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Retrieval of water  
vapor columns

R. Lang et al.

# Application of the Spectral Structure Parameterization technique: retrieval of total water vapor columns from GOME

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Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

## Abstract

We use a recently proposed spectral sampling technique for measurements of atmospheric transmissions called the Spectral Structure Parameterization (SSP) in order to retrieve total water vapor columns (WVC) from reflectivity spectra measured by the Global Ozone Monitoring Experiment (GOME). SSP provides a good compromise between efficiency and speed when performing retrievals on highly structured spectra of narrow-band absorbers like water vapor. We show that SSP can be implemented in a radiative transfer scheme which treats both direct-path absorption and absorption by singly scattered light directly. For the retrieval we exploit a ro-vibrational overtone band of water vapor located in the visible around 590 nm. We compare our results to independent values given by the data assimilation model of ECMWF. In addition, results are compared to those obtained from the more accurate, but slower, Optical Absorption Coefficient Spectroscopy (OACS).

## 1. Introduction

The retrieval of concentrations of narrow-band line absorbers, such as water vapor (WV), from space-borne spectral instruments is often complicated by the relatively low resolution of the detector with respect to the width of an individual absorption line. For example, GOME spectral sampling in the visible region (channel 3) is 0.2 nm which covers up to 12 individual absorption lines of WV in the spectral region around 580 nm (ESA, 1995). To allow accurate modeling of such absorption spectra requires a large number of spectral realization points in order to resolve the narrowest lines, especially at high altitudes where pressure broadening is absent. In principle, band models, exponential sum fitting methods or other opacity sampling techniques, like the well known k-distribution method (Lacis and Oinas, 1991; Kato et al., 1999), and the more recently proposed Optical Absorption Coefficient Spectroscopy technique (OACS) (Maurellis et al., 2000a; Lang et al., 2002) can be used to solve this problem. In practice, both of the

## Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

**Retrieval of water vapor columns**

R. Lang et al.

latter two methods use opacity sampling probability functions to represent the effective absorption within a specific spectral sampling interval. These probabilities are then used either directly (OACS), or indirectly, via a transformation from wavelength into cross-section space (k-distribution), as weighting functions of a set of opacity basis functions, whose summation replaces a computationally expensive spectral sampling integral. Both of these methods have demonstrated their ability to solve the problem of spectral sampling in an accurate and efficient way for atmospheric water vapor absorption, in case of OACS utilizing GOME measurements in the absorption region around 590 nm (Lang et al., 2002) and, in case of the k-distribution method, on synthetic spectra in the IR regions of the SCanning Imaging Absorption CartographY (SCIAMACHY) instrument on ESA's Environmental satellite ENVISAT (Buchwitz et al., 2000).

In contrast, in the Spectral Structure Parameterization (SSP) recently proposed by Maurellis et al. (2000b) and applied here for the first time to retrievals from reflectivity measurements, the number of opacity functions and their weights are reduced to only one basis function with one weighting parameter, called the spectral structure parameter  $w$ , which characterizes the spectral structure of the absorber within a specific wavelength range and at a specific altitude. SSP lends itself therefore more to the basic concept of band models, where the average absorption over a specific wavelength interval is represented by averaged line parameter values. Different to band models, SSP is well suited for relatively small sampling regions containing only a small number of distinct absorption lines, such as those found in the regions covered by the detector pixels of both GOME and SCIAMACHY (both instruments work at relatively high spectral resolution across a wide wavelength region). In addition, the implementation of SSP in a radiative transfer scheme including direct path absorption, as well as the contribution of single-scattering to the measured reflectivities, is relatively simple and requires only a small amount of computational effort. Consequently, SSP retrieval of WVC is fast and may be performed on a global scale in a reasonable time, as demonstrated in this paper .

Retrieval of WVC from GOME has also been demonstrated by Noël et al. (1999)

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

**Retrieval of water  
vapor columns**

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

and Casadio et al. (2000) utilizing absorption bands in the region around 700 and 740 nm, respectively. We are focussing on a ro-vibrational overtone band of WV in the visible between 580 and 605 nm covered by the GOME instrument. This band was first exploited for the retrieval of WVC by Maurellis et al. (2000a), using the OACS sampling technique. SSP is a retrieval method, which is easy to implement, fast, self-contained and based on first principles. The only inputs needed for the retrieval of WVC by SSP are pressure and temperature profiles together with some information on background absorbers, which are, for global retrievals, taken from climatology models or GOME level 2 data. SSP retrieves subcolumn profile over 18 atmospheric levels from which a total column is calculated. The only constraints applied to the fits are three standard upper profile constraints for tropical, mid-latitude and high-latitude cases.

In the following section we briefly summarize the basic concept of SSP for homogeneous atmospheres. Then we show how SSP can be applied to nadir satellite reflectivity measurements probing nonhomogeneous atmospheres, including light paths where photons undergo a single scattering event. Thereafter, we present results of SSP forward-modeled reflectivity spectra, which we compare to both line-by-line (lbl) forward-modeled spectra, as well as GOME measurements. SSP retrieval results from lbl forward-modeled spectra are used to assess the method related bias and to introduce a retrieval correction term for high WVC values. The impact of multiple scattering and aerosol loading on the retrieval values will also be discussed in some detail. We then expand our retrievals to both a single GOME track and a global coverage (i.e. three days of continuous GOME measurements). The results are compared and discussed with respect to co-located WVC given by the European Center of Medium Range Weather Forecast (ECMWF) data assimilation model, after which we present our conclusions.

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**Retrieval of water  
vapor columns**

 R. Lang et al.
 

---

[Title Page](#)
[Abstract](#)
[Introduction](#)
[Conclusions](#)
[References](#)
[Tables](#)
[Figures](#)
[◀](#)
[▶](#)
[◀](#)
[▶](#)
[Back](#)
[Close](#)
[Full Screen / Esc](#)
[Print Version](#)
[Interactive Discussion](#)

© EGS 2002

## 2. Sampling of homogeneous absorption

For homogeneous absorption over a direct path in the absence of any scattering events, the spectrally sampled transmittance may be defined as

$$\langle T \rangle_{\delta\lambda} = \int_{\delta\lambda} \exp(-\sigma(\lambda)N) \frac{d\lambda}{\delta\lambda}, \quad (1)$$

5 where  $\sigma(\lambda)$  is the absorption cross-section,  $N$  is the total column density and  $\delta\lambda$  is the sampling width. The transmittance  $\langle T \rangle_{\delta\lambda}$  may also be expressed as:

$$\langle T \rangle_{\delta\lambda} = 1 + w \exp(-SN) - w. \quad (2)$$

Here,  $S$  has units of cross-section, therefore, determines the height of a box with width  $w$  ( $0 < w \leq 1$ ) in units of  $\delta\lambda^{-1}$ , which covers the area of absorption equivalent to  
 10 the area covered by the individual lines of the absorber within the spectral region  $\delta\lambda$  (see Fig. 1) with

$$\langle \sigma(\lambda) \rangle_{\delta\lambda} = wS, \quad (3)$$

and the pixel-averaged optical depth  $\langle \tau \rangle_{\delta\lambda} = wSN$ .  $w$  and  $S$  are now defined by the simultaneous solution of Eqs. (2) and (3) with the result that they are implicitly dependent on  $N$ . A Taylor expansion to second order of the exponents in Eqs. (1) and (2)  
 15 reveals that

$$w \approx \frac{\langle \sigma \rangle_{\delta\lambda}^2}{\langle \sigma^2 \rangle_{\delta\lambda}} \text{ and } S \approx \frac{\langle \sigma^2 \rangle_{\delta\lambda}}{\langle \sigma \rangle_{\delta\lambda}}. \quad (4)$$

The latter is a good approximation of  $w$  and  $S$  in cases of a wavelength-averaged optical depth of lower than 0.1 (Maurellis et al., 2000b).

