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S. S. Ghosh, A. Sen, G. S. Lakhina. Dromion solutions for nonlinear electron acoustic waves in space plasmas. *Nonlinear Processes in Geophysics*, 2002, 9 (5/6), pp.463-475. hal-00302152

HAL Id: hal-00302152

<https://hal.science/hal-00302152>

Submitted on 18 Jun 2008

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Dromion solutions for nonlinear electron acoustic waves in space plasmas

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Received: 24 January 2002 – Revised: 12 June 2002 – Accepted: 13 June 2002

Abstract. Recent high resolution measurements of POLAR and FAST satellites have revealed two-dimensional coherent structures in the polar cap boundary layer (PCBL) region. Since the ion temperature is often observed to be greater than the electron temperature in this region, electron acoustic waves can exist as a normal mode of the plasma system. It is shown that the nonlinear evolution of such waves can be modelled by the coupled Davey-Stewartson I equations. These equations, which are a generalization of the nonlinear Schrödinger equation to two dimensions, admit exponentially localized solutions called dromions. A detailed parametric characterization of the regions of existence of such solutions is presented in the context of the PCBL region.

1 Introduction

Solitons, which are exponentially localized wave forms in (1+1) dimension, are well-known as exact solutions of certain integrable partial differential equations. In the past few decades they have been extensively studied and applied in a variety of fields such as physics, chemistry, biology, neurophysics, etc. They have also figured prominently in the interpretation of several wave structure observations in space plasmas. However, work related to two- and three-dimensional generalizations of solitons is of more recent origin and consequently is relatively less well-known. One of the earliest work is that of Kadomtsev and Petviashvili (1970), who constructed a two-dimensional generalization of the K-dV equation. Their model equation, known as the Kadomtsev-Petviashvili (KP) equation, admits a variety of nonlinear solutions. Line solitons are the simplest kind and are exponentially localized in a single direction (say x) and propagate in the (x, y) plane. Their equation also admits lump or algebraic solitons which decay polynomially (i.e. as $O(1/x^2, 1/y^2)$) in all directions (Satsuma and Ablowitz, 1979). A host of other independent equations governing sim-

ilar two-dimensional solitary structures have been studied by many authors (Djordjevic and Redekopp, 1977; Manakov, 1976; Ablowitz and Segur, 1979). However, none of these studies dealt with a nonlinear solution that had an exponential decay in both dimensions. This lacuna was removed by the work of Boiti et al. (1989), who found coherent localized structures as a solution of (2+1) dimensional integrable partial differential equations which, unlike a lump or algebraic solitons (Satsuma and Ablowitz, 1979; Janaki et al., 1991), decayed exponentially in both directions in the (x, y) plane. One characteristic feature of these solutions is that they involve two field variables. One of the variables represents the physical entity which is being evolved (the electric field amplitude, for example) and displays a localized structure in space. The other auxiliary variable shows a plane wave character with vanishing physical fields at the boundaries and is often called a ghost soliton (Hietarinta, 1990). The dromion solutions also require special time dependent boundary conditions which are described by those ghost solitons. These ghost solitons are in fact the line solitons (Fokas and Santini, 1989) mentioned earlier and represent the “tracks” (*dromos* in Greek), along which the localized lumps of the physical variable travel. In fact, the lumps are located precisely at the intersection of two such tracks of the auxiliary variable. Since they are “driven” by their boundaries, they were, therefore, named as *dromions* (Fokas and Santini, 1990). Ever since their discovery the mathematical properties of dromions have been investigated by many authors both analytically (Santini, 1990; Hietarinta, 1990; Hietarinta and Hirota, 1990; Radha and Lakshmanan, 1995), as well as numerically (Besse and Bruneau, 1998; Nishinari and Yajima, 1994; Nishinari et al., 1996).

A large class of two-dimensional partial differential equations have been shown to possess dromion solutions (Radha and Lakshmanan, 1997b; Guil and Manas, 1996; Imai, 1997; Ruan and Chen, 1999; Lou and Ruan, 2001). One of the most well-known of such equations is the so-called Davey-Stewartson-I equation (DS-I) (Davey and Stewartson, 1974), which is a two-dimensional generalization of the nonlinear

Schrödinger equation (NLSE). While the NLSE finds extensive applications in plasma physics, including the modelling of several wave phenomena in space plasmas, the DS-I equations have received scant attention from plasma and space scientists. A rare exception is the work of Nishinari et al. (1994), who have shown, with the help of a suitable reductive perturbation method, that the nonlinear evolution of an ion acoustic wave travelling in a magnetized plasma can be modelled by a DS-I type of equation. Similar considerations might, therefore, apply for other modes in a plasma, particularly when their nonlinear stationary states display two- or three-dimensional structures. Our present work is motivated by such a consideration particularly in the context of space plasma observations which provide a rich source of such potential structures. For instance, the GEO-TAIL spacecraft was one of the earliest (Matsumoto et al., 1994a) to detect interesting waveforms consisting of a series of solitary pulses that give rise to the broad band electrostatic noise (BEN) (Scarf et al., 1974; Gurnet et al., 1976) in the plasma sheet boundary layer. These structures were commonly called ESW, i.e. Electrostatic Solitary Waves. However, it was soon noticed that the ESWs do not exhibit the characteristics of solitons or envelope solitons (Matsumoto et al., 1994b). It was further observed that their evolution depended crucially on the ion temperature and electron flow (Kojima et al., 1994). This led Omura et al. (1994) to propose a 1-D BGK mode model to explain these spiky structures. They argued that the BEN/ESW structures can be sustained when the condition $T_i > T_e$ is satisfied (Omura et al., 1994). Further credence to this model was provided by experimental observations of energetic ions in the magnetospheric tail and ring currents (Baumjohann et al., 1989; Buti et al., 1980). However, recent high resolution measurements from POLAR and FAST satellites have revealed 2-D structures for BEN (Franz et al., 1998) which cannot be explained by the previously proposed one-dimensional models (Omura et al., 1996). The observed two-dimensional electrostatic structures consist of both monopolar and bipolar pulses and occur in a region where typically $T_i > T_e$ (Tsurutani et al., 1998). Past theoretical works have pointed out the possibility of exciting linear electron acoustic waves in such a regime (Lashmore-Davies and Martin, 1973) and also discussed their one-dimensional nonlinear properties (Buti, 1980; Bharuthram and Shukla, 1988). There is also reasonable experimental evidence (gleaned from detailed analysis of FAST and Viking data) of the presence of electron acoustic waves associated with BEN activity in auroral plasmas (Pottelette et al., 2001; Dubuloz et al., 1991b). Comparative studies of Singh and Lakhina (2001); Singh et al. (2001) and Dubuloz et al. (1991a) further highlight the consistency between the theoretical models and spacecraft data for both linear and nonlinear regimes. It thus becomes important to examine the nonlinear evolution of these waves and to look for possible two-dimensional saturated stationary structures. We have carried out such an investigation in this paper and have shown that the nonlinear evolution of the electron acoustic wave can be modelled by the system of Davey-Stewartson-I

equations under certain conditions. Further, we have used these conditions to delineate the parametric regions for the existence of dromion solutions for two-dimensional electron acoustic waves.

The paper is organized as follows. In the next section (Sect. 2) we derive a general nonlinear evolution equation for finite amplitude electron acoustic waves which are travelling in an arbitrary direction. The derivation is based on the reductive perturbation method as applied to a two-fluid plasma model. The special case of pure perpendicular wave propagation is considered in Sect. 3 and the conditions necessary for the evolution equations to reduce to the DS-I equations are stated. These conditions are used to numerically delineate the parametric regimes for the occurrence of dromion solutions. An analytic form for the dromion solution using the Hirota bilinear method is discussed as a special case in Appendix B. Section 4 discusses the conditions necessary for the general evolution equation to reduce to the DS-I equations. These conditions are further numerically analyzed. In Sect. 5 some comparisons are made with the actual experimental observations in space and Sect. 6 provides a brief summary and concluding remarks.

