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## Quantifying uncertainty in estimates of mineral dust flux: An intercomparison of model performance over the Bodélé Depression, northern Chad

M. C. Todd,<sup>1</sup> D. Bou Karam,<sup>2</sup> C. Cavazos,<sup>1</sup> C. Bouet,<sup>3</sup> B. Heinold,<sup>4</sup> J. M. Baldasano,<sup>5,6</sup> G. Cautenet,<sup>3</sup> I. Koren,<sup>7</sup> C. Perez,<sup>5</sup> F. Solmon,<sup>8</sup> I. Tegen,<sup>4</sup> P. Tulet,<sup>9</sup> R. Washington,<sup>10</sup> and A. Zakey<sup>11</sup>

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[1] Mineral dust aerosols play an important role in the climate system. Coupled climate-aerosol models are an important tool with which to quantify dust fluxes and the associated climate impact. Over the last decade or more, numerous models have been developed, both global and regional, but to date, there have been few attempts to compare the performance of these models. In this paper a comparison of five regional atmospheric models with dust modules is made, in terms of their simulation of meteorology, dust emission and transport. The intercomparison focuses on a 3-day dust event over the Bodélé depression in northern Chad, the world's single most important dust source. Simulations are compared to satellite data and in situ observations from the Bodélé Dust Experiment (BoDEx 2005). Overall, the models reproduce many of the key features of the meteorology and the large dust plumes that occur over the study domain. However, there is at least an order of magnitude range in model estimates of key quantities including dust concentration, dust burden, dust flux, and aerosol optical thickness. As such, there remains considerable uncertainty in model estimates of the dust cycle and its interaction with climate. This paper discusses the issues associated with partitioning various sources of model uncertainty.

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### 1. Introduction

[2] Mineral dust aerosols are an important component of the Earth's climate system by virtue of (1) their direct influence on the Earth's radiation budget through absorption and scattering of solar and terrestrial radiation, (2) the indirect effect through modification of cloud properties, (3) the effect of items 1 and 2 on atmospheric stability

and circulation, and (4) their role in terrestrial and oceanic biogeochemical cycles. Despite increased research into these processes in recent decades, aerosols remain one of the greatest sources of uncertainty in interpretation and projection of past and future climate change [Forster *et al.*, 2007]. The distribution, properties and climate impact of mineral dust in particular is relatively poorly understood, despite evidence that the radiative impact may be comparable to that of anthropogenic sulphate aerosols [Forster *et al.*, 2007]. In many regions mineral dust is the biggest contribution to atmospheric optical thickness [Tegen *et al.*, 1997], and evidence exists of increasing dust production in recent decades [Goudie and Middleton, 1992; Prospero and Lamb, 2003].

[3] Observations and simulations have demonstrated that in regions of high dust emission such as North Africa dust aerosols strongly influence surface and tropospheric radiation budgets and consequently the atmospheric circulation [e.g., Miller and Tegen, 1998; Haywood *et al.*, 2005; Grini *et al.*, 2006; Mahowald *et al.*, 2006; Yoshioka *et al.*, 2007]. It is likely that numerical weather prediction models benefit from inclusion of dust aerosols in these regions [e.g., Pérez *et al.*, 2006a; Rodwell and Jung, 2008]. As a result, in recent years a number of experimental and operational dust forecast systems have been developed at institutes around the world, for example, the Navy Aerosol Analysis and

<sup>1</sup>Department of Geography, University College London, London, UK.

<sup>2</sup>Service d'Aéronomie, Institut Pierre-Simon Laplace, Université Pierre et Marie Curie, Paris, France.

<sup>3</sup>Laboratoire de Météorologie Physique, Université Blaise Pascal, Aubière, France.

<sup>4</sup>Leibniz Institute for Tropospheric Research, Leipzig, Germany.

<sup>5</sup>Earth Sciences Department, Barcelona Supercomputing Center, Barcelona, Spain.

<sup>6</sup>Laboratory of Environmental Modelling, Polytechnic University of Catalonia, Barcelona, Spain.

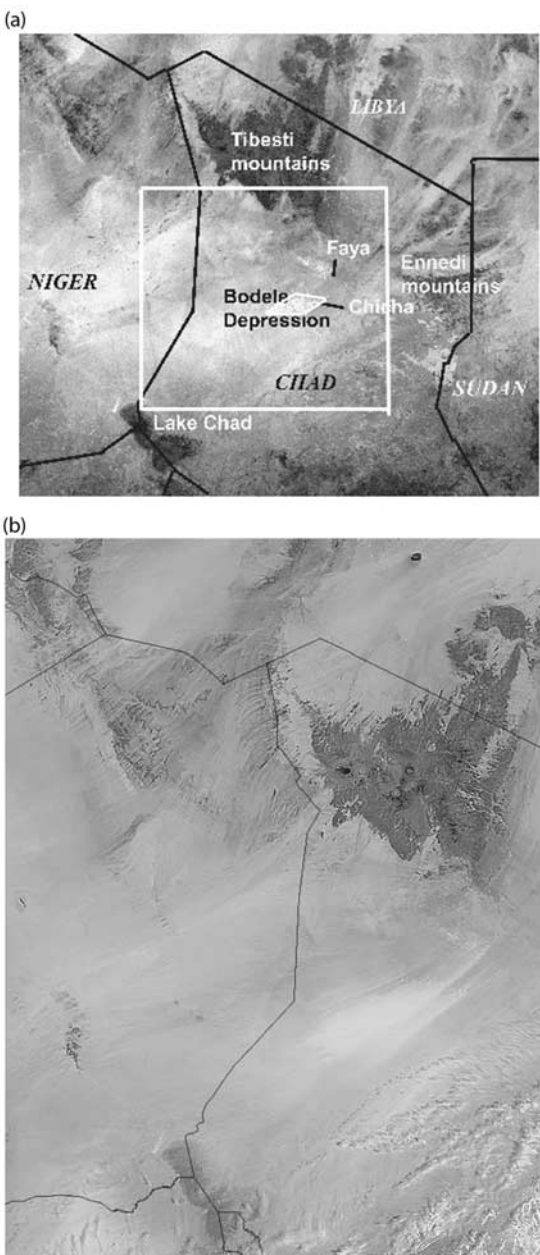
<sup>7</sup>Weizmann Institute, Rehovot, Israel.

<sup>8</sup>Laboratoire d'Aérodynamique, Université Paul Sabatier, Toulouse, France.

<sup>9</sup>GMEI, CNRM, Météo France, Toulouse, France.

<sup>10</sup>Oxford Centre for the Environment, University of Oxford, Oxford, UK.

<sup>11</sup>Abdus Salam International Centre for Theoretical Physics, Trieste, Italy.



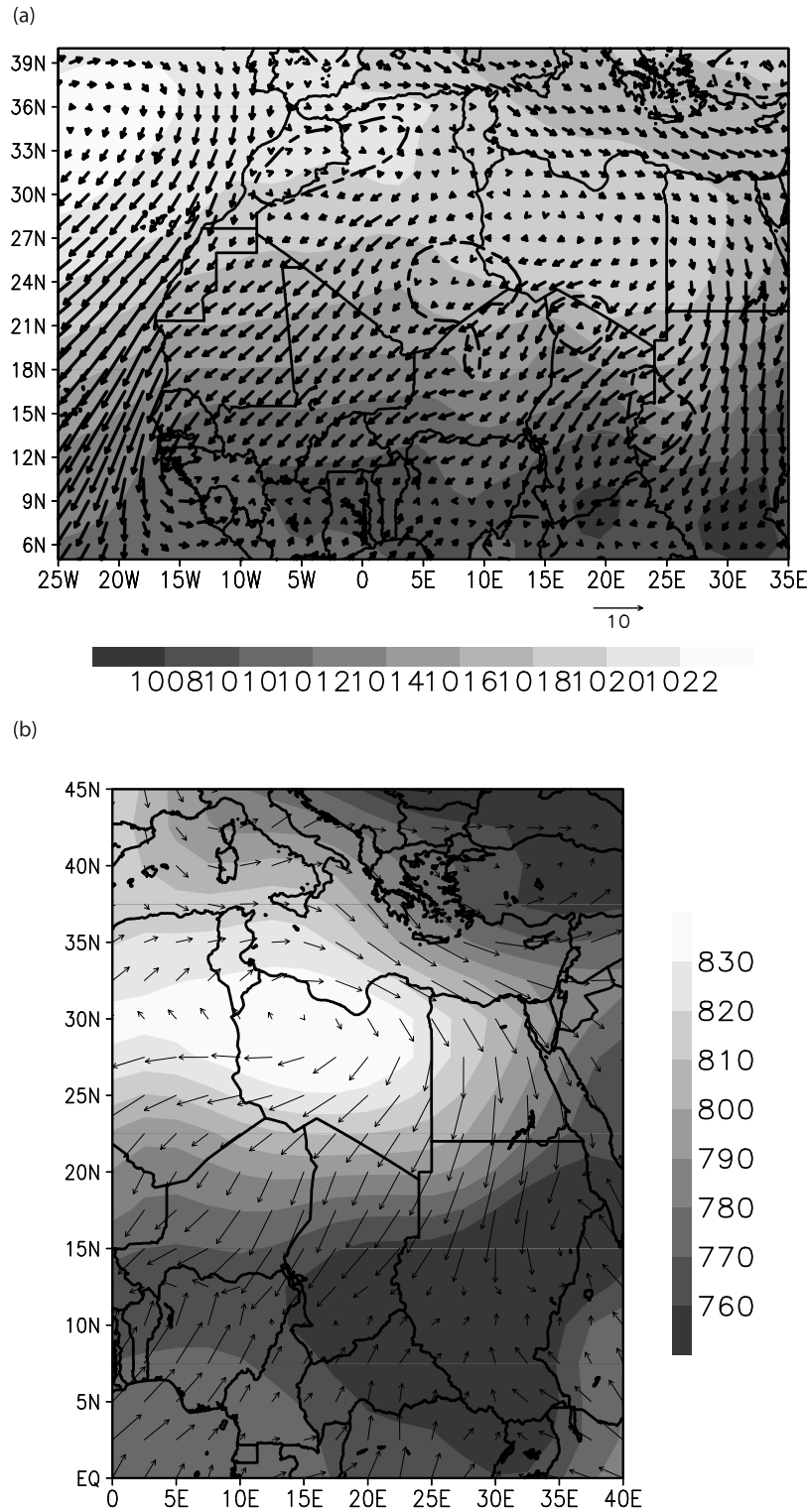
**Figure 1.** (a) MODIS channel 4 surface reflectance (8-day composite, 27 February 2005 to 6 March 2005). (b) MODIS true color composite (channels 3, 4, and 1) image for 0945 UTC 11 March 2005. The location of the BoDEx 2005 field site at Chicha (16.9°N, 18.5°E) and other political and geographical features are indicated. In Figure 1a the domains over which the model quantities are derived in Tables 1–5 are indicated as white lines. The smaller of these encompasses the area of exposed diatomite sediment within the Bodélé Depression, visible as the bright surface immediately east of Chicha.

Prediction System (NAAPS), the Global and regional Earth-system (Atmosphere) Monitoring using Satellite and in situ data (GEMS) at the European Centre for Medium-range Weather Forecasts [Morcrette *et al.*, 2007], and the Dust REgional Atmospheric Model (DREAM) [Nickovic *et al.*,

2001]. At the same time, inclusion of dust aerosols is an important step in the further development of climate and Earth system models. Such dust models can make an important contribution to understanding the role of dust in the climate system. At present, however, there exists considerable uncertainty in model estimates of dust emission and fluxes at global [Zender *et al.*, 2004] and regional scales [Uno *et al.*, 2006]. Therefore, it is important that numerical simulation of dust processes in weather forecast and climate models is improved. A common understanding of model behavior will help reduce this uncertainty in the future and help inform the interpretation of dust forecasts.

