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UV-A/UV-B from UV
irradiance spectra**

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Study of suitability of cheap AvaSpec array spectrometer for solar UV field measurements

I. Ansko^{1,2}, K. Eerme¹, S. Lätt^{1,2}, M. Noorma², and U. Veismann¹

¹Tartu Observatory, 61602, Tõravere, Estonia

²Tartu University, Ülikooli 18, 50090, Tartu, Estonia

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Correspondence to: K. Eerme (kalju@aai.ee)

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Abstract

A system to record the ultraviolet (UV) spectra of atmospheric global irradiance with the miniature fiber optic spectrometer AvaSpec-256 was developed for continuous computer-aided spectrometry at Tartu Observatory in 2005. As a result, the database of spectra recorded with 15-min-interval round 24 h over 300–400 nm, has been developed. The quantities retrieved from the spectra have been compared with those measured by the Scintec erythematous UV-SET sensor and the Kipp&Zonen narrowband 306 nm sensor. Almost clear and overcast days were selected for comparison. Reliable results on the spectral distribution of the UV global irradiance as well as the integrated daily spectral doses could be obtained at least during the bright half-year. The results were compared with the calculations performed by means of the LibRadtran package. The biases in irradiance were significant at SZA above 70–75°. At dominating larger SZA the recorded values need sophisticated corrections and remain less reliable. At lower latitudes than that of the study site (58.3 degrees), the reliability of the spectrometer is expected to increase due to a smaller contribution of data measured at large SZA.

The variations of the ratio of UV-A/UV-B irradiance, retrieved from the spectra, were investigated. Also the covariation of the narrowband 306 nm irradiance and the irradiance integrated over the whole UV-B range was studied. The biases between calculated by means of the LibRadtran package and the measured ratio of UV-A/UV-B irradiance were small at SZA below 70°. At larger SZA the values of the ratio as well as the biases increased, significantly depending on total ozone.

1 Introduction

The importance of recording ground-level solar UV radiation spectra in addition to the broadband and narrowband filter instrument measurements has increased in recent decade (Seckmeyer et al., 2001; WMO, 2007). It is related to the deepened

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research of the health effects of UV radiation (Berwick and Kesler, 2005; Lehmann, 2005; Grant et al., 2005) as well as of its environmental effects in the atmosphere (Brönnimann et al., 2001), in plants and microorganisms (Neale et al., 2007; Sullivan et al., 2007). International collaboration has widened in the recent decades. The European Database for UV Climatology and Evaluation (EDUCE) stores the available UV spectra (<http://www.muk.uni-hannover.de/~seckmeyer/EDUCE/database.html>) for their use by the European UV community. The database was created in the first years of this century along with the quality assurance and quality control methods for spectral measurements (Gröbner et al., 2002; Gröbner et al., 2006). In most cases the Brewer spectrometers and other expensive spectrum scanning instruments are used. As an alternative, the advancement of technology has made available the compact and simple single-monochromator array spectrometers (Basics of Spectral. . . , 2003; Oliver and Moseley, 2002; Ylianttila et. al., 2005). However, the limited dynamic range and intrinsic stray light problems complicate their use.

Taking into account our previous experience in exploiting such kind of instruments at Tartu Observatory (Kutser et al., 1999), a complementary metal-oxide semiconductor (CMOS) array minispectrometer AvaSpec-256, produced by Avantes Inc., was suitable for continuous field measurements in case the necessary auxiliary devices were added. Preparation of a cost-effective system to record regularly the UV spectra of solar global irradiance was started in 2004. Since 2005, the UV spectra in the wavelength range of 300–400 nm have been regularly recorded with a period of 15 min with a spectral resolution of about 1 nm. The aim was to store the data on spectral irradiances in the local database in compliance with the EDUCE standard requirements (<http://www.muk.uni-hannover.de/~seckmeyer/EDUCE/database.html>) and study the most characteristic features of spectra in different weather conditions. More than 40 000 UV spectra have been recorded at the Tartu Observatory site (58°15' N, 26°28' E, 70 m a.s.l.). Located next to Tartu Observatory, Tartu-Tõravere Meteorological Station of the Estonian Meteorological and Hydrological Institute (EMHI), included since 1999 into the Baseline Surface Radiation Network (BSRN), supports the spectral

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measurements with the pyranometer- and pyrliometer-measured broadband radiation data and the UV data measured by filter radiometers. All supporting meteorological and related to the topic necessary information, including the aerosol optical depth (AOD) measured by the AERONET sun photometer, direct sun total ozone and cloud data, is available from the station.