2 Derivation of the nonlinear evolution equations

For $T_i > T_e$, the electron acoustic wave is a normal mode of a magnetized plasma with a typical phase velocity of $\sqrt{(T_i/m_e)}$. This mode is analogous to the ion-acoustic wave with the roles of electrons and ions reversed, that is, the electron mass now provides the inertial restoring force to balance the ion thermal pressure force. An appropriate fluid description of this mode is provided by considering the electron equations of continuity and motion and assuming the ions to obey a Boltzmann distribution. The Boltzmann description for ions is justified on account of their large thermal energy which allows one to treat them as an unmagnetized species. We adopt the following normalized set of equations (Ghosh et al., 1996),

$$\frac{\partial n}{\partial t} + \nabla \cdot (n\mathbf{v}) = 0, \quad (1)$$

$$\frac{\partial \mathbf{v}}{\partial t} + (\mathbf{v} \cdot \nabla) \mathbf{v} + 3\beta n \nabla n - \nabla \phi + \alpha (\mathbf{v} \times \mathbf{b}) = 0, \quad (2)$$

$$\nabla^2 \phi = n_e - n_i = n_e - \exp(-\phi), \quad (3)$$

where $\alpha (= \omega_{ce}/\omega_{pe})$ is the ratio of electron cyclotron and electron plasma frequencies, $\beta (= T_e/T_i)$ is the electron to ion temperature ratio, $\mathbf{b} = (0, 0, 1)$ is normalized by the ambient magnetic field B_0 and $\phi (= e|\phi'|/T_i)$ is the normalized electrostatic potential, with ϕ' being the actual measured potential. All the space variables are normalized by the ion Debye length (λ_i), time by reciprocal of electron plasma frequency (ω_{pe}^{-1}), velocities by electron acoustic speed ($c_s = \sqrt{T_i/m_e}$) and number densities by the ambient plasma density n_0 . We now assume a perturbation of the

form $\sim \exp[\iota(\mathbf{k} \cdot \mathbf{r} - \omega t)]$, where $\mathbf{k} = (0, k_\perp, k_\parallel)$ is the wave vector and we adopt the reductive perturbation expansion technique (Nishinari et al., 1994) to derive the nonlinear evolution equation of the perturbation. We expand the physical quantities as

$$n = 1 + \sum_{n=1}^{\infty} \epsilon^n \sum_{l=-\infty}^{\infty} n_1^{(n)} \exp[\iota l(\mathbf{k} \cdot \mathbf{r} - \omega t)], \quad (4)$$

$$\phi = \sum_{n=1}^{\infty} \epsilon^n \sum_{l=-\infty}^{\infty} \phi_1^{(n)} \exp[\iota l(\mathbf{k} \cdot \mathbf{r} - \omega t)], \quad (5)$$

$$\mathbf{v} = \sum_{n=1}^{\infty} \epsilon^n \sum_{l=-\infty}^{\infty} \begin{pmatrix} v_{x1}^{(n)} \\ v_{y1}^{(n)} \\ v_{z1}^{(n)} \end{pmatrix} \exp[\iota l(\mathbf{k} \cdot \mathbf{r} - \omega t)]. \quad (6)$$

We further introduce the following stretched variables (Nishinari et al., 1994)

$$\begin{aligned} \xi &= \epsilon x, & \eta &= \epsilon(y - M_y t) \\ \zeta &= \epsilon(z - M_z t), & \tau &= \epsilon^2 t, \end{aligned} \quad (7)$$

where M_y, M_z are the respective group velocities in the perpendicular and parallel directions. The choice of the stretched variables and the consequent scaling of the various terms is guided primarily by the nature of dispersion of the linear waves, as has been extensively discussed in the literature (Lokenath, 1997). Transforming all independent variables by Eq. (7), we expand Eqs. (1–3) by using Eqs. (4–6) and carry out a systematic balancing of terms at each order of ϵ .

To the first order in ϵ we obtain

$$\epsilon : l = 1; \quad \phi_1^{(1)} = -K_1 n_1^{(1)}, \quad (8)$$

$$\begin{aligned} v_{x1}^{(1)} &= -\iota a_2 k_\perp K_2 n_1^{(1)}, \\ v_{y1}^{(1)} &= a_1 k_\perp K_2 n_1^{(1)}, \\ v_{z1}^{(1)} &= \frac{k_\parallel}{\omega} K_2 n_1^{(1)}, \end{aligned} \quad (9)$$

where, for convenience, the coefficients K_i s, a_i s ($i, j = 1, 2, 3$, etc.) and those occurring in subsequent expressions below are listed in Appendix A. Combining the above expressions leads to the linear dispersion relation for the electron acoustic wave in a warm electron fluid, namely

$$\omega^4 - \omega^2(\alpha^2 + |k|^2 K_2) + \alpha^2 k_\parallel^2 K_2 = 0. \quad (10)$$

For an unmagnetized plasma ($\alpha = 0$) it reduces to the usual linear electron acoustic dispersion

$$\frac{\omega}{k} = \sqrt{K_2},$$

while for a strong magnetic field and perpendicular propagation it reduces to

$$\frac{\omega}{k} = \sqrt{\frac{\alpha^2}{k^2} + K_2},$$

indicating a stronger dispersion which is consistent with our choice of scaling (Lokenath, 1997).

The next order balance yields

$$\epsilon^2 : l = 0; \quad \phi_0^{(2)} = -\left(n_0^{(2)} - K_1^2 |n_1^{(1)}|^2\right), \quad (11)$$

$$\begin{aligned} v_{x0}^{(2)} &= 0, \\ v_{y0}^{(2)} &= -2c_2 |n_1^{(1)}|^2, \end{aligned} \quad (12)$$

$$\epsilon^2 : l = 1; \quad \phi_1^{(2)} = -K_1 \left[n_1^{(2)} + 2\iota K_1 (\mathbf{k} \cdot \nabla_\xi) n_1^{(1)} \right], \quad (13)$$

$$\begin{aligned} v_{x1}^{(2)} &= -\iota K_2 \left(a_2 k_\perp n_1^{(2)} + a_1 \frac{\partial n_1^{(1)}}{\partial \xi} \right) + 2a_2 k_\perp \left(X_{x\eta} \frac{\partial n_1^{(1)}}{\partial \eta} + X_{x\zeta} \frac{\partial n_1^{(1)}}{\partial \zeta} \right), \\ v_{y1}^{(2)} &= K_2 \left(a_1 k_\perp n_1^{(2)} + a_2 \frac{\partial n_1^{(1)}}{\partial \xi} \right) + \iota a_1 k_\perp \left(X_{y\eta} \frac{\partial n_1^{(1)}}{\partial \eta} + X_{y\zeta} \frac{\partial n_1^{(1)}}{\partial \zeta} \right), \\ v_{z1}^{(2)} &= \frac{k_\parallel}{\omega} \left[K_2 n_1^{(2)} + \iota \left(X_{z\eta} \frac{\partial n_1^{(1)}}{\partial \eta} + X_{z\zeta} \frac{\partial n_1^{(1)}}{\partial \zeta} \right) \right], \end{aligned} \quad (14)$$

$$\begin{aligned} \epsilon^2 : l = 2; \quad n_2^{(2)} &= S_n \left(n_1^{(1)} \right)^2, \\ \phi_2^{(2)} &= -\kappa_1 \left\{ S_n - \frac{1}{2} K_1^2 \right\} \left(n_1^{(1)} \right)^2, \end{aligned} \quad (15)$$

$$\begin{aligned}
v_{x2}^{(2)} &= -ik_{\perp} S_x \left(n_1^{(1)}\right)^2, \\
v_{y2}^{(2)} &= k_{\perp} S_y \left(n_1^{(1)}\right)^2, \\
v_{z2}^{(2)} &= k_{\parallel} S_z \left(n_1^{(1)}\right)^2.
\end{aligned} \tag{16}$$

The above relations from the second order balance give us the relations for the group velocities as

$$M_y = 2k_{\perp} \frac{a_1 K_2 - c_1 K_1^2}{1 + c_3 K_2}, \tag{17}$$

$$M_z = 2k_{\parallel} \frac{\frac{\omega}{\omega} - c_1 K_1^2}{1 + c_3 K_2}. \tag{18}$$

