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# On the ability of chemical transport models to simulate the vertical structure of the $N_2O$ , $NO_2$ and $HNO_3$ species in the mid-latitude stratosphere

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Abstract. In this paper we study the impact of the modelling of  $N_2O$  on the simulation of  $NO_2$  and  $HNO_3$  by comparing in situ vertical profiles measured at mid-latitudes with the results of the Reprobus 3-D CTM (Three-dimensional Chemical Transport Model) computed with the kinetic parameters from the JPL recommendation in 2002. The analysis of the measured in situ profile of N<sub>2</sub>O shows particular features indicating different air mass origins. The measured N<sub>2</sub>O, NO<sub>2</sub> and HNO<sub>3</sub> profiles are not satisfyingly reproduced by the CTM when computed using the current 6hourly ECMWF operational analysis. Improving the simulation of N<sub>2</sub>O transport allows us to calculate quantities of NO<sub>2</sub> and HNO<sub>3</sub> in reasonable agreement with observations. This is achieved using 3-hourly winds obtained from ECMWF forecasts. The best agreement is obtained by constraining a one-dimensional version of the model with the observed N<sub>2</sub>O. This study shows that the modelling of the NO<sub>v</sub> partitioning with better accuracy relies at least on a correct simulation of N<sub>2</sub>O and thus of total NO<sub>v</sub>.

#### 1 Introduction

The partitioning of individual nitrogen species within reactive nitrogen family  $(NO_y)$  is a very important factor in the stratospheric chemistry of ozone. The photo-oxidation of N<sub>2</sub>O is the main source of NO<sub>y</sub> species in the stratosphere:

$$N_2O + O(^1D) \to N_2 + O_2$$
 (1a)

$$\rightarrow 2NO$$
 (1b)

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Reaction (1b) is the dominant path for the production of NO<sub>x</sub> species (e.g. McElroy and McConnell, 1971). NO<sub>x</sub> (NO+NO<sub>2</sub>), the most reactive form of the NO<sub>v</sub> compounds, are involved in the main processes controlling the ozone budget. They both account for a major part of ozone reduction in the middle and upper stratosphere through rapid catalytic cycles (e.g. McElroy et al., 1992; Osterman et al., 1997; Brühl et al., 1998) and moderate the ozone-destroying catalytic cycles involving halogen and hydrogen radicals (e.g. Salawitch et al., 1994; Wennberg et al., 1994). NO<sub>x</sub> are chemically linked to nitric acid HNO3 which is the main reservoir of the  $NO_v$  family below  $\sim 30$  km and are primarily produced by photolysis of HNO<sub>3</sub>. The partitioning of NO<sub>v</sub> species, and in particular NO2 and HNO3, is also affected by heterogeneous reactions on sulphate aerosols that convert active nitrogen to HNO<sub>3</sub> (e.g. McElroy et al., 1992; Fahey et al., 1993).

Nevertheless the quantitative understanding of the partitioning of the NO<sub>v</sub> species in the models remains limited. Several studies have shown significant discrepancies between computed and observed partitioning of NO<sub>v</sub> species depending on the location (mid- or high latitudes) and time (day or night, period of the year) (e.g. Sen et al., 1998; Gao et al., 1999; Payan et al., 1999; Wetzel et al., 2002; Stowasser et al., 2002, 2003). Observations of the various NO<sub>v</sub> species such as HNO<sub>3</sub> and NO<sub>2</sub> are not satisfyingly reproduced by chemistry-transport models at mid-latitudes with in particular recurrent underestimations of NO2 mixing ratios (Wetzel et al., 2002; Stowasser et al., 2003). The different compounds among the NO<sub>v</sub> family are also represented as ratios to separate dynamical from chemical effects on the NOv partitioning and to reduce the influence of missing amounts of total NO<sub>v</sub> in the models. Simulations of these ratios generally appear to be more satisfactory (e.g. Kondo et al., 2000;

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Wetzel et al., 2002) even though Stowasser et al. (2003) have presented  $NO_2/NO_y$  ratios strongly underestimated by model results in the lower stratosphere.

