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Annual variation and global distribution of strato-mesospheric carbon monoxide measured by ground-based Fourier Transform Infrared spectrometry

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Abstract

We present long-term time-series of strato-mesospheric CO vertical columns measured from stations located in Antarctica, mid-latitudes and the Arctic, covering the period from 1997–2005. The instrument and the measurement technique allows the separation of tropospheric and strato-mesospheric contributions to the CO column, therefore providing information on the chemistry and dynamics both at low and high altitudes. Data from polar stations show a similar annual variability of strato-mesospheric CO with a strong maximum in late winter and spring and a small enhancement in late summer for some stations, which we call the “summer bulge”. Generally, the mid-latitude stations show no significant annual variability of strato-mesospheric CO columns. Measurements were compared with a two-dimensional chemistry-transport model of the middle atmosphere. The annual and latitudinal variations of CO are reproduced very well by a model run including thermospheric CO. Comparison with different model scenarios show that the polar winter maximum is due solely to downward transport of thermospheric CO, while the summer maximum is due to CHO_x chemistry in the stratosphere.

1 Introduction

The high latitude regions have a larger seasonal atmospheric variation than anywhere else on earth (Notholt et al., 1997). To study this variation, we use carbon monoxide (CO), which is an excellent tracer for global transport and air mass descent rates in the polar stratosphere and lower mesosphere. Its importance as an indicator of vertical transport in the mesosphere was highlighted in results from ground-based measurements e.g. Künzi and Carlson (1982). At altitudes of about 50 km, the photochemical lifetime of CO is about 7 days, which is comparable to the vertical and horizontal advection time scales at these altitudes (Dupuy et al., 2004, Solomon et al., 1985).

The primary source of carbon monoxide in the upper mesosphere and lower ther-

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mosphere is the photolysis of carbon dioxide. In the stratosphere, CO is produced through the oxidation of atmospheric methane, but OH rapidly destroys it through oxidation. This reaction acts as the main sink of CO. This process does not take place during the polar night, since OH is produced by reactions involving photolysis, e.g. of H₂O, and since the concentration of OH diminishes rapidly during polar night. Carbon monoxide follows the meridional circulation towards the winter hemisphere polar night region. The consecutive downward motion induces a sharp gradient in the CO concentrations down to the stratosphere (Solomon, et al., 1985). In summer, uplifting of air masses with low CO content takes place. Solomon et al. (1985) predicted that very large abundances should accumulate in the polar night mesosphere because of the absence of photochemical destruction processes.

Previous ground-based measurements of mesospheric CO were shown in pioneering works such as that of Künzi and Carlson (1982); Clancy, et al. (1982); Farmer, et al. (1980); Zander et al. (1981); Goldsmith et al. (1979); etc. However, there is a lack of long term ground-based observations of strato-mesospheric CO in the literature. Recently, satellites have been able to measure high-altitude CO (e.g. Dupuy et al., 2004; Clerbaux et al., 2005). But until now, the longest reported time series of CO in the upper atmosphere was done by Forkmann et al. (2003) over the Onsala Space laboratory, Sweden (57.4° N, 12° E). The time series spanned from September 2000 to September 2002.

Kasai et al. (2005) reported the first ground based measurements of strato-mesospheric CO by Fourier Transform Infrared (FTIR) spectrometry over Poker Flat, Alaska (65° N, 147° W). They have established the capability of the FTIR technique to detect the seasonal variability of strato-mesospheric CO.

In this paper we report more than seven years of time series of strato-mesospheric CO measured from stations located in the high latitude regions of the Arctic (79° N) and the Antarctic (78° S). We augment this data set from two additional high-latitude stations and two mid-latitude stations. We also show comparisons with a global two-dimensional photolysis, chemistry and transport model in the stratosphere and meso-

sphere.