[4] It is widely recognized from analysis of satellite data that dust emission occurs primarily in a relatively small number of preferential source regions, mostly located in the world's arid zones, notably the Sahara desert [Herman *et al.* 1997; Prospero *et al.*, 2002; Torres *et al.*, 2002; Washington *et al.*, 2003; Zhang and Christopher, 2003]. Of these source regions the Bodélé depression, in northern Chad is the region with consistently the highest mineral dust aerosol loadings. This region is, therefore, the single most significant dust source, and has been referred to as the “dustiest place on Earth” [Washington *et al.*, 2003]. Unlike other regions of the Sahara it is a major dust source throughout the year. The Bodélé depression is the lowest part of the paleo-lake basin Mega-Chad. The lake last dried up some 5000 years BP [Drake and Bristow, 2006] leaving diatomite sediments exposed at the surface. Today, these sediments cover an area of  $\sim 10,800$  km<sup>2</sup> centered near 17°N, 18°E, clearly apparent from satellite imagery (Figure 1a). The high dust emission from the region is due to the erosive action on these readily erodible sediments of locally strong winds associated with the Bodélé Low Level Jet (LLJ) [Washington and Todd, 2005; Washington *et al.*, 2006; Todd *et al.*, 2007]. This phenomenon is described in more detail in section 3.1 but is associated with the channeling effects of the Tibesti and Ennedi mountains that rise around 2600 m to the north, and 1000 m to the east, respectively, above the Bodélé depression (Figures 1a and 2b). Thus, the coincidence of the LLJ and diatomite sediment produces extensive dust plumes, which extend as singular features up to 1000 km downwind (Figure 1b) and occur with remarkable frequency, around 100 times per year [Koren and Kaufman, 2004; Washington *et al.*, 2006]. Koren *et al.* [2006] estimate that about half of all mineral dust transported from North Africa to South America originates in the Bodélé depression.

[5] A key requirement of dust models, therefore, is that they simulate accurately the dust flux from the Bodélé depression. This represents a key test of model fidelity. This paper presents the results of an intercomparison of model simulations of a large dust event from the Bodélé over 3 days from 10 to 12 March 2005, coincident with the first in situ data acquired from the region during the Bodélé Dust Experiment (BoDEx 2005) [Washington *et al.*, 2006; Todd *et al.*, 2007; Warren *et al.*, 2007] (<http://www.geog.ox.ac.uk/research/climate/projects/bodex/>). The paper aims to (1) determine the degree of uncertainty in estimates of dust emission and transport from a range of model simulations, (2) highlight the sources of uncertainty in these estimates, and (3) point to the key foci for future research to constrain this uncertainty. The aim is not to rank the model



**Figure 2.** (a) Mean (1979–2001) winter JFM season sea level pressure (SLP in hPa, shading) and 925 hPa wind (vectors); surface 1000-m contour is marked with dotted line. (b) Mean 925 hPa height (gpm, shading) and wind ( $\text{m s}^{-1}$  vectors) for 10–12 March 2005. Data for Figure 2a are from ERA-40 reanalysis, and data for Figure 2b are from NCEP reanalysis.

performance but rather highlight the degree of uncertainty in various aspects of model simulation.

[6] The paper is organized as follows. Section 2 provides a description of the models and experimental configuration,

as well as the observational data used for model evaluation. In section 3, an overview of the meteorology of the study region of particular relevance to dust emission is provided. Section 4 includes the primary results of the model inter-

**Table 1.** Dust Model Results for Mean 10-m Height Wind Speed Over Period 10–12 March 2005<sup>a</sup>

Averaging Area	Observations	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Chicha	13.5	8.82	8.42	9.62	11.73	6.78
Area < 250 m height	N/A	8.92	8.08	9.19	11.15	5.98
14°E–21°E, 14°N–21°N	N/A	6.71	6.27	6.0	8.34	4.98

<sup>a</sup>Wind speed given in m s<sup>-1</sup>. N/A denotes not available.

comparison subdivided into sections on meteorology and various components of the dust cycle. Discussion and conclusions follow in section 5.

## 2. Data and Methods

### 2.1. Dust Models

[7] *Kinne et al.* [2003] report that many global dust models underestimate dust emission from key source regions like the Bodélé. In this study we analyze the results of dust flux estimates from dust models configured at the regional scale more appropriate to simulation of this source region. We obtained results from five dust models summarized in Appendix A, Table A1. All the models in this study except DREAM utilize essentially the same dust emission scheme originally developed by *Marticorena and Bergametti* [1995]. In this scheme size resolved dust emission ( $F$ ) is based on parameterizations of soil aggregate saltation and sandblasting processes, in which saltating sand size soil aggregates break soil into smaller particles which are released into the atmosphere (equation (1)). Saltation occurs when the wind friction velocity exceeds a minimum threshold ( $u_{*tri}$ ), itself a function of soil particle size, soil particle density, surface roughness and soil moisture. The size resolved horizontal saltation flux is then largely a function of the third power of wind friction velocity. Finally, the vertical dust flux into the atmosphere results from the sandblasting effect of saltation and is a function of the kinetic energy of the aggregate (proportional to the horizontal flux) and the soil binding energies.

$$F = \alpha \frac{\rho_a}{g} u_*^3 \sum_i \left[ \left( 1 + \frac{u_{*tri}}{u_*} \right) \left( 1 - \frac{u_{*tri}^2}{u_*^2} \right) S_i \right] \cdot (1 - A_{veg})(1 - A_{snow}), \quad (1)$$

where  $\alpha$  is the ratio between the vertical dust flux and the horizontal saltation flux (the sandblasting efficiency), and the rest of the right hand side of equation (1) describes the horizontal saltation flux in which  $\rho_a$  is the air density,  $g$  is

the gravitational constant,  $u_*$  is the surface friction velocity,  $u_{*tri}$  is the threshold friction velocity for size fraction  $i$ ,  $S_i$  is the relative surface area covered by particles of size fraction  $i$ ,  $A_{veg}$  is the part of the area covered by vegetation, and  $A_{snow}$  is the part of the area covered by snow [from *Tegen et al.*, 2006]. The  $\alpha$  parameter expresses the relationship between the kinetic energy of the saltating aggregate impacting on the soil and the binding strength of the soil, but the method by which the values are determined varies between the models. Table A1 in Appendix A shows the values of  $\alpha$  where available. In RegCM3 the  $\alpha$  parameter is calculated explicitly as a function of the kinetic energy of the saltating particles (proportional to the mass flux) and the soil binding energies of the soil particles from *Alfaro and Gomes* [2001], and as such is a complex function of friction velocity, soil texture and emission mode size. The values of  $\alpha$  in the RAMS-DPM are calculated from the empirical relation deduced from the measurements of *Gillette* [1979], and are a function of soil texture, specifically clay content, where higher soil clay content is associated with higher values of  $\alpha$ . In LM-MUSCAT the single  $\alpha$  parameter was actually “tuned” for these particular experiments by optimizing to ensure good match between modeled and observed aerosol optical thickness (AOT) during BoDEX itself. In Meso-NH a single  $\alpha$  parameter is used on the basis of that utilized in global simulations by *Zender et al.* [2003]. This  $\alpha$  value provided acceptable estimates of AOT by Meso-NH compared to observations for model runs over West Africa at relatively low resolution (40–60 km). Interestingly, P. Tulet (personal communication, 2008) notes that the  $\alpha$  values here may be too high by about 33%, as subsequent model calibration experiments suggest a dependence on model resolution, an issue considered further in section 5. The values of  $\alpha$  used in the models here (Table A1 in Appendix A) range over  $\sim 3$  orders of magnitude which is similar to the range of values observed by *Gillette* [1979] associated with variations in soil clay content from 0 to 20%. In most cases over the Sahara the clay content of the soil, like soil texture, is not well constrained [*Zender et al.*, 2003; *Tegen et al.*, 2006] such

**Table 2.** Dust Model Results for Dust Emission Rate and Total Emission From Available Models<sup>a</sup>

Quantity	Averaging Area	RegCM3	LM-MUSCAT	RAMS-DPM
Peak emission (1200 UTC 10 March 2005)	Chicha	$6.93 \times 10^{-4}$	$5.92 \times 10^{-4}$	$3.38 \times 10^{-3}$
	area <250 m height	$9.0 \times 10^{-4}$	$6.17 \times 10^{-4}$	$2.58 \times 10^{-3}$
	14–21°E, 14–21°N	$2.63 \times 10^{-4}$	$2.58 \times 10^{-4}$	$5.61 \times 10^{-4}$
Mean emission (10–12 March 2005)	Chicha	$4.1 \times 10^{-4}$	$4.69 \times 10^{-4}$ <sup>b</sup>	$1.38 \times 10^{-3}$
	area <250 m height	$5.0 \times 10^{-4}$	$5.37 \times 10^{-4}$ <sup>b</sup>	$1.3 \times 10^{-3}$
	14–21°E, 14–21°N	$1.39 \times 10^{-4}$	$2.1 \times 10^{-4}$ <sup>b</sup>	$2.81 \times 10^{-4}$
Total emission (10–12 March 2005)	area <250 m height	1.4	1.5	3.6
	14–21°E, 14–21°N	21.1	31.9	42.5

<sup>a</sup>Dust emission rate given in g m<sup>-2</sup> s<sup>-1</sup>; total emission given in Tg.

<sup>b</sup>Emission for LM-MUSCAT is for 1200 UTC only.

**Table 3.** Dust Model Results for Mean Lowest Level Dust Concentration and Column Integrated Dust Mass Averaged Over 10–12 March 2005<sup>a</sup>

Quantity	Averaging Area	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Lowest level dust concentration	Chicha	3757	15322	26000	19740	2148
	area <250 m height	4500	15700	25200	26800	2700
Column integrated dust mass	Chicha	1.91	11.57	14.92	24.5	2.44
	area <250 m height	1.57	11.58	16.1	26.37	2.87
	14°E–21°E, 14°N–21°N	0.77	5.74	11.36	11.61	1.84

<sup>a</sup>Dust concentration given in  $\mu\text{g m}^{-3}$ , and column integrated dust mass given in  $\text{g m}^{-2}$ .

that empirical “tuning” of the  $\alpha$  parameter is frequently required (see section 5 for further discussion).

[8] The DREAM model uses the dust emission scheme originally developed by *Shao et al.* [1993] in which the dust flux,  $F$ , is, again, a function of the third power of wind friction velocity.

$$F = c \times \delta \times u_*^3 [1 - (u_{*i}/u_*)]. \quad (2)$$

Dust flux is scaled by dust productivity factor  $\delta$  which takes into account effects of soil structure and particle size distribution, and a constant,  $c$  (conceptually similar to the  $\alpha$  parameter in equation (1)). Dust flux is then used to determine the surface dust concentration. However, unlike the other models, DREAM includes a viscous sublayer between the surface and the lowest model layer [*Janjic*, 1994], since there is a physical similarity between mass/heat/momentum exchange over surfaces such as the ocean with that of mobilized dust particle over desert surfaces [*Chamberlain*, 1983; *Segal*, 1990]. This parameterizes the turbulent transfer of dust into the lowest model layer accounting for different turbulent regimes, using the simulated surface dust concentration as the lower boundary. More details on this scheme can be found in *Nickovic et al.* [2001] and *Pérez et al.* [2006b]. As a forecast model DREAM has been extensively “tuned” with respect to observations of AOT over North Africa by optimizing the constant  $c$  in the dust flux equation (2).

[9] This saltation and sandblasting mechanism parameterized in equation (1) is commonly used in dust models. However, observations during BoDEx 2005 showed that in the diatomite sediments of the Bodélé depression the actual process of dust generation is more complex [*Warren et al.*, 2007]. In the Bodélé saltation and sandblasting occur mostly where transport of sand sized quartz particles from the surrounding sand dunes is pronounced, largely at the northeastern end of the diatomite sediment where the dune density is highest. However, over most of the diatomite sediment of the Bodélé an unusual process of “autoabrasion” occurs; in which saltating diatomite aggregates themselves disintegrate into fine dust sized particles [*Warren et al.*, 2007]. This dust generation mechanism has yet to be

parameterized in dust models. Therefore, in this model intercomparison we evaluate the extent to which the *Marticorena and Bergametti* [1995] scheme (equation (1)) and the *Shao et al.* [1993] scheme of DREAM (equation (2)), under various model configurations, can represent these processes.