Finally, in the third order we obtain

$$\epsilon^3 : l = 1; \quad \phi_1^{(3)} = -K_1 [n_1^{(3)} + K_1 \{ \nabla_{\xi}^2 n_1^{(1)} + 2i (\mathbf{k} \cdot \nabla_{\xi}) n_1^{(2)} - n_1^{(1)} n_0^{(2)} \} - 4K_1^2 (\mathbf{k} \cdot \nabla_{\xi})^2 n_1^{(1)} + K_1^3 n_1^{(1)} |n_1^{(1)}|^2]. \tag{19}$$

Using Eqs. (8)–(19) in Eqs. (1)–(3) and solving for the third order of ϵ we obtain the following set of coupled equations

$$-M_y^2 \frac{\partial^2 n_0^{(2)}}{\partial \eta^2} - 2M_y M_z \frac{\partial^2 n_0^{(2)}}{\partial \eta \partial \zeta} + (3\beta + 1 - M_z^2) \frac{\partial^2 n_0^{(2)}}{\partial \zeta^2} + p_{\eta\eta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \eta^2} + p_{\eta\zeta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \eta \partial \zeta} + p_{\zeta\zeta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \zeta^2} = 0, \tag{20}$$

$$-M_y^2 \frac{\partial^2 v_{z0}^{(2)}}{\partial \eta^2} - 2M_y M_z \frac{\partial^2 v_{z0}^{(2)}}{\partial \eta \partial \zeta} + (3\beta + 1 - M_z^2) \frac{\partial^2 v_{z0}^{(2)}}{\partial \zeta^2} + q_{\eta\eta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \eta^2} + q_{\eta\zeta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \eta \partial \zeta} + q_{\zeta\zeta} \frac{\partial^2 |n_1^{(1)}|^2}{\partial \zeta^2} = 0 \tag{21}$$

for $l = 0$ and

$$i \frac{\partial n_1^{(1)}}{\partial \tau} + \mu_{\xi\xi} \frac{\partial^2 n_1^{(1)}}{\partial \xi^2} + \mu_{\eta\eta} \frac{\partial^2 n_1^{(1)}}{\partial \eta^2} + \mu_{\eta\zeta} \frac{\partial^2 n_1^{(1)}}{\partial \eta \partial \zeta} + \mu_{\zeta\zeta} \frac{\partial^2 n_1^{(1)}}{\partial \zeta^2} - \left(\mu_{02} n_0^{(2)} + k_{\parallel} v_{z0}^{(2)} \right) n_1^{(1)} - \mu_{11} |n_1^{(1)}|^2 n_1^{(1)} = 0 \tag{22}$$

for $l = 1$, where all the coefficients are given in Appendix A. Equations (20)–(22) provide a generalized description of the evolution of the nonlinear wave packet in three dimensions. They can be further simplified by assuming special situations which we discuss in the subsequent sections.

3 Perpendicular propagation of the wave

For pure perpendicular propagation of the wave we can set $k_{\parallel} = M_z = 0$. We can further assume $\partial_{\xi} \rightarrow 0$. Due to the symmetry of the system, all the cross derivative terms then vanish. Substituting these conditions, Eqs. (22) and (20), respectively, reduce to

$$i \frac{\partial n_1^{(1)}}{\partial \tau} + \mu_{\eta\eta} \frac{\partial^2 n_1^{(1)}}{\partial \eta^2} + \mu_{\zeta\zeta} \frac{\partial^2 n_1^{(1)}}{\partial \zeta^2} - \left(\mu_{02} n_0^{(2)} + \mu_{11} |n_1^{(1)}|^2 \right) n_1^{(1)} = 0, \tag{23}$$

$$-M_y^2 \frac{\partial^2 n_0^{(2)}}{\partial \eta^2} + (3\beta + 1) \frac{\partial^2 n_0^{(2)}}{\partial \zeta^2} + K_4 \frac{\partial^2 |n_1^{(1)}|^2}{\partial \eta^2} = 0. \tag{24}$$

Substituting

$$n_1^{(1)} = A; \quad n_0^{(2)} = Q,$$

and

$$\mu_{\eta\eta} = d_{\eta}; \quad \mu_{\zeta\zeta} = d_{\zeta},$$

Eqs. (23) and (24) can be rewritten as

$$i A_{\tau} + d_{\eta} A_{\eta\eta} + d_{\zeta} A_{\zeta\zeta} - \left(\mu_{02} Q + \mu_{11} |A|^2 \right) A = 0, \tag{25}$$

$$-M_y^2 Q_{\eta\eta} + (3\beta + 1) Q_{\zeta\zeta} + K_4 |A|_{\zeta\zeta}^2 = 0, \tag{26}$$

where the subscripts of the dependent variables A and Q denote partial derivatives. The two coupled Eqs. (25) and (26) have the form of the DS-I equations, provided the condition

$$d_{\eta}/d_{\zeta} > 0; \quad \mu_{11} > 0 \tag{27}$$

is satisfied, with μ_{02} being always positive. Condition (27) provides us with the necessary criterion to explore the physical parameter space where dromion structures can appear as nonlinear saturated states of the electron acoustic wave. We have carried out this exploration numerically and Figs. 1a–c represent the parameter space at different values of β (e.g. $\beta=0, 0.1$ and 1 , respectively) for which the condition (27) is satisfied. In these parameter regions, the nonlinear evolution for electron acoustic waves is governed by the DS-I equation and hence may admit dromion or breather type solutions. For some special values of the coefficients, the coupled equations (viz. Eqs. 25 and 26) become analytically integrable and thus, when solved for appropriate initial-boundary conditions, give dromions. One such technique is the Hirota bilinear method, as discussed, for example in Radha and Lakshmanan (1997a); Radha and Lakshmanan (1997b). We have illustrated this method in Appendix B and a typical form of

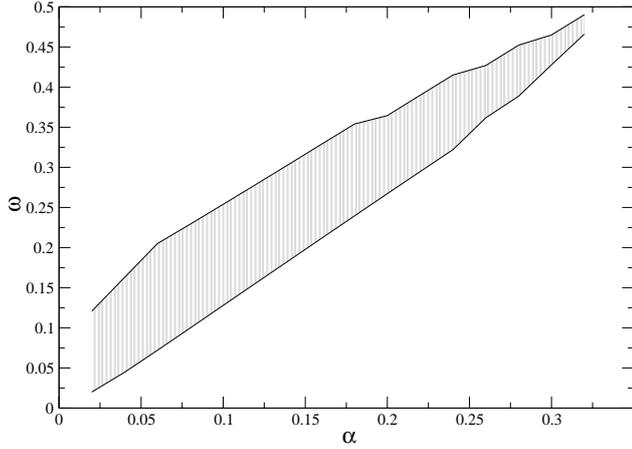


Fig. 1a. Existence domain for DS-I equation ; $\beta = 0$.

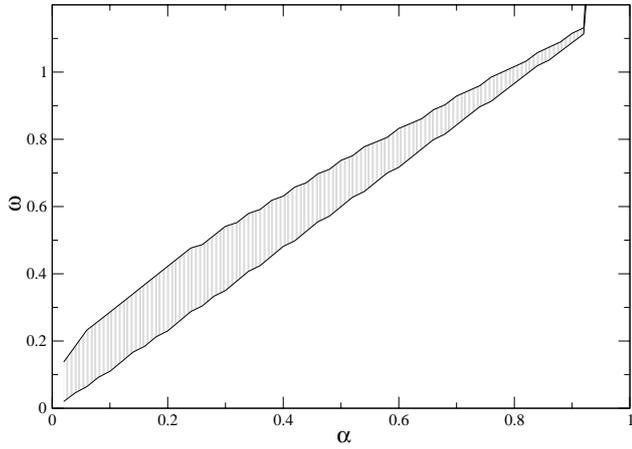


Fig. 1b. Existence domain for DS-I equation ; $\beta = 0.1$.