2 Instruments

Measurements were taken from three Arctic stations (Ny Alesund 79° N, Kiruna 68° N and Poker Flat 65° N), one Antarctic station (Arrival Heights 78° S) and two mid-latitude stations (Bremen, Germany 53° N and Lauder, New Zealand 45° S). All stations are equipped with Bruker 120HR spectrometers except for Arrival Heights (120 M) and Bremen (125 HR). For the polar stations, measurements are limited by the polar night. Solar spectra in Ny-Alesund can be recorded between March and early October. In Kiruna, the polar night is between early December and mid January. At the Arrival Heights station in Antarctica, the polar night is between late April and mid August. Measurements in Poker Flat are possible between early February and mid October.

2.1 Retrieval of strato-mesospheric CO

Partial columns of CO were retrieved from the absorption lines in two steps. First, volume mixing ratio (VMR) profiles from the surface up to about 100 km were retrieved. The profiles were then converted into number densities using height-temperature-pressure profiles and integrated to yield partial columns. The program used for the retrieval of CO profiles for Ny Alesund, Poker Flat, Bremen, Lauder and Arrival Heights is SFIT-2 version 3.8 (see also Rinsland et al. (1998)). The program used to retrieve the Kiruna data is PROFFIT 9 developed by Hase (2000). A detailed description and comparison of both retrievals are shown in Hase et al. (2004). The pressure and temperature profiles necessary for the forward model were obtained from balloon sondes that were launched daily from the stations. Above the altitude limits of the sondes (approximately 30 km), standard pressure and temperature profiles were taken. In Kiruna, pressure and temperature profiles were taken from NCEP (Kanamitsu, 1989).

CO spectra were analyzed in the CO micro-windows based on the previous work

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of Rinsland et al. (1998). For all stations, the HITRAN2k line list plus updates to 2001 (Rothman et al., 2003) was used. Typical averaging kernels for the retrievals in Ny Alesund are shown in Fig. 1. The figure shows that the 2 partial columns from 0.2–18 km and from 18 km to 85 km can be separated. For some cases, the kernel for the 18 km to 85 km is not perfect, i.e. it does not have a maximum of 1. Furthermore, the stratospheric columns and the mesospheric columns are not separable, thus we provide the measurements as strato-mesospheric columns.

2.2 The Chemical Transport Model

The model used is a global two-dimensional photolysis, chemistry and transport model of the stratosphere and mesosphere. It is a coupled chemistry-dynamics model which combines the THIN AIR meteorological code (Kinnery, 1996) and the SLIMCAT chemistry code (Chipperfield, 1999). Temperature, pressure and wind fields are calculated by the THIN AIR code on isentropic surfaces from the bottom up to ~100 km with a vertical spacing of about 3 km.

The model has a horizontal resolution of about 9.5° extending from -85.3° S to 85.3° N in 19 evenly spaced latitude bins. The chemistry module uses JPL-2003 photochemistry data (Sander et al., 2003). Though the SLIMCAT model is not appropriate for the troposphere, it is applied to the entire vertical range of the model. Heating rates are calculated in the THIN AIR module, using O_3 , NO_2 and CH_4 values provided by the chemistry module. CO_2 is also used to calculate heating rates, however, as this is very long-lived in the stratosphere and mesosphere, it is not accounted for in the chemistry code. In the past, the model has been used for a number of studies concerning the composition of the middle atmosphere (Sinnhuber et al, 2003; Chipperfield and Feng, 2003). Tropospheric sources of chlorine, bromine and fluorine compounds as well as greenhouse gases like CH_4 , N_2O and CO_2 are based on the WMO A1 scenario (WMO 2002, 2003).

Model runs started from 1988, and run to 2005. For this study, two model runs were carried out. The “base” scenario uses the original SLIMCAT chemistry which does

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not contain CO₂. This means that CO is produced solely from CH₄ oxidation. In a second model run, called “thermospheric”, CO in the uppermost model box is fixed to the CO₂ value. This means that CO₂ is transported unchanged into the thermosphere, where it is transferred into CO immediately. CO will then be transported down into the mesosphere and stratosphere during polar winter; after polar sunrise, CO will react with OH, re-forming CO₂ in a couple of days.