[10] Various model output fields at 3-hourly resolution were processed for comparative purposes over the period 10–12 March 2005 during a substantial dust event. The variables represent the driving winds (Table 1), the processes of dust emission (Table 2) and resulting dust burden (Tables 3 and 4) and transport (Table 5). In this study no consideration of deposition rates is made. The research teams involved were free to configure their model experiments as they chose. This condition was agreed for logistical reasons. Therefore, the model horizontal and vertical resolution, domain size, driving boundary conditions, land surface conditions, run duration, and dust physical and optical properties vary between the models, as summarized in Table A1 in Appendix A. Although the differences between model configurations mean the model outputs need to be carefully interpreted, the results provide an initial benchmark of model uncertainty. Future model intercomparison projects can then utilize a more tightly constrained model configuration. For display and analysis model output data were remapped to a common  $0.2^\circ$  spatial resolution over the domain of  $12^\circ\text{E}–24^\circ\text{E}$ ,  $12^\circ\text{N}–24^\circ\text{N}$  and a common vertical resolution of 100 m in the lowest 2 km of the atmosphere. All dust quantities except AOT were integrated over the size range of  $0.1–10 \mu\text{m}$  diameter to represent the component involved in long-range transport. Estimates of AOT were provided by the model teams and are based on the dust optical properties and size distributions specified in each model.

## 2.2. Comparative Data

[11] The focus of this intercomparison study is the large dust event of 10–12 March 2005. We have collated observations of local meteorology and estimates of dust burden, emission flux and aerosol optical thickness from various sources, for comparison with model output. Large-scale meteorological fields were obtained from National Centers for Environmental Prediction (NCEP) reanalysis for the

**Table 4.** Dust Model Results for Mean AOT Averaged Over 10–12 March 2005

Averaging Area	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Chicha	0.93	4.29	6.53	3.67	1.82
Area <250 m height	0.85	4.71	6.47	3.95	2.14
14°E–21°E, 14°N–21°N	0.58	2.1	4.23	1.74	1.38

**Table 5.** Dust Model Results for Net Westward Dust Mass Flux Integrated Over 12.5°N–21°N and 10–12 March 2005<sup>a</sup>

Quantity	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Flux at 16°E 10–12 March 2005	0.44	4.59	4.89	12.8	0.79
Flux at 12.5°E 11–12 March 2005	0.29	2.32	2.76	N/A	0.54

<sup>a</sup>Dust mass flux given in Tg d<sup>-1</sup>.

study period and the ERA-40 reanalysis for the long-term mean. Satellite estimates of AOT were obtained from the “Deep Blue” algorithm of *Hsu et al.* [2004] applied to Sea-viewing Wide Field-of-view Sensor (SeaWiFS) and MODerate resolution Imaging Spectroradiometer (MODIS) data (0.1° resolution, daily at local noon) providing estimates even over highly reflective desert surfaces. SeaWiFS observed the study region on 10 and 12 March only, and there is some cloud contamination on 10 March. It should be acknowledged that Deep Blue estimates of AOT are subject to uncertainties, most notably those associated with assumptions about the aerosol size and shape and vertical distribution used in the generation of the look-up tables of satellite radiances associated with AOT and single scattering albedo. Estimates of AOT from the Multiangle Imaging SpectroRadiometer (MISR) system have very limited coverage over the Bodélé during the period of interest and are only used for calculation of dust mass transport downwind where an overpass occurred. In situ data were obtained from the Bodélé Dust Experiment (BoDEx 2005). Observations of 2 m height winds (converted here to 10 m height using a logarithmic wind profile defined with an appropriate roughness length value for desert surfaces) and AOT measured using a microtops Sun photometer were made at “Chicha” (16.9°N, 18.5°E). Observations of AOT were necessarily sporadic owing to very high dust emissions rendering observation problematic.

[12] Model output and observations are compared at three spatial scales: (1) point quantities at Chicha, (2) quantities averaged over an area which approximates the area of exposed diatomite sediment, defined by the grid cells lower than the 250 m surface height contour in the RegCM3 model domain, and (3) quantities averaged over the domain 14°N–21°N, 14°E–21°E, which covers the area from the northern point of lake Chad in the southwest to the Tibesti-Ennedi gap in the northeast, encompassing the dust plumes in their entirety. These domains are approximately indicated on Figure 1a.

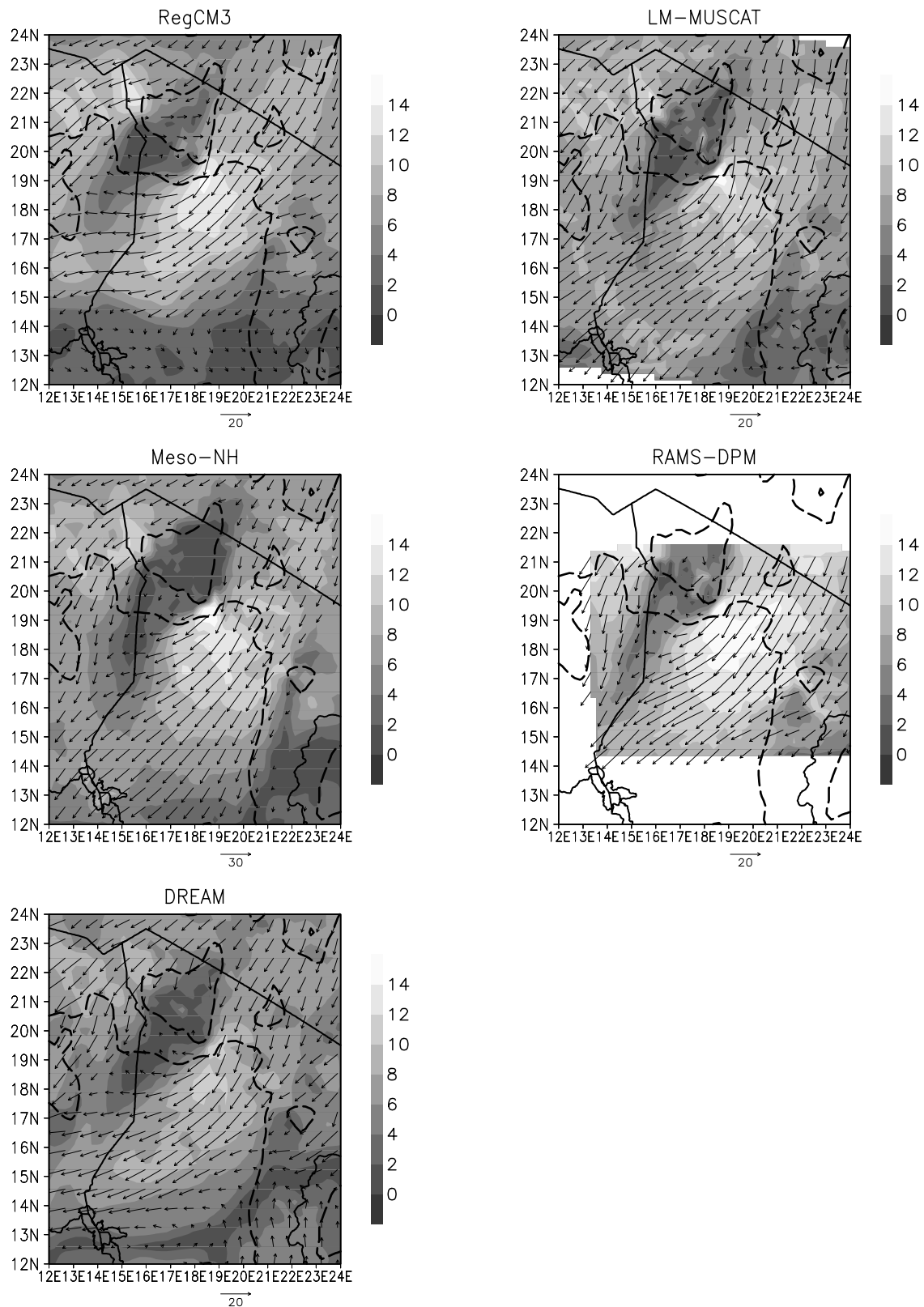
### 3. Overview of the Climate of the Bodélé Depression

[13] In winter, North Africa including the Chadian sector is dominated by the low-level northeasterly Harmattan winds [*Hamilton and Archibald*, 1945; *Hastenrath*, 1988], which occur in the lowest 100 hPa of the atmosphere. These winds are driven by north-south pressure gradient between the mean high-pressure ridge centered over Libya (the Libyan High) and the equatorial trough (Figure 2a). The Bodélé LLJ is embedded in this mean northeasterly Harmattan wind and is oriented NE-SW with a mean maximum of  $\sim 12 \text{ m s}^{-1}$  (at the 925-hPa level in reanalysis) centered near 17°N, 19°E in the ERA-40 reanalysis data (Figure 2a). Mean wind speeds at 925 hPa in excess of  $11 \text{ m s}^{-1}$  extend over a large region from northern Chad to Lake

Chad, a distance of some 1200 km. The high wind speeds to the northwest of the Tibesti mountains (centered near 22°N, 14°E) indicate a split in the Harmattan wind associated with the presence of the mountains.

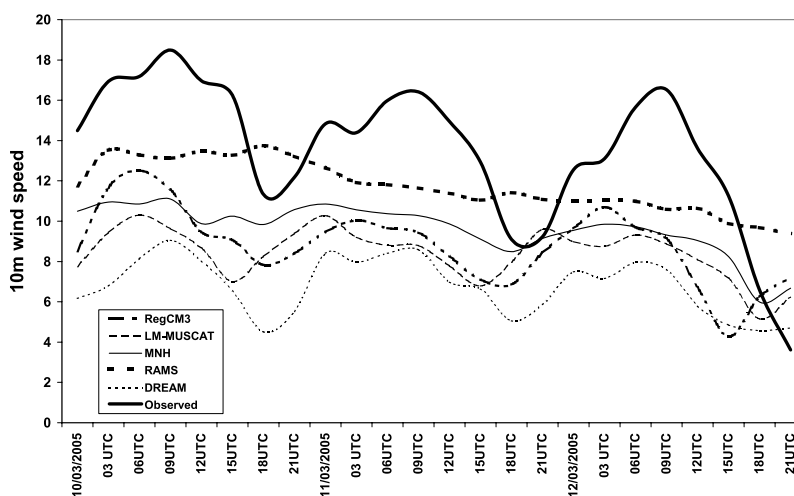
[14] The LLJ exhibits a pronounced diurnal cycle with a nighttime (daytime) maximum (minimum) consistent with the inertial oscillation mechanism [*Blackadar*, 1957; *Washington et al.*, 2006; *Todd et al.*, 2008]. During the day, surface heating leads to a well mixed deep boundary layer that exerts a frictional deceleration of the LLJ. After sunset, radiative cooling of the surface leads to a near-surface temperature inversion and decoupling of the surface and low-level winds. This initiates an inertial oscillation in which the LLJ accelerates through the night, until turbulent mixing starts again the following day. This process leads to an out-of-phase relationship of the diurnal cycle of the LLJ with respect to that of surface winds. At the surface, maximum wind speeds are observed during midmorning when turbulence mixes the momentum of the nocturnal LLJ down to the surface. At night, decoupling of the surface and low levels leads to minimum wind speeds at the surface and LLJ maximum [*Washington et al.*, 2006; *Todd et al.*, 2008]. This pattern is typical of many regions in the tropics that experience a nocturnal LLJ phenomenon [e.g., *May*, 1995; *Parker et al.*, 2005; *Sultan et al.*, 2007; *Knippertz*, 2008]. Accordingly, dust production is ‘locked’ to the diurnal cycle of surface winds and tends to peak during the early mid morning, and weaken during the afternoon [*Washington et al.*, 2006; *Todd et al.*, 2007; *K. Schepanski et al.*, Meteorological processes forcing Saharan dust emission inferred from MSG-SEVIRI observations of subdaily dust source activation, submitted to *Journal of Geophysical Research*, 2008]. This results in discrete daily pulses of dust that are then transported out of the region within the LLJ [*Koren and Kaufman*, 2004; *Todd et al.*, 2007].