The SPIRALE (French acronym for "Spectroscopie In-Frarouge par Absorption de Lasers Embarqués") balloonborne instrument is able to measure in situ profiles of NO2 and HNO<sub>3</sub> and of their precursor N<sub>2</sub>O with a high vertical resolution. Such in situ measurements appear to be more able to characterize local processes than remote sensing observations using lines of sight of a few hundred of kilometres. We report here measurements of these species at mid-latitudes on 2 October 2002 in the frame of the first validation campaign of the ENVISAT satellite. In this paper we try to balance the dynamical and chemical effects on the modelling of the NO<sub>v</sub> partitioning. First we assess the influence of the 3-D winds used in the Reprobus CTM on its ability to reproduce the NO<sub>v</sub> partitioning. Then a one-dimensional (1-D) version of the model is used to test the impact of a correct initialisation of N<sub>2</sub>O on the simulation of the chemistry controlling the NO<sub>v</sub> partitioning. It is the first modelling study of the in situ simultaneous observations by SPIRALE of NO2 and HNO<sub>3</sub> with a CTM.

#### 2 The SPIRALE balloon-borne instrument

A detailed description of the instrumental characteristics of SPIRALE and of its operating mode can be found in Moreau et al. (2005). To summarize, the SPIRALE balloon-borne instrument performs in situ simultaneous measurements of several long-lived and short-lived chemical species from the tropopause up to 40 km. It uses six tunable salt laser diodes in the mid-infrared domain (3  $\mu$ m to 10  $\mu$ m). The six laser diodes are cooled and the beams are injected into a multipass Heriott cell located under the gondola. The cell (3.5 m long) is deployed during the flight above the tropopause. Eighty six reflections occur, giving a 300 m optical path. Species concentrations are retrieved from direct absorption, by fitting experimental spectra with spectra calculated using HITRAN 2001 database (Rothman et al., 2003). The instrument provides measurements with a vertical resolution of a few meters.

The global uncertainties for the volume mixing ratios of  $N_2O$ ,  $HNO_3$  and  $NO_2$  have been assessed by taking into account the random errors and the systematic errors, and combining them as the square root of their quadratic sum (Moreau et al., 2005). In brief, there are two important sources of random errors: (1) the fluctuations of the laser background emission signal and (2) the signal-to-noise ratio. These error sources are the main contributions for  $NO_2$  giving a total uncertainty of 30% at the lower altitudes (around 15 km), reduced to 20% around 20 km, and even to 5% at higher altitudes (around 30 km). For HNO<sub>3</sub> these random errors are less significant but two sources of systematic errors have to be considered: the laser line width (an intrinsic

characteristic of the laser diode) and the non-linearity of the detectors resulting in an uncertainty of 20% on the whole profile. Concerning the N<sub>2</sub>O species which is abundant and measured using an efficient detection system, the overall uncertainty is of 3% over the whole vertical profile. With respect to the above errors, systematic errors on spectroscopic data (essentially molecular line strength and pressure broadening coefficients) are considered to be negligible for these three well studied species.

SPIRALE is used routinely, in particular as part of European and satellite validation campaigns. On 2 October 2002, during the ENVISAT validation campaign, the SPIRALE flight was conducted during the morning at mid-latitude from the Aire sur l'Adour launch base (France,  $43.7^{\circ}$  N,  $0.3^{\circ}$  W). Unfortunately measurements from the instruments onboard ENVISAT were not satisfyingly close in time and space to the SPIRALE observations for direct comparison and validation. In our study we analyse the measurements of N<sub>2</sub>O, HNO<sub>3</sub> and NO<sub>2</sub> profiles obtained during the ascent of the balloon between 07:30 UT and 08:30 UT. The measurement position remained rather constant during the ascent with a displacement of the balloon from  $43.7^{\circ}$  N–0.18° W to  $43.60^{\circ}$  N–0.16° E. The data used in this study are averaged over a vertical range of 250 m.

#### 3 Model calculations

The Reprobus chemical-transport model (Lefèvre et al., 1994, 1998) has been widely used in previous studies of the stratospheric chemistry (e.g. Lefèvre et al., 1994; Deniel et al., 1998; Hoppel et al., 2002). The model is designed to perform annual simulations as well as detailed process studies. It computes the evolution of 55 species by means of about 160 photolytic gas-phase and heterogeneous reactions, with a time step of 15 min in this study. A semi-lagrangian code transports 40 species or chemical families, typically long-lived tracers but also more unstable compounds which may have a long lifetime in darkness. Kinetics parameters used in the present study are based on the most recent data in general taken from Sander et al. (2003). The new laboratory measurements of photodissociation cross-sections of HO<sub>2</sub>NO<sub>2</sub> both in the UV (Knight et al., 2002) and in the near IR (Roehl et al., 2002) have been included in the photodissociation calculations. The heterogeneous chemistry module includes reactions on liquid aerosols. Their surface area densities are inferred from SAGE II measurements (Thomason et al., 1997).