In the present paper the results of measurements of the erythemally weighted irradiances and doses of the Scintec broadband instrument UV-SET and those retrieved from spectra have been compared. Also the spectral irradiance at 306 nm as well as an integrated value over the wavelength region 290–315 nm has been compared with the narrowband spectral irradiance measured by the Kipp&Zonen CUVB1 instrument at 306 nm. The diurnal cycles of the ratio UV-A/UV-B irradiances in sunshine and overcast conditions have been studied. The analysis of the complex spectra (the solar disk is part time opened and part time obscured) remains out of the scope of the present paper.

2 Instrumentation and calibration

2.1 Spectrometer

The main device in the spectrometric system is the miniature fiber optic spectrometer AvaSpec-256. The optical design of the spectrometer is based on the symmetrical Czerny-Turner design with 256 pixel detector array. The CMOS detector HAM 256 is connected to an electronics board with a 14 bit AD converter and USB interface. The grating 600 lines per mm was selected to cover the spectral range 237–444 nm with blaze by 250 nm. Entrance slit width is 50 μm . For irradiance measurements a teflon diffuser of 30 mm diameter and 0.4 mm thickness has been used as an optical input. At zenith angles above 80° the actual response of the system diffuser+spectrometer was more than 20% lower than that of the cosine. At smaller zenith angles the differences did not exceed $\pm 10\%$, mostly remaining within $\pm 5\%$. A quartz fiber of 4 m length

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and 100 μm diameter connects the optical head on the roof to the spectrometer. A Russian UFS-5 color glass filter was installed between the diffuser and fiber to reduce the longwave radiation in the spectrometer and to guarantee the reliable recording of the signal in the whole UV spectral range. For reliable detection of noise level, the optical input is covered by a shutter before and after each measurement cycle.

A schematic diagram of the spectroradiometric system is presented in Fig. 1. The measurement process is fully programmed (Linux). The spectra are stored in MySQL format. The control and data acquisition computer of the spectrometer is connected to the Tartu Observatory web. The measurements can be tracked by any computer of the local network. It is also possible to have access to the archive of spectra. Preliminary quality testing of the spectra is performed using the check UVspec package according to EDUCE roles (including the comparison with model calculations).

Tuning of the total responsivity by recording spectra is realized through automatic selection of integrating time within the interval of 1 to 60 s. Thus a maximum value of the signal reaching approximately 16 000 counts is realized for each recorded spectrum. The effective integration time can be increased by digital summing of up to 30 spectra. The temperature effects on the dark current of the sensor were significant, evoking the necessity of temperature control. The spectrometer is installed in a cooling box and kept at a constant temperature $+7^\circ\text{C}$ to reduce the noise level.

Calibration of UV sensors at Tartu Observatory is based on the tungsten-halogen standard lamps FEL, certificated by the Oriel company traceable to the US National Institute of Standards and Technology (NIST). A schematic setup of the sensor calibration is presented in Fig. 2. As the spectrometer's responsivity changes with time, the system needs frequent recalibration. The changes of responsivity over time of exploitation is presented in Fig. 3. The reason for the decrease of responsivity during the first year of exploitation could be caused by the condensation of water vapour on the optical surfaces within the instrument, not the decrease of array sensitivity. The results show that after a significant drop of responsivity during the first year the variations remain within a few per cent.

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The intrinsic stray-light problems limit the use of array spectrometers in the UV-B region. A program for compensatory calculation of the stray light influence was applied (Kostkowsky, 1997; Brown et al., 2003; Zong et al., 2006). An example of the reduction of stray light is presented in Fig. 4. Besides stray light, another strongly restricting factor in the UV-B region is a narrow dynamic range of the CMOS array. A shortwave threshold of the reliably recorded irradiance depends on the solar elevation angle and cloud cover. During midsummer noon hours it reaches 300 nm. On clear days in mid-winter the noon threshold is around 310 nm. In midwinter overcast conditions no UV-B irradiance has been recorded in most cases. The spectral irradiance at wavelengths below 300 nm remains inaccessible even in the best conditions.