### 3. Application to nadir measurements in nonhomogeneous atmospheres

For the specific scenario of modeling the WV absorption from a nadir viewing geometry, we subdivide the atmosphere into 18 levels  $\ell$ , containing WV subcolumn densities  $N_{\ell}$ . For real measurements, pressure and temperature profiles, as well as oxygen and nitrogen profiles for the calculation of Rayleigh scattering albedo, are taken from the Neutral Atmosphere Empirical Model MSISE90 (Hedin et al., 1991) for a given date, time, and geolocation at fixed altitude levels. We use 18 atmospheric layers  $\ell$  to cover altitudes from the ground up to 9 km and, therefore, more than 99% of the atmospheric water vapor. The WV density drops generally by more than 2 orders of magnitude over the first 10 km (see, also Lang et al., 2002). In general, the  $w$  width parameter or, respectively, the structure of the spectrum, changes with pressure and temperature over altitude due to the pressure and temperature dependent width of the lines. The same holds for the pixel-averaged line strength  $S$ . For the treatment of nonhomogeneous paths traversed by photons undergoing higher orders of scattering we now assume that the absorption over the different light paths throughout the atmosphere is dominated by the level with the maximum absorption, as is the shape of the measured absorption line. Consequently, the choice of a dominant absorbing layer is dependent on the shape of the profile of the absorber of interest. Note, that due to the strong exponential decrease of the WV profile with respect to altitude, we may assume that the maximum impact on the absorption over a specific light path will occur at the lowest point of the path. A solution of the transport equation of scalar radiative transfer in its plane parallel approximation (Lang et al., 2002) reveals that the measured reflectivity  $R$  of a nadir viewing instrument may be separated into three distinctive parts, viz.,

$$R_j = \Lambda \langle R_{\text{surf}} \rangle_j + \langle R_{\text{ss}} \rangle_j + \langle R_{\text{ms}} \rangle_j, \quad (5)$$

where  $j$  is a specific spectral interval with width  $\delta\lambda$ ,  $\Lambda$  is the surface-albedo of a Lambertian surface,  $R_{\text{surf},j}$  is the reflectivity of the direct light path reflected at the earth surface,  $R_{\text{ss},j}$  is the reflectivity of singly scattered and  $R_{\text{ms},j}$  of multiply scattered light. Using the assumption above on the dominant  $w$ -parameter, we define  $w_{\text{max}} = w(N(\bar{\ell}))$ ,

## Retrieval of water vapor columns

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

where  $N(\bar{\ell})$  is the maximum subcolumn density for a specific path at level  $\bar{\ell}$ . Now, we may write the reflected light coming from the surface, utilizing Eq. (2), as

$$\langle R_{\text{surf}} \rangle_j = 1 + w(\bar{\ell}_{\text{surf}})_j \exp \left[ - \sum_{\ell} \tilde{\mu} S(\ell)_j N(\ell) \right] - w(\bar{\ell}_{\text{surf}})_j, \quad (6)$$

where  $\tilde{\mu}$  is the geometric path-length factor for a nadir viewing instrument, with  $\tilde{\mu} \equiv (\frac{1}{\mu_o} + 1)$ ,  $\mu_o$  is the cosine of the solar zenith angle (SZA) and  $w(\bar{\ell}_{\text{surf}})_j$  is the dominant  $w$ -parameter  $w_{\text{max}}$  for the specific case of the direct, surface reflected path. Because, in the case of WV absorption, we expect the maximum optical density at the surface layer we may set  $\bar{\ell}_{\text{surf}} = 1$ . Using the same reasoning for the single-scattering contribution we can then write, by again utilizing Eq. (2),

$$\langle R_{\text{ss}} \rangle_j = \frac{\rho(\Theta)}{4\mu_o} \int_0^{\infty} \langle \beta(\ell(z))_{\text{sca}} \rangle_j \times \left\{ 1 + w(\bar{\ell}_{\text{ss}})_j \exp \left[ - \sum_{\ell=\ell(z)}^{\ell_{\text{top}}} \tilde{\mu} S(\ell)_j N(\ell) \right] - w(\bar{\ell}_{\text{ss}})_j \right\} dz, \quad (7)$$

where  $\bar{\ell}_{\text{ss}} = \ell(z)$ , because  $\ell(z)$  is the lowest altitude point for each single-scattering path and, therefore, in the case of WV, the point with most impact on the absorption. In addition, we assume that Rayleigh scattering is the dominant form of single-scattering and  $\langle \beta(\ell(z))_{\text{sca}} \rangle_j$  is the mean Rayleigh scattering coefficient within  $\delta\lambda_j$  at altitude  $z$ , with a corresponding phase function of  $\rho(\Theta) = \frac{3}{4}(1 + \cos^2(\Theta))$ , where  $\Theta$  denotes the SZA.

The multiple-scattering contribution  $\langle R_{\text{ms}} \rangle_j$  to the reflected light is represented by a first order polynomial with free parameters  $C$  and  $D$ , which accounts for the broadband effect of multiple scattering. In the previous study by Lang et al. (2002) it was shown that the error on the retrieved WVC using this kind of treatment for the multiple-scattering contribution is lower than 20% in cloud-free situations for a worst-case scenario of high WVC of  $1.4 \times 10^{23}$  molec/cm<sup>2</sup>, low surface albedo of 0.03 and 0.1 and

## Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002



**Retrieval of water  
vapor columns**

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

high aerosol loading for maritime and rural scenarios, respectively, utilizing the OACS retrieval technique (see also Sect. 7 for the impact of aerosol loading and multiple-scattering on the SSP retrieval). Similarly, the unknown surface albedo is represented by a first order polynomial with free parameters  $A$  and  $B$ . The parameters  $A$  to  $D$  have to be adjusted to the real measurement when comparing SSP forward model results to the measurements, as well as during the retrieval and fitting of WVC values.

For details concerning the different contributions to the total reflectivity, as well as a detailed discussion about the impact of multiple scattering within this wavelength region, the reader is referred to the aforementioned study by Lang et al. (2002) and, for the specific case of SSP, to section 7 of this work.

#### 4. Construction of look-up-tables for $w$ and $S$

In this study we utilize Eq. (4) in order to construct a look-up-table for each of the parameters  $w$  and  $S$ . For a range of 22 temperatures  $T$  and 31 pressure values  $p$ , covering all  $p$  and  $T$  values thought relevant for our altitude region of interest for all possible geolocations, we then calculate realizations of the cross-section  $\sigma$  utilizing line-parameters from the HITRAN '96 database and Voigt absorption line shapes (Armstrong, 1967). For those realizations we use 50 000 numerical grid points over the whole spectral region between 585 and 600 nm, in order to resolve the narrowest lines with at least 5 points. We use a sampling width  $\delta\lambda$ , which is three times smaller than the sampling width of a GOME detector pixel and more than 4 times smaller than the full-width-half-maximum (FWHM) of the instrumental response function, which satisfies the Nyquist criterion (see also following section) and results in a total of 207 spectral bins. By doing so we create for each parameter  $w$  and  $S$ , a 22 by 31 by 207 look-up-table matrix which we can interpolate over  $p$  and  $T$  for a specific measurement scenario at each altitude level  $\ell$ . Figure 2 shows  $w$  and  $S$  for one wavelength bin  $j$  as a function of the  $p$  and  $T$  reference profiles. From this we see that  $w$  and  $S$  vary smoothly with respect to  $p$  and  $T$ , meaning that interpolation on intermediate  $p$ ,  $T$ -tuples leads to reasonable

**Retrieval of water  
vapor columns**

R. Lang et al.

results (Maurellis et al., 2000b). As the pressure decreases the  $w$  parameter also decreases (upper panel Fig. 2), i.e. the spectrum becomes more structured due to the pressure-related narrowing of the lines. The line width is only weakly dependent on the temperature (Rothman et al., 1998) which can also be seen from the  $w$  contour plot. The  $S$  parameter (lower panel Fig. 2) changes significantly only for the lowest pressures. For low pressures, when the  $w$  parameter decreases, the  $S$  parameter increases, because the area of absorption or the averaged line intensity within a certain wavelength bin has to be conserved. Due to the temperature dependence of the line-intensity value of individual absorption lines (Rothman et al., 1998), the variation in temperature for the  $S$  parameter is stronger than that for the  $w$  parameter.

In general, the  $w, S$ -tables are smaller than the usual tables used for the k-distribution method, and significantly smaller than the look-up-tables used for the OACS method, because the extra dimension for the summation over cross-section probability, or cross-section bin space, necessary in the latter two methods, is not required for SSP. Therefore, the interpolation on the look-up-tables can be performed very rapidly even for small  $\Delta\lambda$ , and thus, for high wavelength resolution.