Reprobus was integrated from 1 April 2002 to 15 October 2002. The model extends from the surface up to 0.1 hPa on 42 levels, resulting in a vertical resolution of about 1.3 km in the lower stratosphere. The horizontal resolution is  $2^{\circ}$  latitude by  $2^{\circ}$  longitude. Temperature, winds and ground pressure are specified from the three-dimensional (3-D) European Centre for Medium-Range Weather Forecast (ECMWF)

meteorological data. The ozone field was initialized on 1 April 2002 from the ECMWF ozone analysis. Other species were initialized from an April zonal mean computed from a long-term simulation of Reprobus.

#### 4 Comparisons between SPIRALE measurements and 3-D calculations

#### 4.1 N<sub>2</sub>O profile

#### 4.1.1 Measurement

The in situ profile of  $N_2O$  measured by SPIRALE (Fig. 1) presents different features which suggest that the instrument observed air masses of different origins. In particular a 5 km layer with a minimum value of 50 ppbv at 27 km is clearly present between 23 and 29 km.

A detailed study based on the SPIRALE observations of N<sub>2</sub>O and CH<sub>4</sub> has been made to investigate the dynamical state of the stratosphere during this flight (Huret et al., 2006). Using a combination of [CH<sub>4</sub>]:[N<sub>2</sub>O] correlations and potential vorticity (PV) maps calculated by the MIMOSA contour advection model (Hauchecorne et al., 2002) this study shows a perturbed situation probably due to the vertical wind shear occurring during the first stage of the polar vortex formation. From the [CH<sub>4</sub>]:[N<sub>2</sub>O] correlations, mid-latitude air was diagnosed between about 16 and 23 km whereas the air masses sampled above 29 km mainly originated from tropical latitudes. This is consistent with the location of the instrument with respect to the dynamical barrier existing in this region: the MIMOSA model calculates potential vorticity values in the 43–50 pvu range at 550 K (about 22.5 km) and close to 149 pvu at 810 K (about 30 km) at the location of the measurement at 06:00 UT (Figs. 2a and c). Between 23 and 29 km the N<sub>2</sub>O profile is clearly non-monotonic (Fig. 1) and the measurements appear to be located very close to the maximum PV gradient region such as at 625 K (Fig. 2b). The vertical profiles of the CH<sub>4</sub> and N<sub>2</sub>O tracers are associated with a significantly different [CH<sub>4</sub>]:[N<sub>2</sub>O] correlation with respect to Michelsen et al. (1998) correlation curves (Huret et al., 2006). This gives an indication that mixing processes have probably occurred, similarly to the mixing event processes occurring inside the Arctic vortex and across the vortex edge (Plumb et al., 2003; Konopka et al., 2004) or to the mixing of polar vortex air into middle latitudes detected from tracer-tracer scatter plots (Waugh et al., 1997). The analysis of such a mixing event is out of the scope of our paper.

Similar conclusions can be derived from the comparisons performed by Urban et al. (2005) between SPIRALE measurements and the observations of the SMR instrument onboard the Odin satellite (e.g. Murtagh et al., 2002). This study shows that averaging Odin/SMR measurements over  $27.5-42.5^{\circ}$  N (mostly tropical air) and  $42.5-55^{\circ}$  N (mostly mid-latitude air) gives N<sub>2</sub>O amounts similar to those ob-



Fig. 1. In situ volume mixing ratios (VMRs) of  $N_2O$  measured by the SPIRALE instrument on 2 October 2002 at mid-latitudes (black line) and compared to the VMRs simulated by the Reprobus 3-D CTM at the time of SPIRALE observations (between 07:30 UT and 08:30 UT). The profiles simulated at the two model grid points closest to the measurement positions have been averaged (red line). The model is computed using ECMWF operational analysis data. For more clarity the measurement error bars are represented every 100 m.

served by SPIRALE above 29 km and below 23 km, respectively. Between 23 and 29 km the  $N_2O$  profile measured by SPIRALE is located between the two averaged profiles observed by Odin/SMR, suggesting a mixing event.

#### 4.1.2 Model calculations

CTM calculations were driven using the 3-D 6-hourly ECMWF operational analysis for winds, temperature and ground pressure recurrently employed in modelling studies (e.g. Lefèvre et al., 1998; Wetzel et al., 2002; Stowasser et al., 2003; Ricaud et al., 2005). The SPIRALE measurements were located between the  $42^{\circ}$  N–0° E and  $44^{\circ}$  N–0° E model grid points and remain close to these positions during the balloon ascent (from  $43.7^{\circ}$  N–0.18° W to  $43.60^{\circ}$  N–0.16° E). The profiles simulated at these two positions have been averaged for comparisons with the observations.