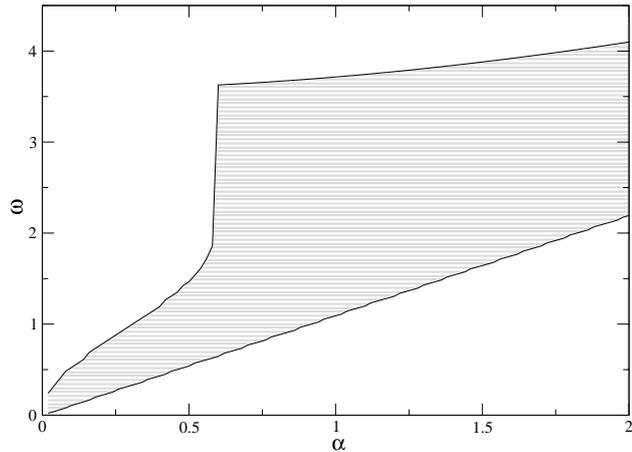


Fig. 1c. Existence domain for DS-I equation ; $\beta = 1$.

the solution is obtained (Eq. B19). Using an appropriate realistic set of parameters ($\alpha=0.9$, $\omega=1.098$ from Fig. 1b) which allows the analytic solution for DS-I equations (see Appendix B) and putting them into Eq. (B19), we have obtained an actual analytical solution for dromion, as shown in

Fig. 2. For convenience, the shape of the dromion is assumed to be spherical (i.e. equal width ratio) while the time is taken to be $t = 0$. However, the analytic method is restricted only to very specific sets of parameters and for a more general case, one needs to solve the DS-I equations numerically.

In the present paper our primary aim has been to delineate the parametric regions where the evolution equations take the form of the DS-I equations and to further demonstrate from available experimental data that these parametric regimes are physically relevant in the PCBL region. At the moment we note from Figs. 1a–c that the area of the parametric region where condition (27) is satisfied is larger for larger values of β (i.e. it increases as $\beta \rightarrow 1$). Since space plasma observations show an appreciably large T_e for the $T_i \geq T_e$ (Gage and Reid, 1981), our analysis seems to support the possibilities of finding dromion solutions for those regimes. It is to be noted that for the case of $\beta \sim 1$, the electron acoustic wave tends to become damped due to Landau damping in a pure proton-electron (single species) plasma (Lashmore-Davies and Martin, 1973). However, since the electron component of the auroral plasma often displays a two-temperature character (i.e. a hot and a cold electron species), the electron acoustic wave may still survive in that region without any appreciable damping. Our fluid model treatment is strictly valid when such kinetic effects can be ignored, e.g. in the region of small β (say, $\beta_{\max} \approx 0.1$), where electron temperatures are not too large. Our further analysis is primarily restricted to such regimes.

4 Arbitrary propagation

We now discuss the case of arbitrary propagation, i.e. the case when $k_{\parallel} \neq 0$. Combining Eqs. (20)–(22) we obtain

$$-M_y^2 \frac{\partial^2 Q'}{\partial \eta^2} - 2M_y M_z \frac{\partial^2 Q'}{\partial \eta \partial \zeta} + (3\beta + 1 - M_z^2) \frac{\partial^2 Q'}{\partial \zeta^2} + r_{\eta\eta} \frac{\partial^2 |A|^2}{\partial \eta^2} + r_{\eta\zeta} \frac{\partial^2 |A|^2}{\partial \eta \partial \zeta} + r_{\zeta\zeta} \frac{\partial^2 |A|^2}{\partial \zeta^2} = 0, \quad (28)$$

$$i \frac{\partial A}{\partial \tau} + \mu_{\eta\eta} \frac{\partial^2 A}{\partial \eta^2} + \mu_{\eta\zeta} \frac{\partial^2 A}{\partial \eta \partial \zeta} + \mu_{\zeta\zeta} \frac{\partial^2 A}{\partial \zeta^2} - \mu_{02} Q' A - \mu_{11} |A|^2 A = 0, \quad (29)$$

where

$$Q' = n_0^{(2)} + \frac{k_{\parallel}}{\mu_{02}} v_{z0}^{(2)} ; \quad r_{\alpha\beta} = p_{\alpha\beta} + \frac{k_{\parallel}}{\mu_{02}} q_{\alpha\beta}.$$

The conditions under which the above equations can reduce to the DS-I equations are not as simple as was the case for the perpendicular propagation limit. In addition, in order to satisfy a criterion analogous to Eq. (27), we will also need to ensure that all three cross derivative terms of Eqs. (28) and (29) vanish simultaneously. To derive these conditions it is convenient to adopt the following linear transformation of the independent variables

$$\partial_{\eta} = p_y \partial_y + p_z \partial_z ; \quad \partial_{\zeta} = q_y \partial_y + q_z \partial_z,$$

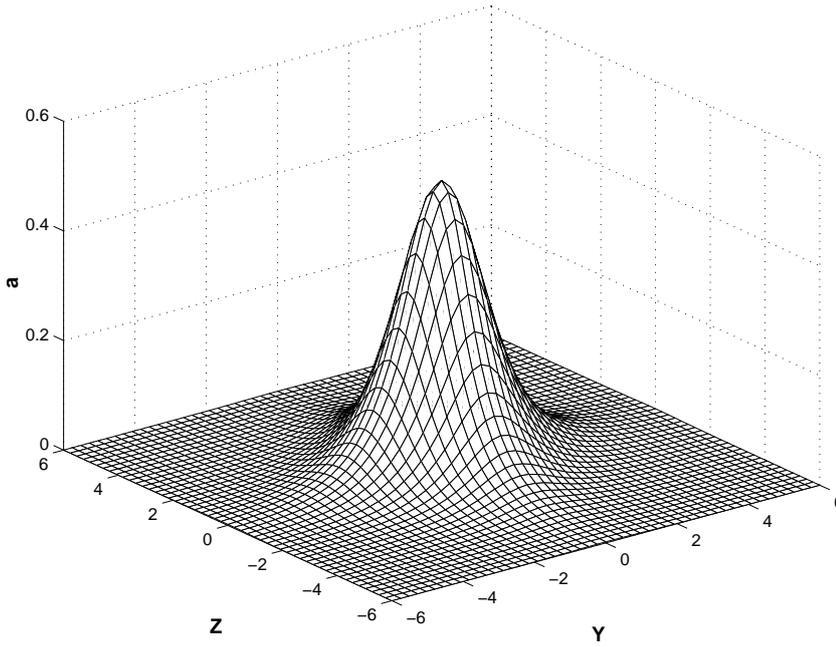


Fig. 2. Analytic solution for a dromion.

where p_y, p_z, q_y and q_z are real numbers. The condition for all the terms containing $\partial_y \partial_z$ to be simultaneously zero is given by

$$\hat{\mathbf{M}} \cdot \hat{\mathbf{T}} = 0, \quad (30)$$

where $\hat{\mathbf{M}}$ is a 3×3 matrix consisting of the coefficients of Eqs. (28) and (29) and $\hat{\mathbf{T}}$ is a three-component column vector consisting of the transform coefficients, viz. $p_{y,z}, q_{y,z}$,

and are given in Appendix A. Equation (30) can have non-trivial solutions if and only if $\det |\hat{\mathbf{M}}| = 0$, which leads to the condition

$$\mu_{\eta\eta} + \mu_{\zeta\zeta} (e_1 - e_2 e_3) - e_3 \mu_{\eta\zeta} = 0. \quad (31)$$