## 5. Forward modeling and retrieval

The spectral region between 585 and 600 nm is covered by 69 detector pixels of the GOME instrument. Within this region absorptions of the additional absorbers  $(\text{O}_2)_2$ ,  $\text{O}_3$ ,  $\text{NO}_2$  and Sodium have to be taken into account, together with the loss of light scattered out of the light path by means of Rayleigh scattering. Apart from Sodium and  $\text{NO}_2$ , which contributes very little to the total absorption in this spectral region, all of these remaining absorbers contribute to a smooth background absorption. The contribution of the Fraunhofer line absorption to the so called  $F_0$  effect and the contribution of the Ring effect, from rotational Raman-scattering, to the total reflectivity is also assumed to be rather small within this spectral region (for a detailed discussion the reader is referred to Lang et al., 2002).

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[|◀](#)[▶|](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

## Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

GOME measures the earth radiance  $I$  in nadir viewing geometry for a footprint of 40 by 320 km. The solar irradiance  $F_0$  is measured once for every orbit (for a detailed description of the instrument see [ESA, 1995](#); [Burrows et al., 1999](#)). The instrumental function  $\mathcal{H}(\lambda, \lambda'; \Sigma)$  of channel 3 is represented here by a Gaussian function with a FWHM  $\Sigma$ . The reflectivity measured by the  $k$ th detector pixel is then defined as

$$R_k^{\text{GOME}} = \int_{\Delta\lambda_k} \int_{-\infty}^{+\infty} \frac{\pi I(\lambda)}{\mu_o F_0} \mathcal{H}(\lambda, \lambda'; \Sigma) d\lambda' \frac{d\lambda}{\Delta\lambda_k}, \quad (8)$$

where  $\Delta\lambda_k$  is the wavelength coverage of a GOME detector pixel of about 0.21 nm in case of channel 3. Here, it is assumed that the solar irradiance  $F_0$  is constant within  $\Delta\lambda_k$ . The modeled reflectivity  $R_k$  can be approximated by

$$R_k = \sum_j^K \sum_i^M R_j \mathcal{H}(\lambda_j, \lambda_i; \Sigma) \delta\lambda_i \frac{\delta\lambda_j}{\Delta\lambda_k}, \quad (9)$$

where  $M$  is the number of sampling bins  $j$  over the total spectral region of interest,  $K$  is the number of sampling bins  $j$  within the spectral region covered by one GOME detector pixel and  $|\delta\lambda_{i,j}| = K^{-1} \Delta\lambda_k$ .  $R_j$  is the sampled reflectivity from Eq. (5) with

$$R_j = (A\lambda_j + B)\langle R_{\text{surt}} \rangle_j + \langle R_{\text{ss}} \rangle_j + (C\lambda_j + D). \quad (10)$$

Here, it must be noted that Eq. (9) is only a good representation of Eq. (8) when  $\delta\lambda \ll \Sigma$ , satisfying the Nyquist criterion. In our case  $\delta\lambda$  (around 0.07 nm) is a third of the spectral width covered by a single GOME detector pixel ( $K = 3$ ), whilst  $\Sigma$  is about 0.29 nm.

In Fig. 3 we present a direct comparison between sample GOME measurements and forward calculations using Eq. (9) for two representative measurement geometries and geolocations with both high and low WVC. Realistic WV profiles used for the forward calculations are taken from the ECMWF database, with the resulting fit residuals being smaller than 1%, except for the region near the solar sodium Fraunhofer lines which is not explicitly modeled here.

For the SSP retrievals we fit Eq. (9) to GOME reflectivity spectra utilizing a robust, non-linear, large-scale trust-region method (Byrd et al., 1988), which solves the optimization problem

$$\min_{N_{\ell}, (A \dots D)} \sum_k \left[ R_k^{\text{GOME}} - R_k(N_{\ell}, (A \dots D)) \right]^2. \quad (11)$$

Each SSP fit is initialized with a flat WV subcolumn profile  $N_{\ell,0} = 10^{16}$ , where the  $N_{\ell}$  are given in units of subcolumn density. Each fit is constrained by an upper limit  $N < N_{\ell, \max}$  in the form of a step function.  $N_{\ell, \max}$  consists out of three steps over the total number of levels with high values over the first three kilometers, medium values between 3 and 5 km, and low values for the higher levels. The lower profile constraint is set to  $N_{\ell, \text{lower}} = 0$ . The constraints prevent the fit from giving too much weight to the higher altitude levels, which otherwise would increase the relative contribution of the singly scattered photons to unrealistic values by setting the surface albedo to zero and, in doing so, decrease the total mean free path length. We scale the upper profile constraint differently for geolocations between 30° N and 30° S latitude, between 30° and 60° latitude and latitudes above or below 60°.

## 6. Method accuracy and bias adjustment

Two important assumptions were made in the derivation of Eqs. (6) and (7): (i) the analytical derivation of the  $w$  and  $S$  parameter (Eq. 4) which is only accurate for optical densities lower than 0.1 and (ii) the assumption on the dominant spectral structure parameter  $w_{\max}$  for individual light path of single-scattered photons (Sect. 3). The bias introduced due to these assumptions we call the method-related bias of SSP.

In principle, by reducing the thickness of the individual layers (i.e. utilizing more altitude layers) the optical thickness per layer may be always reduced below the 0.1 limit, where analytical derivation for  $w$  and  $S$  are a good approximation of the real value and the non-linear nature of the absorption of individual lines per layer becomes

## Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

**Retrieval of water vapor columns**

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

weakened. However, increasing the number of altitude layers increases the number of fit parameters and makes the optimization procedure computationally more expensive without gaining much additional information about the profile. Figure 4 shows that, by using 18 altitude layers, the spectrally averaged optical density per layer is significantly below the 0.1 using a high WVC of  $1.84 \times 10^{23}$  molec/cm<sup>2</sup>. Therefore, in our case, the analytical derivation of  $w$  and  $S$  is a good assumption and does not contribute to the method-related bias. Consequently, our assumption of the dominant spectral structure parameter  $w_{\max}$ , which is equal to the  $w$  value for the layer with the highest optical density in the path, creates a bias. The fixed  $w_{\max}$  is usually higher with respect to what the real  $w$  parameter per layer would be. From Eq. (2), an underestimation follows for high WVC.

In order to quantify this bias and to correct for it, we perform SSP retrievals of 26 lbl forward-modeled reflectivity spectra on the basis of Eq. (10) from a range of WV profiles, with the total WVC ranging between  $8 \times 10^{21}$  and  $1.8 \times 10^{23}$  molec/cm<sup>2</sup>. These profiles were taken from the ECMWF database and chosen to be representative of a wide range of significantly different geolocations. The contributions by the additional background absorbers were also accounted for. In order to make a direct comparison, realistic values for the variables A to D were obtained from the best fit of the forward-lbl modeling results to GOME measurements taken at corresponding geolocations.

The results are summarized in Fig. 5, which shows the relative and absolute differences between WVCs retrieved using SSP and that used for the forward-lbl modeling, respectively. For low- and mid-range columns ( $\leq 1 \times 10^{23}$  molec/cm<sup>2</sup>) the maximal differences between the retrieved and the true values is about 10%. For instances of very low WVC ( $< 1 \times 10^{22}$  molec/cm<sup>2</sup>) which occur predominantly at high latitudes, the differences are larger due to the high SZA which are usually associated with such GOME measurements. Note that together with the lowest WVC (i.e.  $8 \times 10^{21}$  molec/cm<sup>2</sup>) utilized in the forward-lbl calculations, a SZA of 61° was used, which corresponds roughly to latitudes  $> 70^\circ$ . The accuracy (dotted line Fig. 5a) for low and mid columns up to  $1 \times 10^{23}$  molec/cm<sup>2</sup>, calculated from the mean relative differences, is  $-1.1\%$ .

**Retrieval of water  
vapor columns**

R. Lang et al.

For instances of high WVC ( $>1 \times 10^{23}$  molec/cm<sup>2</sup>) the systematic retrieval bias increases up to a maximum of  $-20\%$ , which is referred to as the underestimation of the WVC by the dominant layer assumption. An empirical correction term was found by fitting a second-order polynomial (solid line in Fig. 5b) to the absolute differences, which results in

$$WVC_{\text{corr}} = -0.018 * (WVC - 0.096)^2 + 0.164 * (WVC - 0.096) - 0.283, \quad (12)$$

where the WVC is in units of  $10^{22}$  molec/cm<sup>2</sup> (Fig. 5).