[15] Although the nocturnal LLJ is a widespread phenomenon in the region, it is strongest and most pronounced over the Bodélé depression as a result of the local orographic configuration. The northeasterly flow is channeled through the Tibesti-Ennedi gap (Figure 2a) and the combination of a gap wind and downslope wind-forcing acts to accelerate the LLJ by up to 50% [*Todd et al.*, 2008]. Dust events occur when surface winds exceed the threshold velocity for dust emission (about  $10 \text{ m s}^{-1}$  at 2 m height [*Todd et al.*, 2007]). Outside of summer, variability in surface winds is associated with day-to-day synoptic-scale variability in the strength of the north-south pressure gradient across the region, most notably controlled by ridging of the Libyan High [*Washington and Todd*, 2005]. Thus, the temporal characteristics of dust emission from the Bodélé during winter results from the interaction of the basic state of the atmosphere over North Africa, local/regional topography, synoptic variability and the coupled diurnal cycle of the LLJ and surface winds.



**Figure 3.** Model simulated 925 hPa winds ( $\text{m s}^{-1}$ , vectors) and 10-m scalar wind speed ( $\text{m s}^{-1}$ , shading) at 0900 UTC on 10 March 2005. The 500-m and 1000-m surface elevation (dashed line) contours are shown.





**Figure 4.** Time series of in situ observed and model simulated 10 m height wind speed at Chicha ( $\text{m s}^{-1}$ ).

[16] The large dust event from 10 to 12 March 2005 occurred when a blocking anticyclone over the northeastern Atlantic suddenly migrated eastward and extended a pronounced ridge across North Africa, drawing an anomalously strong north-northeasterly flow across the Bodélé region (Figure 2b). The structure and duration of this circulation pattern is quite typical condition for wintertime Bodélé dust events [Washington and Todd, 2005], resulting in pronounced dust emission over 3 days. Typically, the diurnal cycle of winds resulted in emission occurring in pulses that peaked after sunrise and ceased during the late afternoon/evening in each day.

## 4. Results and Discussion

### 4.1. Simulation of Regional Winds

[17] The models under study here all provide similar representation of the spatial structure of the LLJ and 10 m height winds (Figure 3). Key features include a split jet north and south of the Tibesti Mountains with peak wind speeds located immediately south of Tibesti, both in the 3-day time mean and instantaneous fields (Figure 3). All models, except RAMS-DPM, also simulate well the day-to-day variability at the location of Chicha (Figures 4 and 5). All models have a nocturnal LLJ maximum peaking below 1000 m elevation, but this is rather weakly defined in RAMS-DPM. The models LM-MUSCAT, RegCM3 and DREAM all simulate a nocturnal LLJ of similar structure and magnitude, with peak wind speeds of 20–24  $\text{m s}^{-1}$  on 10 March 2005, likely to be lower than the real LLJ wind speeds [Todd *et al.*, 2008]. In addition, all models underestimate mean (in situ) 10 m height winds by 13–50% (Table 1 and Figure 4). RAMS-DPM simulates the highest mean wind speeds, but with peak wind speeds  $\sim 30\%$  lower than in situ observations. It is interesting to note that in the LM-MUSCAT model when the dust-radiative feedback is not activated surface wind speed underestimation is reduced (I. Tegen, personal communication, 2008). All models except RAMS-DPM have pronounced diurnal cycle in surface winds and LLJ but have smaller amplitude than observations (Figures 4 and 5). However, only the DREAM model captures the diurnal phase of wind speed maxima at

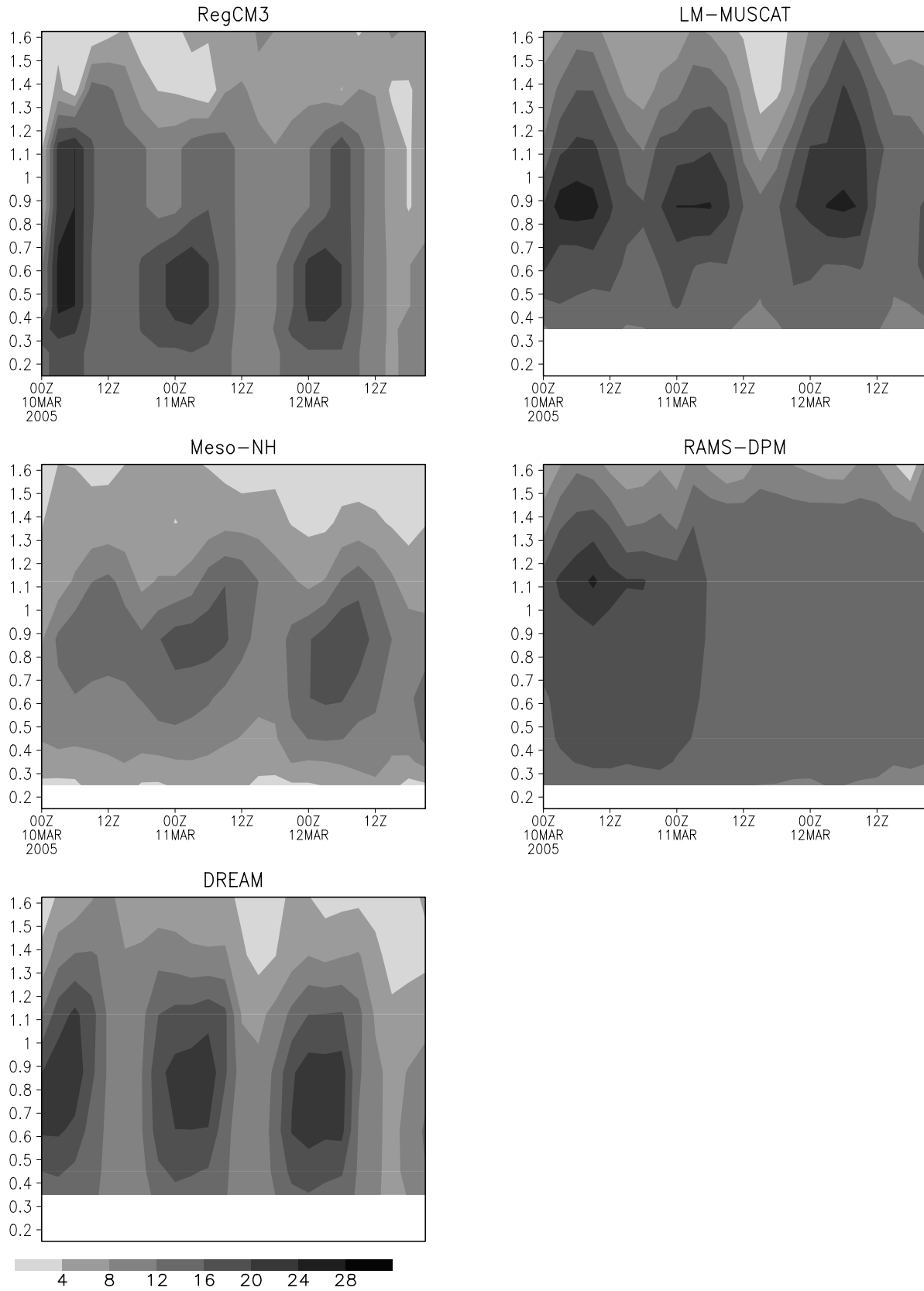
0900 UTC consistent with observations. All other models tend to produce the peak rather too early by  $\sim 3$  h typically. In these models the diurnal cycle of near-surface wind speed is in phase with LLJ jet core such that downward mixing of nocturnal LLJ may not be well simulated. There does not appear to be a relationship between model horizontal or vertical resolution and the accuracy of near-surface wind speed estimates.

[18] Todd *et al.* [2008] undertook a sensitivity study using the MM5 mesoscale model for near-surface and boundary layer winds over the BoDEx period. They note that there is a factor of  $\sim 2$  variation in mean and peak wind speeds over the BoDEx period using MM5 under various configurations. Simulated winds are most sensitive to the model Planetary Boundary Layer (PBL) scheme and vertical resolution in the lowest 200 m. Ten-meter wind speeds from the models in the present study lie within the range simulated using the MM5 model of Todd *et al.* [2008]. Under an optimized configuration (Eta PBL scheme, 15 vertical levels in lowest 1000 m with lowest level at 2 m, roughness length = 0.001 m) the MM5 estimated LLJ and surface wind speeds and phase are close to observed quantities. This suggests that improved representation of the surface and low-level winds responsible for emission and transport of dust in these models would be possible with careful model configuration. Indeed, Laurent *et al.* [2008a] report that a more recent version of RAMS (v6.0) that includes an improved soil scheme provides a more realistic simulation of the diurnal cycle of surface winds than the version used in this study and by Bouet *et al.* [2007].

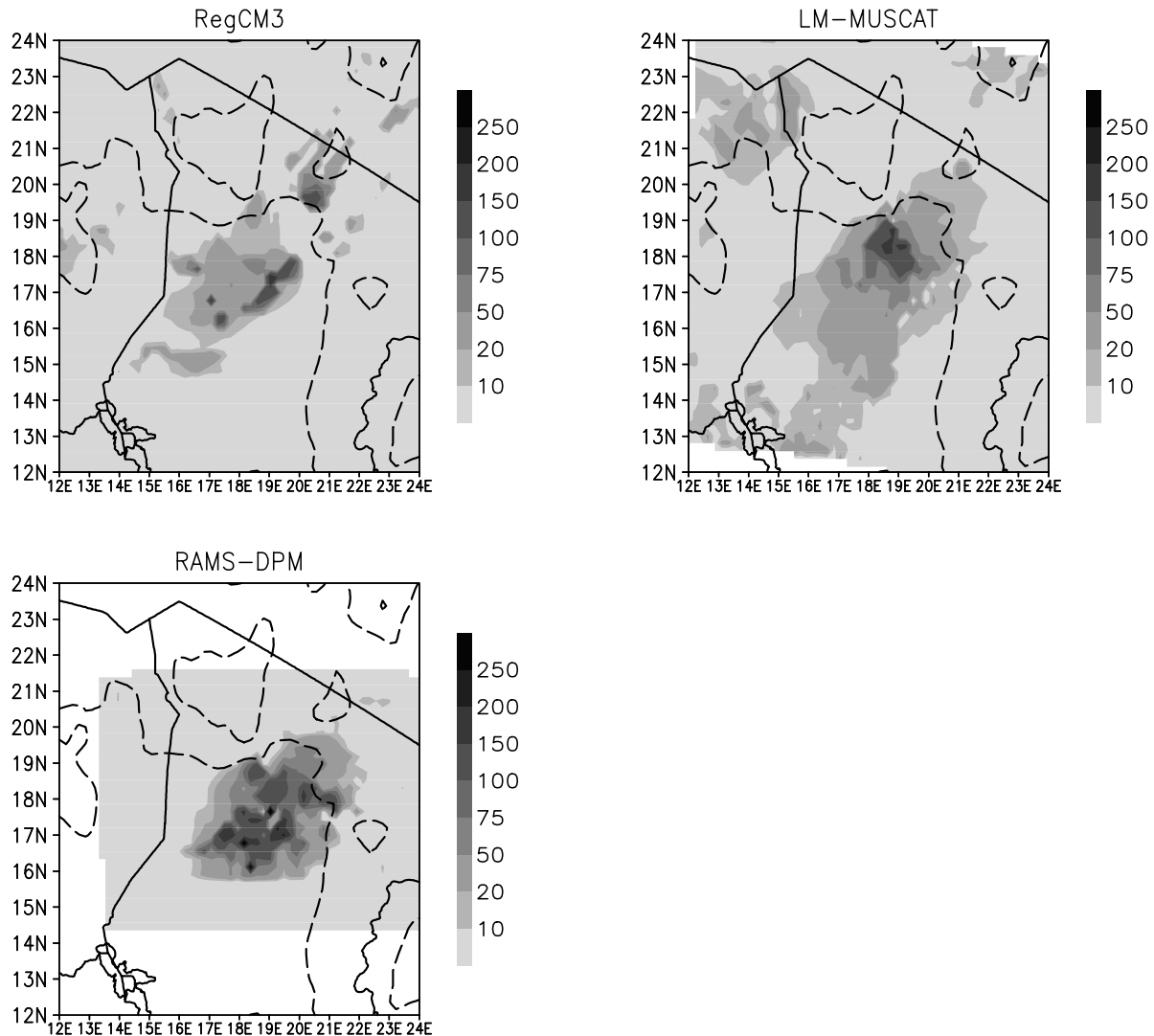
### 4.2. Simulation of Dust Emission and Transport

#### 4.2.1. Dust Emission

[19] Emission estimates are available for all models except Meso-NH. Estimates of dust emission from DREAM are not directly comparable to those of the other models, owing to the use of a viscous sublayer close to the surface (see section 2.1), and are not presented here. The spatial structure in emission (Figure 6) is broadly similar in RAMS-DPM, LM-MUSCAT and RegCM3 with emission restricted to areas of low elevation close the Bodélé depression. RegCM3 produces localized emission sources which



**Figure 5.** Time series of model simulated vertical profile of wind speed at Chicha ( $\text{m s}^{-1}$ ). The y axis is height above the surface in kilometers. Note that for Meso-NH the location is  $16^\circ\text{E}$ ,  $17^\circ\text{N}$ .