The elliptic-hyperbolic nature of the Eqs. (28) and (29) is ensured by

$$\mu_{\eta\eta} \mu_{\zeta\zeta} (1 + E^2 t^4) + \mu_{\eta\zeta} (\mu_{\eta\eta} + \mu_{\zeta\zeta} E) (ET^3 + t) + (\mu_{\eta\eta}^2 + \mu_{\eta\zeta}^2 E + \mu_{\zeta\zeta}^2 E^2) t^2 > 0, \quad (32)$$

$$e_1 (1 + E^2 t^4) + e_2 (E - e_1) (ET^3 + t) - (e_1^2 + e_2^2 E + E^2) t^2 > 0, \quad (33)$$

where $t = p_z/q_z$ and $E = e_1 - e_2 e_3$ and the interrelations among the transformation coefficients, viz. $p_{y,z}, q_{y,z}$, have been taken into account. Provided Eq. (31) is satisfied, t can be determined from the following second order equation (see Appendix A)

$$Et^2 + 2e_3 t + 1 = 0. \quad (34)$$

Equations (31) to (34), together with the relation $\mu_{11} > 0$, complete the set of conditions necessary for ensuring that the generalized (oblique) propagation equations reduce to the DS-I equations. It is to be noted that the Eq. (31) has necessarily assumed that $k_{\parallel} \neq 0$. The existence of the DS-I equations is satisfied only for a specific set of transformations satisfying Eq. (34). We have once again carried out a numerical exploration in parameter space of these conditions. Figures 3a and b present the parametric domain of the existence of the DS-I equations in the three-dimensional parameter space of α, k_y and ω for $\beta = 0$ and 0.1, respec-

tively. Figures 4a and b represent the corresponding two-dimensional projections in the α - ω plane, respectively, for different k_y and k_z , whereas Figs. 5a and b represent similar projections in the α - θ plane, with θ being the angle of propagation with respect to the ambient magnetic field. We note that, analogous to the perpendicular propagation case, for oblique propagation also, an increasing β favors the conditions of DS-I equations (e.g. Figs. 3a and b, respectively). But in contrast to the previous case, the allowed range of β values is found to be more restricted for an oblique propagation. In fact, there appears to be an upper cutoff in β (i.e. $\beta \leq \beta_c$, where $\beta_c \approx 0.17$) beyond which the equations cannot be reduced to the DS-I form.

5 Comparison with the space plasma observations

Our numerical results presented in the previous sections display the regions in parameter space where the propagation

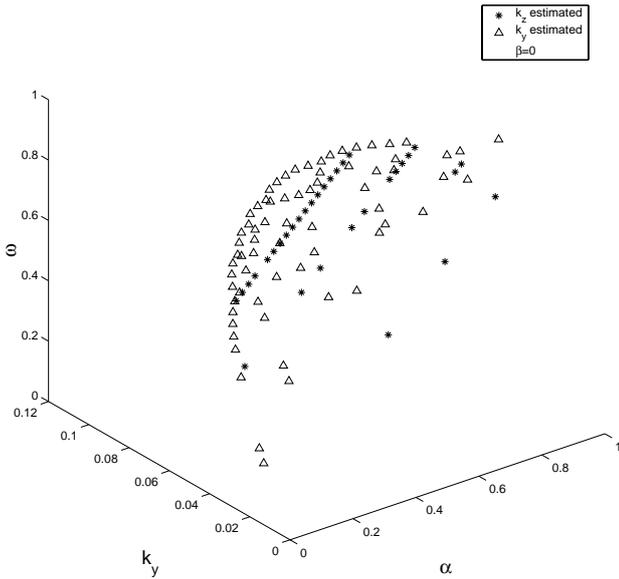


Fig. 3a. Existence domain for DS-I equation for oblique propagation; $\beta = 0$.

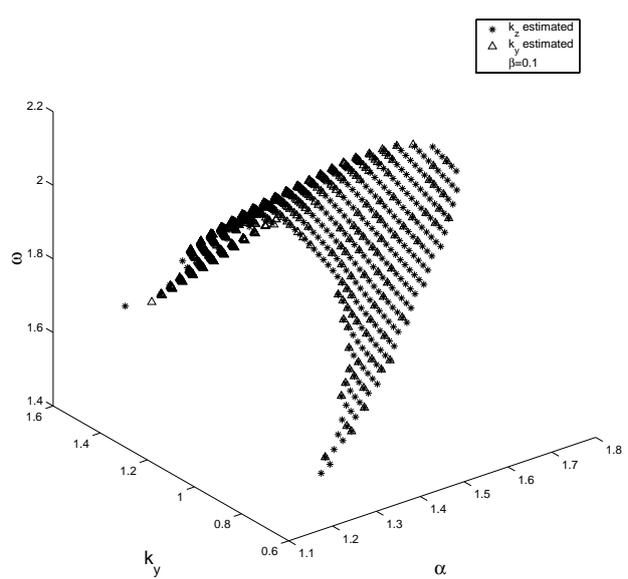


Fig. 3b. Existence domain for DS-I equation for oblique propagation; $\beta = 0.1$.

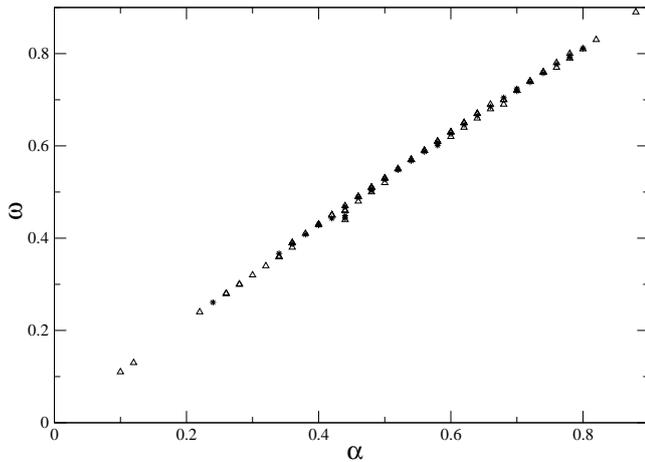


Fig. 4a. α vs. ω for oblique propagation $\beta = 0$.

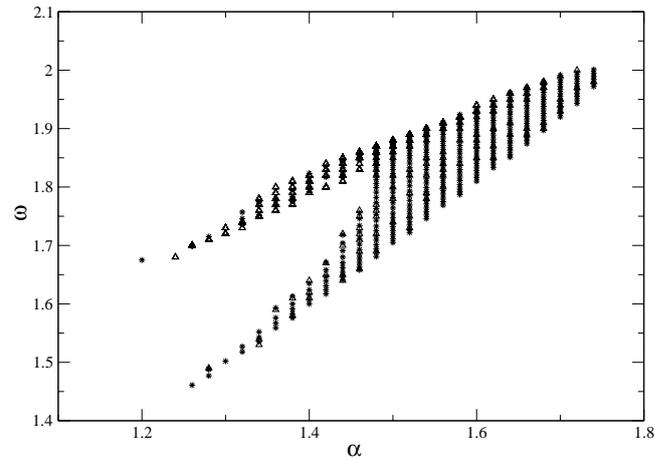


Fig. 4b. α vs. ω for oblique propagation; $\beta = 0.1$.

equation can be represented by the DS-I equations. For the case of perpendicular propagation, where condition (27) needs to be satisfied, the parameter space is represented by α and ω , with ω being calculated from the dispersion rela-

tion (10). Figure 1a displays the results for cold electrons (i.e. $\beta = 0$), while Figs. 1b and 1c represent $\beta = 0.1$ and $\beta = 1$, respectively. We now check to see if these parametric regions include the physical parameter range corresponding

Table 1. Wave frequencies for different f_{ce}

f_{ce} (kHz)	α (normalized)	ω (normalized)	$1/f\omega$ (ms)
2	0.2	0.3	0.33
5	0.5	0.6	0.17
8	0.8	0.9	0.11

Table 2. Width ratios for different α

α	k	ω	ρ
0.2	0.2	0.301	0.238
0.5	0.35	0.629	0.635
0.8	0.55	0.981	0.849

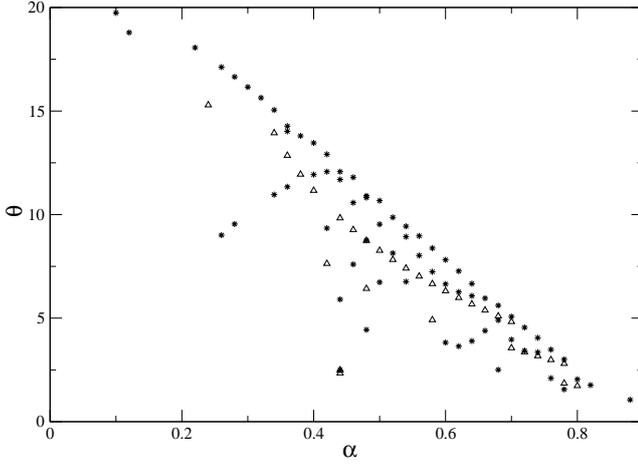


Fig. 5a. α vs. θ for oblique propagation; $\beta = 0$.