Figure 1 compares the observed profile of  $N_2O$  to the model result obtained at the grid point closest to the measurement position. The model slightly underestimates the observations below 20 km but discrepancies are mainly observed above 25 km: Reprobus overestimates SPIRALE measurements of  $N_2O$  above 30 km, and does not reproduce the structure at 27 km. The Reprobus overestimation of  $N_2O$  above 30 km is in contrast to former KASIMA 3-D CTM comparisons with the MIPAS-B balloon-borne instrument showing a distinct underestimation of  $N_2O$  above 28 km (see Stowasser et al., 2002). Other measurements-model comparisons have revealed overestimations of the calculated amounts of  $N_2O$ .  $N_2O$  mixing ratios measured by the MIPAS-B instrument at



Fig. 2. Potential Vorticity (PV) isocontours calculated by the MIMOSA high-resolution advection model at 06:00 UT on 2 October 2002 at: (a) 550 K (about 22.5 km); (b) 625 K (about 25 km); and (c) 810 K (about 30 km). The red regions denote air masses from tropical latitudes. The black cross denotes the SPIRALE instrument location.

mid-latitudes during summer time are overestimated by the KASIMA model below 30 km (Wetzel et al., 2002). Comparisons between Reprobus and the Odin satellite measurements have shown that the model overestimated the N2O amounts observed by Odin/SMR at mid-latitudes in mid-September 2002 (Ricaud et al., 2005). Similar disagreements are observed with preliminary measurements of the MIPAS instrument on board the ENVISAT satellite for the same period (R. Runhke, personal communication). This notorious problem appears to be related to a too strong vertical transport above the Equator when using ECMWF operational analysis winds which provide too large quantities of N<sub>2</sub>O in the tropical middle stratosphere (Ricaud et al., 2005). Then these amounts reach mid-latitudes by the Brewer-Dobson circulation and quasi-horizontal transport from the tropical latitudes (Holton et al., 1995).

#### 4.2 NO<sub>y</sub> partitioning

The diurnal variation of the species has been taken into account in the simulated profiles hereafter described. The solar zenith angle varies from about  $64^{\circ}$  to  $74^{\circ}$  between 07:30 and 08:30 UT at the studied locations resulting in a variation of 3–6% for NO<sub>2</sub> above 25 km. Figure 3 shows the comparisons between the observed profiles of NO<sub>2</sub> and HNO<sub>3</sub> and the model calculations. Significant underestimation is visible on

the simulated profile of NO<sub>2</sub> above 22 km (30, 37 and 15% lower than the observation at 23, 27 and 31 km, respectively, if not accounting for error bars) where N<sub>2</sub>O is overestimated by the model. The model underestimates the measured mixing ratios of HNO<sub>3</sub> above 27 km or hardly reaches the error bars (13, 16 and 24% lower than the observation at 27, 28 and 31 km respectively). Simulated HNO<sub>3</sub> is overestimated below 22 km (70 and 25% higher than the observation at 18 and 20 km, respectively) where N<sub>2</sub>O is underestimated by the model. For other altitudes Reprobus results stay within the error bars of the SPIRALE measurements.

One must examine whether such underestimation of the mixing ratios of these two major species among the NO<sub>y</sub> family is related to the ability of the model to calculate the total amounts of NO<sub>y</sub>. Indeed 3-D CTMs may underestimate the total amount of NO<sub>y</sub> at mid-latitudes, as shown by the comparison between the KASIMA CTM and observations by the MIPAS-B instrument (Wetzel et al., 2002), which affects the simulation of the various NO<sub>y</sub> species. This result is supported by recent comparisons between different European models with preliminary MIPAS observations (Roland Ruhnke, personal communication). Since SPIRALE does not measure all of the NO<sub>y</sub> species (NO for this flight, N<sub>2</sub>O<sub>5</sub>, CIONO<sub>2</sub>, NO<sub>3</sub> and HO<sub>2</sub>NO<sub>2</sub> are missing), we have reconstructed the in situ profile of total NO<sub>y</sub> (hereafter NO<sub>y</sub>\*)



**Fig. 3.** Comparison between measured vertical profiles of NO<sub>2</sub> (**a**) and HNO<sub>3</sub> (**b**) on 2 October 2002 at mid-latitudes (black lines) and profiles simulated by the Reprobus 3-D CTM between 07:30 UT and 08:30 UT with a solar zenith angle varying from  $64^{\circ}$  to  $74^{\circ}$  (red lines). The model is computed using ECMWF operational analysis data.