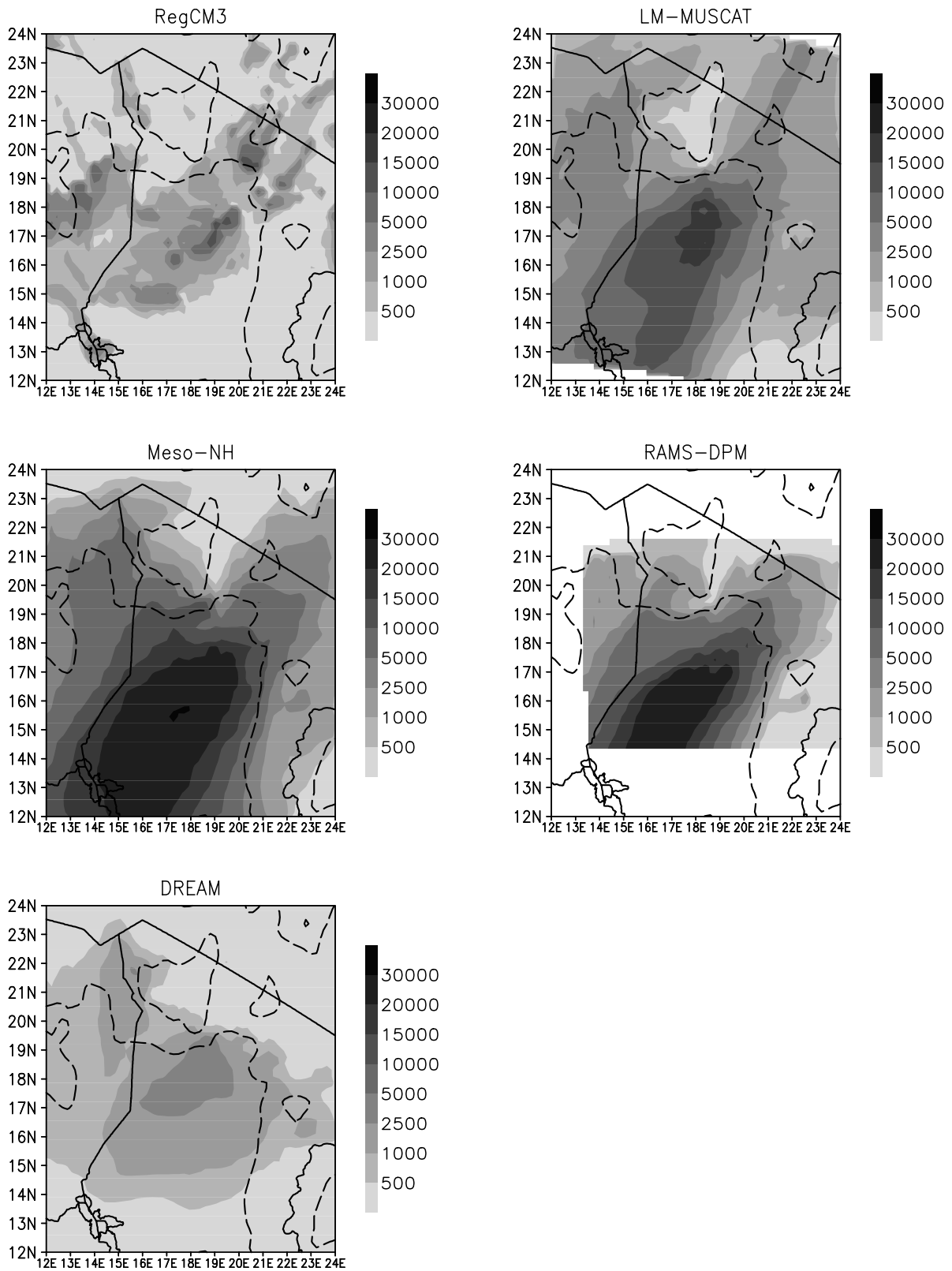


**Figure 6.** Model simulated mean rate of surface dust emission 10–12 March 2005 ( $\text{g m}^{-2} \text{s}^{-1}$ ). The 500-m and 1000-m surface elevation (dashed line) contours are shown.

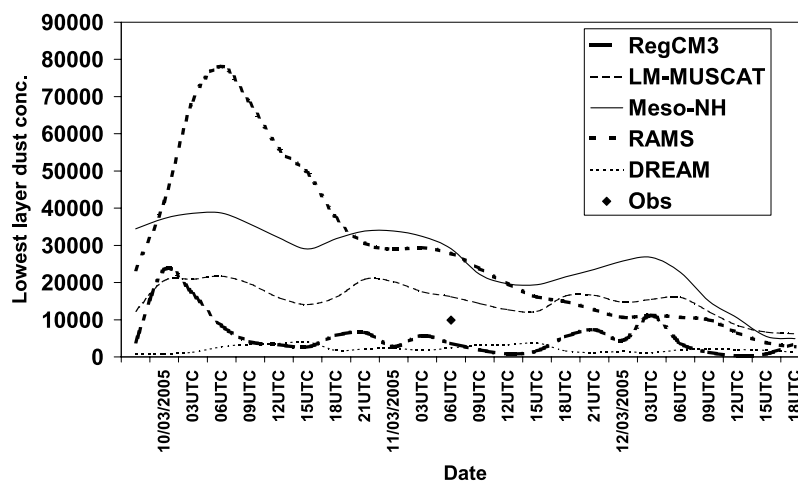
represent well the areal extent of the localized outcrops of diatomite sediment identified as the dust source regions in satellite observations (see below), although not in precisely the same location. In the model these locations correspond to the areas where the soil texture (derived from the FAO/USDA (Food and Agriculture Organization of the United Nations/United States Department of Agriculture) map) is 100% sand, rather than a representation of diatomite, thus activating dust emission by saltating sand particles via the *Martcorena and Bergametti* [1995] emission parameterization. In LM-MUSCAT emission is highest from regions north of  $17^\circ$ . In this model preferential dust sources are parameterized to ensure that areas of high dust productivity are collocated with enclosed topographic depressions including the Bodélé depression [Teegen *et al.*, 2006].

[20] At the location of Chicha, dust emission closely matches the diurnal cycle of near-surface winds, such that the models (except RAMS-DPM) do simulate a diurnal “pulsing” of emission (not shown). Mean emission from the three models available (Table 2) show an approximately two to threefold range in mean and peak dust emission at

the location of Chicha and averaged over the wider domains. Estimates of total emission from the exposed diatomite area within the Bodélé depression were calculated from the product of the mean emission rate over the area less than 250 m elevation (which approximates the diatomite location) and the area of diatomite ( $10,800 \text{ km}^2$ , derived from MODIS surface reflectance data, Figure 1a). Over the 3 days 10–12 March 2005, estimates range between 1.4 Tg for RegCM3 to 3.6 Tg for RAMS-DPM (Table 2). *Todd et al.* [2007] estimated the total dust flux from the Bodélé during the 3-day dust event of 10–12 March 2005 using a simple relationship of surface wind speed and vertically integrated dust flux relationship derived empirically from observations at Chicha, then scaled to the entire area of diatomite sediment. Their estimate was 3.54 Tg, which falls at the upper end of the range of model estimates. It should be noted, however, that the models do not all show peak emission collocated with the region  $< 250 \text{ m}$  height (Figure 6). The equivalent values for the entire domain ( $12^\circ\text{E}–21^\circ\text{E}$ ,  $12^\circ\text{N}–21^\circ\text{N}$ ) range from 21.1 Tg for RegCM3 to 42.5 Tg for RAMS-DPM (Table 2).



**Figure 7.** Model simulated mean dust concentration in model lowest layer 10–12 March 2005 ( $\mu\text{g m}^{-3}$ ). The 500-m and 1000-m surface elevation (dashed line) contours are shown.



**Figure 8.** Time series of model simulated dust concentration in model lowest layer ( $\mu\text{g m}^{-3}$ ) at the location of Chicha.

#### 4.2.2. Dust Concentration

[21] Mean dust concentration ( $0.1\text{--}10\ \mu\text{m}$  diameter) in the lowest model layer show a wide range between the models (Figure 7). Over the study domain as a whole, Meso-NH has the highest mean value, up to  $30,000\ \mu\text{g m}^{-3}$  and DREAM the lowest, about one tenth of that of Meso-NH. RAMS-DPM has similarly high peak values but over a more restricted dust plume area. At Chicha, model mean dust concentration values range over about 1 order of magnitude from 2100 (DREAM) to  $26,000\ \mu\text{g m}^{-3}$  (Meso-NH) (Table 3). Similar values are estimated over the region with elevation  $< 250\ \text{m}$  with RAMS and Meso-NH producing dust concentration of about 1 order of magnitude greater than DREAM. In most models dust concentration at Chicha shows the diurnal pulsing and peaks at 0900 UTC on 10 March when RAMS simulates the highest value of  $78,000\ \mu\text{g m}^{-3}$  (Figure 8). In situ observations at Chicha are limited but a maximum value of  $9900\ \mu\text{g m}^{-3}$  at 2 m height was observed at 1000 UTC on 11 March (Figure 8). Observed dust concentrations therefore lie toward the lower end of the range simulated in the dust models. Caution must be exercised in comparing the models owing to the varying heights of the lowest model layers. As noted above dust emission fields were not available from Meso-NH and DREAM. However, as Meso-NH produces the highest dust concentration of all models it is likely to have the highest emission. This may be related to the soil data used whereby the soil texture is dominated by sand across the entire domain  $14^{\circ}\text{N} - 17^{\circ}\text{N}$  leading to a high horizontal saltation flux and subsequent emission.