## 7. Aerosol loading and multiple scattering

Even in the absence of clouds, the background aerosol may cause a significant error in the retrieved WVC value. Aerosols may affect the net absorption, as well as the path length of the scattered light and, therefore, affect the assumptions made with respect to scattering and surface albedo.

The full effect of multiple scattering and aerosols on the retrieved WVC has been estimated by solving the full scalar radiative transfer equation, including the multiple scattering source term, by employing a doubling-adding model (DAM) (de Haan et al., 1987), which has been reduced to the scalar representation of the radiation field. The reflectance  $R(\lambda)$  is calculated in a line-by-line mode for a spectral resolution of  $0.01 \text{ cm}^{-1}$ .

We investigate four atmospheric scenarios: a maritime case and a rural boundary layer case both with and without aerosol loading (clear-sky scenario). The clear sky scenario solely quantifies the contribution of multiple Rayleigh scattering and its impact on the retrieved WVC. All scenarios include the effects of Rayleigh scattering, ozone absorption and Lambertian surface reflection. In the case of the maritime scene we use a surface albedo of 0.03, whereas for the rural scene a surface albedo of 0.1 is used. For maritime aerosol loading a constant particle density of 4000 particles per cm<sup>3</sup> is assumed. For the middle and upper troposphere, we assume a tropospheric

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[I◀](#)[▶I](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

**Retrieval of water  
vapor columns**

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

background aerosol, for which the particle density decreases with the third power in pressure with the optical properties of the aerosols being taken from [Shettle and Fenn \(1979\)](#). In the rural case we have chosen a constant, but much higher, particle density of 15 000 particles per  $\text{cm}^3$ .

Table 1 lists the relative contribution of ground-reflected, single and multiple-scattered photons as a percentage of the total reflectivity using the DAM-model for the four different aerosol loading cases described above and with a medium WVC of  $7.72 \times 10^{22}$  molec/ $\text{cm}^2$ , for which the systematic method related bias may be neglected. Without aerosols, the maximum contribution of multiply-scattered photons is about 8% at 590 nm. Single scattering is the dominant source of reflectivity for the maritime aerosol scenario (47%) because of low surface albedo, with the ground-reflected component being comparable or lower than the contribution due to multiple scattering. This may be contrasted with the rural aerosol scenario, in which multiple scattering is the dominant source of scattered light (51%) due to the high aerosol optical density. Aerosols reduce the ground-reflected component down to 13%.

The first-order polynomial utilized in this retrieval to implement multiple-scattering (Eq. 5) does not account for a differential contribution reflecting the absorption by water vapor, but only for the broad-band effect. The retrieval of parameters C and D may, therefore, influence the retrieval of the surface albedo (parameters A and B) and the WVC.

Table 1 also includes the SSP retrieved WVC and surface albedo values from the DAM forward-modeled reflectivity spectra for the four scenarios. The results show that SSP retrieves accurate values for both of the clear sky situations. It underestimates the WVC by 12%, in the case of maritime aerosol loading, and overestimates it by 18% in case of rural aerosol loading. The error in the retrieved albedo is generally of the order of 20%, except for the maritime case, where the aerosol layer above the surface alters the retrieved surface albedo value significantly resulting in a decrease in the retrieved WVC.

For high WVC, the retrieval using SSP is affected by the method-related bias (see

previous section). After application of the empirical correction (Table 2), the error due to multiple scattering for the maritime and rural case is less than 6 and 2%, respectively (clear sky cases). The impact due to maritime and rural aerosol loading on the retrieved column is less than 9 and 14%, respectively, after correction for the method-related bias.

## 8. Results for single GOME pass

Figure 6a shows uncorrected WVC results (stars) using SSP for the GOME pass on 23 October 1998, between 80° S and 80° N in latitude and longitudes between 90° W and 150° W, which contains a total of 476 ground-pixel nadir measurements. For comparison, we show values from the OACS method utilizing the same GOME measurements (solid line) and exploiting the same spectral region, together with co-located values from ECMWF (solid line with dots).

In general, the SSP values compare well with the ECMWF values, except in situations where the percentage cloud cover significantly exceeds 10%. This may be explained by an enhanced photon path length, as is modeled by SSP, due to an enhanced contribution of multiple-scattered photons in case of the presence of thick clouds. Moreover, the presence of a significant cloud cover will reduce the transmission of light to the lower layers of the atmosphere where most of the water vapor is located. These two effects occasionally tend to cancel each other, as for the first the SSP model underestimates the absorption resulting in an overestimation of the WVC by the fitting procedure, whilst for the second SSP overestimates the absorption resulting in an underestimation of the WVC.

In comparison to the similarly implemented but more accurate OACS sampling method, the uncorrected SSP results systematically underestimate the higher columns, as would be expected considering the method related bias of SSP (see Sect. 6). The highest WVC in the tropics are systematically 15 to 20% lower than those of the other datasets. The circles in Fig. 6 show SSP retrieved WVC after the bias-correction (Eq. 12) for val-

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion



ues  $>1 \times 10^{23}$  molec/cm<sup>2</sup> were applied.

A critical parameter which governs the accurate retrieval of WVC is the shape of the resulting subcolumn profile retrieved by SSP. Due to the effect of the altitude dependence of both  $\rho$  and  $T$  on the spectral structure of WV, the effective absorption is governed by the specific atmospheric paths through which the photons travel. The retrieval of an unrealistic shape of the retrieved subcolumn profile may, therefore, unrealistically alter the weights of various atmospheric paths, i.e. the relative contribution between ground-reflected, singly and multiply scattered light. This yields an error in the retrieved surface albedo, which weights the direct, ground-reflected, light path contribution (Eq. 5) or an error in the retrieved broad-band multiple scattering contribution. This error propagates into the final retrieved column. From Fig. 7 we can see that the differences between the WVC values retrieved by OACS and SSP, and the differences between the surface albedo values retrieved by both methods, are very well correlated. Introducing upper fit constraints assists both methods in finding the accurate profile shapes but the OACS method is, due to its more accurate treatment of the representation of the spectral structure at each altitude level, superior to the SSP method, resulting in higher sensitivity and better agreement with the independent data-set from ECMWF, especially for high WVC. However, a big benefit of SSP is that it is about 4 to 5 times faster than the OACS method even though it uses a higher sampling resolution than the one chosen for OACS in Lang et al. (2002).

## 9. Global retrieval results

Finally we present the results of a global WVC retrieval utilizing 47 000 ground pixels from GOME between 70° S and 70° N latitude on 22–24 October 1998 using the SSP method. The globally corrected results (Eq. 12) are shown in the global plot in Fig. 8a. For comparison we show the global WVC, as given by ECMWF, on 23 October 1998, 18:00 UTC (Fig. 8b). Figure 8c shows the global WVC as retrieved by SSP for cloud-free pixels, i.e. where the cloud-fraction is reported to be lower than 10%. The cloud

### Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

**Retrieval of water vapor columns**

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

cover fraction was taken as reported by the GOME Data Processor (GDP) level 2 data (Balzer and Loyola, 1996; DLR, 1999). This fraction is derived by the initial cloud fitting algorithm (ICFA) with the cloud top pressure being taken from the International Satellite Cloud Climatology project (ISCCP) (Koelemeijer et al., 1999).

The general latitudinal dependence of WV concentration, with high WVC in the equatorial regions and low WVC towards the poles, is well represented by the SSP retrieval results. Individual features common to both the ECMWF model results and the SSP results may also be identified as, for example, the low WVC above north-west China and the Gobi desert as well as in the United States at this time above North-west Carolina. A similar transition from low to high WVC can be observed from the southern part of the United States to Mexico and the Caribbean Sea. However, GOME-SSP WVC results show significantly lower values compared to ECMWF above the Sahara, Saudi Arabia and the Andes around 40° S latitude. SSP retrieves high WVC above the Philippine Sea north of New Guinea where the ECMWF model reports only medium water vapor concentrations.

Figure 9 shows a scatter plot between SSP retrieved WVC from GOME and collocated WVC data reported by ECMWF, where the GOME measurements coincide with ECMWF data within a 1 hour time window at 23 October 1998. We compare 420 GOME measurements with a GDP level 2 data reported cloud fraction of less than 10%. The gradient of the scatter data is 0.98 with a maximum scatter of 50%.