2.2 Filter radiometers and auxiliary instruments

The quantities retrieved from the spectra have been compared with these obtained by the broadband erythemally weighted Scintec sensor UV-SET and with the spectral irradiances of Kipp&Zonen narrowband sensor CUVB1. The latter was centered at the wavelength 306 ± 0.2 nm with the respective bandwidth 2 ± 0.5 nm. The distance between the location of the spectrometer and the filter radiometers installed at the EMHI meteorological station is about 250 m.

The temperature of the sensor CUVB1 is stabilized at 40° C. The narrowband measurements of UV-B irradiance at Tartu-Tõravere Meteorological Station have been performed since February 2002. The sensor CUVB1 can be relatively easily recalibrated in units of spectral irradiance by using a standard radiation source. In 2002–2007 six calibrations of the sensor CUVB1 were performed in radiometric laboratory. The average value of the instrument's responsivity exceeded the value established by the producer only by 0.8%. Taking into consideration that uncertainty of the standard lamp flux at 300 nm is approximately 2%, we continue to use in processing of the measurement data the original value of 27.17 volts per $\text{W}\cdot\text{m}^{-2}\cdot\text{nm}^{-1}$.

The local calibration of the Scintec UV-SET sensor is more complicated unless a Brewer or some other spectroradiometer, allowing reliable measurements of spectral

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irradiances at the shortest wavelengths up to 290 nm and detection of the spectral responsivity in the wavelengths region of its steep decrease, is used. The difference in the spectrally integrated irradiance due to the deviation of the spectral responsivity of the instrument from the CIE weighted has been estimated to be about 3%. The total response of the UV-SET sensor has been regularly checked against the standard FEL lamp. The coefficient of radiometer response has been corrected for the decrease of total sensitivity over time, trying to maintain the initial producers calibration.

The total ozone daily values have been estimated, using the MICROTOPS-II instrument preferably during hours close to the noon; also the Earth Probe satellite TOMS and the Aura satellite OMI instrument data have been used when the local direct sun measurements were impossible. The local total ozone measurements were based on the producer calibration and the results were regularly compared with the results obtained by the OMI instrument. The average ratio of the MICROTOPS-II to OMI total ozone daily values has been 1.002 with a standard deviation of 2.3%, while the extreme differences remain within $\pm 6\%$. The AERONET Cimel-18 sun-photometer, providing the data on aerosol optical depth (AOD) in the range 340 to 1020 nm, is also located at the EMHI meteorological station. In most cases the ratios of AOD at 340 nm to those at 380 nm and 500 nm occurred stable (Eerme et al., 2006). The AOD in the UV-B region was expected to be proportional to the value at 340 nm. Reliance on the proportionality in the spectral AOD may be the reason for biases between the calculated and real spectral irradiances.

The spectral distribution of the UV irradiance is strongly influenced by cloud cover. The hourly cloud data have been detected visually in tenths at all three basic levels by the staff of the meteorological station. A wide-angle videocamera is used aside the spectrometer for recording the current cloud situation during the measurement cycle.

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3 Results and agreement with the filter instruments

The primary set of quantities retrieved from the spectra contains the erythral UV Index (UVI), spectral irradiance at 306 nm, the integrated UV-A (the wavelengths 315–400 nm) and UV-B (the wavelengths below 315 nm) irradiances weighted by the rectangular boxes as well as the ratio of UV-A/UV-B. The daily doses of erythral, 306 nm spectral, UV-A and UV-B irradiances are integrated, interpolating them over all the recorded spectra which satisfy the quality needs. The measurement data of the instrument CUVB1 as well as these of the erythemally weighted broadband instrument UV-SET have been recorded together with other radiation data with a one min time resolution. The daily doses in $\text{J}\cdot\text{m}^{-2}\cdot\text{nm}^{-1}$ and $\text{J}_{\text{eff}}\cdot\text{m}^{-2}$ have been integrated, considering all the recorded values. The weather conditions often change rapidly. During the integration of a spectrum the solar disk may be part-time cloud-free and part-time covered with clouds. As mentioned already in the Introduction, such spectra have not been considered in the present work.