3 Results

A comparison of the partial column densities measured by the FTIRs in molecules/cm² from 18 km to the top of the atmosphere and the model results are shown in Fig. 2 for the period of 2003–2004. The partial column densities measured by the FTIRs are shown by the blue dots. The cyan curve represents the base run from the model, where thermospheric CO is neglected. Clearly this run alone cannot reproduce the measurements. The green curve represents the complete run with the thermospheric CO. The model run smoothed by the typical averaging kernels of the FTIR is represented by magenta curve. The smoothing was done according to the formalism described in Rodgers and Connor (2003) and shown in Velazco et al. (2005) for FTIR, model and satellite data comparison of CO profiles. The smoothed curves represent what the FTIR should “see” if the model were to represent the true CO. Measurements and smoothed model results compare very well in the northern hemisphere polar stations. For the Arrival heights station, the annual variation of CO is captured by the model qualitatively, however, actual values of the winter maxima are significantly higher in the model results. This might be a result of a wrong estimation of the downward transport in the model. In the mid-latitudes, the model predicts a slight enhancement of the strato-mesospheric CO columns during winter times over Bremen and lesser values over Lauder. In both places, the quantitative agreement between model values and measurements are very good.

Long term time series of CO partial column densities from 18 km to the top of the

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atmosphere measured from FTIR (blue dots) and compared with the model (magenta dashes) are shown in Fig. 3. The measurements were taken from the three polar stations (Ny Alesund and Kiruna from the Arctic and Arrival Heights in Antarctica). The gray shaded areas represent the polar night where solar absorption measurements are not possible. Unlike the seasonal cycles of CO in the troposphere, the seasonal cycles in the strato-mesosphere show very steep gradients, with maximum values occurring in January in the Arctic and in June–July in the Antarctic. As shown by the model (and partially by the measurements), the CO column above 18 km increases from about 4.0×10^{16} molecules/cm² in summer to about 14×10^{16} molecules/cm² in winter (an increase of about 3.5 times). This rapid increase is followed by a rapid decrease as soon as the sun re-appears in spring.

Although the annual variability in both Arctic and Antarctic are similar, there is a slight difference in amplitude. The measurements show that the values of the CO column above 18 km in the Arctic is slightly higher than in the Antarctic. This can be seen better in the average curves on Fig. 4. The average curves were calculated from the whole time series of each station. Values over 15 days were averaged to yield one data point in the curve.

A small enhancement of CO columns can be observed in late summer from the measurements at the high-latitude stations. This could be seen as a small “bulge” from the average curves from Ny Alesund and Poker Flat in Fig. 4. The Kiruna and Arrival Heights data exhibit this “summer bulge” for some years (e.g. 1998, 2002 and 2004 in Arrival Heights) but it tends to be averaged out due to the changing patterns in the data. Note also that Kiruna is often at the edge of the polar vortex. This summer maximum is also seen in the model, it is much clearer in the “base” model run. In the thermospheric model run, it is superimposed by the much sharper signal of mesospheric CO. This summer maximum is produced by methane oxidation in the stratosphere, which occurs faster in summer. CO is an intermediate product of methane, it is then slowly transformed into CO₂.

The agreement between model and Ny Alesund data is very good and estimates

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from the model for this site indicates that, generally about 20%–80% of the column above 18 km comes from the stratosphere (18–26 km), shown on Fig. 5. This estimate is based on the model data for Ny Alesund for one year. In winter, a significant portion of the column comes from above 56 km, i.e., the mesosphere and lower thermosphere.

5 The summer maximum originates in the layer between 18–26 km. It cannot be seen above 56 km. This means that the “summer bulge” detected by the FTIR really originates from the stratosphere

The Lauder data (45° S) do not show the very high values of strato-mesospheric CO (Fig. 6). Although there is a strong variability in the columns below 18 km mostly due to biomass burning (Jones et al., 2001), the strato-mesospheric columns show almost no variability. This shows that the columns below 18 km do not influence the columns above 18 km and that the retrieval could clearly separate both columns. The only exceptions to the monotonous strato-mesospheric CO trends in the mid-latitudes are the values measured in the winters of 2002–2003 and 2004–2005 over Bremen (53° N).

4 Conclusions

The downward transport of strato-mesospheric CO above 18 km in the winter polar regions, which is strongly influenced by the meridional circulation can be seen in the data. The strong gradient showing a maximum in winter and minimum in summer are well captured by the measurements and verified by the model. It was shown that this feature is generally not observed in mid-latitude stations.