using the compact N<sub>2</sub>O-NO<sub>v</sub> relations measured in different latitude bands. Figure 4 presents the NOv\* profile inferred from the 28–46° N correlation given by Michelsen et al. (1998). The profile from the 3-10° N latitude band correlation is also represented above 29 km where air masses appear to mainly originate from low latitudes (see Sect. 4.1.1). N<sub>2</sub>O and NO<sub>v</sub> are anticorrelated for N<sub>2</sub>O mixing ratios larger than about 50 ppbv. Subsequently, the comparison between the reconstructed and simulated profiles of NO<sub>v</sub> (Fig. 4) reveals a significant underestimation by Reprobus above 25 km where  $N_2O$  is overestimated by the model. Above 30 km the model particularly underestimates the NOv\* profile inferred from the 3-10° N N<sub>2</sub>O-NO<sub>v</sub> correlation. A slight overestimation is observed at some altitudes below 20 km where N<sub>2</sub>O is underestimated by the model. These two behaviours are the result of the discrepancies observed in Fig. 1. Any problem in the modelling of N<sub>2</sub>O has a direct impact on the calculated quantities of total NO<sub>v</sub>. The transport of too large quantities of the N2O precursor species at mid-latitudes stated above can result in a too weak production of NO<sub>v</sub> after its transport from tropical latitudes (e.g. Ricaud et al., 2005) which explains the low values for the upper part of the profile in Fig. 4.

To minimize the effect of the uncertainty in the total  $NO_y$  modelling we have compared the  $NO_2/HNO_3$  ratios rather than individual profiles (Fig. 5). In that case better results are obtained: the  $NO_2/HNO_3$  ratio is well reproduced by Reprobus above 27 km and between 17 and 22 km. However the simulated  $NO_2/HNO_3$  ratio barely reaches the error bars between 22 and 27 km (29 and 26% lower than the observation at 23 and 27 km, respectively) as a result of  $NO_2$  underestimation. From the  $NO_2/HNO_3$  ratio we can see in this altitude range that the particular structure attributed to a mixing event affects our ability to reproduce the observed  $NO_y$  partitioning. We conclude from Fig. 5 that the  $NO_2/HNO_3$ 



**Fig. 4.** Vertical profiles of reconstructed NO<sub>y</sub>\* compared to the Reprobus model output on 2 October 2002 (red line). Black curve: NO<sub>y</sub>\* profile inferred from the N<sub>2</sub>O-NO<sub>y</sub> correlation determined by Michelsen et al. (1998) in the 28–46° N latitude band. Orange curve: NO<sub>y</sub>\* profile inferred from the correlation valid in the 3–10° N latitude band (only represented above 29 km where the air seems to mostly originate from low latitudes).

partitioning seems to be correctly reproduced by the CTM at all altitudes except at those of the particular vertical structure. Note that the strong peaks of the  $NO_2/HNO_3$  ratio at 15 and 16.5 km cannot be reproduced by the model which could be due to its too low vertical resolution.

From this result it is worth trying to improve the simulation of the total  $NO_y$  profile. The modelling of  $N_2O$  relies mostly on transport. Thus, one can expect that a better simulation of the transport of  $N_2O$  can improve the agreement between measurements and calculations of the  $NO_2$  and  $HNO_3$ quantities.



Fig. 5.  $NO_2/HNO_3$  ratio measured by SPIRALE (black line) and calculated by Reprobus 3-D CTM between 07:30 UT and 08:30 UT (red line). The model is driven by ECMWF 6-hourly operational analysis.

#### 5 Sensitivity to the wind fields

#### 5.1 6-hourly and 3-hourly ECMWF winds

In the work of Legras et al. (2004) in situ measurements of N<sub>2</sub>O performed from the ER-2 aircraft are compared both to Reprobus 3-D outputs (computed at the time/positions of the flight track) and to reconstructed trajectories. Significant discrepancies are observed between the CTM and the reconstructed trajectories when computed with 6-hourly ECMWF operational analysis. These are characterized by strong variations and unrealistic structures apparently as a result of highfrequency fluctuations in the winds. A dramatic improvement is obtained by Legras et al. (2004) when using 3-hourly analysis to compute air mass trajectories and the resulting N<sub>2</sub>O distribution. We have tested this result by performing a second simulation using 3-hourly wind fields obtained by interleaving 3-h and 9-h forecasts between the ECMWF 6hourly operational analysis. Figure 6 compares N<sub>2</sub>O zonal means calculated by Reprobus when driven by 6-hourly and 3-hourly winds on 2 October 2002, at the time of SPIRALE observations. It can be seen that using 3-hourly winds leads to a slower circulation and a reduced vertical diffusion. This result is consistent with the analysis of Legras et al. (2004). It also supports the work of Stohl et al. (2004) who show that forecast winds are less diffusive than operational analysis.