## 10. Discussion

Differences in the surface albedo retrieved by the OACS sampling technique and SSP, both exploiting the same WV absorption band, lead to overall differences between the retrieved WVC. The difficulties in the retrieval of the correct surface albedo is a general problem for nadir remote-sensing techniques, especially during instances of high aerosol optical depth. This is due to the uncertain contribution to the backscattered light introduced by aerosol layers, with the aerosol profile being a critical parameter.

**Retrieval of water vapor columns**

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

Moreover, surface albedo is also dependent on wavelength, geolocations and season. In our wavelength region of interest, the utilized first-order polynomial is a good approach for the wavelength dependence of the albedo for most cases (Koelemeijer et al., 1997). We find, that the accuracy of the retrieved values and the gradients of the surface albedo are correlated with the shape of the retrieved WVC profile and the retrieved contribution of multiply scattered photons, which is also accounted for by a first order polynomial. For this reason, we introduced three different step functions, depending on geolocation, as upper constraints for the retrieved profile, in order to prevent the optimization method occasionally finding a local fit minima by setting the ground albedo to zero and unrealistically increasing the scattering components.

An additional concern for any WV retrieval approach is the accuracy and completeness of the cross-section database used for the construction of the  $w$ - $S$  look-up table. From recent measurement studies of water vapor absorption bands other than the one used in this study, Learner et al. (2000) recently showed discrepancies ranging from about 100% for small lines to about 20% for strong lines for the main water vapor absorption bands in the region between 1110 and 685 nm. They also found systematic differences in various bands ranging from 6% to 33%. Even though measurements by Learner et al. (2000) in the wavelength range between 585 and 600 nm have not been studied, large differences in some of the water absorption bands between 1110 and 685 nm suggest the presence of potentially large uncertainties in the reference cross-sections of the HITRAN'96 database used in this study. In addition, the presence of many weak absorption lines, not accounted for in the HITRAN'96 database, may contribute to an additional background absorption, which may affect the retrieval of both surface albedo and broad-band multiple-scattering contribution and, therefore, also the retrieval of WVC.

The contribution of aerosols and clouds to the error in the retrieved column, as given in Sect. 7, can only be estimated. In the case of aerosols, this is due to their high variability in optical depth, concentration profile and scattering probabilities, all of which depend on geolocation and season. In the case of clouds, the uncertainties arise from

the high variability with respect to cloud-top heights, form and density of clouds. Furthermore, the scatter between SSP-retrieved WVC values and those given by ECMWF may also be affected by systematic errors for cases, where the cloud-fraction is reported to be less than 10%. Differences in the reported cloud-fraction by GDP level 2 data with respect to the co-located Along Track Scanning Radiometer-2 (ATSR-2) can be as much as 18% (Koelemeijer et al., 1999).

## 11. Conclusion

In this study the spectral structure parameterization technique has been implemented and tested for retrieval of total WVC from GOME measurements utilizing a radiative transfer scheme, which includes molecular clear-sky direct-path absorption as well as molecular clear-sky single-scattering, together with a broadband approximation of higher orders of scattering. The radiative transfer scheme was previously used for retrieval of total WVC using the Optical Absorption Coefficient Spectroscopy. SSP is less accurate and less sensitive than the OACS retrieval method but significantly faster and easier to implement. Retrieval of WVC using SSP for one ground pixel takes less than 30 seconds on a Pentium III, 800 MHz, using an uncompiled MATLAB retrieval code, and requires less than 128 MB internal memory.

Whereas OACS utilizes a spectral sampling technique based on cross-section probability density distribution functions, SSP utilizes average line parameters within the sampling region, and is, therefore, conceptually closer to band models. From retrieval studies of synthetic spectra we conclude that the sensitivity of SSP is better than 10% and the accuracy is about 1% for WVC lower than  $1 \times 10^{23}$  molec/cm<sup>2</sup>. For WVC above this limit, a systematic offset of up to 20% is introduced by using an analytical second-order approach for the determination of the average line parameters  $w$  and  $S$  (Eq. 4). Moreover, the assumption of a dominant absorbing layer applied to the modeling of nonhomogeneous paths (Sect. 3 and Eq. 7) also introduces a bias. This bias has been corrected for in the current study by applying an empirical second-order polynomial

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

(Sect. 6).

We studied the impact of aerosols and the differential contribution of multiple scattering for four different scenarios utilizing synthetic spectra from a doubling adding radiative transfer method. For medium WVC, for which no correction is needed, the multiple-scattering impact on the SSP retrieved WVC is below 1%. Neglecting maritime aerosols and rural aerosols results in a retrieval error of less than 13 and 19% respectively. For corrected high WVC (Table 2) the error due to multiple scattering for maritime and rural scenario is less than 6 and 2%, respectively (clear sky cases). Here, the impact due to maritime and rural aerosols on the retrieved column is less than -9 and +14%, respectively, after correction for the method-related bias.

A global comparison between SSP retrieved WVC and data from ECMWF shows good general agreement in the latitudinal dependence of the water vapor concentration. However, significant regional differences, for example, over the Sahara and the Philippine Sea are found. The higher spatial resolution of the GOME measurements over sea and land with respect to ECMWF, where the density of ground measurements is low, is clearly an advantage of this kind of SSP retrieval from GOME. However, ECMWF data is provided for a specific time and day, whereas global GOME retrieval results, as shown in Fig. 8, are collected over three days. Wide parts of the differences in the global data set may, therefore, also be related to changes in the global water vapor distribution over the analysis period. The scatter plot comparison between SSP and ECMWF (Fig. 9), as well as the single GOME pass comparison (Fig. 6), are correlated in time and geolocation. The scatter plot shows good correlation between the two data-sets of correlated cloud-free ocean pixel with a gradient of 0.98. However, in some cases, the scatter can be as high as 50%, which points to local differences between the two data products.

Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

## 12. Outlook

In general, SSP is well suited for fast clear-sky retrieval of narrowband absorbers from nadir satellite instruments like GOME. The accuracy in the retrieval of the surface albedo and the impact of aerosols and multiple scattering limits the accuracy of the retrieved WVC. The latter may be improved by introducing surface albedo from global databases once available and tested. The current implementation of SSP in a direct path and single-scattering radiative transfer scheme provides the possibility for implementation of specific aerosol optical properties. The treatment of clouds for such nadir-viewing measurements in the visible is a yet unsolved problem. SSP provides WVC (and potentially profile) retrieval from instruments like GOME and SCIAMACHY, which were not originally intended for WV retrieval. SSP retrieved WV values may be used to refine other retrievals from such satellite data because WV “contaminates” the entire spectrum. In addition, SSP is, in principle, also suitable for the retrieval of spectrally overlapping narrowband absorptions from different species. This situation occurs in the IR regions around  $2\ \mu\text{m}$  covered by the SCIAMACHY instrument. Multi-species retrieval may be performed by introducing an additional averaged line parameter, which accounts for the degree of overlap between the different absorbers within a specific spectral sampling width.

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### Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

## References

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### Retrieval of water vapor columns

R. Lang et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

## Retrieval of water vapor columns

R. Lang et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

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[1110](#)



Retrieval of water  
vapor columns

R. Lang et al.

**Table 1.** Relative contribution of ground-reflected, singly-scattered and multiply-scattered photons to the total reflectivity calculated by the DAM model assuming a SZA of 40° and a WVC of  $7.72 \times 10^{22}$  [molec/cm<sup>2</sup>] for each case

	Model Input		Model Output			SSP Retrieval Results			
	AOD <sup>a</sup>	$\Lambda$	GR <sup>b</sup>	SS <sup>c</sup>	MS <sup>d</sup>	WVC <sup>e</sup>	$\Delta$ WVC [%]	$\Lambda$	$\Delta\Lambda$ [%]
Maritime Clear Sky	0	0.03	42.2	49.6	8.2	7.68	-0.33	0.036	+20.0
Maritime Aerosol	0.30	0.03	17.9	46.9	35.2	6.76	-12.3	0.052	+73.3
Rural Clear Sky	0	0.10	67.8	23.9	8.3	7.69	-0.26	0.110	+10.0
Rural Aerosol	0.63	0.10	13.0	25.1	61.9	9.14	+18.6	0.121	+21.0

<sup>a</sup> Aerosol Optical Depth<sup>b</sup> Percentage of ground reflected light at 592 nm<sup>c</sup> Percentage of single scattering at 592 nm<sup>d</sup> Percentage of multiple scattering at 592 nm<sup>e</sup>  $10^{22}$  molec/cm<sup>2</sup>