CO in the mesosphere is influenced by the competition between downward transport from the thermosphere and OH oxidation. The measurements show that the pattern of the strato-mesospheric CO columns for all years in one station are almost similar. We assign this to CO having a shorter lifetime compared to the downward vertical transport in the mesosphere-stratosphere during the measurement periods. On the other hand, there is a slight difference in column amounts of the strato-mesospheric CO between

the two poles, i.e. in spring, there is relatively more CO above 18 km in Ny Alesund than in Arrival Heights based on measurements and on the average curves. We speculate that this is due to the presence of more subsidence in the Arctic compared to the Antarctic.

5 Comparisons with a global two-dimensional photolysis, chemistry and transport model in the stratosphere and mesosphere were shown. The assumption in the model that all the CO₂ in the thermosphere gets converted into CO via photolysis has been shown to be very reasonable. There is a good agreement between the model results and the measurements. According to the model, the production of CO from CH₄ oxidation in the mid to upper stratosphere results to a signal indicated by the “summer bulge” in the FTIR data seen in Ny Alesund, Poker Flat and occasionally in Kiruna and Arrival Heights.

15 *Acknowledgements.* V. Velazco wishes to thank M. Palm and B. Martin Sinnhuber for helpful discussions. S. W. Wood acknowledges the New Zealand Foundation for Research Science and Technology and Antarctica New Zealand. I. Kramer, T. Blumenstock and F. Hase gratefully acknowledge the Goddard Space Flight Center for providing temperature and pressure profiles via the auto mailer system and U. Raffalski from IRF Kiruna for support of FTIR measurements. Y. Kasai acknowledges Y. Muaraya (NICT), A. Kagawa (Fujitsu FIP), and T. Koshiro (Kyoto University) for the analysis of the FTIR data and useful discussions. Financial supports from the European Commission via the project UFTIR (contract EVK2-2002-000159) and from the Helmholtz association within the virtual institute PEP are gratefully acknowledged.

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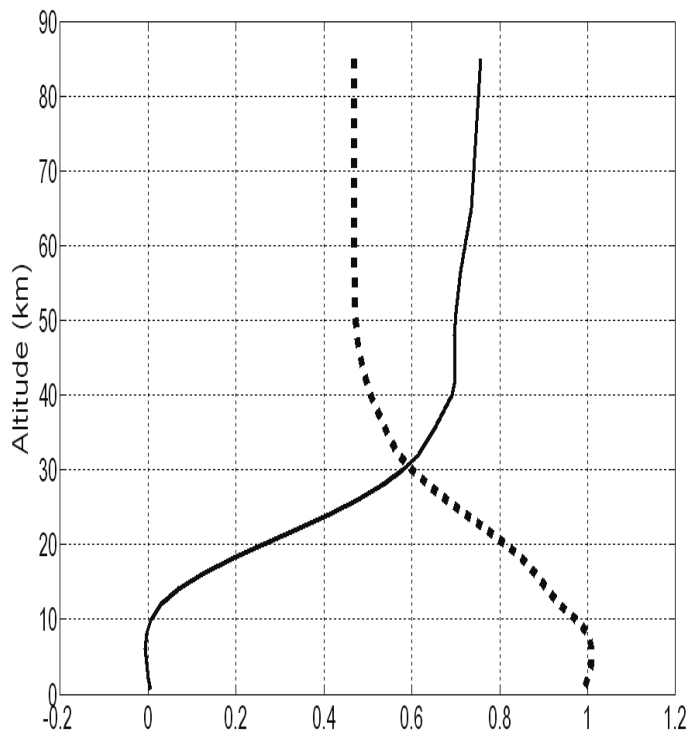


Fig. 1. Typical averaging kernels for the retrieval of CO from 18 km to the top most layer of the retrieval are shown by the solid curve. The dotted curve is the averaging kernel for the retrieval of the CO column from the ground to 18 km. (Calculated for Ny Alesund for solar zenith angle of 58° , optical path difference of 180 cm and a signal to noise ratio of 200).