#### 4.2.3. Aerosol Optical Thickness

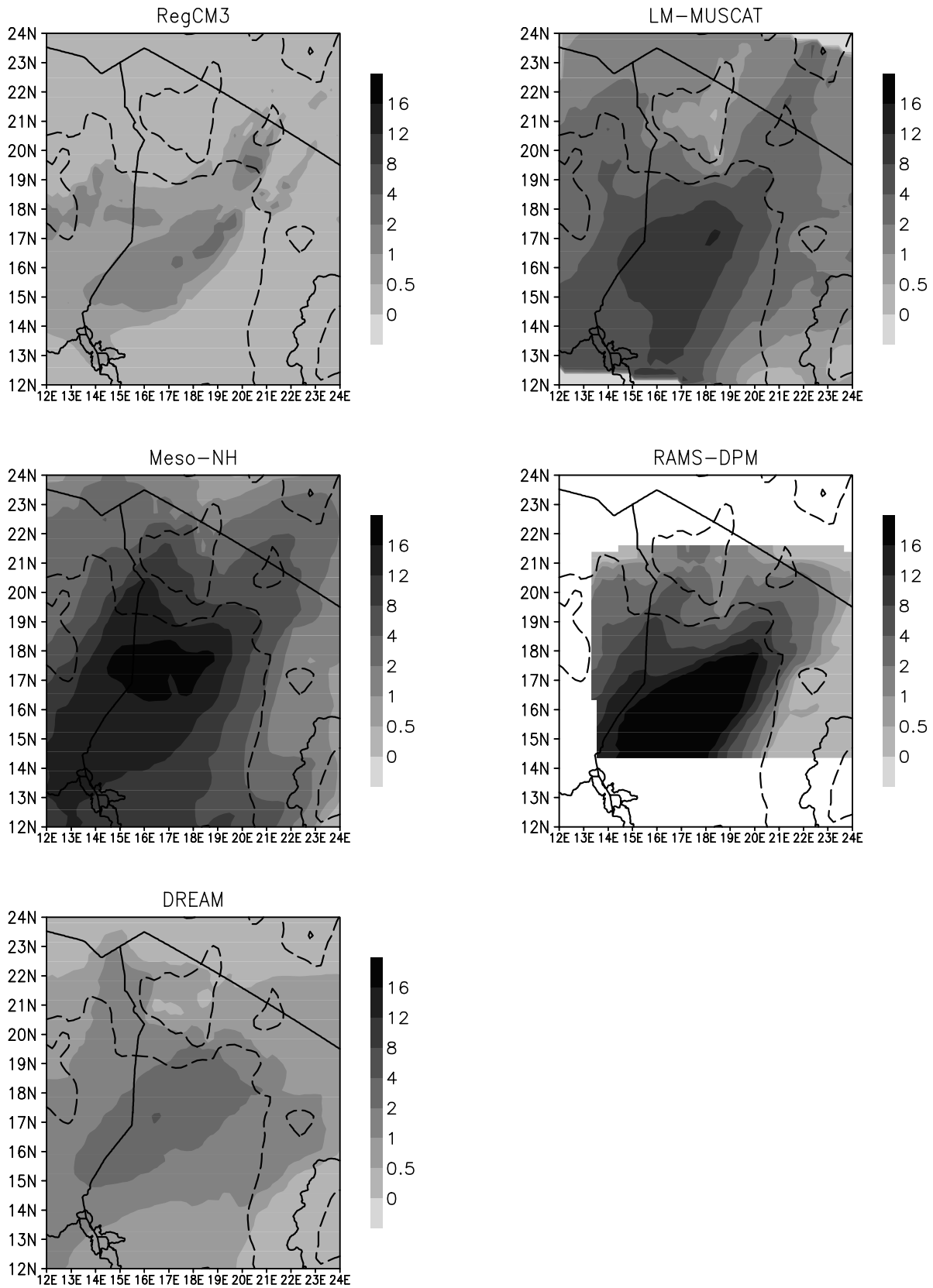
[22] Dust emitted from the surface is transported in the low-level northeasterly winds in all models. As a result, all models simulate NE-SW oriented dust plumes (centered below 1000 m elevation, not shown), which have the fundamental characteristics of the observed data (Figures 7, 9, 10, and 11). Thus, despite small-scale differences in the location of dust source regions the coherent northeasterly LLJ results in dust plumes that are generally good representations of reality. Nevertheless, most models produce a tongue of high

atmospheric dust loadings extending to the north of the main NE-SW oriented plume, centered close to the Chad-Niger border north of about  $17^{\circ}\text{N}$ . Satellite estimates of AOT do not show elevated AOT values there (Figures 10 and 11). The time evolution of the dust plumes shows rapid transport to the southwest in all cases, with the dust transported in distinct diurnal “packets” in most cases. The speed of propagation differs however and the time taken for dust transport to  $14^{\circ}\text{E}$  (the longitude of Lake Chad) varies between about 9 h for Meso-NH and about 15 h for DREAM.

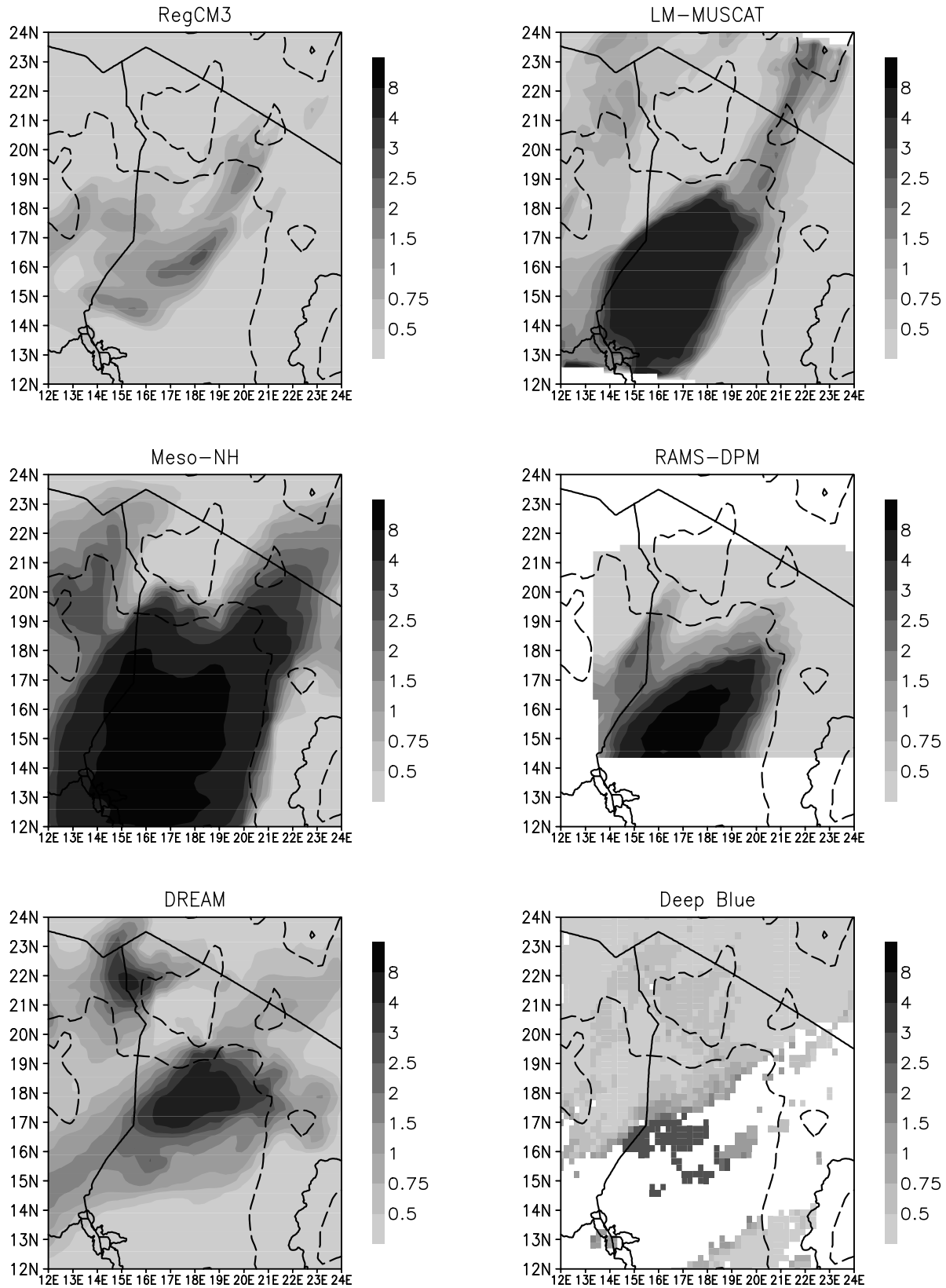
[23] The time-mean column dust mass ( $0.1\text{--}10\ \mu\text{m}$  diameter) at Chicha varies from  $1.91\ \text{g}^{-2}$  in RegCM3 to  $24.5\ \text{g}^{-2}$  in RAMS-DPM (Table 3). At Chicha the time distribution of dust mass is similar in most models with a peak near 1200 UTC on 10 March and subsequent decline (Figure 12). Estimates based on in situ data from Todd *et al.* [2007] fall toward the lower end of the model range closest to the estimates by the RegCM3 and DREAM models. There is about a 15-fold difference in mean column dust mass between RAMS-DPM and RegCM3 averaged over the whole domain (Table 3).

[24] While there is generally close agreement between the models in terms of the spatial structure of AOT there is an approximately sevenfold range in simulated AOT at all spatial scales (Figures 10 and 11 and Table 4). Meso-NH estimates the highest AOT and RegCM3 the lowest. Satellite estimates of AOT from the “Deep Blue” algorithm applied to SeaWiFS and AI from OMI data indicate the distinct dust NE-SW oriented plumes. The spatial structure is particularly well represented by RegCM3 and RAMS-DPM, with Meso-NH and LM-MUSCAT overestimating the areal extent of the plume and DREAM locating the plume rather too far north, a consequence of the source areas and plume propagation speed. We should be aware of potential errors in the Deep Blue aerosol product but the Aerosol Index product [Herman *et al.*, 1997] from the Ozone Monitoring Instrument (OMI, not shown) presents a very similar spatial structure to the Deep Blue AOT.

[25] A quantitative comparison of AOT is problematic as the “Deep Blue” algorithm saturates at a maximum AOT of



**Figure 9.** Model simulated mean vertically integrated dust mass 10–12 March 2005 ( $\text{g m}^{-2}$ ). The 500-m and 1000-m surface elevation (dashed line) contours are shown.



**Figure 10.** AOT estimated by models and “Deep Blue” satellite algorithm on 1200 UTC 10 March 2005. The 500-m and 1000-m surface elevation (dashed line) contours are shown. Note cloud contamination (white color) over dust plume in Deep Blue product.