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

I◀

▶I

◀

▶

Back

Close

Full Screen / Esc

Print Version

Interactive Discussion

© EGS 2002

Retrieval of water  
vapor columns

R. Lang et al.

**Table 2.** Same as Table 1 but for a high WVC of  $1.43 \times 10^{23}$  molec/cm<sup>2</sup>. The relative difference  $\Delta$  WVC is calculated between the model value and the corrected retrieved WVC

	Model Input		Model Output			SSP Retrieval Results				
	AOD <sup>a</sup>	$\Lambda$	GR <sup>b</sup>	SS <sup>c</sup>	MS <sup>d</sup>	WVC	corr. WVC <sup>e</sup>	$\Delta$ WVC [%]	$\Lambda$	$\Delta\Lambda$ [%]
Maritime Clear Sky	0	0.03	43.8	48.2	8.0	12.5	13.5	-5.2	0.039	+30.0
Maritime Aerosol	0.21	0.03	20.0	53.2	26.8	14.0	15.5	+8.5	0.038	+26.6
Rural Clear Sky	0	0.10	69.1	22.8	8.1	13.3	14.5	+1.6	0.102	+2.0
Rural Aerosol	0.51	0.10	17.9	30.9	51.2	14.6	16.3	+14.1	0.112	+12.0

<sup>a</sup> Aerosol Optical Depth<sup>b</sup> Percentage of ground reflected light at 592 nm<sup>c</sup> Percentage of single scattering at 592 nm<sup>d</sup> Percentage of multiple scattering at 592 nm<sup>e</sup>  $10^{22}$  molec/cm<sup>2</sup>

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

I◀

▶I

◀

▶

Back

Close

Full Screen / Esc

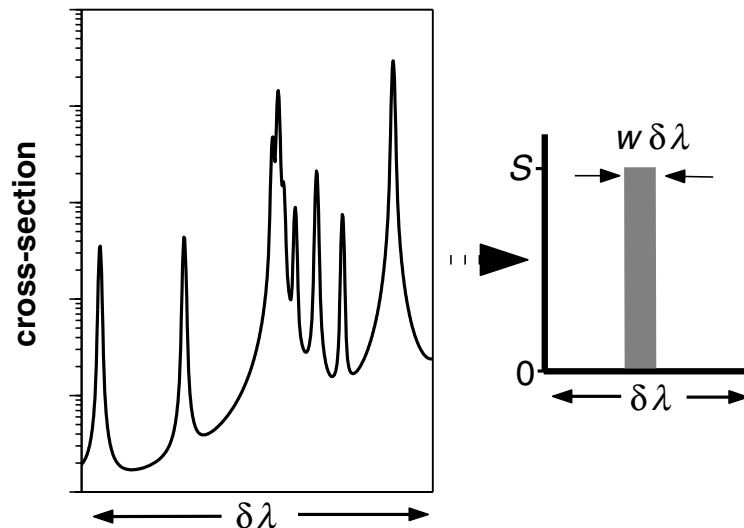
Print Version

Interactive Discussion

© EGS 2002

Retrieval of water  
vapor columns

R. Lang et al.



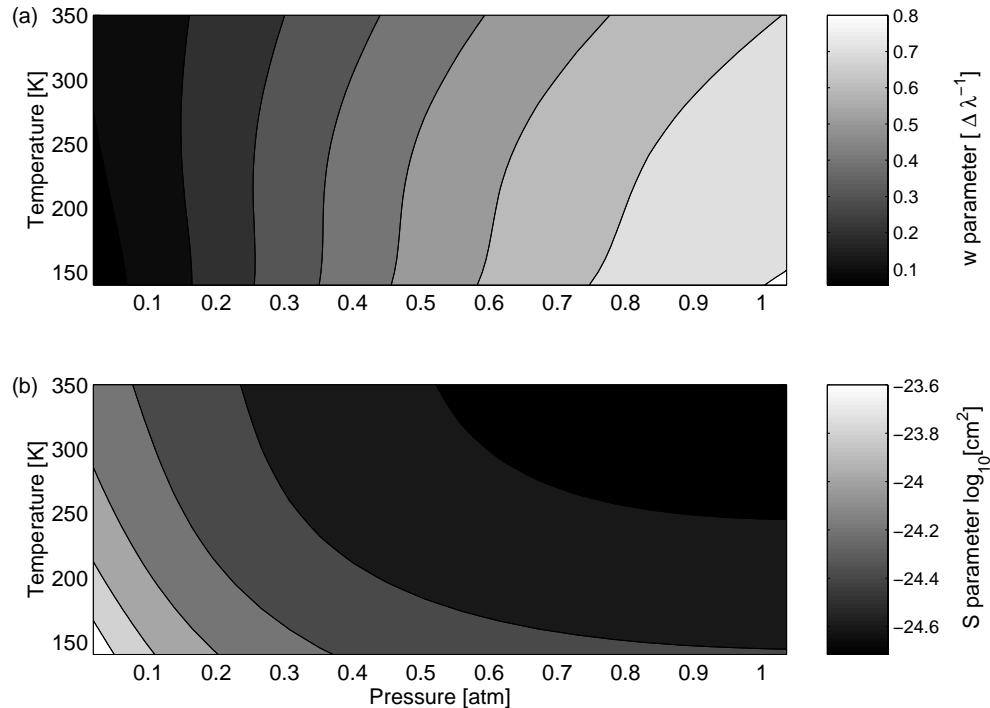
**Fig. 1.** Schematic of the construction of an effective pixel coverage parameter  $w$  and effective mean value for the absorption cross-section  $S$  (right hand panel) from a typical absorption spectrum which has been sampled over a wavelength range  $\delta\lambda$  (left hand panel).

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[I◀](#)[▶I](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

© EGS 2002

Retrieval of water  
vapor columns

R. Lang et al.



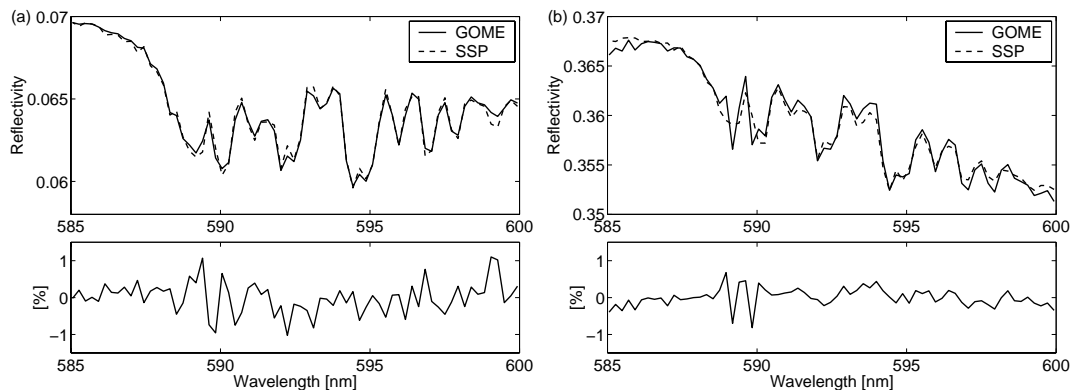
**Fig. 2.** Contour plot of the  $w$ - (a) and the  $S$ -parameter (b) look-up-table for one spectral sampling bin  $j$  in dependence of temperature and pressure. A standard pressure and temperature grid is used for the construction of the look up tables, covering all possible pressure and temperature ranges within the altitude region of interest.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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Retrieval of water  
vapor columns