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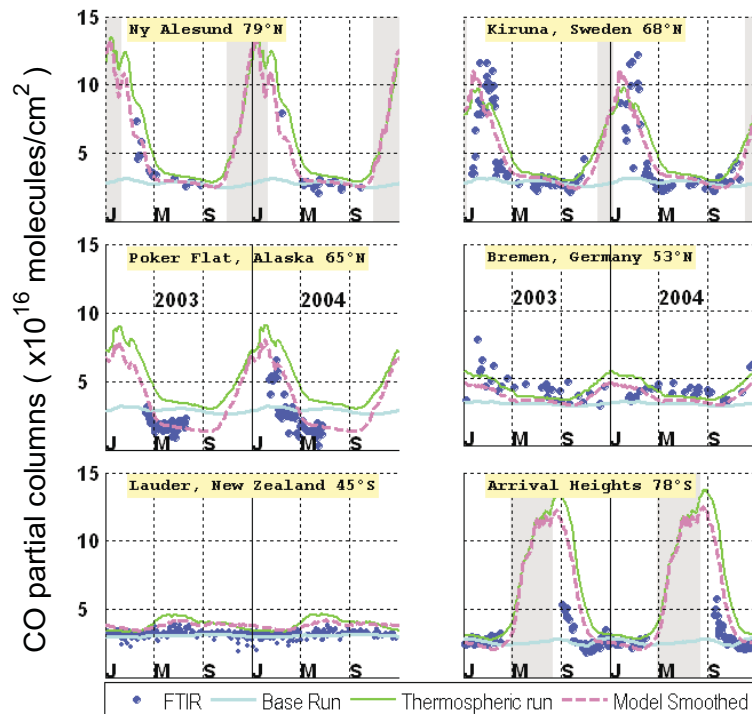


Fig. 2. A comparison of the partial column densities measured by the FTIRs in molecules/cm² from 18 km to the top of the atmosphere and the model results for the period of 2003–2004. Cyan line: base model run without thermospheric CO. Green line: Model run with thermospheric CO. Blue dots: Retrieved columns from the FTIR. Magenta line (dashed): model run smoothed with the averaging kernels of the FTIR according to Rogers and Connor (2003). The shaded area approximately represents polar night at the sites.

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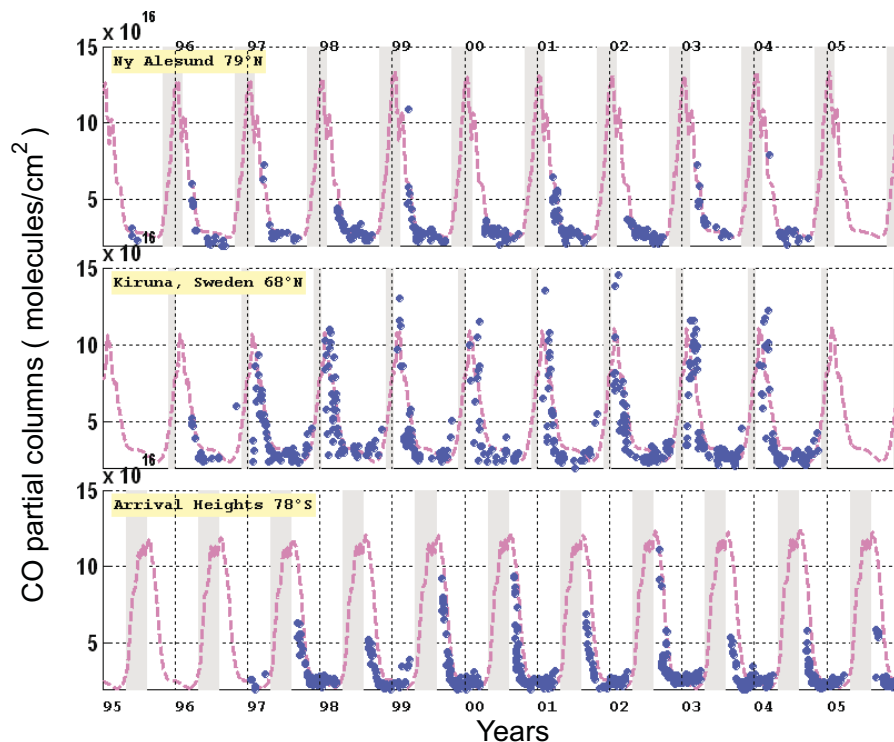


Fig. 3. Long-term FTIR measurements from the three polar stations; Ny Alesund and Kiruna from the Arctic and Arrival Heights in Antarctica (blue dots) compared to the smoothed model data (dashed magenta curves). The measurements and model data are in partial column densities from 18 km to the top of the atmosphere.