The agreement of CIE weighted integrated daily doses and those recorded by the Scintec UV-SET instrument in the sunshine as well as under overcast conditions, is presented in Fig. 5. The linear correlation between the doses in both situations was 0.965. Partly the disagreement is caused by the different temporal resolution in integration of the doses. Another reason is related to differences in angular responses of the instruments. The UVI values as well as the doses measured by the Scintec UV-SET tend to be higher than those retrieved from spectra. The reason could be explained by the difference in irradiance scales. As noted above, the total response of the Scintec UV-SET has been regularly checked and corrected to keep the producers' scale; one cannot exclude that the manufacturers scale overestimates the UVI. The results of our previous calculations of the UVI and the daily dose values performed by means of the LibRadtran package (Mayer and Kylling, 2005) also were nearly 10% lower than those measured by the UV-SET (Eerme et al., 2006). In Fig. 6 an example of the measured to calculated ratios of summertime clear weather UV Index retrieved from the spectra

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as well as measured simultaneously by the UV-SET are presented versus SZA. The bias between the UV-SET measured and retrieved from the spectra UVI reaches 30%. The UV-SET ratio is more stable than the one retrieved from the spectra; however the latter is much closer to the calculated value.

5 The values of spectral irradiance at 306 nm retrieved from the spectra agreed with the CUVB1 data within $\pm 10\%$ at SZA values below 70° . The day-to-day variations of the average ratio were within a few per cent. An example of the daily cycles of both spectral irradiances relative to the calculated by means of the LibRadtran package are presented in Fig. 7. Mutual agreement between the measured values is good; 10 the biases between the measured and calculated values, however, reach 20%. The reason is not well understood yet. Significant differences between both measured values were also found at SZA above 70° . The linear correlation between the daily doses of irradiance in the full UV-B range and the daily doses of spectral irradiance at 306 nm were as high as 0.986. Their covariance is illustrated in Fig. 8. One can see 15 that the narrowband spectral irradiance at 306 nm could be considered as a good proxy for the whole UV-B irradiance.

In Fig. 9 the ratios of integrated over spectra irradiances UV-A/UV-B together with the model calculated ratios versus SZA in the sunshine and small cloud amount conditions are presented for different available total ozone values in spring-autumnal and 20 summer conditions. The UV-A contributions are higher at larger and lower at smaller total ozone values. At SZA below 70° the biases between the calculated and measured values occurred small for all total ozone values. In the SZA range of 35° to $65\text{--}70^\circ$ the average ratio of UV-A/UV-B increases slowly from about 50 to about 100. At larger SZA the surplus of UV-A as well as the bias between the calculated and measured value 25 strengthens. One can see that the ratio is much larger in spring high ozone conditions (more absorptance of UV-B) than in autumnal low ozone conditions. The maximum bias between the calculated and measured values, strongly dependent on total ozone, was found at SZA around 85° .

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4 Conclusions

The experience of nearly three-year-exploitation of the cheap CMOS array minispec-
trometer AvaSpec-256 at Tartu Observatory confirms that realistic spectral distributions
of global UV irradiance can be obtained. The quantities retrieved from the spectra us-
ing different weighting functions are comparable with those measured with filter instru-
ments. The variations of the values are highly correlated. At SZA below 70–75° the
ratios are stable and vary within a few per cent. At larger SZA the biases grow sig-
nificantly due to deviations of the instrumental angular response from the cosine law;
special correction methods are necessary to increase the reliability of results.

Significant biases between the measured values and those calculated using LibRad-
tran codes also appear at SZA above 70–75°. At smaller SZA the ratios between the re-
trieved from the spectra and calculated values are stable. Systematic biases between
the calculated and measured quantities arising also in the case of filter instruments
reaching in some cases for 20% and more are not well understood yet.

The agreement between the daily UV doses integrated from the spectra, recorded
by the filter instruments, and by the LibRadtran codes at the study site (latitude
58.3 degrees) in the summer half-year period is satisfactory. At the study site at latitude
58.3 degrees in the summer half-year period the integrated from the recorded spectra
daily UV doses weighted by different response functions are in satisfactory agreement
with the calculated values and those recorded by filter instruments. At lower latitudes
the reliability of the results obtained by means of the array spectrometers should in-
crease due to the smaller contribution of large SZA in daily doses.

Aside the cosine correction problems the major complications are related to the rela-
tively restricted dynamic range of the instrument, the changes of responsivity over time
and the stray light in the instrument. The instrument needs recalibration at least every
two-three months. The stray light correction of spectra is necessary.

The shortwave threshold of reliable recording depends on solar elevation and cloudi-
ness. At the study site it reaches the wavelength 300 nm in midsummer noon sunshine

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conditions and is limited to only 310 nm in noon sunshine around midwinter. In midwinter overcast conditions recording of the UV-B radiation is quite rare.