#### 5.2 3-D CTM calculations using 3-hourly winds

SPIRALE observations are now compared to two additional Reprobus 3-D CTM simulations using different wind sources: 1) a simulation considering 3-hourly wind fields, interleaving operational analysis at 00:00, 06:00, 12:00, and 18:00 UT with forecasts valid at 03:00, 09:00, 15:00 and 21:00 UT, as in the work of Legras et al. (2004) (hereafter named Reprobus-3h-ope-for). 2) a simulation only considering forecasts at 3-h interval, based on the analysis at 00:00 and 12:00 UT (hereafter Reprobus-3h-for). The "baseline" simulation using 6-hourly operational analysis is referred as Reprobus-ope in the following. Initialising the transport every 3h has clearly allowed us to model more accurately the vertical profile of N<sub>2</sub>O tracer species at the time of the SPI-RALE measurement (Fig. 7). Below 20 km the 3-hourly profiles of N<sub>2</sub>O are characterized by slightly higher values with a better agreement with Reprobus-3h-for. Above 20 km the N<sub>2</sub>O quantities are decreased towards the profile obtained from the operational analysis. A slight underestimation of the measurements is observed for the upper part of the profile but the decrease of N<sub>2</sub>O amounts as altitude increases appears to be more consistent than the Reprobus-ope simulation results. However, the vertical structure observed by SPIRALE around 27 km cannot be reproduced with the new simulations.

The Reprobus-3h-ope-for and Reprobus-3h-for simulations result in enhanced total mixing ratios of  $NO_y$  above 20 km where N<sub>2</sub>O calculated values are lower in comparison with the typical Reprobus-ope simulation (Fig. 8). Below these altitudes, the increased simulated amounts of N<sub>2</sub>O give lower NO<sub>y</sub> mixing ratios in better agreement with the NO<sub>y</sub>\* in situ profile deduced from Michelsen et al.'s correlation. Above 30 km however, the modelled total amounts of NO<sub>y</sub> from the two additional simulations do not appear to be sufficient when compared to the two NO<sub>y</sub>\* profiles (cf Sect. 4.2). This could be partly considered as a limitation of the use of the Michelsen et al.'s correlation curve between N<sub>2</sub>O and NO<sub>y</sub>.

NO<sub>2</sub> and HNO<sub>3</sub> profiles calculated by the 3-hourly wind simulations have increased mixing ratios which are in better quantitative agreement with SPIRALE observations (Fig. 9). For example, NO<sub>2</sub> quantities calculated with the Reprobus-3h-for simulation above 22 km are increased by 28, 28 and 10% at 23, 27 and 31 km, respectively, in comparison with the Reprobus-ope simulation (Fig. 9a). Nevertheless, underestimation is still visible around the 27-km structure and also above 30 km as a result of the apparent underestimation of total NO<sub>v</sub>\* quantities (19 and 6% lower than the observation at 27 and 31 km, respectively). Below 22 km the disagreement remains roughly similar between the three model experiments. For HNO<sub>3</sub>, most of the Reprobus-3h-for simulation values are located within the error bars of the SPIRALE profile (Fig. 9b), even though some fine specific structures cannot be reproduced, such as at 21 km (27% higher than the observation). An improvement is obtained in comparison with the Reprobus-3h-ope-for experiment below 22 km where Reprobus-3h-for simulation gives the best agreement with the NO<sub>v</sub>\* in situ profile. The results of the two 3-hourly simulations of the NO2/HNO3 ratio fall now within the error bars of the in situ observation (Fig. 9c) which shows that the mid-latitude partitioning between these two species can be modelled correctly provided that the simulated NO<sub>y</sub> total



Fig. 6.  $N_2O$  zonal means simulated by Reprobus on 2 October 2002 at 08:00 UT. (a) Results obtained using 6-hourly ECMWF operational analysis data; (b) Results obtained using 3-hourly analysed and forecast data (see text).



**Fig. 7.** Same as Fig. 1 but the comparison is made with the results of three model simulations. Red line: 6-hourly operational analysis (Reprobus-ope); green line: 3-hourly data obtained by interleaving operational analysis with forecasts (Reprobus-3h-ope-for, see text); blue line: 3-hourly forecasts available since 2002 (Reprobus-3h-for).

amounts are well taken into account. Note that we clearly see in Fig. 9c that the  $NO_2/HNO_3$  ratio depends on the meteorological data used to drive the model. In the perturbed dynamical situation studied here (as shown in Fig. 2) the geographical positions of the air-mass trajectories computed in Reprobus could depend on the chosen ECMWF data. Different temperatures and solar zenith angles along the computed trajectories between the 6-hourly and 3-hourly simulations would then affect the calculated quantities of the  $NO_y$  species and especially of  $NO_2$  and NO.