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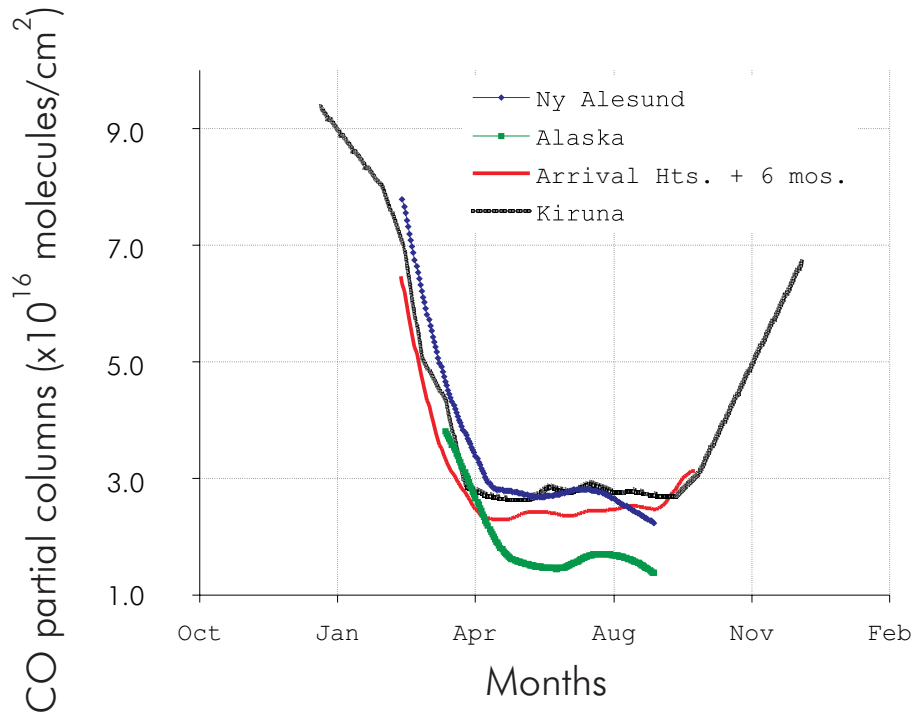


Fig. 4. Average curves calculated from the time series of the four Polar stations: Ny-Alesund, Poker Flat, Kiruna and Arrival Heights. The summer bulge in July–August can be clearly distinguished from the Poker Flat and Ny Alesund data. The Arrival heights data (shifted by six months) only shows small “lumps” in June and August. The variability in the data in Arrival heights averages out the summer bulge. The average curves were calculated from the whole time series of each station. Values over 15 days were averaged to yield one bin in the curve.