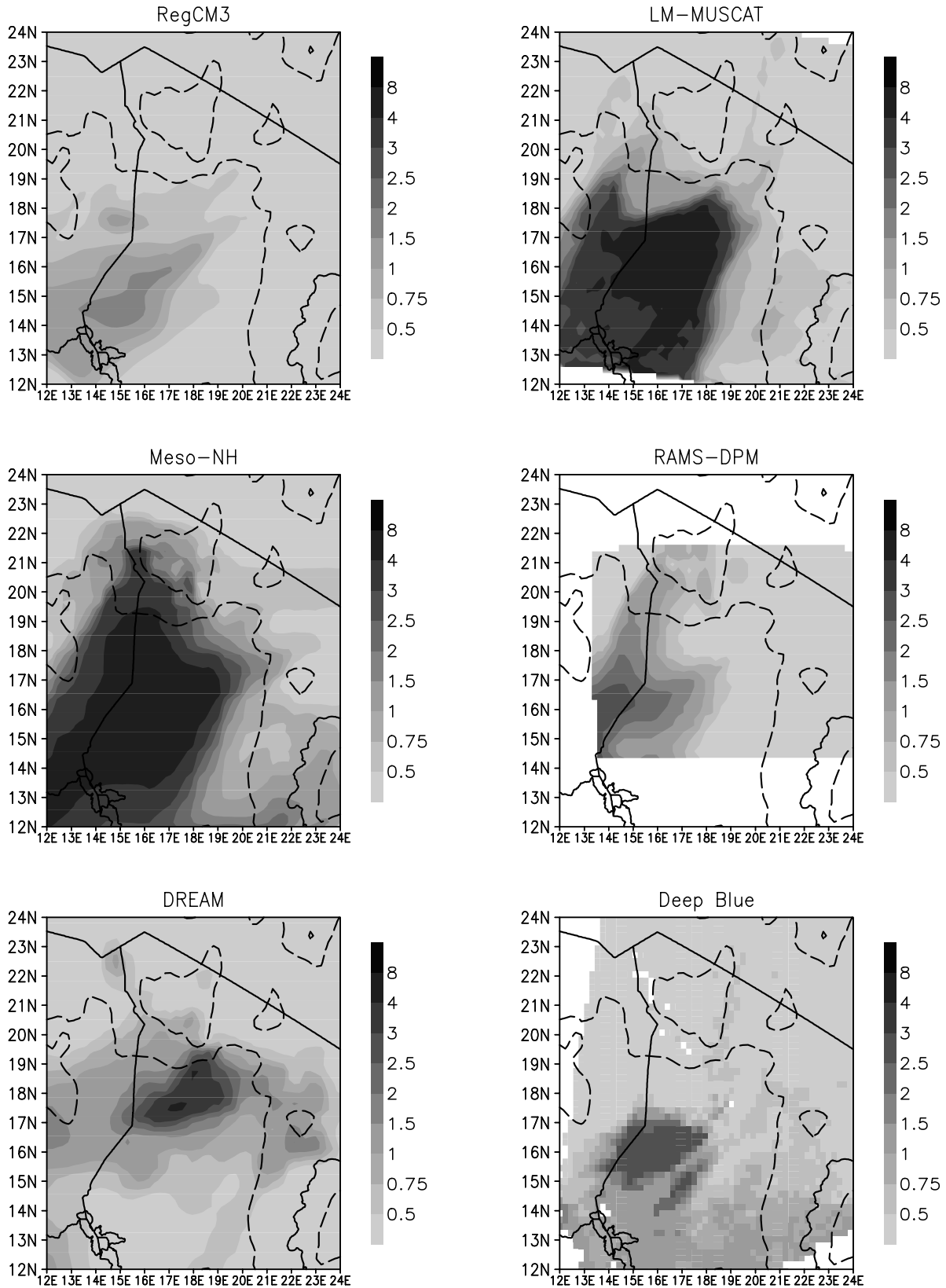
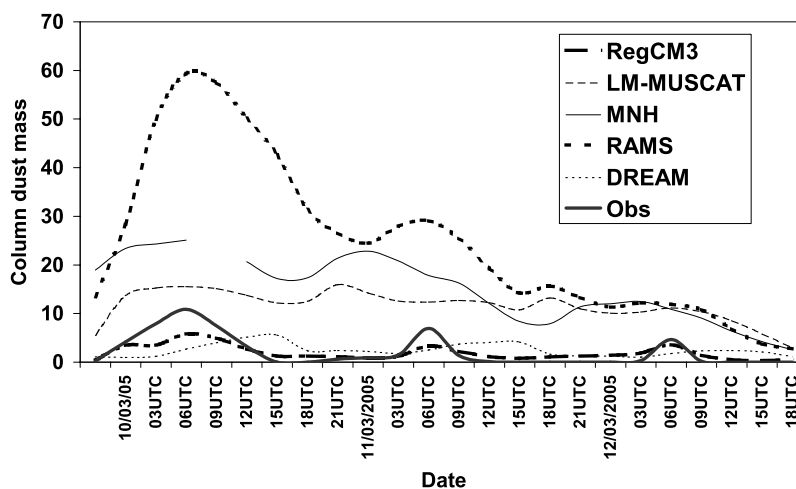


Figure 11. As in Figure 10 except for 1200 UTC on 12 March 2005.





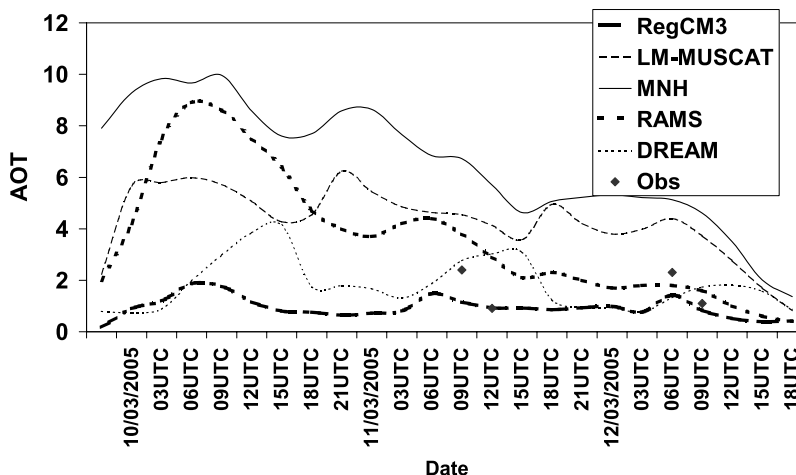
**Figure 12.** Time series of model simulated vertically integrated dust mass ( $\text{g m}^{-2}$ ) at the location of Chicha.

3.0. At Chicha limited observations of AOT were made during the afternoons of 11 and 12 March. On both days the decline in wind speeds after midday was such that transport exceeded local emission (which ceased in early afternoon) and AOT declined rapidly (Figure 13). All models except DREAM simulate this decline but the LM-MUSCAT and Meso-NH models overestimate AOT by about a factor of 2 while RegCM3 AOT is about 50% of observed. DREAM and RAMS-DPM provide a close agreement with observed AOT at Chicha during 11–12 March. Overall, the combined in situ and satellite observations indicate overestimation of AOT by Meso-NH and LM-MUSCAT and underestimation by RegCM3. Some of the overestimation by Meso-NH may result from use of a suboptimal  $\alpha$  parameter derived from model “tuning” experiments conducted at lower resolution. Application of an updated  $\alpha$  parameter would result in AOT estimates about one third lower (P. Tulet, personal communication, 2008). It is interesting to note that the ratio of model minimum to model maximum mean column dust mass (Table 3) is about twice that for AOT (Table 4) reflecting differences in the size distribution and the optical properties of the dust aerosols in the models. AOT calcu-

lations in the models are based on the full size range of the dust particles.

**4.2.4. Long-Range Dust Transport**

[26] Previous studies, notably that of *Koren et al.* [2006], have clearly indicated substantial long-range transport of dust from the Bodélé. The impact of this on climate and remote ecosystems will depend on the mass flux from the region. Here, we derive the mass flux of dust from the region estimated by each model from the product of the wind speed and dust concentration in each model layer, summed over the atmospheric column. Dust transport is dominated by a westward flux in the Harmattan northeasterly flow such that in the vertical integration calculation at levels where the zonal wind component is eastward the dust flux is considered to be negative. The net flux is summed over the latitude range  $12.5^{\circ}\text{N}–21^{\circ}\text{N}$ . Therefore, the result represents the net westward export of dust mass from the entire region toward West Africa and the Gulf of Guinea. Estimates of westward dust flux over the 3-day period leaving Chad at  $16^{\circ}\text{E}$  range widely from  $0.44 \text{ Tg d}^{-1}$  from RegCM3 to  $12.8 \text{ Tg d}^{-1}$  from RAMS-DPM (Table 5).



**Figure 13.** Time series of model simulated and observed AOT at the location of Chicha.

[27] *Koren et al.* [2006] estimate the long-range dust flux from the Bodélé by identifying dust plumes from MODIS satellite imagery and determining their average dimensions after 1 day when the plume lies to the west of Lake Chad. The dust mass within the plume was then derived from the product of plume area and column integrated dust mass, estimated from the satellite MISR-derived AOT. The average plume mass over the period October 2003 to October 2004 was estimated to be  $0.77 \text{ Tg d}^{-1}$ . The same method is applied to the BoDEx dust event except that we consider a greater range of uncertainty in the estimates associated with (1) satellite AOT, by considering contrasting estimates derived from MISR and the “Deep Blue” algorithm applied to MODIS data, and (2) the conversion of AOT to dust mass, by using the contrasting values of *Koren et al.* [2006] ( $2.7 \text{ g}^{-3}$ ) and *Todd et al.* [2007] ( $1.9 \text{ g}^{-3}$ , updated using Sun photometer observations at the AERONET (AERosol RObotic NETwork, *Holben et al.* [1998]) station at Maine Soroa ( $12.02^\circ\text{E}$ ,  $13.22^\circ\text{N}$ ), downwind from the Bodélé). The resulting estimated range of the dust mass flux for the emission events of 10–11 March 2005 is  $0.9\text{--}2.9 \text{ Tg d}^{-1}$ , not accounting for uncertainty in the estimates of dust area. Assuming these “1-day-old” plumes identified from MODIS are centered on  $13^\circ\text{E}$  then the equivalent estimates for the models are  $0.299$ ,  $2.31$ ,  $2.76$  and  $0.54 \text{ Tg d}^{-1}$  from the RegCM3, LM-MUSCAT, Meso-NH and DREAM models, respectively (Table 5), and we might expect higher values from RAMS-DPM had data been available at  $13^\circ\text{E}$ . As such, the “observational” estimates using an updated method of *Koren et al.* [2006] lie within the model range.

## 5. Discussion and Conclusions

[28] A dust model intercomparison has been undertaken for a 3-day dust event over the Bodélé depression, the world’s greatest single dust source region. The meteorology during the event was typical of wintertime dust events. The results here are not strictly intended to determine which models are “better” than others, and no single model performs better than the others in all aspects. Results show that near-surface wind speeds are underestimated by up to 50% and the out-of-phase relationship of the diurnal cycles of the LLJ and surface winds is not well resolved. The pronounced diurnal cycle in surface and lower level winds is present in most models and results in pulsed dust emission as in observations, but peak winds exhibit slight phase shift relative to observations in all but one model. In a comparable study of dust models over Asia, *Uno et al.* [2006] noted that for most locations wind speed observations lie within the range of model results. Underestimation of winds over the Bodélé may result from poor representation of surface and boundary layer conditions and processes highlighting the need for improved configuration of Regional Climate Models [*Todd et al.*, 2008]. Given that in most dust schemes the dust flux is proportional to the third power of friction velocity then small errors in wind speeds can result in large dust flux uncertainties. In this study, generally, the models with highest wind speeds (RAMS and Meso-NH) do tend to produce the highest dust fluxes. However, the substantial differences in dust fluxes between models with similar wind magnitudes (e.g., Meso-NH, LM-MUSCAT and RegCM3) serve to highlight the

importance of surface factors such as soil texture, soil size distribution and surface roughness in determining emission, as well as the calibration of the  $\alpha$  factor in the emission equation (equation (1)).

[29] Models generally show good representation of Bodélé dust plumes over the period 10–12 March 2005 and the temporal variability. All have emission in the wider Bodélé region. At local scales, however, the sources of dust emission show variability between models and RAMS-DPM and LM-MUSCAT show emission that is rather too spatially widespread, relative to observations. RegCM3 produces more localized emission from regions where, coincidentally, the soil texture map indicates pure sand soils. In reality, however, within the Bodélé depression the diatomite sediment has a range of size distributions and the soil data sets in the models do not represent this feature well. Moreover, the sandblasting emission scheme of the models is only partially applicable to the diatomite sediment. As such, the model representations of dust plumes in the region are rather crude representations of the complexity of actual emission process in the diatomite sediment of the Bodélé. It may be that the models based on the *Marticorena and Bergametti* [1995] scheme simulate reasonable dust emission from the Bodélé for the wrong reasons, i.e., because the soil maps erroneously indicate sandy soils. Nevertheless, for the purposes of simulating long-range transport and climate impact the models may be acceptable.

[30] The range of model estimates of quantities (for which we have estimates from all models, specifically surface dust concentration, vertically integrated dust mass and AOT) is approximately 1 order of magnitude. In terms of the total dust mass exported from the Bodélé depression the model estimates range over a factor of about 30. It is reassuring, however, to note that observational estimates of the equivalent quantities from a range of sources do lie within the model range. It should be noted that the model estimates presented here do refer to a common dust particle size range ( $0.1\text{--}10 \mu\text{m}$  diameter) such that these differences between the models do not result from certain models having a wider or smaller range of particle sizes.