R. Lang et al.



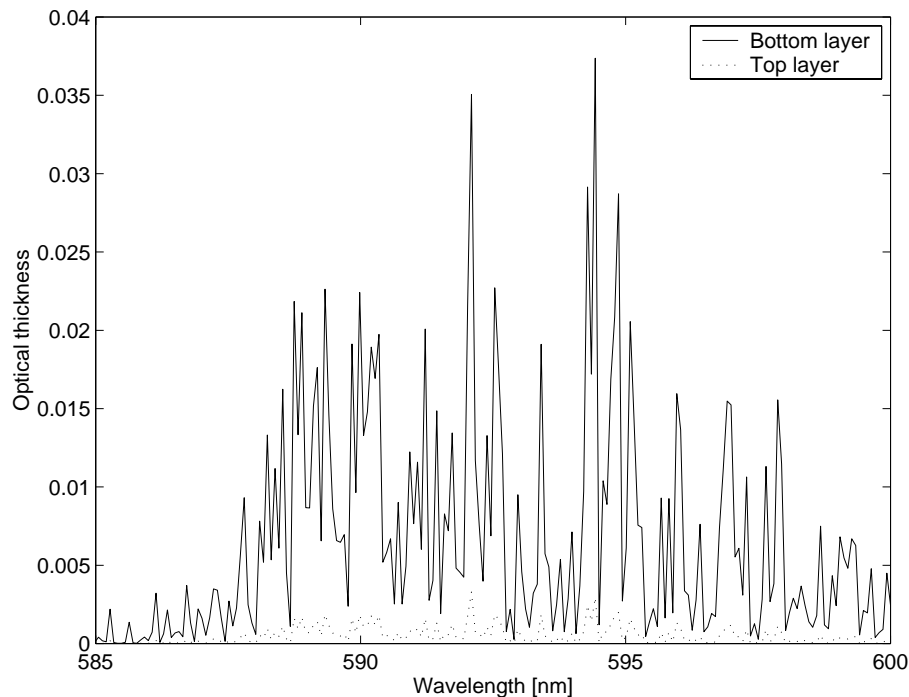
**Fig. 3.** The left panels **(a)** show a GOME measurement (solid line) at  $34^\circ$  N latitude and  $110^\circ$  W longitude over the pacific with a SZA of  $23^\circ$  together with the result of a SSP forward model (dashed line) and their residual (GOME-SSP)/GOME). The albedo was adjusted using a first order polynomial, resulting in a value of 0.04 at 590 nm. The WV density profile was taken from ECMWF with a high WVC of  $1.34 \times 10^{23}$  [molec/cm<sup>2</sup>]. The right panels **(b)** show the same but than for a rural GOME measurement at  $62^\circ$  N latitude and  $100^\circ$  W longitude with a high SZA of  $73^\circ$  and a low WVC of  $8.9 \times 10^{21}$ . For this pixel the albedo was adjusted resulting in a value of 0.32 at 590 nm.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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**Retrieval of water  
vapor columns**

R. Lang et al.



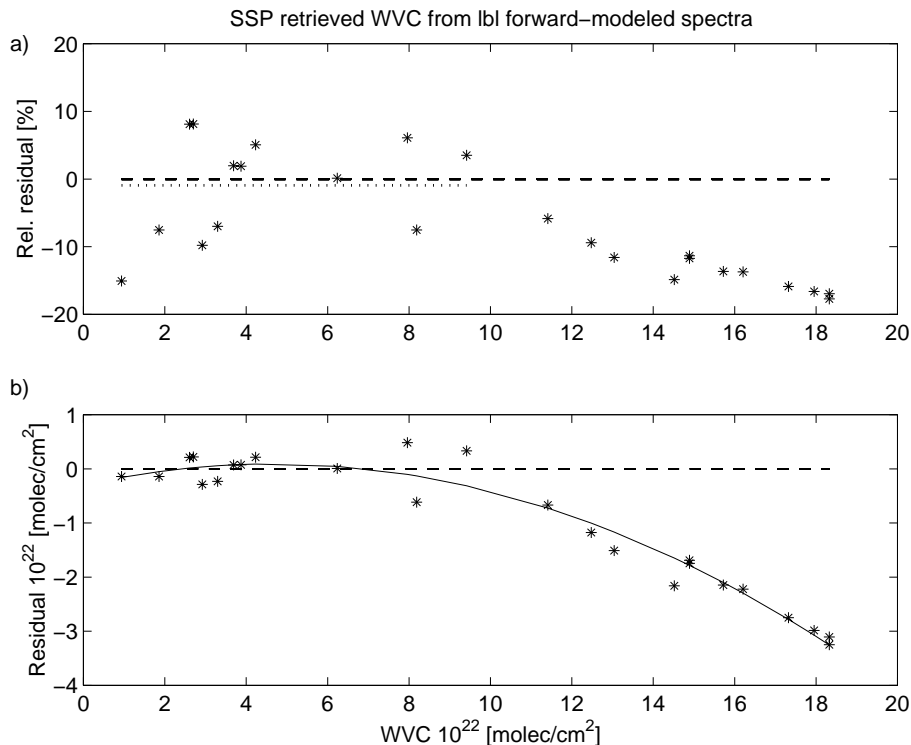
**Fig. 4.** Spectral averaged optical thickness for the bottom and top atmospheric layer used in the retrieval with subcolumn densities of  $1.8 \times 10^{22}$  and  $1.3 \times 10^{21}$  molec/cm<sup>2</sup>, respectively. Here, we selected a high total WVC of  $1.83 \times 10^{23}$  molec/cm<sup>2</sup>.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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Retrieval of water  
vapor columns

R. Lang et al.



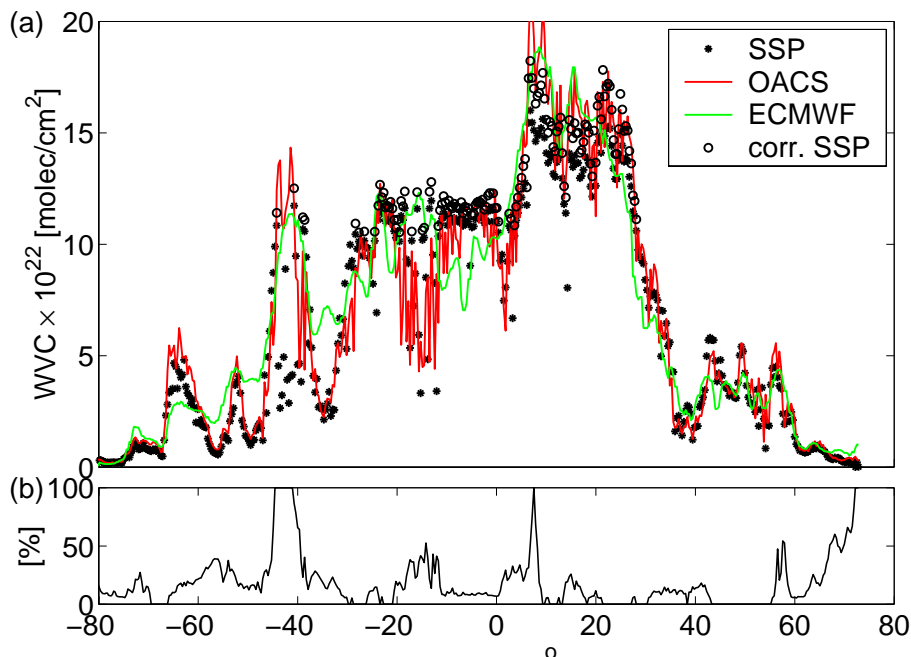
**Fig. 5.** Relative (a) and absolute (b) differences between the retrieved WVC by SSP and the WVC utilized in the lbl forward model. The solid line in the lower panel (b) shows a second order polynomial fit through the absolute difference values. The polynomial is used to adjust for the systematic method related bias of the SSP retrieval. The upper panel (a) also shows the accuracy of the method (dotted line) for low and mid-range WVC up to  $1 \times 10^{23}$  molec/cm<sup>2</sup>.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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Retrieval of water  
vapor columns

R. Lang et al.



**Fig. 6.** SSP uncorrected WVC fit results over latitude ((a), filled circles) for a GOME pass on 23 October 1998, covering longitudes from  $90^\circ$  W at  $73^\circ$  N latitude to  $150^\circ$  W at  $80^\circ$  S latitude. For validation, the WVC values from OACS using GOME measurements (red curve) and from ECMWF given at 18:00 UTC are shown (green curve). The open circles denote SSP corrected values for WVC higher than  $1 \times 10^{23}$  molec/cm<sup>2</sup>. The cloud coverage in percentage per GOME observation is taken from GOME GDP level-2 data and is indicated by the solid curve in the lower panel (b).

