with the exact result Eq. (76) for  $N=2000$ .

Near  $z=1$  we have  $h(1+x) = x^2/6 + O(x^3)$ . Therefore, from Eq. (87) we find that  $Q_N(E)$  has a Gaussian form close to the mean  $\langle E \rangle \sim (1/2) \ln N$ , with a variance  $\langle E^2 \rangle - \langle E \rangle^2 \sim (3/4) \ln N$ ;—we note that by taking the large  $N$  limit in Eqs. (78) and (79), respectively, we independently recover these mean and variance. However, far away from the mean, the PDF is quite asymmetric as clearly seen in Fig. 5.

## VI. SUMMARY AND CONCLUSION

In summary, we have studied analytically various statistical properties in the final fan state of a one-dimensional sticky gas of  $N$  particles. The particles are initially distributed randomly in space, each with unit mass and with an initial velocity drawn independently from an identical distribution  $\phi(v)$ . Each particle moves ballistically and when two particles collide, they form a single cluster with the total mass and total momentum conserved in the collision process.



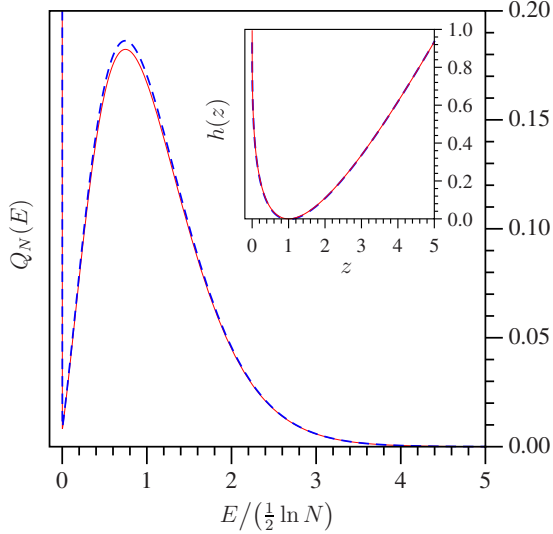


FIG. 5. (Color online) Main: Distribution of the total energy for  $N=2000$ . The solid (red) line plots the saddle-point approximation (87), and the dashed (blue) line plots the exact form of Eq. (76), evaluated using MATHEMATICA. Inset: The solid (red) line plots the large deviation function (88), and the dashed (blue) line plots  $-(\ln N)^{-1} \ln Q_N(\frac{z}{2} \ln N) - (\ln N)^{-1} \ln[(\frac{3\pi}{2} \ln N)^{1/2} z^{5/6} \Gamma(z^{1/3})]$  as a function of  $z$ , using the exact form of  $Q_N(E)$  given in Eq. (76).

In the long time limit, the system reaches a fan state where the system consists of a finite number of clusters with their velocities increasing from left to right and there is no further collision. We have shown that the sizes and the number of clusters are distributed universally in the fan state, independent of  $\phi(v)$  and this distribution is identical to that of the cycles lengths in random permutation problem. We have also computed exactly the distribution of the size and the velocity of the right-most cluster (leader) moving with the largest velocity. Furthermore, the distribution of the total energy in the fan state is also computed exactly for Gaussian  $\phi(v)$ . Our results provide an exactly solvable case of a many-body system (of finite size  $N$ ) with dissipative dynamics and in particular, brings out the universal features of the final stationary state in an explicit way.

There are interesting open issues for further research. Here we have focused solely on the final fan state. It would be interesting to study the dynamics, i.e., the approach to this fan state and investigate to what extent the universal features exist away from the fan state.

There are also immediate generalizations of the problem studied here that are open for future research. For example, it would be interesting to know to what extent the universal features in the fan state remain valid for an inhomogeneous sticky gas. The inhomogeneity can arise either from unequal initial masses or nonidentical velocity distributions in the initial condition. For an infinite system, i.e., with a finite density of initial particles, the ballistic aggregation model with inhomogeneous initial condition has been studied [32]. It would be interesting to extend this study to the case of a finite number  $N$  of particles.

Finally, it would be interesting to study this finite system of  $N$  particles when the collisions are inelastic, but not nec-

essarily sticky, i.e., with a nonzero coefficient of restitution. This problem has been studied for an infinite system and many asymptotic properties such as the energy decay and the velocity distribution were found to be similar to the sticky gas limit [33,34]. It would be interesting to see to what extent the universal features in the stationary state for finite  $N$  are retained when one introduces a nonzero coefficient of restitution.

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## APPENDIX A: PROOF OF EQUATION (22)

In this appendix, we derive the distribution of the cluster sizes (regardless of their positions) starting from joint probability distribution of the cluster sizes and velocities given in Eq. (5). To calculate  $P_N(c_1, c_2, \dots, c_N) = \text{Prob}\{c_1, c_2, \dots, c_N\}$  from the joint-PDF  $p_N(k; \{n_i, u_i\})$ , we must (i) integrate out the velocities, (ii) sum over all configurations with the same cluster sizes (but placed in different orders), (iii) change the variables from the list  $(n_1, \dots, n_k)$  that represents the sizes of the  $k$  consecutive clusters to the set  $\{c_1, \dots, c_N\}$  that encodes the number of clusters of a given size. We thus have

$$\text{Prob}\{c_1, c_2, \dots, c_N\} = \sum_{\substack{\text{Permutations of} \\ \text{unequal size clusters}}} p_N(n_1, \dots, n_k), \quad (\text{A1})$$

where the sum is restricted to permutations that exchange clusters having different sizes, i.e., permutations of the list  $(n_1, \dots, n_k)$  that exchange  $n_i$ 's having different values. Using Eq. (6) and the fact that  $\prod_{i=1}^k n_i = \prod_{j=1}^N j^{c_j}$ , we find that  $p_N(n_1, \dots, n_k)$  is given by

$$p_N(n_1, \dots, n_k) = \frac{\delta(N - \sum_{j=1}^N j c_j)}{\prod_{j=1}^N j^{c_j}} \int \prod_{l=1}^k du_l \mathcal{P}(n_l, u_l) \prod_{l=1}^{k-1} \theta(u_{l+1} - u_l). \quad (\text{A2})$$