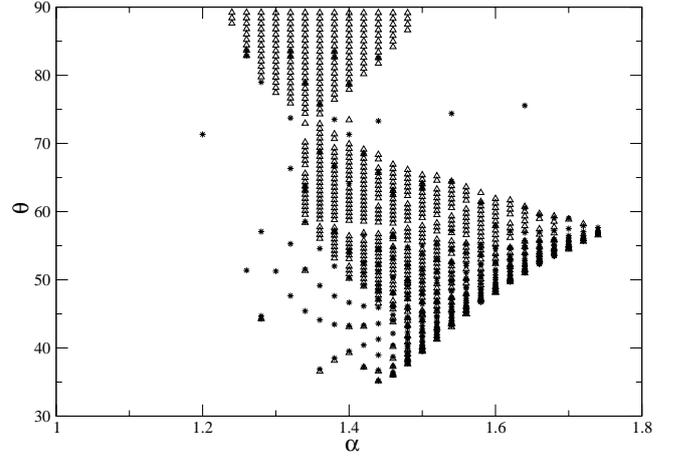


Fig. 5b. α vs. θ for oblique propagation; $\beta = 0.1$.

to experimental observations. The typical measured parameters in the region of the satellite observations (at a height of 5 to 7 Earth radii) are $T_i = 100$ eV, $T_e = 10$ eV (i.e. $\beta = 0.1$), $n \simeq 1$ cm $^{-3}$ ($f_{pe} \simeq 10$ kHz) and $f_{ce} \simeq 5$ to 8 kHz (i.e. $\alpha = 0.5$ to 0.8) (Tsurutani et al., 1998). We have chosen three values of f_{ce} (α) and computed the corresponding ω values from Fig. 1. The result is summarized in Table 1. Since the experimentally observed duration of the nonlinear structures is of the order of 1 ms or less (Tsurutani et al., 1998), our calculated values of ω appear to be consistent with the observations. The narrowing of the electronic pulse with increasing α observed in Table 1 is maintained for other β values, as indicated in Figs. 1a–c. The same trend continued for arbitrary (oblique) propagation as well which is evident from Figs. 4a and b, respectively.

In Fig. 2, we have shown a possible dromion solution for a chosen set of parameters (i.e. $\alpha = 0.9$, $\omega = 1.098$, with β being 0.1) at time $t = 0$ (Appendix B). For simplicity, we have assumed equal width for both directions. We have compared the shape and size of our solution with that of the experimentally observed structures where the physical conditions are assumed to remain the same (i.e. $f_{pe} = 10$ kHz, $T_i = 100$ eV, β being 0.1). According to Franz et al. (1998), the typical parallel scale sizes of the coherent electric field structures corresponding to recent POLAR observations are of the order of 100–1000 m, whereas the calculated half width of the dromion in Fig. 2 is around 166.8 m (2.5 Debye length), which agrees fairly well with the experimental observations. Regarding the perpendicular scale size, we have calculated the width ratio ($\rho = L_{\parallel}/L_{\perp}$, L_{\parallel} (L_{\perp}) being the parallel (perpendicular) scale sizes, respectively) from the following analytic relation derived by Franz et al. (2000)

$$\rho \simeq \left(1 + r_L^2\right)^{-1/2}, \quad (35)$$

where r_L is the normalized electron gyroradius. Setting $r_L = M_y/\alpha$ and using Eq. (18), we have obtained $\rho = 1.035$, indicating a near spherical structure (i.e. $L_{\parallel} \approx L_{\perp}$) for our chosen set of parameters. This also confirms the validity

of our assumption of a spherical dromion in Fig. 2. In order to find out the amplitude ($a = 0.52$ in Fig. 2) we have used Eq. (8) and the corresponding gradient is numerically estimated from Fig. 2. The calculated amplitude is around 0.75 mV/m while the typical amplitudes of the structures are reported as being of the order of 1 mV/m (Franz et al., 2000). Hence, our calculated estimation seems to agree quite satisfactorily with the recent spacecraft observations in the PCBL regions. The POLAR observations have further indicated that the shapes of the two dimensional structures depend crucially on α (Franz et al., 2000), being nearly spherical for $\alpha > 1$ while becoming more oblate ($L_{\parallel} < L_{\perp}$) with decreasing α (i.e. $\alpha < 1$). In Table 2 we have calculated ρ by using Eq. (35) for three different values of α and k which satisfy the condition for DS-I equations (Fig. 1b). The decreasing value of ρ with decreasing α indicates an expected oblate shape for the possible solutions. Our theoretical model thus remain consistent with present observational data.

The question of the direction of propagation of the nonlinear pulses has not been satisfactorily established yet. Earlier studies by Gurnet et al. (1976) and Onsager et al. (1993) indicated the possibilities of perpendicular propagation but later observations (Kojima et al., 1997) found inconsistencies with this conclusion. In fact, most of the current experimental observations (Franz et al., 1998; Ergun et al., 1998) and some of the past analytic studies (Goldman et al., 1999; Muschietti et al., 1999) emphasize that propagation of these nonlinear pulses is near parallel. Our theoretical analysis for oblique propagation provides a framework for looking at such a possibility. From Fig. 5a we see that dromion solutions with near parallel propagation ($\theta < 10^\circ$) are also possible in the region of $\alpha \sim 0.5$ for very low electron temperatures. With increasing α ($\alpha \geq 0.7$), the propagation angle seems to reduce further ($\theta \leq 5^\circ$; Fig. 5a), strengthening the possibilities of such a near parallel propagation. However, in the absence of more definitive data on the direction of propagation of the pulses, it is not very meaningful to define any realistic quantitative bounds on the existence domain of these structures

on the basis of our model.

6 Conclusion

In the present work, we have shown that the nonlinear evolution of two-dimensional electron acoustic waves can be modelled by the Davey-Stewartson-I equations. These equations admit dromion solutions which are generalizations of soliton solutions to two dimensions and exhibit exponential spatial decay in both directions. Since experimentally observed potential structures in the PCBL region appear to have a distinctly two-dimensional structure, it would be meaningful to characterize them in terms of such two-dimensional nonlinear entities. The basis for the present description rests on the primary assumption of the excitation and sustenance of electron acoustic waves in the PCBL region. Such a proposition for one-dimensional wave has been investigated by many authors (Buti, 1980; Dubuloz et al., 1991a,b; Singh et al., 2001). Using high resolution data of FAST, Pottelette et al. (2001) presented evidence for the presence of nonlinear electron acoustic waves in the auroral kilometric radiation source region. Experimental evidence pointing to the existence of energetic ions ($T_i > T_e$) in the PCBL region (Tsurutani et al., 1998) further supports the assumption of sustaining electron acoustic wave in that region (Buti, 1980; Omura et al., 1994). In fact, there is considerable past work devoted to the question of the generation mechanism of this wave (Bharuthram and Shukla, 1988; Singh and Lakhina, 2001) and its stability (Lashmore-Davies and Martin, 1973). Our present work can be viewed as a natural extension of these investigations towards the description of a nonlinear saturated state of the wave in two spatial dimensions. One possible saturation mechanism of such two-dimensional structures may be the modulational instability as observed experimentally by Pottelette et al. (1999). This can be conjectured further in the context of the stability analysis presented by Nishinari et al. (1994) for the ion acoustic wave. It should be mentioned that the physical saturation mechanisms for a two-dimensional