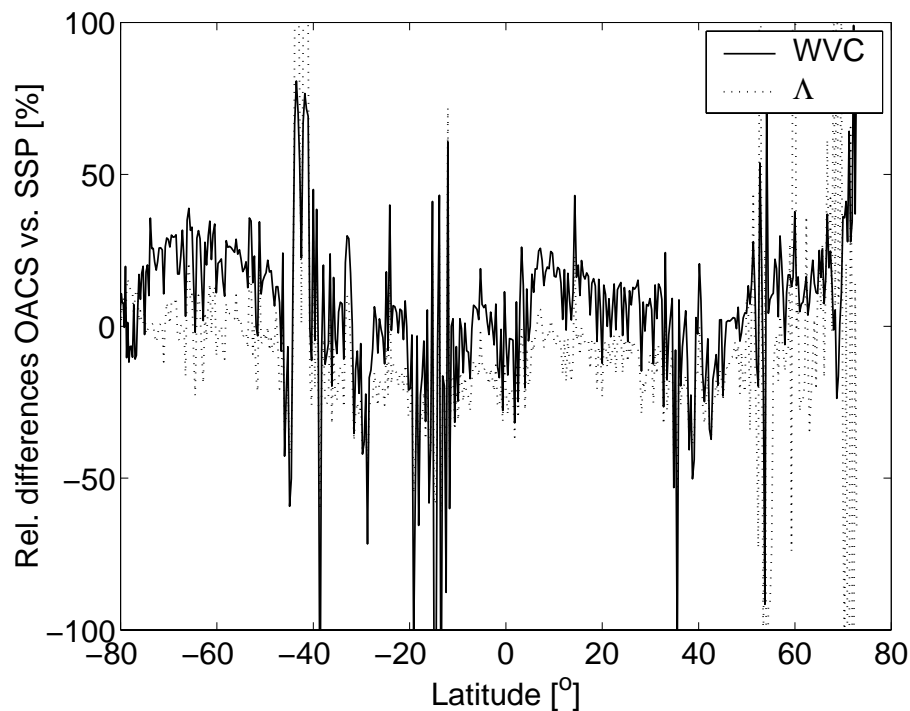
[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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Retrieval of water  
vapor columns

R. Lang et al.



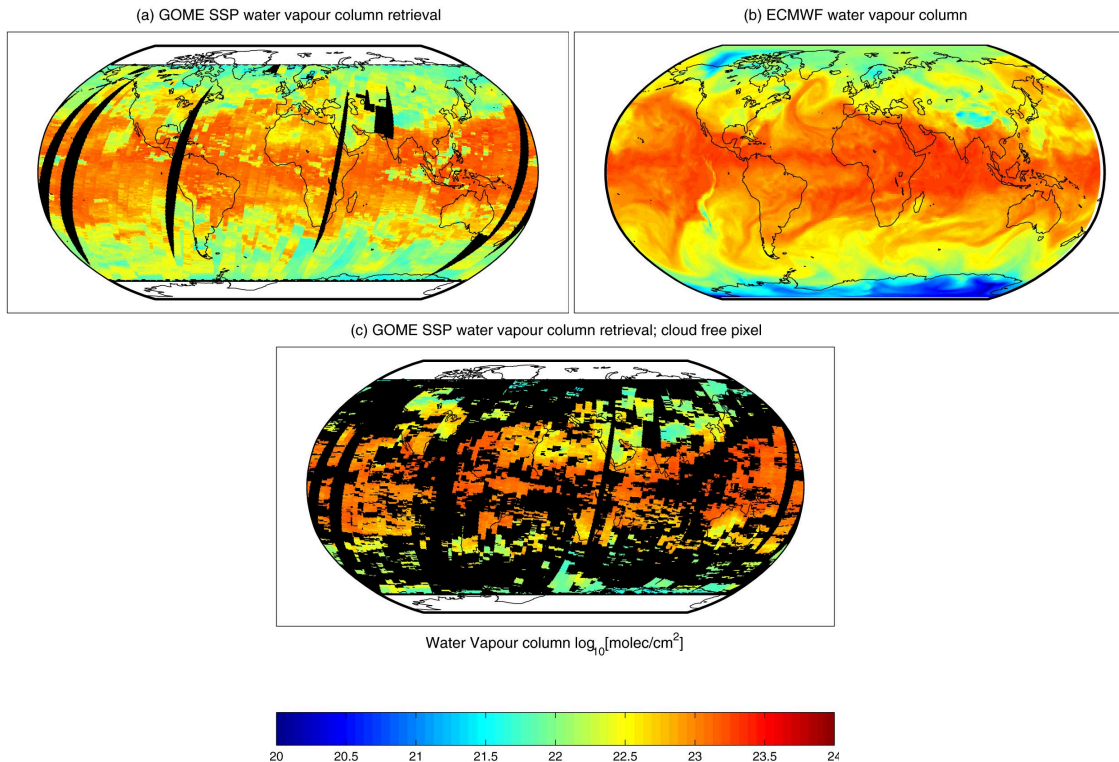
**Fig. 7.** Relative differences between SSP and OACS for retrieved values of WVC ( $[\text{OACS}-\text{SSP}]/\text{OACS}$ ; solid line) and surface albedo  $\Lambda$  ( $[\text{SSP}-\text{OACS}]/\text{OACS}$ ; dotted line). An underestimation of the surface albedo leads to an overestimation in the retrieved WVC. The relative differences result from the same GOME track as shown in Fig. 6 including all ground pixel.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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## Retrieval of water vapor columns

R. Lang et al.



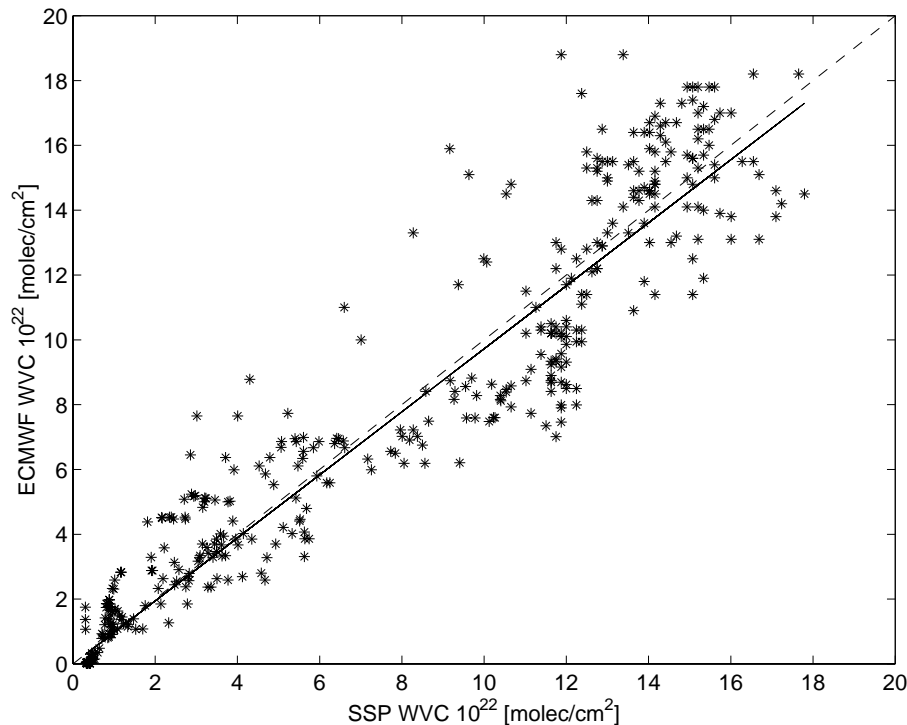
**Fig. 8.** Global total water vapor column plots for all retrieved GOME ground pixel using SSP (a). The measurements were performed from 22–24 October 1998. The right panel (b) shows global WVC as given by ECMWF at 23 October, 1998, 18:00 UTC. The lower panel (c) shows SSP columns for GOME observations where the cloud fraction reported by GOME level-2 data is lower than 10%. Data gaps for GOME retrievals are due to calibration periods or data processing failures. At 24 October from 14:00 UTC GOME was in Narrow Swath Mode until 11:00 UTC the next day. For these tracks the data was interpolated to the standard swath width.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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Retrieval of water  
vapor columns

R. Lang et al.



**Fig. 9.** Scatter plot between ECMWF total WVC and SSP retrievals for GOME measurements with a GDP level-2 data reported cloud fraction lower than 10%. The GOME measurements are performed between 70° S and 70° N in latitude: on 30 September 1999 from 0:09 to 1:05 UTC between 120° E and 180° E longitude and on 23 October 1998 from 17:52 to 18:32 UTC between 90° W and 150° W longitude. Co-located ECMWF data is given for 00:00 UTC and 18:00 UTC, respectively. The solid line denotes the best fit of a linear polynomial to the retrieved columns.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Print Version](#)[Interactive Discussion](#)

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