[31] The degree of model uncertainty reported here is actually very similar to the range of mean emission and typical dust concentration from nine models compared for two dust periods over East Asia by *Uno et al.* [2006]. For subregions within East Asia the range is about 1 order of magnitude. It seems, therefore, that the current degree of uncertainty of dust models in regional studies over relatively short periods of days to weeks may be about 1 order of magnitude. This is somewhat greater than the uncertainty associated with model estimates of long-term mean global dust emission [*Zender et al.*, 2004], suggesting that model over and underestimation in various regions may cancel out to a certain degree. All the models here have been “calibrated” in previous studies to produce sensible results, usually with respect to satellite or surface based estimates of AOT. This calibration tends to be conducted at the larger subcontinental scale, over North Africa in most cases. Therefore, the degree of model divergence in this case study may be surprising. As such, it is logical to conclude from this paper that model divergence may increase at smaller space/timescales. The BoDEx case study provides a particularly rigorous test of model performance, especially

Table A1. Dust-Climature Model Details<sup>a</sup>

	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Reference	Zakay <i>et al.</i> [2006]	Tegen <i>et al.</i> [2006]; Heinold <i>et al.</i> [2007]	Lafore <i>et al.</i> [1997]; Tulet <i>et al.</i> [2005]; Grini <i>et al.</i> [2006]	Cotton <i>et al.</i> [2003]; Bouet <i>et al.</i> [2007]	Nickovic <i>et al.</i> [2001]
Boundary conditions	NCEP reanalysis	global forecast model of the Deutscher Wetterdienst (DWD) corners at 12.60°N, 11.63°E; 10.30°N, 25.00°E; 22.17°N, 29.08°E; 25.09°N, 12.10°E	ECMWF	ECMWF	NCEP analysis 1°
Domain	3°N–25°N, 3°E–25°E	8.3°N–25.3°N, 9.1°E–26.9°E	8.3°N–25.3°N, 9.1°E–26.9°E	14.4°N–21.5°N; 13°E–25°E	0°–28°N; 0°–28°E
Dust-radiative feedback?	yes	yes	yes	no	no (for this version)
Simulation period	8–13 March 2005	1–13 March 2005	1–13 March 2005	5–15 March 2005	1–16 March 2005
Horizontal resolution (km)	20	7	20	10	~26
Vertical levels	23	40	72	30	24
Height of lowest layer (m)	40	68	10	10	86
PBL and surface schemes	K-theory eddy diffusion with nonlocal closure [Holltag and Boville, 1993]	Prognostic TKE-based turbulence closure scheme (order 2.5) and surface layer scheme [Raschendorfer, 2001]	The 1.5-order closure turbulence scheme [Cuxart <i>et al.</i> , 2000] and the Interaction Soil Biosphere Atmosphere (ISBA) scheme [Noilhan and Mahouf, 1996]	K-theory eddy diffusion [Smagorinsky, 1963] modified using Lilly [1962] and Hill [1974]; surface scheme of Louis [1979]	prognostic TKE - Mellor-Yamada 2.5, and Monin-Obukhov similarity theory in the surface layer
Emission scheme <sup>b</sup>	U* <sup>3</sup> (MB 95)	U* <sup>3</sup> (MB 95)	U* <sup>3</sup> (MB 95)	U* <sup>3</sup> (MB 95)	U* <sup>3</sup> [Shao <i>et al.</i> , 1993] with viscous sublayer concept [Janjic, 1994; Nickovic <i>et al.</i> , 2001] <sup>c</sup>
$\alpha$ parameter in equation (1) (cm <sup>-1</sup> )	N/A	$1.8 \times 10^{-6}$	$1.8 \times 10^{-3}$	$1.0 \times 10^{-6}$ – $7.5 \times 10^{-6}$	N/A
Land use data	18 classes on 1° grid based on work of Mathews [1984] and USGS	27 potential vegetation types from equilibrium terrestrial biogeography model BIOME4 of Kaplan [2001]	ECOLIMAP [Masson <i>et al.</i> , 2003]	LEAF-2 (Land Ecosystem- Atmosphere Feedback model [Walke <i>et al.</i> , 2000])	Olson World Ecosystems/ (SSIB) vegetation types [Wilson and Henderson-Sellers, 1985]
Surface roughness	one value per land class (0.1m)	from POLDER satellite data 1/15° [Marricorena <i>et al.</i> , 2004]	one value per land class ( $1 \times 10^{-4}$ m)	from POLDER satellite data 1/15° [Marricorena <i>et al.</i> , 2004] Laurent <i>et al.</i> [2008b]	Zoumakis and Kelessis [1991]
Soil texture data	12 texture classes based on FAO [Zobler, 1986] and USDA data [Dickinson <i>et al.</i> , 1993]	Texture classes from Zobler [1986] except in preferential source regions where uniform texture is applied	IGBP Global Soil Data Task		From FAO and Zobler [1986]

Table A1. (continued)

	RegCM3	LM-MUSCAT	Meso-NH	RAMS-DPM	DREAM
Soil aggregate size distribution	3-mode, lognormal distribution associated with each texture class	4-mode, lognormal particle size distribution of clay, (2 $\mu\text{m}$ ), silt (15 $\mu\text{m}$ ), fine sand (158 $\mu\text{m}$ ), and coarse sand (720 $\mu\text{m}$ ) associated with each texture class	3-mode, lognormal distribution associated with each soil type.	3-mode, lognormal distribution associated with each soil type.	4-mode size distribution (0.73, 6.1, 18 and 38 $\mu\text{m}$ radii) associated with each texture class
Preferential dust sources	no	Yes <sup>d</sup>	no	no	no
Soil particle density ( $\text{g m}^{-3}$ )	2.65	2.1 <sup>e</sup>	2.5	2.65	2.5–2.65
Number of dust bins	4 size bins for dust transport derived from 3 lognormal dust emission modes	5 size bins for dust transport with lognormal distribution	3 size modes with lognormal distribution	20 size bins with uniform distribution	4 size bins with lognormal distribution
Dust particle size range ( $\mu\text{m}$ , diameter) <sup>f</sup>	0.1–20	0.2–50	0.1–10	0.1–26	1.46–76

<sup>a</sup>N/A denotes not applicable.

<sup>b</sup>In emission scheme,  $u^*$  is the wind friction velocity and MB95 refers to *Marticorena and Bergametti* [1995].

<sup>c</sup>DREAM derives the dust fluxes between the surface and a thin viscous sublayer and then between this sublayer and the lowest model layer, using concentration as a lower boundary condition. As such, DREAM dust fluxes are not directly comparable emission fluxes from the other models.

<sup>d</sup>In LM-MUSCAT, preferential source regions correspond to the distribution of paleo and temporal lake beds computed with the water routing and storage model HYDRA [Coe, 1998]. For the Bodélé depression this area corresponds to the extent of the paleo lake Mega-Chad, whose deposits account for today's strong dust activity. In these preferential source regions, soil texture is set to a uniform 100  $\mu\text{m}$  to ensure high emission.

<sup>e</sup>This value was chosen to be consistent with field observations of surface sediment samples from BoDEx 2005.

<sup>f</sup>In the results presented here the dust quantities are derived over a common size range of 0.1–10  $\mu\text{m}$ .

given the particular soil and emission conditions in the region. The results highlight that model calibration which optimizes performance at larger scales may not be appropriate for specific regions.

[32] Further model intercomparison is necessary to determine the extent to which model divergence is associated with difference calibration of emission equation coefficients. In many cases, the model calibration or “tuning” is conducted by varying the  $\alpha$  parameter in equation (1) such that model estimates of AOT are close to observations. This is not unreasonable since lack of accurate soil information means that direct application of the empirical relationship between clay fraction and  $\alpha$  [Gillette, 1979] is problematic [Zender *et al.*, 2003]. However, it should be noted that model simulations over the Sahara by Laurent *et al.* [2008b] in which  $\alpha$  is derived from soil texture provide acceptable results. In any case, model “tuning” by varying the  $\alpha$  parameter (or the  $c$  constant in equation (2) for DREAM) to optimize model AOT represents a “catch-all” correction which accounts for uncertainty in all aspects of the model, including meteorology, surface roughness, soil texture, emission size distribution and optical properties of the dust as well as the deposition schemes.

[33] Model intercomparison can be more useful to the research community when we are able to partition model error associated with various sources within the model. This process is hindered in the present study by differences in model configurations and model “tuning.” Thus, a future model intercomparison exercises should involve much more tightly constrained model configurations such that the sensitivity of model estimates to individual factors such as meteorological model performance, model resolution, boundary conditions, soil characteristics, dust physical and optical properties and the various parameterizations within the emission (and deposition) schemes can be quantified. This requires a coordinated set of multiple model experiments using a consistent set of predefined model configurations in which the domain, horizontal and vertical resolution, lateral boundary conditions, soil data, emission scheme coefficients, and aerosol size and optical properties are carefully defined.

[34] In addition, this program requires adequate observations for comparison. At present, the observational data on terms of dust emission, concentration, mass and flux are neither sufficiently accurate nor extensive enough in space and time to provide a suitable estimate of “truth” for absolute comparison with models. Therefore, the study highlights the pressing need for more comprehensive in situ observations in dust source regions like the Bodélé depression covering dust concentration, physical and optical properties of dust, as well and boundary layer meteorology, which are more extensive in space and time. Further detailed observations, perhaps using wind tunnel experiments, of the emission processes in the Bodélé are also required. Only when such data are available will we be able to really constrain model uncertainty and to identify individual sources of error. In the meantime, however, model experiments can comprehensively consider model sensitivity.

[35] Moreover, there is a clear need for further research into other source regions in the Sahara, and elsewhere. This should include identification of the key preferential dust

source regions in the Sahara, observation of the dust emission processes and determination of the associated soil characteristics at a resolution higher than that available from existing soil data sets. While satellite-derived estimates of surface roughness are available [e.g., Marticorena *et al.*, 2004] used here in LM-MUSCAT and RAMS-DPM dust models tend to rely on maps of soil texture (from which the dry soil size distribution is derived) with questionable accuracy in remote regions of the Sahara desert. There is a clear requirement for improved soil texture data for the Sahara that can resolve features like the diatomite deposits in the Bodélé depression. The results here demonstrate that the model dust emission is strongly associated with sandy soils such that the emission scheme itself may require modification in preferential source regions with similar characteristics to the Bodélé. It is important to determine to what extent the Bodélé is unusual among the source regions of the Sahara in this regard.

## Appendix A

[36] This study involves a model intercomparison “of opportunity” such that the dust models have differing physics options and configurations. In Table A1 we provide details of the salient model features.

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- J. M. Baldasano and C. Perez, Earth Sciences Department, Barcelona Supercomputing Center, Barcelona, E-08034 Spain. (Jose.baldasano@bsc.es; carlos.perez@bsc.es)
- D. Bou Karam, Service d'Aéronomie, Institut Pierre-Simon Laplace, Université Pierre et Marie Curie, Paris, France. (diana@aero.jussieu.fr)
- C. Bouet and G. Cautenet, Laboratoire de Météorologie Physique, Université Blaise Pascal, F-63170 Aubière, France. (bouet@lisa.univ-paris12.fr; G.Cautenet@opgc.univ-bpclermont.fr)
- C. Cavazos and M. C. Todd, Department of Geography, University College London, Gower Street, London WC1E 6BT, UK. (ccavazos@geog.ucl.ac.uk; m.todd@geog.ucl.ac.uk)
- B. Heinold and I. Tegen, Leibniz Institute for Tropospheric Research, D-04318 Leipzig, Germany. (heinold@mail.tropos.de; itegen@tropos.de)
- I. Koren, Weizmann Institute, Rehovot 76100, Israel. (ilan.koren@weizmann.ac.il)
- F. Solmon, Laboratoire d'Aérodologie, Université Paul Sabatier, F-31077 Toulouse, France. (Fabien.solmon@aero.obs-mip.fr)
- P. Tulet, GMEI, CNRM, Météo France, F-31057 Toulouse, France. (Pierre.Tulet@meteo.fr)
- R. Washington, Oxford Centre for the Environment, University of Oxford, Oxford OX1 3QY, UK. (Richard.washington@geog.ox.ac.uk)
- A. Zakey, Abdus Salam International Centre for Theoretical Physics, I-34014 Trieste, Italy. (azakey@ictp.it)