From the definition of  $\mathcal{P}(n_l, u_l)$  given in Eq. (7), the integral on the RHS of Eq. (A2) is found to be

$$\int \prod_{i=1}^N dv_i \phi(v_i) \prod_{l=1}^{k-1} \theta\left(\frac{1}{n_{l+1}} \sum_{j=N_l+1}^{N_{l+1}} v_j - \frac{1}{n_l} \sum_{j=N_{l-1}+1}^{N_l} v_j\right), \quad (\text{A3})$$

where  $N_l = n_1 + \dots + n_l$  for  $l = 1, \dots, k$ . The proof of Eq. (22) hence reduces to showing the following identity:



$$\begin{aligned}
& \sum_{\text{Permutations of unequal clusters}} \int \prod_{i=1}^N dv_i \phi(v_i) \\
& \times \prod_{l=1}^{k-1} \theta \left( \frac{1}{n_{l+1}} \sum_{j=N_l+1}^{N_{l+1}} v_j - \frac{1}{n_l} \sum_{j=N_{l-1}+1}^{N_l} v_j \right) \\
& = \frac{1}{c_1! c_2! \dots c_N!}, \quad (\text{A4})
\end{aligned}$$

where we recall that  $c_1$  is equal to the total number of 1's in the list  $(n_1, \dots, n_k)$ ,  $c_2$  is equal to the total number of 2's in this list, etc. This identity is proved by the following argument: consider a permutation that exchanges two clusters of the same size  $p$ ; then, by relabeling the  $v_i$  (which are dummy variables) we observe that the integral on the left-hand side (LHS) of Eq. (A4) is invariant. There are  $c_p!$  possible permutations of the  $c_p$  clusters of size  $p$ . Thus we have formally

$$c_1! c_2! \dots c_N! \sum_{\text{Permutations of unequal clusters}} = \sum_{\text{All permutations}}. \quad (\text{A5})$$

If we have  $k$  distinct numbers  $u_1, \dots, u_k$  there is only one permutation that orders them in increasing order. Therefore, we have

$$\sum_{\sigma \in \Sigma_k} \prod_{l=1}^{k-1} \theta(u_{\sigma(l+1)} - u_{\sigma(l)}) = 1, \quad (\text{A6})$$

where  $\Sigma_k$  is the symmetric group of  $k$  elements. After substituting this identity in Eq. (A5) and using the fact that the PDF  $\phi(v_i)$  is normalized, Eq. (A4) is finally proved. We note that the method of proof used here is closely related to the derivation by Spitzer of a generalization of the Sparre-Andersen theorem [16].

## APPENDIX B: LIMITING BEHAVIOR OF $F(z)$ AS $z \rightarrow 0$ IN EQ. (57)

Let us write  $F(z)$  in Eq. (57) as  $F(z) = \exp[-S(z)]$  where the sum

$$S(z) = \sum_{n=1}^{\infty} \frac{1}{2n} \operatorname{erfc} \left( \sqrt{\frac{n}{2}} z \right). \quad (\text{B1})$$

We want to compute the behavior of  $S(z)$  as  $z \rightarrow 0$ . We first note the identity

$$\sum_{n=1}^{\infty} \frac{1}{2n} \exp(-nz^2/2) = -\frac{1}{2} \ln[1 - e^{-z^2/2}]. \quad (\text{B2})$$

Subtracting Eq. (B2) from Eq. (B1) gives

$$S(z) + \frac{1}{2} \ln[1 - e^{-z^2/2}] = \sum_{n=1}^{\infty} \frac{1}{2n} \left[ \operatorname{erfc} \left( \sqrt{\frac{n}{2}} z \right) - e^{-nz^2/2} \right]. \quad (\text{B3})$$

Now, let us define a new variable  $y = nz^2/2$ . As  $n$  changes by 1,  $y$  changes by  $\Delta y = z^2/2$ . Now, in the limit  $z \rightarrow 0$ , this increment  $\Delta y$  is very small. Thus, one can replace the sum over  $n$  on the RHS of Eq. (B3) by an integral over  $y$  to leading order for small  $z$ , giving

$$S(z) + \frac{1}{2} \ln[1 - e^{-z^2/2}] \approx I = \frac{1}{2} \int_0^{\infty} \frac{dy}{y} [\operatorname{erfc}(\sqrt{y}) - e^{-y}]. \quad (\text{B4})$$

The integral  $I$  on the RHS is just a constant and can be evaluated by parts

$$I = -\frac{1}{2} \left[ \int_0^{\infty} e^{-y} \ln(y) dy - \frac{1}{\sqrt{\pi}} \int_0^{\infty} \frac{\ln(y)}{\sqrt{y}} e^{-y} dy \right]. \quad (\text{B5})$$

The two integrals on the RHS are elementary ones and can be easily evaluated to give finally,  $I = -\ln(2)$ . Substituting this in Eq. (B3) and taking the  $z \rightarrow 0$  limit gives

$$S(z) \rightarrow -\ln(\sqrt{2}z), \quad (\text{B6})$$

implying as  $z \rightarrow 0$  to leading order

$$F(z) = \exp[-S(z)] \rightarrow \sqrt{2}z. \quad (\text{B7})$$

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