Acknowledgements. The work has been supported by grants No 5348 and 7137 of the Estonian Science Foundation. The authors thank the Tartu-Tõravere Meteorological station of the EMHI and especially E.-M. Maasik for providing the careful measurements of total ozone.

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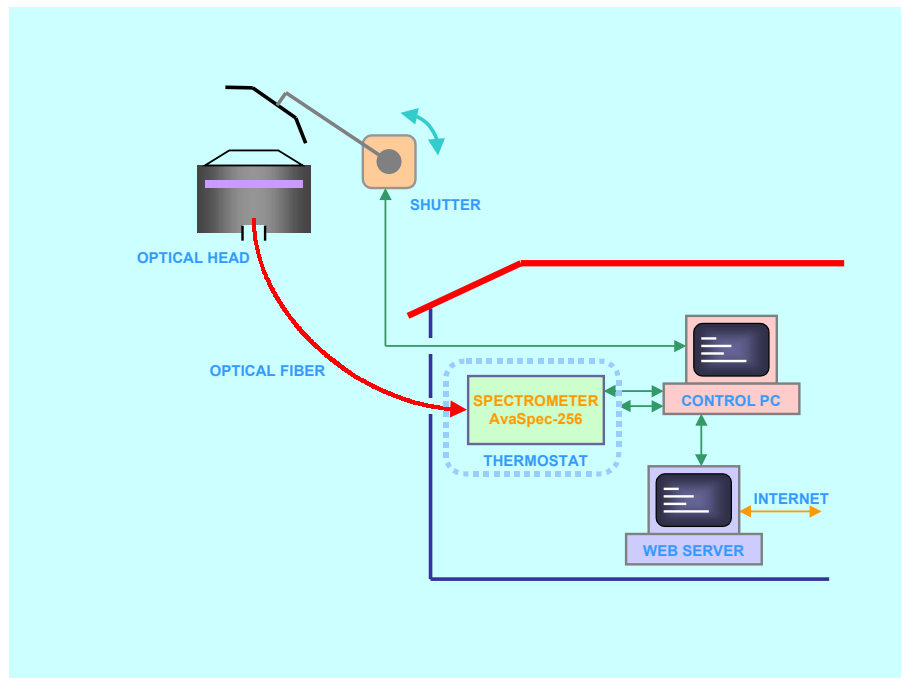
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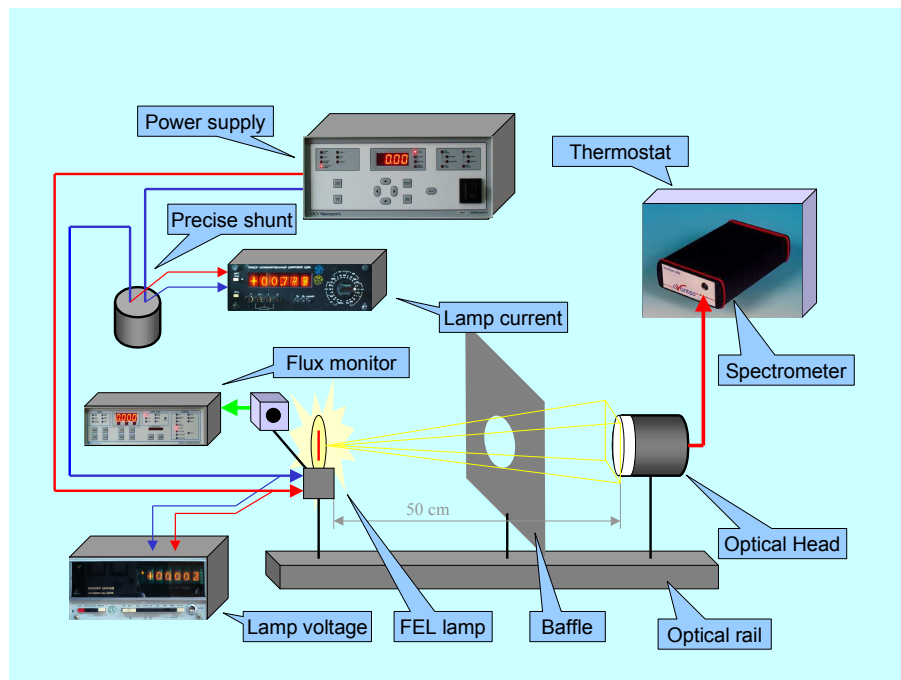
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**Fig. 1.** Schematic diagram of the spectroradiometric system.[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

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**Fig. 2.** Schematic setup of the sensor calibration.