**Fig. 8.** Same as Fig. 4 but the comparison is made with the results of the three model simulations. Red line: 6-hourly operational analysis (Reprobus-ope); green line: 3-hourly data from operational analysis and forecasts (Reprobus-3h-ope-for, see text); blue line: 3-hourly forecasts (Reprobus-3h-for).

Ignoring the specific problem of mixing, the Reprobus 3-D CTM underestimates the NO<sub>2</sub> and HNO<sub>3</sub> amounts at midlatitudes due to the underestimation of total NO<sub>y</sub> which is in turn the result of transport uncertainties in the model that affect the simulation of N<sub>2</sub>O. ECMWF 6-hourly operational analysis seem to be characterized by some noisy and unrealistic features which are not present in the forecasts (Legras et al., 2004) and apparently affecting the transport calculations in Reprobus and in other CTMs. The strict reasons for the significant improvements of the NO<sub>y</sub> comparisons using the



**Fig. 9.** NO<sub>2</sub> (**a**) and HNO<sub>3</sub> (**b**): same as Fig. 3 but the comparison is made with the results of the three model simulations (Reprobus-ope, Reprobus-3h-ope-for and Reprobus-3h-for). NO<sub>2</sub>/HNO<sub>3</sub> ratio (**c**): same as Fig. 5 but the comparison is made with the results of the three model simulations.

3-hourly winds are still unclear. Nevertheless it is thought that using 3-hourly wind data gives a better sampling for the interpolation of the positions between two consecutive wind records in Reprobus.

However discrepancies between the SPIRALE measurements and the model results in Fig. 9 still remain especially for NO<sub>2</sub> above 30 km as a result of the underestimation of the total amounts of NO<sub>y</sub>\* at these altitudes. Recent studies have shown that satisfying comparisons can be obtained between observations and model calculations when those are constrained by observations of total NO<sub>y</sub> (e.g. Küll et al., 2002). This points out the necessity of an adequate simulation of total amounts of NO<sub>y</sub> to reproduce the partitioning of the individual species.

#### 6 1-D calculations

To get rid of the influence of the problem of transport and hence of the missing quantities of total NO<sub>v</sub>, we have used the Reprobus model in its one-dimensional form. Constraining 1-D-Reprobus calculations by observed total NOv has significantly improved the modelled partitioning of NO<sub>v</sub> species compared to balloon measurements in the summer polar stratosphere (Dufour et al., 2005). Similar conclusions have been obtained from comparisons between balloon observations and lagrangian simulations in January 1999 inside a warm Arctic vortex (S. Bausch, private communication). In the present study the 1-D-Reprobus initialisation is constrained by the N<sub>2</sub>O profile measured by SPIRALE and by the corresponding NOy\* profiles. However the N2O-NOy correlation curves from Michelsen et al. (1998) have been determined on given latitude bands and it is not straightforward to choose the adequate correlation values that have



**Fig. 10.** NO<sub>2</sub> (**a**), HNO<sub>3</sub> (**b**) and NO<sub>2</sub>/HNO<sub>3</sub> ratio (**c**) measured by SPIRALE (black curve) and simulated by Reprobus. The model is here constrained by the measured profile of N<sub>2</sub>O and the NO<sub>y</sub>\* deduced from the N<sub>2</sub>O-NO<sub>y</sub> correlation. Blue curve: Reprobus-1-D-midlat experiment, NO<sub>y</sub>\* profile inferred from the correlation valid in the 28–46° N latitude band. Orange curve: Reprobus-1-D-midlat-trop experiment, NO<sub>y</sub>\* profile inferred from the correlation valid in the 3–10° N latitude band, only shown above 29 km where the air seems to mostly originate from low latitudes.

to be considered above 29 km where air seems to originate from low latitudes. Two simulations are performed in our study. The first one (referred as Reprobus-1-D-midlat) is constrained with the NO<sub>y</sub>\* profile inferred from the N<sub>2</sub>O-NO<sub>y</sub> correlations in the 28–46° N latitude band. In the second simulation (hereafter Reprobus-1-D-midlat-trop) the computed NO<sub>y</sub>\* profile is defined by the N<sub>2</sub>O-NO<sub>y</sub> correlation values in the 28–46° N and 3–10° N latitude ranges below and above 29 km, respectively (cf Sect. 4.2). These two model experiments allow us to test the sensitivity of the choice of the N<sub>2</sub>O-NO<sub>y</sub> correlation law on the modelled partitioning of NO<sub>2</sub> and HNO<sub>3</sub>. The other species are initialised from the 3-D-Reprobus simulation as in the work of Dufour et al. (2005).