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Interactive Discussion

**Strato-mesosphere
CO measurements by
ground-based FTIR**

V. Velazco et al.

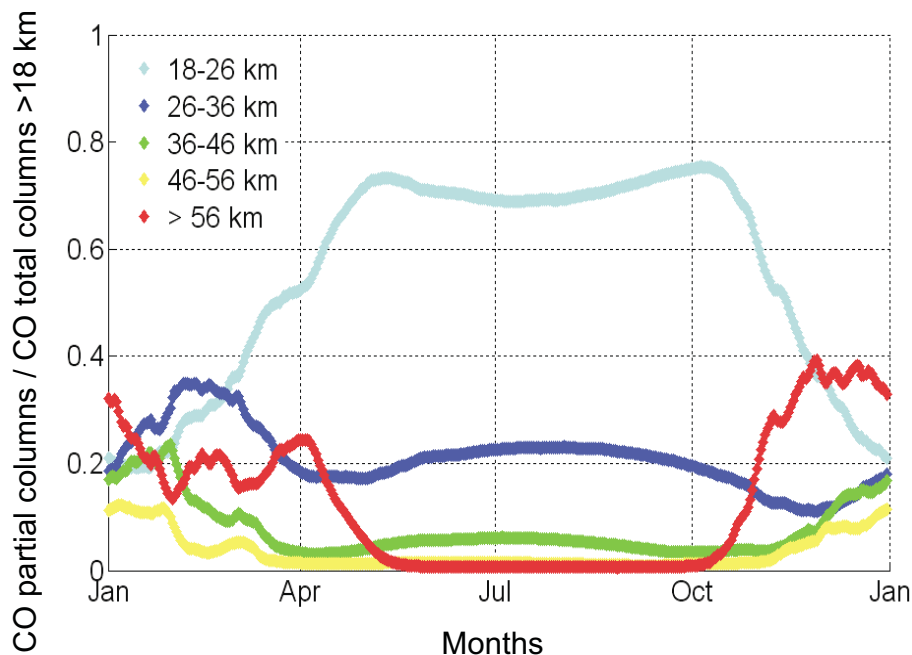


Fig. 5. Relative contributions of each layer to the column above 18 km calculated from the model. The contributions are calculated from the CO partial columns coming from a certain layer divided by the total columns above 18 km. These curves indicate where the signal should come from. In winter, most of the CO partial column above 18 km is dominated by the CO coming from above 56 km. And from the 26–36 km layer. This is also the time when the downward transport from the mesosphere is strongest. In summer, the column above 18 km is dominated by CO at 18–26 km where CH₄ oxidation plays a significant role in the production of CO.

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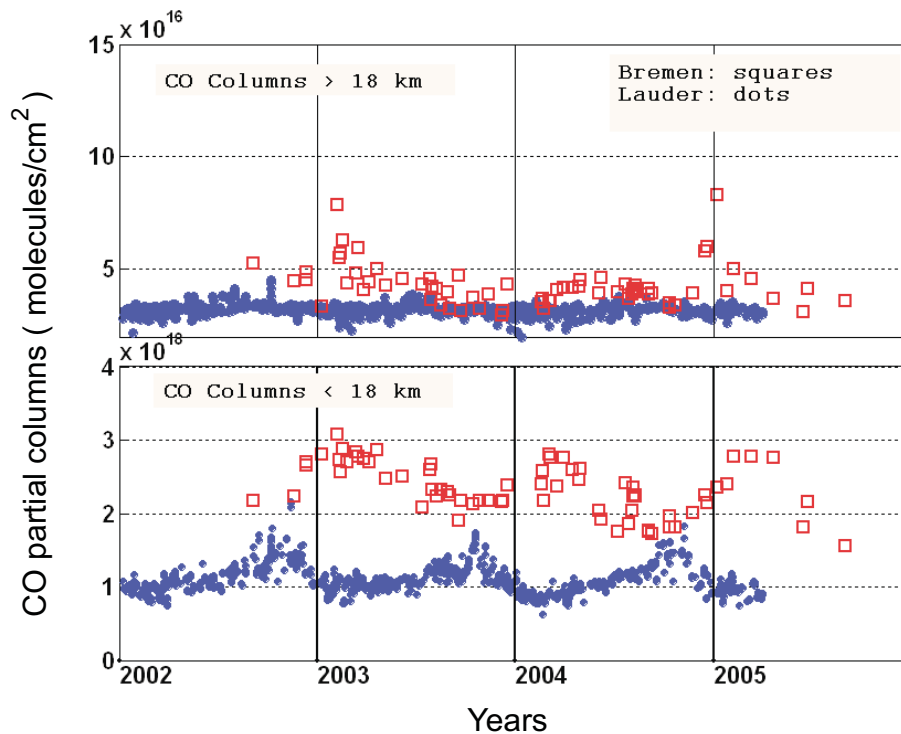


Fig. 6. CO columns above and below 18 km from the two mid-latitude stations: Bremen 53° N (red squares) and Lauder 45° S (blue dots).

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