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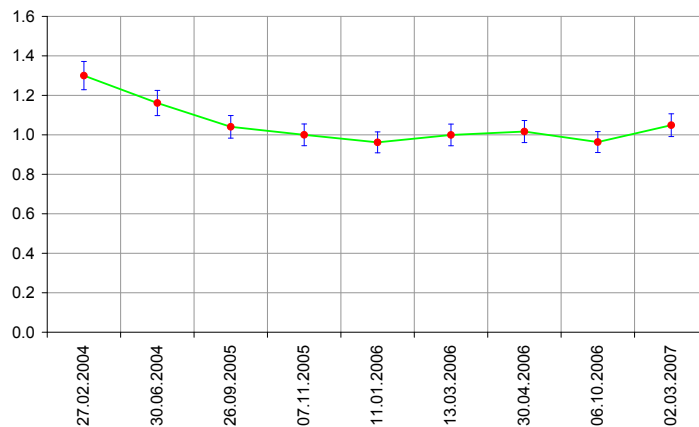
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**Fig. 3.** Changes of the responsivity of AvaSpec-256 over time.

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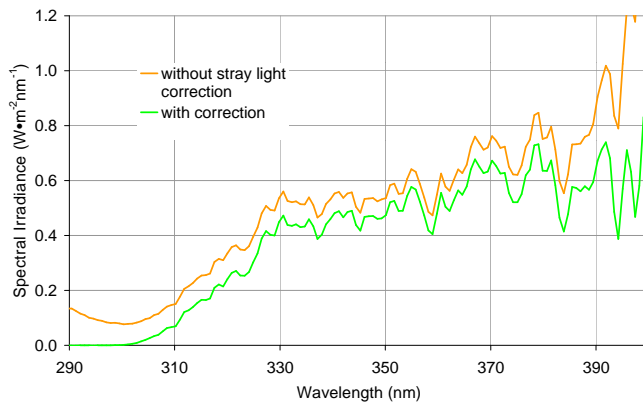
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**Fig. 4.** Example of the stray light correction of spectra.[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

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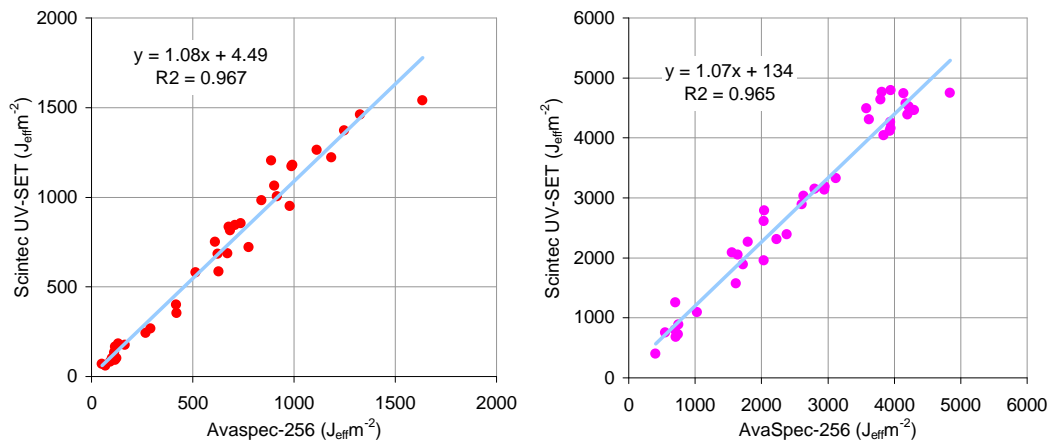


Fig. 5. Erythemal daily doses retrieved from AvaSpec-256 spectra versus Scintec UV-SET measured values in overcast (left) and sunshine (right) conditions.