electron acoustic wave would be quite analogous and similar in nature to what has been proposed for the case of the ion acoustic wave (Nishinari et al., 1994). In other words, a two-dimensional localization can arise due to the simultaneous onset of a self-focusing instability in the transverse direction and a modulational instability in the longitudinal direction. The evolution and saturation of such an unstable fluctuation into a dromion state can happen if the DS-I condition is satisfied. In our present work we have used this condition to define a parametric space relating the plasma equilibrium quantities and wave propagation characteristics. We have delineated this parameter region by a detailed numerical investigation of the constraint conditions both for the special case of perpendicular propagation and for the general case of oblique propagation. Comparisons with available experimental data show that dromion solutions can exist both for perpendicular propagation, as well as for nearly parallel propagation of the electron acoustic mode. The shape and size of our analytical solution was found to agree well with those observed in POLAR spacecraft measurements. This further strengthens the possibilities of the physical existence of dromions in the auroral plasmas. However, for a more quantitative understanding of the shape and finer features of the observed structure, one needs to carry out a detailed numerical investigation of the evolution equation with appropriate initial conditions. Such an investigation is in progress and will be reported elsewhere. Our present motivation was to provide an alternate paradigmatic model for the observed structures and to demonstrate the basis of its applicability. Further, it should also be pointed out that there can be competing processes in the PCBL region involving other low frequency waves, such as the lower hybrid waves (Dubuloz et al., 1991a). A nonlinear analysis of such waves, along the present lines, to examine two-dimensional solutions would be very worthwhile and to the best of our knowledge has not been accomplished to date. Thus, dromions offer a rich and new paradigm for understanding plasma wave phenomena in the auroral plasma and need to be investigated more intensely.

Appendix A Coefficients for the equations

$$\begin{aligned}
 a_1 &= \frac{\omega}{\omega^2 - \alpha^2 b_z^2}, & a_2 &= \frac{\alpha b_z}{\omega^2 - \alpha^2 b_z^2}, & a_3 &= \frac{\omega^2 + \alpha^2 b_z^2}{\omega^2 - \alpha^2 b_z^2}, & a_{12}^2 &= a_1^2 + a_2^2, & a_4 &= a_1^2 + 3a_2^2 \\
 K_1 &= \frac{1}{1 + |k|^2}, & K_2 &= 3\beta + K_1, & K_3 &= 3\beta - K_1^2, & K_4 &= K_2 + K_3, & K_5 &= 3\beta + K_1^2 \\
 c_1 &= a_1 k_{\perp}^2 + \frac{k_{\parallel}^2}{\omega}, & c_2 &= a_1 k_{\perp} K_2, & c_3 &= a_{12}^2 k_{\perp}^2 + \frac{k_{\parallel}^2}{\omega^2}, & c_4 &= a_1 a_4 k_{\perp}^2 + \frac{k_{\parallel}^2}{\omega^3} \\
 a_{1\omega} &= \frac{2\omega}{4\omega^2 - \alpha^2 b_z^2}, & a_{2\omega} &= \frac{\alpha b_z}{4\omega^2 - \alpha^2 b_z^2}, & a_{3\omega} &= \frac{\omega^2 + 2\alpha^2 b_z^2}{\omega^2 - \alpha^2 b_z^2} \\
 a_{12\omega}^2 &= a_1 a_{1\omega} + a_2 a_{2\omega}, & a'_{12} &= 2\omega a_1 + \alpha b_z a_2, & a'_{21} &= \alpha b_z a_1 + 2\omega a_2, & a_{4\omega} &= a_{1\omega} a_{3\omega}
 \end{aligned}$$

$$\kappa_1 = \frac{1}{1 + 4|k|^2}, \quad \kappa_2 = 3\beta + \kappa_1, \quad \kappa_3 = 3\beta - \kappa_1 K_1^2, \quad \kappa_4 = K_2 + \kappa_3$$

$$c_{1\omega} = a_{1\omega} k_{\perp}^2 + \frac{k_{\parallel}^2}{2\omega}, \quad c_{2\omega} = a_{1\omega} a'_{12} k_{\perp}^2 + \frac{k_{\parallel}^2}{\omega}, \quad c_{3\omega} = a_{12\omega}^2 k_{\perp}^2 + \frac{k_{\parallel}^2}{2\omega^2}, \quad c_{4\omega} = a_1 a_{4\omega} k_{\perp}^2 + \frac{k_{\parallel}^2}{\omega^2}$$

$$X_{x\eta} = a_1 K_2 M_y + k_{\perp} K_1^2 - \frac{1}{2k_{\perp}} K_2, \quad X_{x\zeta} = a_1 K_2 M_z + k_{\parallel} K_1^2,$$

$$X_{y\eta} = a_3 K_2 \frac{M_y}{\omega} + 2k_{\perp} K_1^2 - \frac{1}{k_{\perp}} K_2, \quad X_{y\zeta} = a_3 K_2 \frac{M_z}{\omega} + 2k_{\parallel} K_1^2,$$

$$X_{z\eta} = K_2 \frac{M_y}{\omega} + 2k_{\perp} K_1^2, \quad X_{z\zeta} = K_2 \frac{M_z}{\omega} + 2k_{\parallel} K_1^2 - \frac{1}{k_{\parallel}} K_2$$

$$S_x = 2a_{2\omega} \kappa_2 S_n + f_x, \quad S_y = 2a_{1\omega} \kappa_2 S_n + f_y, \quad S_z = \frac{\kappa_2}{\omega} S_n + f_z, \quad S_n = \frac{\omega + f_{yz}}{\omega - 2c_{1\omega} \kappa_2}$$

$$f_x = \frac{1}{2} a_{1\omega} a'_{21} K_2 + a_{2\omega} \kappa_3, \quad f_y = \frac{1}{2} a_{1\omega} a'_{12} K_2 + a_{1\omega} \kappa_3, \quad f_z = \frac{1}{2\omega} (K_2 + \kappa_3) = \frac{\kappa_4}{2\omega},$$

$$f_{yz} = k_{\perp}^2 f_y + k_{\parallel}^2 f_z = \frac{1}{2} c_{2\omega} K_2 + c_{1\omega} \kappa_3, \quad f_k = c_{4\omega} K_2^2 + 2c_{3\omega} K_2 (2\kappa_2 S_n + \kappa_3) + \kappa_4 S_n - \kappa_1 K_1^2$$

$$p_{\eta\eta} = 0, \quad p_{\eta\zeta} = 2M_y K_2 \frac{k_{\parallel}}{\omega} (1 + a_2^2 k_{\perp}^2 K_2), \quad p_{\zeta\zeta} = 2M_z K_2 \frac{k_{\parallel}}{\omega} (1 + a_2^2 k_{\perp}^2 K_2) + K_4$$

$$q_{\eta\eta} = 2 \frac{k_{\parallel}}{\omega} (a_2 k_{\perp} K_2)^2 M_y^2, \quad q_{\eta\zeta} = 4 \frac{k_{\parallel}}{\omega} (a_2 k_{\perp} K_2)^2 M_y M_z + M_y K_4, \quad q_{\zeta\zeta} = 2 \frac{k_{\parallel}}{\omega} [(a_2 k_{\perp} K_2)^2 M_z^2 + (3\beta + 1) K_2] + K_4 M_z$$

$$\mu_{02} = \frac{c_1 K_4}{1 + c_3 K_2}, \quad \mu_{11} = \frac{c_1 f_k}{1 + c_3 K_2} - 2c_2 k_{\perp}, \quad \mu_{\xi\xi} = \frac{a_1 K_2 - c_1 K_1^2}{1 + c_3 K_2},$$

$$\mu_{\eta\eta} = \frac{1}{1 + c_3 K_2} [c_4 K_2 M_y^2 + 2k_{\perp} M_y (c_3 K_1^2 - a_{12}^2 K_2) + 4k_{\perp}^2 K_1^2 (c_1 K_1 - a_1) + (a_1 K_2 - c_1 K_1^2)]$$

$$\mu_{\eta\zeta} = \frac{2}{1 + c_3 K_2} [c_4 K_2 M_y M_z + c_3 K_1^2 (k_{\perp} M_z + k_{\parallel} M_y) - K_2 (a_{12}^2 k_{\perp} M_z + \frac{1}{\omega^2} k_{\parallel} M_y) - 2k_{\perp} k_{\parallel} K_1^2 (a_1 + \frac{1}{\omega} - 2c_1 K_1)]$$