For both experiments the calculated mixing ratios of  $NO_2$  and  $HNO_3$  are close to the measurements (Fig. 10). For

NO<sub>2</sub> significant improvement is obtained above 22 km: the Reprobus-1-D-midlat experiment provides values that are only 17% and 3% lower than the SPIRALE profile at 23 and 27 km, respectively. Above 29 km the measured quantities lie between the results of the two simulations (Fig. 10a). Similar conclusions can be derived from the HNO<sub>3</sub> profile which is satisfyingly simulated for the whole altitude range (24% and 10% lower than the observation at 18 and 20 km, respectively, 6% and 24% higher than the observation at 23 km and 28 km, respectively) except in the layer around 27 km (41% higher than the observation at 27 km). The calculation of the NO<sub>2</sub>/HNO<sub>3</sub> ratio is not significantly improved between 22 and 27 km (22% and 31% lower than the observation at 23 km and 27 km respectively), the computed ratio being underestimated at 27 and 28 km as a result of HNO3 overestimation. Taking into account the N<sub>2</sub>O-NO<sub>v</sub> correlation

uncertainty of Michelsen et al. (1998) the 1-D-simulation cannot reproduce simultaneously  $NO_2$  and  $HNO_3$  in the particular layer of a possible previous mixing event. Actually the correlation curves of Michelsen et al. (1998) determined on given latitude bands cannot account for a full description a  $N_2O$ - $NO_y$  correlation within a layer containing mixed air from tropical and mid- latitudes.

#### 7 Conclusion

We have used in situ observations of N<sub>2</sub>O, NO<sub>2</sub> and HNO<sub>3</sub> by the SPIRALE balloon-borne instrument to test our ability to reproduce the observed NO<sub>v</sub> partitioning at mid-latitudes. In a first part we have used the Reprobus 3-D CTM computed with the most recent kinetics data from the JPL recommendation and driven by the meteorological parameters from ECMWF. Discrepancies appear between the simulation and the measurement of N2O when transport calculations are conducted with the ECMWF 6-hourly operational analysis widely used in the scientific community. This affects the modelling of NO<sub>2</sub> and HNO<sub>3</sub> with in particular an underestimation of the NO<sub>2</sub>/HNO<sub>3</sub> ratio. The comparisons for N<sub>2</sub>O are improved at all altitudes when transport calculations are driven by 3-hourly wind data. This result is consistent with the analysis of Legras et al. (2004) who shows much better agreement between ER-2 in situ observations and distributions of N<sub>2</sub>O calculated using such 3-hourly wind fields. From this improved simulation, more realistic N<sub>2</sub>O, NO<sub>2</sub>, and HNO<sub>3</sub> profiles are calculated by Reprobus. However, the vertical feature observed by SPIRALE around 27 km in the N<sub>2</sub>O profile cannot be reproduced by the 3-D CTM. This feature is attributed to a previous mixing event which is beyond the capabilities of the model at its current horizontal and vertical resolution.

The best results regarding the partitioning of NO<sub>y</sub> are obtained by constraining the 1-D version of model with the observed profile of N<sub>2</sub>O and the corresponding total NO<sub>v</sub> deduced from established N<sub>2</sub>O-NO<sub>v</sub> correlations. In these conditions, our model simulations show that the observed midlatitude NO<sub>2</sub>/HNO<sub>3</sub> ratio can be reproduced with excellent agreement at all stratospheric altitudes. Modelling the NO<sub>v</sub> partitioning with the best accuracy relies at least on a satisfying simulation of N<sub>2</sub>O and thus of total NO<sub>v</sub>. This result is consistent with the conclusions of Dufour et al. (2005) who constrained their model for the Arctic summer period with the sum of the measured NO<sub>v</sub> species. It also suggests that discrepancies found between observations and modelling of the NO<sub>v</sub> partitioning at mid-latitudes mainly originate from transport effects rather than general knowledge of the stratospheric chemistry.

The problem of transport calculation may also affect the modelling of diabatic descent inside the polar vortex where disagreements appear when comparisons are performed between models and satellite observations (e.g. Ricaud et al., 2005). A next step of our work is to make comparisons at a global scale and for long time series using satellite data such as those of ODIN and ENVISAT.

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