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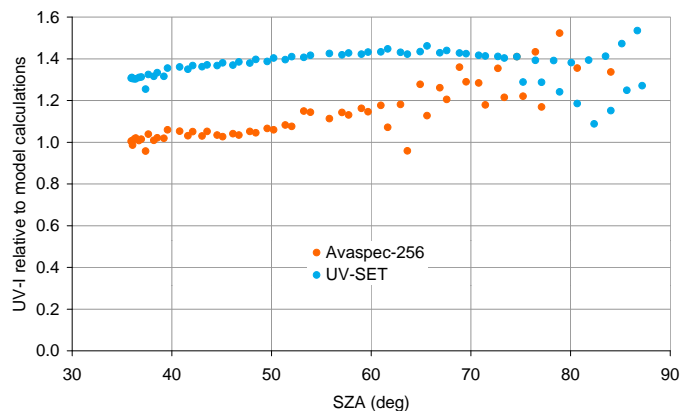


Fig. 6. Ratio of UV index measured (by UV-SET and retrieved from spectra on 3 June 2006, total ozone 373 DU, AOD at 500 nm 0.05) to LibRadtran calculated versus SZA.

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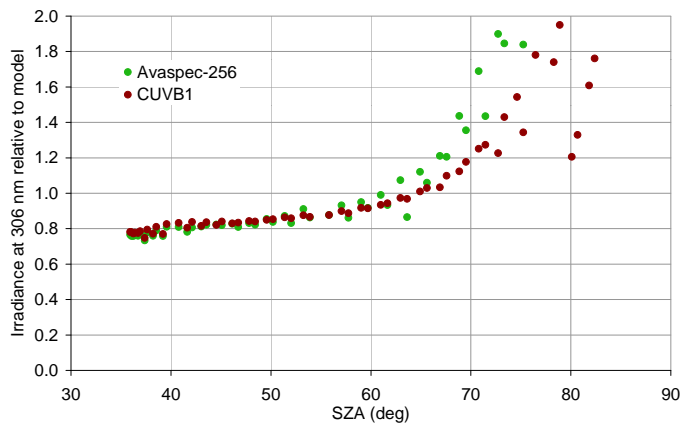


Fig. 7. Ratios of the AvaSpec-256 and the CUVB1 measured 306 nm spectral irradiance to the LibRadtran calculated versus SZA.

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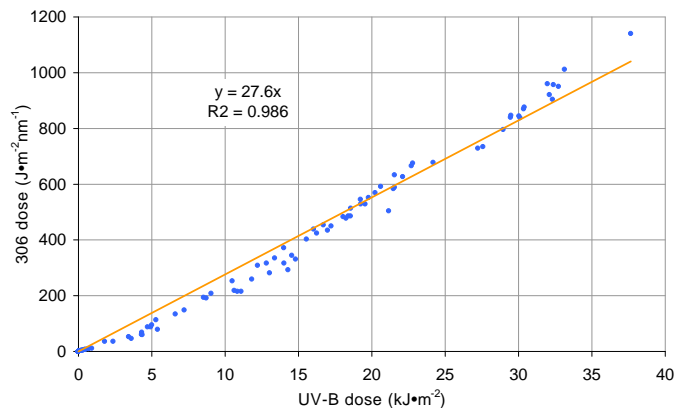


Fig. 8. Covariance of the AvaSpec-256 UV-B and narrowband filter instrument CUVB1 306 nm measured spectral irradiance daily doses.

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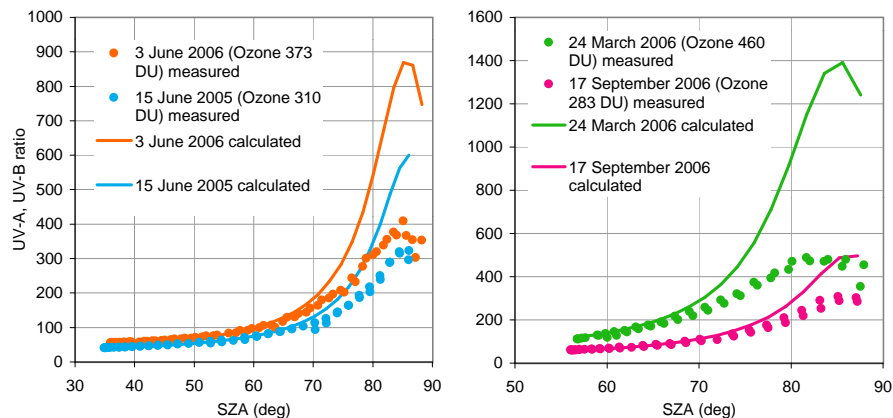


Fig. 9. Measured and calculated ratios UV-A/UV-B versus solar elevation angle in the sunshine conditions at different atmospheric total ozone values in summer (left) and spring-autumn (right).

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