$$\mu_{\zeta\zeta} = \frac{1}{1 + c_3 K_2} [c_4 K_2 M_z^2 + 2k_{\parallel} M_z (c_3 K_1^2 - \frac{K_2}{\omega^2}) + 4k_{\parallel}^2 K_1^2 (c_1 K_1 - \frac{1}{\omega}) + (\frac{K_2}{\omega} - c_1 K_1^2)]$$

$$e_1 = \frac{M_y^2}{3\beta + 1 - M_z^2}, \quad e_2 = \frac{2M_y M_z}{3\beta + 1 - M_z^2}, \quad e_3 = \frac{R_{\eta\eta} + e_1 R_{\zeta\zeta}}{R_{\eta\zeta} + e_2 R_{\zeta\zeta}}$$

$$R_{\alpha\beta} = \frac{r_{\alpha\beta}}{1 + 3\beta - M_z^2}$$

$$\hat{M} = \begin{pmatrix} \mu_{\eta\eta} & \mu_{\eta\zeta} & \mu_{\zeta\zeta} \\ R_{\eta\eta} & R_{\eta\zeta} & R_{\zeta\zeta} \\ -e_1 & -e_2 & 1 \end{pmatrix}; \quad \hat{T} = \begin{pmatrix} 2p_y p_z \\ p_y q_z + p_z q_y \\ q_y q_z \end{pmatrix}$$

Provided $\det|\hat{M}|=0$

$$\frac{q_y}{p_y} = E \frac{p_z}{q_z}; \quad \frac{q_z}{p_z} + \frac{q_y}{p_y} = -2e_3$$

Assuming $\frac{p_z}{q_z} = t$,

$$t = \frac{-e_3 \pm \sqrt{e_3^2 - E}}{E}$$

$$l_y = \frac{2\mu_{11}}{K_4 \mu_{02}} \frac{d_{\zeta}}{d_{\eta}} M_y^2, \quad l_z = \frac{2\mu_{11}}{K_4 \mu_{02}} (3\beta + 1)$$

Appendix B Hirota bilinear method

We briefly outline the Hirota bilinear method here and apply it to obtain an exact analytic one-dromion solution for the perpendicular propagation case. We first rescale the dependent and independent variables as

$$\eta = \sqrt{\frac{2d_\eta}{\mu_{11}}} Y; \quad \zeta = \sqrt{\frac{2d_\zeta}{\mu_{11}}} Z; \quad \tau = \frac{1}{\mu_{11}} T, \\ \frac{\mu_{02}}{\mu_{11}} Q = q; \quad A = a. \quad (\text{B1})$$

Using Eq. (B1), Eqs. (25) and (26) transform to

$$i a_T + \frac{1}{2} (a_{YY} + a_{ZZ}) - (q + |a|^2) a = 0, \quad (\text{B2})$$

$$l_y q_{YY} - l_z q_{ZZ} - 2|a|^2_{ZZ} = 0, \quad (\text{B3})$$

where l_y and l_z are given in Appendix A.

Putting the usual substitutions (Radha and Lakshmanan, 1997a)

$$a = \frac{g}{f}; \quad q = -2\partial_z^2 \ln f, \quad (\text{B4})$$

with the assumption $l_y, l_z \approx 1$ and rotating the coordinate axis by 45° , Eqs. (B2) and (B3) reduce to the bilinear form as

$$\left[i D_T + \frac{1}{2} (D_Y^2 + D_Z^2) \right] g \cdot f = 0, \\ D_Y D_Z f \cdot f = -2 g \cdot g^*, \quad (\text{B5})$$

where $D_{Y,Z}$ are the Hirota bilinear operators. Expanding the new variables as

$$g = \epsilon g^{(1)} + \epsilon^3 g^{(3)} + \dots, \\ f = 1 + \epsilon^2 f^{(2)} + \epsilon^4 f^{(4)} + \dots, \quad (\text{B6})$$

with ϵ being a small quantity, and equating for different orders we obtain

$$\epsilon : i g_T^{(1)} + \frac{1}{2} (g_{YY}^{(1)} + g_{ZZ}^{(1)}) = 0, \quad (\text{B7})$$

$$\epsilon^2 : f_{YZ}^{(2)} = g^{(1)} g^{*(1)}, \quad (\text{B8})$$

$$\epsilon^3 : i g_T^{(3)} + \frac{1}{2} (g_{YY}^{(3)} + g_{ZZ}^{(3)}) = \\ - \left[i D_T + \frac{1}{2} (D_Y^2 + D_Z^2) \right] g^{(1)} \cdot f^{(2)}, \quad (\text{B9})$$

$$\epsilon^4 : f_{YZ}^{(4)} + \frac{1}{2} D_Y D_Z (f^{(2)} \cdot f^{(2)}) = \\ (g^{(1)} \cdot g^{*(3)} + g^{(3)} \cdot g^{*(1)}) \quad (\text{B10})$$

and so on. Solving Eq. (B7) we obtain

$$g^{(1)} = \sum_{j=1}^N \exp \chi_j, \quad \chi_j = p_j Y + s_j Z - i \omega T + b_j, \quad (\text{B11})$$

where p_j, s_j and b_j are complex constants satisfying the condition

$$\frac{1}{2} (p_j^2 + s_j^2) - \omega = 0. \quad (\text{B12})$$

To construct a one-dromion solution, we set $N = 1$. Substituting Eq. (B11) in Eq. (B8) we obtain

$$f^{(2)} = \exp(\chi_1 + \chi_1^* + 2\psi), \quad \exp(2\psi) = \frac{1}{4p_{1R}s_{1R}}, \quad (\text{B13})$$

with p_{1R}, s_{1R} being the real part of p_1, s_1 , respectively. Using $g^{(1)}, f^{(2)}$ from Eqs. (B7) and (B8) into Eqs. (B9) and (B10), one can show that $g^{(j)} = 0$ for $j \geq 3$ and $f^{(j)} = 0$ for $j \geq 4$, respectively. Now using Eqs. (11) and (13) in Eq. (B4), the physical quantities a and q will assume the following form

$$a = \sqrt{p_{1R}s_{1R}} \operatorname{sech}(\chi_{1R} + \psi) \exp(i\chi_{1I}), \quad (\text{B14})$$

$$q = -2s_{1R}^2 \operatorname{sech}^2(\chi_{1R} + \psi). \quad (\text{B15})$$

To obtain an analytic form for a dromion solution, we use the ansatz (Radha and Lakshmanan, 1997a)

$$f = 1 + \exp(\chi_1 + \chi_1^*) + \exp(\chi_2 + \chi_2^*) \\ + \exp(\chi_1 + \chi_1^* + \chi_2 + \chi_2^*), \quad (\text{B16})$$

where

$$\chi_1 = p_1 Y + i \frac{p_1^2}{2} T + b_1 \\ \chi_2 = s_1 Y + i \frac{s_1^2}{2} T + b_2. \quad (\text{B17})$$

Substituting Eq. (B16) with Eq. (B17) and putting it into Eq. (B8), we obtain

$$g = \gamma \exp(\chi_1 + \chi_2); \quad |\gamma|^2 = 4p_{1R}s_{1R} \quad (\text{B18})$$

For simplicity, we have assumed $p_1 = s_1 = p$. For $T=0$ the dromion solution for a is obtained as

$$a(Y, Z, T = 0) \\ = \frac{2p \exp[p(Y + Z)]}{1 + \exp(2pY) + \exp(2pZ) + \exp[2p(Y + Z)]}. \quad (\text{B19})$$

Numerically we have obtained $l_y \approx l_z \approx 0.825$ for $\alpha = 0.9$ and $k_y = 0.6234$ ($\omega = 1.098$). Calculating p from Eq. (B12) and putting it into Eq. (B19), we obtain the solution as presented in Fig. 2.

Acknowledgements. One of the authors, S. S. Ghosh, would like to thank the Council of Scientific and Industrial Research, India, for its financial support.

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